Simulation in Fermilab Neutrino Experiments

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Introduction

Organized around the stages of simulation in v exps:

- Neutrino Beams
- Neutrino Interaction Generation
- Particle Tracking
 - Notes on G4 interfaces
 - Special cases: optical photons
- Looking Forward:
 - Features
 - Performance
 - Architectures

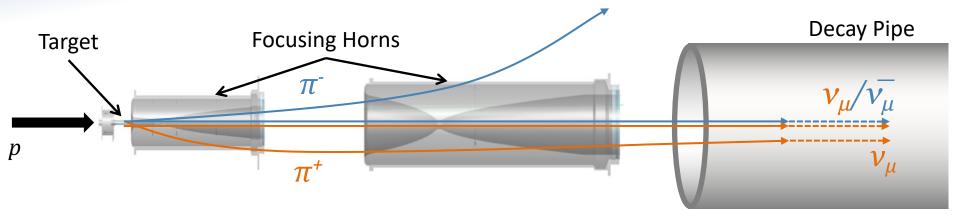
Which experiments, exactly, are we talking about?

- Segmented organic scintillator
 NOvA
- Liquid Argon
 - MicroBooNE
 - SBND
 - ICARUS
 - DUNE FD, Single/Dual
 - ProtoDUNEs
 - LArIAT
- Water Cherenkov
 ANNIE
- Multi-component
 DUNE ND

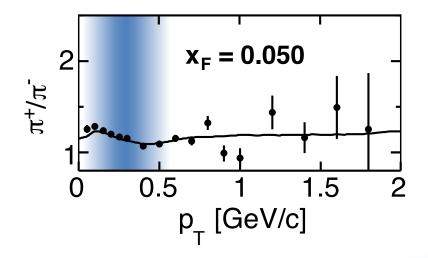
Introduction

- Who am I?
 - DUNE Far Detector Simulation+Reconstruction Convener
 - DUNE Single Phase Photon Detection Simulation+Physics Convener
 - NOvA Analysis Coordinator, previously Computing Coordinator
 - Former simulation expert on Super-K/Hyper-K.
- So, I can speak about NOvA and DUNE FD with more authority, but I sought input from across the program.

Neutrino Beam Simulation



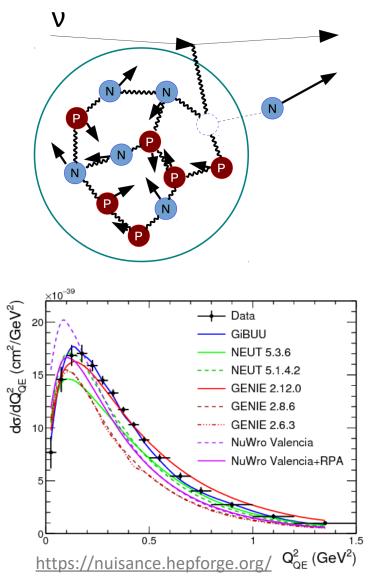
- Everyone uses GEANT4
- Big uncertainties in hadroproduction, no well-tuned physics list.
- Largely mitigated in multi-detector experiments, but a major concern in cross-section measurements.
 - In NOvA, 10% before extrapolation,
 < 2% after extrapolation.
- NOvA uses PPFX to reweight to hadroproduction data
 - Developed by MINERVA
 - Also creates systematic variations.



NA49, Eur. Phys. J. C 49 897 (2007)

Stages of Simulation in Neutrino Experiments

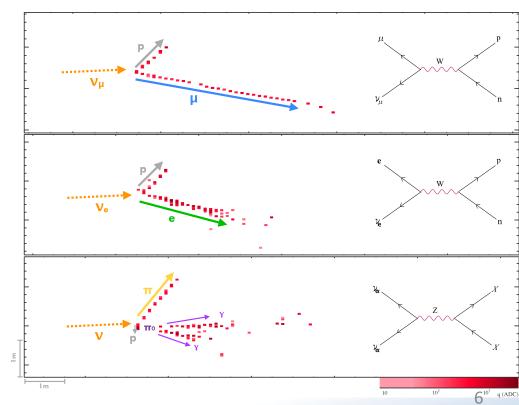
- Neutrino Event Generation: GENIE
- GENIE is the primary generator used at FNAL, but there a bunch of others out there
 - NEUT, NuWro, GiBUU, MARLEY, etc.
- Even within GENIE, there are many choices of models available for use.
- This is another area where there are significant uncertainties, which are partially reduced in multi-detector experiments.
 - In NOvA, extrapolation reduces uncertainty from ~15% to ~5%.
 - The goal for DUNE is ~1%, so additional work is needed here, but that's out of scope of this group.



Tracking Particles in the Detector

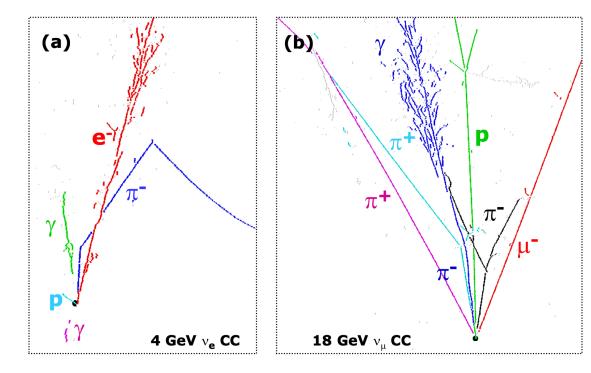
• Everyone uses Geant4

- Though we often use different versions across experiments.
 - Upgrading major externals like G4 is a lot of work, so the algorithm is usually to get current when the opportunity arises every 1-2 years.
- NOvA: 4.10.4 p02b
 - Probably for a while
- LArSoft: 4.10.3 p03e
 - DUNE, SBN, μBooNE...
 - Discussing 4.10.6 p01 now
- ANNIE: 4.10.1/2
- We use G4 to track final state particles which leave the nucleus.
- Track through any interactions and record energy deposits.



What are we tracking?

- Which particles?
 μ, e, p, n, π, γ, K, τ
- Which processes?
 - Ionization
 - EM showering
 - Scattering
 - hard, MCS
 - Decays (*μ*, *π*, *K*)
 - Captures (μ , n)
 - Nuclear de-excitation
 - particularly important for Supernova v
 - Neutron thermalization
- Energy range:
 - Lowest: few hundred KeV (e.g. de-excitation gammas)
 - Typical: 10's of MeV to a few GeV
 - Highest: a few TeV (cosmic ray muons)



Geant4 Interfaces

- We all use a common analysis framework called ART, which has a built-in interface to Geant4.
 - Coordinates input from flux and event generator, output to custom parts of simulation.
 - Everyone uses GDML to describe detector geometries.
- Specific example: NOvA
 - Treat the bulk of scintillator in each cell as a "Sensitive Detector" and record any energy deposits.
 - Custom code then handles scintillation and Cherenkov light production, transport and collection of photons, electronics, etc.

Far Detector

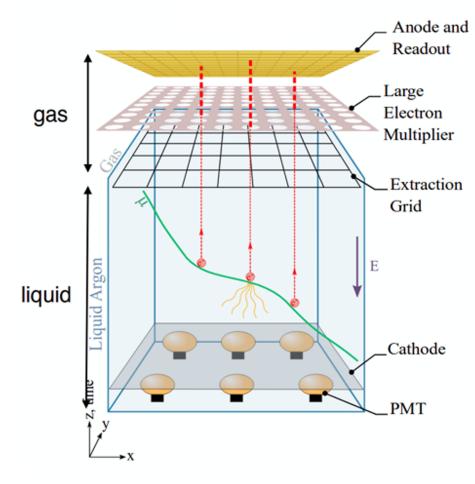
896 planes 344,064 channels

^{3.9} cm

6.0 cn

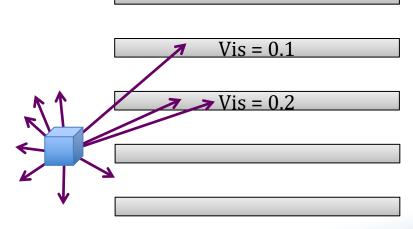
Geant4 Interfaces

- Specific example: liquid argon
 - Note: describing what we call "new LArG4," a new interface developed by Hans Wenzel based on <u>artg4tk</u>.
- Treat the whole LAr volume as the sensitive detector.
 - Record "SimEnergyDeposits" frequently with limited step lengths.
- These energy deposits then passed to other framework modules to handle:
 - Scintillation and ionization
 - Photons and drift electrons
 - Box model, NEST, etc.
 - Drift of electron to readout plane
 - Including simulation of field response (see arXiv:1802.08709)
 - Fast simulation of optical photons.



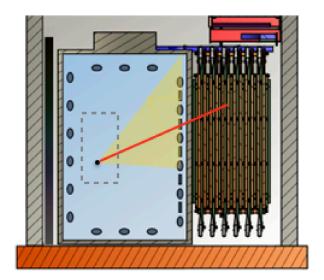
Special Case: Optical Photons in LAr

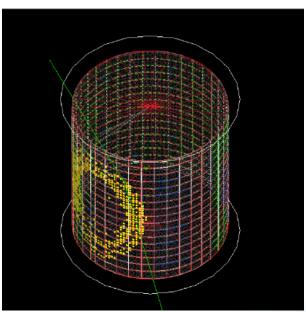
- Photons in LAr
 - LAr is a strong scintillator $40k \gamma$ /MeV.
 - It is prohibitively slow (minutes/event) to track every photon
- A variety of fast simulation approaches, all trained on full Geant4.
 - Take advantage of scintillation light being isotropic to generate "bombs" of photons distributed in the geometry.
 - Most LAr experiments have used a 4D look-up library.
 - DUNE geometries are too large for the library, so we're now using a semiempirical parameterization, exploring ML solutions.
- These tricks all break for directional Cherenkov light.
 - So far, we are mostly ignoring this few-percent contribution.
- If fast (GPU?) photon tracking became available in Geant4, I think there would be significant interest in adopting it.



Special Case: Optical Photons in ANNIE

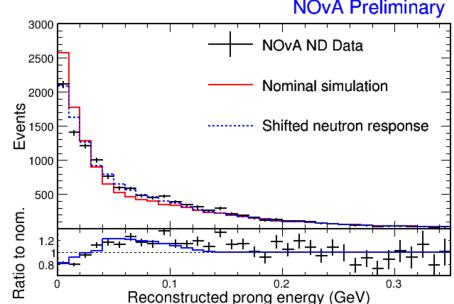
- ANNIE is a Water Cherenkov experiment, so photon tracking is vital to the simulation. They use 2 simulations:
- WCSim
 - Generic Geant4-based simulation package for WC detectors.
 - Uses full G4 optical photon tracking, though scaled down to detector QE to save computation time.
 - Currently the primary simulation for near-term analyses.
- RAT-PAC
 - Ray-tracing simulation developed for low energy (dark matter, $0\nu\beta\beta$ experiments).
 - Currently being used to study the eventual addition of water-based liquid scintillator, since this functionality is already available in this package.
- All simulation so-far is CPU-based, but obviously would be interested in GPU simulation if it becomes available.





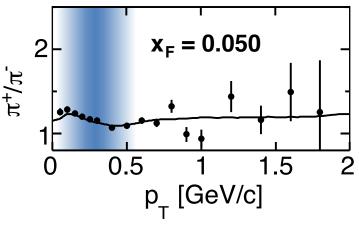
Looking Forward: Features

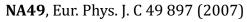
- (Largely my opinions)
- Improved neutron physics, or the ability to change models or adjust the behavior of the current model.
 - NOvA is putting a lot of effort into understanding the behavior of neutrons in our detector, and there are some clear discrepancies.
 - ANNIE's is explicitly measuring neutrons from neutrinos.
 - Useful both for tuning simulation to data, and for applying more physics-based systematic uncertainties.

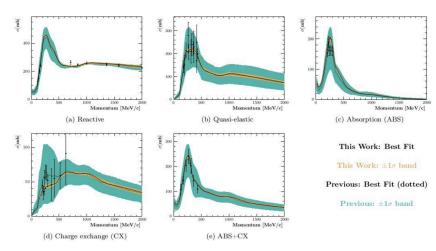


Looking Forward: Features

- Physics list designed for neutrino experiments.
 - Separate regimes: neutrino beam production, charged particles in detector, etc.
 - Would love to avoid model transitions in the middle of our energy regimes if we can.
 - If tuning made available to experiments, could tune against test beam experiments.
- Handles for creating hadronic interaction systematics.
 - Work is on-going now to introduce reweightable systematics for π^{\pm} interactions in the detector.
 - Especially important for cross-section measurements with exclusive final states.







P. Guerra et. al, PRD99, 052007 (2019)

Looking Forward: Performance

- Faster is obviously better, but how important speed is varies a lot with experiment and context.
 - Liquid argon:
 - The time spent in Geant4 is a relatively small part of the whole simulationreconstruction chain.
 - Signal generation and processing and reconstruction tasks take longer by an order of magnitude.
 - NOvA:
 - NOvA has recently significantly sped up reconstruction, so now simulation is the dominant part of the chain.
 - However, the stage of the experiment is such that it may not be well-suited to pick up major speed-ups.
- Memory
 - Memory use is a significant concern for everyone since the way "slots" are accounted in grid systems charge penalties for using > 2 GB of RAM.
 - Supporting multi-threading could help this by allowing the extra CPUs which come with higher memory requests to be put to use.

Looking Forward: Architectures

- It seems clear that future computing is going to requires us to be versatile about the kinds of computing resources we use.
 - High-throughput computing (e.g. Open Science Grid)
 - High-performance computing (e.g. NERSC, ORNL, etc.)
 - Co-processors (GPUs, FPGAs, etc.)
- Ideally, we'd like simulation to be able to use the advantages provided by different environments (e.g. GPUs if they're available), but we want to avoid *requiring* them.
 - This allows us to remain flexible, being able to assign the right sorts of tasks to the right sorts of resources, but not be stuck if a particular resource isn't available.
- Special case: If GPU-accelerated photon tracking were available, I think there would be significant interest in putting it to use ASAP.

Conclusions

- Many experiments, but there's similarity in how the simulation runs.
 - Doubly so in the LAr program thanks to a shared framework.
- So far, most simulation "attention" in neutrino expeirments has gone to flux and cross section simulation and systematics.
 - Flux is, in principle, a GEANT situation, but expectations for the hadronic modeling have been low and the focus is post-hoc reweighting to data.
 - IMHO, having that tuning work feed back into the underlying GEANT models would be broadly beneficial.
- But, as flux and cross section systematics come down, the detector systematics become more important.
 - Work is beginning on how to develop post-hoc hadronic interaction systematics.
 - However, "first party" systematics (reweightable or not) from GEANT would be a very useful guide.