



29TH INTERNATIONAL
CONFERENCE ON ULTRARELATIVISTIC
NUCLEUS - NUCLEUS COLLISIONS
APRIL 4-10, 2022
KRAKÓW, POLAND



Experimental overview of electromagnetic probes

Klaus Reygers, Heidelberg University

Why electromagnetic probes?

Direct information about the medium, not affected by hadronization

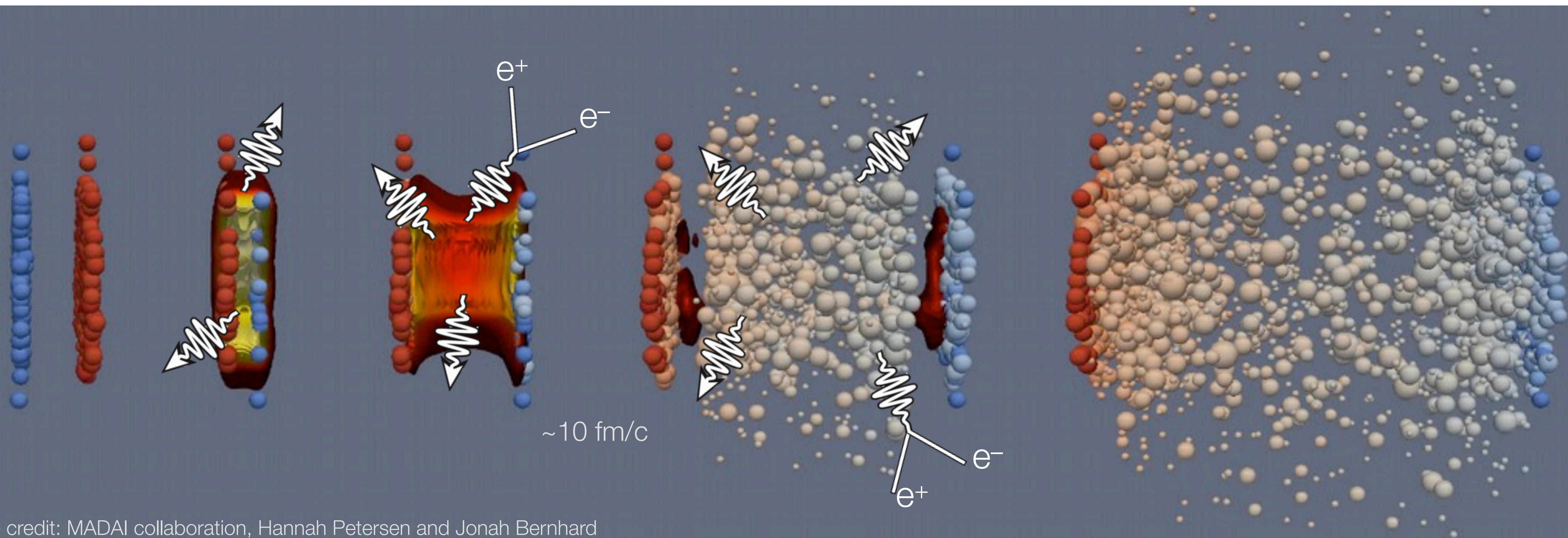


figure credit: MADAI collaboration, Hannah Petersen and Jonah Bernhard

γ , γ^* , W , Z :

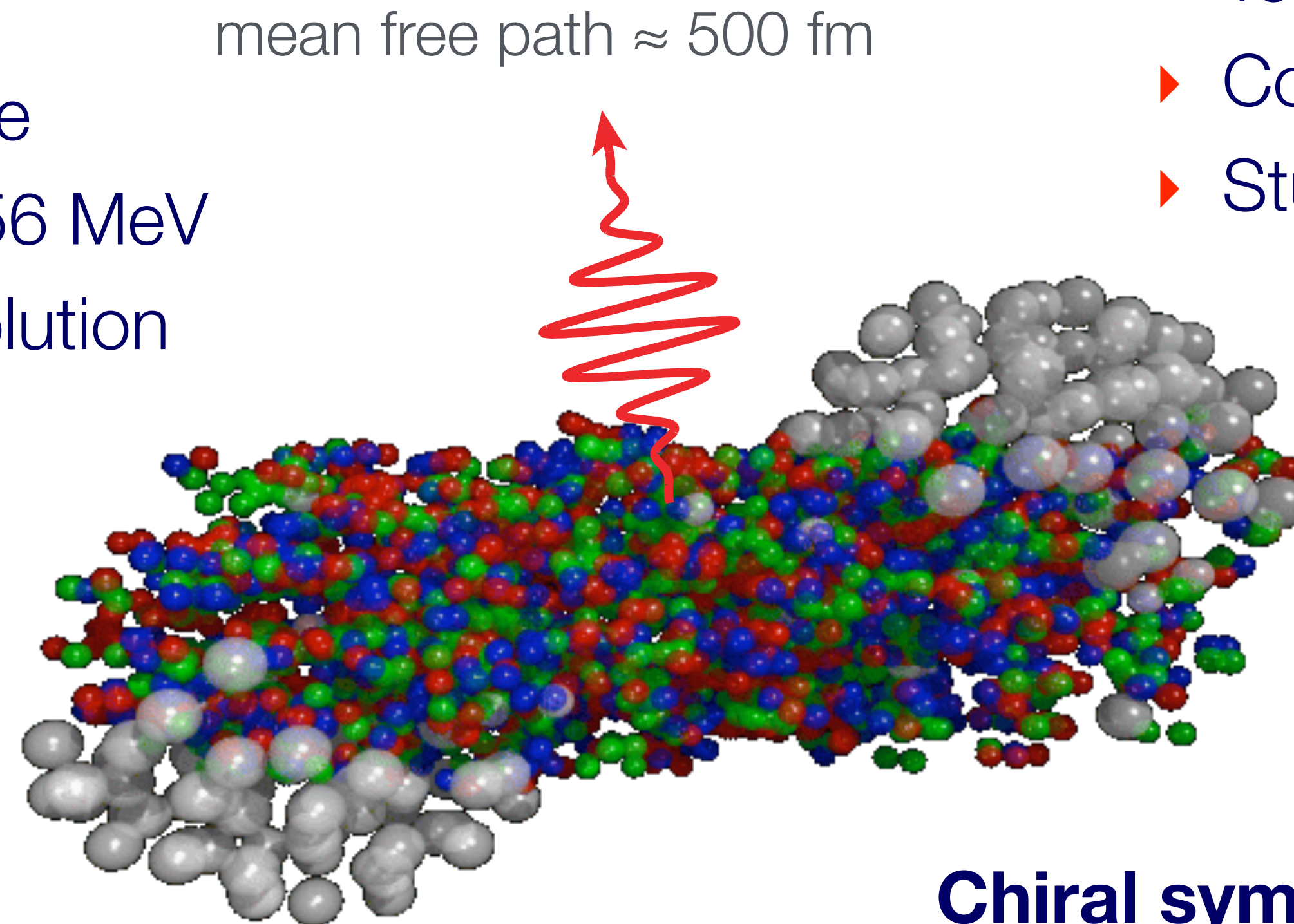
different information about the medium
than from light-quark and heavy-quark observables

Why electromagnetic probes?

Different aspects of electromagnetic probes

Thermal radiation

- ▶ effective QGP temperature
- ▶ Demonstrate $T > T_{pc} = 156$ MeV
- ▶ Constrain space-time evolution



Preequilibrium phase

- ▶ mechanism of equilibration
- ▶ transition to hydro phase

Initial hard scattering

- ▶ Test N_{coll} scaling
- ▶ Constrain nuclear PDFs
- ▶ Study energy loss: γ/Z -tagged jets

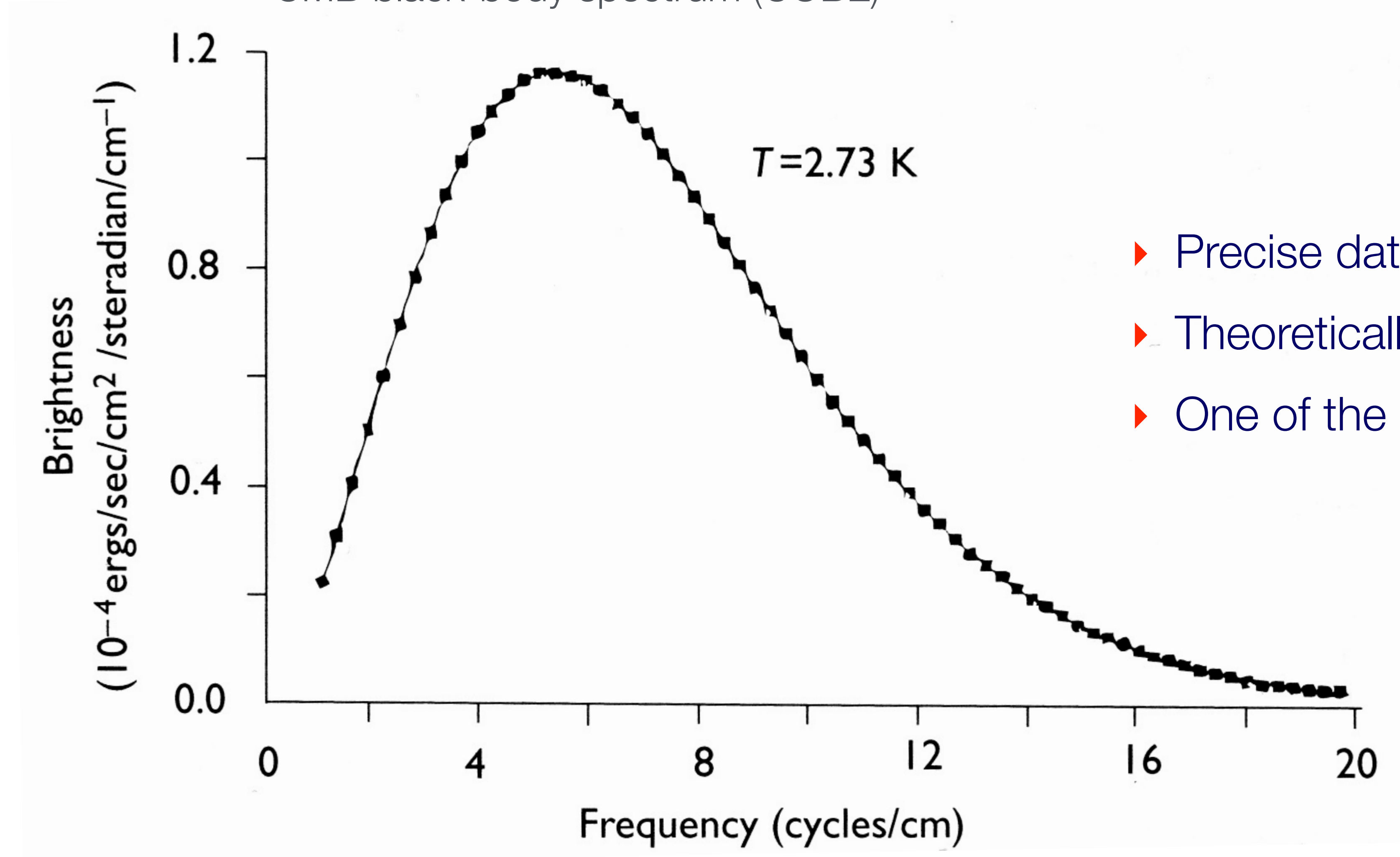
Chiral symmetry restoration (dileptons)

- ▶ ρ broadening
- ▶ constrain mechanism

A central goal ...

Temperature through the measurement of thermal radiation

CMB black-body spectrum (COBE)



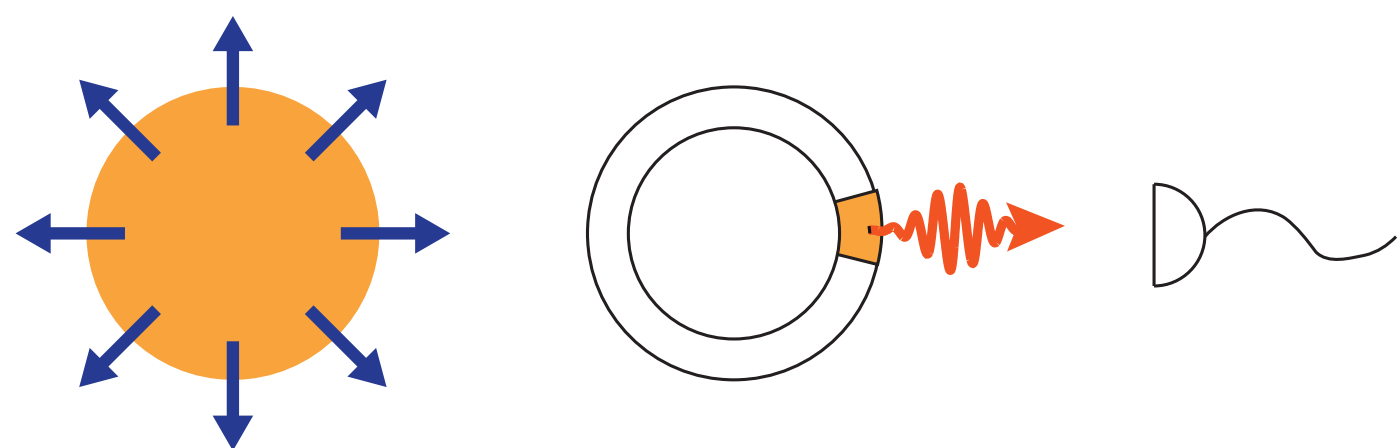
- ▶ Precise data
- ▶ Theoretically understood
- ▶ One of the pillars of cosmology

Thermal radiation

Basic expectations: direct photons

Thermal radiation expected to dominate for $p_T < 2-3 \text{ GeV}/c$

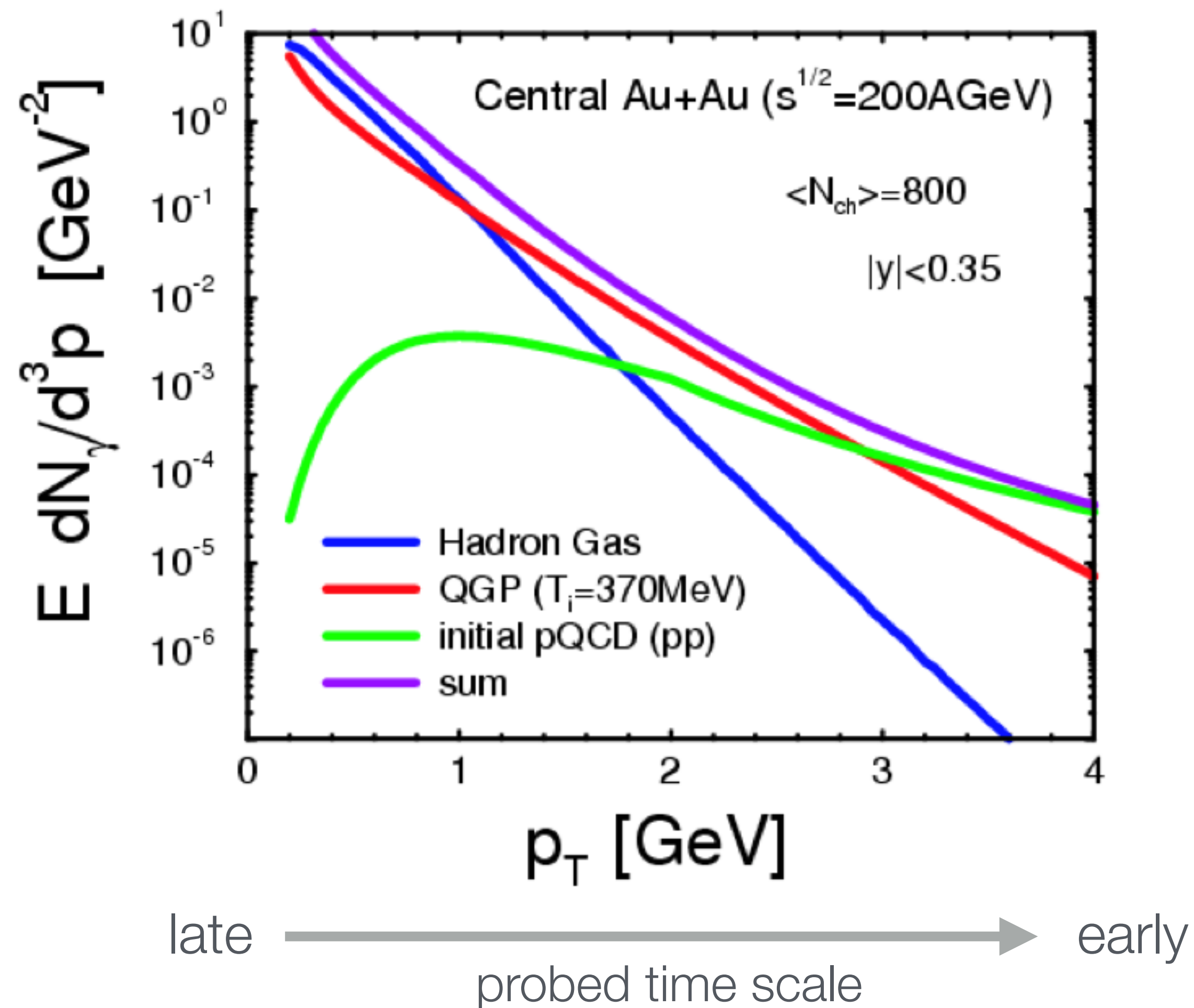
Inverse slope parameter T_{eff} affected by blueshift: $T_{\text{eff}} > T_{pc}$ does not necessarily indicate QGP formation



$$E_\gamma \frac{d^3 N_\gamma}{d^3 p_\gamma} \propto e^{-E_\gamma/T_{\text{eff}}} \quad T_{\text{eff}} = \underbrace{\sqrt{\frac{1 + \beta_{\text{flow}}}{1 - \beta_{\text{flow}}}}}_{2 \text{ for } \beta_{\text{flow}}=0.6} \times T$$

Possibly significant contribution of preequilibrium photons for $p_T \gtrsim 3 \text{ GeV}/c$

Turbide, Rapp, Gale,
Phys. Rev. C 69 (014902), 2004



G. David, Rept.Prog.Phys. 83 (2020) 4, 046301
 A. Monnai, 2203.13208

Thermal radiation

Basic expectations: dileptons

ρ meson sensitive to surrounding medium
($\tau_\rho = 1.3 \text{ fm}/c < \tau_{\text{fireball}}$)

ρ broader in hot baryon-rich medium

Connection between chiral symmetry restoration and ρ melting

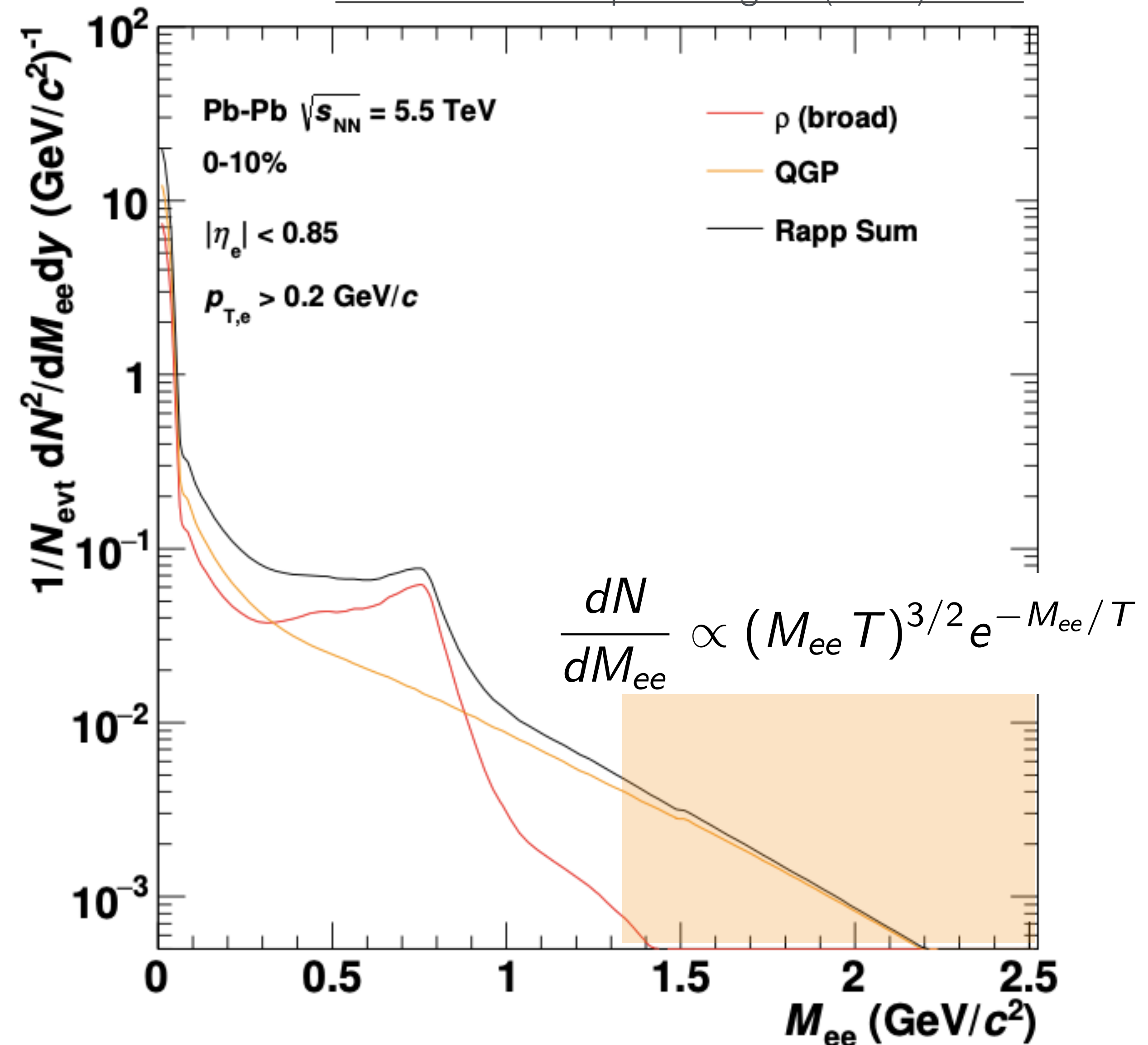
Mass region $1.5 < M_{ee} < 2.5 \text{ GeV}$ sensitive to effective QGP temperature T (not affected by blueshift)

NA60: $T = 205 \pm 12 \text{ MeV}$

In-In, fit range 1.2–2.0 GeV, [AIP Conf.Proc. 1322 \(2010\)](#)

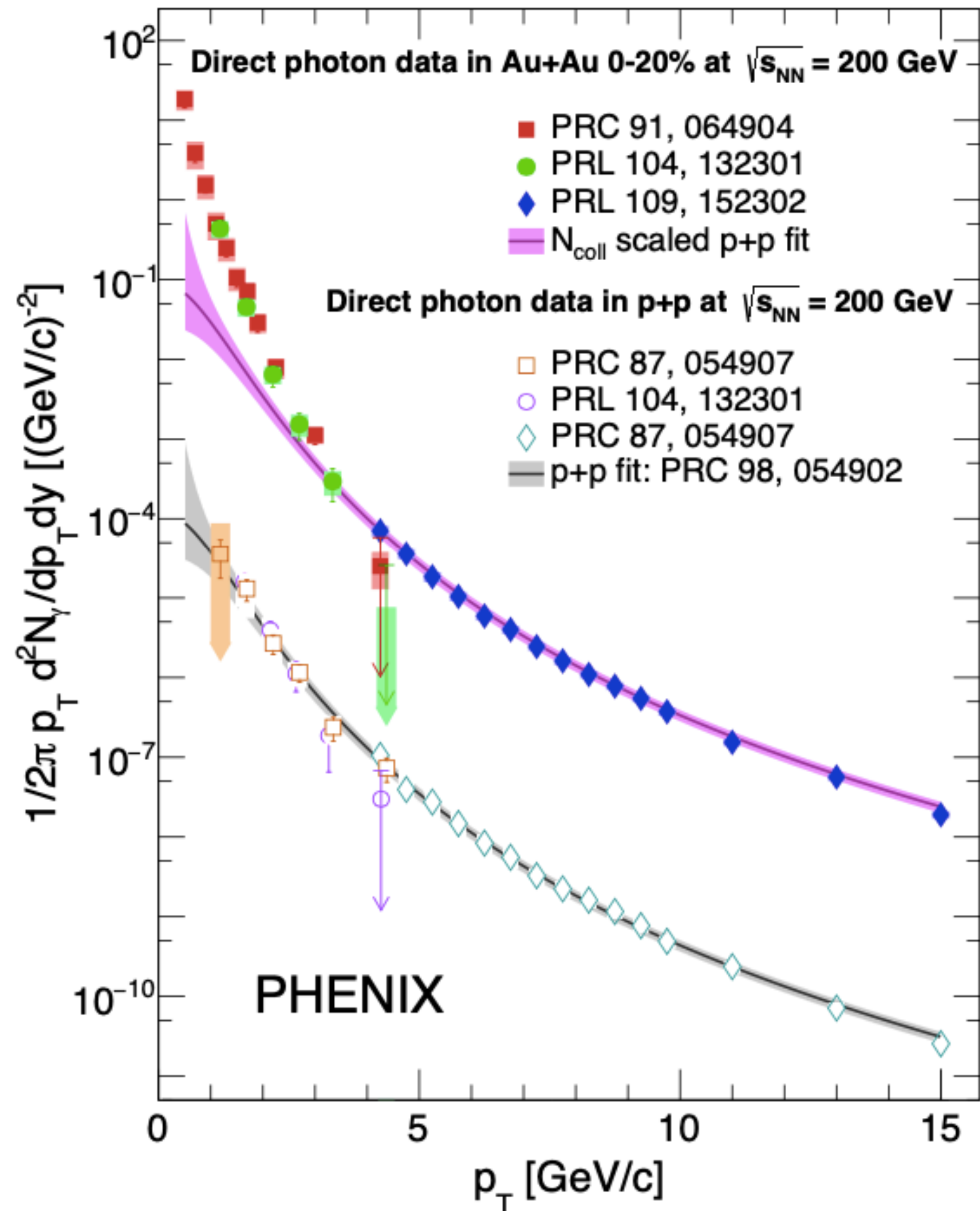
Preequilibrium contribution could affect temperature extraction in $1.5 \lesssim M_{ee} \lesssim 2.5 \text{ GeV}$

CERN Yellow Rep.Monogr. 7 (2019) 1159



Low- p_T direct photons

Photon excess above pp/pQCD at low p_T



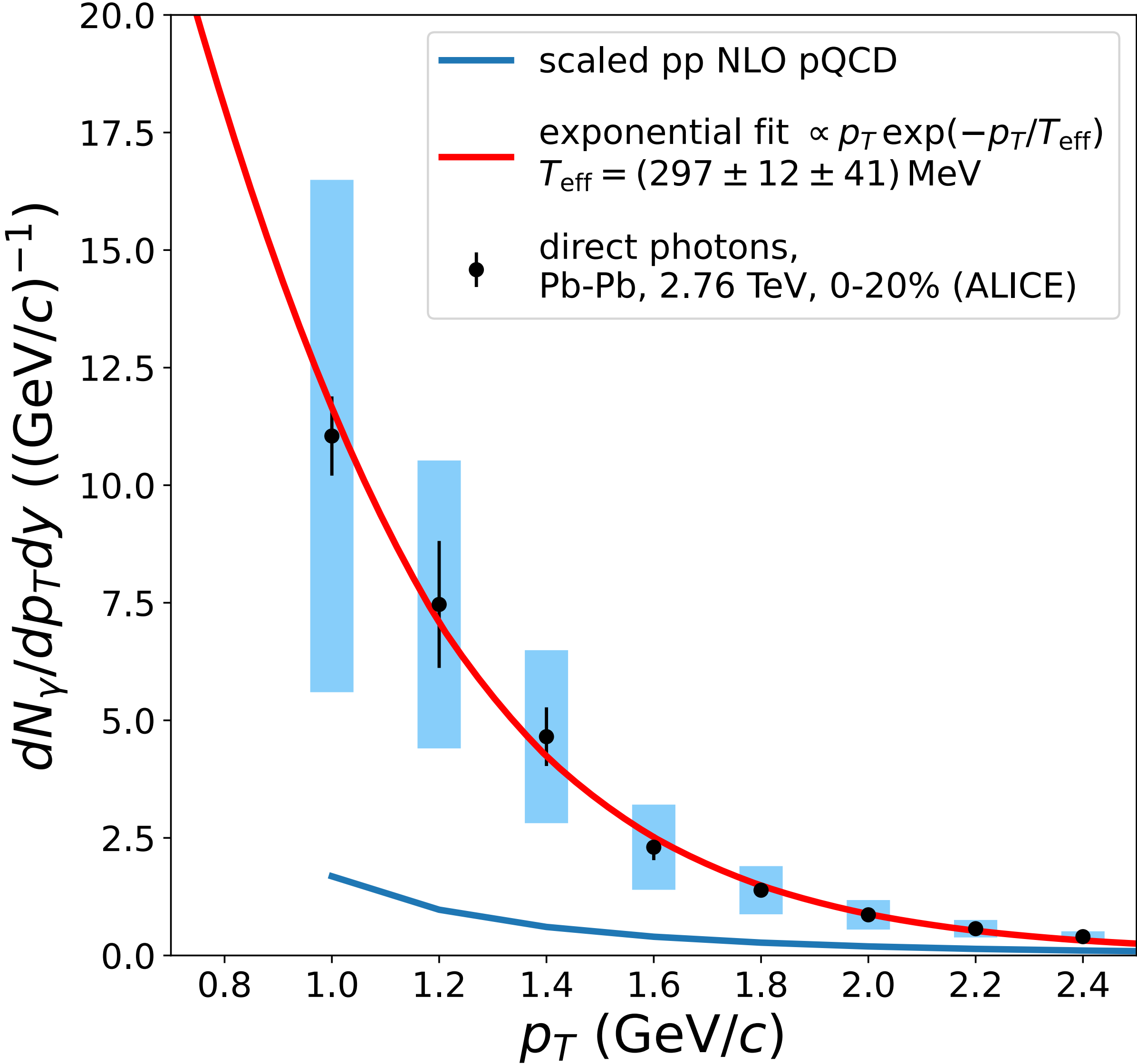
Various direct photon measurements in different systems from PHENIX

Excess above pQCD / scaled pp reference for $p_T < 4$ GeV/c

Excess above pQCD has exponential shape

$T_{\text{eff}} \sim (240 \pm 25)$ MeV in central Au-Au at 200 GeV ($0.6 < p_T < 2$ GeV/c)

Same trend as at RHIC, relatively large uncertainties



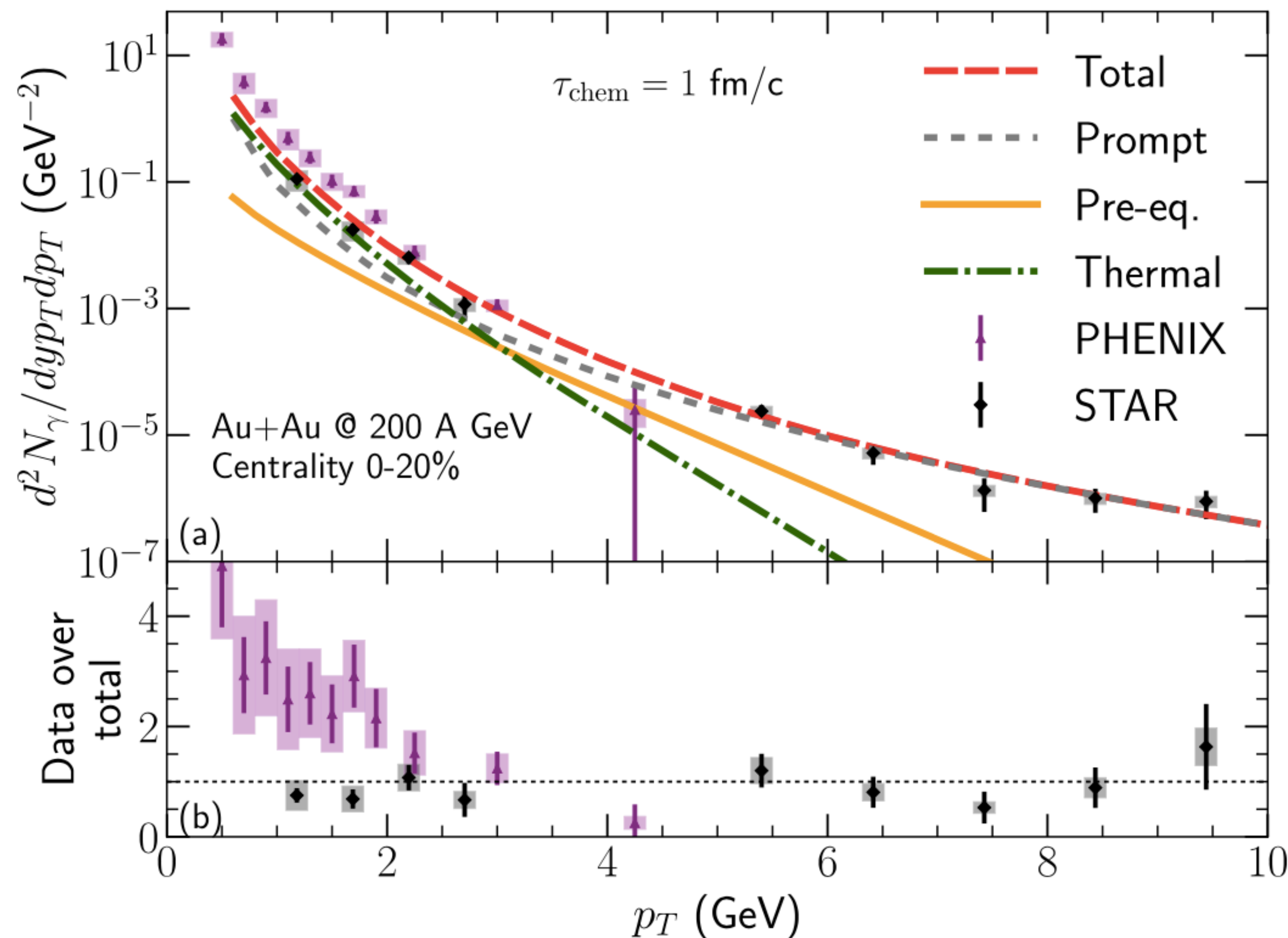
$T_{\text{eff}} \sim (297 \pm 12^{\text{stat}} \pm 41^{\text{syst}}) \text{ MeV}$
($0.9 < p_T < 2.1 \text{ GeV}/c$, pQCD subtracted)

γ_{dir} significance $< 3\sigma$ (a start ...)

Smaller uncertainties highly desirable

ALICE, *Phys. Lett. B* 754 (2016) 235

Part I: Low- p_T yields not described by state-of-the-art models



Yield for $p_T < 2$ GeV/c factor 2–3 above model predictions

Discrepancy between PHENIX and STAR data

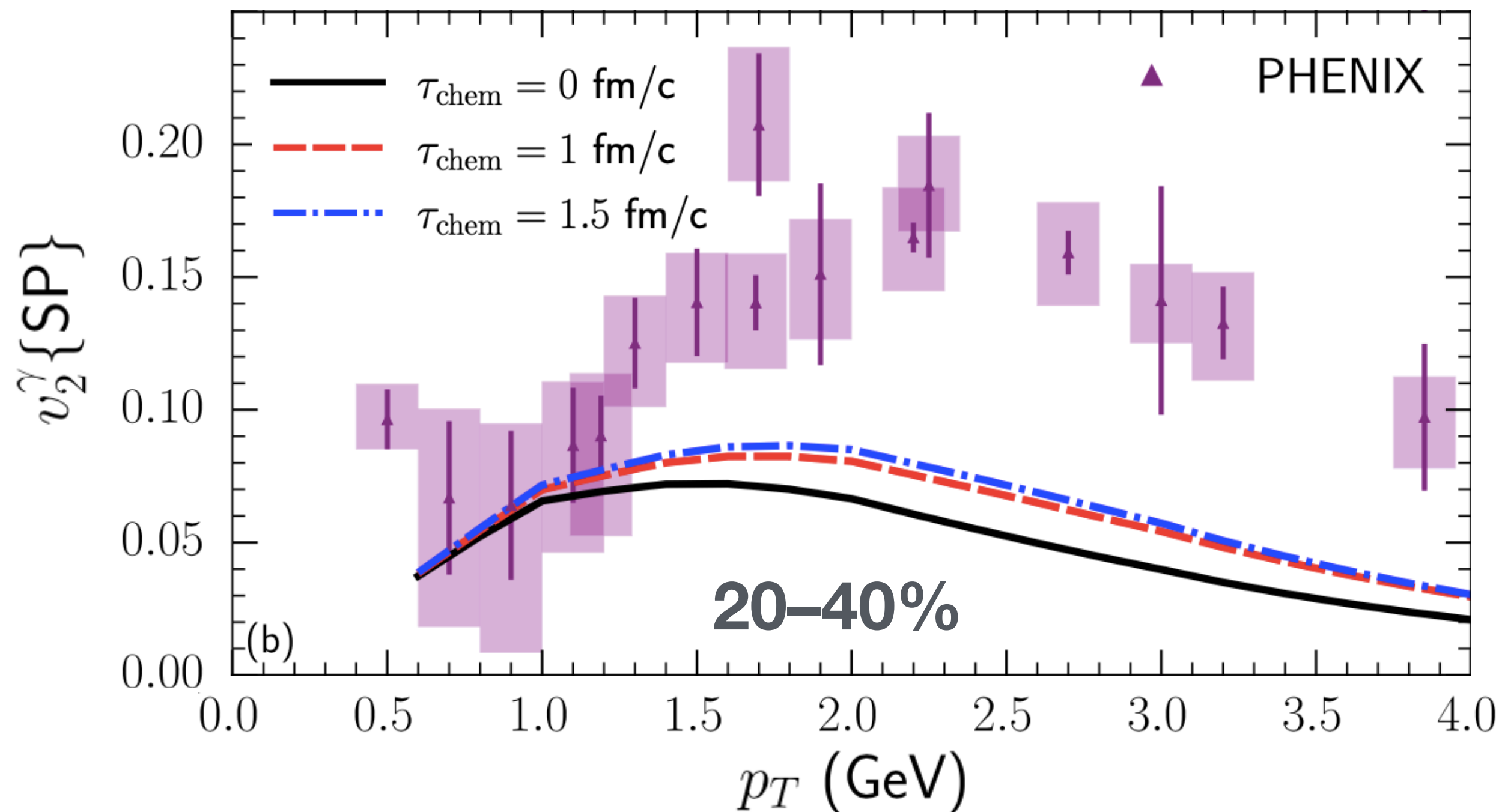
Yield at the LHC (ALICE, 2.76 TeV):

- ▶ similar trend as PHENIX
- ▶ but no puzzle within (large) exp. uncertainties

C. Gale, J-F. Paquet, B. Schenke, C. Shen
 Phys.Rev.C 105 (2022) 1, 014909

Part II: Direct photon $v_2 \approx$ pion v_2 , not described by models

C. Gale, J-F. Paquet, B. Schenke, C. Shen
Phys.Rev.C 105 (2022) 1, 014909



RHIC (PHENIX):

Large $v_{2,\text{dir}} (\approx v_{2,\pi})$ not reproduced by state-of-the-art hydro models

LHC (ALICE):

$v_{2,\text{dir}} \approx v_{2,\pi}$, but no puzzle within (large) exp. uncertainties

Direct photon production dominated by late stage (cross over, hadron gas)?

Particularly challenging:

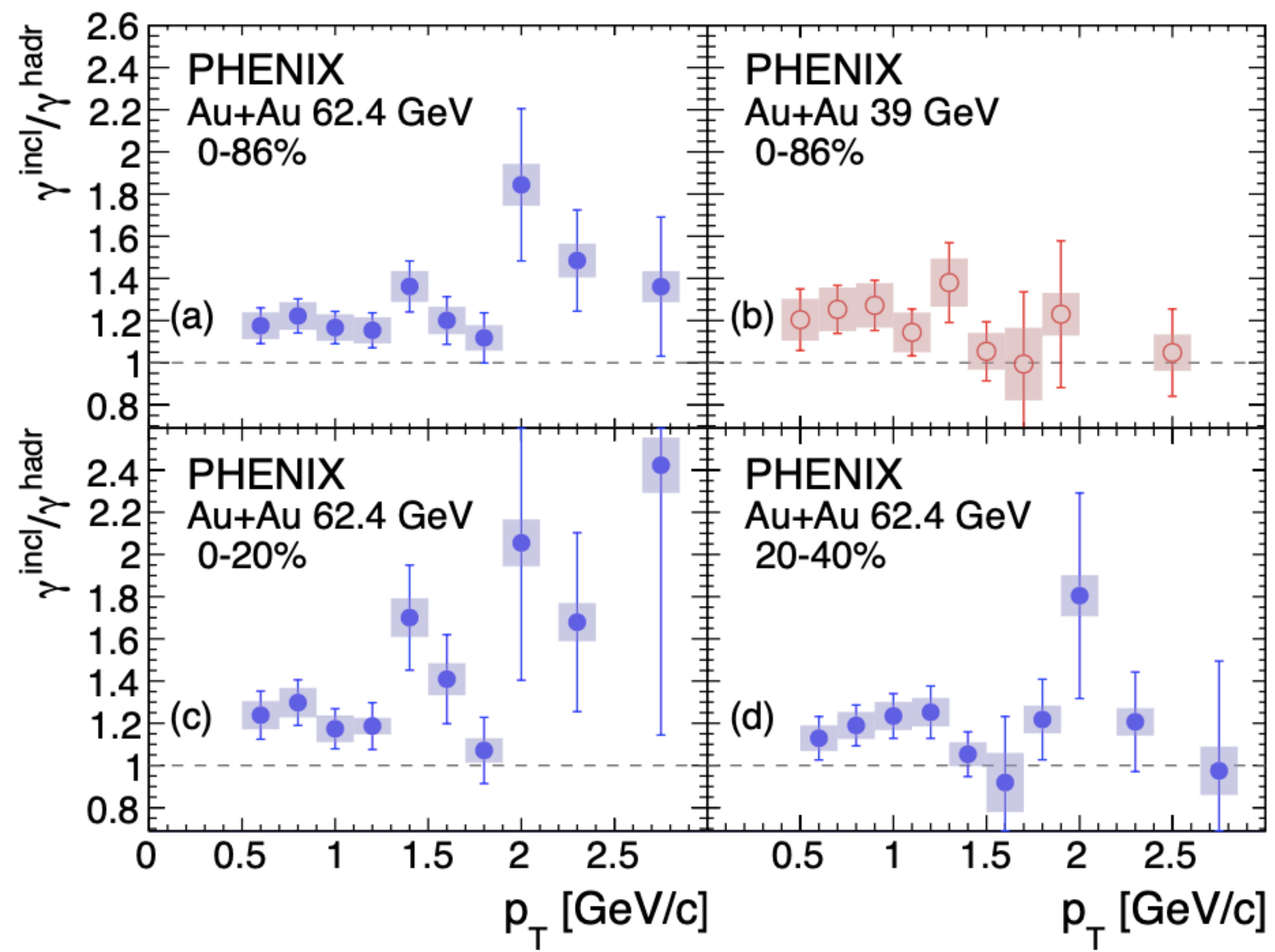
Simultaneously description of yield and v_2



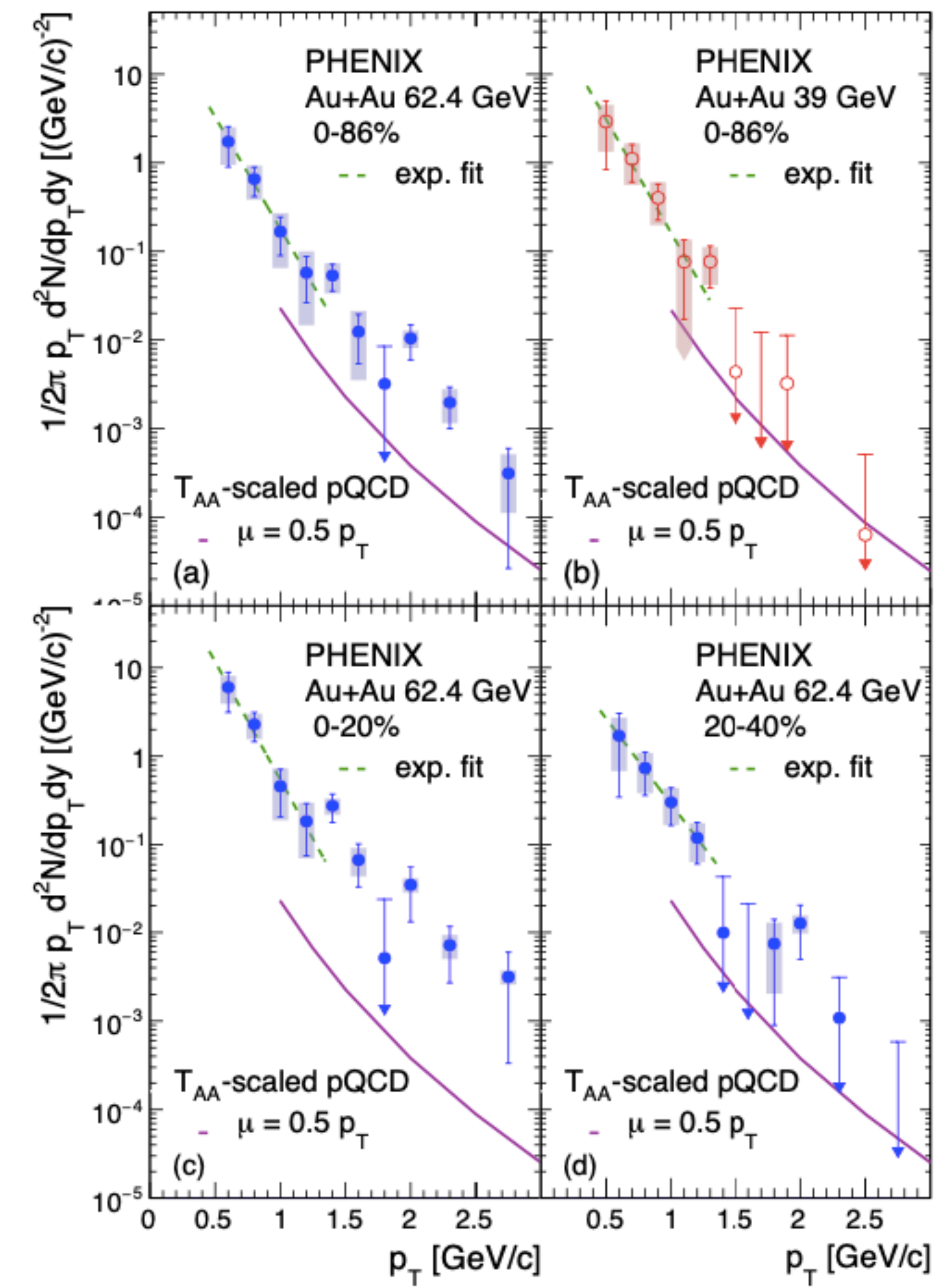
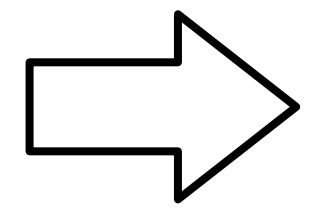
“Direct photon puzzle” (2020, postponed)
EMMI Rapid Reaction Task Force (RRTF)
A. Marin, J. Stachel, K.R.

New PHENIX direct-photon data (1): Au-Au at 39 and 62.4 GeV

~20% direct photon excess above decay photon background

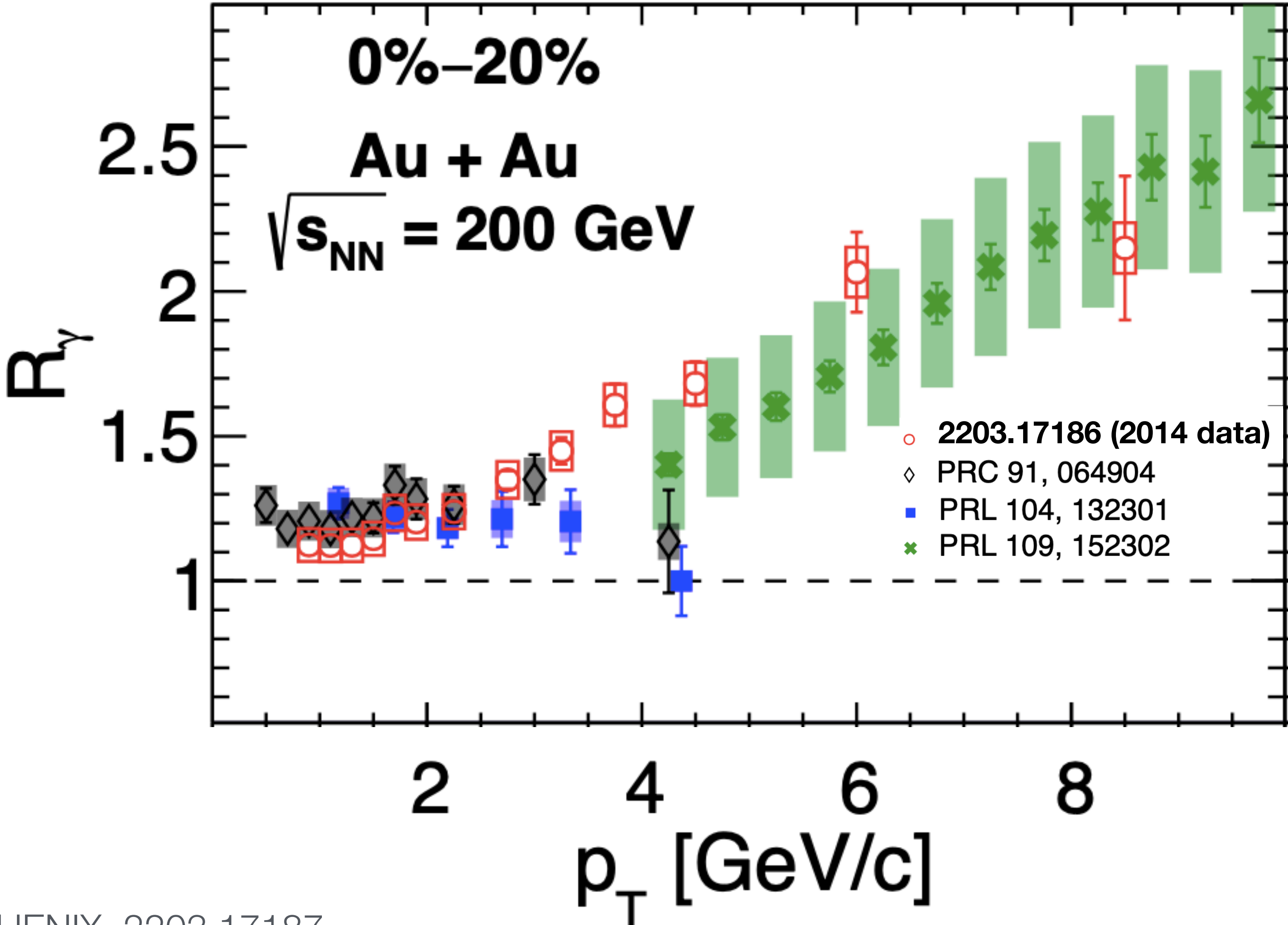


photon conversion method



Different methods and results for different datasets agree

$$R_\gamma = \gamma_{\text{incl}}/\gamma_{\text{decay}} = 1 + \gamma_{\text{direct}}/\gamma_{\text{decay}}$$



Methods

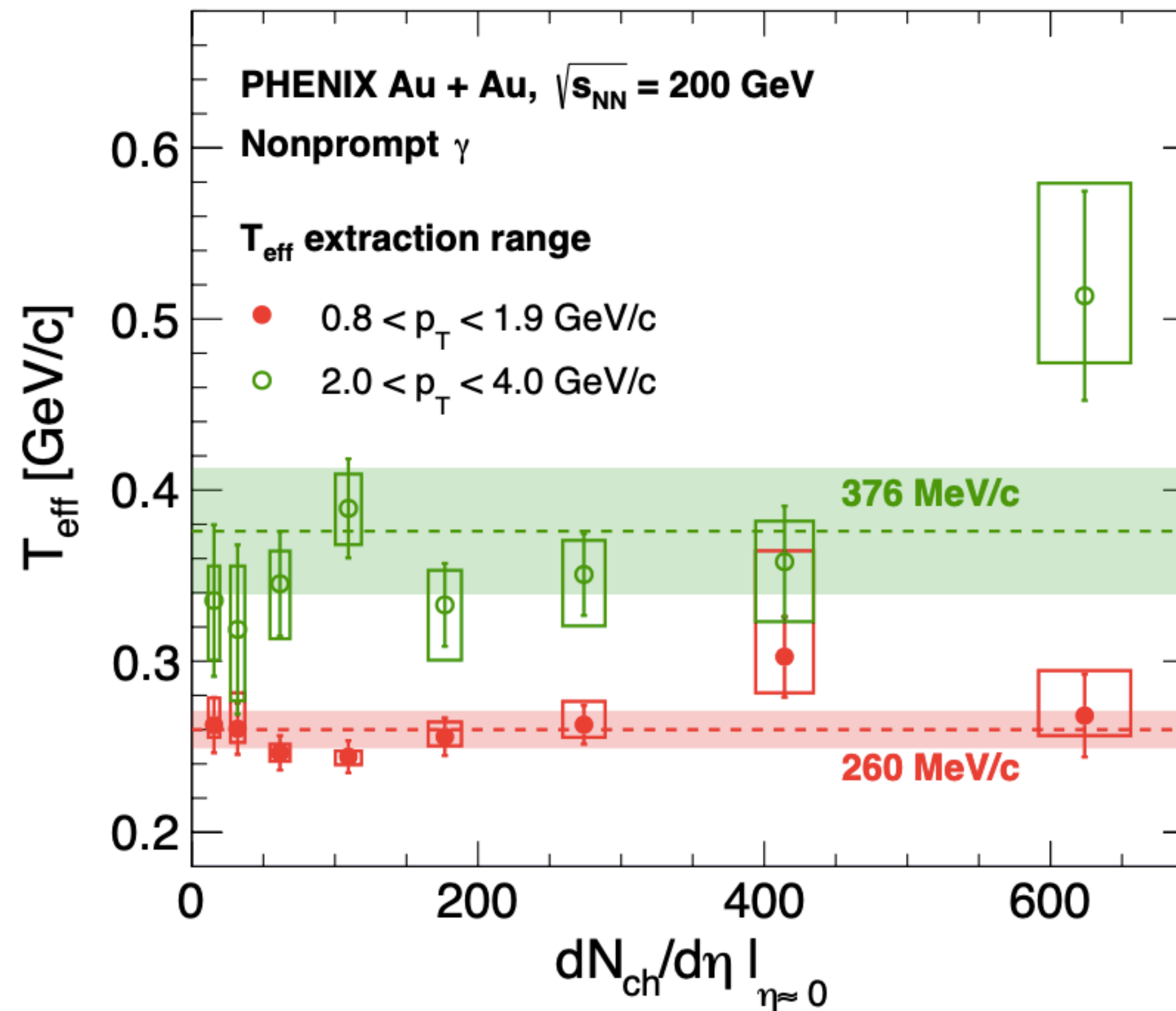
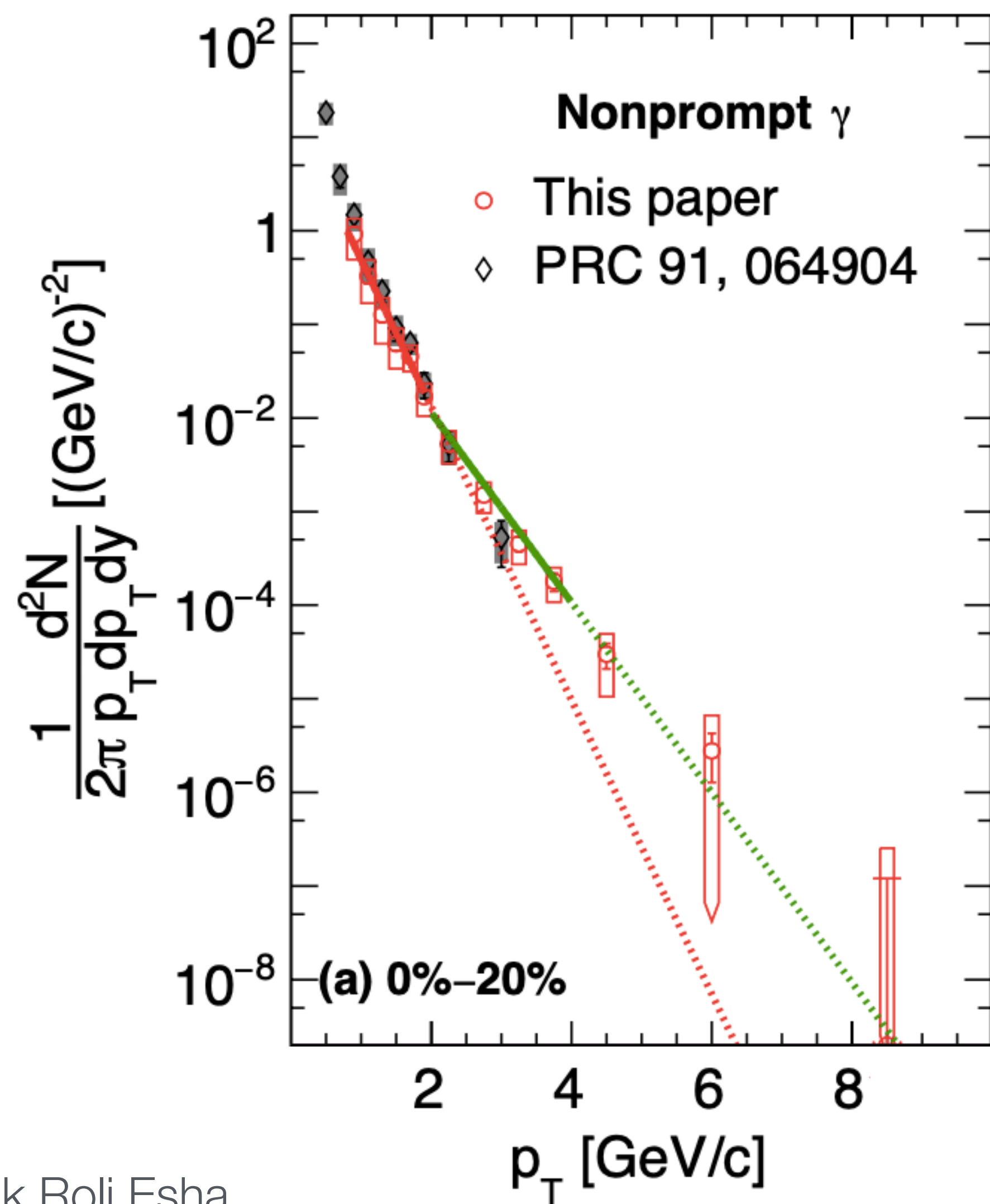
- ▶ calorimeter'
- ▶ virtual photon method
- ▶ conversions

2014 data

- ▶ conversion method
- ▶ x10 more statistics

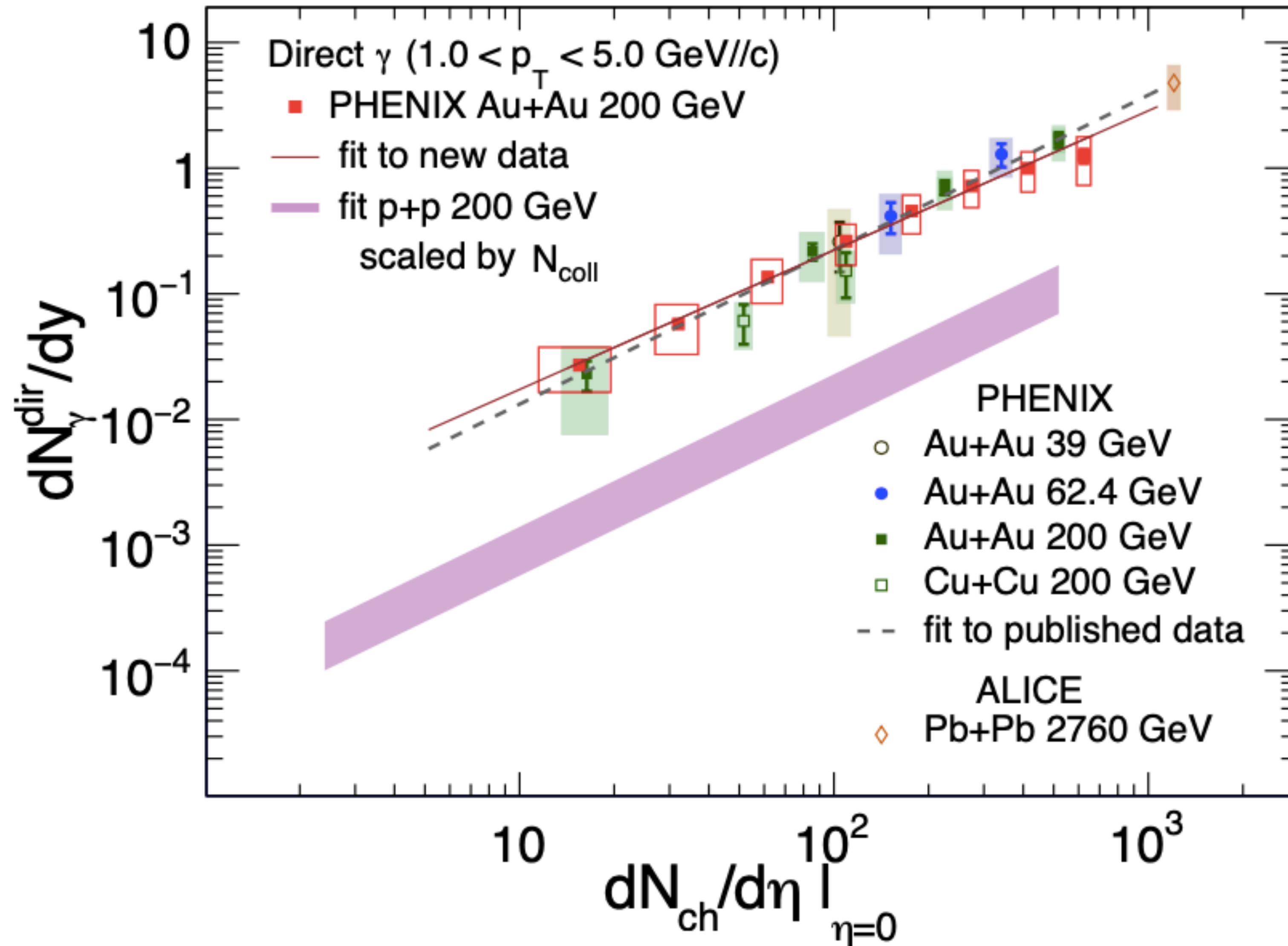
T_{eff} from fit with exponential

T_{eff} increases with p_T : in line with a larger early-time contribution at higher p_T



“nonprompt” = direct – $N_{\text{coll}} \times \text{pp}$

Universal scaling of $dN_{\gamma,dir}/dy$ with $dN_{ch}/d\eta$



$$\frac{dN_{\gamma}}{dy} = A \left(\frac{dN_{ch}}{d\eta} \right)^{\alpha}$$

$$\alpha = 1.11 \pm 0.02 \text{ (stat)}_{-0.08}^{+0.09} \text{ (syst)}$$

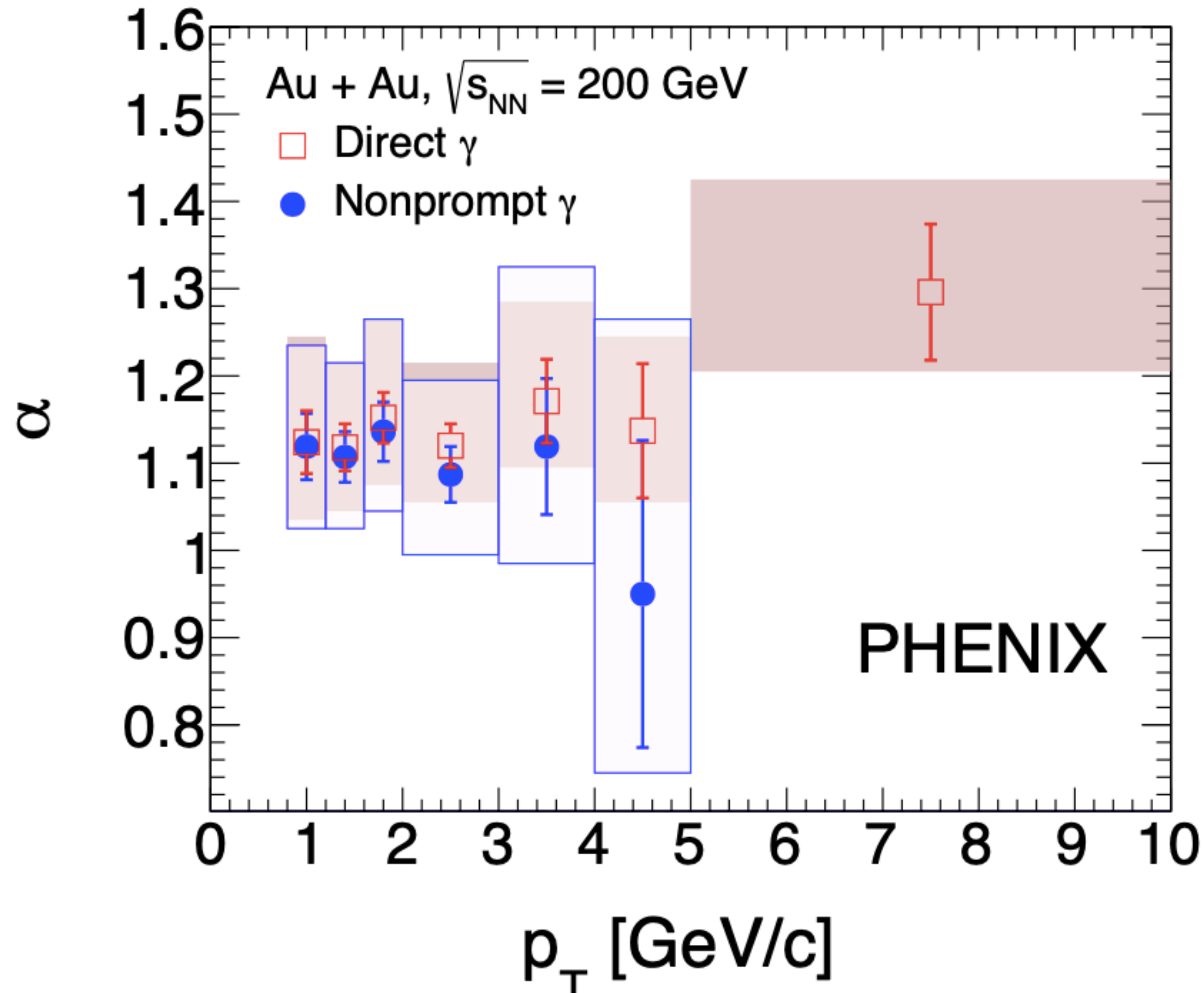
Hydro prediction for thermal photons:

$$\alpha \approx 1.6$$

Shen, Heinz, Paquet, Gale, [PRC 89, 044910 \(2014\)](#)

Constraining production mechanisms by studying scaling patterns (2)

$\alpha(p_T) \approx \text{constant}$: Expected p_T dependence of α not observed

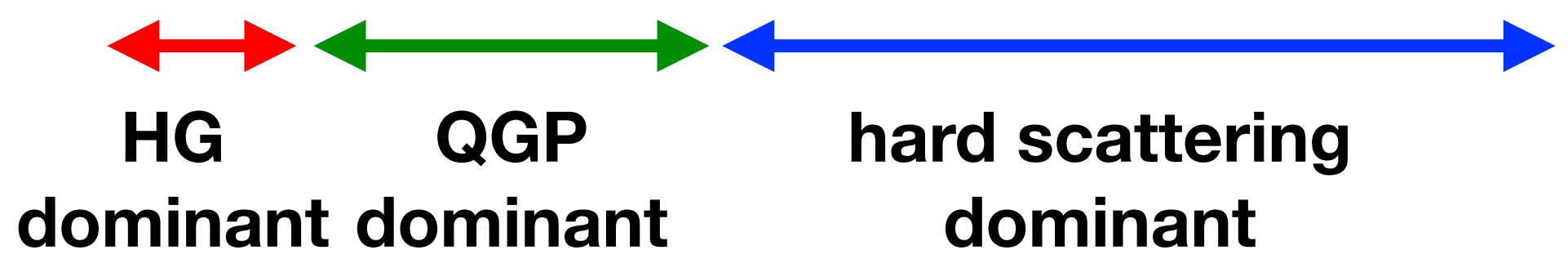


$$\frac{dN_\gamma}{dy} = A(p_T) \left(\frac{dN_{ch}}{d\eta} \right)^{\alpha(p_T)}$$

photon sources: Shen, Heinz, Paquet, Gale, PRC 89, 044910 (2014)

HG	QGP	pQCD
$\alpha_{HG} \approx 1.23$	$\alpha_{QGP} \approx 1.83$	$\alpha_{pQCD} \approx 1.25$

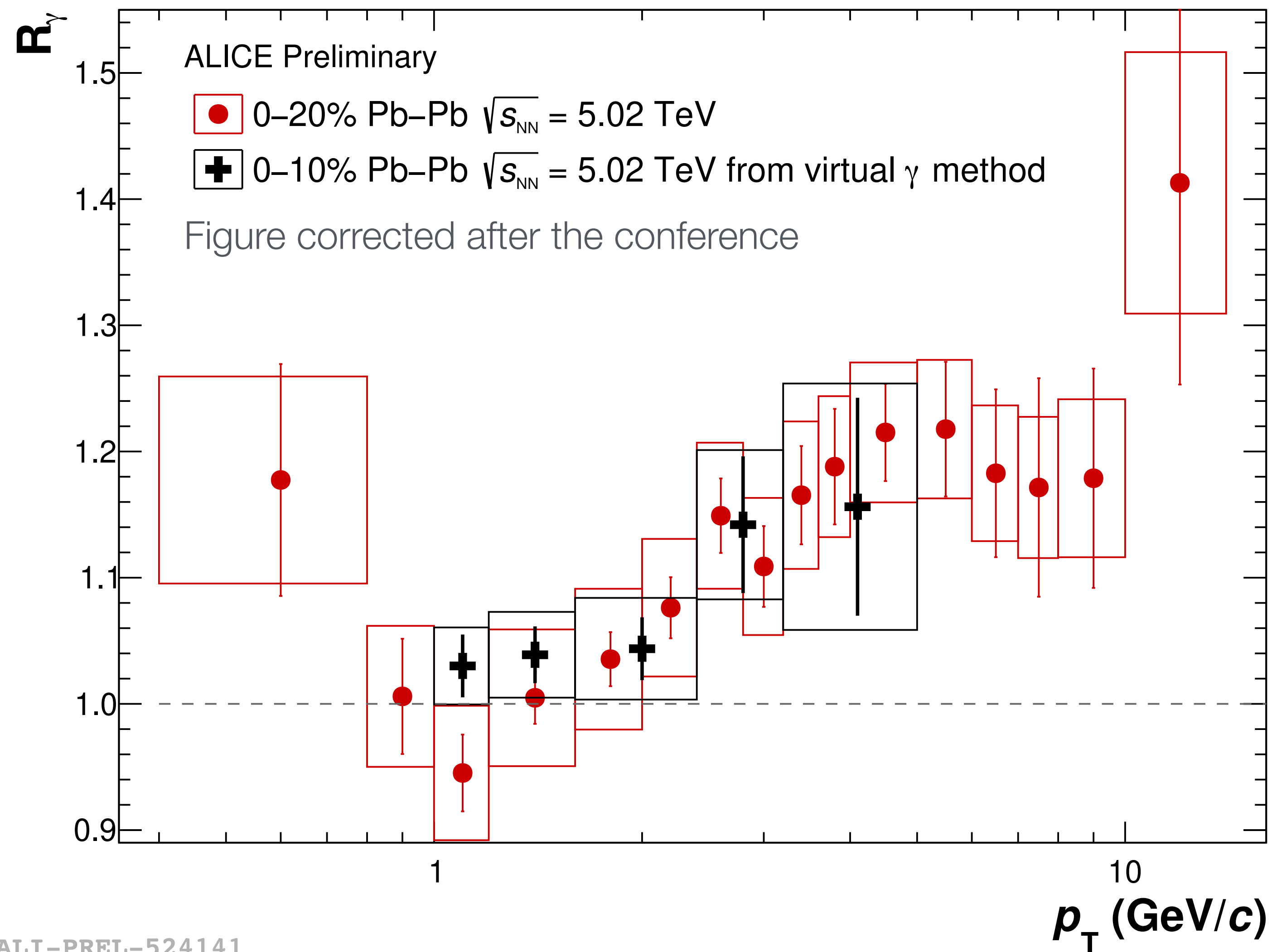
relative contribution varies with p_T



poster Wenqing Fan
PHENIX, 2203.17187

Conversion method and virtual photon method agree

$$R_\gamma = \gamma_{\text{incl}}/\gamma_{\text{decay}} = 1 + \gamma_{\text{direct}}/\gamma_{\text{decay}}$$



Photon conversion method

- ▶ New data-driven approach to “measure” the material budget
- ▶ Material budget uncertainty reduced from **4.5% to 2.5%**

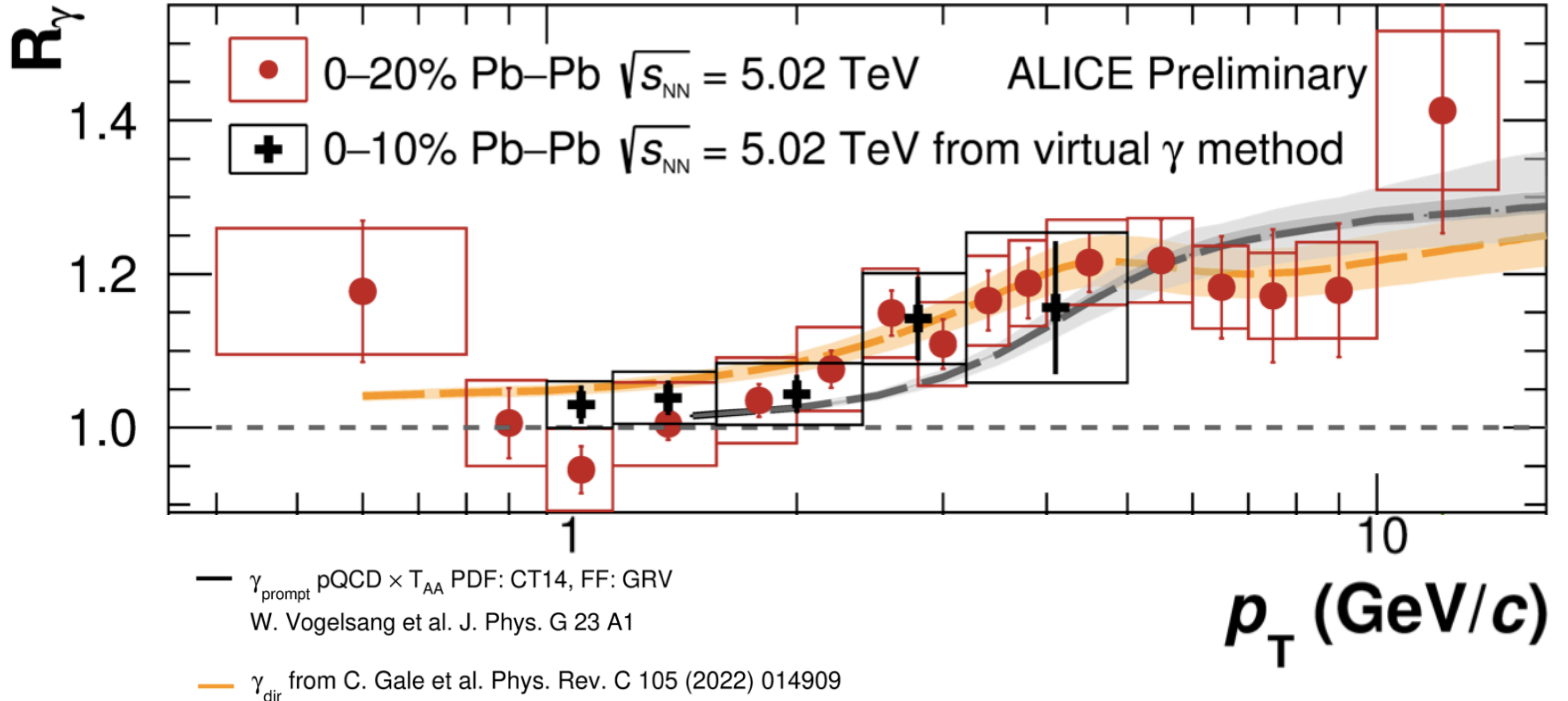
Virtual photon method

- ▶ Dielectron excess measured above pion mass
- ▶ Extrapolated to $m = 0$ with Kroll-Wada formula

Both pQCD and pQCD+PE+thermal consistent with the data

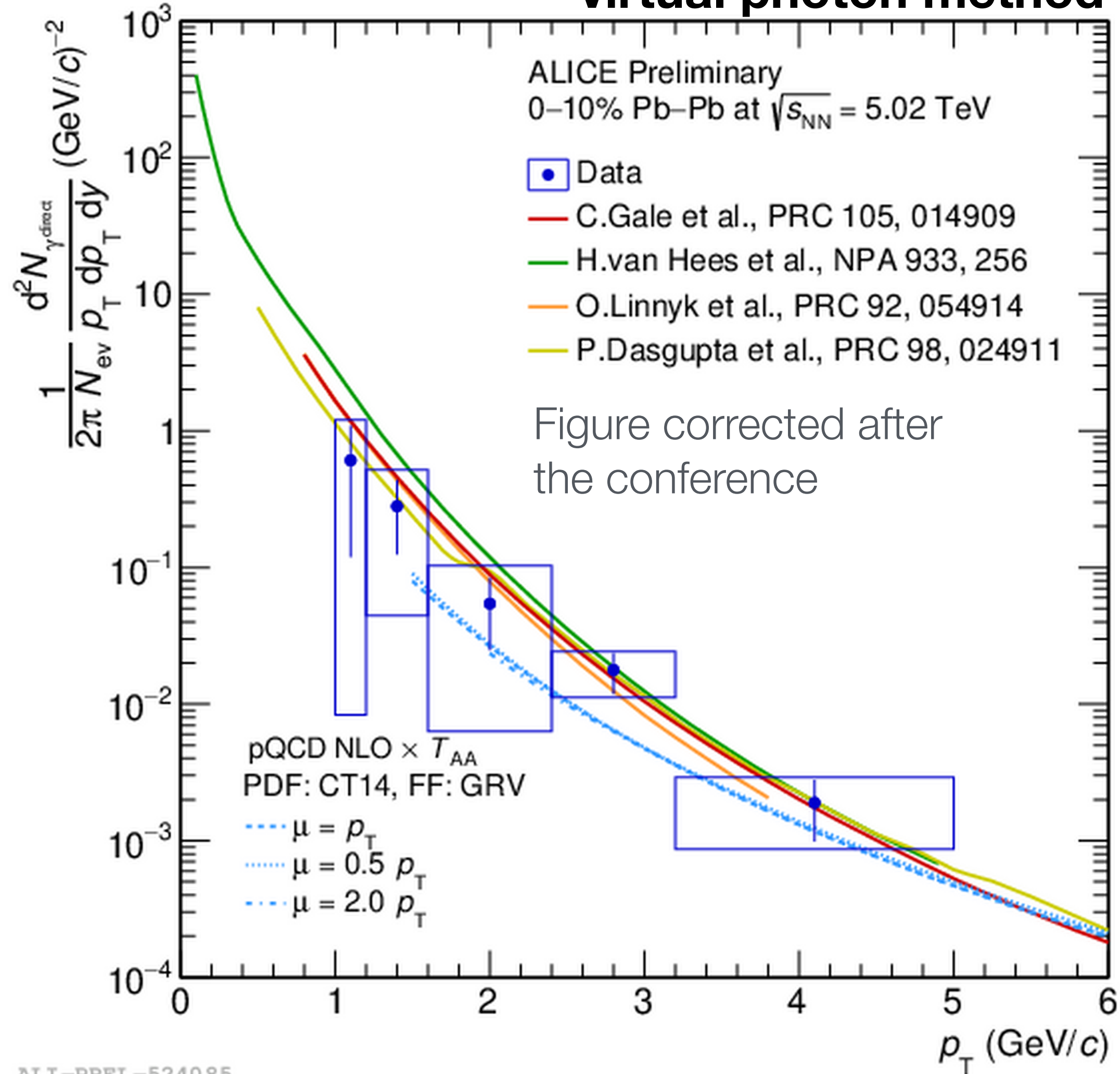
$$R_\gamma^{\text{theory}} = (\gamma_{\text{dir}}^{\text{theory}} + \gamma_{\text{decay}}^{\text{ALICE}}) / \gamma_{\text{decay}}^{\text{ALICE}}$$

Figure corrected after the conference



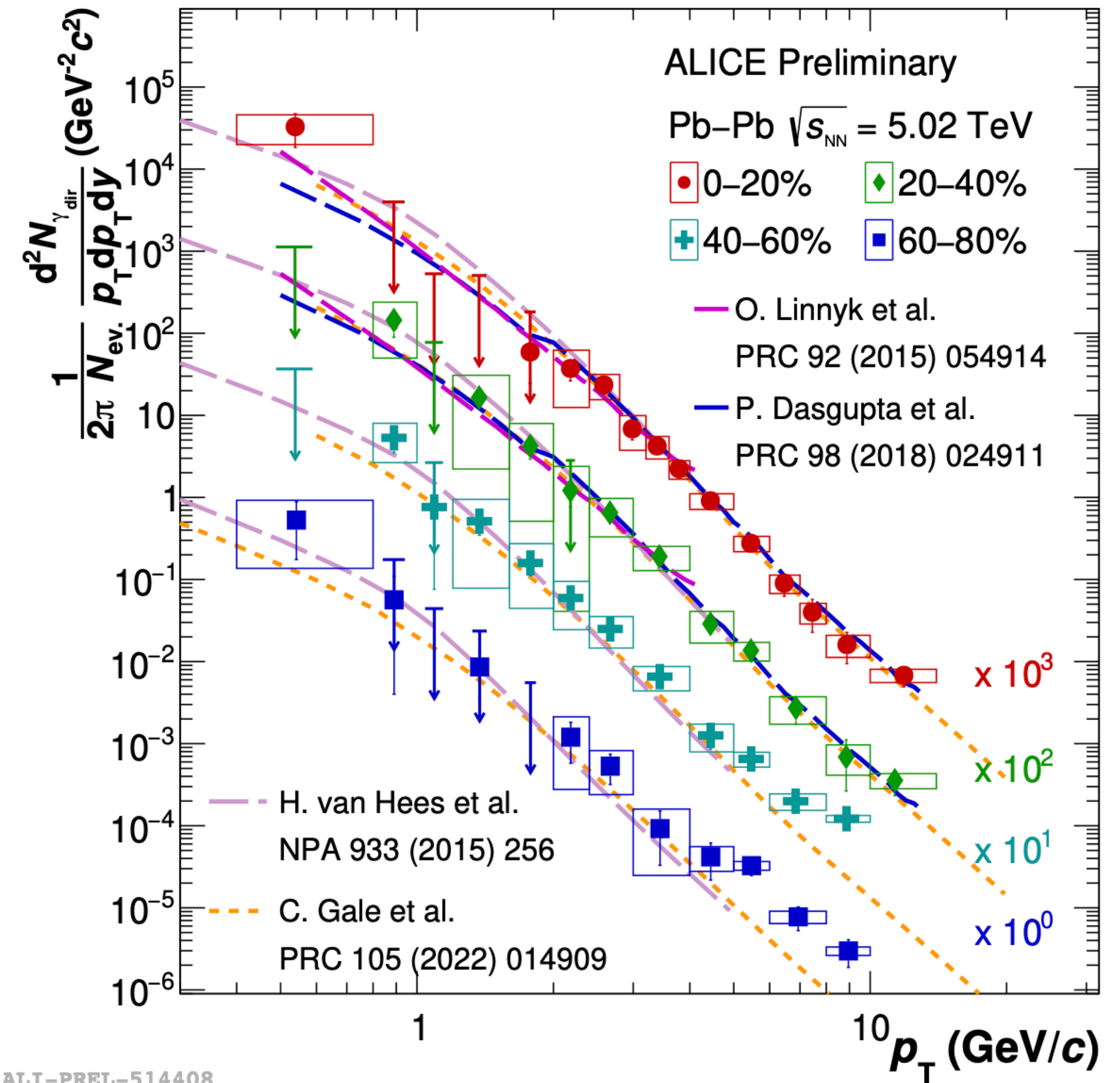
All considered models in fair agreement with the data

virtual photon method



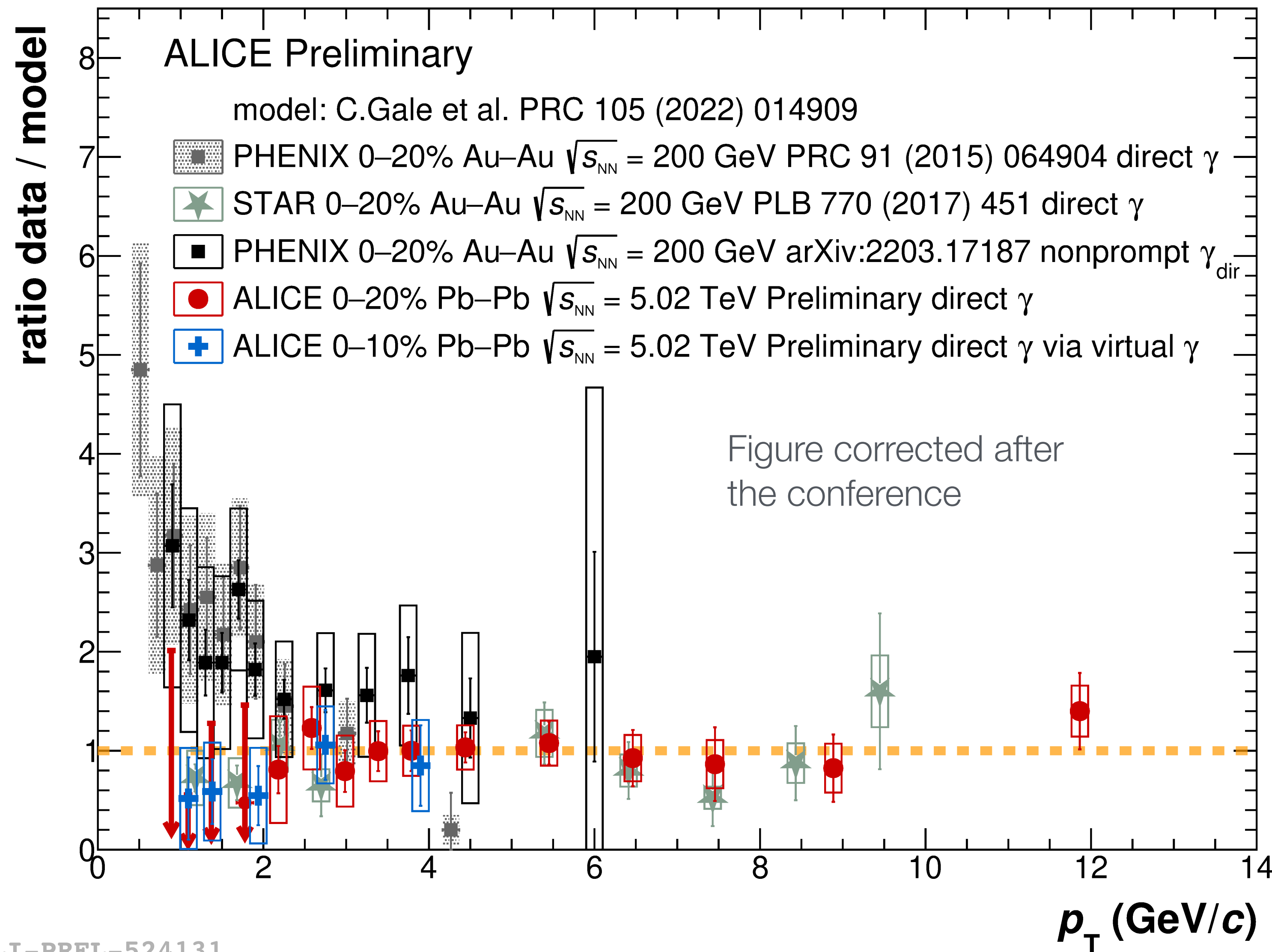
ALI-PREL-524085

conversion method



ALI-PREL-514408

Direct photon puzzle: Not too much of a puzzle left for yields



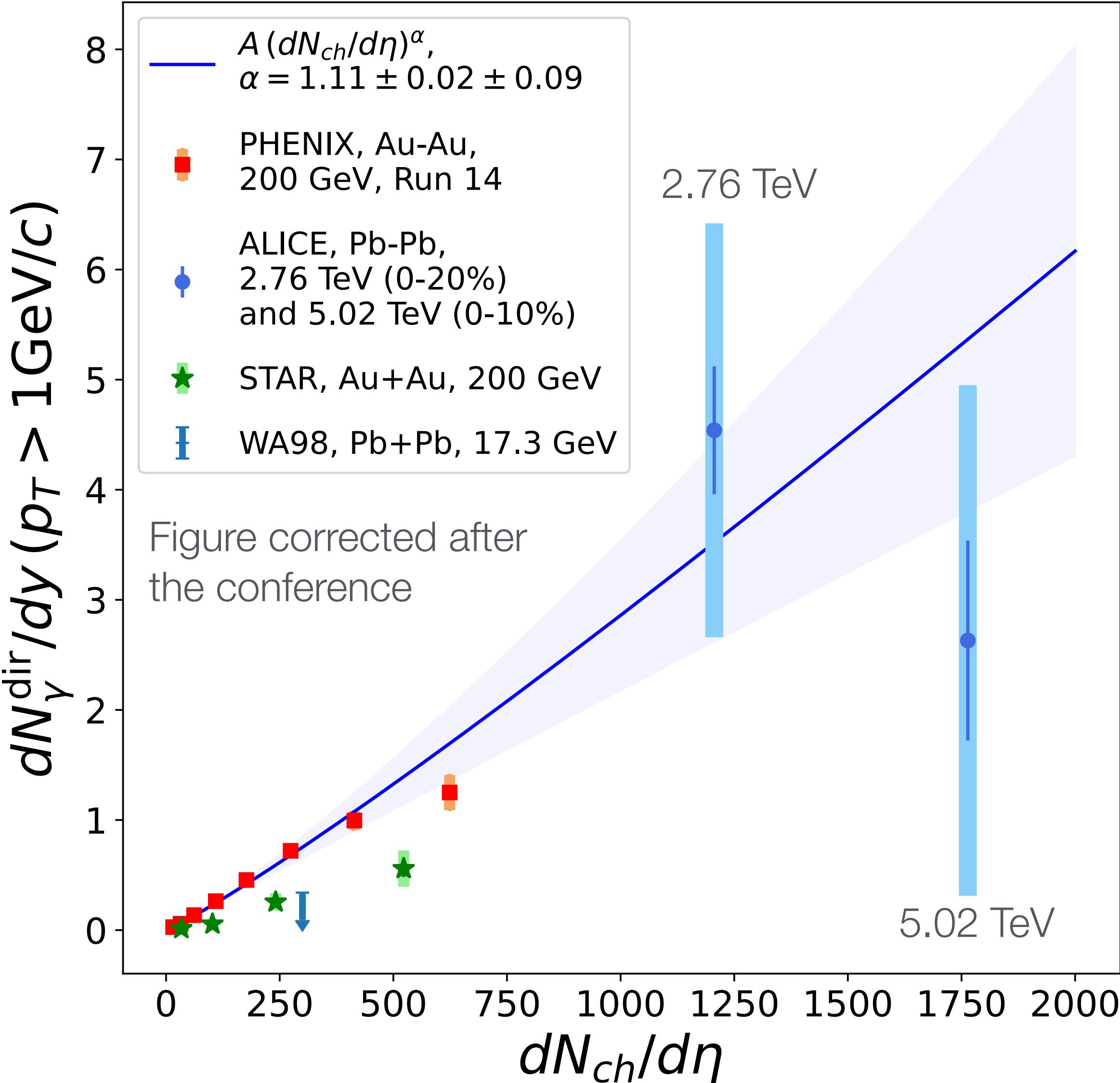
New PHENIX data:
no big puzzle for yields

New ALICE data:
fair agreement with theory

Level of agreement with theory
appears similar for PHENIX and
ALICE virtual photon result

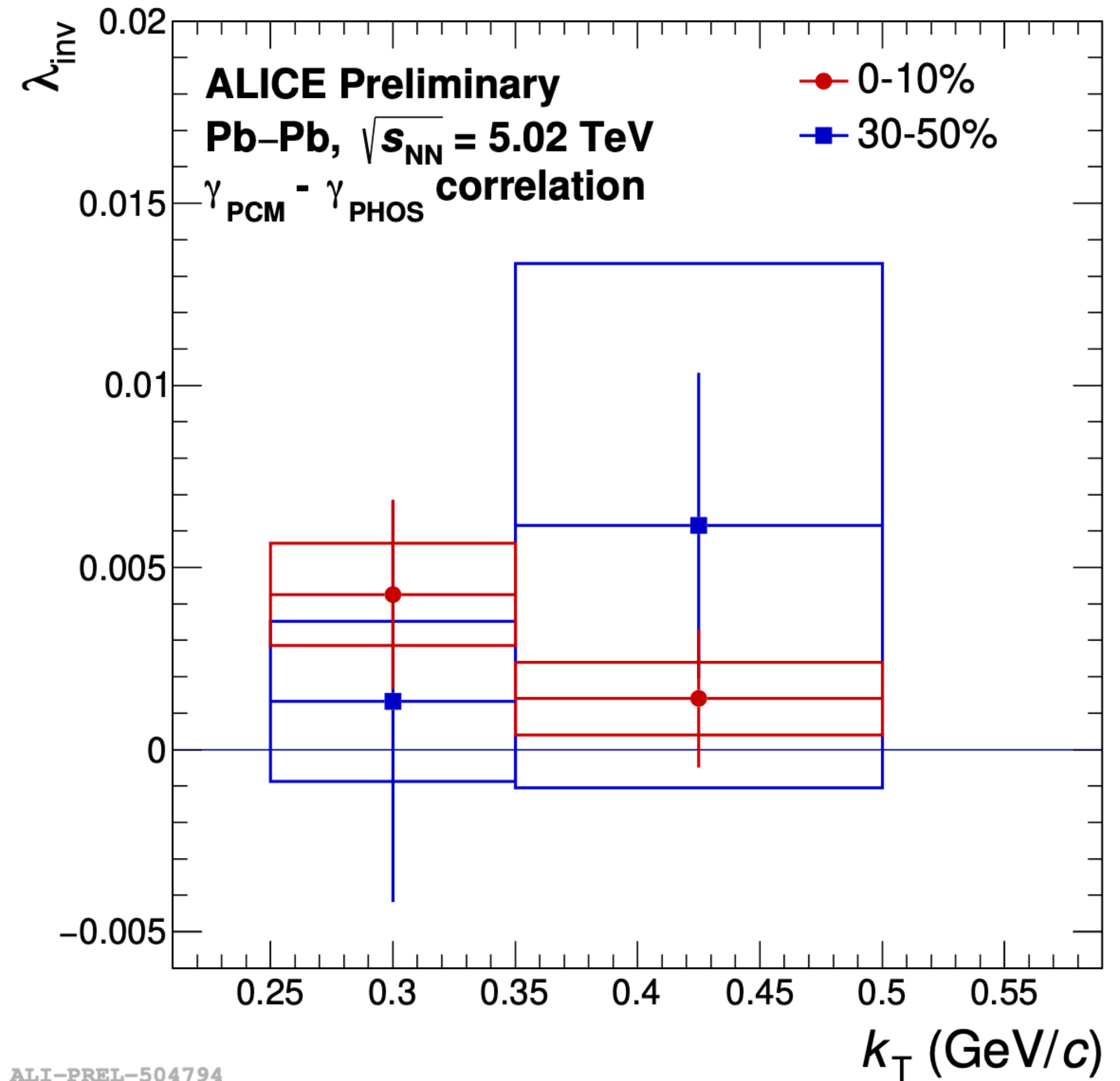
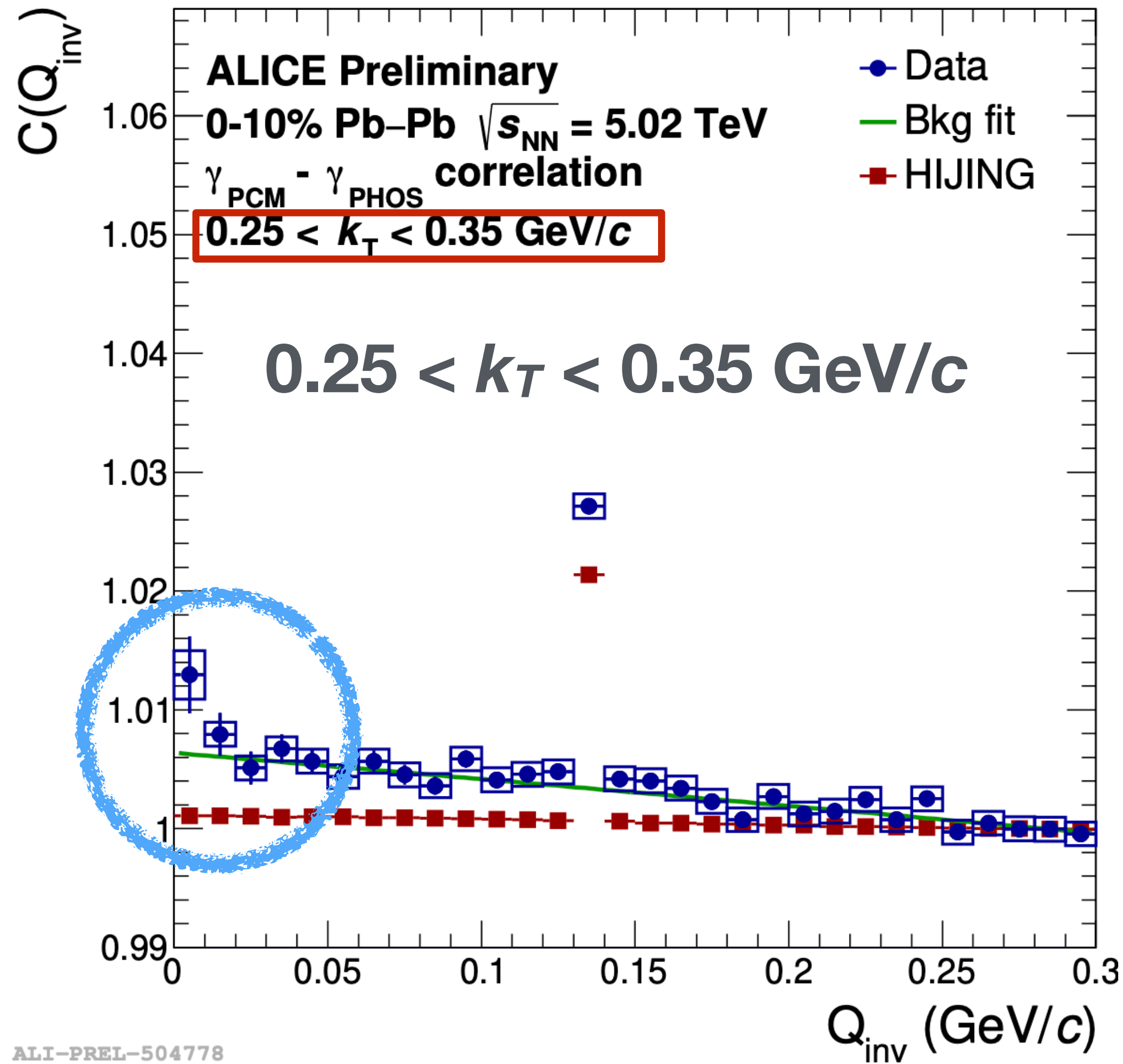
Universal scaling of direct photon yields

LHC direct photon yields consistent with PHENIX scaling



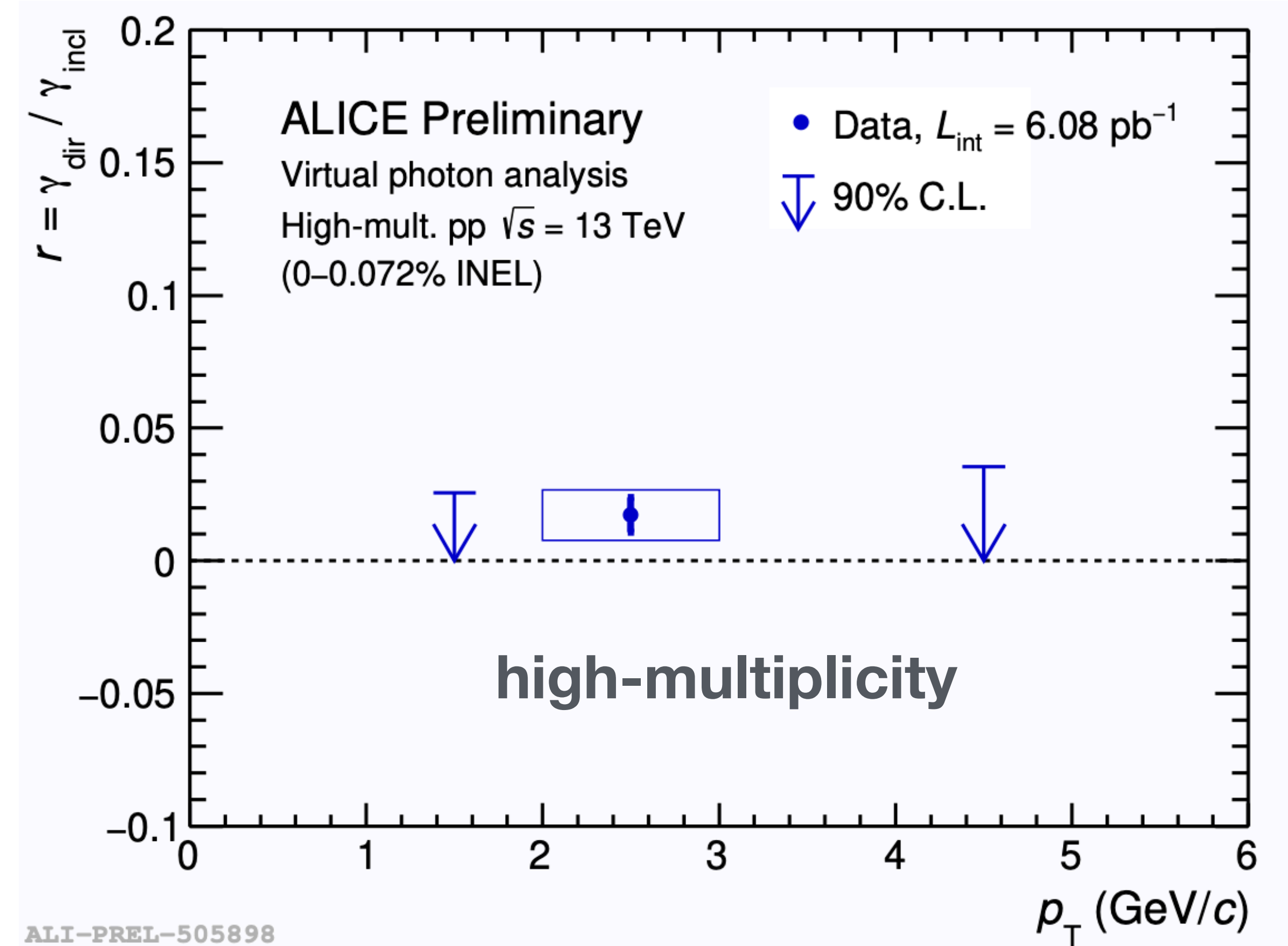
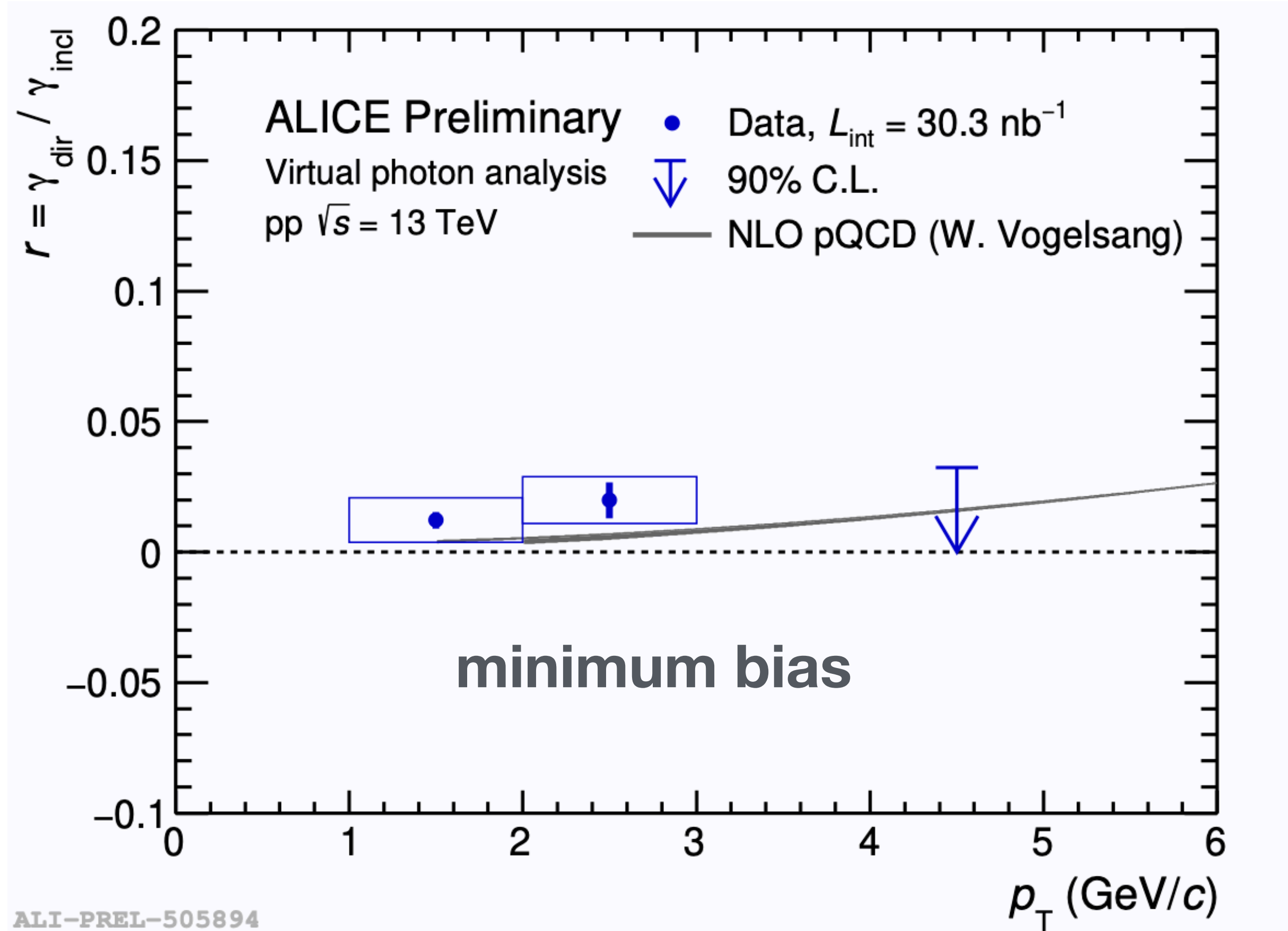
LHC points appear to be consistent with extrapolation of STAR data as well

Hints of a correlation peak, somewhat narrower than expected



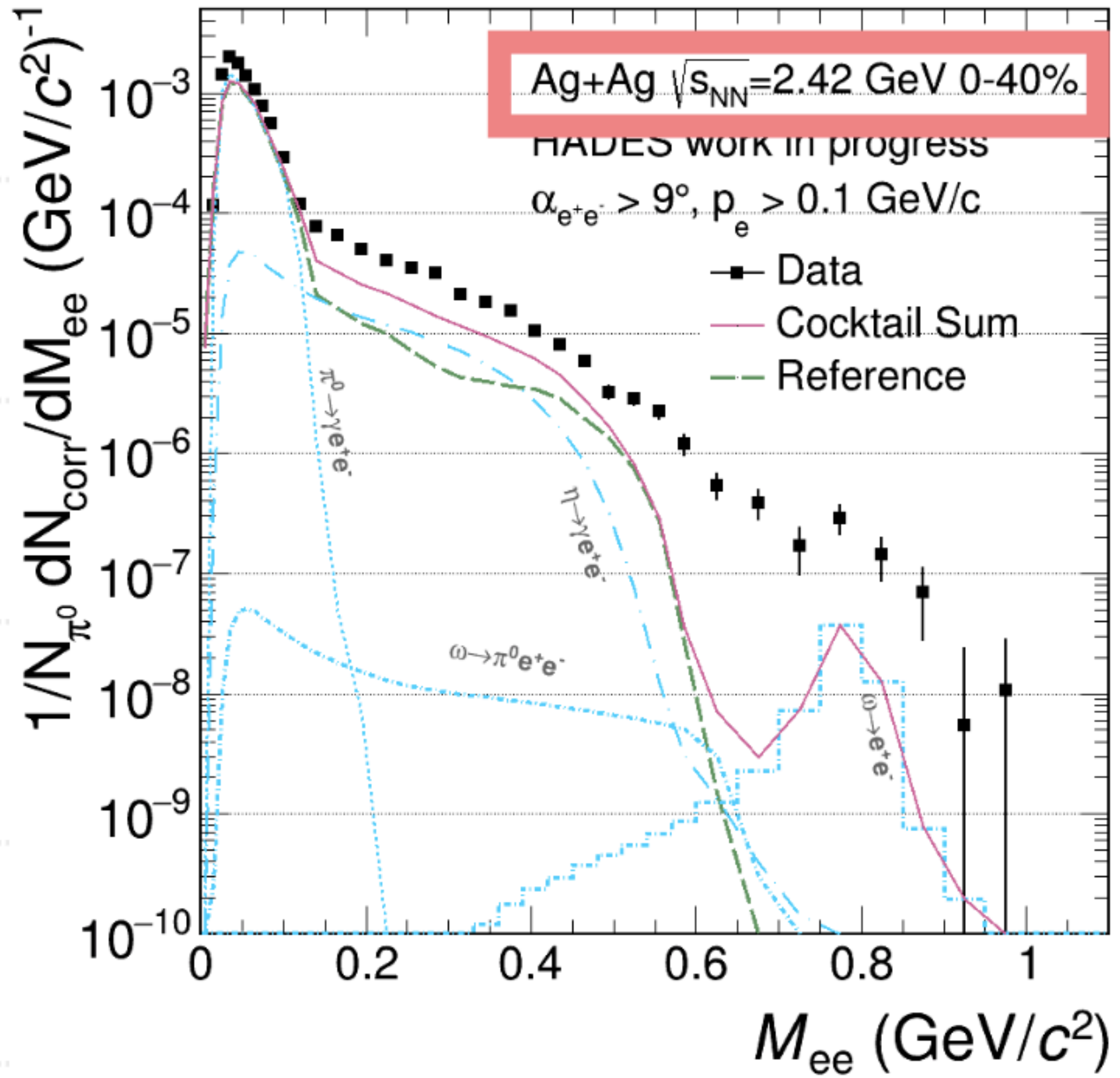
$\lambda_{\text{inv}} \rightarrow$ direct photon yield ... but uncertainties currently large

No indications for thermal photons in pp collisions



Dileptons

Clear dielectron excess above cocktail also in Ag-Ag

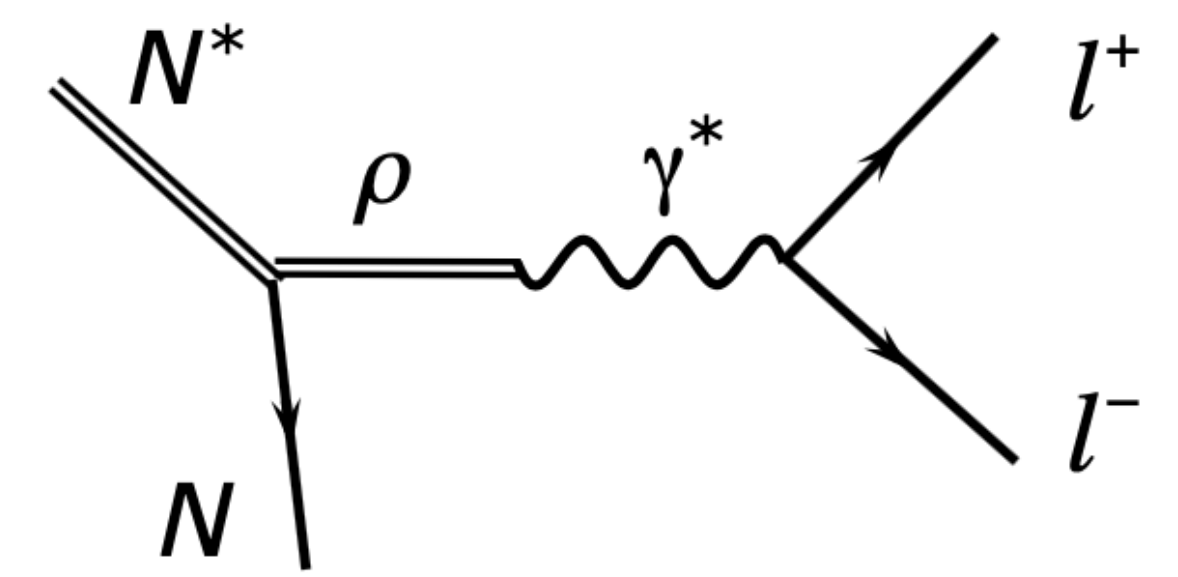


2012 Au-Au at $\sqrt{s_{NN}} = 2.42 \text{ GeV}^*$

2019 Ag-Ag at $\sqrt{s_{NN}} = 2.42 \text{ GeV}$ and 2.55 GeV

^{197}Au : $Z/A = 0.40$, ^{107}Ag : $Z/A = 0.44$

Understanding of e^+e^- production via resonances crucial

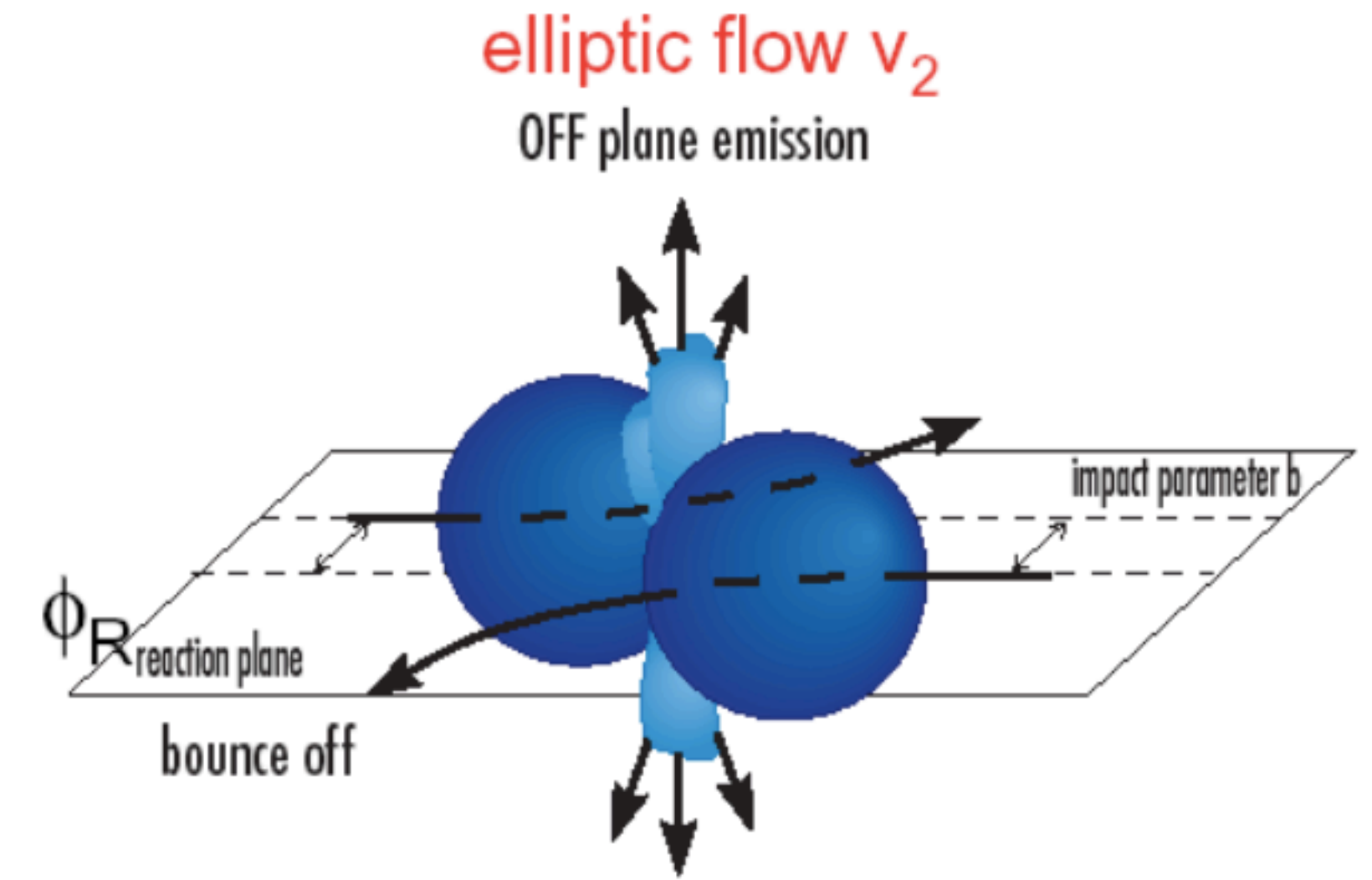
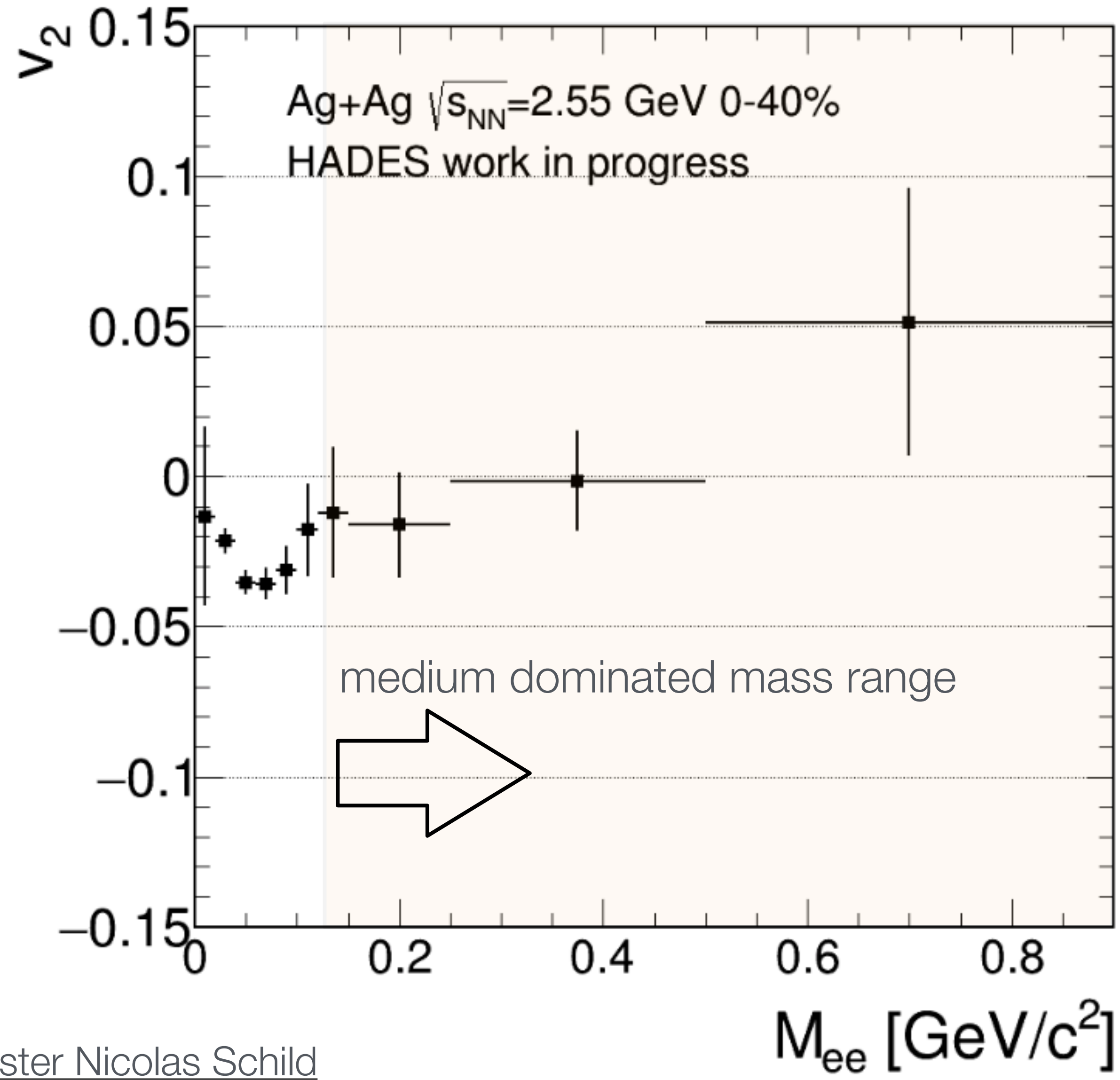


HADES measured effective transition form factor in $\pi^- p \rightarrow n e^+ e^-$. Important role of $N^*(1520)$.

Excess yield for same energy and N_{part} larger in Au-Au than in Ag-Ag. Why?

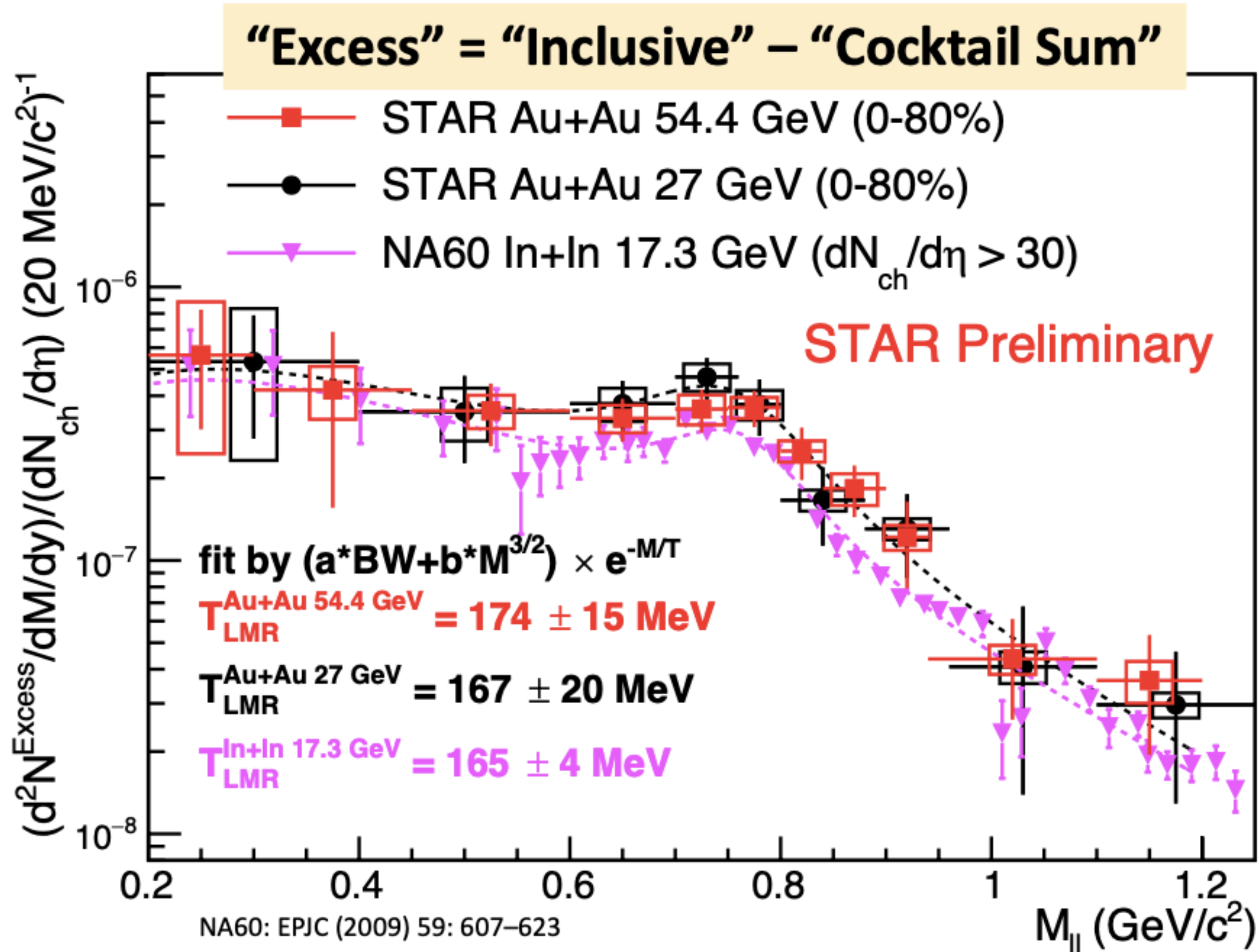
Dielectron elliptic flow consistent with zero in medium dominated mass range

Confirms dileptons as penetrating probes of hot and dense medium



v_2 due to spectator shadowing at SIS energies ("squeeze-out")

Temperature from excess dielectron mass spectrum: 1. LMR



Clear enhancement w.r.t. cocktail in both low mass region (LMR) and intermediate mass region (IMR)

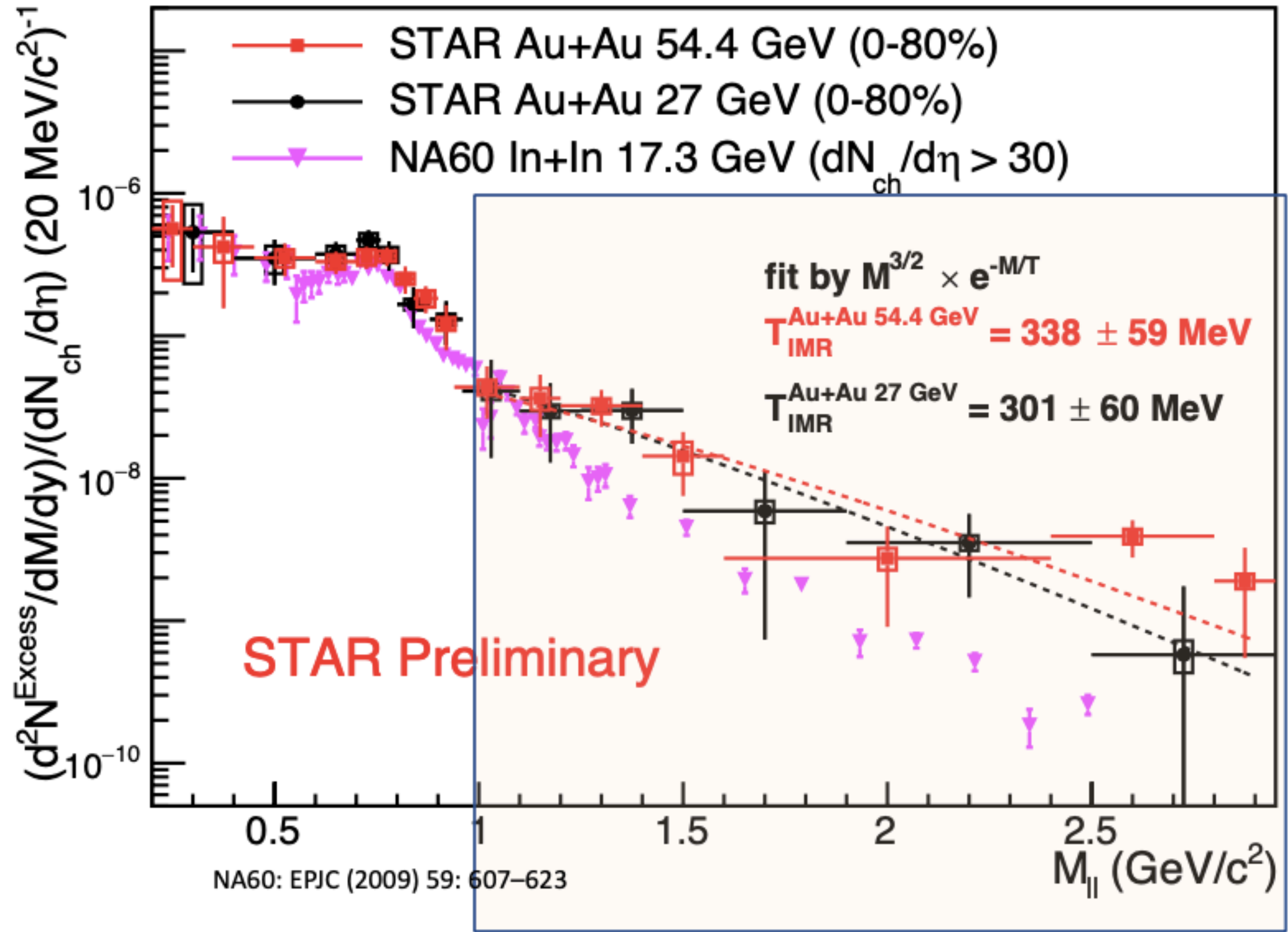
Fit to excess yield:
 ρ lineshape (Breit-Wigner) + thermal

Fit also done for NA60 data

Extracted temperatures close to $T_{pc} = 156$ MeV

talk Zaochen Ye

Temperature from excess dielectron mass spectrum: 2. IMR



$T_{IMR} > T_{pc} = 156 \text{ MeV}$

$T_{IMR}(\text{STAR}) > T_{IMR}(\text{NA60})$

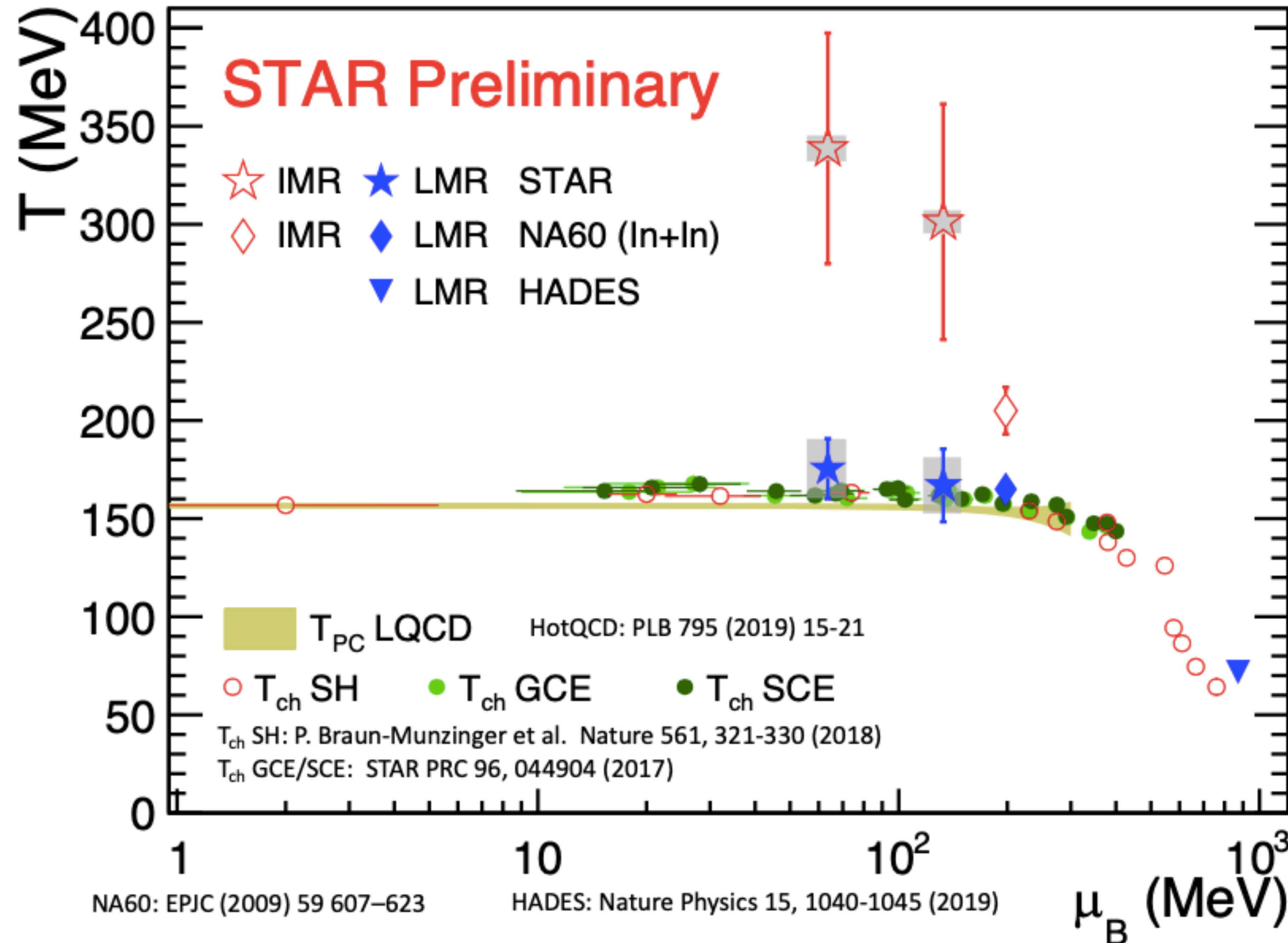
$\perp 205 \pm 15 \text{ MeV}$

Statistical uncertainties dominant

talk Zaochen Ye

New dielectron spectra in 27 GeV and 54.4 GeV Au-Au from STAR (3)

$T_{\text{eff}} > T_{\text{pc}} = 156 \text{ MeV}$ in intermediate mass range



Interpreted as space-time averaged QGP temperature of $T \approx 300 \text{ MeV}$

Back-on-the-envelope estimate

- multiplicity $\propto s^{0.155}$
- initial temperature $T^3 \propto dN/dy$

Expected temperature increase from $\sqrt{s_1} = 17.3 \text{ GeV}$, $T_1 \approx 200 \text{ MeV}$:

- 4% at 27 GeV
- 13% at 54.4 GeV

Measured T surprisingly large

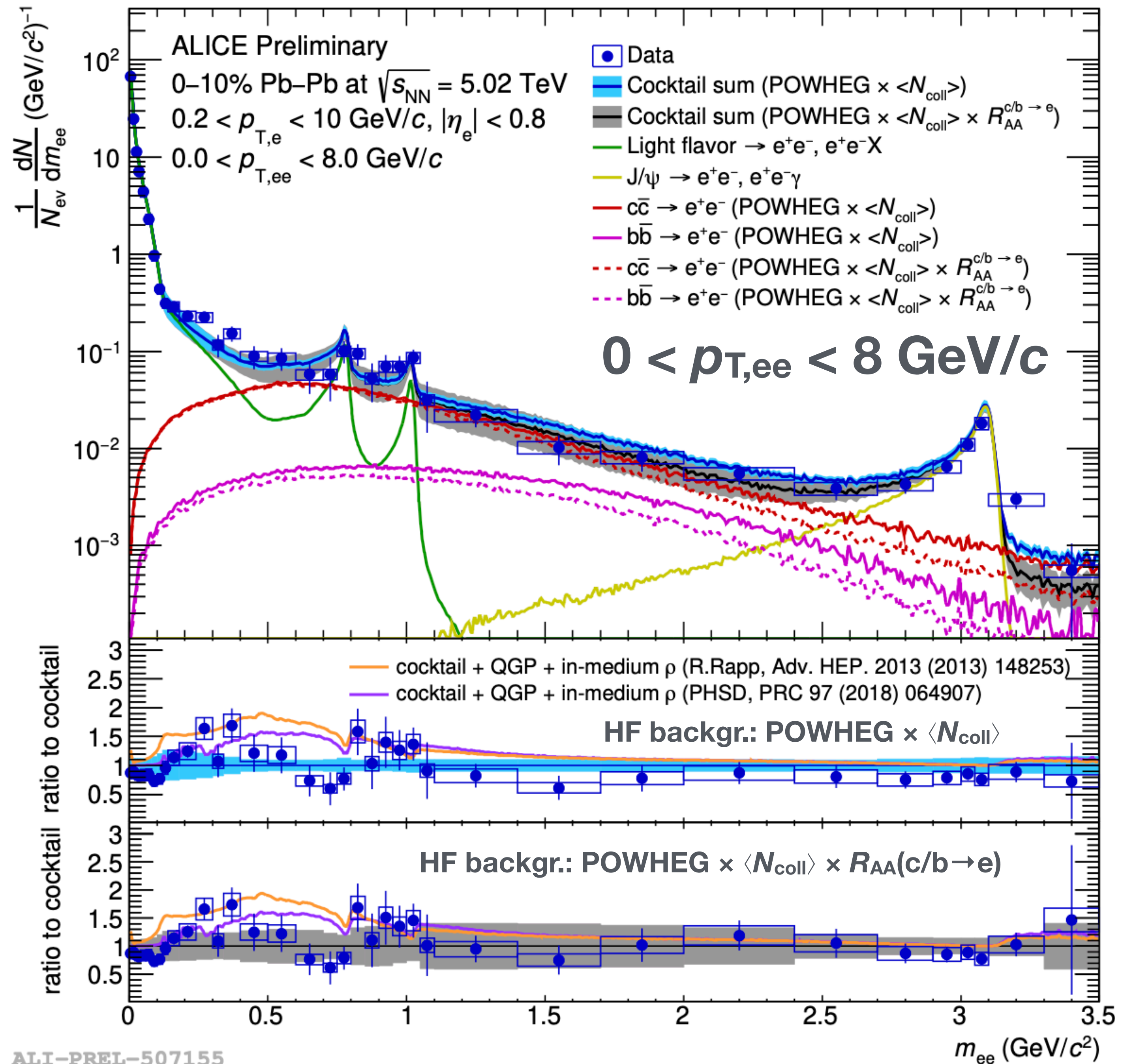
Dielectrons at 5.02 TeV Pb-Pb from ALICE

No significant excess within uncertainties

Handling of $c\bar{c}, b\bar{b} \rightarrow e^+e^-$ background is key for measuring thermal dielectrons from the QGP

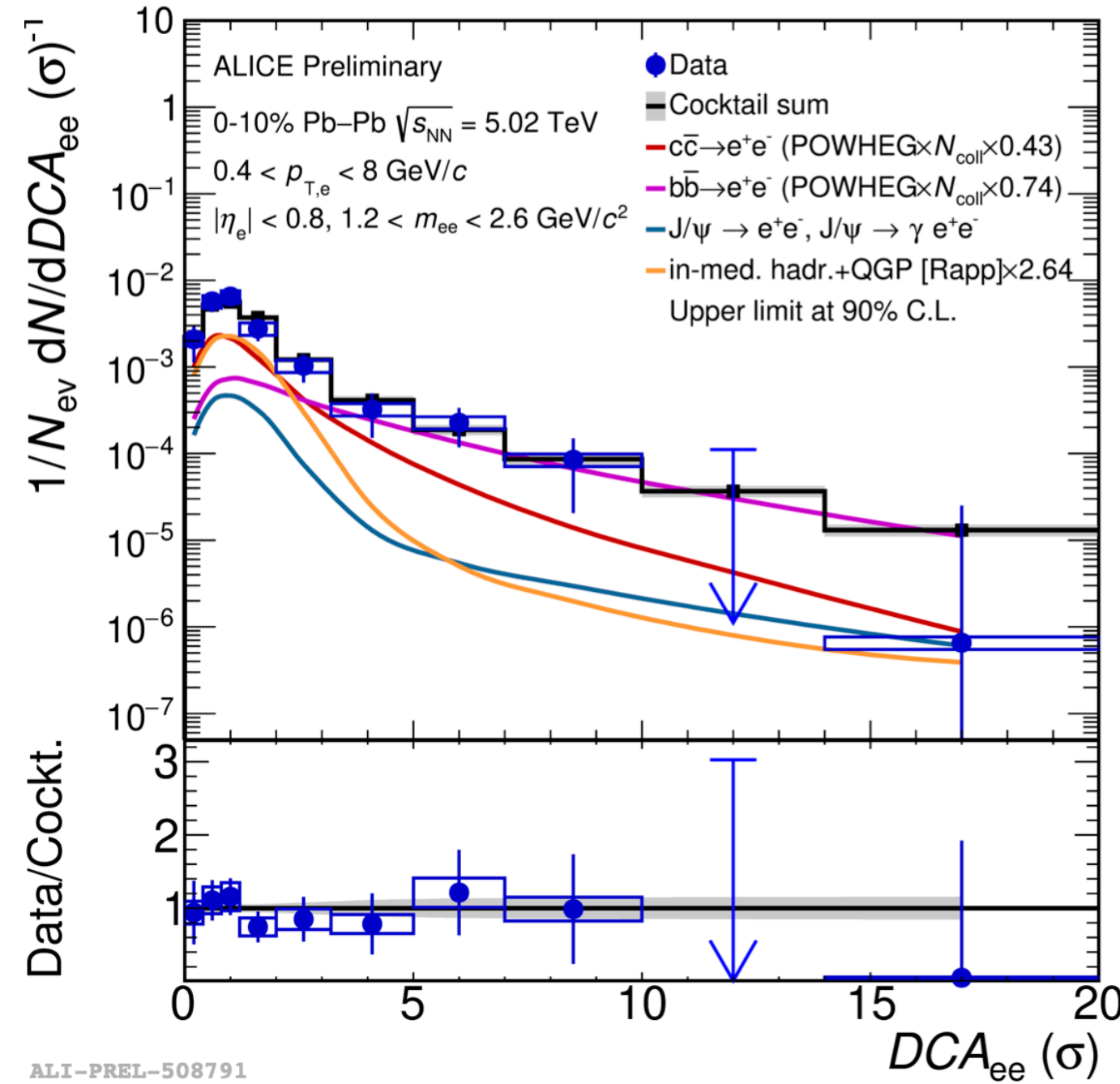
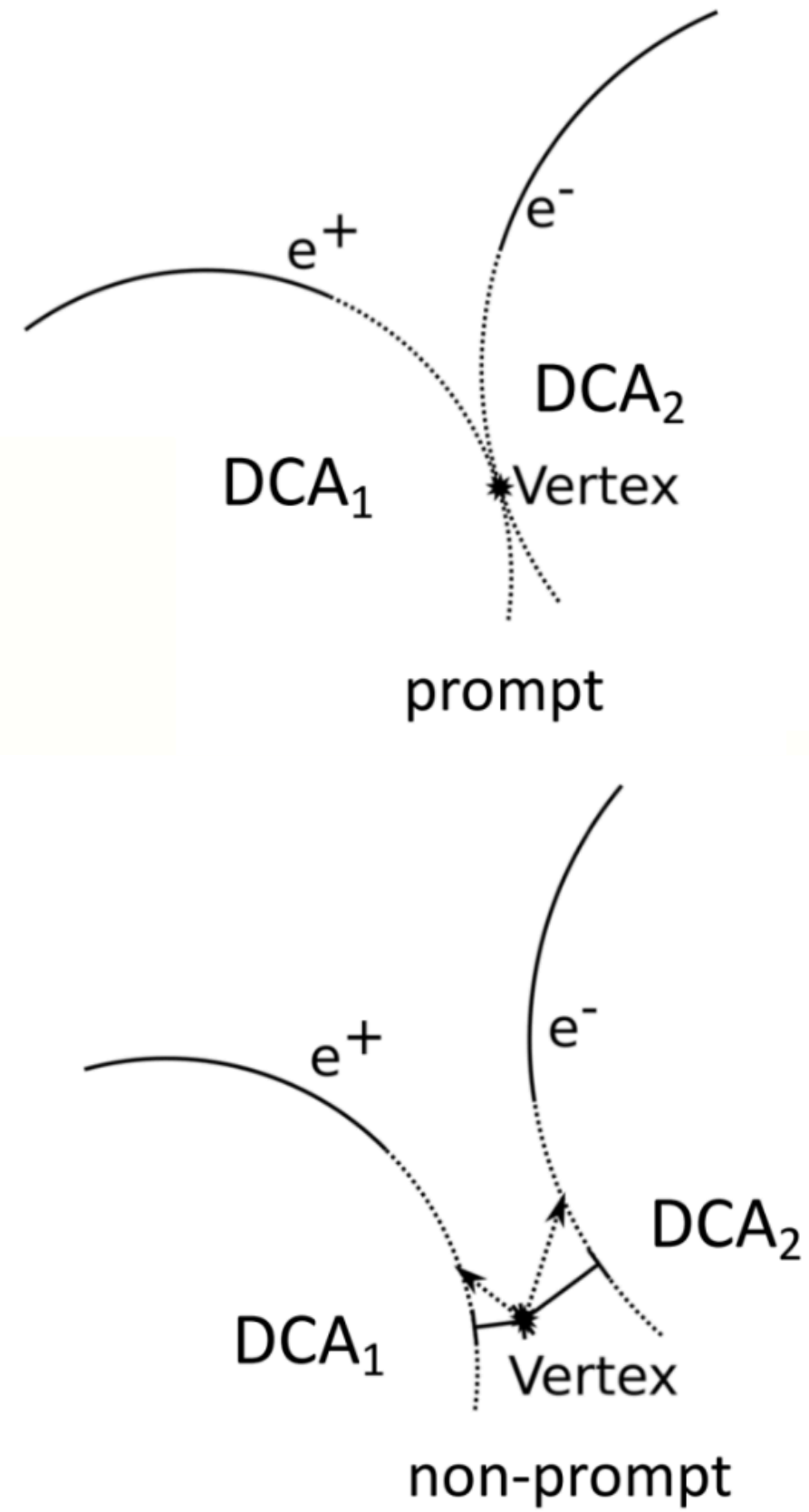
Two different cocktails

- ▶ N_{coll} -scaled HF in pp
- ▶ Include R_{AA} of $c/b \rightarrow e^{+/-}$

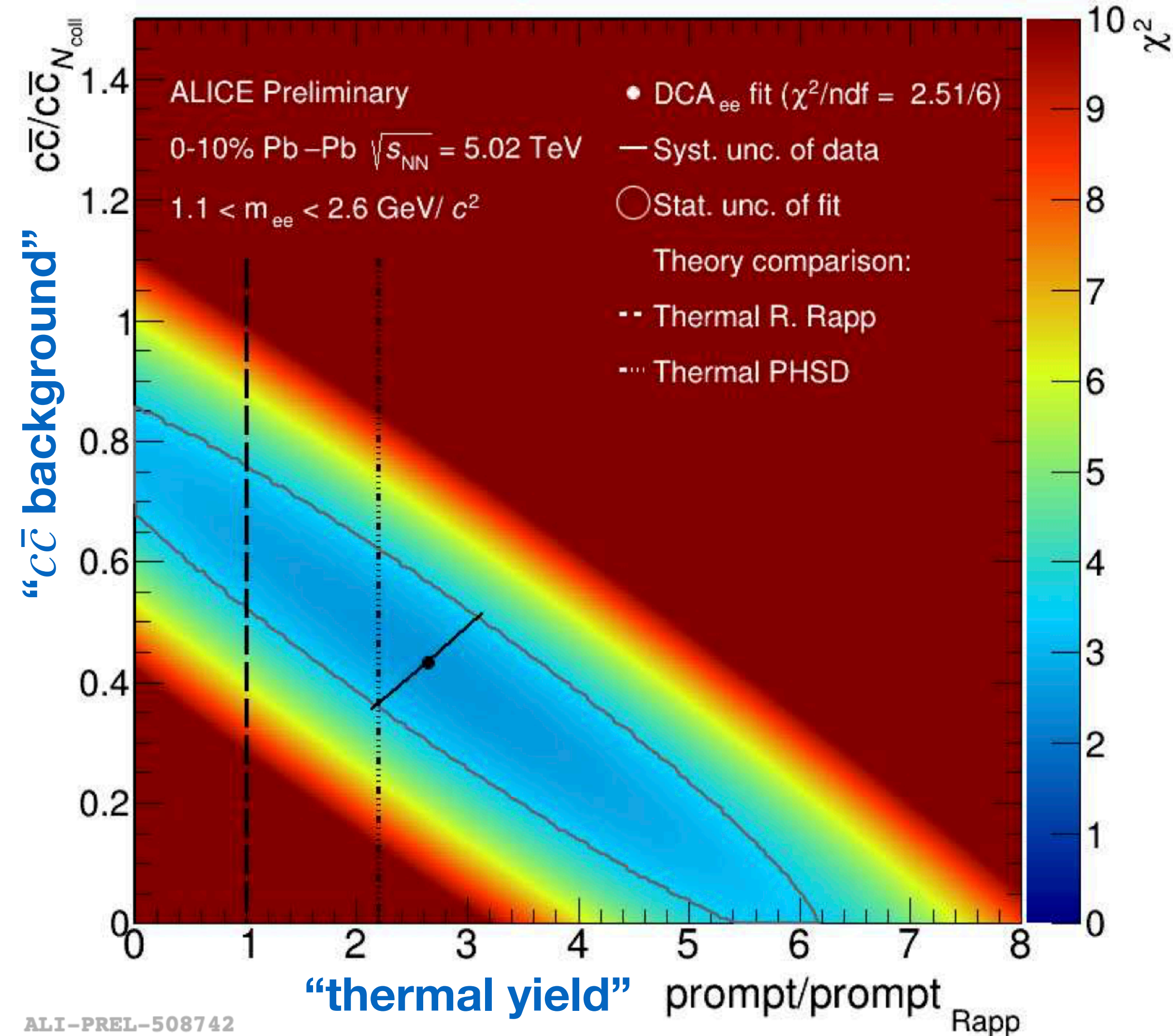


ALI-PREL-507155

High statistics and very good DCA resolution crucial



templates:
 $c\bar{c} \rightarrow e^+e^-$, $b\bar{b} \rightarrow e^+e^-$, prompt



Currently limited by statistics and DCA resolution:

→ **ALICE 2** (Run 3 & 4, **ITS2**, **ITS3**)

→ **ALICE 3** (Run 5+)

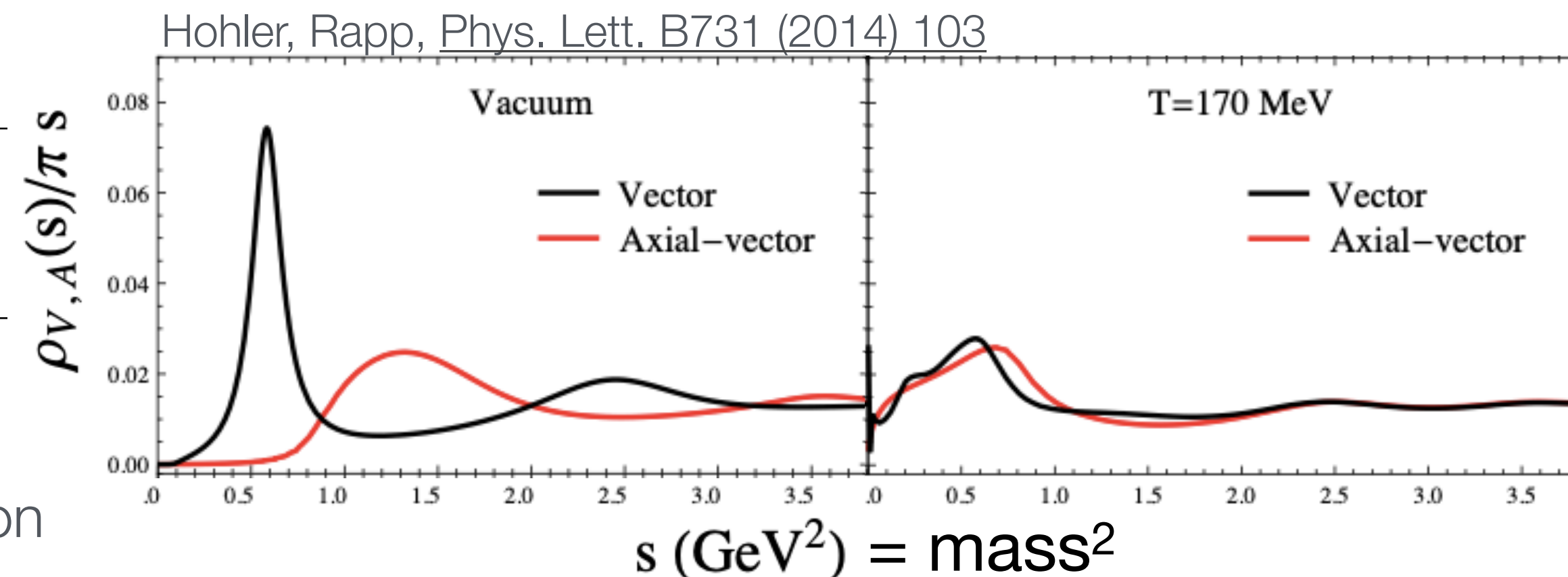
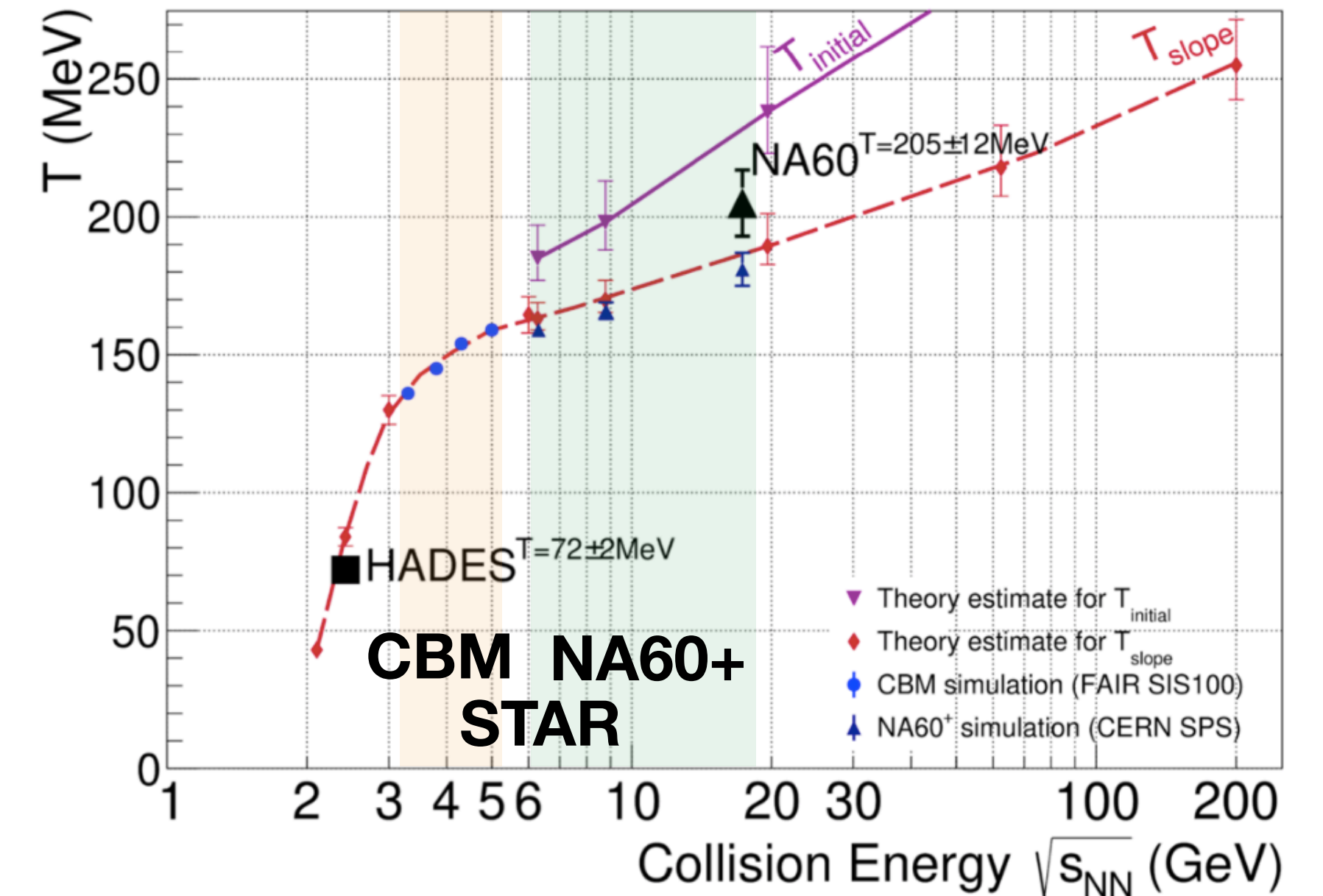
Future measurements

ALICE 3, [CERN-LHCC-2022-009](#)
 NA60+, [CERN-SPSC-2019-017](#)

Tetyana Galatyuk et al.

	High μ_B	Vanishing μ_B
Thermal radiation (γ , γ^* spectra)	STAR, NA60+, CBM, NICA	PHENIX, STAR, ALICE 2, ALICE 3
Caloric curve (T_{eff} in IMR vs. $\sqrt{s_{\text{NN}}}$)	STAR, NA60+, CBM, NICA	
Mechanism of CSR, ρ - a_1 mixing	CBM, NA60+	STAR, ALICE 3
Dilepton + direct photon v_n	CBM, HADES, NA60+	STAR, (ALICE 2), ALICE 3
Time dependence of T_{eff} , preequilibrium contribution		ALICE 3
Ultra-soft photons (test of Low's theorem)		ALICE 3
Electrical conductivity	HADES	STAR, ALICE 3*

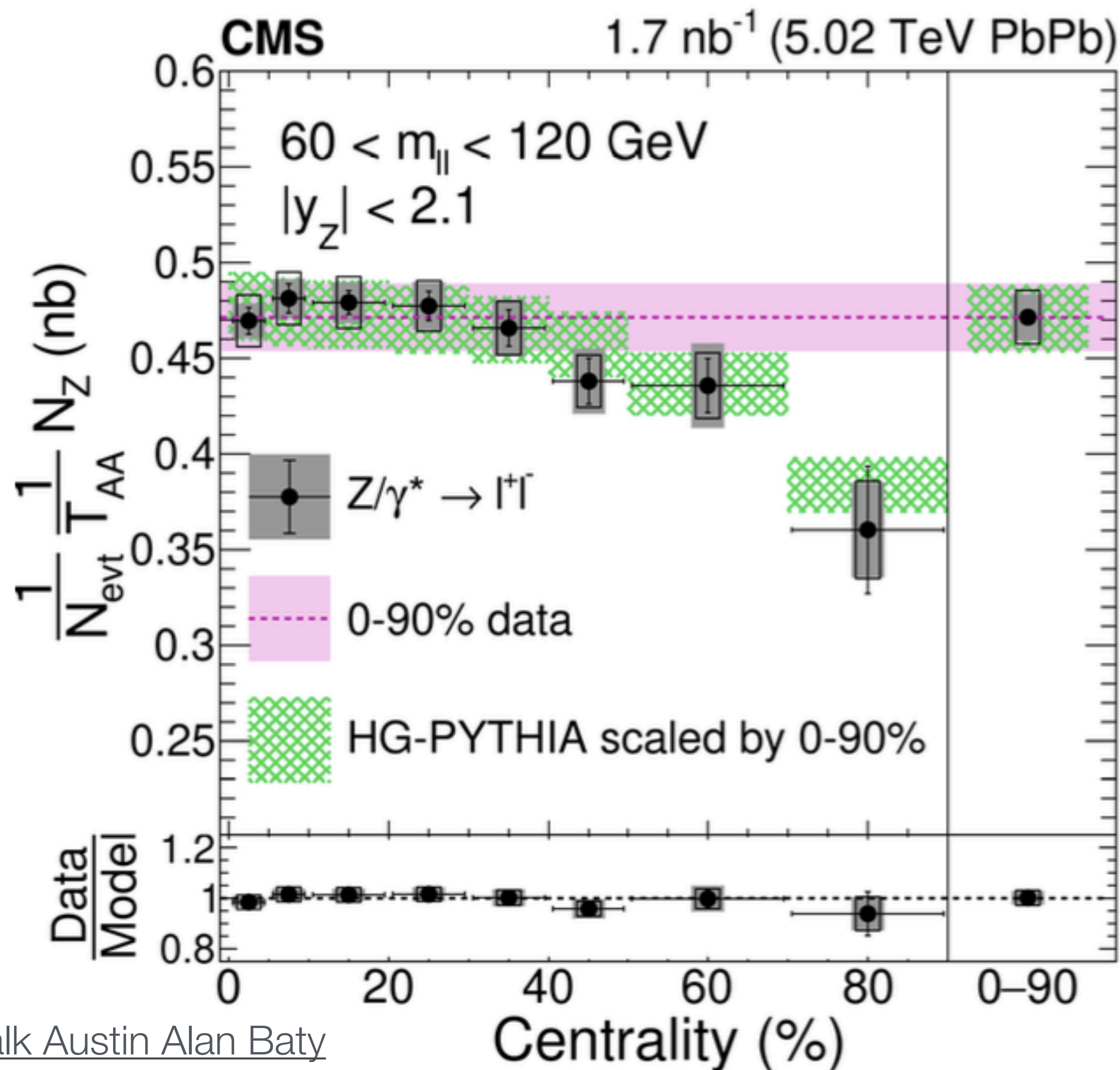
*under investigation



Initial state

Experimental T_{AA} from weak boson yield in CMS

HG-PYTHIA: C.Loizides,
A. Morsch, PLB B773 (2017) 408



Deviation of Z boson yield from Glauber T_{AA} scaling reproduced by HG-PYTHIA
→ interpreted as selection bias

Replace Glauber T_{AA} with $T_{AA}^{\text{exp}} = \frac{N_Z}{N_{\text{evt}} \cdot \sigma_Z^{\text{NN}}}$

Similar approach in PHENIX:
Long-standing puzzle of peripheral $R_{dAu}^{\pi^0} > 1$ for $p_T > 10$ GeV/c resolved by using direct γ 's:

$$\langle N_{\text{coll}}^{\text{exp}} \rangle = \frac{\left(\frac{d^2 N^\gamma}{dp_T d\eta} \right)_{dAu}}{\left(\frac{d^2 N^\gamma}{dp_T d\eta} \right)_{pp}} \quad \text{talk Niveditha Ramasubramanian}$$

Conclusions

Low- p_T direct photons

PHENIX data qualitatively in line with thermal QGP radiation. Yield puzzle appears less puzzling with 2014 data.

Important aspects not fully understood (e.g., direct photon v_2 , constant $\alpha(p_T)$)

ALICE: still no yield puzzle with new 5.02 TeV Pb-Pb data and improved analysis methods

Dileptons

So far only one measurement of QGP temperature from dileptons in IMR:

$T = 205 \pm 15$ MeV from NA60

New STAR data:

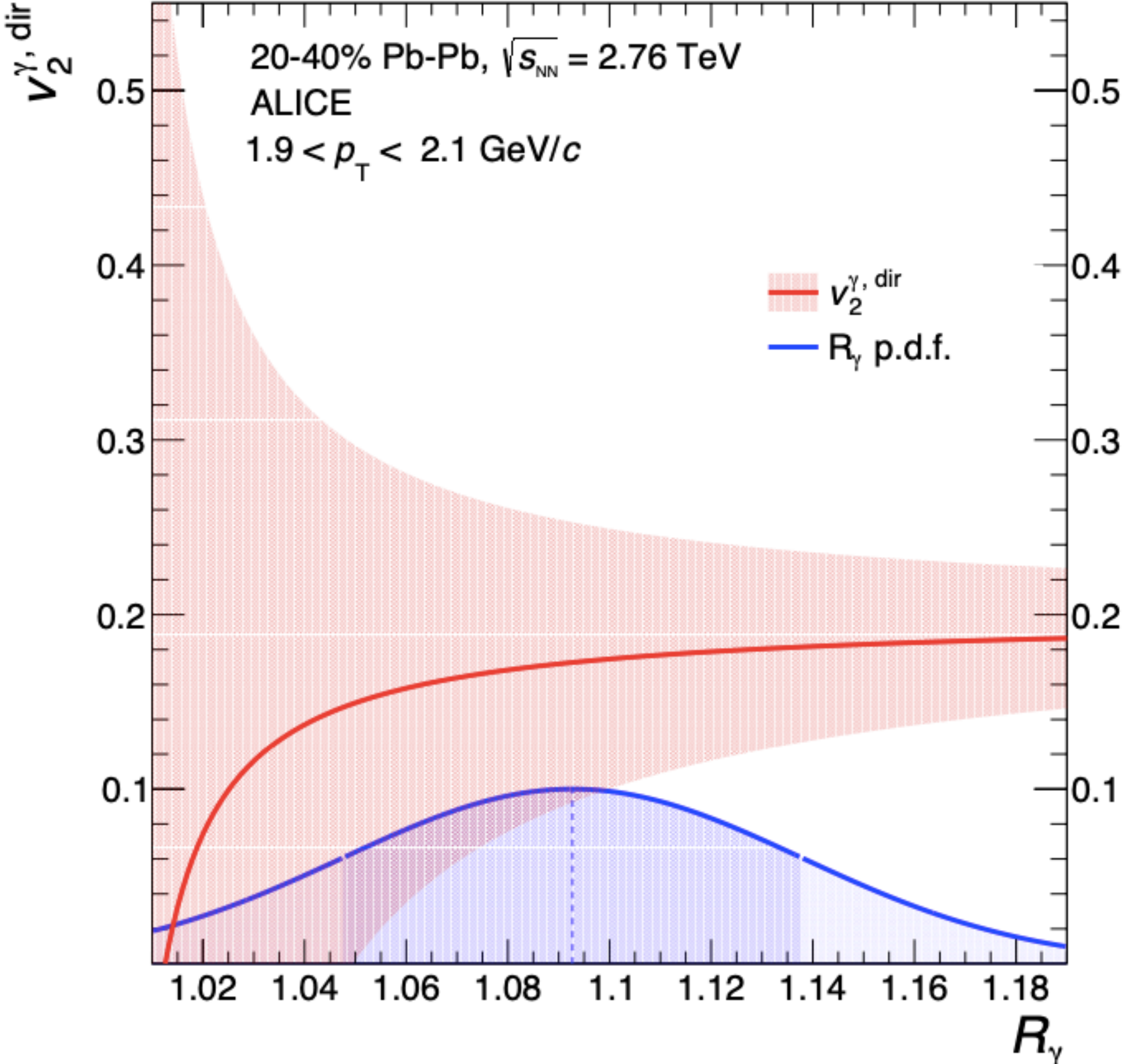
Surprisingly large T_{eff} ($\sim 300 \pm 60$ MeV) from fit in IMR in Au-Au at 27 and 54.4 GeV

Many thanks to: 🙏

Raphaëlle Bailhache, Peter Braun-Munzinger, Meike Danisch, Gabor David, Axel Drees, Tetyana Galatyuk, Frank Geurts, Szymon Harabasz, Christian Klein-Bösing, Aleksas Mazeliauskas, Norbert Novitzky, Jean-François Paquet, Mike Sas, Sören Schlichting, Daiki Sekihata, Johanna Stachel

Extra slides

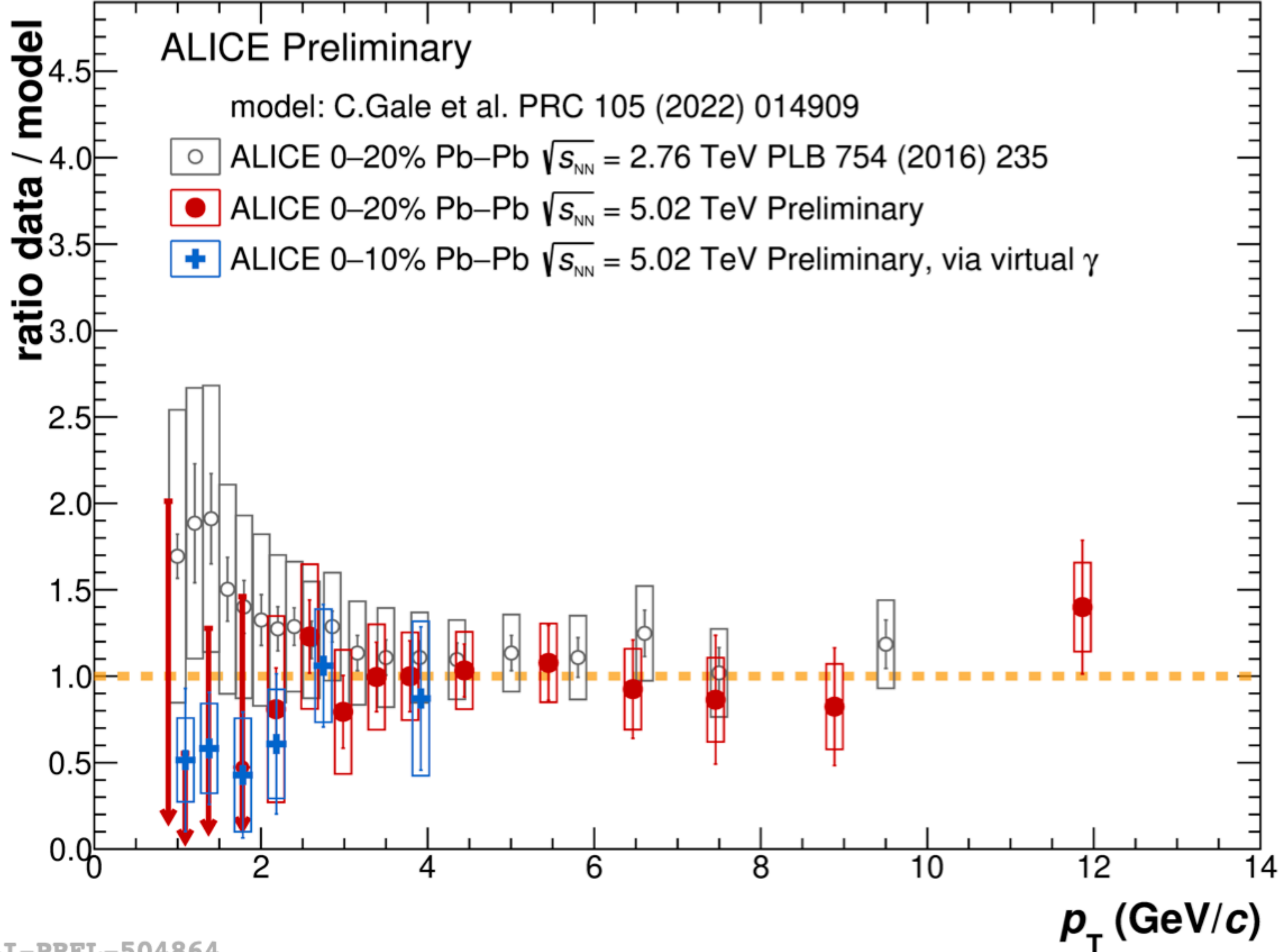
v_2 uncertainty vs R_γ



ALICE, Phys. Lett. B 789 (2019) 308

Data / (prompt + PE + thermal)

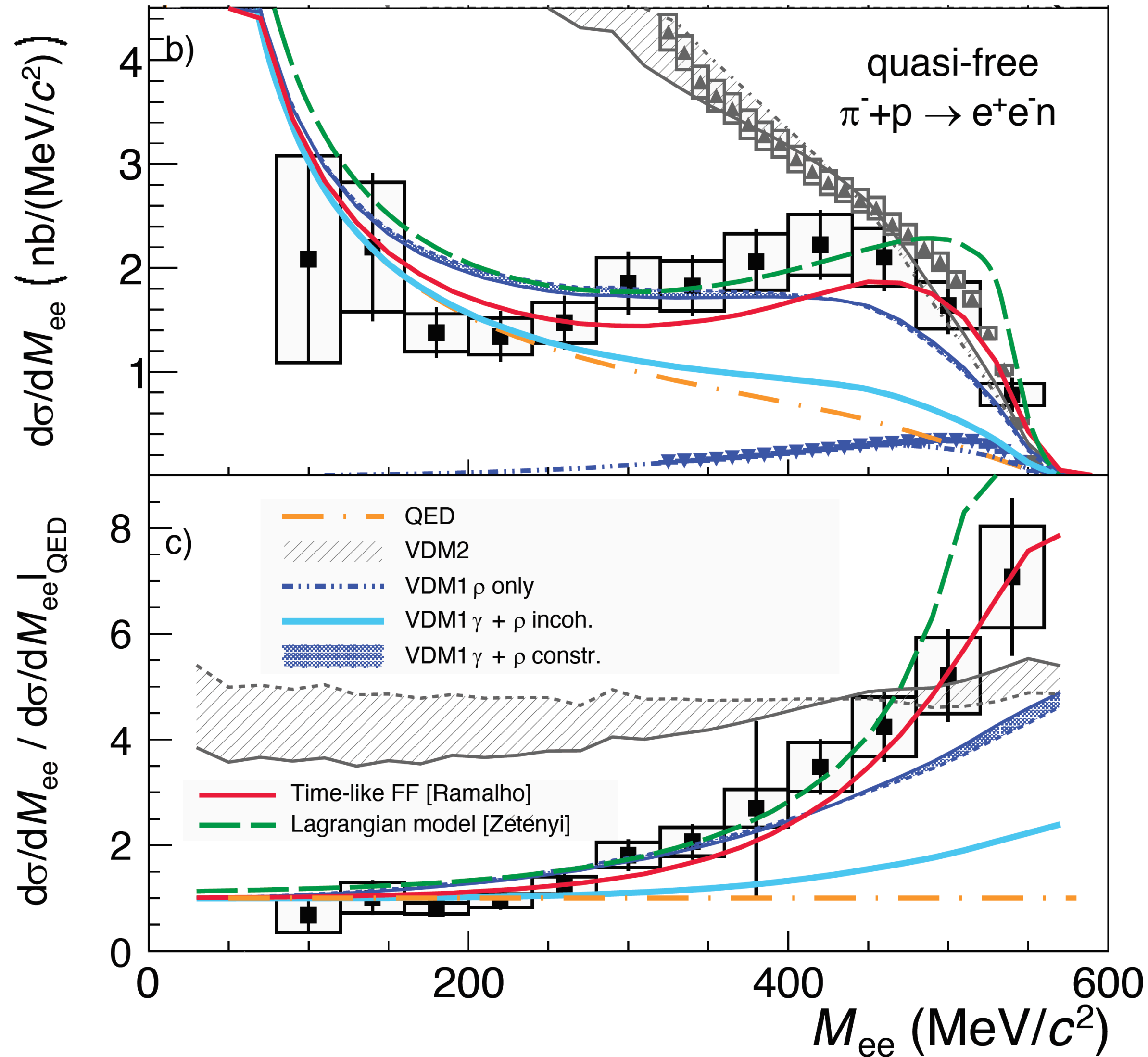
2.76 TeV and 5.02 TeV



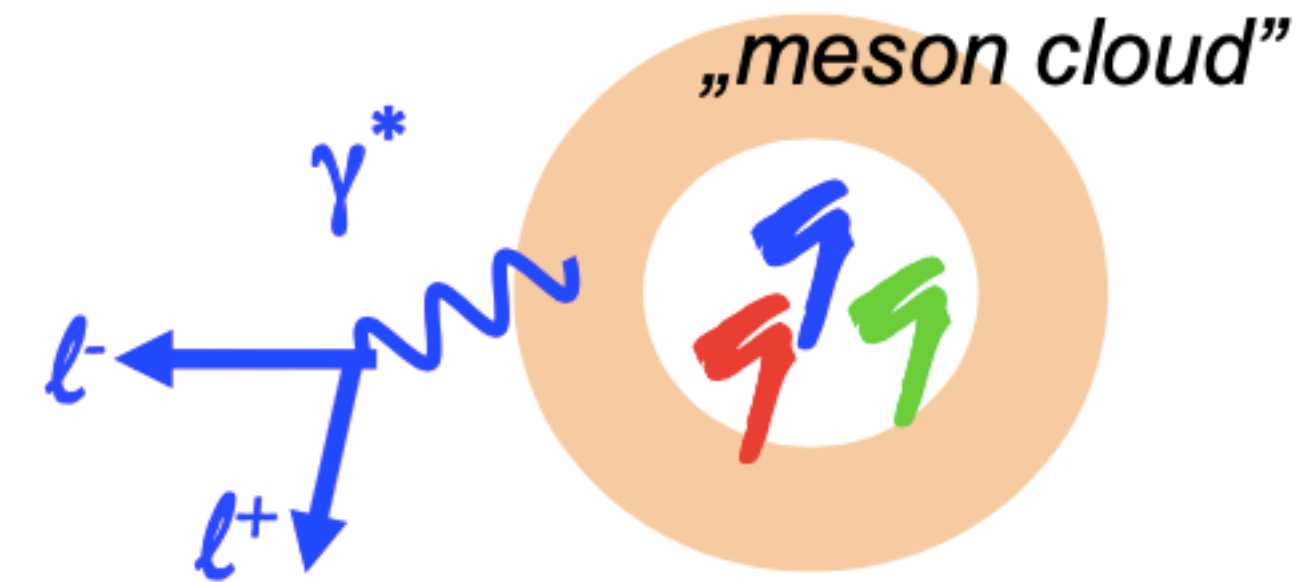
ALI-PREL-504864

Meson cloud exclusive analysis of $\pi^- p \rightarrow n e^+ e^-$

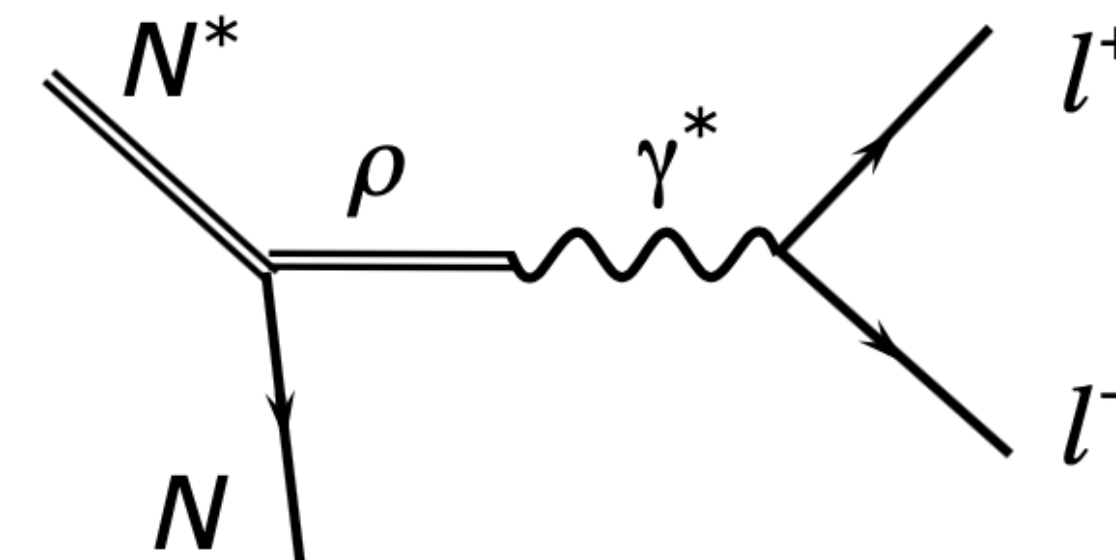
Dominance of the $N^*(1520)$ resonance



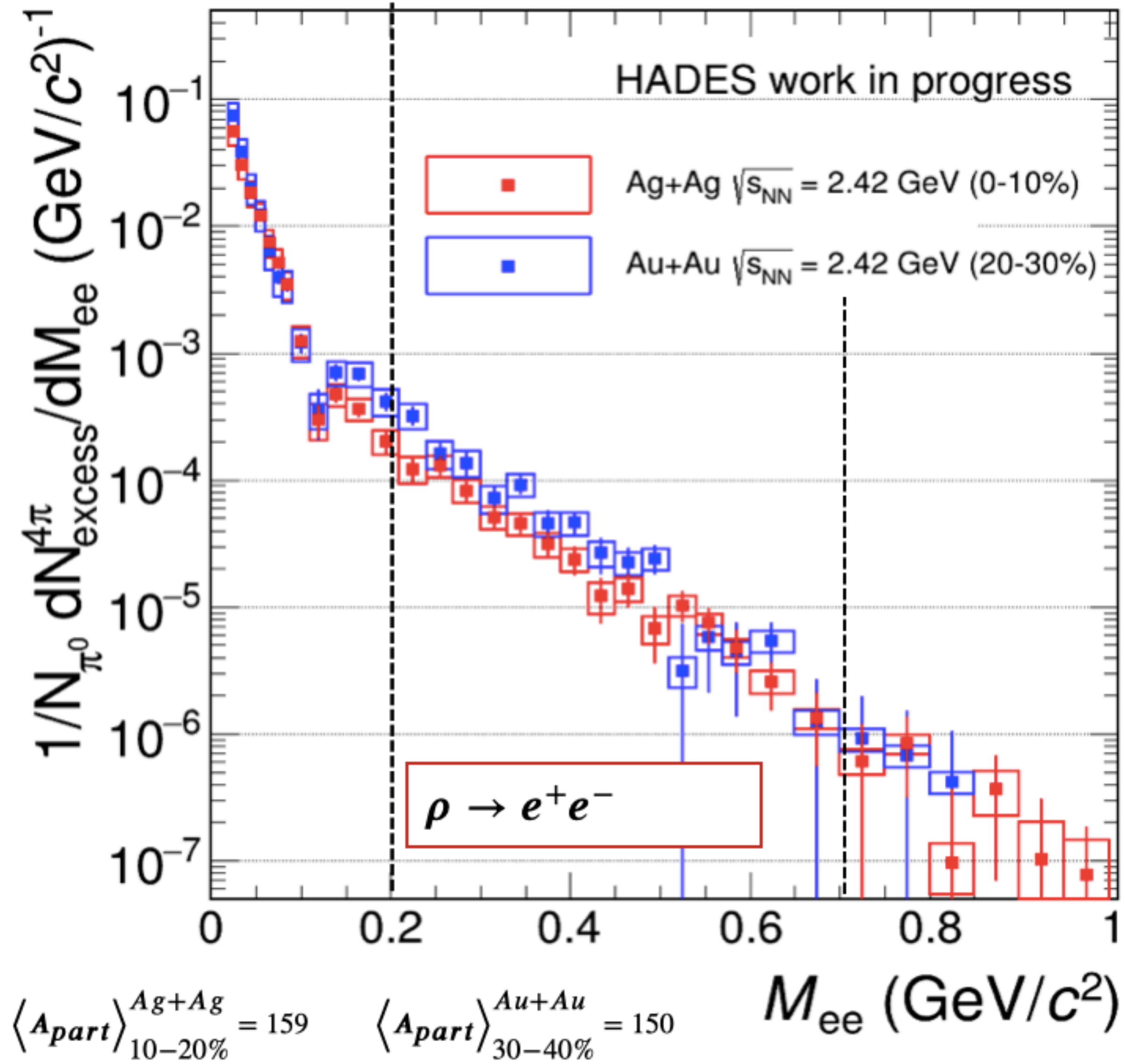
Excitation of a baryon can be carried by the meson cloud



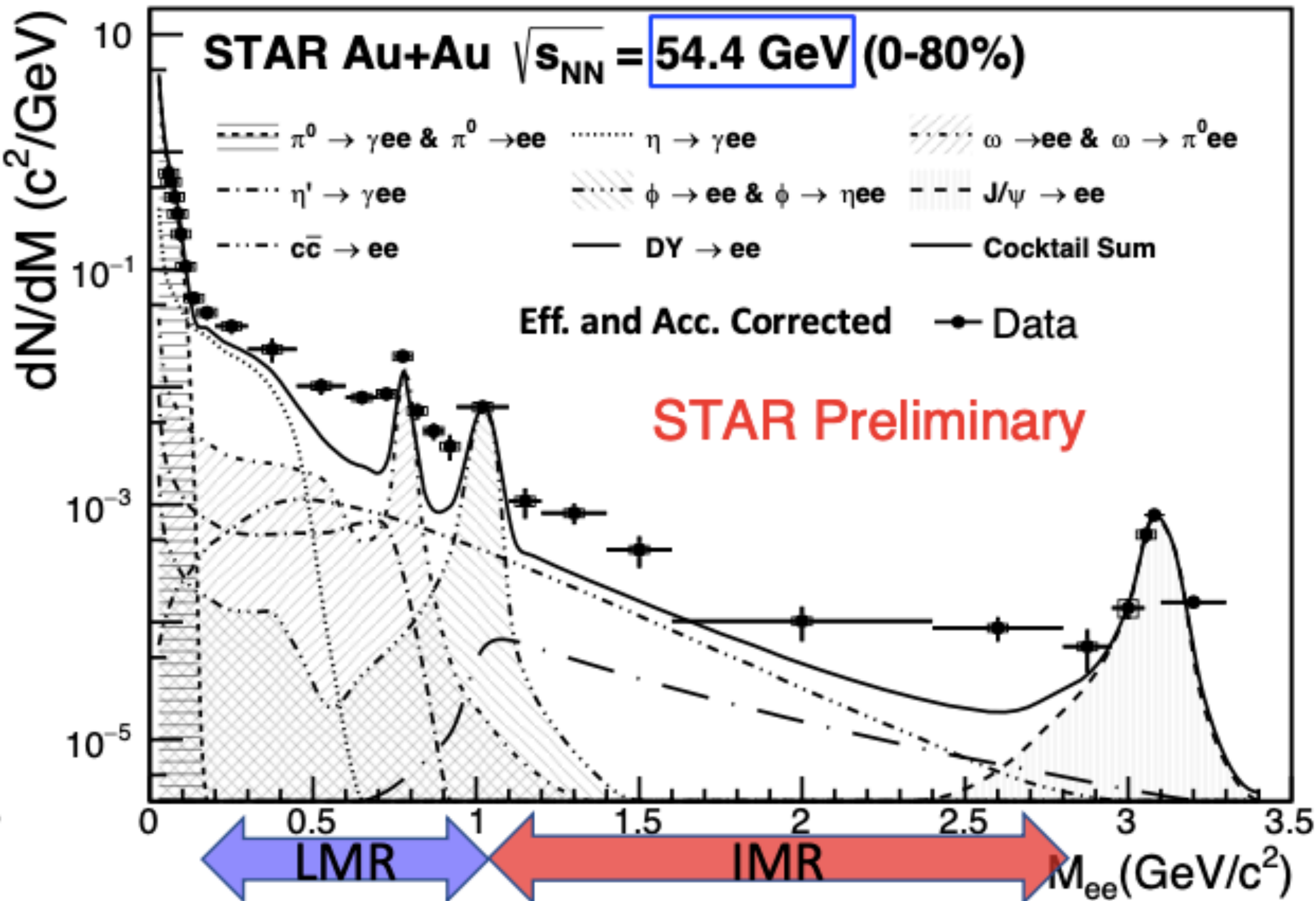
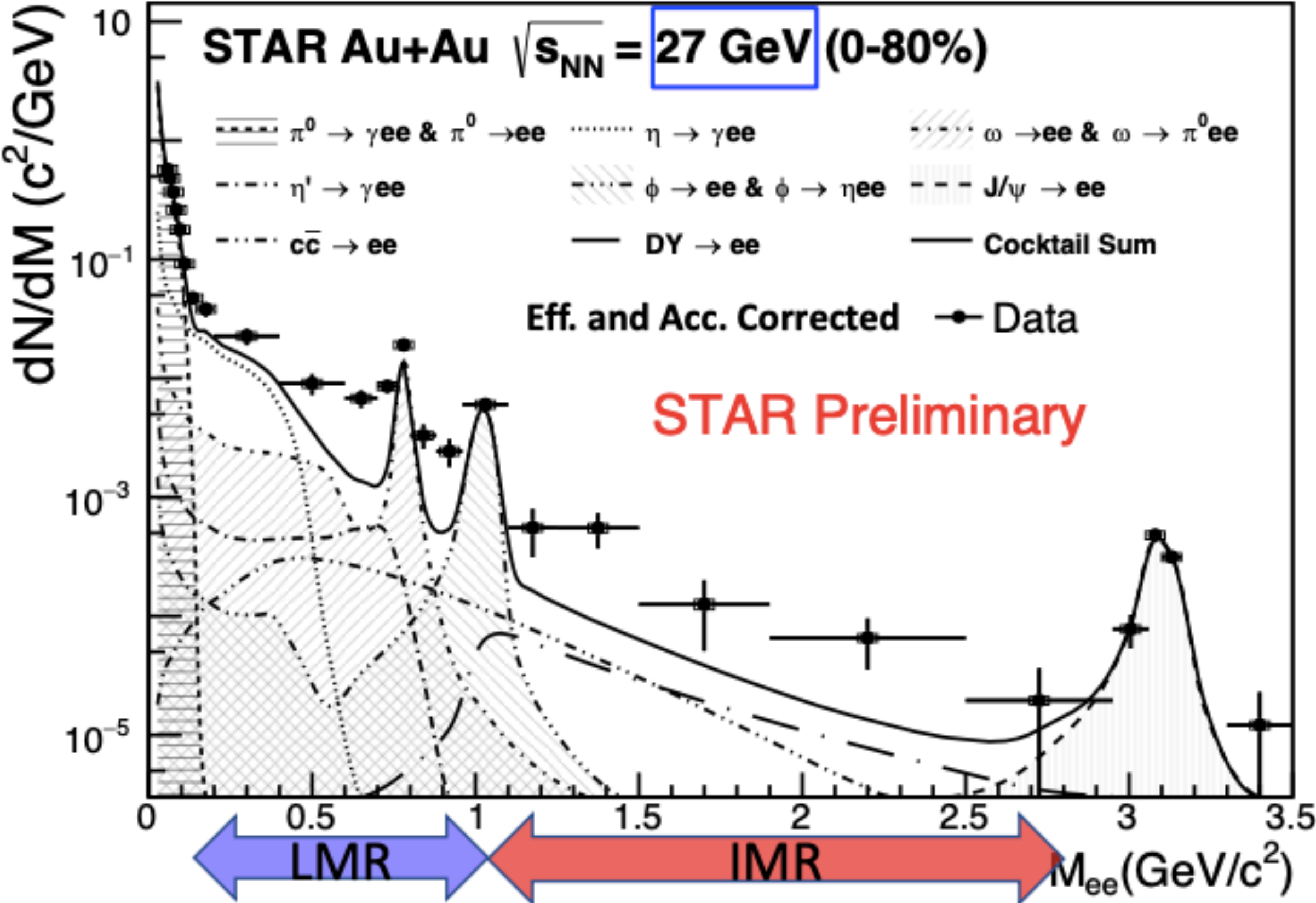
Vector Meson Dominance: the basis of emissivity calculations for QCD matter



Excess yield for same energy and N_{part} larger in Au-Au than Ag-Ag. Why?

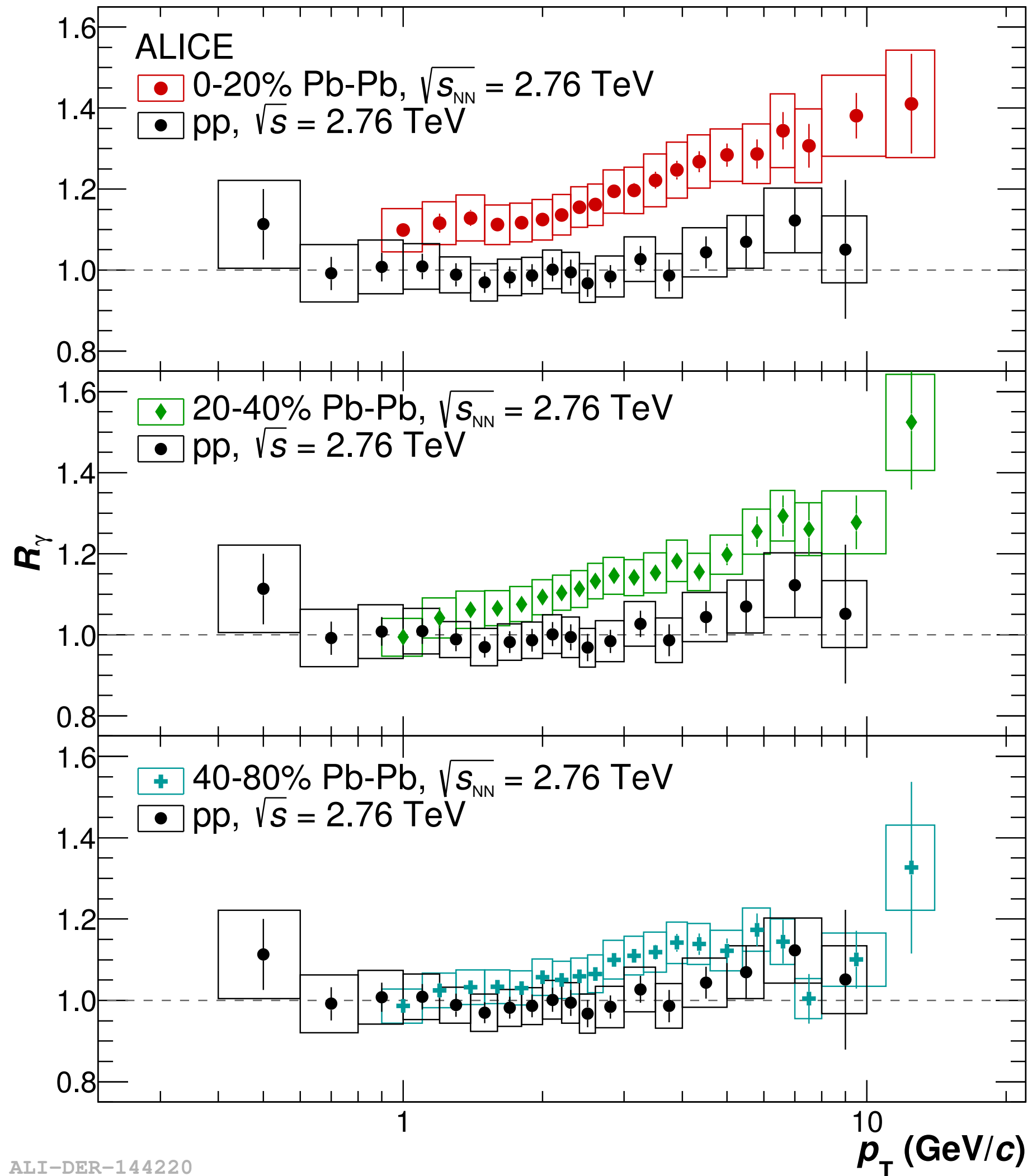


STAR dielectron spectra in Au-Au at 27 GeV and 54.4 GeV

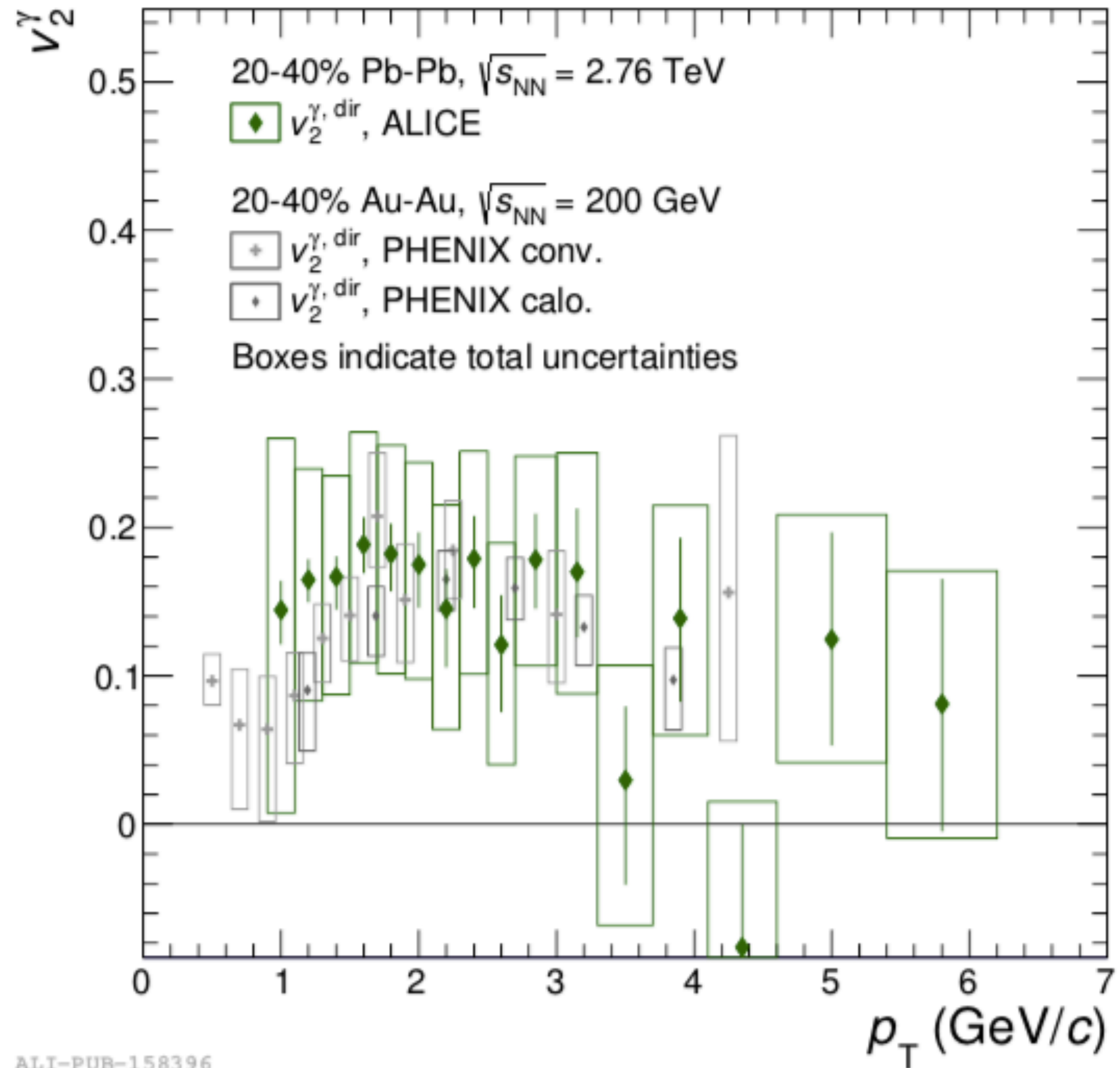


ALICE direct photon results in Pb-Pb at 2.76 TeV

R_γ and v_2^γ



ALI-DER-144220



ALI-PUB-158396

Upper limits on dielectrons
excess consistent with models

QGP radiation and ρ broadening consistent with the data

Expect temperature measurement via dielectrons from ALICE 2 (Run 3 + 4)

Time dependence of T from ALICE 3 (Run 5+)

