

# Experimental and theory overview of QCD phase diagram

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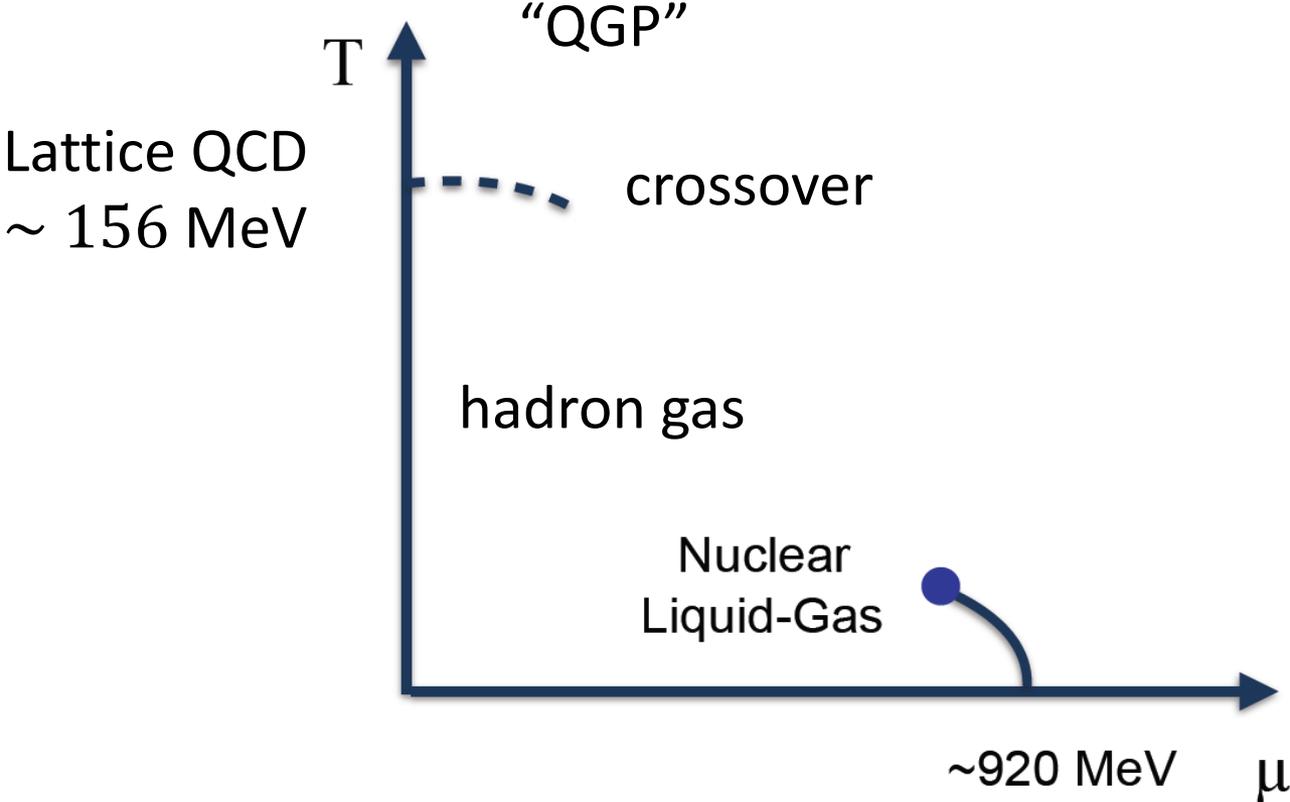
Supported by NCN  
2018/30/Q/ST2/00101



## Outline

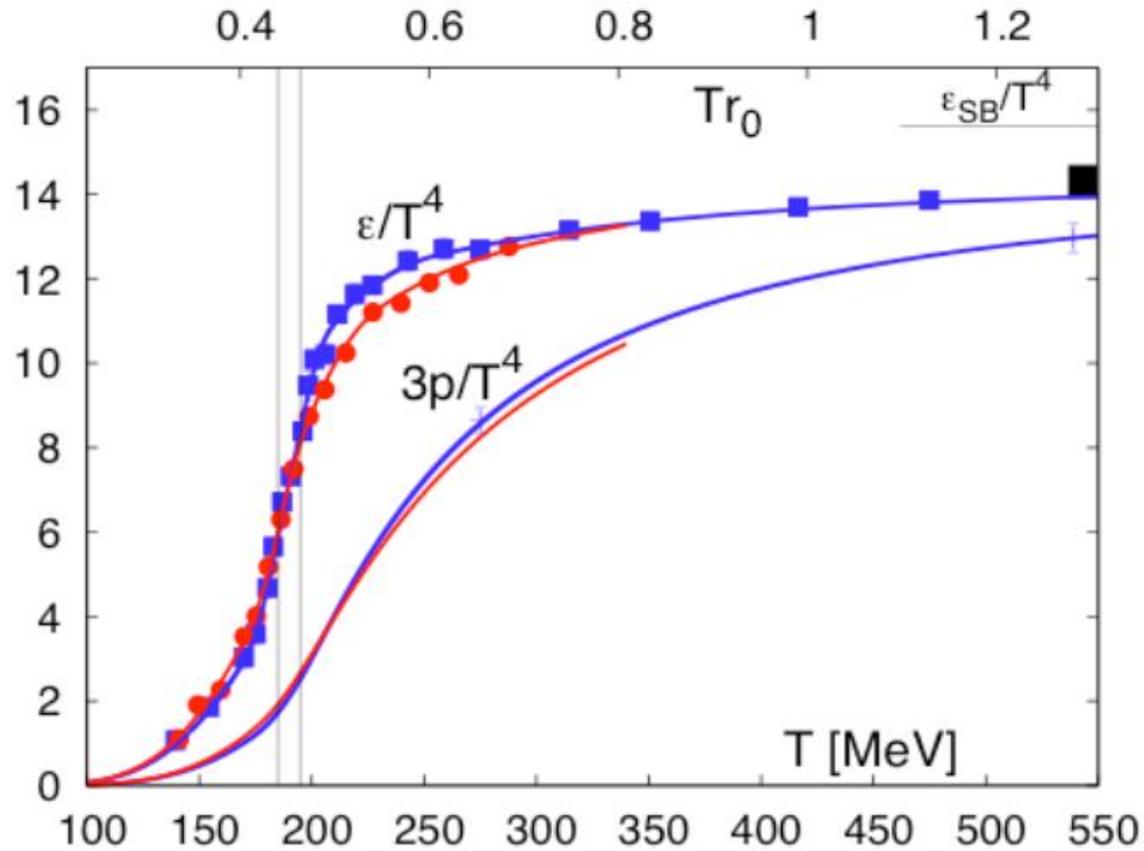
- what we know
- theory vs. experiment
- cumulants, factorial cumulants, factorial moments
- long-range and short-range correlations
- expectations
- measurements and interpretation
- summary

# The QCD phase diagram



Figured taken from V.Koch

# Lattice QCD



A smooth and wide crossover

# Hopes

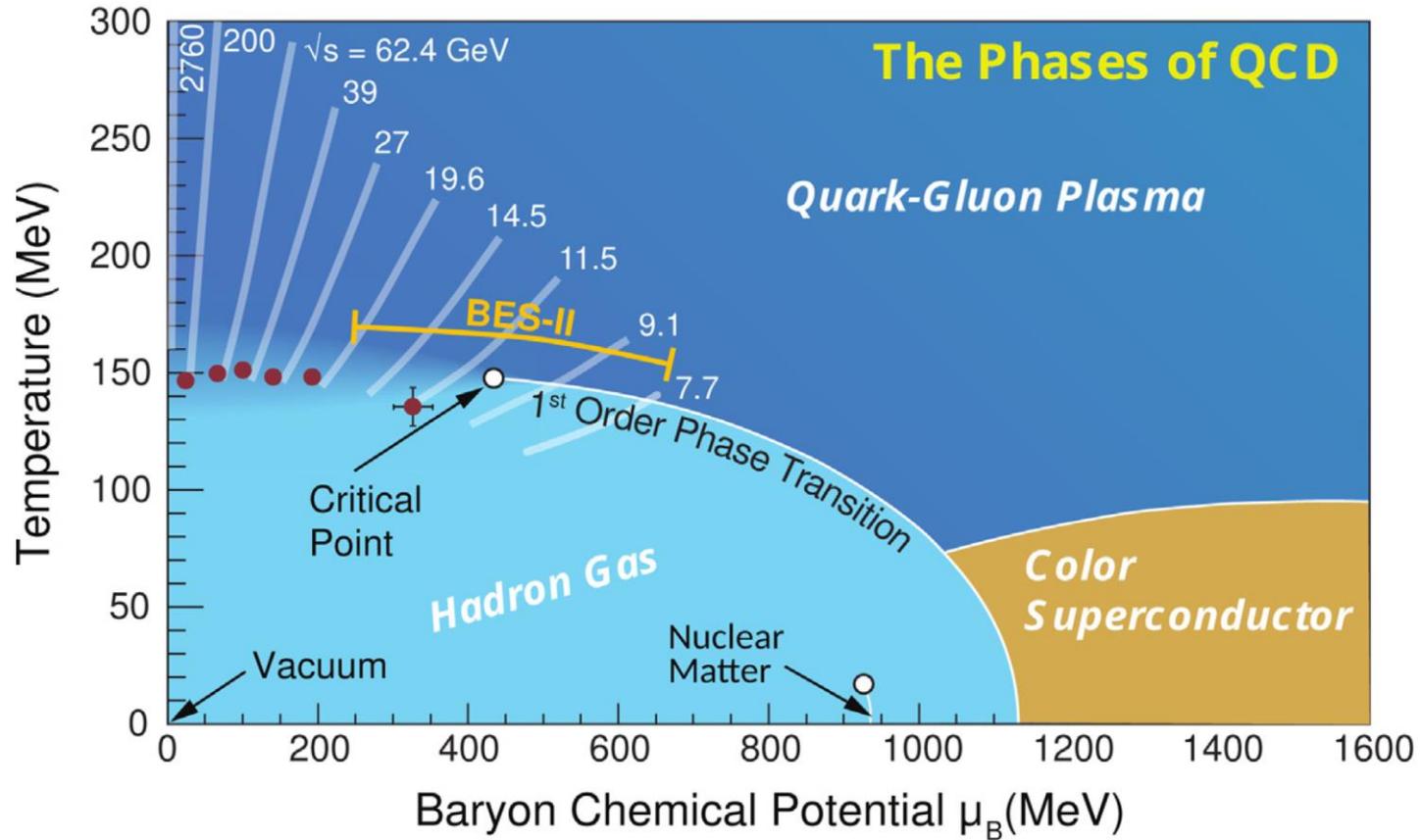
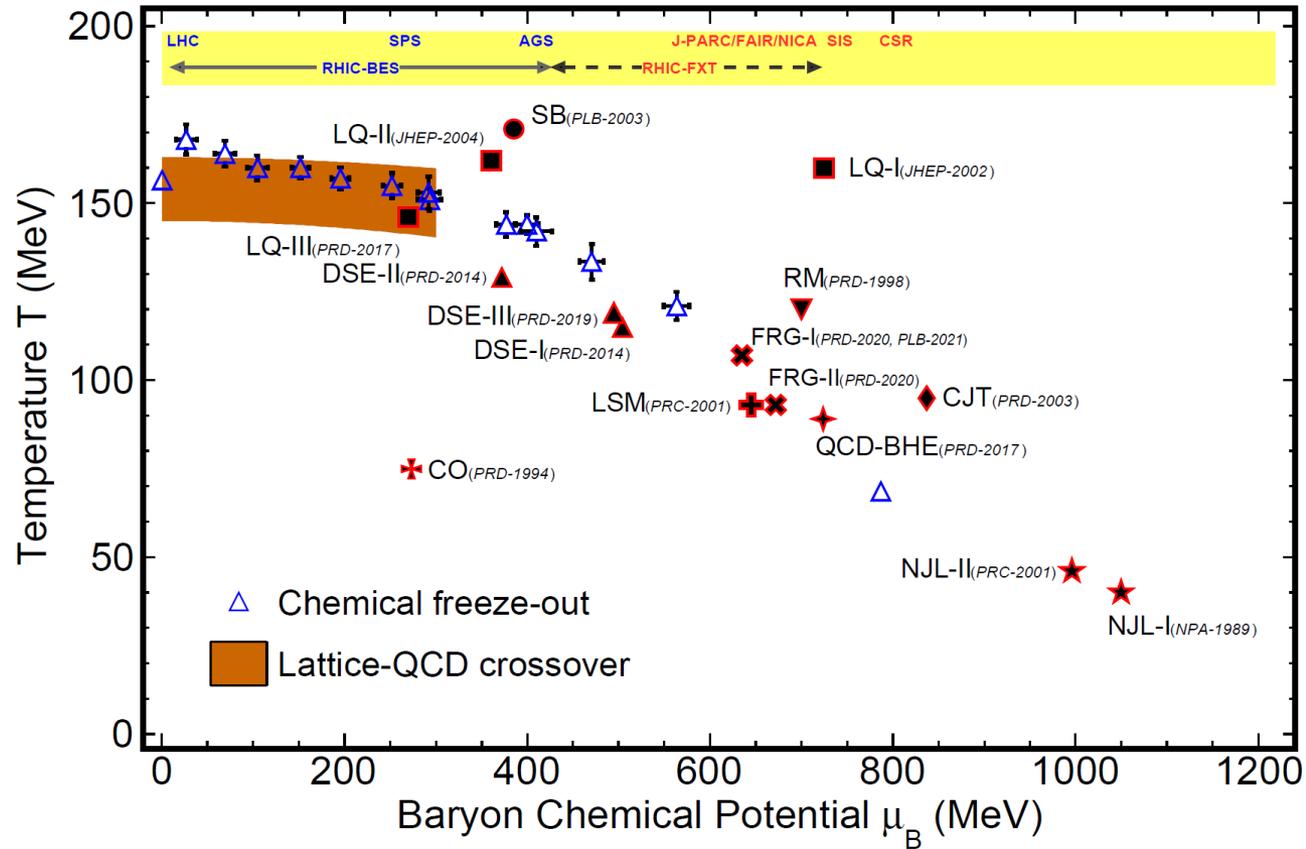


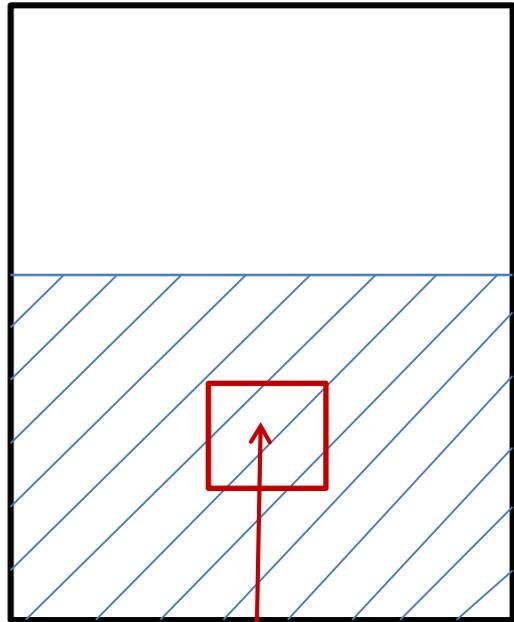
Figure from Phys. Rept. 853 (2020) (AB, S.Esumi, V.Koch, J.Liao, M.Stephanov, N.Xu)

# Critical point?



A. Pandav, D. Mallick, B. Mohanty, 2203.07817  
M. Stephanov, hep-lat/0701002

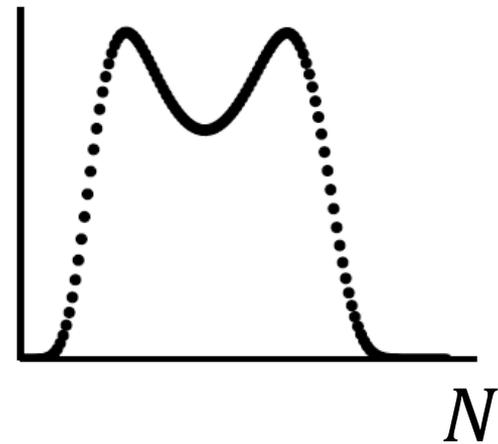
How to approach this problem?  
Consider water vapour transition



$P(N)$

right at the phase transition

$P(N)$



number of  $H_2O$  molecules

so we measure multiplicity distributions

In QCD we use, e.g., net-baryon, net-charge, net-strangeness

# Theory vs. experiment

## **Theory**

Coordinate space

Fixed volume

Long-lived

Conserved charges

## **Experiment**

Momentum space

Expanding and fluctuating volume

Extremely short-lived

Non-conserved numbers

Lots of detector problems + various corrections (e.g. volume fluctuation)

Some progress towards dynamical models.

## HYDRO+ (hydrodynamics with critical modes)

M. Stephanov, Y. Yi, PRD 98, 036006 (2018)

K. Rajagopal, G. Ridgway, R. Weller, Y. Yin, PRD 102, 094025 (2020)

*We consider ... a generic extension of hydrodynamics by a parametrically slow mode or modes ("Hydro+") and a description of fluctuations out of equilibrium.*

## Equation of state with critical point

P.Parotto, M.Bluhm, D.Mroczek, M.Nahrgang, J. Noronha-Hostler, K.Rajagopal, C.Ratti, T.Schäfer, M. Stephanov, PRC 101, 034901 (2020)

*We construct a family of equations of state for QCD in the temperature range  $30 \text{ MeV} \leq T \leq 800 \text{ MeV}$  and in the chemical potential range  $0 \leq \mu_B \leq 450 \text{ MeV}$ . These equations of state match available lattice QCD results up to  $O(\mu_B^4)$  and in each of them we place a critical point in the three-dimensional (3D) Ising model universality class.*

## Molecular dynamics

[V.A. Kuznietsov, O. Savchuk, M.I. Gorenstein, V. Koch, V. Vovchenko, 2201.08486](#)

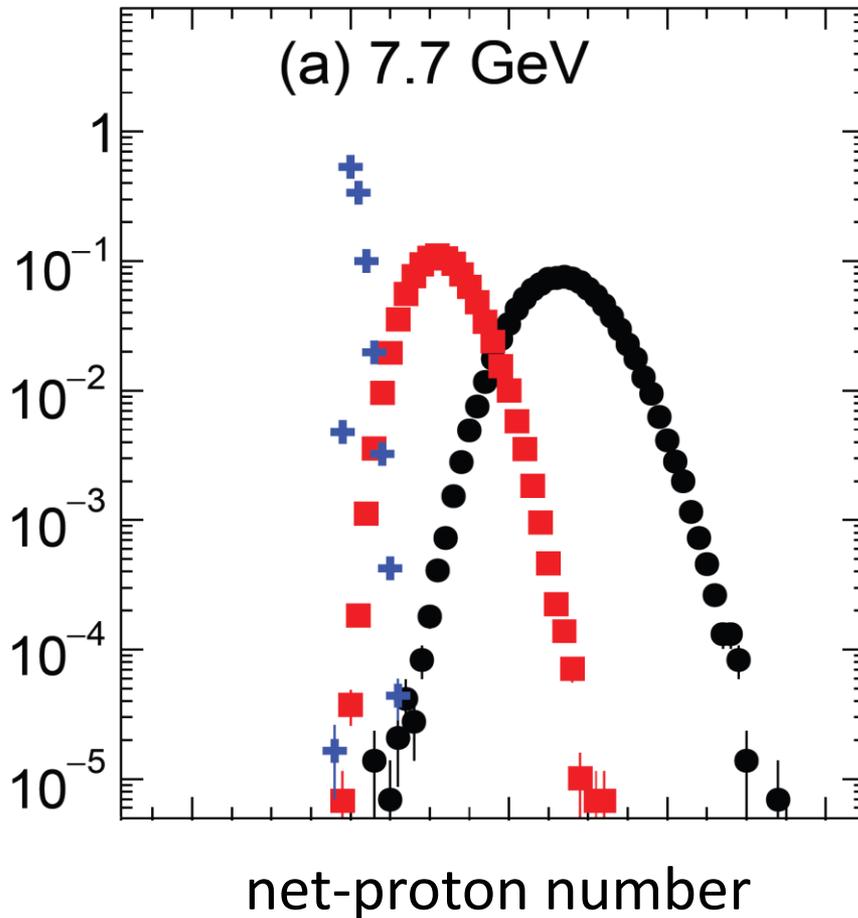
*We find that large fluctuations associated with the critical point are observed when measurements are performed in coordinate subspace, but, in the absence of collective flow and expansion, are essentially washed out when momentum cuts are imposed instead.*

## Hydrodynamics with many non-critical contributions

[V.Vovchenko, V.Koch, C.Shen, PRC 105 \(2022\) 1, 014904](#)

*The experimental data of the STAR Collaboration are consistent at  $\sqrt{s_{NN}} \gtrsim 20$  GeV with simultaneous effects of global baryon number conservation and repulsive interactions in the baryon sector...*

So we measure multiplicity distributions



## Au+Au Collisions

$0.4 < p_T < 2.0$  (GeV/c)

$|y| < 0.5$

● 0-5%

■ 30-40%

+ 70-80%

raw distributions  
(not corrected)

For baryons absolutely minimal goal is to see any deviations from Poisson (Skellam) distribution.

It is difficult to see something in multiplicity distributions. If there is any signal, it is likely very tiny.

We usually characterize  $P(N)$  by:

- cumulants  $\kappa_n$
- factorial cumulants,  $C_n$  (or  $\hat{C}_n$ )
- factorial moments  $F_n$  (mean number of pairs, triplets, etc.)

**Warning.** STAR uses opposite notation  $\kappa_n \leftrightarrow C_n$

On the experimental side we need to measure various fluctuation observables and hope to see some nontrivial energy or/and system-size dependence

see, e.g.,

Stephanov, Rajagopal, Shuryak, PRL (1998)

Stephanov, PRL (2009)

Skokov, Friman, Redlich, PRC (2011)

There are many results:

ALICE, STAR, HADES

Cumulants, factorial cumulants

Proton  $v_1$  (STAR)

HBT radii (STAR)

R.A. Lacey, PRL 114 (2015) 142301

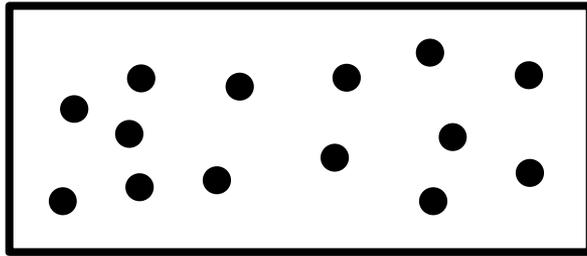
NA61/SHINE

Intermittency, cumulants

Scaled variance

Strongly intensive variables

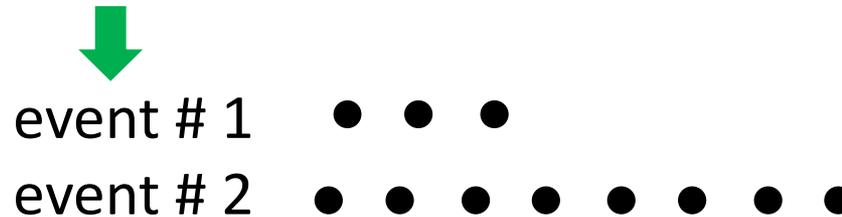
# Poisson distribution (no correlations)



$$N = 10^{10}$$

$$p = 10^{-9}$$

$$\langle n \rangle = Np = 10$$



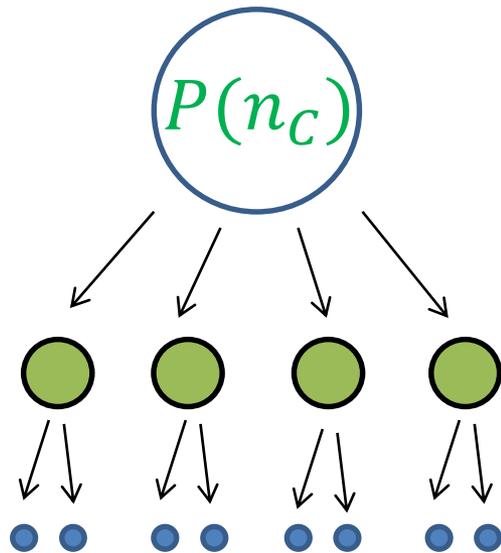
$P(n) = \text{Poisson}$  if  $N \rightarrow \infty$ ,  $p \rightarrow 0$ ,  $Np = \langle n \rangle$

cumulants  $\kappa_i = \langle n \rangle$

factorial cumulants  $C_i = 0$

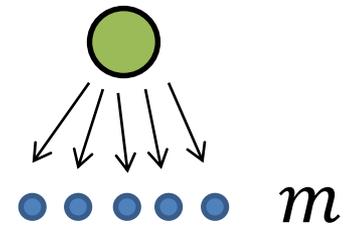
factorial moments  $F_i = \langle n \rangle^i$

# Factorial cumulants – example



Poisson

$m$  particle cluster



$$C_2 \neq 0$$

$$C_k = 0, k > 2$$

$$C_{2,3,\dots,m} \neq 0$$

$$C_k = 0, k > m$$

factorial  
cumulants

$$C_k = \frac{d^k}{dz^k} \ln \left( \sum_n P(n) z^n \right) \Big|_{z=1}$$

## Two-particle correlation function

$$\rho_2(y_1, y_2) = \rho(y_1)\rho(y_2) + C_2(y_1, y_2)$$

Integrating both sides over some bin in rapidity

$$\langle n(n - 1) \rangle = \langle n \rangle^2 + C_2$$

$$C_2 = \int C_2(y_1, y_2) dy_1 dy_2$$

Same with multiparticle correlations.

Factorial cumulants are integrated multiparticle correlation functions

## Factorial cumulants vs cumulants

factorial  
cumulant

$$C_i = \frac{d^i}{dz^i} \ln \left( \sum_n P(n) z^n \right) \Big|_{z=1}$$

cumulant

$$\kappa_i = \frac{d^i}{dt^i} \ln \left( \sum_n P(n) e^{tn} \right) \Big|_{t=0}$$

Poisson

$$C_i = 0, \kappa_i = \langle n \rangle$$

cumulants naturally appear  
in statistical physics

$$\ln(Z) = \ln \left( \sum_i e^{-\beta(E_i - \mu N_i)} \right)$$

## Cumulants (one species of particles)

$$\kappa_2 = \langle N \rangle + C_2$$

$$\kappa_3 = \langle N \rangle + 3C_2 + C_3$$

$$\kappa_4 = \langle N \rangle + 7C_2 + 6C_3 + C_4$$

Cumulants mix integrated correlation functions of different orders

They might be dominated by  $\langle N \rangle$ .

See, e.g.,

B. Ling, M. Stephanov, PRC 93 (2016) 034915

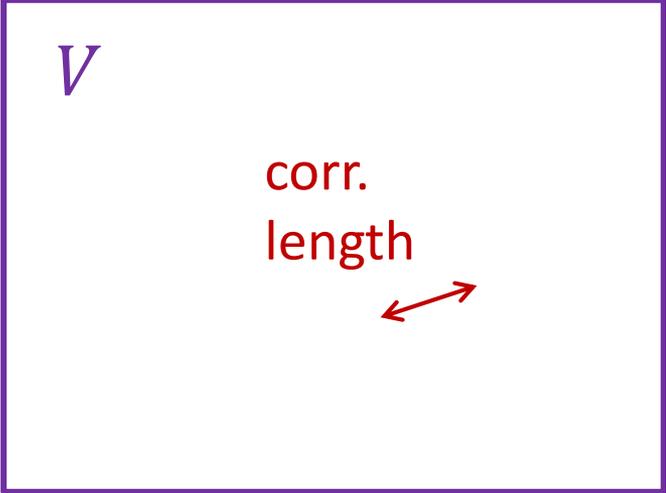
AB, V.Koch, N.Strodthoff , PRC 95 (2017) 054906

“Cumulant ratios do not depend on volume”

but depend on volume fluctuation

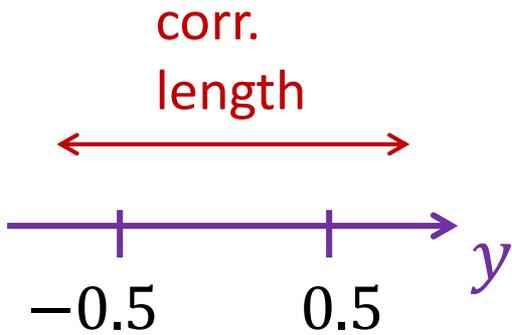
It is true if a correlation length is much smaller than the system size

coordinate space



Here this condition is satisfied

momentum rapidity space



Correlation length is usually larger than one unit of rapidity.

Cumulant ratios are expected to depend on acceptance in rapidity

## Short-range correlations

$$C_i \sim \langle N \rangle \sim \Delta y$$

$$\kappa_i \sim \langle N \rangle \sim \Delta y$$

## Long-range correlations (expected in rapidity)

$$C_i \sim \langle N \rangle^i \sim (\Delta y)^i$$

$\kappa_i$  is complicated, for example

$$\kappa_4 = \langle N \rangle + (\sim \langle N \rangle^2) + (\sim \langle N \rangle^3) + (\sim \langle N \rangle^4)$$

$$\kappa_2 = \langle N \rangle + (\sim \langle N \rangle^2)$$

polynomial in  $\Delta y$

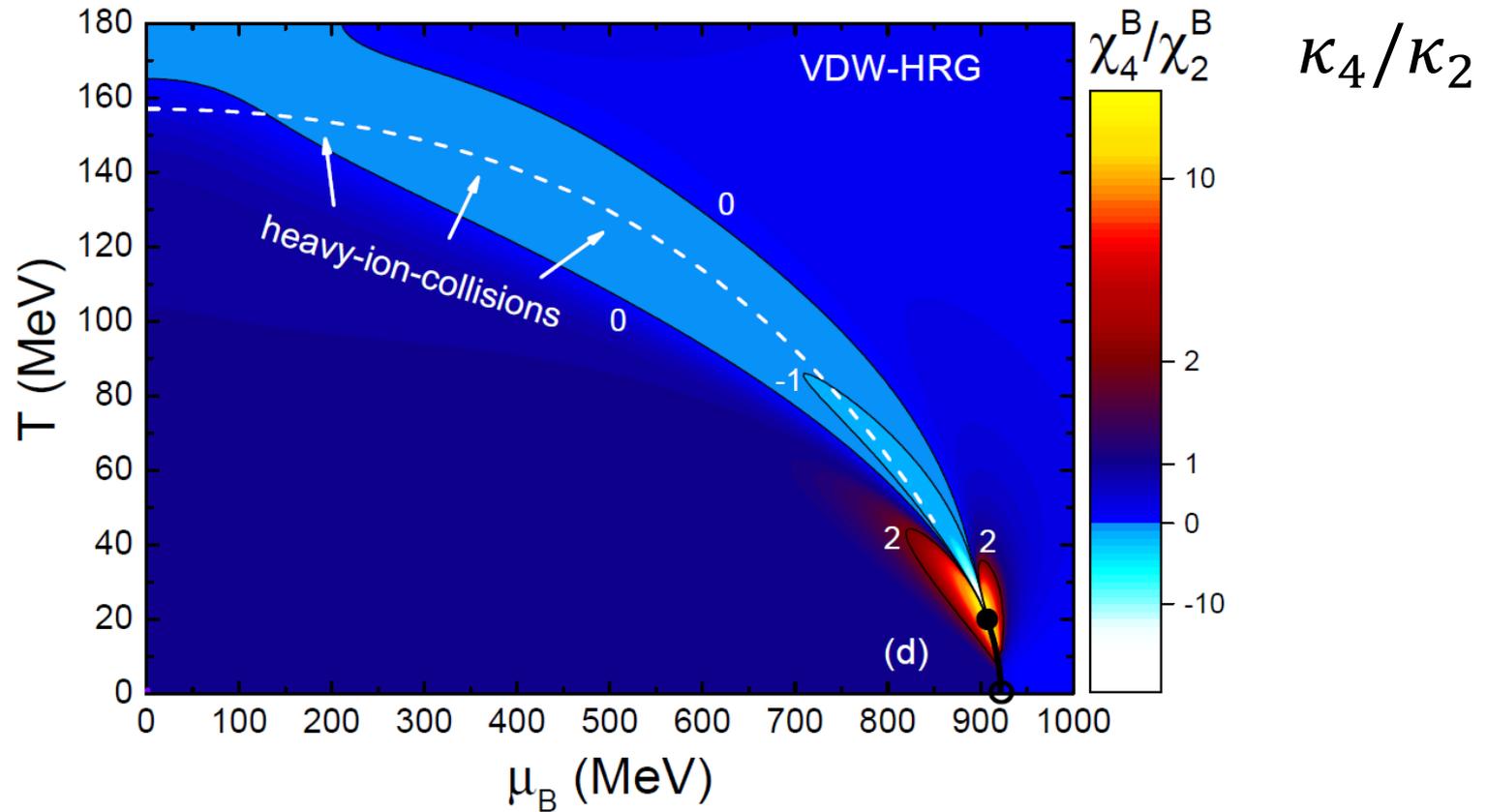
Cumulant ratios may strongly depend on acceptance in rapidity and in transverse momentum

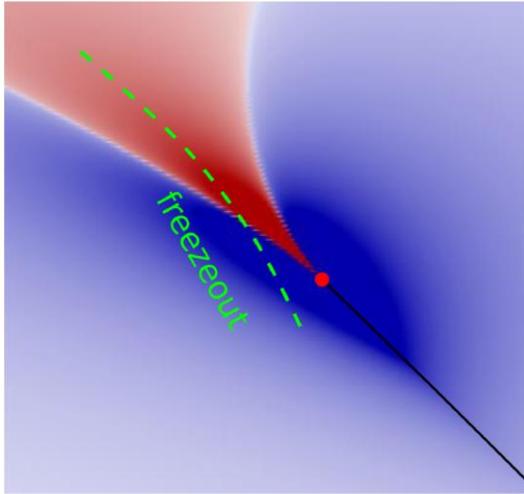
Comparison with models which do not have experimental acceptance is questionable

Comparison with lattice QCD calculations is very tricky

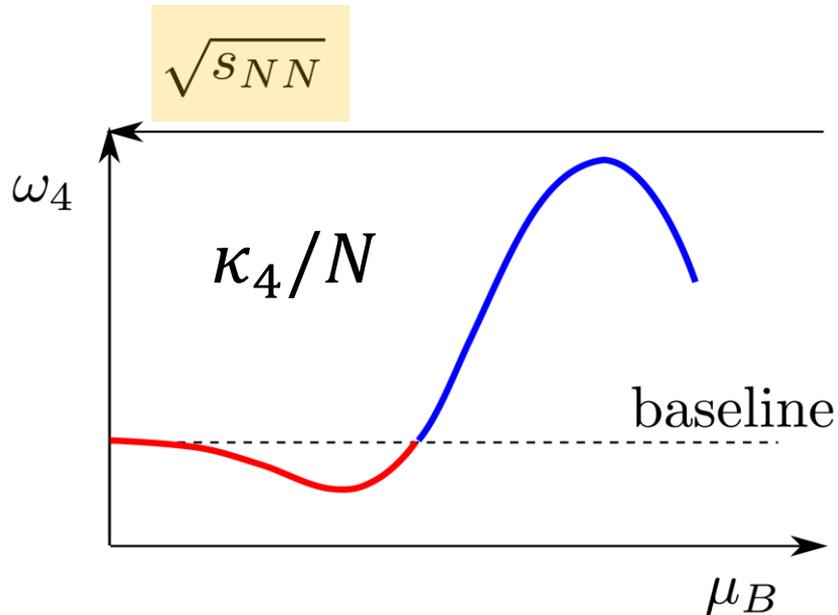
With long-range rapidity correlations the cleanest observable is  $\frac{C_i}{\langle N \rangle^i}$

# HRG with attractive and repulsive Van der Waals interactions between (anti)baryons

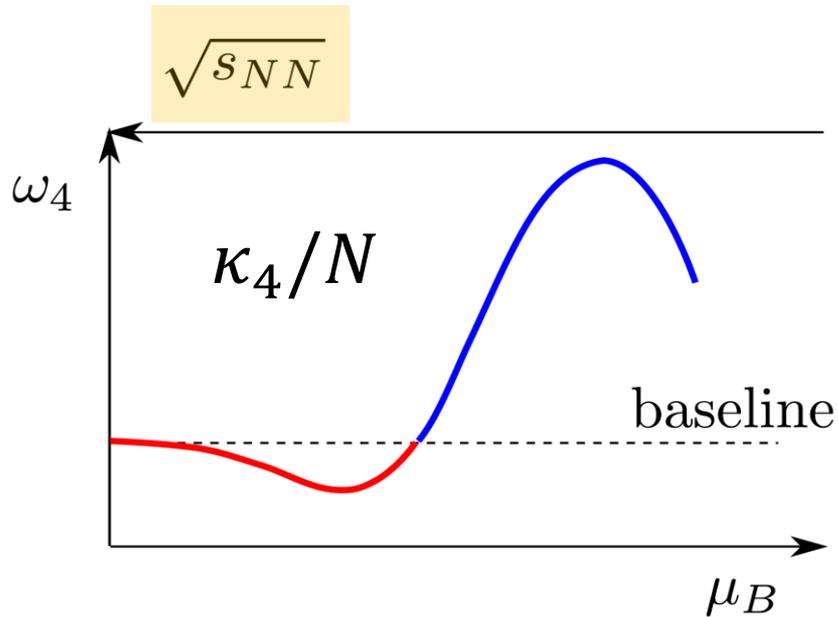




Density plot of the quartic cumulant obtained by mapping the Ising model into QCD.  
Freezeout line is for demonstration only.



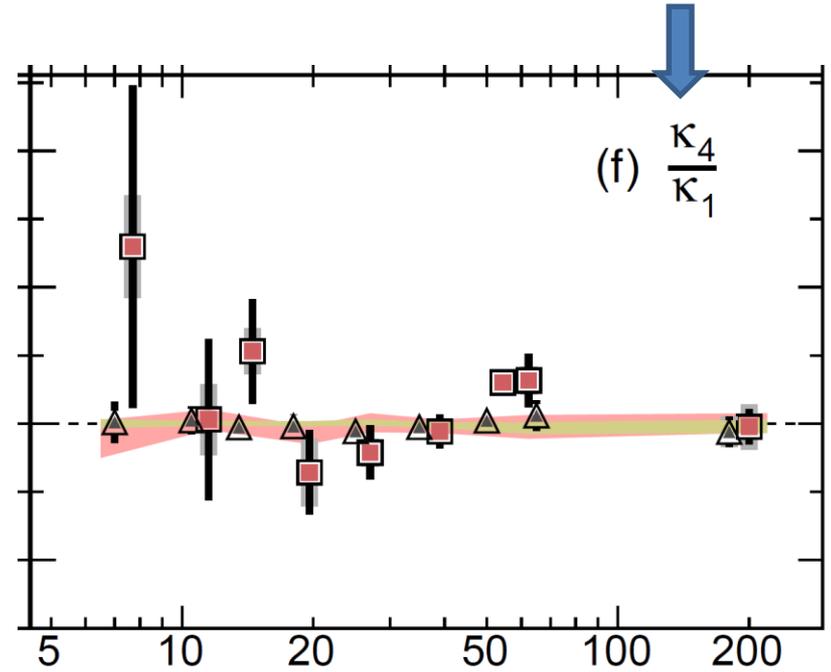
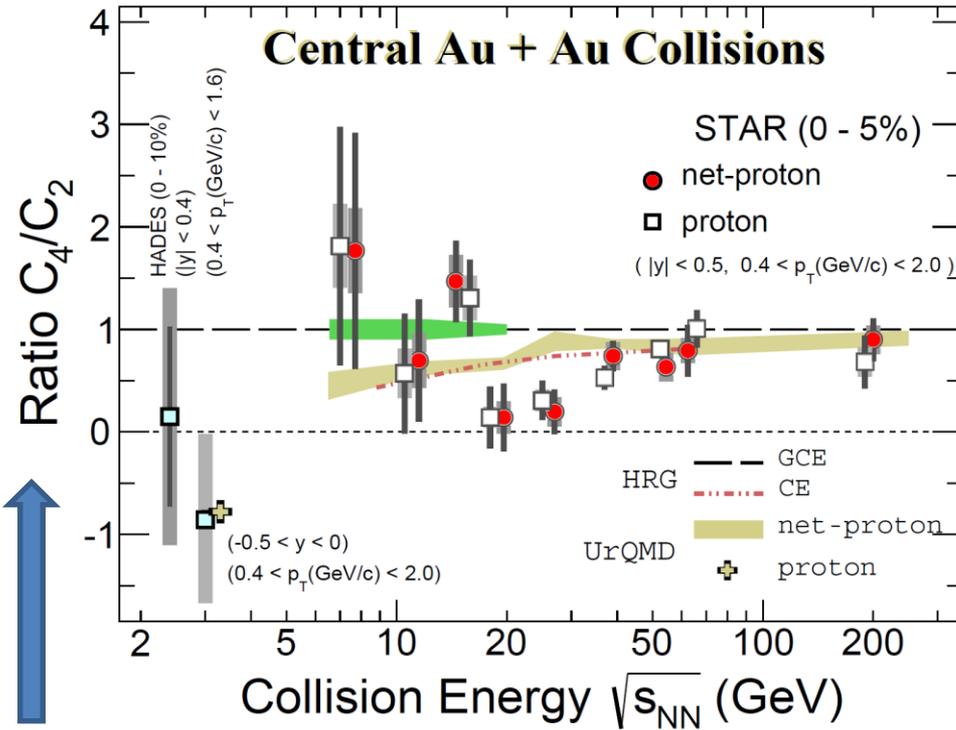
Normalized quartic cumulant of proton multiplicity



D. Mroczek, A.R. Nava Acuna, J. Noronha-Hostler, P. Parotto, C. Ratti, M.A. Stephanov, PRC 103 (2021) 3, 034901

*We find that, while the peak remains a solid feature, the presence of the critical point does not necessarily cause a dip in  $\chi_4^B$  on the freezeout line below the transition temperature.*

kindly read  $C_4/C_1$

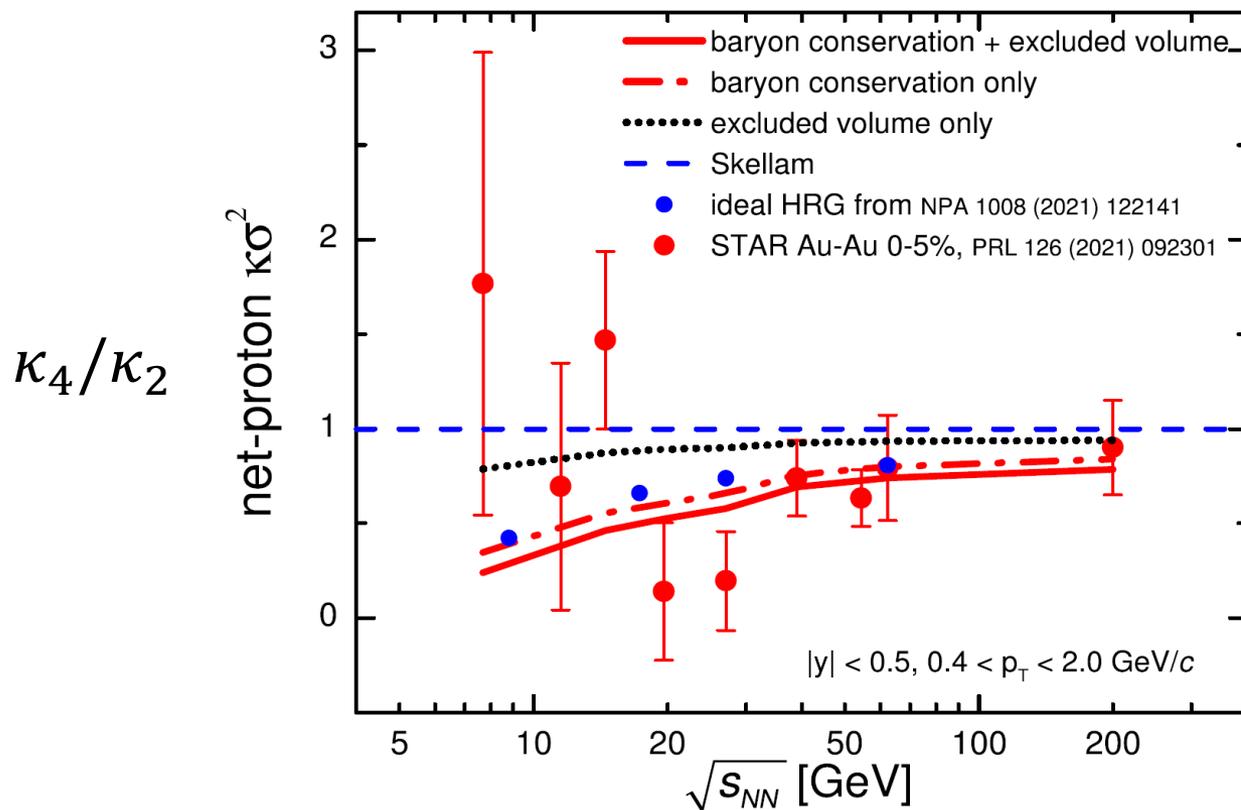


kindly read  $\kappa_4/\kappa_2$

Visible four-proton correlations at 7.7 GeV (large errors)

A hint of non-monotonic dependence

# STAR data vs. hydrodynamics with baryon conservation and excluded volume

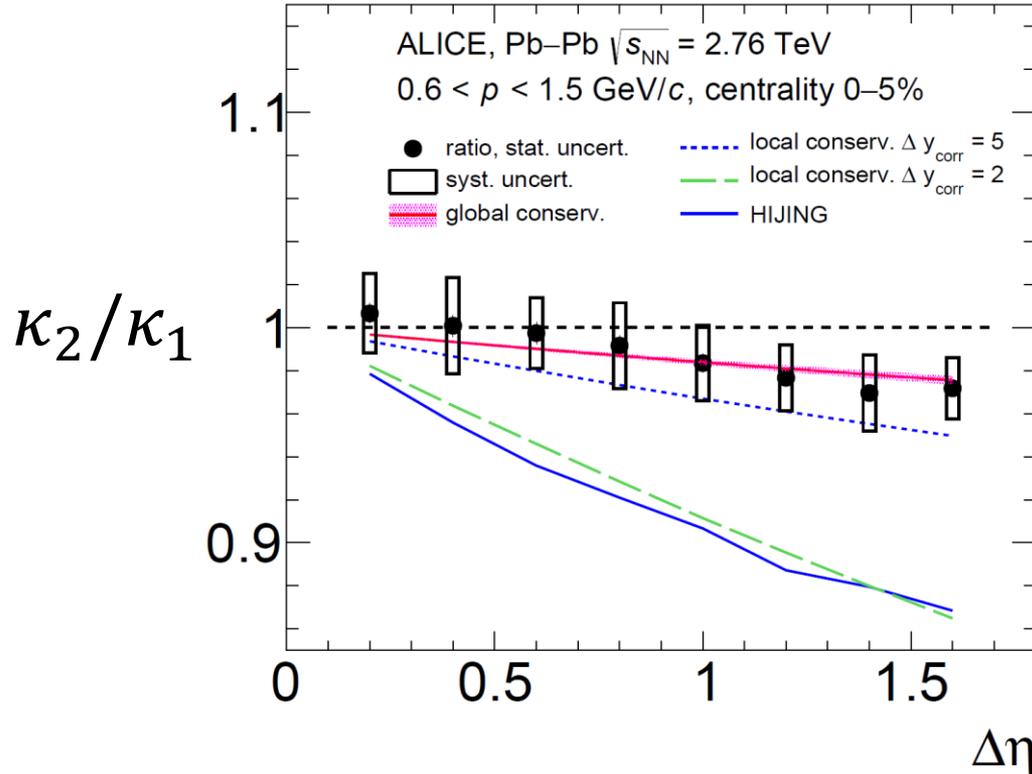


Baryon conservation for  $\sqrt{s} > 20 \text{ GeV}$

V.Vovchenko, V.Koch, C.Shen, PRC 105, 014904 (2022)

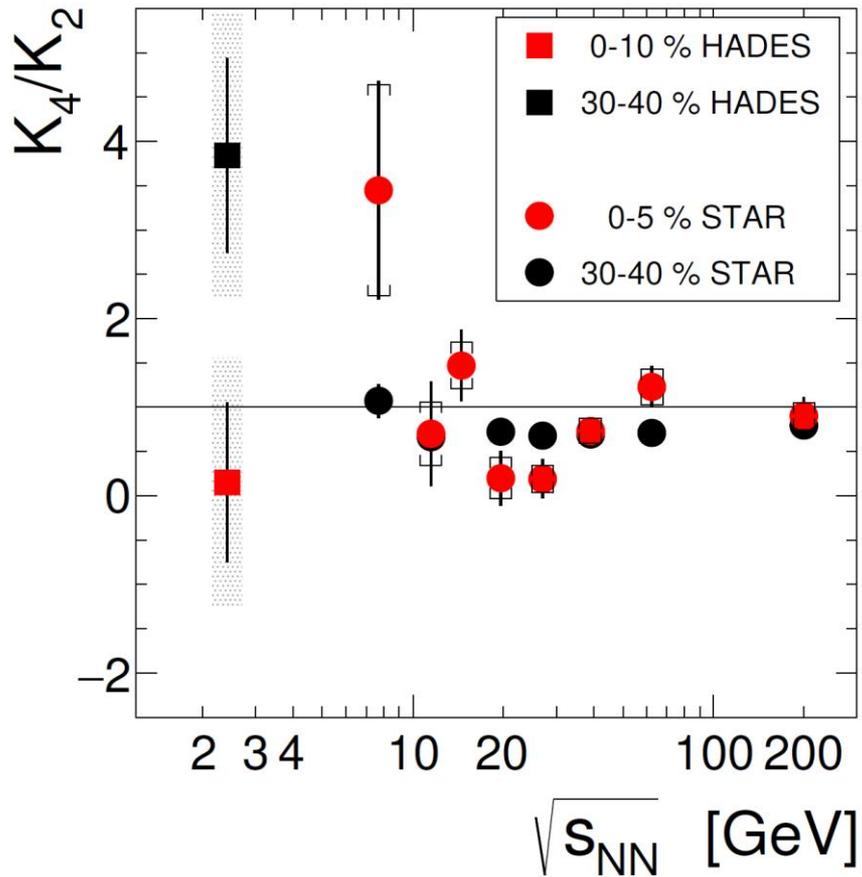
P.Braun-Munzinger, B.Friman, K.Redlich, A.Rustamov, J.Stachel, NPA 1008 (2021) 122141

AB, V.Koch, V.Skokov, PRC 87 (2013) 1, 014901

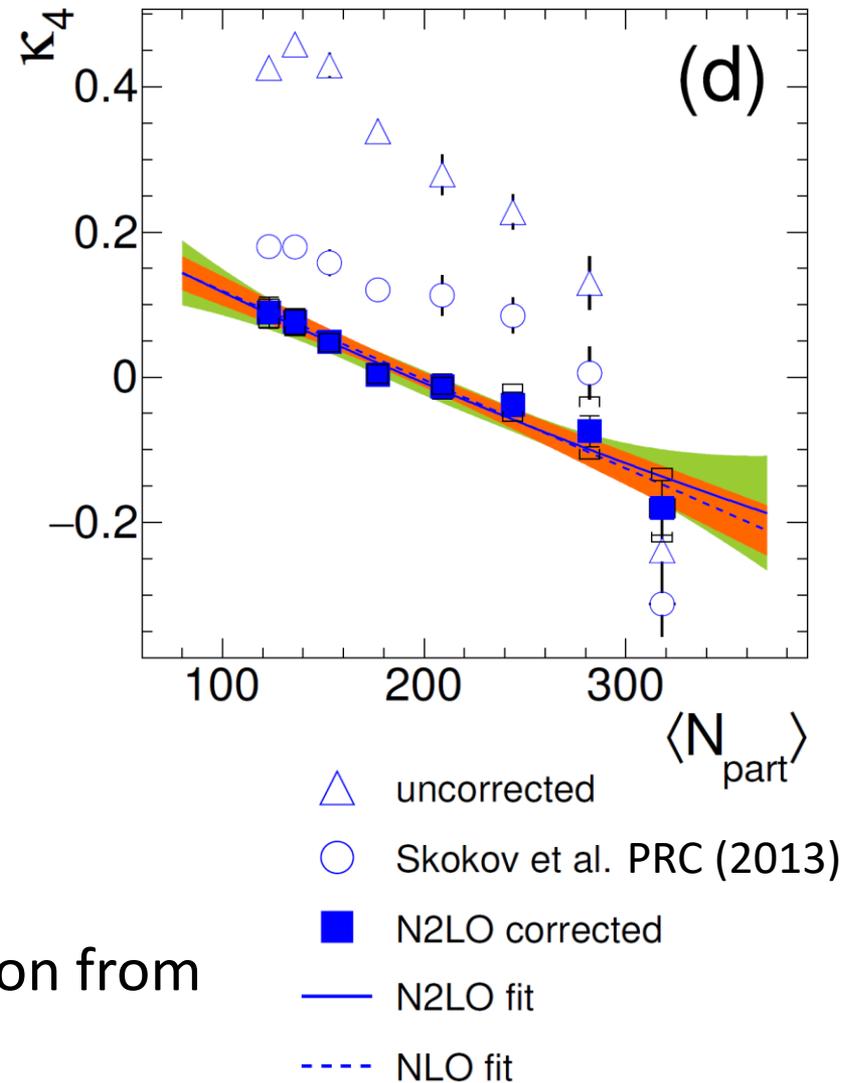


**Global** (not local) baryon conservation! Something to understand.

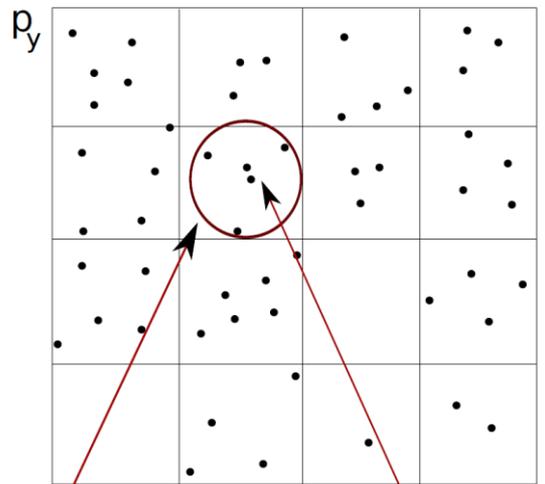
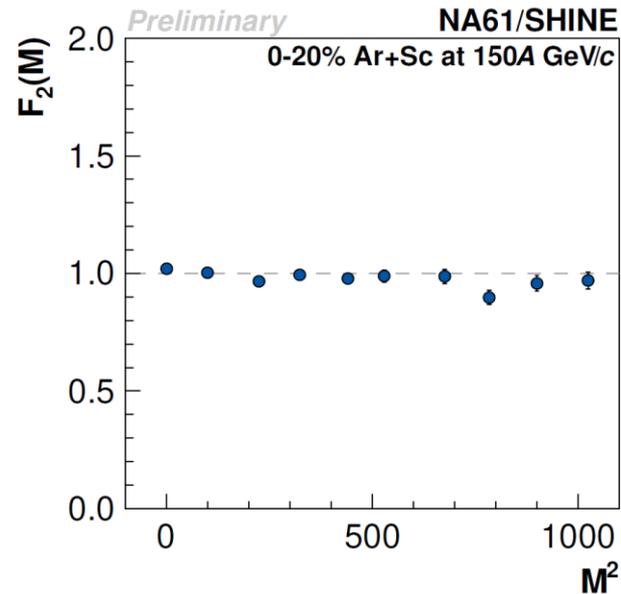
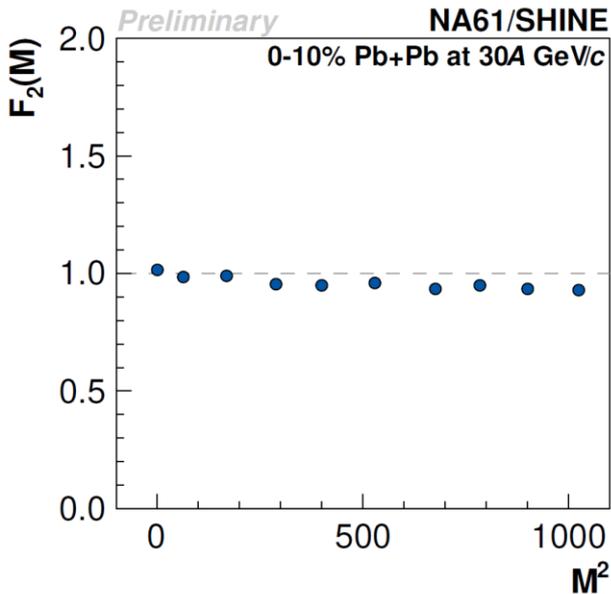
It would be good to measure proton, antiproton and mixed proton-antiproton factorial cumulants. See a poster by **Michał Barej**



Significant correction from volume fluctuation



# NA61/SHINE Collaboration



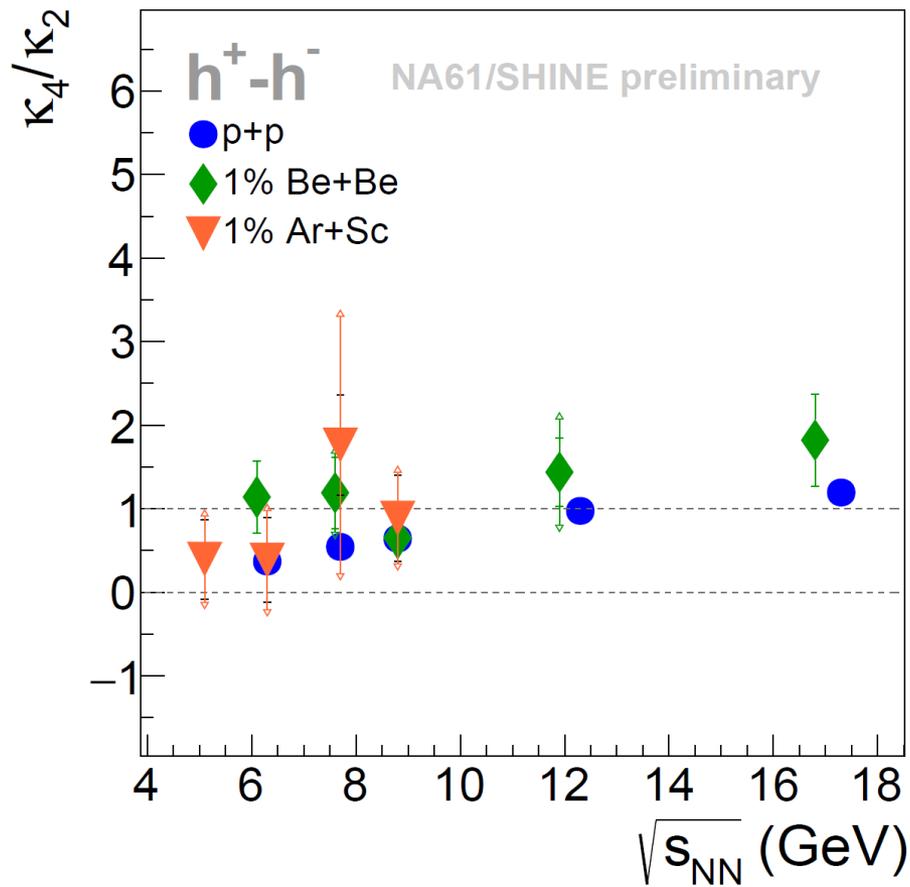
statistical uncertainties only

$$F_2(M) = \frac{\left\langle \frac{1}{M^2} \sum_{i=1}^{M^2} n_i (n_i - 1) \right\rangle}{\left\langle \frac{1}{M^2} \sum_{i=1}^{M^2} n_i \right\rangle^2}$$

$$F_2(M) \sim (M^2)^{5/6}$$

N.G. Antoniou, F.K. Diakonou, A.S. Kapoyannis,  
K.S. Kousouris, PRL 97, 032002 (2006)

# NA61/SHINE Collaboration



No critical signal. Consistent with p+p.

## Conclusions

Interesting and important physics but so far no success

Clear signal of (global?) baryon conservation

Interesting STAR point at 7.7 GeV. Four-proton correlations (physics?)

We definitely need better statistics

Hopefully some progress will come from lattice QCD