



Orbit corrector budget for HL-LHC v1.5

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General assumptions

- **HL-LHC Optics V1.5**
 - **7 TeV; $\beta^* = 15\text{cm}$ round; $\epsilon_N = 2.5 \mu\text{m}$; $\delta_p = 1.1\text{e-}4$**
- Here shown only at **right side of IP5**
 - **V crossing** and H separation
 - *for H crossing (IP1) and left side symmetries apply*
 - **Residual orbit** and corrector strength given in **2*r.m.s.**
 - *Standard approach considered to fit LHC experience ([Chamonix08](#))*
- **Note:** results obtained with **Python framework** by Joel:
 - Assuming **fully linear optics**
 - Most computations using **SVD inversion of response matrices** generated from Twiss functions
 - Orbit correction at **BPMs only**
 - **Framework source code** and **examples:** [POCKPy on GitLab](#)
 - Also **documented** in Joel's **master thesis:** *A Linear Framework for Orbit Correction in the High-Luminosity Large Hadron Collider* - [link](#)

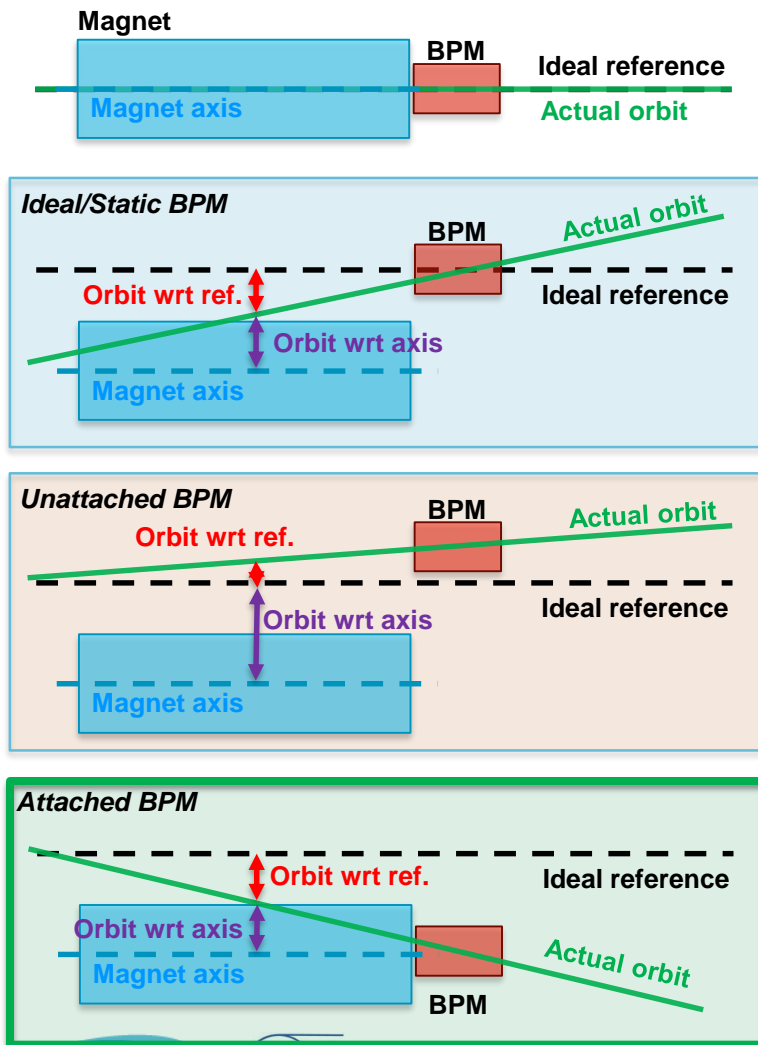
Orbit Correction and Corrector Budget

Orbit correction due to errors

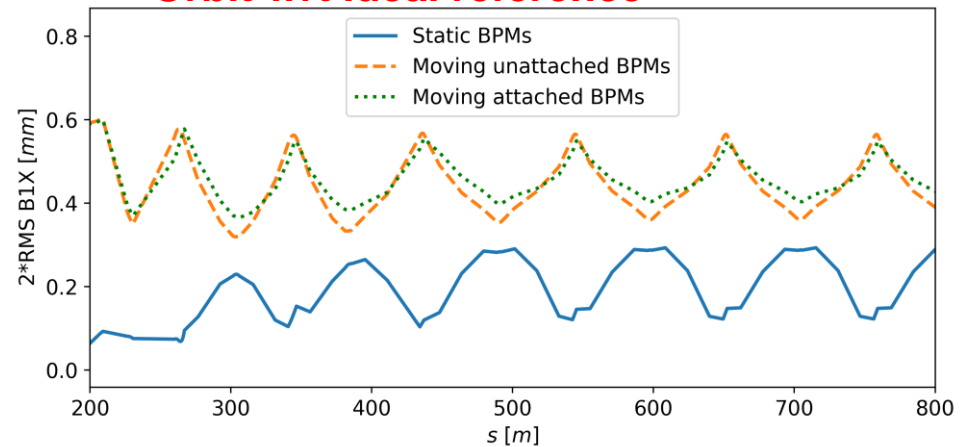
- **Considered element errors** (all square distributions):
 - **Quadrupoles:**
 - Transverse offset: $\pm 0.5 \text{ mm}$
 - Rotation (DPSI): $\pm 1 \text{ mrad}$
 - Relative field strength error: $\pm 0.2\%$
 - **Dipoles:**
 - (NEW) Transverse offset: $\pm 0.5 \text{ mm}$
 - *Used only to give orbit w.r.t. center of magnet and nearby BPM*
 - Rotation (DPSI): $\pm 0.5 \text{ mrad}$
 - Relative field strength error: $\pm 0.2\%$
 - **BPMs:**
 - (NEW) Transverse offset: $\pm 0.5 \text{ mm}$
 - Several cases considered, **finally assumed to move with nearby quadrupole**
- **Missing errors** w.r.t. previous studies:
 - **Longitudinal misalignment**
 - Not easy to implement in present (analytical) framework
 - **Deemed to be negligible** in previous studies
- **Important remark:**
 - Still using “**standard**” numbers for **expected errors**
 - **Update numbers** may come from **WG Alignment** ([espace](#))
 - **Natural entanglement** between what is **desirable** and what is **achievable!**

Residual orbit post-correction (in the arc)

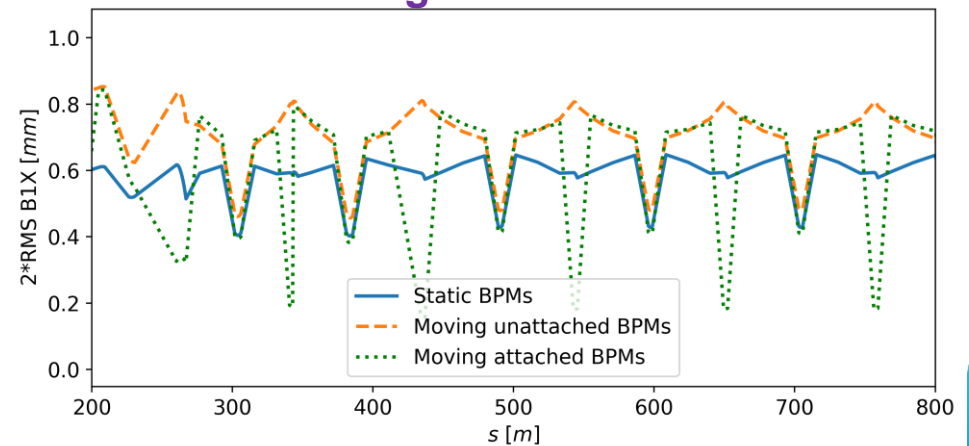
- Depending on assumption on BPM behavior, one gets different results.
 - Note:** correcting such to minimize orbit wrt center of all BPMs with the same weight



Orbit wrt ideal reference



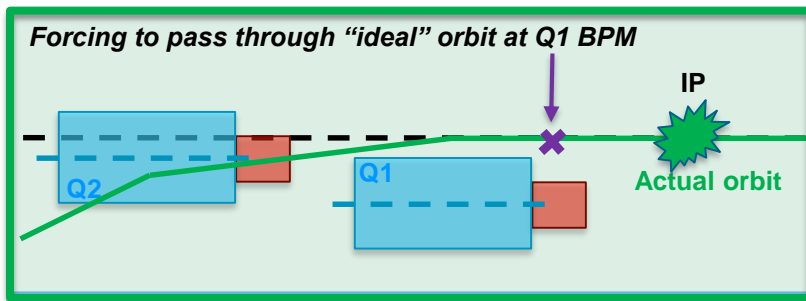
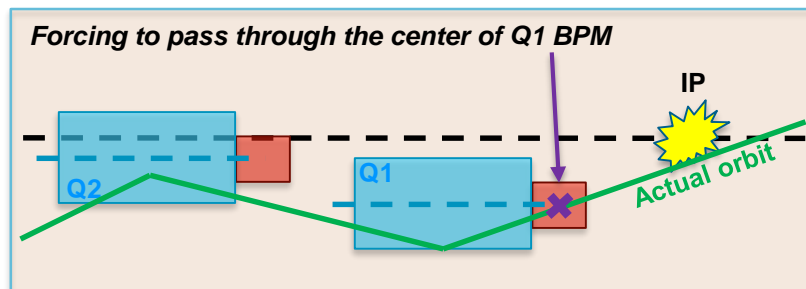
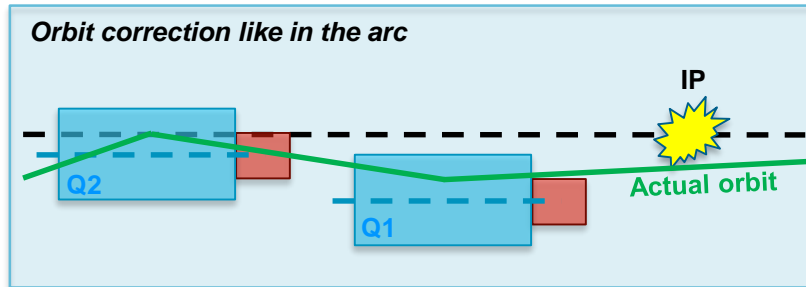
Orbit wrt magnet axis



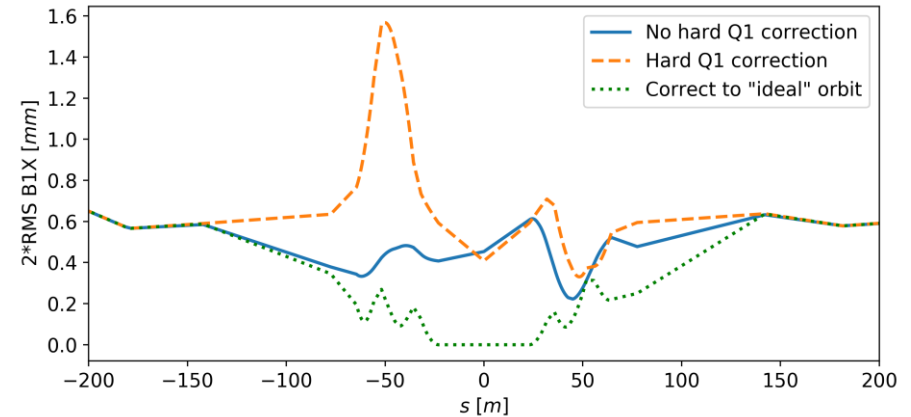
Assumption used in the following slides

Orbit correction strategy in the triplet

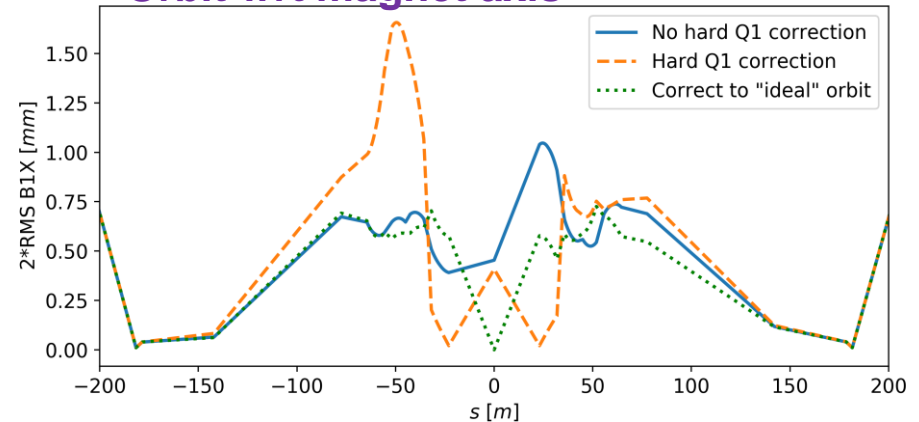
- Need to assume a strategy to “define”/ “find” ideal IP position



Orbit wrt ideal reference



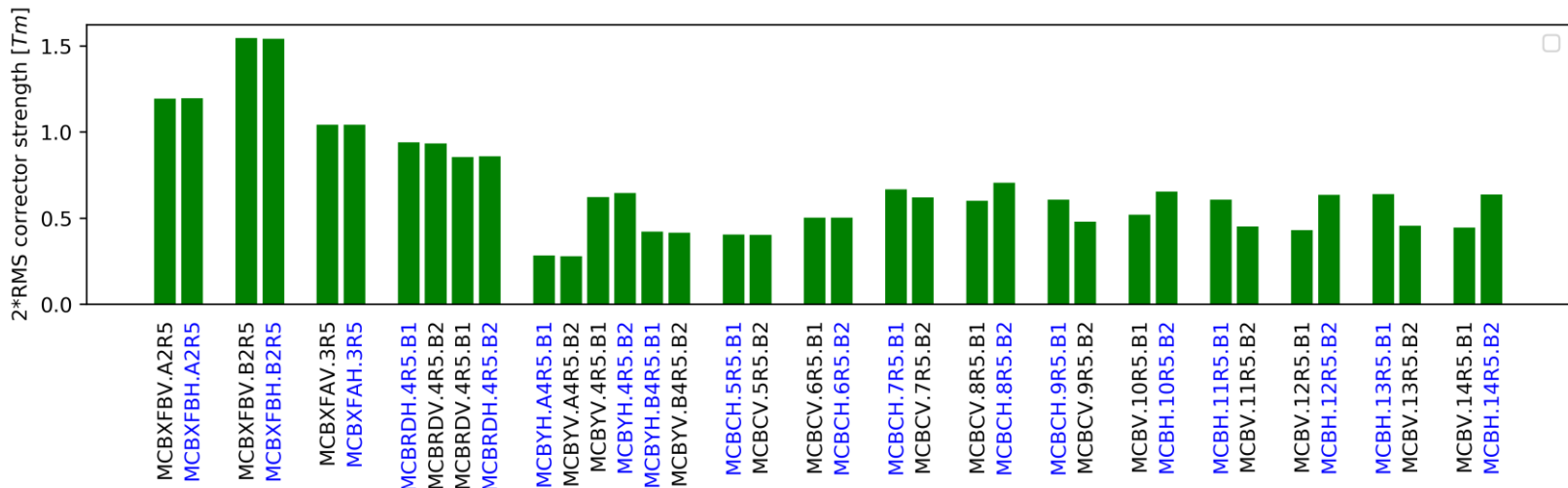
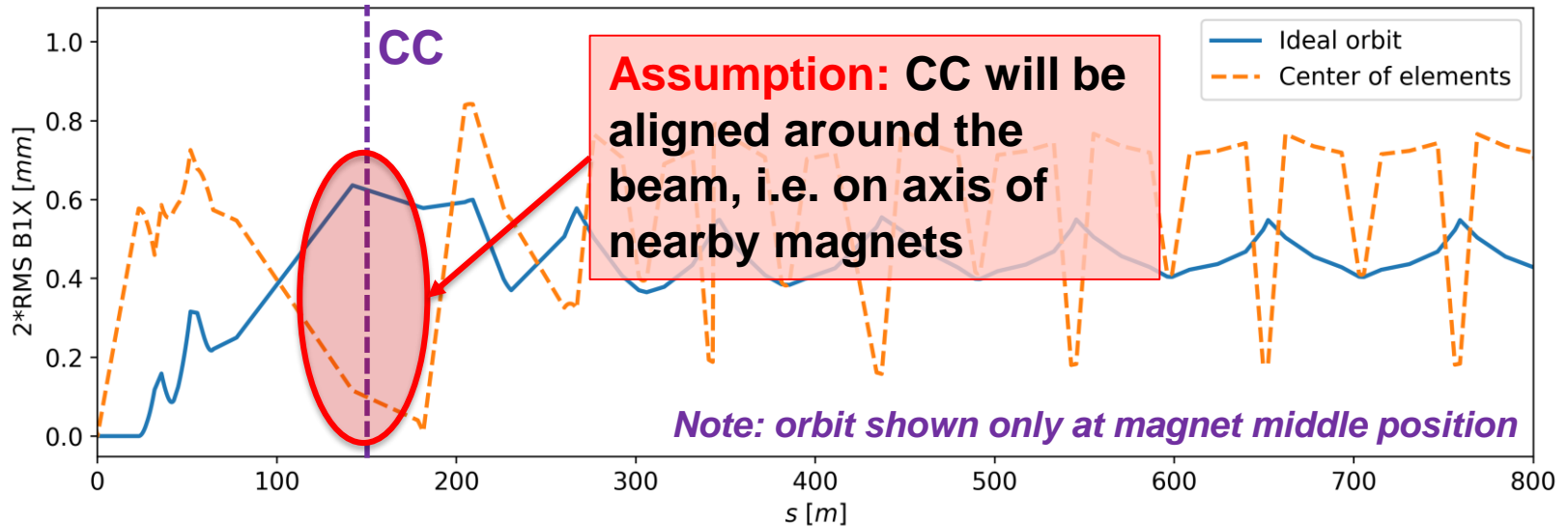
Orbit wrt magnet axis



Imagining that ballistic optics could give us a good “ideal” reference at Q1 BPM

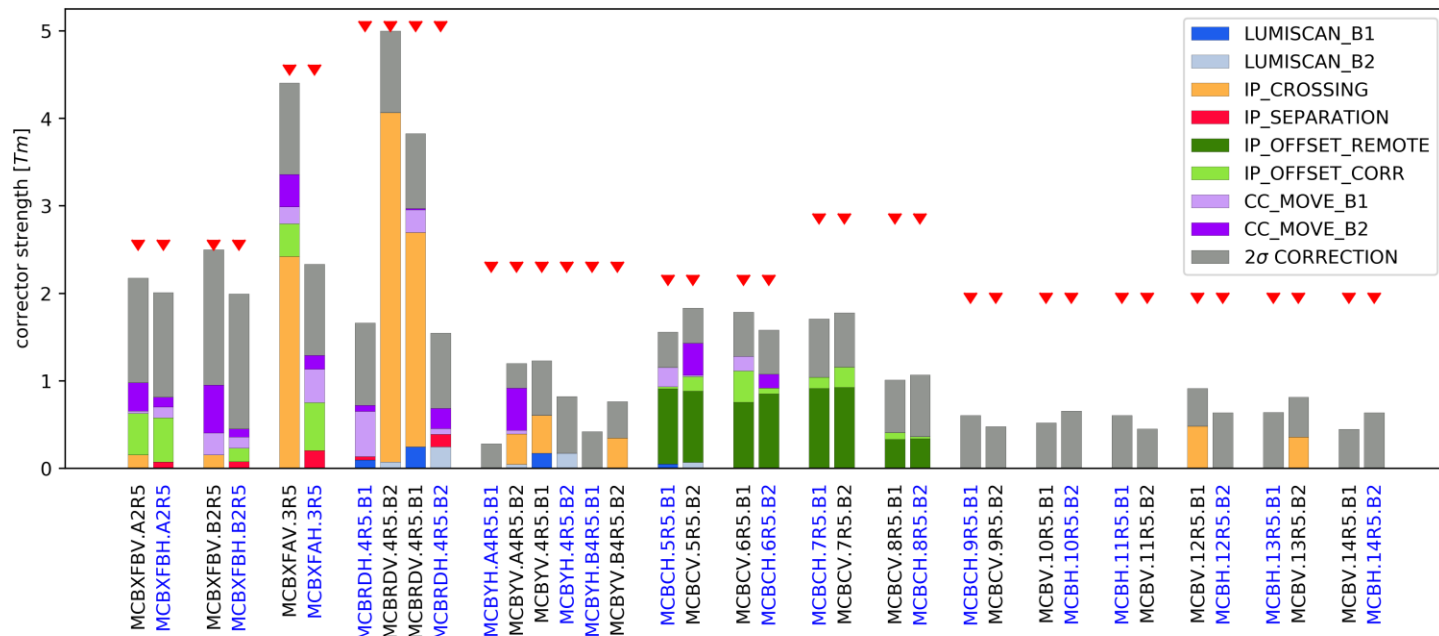
Resume: Residual Orbit and Correctors Usage

- BPMs move with nearby magnet + “strong” correction to get beam at ideal IP



Knobs implementation

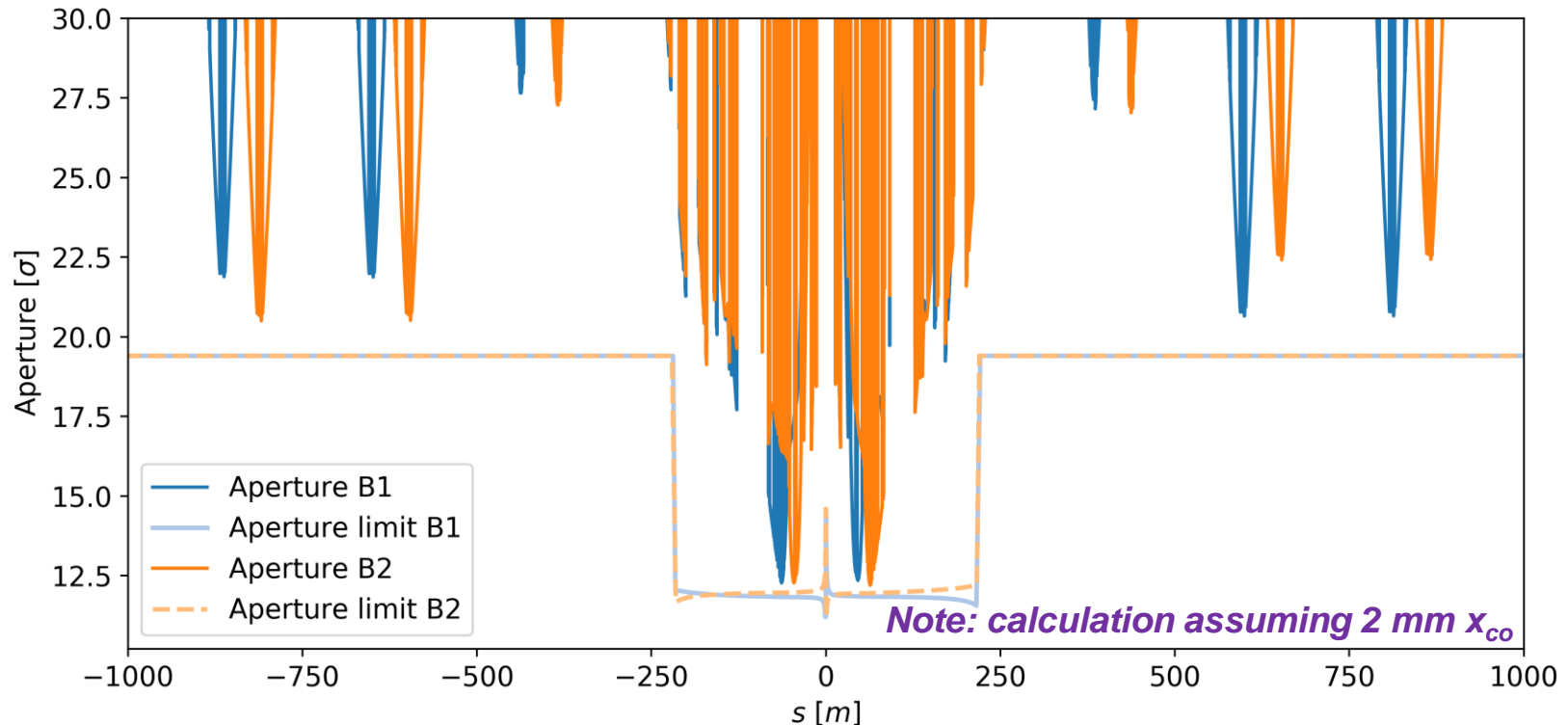
- 295 μrad crossing angle in V plane
 - (Made of 80% “short” + 20% “long” official versions - (**~ 0.66 mm at CC**))
- ± 0.75 mm separation in H plane
- 100 μm IP movement independent for B1/B2 for **lumiscan**
- 2 mm IP offset with correctors + remote alignment
 - Q1-Q4 displaced by 2 mm; Q5 displaced by 1 mm**
- ± 500 μm IP offset with orbit corrector only (requires **~ 1 mm CC re-alignment**)
- ± 500 μm movement independent for B1/B2 for **CC alignment**



Failure Scenarios

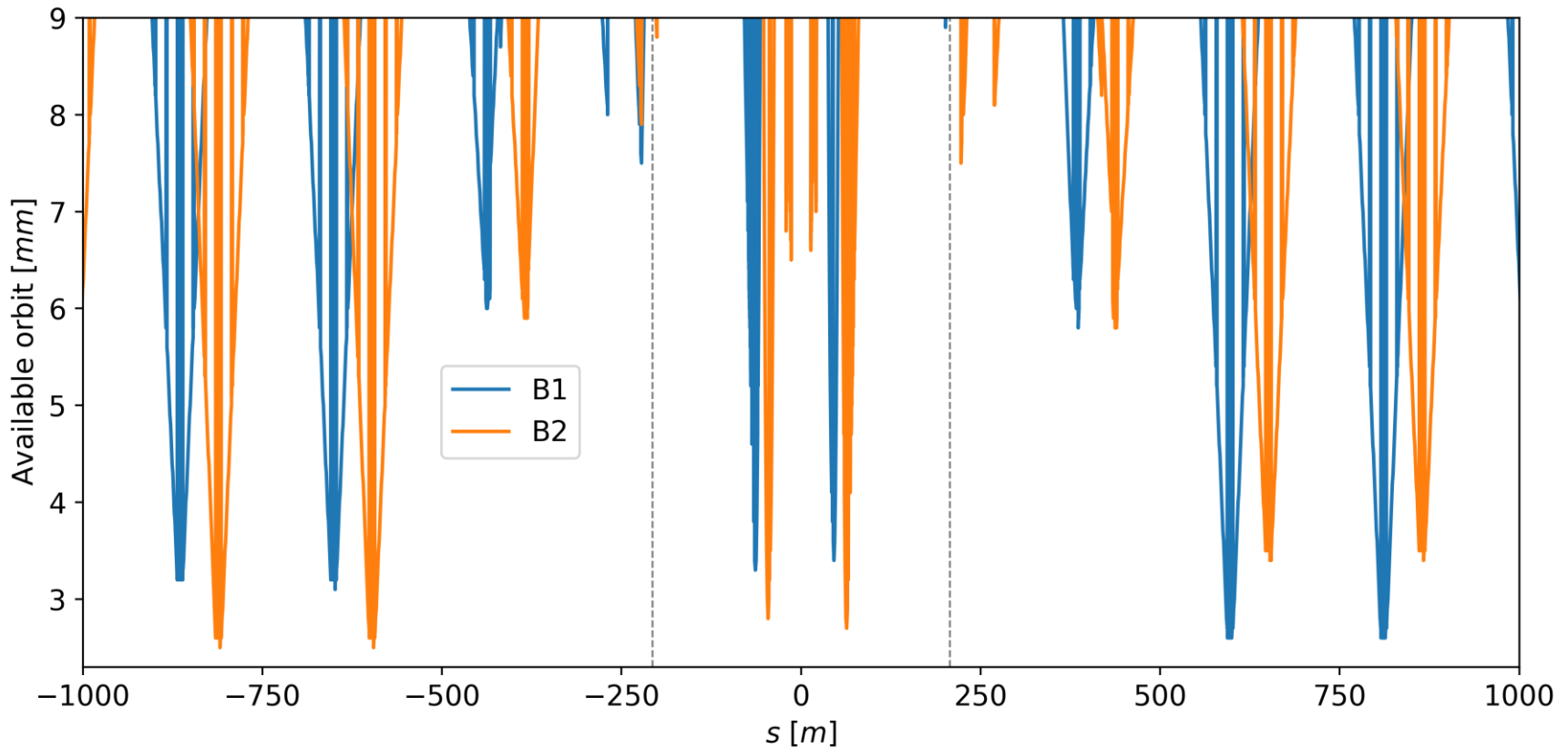
Aperture considerations

- Previously, e.g. [162th WP2](#), required aperture limits were $20/13.2\sigma$ in the arc/triplet.
- Here, using 19.4σ in the arc and modulated ($\sim 12\sigma$) limit in triplet according to [CERN-ACC-2017-0051](#)
 - Could probably apply also to the arc, but to be crosschecked.



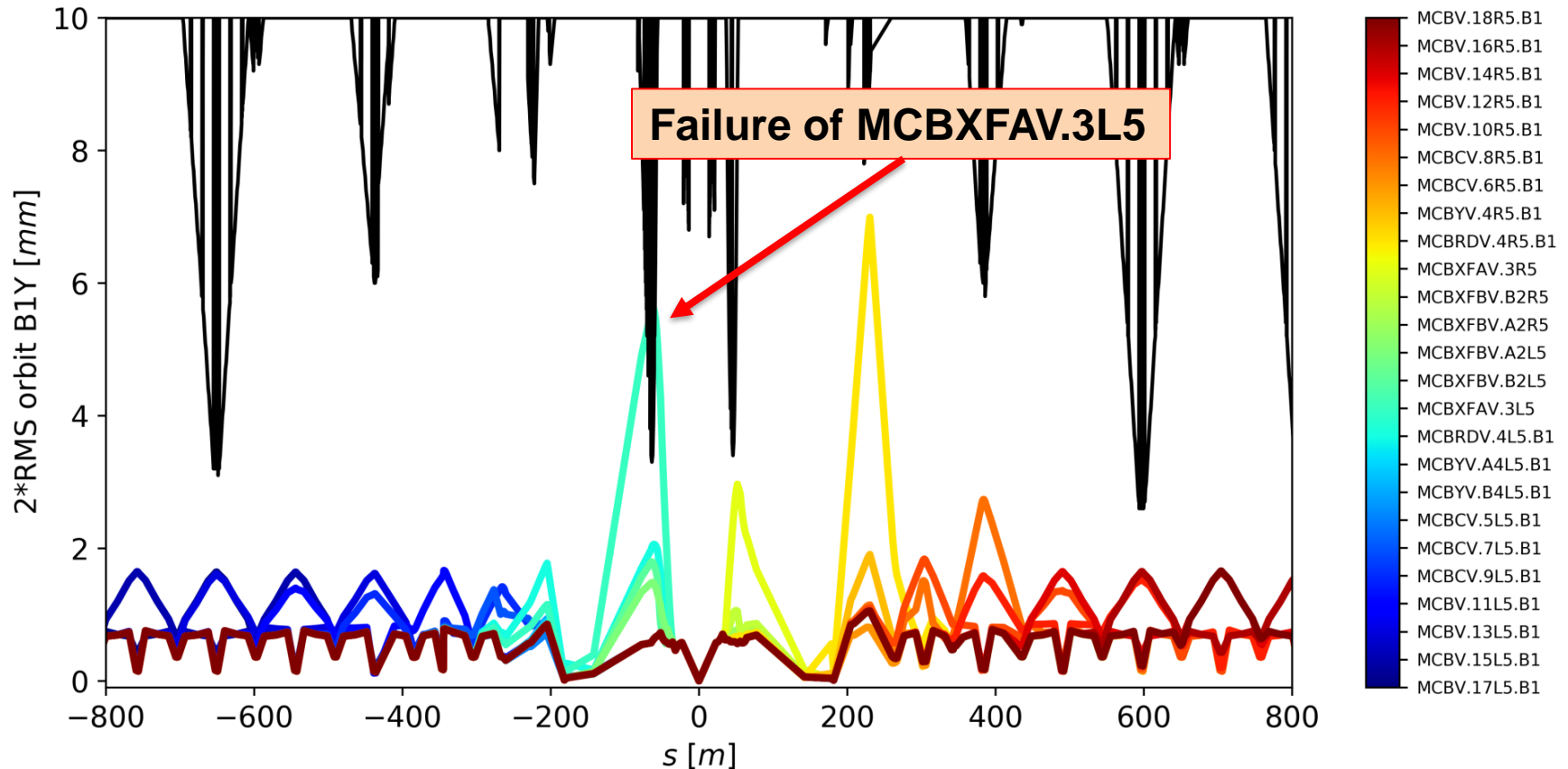
Available “aperture” for orbit

- **Scanning over x_{co}** (default 2 mm) one can get the **radial orbit clearance**, wrt to target aperture.
 - **Conservative approach**, but not too far from reality.



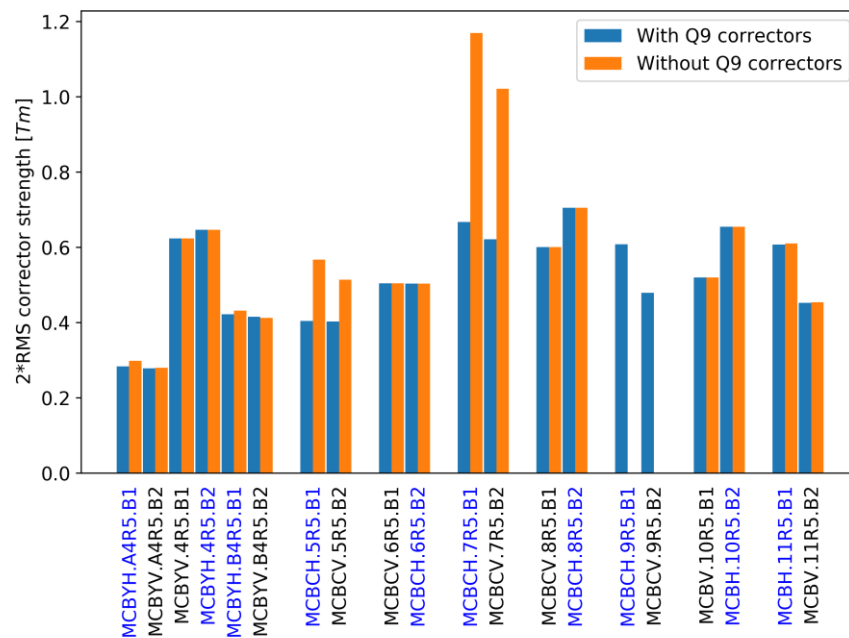
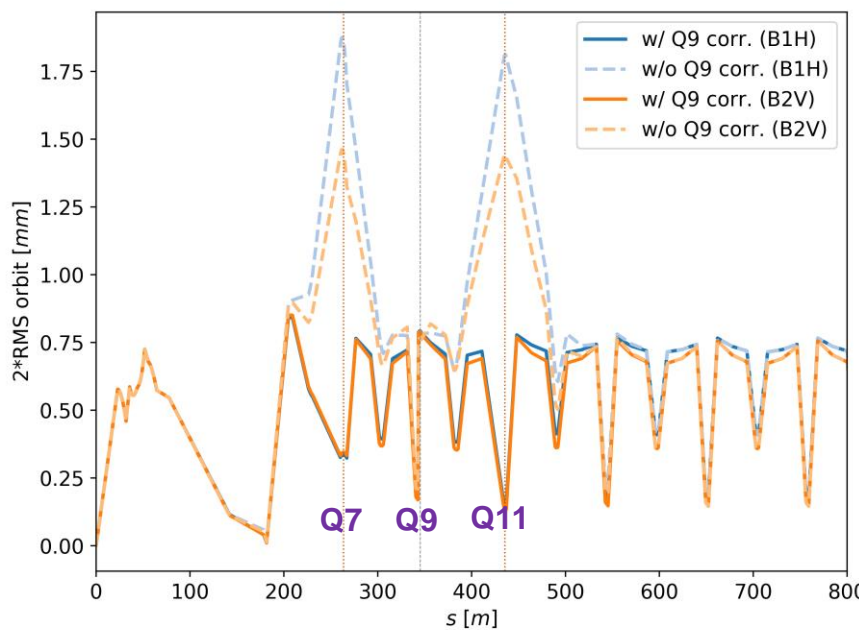
Failure Scenarios and Orbit Correction

- Each color represent **one orbit corrector failing**
 - but **still correcting** for misalignments with all other correctors
- In this respect, **only MCBXFA.3 seems to be fundamental!**



Failure Scenarios and Knobs Implementation

- Technically, for a generic knob implementation, we cannot fail:
 - MCBXFA.3**; **any MCBRDs**; the non-redundant **MCBYs**
 - Strongly used for crossing knob implementation.*
- However, also in other cases one should carefully verify all knob implementations on a case-by-case basis.
- In practice**: main interest is to verify failure of MCBC Q9 (e.g. [162th WP2](#))
 - Corrector not used for any knob implementation

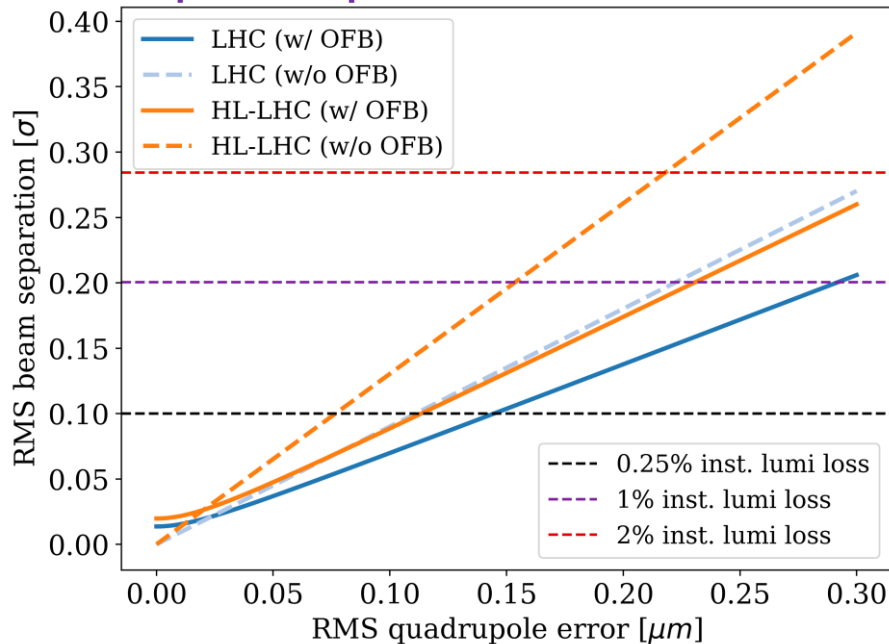


Orbit Feedback considerations

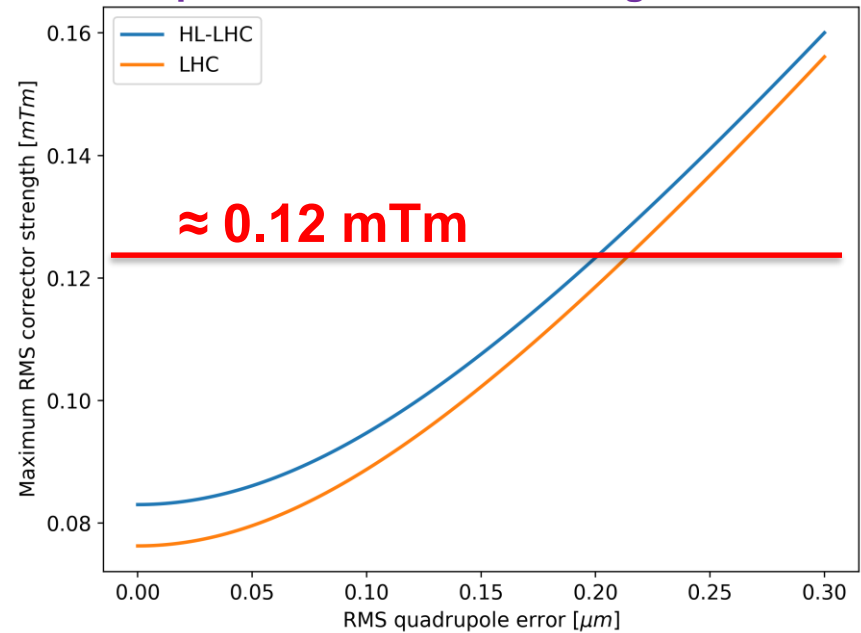
Orbit Correction during Stable Beam

- Assuming **typical** use of **orbit feedback** as today (LHC - 40/~500 singular values per plane), also including **triplet orbit correctors** (HL-LHC only).
 - From study of Joel presented at [164th](#) (and [156th](#)) WP2 Meetings

Expected separation at IP



Expected rms corrector usage



IP5; assumed RMS BPM error = $5 \mu\text{m}$

Required speed for MCBX orbit correctors

- 0.12 mTm = **5e-5 corrector usage** wrt nominal strength
 - Corresponds to about **80 mA rms orbit corrector usage**
 - Assuming 1 Hz oscillation, max derivative **about 0.7 A/s**
- Required performance of PC:**

Table 15: Comparison of the relevant orbit correctors and separation dipoles [1].

	MCBXFA	MCBXFB	MCBRD	MCBY	MBXF	MBRD
Nom. Int. field [Tm]	4.50	2.50	5.00	2.79	35.00	35.00
Nom. Current [A]	1600	1600	430	88	12000	12000
Ramp rate [A/s]	15.00	15.00	2.00	0.67	12.00 ^a	12.00 ^a
Field Rate [mTm/s]	42.19	23.44	23.26	21.15	35.00	35.00
Angle Rate [μrad/s @ 7TeV]	1.81	1.00	1.00	0.91	1.50	1.50
Ramp Acc. [A/s²]	5.00	5.00	1.00	0.25	2.00	2.00
Field Acc. [mTm/s²]	14.06	7.81	11.63	7.93	5.83	5.83
Angle Acc. [μrad/s² @7TeV]	0.60	0.33	0.50	0.34	0.25	0.25
Time to nom. rate [s]	3.00	3.00	2.00	2.67	6.00	6.00

^a In [1] it was specified 20 A/s as a first estimation.

Table from [CERN-ACC-2017-0101](https://cds.cern.ch/record/2002347)

- Concerns** that Quench Protection System (**QPS**) of **LHC MCBX** does not allow for high **di/dt** (false-positive quench detection)
 - Not an issue for HL-LHC** MCBXF as they will have middle voltage tap (EDMS [2002347](https://cds.cern.ch/record/2002347), R. Denz – HL-LHC Coll. Meeting 2018 [indico](#))

Conclusions

- A **generic tool** to quickly check correctors budget and residual orbit has been implemented (*by Joel - Thanks!*)
- **HL-LHCv1.5 $\beta^*=15$ cm** round optics verified:
 - Can **safely implement** all standard **knobs**
 - **Residual orbit** (wrt magnet axis) **<1 mm (2*rms)**
 - It can **sustain loss of Q9 MCBX** in case of radiation damage
 - **IP orbit stabilisation** during stable beam is expected to require **<0.1 mTm** (assuming 0.1 μ m quadrupole-displacement-equivalent errors, 5 μ m BPMs error) keeping the **luminosity loss below 0.25%**
 - **Compatible** with required **orbit corrector speed**.
- **Not covered here: flat optics** has been also analyzed
 - **No major differences, but tighter aperture**
- All results are **being summarized** in a **detailed note**.

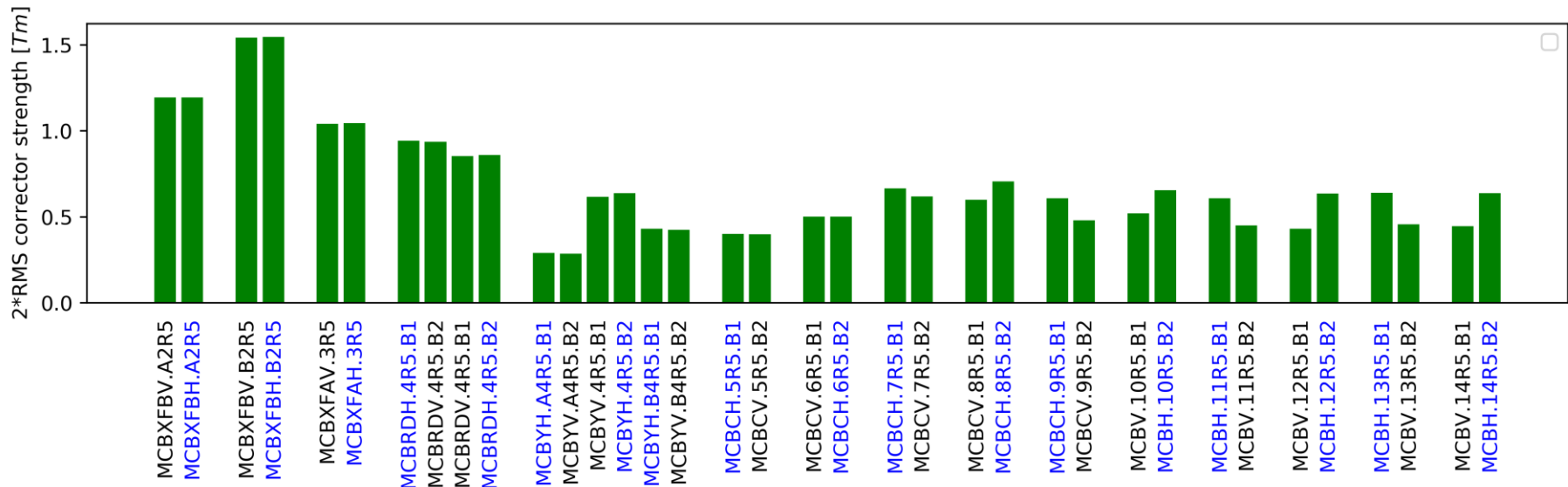
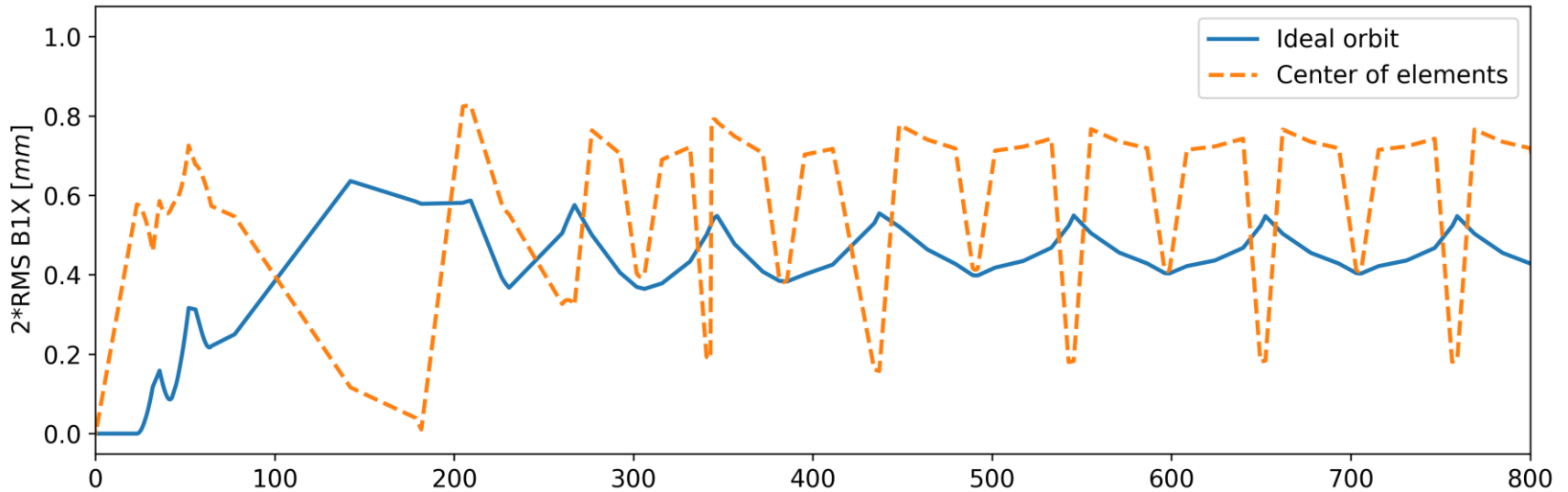
Thank you for your attention!

Backup

Flat Optics

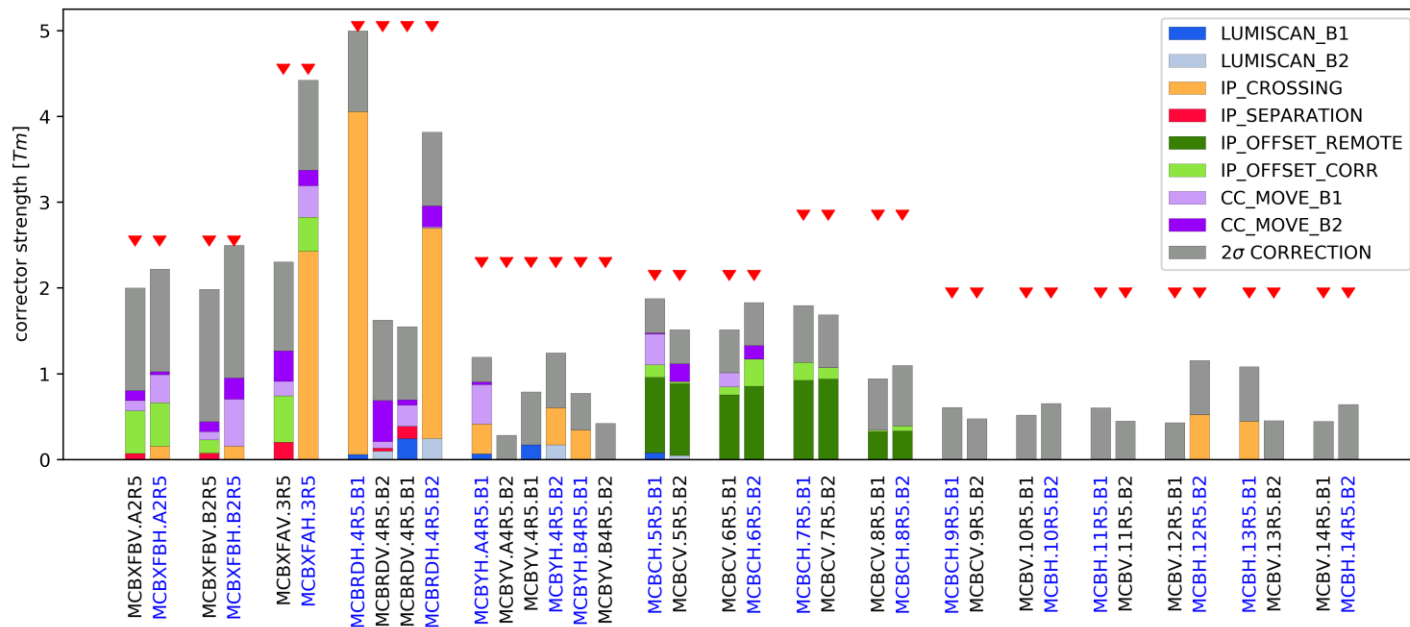
Flat Optics: Residual Orbit and Correctors Usage

- No difference! – strengths of elements in this region is basically unchanged!



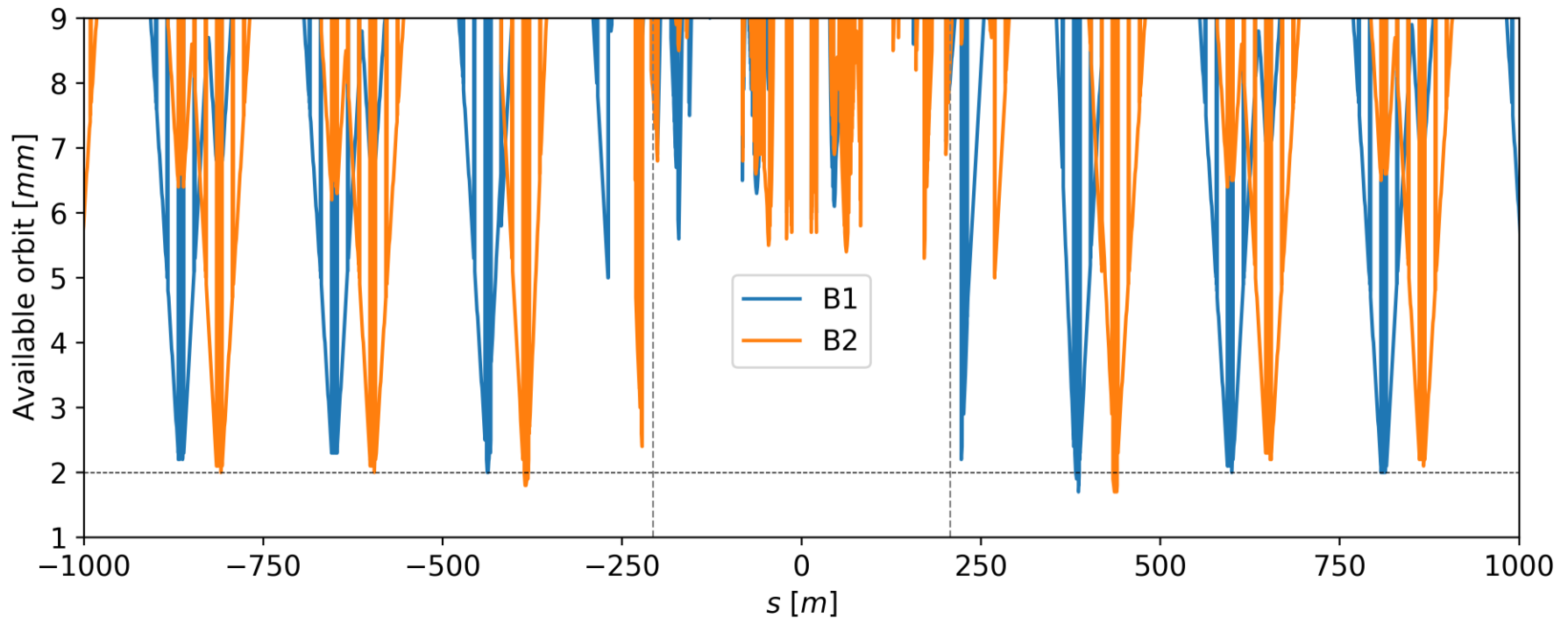
Flat Optics: Knobs implementation

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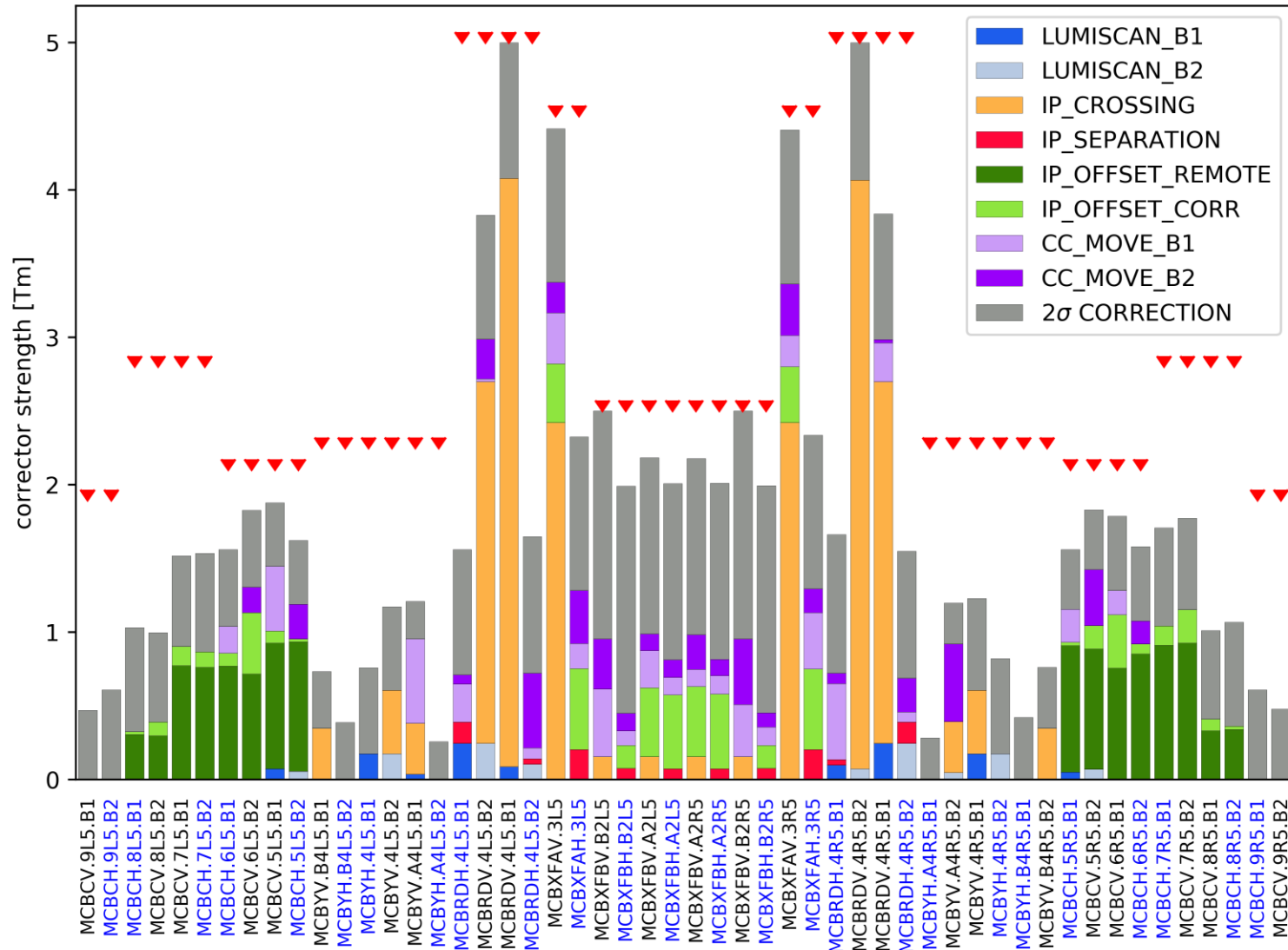
Flat optics: Available “aperture” for orbit

- Slightly less aperture, touching in a few points.
 - Still well compatible with expected residual orbit ($<1 \text{ mm } 2^* \text{r.m.s.}$)



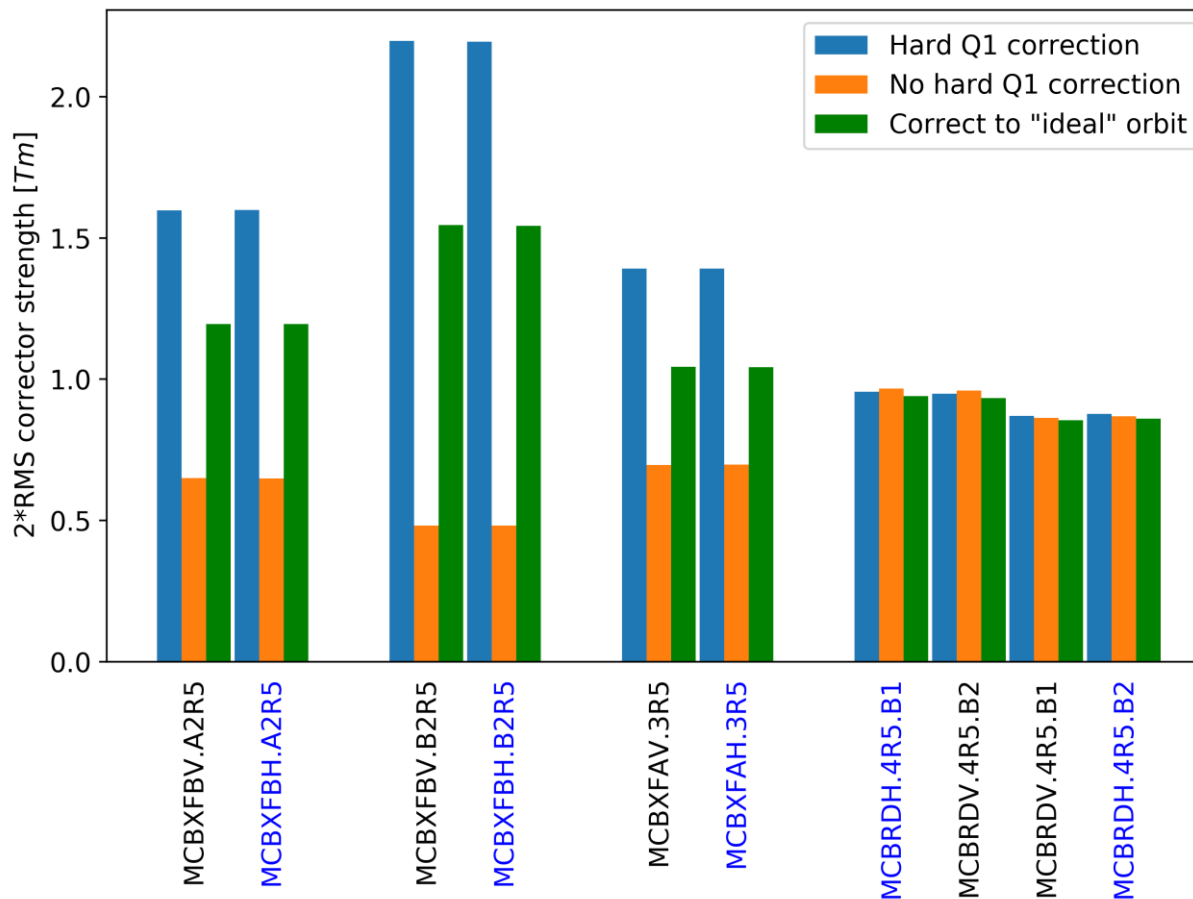
Orbit Corrector Budget

Corrector budget – complete up to Q9



Corrector strength expenditure for correction

- Comparison between orbit corrector budget use for different correction strategies in the triplet:
 - “Hard Q1 correction” = overcorrection at Q1 BPM
 - “Correct to ideal orbit” = orbit correction at ideal orbit at Q1 BPM
 - “No hard Q1 correction” = simple orbit correction like in the arc.



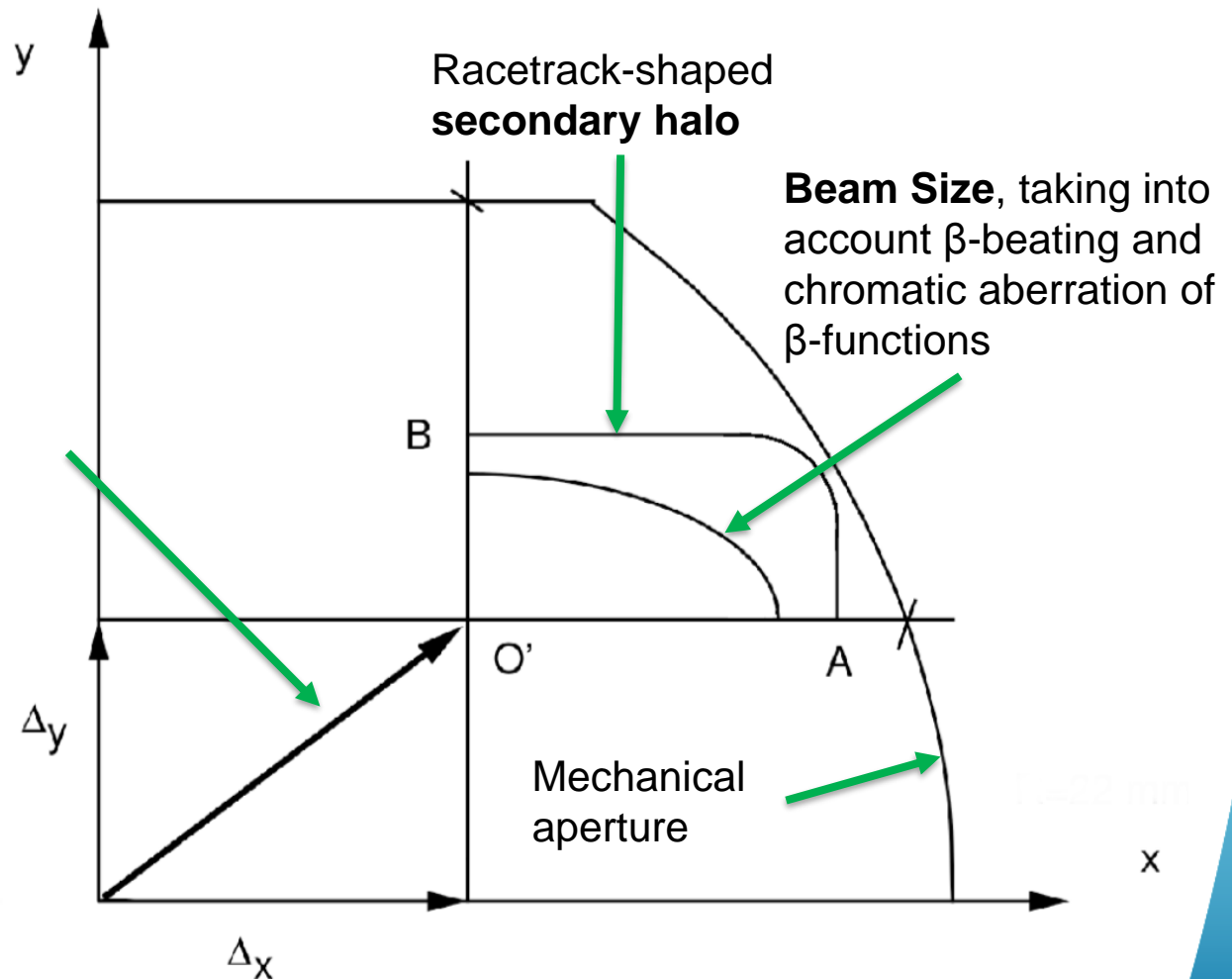
Apertures and Failures

Aperture considerations: how to compute?

- Present baseline for aperture computation ([CERN-ACC-2017-0051](#))

Orbit tolerances, including:

- Closed orbit deviation ($x_{co} = 2 \text{ mm}$);
- Mechanical alignment tolerance (Δ_{al});
- Beam screen alignment (Δ_{ba});
- Cold bore alignment (Δ_{cb});
- Off-momentum component ($D\delta_p$; taking into account dispersion beating)



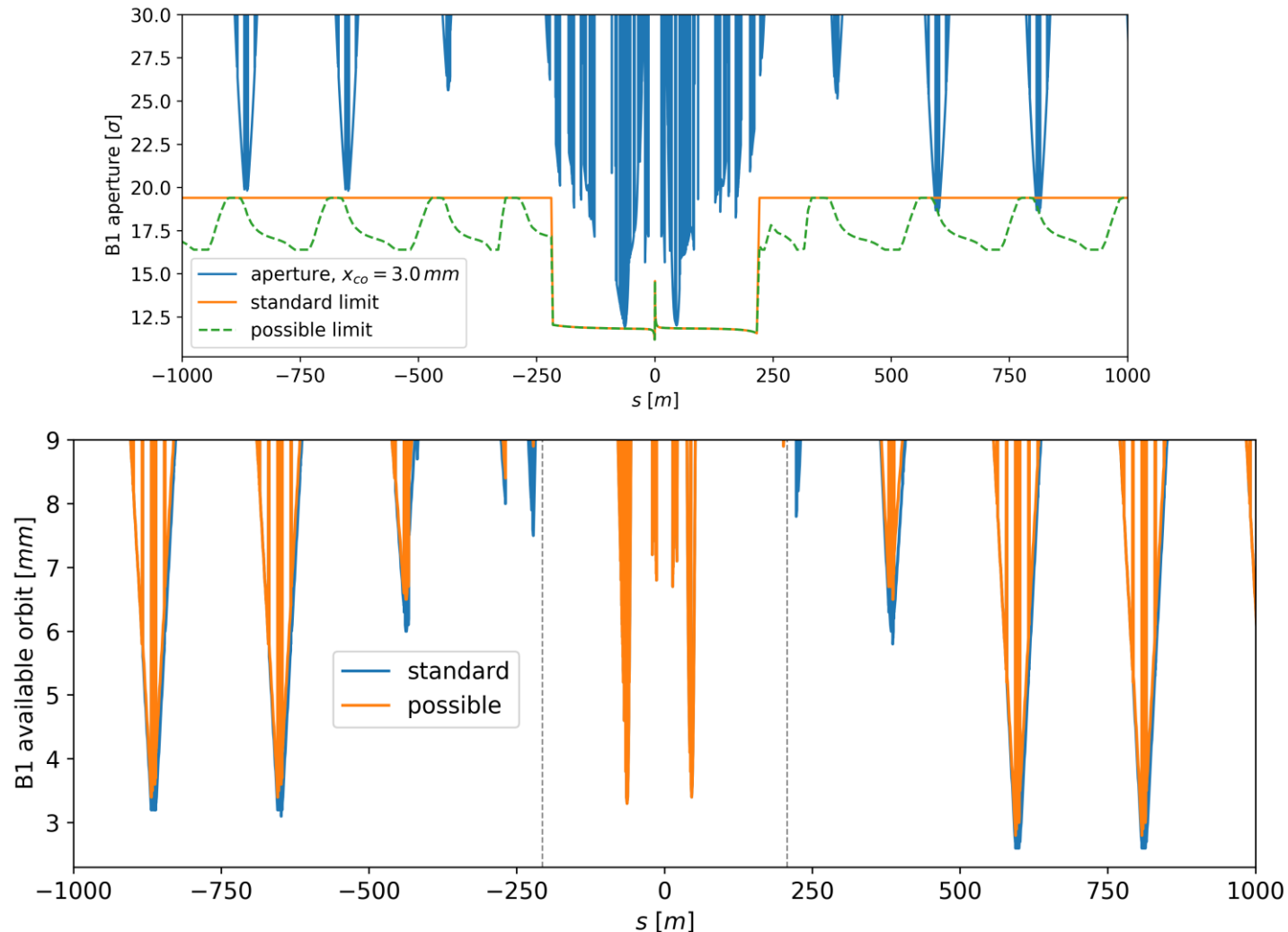
Aperture considerations: how to compute?

- Present baseline for aperture computation ([CERN-ACC-2017-0051](#)) assumes:

Parameter	Injection <u>Note (Example)</u>	Top energy <u>Note (Example)</u>	Description
Halo(s)	6σ	6σ	Primary and secondary halo extensions
ϵ_n	2.5 (2.5) μm	2.5 (2.5) μm	Normalized emittance.
dPMax	8.6e-4 (8.6e-4)	2e-4 (2e-4)	“Bucket edge at the current beam energy.” -> to be set to 0 for TWISS_DELTAP != 0
x_{co}	2 (2) mm	2 (2) mm	Max closed orbit deviation - radial
k_β	1.05 (1.05)	1.1 (1.1)	β beating
f_{arc}	0.14 (0.14)	0.1 (0.1)	Relative parasitic dispersion (scaling from arc to local dispersion) (DPARX/DPARY in MAD-X)
δ_p	8.6e-4 (6e-4?)	2e-4 (2e-4?)	Momentum offset used to compute off-momentum β beating by executing 3 separate Twiss $-\delta_p$; 0; $+\delta_p$
σ_p	(4.5e-4)	(4.5e- 4*sqrt(450/7000))	Beam energy spread, used in beam definition -> not being used by aperture calculation
Interval	n.a. (1.0)	n.a. (1.0)	Approximate length in meters between measurements.
SPECIF	(12.6)	(14.6)	Aperture spec, for plotting only.
VMAXI	(30)	(30)	??

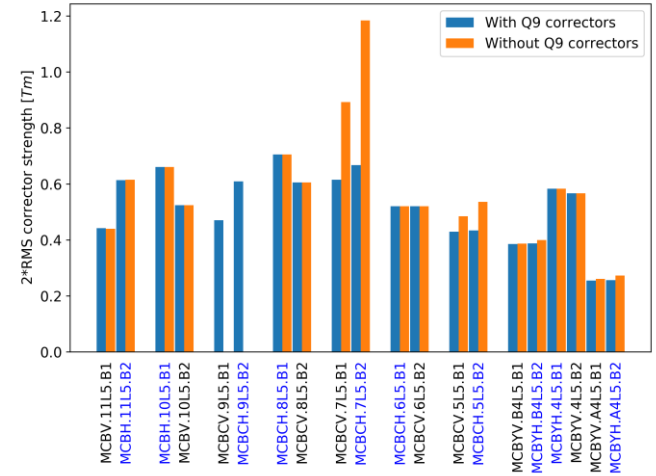
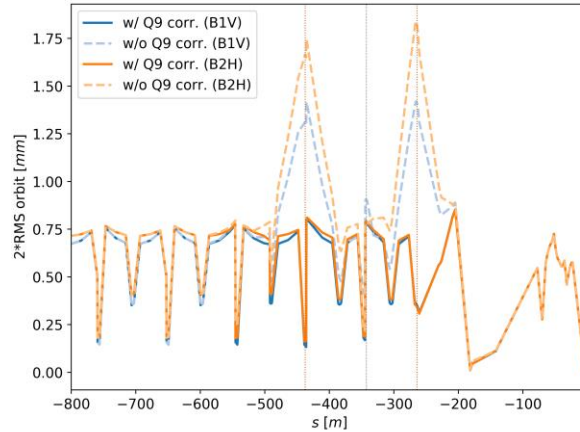
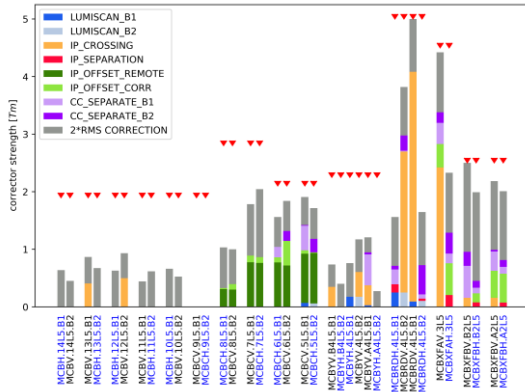
Aperture B1 comparison (round optics)

- Using [CERN-ACC-2017-0051](#) as aperture limit also in the arc

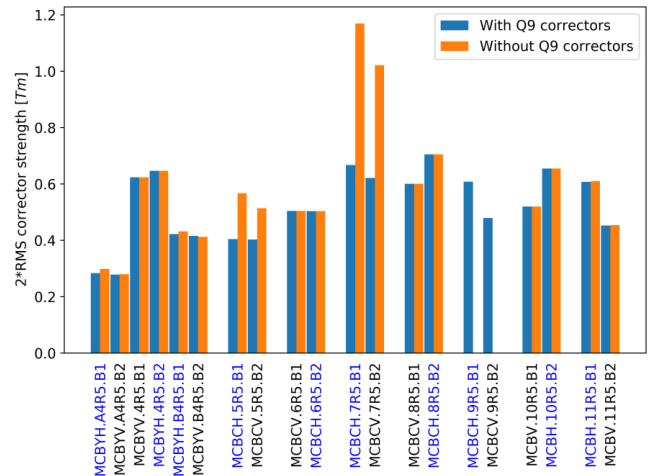
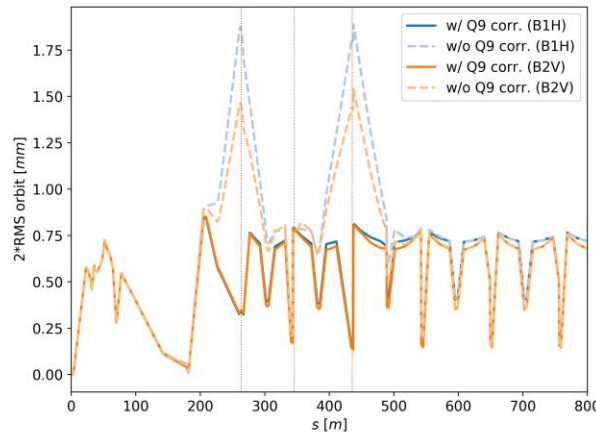
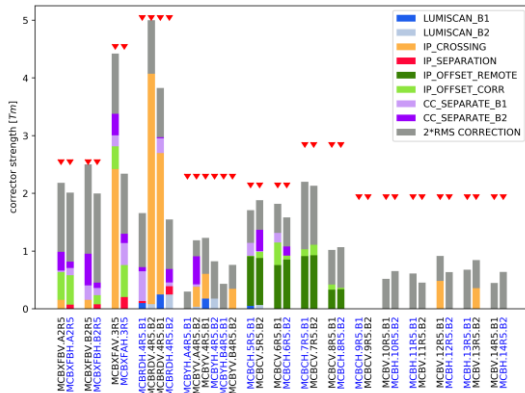


Full case for failures of Q9 orbit correctors

Nominal/failure case around Q9 Left H/V

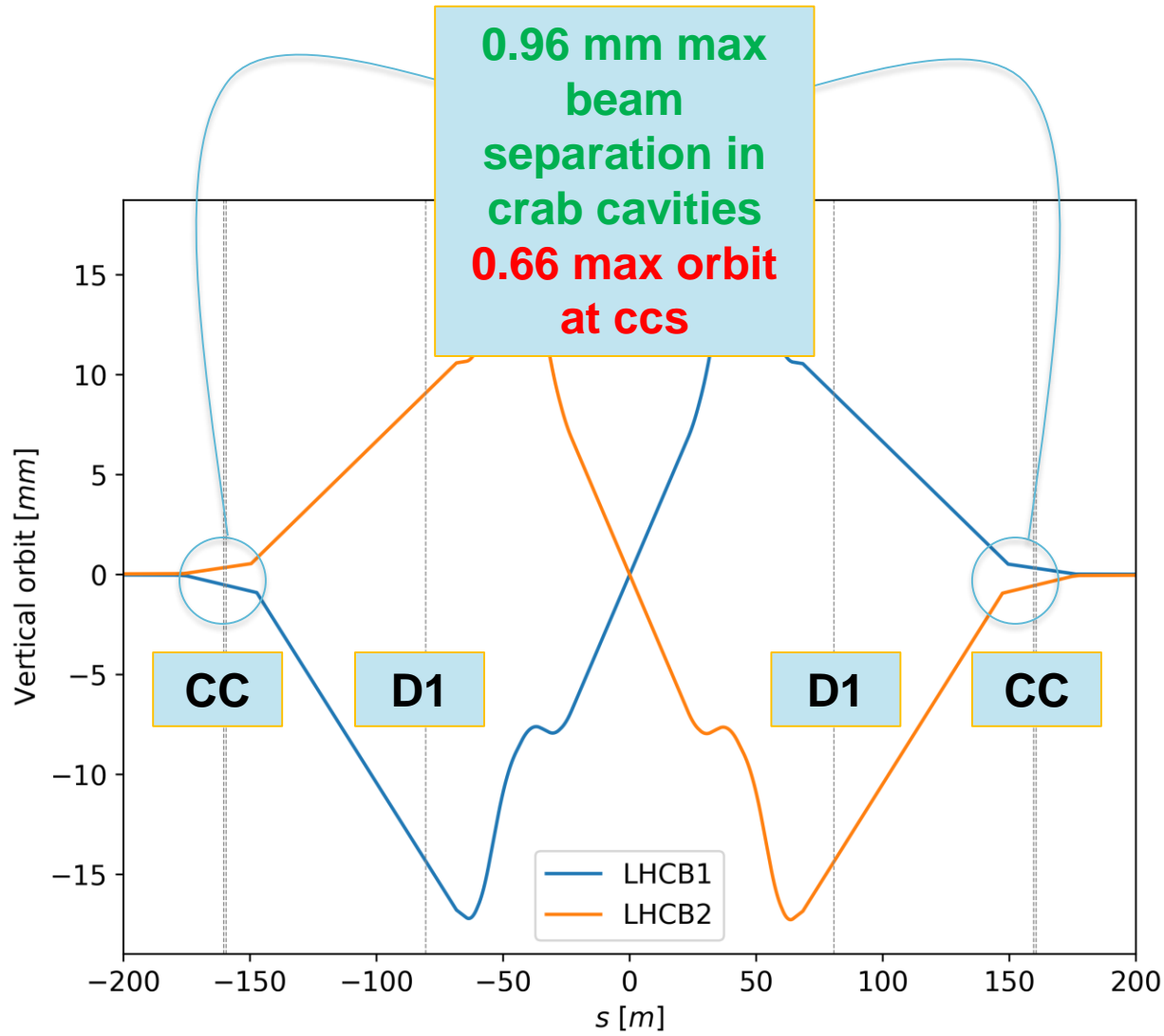


Nominal/failure case around Q9 Right H/V

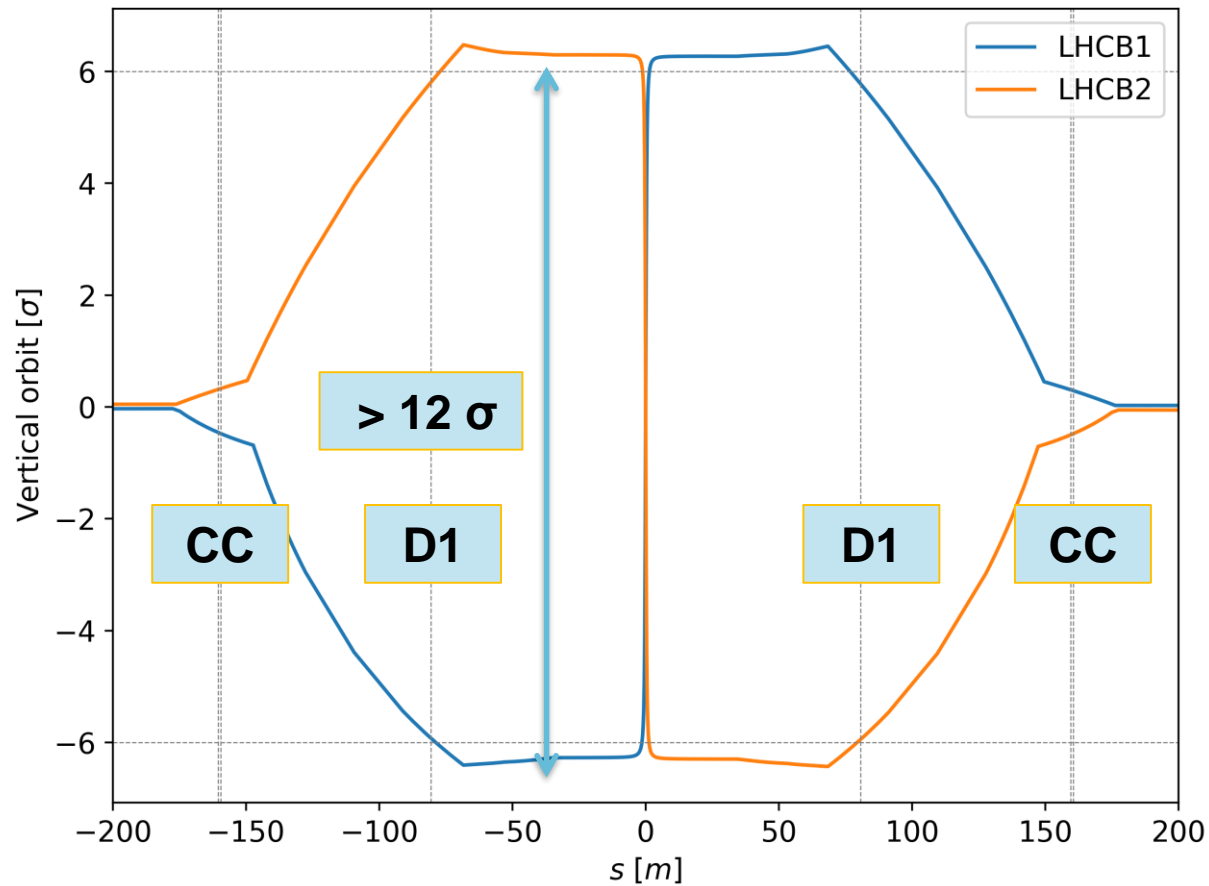


Knobs

Crossing knob

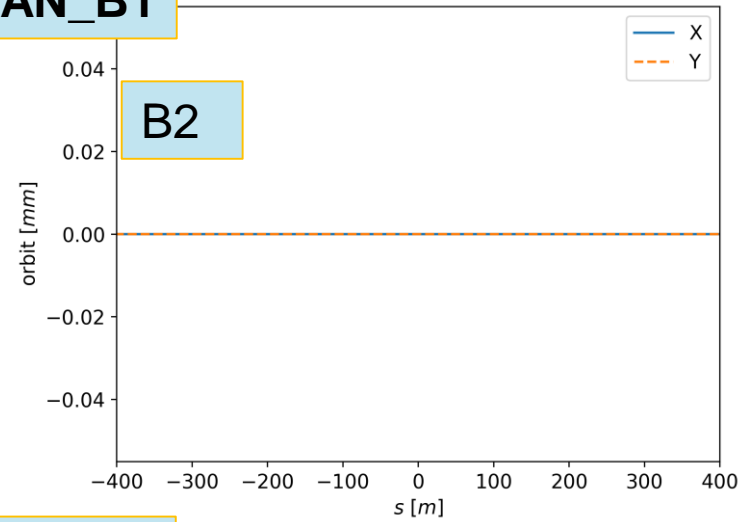
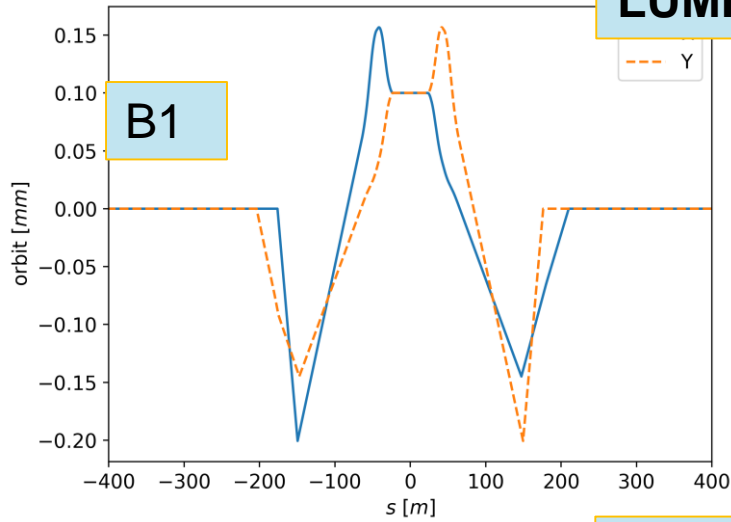


Crossing knob (in beam sigmas)

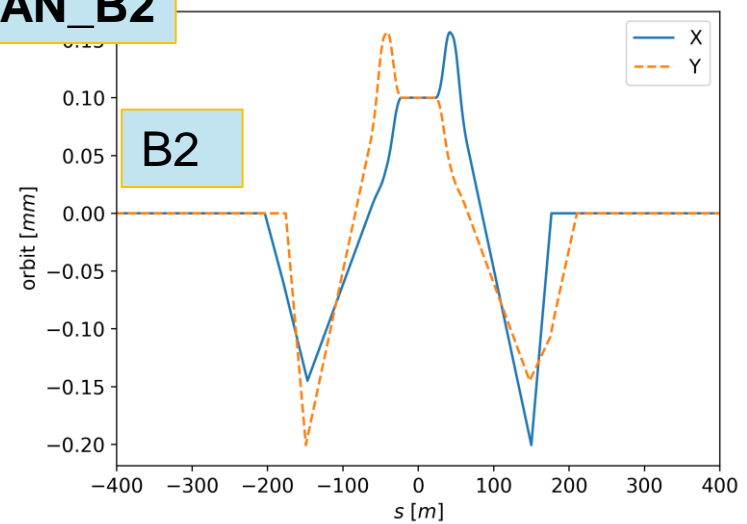
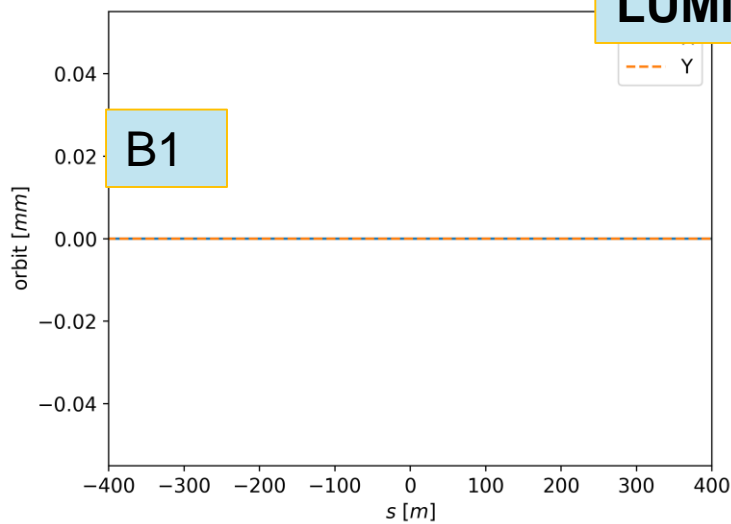


Knob orbits

LUMISCAN_B1

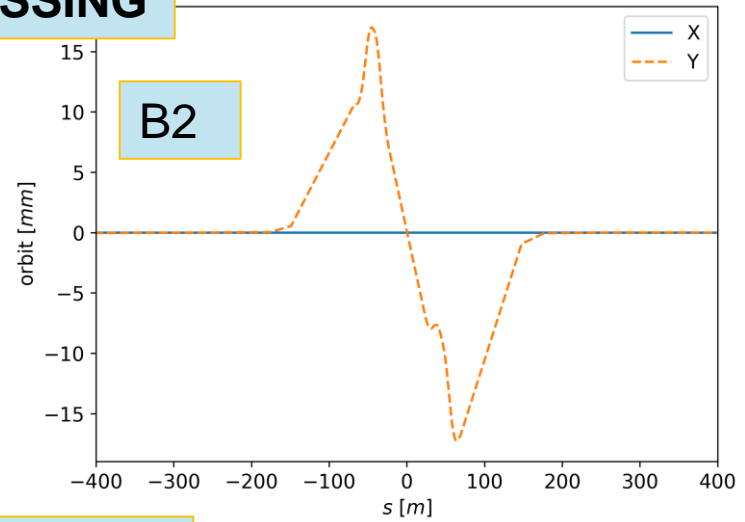
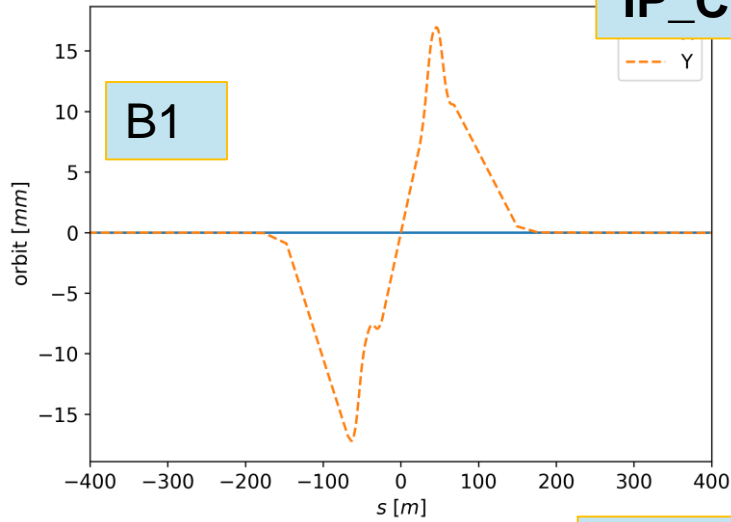


LUMISCAN_B2

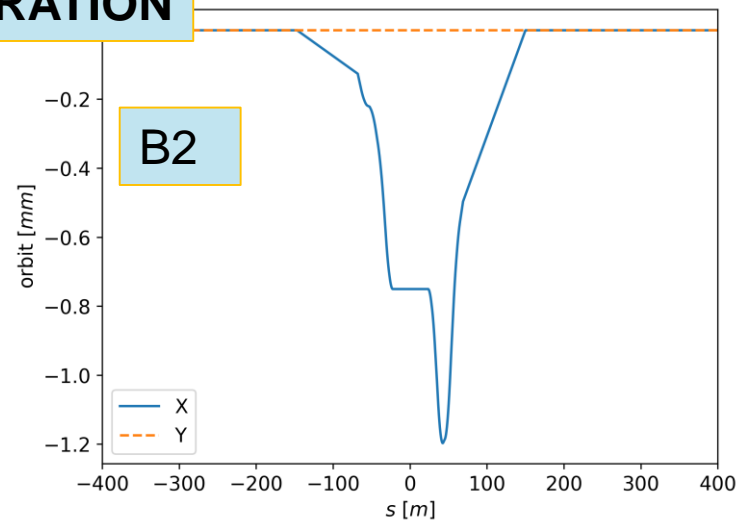
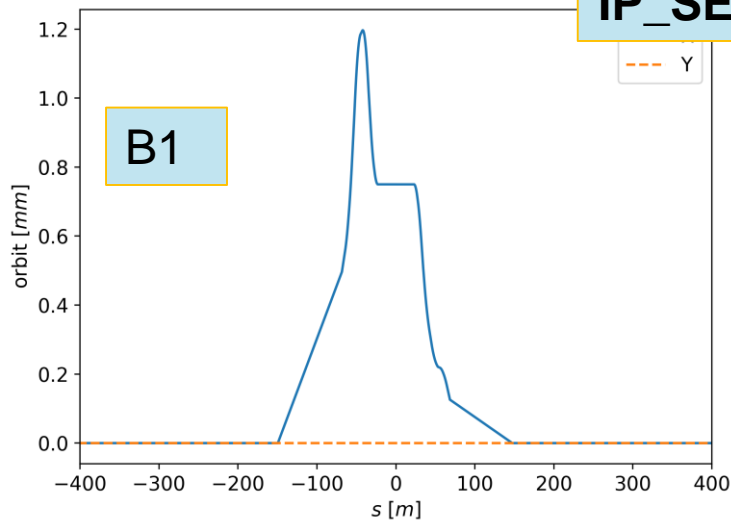


Knob orbits

IP_CROSSING

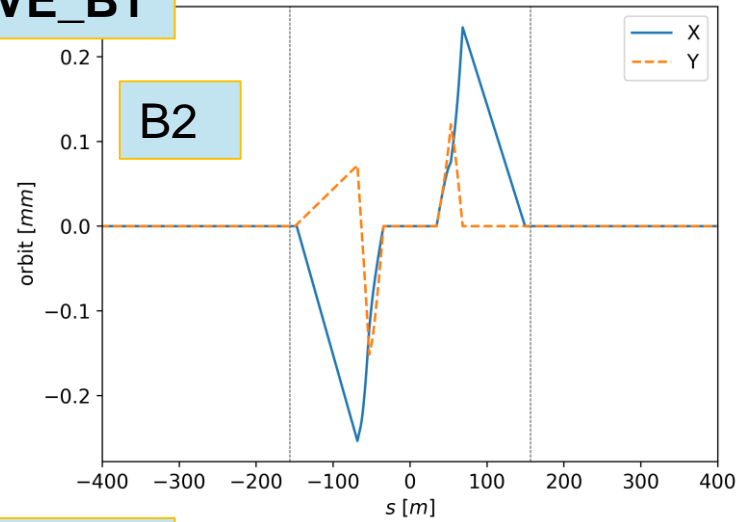
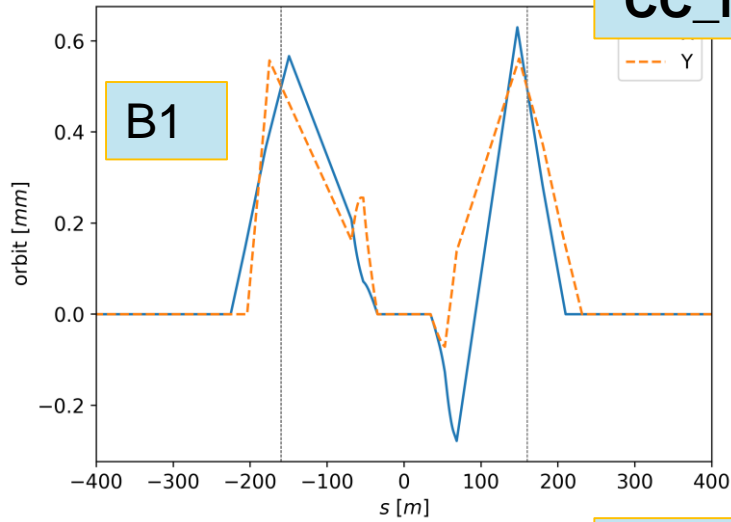


IP_SEPARATION

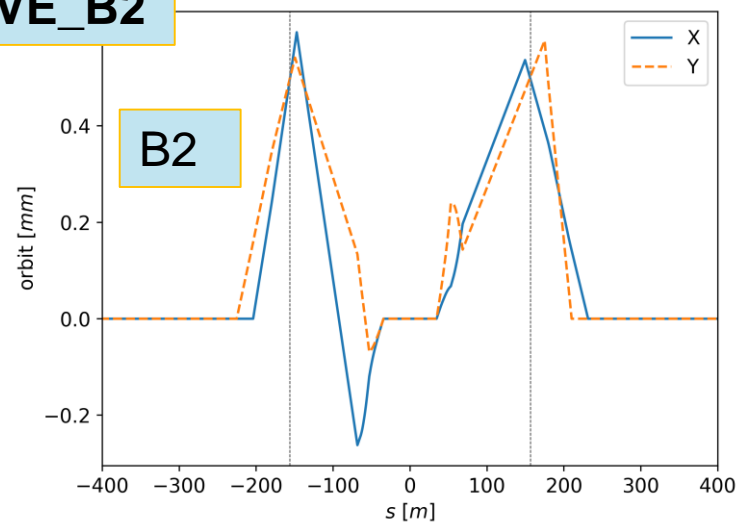
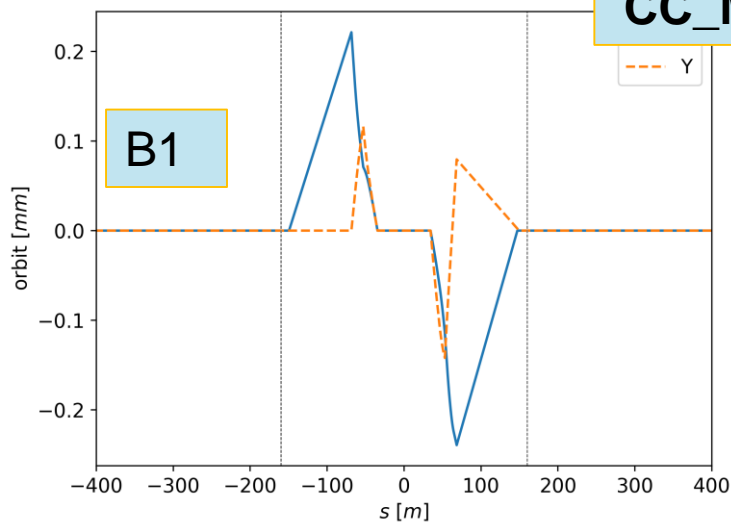


Knob orbits

CC_MOVE_B1

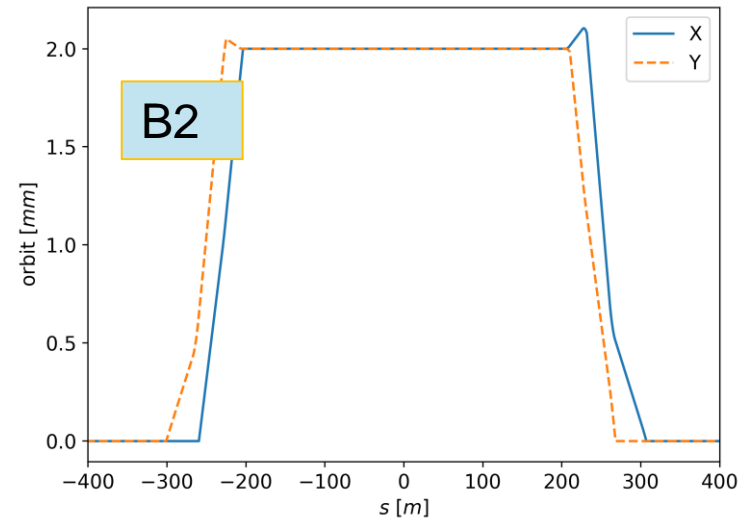
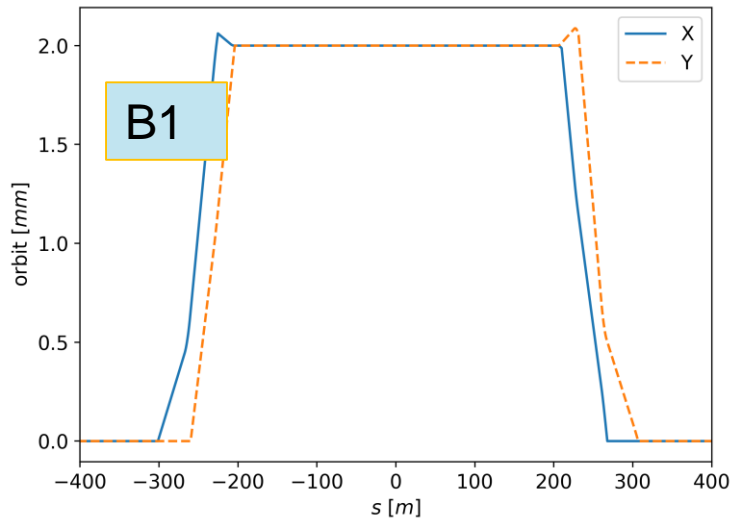


CC_MOVE_B2

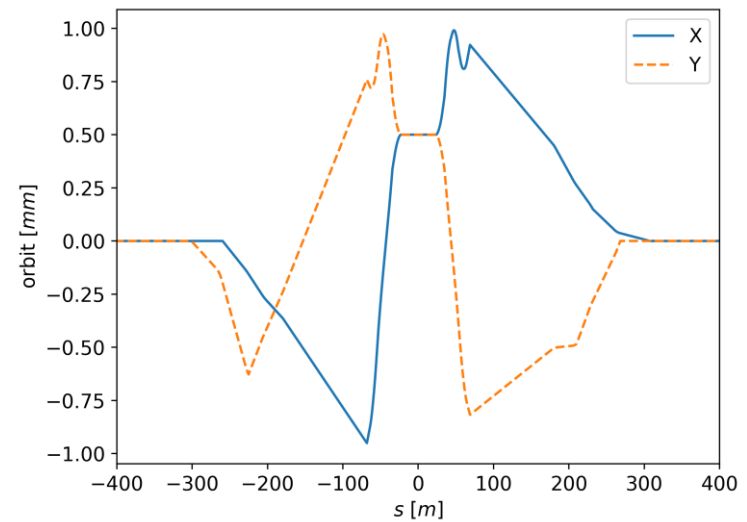
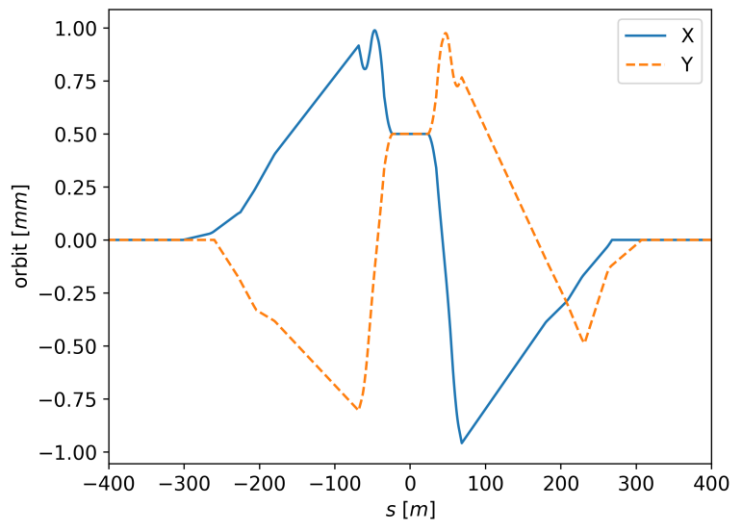


Knob orbits

IP_OFFSET_REMOTE

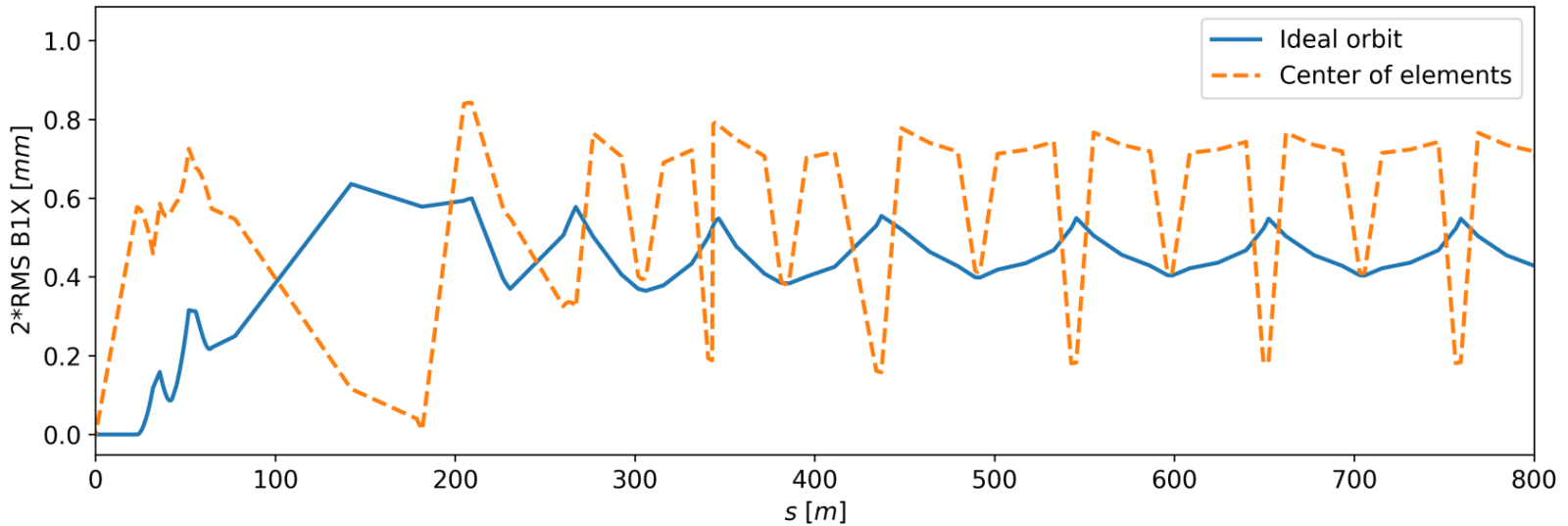
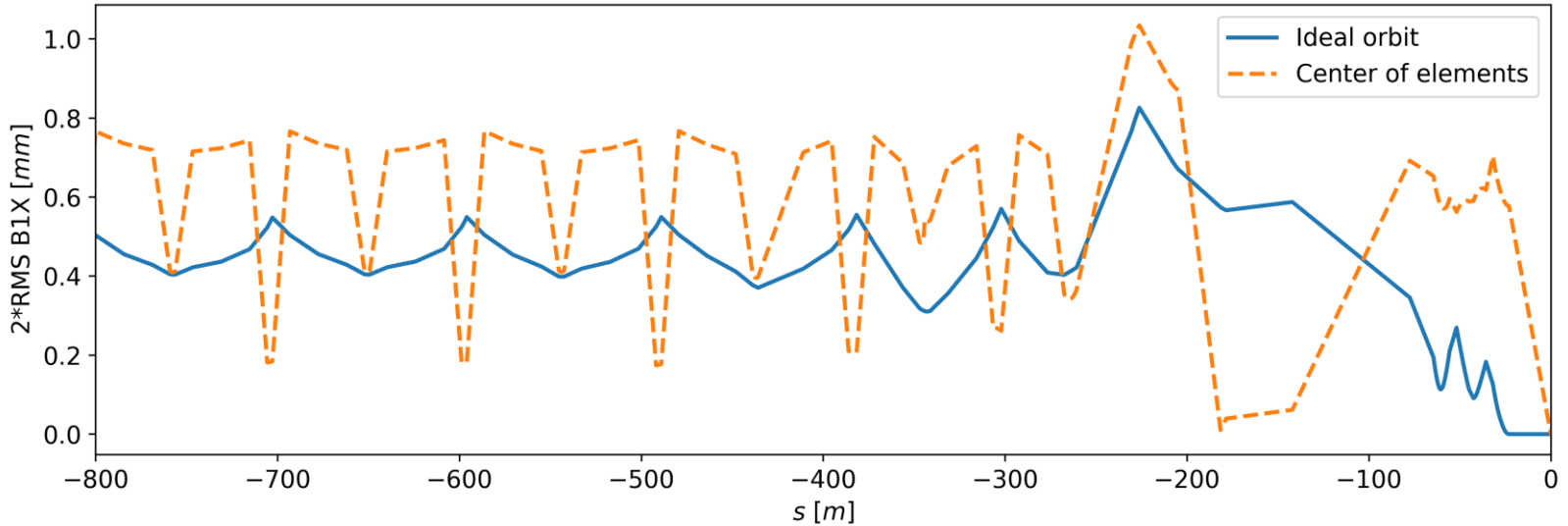


IP_OFFSET_CORR



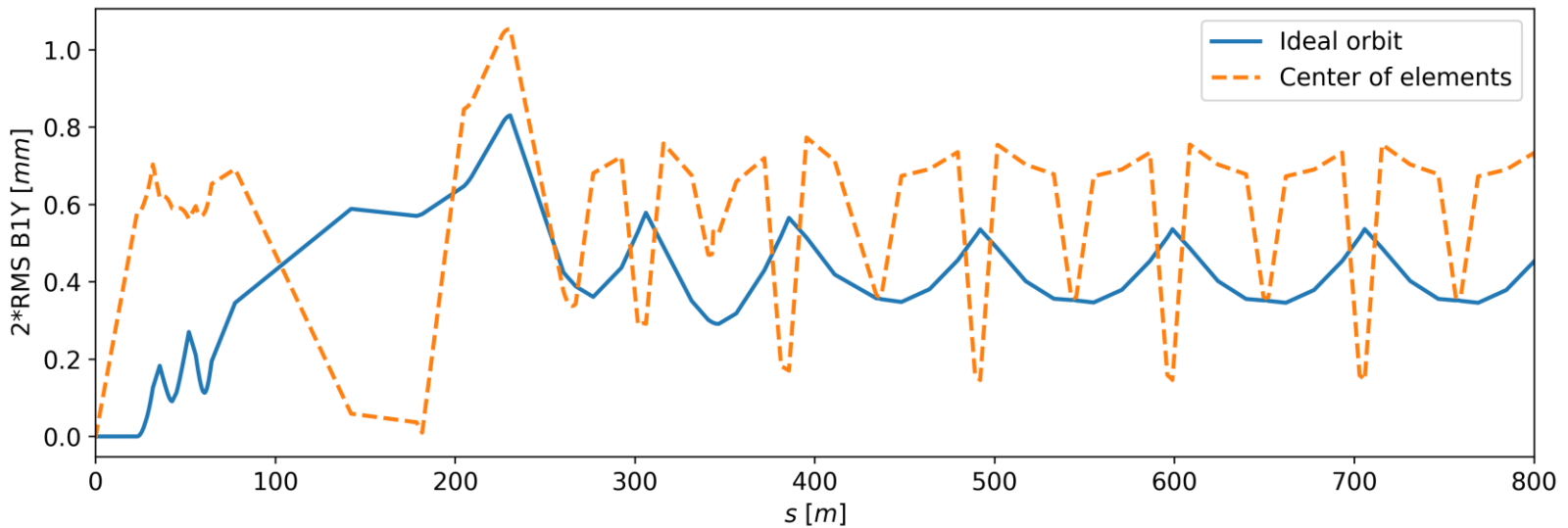
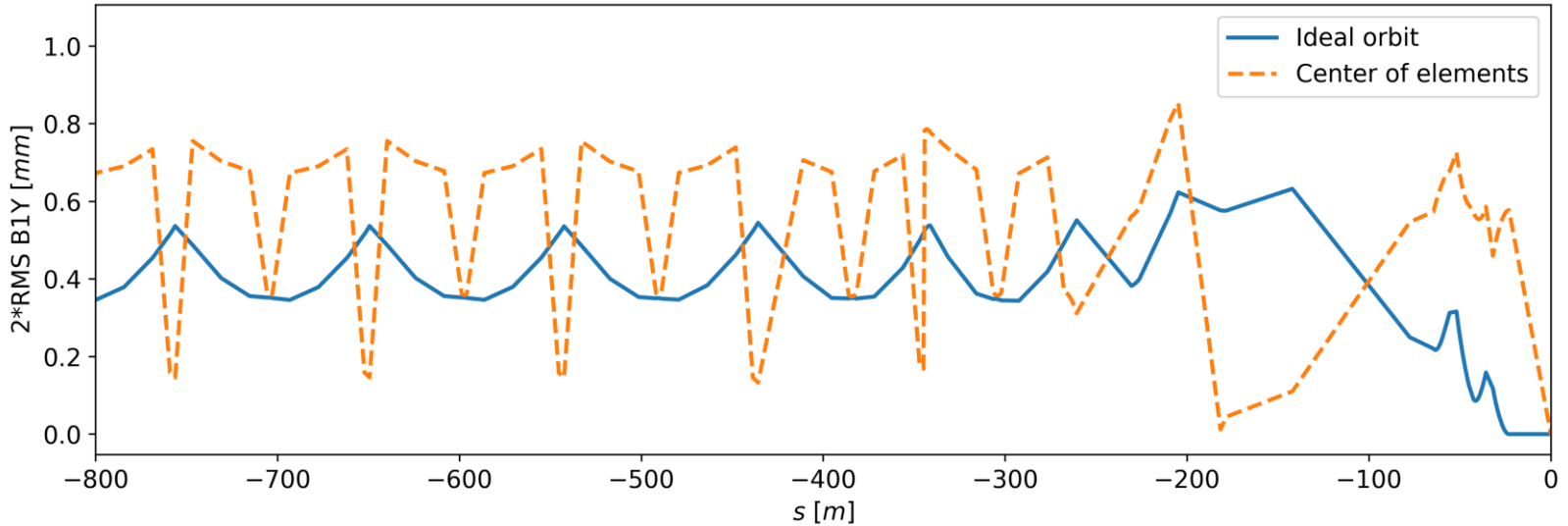
Residual orbit B1X

- **BPMs move with nearby magnet + “strong” correction to get beam at ideal IP**



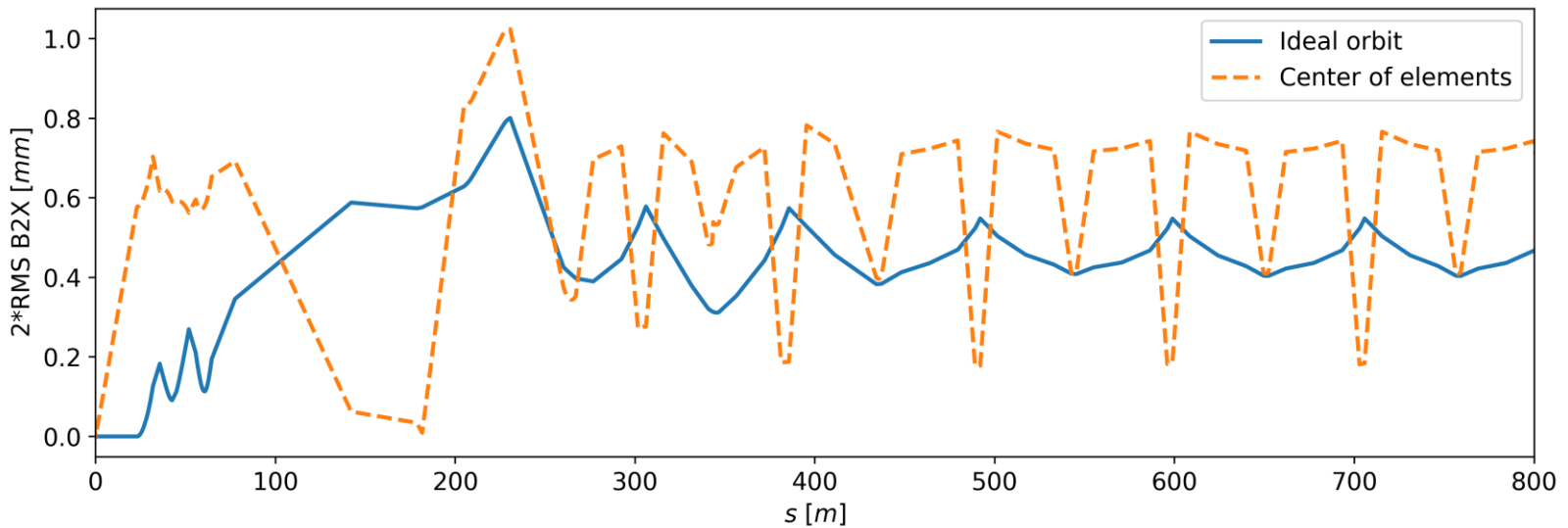
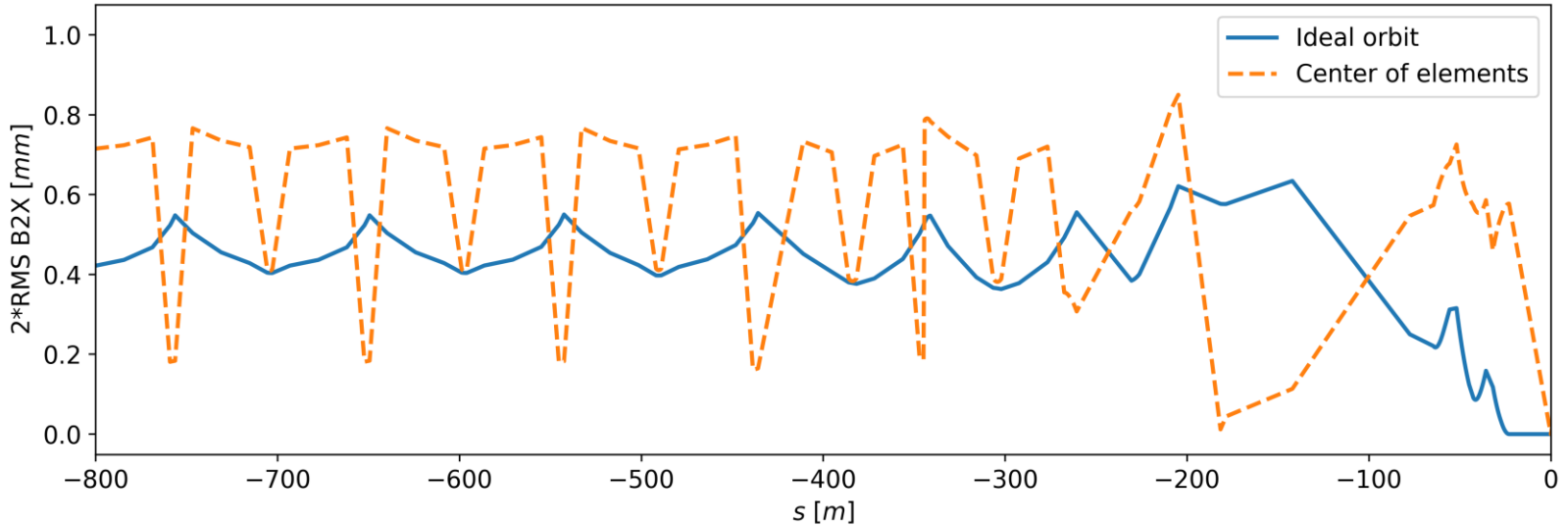
Residual orbit B1Y

- **BPMs move with nearby magnet + “strong” correction to get beam at ideal IP**



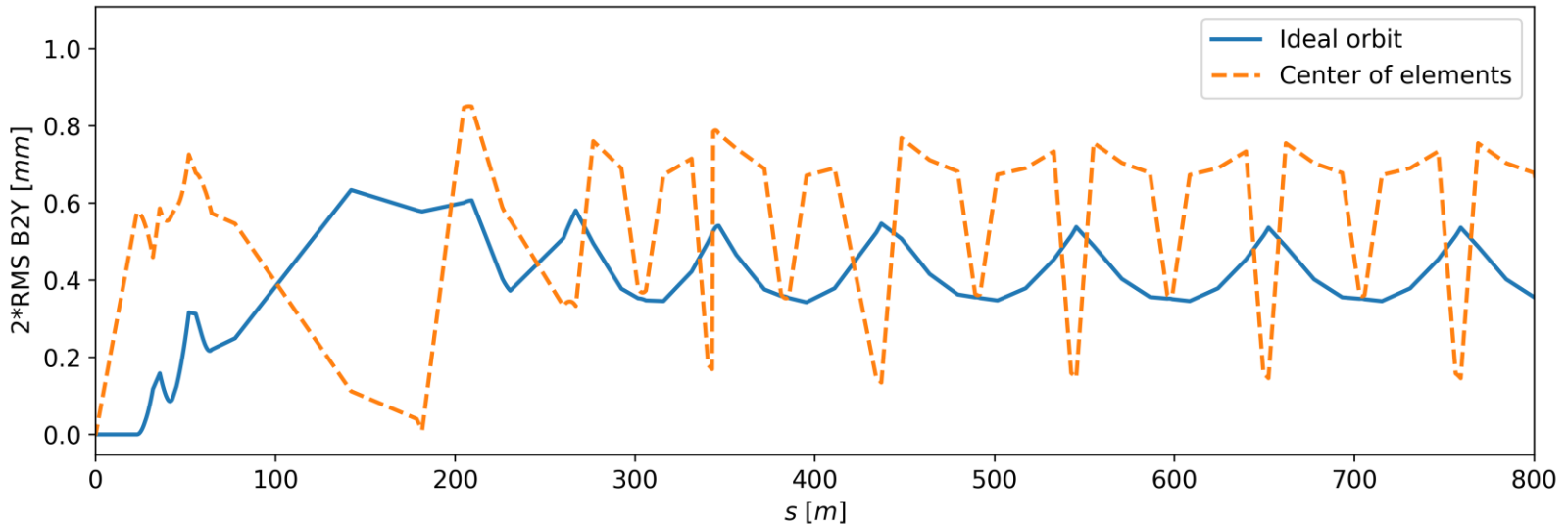
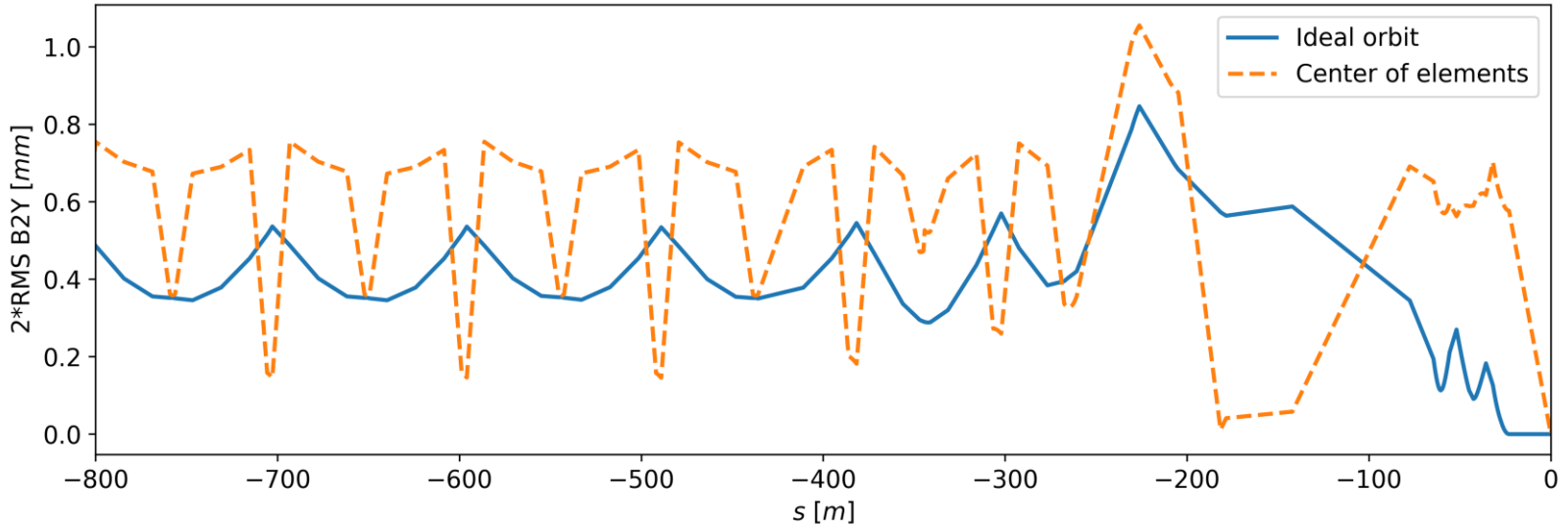
Residual orbit B2X

- **BPMs move with nearby magnet + “strong” correction to get beam at ideal IP**

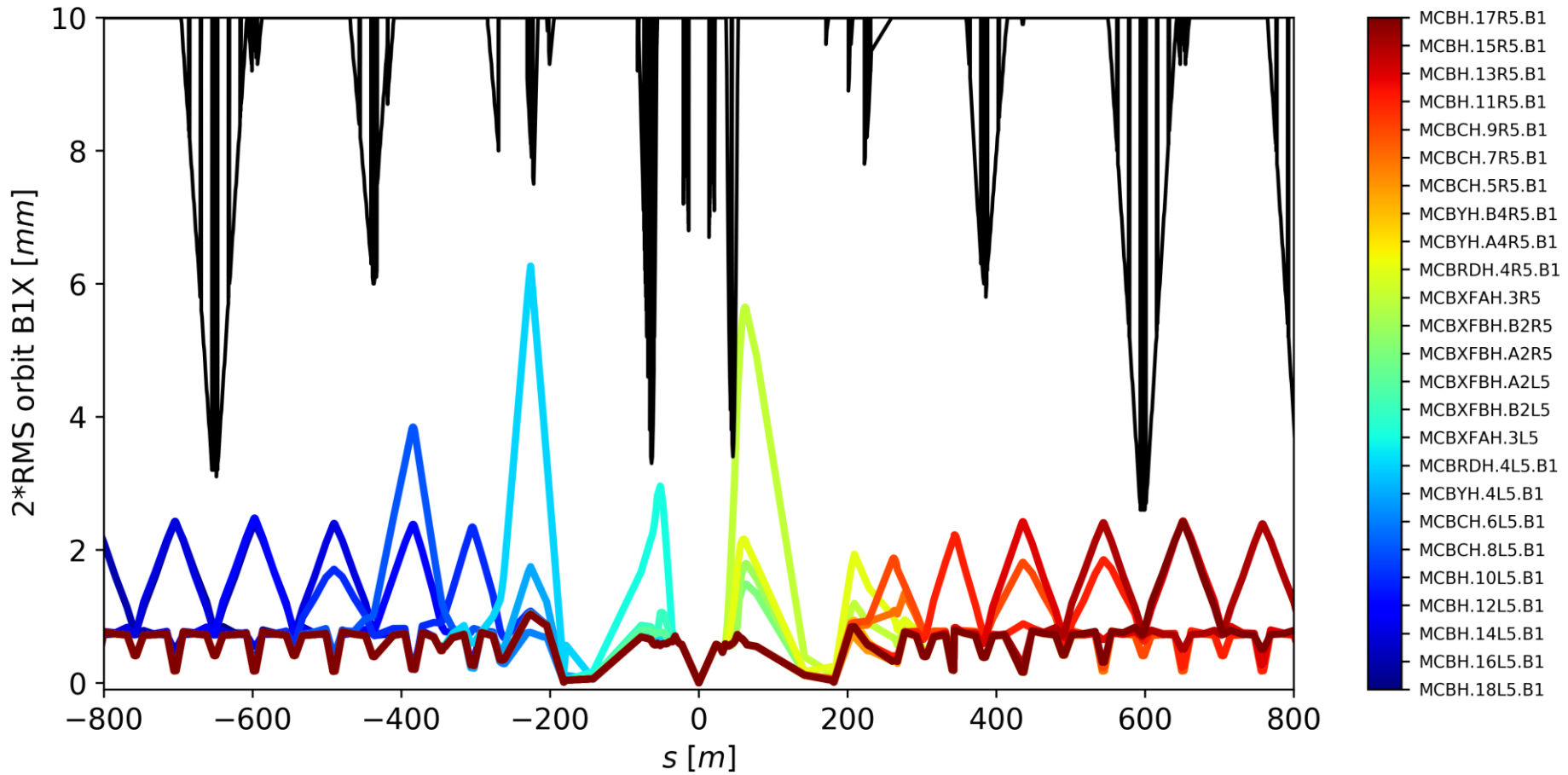


Residual orbit B2Y

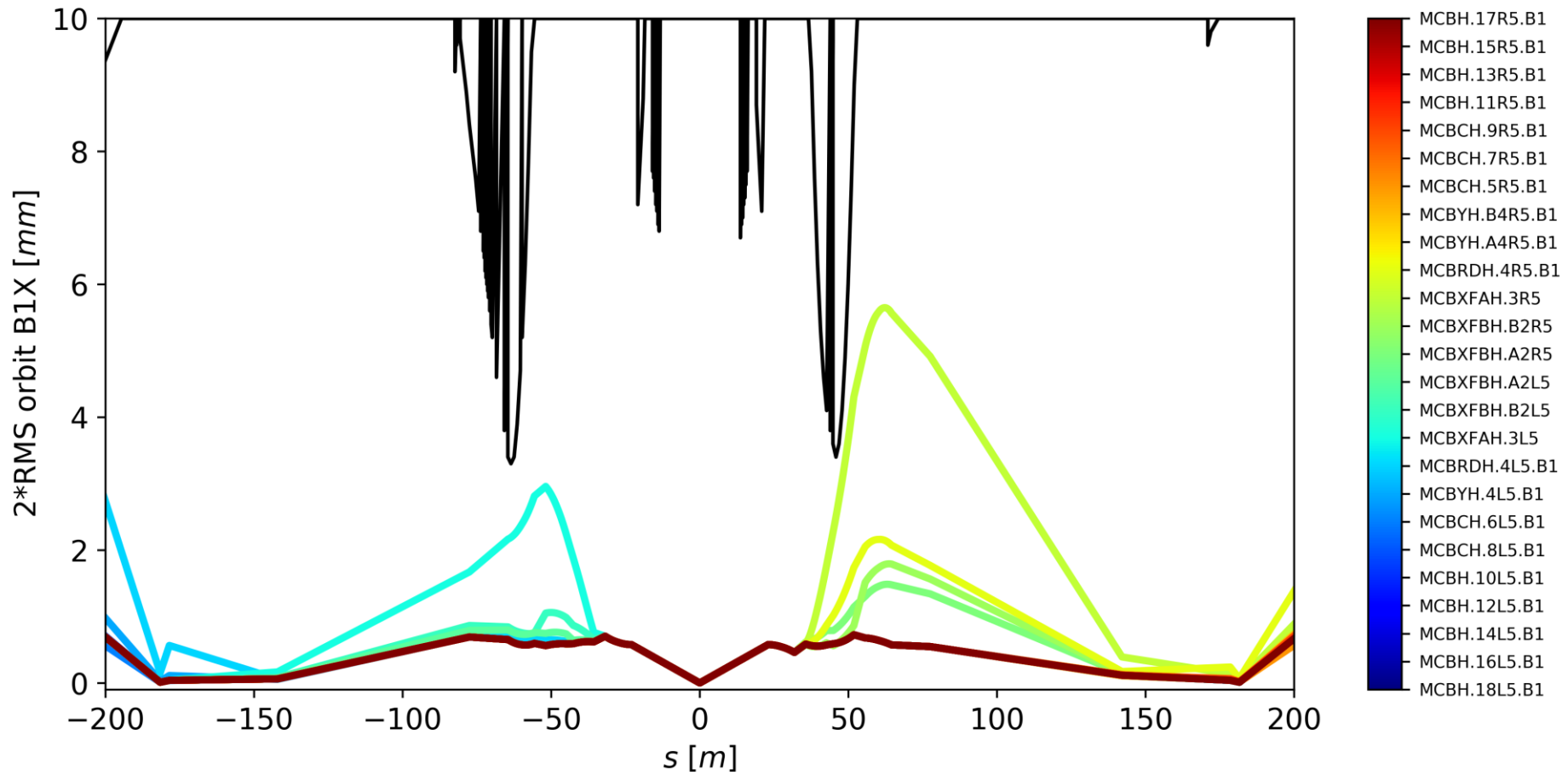
- BPMs move with nearby magnet + “strong” correction to get beam at ideal IP



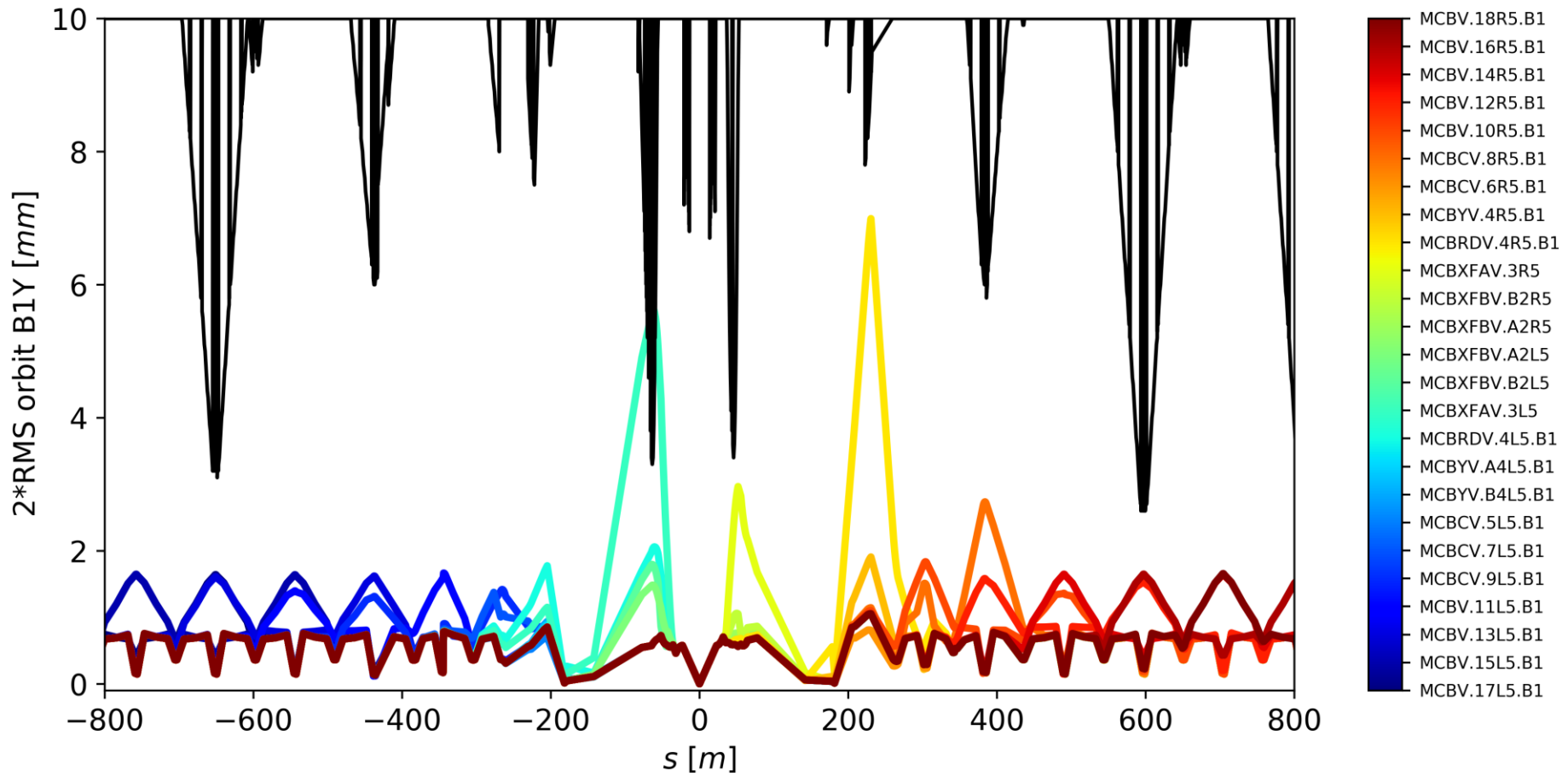
Failure scenario: B1X



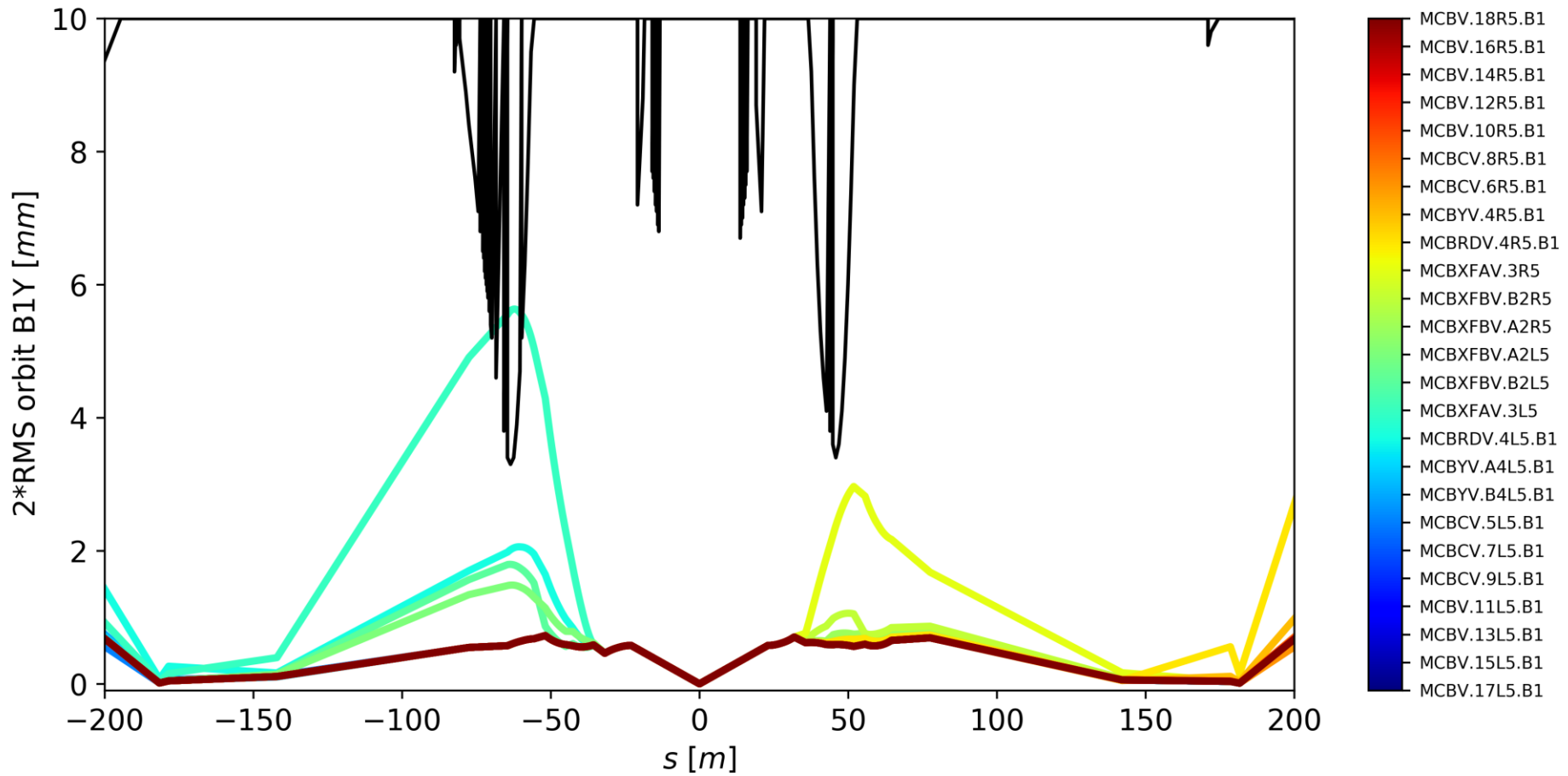
Failure scenario: B1X (zoomed)



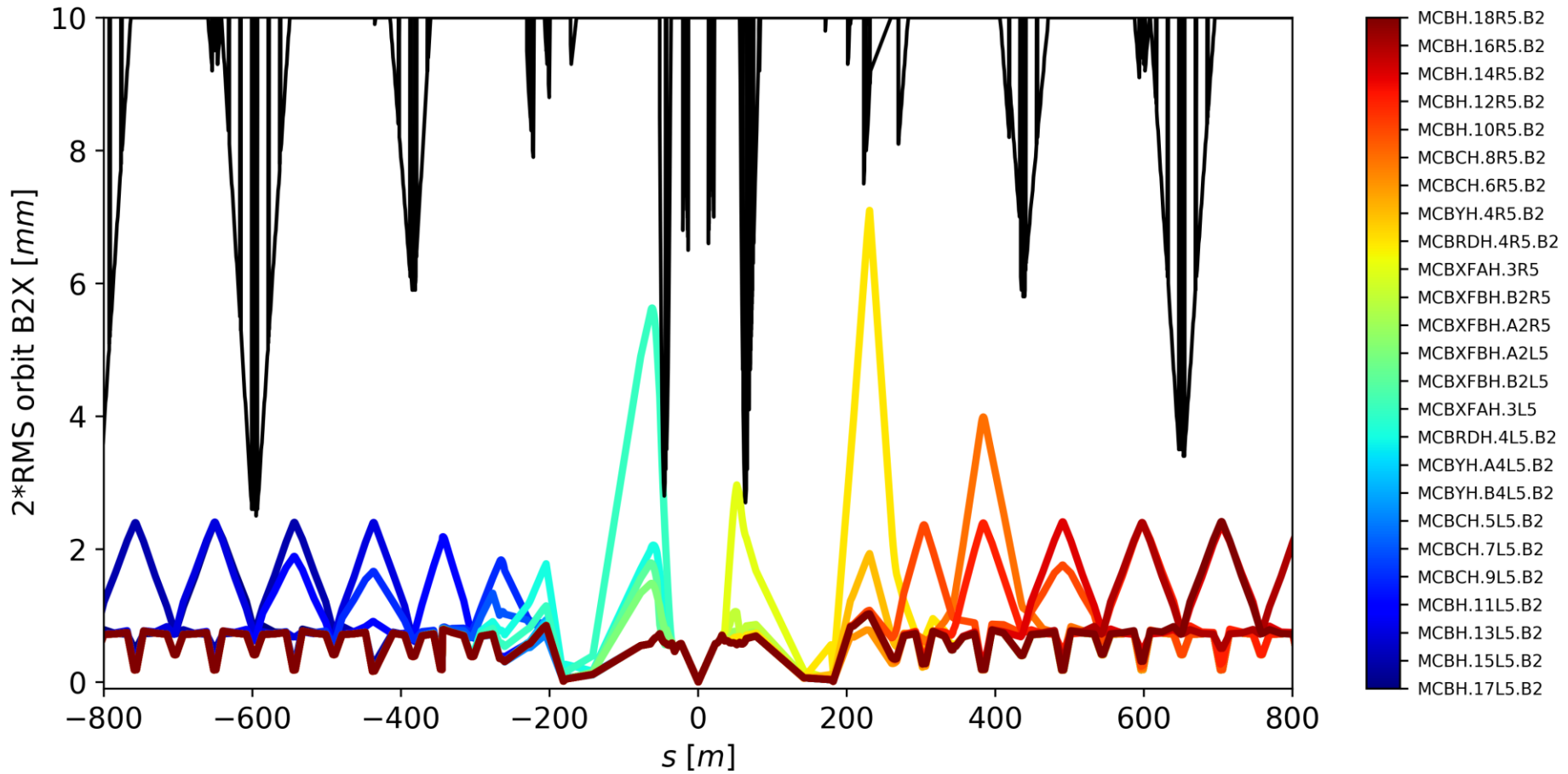
Failure scenario: B1Y



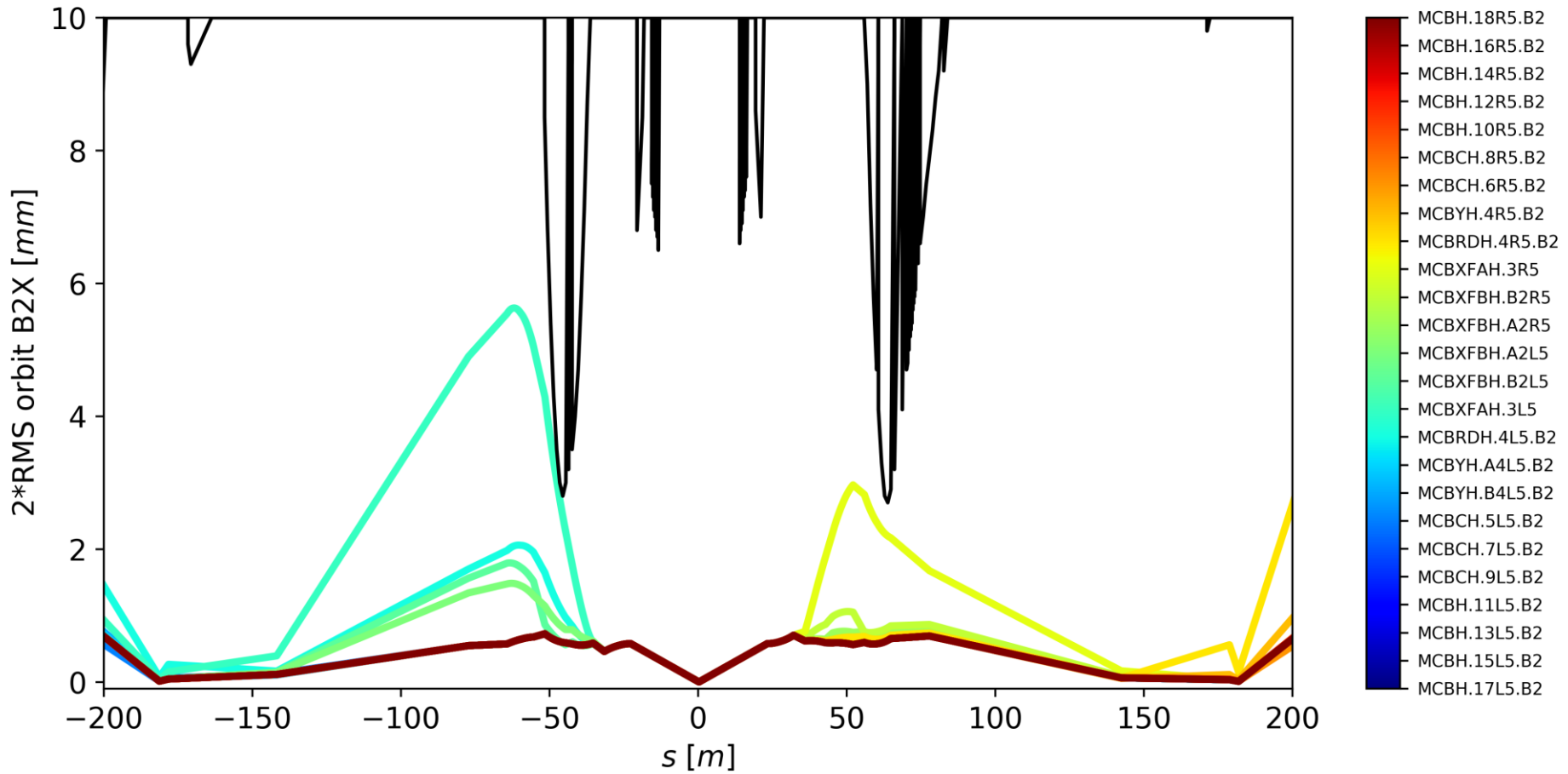
Failure scenario: B1Y



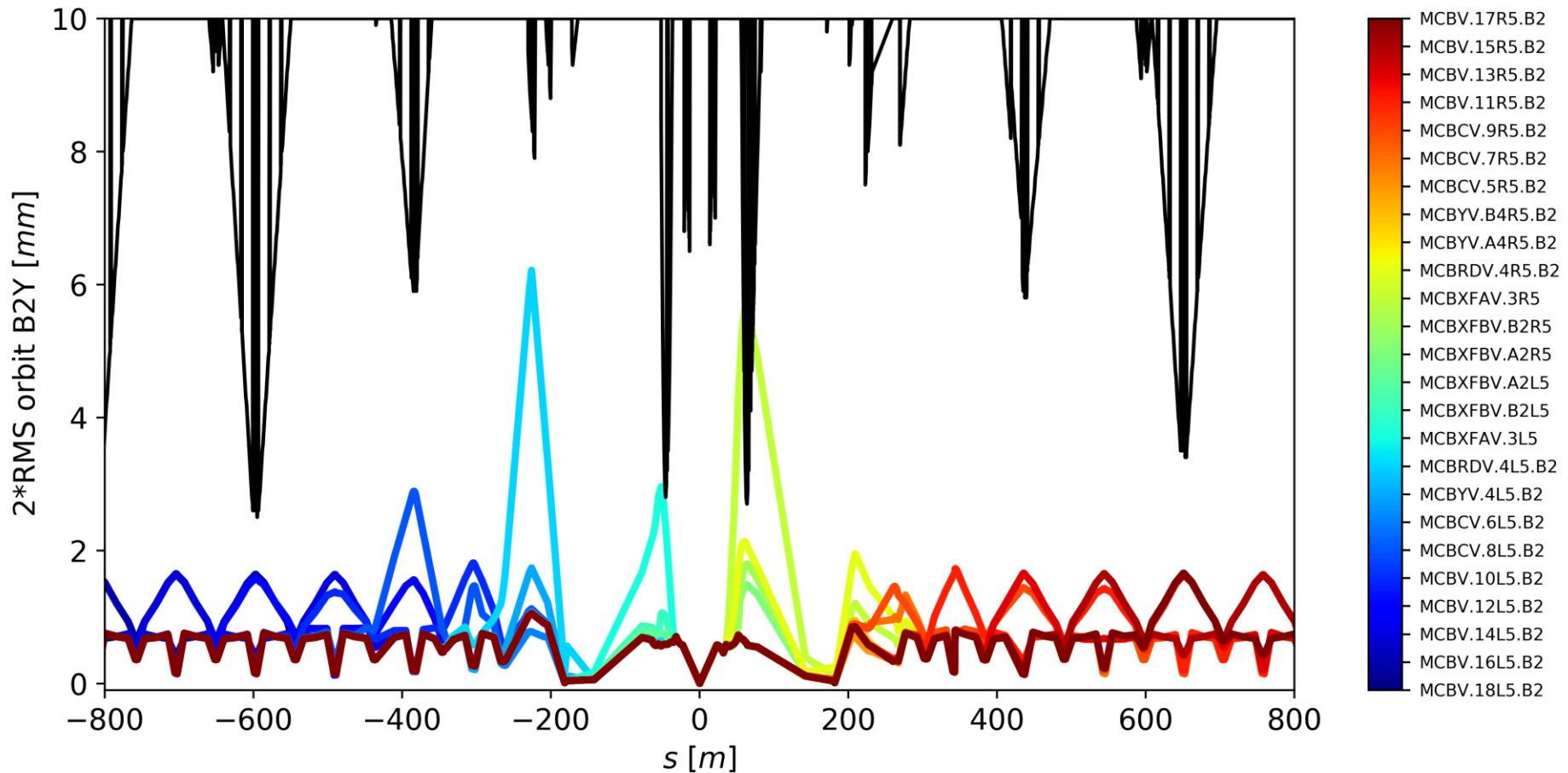
Failure scenario: B2X



Failure scenario: B2X



Failure scenario: B2Y



Failure scenario: B2Y (zoomed)

