

Double-weak decays of ^{124}Xe and ^{136}Xe in the XENON1T and XENONnT experiments

Group C (Neutrino Physics)

Featuring results from [Phys. Rev. C 106, 024328](#)

Group Members:

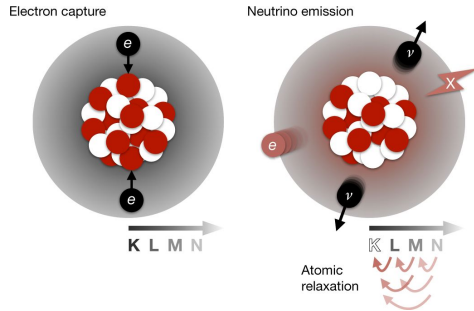
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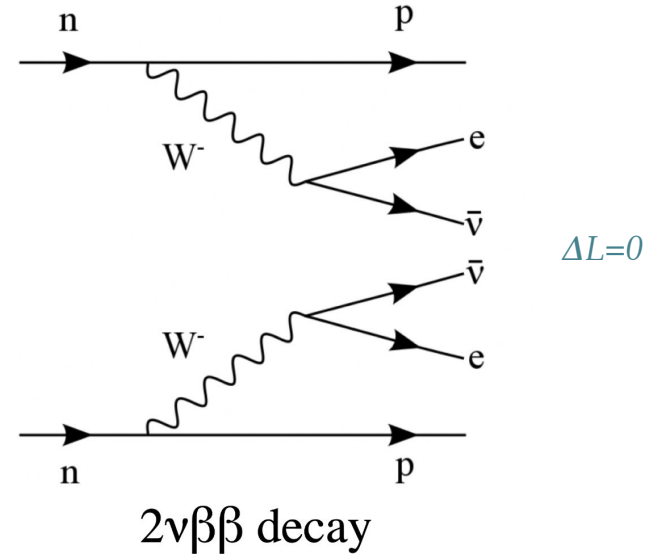
← A neutrino was here
(and trillions more! :P)

Two-neutrino double electron capture ($2\nu\text{ECEC}$)

- Important for nuclear structure models
- Second-order weak-interaction process.



- Nuclear binding energy released via the 2ν , which cannot be detected.

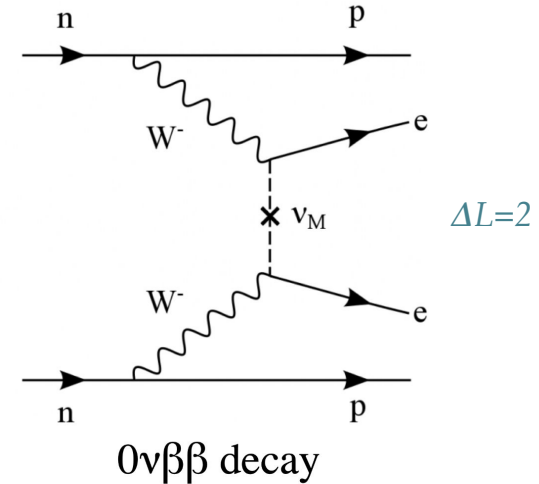


Neutrinoless double- β decay ($0\nu\beta\beta$)

- Possible if neutrino is Majorana
- Sensitive to
 - Neutrino mass
 - Lepton number violation
- Best lower limit on the ^{136}Xe half-life up-to-date from KamLAND-Zen

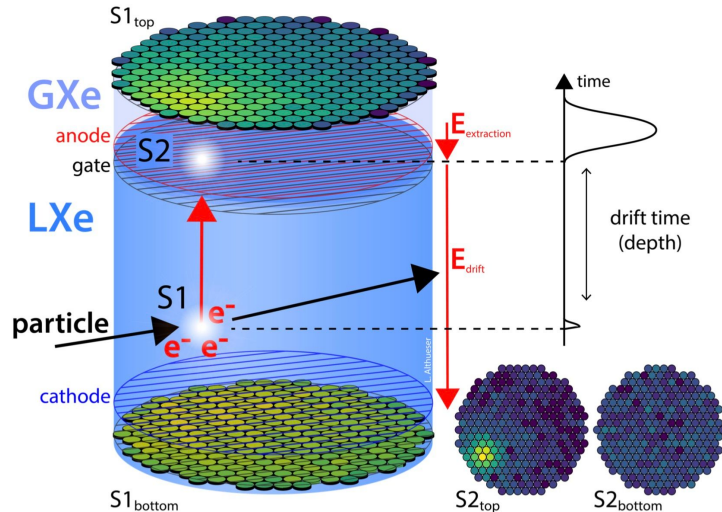
$$T_{1/2}^{0\nu} > 2.3 \times 10^{26} \text{ yr at 90\% C.L.}$$

- Requirements for a $0\nu\beta\beta$ decay of ^{136}Xe
 - Low background rate
 - Good energy reconstruction and resolution



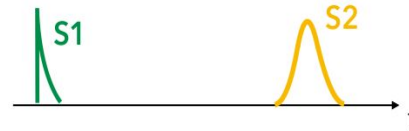
- **Signature in Xenon1T:**
energy deposits on atomic electrons in xenon:
electronic recoil (ER)

- **Main goal:** Direct detection of WIMPs (DM candidates)
- Location: Laboratori Nazionali del Gran Sasso (LNGS), Italy,
 - ~1300m rock-equivalent wall
- 3.2 t detector \Rightarrow Target mass of 2.0 t
- **(Dual-phase) Time Projection Chamber (TPC) [1]**



$$\frac{S2}{S1}(ER) > \frac{S2}{S1}(NR)$$

Single-Site - $0\nu\beta\beta$ signal

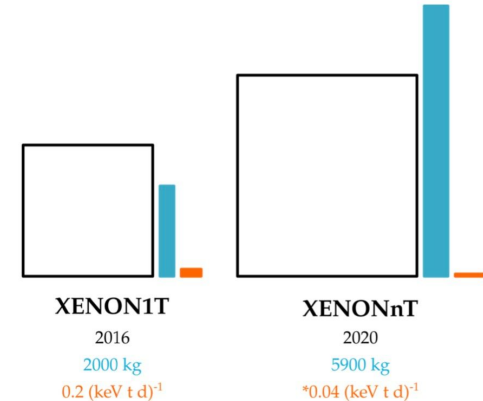


Multiple-Site - Bkg



Upgrades for XENONnT:

- **More Xe:** 5.9 t LXe with a 8.4 t capacity
- **Less background:** Radon removal, LXe purification, neutron veto, ...

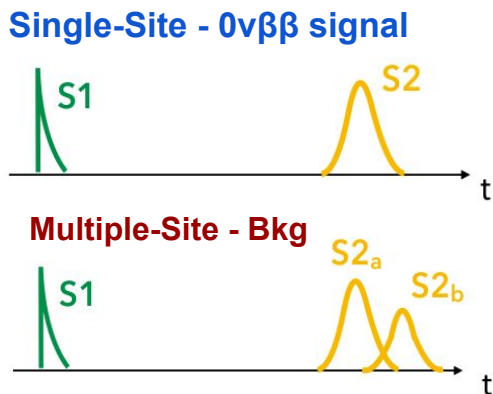


Dataset:

- Scan energies between 1600 and 3200 keV
 - $0\nu\beta\beta$ peak at $Q_{\beta\beta} = (2457.83 \pm 0.37)$ keV
- Isotope exposure of 36.16 kg yr

Analysis Strategy:

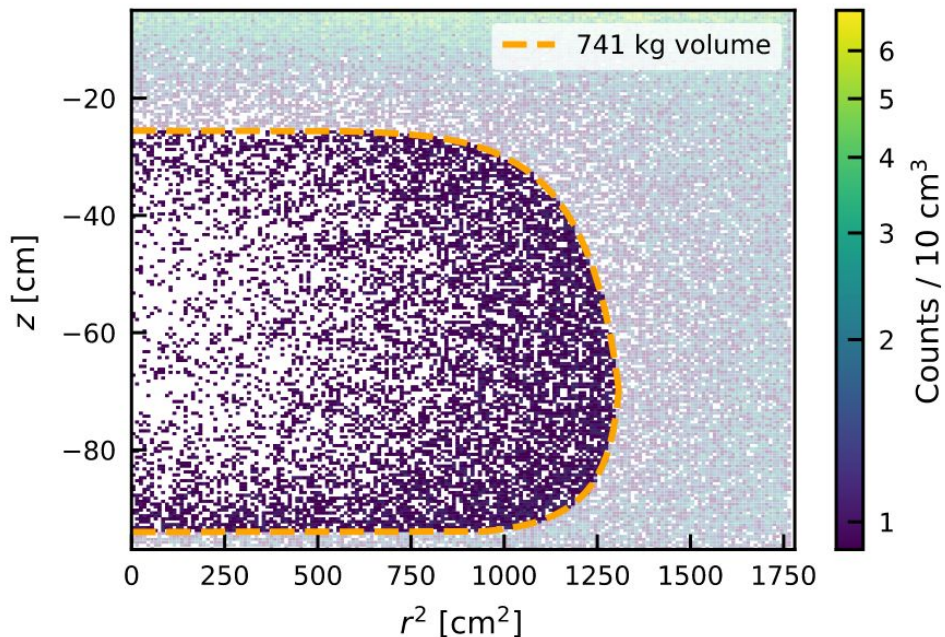
- Select events that have one S1 signal and one S2 signal with the correct shape pattern of the S2 signal



Ref: [Xenon1T Public Talk](#)

Signal Acceptance Optimization

- An FV (Fiducial Volume) in the detector is chosen that maximises signal-to-noise ratio:

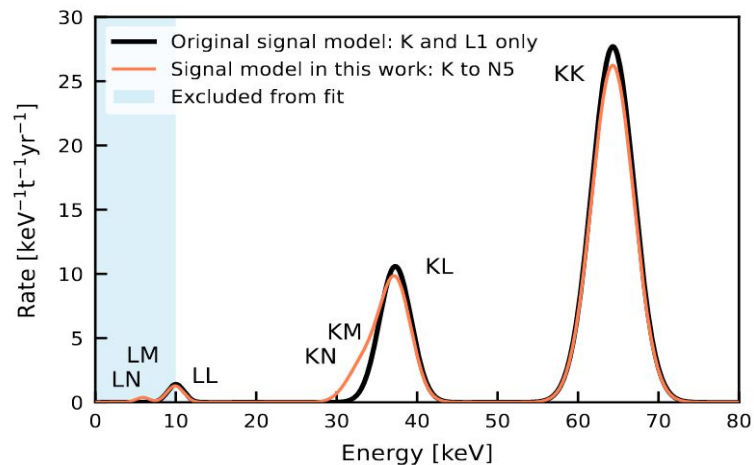
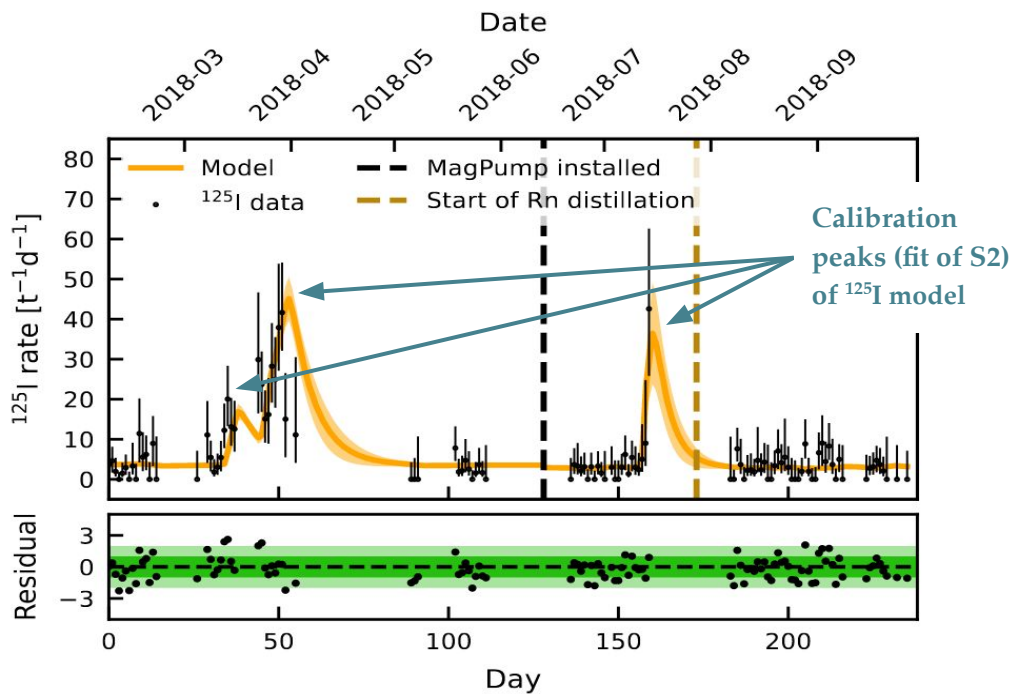


Background(s):

- Key background: ^{125}I coming from neutron activation of ^{124}Xe
 \Rightarrow Peak at 67.3 keV. $^{125}\text{I} + e^- \rightarrow ^{125}\text{Te}^* + \nu_e$
- Other backgrounds com from ^{214}Pb , ^{85}Kr and Elastic scattering of solar neutrinos off atomic electrons.

Signal model:

Account for additional shells (not only K+L)
 \Rightarrow Relative fraction of KK, KL, and LL captures slightly decreased.



Fitting strategy \Rightarrow Binned log-likelihood:

$$\mathcal{L}(\mu_s, \vec{\theta}) = \prod_i^{\text{bins}} \text{Poisson}(N_i, \lambda_i(\vec{\theta}) + n_i^s(\mu_s, \vec{\theta})) \times \prod_j^{\text{constraints}} \text{Gauss}(\theta_j, \mu_j, \sigma_j).$$

Parameter of interest: $2\nu\text{ECEC}$ event rate $A_{\beta\beta}$

$$T_{1/2}^{\beta\beta} = \ln 2 \times \frac{N_A \times \eta_{\text{Xe}} \times \epsilon_{\text{SS}}}{A_{\beta\beta} \times M_A}$$

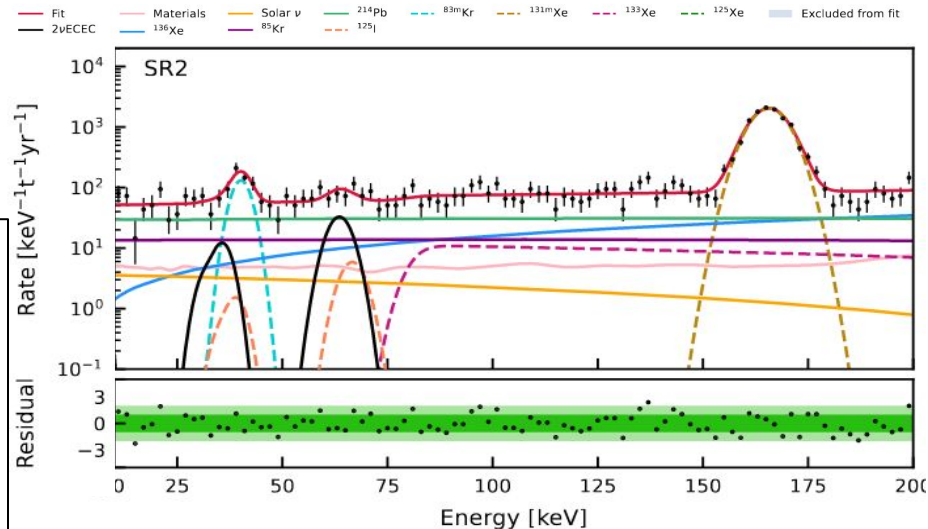
Fit Results:

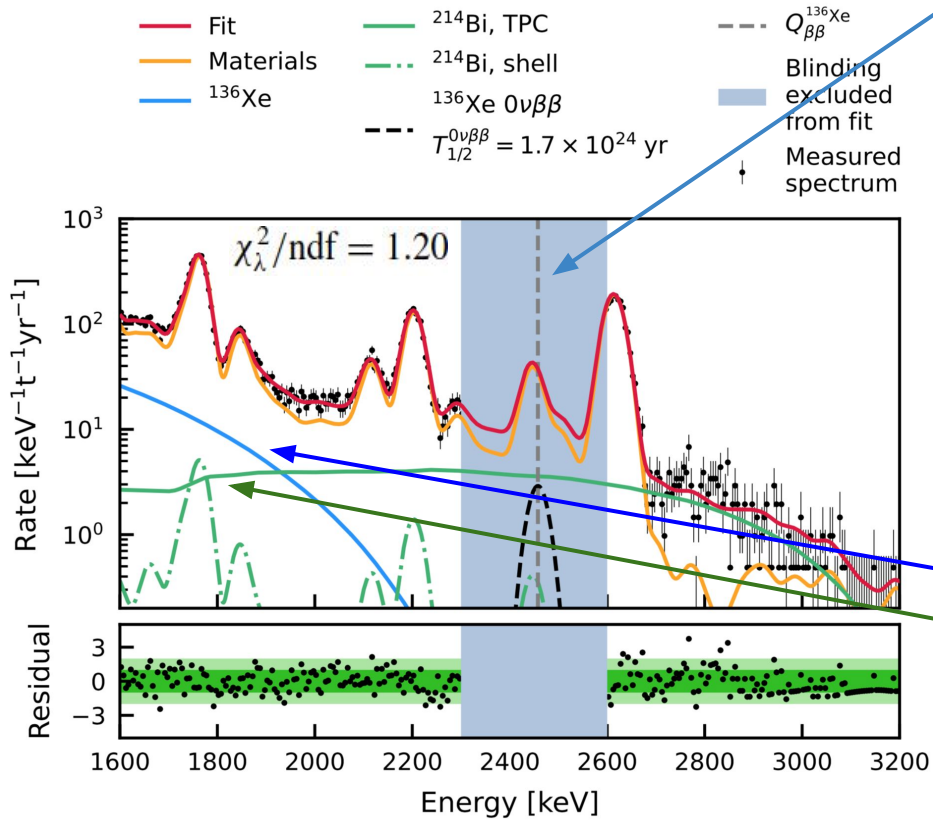
The best-fit double-electron capture rate:

$$A_{2\nu\text{ECEC}} = (300 \pm 50) \text{ t}^{-1} \text{ yr}^{-1}$$

The resulting $2\nu\text{ECEC}$ half-life (first $>5\sigma$ in any isotope):

$$T_{1/2}^{2\nu\text{ECEC}} = (1.1 \pm 0.2_{\text{stat}} \pm 0.1_{\text{sys}}) \times 10^{22} \text{ yr}$$





Background validation performed by excluding the blinded signal region (2300–2600 keV)

Materials (main) background

Originating from ^{238}U and ^{232}Th natural radioactive decays of the material

^{238}U
 $T_{1/2} = 4.5$ Gyr

↓ α BR=100%
Q=4.3MeV

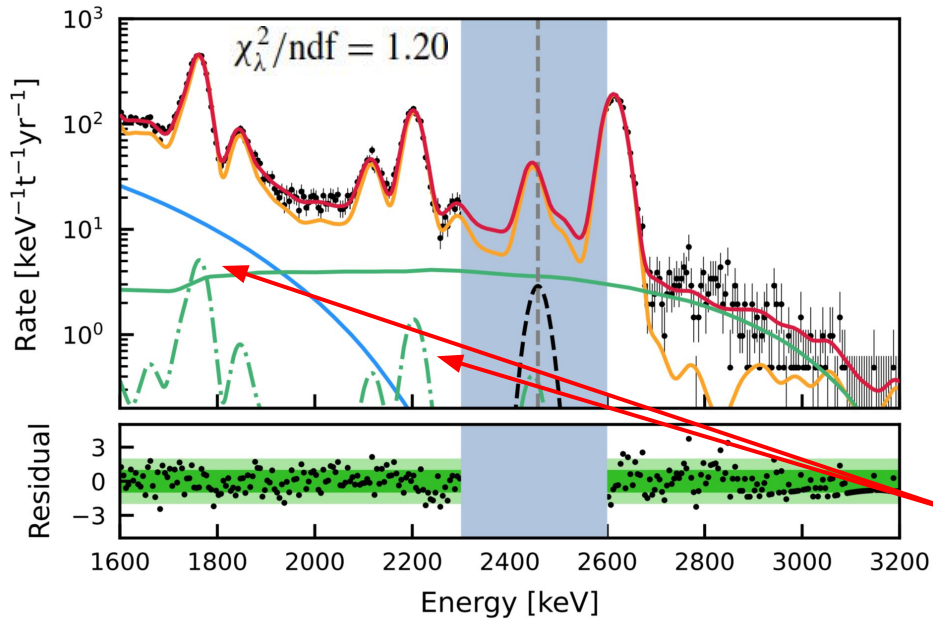
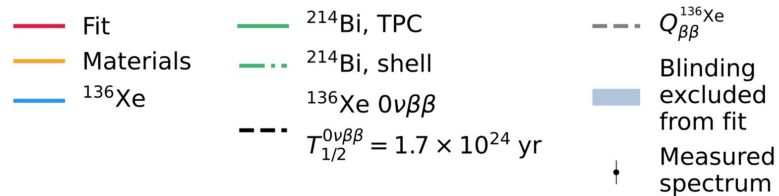
^{234}Th
 $T_{1/2} = 24.1$ d

~ γ BR=2.1% E=96.4keV
~ γ BR=2.1% E=92.8keV

In the form of Compton scatter from radiated γ -rays
 In particular, absorption peaks for

- ^{208}Tl ,
- ^{214}Bi , and
- ^{60}Co γ -lines

Contributions obtained using radioassay



$2\nu\beta\beta$ of ^{136}Xe

Continuous spectrum ending at signal peak
 Modelled from theoretical calc.

^{222}Rn contamination

Rn emanation from materials to LXe target

Internal TPC ^{214}Bi β -decays to ^{214}Po are easily tagged and have >99% rejection efficiency

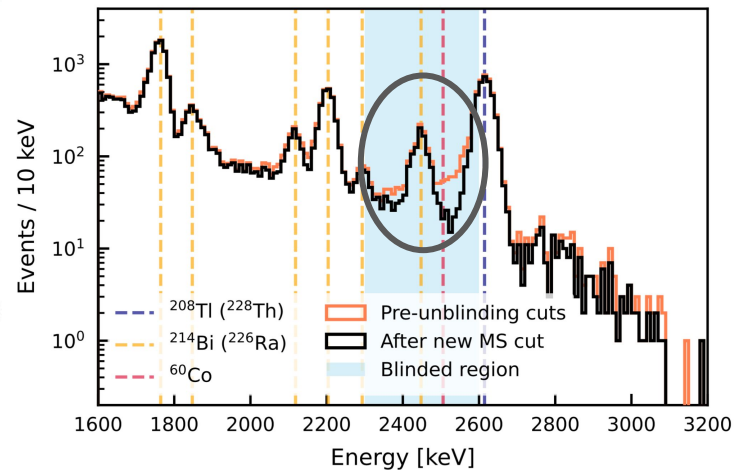
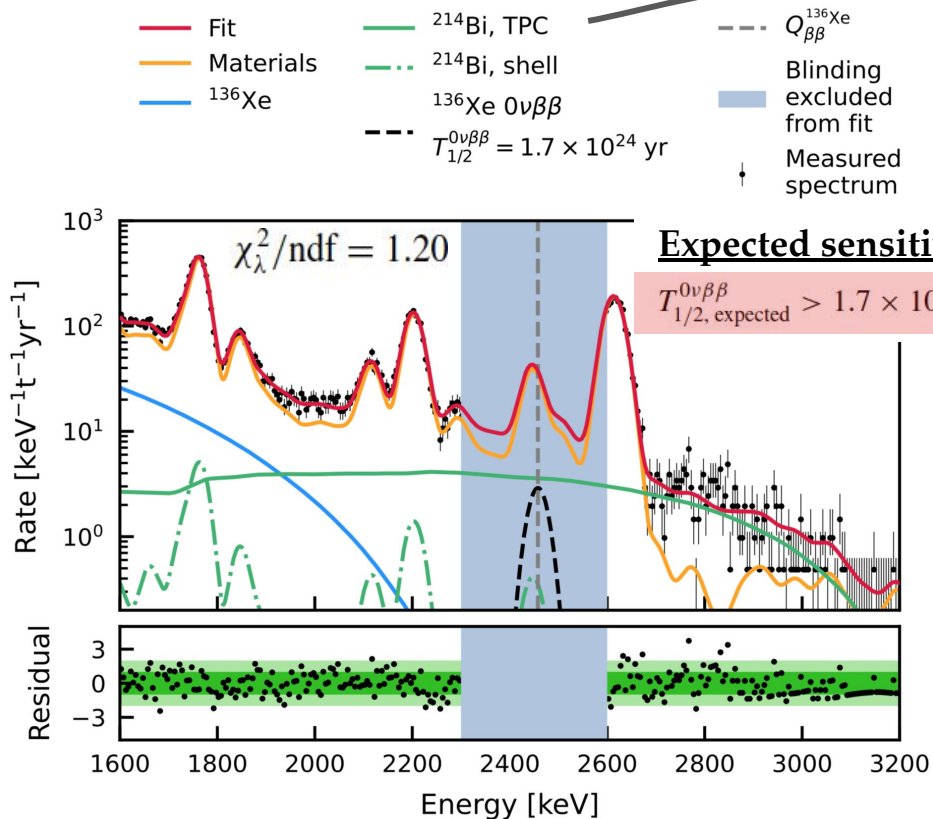
- ^{214}Bi spectrum is continuous

External (shell) ^{214}Bi β -decays with γ -rays in the active volume

- Several ^{214}Bi absorption peaks

Fit to blinded Data

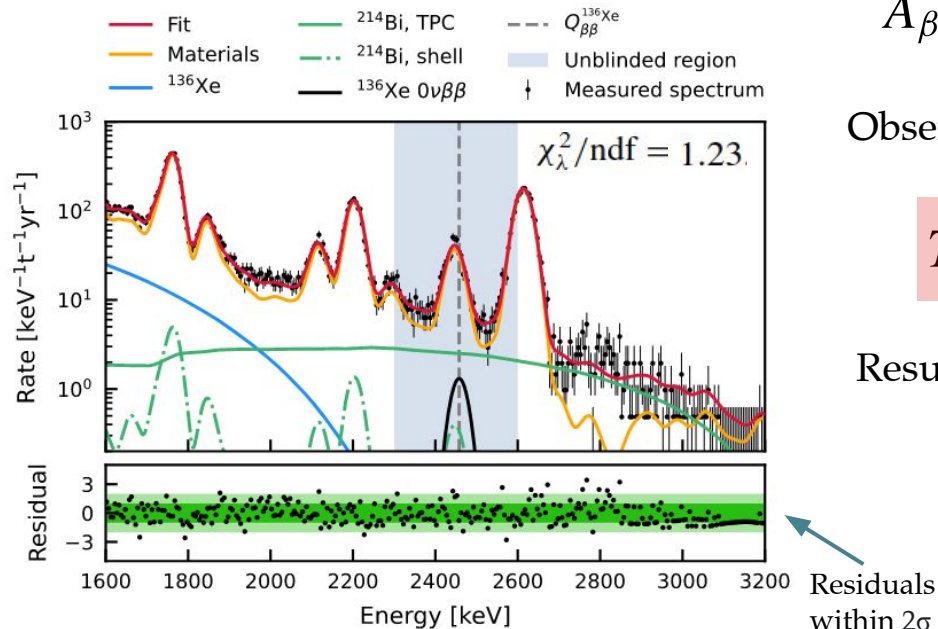
Unblinding



Unexpected background leakage observed in blinded region (~2550 keV).

- **Source:** MS events not rejected by the cuts. \Rightarrow Misidentification as SS events
- An stricter cut on PMT signal collection time cut applied.

Final unblinded result:



Best-fit $0\nu\beta\beta$ rate:

$$A_{\beta\beta} = (65 \pm 87) \text{ t}^{-1}\text{yr}^{-1}$$

$$\langle m_{\beta\beta} \rangle \equiv |\sum_i U_{ei}^2 m_{\nu_i}|$$

$$\langle m_{\beta\beta} \rangle^2 = \frac{m_e^2}{G_{0\nu} |M_{0\nu}|^2 T_{1/2}^{0\nu}}$$

Observed exclusion limit:

$$T_{1/2}^{0\nu\beta\beta} > 1.2 \times 10^{24} \text{ yr at 90\% CL}$$

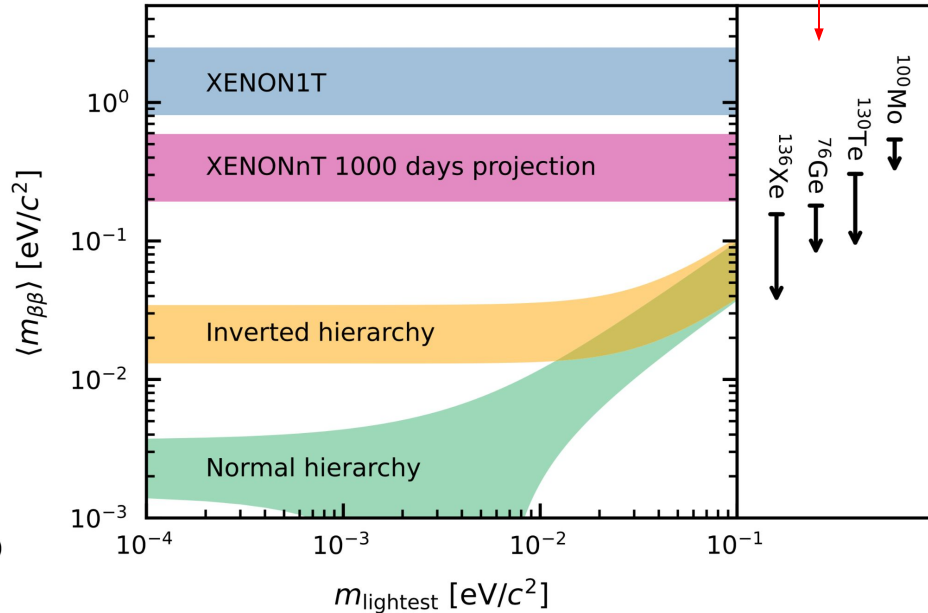
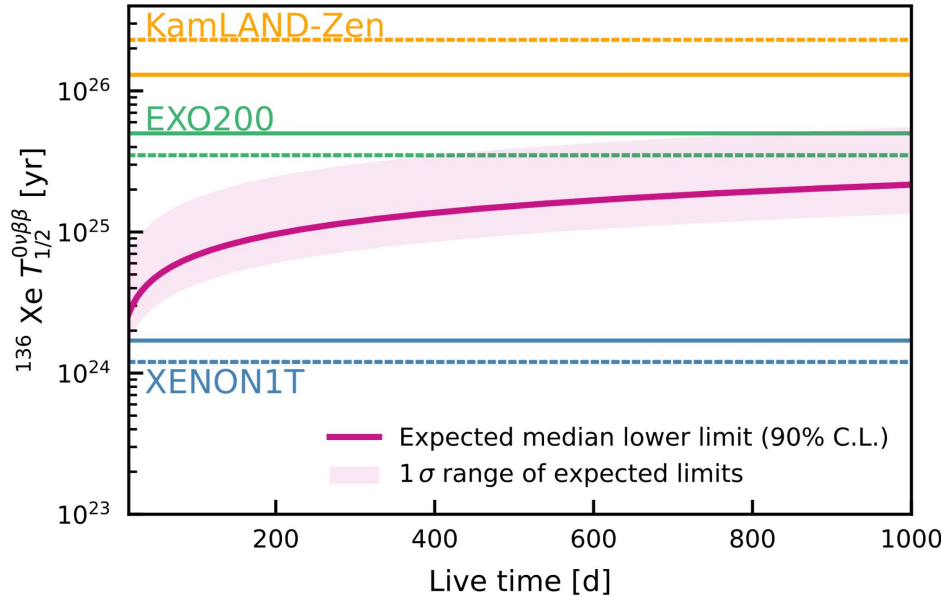
Resulting neutrino mass range:

$$\langle m_{\beta\beta} \rangle < (0.8\text{--}2.5) \text{ eV}/c^2$$

Large uncertainty due to nuclear ME

EXO-200 [2] and KamLAND-Zen [3] supersede this result by 2 orders of magnitude.

Best experimental limits for different isotopes



- XENONnT is not yet competitive with dedicated experiments
- However, future Xe DM detectors can also enter the game

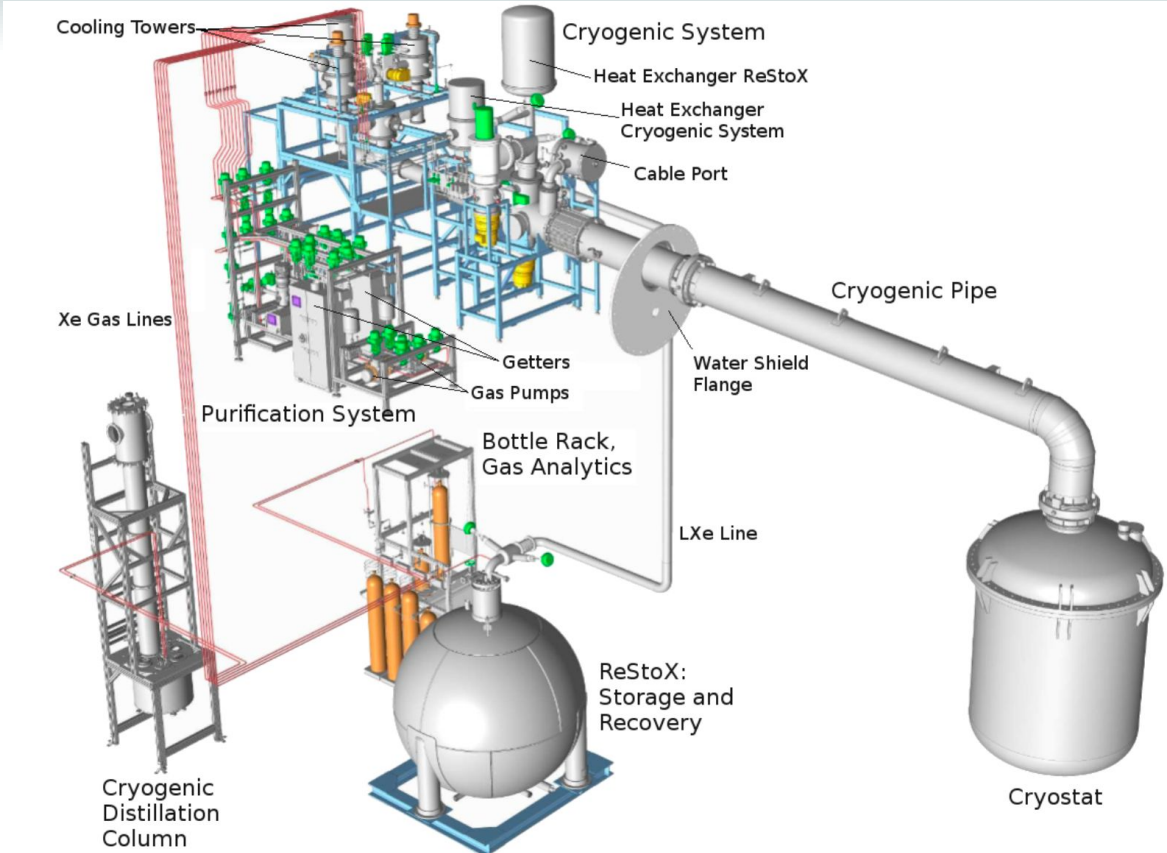
- XENON1T is awesome
 - **Main goal**: Direct detection of Dark Matter
 - Nevertheless, **broader physics program** such as
 - First significant measurement of $2\nu\text{ECEC}$
 - **Searches of $0\nu\beta\beta$** for experimental evidence of Majorana neutrinos
- XENONnT is n times more awesome!
 - Sensitivity studies to $0\nu\beta\beta$ decay \Rightarrow Closer to dedicated experiments
 - Exciting prospects for XENONnT, low background and optimized detection



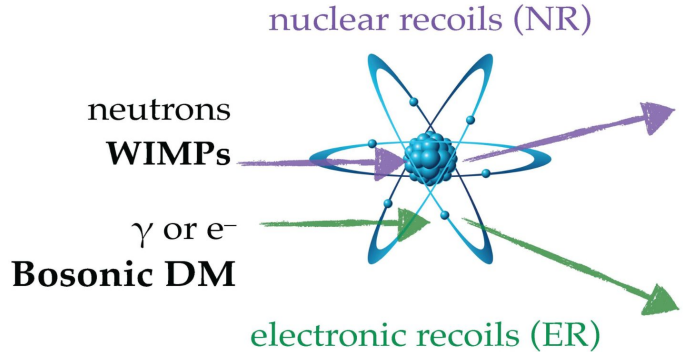
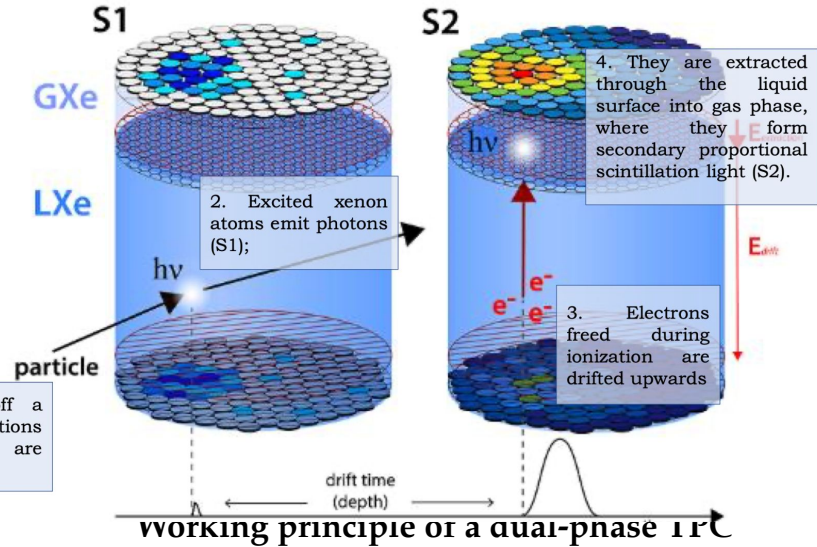
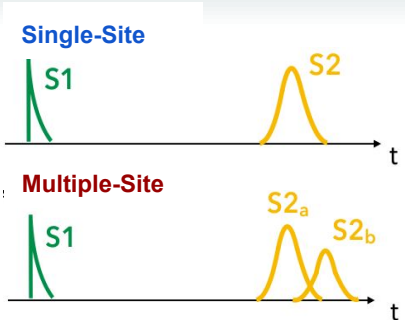
**Thank you very much
for your attention!**

Backup Slides

- [1] E. Aprile et al. (XENON Collaboration), [Eur. Phys. J. C 77, 881 \(2017\)](#).
- [2] G. Anton et al. (EXO-200 Collaboration), [Phys. Rev. Lett. 123, 161802 \(2019\)](#).
- [3] S. Abe et al. (KamLAND-Zen Collaboration), [arXiv:2203.02139](#)
- [4] Murra, M., et al. "Design, construction and commissioning of a high-flow radon removal system for XENONnT." arXiv preprint arXiv:2205.11492 (2022).

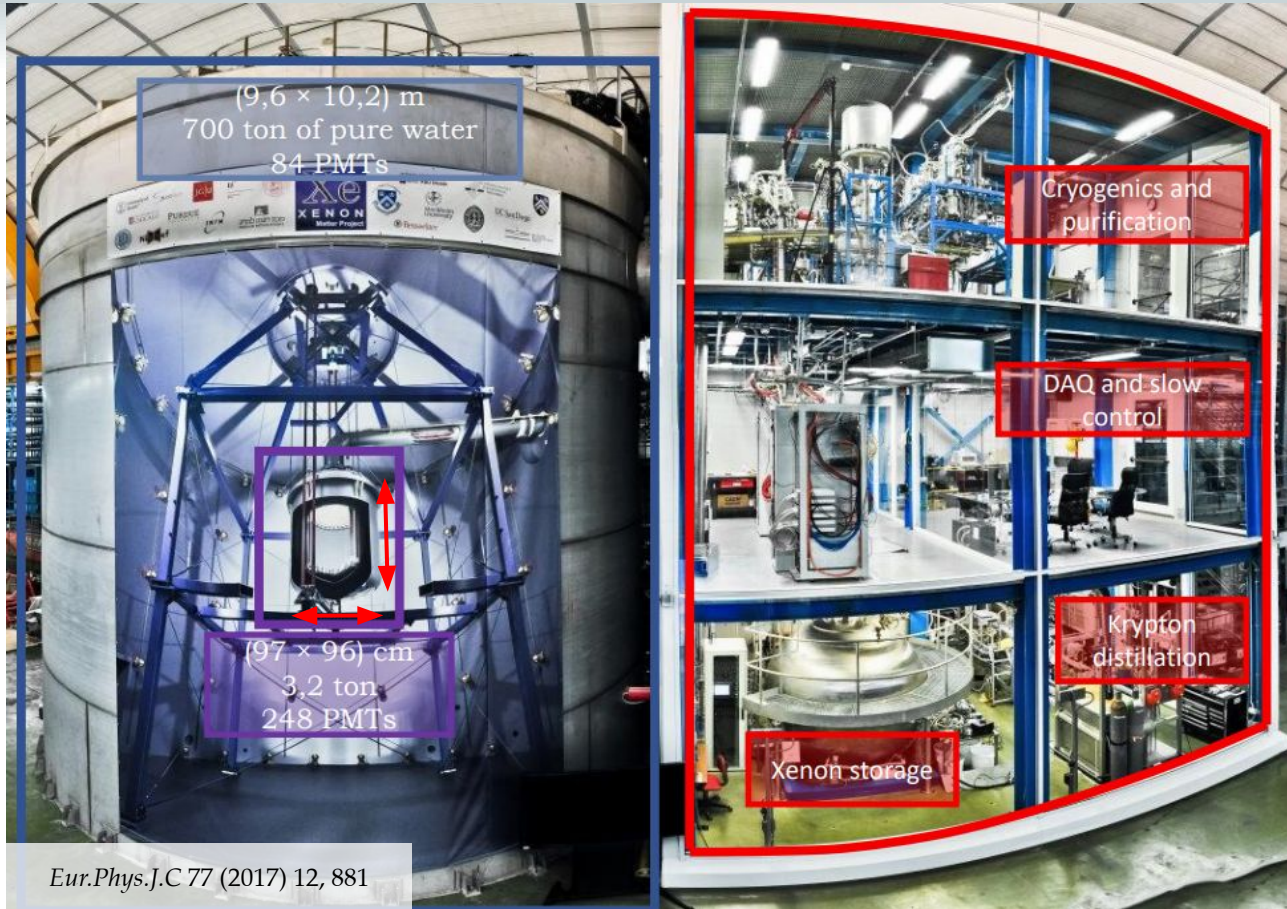


1. Recoil of a *particle* with the Xe atom excite and ionize which causes a release of photons and e^-
 - 1.1. *WIMP, neutron* or ν interacts with LXe: Nuclei scatter \rightarrow Nuclear recoil (NR)
 - 1.2. γ rays, charged particles or ν_e interacts with LXe: Atomic electrons \rightarrow Electronic recoil (ER)
2. First scintillation signal S1 is emitted from the first photons of the excited Xe atom that reach on the top and the bottom PMTs
3. Ionized electrons go upwards with the electric field and drift to the liquid-gas
4. Second scintillation signal S2 is emitted in GXe
5. Events can be reconstructed in 3D using the drift time between S1 and S2 to reconstruct the z coordinate and the S2 hit pattern on the top PMT array for the (x,y) coordinates.



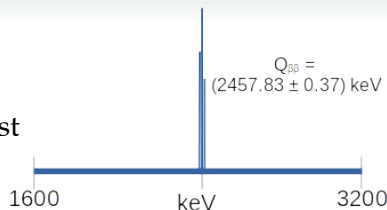
$$\frac{S2}{S1} (ER) > \frac{S2}{S1} (NR)$$

A particle interacting with LXe



Dataset:

- Scan energies between 1600 and 3200 keV
 - $0\nu\beta\beta$ peak at $Q_{\beta\beta} = (2457.83 \pm 0.37)$ keV
- Isotope exposure of 36.16 kg yr (current most sensitive – KamLand-ZEN 89.5 kg yr)



Cuts:

- Data cleaning cuts (busy periods, light emission etc.)
- Compton background removal
 - Shows up as **MS** events as opposed to $0\nu\beta\beta$ as **SS** events.
- Remove pileup events
 - example: $^{214}\text{BiPo}$ decay - shows up as multiple S1 signals
- Apply position reconstruction algorithm - Neural Net + S2 shape fit
 - removes events near edge of TPC or where bad PMTs were located
- Electron Recoils removed - 98% efficiency.
 - Done using S1/S2 ratio
- Restrict to events where S2 shape matches what we expect from drifted electron diffusion

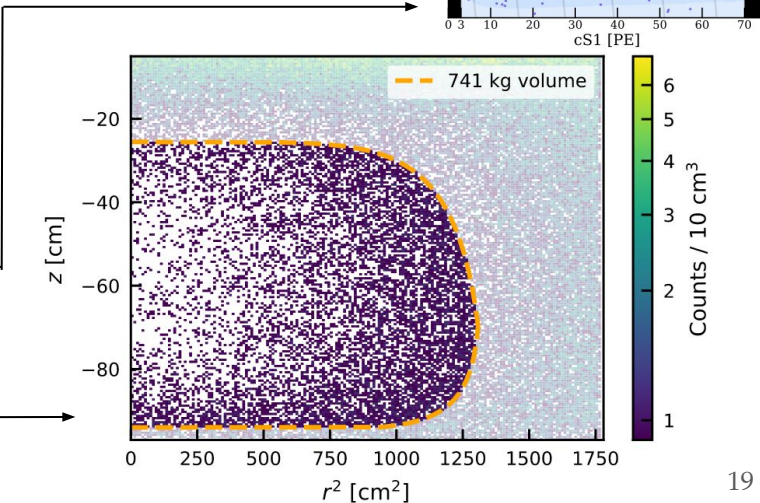
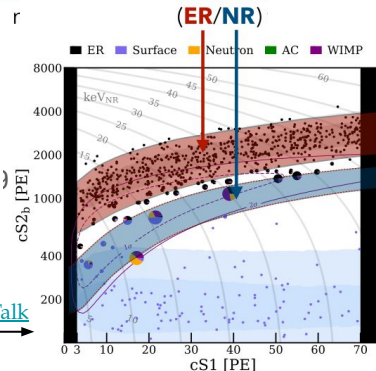
Single-Site



Multiple-Site



Ref: [Xenon1T Public Talk](#)



Signal Acceptance Optimization

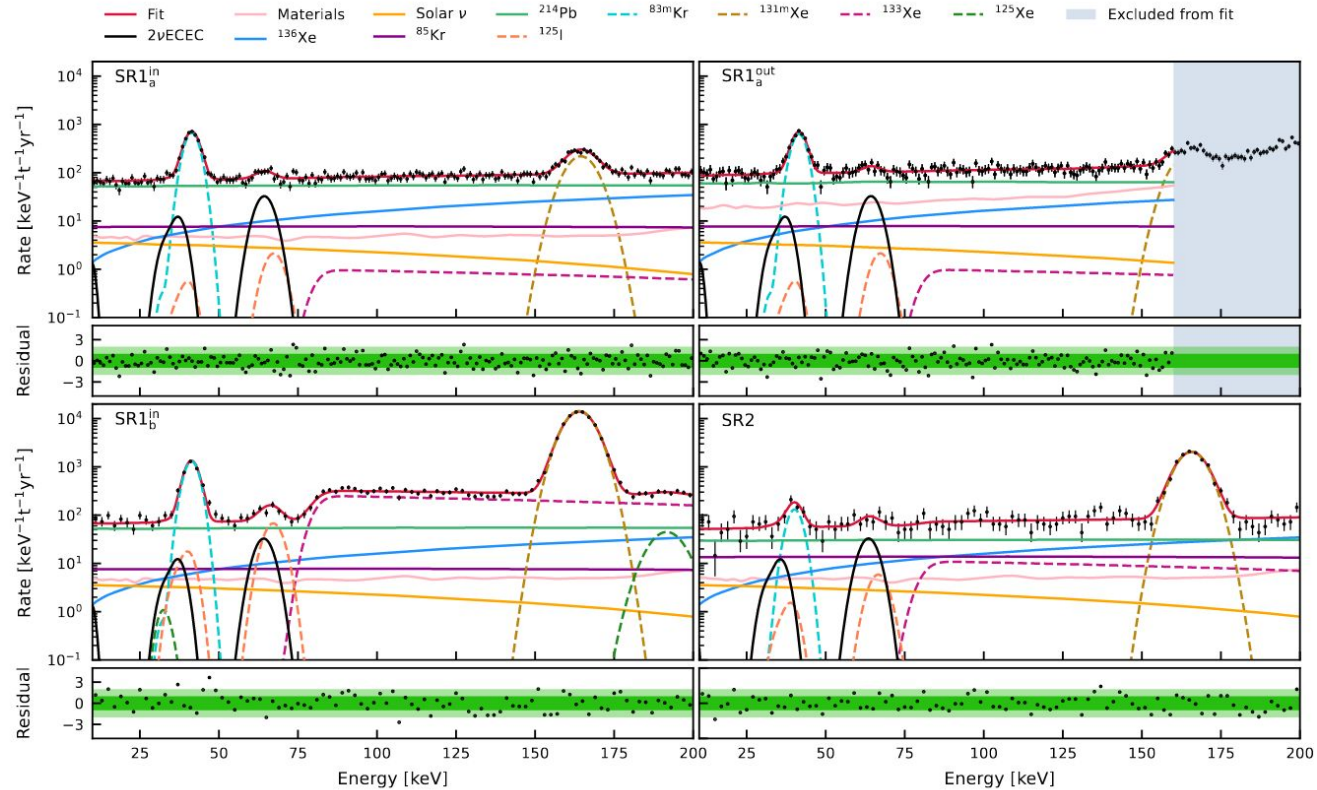
- An FV (Fiducial Volume) in the detector is chosen that maximises signal-to-noise ratio:

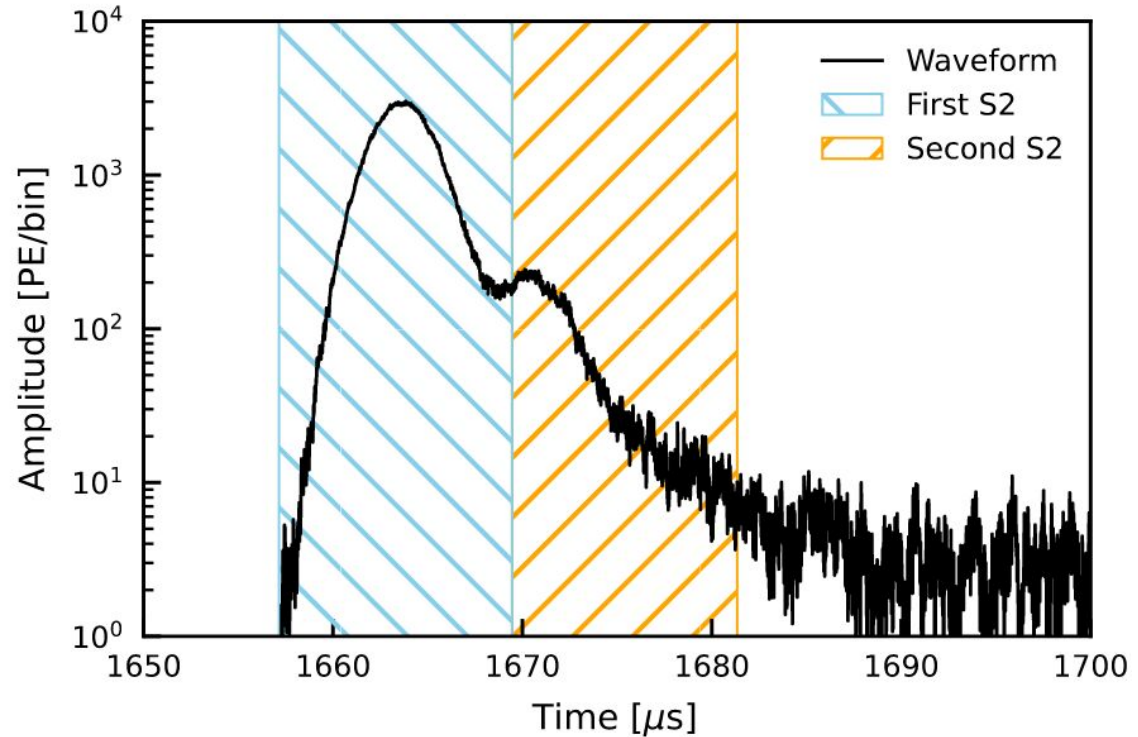
Mass of isotopes

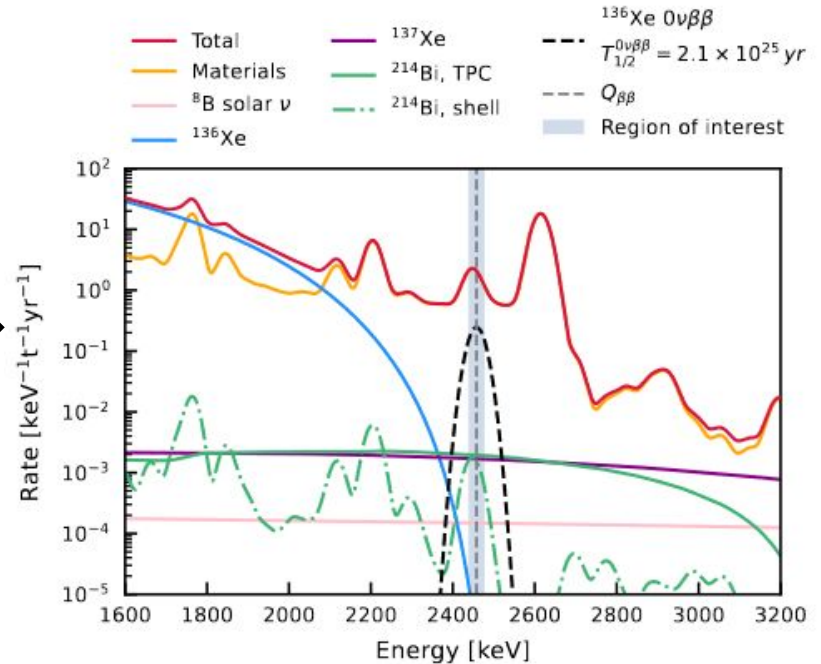
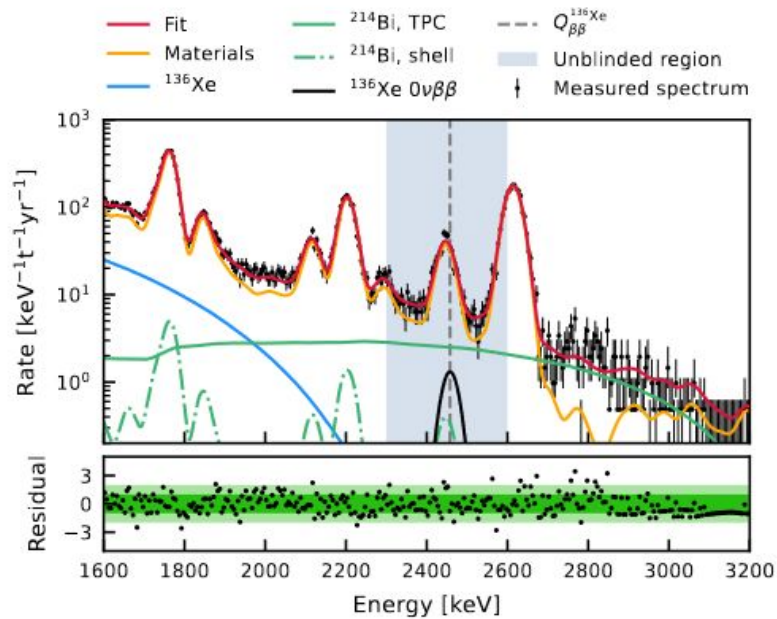
	mass (g/mol)	delta m (keV)		mass (g/mol)	delta m (keV)
Te124	123.9028179	-2864.437522	Te136	135.9201	11998.57557
I124	123.9062099	295.1904819	I136	135.91465	6921.932695
Xe124	123.905893	0	Xe136	135.907219	0
Cs124	123.912258	5928.959979	Cs136	135.9073116	86.25635414
Ba124	123.915094	8570.677261	Ba136	135.9045759	-2462.032069

EXO-200 and KamLAND-Zen supersede this result
by 2 orders of magnitude

- EXO-200 [2]: $T_{1/2} > 3.5 \times 10^{25}$ yr $\langle m_{\beta\beta} \rangle < (93 - 286)$ meV
- KamLAND-Zen [3]: $T_{1/2}^{0\nu} > 2.3 \times 10^{26}$ yr at 90% C.L. $\langle m_{\beta\beta} \rangle < (36 - 156)$ meV



Limits $0\nu\beta\beta$ XENON1T: Background leakage into unblinded region



Background sources

- Main background: radioactive impurities in the detector material, comprised of several radioactive decay chains
 - ^{238}U , ^{232}Th , ^{40}K , ^{60}Co
 - Their contributions are obtained using radioassay
- Intrinsic LXe backgrounds also play a (subdominant) role
 - Comes from ^{222}Rn emanation from detector materials to LXe target
 - Composed of ^{214}Pb and ^{214}Bi isotopes, >99% rejection using timing information
- Background modelling is validated in a fit that excludes the blinded signal region (2300–2600 keV)

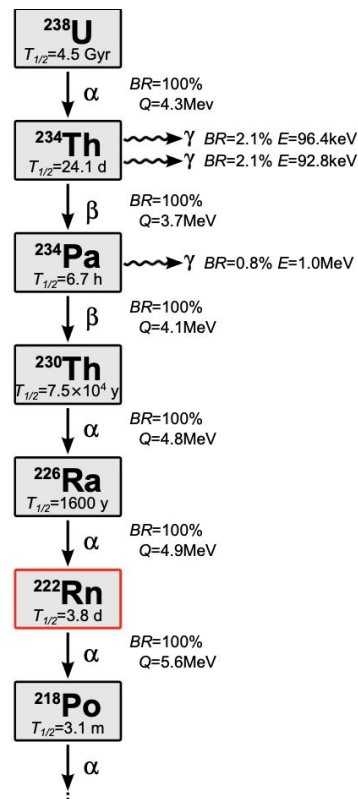


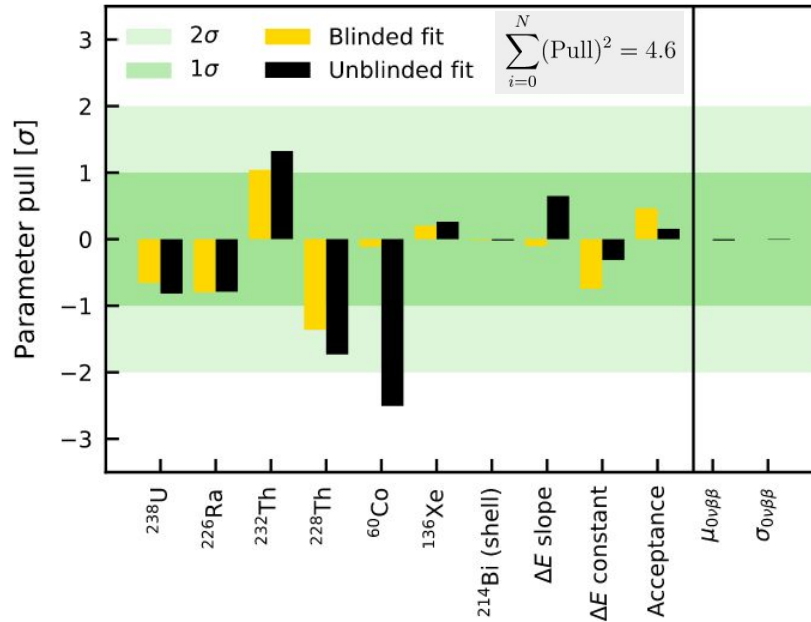
TABLE VII. Best-fit signal and background model parameters for the XENON1T 2ν ECEC search that were shared among all datasets. Unitless parameters state the relative change to the expected value. The meaning of the parameters is given in Sec. III D.

Parameter	Fit value	Constraint	Unit
$A_{2\nu\text{ECEC}}$	300 ± 50	–	$\text{t}^{-1}\text{yr}^{-1}$
Solar ν	1.00 ± 0.02	1.00 ± 0.02	–
^{136}Xe	0.99 ± 0.03	1.00 ± 0.03	–
^{238}U	1.0 ± 0.5	1.0 ± 0.6	–
^{226}Ra	0.5 ± 0.3	1.0 ± 0.5	–
^{232}Th	0.9 ± 0.6	1.0 ± 0.6	–
^{228}Th	0.9 ± 0.6	1.0 ± 0.6	–
^{60}Co	0.6 ± 0.3	1.0 ± 0.4	–
^{40}K	1.0 ± 0.3	1.0 ± 0.3	–
$\mu_{83\text{mKr,misID}}$	32.1 ± 0.6	32.1 ± 0.6	keV
$\sigma_{83\text{mKr,misID}}$	1.3 ± 0.2	1.3 ± 0.2	keV
$f_{83\text{mKr,misID}}$	2.6 ± 0.4	2.6 ± 0.4	10^{-4}

TABLE IX. Parameter constraints and best-fit parameters for the XENON1T $0\nu\beta\beta$ search. Best-fit parameters are given for the blinded and unblinded data. Unitless parameters state the relative change to the expected value. For the acceptance, the constrained and best-fit acceptances are given instead of the fit parameters. Due to the implementation of the acceptance, the Gaussian constraint on the acceptance scaling parameter yields an asymmetric acceptance range. Due to the internal processing of the thorium chain, the respective best-fit parameters cannot be translated directly to ^{228}Th event counts in Tab. V.

Parameter	Unit	Constraints	Blinded fit value	Unblinded fit value
^{238}U	–	1.0 ± 0.6	0.6 ± 0.1	0.5 ± 0.1
^{226}Ra	–	1.0 ± 0.5	0.620 ± 0.008	0.624 ± 0.008
^{232}Th	–	1.0 ± 0.6	1.6 ± 0.1	1.8 ± 0.1
^{228}Th	–	1.0 ± 0.6	0.2 ± 0.1	0.0 ± 0.1
^{60}Co	–	1.0 ± 0.4	1.0 ± 0.3	-0.1 ± 0.1
^{136}Xe	–	1.00 ± 0.03	1.01 ± 0.03	1.01 ± 0.03
^{214}Bi , TPC	$\mu\text{Bq/kg}$	–	1.7 ± 0.1	1.25 ± 0.09
^{214}Bi , LXe shell	$\mu\text{Bq/kg}$	10 ± 5	10 ± 5	10 ± 5
ΔE_{slope}	–	$(1.5 \pm 0.2) \times 10^3$	$(1.5 \pm 0.1) \times 10^3$	$(1.6 \pm 0.1) \times 10^3$
ΔE_{offset}	keV	-4.4 ± 0.3	-4.6 ± 0.3	-4.5 ± 0.2
ϵ	%	$88.6^{+8.9}_{-0.3}$	$88.3^{+0.6}_{-0.1}$	$85.2^{+4.6}_{-0.3}$
$\mu_{0\nu\beta\beta}$	keV	2457.8 ± 0.4	–	2457.8 ± 0.4
$\sigma_{0\nu\beta\beta}$	keV	19.7 ± 0.3	–	19.7 ± 0.3
$A_{0\nu\beta\beta}$	$\text{t}^{-1}\text{yr}^{-1}$	–	–	60 ± 90

Fit to blinded data



Pulls:

- No pulls of the blinded fit $>2\sigma$.
- The pull on ^{60}Co - close to 0.
- No notable pulls on the systematic uncertainty parameters.