Karlsruhe Institute of Technology

Third order corrections to the semileptonic $b \to c$ decay rate and the kinetic quark mass relation

11th International Workshop on the CKM Unitarity Triangle, 2021

Kay Schönwald | November 23, 2021

TTP KARLSRUHE

[based on: Fael, Schönwald, Steinhauser PRL 125 (2020); JHEP 10 (2020); PRD 103 (2021); PRD 104 (2021)]

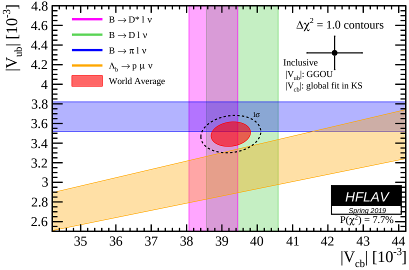
Institute for Theoretical Particle Physics

TRR 257 - Particle Physics Phenomenology after the Higgs Discovery

Outline

Motivation

- $b \to c\ell\nu$ is an important ingredient in the inclusive determination of $|V_{cb}|$:
  - Currently there is a tension between inclusive and exclusive determination of $|V_{cb}|$.
  - Errors are mostly theory dominated.
  - Precise measurements of the CKM matrix elements $|V_{ib}|$ are among main goals of Belle II and LHCb.
  - The semi-leptonic decay rate is an important ingredient in the global fit for the inclusive determination.
  - The global fits are performed in the kinetic scheme.

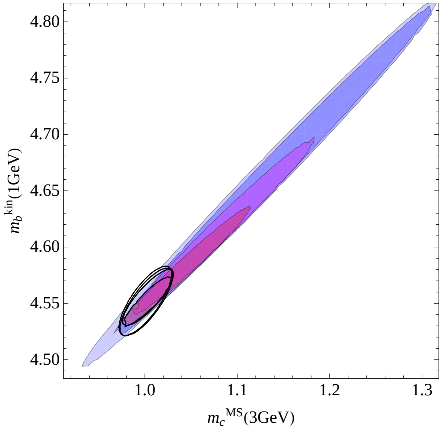


Motivation

The Heavy Quark Expansion (HQE):

$$\mathrm{d}\Gamma = \mathrm{d}\Gamma_0 + \mathrm{d}\Gamma_{\mu_\pi}\frac{\mu_\pi^2}{m_b^2} + \mathrm{d}\Gamma_{\mu_G}\frac{\mu_G^2}{m_b^2} + \mathrm{d}\Gamma_{\rho_D}\frac{\rho_D^3}{m_b^3} + \mathrm{d}\Gamma_{\rho_{LS}}\frac{\rho_{LS}^3}{m_b^3} + \ldots$$

- $\mathrm{d}\Gamma_i$ are computed in perturbative QCD
- dependece on non-perturbative HQE parameters: $\mu_\pi, \mu_G, \rho_D, \rho_{LS}, \ldots$
- Perturbative corrections to $b \to c\ell\nu$ exhibit a bad convergence in the on-shell and the $\overline{\mathrm{MS}}$ scheme for the heavy quark masses.
- Better knowledge of scheme conversion can further constrain the global fit.



[Gambino, Schwanda (PRD 89 (2014))]

	tree	α_s	α_s^2	α_s^3
1	✓	✓	✓	this talk
$1/m_b^2$	✓	✓		
$1/m_b^3$	✓	✓		
$1/m_b^{4,5}$	✓			
$\overline{m}_b - m_b^{\text{kin}}$		✓	✓	this talk

[Jezabek, Kühn, NPB 314 (1989); Gambino et al., NPB 719 (2005)]

[Melnikov, PLB 666 (2008); Pak, Czarnecki, PRD (2008)]

[Fael, KS, Steinhauser, PRD 104 (2021)]

[Alberti, Gambino, Nandi, JHEP 1401 (2014)]

[Mannel, Pivovarov, Rosenthal, PRD 92 (2015)]

[Becher, Boos, Lunghi, JHEP 0712 (2007)]

[Mannel, Pivovarov, PRD 100 (2019)]

[Dassinger, Mannel, Turczyk, JHEP 0703 (2007); JHEP 1011 (2010)]

[Fael, Mannel, Vos, JHEP 02 (2019); JHEP 12 (2019)]

[Bigi et al., PRD 56 (1997); Czarnecki, Melnikov, Uraltsev, PRL 80 (1998)]

[Fael, Steinhauser, KS, PRL 125 (2020); PRD 103 (2021)]

The Kinetic Heavy Quark Mass

The Kinetic Heavy Quark Mass

- The kinetic mass scheme is tailored for $b \to c$ transitions.
- It is based on the heavy quark – hadron mass relation in HQET:

Definition

$$M_B = m_b - \overline{\Lambda}(\mu) - \frac{\mu_\pi^2(\mu)}{2m_b} + \dots$$

[Bigi et al., PRD 56 (1997); Czarnecki, Melnikov, Uraltsev, PRL 80 (1998)]

- Key mass scale is m_b, not M_B:

$$\Gamma_{\mathrm{sl}} \simeq \frac{G_F^2 |V_{cb}|^2}{192\pi^3}(M_B - \overline{\Lambda})^5$$

The Kinetic Heavy Quark Mass

- The kinetic mass scheme is tailored for $b \to c$ transitions.
- It is based on the heavy quark – hadron mass relation in HQET:

Definition

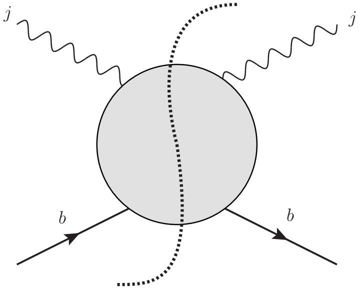
$$m^{\text{kin}} = m^{\text{OS}} - \overline{\Lambda}(\mu)\Big|_{\text{pert}} - \frac{\mu_\pi^2(\mu)\Big|_{\text{pert}}}{2m^{\text{kin}}} + \dots$$

[Bigi et al., PRD 56 (1997); Czarnecki, Melnikov, Uraltsev, PRL 80 (1998)]

- Key mass scale is m_b, not M_B:

$$\Gamma_{\text{sl}} \simeq \frac{G_F^2 |V_{cb}|^2}{192\pi^3} (M_B - \overline{\Lambda})^5$$

The Small Velocity Sum Rules

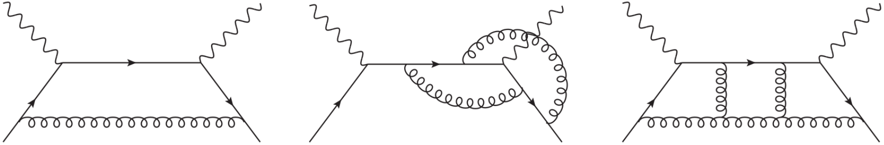


- $W(\omega, \vec{v}) = 2\,\mathrm{Im}\left[\frac{i}{2m}\int \mathrm{d}^4x\, e^{-iqx}\langle Q|\, TJ(x)J(0)\,|Q\rangle\right]$
  (J is an arbitrary current)
- $\vec{v}$: velocity of the heavy quark
- ω: excitation energy of the heavy quark

$$\overline{\Lambda}(\mu)\Big|_{\text{pert}} = \lim_{\vec{v}\to 0}\lim_{m\to\infty}\frac{2}{\vec{v}^2}\frac{\int_0^\mu \mathrm{d}\omega\,\omega\, W(\omega,\vec{v})}{\int_0^\mu \mathrm{d}\omega\, W(\omega,\vec{v})}, \qquad \mu_\pi^2(\mu)\Big|_{\text{pert}} = \lim_{\vec{v}\to 0}\lim_{m\to\infty}\frac{3}{\vec{v}^2}\frac{\int_0^\mu \mathrm{d}\omega\,\omega^2\, W(\omega,\vec{v})}{\int_0^\mu \mathrm{d}\omega\, W(\omega,\vec{v})}$$

Calculation Strategy

- We have to calculate the imaginary part of the forward scattering amplitudes:



$$W(\omega, \vec{v}) = W_{\rm el}(\vec{v})\delta(\omega) + \frac{\vec{v}^2}{\omega} W_{\rm real}(\omega)\theta(\omega) + \mathcal{O}(v^4, \frac{\omega}{m_b})$$

- Virtual corrections: only contribute to the denominator
  - given by the static limit of the massive form factor
    [Ravidran, Neerven, PLB 445 (1998),. . . , Ablinger et al. PRD 97 (2018), . . .]
- real corrections: only contribute to the numerator
  - needs to be computed in the appropriate limit

Calculation Strategy

- We can translate to a kovariant expansion:
  (p: momentum of the quark, q: momentum of the current)

$$y = m_b^2 - s = m_b\omega(2 + v^2) + \mathcal{O}(\omega^2, v^4)$$
$$q^2 = -m_b v^2(m_b - \omega) + \mathcal{O}(\omega^2, v^4)$$
$$2p \cdot q = y - q^2$$

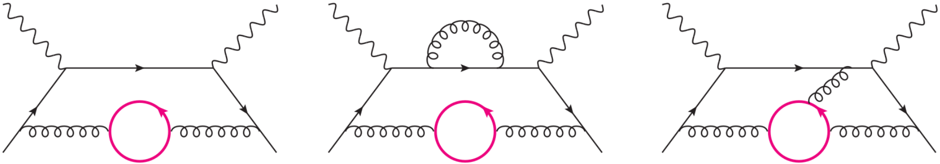
- We realize the threshold expansion via expansion by regions:
  [Beneke, Smirnov (Nucl. Phys. B (1998))]
  - the loop momenta can scale: hard (h) $k_i \sim m_b$ or ultrasoft (u) $k_i \sim y/m_b$
  - not all k_i can be hard, otherwise no imaginary part is produced.
  - we checked that we reproduce all regions with these scalings with `Asy.m`
    [Pak, Smirnov (Eur. Phys. J. C (2011))]
- expansion in $\vec{v}$ reduces to a Taylor expansion in q.

Calculation Strategy

Different kinds of master integrals appear in hard or ultra-soft regions:

- hard regions: up to two-loop on-shell master integrals.
  [Melnikov, van Ritbergen (Nucl. Phys. B (2000) ; Lee, Smirnov (JHEP (2011))]
- soft regions: three-loop ultra-soft master integrals with (massive) eikonal propagators
  $\Rightarrow$ new calculation necessary

Charm Mass Dependence



- Previously no charm mass effects for the kinetic mass had been known.
- We assume $|y| \ll m_c^2, m_b^2$, i.e. no cuts through a charm loop.
- The bare result has a non-trivial dependence on m_c.
- After renormalization only decoupling effects are left.

$\Rightarrow$ The kinetic mass conversion has no explicit m_c dependence if parametrized in terms of $\alpha_s^{(3)}$.

Results

$$
\begin{aligned}
\frac{m^{\mathrm{kin}}}{m^{\mathrm{OS}}} &= 1 - \frac{\alpha_s^{(n_l)}}{\pi} C_F \left(\frac{4}{3}\frac{\mu}{m^{\mathrm{OS}}} + \frac{1}{2}\frac{\mu^2}{(m^{\mathrm{OS}})^2} \right) + \left(\frac{\alpha_s^{(n_l)}}{\pi} \right)^2 C_F \left\{ \frac{\mu}{m^{\mathrm{OS}}} \left[C_A \left(-\frac{215}{27} + \frac{2\pi^2}{9} + \frac{22}{9} l_\mu \right) + n_l T_F \left(\frac{64}{27} - \frac{8}{9} l_\mu \right) \right] \right. \\
&\left. + \frac{\mu^2}{(m^{\mathrm{OS}})^2} \left[C_A \left(-\frac{91}{36} + \frac{\pi^2}{12} + \frac{11}{12} l_\mu \right) + n_l T_F \left(\frac{13}{18} - \frac{1}{3} l_\mu \right) \right] \right\} + \left(\frac{\alpha_s^{(n_l)}}{\pi} \right)^3 C_F \left\{ \frac{\mu}{m^{\mathrm{OS}}} \left[C_A^2 \left(-\frac{130867}{1944} \right. \right. \right. \\
&+ \frac{511\pi^2}{162} + \frac{19\zeta_3}{2} - \frac{\pi^4}{18} + \left(\frac{2518}{81} - \frac{22\pi^2}{27} \right) l_\mu - \left. \frac{121}{27} l_\mu^2 \right) + C_A n_l T_F \left(\frac{19453}{486} - \frac{104\pi^2}{81} - 2\zeta_3 \right. \\
&+ \left. \left(-\frac{1654}{81} + \frac{8\pi^2}{27} \right) l_\mu + \frac{88}{27} l_\mu^2 \right) + C_F n_l T_F \left(\frac{11}{4} - \frac{4\zeta_3}{3} - \frac{2}{3} l_\mu \right) + \left. n_l^2 T_F^2 \left(-\frac{1292}{243} + \frac{8\pi^2}{81} + \frac{256}{81} l_\mu - \frac{16}{27} l_\mu^2 \right) \right] \\
&+ \frac{\mu^2}{(m^{\mathrm{OS}})^2} \left[C_A^2 \left(-\frac{96295}{5184} + \frac{445\pi^2}{432} + \frac{57\zeta_3}{16} - \frac{\pi^4}{48} + \left(\frac{2155}{216} - \frac{11\pi^2}{36} \right) l_\mu - \frac{121}{72} l_\mu^2 \right) + C_A n_l T_F \left(\frac{13699}{1296} - \frac{23\pi^2}{54} \right. \right. \\
&- \left. \frac{3\zeta_3}{4} + \left(-\frac{695}{108} + \frac{\pi^2}{9} \right) l_\mu + \frac{11}{9} l_\mu^2 \right) + C_F n_l T_F \left(\frac{29}{32} - \frac{\zeta_3}{2} - \frac{1}{4} l_\mu \right) + \left. \left. n_l^2 T_F^2 \left(-\frac{209}{162} + \frac{\pi^2}{27} + \frac{26}{27} l_\mu - \frac{2}{9} l_\mu^2 \right) \right] \right\}, \quad (4)
\end{aligned}
$$

Results

Using as inputs $\overline{m}_b(\overline{m}_b) = 4163\,\text{MeV}$, $\alpha_s^{(5)}(M_Z) = 0.1179$:

$m_c = 0$:

$n_l = 3$: $m_b^{\text{kin}}(1\,\text{GeV}) = (4163 + 248 + 81 + 30)\,\text{MeV} = 4521(15)\,\text{MeV}$

$n_l = 4$: $m_b^{\text{kin}}(1\,\text{GeV}) = (4163 + 259 + 77 + 25)\,\text{MeV} = 4523(12)\,\text{MeV}$

$m_c \neq 0$:

$n_l = 3$: $m_b^{\text{kin}}(1\,\text{GeV}) = (4163 + 248 + 80 + 30)\,\text{MeV} = 4520(15)\,\text{MeV}$

$n_l = 4$: $m_b^{\text{kin}}(1\,\text{GeV}) = (4163 + 259 + 78 + 26)\,\text{MeV} = 4526(12)\,\text{MeV}$

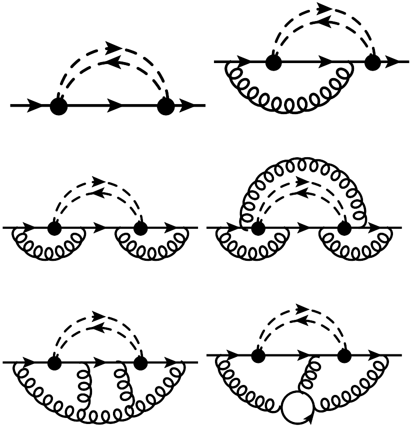
To be compared with:

- scheme conversion uncertainty at two loops: $\delta m_b^{\text{kin}} = 30\,\text{MeV}$
  [Gambino, JHEP 09 (2011)]
- m_b from $b \to c\ell\nu$ global fit: $m_b^{\text{kin}}(1\,\text{GeV}) = 4554 \pm 18\,\text{MeV}$
  [HFLAV, EPJC 81 (2021)]

The Inclusive Semileptonic Decay Rate

Method of Calculation

- We calculate the inclusive decay rate to third order via the optical theorem, i.e. we consider the imaginary part of 5-loop forward scattering diagrams.
- We consider massless leptons, i.e. we have two dimensionful scales, the bottom mass m_b and the charm mass m_c.
- Analytical dependence on charm and bottom mass seems out of reach:
  $\Rightarrow$ consider expansion in mass difference

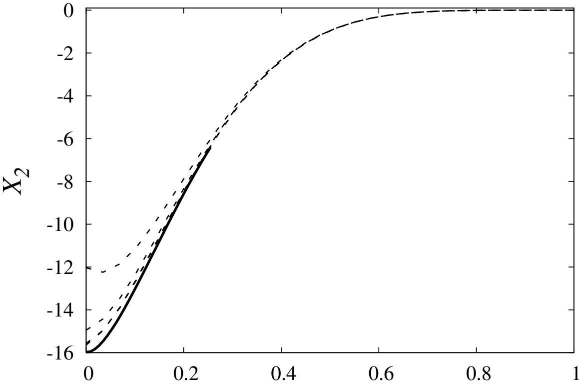


The Heavy-Daughter Expansion

- Perform the expansion in the limit $m_c \sim m_b$: $\delta = 1 - \rho = 1 - \frac{m_c}{m_b} \ll 1$
- Limit has been shown to converge well down to $m_c/m_b \to 0$ at 2-loop order.
  [Czarnecki, Dowling, Piclum (Phys. Rev. D 78 (2008))]

$$\Gamma(b \to c\ell\nu) = \Gamma_0 \left[X_0 + C_F \sum_{n\geq 1} \left(\frac{\alpha_s}{\pi}\right)^n X_n \right]$$

with $\Gamma_0 = G_F^2 m_b^2 |V_{cb}|^2 / 192\pi^3$



Details on the Calculation

- Integrate out the $(\ell\overline{\nu})$-loop.
- Loop momentum through $(\ell\overline{\nu})$-loop q:
  - In the asymptotic expansion q has to be ultra-soft (u), i.e. $q \sim \delta \cdot m_b$.
  - q factorizes out of other loops, so 1-loop tensor integral can be performed.
- The remaining loop integration have the following scalings:

	scaling	n. regions
$\mathcal{O}(\alpha_s)$	h, u	2
$\mathcal{O}(\alpha_s^2)$	hh, hu, uu	4
$\mathcal{O}(\alpha_s^3)$	hhh, huu, hhu, uuu	8

- In case a single region with either hard or ultra-soft scaling remains we can also integrate it out analytically.
- The remaining two- or three-loop integrals have integer powers of the propagators and can be reduced to master integrals via IBP reduction.

Details on the Calculation

- We have to consider 1450 five-loop diagrams.
- Several subtleties with `FORM`:
  - Automated partial fractioning and sector/families mapping.
    [LIMIT, F. Herren, PhD Thesis, KIT, 2020]
  - Major obstacle is to keep the size of `FORM` expressions as low as possible.
  - Intermediate `FORM` expressions of $O(100)$ GB.
- About 25M three-loop integrals with positive and negative indices up to 12 had to be reduce. We used (a private version of) `FIRE` together with `LiteRed` (also standalone).

Different kinds of master integrals appear in hard or ultra-soft regions:

- hard regions: up to three loop on-shell master integrals.
  [Melnikov, van Ritbergen (Nucl. Phys. B (2000) ; Lee, Smirnov (JHEP (2011))]
- soft regions: three-loop ultra-soft master integrals with (massive) eikonal propagators

Details on the Calculation

- We have to consider 1450 five-loop diagrams.
- Several subtleties with `FORM`:
  - Automated partial fractioning and sector/families mapping.
    [LIMIT, F. Herren, PhD Thesis, KIT, 2020]
  - Major obstacle is to keep the size of `FORM` expressions as low as possible.
  - Intermediate `FORM` expressions of $O(100)$ GB.
- About 25M three-loop integrals with positive and negative indices up to 12 had to be reduce. We used (a private version of) `FIRE` together with `LiteRed` (also standalone).

Different kinds of master integrals appear in hard or ultra-soft regions:

- hard regions: up to three loop on-shell master integrals.
  [Melnikov, van Ritbergen (Nucl. Phys. B (2000) ; Lee, Smirnov (JHEP (2011))]
- soft regions: three-loop ultra-soft master integrals with (massive) eikonal propagators

Renormalization

- For the renormalization of the decay width we need
  - the wave function renormalization constant Z_2
  - the mass renormalization constant Z_m

  with two massive quarks in the expansion $m_c \sim m_b$ up to $O(\alpha_s^3)$.
- Previously they were only known in the expansion $x = m_c/m_b \sim 0$ and numerically for larger values of m_c [Bekavac, Grozin, Seidel, Steinhauser, JHEP 10 (2007)] .

$\Rightarrow$ new calculation necessary

	one mass-scale	two mass-scales
$\mathcal{O}(\alpha_s)$	✓	–
$\mathcal{O}(\alpha_s^2)$	✓ [Tarrach, NPB 183 (1981)]	✓ [Broadhurst et al, ZPC 48 (1990); 52 (1991)]
$\mathcal{O}(\alpha_s^3)$	✓ [Chetyrkin et al., PRL 83 (1999); ...]	✓ [Bekavac et al. JHEP(2007)] ✓ [Fael, KS, Steinhauser, JHEP(2020)]
$\mathcal{O}(\alpha_s^4)$	✓ [Marquard et al., PRL 114 (2015); ...]	✗ [Fael, Lange, KS, Steinhauser, JHEP(2021)]

Renormalization

- For the renormalization of the decay width we need
  - the wave function renormalization constant Z_2
  - the mass renormalization constant Z_m

  with two massive quarks in the expansion $m_c \sim m_b$ up to $O(\alpha_s^3)$.
- Previously they were only known in the expansion $x = m_c/m_b \sim 0$ and numerically for larger values of m_c [Bekavac, Grozin, Seidel, Steinhauser, JHEP 10 (2007)] .

$\Rightarrow$ new calculation necessary

	one mass-scale	two mass-scales
$\mathcal{O}(\alpha_s)$	✓	–
$\mathcal{O}(\alpha_s^2)$	✓ [Tarrach, NPB 183 (1981)]	✓ [Broadhurst et al, ZPC 48 (1990); 52 (1991)]
$\mathcal{O}(\alpha_s^3)$	✓ [Chetyrkin et al., PRL 83 (1999); ...]	✓ [Bekavac et al. JHEP(2007)] ✓ [Fael, KS, Steinhauser, JHEP(2020)]
$\mathcal{O}(\alpha_s^4)$	✓ [Marquard et al., PRL 114 (2015); ...]	✗ [Fael, Lange, KS, Steinhauser, JHEP(2021)]

Results - Decay Width

$$\Gamma(b \to c\ell\nu) = \Gamma_0 \left[X_0 + C_F \sum_{n\geq 1} \left(\frac{\alpha_s}{\pi}\right)^n X_n \right]$$

$$X_3 = \delta^5 \left(\frac{266929}{810} - \frac{5248 a_4}{27} + \frac{2186\pi^2\zeta_3}{45} - \frac{4094\zeta_3}{45} - \frac{1544\zeta_5}{9} - \frac{656 l_2^4}{81} + \frac{1336}{405}\pi^2 l_2^2 + \frac{44888\pi^2 l_2}{135} - \frac{9944\pi^4}{2025} - \frac{608201\pi^2}{2430} \right) + \delta^6 \left(-\frac{284701}{540} + \frac{2624 a_4}{9} - \frac{1093\pi^2\zeta_3}{15} + \frac{391\zeta_3}{3} + \frac{772\zeta_5}{3} + \frac{328 l_2^4}{27} - \frac{668}{135}\pi^2 l_2^2 - \frac{1484\pi^2 l_2}{3} + \frac{4972\pi^4}{675} + \frac{591641\pi^2}{1620} \right) + \mathcal{O}(\delta^7 \ln^2(\delta)) ,$$

with $l_2 = \ln(2)$, $a_4 = \mathrm{Li}_4(1/2)$, $\zeta_i = \sum_{j=1}^{\infty} j^{-i}$ and $\mu_s = m_b$.

- We have calculated the expansion up to δ^{12} (for general color factors).
- A subset of color factors has been independently been computed up to δ^9.

[Czakon, Czarnecki, Dowling (Phys.Rev.D 103 (2021))]

Results - Decay Width

$$\Gamma(b \to c\ell\nu) = \Gamma_0 \left[X_0 + C_F \sum_{n\geq 1} \left(\frac{\alpha_s}{\pi}\right)^n X_n \right]$$

$$X_3 = \delta^5 \left(\frac{266929}{810} - \frac{5248 a_4}{27} + \frac{2186\pi^2\zeta_3}{45} - \frac{4094\zeta_3}{45} - \frac{1544\zeta_5}{9} - \frac{656 l_2^4}{81} + \frac{1336}{405}\pi^2 l_2^2 + \frac{44888\pi^2 l_2}{135} - \frac{9944\pi^4}{2025} - \frac{608201\pi^2}{2430} \right) + \delta^6 \left(-\frac{284701}{540} + \frac{2624 a_4}{9} - \frac{1093\pi^2\zeta_3}{15} + \frac{391\zeta_3}{3} + \frac{772\zeta_5}{3} + \frac{328 l_2^4}{27} - \frac{668}{135}\pi^2 l_2^2 - \frac{1484\pi^2 l_2}{3} + \frac{4972\pi^4}{675} + \frac{591641\pi^2}{1620} \right) + \mathcal{O}(\delta^7 \ln^2(\delta)) \,,$$

with $l_2 = \ln(2)$, $a_4 = \mathrm{Li}_4(1/2)$, $\zeta_i = \sum_{j=1}^{\infty} j^{-i}$ and $\mu_s = m_b$.

- We have calculated the expansion up to δ^{12} (for general color factors).
- A subset of color factors has been independently been computed up to δ^9.

  [Czakon, Czarnecki, Dowling (Phys.Rev.D 103 (2021))]

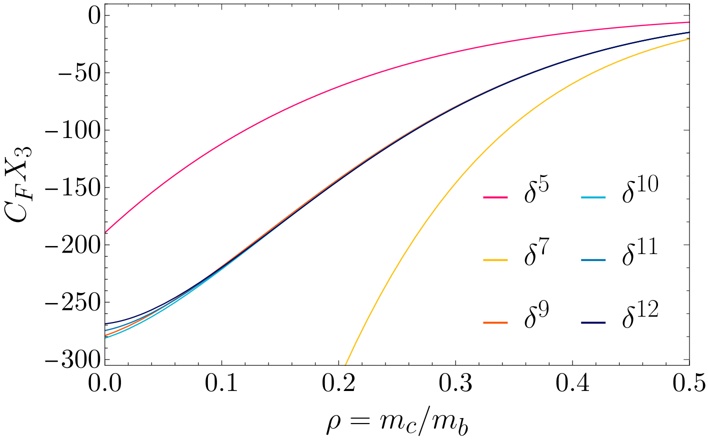
Convergence – Quark Decays

- We see a good convergence at the physical point of $\rho = m_c/m_b \approx 0.28$.
- We find:

$$X_3(\rho = 0.28) = -68.4 \pm 0.3$$

- We use the difference of the last two expansion terms to estimate the uncertainty.
- For $\rho \to 0$ we can extract values for $b \to u\ell\nu$:

$$X_3^u = -202 \pm 20$$



Different Renormalization Schemes

Karlsruhe Institute of Technology

- The total decay rate of quarks expressed in terms of on-shell masses converges poorly:

$$\Gamma_{\mathrm{sl}} \sim 1 - 1.72\frac{\alpha_s(m_b)}{\pi} - 13.1\left(\frac{\alpha_s(m_b)}{\pi}\right)^2 - 163\left(\frac{\alpha_s(m_b)}{\pi}\right)^3$$

- Also the $\overline{\mathrm{MS}}$ scheme usually behaves poorly, since the scale has to be chosen rather low.
- Different threshold masses like the PS [Beneke (1998)] , $1\mathrm{S}$ [Hoang, Ligeti, Manohar (1998)] or kinetic mass [Bigi, Shifman, Uraltsev, Vainshtein (1996)] have been proposed to improve the convergence.
- We see a much better behavior in the convergence for the schemes used for the global fits of inclusive quantities.
- E.g. for the kinetic mass:

$$m_b^{\mathrm{kin}}, m_c^{\mathrm{kin}}: \quad \Gamma(b \to c\ell\nu)/\Gamma_0 = 0.633\,(1 - 0.066 - 0.018 - 0.007) \quad \approx 0.575$$

$$m_b^{\mathrm{kin}}, \overline{m}_c(3\text{ GeV}): \quad \Gamma(b \to c\ell\nu)/\Gamma_0 = 0.700\,(1 - 0.116 - 0.035 - 0.010) \quad \approx 0.587$$

Different Renormalization Schemes

$$\Gamma(b \to c\ell\nu) = \frac{G_F^2 m_b^5 |V_{cb}|^2}{192\pi^3} X_0 \left[1 + C_F \sum_{n\geq 1} \left(\frac{\alpha_s}{\pi}\right)^n Y_n \right]$$

BLM and non-BLM part

$$Y_2 = \beta_0 Y_2^{\beta_0} + Y_2^{\mathrm{rem}}$$
$$Y_3 = \beta_0^2 Y_3^{\beta_0} + Y_3^{\mathrm{rem}}$$

	Y_1	Y_2^{rem}	$\beta_0 Y_2^{\beta_0}$	Y_3^{rem}	$\beta_0^2 Y_3^{\beta_0^2}$
$m_b^{\mathrm{OS}}, m_c^{\mathrm{OS}}$	-1.72	3.08	-16.17	48.8	-212.1
$m_b^{\mathrm{kin}}, m_c^{\mathrm{kin}}$	-0.94	0.33	-4.08	-5.4	-15.4
$m_b^{\mathrm{kin}}, \overline{m}_c(3\text{ GeV})$	-1.67	-3.39	-3.85	-97.7	69.1
$m_b^{\mathrm{kin}}, \overline{m}_c(2\text{ GeV})$	-1.25	-1.21	-2.43	-68.8	67.9
$\overline{m}_b(\overline{m}_b), \overline{m}_c(3\text{ GeV})$	3.07	-21.81	35.17	-56.7	119.4
$m_b^{\mathrm{PS}}, \overline{m}_c(2\text{ GeV})$	-0.47	-6.10	-2.31	-93.1	-7.19
$m_b^{\mathrm{1S}}, \overline{m}_c(2\text{ GeV})$	-3.59	-0.98	-19.39	-39.83	-80.22
m_b^{1S}, m_c via HQET	-1.38	0.73	-7.05	5.04	-38.09

Different Renormalization Schemes

BLM and non-BLM part

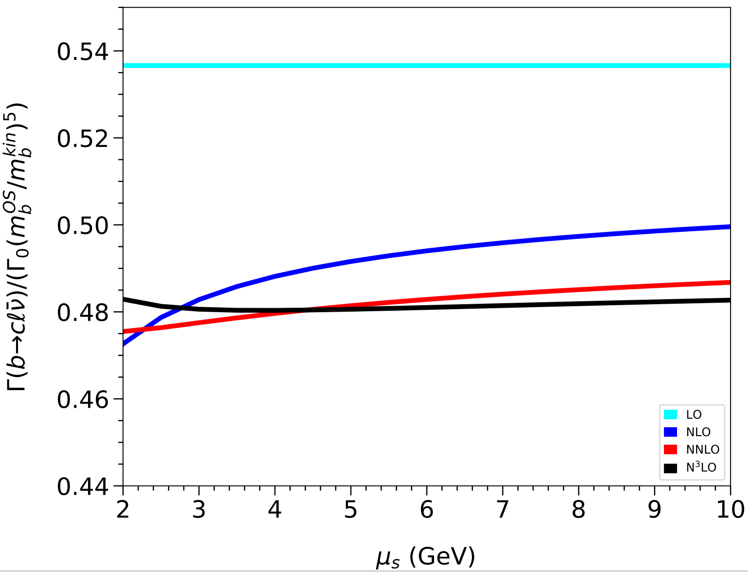
$$\Gamma(b \to c\ell\nu) = \frac{G_F^2 m_b^5 |V_{cb}|^2}{192\pi^3} X_0 \left[1 + C_F \sum_{n\geq 1} \left(\frac{\alpha_s}{\pi}\right)^n Y_n \right]$$

$$Y_2 = \beta_0 Y_2^{\beta_0} + Y_2^{\mathrm{rem}}$$
$$Y_3 = \beta_0^2 Y_3^{\beta_0} + Y_3^{\mathrm{rem}}$$

	Y_1	Y_2^{rem}	$\beta_0 Y_2^{\beta_0}$	Y_3^{rem}	$\beta_0^2 Y_3^{\beta_0^2}$
$m_b^{\mathrm{OS}}, m_c^{\mathrm{OS}}$	-1.72	3.08	-16.17	48.8	-212.1
$m_b^{\mathrm{kin}}, m_c^{\mathrm{kin}}$	-0.94	0.33	-4.08	-5.4	-15.4
$m_b^{\mathrm{kin}}, \overline{m}_c(3\text{ GeV})$	-1.67	-3.39	-3.85	-97.7	69.1
$m_b^{\mathrm{kin}}, \overline{m}_c(2\text{ GeV})$	-1.25	-1.21	-2.43	-68.8	67.9
$\overline{m}_b(\overline{m}_b), \overline{m}_c(3\text{ GeV})$	3.07	-21.81	35.17	-56.7	119.4
$m_b^{\mathrm{PS}}, \overline{m}_c(2\text{ GeV})$	-0.47	-6.10	-2.31	-93.1	-7.19
$m_b^{\mathrm{1S}}, \overline{m}_c(2\text{ GeV})$	-3.59	-0.98	-19.39	-39.83	-80.22
m_b^{1S}, m_c via HQET	-1.38	0.73	-7.05	5.04	-38.09

Different Renormalization Schemes – Kinetic Mass

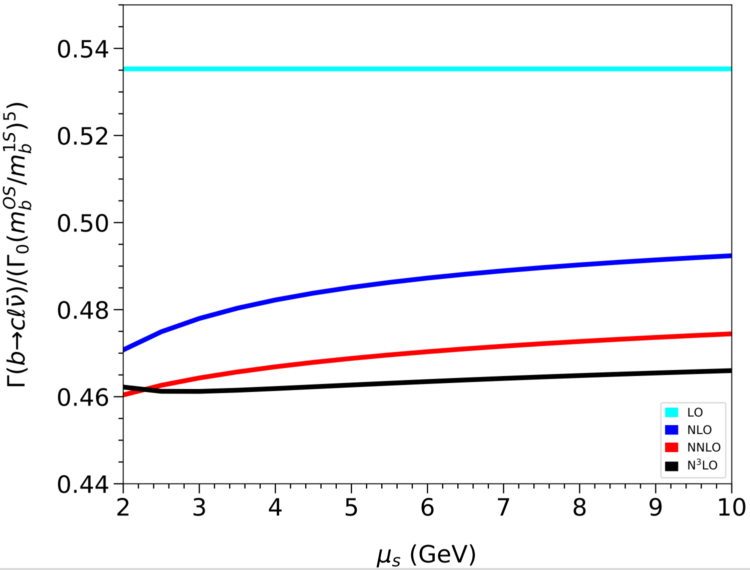
- m_b is expressed in the kinetic scheme.
- m_c is expressed in the $\overline{\mathrm{MS}}$ scheme at 2 GeV.



Different Renormalization Schemes – $1\mathrm{S}$ Mass

- m_b is expressed in the $1\mathrm{S}$ scheme.
- m_c is expressed through meson masses via the HQET relation:

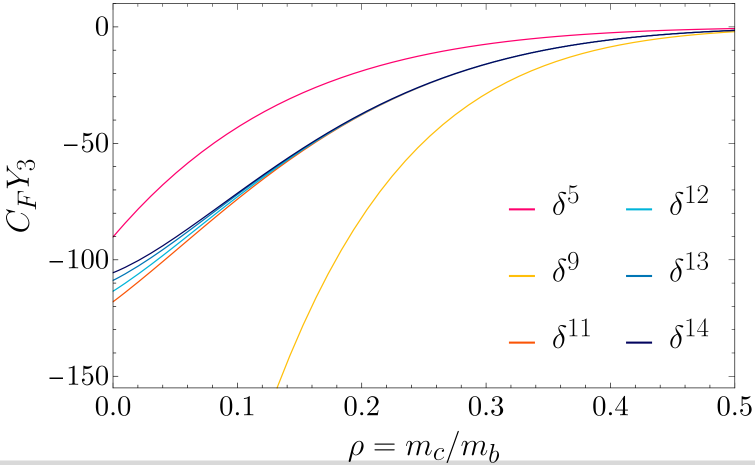
$$m_b^{\mathrm{OS}} - m_c^{\mathrm{OS}} = \overline{m}_B - \overline{m}_D + \ldots$$



Moments of Differential Distributions

- The method can be used to calculate inclusive moments of differential distributions.
- For example we can calculate-q^2 moments:

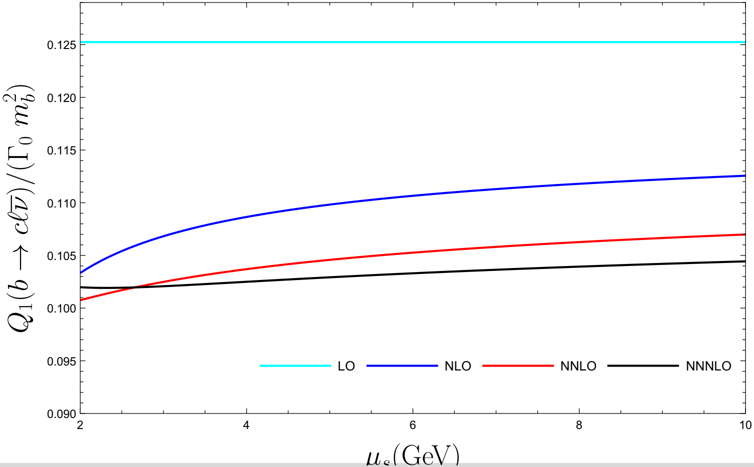
$$Q_n = \frac{1}{\Gamma_0} \int dq^2 \left(\frac{q^2}{m_b^2}\right)^n \frac{\mathrm{d}\Gamma}{\mathrm{d}q^2}$$



Moments of Differential Distributions

- The method can be used to calculate inclusive moments of differential distributions.
- For example we can calculate-q^2 moments:

$$Q_n = \frac{1}{\Gamma_0} \int dq^2 \left(\frac{q^2}{m_b^2}\right)^n \frac{\mathrm{d}\Gamma}{\mathrm{d}q^2}$$



Convergence – Muon Decays

- Specifying the color factor to QED and setting $\rho = m_e/m_\mu \approx 0$ we get the 3-loop contributions to the muon decay.
- We find:

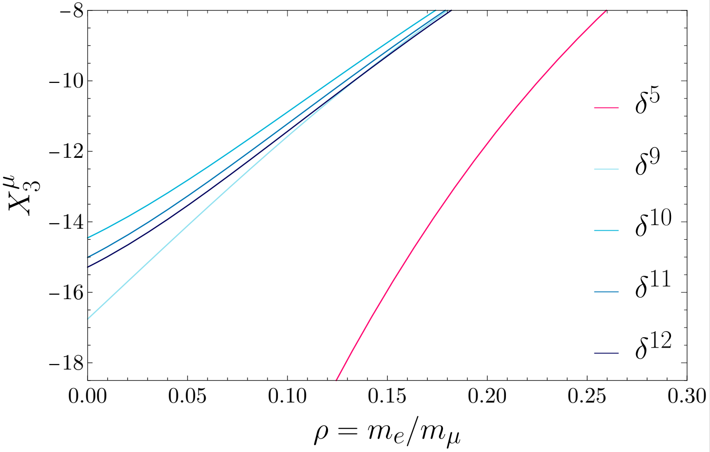
$$X_3^\mu = -15.3 \pm 2.3$$

- This leads to the shift:

$$\Delta\tau_\mu = (-9 \pm 1) \cdot 10^{-8}\,\mu s$$

- The current experimental value reads:

$$\tau_\mu = (2.1969811 \pm 0.0000022)\,\mu s$$



Conclusions and Outlook

Conclusions

- We have calculated the relation between the kinetic and on-shell mass up to $\mathcal{O}(\alpha_s^3)$.
- We have computed the α_s^3 corrections to the width of $b \to c\ell\nu$.
- We performed an expansion in the limit $1 - m_c/m_b \ll 1$ and demonstrated its good convergence.
- The result is one of the few third order corrections involving two mass scales.
- The results have been used to improve the inclusive determination of $|V_{cb}|$.
- The results are also relevant for $b \to u\ell\nu$ and the muon decay.

Outlook

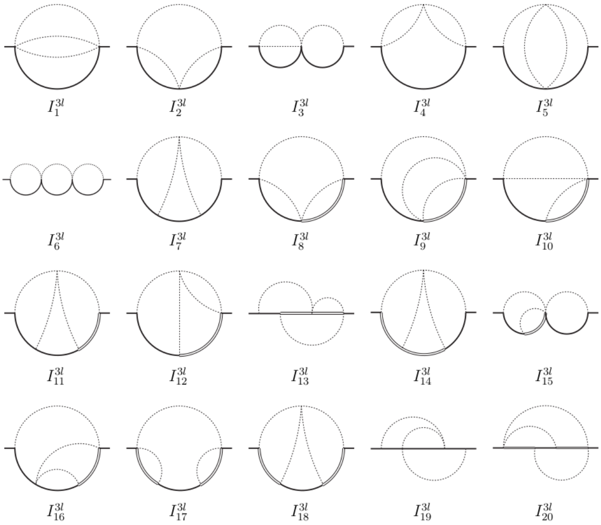
- The method of calculation can be applied for the calculation of moments of the differential distributions.

Backup

Calculation of the Master Integrals – ultrasoft regions

Three types of propagators left:

- massless propagators:
  k_i^2 (dotted lines)
- eikonal propagators:
  $2k_i.p$ (double lines)
- 'massive' eikonal propagators:
  $2k_i.p - y$ (solid lines)



Asymptotic Expansion

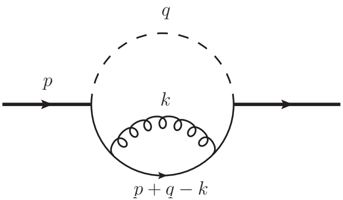
The expansion in $\delta = 1 - m_c/m_b$ is very similar to the threshold expansion for the kinetic mass:

- We use the method of regions to perform the expansion.
  [Beneke, Smirnov (Nucl. Phys. B (1998))]
- Loop momenta can either scale hard $k_i \sim m_b$ or ultra-soft $k_i \sim \delta m_b$ (regions have been cross-checked with `Asy`). [Pak, Smirnov (Eur. Phys. J. C (2011))]
- The momentum of the electron-neutrino loop can be integrated trivially.
- The properties of the asymptotic expansion allow to factorize the leptonic system completely.

$\Rightarrow$ We can reduce our 5-loop to 3-loop integrals with integer powers without any integration-by-parts.

Asymptotic Expansion – Example

Look at the 1-loop integral (we already integrated out the electron-neutrino loop):



$$\sim \int \frac{\mathrm{d}q\mathrm{d}k}{[q^2]^\alpha[(p+q)^2 - m_c^2]^2[k^2][(p+q+k)^2 - m_c^2]}$$

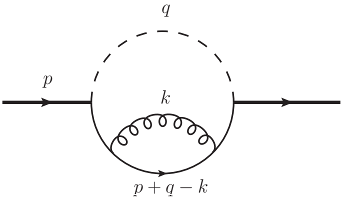
- case 1: q has to be ultra-soft, k is hard ($q \sim \delta m_b$, $k \sim m_b$);

$$\to \int \frac{\mathrm{d}q}{[q^2]^\alpha[2p\cdot q + 2m_b^2\delta]^2} \times \int \frac{\mathrm{d}k}{[k^2][(k+p)^2 - m_b^2]}$$

Asymptotic Expansion – Example

Karlsruhe Institute of Technology

Look at the 1-loop integral (we already integrated out the electron-neutrino loop):



$$\sim \int \frac{\mathrm{d}q\mathrm{d}k}{[q^2]^\alpha[(p+q)^2-m_c^2]^2[k^2][(p+q+k)^2-m_c^2]}$$

- case 2: q and k are ultra-soft ($q, k \sim \delta m_b$);

$$\to \int \frac{\mathrm{d}q\mathrm{d}k}{[q^2]^\alpha[2p\cdot q+2m_b^2\delta]^2[k^2][2p\cdot k+\underbrace{2p\cdot q+2m_b^2\delta}_{\text{fixed combination}}]}$$

$$= \int \frac{\mathrm{d}q}{[q^2]^\alpha[2p\cdot q+2m_b^2\delta]^\beta} \times \int \frac{\mathrm{d}k}{[k^2][2p\cdot k+1]}$$

The Updated Fit

- experimental moments from 2014 [Belle,Babar,CDF,CLEO,DELPHI]
- $\mathcal{O}(\alpha_s^3)$ corrections to $\Gamma(B \to X_c \ell \bar{\nu})$ [Fael,Schönwald,Steinhauser (2020)]
- $\mathcal{O}(\alpha_s^3)$ corrections to $\overline{m}_b - m_b^{\mathrm{kin}}$ relation [Fael,Schönwald,Steinhauser (2020)]
- $\overline{m}_c(3\text{ GeV}) = 0.988(7)\text{ GeV}$, $\overline{m}_b(\overline{m}_b) = 4.198(12)\text{ GeV} \longrightarrow m_b^{\mathrm{kin}} = 4.565(19)\text{ GeV}$ [FLAG (2019)]

$$|V_{cb}| = 42.16(30)_{\mathrm{th}}(32)_{\mathrm{exp}}(25)_{\Gamma} \times 10^{-3}$$

- $\Gamma(B \to X_c \ell \bar{\nu})_{\mathcal{O}(\alpha_s^3)}$:
  shift: $|V_{cb}|$ by $+0.6\%$
  reduce uncertainty: $(50)_\Gamma$ ➪ $(25)_\Gamma$
- 1.2% uncertainty
- $(32)_{\mathrm{exp}}$ ➪ improvements from Belle II

The Updated Fit

- experimental moments from 2014 [Belle,Babar,CDF,CLEO,DELPHI]
- $\mathcal{O}(\alpha_s^3)$ corrections to $\Gamma(B \to X_c \ell \bar{\nu})$ [Fael,Schönwald,Steinhauser (2020)]
- $\mathcal{O}(\alpha_s^3)$ corrections to $\overline{m}_b - m_b^{\mathrm{kin}}$ relation [Fael,Schönwald,Steinhauser (2020)]
- $\overline{m}_c(3\text{ GeV}) = 0.988(7)$ GeV , $\overline{m}_b(\overline{m}_b) = 4.198(12)$ GeV $\longrightarrow m_b^{\mathrm{kin}} = 4.565(19)$ GeV [FLAG (2019)]

$$\boxed{|V_{cb}| = 42.16(30)_{\mathrm{th}}(32)_{\mathrm{exp}}(25)_{\Gamma} \times 10^{-3}}$$

- $\Gamma(B \to X_c \ell \bar{\nu})_{\mathcal{O}(\alpha_s^3)}$:
  shift: $|V_{cb}|$ by $+0.6\%$
  reduce uncertainty: $(50)_\Gamma$ ➪ $(25)_\Gamma$
- 1.2% uncertainty
- $(32)_{\mathrm{exp}}$ ➪ improvements from Belle II