

Affects of criticality on the light nuclei yields

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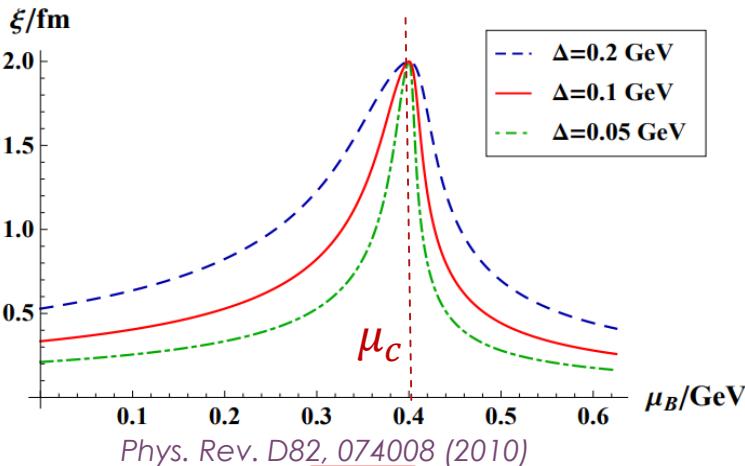
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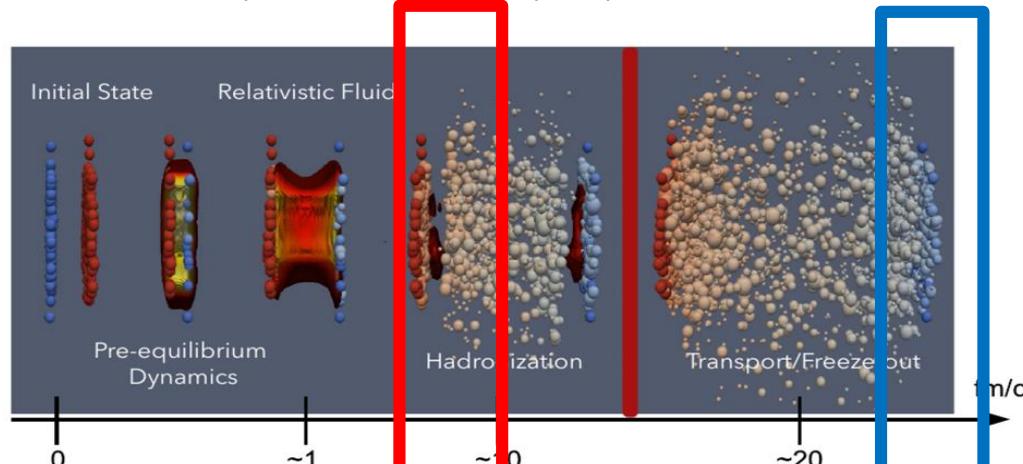


Main Logic

At chemical freeze-out:



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- Enhancement of correlation length (of order parameter) around CEP
- Larger correlation length leads to larger light nuclei yield ratio
- **Assumption:** the large correlation length survives the hadronic evolution
- Appearance of CEP - \rightarrow Enhancement of light nuclei yield ratio

Approach CEP

Light Nuclei Formation

Coalescence Model

- Deuteron abundance:

$$N_d = g_d \int dx_1 dx_2 dp_1 dp_2 f_{np}(x_1, p_1; x_2, p_2) W_d\left(\frac{x_1 - x_2}{\sqrt{2}}, \frac{p_1 - p_2}{\sqrt{2}}; \sigma_d\right)$$

Joint probability: $f_{np}(x_1, p_1; x_2, p_2) \sim \rho_{np}(x_1, x_2) e^{-\frac{p_1^2 + p_2^2}{2mT}}$

$$\rho_{np}(x_1, x_2) = \rho_n(x_1)\rho_p(x_2) + C_2(|x_1 - x_2|)$$

Long-range Cor.: $C_2^*(r) = \langle \delta\rho_B(r)\delta\rho_B(0) \rangle \approx \left(\frac{\partial\rho_B}{\partial\phi}\right)^2 \langle \delta\phi(r)\delta\phi(0) \rangle$
 $\sim \frac{\zeta_2 \langle \rho_n \rangle \langle \rho_p \rangle e^{-r/\xi}}{r} \quad \zeta_2 = \frac{T}{\pi\rho_B^2} \left(\frac{\partial\rho_B}{\partial\phi}\right)^2 > 0$

- ① Nucleons are emitted from a thermalized source
- ② Neglect the correlation among momenta
- ③ Dominated by the singular term around CEP

- Triton abundance:

$$N_t = g_t \int dx_1 dx_2 dx_3 dp_1 dp_2 dp_3 f_{nnp}(x_1, p_1; x_2, p_2; x_3, p_3) W_t(\sigma_t)$$

Joint probability: $f_{nnp}(x_1, p_1; x_2, p_2; x_3, p_3) \sim \rho_{nnp}(x_1, x_2, x_3) e^{-\frac{p_1^2 + p_2^2 + p_3^2}{2mT}}$

$$\rho_{nnp}(x_1, x_2, x_3) = \rho_n(x_1)\rho_n(x_2)\rho_p(x_3) + \rho_n(x_1)C_2(x_2, x_3) + \rho_n(x_2)C_2(x_1, x_3) + \rho_p(x_3)C_2(x_1, x_2) + C_3(x_1, x_2, x_3)$$

Long-range Cor.: $\widetilde{C}_3^*(k_1, k_2) \approx \frac{\zeta_3}{(k_1^2 + \xi^{-2})(k_2^2 + \xi^{-2}) \left((\vec{k}_1 + \vec{k}_2)^2 + \xi^{-2} \right)} \quad \zeta_3 = \frac{-3g_3 T^2}{\pi^4 \rho_B^3} \left(\frac{\partial\rho_B}{\partial\phi}\right)^3 > 0$

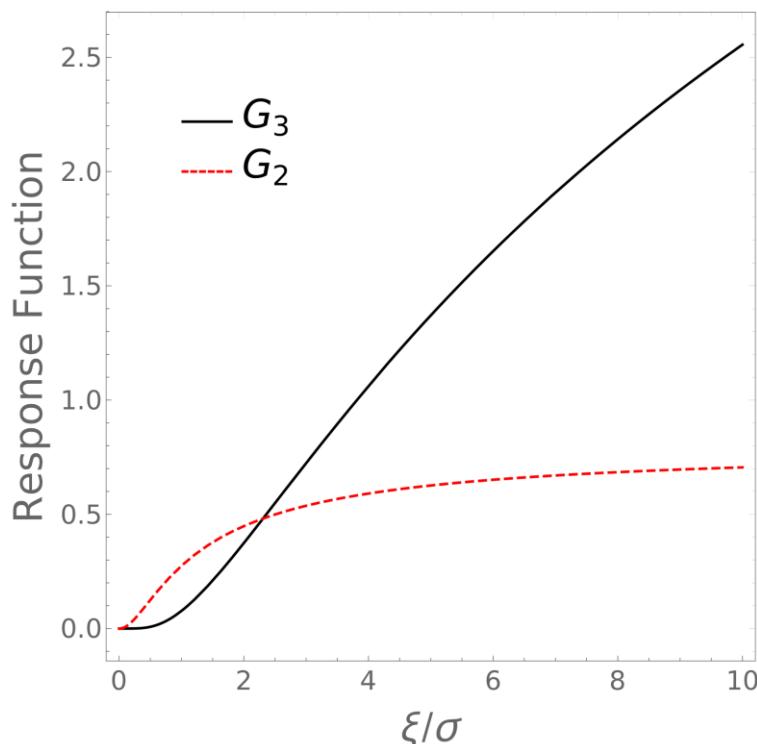
Yields & Correlation Length

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$$N_d \approx N_d^{(0)} \left[1 + C_{np} + \frac{\lambda_2}{\sigma_d} G_2 \left(\frac{\xi}{\sigma_d} \right) \right]$$

$$N_t \approx N_t^{(0)} \left[1 + \Delta\rho_B + 2C_{np} + \frac{3\lambda_2}{\sigma_t} G_2 \left(\frac{\xi}{\sigma_t} \right) + \lambda_3 G_3 \left(\frac{\xi}{\sigma_t} \right) \right]$$

$$\frac{N_t N_p}{N_d^2} \approx \frac{1}{2\sqrt{3}} \left[1 + \Delta\rho_B + \frac{\lambda_2}{\sigma_t} G_2 \left(\frac{\xi}{\sigma_t} \right) + \lambda_3 G_3 \left(\frac{\xi}{\sigma_t} \right) \right]$$

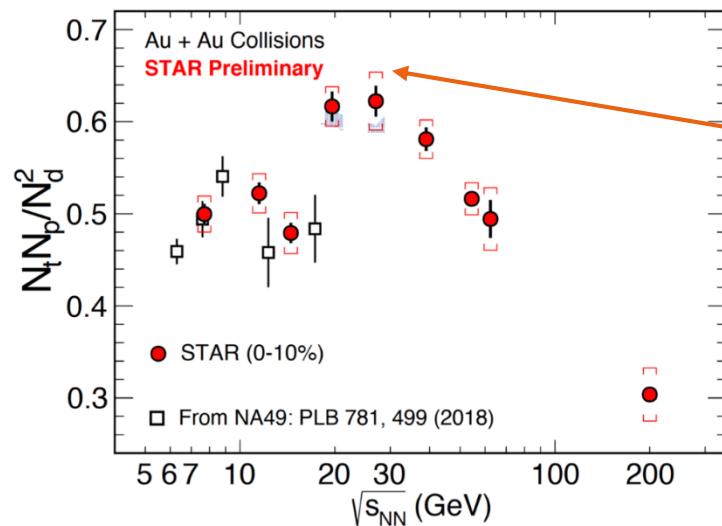
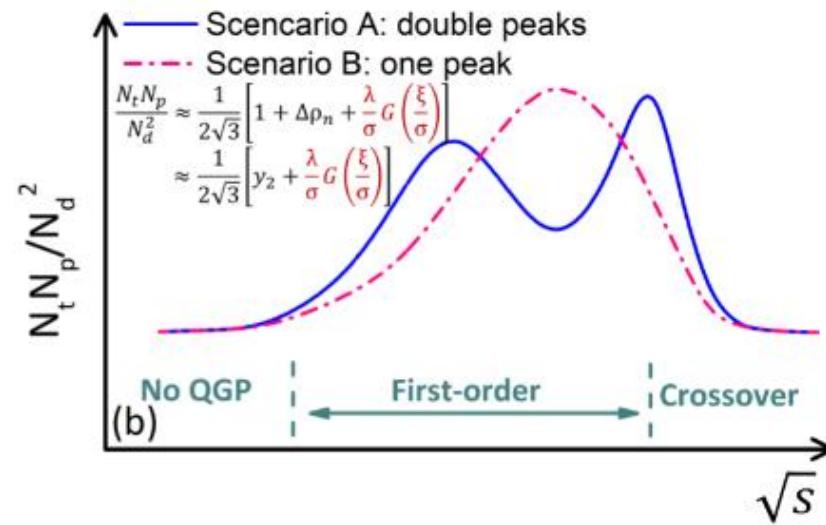
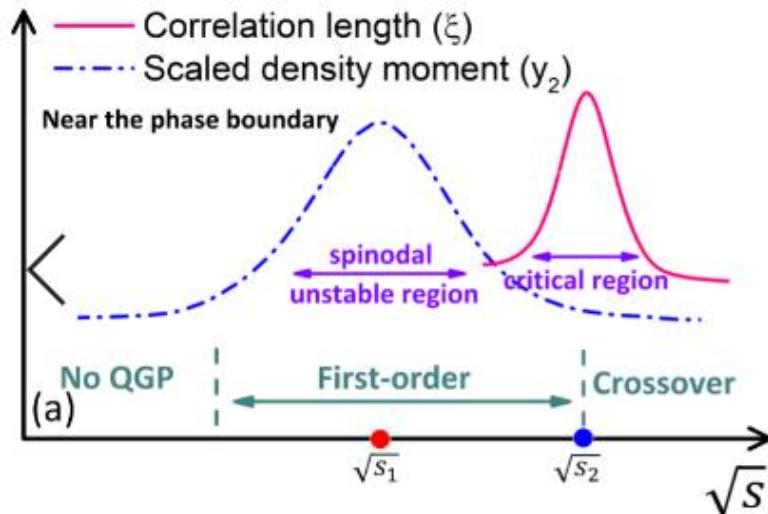


- Light nuclei yields (and their ratios) increase with the correlation length

can thus be employed as a 'ruler' measuring the correlation length

- The resolution of such a 'ruler' is the size of the light nuclei (2-3 fm)
close to ξ_{\max}
- G-functions saturates at large ξ
a broad peak in BES

Two Scenarios in BES



Signal from critical point?

A dynamical simulation (maybe via hydro+) is necessary to see whether the long-range correlation survives during the hadronic evolution

Appendix: Sign of ζ_3

Quark-Meson Model: $H = H_{kin} + H_{q-m} + \frac{\lambda}{4}(\sigma^2 + \pi^2 - \nu^2)^2 - c\sigma$

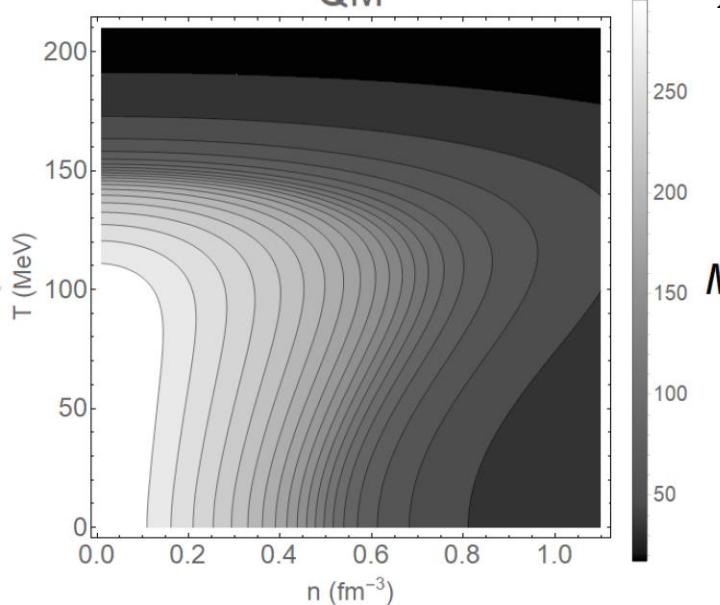
$$\langle \delta\rho_B(x)\delta\rho_B(y)\delta\rho_B(z) \rangle \sim \left(\frac{\partial\rho_B}{\partial\sigma^*} \right)_T^3 \langle \sigma'(x)\sigma'(y)\sigma'(z) \rangle$$

Order parameter for chiral phase transition: $\sigma = \sigma^* + \sigma'$

Expectation value of q-condensate: $\sigma^* = \frac{M_Q}{g_{Q\sigma}} > 0$

$$H = \dots + \lambda\sigma^*\sigma'^3 + \dots$$

$$g_3 \approx \lambda\sigma^* + O(\hbar) > 0$$



$$M_u \text{ MeV}$$

$$\left(\frac{\partial\rho_B}{\partial\sigma^*} \right)_T < 0$$

$$\zeta_3 = \frac{-3g_3T^2}{\pi^4\rho_B^3} \left(\frac{\partial\rho_B}{\partial\phi} \right)^3 > 0 !$$