Pointing with Anisotropic Interactions



Kate Scholberg, Duke University SNEWS 2.0 Workshop June 2020

Neutrino Pointing Methods

□ Anisotropic neutrino interactions

combined with detector technology that can exploit it, using the burst neutrino signal

Triangulation

using inter-detector timing

Oscillation pattern pointing

in high-energy resolution detectors

- High-energy (~GeV) neutrino follow-on pointing in directional detectors, using later neutrinos
- □ All of the above!

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Supernova-relevant neutrino interactions

	Electrons	Protons	Nuclei
	Elastic scattering $\nu + e^- \rightarrow \nu + e^-$	Inverse beta decay $\bar{\nu}_e + p \rightarrow e^+ + n$	$ \nu_e + (N, Z) \to e^- + (N - 1, Z + 1) $ $ \bar{\nu}_e + (N, Z) \to e^+ + (N + 1, Z - 1) $
Charged current	[[] √ _e ► ♥ e ⁻	$ \begin{array}{c} \gamma \\ e^+ & \gamma \\ \overline{\nu}_e \\ \end{array} $	r_{v_e} $r_{e^{+/-}}$ Various possible
Neutral current	ve	Elastic scattering v	$ \nu + A \rightarrow \nu + A^* $ ejecta and deexcitation products $ \gamma \qquad n \qquad \gamma \qquad A \qquad A $
	Useful for pointing	very low energy recoils	$ u + A \rightarrow \nu + A $ Coherent elastic (CEvNS)

Most of these have at least *some* intrinsic anisotropy

	Electrons	Protons	Nuclei
Charged current	Elastic scattering $\nu + e^- \rightarrow \nu + e^-$ $[\nabla_e] \cdots \qquad e^-$	Inverse beta decay $\bar{\nu}_e + p \rightarrow e^+ + n$ γ $e^+ \gamma$ $\overline{\nu}_e$	$\nu_{e} + (N, Z) \rightarrow e^{-} + (N - 1, Z + 1)$ $\bar{\nu}_{e} + (N, Z) \rightarrow e^{+} + (N + 1, Z - 1)$ $\overset{\gamma}{\vdash} \overset{n}{\downarrow} $
Neutral current	ve Useful for pointing	Elastic scattering v v v v v v v v v v v v v v v v v v v	$\nu + A \rightarrow \nu + A^{*}$ $\nu + A \rightarrow \nu + A$

Anisotropic Neutrino Interactions

 25°

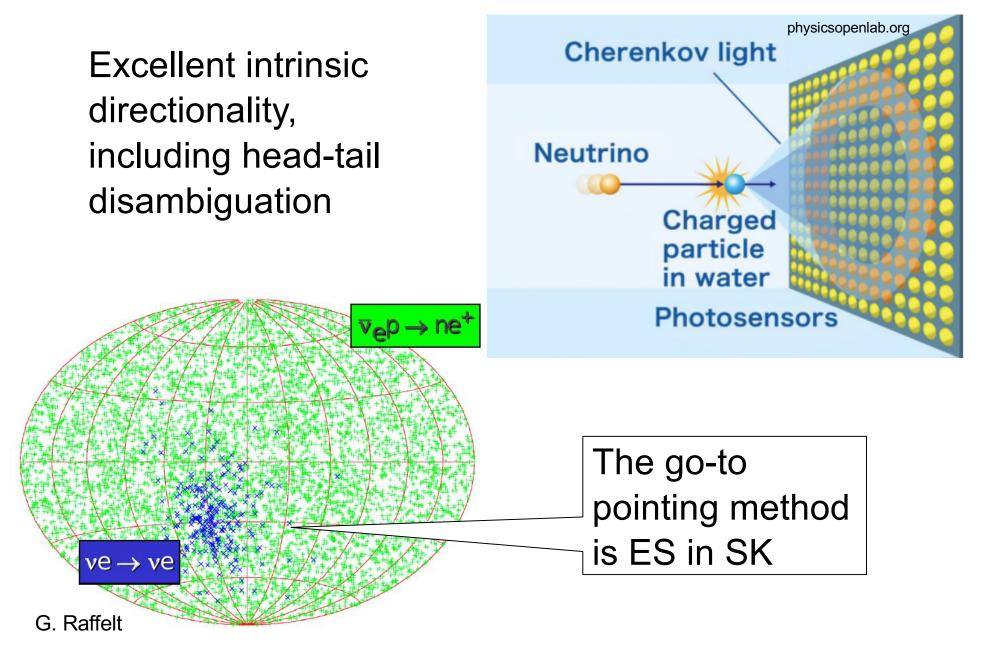
e+/-

 v_{e}

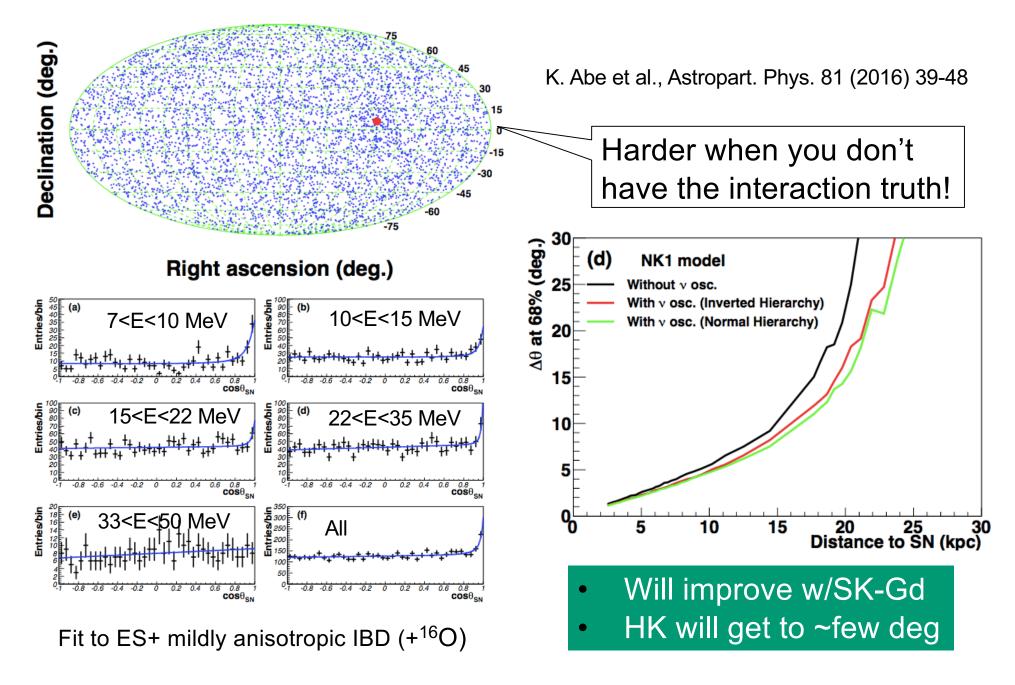
- **Neutrino-electron scattering (ES)**
 - Every detector has electrons
 - Good forward pointing
 - Low cross section, subdominant signal
- Inverse beta decay (IBD)
 - Dominant in water and scintillator
 - Weak energy-dependent anisotropy for lepton
 - Neutron capture position can help
- **Neutrino-nucleus charged-current interactions** $\left(\frac{1}{v_{e}}\right)$
 - Some anisotropy of lepton wrt neutrino
 - Nuclear structure uncertainties $\sim 1 + a \cos \theta$
- Elastic scattering on protons or nuclei
 - Well-defined anisotropy [experimentally hard...]

In all cases, the *detector* must be capable of exploiting directionality via tracking of some kind

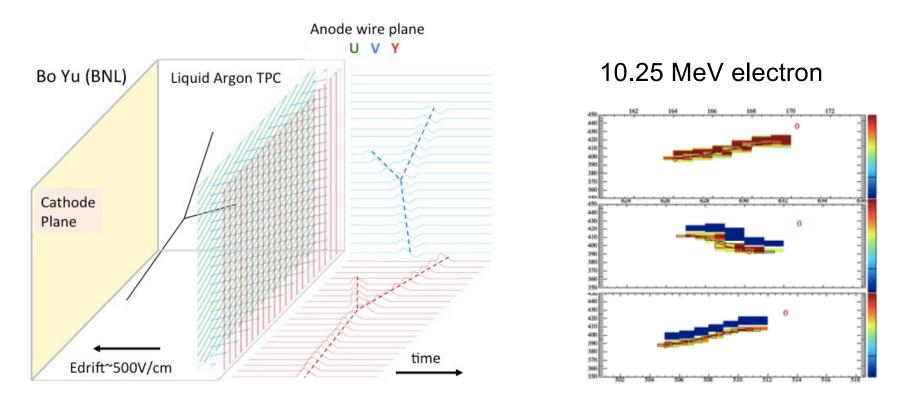
Water Cherenkov Detectors



Pointing in Water Cherenkov: Super-K



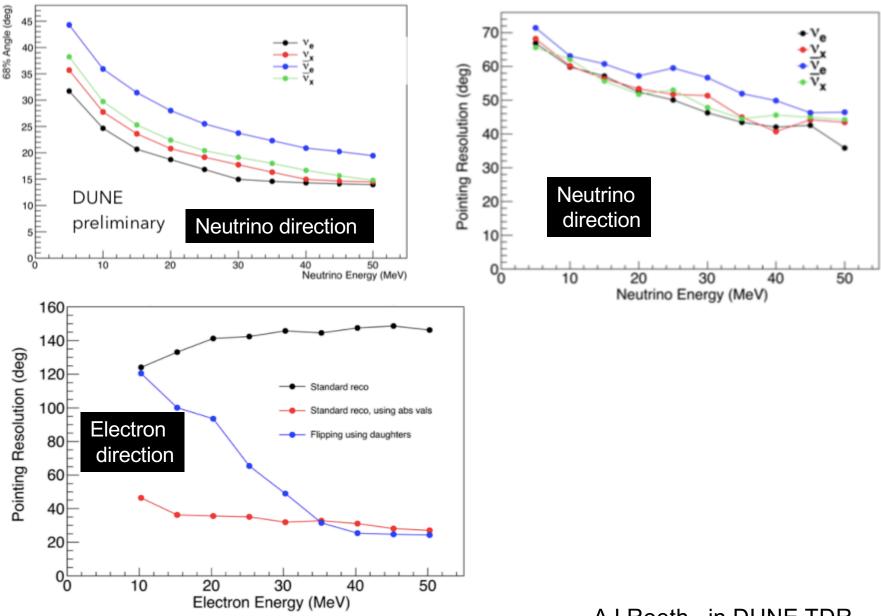
Pointing to the supernova with LArTPCs



Fine-grained tracking...

- but note direction ambiguity, unlike Cherenkov!
- ... but can resolve statistically using bremsstrahlung directionality and multiple scattering

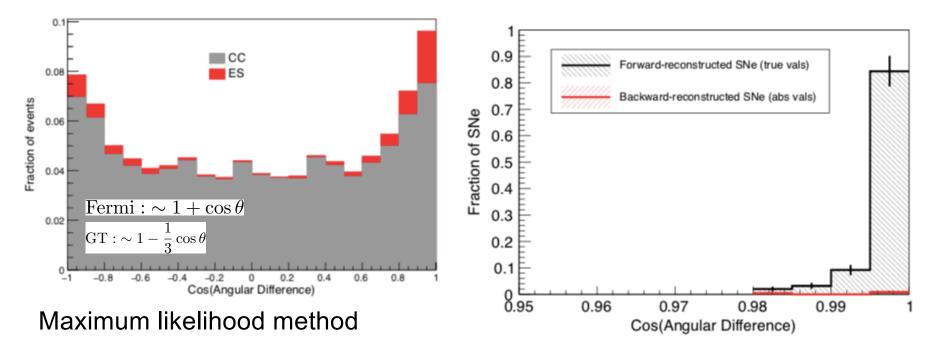
Pointing to the supernova with ES in DUNE

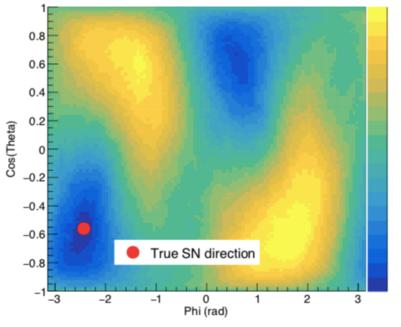


68% Angle vs Neutrino Energy

AJ Roeth, in DUNE TDR

Can make use of both ES and $\nu_e\text{CC}$ directionality





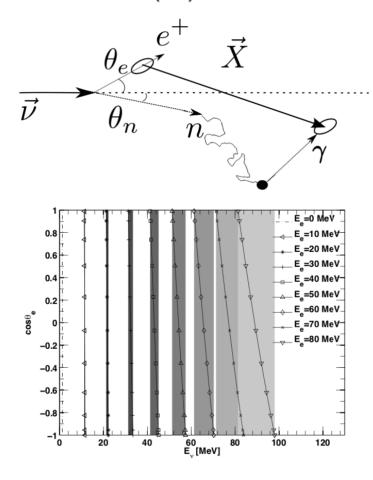
Overall pointing using an ensemble of events from a
 ~10 kpc supernova) → ~5°
 ...improvements still possible

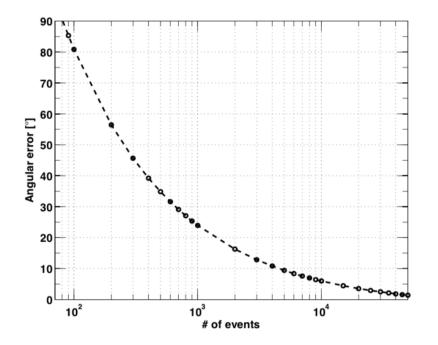
AJ Roeth, DUNE TDR

Pointing with Liquid Scintillator

This is hard, as produced photons get quasi-isotropized... BUT, some statistical prospects using IBD kinematics \rightarrow positron energy + reconstructed vertices of e⁺ and n

Prompt directional detection of galactic supernova by combining large liquid scintillator neutrino detectors V. Fischer (IRFU, Saclay) *et al.*. Apr 21, 2015. 25 pp. Published in JCAP 1508 (2015) 032





Needs good statistics!

Comments

- anisotropic interactions will likely be the best information from neutrinos
 - currently SK has best pointing
- unknown latency... may not be prompt
- we need to consider how to combine with (likely more prompt, but lower quality) triangulation information
 - need tiered strategy