# CALET Ultra-Heavy Cosmic-Ray Observations Incorporating Trajectory Dependent Geomagnetic Rigidities

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# Ultra Heavy Cosmic Ray Analysis

- CALET has a special UH CR trigger utilizing the CHD and the top 4 layers of the IMC that:
  - has an expanded geometry factor of ~4000 cm<sup>2</sup>sr
  - has a very high duty cycle due to low event rate
  - ISS obstructions in FOV reduce benefit and complicate analysis
- Analysis presented here uses data with UH triggers and good trajectories
- Relative abundances of elements below 14 Si impacted as they only trigger at higher incidence angles
- UH analysis requires specialized data corrections and selections optimized for UH range using <sub>26</sub>Fe



# **CALET UH Analysis Status**

- Using ~3 years of CALET Level 2 PASS03.1 UH data
  - Analysis developed on previous 17 month data set applied
  - UH analysis CHD paddle time corrections
  - UH analysis CHD paddle position dependent corrections
  - Data selections for incidence angle, vertical cutoff rigidity, charge consistency, etc. applied
- Abundances fit for previous data sets agree within statistics with other UH measurements (SuperTIGER and ACE-CRIS)
- Work continues on trajectory dependent rigidity thresholds and ISS obstruction identification
- Analysis planned for CALET HE trigger data set with energy reconstruction in TASC for Wolfgang Zober's PhD thesis project.

# CHD 26 Fe Time Corrections

- CHD time step histograms filled until at least 500 <sub>26</sub>Fe range events in each CHD paddle
- In each time step 26 Fe peaks fit with a Gaussian for each paddle and paddle average time steps calculated
- CHD paddle signals multiplied by the ratio of the mean of both layers over the full dataset to the paddle time step mean



# CHD 26 Fe Time Contours



### **Corrected Time Contours**



### **CHD** Position Correction Method

- CHD paddles divided into thirds of the paddle width (1.07 cm) segments
- <sub>14</sub>Si and <sub>26</sub>Fe peaks fit with Gaussian for each segment
- CHD paddle signal multiplied by the ratio of each layer mean to the segment mean



# CHD<sub>26</sub>Fe Position Dependence CHDX CHDY



# CHD <sub>26</sub>Fe After Position Correction

CHDX

CHDY



# CHDX Si Position Dependence



# CHD 14Si After Position Correction

CHDX

#### CHDY



# CHDX Si and 26 Fe Peak Means CHDX CHDY



# Charge Consistency Selections

- Selection cut is made for charge estimate consistency between CHDX and CHDY
- $Z_{est} \propto CHD^{1/1.7}$
- $\Delta Z = (Z_{CHDX} Z_{CHDY}) / (Z_{CHDX} + Z_{CHDY})$  for  $\rm Z_{CHDX}$  and  $\rm Z_{CHDY}$  total layer signals
- $\Delta Z_{1,2}$  uses  $Z_{CHDX}$  and  $Z_{CHDY}$  for sum of signals from two highest layer paddles
- $\pm 2\sigma$  selections applied



 $\Delta Z$  selection Includes more signal from backscatter  $\Delta Z_{1,2}$  selection focused on primary particle track

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# Paddle Dominance Selections

- Best charge estimate uses CHDX and CHDY signals from the two highest paddles
- Events with disproportionately high third paddle signals are selected
- CHDX<sub>3</sub>/(CHDX<sub>1</sub>+CHDX<sub>2</sub>) < 0.04
- CHDY<sub>3</sub>/(CHDY<sub>1</sub>+CHDY<sub>2</sub>) < 0.04



**CHDX** 



CHDY

#### **Current Analysis Charge Histogram**

- Selections on ~3 year dataset:
  - Zest > 24
  - Theta < 45 deg</li>
  - STRM > 4.0 GV
  - Z Consistency
  - Paddle dominance
  - IMC minimum
- We can clearly see well resolved peaks for <sub>32</sub>Ge, <sub>34</sub>Se, and <sub>38</sub>Sr.
- <sub>30</sub>Zn is more than a shoulder, but is not clearly resolved. Even a small improvement in resolution would help a lot here.
- More statistics should be a major help in better defining the peaks
- Geomagnetic cutoff for each trajectory should help in rejecting low energy particles that are very likely broadening the distributions.



#### **Reduced Dataset Charge Histogram**

- Selections on 17 month dataset:
  - Zest > 19
  - Theta < 45 deg</li>
  - STRM > 4.5 GV
  - Z Consistency
  - IMC Energy Correction
- We can clearly see well resolved peaks for <sub>32</sub>Ge, <sub>34</sub>Se, and <sub>38</sub>Sr.
- <sub>30</sub>Zn is more than a shoulder, but is not clearly resolved. Even a small improvement in resolution would help a lot here.
- More statistics should be a major help in better defining the peaks
- Geomagnetic cutoff for each trajectory should help in rejecting low energy particles that are very likely broadening the distributions.



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# **Event Distribution**

- To estimate the abundances detected, we used a maximum likelihood fitting routine to fit the data.
- Fits reasonably good up to <sub>34</sub>Se.
- For higher charges, the low statistics resulted in poor fits.
- For even-Zs above <sub>34</sub>Se (<sub>36</sub>Kr & <sub>38</sub>Sr) the abundances were initially estimated by taking cuts in the valleys.
- Using SuperTIGER abundances, half of the odd-Zs on either side of the even-Z charge was subtracted off of the <sub>36</sub>Kr & <sub>38</sub>Sr numbers to estimate their abundances.



CALETIData Brian 12.17.2018 interpolatedisi fe binwise correction/4th try selections/5th try at fitting/Fit figur

## **Comparing Relative Abundances**

- The ACE and ST data are "in-space" abundances.
- The CALET data have not yet been corrected to the top of the instrument.
  - Those corrections will be small, so they will not change things materially.
- The agreement with ST and ACE-CRIS appears to be quite good.
- Additional data and anticipated improved resolution should result in reduced error bars.



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#### 2017 CALET UH ICRC Results

Selections on ~13 month dataset:

- Zest > 24
- Theta < 60 deg
- STRM > 4.0 GV
- Z Consistency Abundances fit to integer centered charges with fixed  $\sigma = 0.35$



### **Trajectory Based Rigidity Threshold**

Work is ongoing on determining event trajectory based geomagnetic rigidity cutoffs. These will allow a more targeted energy threshold selection that will maximize statistics.





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# Trajectory Dependent Rigidity!

- Trajectory dependent rigidity using Wolfgang Zober's approximate geomagnetic model works!
- Can be used instead of vertical rigidity selection.



# Trajectory Dependent Rigidity!

- Trajectory dependent rigidity using Wolfgang Zober's approximate geomagnetic model works!
- Can be used instead of vertical rigidity selection.
- Resolution better at higher rigidities.



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# Trajectory Dependent Rigidity 2-4 GV

- UH Trigger histograms for events above 2 to 4 GV in 0.1 GV steps.
- Need to optimize rigidity cut to balance statistics with tail spillover.
- No other selections here.



# Trajectory Dependent Rigidity 2-4 GV

- UH Trigger histograms for events above 2 to 4 GV in 0.5 GV steps.
- Applied most of Bob's selections:
  - Theta < 45 deg</p>
  - STRM > 4.0 GV
  - Z Consistency
  - Paddle dominance
  - Still major tails



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# Trajectory Dependent Rigidity 2-4 GV

- UH Trigger histograms for events above 2 to 4 GV in 0.5 GV steps.
- Applied most of Bob's selections:
  - Theta < 45 deg</p>
  - STRM > 4.0 GV
  - Z Consistency
  - Paddle dominance
  - IMC minimum
- Strong cut on tails and on lower charges.
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#### Trajectory Dependent 2-4 GV: UH Range

- UH Trigger histograms for events above 2 to 4 GV in 0.5 GV steps.
- Applied most of Bob's selections:
  - Theta < 45 deg</p>
  - STRM > 4.0 GV
  - Z Consistency
  - Paddle dominance
  - IMC minimum
- Need refined IMC selection that reduces tails without charge bias.



# **UH Sensitivity to Obstructions**

#### Incidence Angle > 45°

#### Incidence Angle < 45°



# Future Work

- Acquire and analyze Level 2 Pass 4 data. Generate new:
  - CHD <sub>26</sub>Fe peak based time corrections
  - CHD  $_{14}$ Si and  $_{26}$ Fe peaks position corrections
  - Charge assignments  $Z(\Delta CHD, \theta)$
  - CHD charge consistency selections
  - IMC selection.
- Implement selection cuts to eliminate ISS obstructions.
- With Wolfgang Zober:
  - Trajectory based rigidity selections using individual event raytracing.
  - UHCR analysis using HE trigger events with TASC information.

# **Backup Slides**



# $\Delta CHD/\langle CHD \rangle$ vs $\Theta$ Dataset

- Partition UH dataset
  - Limit to 14 Si and up to limit incidence angle dependence by selecting:
    CHDX > 150 and CHDY > 150
  - 30 equal statistics bins in incidence angle:
    0° < Θ < 68°</li>
  - 30 equal statistics bins in relative CHD signal:  $-0.076 < \frac{\Delta CHD}{\langle CHD \rangle} < 0.076$



# **CHDX Selected Even Peak Fitting**



# **CHDY Selected Even Peak Fitting**



# Selected Charge Models

- Power Law CALET NIM: CHD = A + BZ<sup>C</sup>
- Voltz Model TIGER/SuperTIGER analysis:  $\frac{dL}{dx} = As \frac{dE}{dx} (1 - Fs)e^{-B_s(1 - Fs)\frac{dE}{dx}} + A_s \frac{dE}{dx}F_s$ Assuming constant energy: CHD = AZ<sup>2</sup> e<sup>BZ<sup>2</sup></sup> + CZ<sup>2</sup>
- BTV Model CALET NIM/SuperTIGER analysis:  $\frac{dL}{dx} = As \frac{dE}{dx} (1 - Fs)/(1 + Bs \frac{dE}{dx} (1 - Fs)) + A_s \frac{dE}{dx} F_s$ Assuming constant energy: CHD = AZ<sup>2</sup>/(1 + BZ<sup>2</sup>) + CZ<sup>2</sup>



# **Comparing Charge Assignments**



- Three charge assignments agree well within the range of the peaks fit for the models, but diverge outside of this region.
- None of the models has peaks aligned with the appropriate low-Z charges.
- Voltz model charge assignment has best low-Z resolution, which is why it has been used previously in TIGER and SuperTIGER UH analyses.

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Counts

# Handling Cross Paddle Events



cross paddle events.

# Third highest paddle versus first showing background.

# **Next Corrections and Selections**

Z<sub>CHDY<sup>1/1.7</sup></sub>



There is some IMC dependence in the latest charge assignment that might be corrected.

 $$z_{\rm CHDX^{\rm M37}}$$  Use charge consistency selections earlier in the analysis.

25

30

35

20

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 $10^{4}$ 

 $10^{3}$ 

 $10^{2}$ 

10

50