

Gauged B-L interacting sterile neutrino dark matter revisited

Osamu Seto (Hokkaido University)

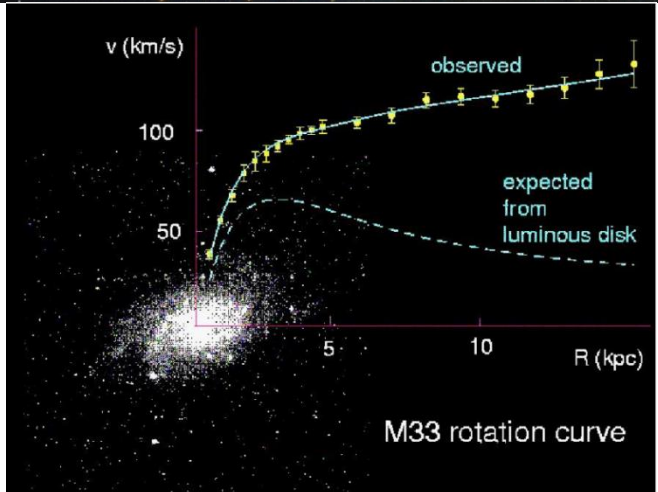
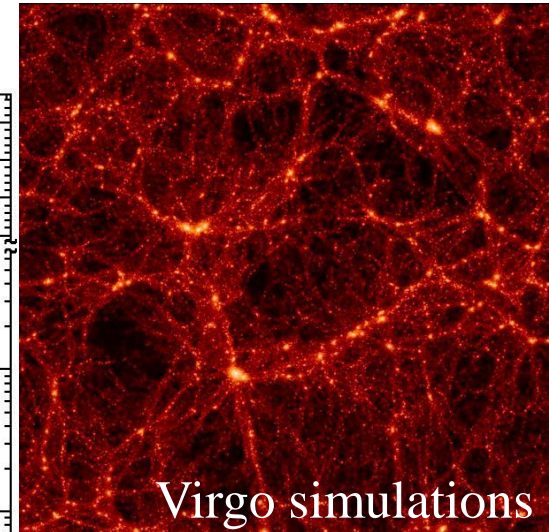
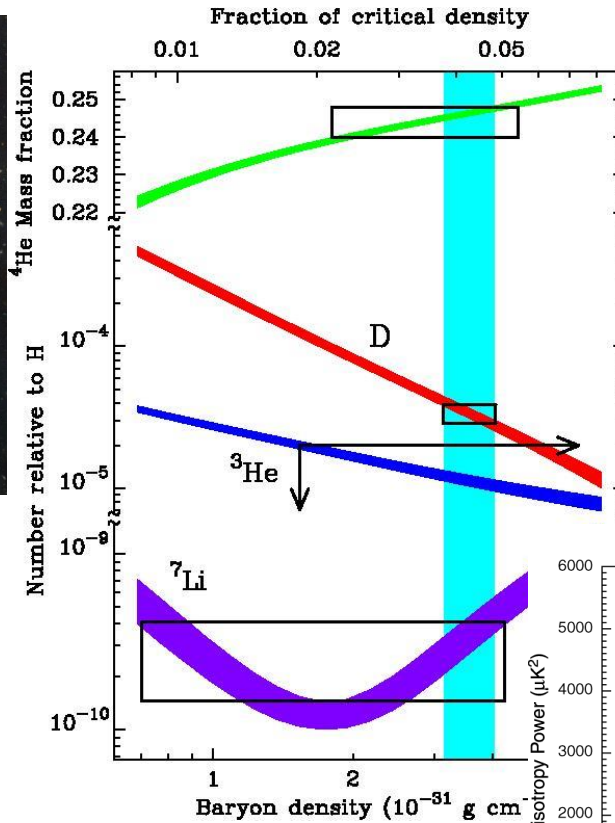
With Shintaro Eijima (ICRR, U. of Tokyo) and Takashi Shimomura (Miyazaki U., Kyushu U.)

Ref 2207.01775 (2022)

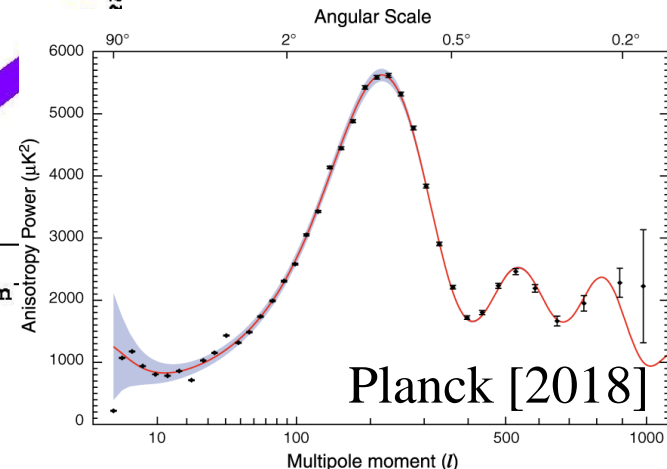
§ Introduction

Dark matter

Convincing evidences



Burles et al [1999]



Dark matter candidates

Hypothetical candidates

Sterile neutrino

- $\nu_S \cong \theta \nu_L + \nu_R^C$

Almost RH

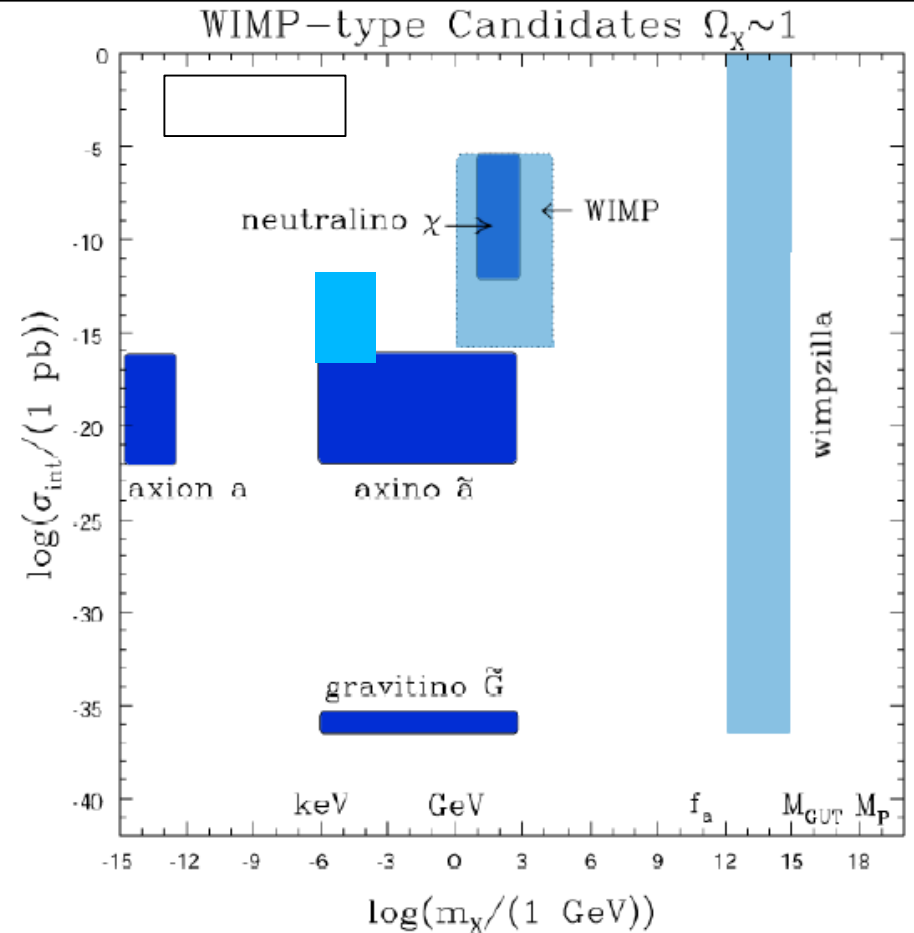
- $\theta = m_D / M_M \ll 1$

Tiny active-sterile mixing

Axion

WIMP

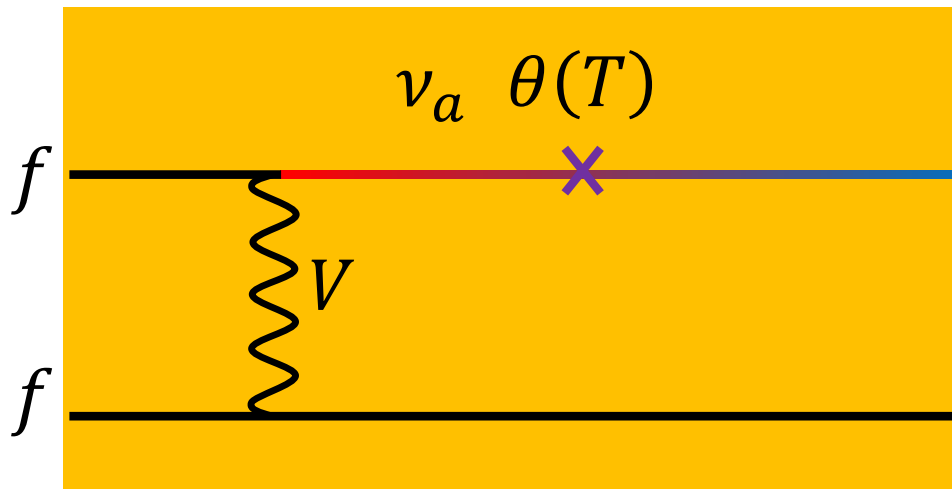
etc



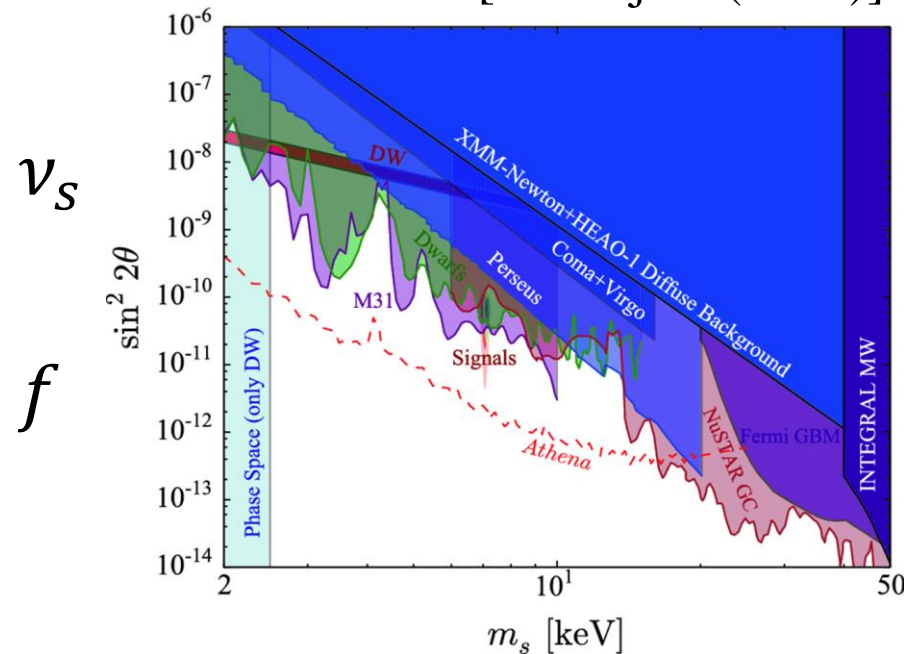
modified L. Roszkowski's diagram

Sterile neutrino is decaying DM

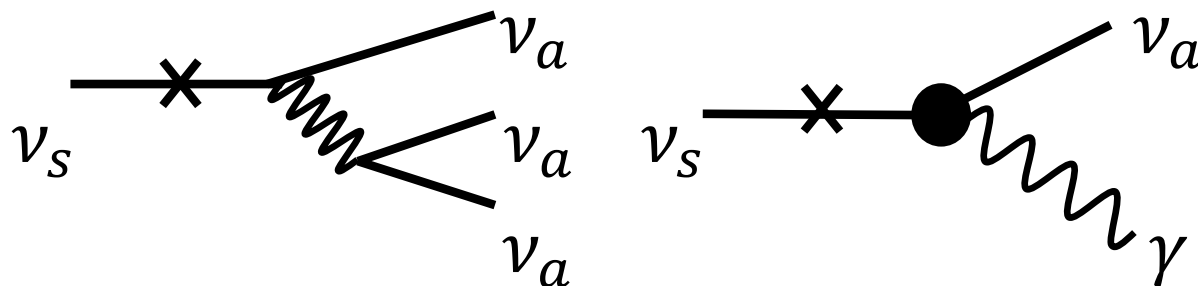
- **Production: Dodelson-Widrow mechanism** [Dodelson and Widrow (1994)]



[Abazajian (2019)]



- **Decay** [Pal and Wolfenstein (1982)]



Sterile neutrino is decaying DM

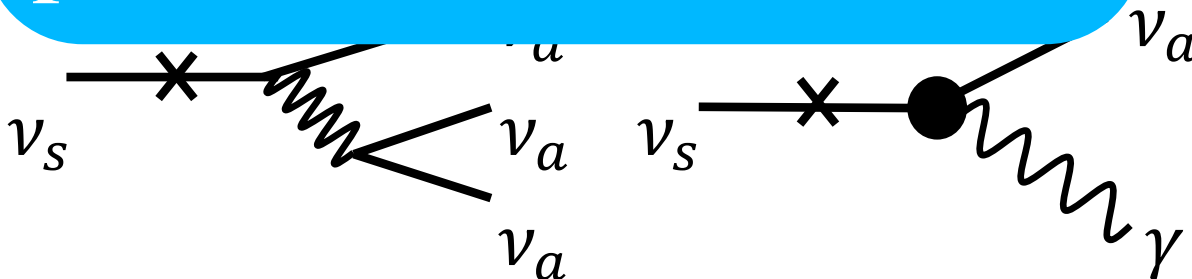
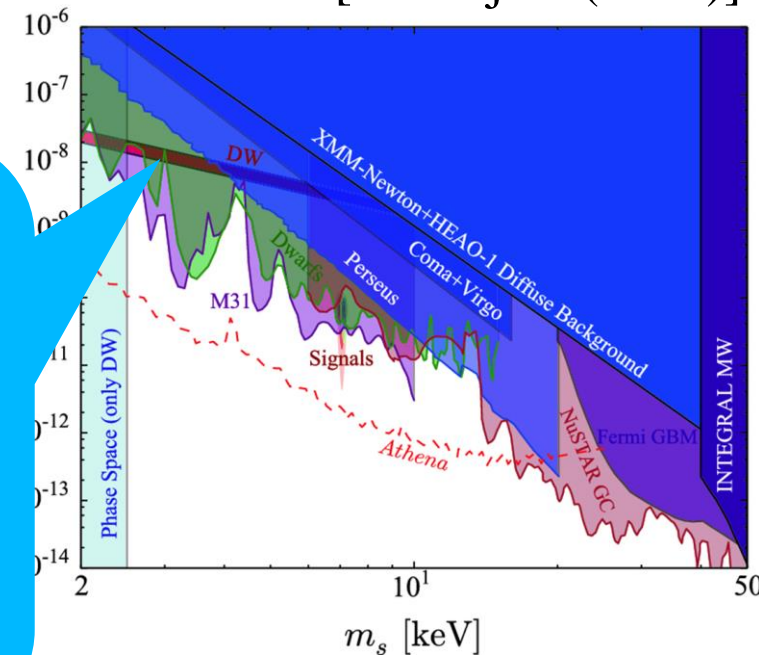
- **Production: Dodelson-Widrow mechanism** [Dodelson and Widrow (1994)]

[Abazajian (2019)]

$$\nu_a \theta(T)$$

Not compatible

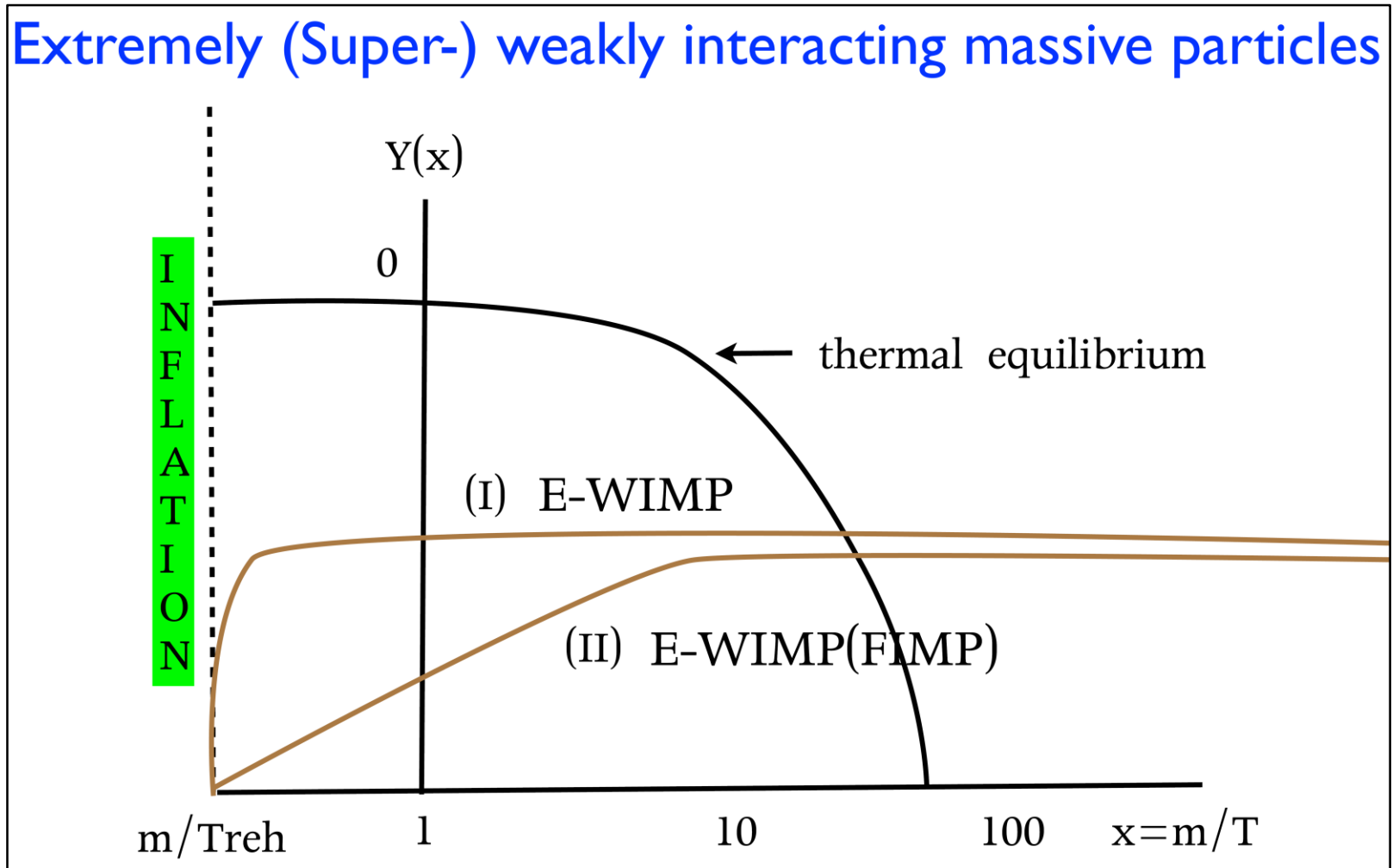
We need an alternative production mechanism



§ Freeze in Production

- A kind of non-thermal production

Extremely (Super-) weakly interacting massive particles



By courtesy of K.Y. Choi

§ Model

- Sterile neutrino DM in feeble gauged U(1) extended model
- Gauged U(1) extension
 - $U(1)_{B-L} : +1$ for baryon, -1 for lepton [Davidson (1979), Mohapatra and Marshak (1980), ...]

- Particle content

	$SU(3)_C$	$SU(2)_L$	$U(1)_Y$	$U(1)_{B-L}$
Q^i	3	2	$\frac{1}{6}$	$\frac{1}{3}$
u_R^i	3	1	$\frac{2}{3}$	$\frac{1}{3}$
d_R^i	3	1	$-\frac{1}{3}$	$\frac{1}{3}$
L^i	1	2	$-\frac{1}{2}$	-1
e_R^i	1	1	-1	-1
ν_R^i	1	1	0	-1
Φ_H	1	2	$\frac{1}{2}$	0
Φ_{B-L}	1	1	0	2

- Masses

- $m_{Z'}^2 = 4g_{B-L}^2 v_{B-L}^2$

- $m_{\nu_R^i} = \frac{y_{\nu_R^i}}{\sqrt{2}} v_{B-L}$

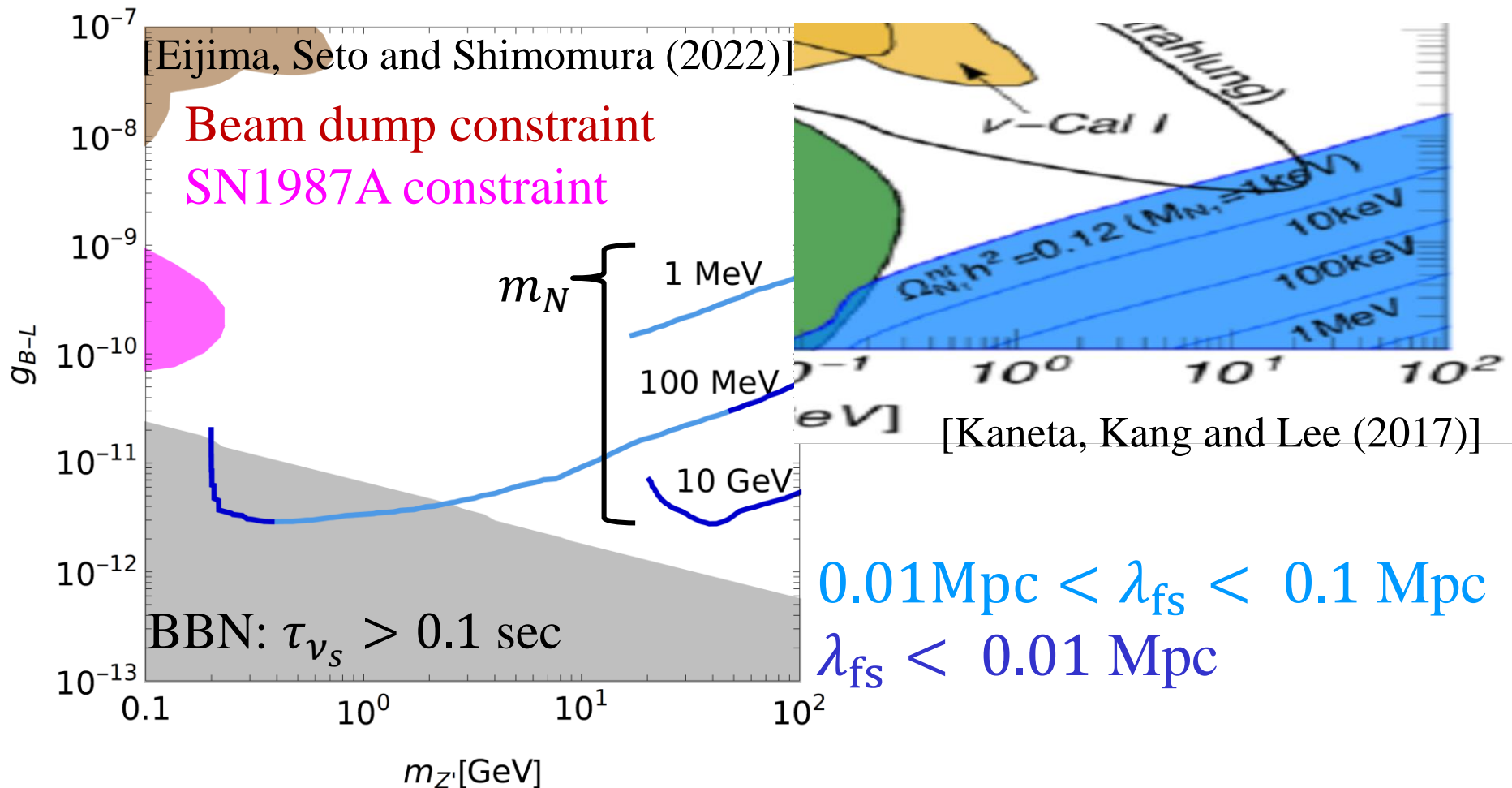
- Singlet-like ϕ

- SM-like h

- The $h - \phi$ mixing α

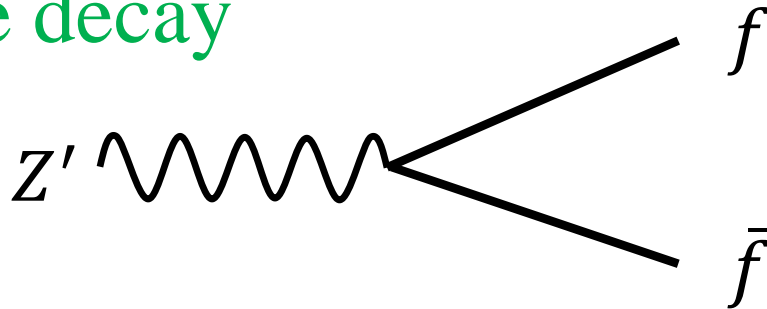
§ § Heavy Z' : $2m_N < m_{Z'}$

- Production by decay: $Z' \rightarrow 2N$ (hereafter $N = \nu_s$)
- The free streaming length λ_{fs} bound

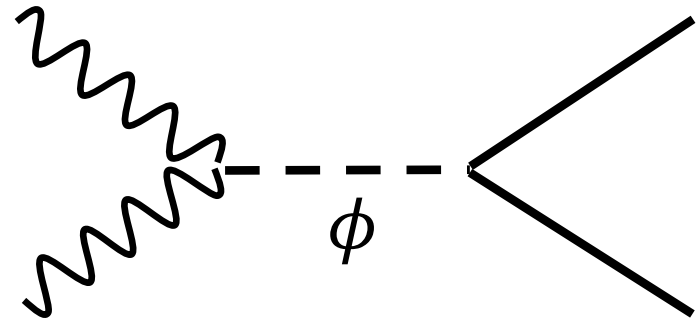
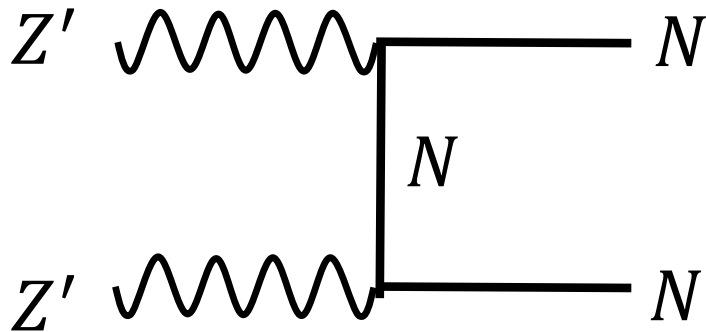


§ § Light Z' : $1\text{MeV} < m_{Z'} < 2m_N$

- Z' can be thermalized by the decay and the inverse decay



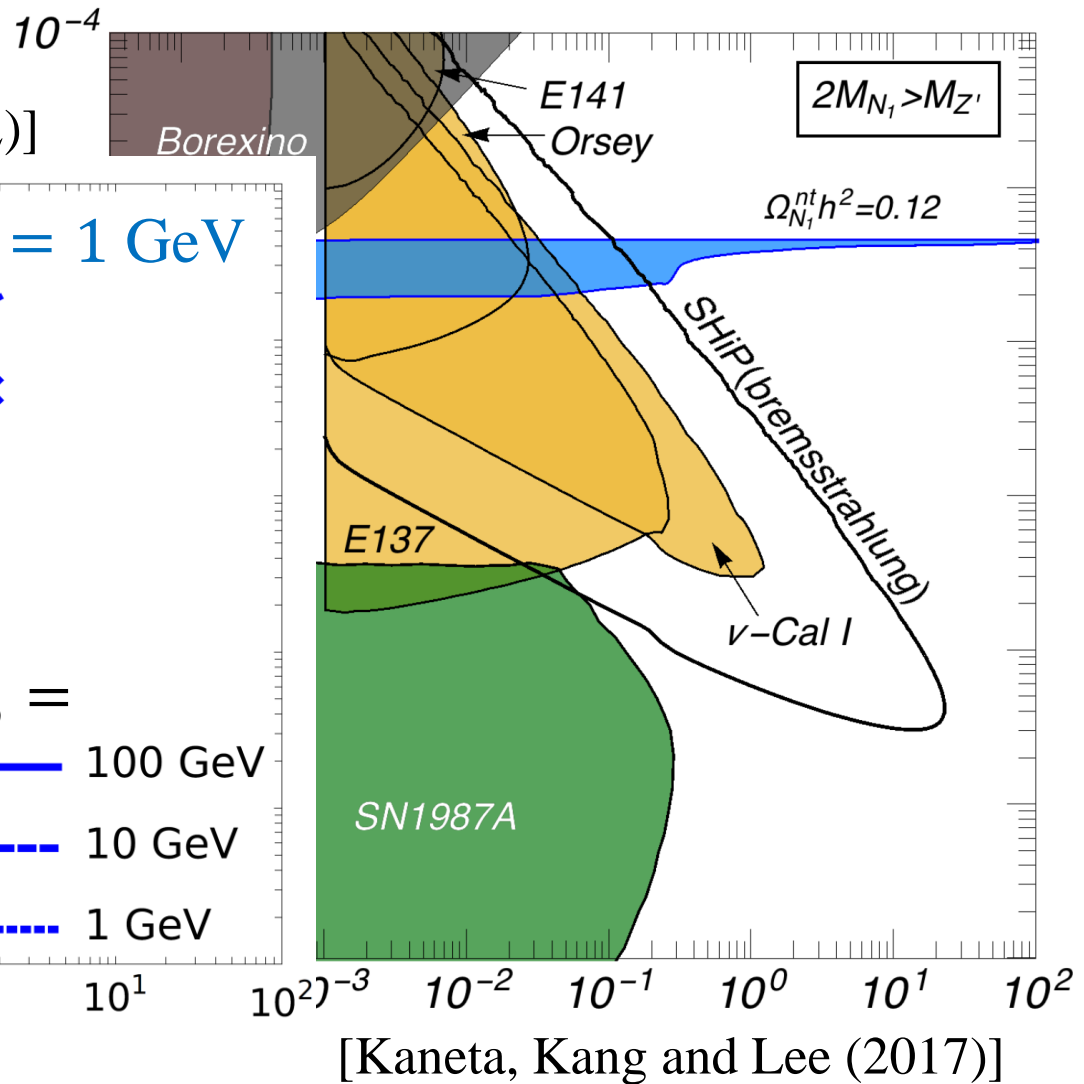
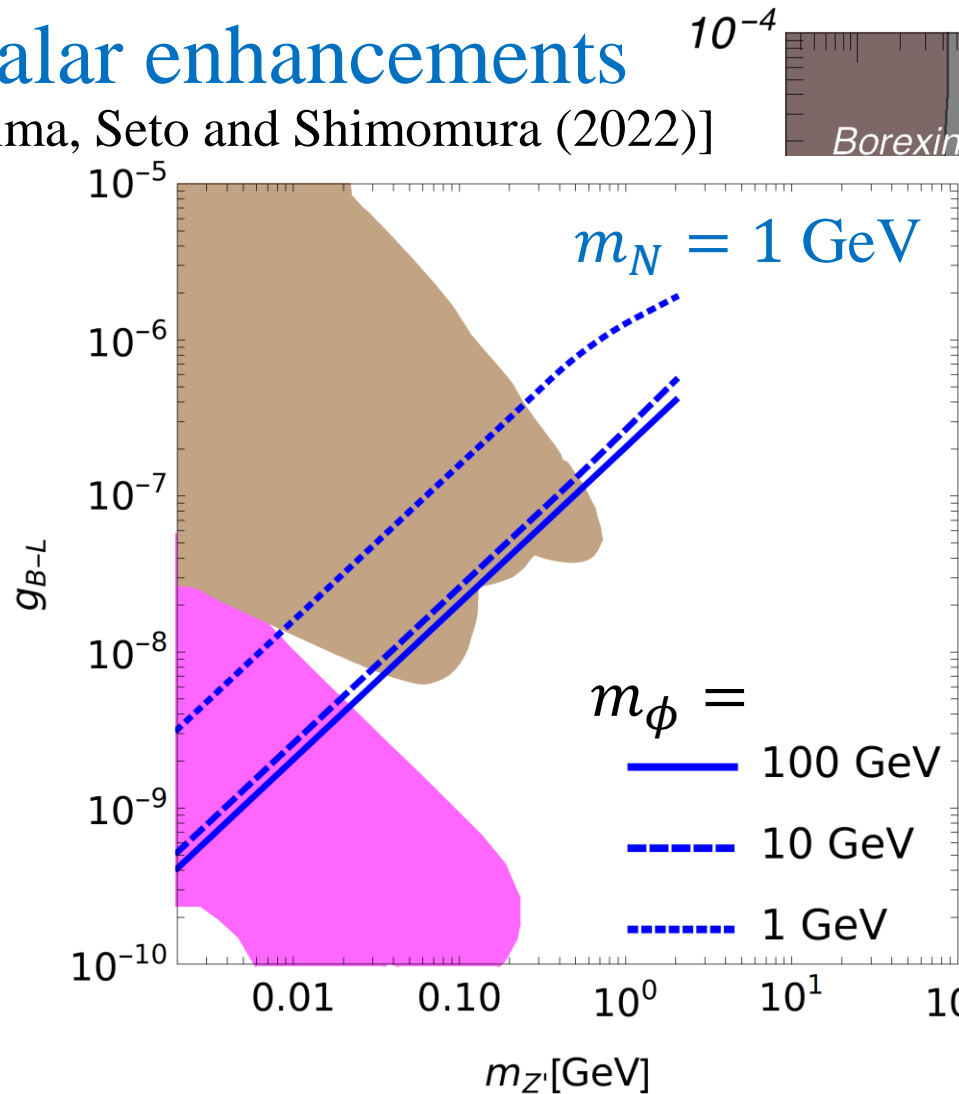
- The dominant production mode from Z'



§ § Light Z' : $1\text{MeV} < m_{Z'} < 2m_N$

longitudinal mode and
scalar enhancements

[Eijima, Seto and Shimomura (2022)]



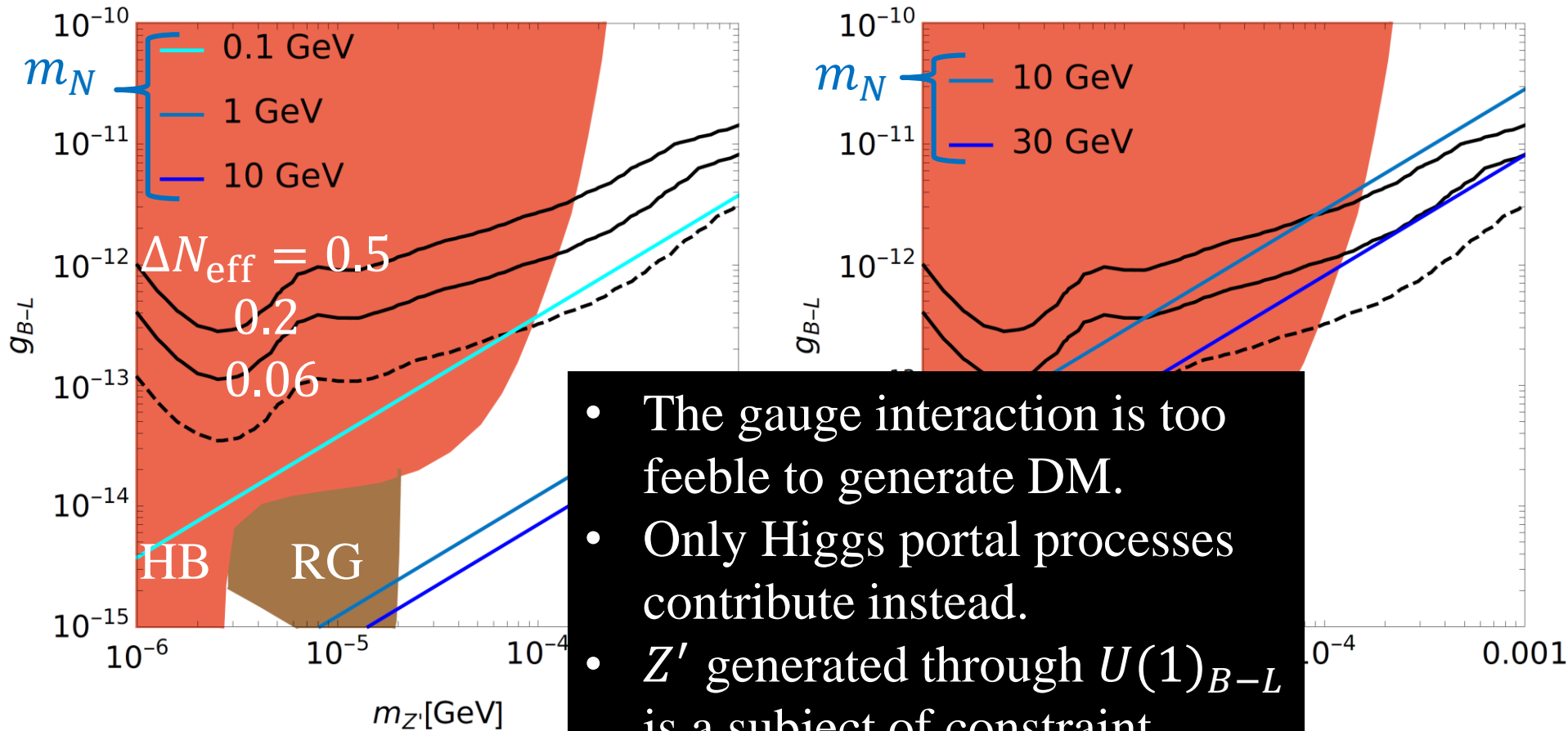
§ § Very light Z' :

$m_{Z'} < 1\text{MeV} < 2m_N$

- Z' and the decay products become dark radiation

$\alpha=0.04, m_\phi=10\text{GeV}$

$\alpha=10^{-4}, m_\phi=2\text{GeV}$



- The gauge interaction is too feeble to generate DM.
- Only Higgs portal processes contribute instead.
- Z' generated through $U(1)_{B-L}$ is a subject of constraint

§ Summary

- We reinvestigated sterile neutrino DM in gauged B-L model
- Production
 - Heavy Z' : free streaming constraints
 - Light Z' : longitudinal mode and scalar enhancements
 - Very light Z' : only Higgs portal viable
- Mass
 - $\gtrsim 1$ MeV
 - $\nu_s \rightarrow \nu_a l \bar{l}$, hadronic modes
 - ✓ c.f. X-ray from radiative decay for keV ν_s