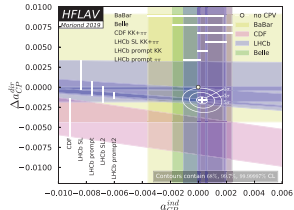
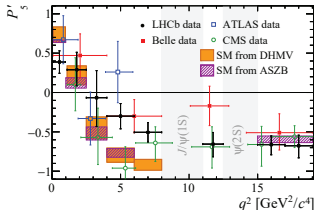
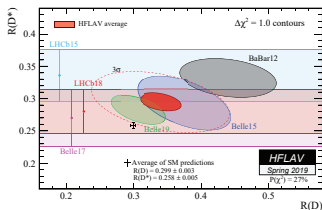


Flavour Anomalies

Antonio Pich

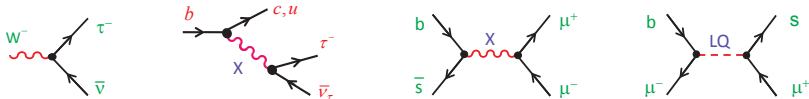
IFIC, U. Valencia – CSIC



Many Interesting Flavour Anomalies

$b \rightarrow c\tau\nu$, $b \rightarrow s\mu^+\mu^-$, $(g-2)_{\mu,e}$, $\tau^\pm \rightarrow \pi^\pm K_S\nu$, Δa_{CP} , V_{ub} , V_{ud} , ...

Some already gone: $B \rightarrow \tau\nu$, $W \rightarrow \tau\nu$, ϵ'_K/ϵ_K , ϵ_K , V_{cb} , ...



- Evidence for New Physics
- Statistical fluctuation
- Underestimated systematics
- Incorrect SM prediction or measurement

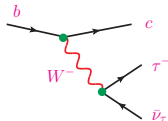


Not easy common explanation (within appealing BSM models)

Separate analyses are (perhaps) more enlightening

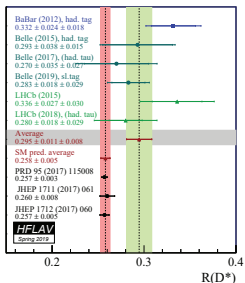
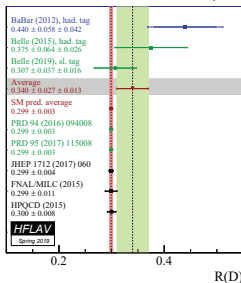
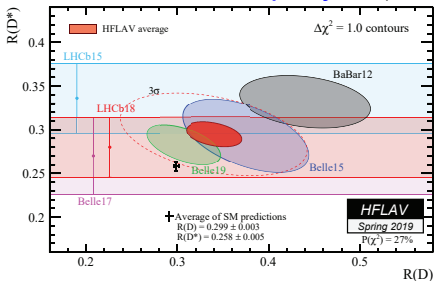
$$\mathcal{R}_{D^{(*)}} \equiv \frac{\mathcal{B}(\bar{B} \rightarrow D^{(*)} \tau^- \bar{\nu}_\tau)}{\mathcal{B}(\bar{B} \rightarrow D^{(*)} \ell^- \bar{\nu}_\ell)}$$

Tree-level process



3.08 σ discrepancy

(3.2 σ with more recent predictions)



LHCb, 1711.05623: $\mathcal{R}_{J/\psi} \equiv \frac{\mathcal{B}(B_c \rightarrow J/\psi \tau \bar{\nu}_\tau)}{\mathcal{B}(B_c \rightarrow J/\psi \mu \bar{\nu}_\mu)} = 0.71 \pm 0.17 \pm 0.18$ (1.7 σ) $\mathcal{R}_{J/\psi}^{\text{SM}} \approx 0.26 - 0.28$

Belle, 1903.03102: $F_L^{D^*} = 0.60 \pm 0.08 \pm 0.04$ (1.6 σ) $F_{L,\text{SM}}^{D^*} = 0.455 \pm 0.003$

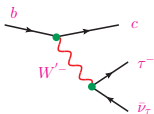
Belle, 1612.00529: $\mathcal{P}_\tau^{D^*} = -0.38 \pm 0.51_{-0.16}^{+0.21}$ $\mathcal{P}_{\tau,\text{SM}}^{D^*} = -0.499 \pm 0.003$

Effective Field Theory

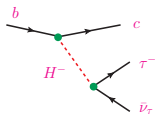
$$C_{AB}^X|^{SM} = 0$$

$$\mathcal{H}_{\text{eff}}^{b \rightarrow c \tau \nu} = \frac{4G_F}{\sqrt{2}} V_{cb} \left\{ \mathcal{O}_{LL}^V + \sum_{A,B=L,R} [\mathcal{C}_{AB}^V \mathcal{O}_{AB}^V + \mathcal{C}_{AB}^S \mathcal{O}_{AB}^S + \mathcal{C}_{AB}^T \mathcal{O}_{AB}^T] + \text{h.c.} \right\}$$

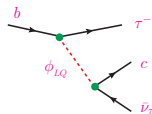
$$\mathcal{O}_{AB}^V = (\bar{c} \gamma^\mu \mathcal{P}_{Ab}) (\bar{\tau} \gamma_\mu \mathcal{P}_{B\nu}), \quad \mathcal{O}_{AB}^S = (\bar{c} \mathcal{P}_{Ab}) (\bar{\tau} \mathcal{P}_{B\nu}), \quad \mathcal{O}_{AB}^T = \delta_{AB} (\bar{c} \sigma^{\mu\nu} \mathcal{P}_{Ab}) (\bar{\tau} \sigma_{\mu\nu} \mathcal{P}_{A\nu})$$



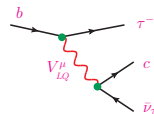
$$C_{LL}^V$$



$$C_{LL}^S, C_{RL}^S \quad (C_{LL}^T)$$



$$C_{LL}^V, C_{LL}^S, C_{LL}^T$$



$$C_{LL}^V, C_{RL}^S$$

Many analyses (usually with single operator/mediator and partial data information)

Freytsis et al, Bardhan et al, Cai et al, Hu et al, Celis et al, Datta et al, Bhattacharya et al, Alonso et al, ...

Global fit to all data: (q^2 distributions included) ν_L Murgui-Penúelas-Jung-Pich, 1904.09311
 ν_R Mandal-Murgui-Penúelas-Pich, 2004.06726

Assumptions

- $C_{AB}^X \neq 0$ for 3rd fermion generation only
- EWSB linearly realized $\rightarrow C_{RL}^V = 0$
- CP symmetry \rightarrow Real Wilson coefficients

$F_L^{D*}, \mathcal{B}_{10}$	Min 1	Min 2
$\chi^2/\text{d.o.f.}$	37.4/54	40.4/54
C_{LL}^V	0.09 ± 0.13 -0.12	0.34 ± 0.05 -0.07
C_{RL}^S	0.09 ± 0.12 -0.61	-1.10 ± 0.48 -0.07
C_{LL}^S	-0.14 ± 0.52 -0.07	-0.30 ± 0.11 -0.50
C_{LL}^T	0.008 ± 0.046 -0.044	0.093 ± 0.029 -0.030

 $\mathcal{B}(B_c \rightarrow \tau \bar{\nu}) < 10\%$
 F_L^{D*} included

- **Strong preference for New Physics** ($\chi_{SM}^2 - \chi^2 = 31.4$)
- **No clear preference for a particular Wilson coefficient in the global minimum**
- **Min 1 compatible with a global modification of the SM**
(Fitting only C_{LL}^V just increases χ^2 by 1.4)
- **Min 2 is further away from the SM & involves large scalar contributions**
- F_L^{D*} **difficult to accommodate at 1σ**
- Complex C_{AL}^X do not improve the χ^2 , but open many more solutions
- Including C_{RL}^V slightly improves the agreement with data ($\chi^2/\text{d.o.f.} = 32.5/53$).
Two additional fine-tuned solutions with $C_{LL}^V \sim -0.9$

Global Fit within ν_R Scenarios

Mandal-Murgui-Peñuelas-Pich, 2004.06726

Sc 1: $\mathcal{O}_{LR}^V, \mathcal{O}_{RR}^V, \mathcal{O}_{LR}^S, \mathcal{O}_{RR}^S, \mathcal{O}_{RR}^T, \mathcal{O}_{LL}^V$

Sc 2: $\mathcal{O}_{LR}^V, \mathcal{O}_{RR}^V, \mathcal{O}_{LR}^S, \mathcal{O}_{RR}^S, \mathcal{O}_{RR}^T$

Sc 3, V^μ : \mathcal{O}_{RR}^V

Sc 4, Φ : $\mathcal{O}_{LR}^S, \mathcal{O}_{RR}^S$ [b: + $\mathcal{O}_{LL}^S, \mathcal{O}_{RL}^S$]

Sc 5, U_1^μ : $\mathcal{O}_{RR}^V, \mathcal{O}_{LR}^S$ [b: + $\mathcal{O}_{LL}^V, \mathcal{O}_{RL}^S$]

Sc 6, \bar{R}_2 : $\mathcal{O}_{RR}^S, \mathcal{O}_{RR}^T$

Sc 7, S_1 : $\mathcal{O}_{RR}^V, \mathcal{O}_{RR}^S, \mathcal{O}_{RR}^T$ [b: + $\mathcal{O}_{LL}^V, \mathcal{O}_{LL}^S, \mathcal{O}_{LL}^T$]

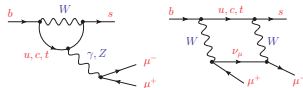
Sc 8, \bar{V}_2^μ : \mathcal{O}_{LR}^S

Scenario	$\mathcal{B}(B_c \rightarrow \tau \bar{\nu})$	$\chi^2/\text{d.o.f}$	Pull _{SM}			Pull _{SM}	p-value
			$\bar{P}_\tau^{D^*}, F_L^{D^*}$	\mathcal{R}_{D,D^*}	$d\Gamma/dq^2$		
SM	2.16%	52.87/59					69.95%
Scenario 1, Min 1	< 10%	37.26/53	0.007	2.08	0.0414	2.4	95.02%
Scenario 1, Min 2	< 10%	38.86/53	0.001 ✗	2.08	0.0006	2.2	92.68%
Scenario 1, Min 1	< 30%	36.42/53	0.022	2.08	0.0866	2.5	96.00%
Scenario 1, Min 2	< 30%	38.54/53	0.011	2.08	0.000	2.2	93.21%
Scenario 2, Min 1	< 10%	38.54/54	0.006 ✗	2.32	0.0113	2.5	93.20%
Scenario 2, Min 2	< 10%	39.05/54	0.004 ✗	2.32	0.0003	2.4	93.73%
Scenario 2, Min 1	< 30%	38.33/54	0.035 ✗	2.32	0.0023	2.5	94.73%
Scenario 2, Min 2	< 30%	38.80/54	0.025 ✗	2.32	0*	2.4	94.09%
Scenario 3	< 10%	39.50/58	0.150 ✗	3.65	0.0835	3.7 ✓	97.00%
Scenario 4a, Min 1	< 10%	49.93/57	0.079 ✗	2.34 ✗	0*	1.2	73.52%
Scenario 4a, Min 2	< 10%	49.93/57	0.079 ✗	2.34 ✗	0*	1.2	73.52%
Scenario 4a, Min 1	< 30%	44.49/57	0.311 ✗	2.66 ✗	0*	2.4	88.62%
Scenario 4a, Min 2	< 30%	44.49/57	0.311 ✗	2.66 ✗	0*	2.4	88.62%
Scenario 4b	< 10%	43.56/55	0.054 ✗	2.07 ✗	0*	1.9	86.70%
Scenario 4b	< 30%	40.03/55	0.218	2.52	0*	2.5	93.54%
Scenario 5a	< 10%	39.39/57	0* ✗	3.22	0.0981	3.2 ✓	96.36%
Scenario 5b	< 10%	39.37/55	0* ✗	3.34	0.0060	2.6	94.47%
Scenario 6	< 10%	44.20/58	0* ✗	3.34	0*	2.9	90.93%
Scenario 7a	< 10%	39.21/57	0.126 ✗	3.22	0.0616	3.3 ✓	96.53%
Scenario 7b	< 10%	39.06/55	0.014 ✗	2.56	0.0112	2.7	94.87%
Scenario 8	< 10%	47.32/57	0.259 ✗	2.56 ✗	0*	1.9	81.60%

- $F_L^{D^*}$ difficult to fit at 1σ
Only possible in Sc 1 and 4b (< 30%)
- Scalar solution \Rightarrow larger $\text{Br}(B_c \rightarrow \tau \bar{\nu})$
- Higher pulls: V^μ, S_1, U_1^μ

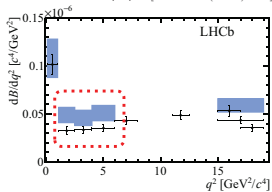


$$b \rightarrow s \mu^+ \mu^-$$

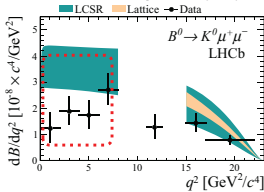


SM loop process

LHCb $B^0 \rightarrow K^{*0} \mu^+ \mu^-$ [JHEP 11 (2016) 047]



LHCb $B^0 \rightarrow K^0 \mu^+ \mu^-$ [JHEP 06 (2014) 133]



Data below SM predictions

$$B^0 \rightarrow K^{(*)0} \mu^+ \mu^-, B^+ \rightarrow K^{(*)+} \mu^+ \mu^-,$$

$$B_S^0 \rightarrow \phi \mu^+ \mu^-, \Lambda_b \rightarrow \Lambda \mu^+ \mu^-, \dots$$

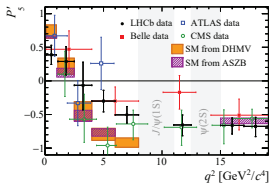
(1-3 σ tensions)

Large hadronic uncertainties

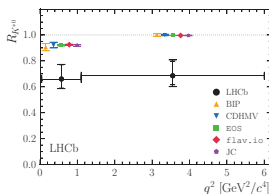
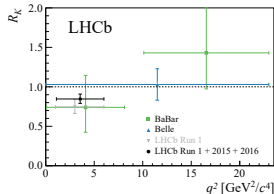
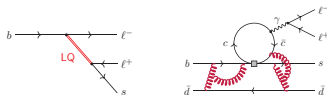
Optimized observables

(angular distribution)

Descotes-Genon et al, 1207.2753



New Physics or SM $\bar{c}c$ loop?



Lepton Flavour Violations

$$\mu^+ \mu^- \text{ versus } e^+ e^-$$

Not confirmed/refuted by Belle

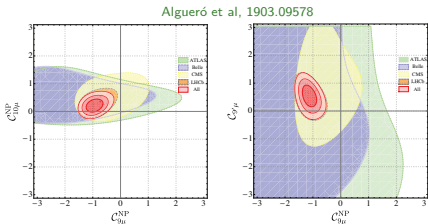
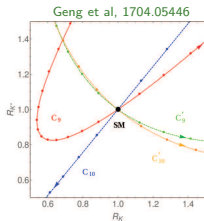
Global 2D Fits:

$$C_{9,\mu}^{\text{NP}} \sim -0.2 C_{9,\mu}^{\text{SM}}$$

$$\mathcal{H}_{\text{eff}}^{b \rightarrow s \ell \ell} = -\frac{4G_F}{\sqrt{2}} V_{tb} V_{ts}^* \frac{\alpha}{4\pi} \sum_i C_i O_i + \text{h.c.}$$

$$O_9 = (\bar{s}_L \gamma_\mu b_L)(\bar{\ell} \gamma^\mu \ell) \quad , \quad O'_9 = (\bar{s}_R \gamma_\mu b_R)(\bar{\ell} \gamma^\mu \ell)$$

$$O_{10} = (\bar{s}_L \gamma_\mu b_L)(\bar{\ell} \gamma^\mu \gamma_5 \ell) \quad , \quad O'_{10} = (\bar{s}_R \gamma_\mu b_R)(\bar{\ell} \gamma^\mu \gamma_5 \ell)$$



$$C_{9,\mu}^{\text{NP}} \sim -0.95$$

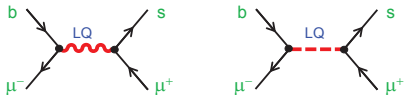
$$C_{10,\mu}^{\text{NP}} \sim 0.20$$

(5.7 σ pull)

- $B_s^0 \rightarrow \mu^+ \mu^-$ strongly constrains pseudoscalar operators and bounds $C_{10,\mu}^{\text{NP}}$
- Preferred solutions: $C_{9,\mu}^{\text{NP}} \neq 0$ or $C_{9,\mu}^{\text{NP}} \approx -C_{10,\mu}^{\text{NP}} \neq 0$
- Slight tension (2 σ) in current $B_s^0 \rightarrow \mu^+ \mu^-$ world average favours $C_{9,\mu}^{\text{NP}} - C_{10,\mu}^{\text{NP}}$
- Recent data allow more space for right-handed currents
- Additional solutions with LFU components (Alguero et al, 1809.08447)
- SMEFT: $b \rightarrow c \tau \nu$ and $b \rightarrow s \ell \ell$ anomalies \rightarrow Large $b \rightarrow s \tau \tau$

$$(\bar{Q}_2 \gamma^\mu Q_3)(\bar{L}_3 \gamma_\mu L_3) + (\bar{Q}_2 \gamma^\mu \sigma^I Q_3)(\bar{L}_3 \gamma_\mu \sigma^I L_3) \approx 2[(\bar{c}_L \gamma_\mu b_L)(\bar{\tau}_L \gamma^\mu \nu_{\tau L}) + (\bar{s}_L \gamma_\mu b_L)(\bar{\tau}_L \gamma^\mu \tau_L)]$$

Leptoquark Solutions



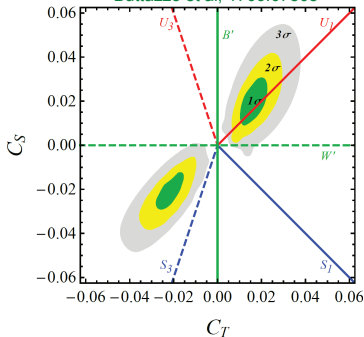
$$\mathcal{L}_{\text{eff}} = -\frac{1}{v^2} \lambda_{ij}^q \lambda_{\alpha\beta}^\ell \left[C_T (\bar{Q}_L^i \gamma_\mu \sigma^a Q_L^j) (\bar{L}_L^\alpha \gamma^\mu \sigma^a L_L^\beta) + C_S (\bar{Q}_L^i \gamma_\mu Q_L^j) (\bar{L}_L^\alpha \gamma^\mu L_L^\beta) \right]$$

$U(2)_q \otimes U(2)_\ell$ Family Symmetry

Angelescu et al, 1808.08179

Model	$R_{D^{(*)}}$	$R_{K^{(*)}}$	$R_{D^{(*)}}$ & $R_{K^{(*)}}$
$S_1 = (\bar{3}, 1, 1/3)$	✓	✗*	✗*
$R_2 = (3, 2, 7/6)$	✓	✗*	✗
$S_3 = (\bar{3}, 3, 1/3)$	✗	✓	✗
$U_1 = (3, 1, 2/3)$	✓	✓	✓
$U_3 = (3, 3, 2/3)$	✗	✓	✗

Buttazzo et al, 1706.07808



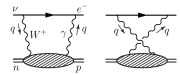
Possible UV completions:

- 4321 model Di Luzio et al
- (Pati-Salam)³ Bordone et al
- PS + VLF Calibbi et al
- Warped PS Blanke-Crivellin
- SU(5) GUT (R_2 & S_3) Becirevic et al
- S_1 & S_3 Crivellin et al, Buttazzo et al, Marzocca
- ...

Superallowed Nuclear β Transitions ($0^+ \rightarrow 0^+$)

$$|V_{ud}|^2 = \frac{\pi^3 \ln 2}{ft G_F^2 m_e^5 (1 + \delta_{RC})} = \frac{(2984.48 \pm 0.05) \text{ s}}{ft (1 + \delta_{RC})}$$

$$\delta_{RC} = \Delta_R^V + \Delta_{\text{Nucl}} \quad , \quad \mathcal{F}t = ft (1 + \Delta_{\text{Nucl}}) = 3072.27 (72) \text{ s}$$

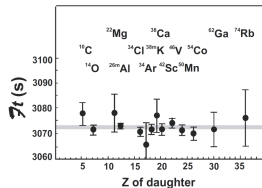
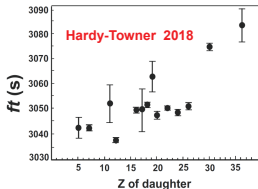


Nucleus-independent radiative correction

$\Delta_R^V = \begin{cases} 0.02361 (38) \\ 0.02467 (22) \end{cases}$	Marciano-Sirlin, 2006 Seng et al, 1807.10197
↓	
$ V_{ud} = \begin{cases} 0.97420 (21) \\ 0.97366 (15) \end{cases}$	Marciano-Sirlin Seng et al

Independent check needed!

$$f_+(0) = 1 + O[(m_u - m_d)^2]$$



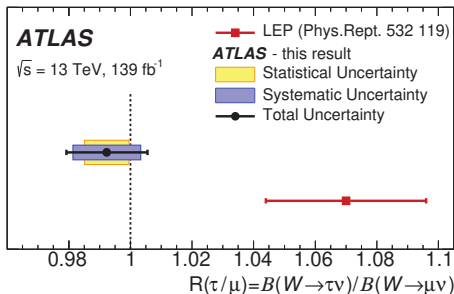
PDG 2020: $|V_{ud}| = 0.97370 (14)$

$$|V_{us}| = \begin{cases} 0.2231 (7) & (K \rightarrow \pi \ell \nu) \\ 0.2245 (8) & (K \rightarrow \ell \nu, \pi \ell \nu) \end{cases}$$



$$1 - |V_{ud}|^2 - |V_{us}|^2 - |V_{ub}|^2 = \begin{cases} 0.00212 (41) & 5.1 \sigma \\ 0.00149 (45) & 3.3 \sigma \end{cases}$$

Lepton Flavour Universality in W Decays



2007.14040



$$\left| \frac{g_\tau}{g_\mu} \right| = \begin{cases} 0.996 \pm 0.007 & \text{ATLAS} \\ 1.034 \pm 0.013 & \text{LEP} \\ 1.004 \pm 0.016 & \text{Average} \end{cases}$$

Summary

- Flavour structure and C/P are major pending questions
- Related to SSB \rightarrow Scalar Sector (Higgs)
- Important cosmological implications (Baryogenesis)
- Sensitive to New Physics: Flavour Anomalies!

Intriguing signals (Most anomalies related to 3rd family)

Many questions. Higher statistics & better systematics (QCD) needed

Eagerly awaiting new experimental results

Backup



Possible Caveats / Constraints:

① Saturation of inclusive width: $\mathcal{B}(B \rightarrow D^{**} \tau \nu) > 0.5\%$ Freytsis et al, 1506.08896

• $\mathcal{R}_{D^{(*)}}$ \rightarrow $\mathcal{B}(B \rightarrow D \tau \nu) + \mathcal{B}(B \rightarrow D^* \tau \nu) = (2.39 \pm 0.13)\%$

• $\frac{\mathcal{B}(B \rightarrow X_c \tau \nu)}{\mathcal{B}(B \rightarrow X_c e \nu)} \Big|_{\text{OPE}} = (0.222 \pm 0.007)$ Not a problem of form factors

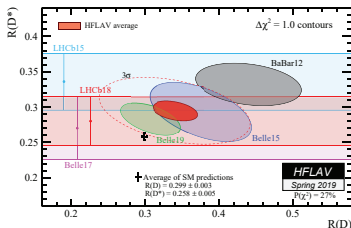
$\mathcal{B}(B \rightarrow X_c e \nu) = (10.65 \pm 0.16)\%$ \rightarrow $\mathcal{B}(B \rightarrow X_c \tau \nu) = (2.36 \pm 0.08)\%$

• LEP: $\mathcal{B}(b \rightarrow X_c \tau \nu) = (2.41 \pm 0.23)\%$

② $b \rightarrow c \tau \nu \leftrightarrow b \bar{c} \rightarrow \tau \nu$: $\mathcal{B}(B_c \rightarrow \tau \nu) < 10\%$ (30%) Akeroyd-Chen
Alonso et al, Celis et al

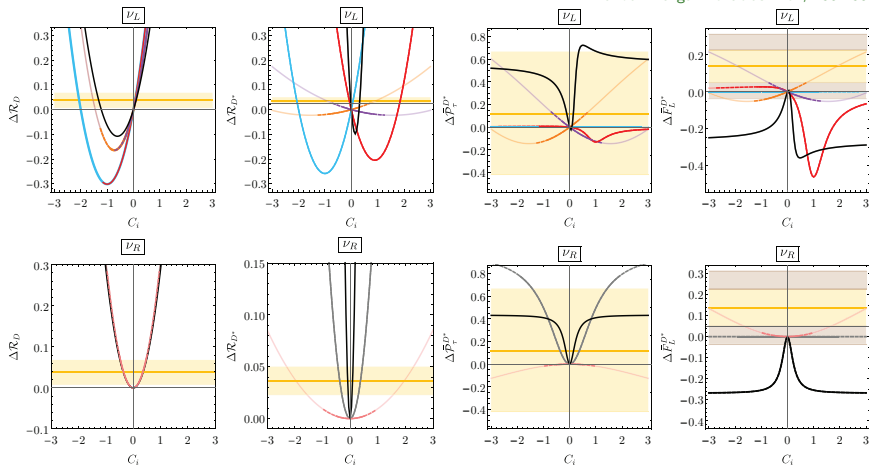
③ Differential distributions. Polarizations: Data self-consistency

④ Time evolution of data:



Sensitivity to individual Wilson coefficients

Mandal-Murgui-Penüelas-Pich, 2004.06726

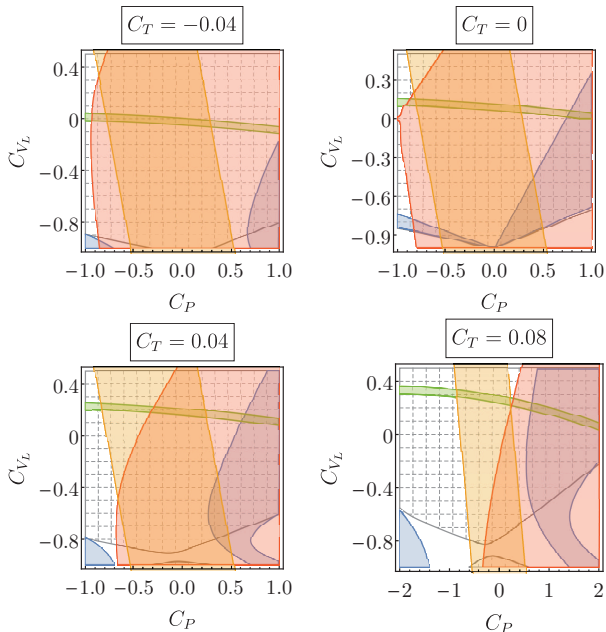


ν_L	C_{LL}^V	C_{RL}^V	C_{LL}^S	C_{LL}^T
ν_R	$C_{LR}^V = C_{RR}^V$	$C_{RR}^S = C_{LR}^S$	C_{RR}^T	

F_L^{D*} always below
the exp. 1σ region

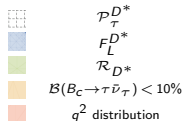
Solid (dashed) lines indicate ranges satisfying $\text{Br}(B_c \rightarrow \tau\nu) < 10\%$ (30%). Faint lines do not fulfil this constraint

D* Observables

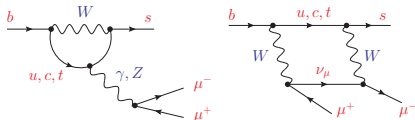


It is not possible to accommodate all D* data at 1σ

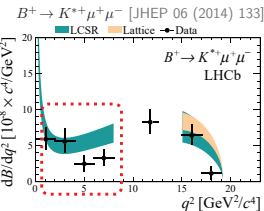
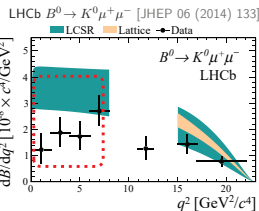
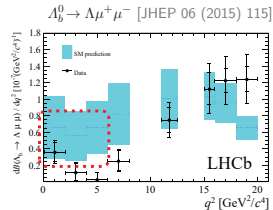
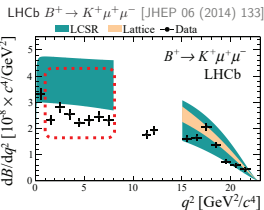
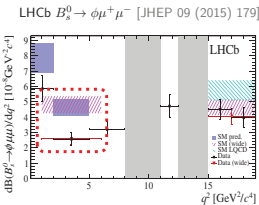
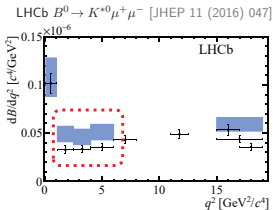
$$C_P \equiv C_{S_R} - C_{S_L}$$



$$b \rightarrow s \mu^+ \mu^-$$



SM loop process



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Data consistently below SM predictions

(1-3 σ tensions)

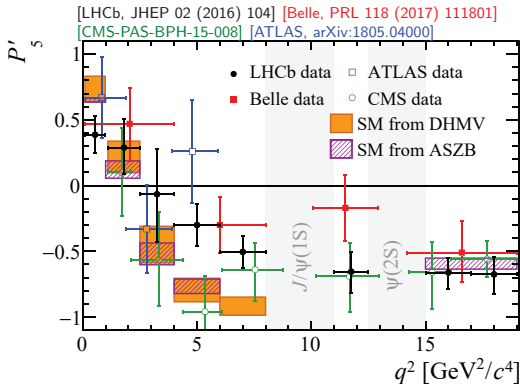
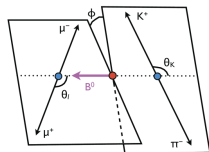
Large hadronic uncertainties

$B \rightarrow K^* \mu^+ \mu^- \rightarrow K \pi \mu^+ \mu^-$

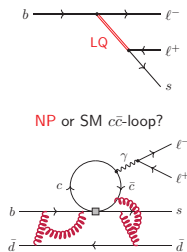
$$\frac{1}{d\Gamma/dq^2} \frac{d^4\Gamma}{d\cos\theta_\ell d\cos\theta_K d\phi dq^2} = \frac{9}{32\pi} \left[\frac{3}{4}(1-F_L)\sin^2\theta_K + F_L\cos^2\theta_K + \frac{1}{4}(1-F_L)\sin^2\theta_K\cos 2\theta_\ell \right. \\ \left. - F_L\cos^2\theta_K\cos 2\theta_\ell + S_3\sin^2\theta_K\sin^2\theta_\ell\cos 2\phi + S_4\sin 2\theta_K\sin 2\theta_\ell\cos\phi \right. \\ \left. + S_5\sin 2\theta_K\sin\theta_\ell\cos\phi + S_6\sin^2\theta_K\cos\theta_\ell + S_7\sin 2\theta_K\sin\theta_\ell\sin\phi \right. \\ \left. + S_8\sin 2\theta_K\sin 2\theta_\ell\sin\phi + S_9\sin^2\theta_K\sin^2\theta_\ell\sin 2\phi \right]$$

$$q^2 = s_{\mu\mu}$$

$$P'_{i=4,5,6,8} = \frac{S_{i=4,5,7,8}}{\sqrt{F_L(1-F_L)}}$$

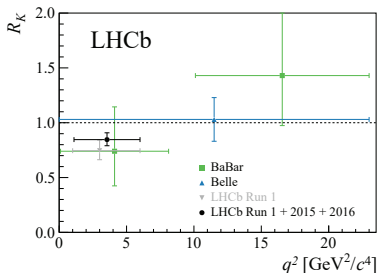
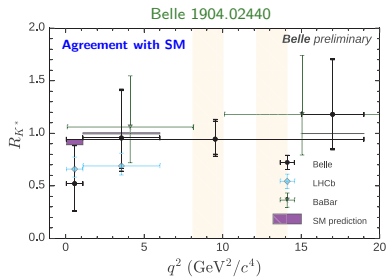
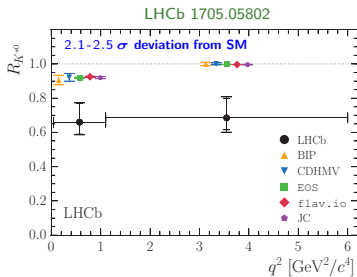


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Violations of Lepton Flavour Universality

$$\frac{\Gamma(B \rightarrow K^* \mu^+ \mu^-)}{\Gamma(B \rightarrow K^* e^+ e^-)}$$



LHCb, 1903.09252

$q^2 \in [1.1, 6.0] \text{ GeV}^2$

$$R_K \equiv \frac{\Gamma(B^+ \rightarrow K^+ \mu^+ \mu^-)}{\Gamma(B^+ \rightarrow K^+ e^+ e^-)} = 0.846^{+0.060}_{-0.054} + 0.016_{-0.014}$$

2.5 σ below the SM

Not confirmed/refuted by Belle 1908.01848

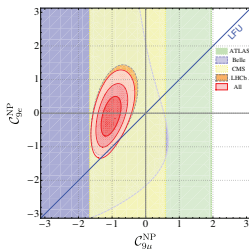
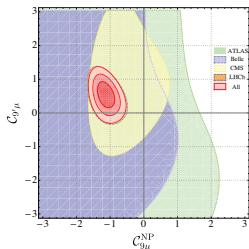
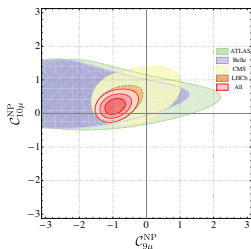
Global 2D Fits:

$$C_{9,\mu}^{\text{NP}} \sim -0.2 C_{9,\mu}^{\text{SM}}$$

$$\mathcal{H}_{\text{eff}}^{b \rightarrow s \ell \ell} = -\frac{4G_F}{\sqrt{2}} V_{tb} V_{ts}^* \frac{\alpha}{4\pi} \sum_i C_i O_i + \text{h.c.}$$

$$O_9 = (\bar{s}_L \gamma_\mu b_L)(\bar{\ell} \gamma^\mu \ell) \quad , \quad O'_9 = (\bar{s}_R \gamma_\mu b_R)(\bar{\ell} \gamma^\mu \ell)$$

$$O_{10} = (\bar{s}_L \gamma_\mu b_L)(\bar{\ell} \gamma^\mu \gamma_5 \ell) \quad , \quad O'_{10} = (\bar{s}_R \gamma_\mu b_R)(\bar{\ell} \gamma^\mu \gamma_5 \ell)$$

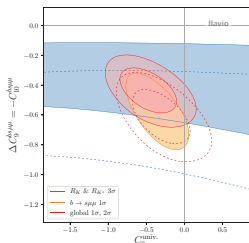
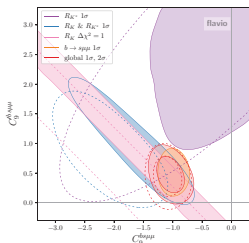
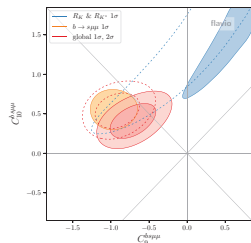


Algueró et al
1903.09578

$$C_{9,\mu}^{\text{NP}} \sim -0.95$$

$$C_{10,\mu}^{\text{NP}} \sim 0.20$$

(5.7 σ pull)



Aebischer et al
1903.10434

$$C_{9,\mu}^{\text{NP}} \sim -0.72$$

$$C_{10,\mu}^{\text{NP}} \sim 0.40$$

(6.2 σ pull)

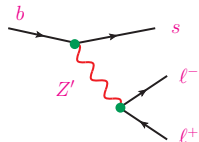
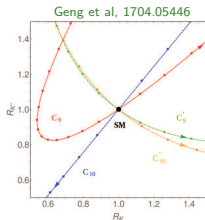
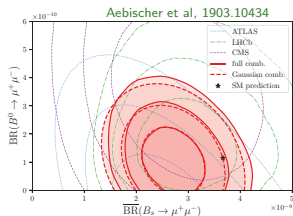
Global 2D Fits:

$$C_{9,\mu}^{\text{NP}} \sim -0.2 C_{9,\mu}^{\text{SM}}$$

$$\mathcal{H}_{\text{eff}}^{b \rightarrow s \ell \ell} = -\frac{4G_F}{\sqrt{2}} V_{tb} V_{ts}^* \frac{\alpha}{4\pi} \sum_i C_i O_i + \text{h.c.}$$

$$O_9 = (\bar{s}_L \gamma_\mu b_L)(\bar{\ell} \gamma^\mu \ell) \quad , \quad O'_9 = (\bar{s}_R \gamma_\mu b_R)(\bar{\ell} \gamma^\mu \ell)$$

$$O_{10} = (\bar{s}_L \gamma_\mu b_L)(\bar{\ell} \gamma^\mu \gamma_5 \ell) \quad , \quad O'_{10} = (\bar{s}_R \gamma_\mu b_R)(\bar{\ell} \gamma^\mu \gamma_5 \ell)$$

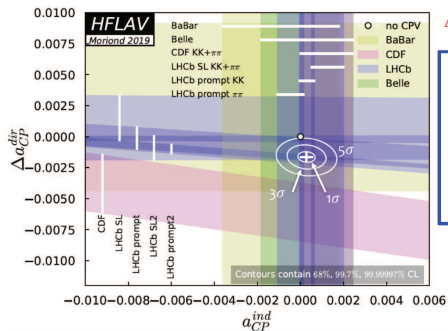


- $B_s^0 \rightarrow \mu^+ \mu^-$ strongly constrains pseudoscalar operators and bounds $C_{10,\mu}^{\text{NP}}$
- Preferred solutions: $C_{9,\mu}^{\text{NP}} \neq 0$ or $C_{9,\mu}^{\text{NP}} \approx -C_{10,\mu}^{\text{NP}} \neq 0$
- Slight tension (2σ) in current $B_s^0 \rightarrow \mu^+ \mu^-$ world average favours $C_{9,\mu}^{\text{NP}} - C_{10,\mu}^{\text{NP}}$
- Moriond data allows more space for right-handed currents
- Additional solutions with LFU components (Algueró et al, 1809.08447)
- SMEFT: $b \rightarrow c \tau \nu$ and $b \rightarrow s \ell \ell$ anomalies \Rightarrow Large $b \rightarrow s \tau \tau$

$$(\bar{Q}_2 \gamma^\mu Q_3)(\bar{L}_3 \gamma_\mu L_3) + (\bar{Q}_2 \gamma^\mu \sigma^I Q_3)(\bar{L}_3 \gamma_\mu \sigma^I L_3) \approx 2 [(\bar{c}_L \gamma_\mu b_L)(\bar{\tau}_L \gamma^\mu \nu_{\tau L}) + (\bar{s}_L \gamma_\mu b_L)(\bar{\tau}_L \gamma^\mu \tau_L)]$$

First evidence of C/P in charm decays (5.3σ)

LHCb 1903.08726 $\Delta a_{CP} = (-15.4 \pm 2.9) \cdot 10^{-4}$, $\Delta a_{CP}^{\text{dir}} = (-15.7 \pm 2.9) \cdot 10^{-4}$



$$\Delta a_{CP} = a_{CP}(K^+K^-) - a_{CP}(\pi^+\pi^-)$$

HFLAV combination
 $a_{CP}^{\text{ind}} = (0.028 \pm 0.026)\%$
 $\Delta a_{CP}^{\text{dir}} = (-0.164 \pm 0.028)\%$
 Consistency with NO CPV hypothesis: 5×10^{-8}

$$a_{CP} \equiv \frac{\Gamma(D^0 \rightarrow f) - \Gamma(\bar{D}^0 \rightarrow f)}{\Gamma(D^0 \rightarrow f) + \Gamma(\bar{D}^0 \rightarrow f)}$$

Large uncertainty in SM prediction:

- Naive perturbative QCD (+ LCSR) $\Rightarrow |\Delta a_{CP}^{\text{dir}}| \leq 3 \cdot 10^{-4}$ Chala et al, 1903.10490
- Re-scattering: $\Delta a_{CP}^{\text{dir}}$ $\Rightarrow \Delta U = 0$ rule in charm Grossman-Schacht, 1903.10952