

Studies on strongly interacting matter within the FAIR Phase-0 program

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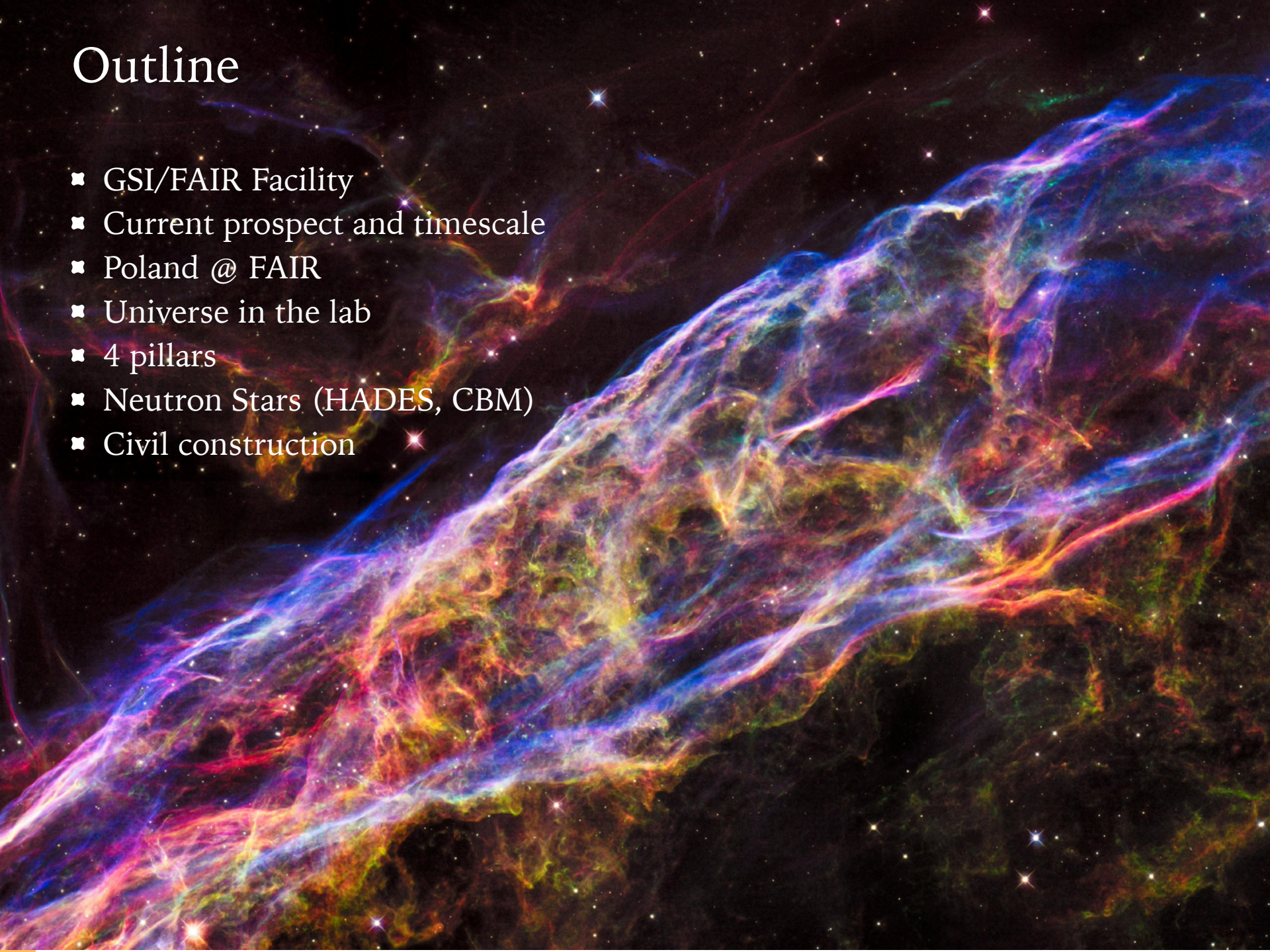
XVI Polish Workshop on Relativistic Heavy-Ion Collisions

December 3rd, 2023



Outline

- ❖ GSI/FAIR Facility
- ❖ Current prospect and timescale
- ❖ Poland @ FAIR
- ❖ Universe in the lab
- ❖ 4 pillars
- ❖ Neutron Stars (HADES, CBM)
- ❖ Civil construction





- Existing facility: GSI Darmstadt (Foundation: 1969)
- Future facility: FAIR (Foundation: 2010)
- Landmark in the European research roadmap (ESFRI)
- Employees on location: approx. 1580

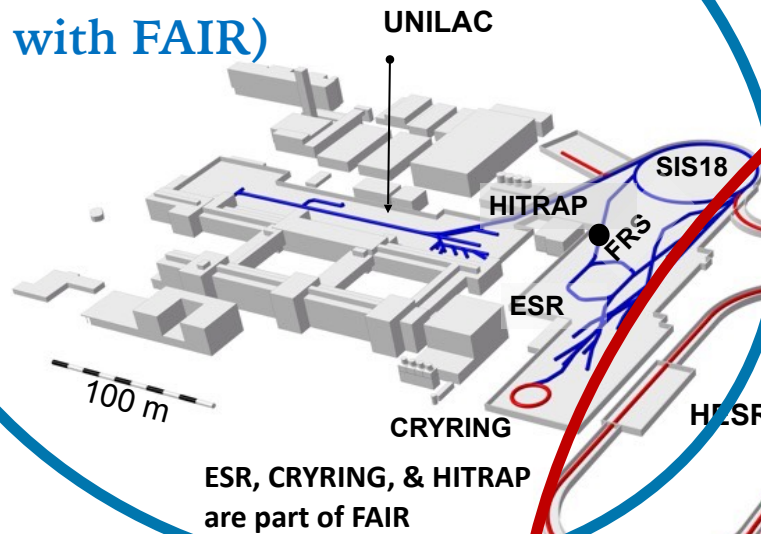
A periodic table of elements. A callout box highlights the elements Bh (107), Hs (108), Mt (109), Ds (110), Rg (111), and Cn (112), which are the elements produced at GSI. The callout box is orange and has a speech bubble shape.

107	108	109	110	111	112
Bh	Hs	Mt	Ds	Rg	Cn
Bohrium	Hassium	Mitlerium	Darmstadtium	Röntgenium	Copernicium

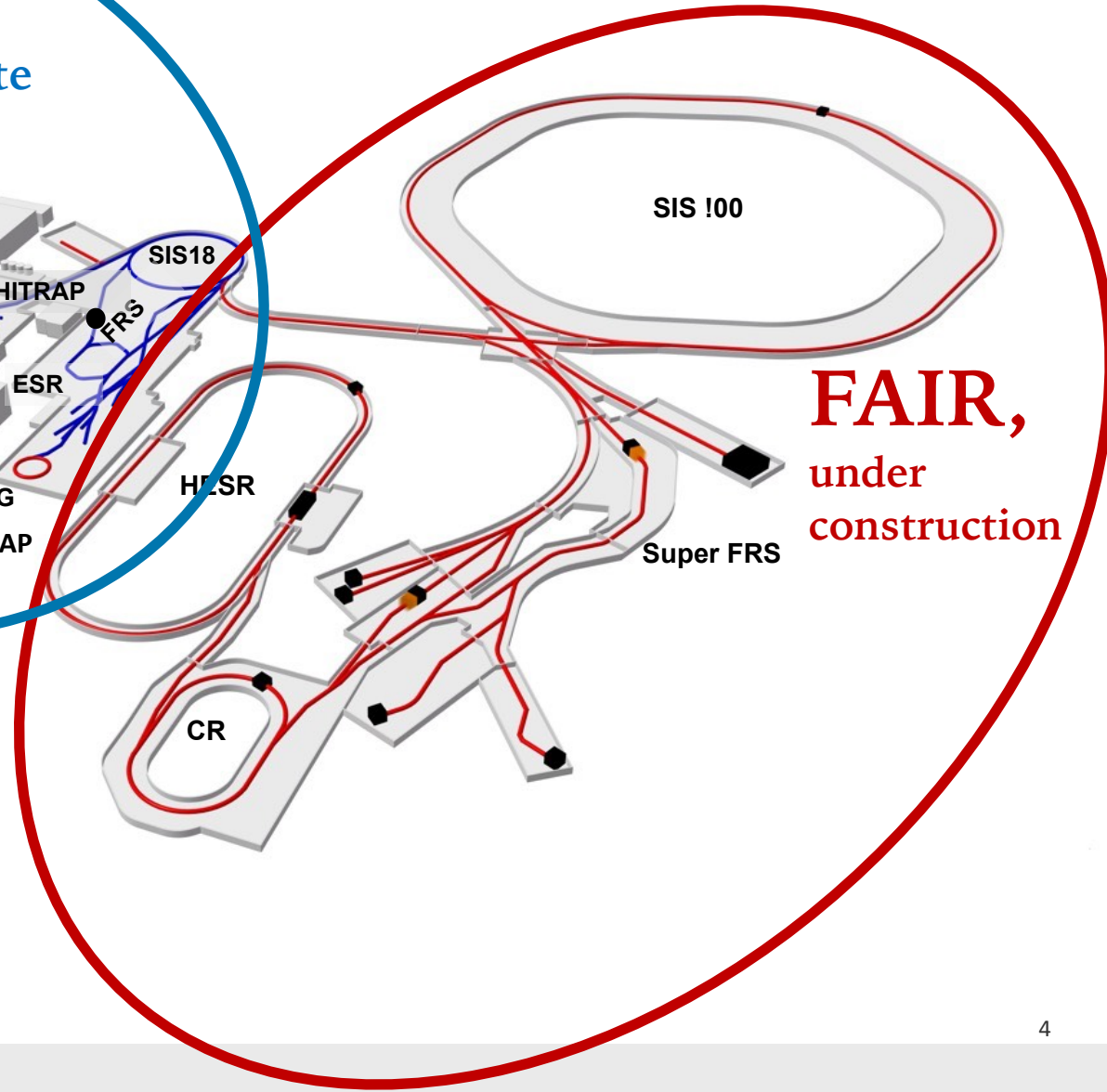
GSI and FAIR – The Facility



GSI, existing
(upgraded to integrate
with FAIR)

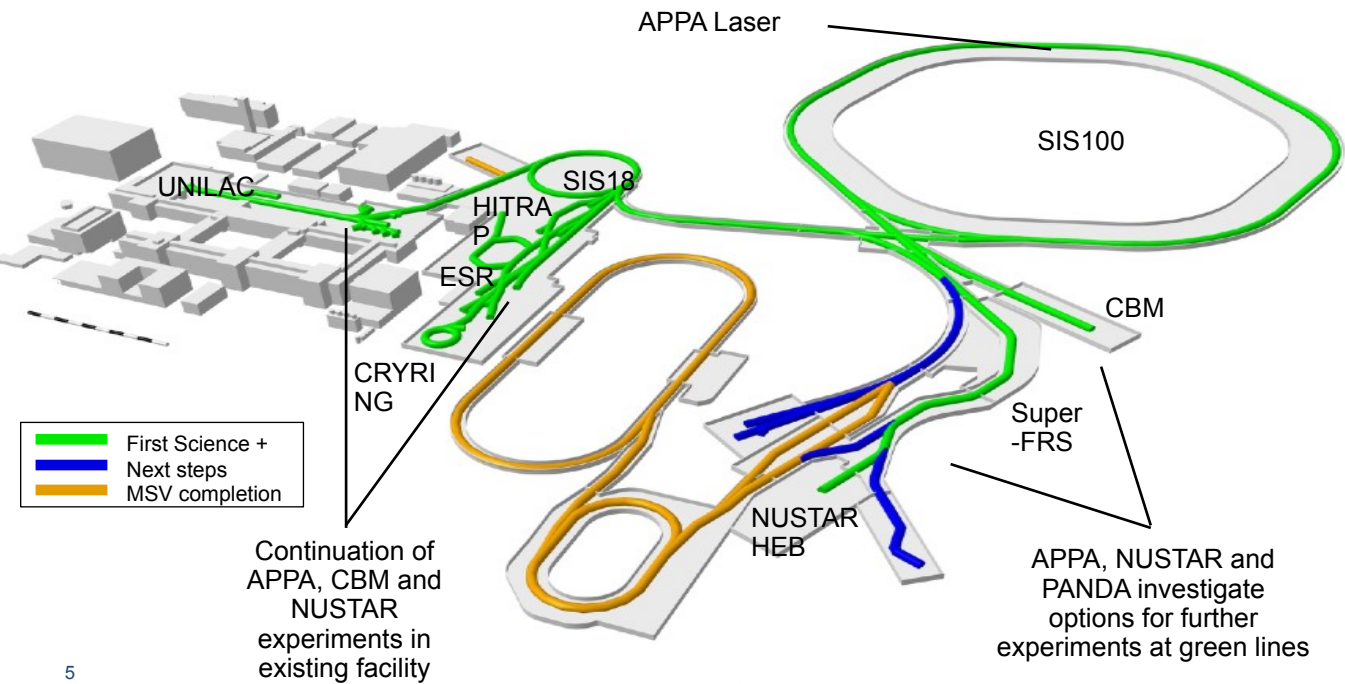


**FAIR,
under
construction**



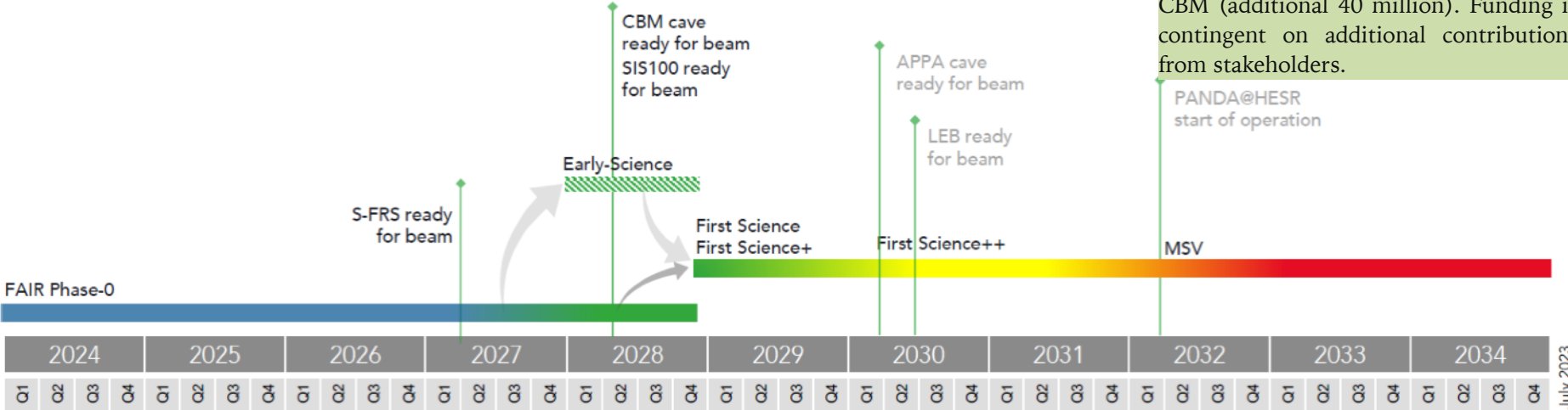
- Intensity
- Precision
- Antiproton beams

Current prospects and timeline














Staged implementation recommended by the Heuer/Tribble Commission's report (2022) with the First Science stage endorsed by the FAIR Council as "the most appropriate starting scenario to achieve world-leading science."

A budget of 2.8 billion Euros is available (including a new contribution of 0.58 billion from Germany), enabling the implementation of First Science without CBM (additional 40 million). Funding is contingent on additional contributions from stakeholders.





- FAIR governed by international convention
 - 9 shareholders:         
 - + 1 associated partner: 
 - + 1 aspirant partner: 
 - Over 3000 Scientists and Engineers from all over the world
- Scientists from More than 200 institutions from 53 countries (orange + blue)

Poland's Participation and Contribution to the FAIR Project



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APPA Physics — Atomic, Plasma Physics and Applications
CBM — Compressed Baryonic Matter experiment

NUSTAR — Nuclear STRUCTure, Astrophysics and Reactions
PANDA — Antiproton Annihilation at Darmstadt experiment

Shareholder ★
Collaborators and Contributors to FAIR ●
Collaborating Researchers and Scientists ●

Poland holds a 2.3% share, is represented by Jagiellonian University, coordinating Polish in-kind contributions to FAIR, funded by the Ministry of National Education (approximately 23.7 million Euros in 2005).

Over 95% of the funds allocated to in-kind contributions to research infrastructure and experiments at FAIR. More at <https://fair.uj.edu.pl/>.

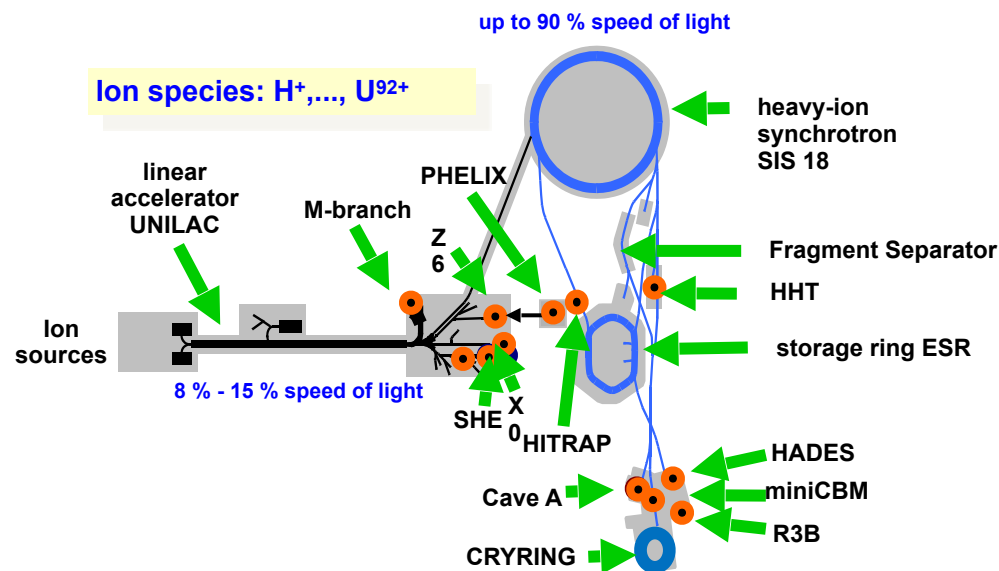
The National Consortium of Femtophysics comprises 12 Polish universities and research institutes (<https://fair.uj.edu.pl/konsorcjum>).

FAIR included in the roadmap of European and Polish research infrastructure.

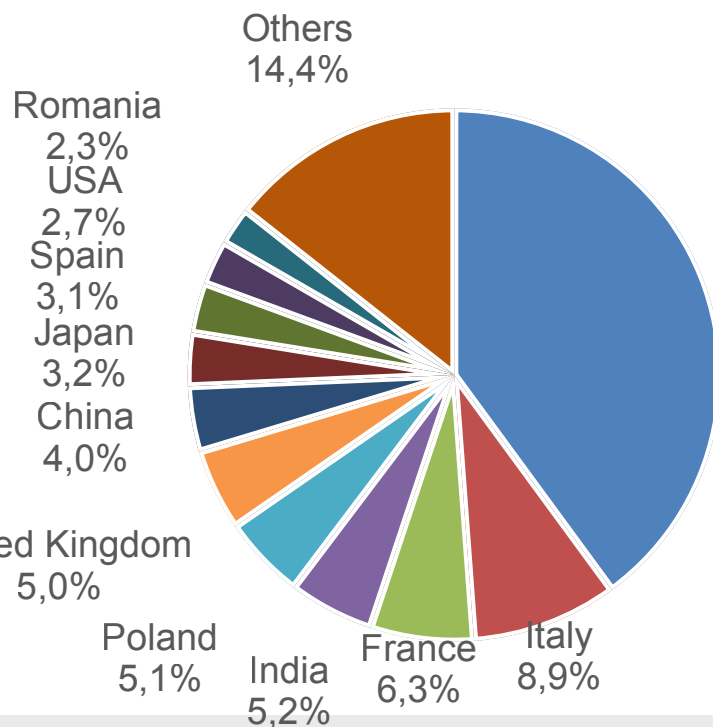
Early science program FAIR Phase-0



- Started in 2019, annual runs of ~110 days until FAIR operation

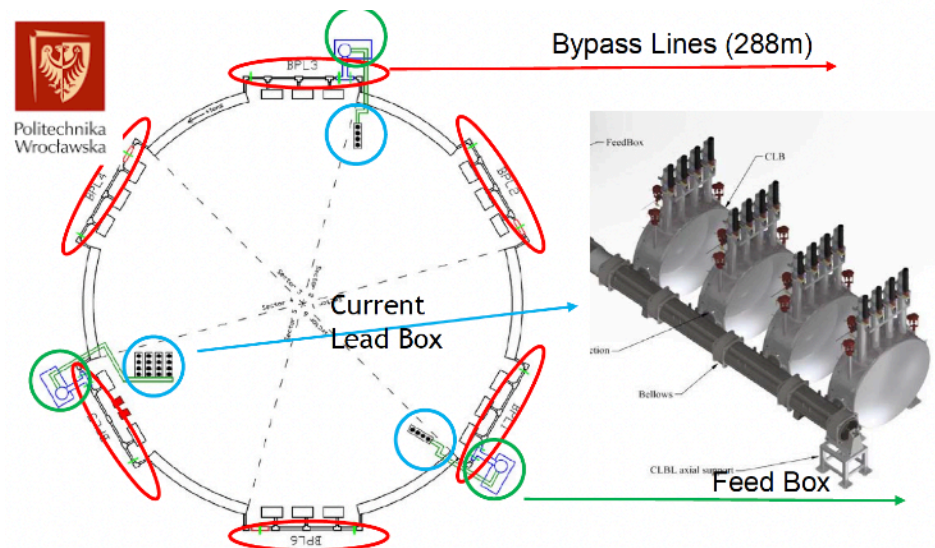


Germany
39,9%



- Science while commissioning FAIR
- latest call: 124 proposals submitted
- 1729 participants of proposals
- From institutes in 45 countries

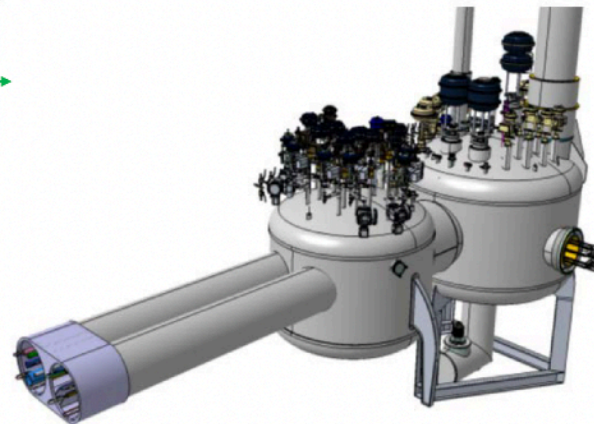
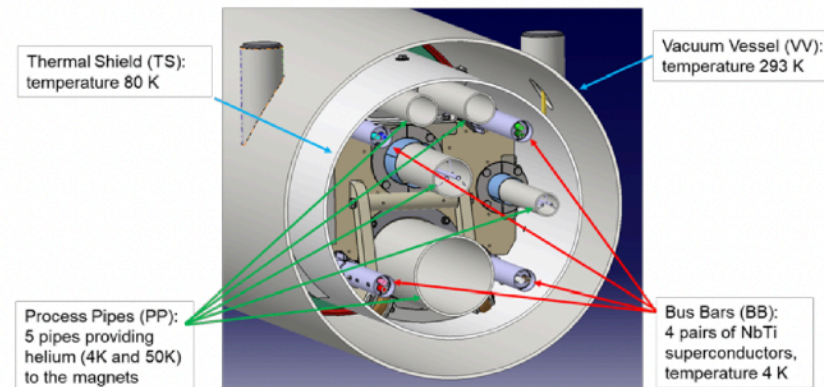
Example of Polish in-kind contributions to SIS-100



Wrocław University of Science and Technology:
Design and implementation of the power (current) and cooling (He) distribution system for the SIS100 magnets. Unique competencies combining low temperatures with superconductivity (transport of helium and current in a single vacuum insulation). A crucial contribution to First Science+.

Polish in-kind contribution to SIS100 includes Bypass Lines (288m), Feed Box, and Current Lead Box.

Cross-section of the Cryogenic Bypass Line



Gdańsk University of Technology:
Test system for SIS100 magnets.

Creating extreme
conditions existing in the
universe with
heavy ion accelerators



Foto: NASA, ESA, G. Dubner (IAFE, CONICET-University of Buenos Aires) et al.; A. Loll et al.; T. Temim et al.;
F. Seward et al.; VLA/NRAO/AUI/NSF; Chandra/CXC; Spitzer/JPL-Caltech; XMM-Newton/ESA; and Hubble/
STScI (oben), Penn State University (unten)

To find answers to fundamental questions about the Universe :
The Universe in the lab ...

Where are heavy
elements created?

NUSTAR



What is in the
interior of a
neutron star?

CBM

PANDA

Glueballs:
What are protons and
neutrons made of?
What is the structure of
hadrons?

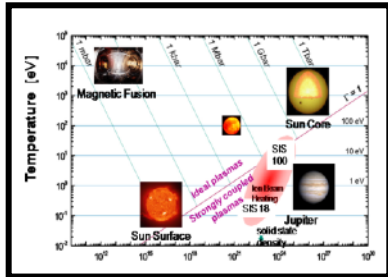


APPA

How do materials
behave under high
pressure?

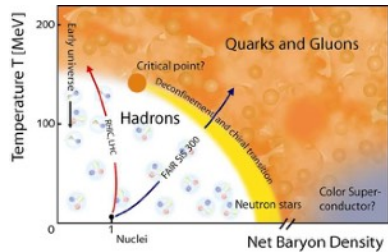


The FAIR science: four pillars



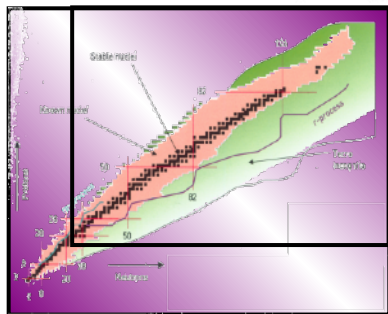
atomic physics, biophysics,
plasma physics, material research

APPA



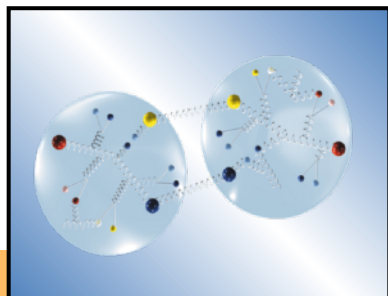
nuclear- and quark-matter

CBM



nuclear structure and
nuclear astrophysics

NuSTAR

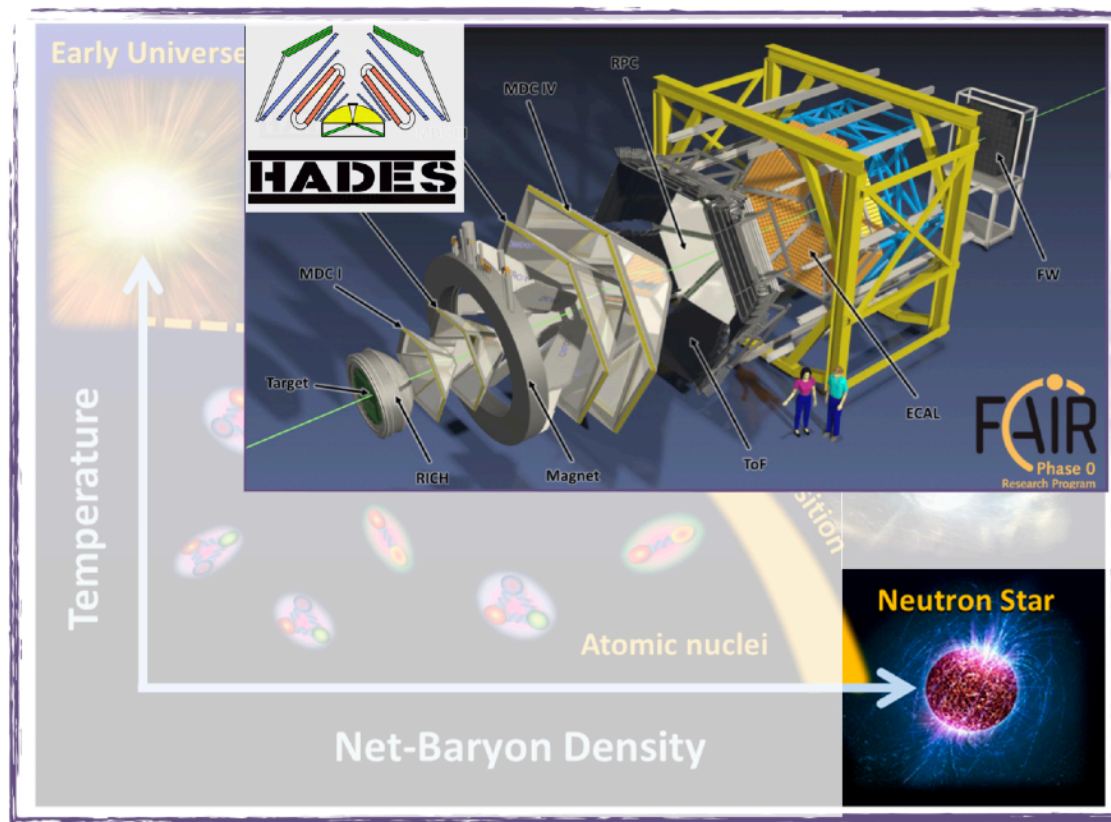


hadron structure and dynamics

PANDA



Exploration of unknown QCD territory: high μ_B



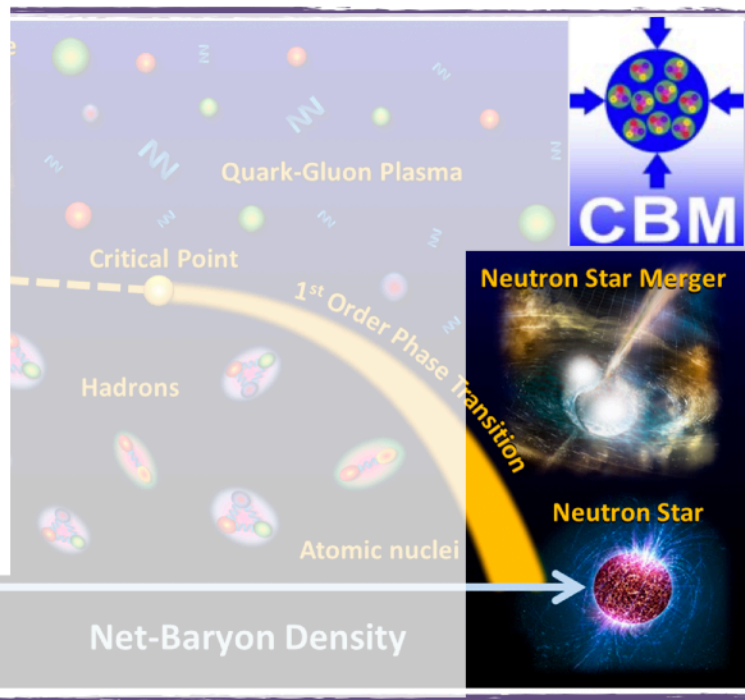
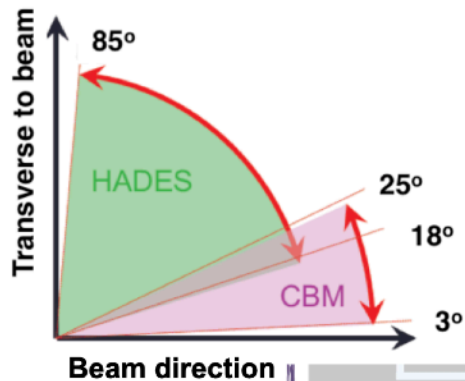
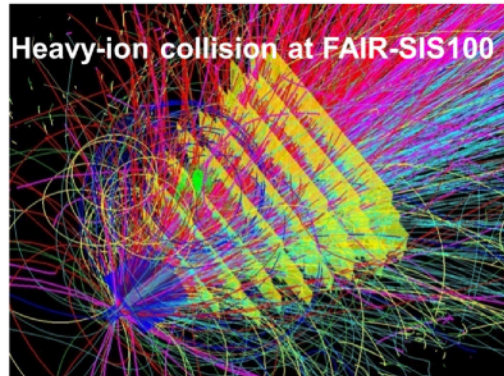
<https://www.researchgate.net>

Temperature
 $T < 10 \text{ MeV}$

Density
 $n < 3n_0$

Lifetime
 $t \sim \text{long}$

CBM and HADES future



Temperature
 $T < 50 \text{ MeV}$

Density
 $n < 2 - 6n_0$

Reaction time
 $t \sim 10 \text{ ms}$
(GW170817)

Temperature
 $T < 10 \text{ MeV}$

Density
 $n < 10n_0$

Lifetime
 $t \sim \text{infinity}$

SIS-100: Temperature $T < 120 \text{ MeV}$ Density $n < 8n_0$ Reaction time $t \sim 10^{-23} \text{ s}$

<https://www.researchgate.net>

NSM and HIC

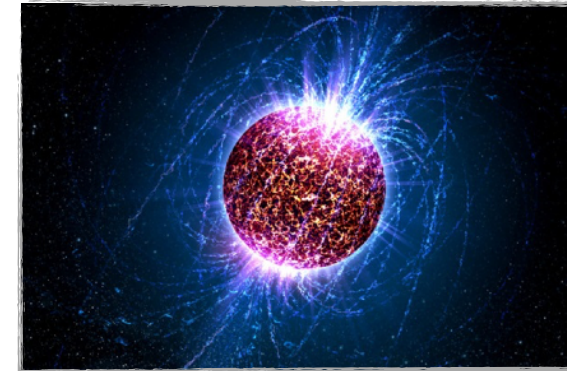
Top row: simulation of neutron stars mergers

2 neutron stars of $1.35 M_{\odot}$ each,

merging into a single object ($2R \sim 10 \text{ km}$, $n \simeq 5n_0$, $T \leq 20 \text{ MeV}$).

Overlap region: $t \simeq 20 \text{ ms}$, $n \simeq 2n_0$, $T \simeq 75 \text{ MeV}$

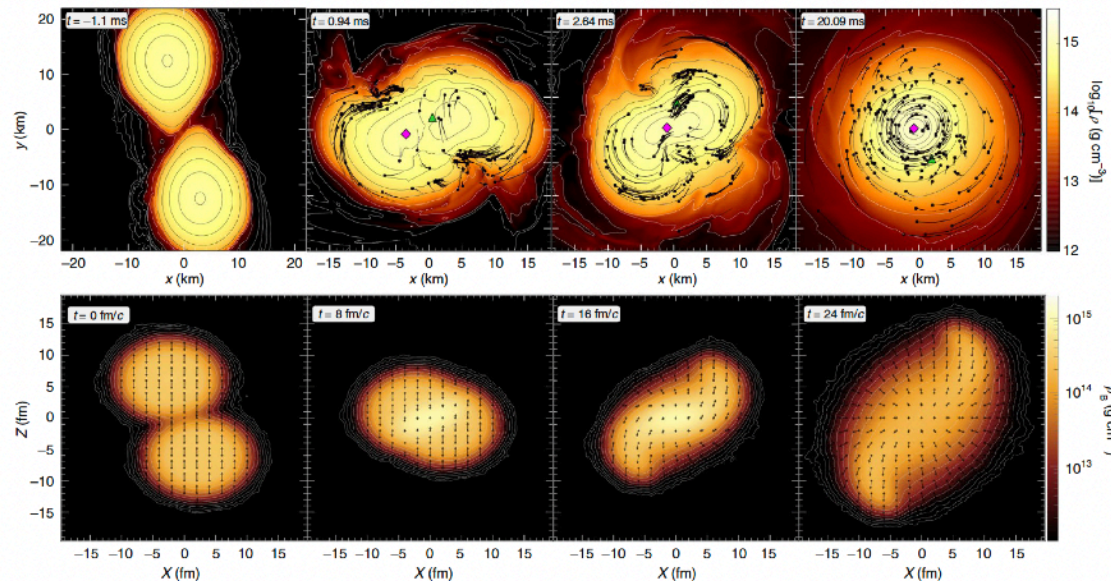
- max. temperature
- max. density



Bottom row: non-central-collision Au+Au at $\sqrt{s_{NN}} = 2.42 \text{ GeV}$

$n \simeq 3n_0$, $T \simeq 80 \text{ MeV}$

HADES, *Nature Phys.* 15, 1040–1045 (2019)



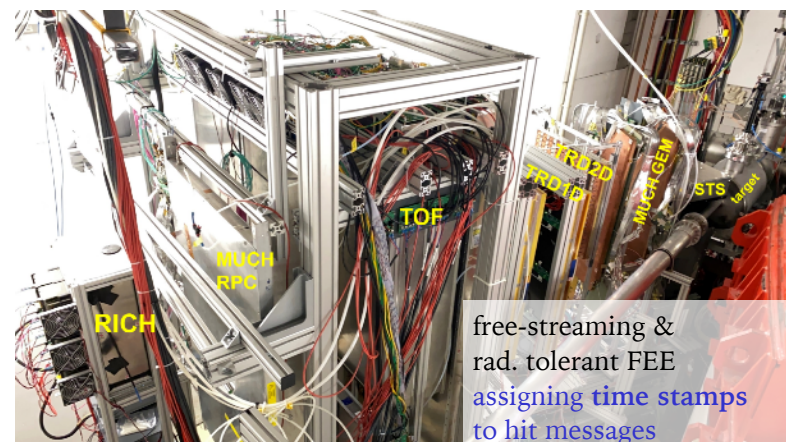
Similar **densities** and **temperatures** are achieved.

Space and **time** scales are vastly different

(km - NS, fm - HIC).

The collision events differ in duration by 20 orders of magnitude.

With mCBM@SIS18 towards CBM



free-streaming &
rad. tolerant FEE
assigning time stamps
to hit messages



FLES entry nodes
CRI FPGA
 μ Slice building
(DAQ container)



optical
fibers
50m

optical
fibers
300 m

FLES processing
nodes
time slice building
event reconstruction
& selection
archiving

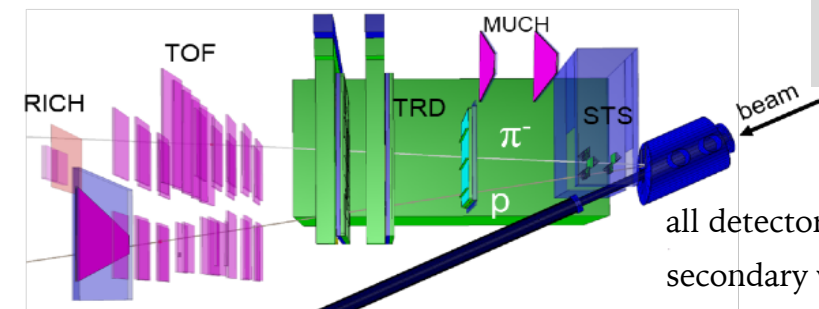


**Free-streaming
CBM data transport**

Pre-series productions
of all CBM detector
systems

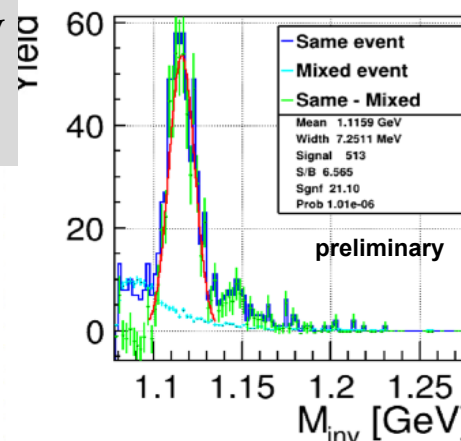
High-rate studies
up to 10 MHz coll. rate
in nucleus-nucleus
collisions

Rare signal reconstructed: $\Lambda \rightarrow p \pi^-$



all detector systems involved
secondary vertex
velocity windows for p and π^-
candidate

Ni+Ni 1.93 AGeV
run 2391 (May '22):
 10^9 collisions, 1:57h
400 kHz av. coll. rate



Campaign 2024:

high-rate studies
online reconstruction
& selection
 Λ baryons in Ni+Ni at
1.0 - 1.93 AGeV

Key observables – systematic measurements!

- **Dileptons**

→ Emissivity of dense baryonic matter: lifetime, temperature, density, in-medium properties

- **Fluctuations**

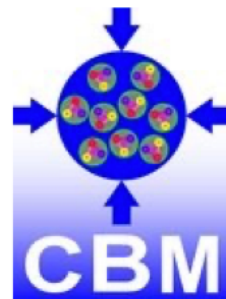
→ System transition via first-order phase transition line, CEP

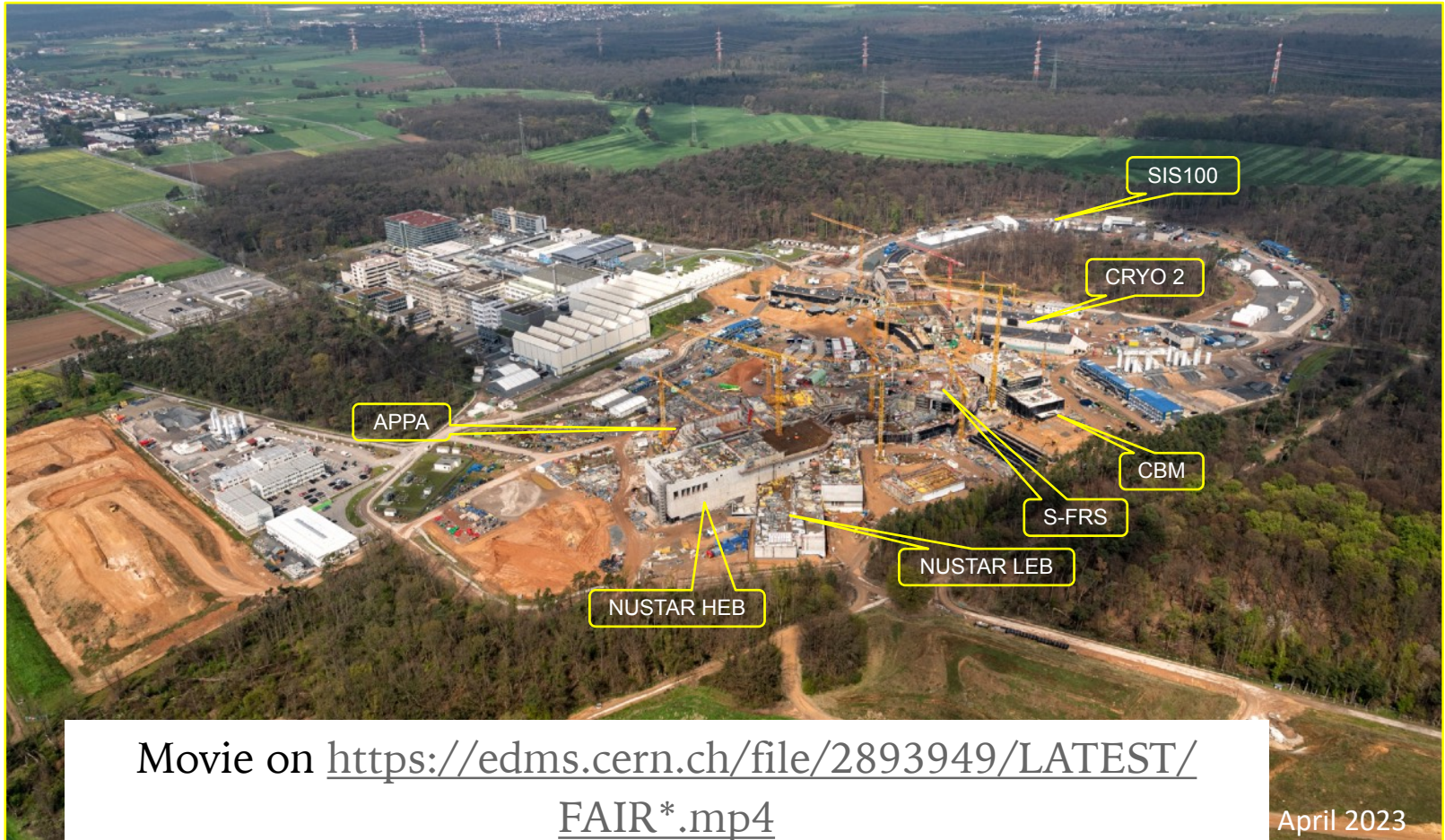
- **Hadrons/ Strangeness/ Charm**

→ System in equilibrium, Hypernuclei, Vorticity, Flow, EOS

- **Correlations**

→ Flow, Vorticity, YN & YNN interactions





FAIR SIS100 accelerator tunnel



FAIR SIS 100 supply tunnel

April 2023



FAIR Area South



FAIR visit

July 2023





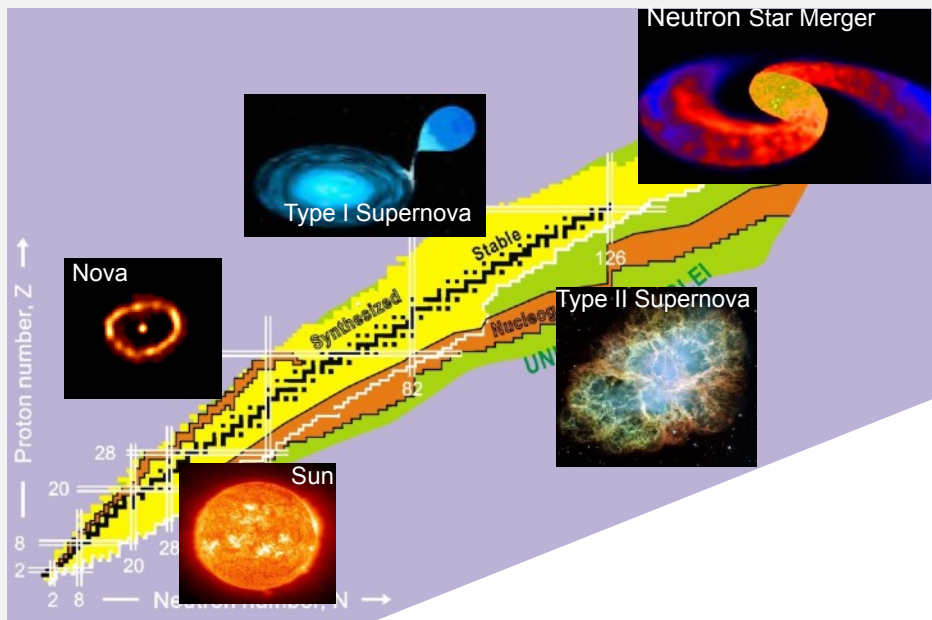
FAIR CBM Cave

FAIR: Unique Opportunities . . . & Challenges





NUSTAR - Origin of Elements in the Universe



„Nucleosynthesis sites“ in the universe

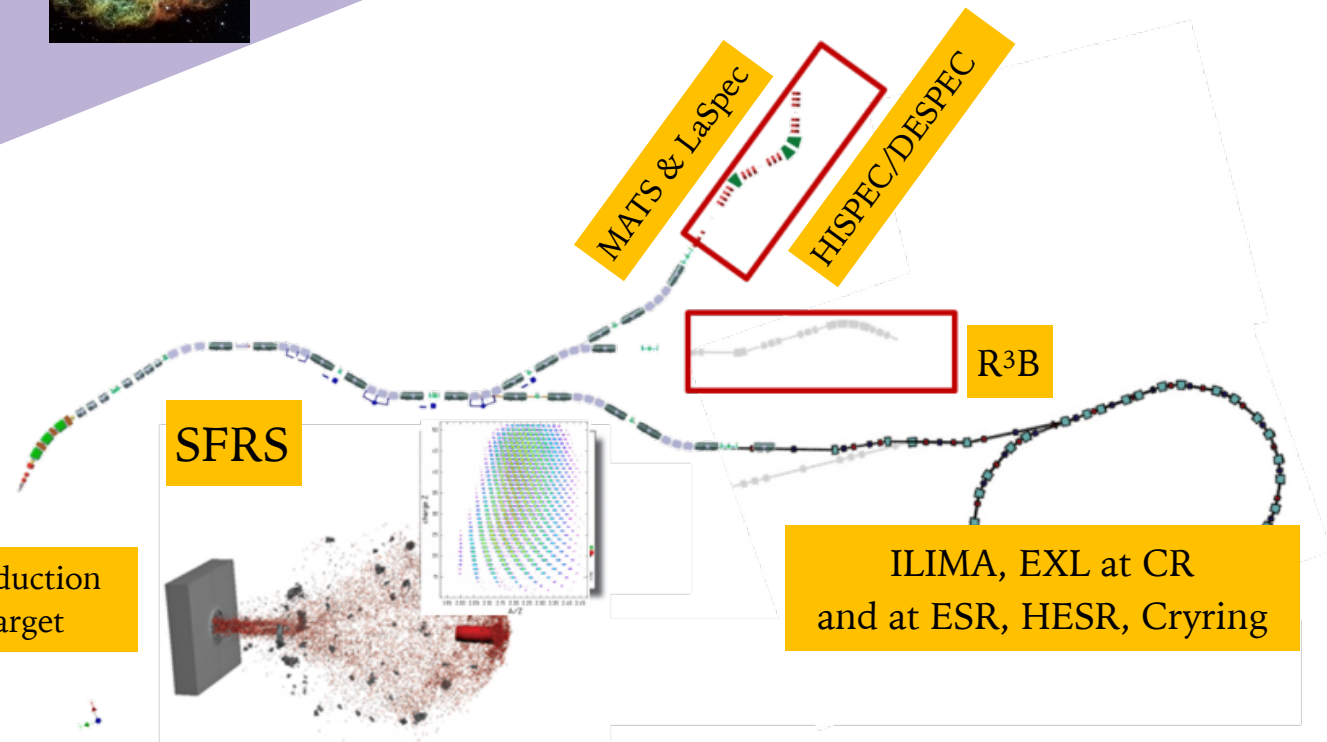
„Nucleosynthesis sites“ at FAIR

Primary intensities
vs. GSI:
x 100

SIS
100



production
target



... with direct applications



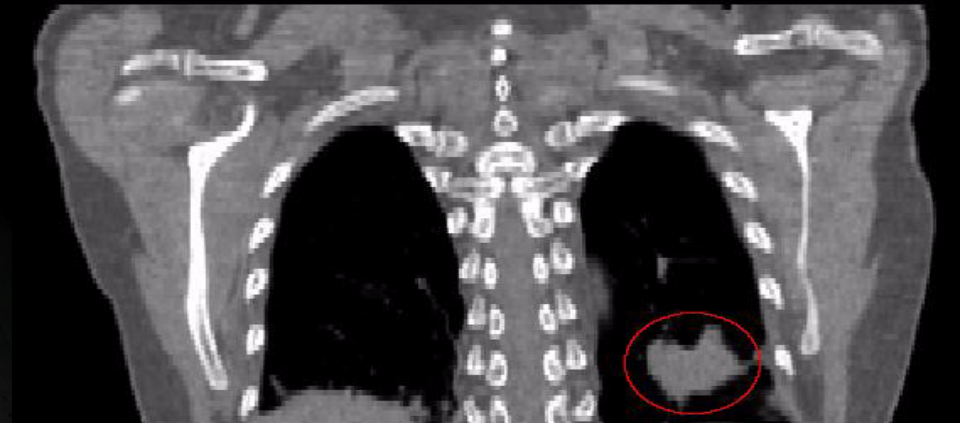
High-performance and scientific computing, big data, green IT



Space radiation protection, unique facility for simulation, collaboration with ESA



Development of nuclear clock:
Promising candidate thorium-229



Novel applications for tumor and non-tumor diseases