Stability of Initial Glasma Fields

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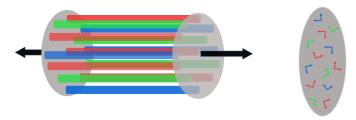
XVI Polish Workshop on Relativistic Heavy-Ion Collisions

Overview

- ► Motivation and glasma fields
- Stability analysis in Minkowski coordinates
- ▶ Abelian and nonAbelian configurations of glasma fields
- Stability analysis in Milne coordinates
- Summary

Motivation and glasma fields

▶ The earliest phase of heavy-ion collisions is described in terms of classical fields.



- Early configuration is unstable, but a mechanism of the instabilities is not fully understood
 - (P. Romatschke and R. Venugopalan, Phys. Rev. Lett. 96, 062302 (2006),
 - P. Romatschke and R. Venugopalan, Phys. Rev. D 74, 045011 (2006).)
- We study the problem systematically starting with the uniform chromoelectric and chromomagnetic fields.

Stability analysis in Minkowski coordinates

Yang-Mills equations in adjoint representation

$$\partial_\mu F_a^{\mu\nu} + g f^{abc} A^b_\mu F_c^{\mu\nu} = j^\nu_a, \qquad F_a^{\mu\nu} = \partial^\mu A^\nu_a - \partial^\nu A^\mu_a + g f^{abc} A^\mu_b A^\nu_c$$

Linearized QCD

$$A_a^{\mu}(t,\mathbf{r}) = \bar{A}_a^{\mu}(t,\mathbf{r}) + a_a^{\mu}(t,\mathbf{r}), \quad \text{where } |\bar{A}(t,\mathbf{r})| \gg |a(t,\mathbf{r})|$$

Stability analysis in Minkowski coordinates

Yang-Mills equations in adjoint representation

$$\partial_\mu F_a^{\mu\nu} + g f^{abc} A^b_\mu F_c^{\mu\nu} = j_a^\nu, \qquad F_a^{\mu\nu} = \partial^\mu A_a^\nu - \partial^\nu A_a^\mu + g f^{abc} A_b^\mu A_c^\nu$$

Linearized QCD

$$A_a^{\mu}(t,\mathbf{r}) = \bar{A}_a^{\mu}(t,\mathbf{r}) + a_a^{\mu}(t,\mathbf{r}), \quad \text{where } |\bar{A}(t,\mathbf{r})| \gg |a(t,\mathbf{r})|$$

Generation of constant fields in one direction is possible in two ways:

- Abelian configuration single color potential linearly dependent on coordinates,
- ▶ nonAbelian configuration multicolor, constant potential.

Calculations done in SU(2) group: $f^{abc} \rightarrow \epsilon^{abc}$.



Stability analysis in Minkowski coordinates

- We considered Abelian and nonAbelian configurations of chromoelectric and chromomagnetic fields.
- Assuming that a_a^μ depends on x^μ through $e^{ik\cdot x}$, the equation of motion is fully algebraic and general solutions were found.
- In each configuration there are unstable modes.

More detailes can be found in:

- S. Bazak and S. Mrówczyński, Phys. Rev. D 105, 034023 (2022)
- S. Bazak and S. Mrówczyński, Phys. Rev. D 106, 034031 (2022)

Glasma fields

Ansatz

- $A^+(x) = \Theta(x^+)\Theta(x^-)x^+\alpha(\tau, x_\perp),$
- $A^{-}(x) = -\Theta(x^{+})\Theta(x^{-})x^{-}\alpha(\tau, x_{\perp}),$
- $\qquad \qquad A^i(x) = \Theta(x^+)\Theta(x^-)x^+\alpha_\perp^i(\tau,x_\perp) + \Theta(-x^+)\Theta(x^-)\beta_1^i(x_\perp) + \Theta(x^+)\Theta(-x^-)\beta_2^i(x_\perp),$

where $\beta_1^i(x_\perp)$ and $\beta_2^i(x_\perp)$ - pre-collision potentials, $\alpha(\tau,x_\perp)$ and $\alpha_\perp^i(\tau,x_\perp)$ - post-collision potentials, i,j=x,y.

Boundary conditions

- $\label{eq:alpha} \bullet \ \alpha(0,x_\perp) = -\tfrac{ig}{2} [\beta_1^i(x_\perp),\beta_2^i(x_\perp)].$

Gauge condition

$$x^+A^- + x^-A^+ = 0$$



Glasma fields

Non-zero components of E and B fields at $\tau = 0$

- $E(x_{\perp}) \equiv E^{z}(0, x_{\perp}) = -2\alpha(0, x_{\perp}),$
- $\blacktriangleright \ B(x_\perp) \equiv B^z(0,x_\perp) = -\partial_y \alpha_\perp^x(0,x_\perp) + \partial_x \alpha_\perp^y(0,x_\perp) ig[\alpha_\perp^y(0,x_\perp),\alpha_\perp^x(0,x_\perp)].$

Glasma initial potentials

- $A_a^t = \frac{g}{2} z f^{abc} \beta_{1b}^i \beta_{2c}^i,$
- $\qquad \qquad A^z_a = \tfrac{g}{2} t f^{abc} \beta^i_{1b} \beta^i_{2c} \text{,}$
- $A_a^i = \beta_{1a}^i + \beta_{2a}^i.$

Uniform chromoelectric field is in the Abelian configuration but uniform chromomagnetic field in the nonAbelian configuration.

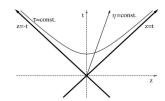


Yang-Mills equations in Milne coordinates

Curvilinear coordinates - Milne coordinates $\tilde{x}^{\mu} = (\tau, x, y, \eta)$ are defined

$$\begin{cases} \tau = \sqrt{t^2 - z^2}, \\ \eta = \frac{1}{2} \ln \frac{t+z}{t-z}, \end{cases} \qquad \begin{cases} t = \tau \cosh \eta, \\ z = \tau \sinh \eta, \end{cases}$$

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Yang-Mills equations:

$$(\nabla_{\mu}F^{\nu\mu})_{a} = \frac{1}{\tau}\partial_{\mu}(\tau g^{\mu\rho}g^{\nu\sigma}F_{\rho\sigma})_{a} + gg^{\mu\rho}g^{\nu\sigma}f^{abc}A^{b}_{\mu}F^{c}_{\rho\sigma} = j^{\nu}_{a},$$

where $g^{\mu\nu} = \text{diag}(1, -1, -1, -\frac{1}{r^2})$, and $F^a_{\mu\nu} = \partial_{\mu}A^a_{\nu} - \partial_{\nu}A^a_{\mu} + gf^{abc}A^b_{\mu}A^c_{\nu}$,

- $A_a^{\mu} = \bar{A}_a^{\mu} + a_a^{\mu}$.
- lacktriangle Yang-Mills equations linearized in a_a^μ (not algebraic for $a^\mu \sim e^{ik\cdot x}$).



$$E^{z} = \frac{1}{\tau} (\partial_{\tau} \bar{A}_{\eta} - \partial_{\eta} \bar{A}_{\tau}) = -\frac{1}{\tau} (\partial_{\tau} (\tau^{2} \bar{A}^{\eta}) + \partial_{\eta} \bar{A}^{\tau}).$$

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Constant and uniform chromoelectric field

$$E^z = -2\bar{A}^\eta - \tau \partial_\tau \bar{A}^\eta$$

$$\bar{A}^{\eta} = -\frac{1}{2}E^z$$

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Constant and uniform chromoelectric field

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solutions for special cases

Special case
$$a_a^x \neq 0$$
, $a_a^y \neq 0$ and $a_a^{\eta} = 0$

Assumptions

- lacktriangledown mixing: $X^\pm=a_2^x\pm ia_3^x$ and $Y^\pm=a_2^y\pm ia_3^y$,
- $X^{\pm}(\tau, x, y, \eta) = e^{i(k_x x + k_y y + \nu \eta)} X^{\pm}(\tau)$

$$Y^{\pm}(\tau, x, y, \eta) = e^{i(k_x x + k_y y + \nu \eta)} Y^{\pm}(\tau)$$

Final equations

$$\left(\partial_{\tau}^2 + \frac{1}{\tau}\partial_{\tau} + k_x^2 + k_y^2 + \left(\frac{\nu}{\tau} \pm \frac{gE\tau}{2}\right)^2\right)X^{\pm} = 0,$$

$$\left(\partial_{\tau}^2 + \frac{1}{\tau}\partial_{\tau} + k_x^2 + k_y^2 + \left(\frac{\nu}{\tau} \pm \frac{gE\tau}{2}\right)^2\right)Y^{\pm} = 0.$$

In the short time limit $(\tau^2\ll 2\nu/(gE))$ - Bessel equation of imaginary order. In the long time limit $(\tau^2\gg 2\nu/(gE),\tau^2\gg 1/(gE))$ - oscillatory running-away solutions



$$B_1^z = F_1^{yx} = \nabla^y \bar{A}_1^x - \nabla^x \bar{A}_1^y + g \bar{A}_2^y \bar{A}_3^x - g \bar{A}_3^y \bar{A}_2^x$$

Constant homogeneous chromomagnetic field

$$\bar{A}^{\mu}_{a} = \left[\begin{array}{ccc} 0 & 0 & 0 & 0 \\ 0 & 0 & \lambda \sqrt{B/g} & 0 \\ 0 & \frac{1}{\lambda} \sqrt{B/g} & 0 & 0 \end{array} \right], \ j^{\nu}_{a} = \left[\begin{array}{ccc} 0 & 0 & 0 & 0 \\ 0 & 0 & \frac{1}{\lambda} \sqrt{gB^{3}} & 0 \\ 0 & \lambda \sqrt{gB^{3}} & 0 & 0 \end{array} \right].$$

The chromomagnetic field is independent of λ , but potentials of different λ are physically nonequivalent.



nonAbelian Chromomagnetic configuration

In special case $a_a^{\eta} = 0$, $a_a^x \neq 0$, $a_a^y \neq 0$

where $\Box = \partial_{\tau}^2 + \frac{1}{\tau}\partial_{\tau} - \partial_x^2 - \partial_y^2 - \frac{1}{\tau^2}\partial_{\eta}^2$ and $V_a^{\pm} = a_a^x \pm ia_a^y$.

For $k_x = 0$ and $k_y = 0$

$$\left(\partial_{\tau}^{2} + \frac{1}{\tau} \partial_{\tau} + \frac{\nu^{2}}{\tau^{2}} + b_{1}^{2} \right) f_{1}(\tau) = 0,$$

$$\left(\partial_{\tau}^{2} + \frac{1}{\tau} \partial_{\tau} + \frac{\nu^{2}}{\tau^{2}} \pm b_{\pm}^{2} \right) f_{\pm}(\tau) = 0.$$

- ▶ The Bessel equation or the modified Bessel equation.
- ▶ For the modified Bessel equation (with "-") there are unstable solutions.
- ▶ The fastest mode grows as $e^{\sqrt{gB}\tau}$.



nonAbelian Chromomagnetic configuration

- ▶ Generalized case: $k_x \neq 0$ and $k_y \neq 0$
- Solutions found using perturbation theory

$$\Lambda^{(1)} = \Lambda^{(0)} + c_x k_x^2 + c_y k_y^2,$$

where $\Lambda^{(0)}$ is the eigenvalue of matrix for $k_x = k_y = 0$.

$$\left(\partial_{\tau}^{2} + \frac{1}{\tau} \partial_{\tau} + \frac{\nu^{2}}{\tau^{2}} + b_{1}^{2} + \left(c_{x}^{1} + 1 \right) k_{x}^{2} + \left(c_{y}^{1} + 1 \right) k_{y}^{2} \right) f_{1}(\tau) = 0,$$

$$\left(\partial_{\tau}^{2} + \frac{1}{\tau} \partial_{\tau} + \frac{\nu^{2}}{\tau^{2}} \pm b_{\pm}^{2} + \left(c_{x}^{\pm} + 1 \right) k_{x}^{2} + \left(c_{y}^{\pm} + 1 \right) k_{y}^{2} \right) f_{\pm}(\tau) = 0.$$

 \blacktriangleright The fastest mode grows as $e^{\sqrt{gB-k_T^2/3}\;\tau}$, $k_T^2\equiv k_x^2+k_y^2$



Temporal evolution of glasma background field

- The glasma fields evolve in time.
- Our stability analysis is reliable if a rate of change of the background is significantly smaller than the one of instabilities, we found.
- In the proper time expansion

$$\begin{array}{rcl} \alpha(\tau,\mathbf{x}_{\perp}) & = & \alpha^{(0)}(\mathbf{x}_{\perp}) + \tau\alpha^{(1)}(\mathbf{x}_{\perp}) + \tau^2\alpha^{(2)}(\mathbf{x}_{\perp}) + \dots, \\ \alpha_{\perp}(\tau,\mathbf{x}_{\perp}) & = & \alpha_{\perp}^{(0)}(\mathbf{x}_{\perp}) + \tau\alpha_{\perp}^{(1)}(\mathbf{x}_{\perp}) + \tau^2\alpha_{\perp}^{(2)}(\mathbf{x}_{\perp}) + \dots. \end{array}$$

For initial uniform chromomagnetic fields one finds

$$\alpha_a = \mathcal{O}(\tau^4),$$

$$\alpha_{\perp a}^x = \delta^{a3} \lambda^{-1} \sqrt{B/g} \Big(1 - \frac{1}{4} \lambda^2 g B \tau^2 + \mathcal{O}(\tau^4) \Big),$$

$$\alpha_{\perp a}^y = \delta^{a2} \lambda \sqrt{B/g} \Big(1 - \frac{1}{4} \lambda^{-2} g B \tau^2 + \mathcal{O}(\tau^4) \Big).$$

Rate of change of background potential can be smaller than the unstable mode growth rate.



Discussion

- ▶ Comparing the growth rate of the unstable mode (\sqrt{gB}) with the rate of change of background field $(gB\tau)$, one finds that the condition of constancy of the background field is satisfied only marginally.
- Fields are spatially uniform in the transverse plane at a scale L (between Q_s^{-1} and $\Lambda_{\rm QCD}^{-1}$, $Q_s \approx 2 {\sf GeV}$, $\Lambda_{\rm QCD} \approx 0.2 {\sf GeV}$).
- Assuming that the domain of uniform field is a square centered in ${\bf r}_\perp=0$ one should replace $(k_x,k_y) \to (2l_x+1,2l_y+1)\frac{\pi}{L}$, where l_x and l_y are integer numbers. The spectrum of eigenmodes becomes discrete.
- From the growth rate of the instability $\sqrt{gB-\frac{k_T^2}{3}}$, one finds that the instability appears if $gB > \frac{2\pi^2}{3L^2}$.
- ▶ For $g\approx 1$, $B\approx Q_s^2\approx 4{\rm GeV}^2$, the upper condition is satisfied for $L^{-1}\approx \Lambda_{\rm OCD}\approx 0.2{\rm GeV}$.

Summary and conclusions

- We found the spectra of eigenmodes for constant and uniform Abelian chromoelectric and nonAbelian chromomagnetic fields configurations.
- The initial chromomagnetic field is unstable, but the conditions of the space-time uniformity of the field are satisfied only marginally.
- The characteristic growth rate of the unstable mode of initial glasma field is of order Q_s and it is much bigger than the growth rate found in the simulations (P. Romatschke and R. Venugopalan, Phys. Rev. Lett. 96, 062302 (2006) and Phys. Rev. D 74, 045011 (2006)).
- The instability found in the simulations (P. Romatschke and R. Venugopalan, Phys. Rev. Lett. 96, 062302 (2006) and Phys. Rev. D 74, 045011 (2006)) is not the instability of glasma initial field, as claimed by some authors. It is rather the Weibel mode, known in plasma physics.

Thank you for attention!