

Constraint on the matter- antimatter symmetry-violating phase in neutrino oscillations by the T2K experiment

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Outline



- Neutrino Oscillations in a nutshell
- The T2K experiment
- Analysis procedure
- Cross-section analysis
- Results
- Next steps and beyond

Neutrino Oscillations in a nutshell

Neutrino oscillations



- Neutrino flavour eigenstates are not the same than the neutrino Lorentz eigenstates.
- Eigenstates are related through a rotation matrix.

Flavour eigenstates

$$(\nu_e, \nu_\mu, \nu_\tau)$$

state of the neutrino interactions

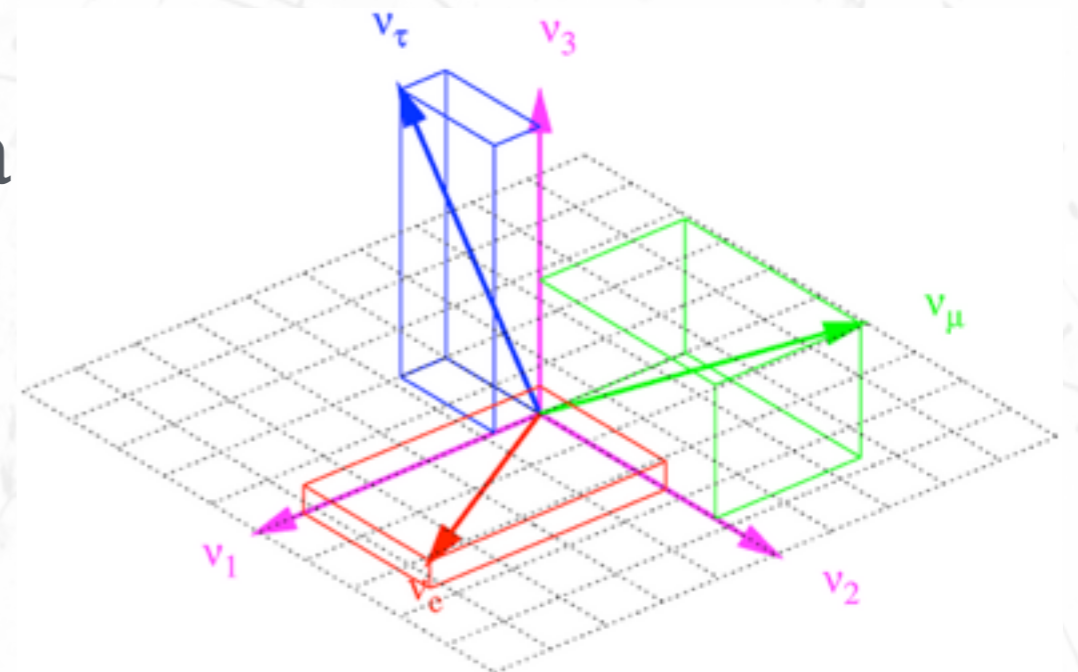
Lorentz eigenstates

$$(\nu_1, \nu_2, \nu_3)$$

states of the neutrino propagation in space

Pontecorvo–Maki–Nakagawa–Sakata
(PMNS) matrix

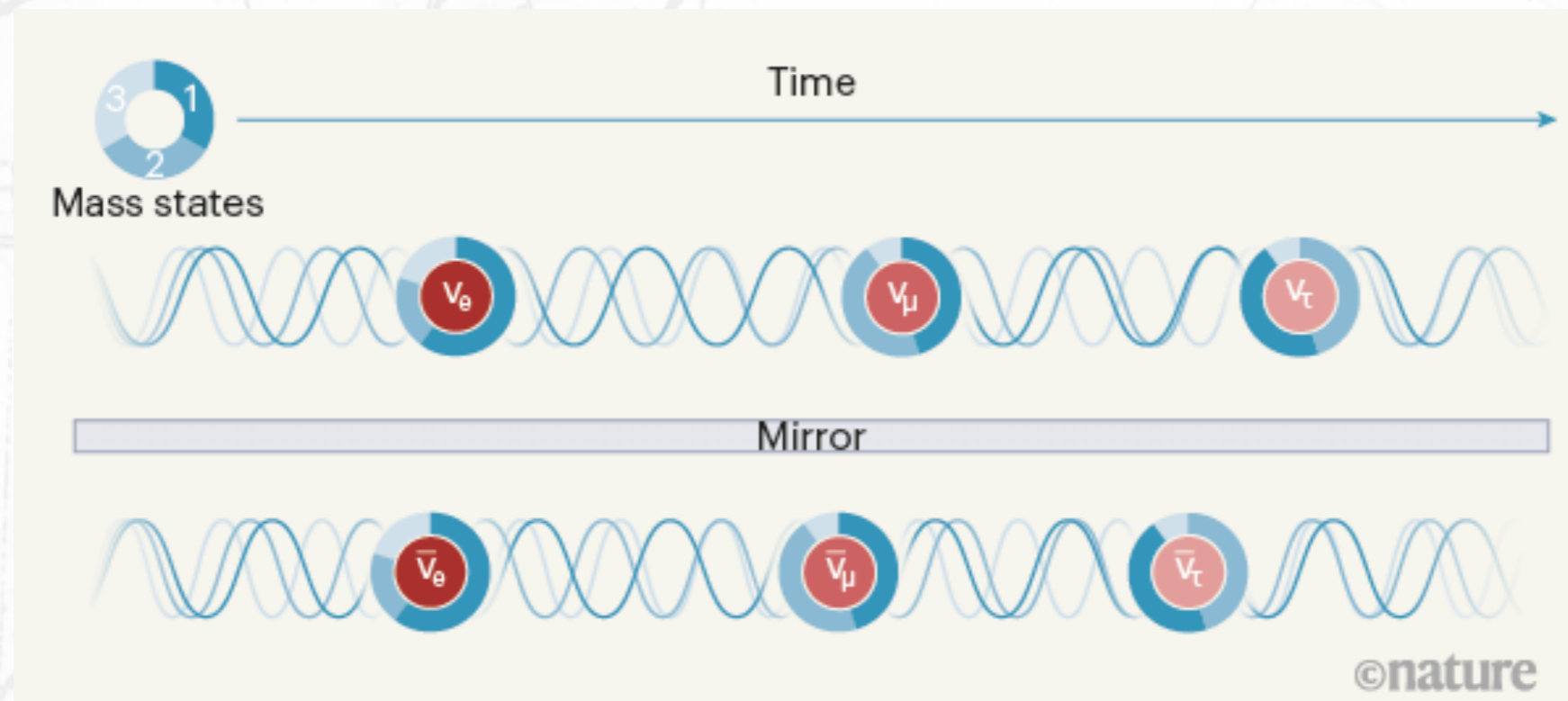
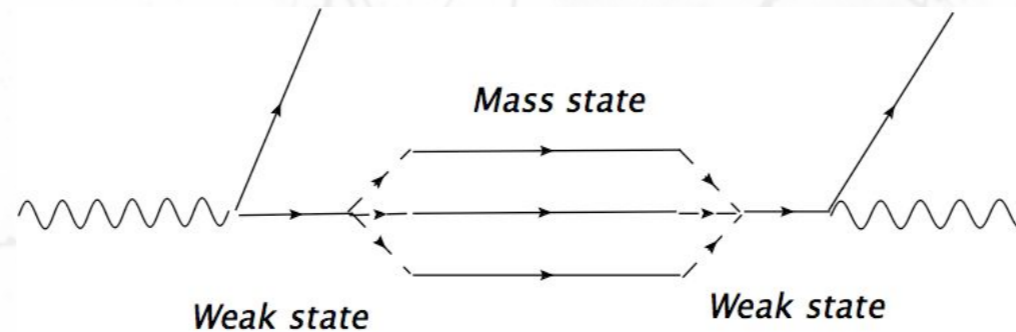
$$(\nu_e \quad \nu_\mu \quad \nu_\tau) = U_{PMNS} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$



Neutrino oscillations

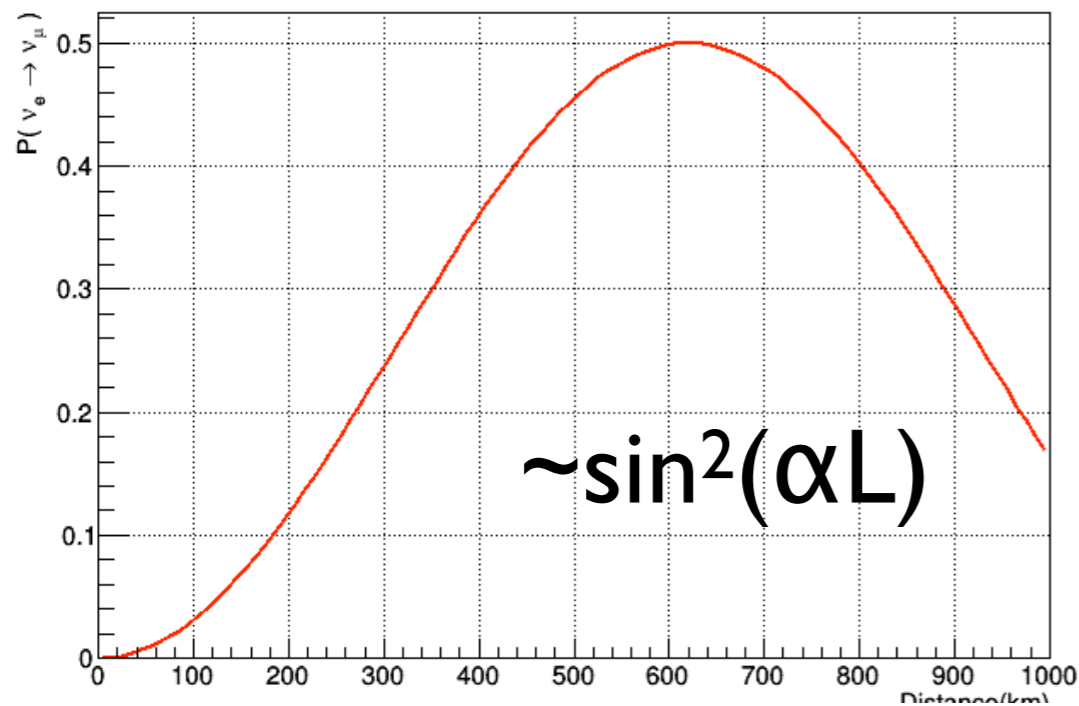


- Neutrinos are produced always as a flavour neutrino (electron, muon, tau) but they propagate in vacuum as mass states (they do not interact)



- Neutrinos propagate at different speeds (mass) keeping the coherence, at the interaction point the proportions change and other neutrino flavour might appear.

Oscillations with 2ν

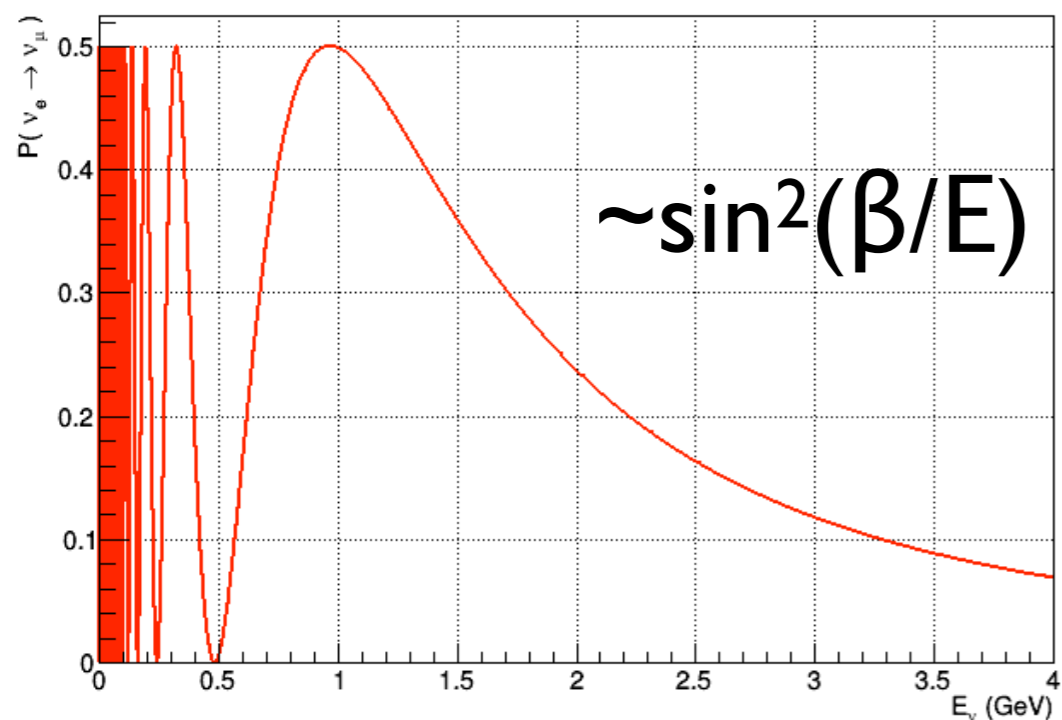


$$\theta = \pi/2$$

$$\Delta m^2 = 2. \times 10^{-3} \text{ eV}^2$$

Simplified
2ν formula

$$| \langle \nu_\mu | \nu_e; t \rangle |^2 = \sin^2 \frac{\theta}{2} \sin^2 1.267 \frac{\Delta m^2 L}{E} \frac{\text{GeV}}{\text{eV}^2 \text{ km}}$$



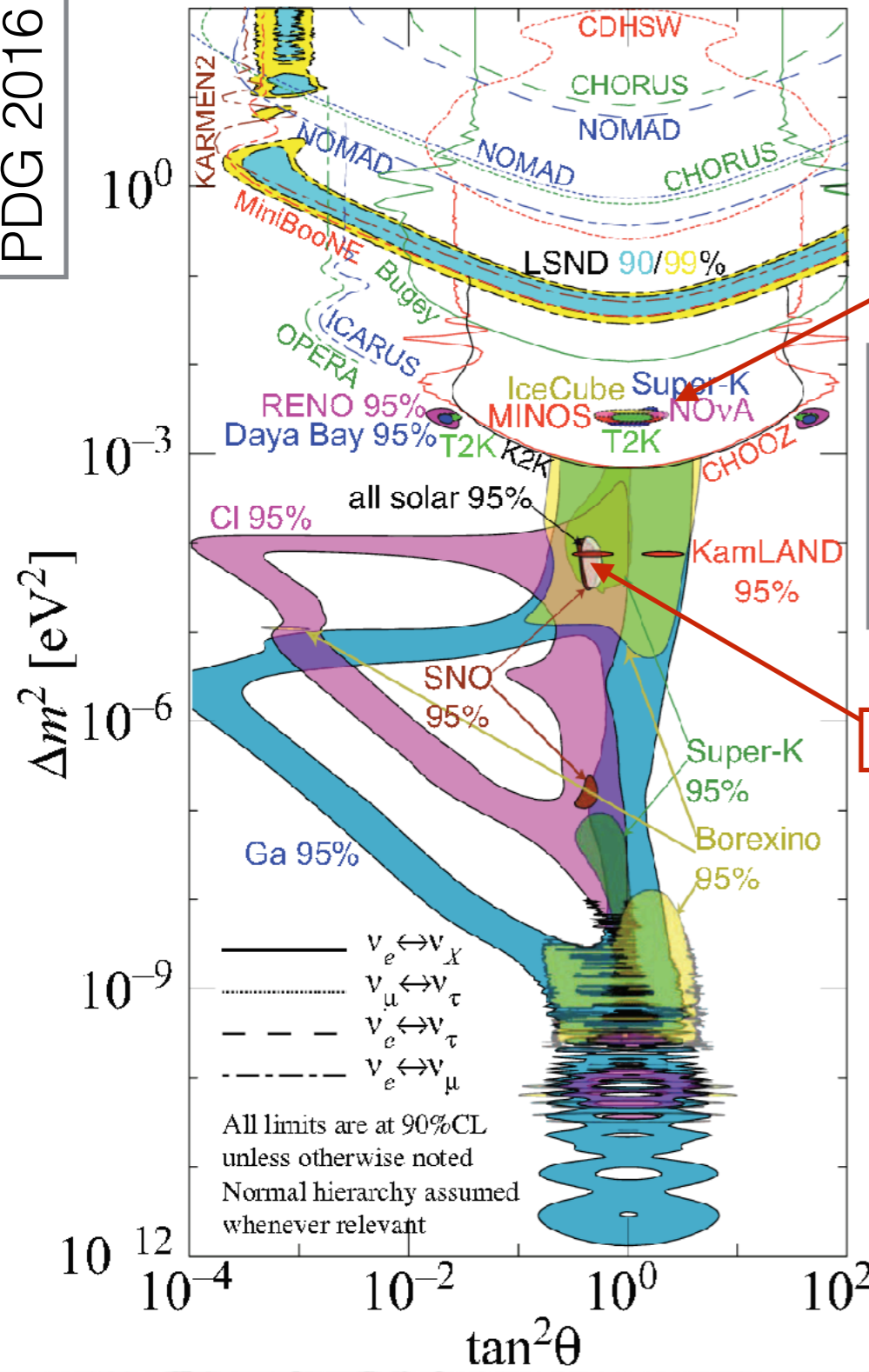
Oscillations are seen as
change of ν flavour
composition as function of:
Energy & Distance

$$P(\nu_\alpha \rightarrow \nu_\beta)$$

Oscillation parameters



PDG 2016



PNMS Matrix

$$U_{PNMS} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos \theta_{23} & \sin \theta_{23} \\ 0 & -\sin \theta_{23} & \cos \theta_{23} \end{pmatrix} \begin{pmatrix} \cos \theta_{13} & 0 & e^{-i\delta_{CP}} \sin \theta_{13} \\ 0 & 1 & 0 \\ -e^{i\delta_{CP}} \sin \theta_{13} & 0 & \cos \theta_{13} \end{pmatrix} \begin{pmatrix} \cos \theta_{12} & \sin \theta_{12} & 0 \\ -\sin \theta_{12} & \cos \theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

PDG 2018

atmospheric

solar

Parameter	best-fit	3σ
~4% $\Delta m_{21}^2 [10^{-5} \text{ eV}^2]$	7.37	6.93 – 7.96
~3% $\Delta m_{31(23)}^2 [10^{-3} \text{ eV}^2]$	2.56 (2.54)	2.45 – 2.69 (2.42 – 2.66)
~11% $\sin^2 \theta_{12}$	0.297	0.250 – 0.354
~15% $\sin^2 \theta_{23}, \Delta m_{31(32)}^2 > 0$	0.425	0.381 – 0.615
~15% $\sin^2 \theta_{23}, \Delta m_{32(31)}^2 < 0$	0.589	0.384 – 0.636
~7% $\sin^2 \theta_{13}, \Delta m_{31(32)}^2 > 0$	0.0215	0.0190 – 0.0240
~7% $\sin^2 \theta_{13}, \Delta m_{32(31)}^2 < 0$	0.0216	0.0190 – 0.0242
~31% δ/π	1.38 (1.31)	2σ: (1.0 - 1.9) (2σ: (0.92-1.88))

Most of the parameters measured with <10% precision
 θ_{23} is known with 15% precision
 Remaining parameters are δ_{CP} and the hierarchy

Mass hierarchy



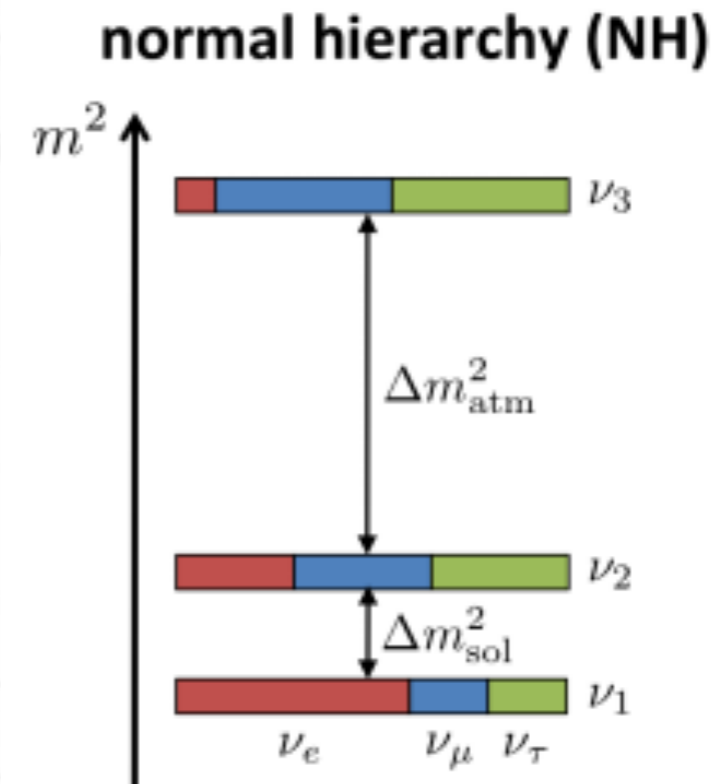
- Oscillations is a quantum interference phenomenon that depends on the (quadratic) mass difference:

$$\Delta m^2_{ij} = m^2_i - m^2_j$$

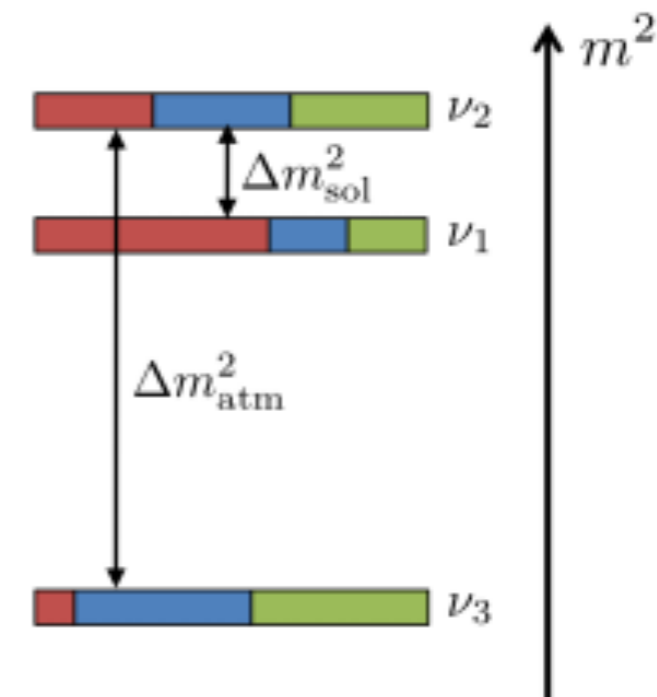
- Due to matter effects in solar neutrinos we know:

$$\Delta m^2_{12} > 0$$

- Hierarchy determines the ordering of the masses. Traditionally:
 - Normal: $m_1 < m_2 < m_3$
 - Inverted: $m_3 < m_1 < m_2$



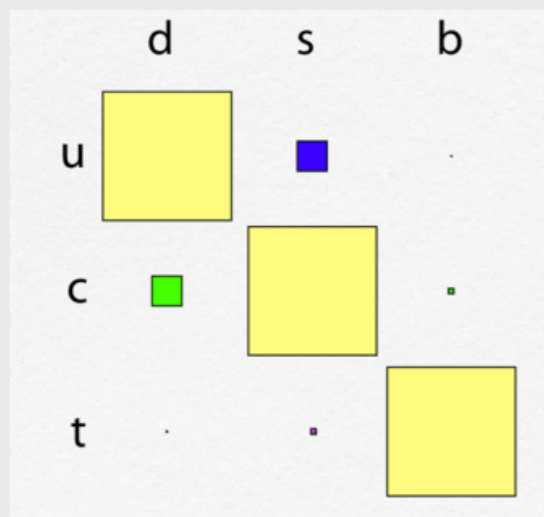
inverted hierarchy (IH)



PNMS vs CKM

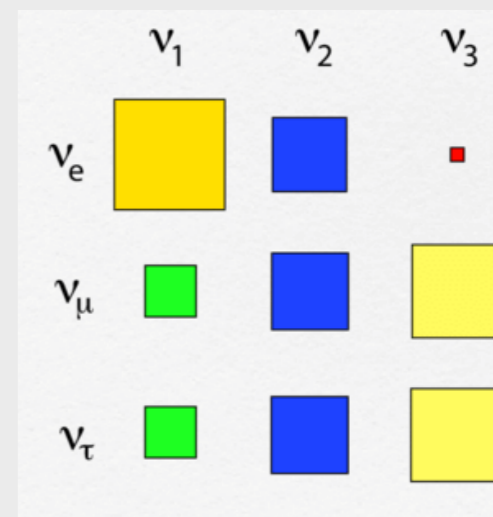


CKM



Mostly diagonal

PNMS



Flat"ish"

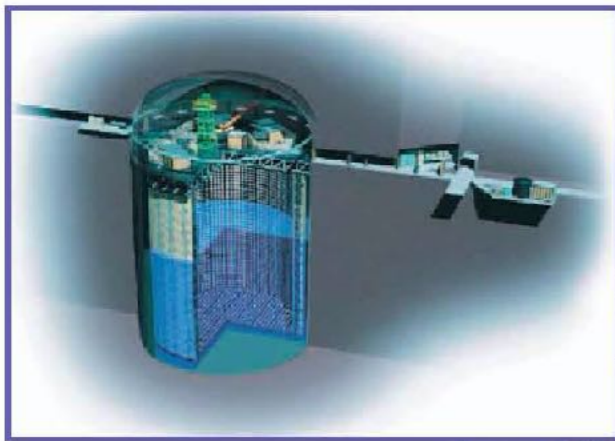
$$\begin{pmatrix} 1 - \lambda^2/2 & \lambda & A\lambda^3(\rho - i\eta) \\ -\lambda & 1 - \lambda^2/2 & A\lambda^2 \\ A\lambda^3(1 - \rho - i\eta) & -A\lambda^2 & 1 \end{pmatrix}$$

Unitarity is not enforced by construction

$$\begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos \theta_{23} & \sin \theta_{23} \\ 0 & -\sin \theta_{23} & \cos \theta_{23} \end{pmatrix} \begin{pmatrix} \cos \theta_{13} & 0 & e^{-i\delta_{CP}} \sin \theta_{13} \\ 0 & 1 & 0 \\ -e^{i\delta_{CP}} \sin \theta_{13} & 0 & \cos \theta_{13} \end{pmatrix} \begin{pmatrix} \cos \theta_{12} & \sin \theta_{12} & 0 \\ -\sin \theta_{12} & \cos \theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

Unitarity enforced by construction following $N_\nu=3$ from LEP

T2K experiment



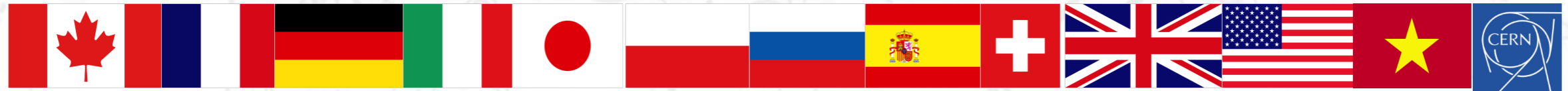
Super-Kamiokande
(ICRR, Univ. Tokyo)



J-PARC Main Ring
(KEK-JAEA, Tokai)



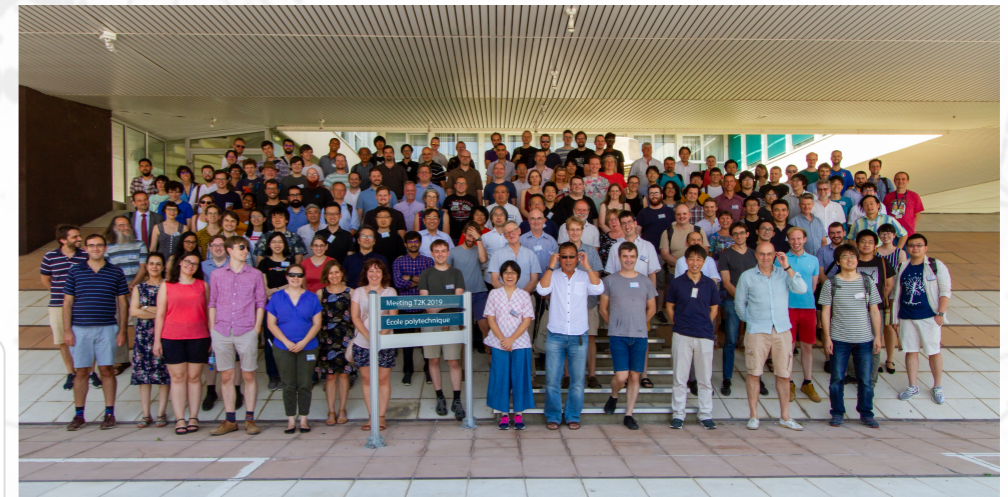
T2K Collaboration



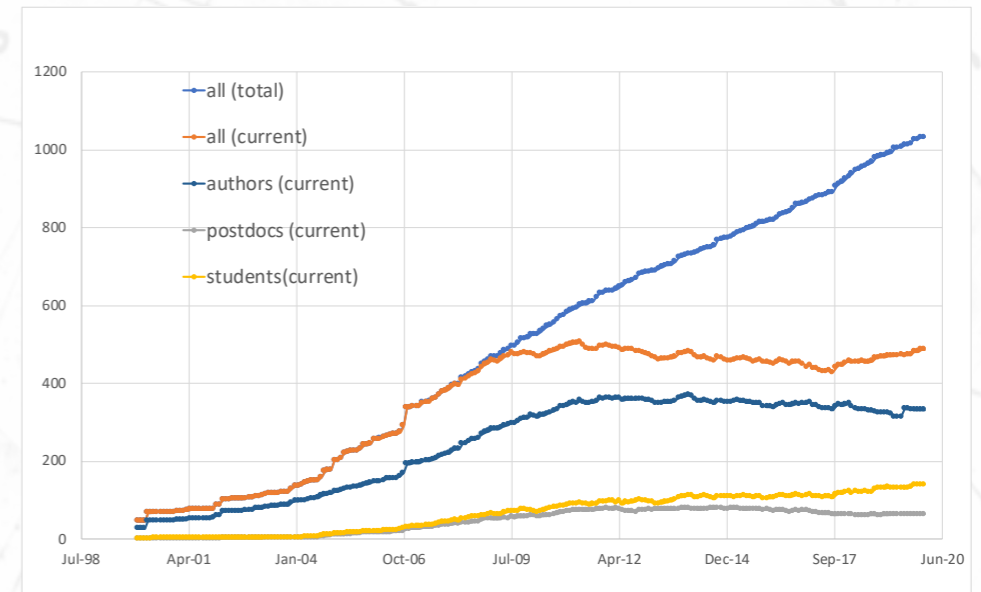
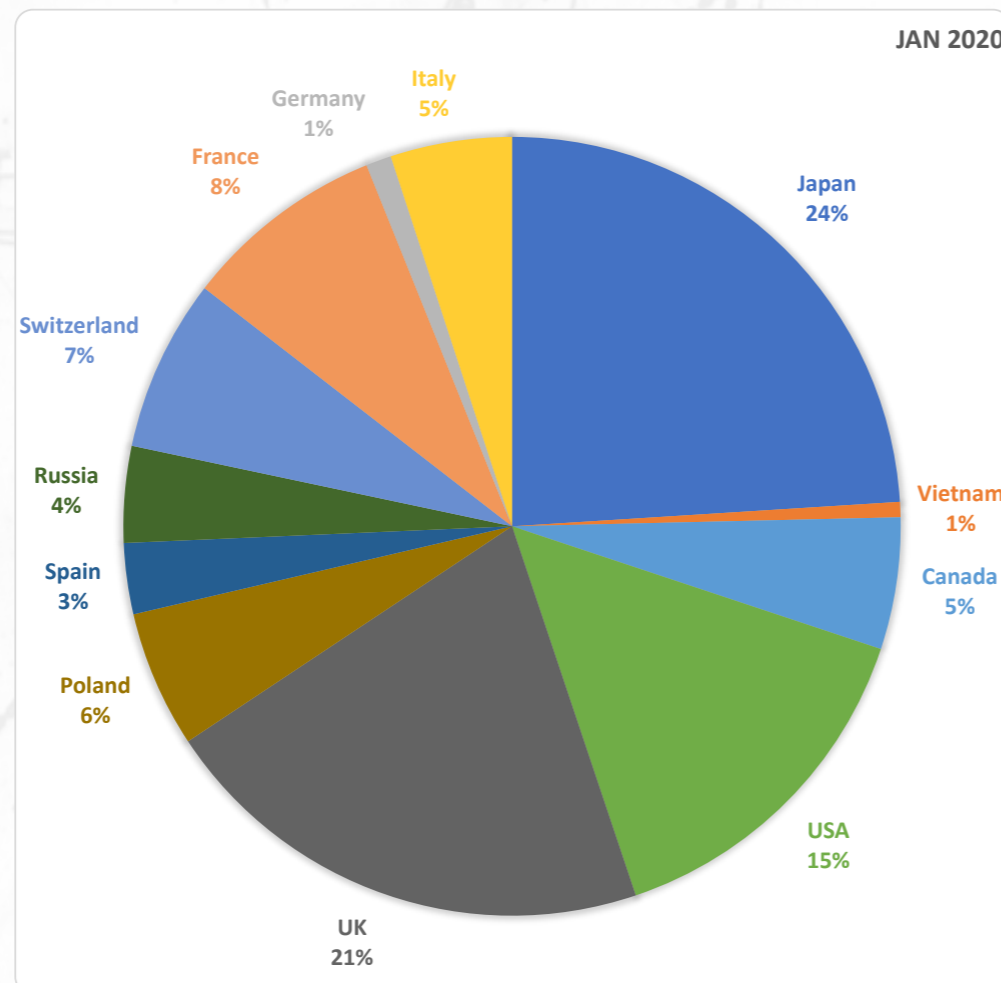
~500 members, 69 Institutes, 12 countries

Asia	117
Japan	114
Vietnam	3

Americas	96
Canada	26
USA	70

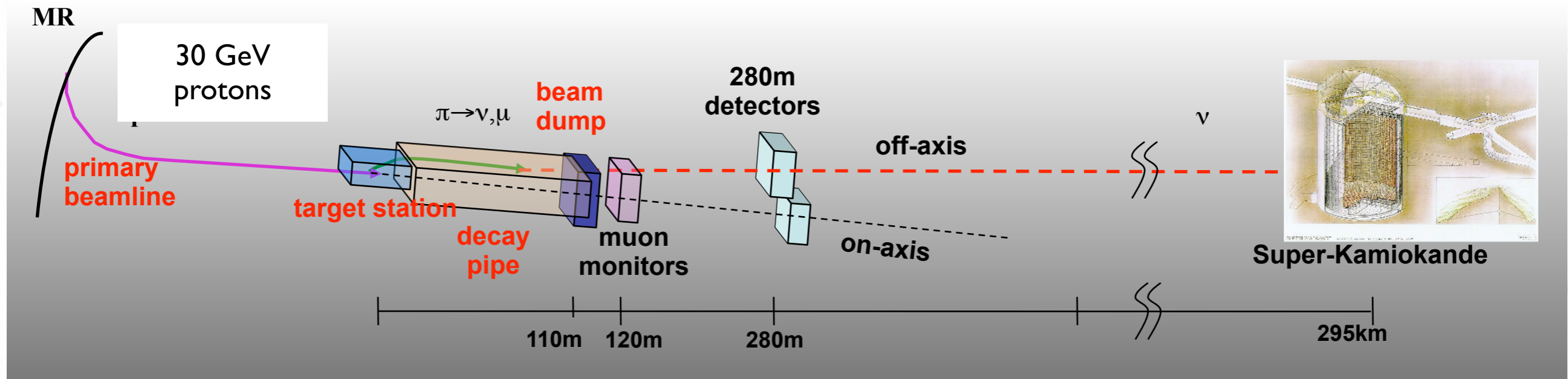


Europe	262
France	40
Germany	5
Italy	24
Poland	27
Russia	19
Spain	14
Switzerland	34
UK	99



Very strong European contribution including CERN

T2K experiment



Neutrinos produced in a particle accelerators or nuclear reactors.

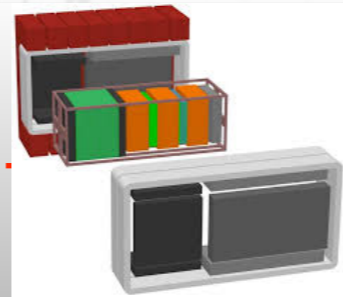
Neutrino flux properties

ν oscillations

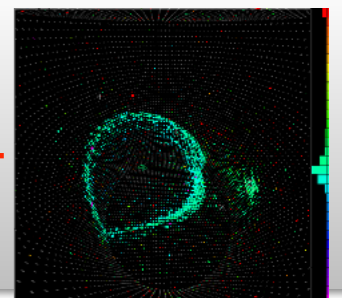
Neutrino flux & flavour



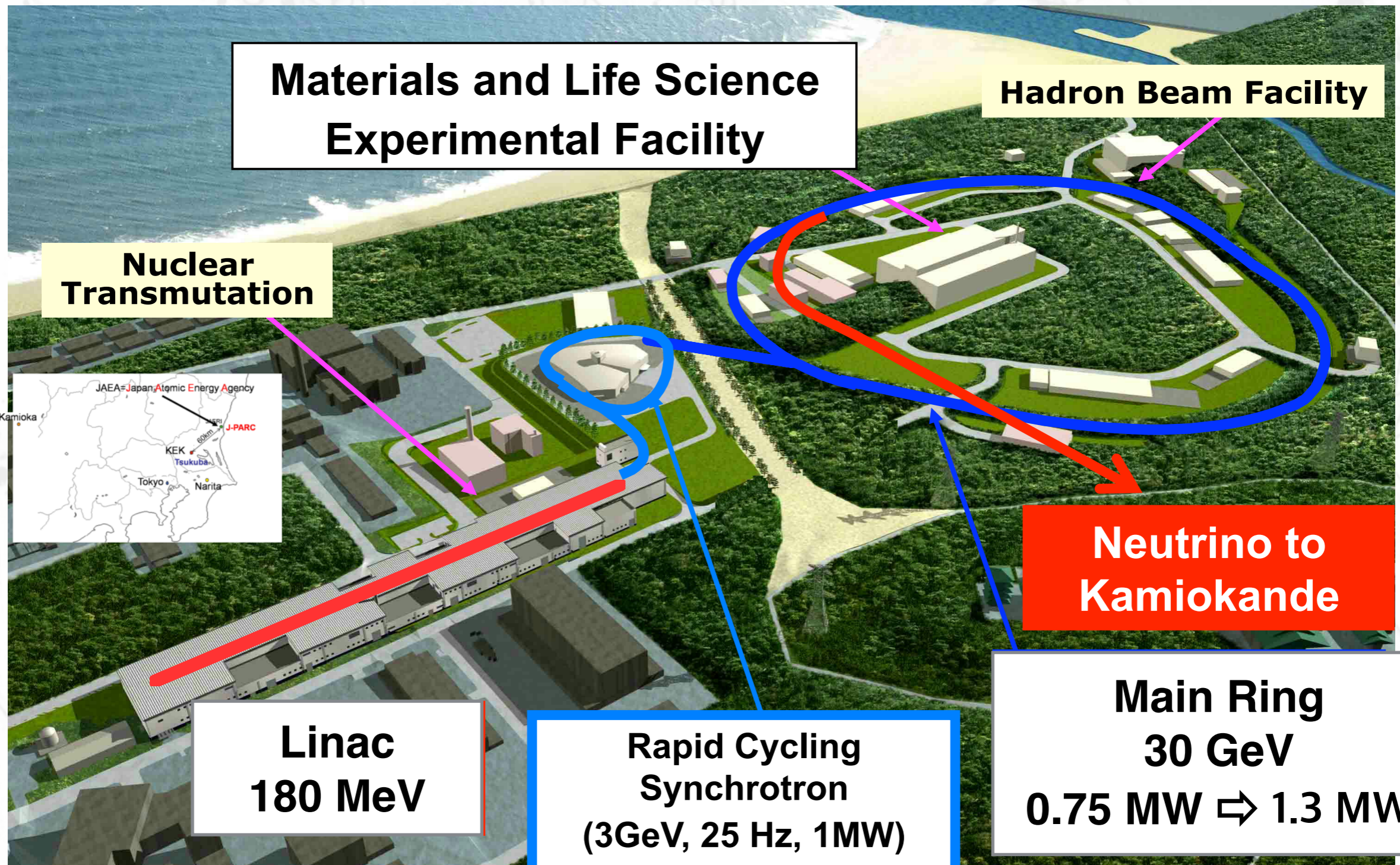
$$\nu_{\mu}, \bar{\nu}_{\mu}$$



$$\begin{aligned} &\nu_e \bar{\nu}_e \\ &\ll \\ &\nu_{\mu} \bar{\nu}_{\mu} \end{aligned}$$



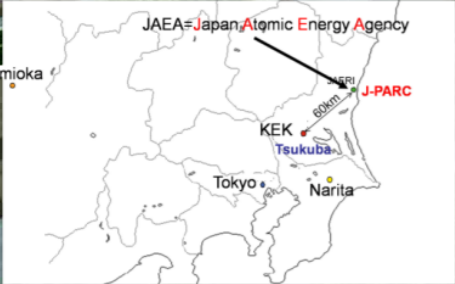
JPARC



**Materials and Life Science
Experimental Facility**

Hadron Beam Facility

**Nuclear
Transmutation**



**Neutrino to
Kamiokande**

**Linac
180 MeV**

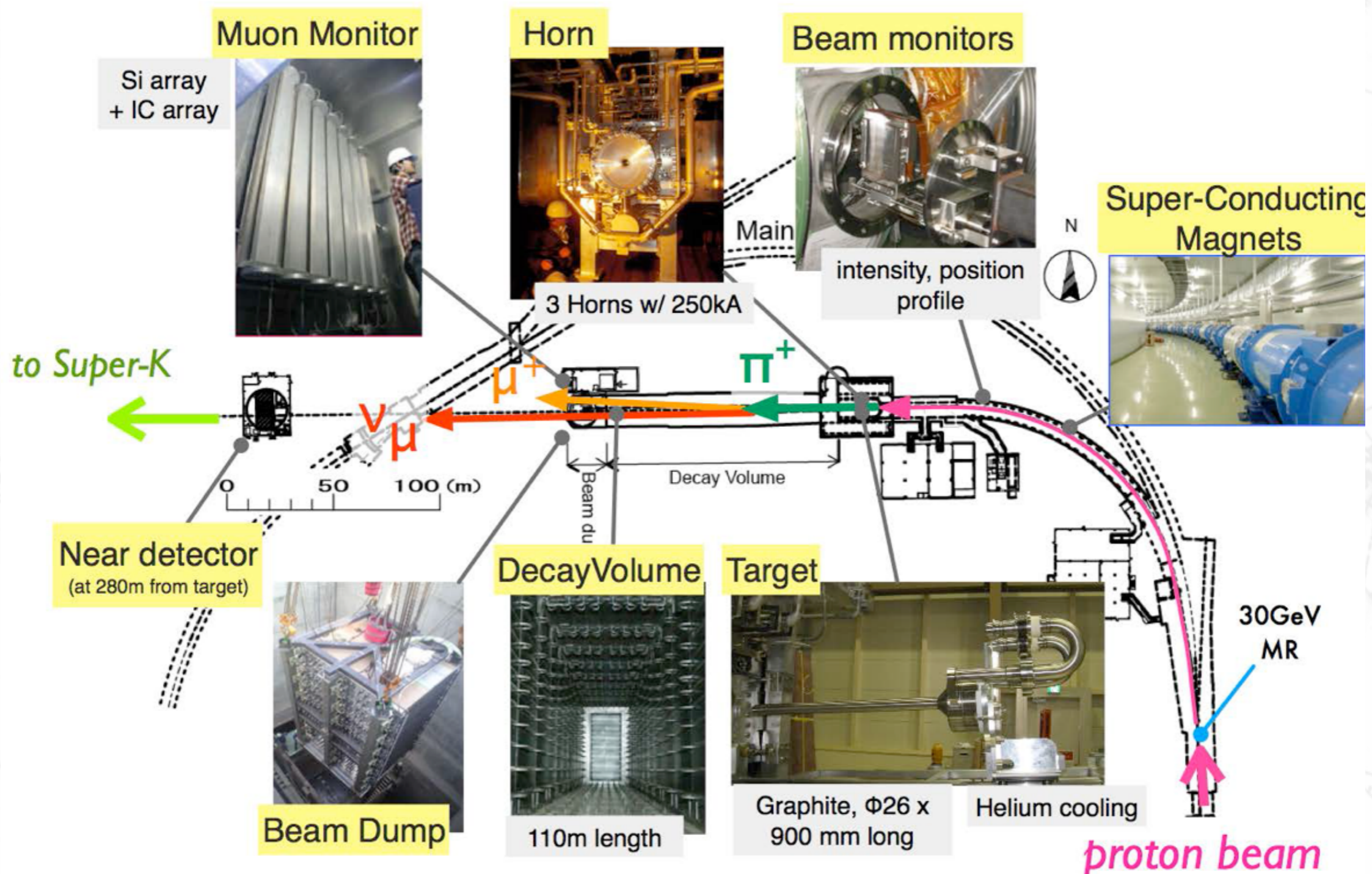
**Rapid Cycling
Synchrotron
(3GeV, 25 Hz, 1MW)**

**Main Ring
30 GeV
0.75 MW ⇔ 1.3 MW**

J-PARC = Japan Proton Accelerator Research Complex

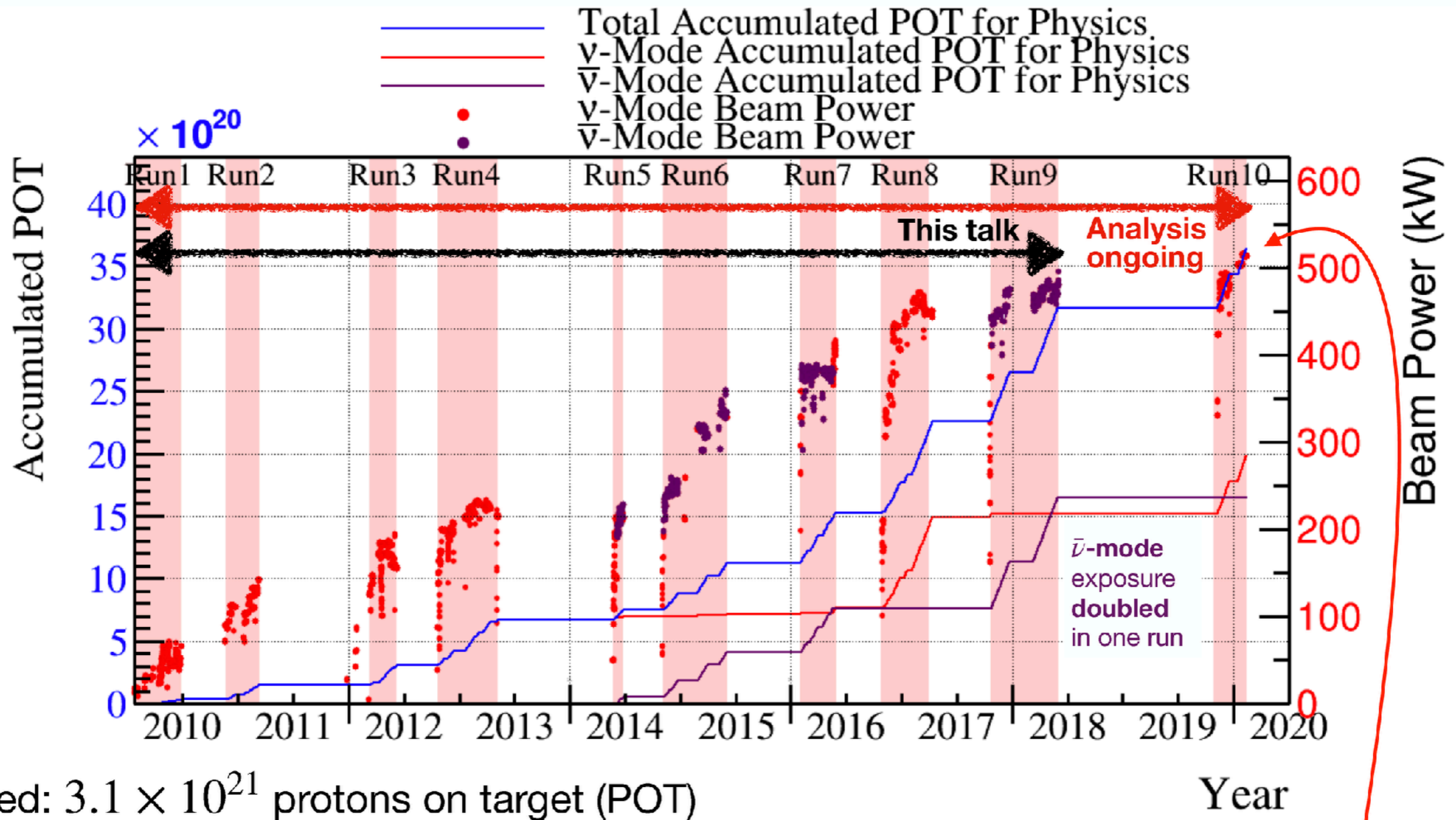
Joint Project between KEK and JAEA

ν beam @ JPARC



Fast extraction with beam pulse every 2 sec.

Data Set



Analyzed: 3.1×10^{21} protons on target (POT)

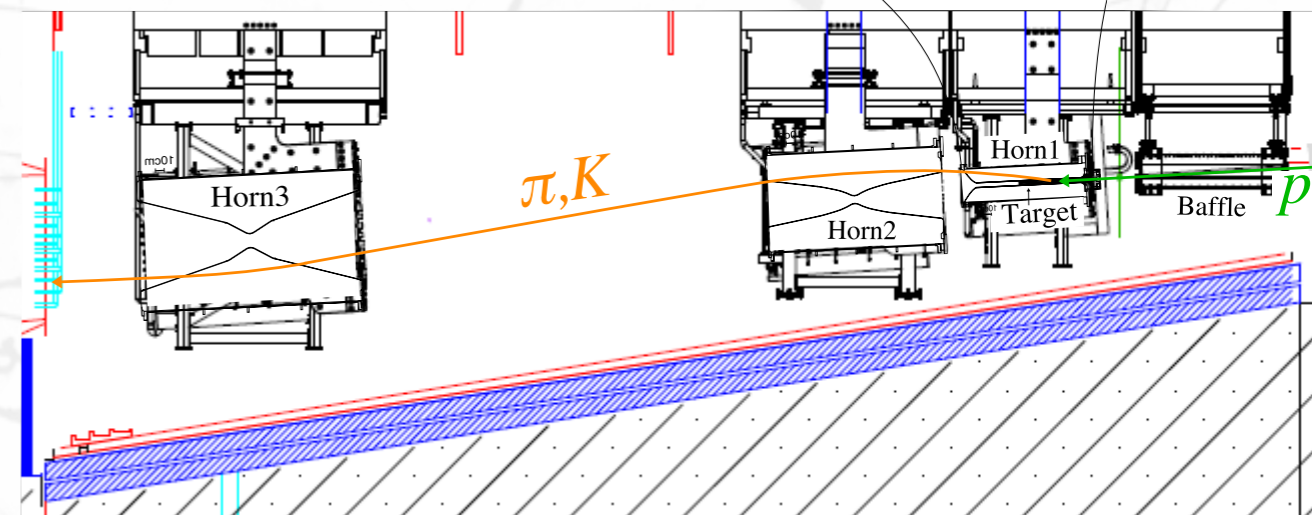
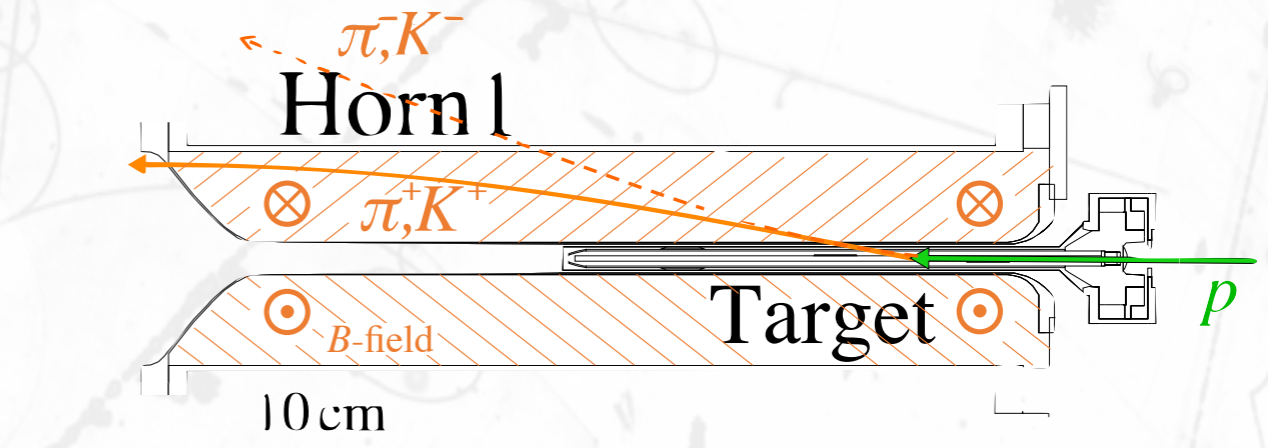
approx 50% ν - 50% $\bar{\nu}$

515 kW stable operation in 2019
 + 33% of ν -mode for next analysis

Neutrino beam

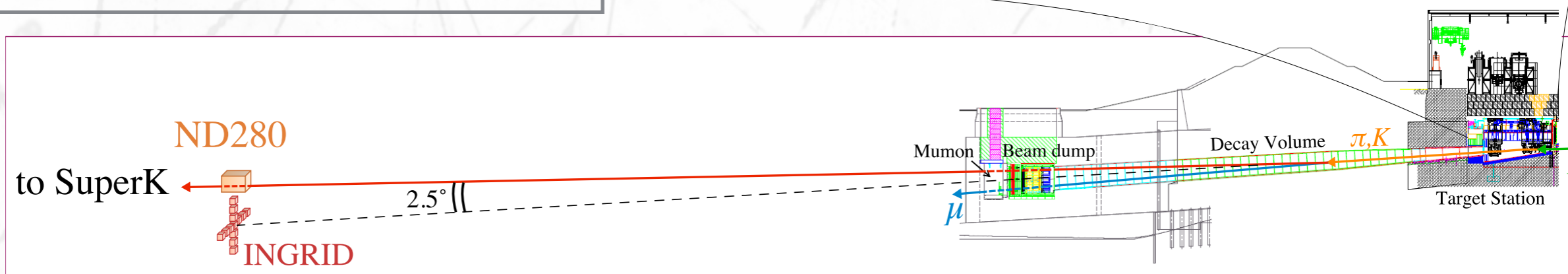


- 3 Horns system with 250 kA current sinusoidal $\sim 3\text{ms}$ pulse.
- Forward (neutrino enhanced) and Reversed (anti-neutrino enhanced) modes.
- The beam is slightly tilted towards the earth.

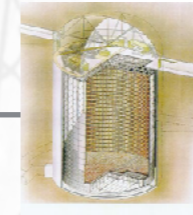
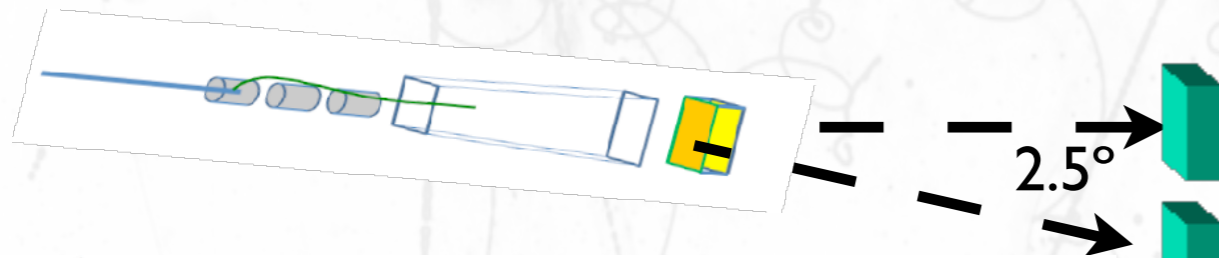


planned upgrade to reach 320kA

→ + $\sim 20\%$ ν flux

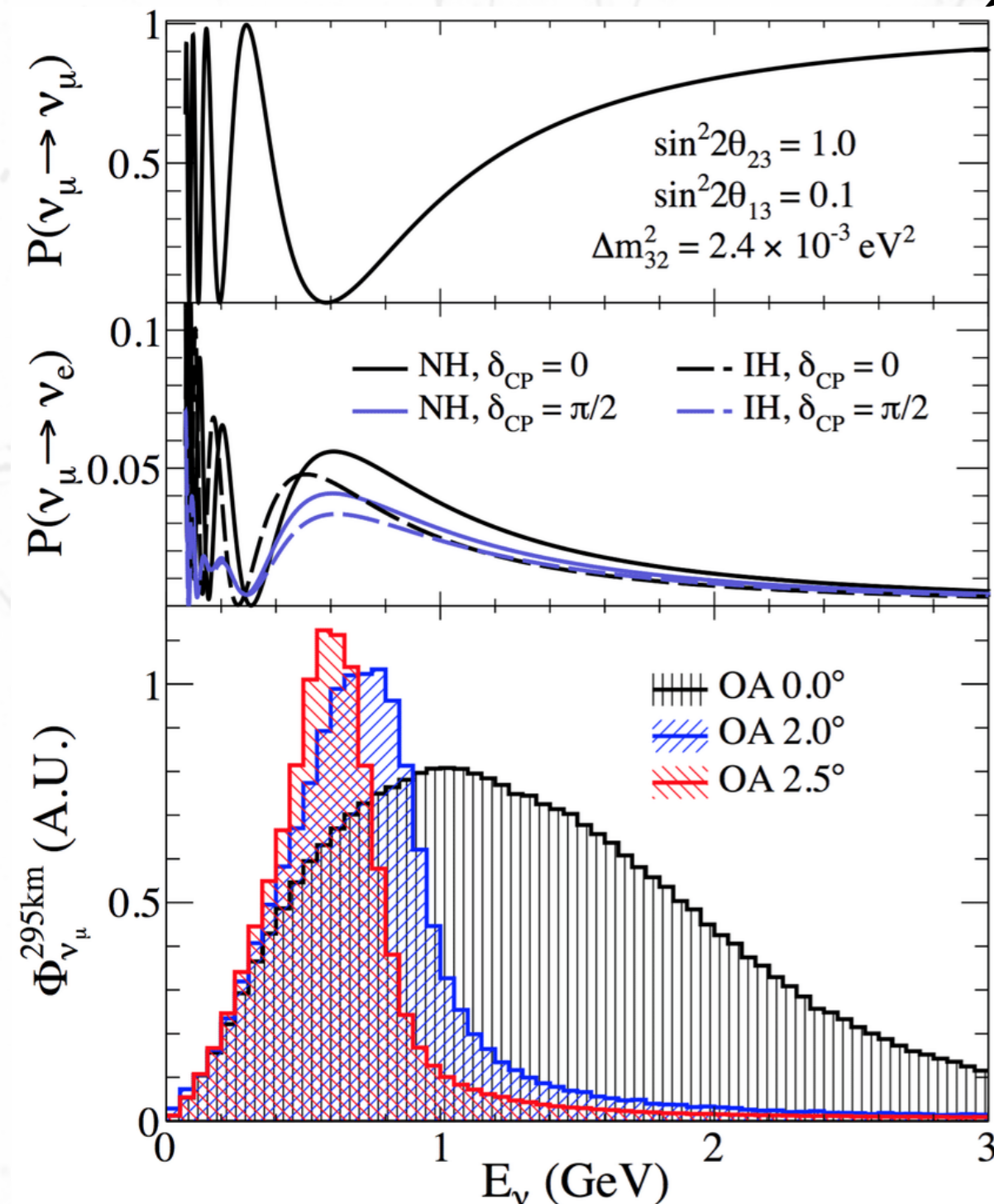


Off-axis beam



Off-axis

- off-axis optimises the flux at the maximum of the oscillation.
- Only one oscillation maximum can be measured at a fixed distance.
- Narrow beam less dependent on beam uncertainties but more on beam pointing.
- Lower energies achieved.

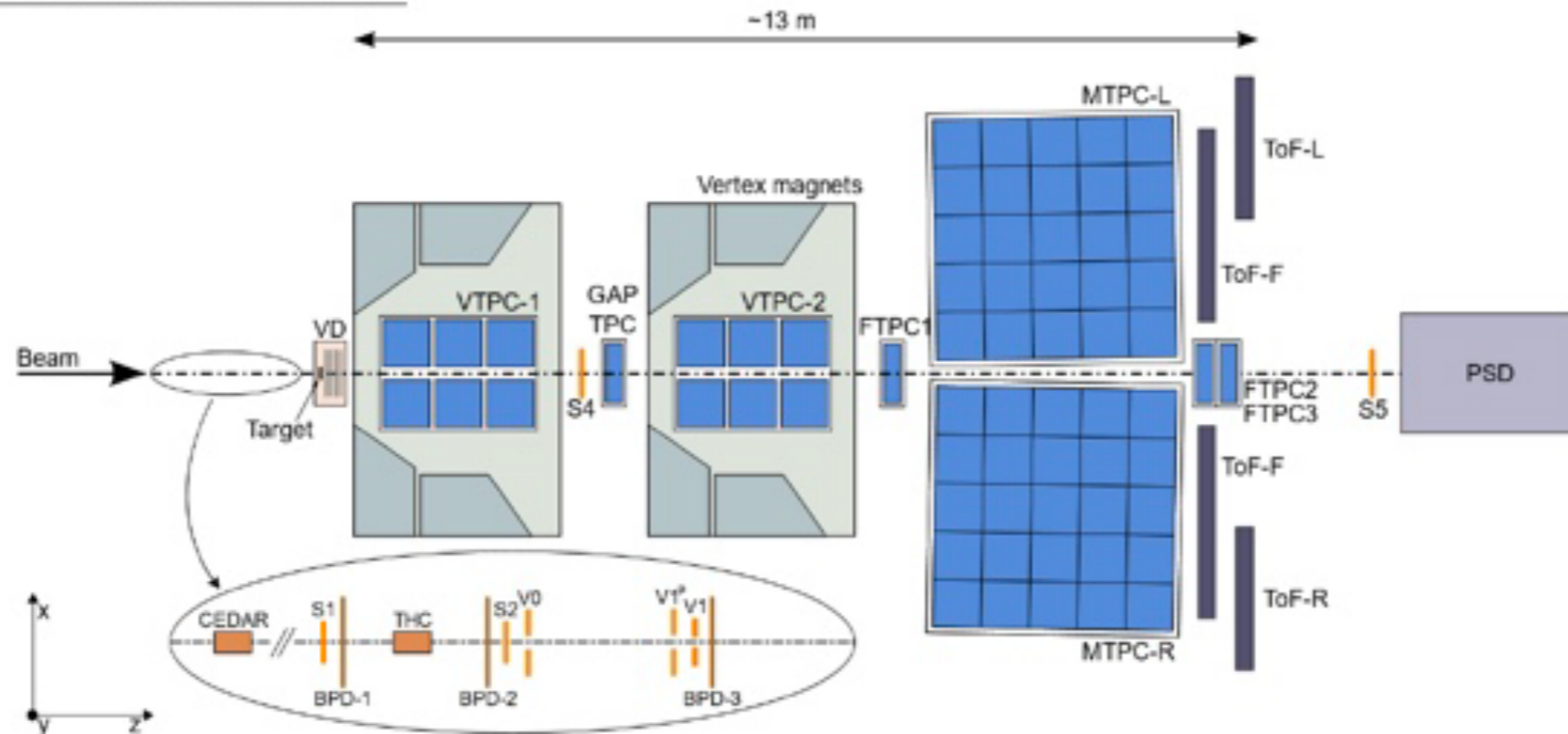


On-axis

- on-axis optimises the total integrated flux.
- Spectrum with higher neutrino energy (longer oscillation distances)
- If broad enough, more than one oscillation maximum can be measured at a fixed distance.

SPS Heavy Ion and Neutrino Expt (SHINE)

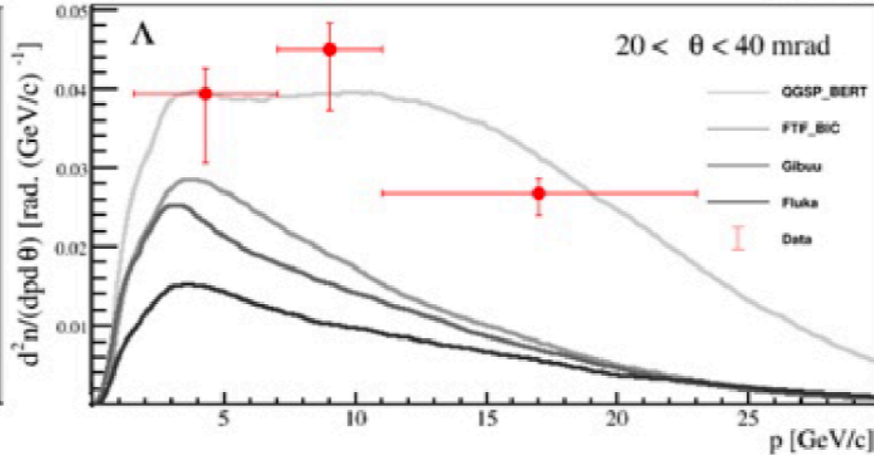
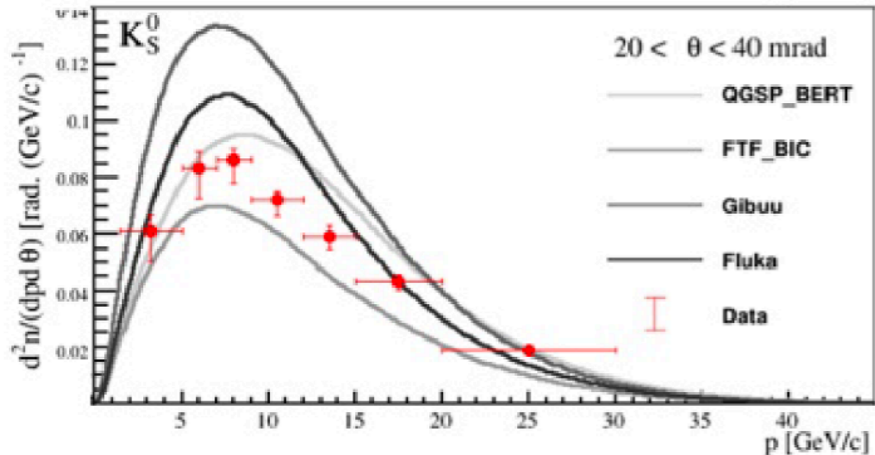
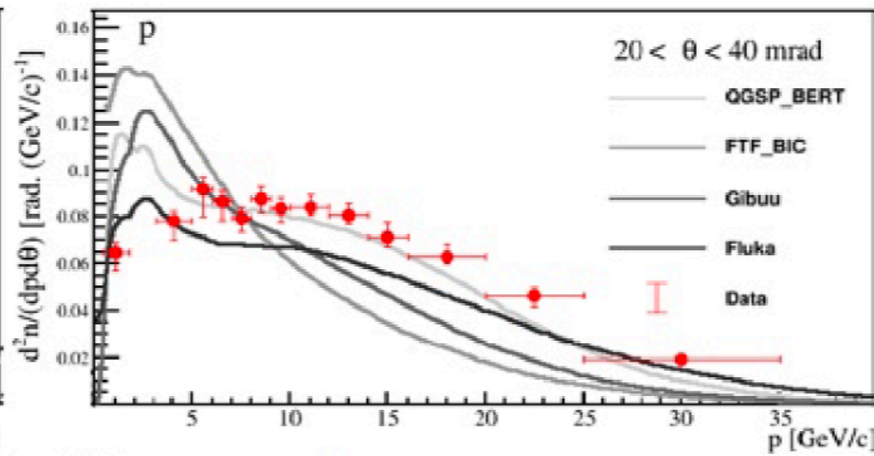
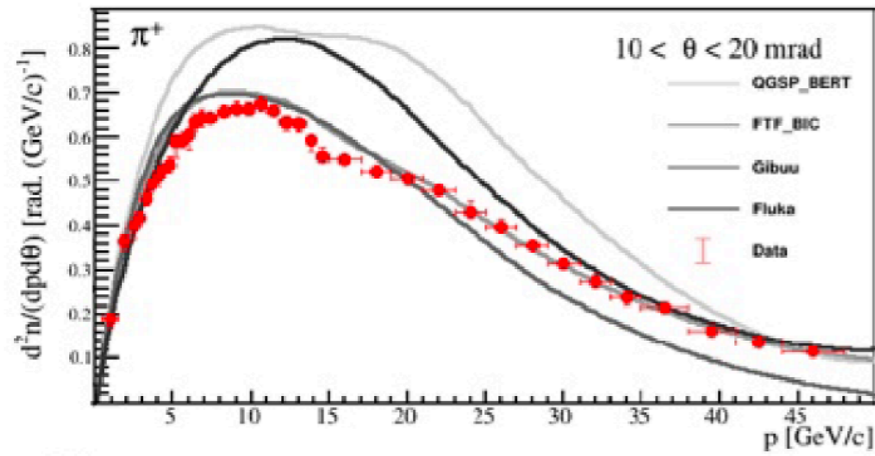
NA61/SHINE



NA61/Shine measures the production of pions and kaons as function of the momentum and angle for protons interacting with carbon.

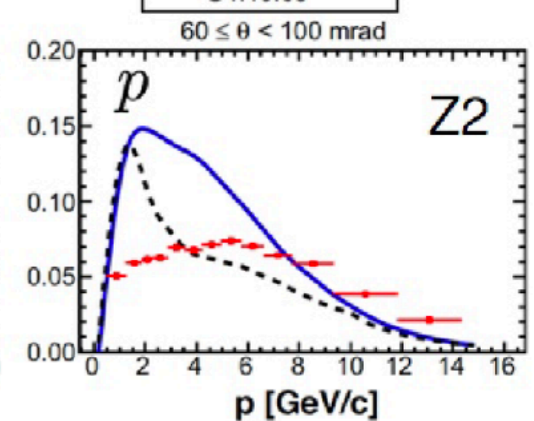
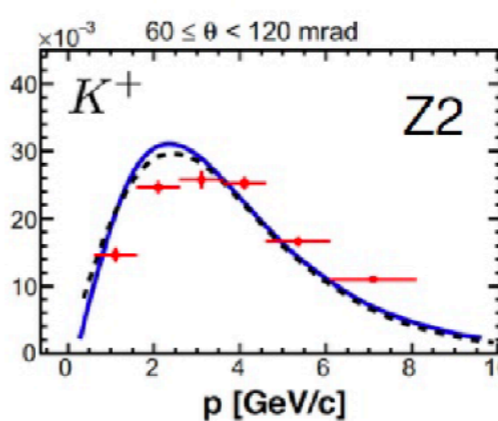
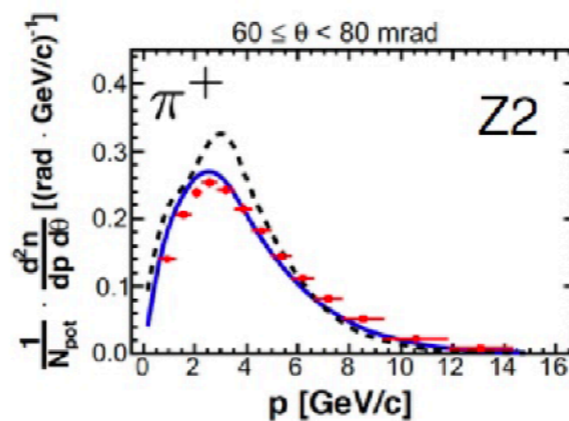
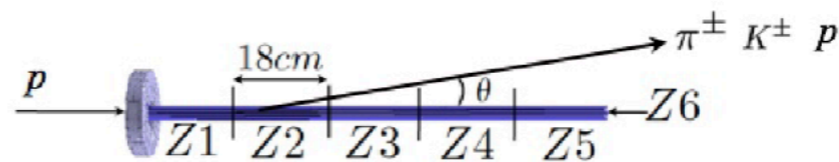
Hadro production experiments carried in equal conditions to ν beam experiments are critical!

Latest measurements made with exact T2K replica target



Measurement of pions, proton, K and Lambdas in a 30 GeV proton-Carbon interactions

Measurement of production with exact replica of the target to account for re-interactions inside the target



+ NA61/SHINE
 — NuBeam G4.10.03
 - - - QGSP_BERT G4.10.03

Beam monitors



Proton beam monitors are essential for protecting beam-line equipment, as well as for understanding and predicting the **neutrino flux**

Beam intensity

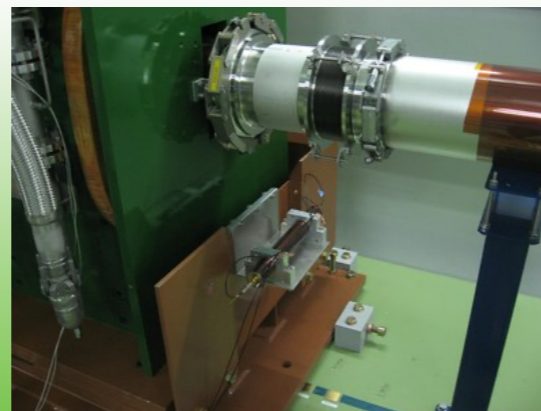
Current transformers

<2.7% precision



Beam Loss

Sensitive down to 16mW loss



Beam position and profile

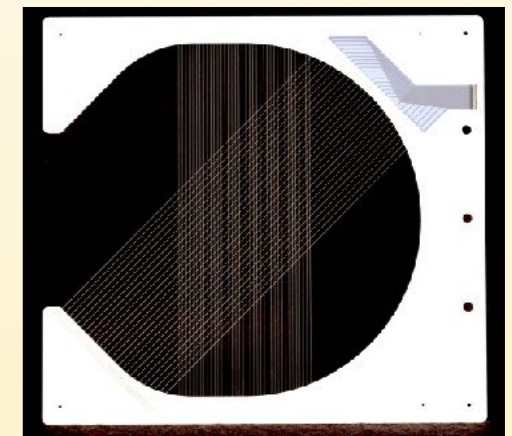
Segmented Secondary Emission Monitors

100 μ m position
200 μ m width



Wire Secondary Emission Monitors

100 μ m position
200 μ m width



Beam position

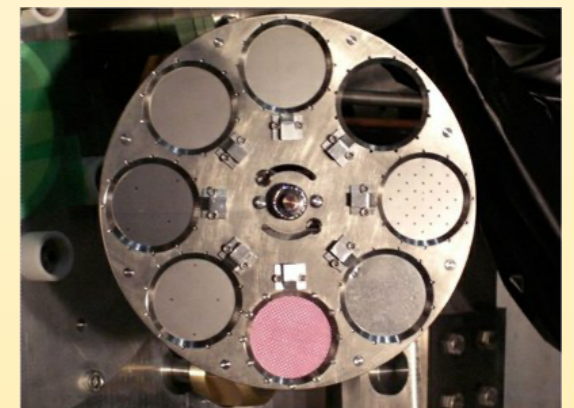
Electro Static monitors

450 μ m precision

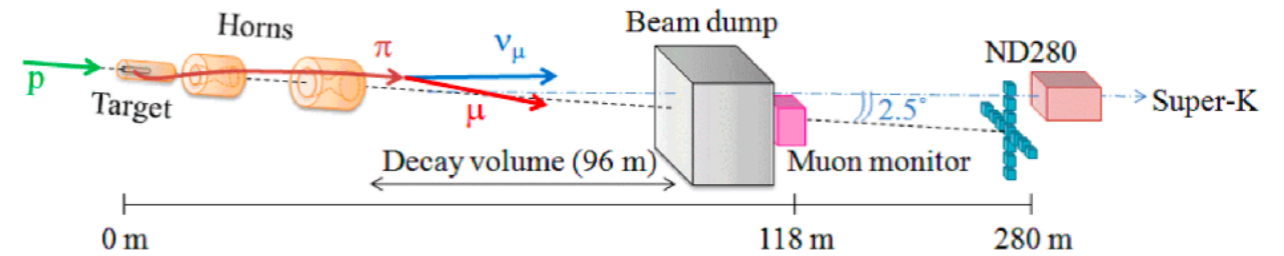


Optical transition radiation

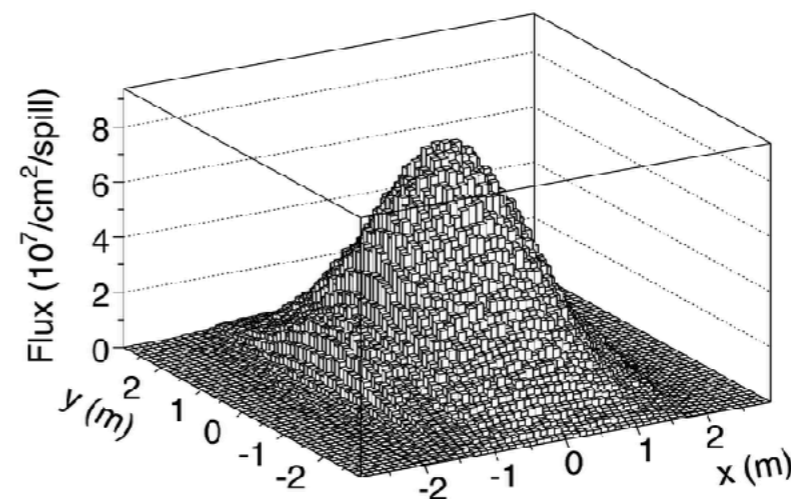
<500 μ m precision



Muon monitors

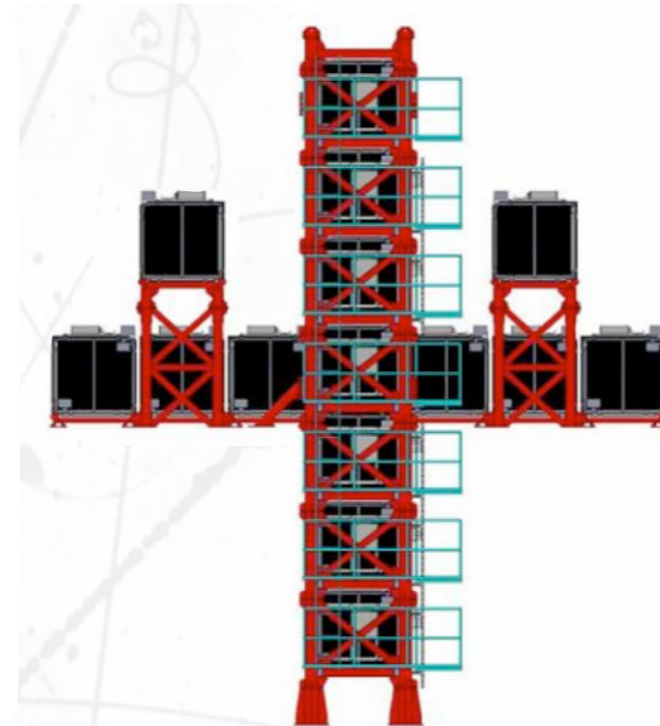
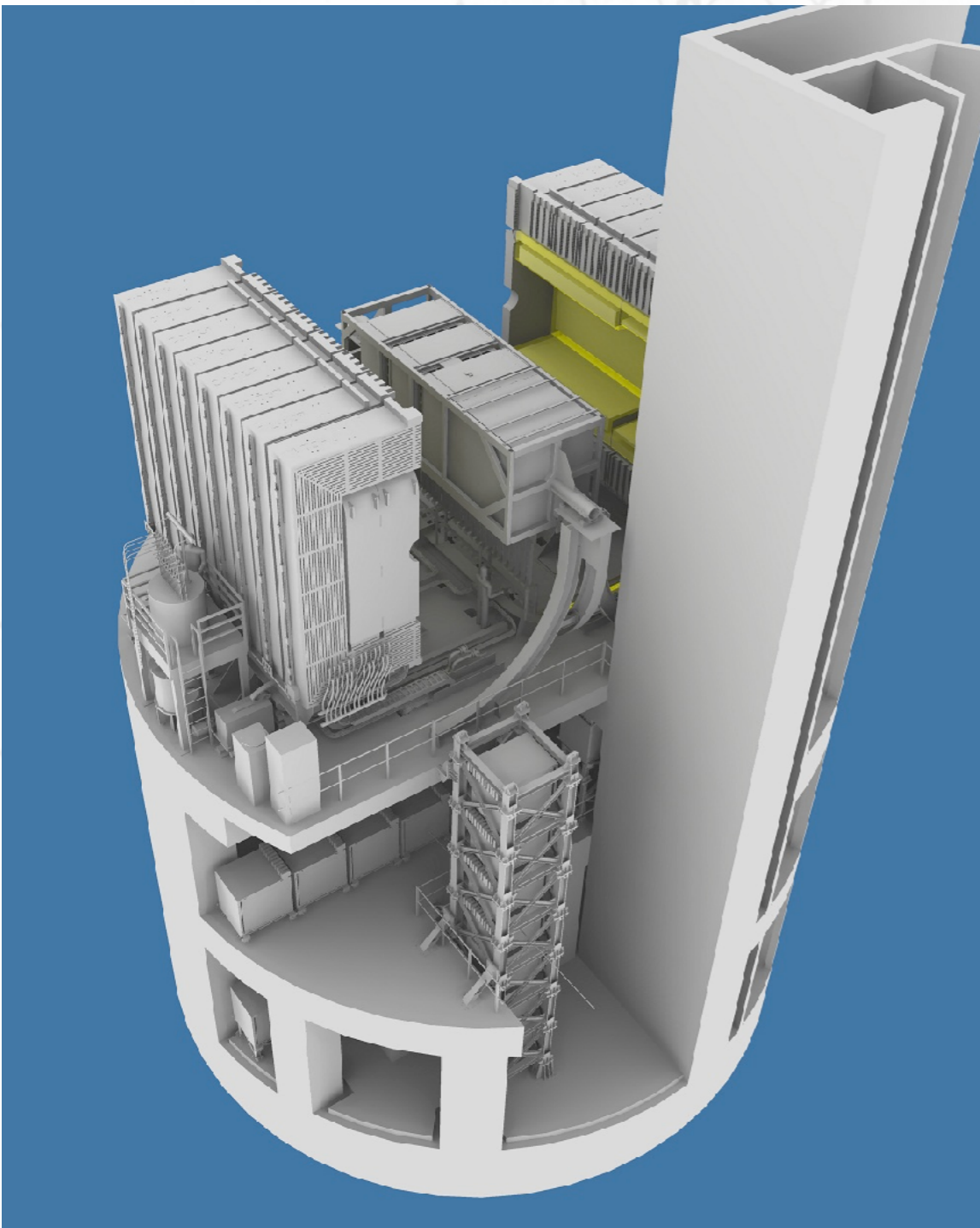


- Monitors the beam direction from the μ produced in π decays.
- Embedded in the beam dump samples the high energy muons.
- ionisation chambers and silicon PIN diodes.
- High irradiation area: $\sim 10^{14}$ electrons/cm/month at 750 KW.



Simulation
of fluence

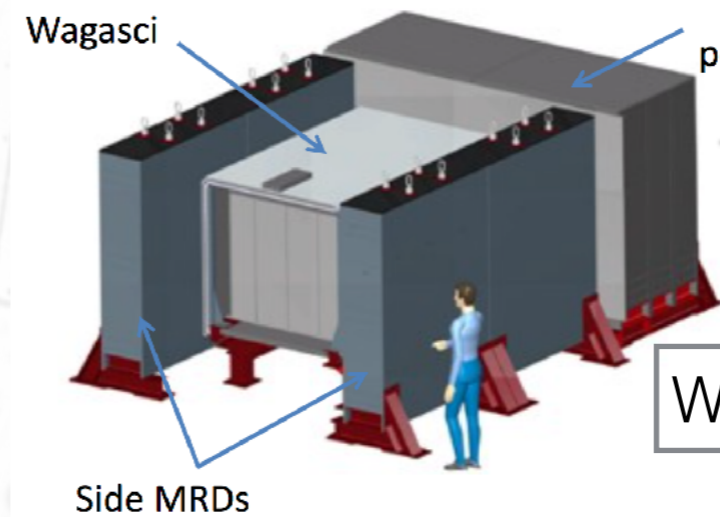
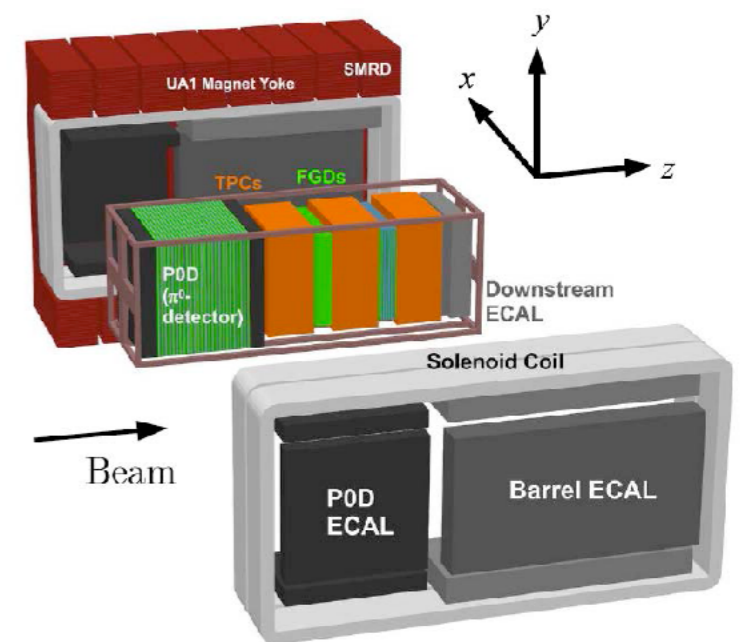
Near Detector Site



INGRID: On-axis



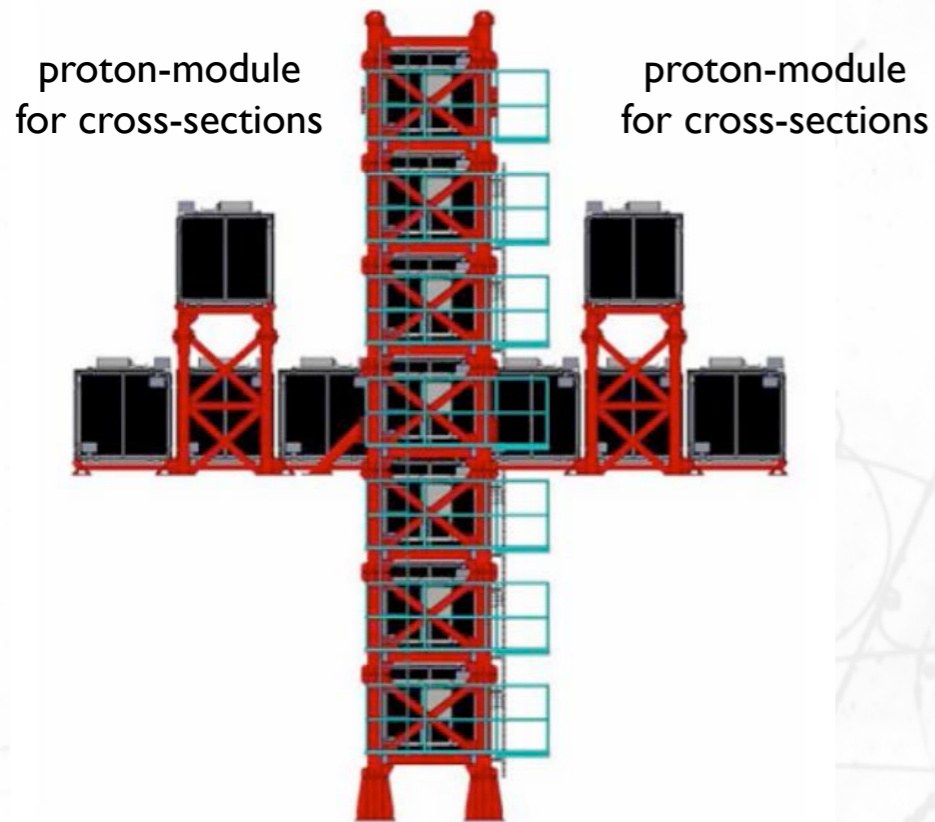
ND280: Off-axis



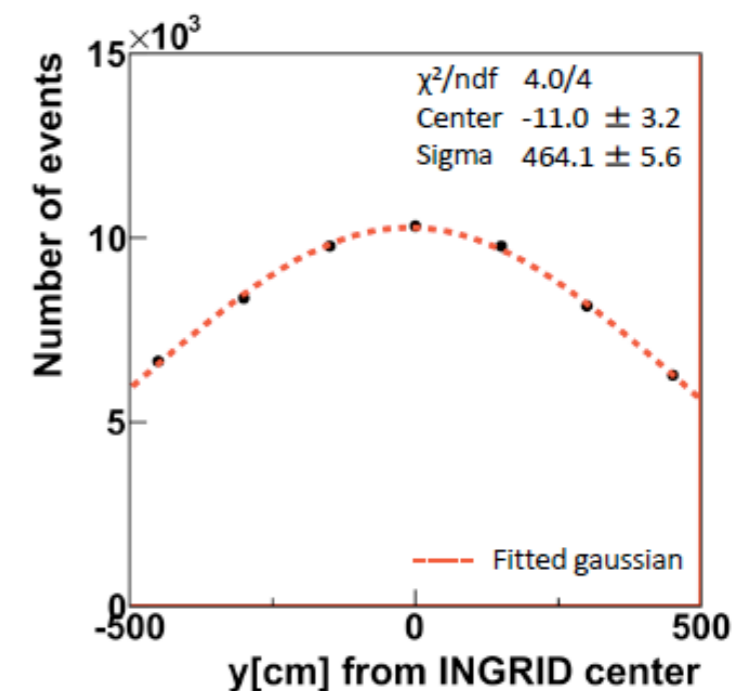
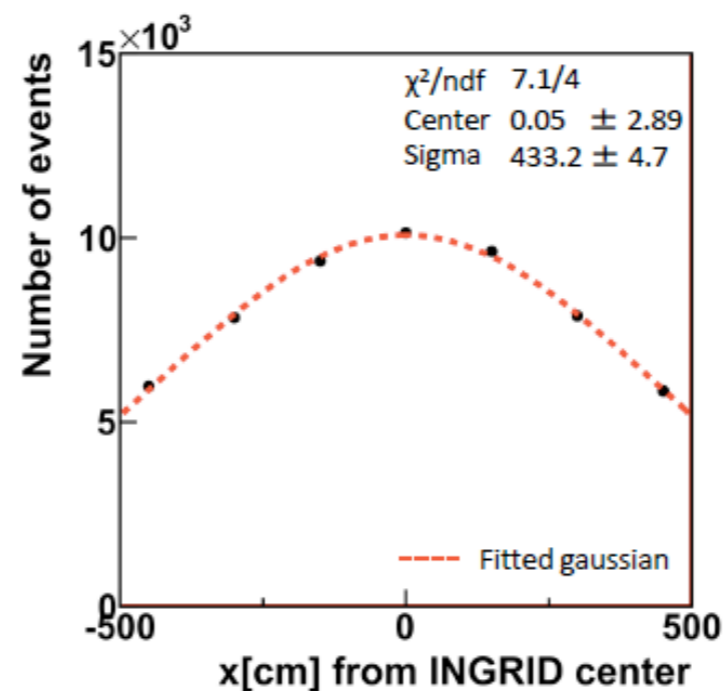
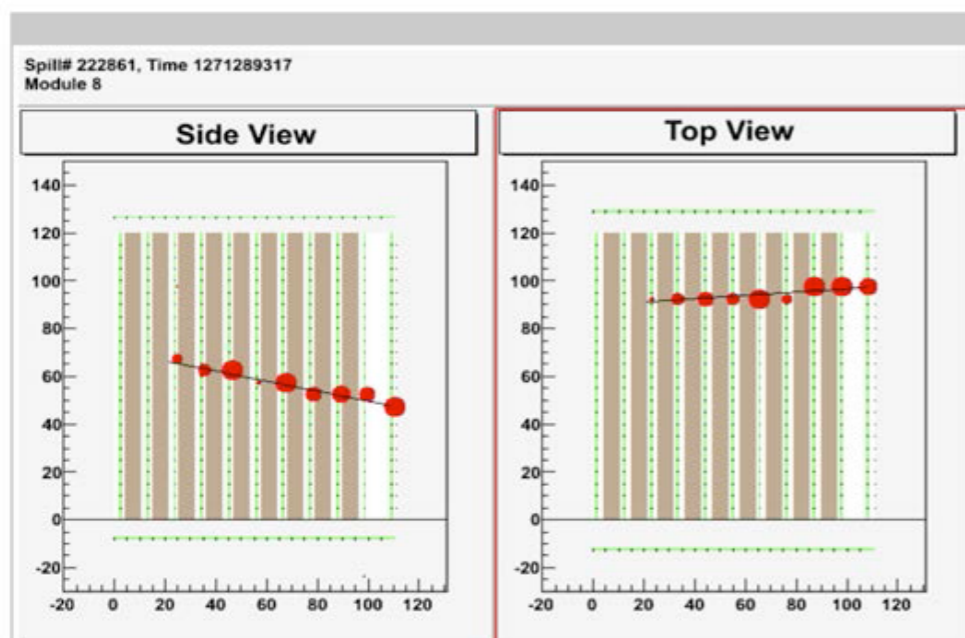
Wagasci/BabyMind: Off-axis

New!

On-Axis ND



- INGRID counts $v(\bar{\nu})$ CC events in a cross of 13 identical detectors:
- total rate monitors beam intensity stability with respect to proton on target counting.
- The relative event counts between modules monitor the beam direction stability.



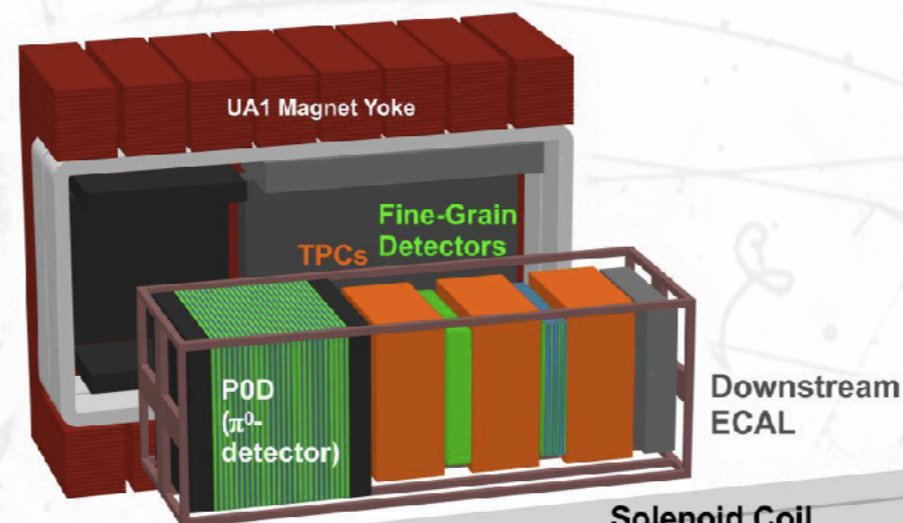
Off-Axis ND



- Same off-axis angle as SuperKamiokande (2.5 degrees)
- Measure ν_μ and ν_e spectrum before the oscillation \rightarrow TPCs + FGDs
- Measure background processes to oscillation (NC π^0 , NC π , CC π ...)
- Compare Carbon and Oxygen interactions (FGD2 and POD)

ND280 installed in ex-UA1 magnet (0.2 T) 3.5x3.6x7.3 m

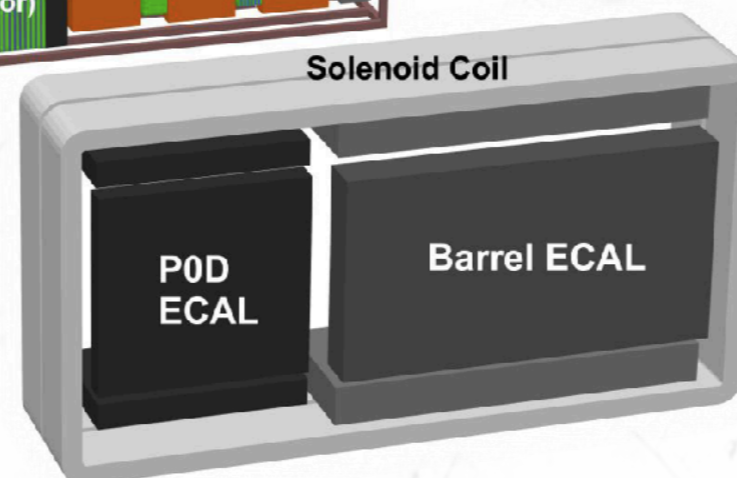
SMRD (Side Muon Range Detector):
scintillator planes in magnet yokes.
Measure high angle muons



2 FGDs (Fine Grained Detector):
active target mass for the tracker, optimized for p/ π separation
Carbon+Water target in FGD2

3 TPCs (Time Projection Chambers):
measure momentum and charge of particles from FGD and POD, PID capabilities through dE/dx

POD (π^0 detector):
scintillator bars interleaved with fillable water target bags and lead and brass sheets.
Optimised for γ detection

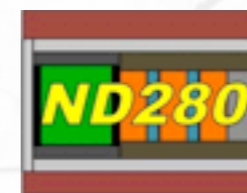


POD, Barrel and Downstream ECAL:
scintillator planes with radiator to measure EM showers

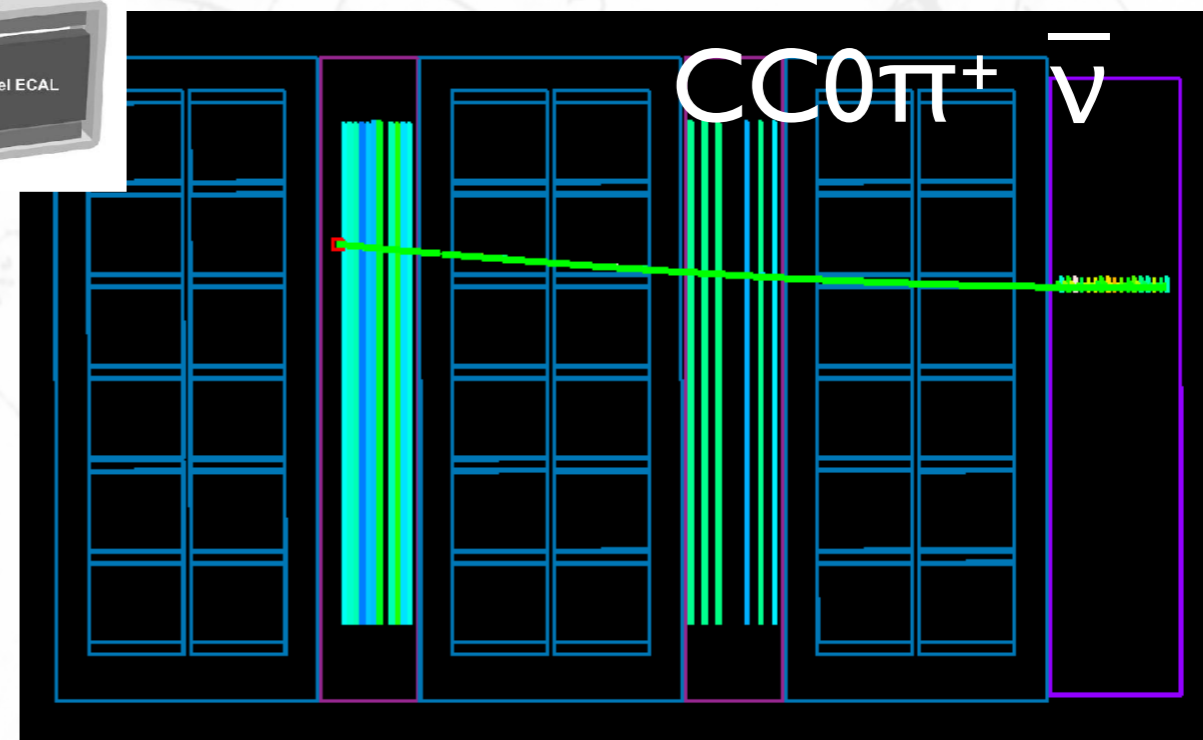
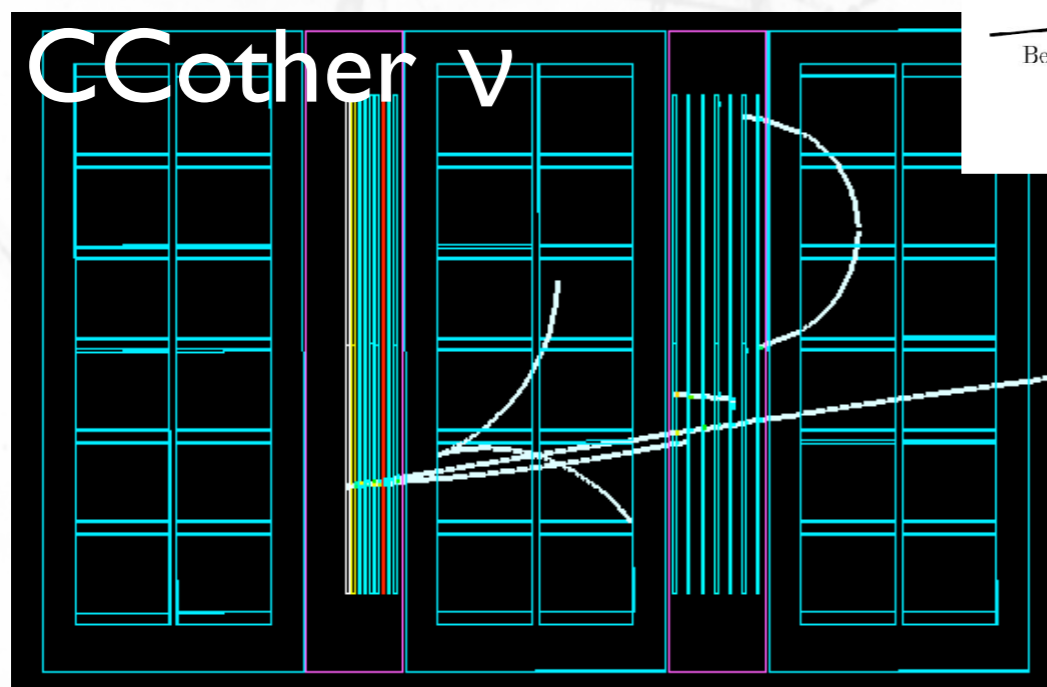
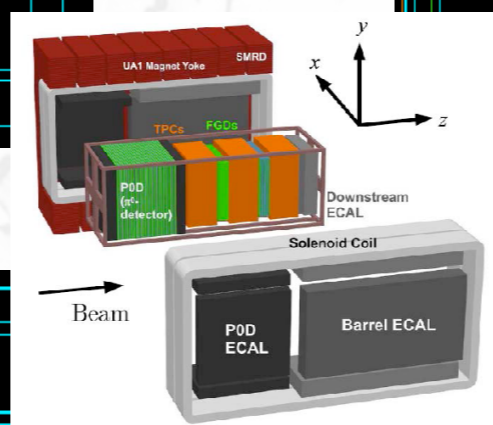
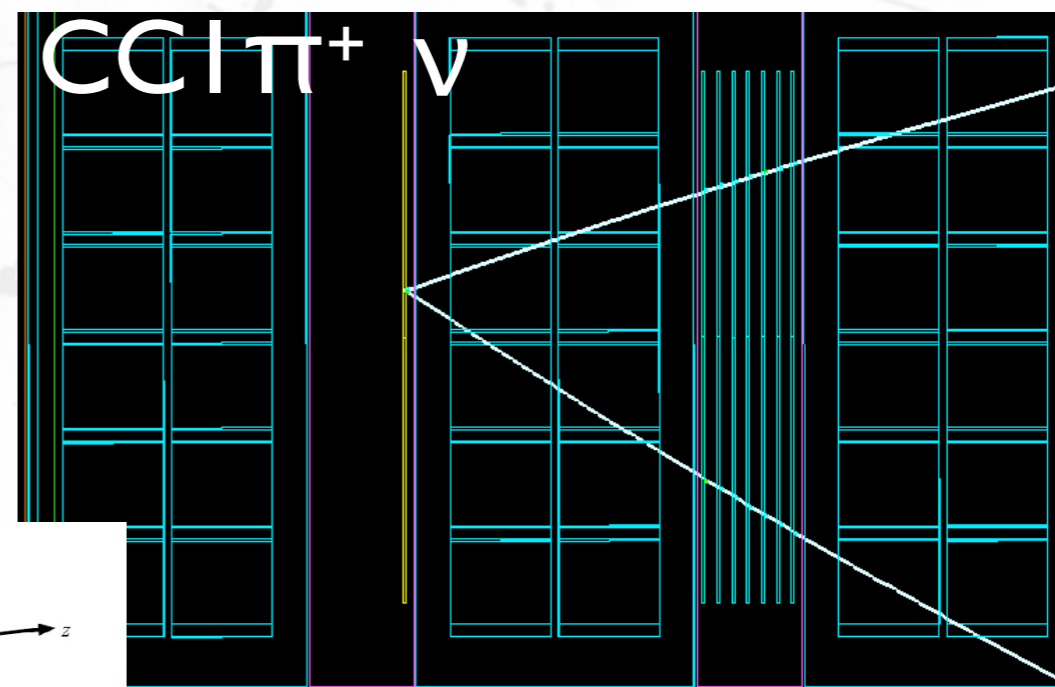
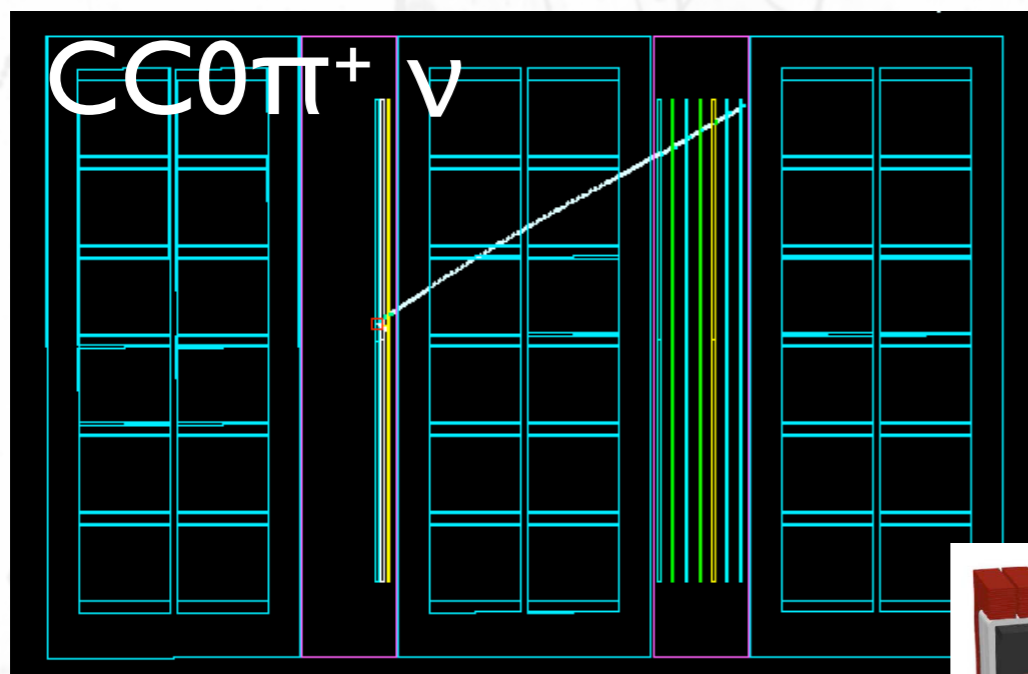


Magnet was granted by CERN

Off-Axis ND



Off-axis ND280 analysis real events

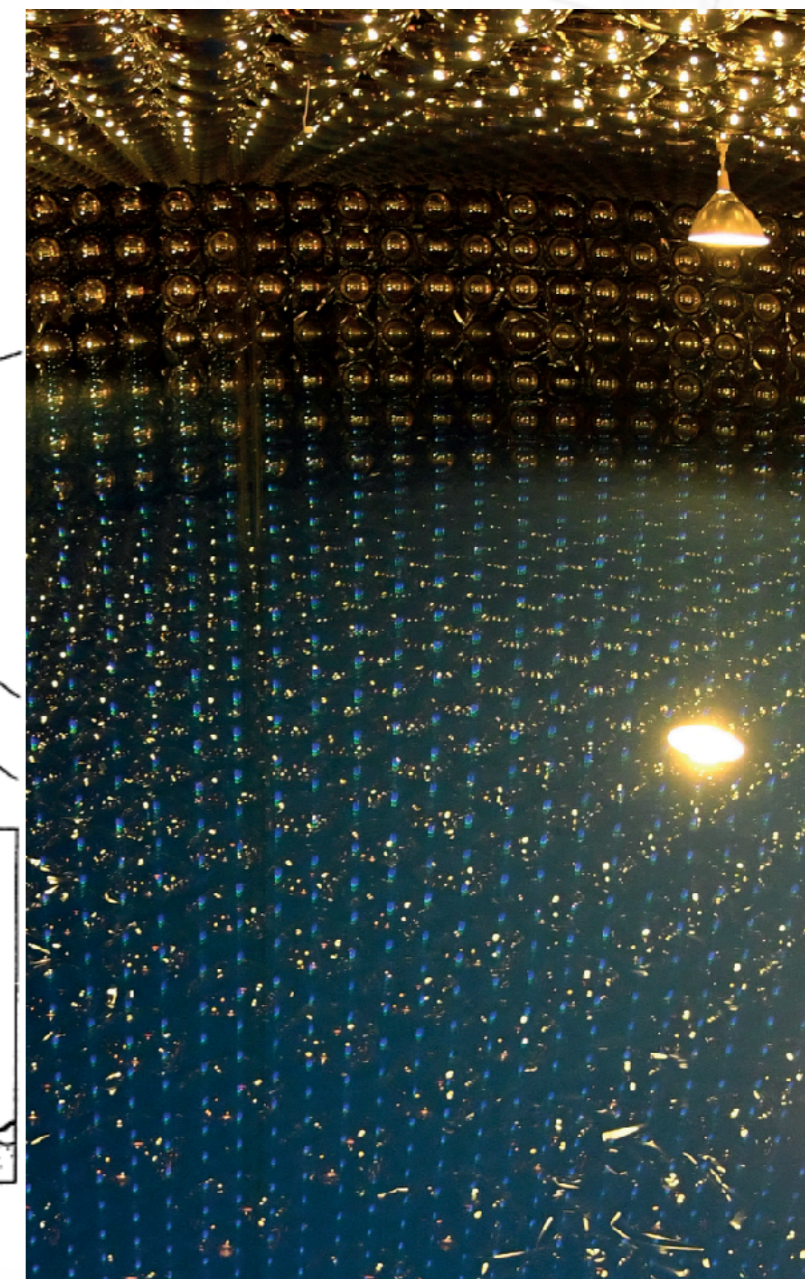
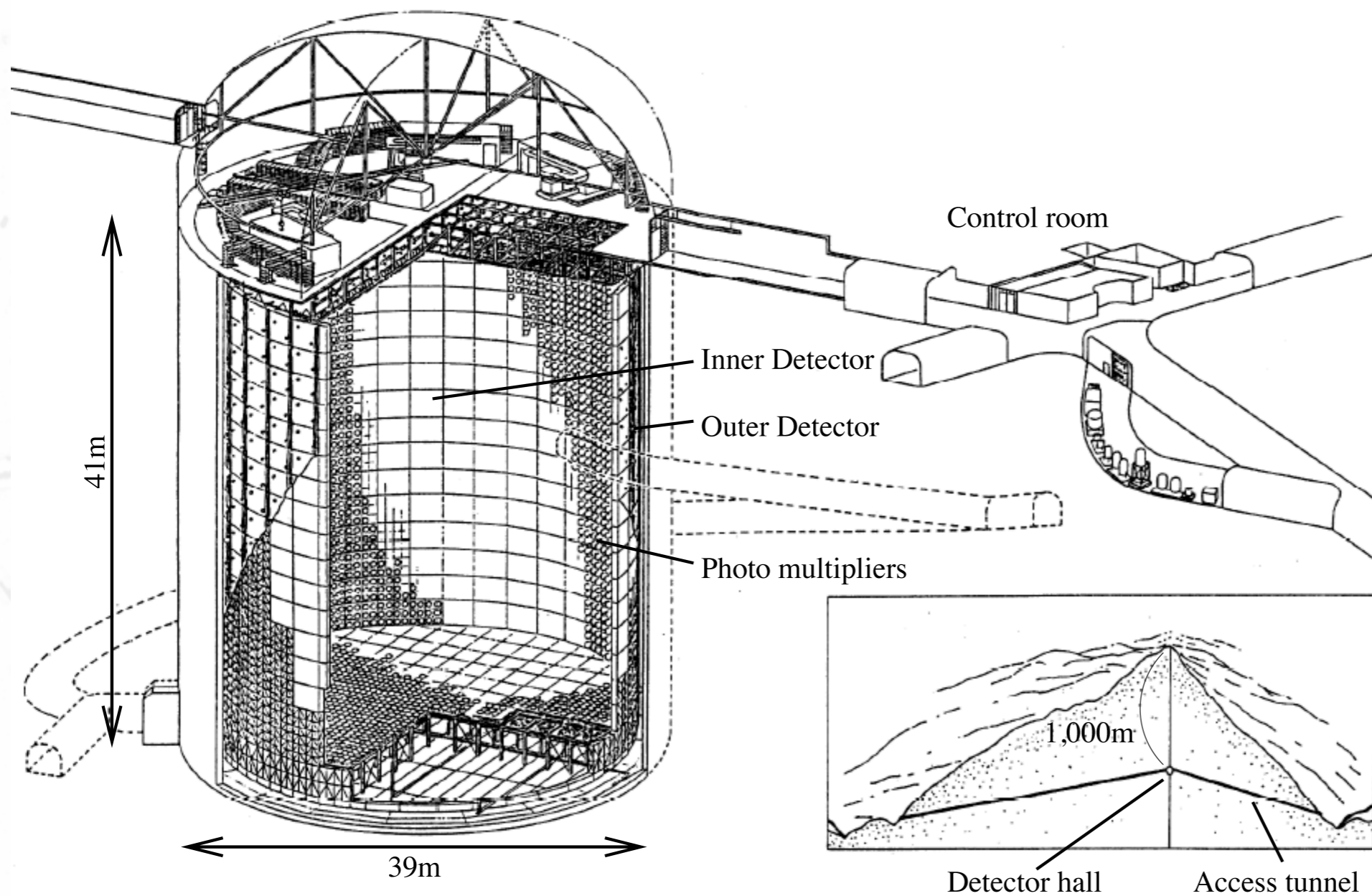


Far detector

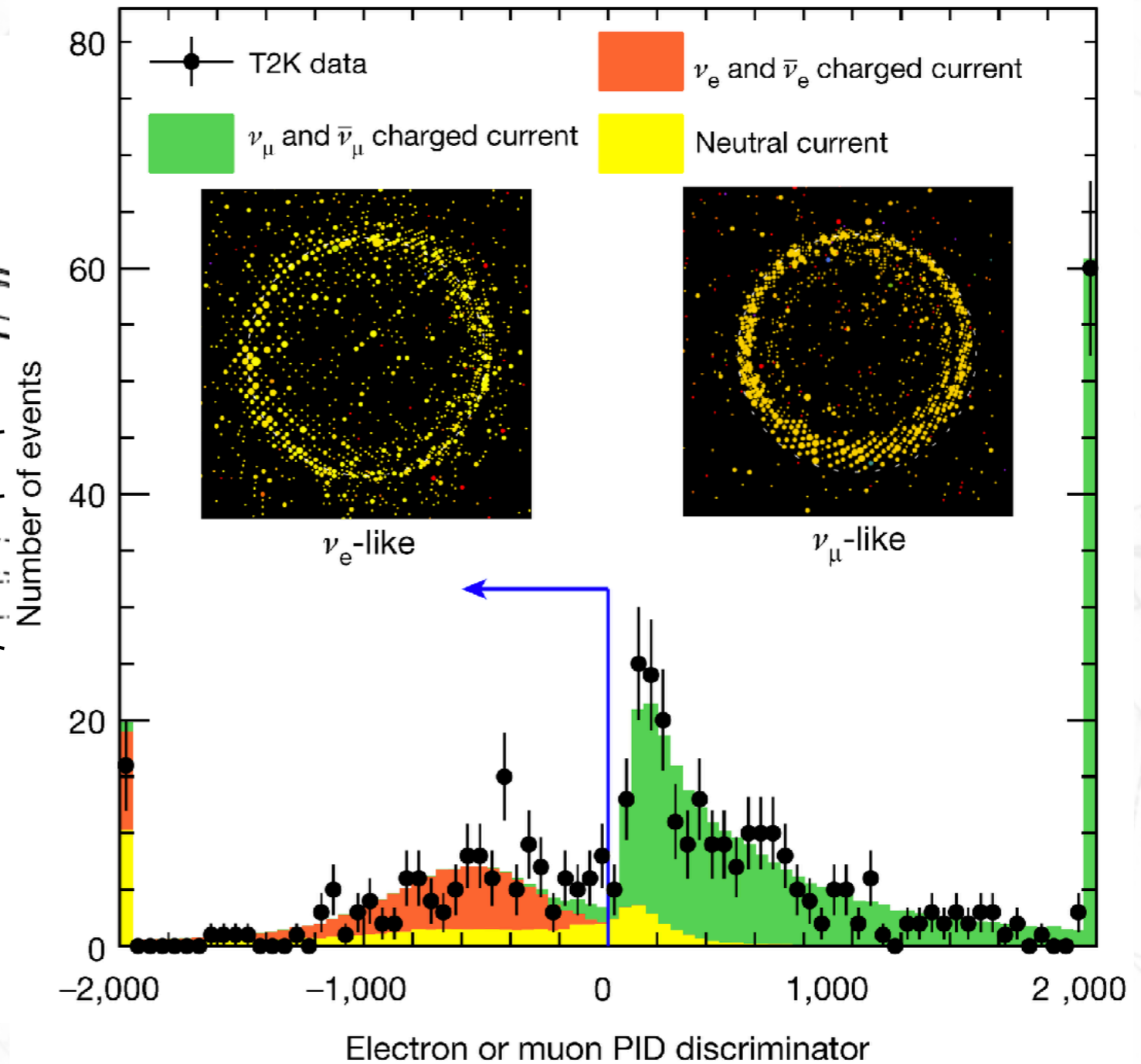
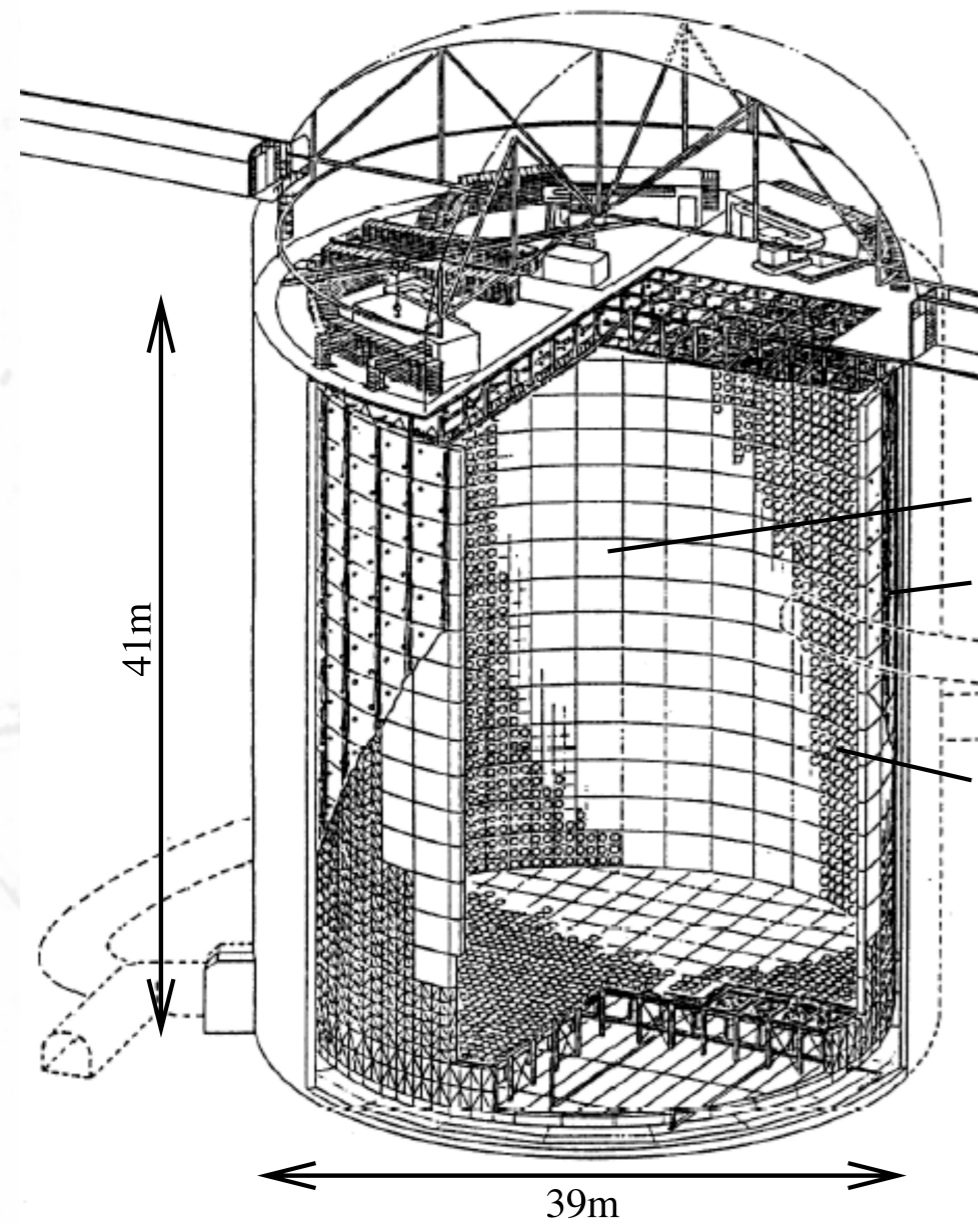


50 kton of ultra-pure water

~11000 20" PMT's



Far detector



Particle identification

Interaction vertex reconstruction

Track Multiplicity

Particle range

Electromagnetic energy reconstruction

Hadronic interactions

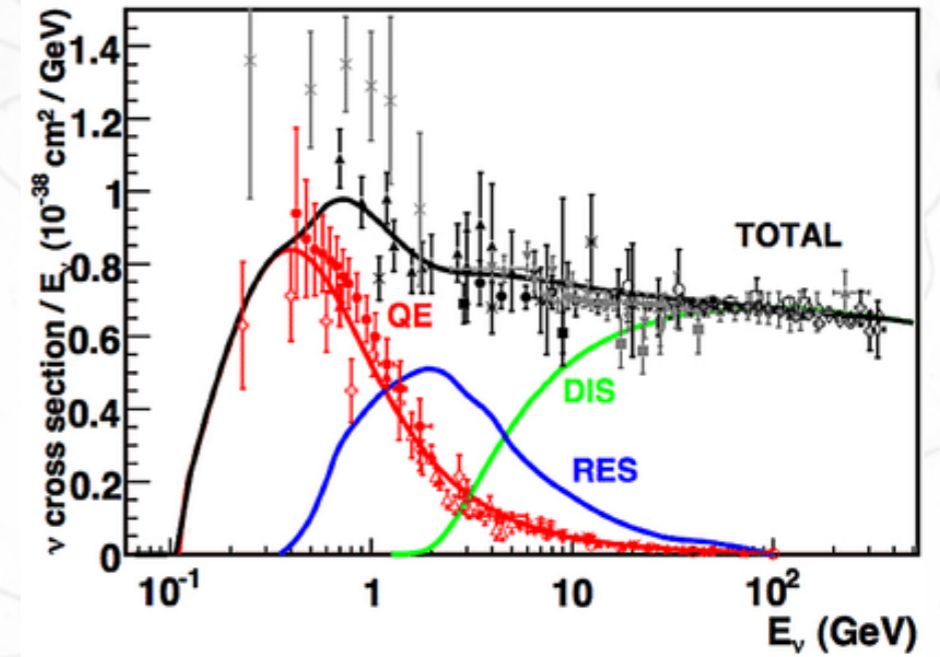
Neutrino-nucleus cross-sections

νA cross-sections

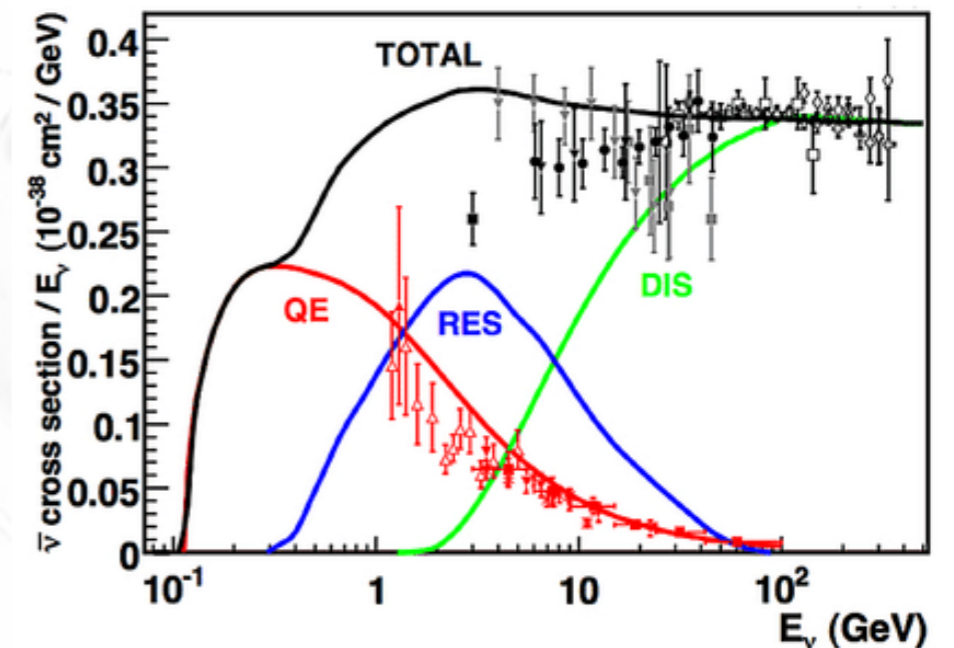
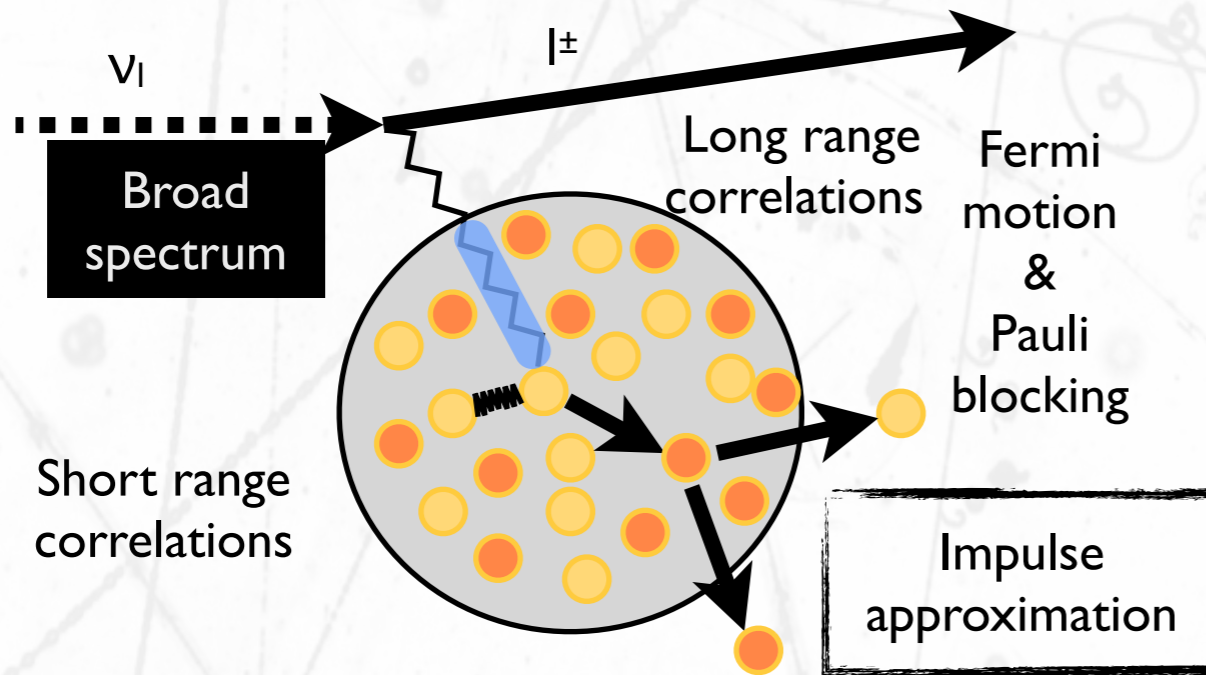


@ the nucleon level !

<i>CCQE</i>	$\nu_{\mu} n \rightarrow \mu^{-} p$
<i>CC1π</i>	$\nu_{\mu} p \rightarrow \mu^{-} \Delta^{++} \rightarrow \mu^{-} \pi^{+} p$ $\nu_{\mu} n \rightarrow \mu^{-} \Delta^{+} \rightarrow \mu^{-} \pi^{+} n$ $\nu_{\mu} n \rightarrow \mu^{-} \Delta^{+} \rightarrow \mu^{-} \pi^{0} p$
<i>CCNπ</i>	$\nu_{\mu} N \rightarrow \mu^{-} \Delta^{+,++} \rightarrow \mu^{-} N' \pi \pi \dots$
<i>CCDis</i>	$\nu_{\mu} N \rightarrow \mu^{-} N' \pi, \pi, \dots$



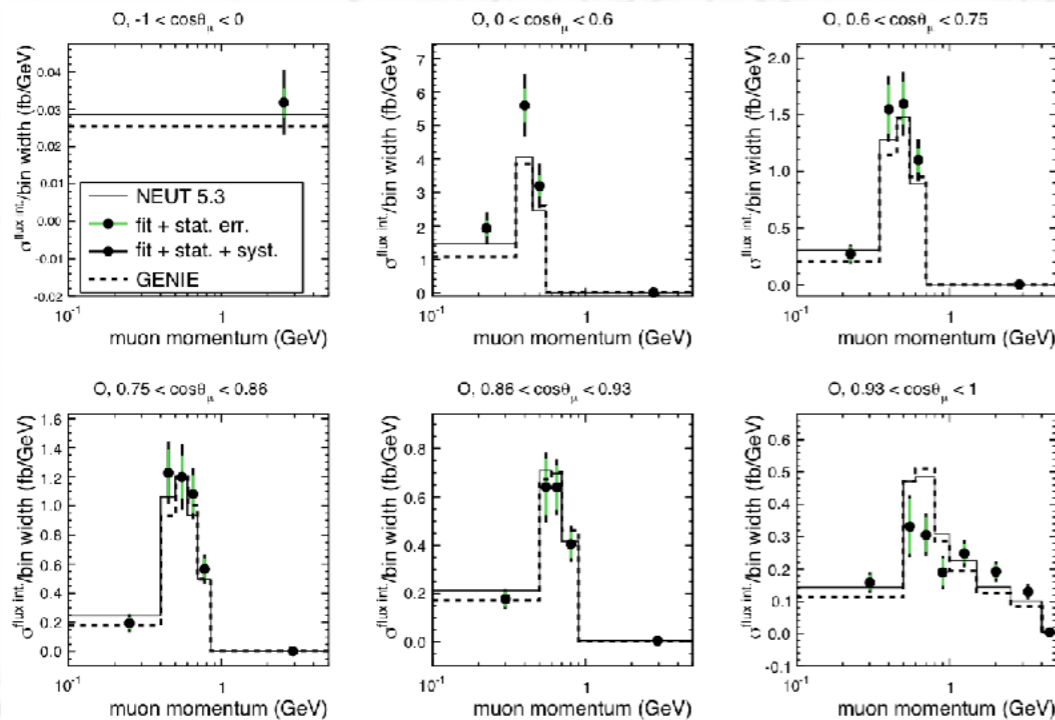
@ the nucleus level !



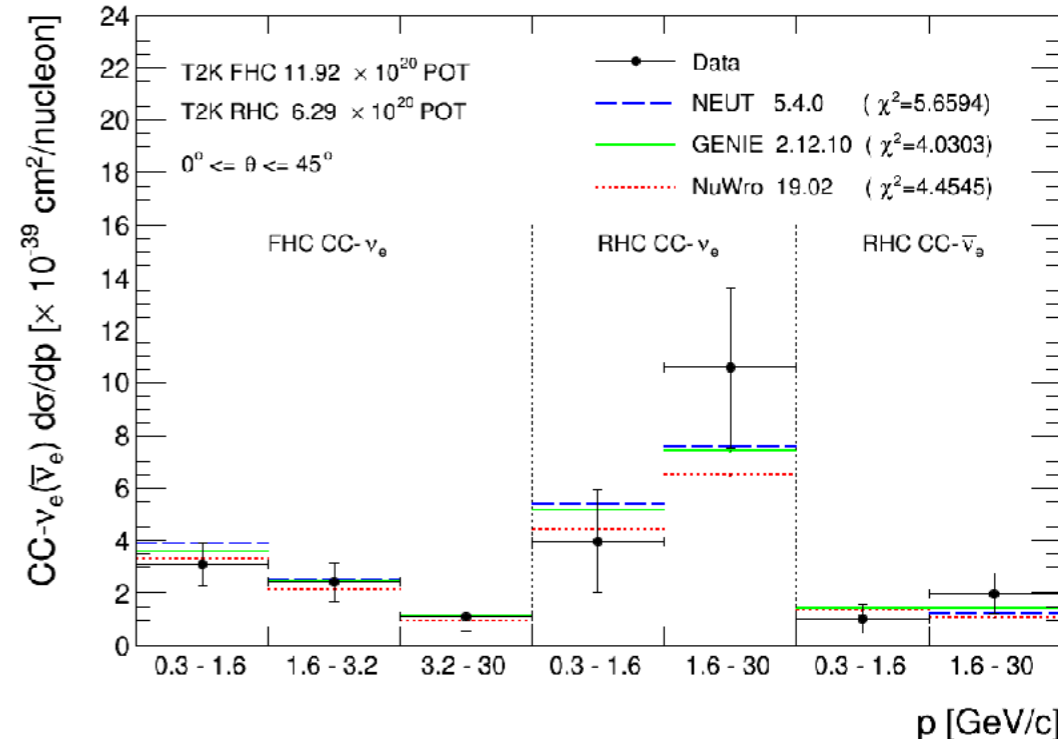
νA cross-sections



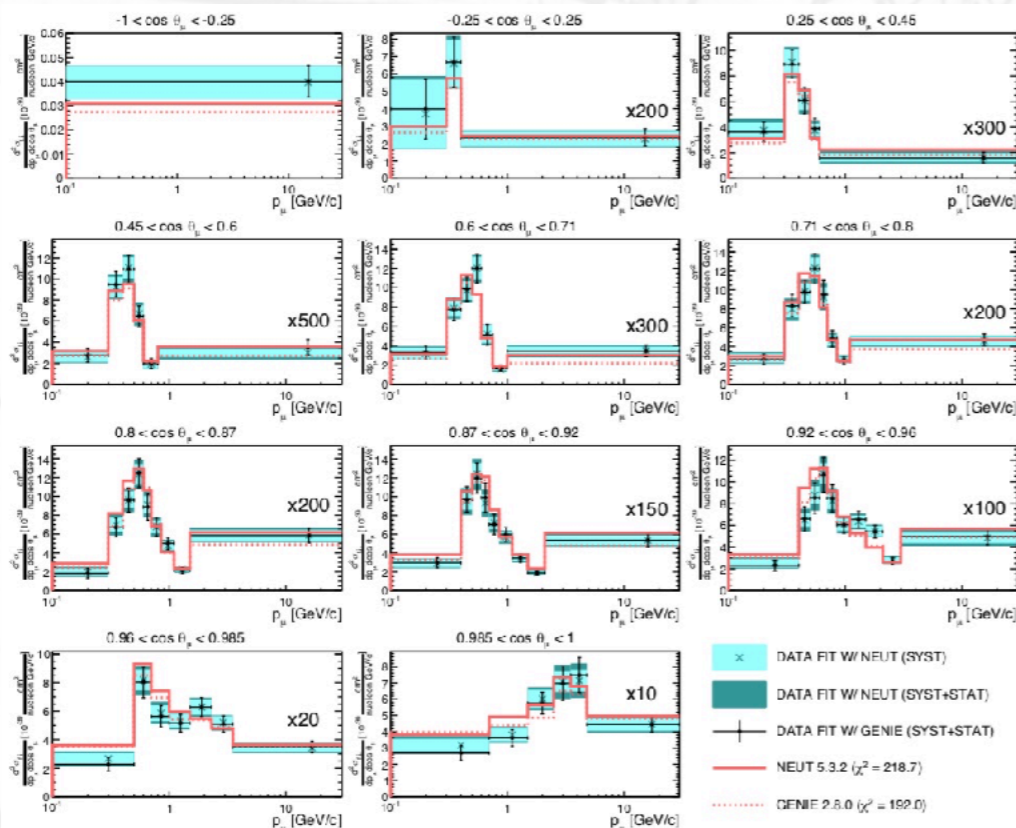
ν_μ CC π^0 on Oxygen



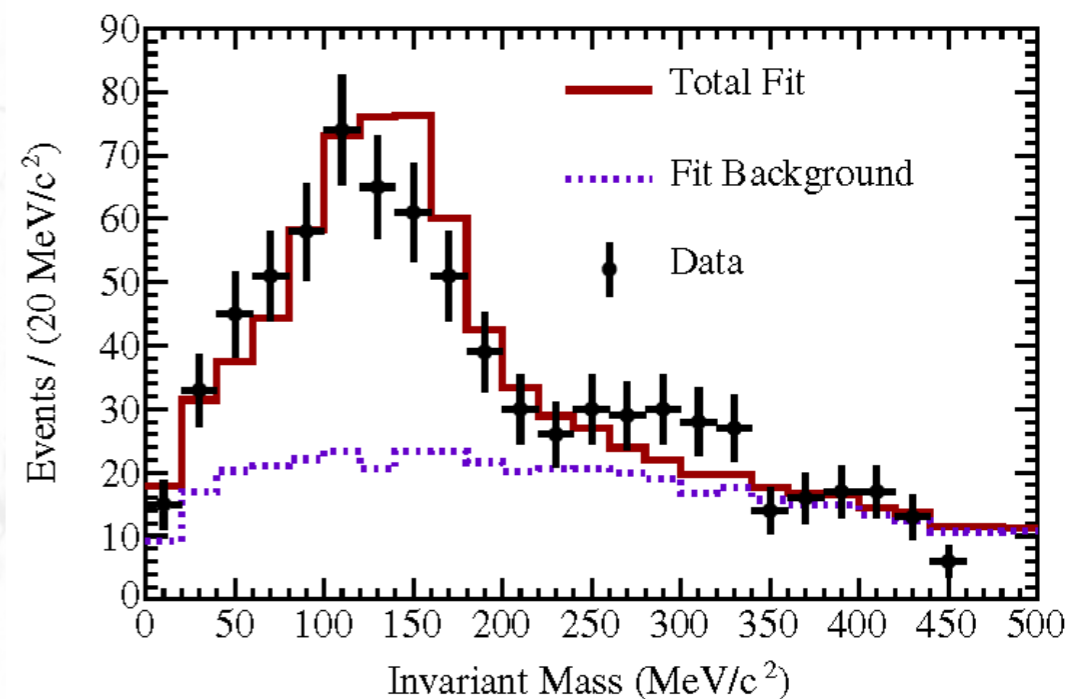
ν_e CC on Carbon



inclusive ν_μ CC on C



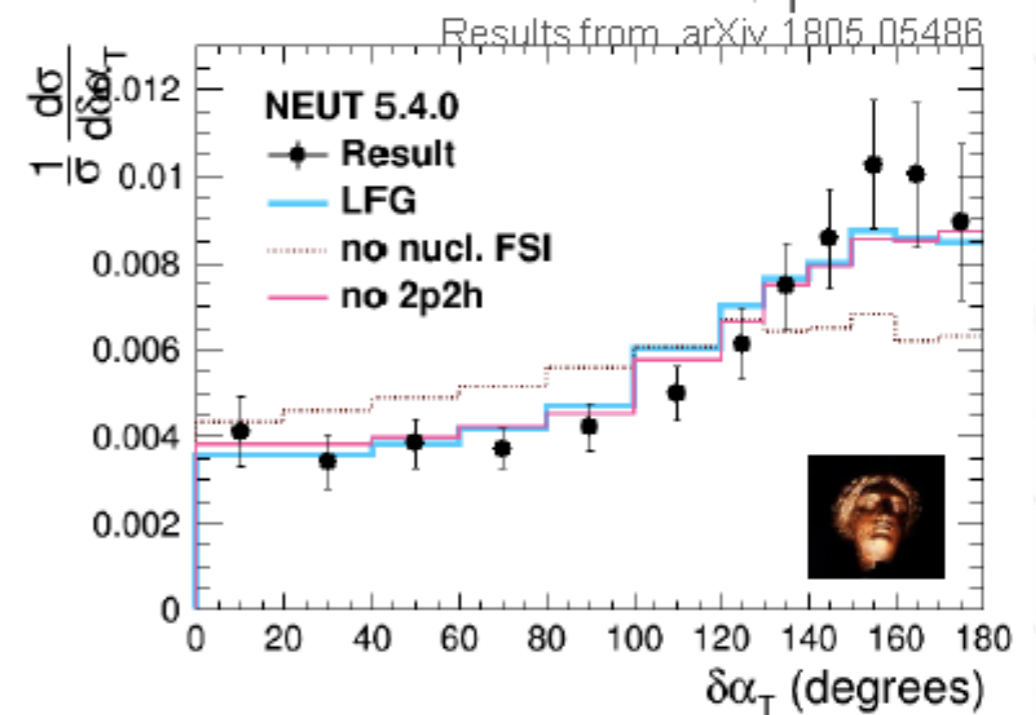
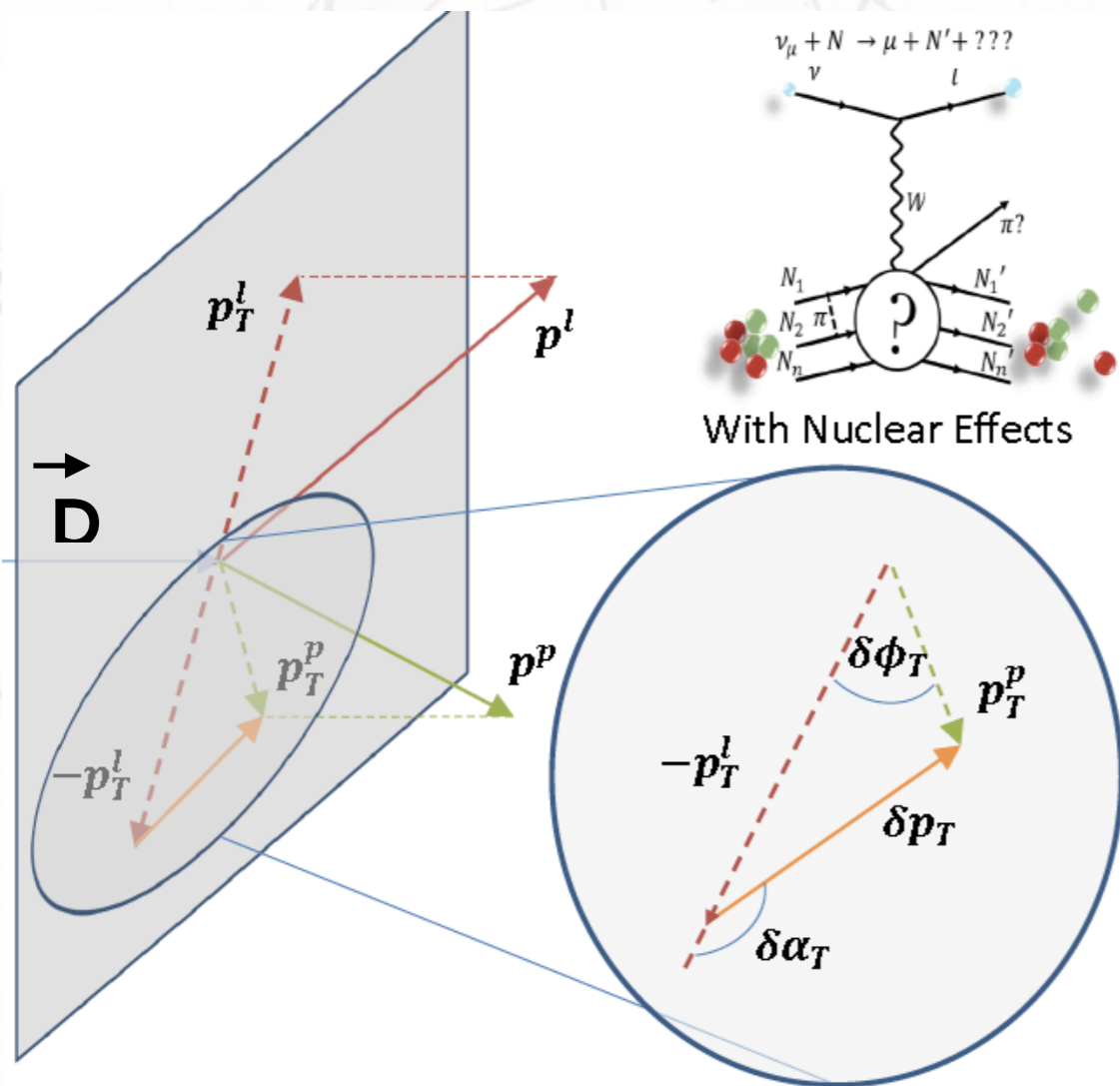
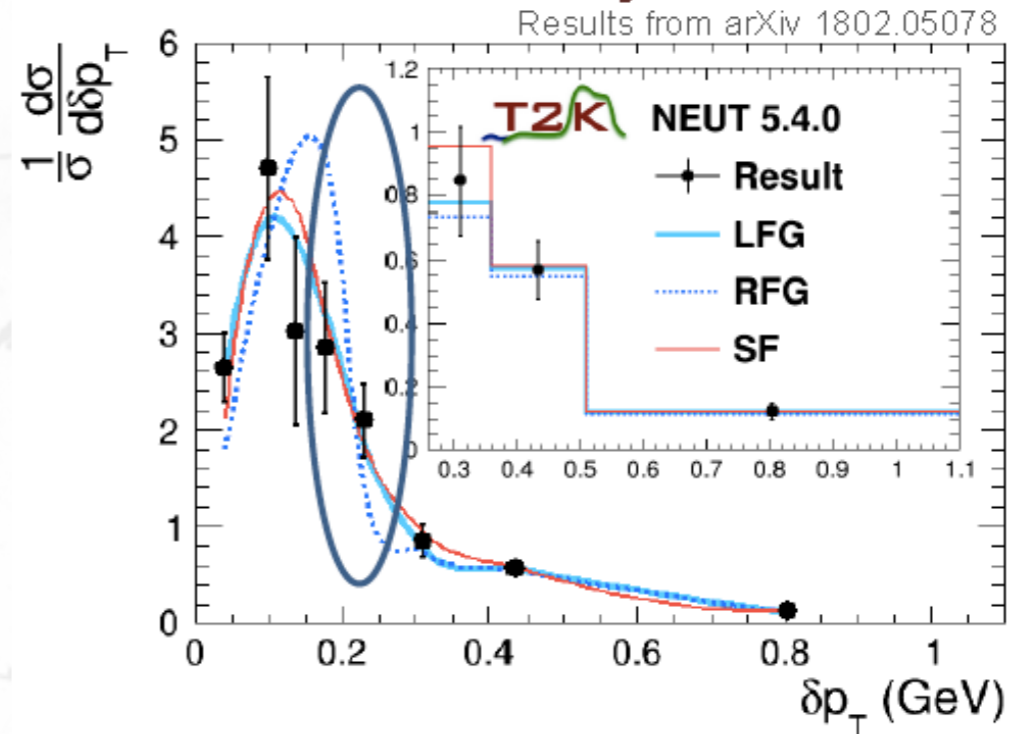
ν_μ NC π^0 on Water



νA cross-sections



Low energy recoil proton allow us to measure transverse momentum imbalance to access nuclear effects: Fermi momentum and re-interactions.



Analysis procedure

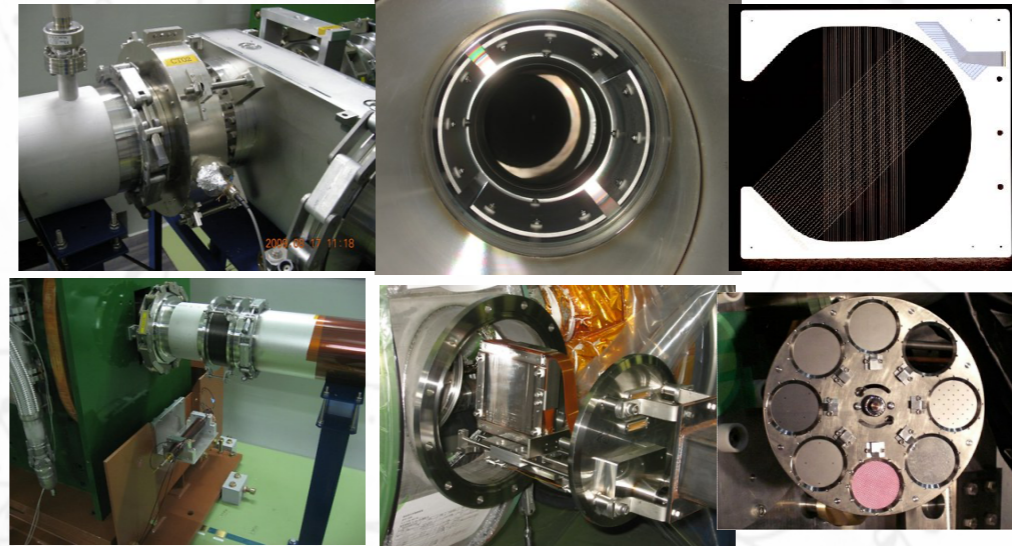
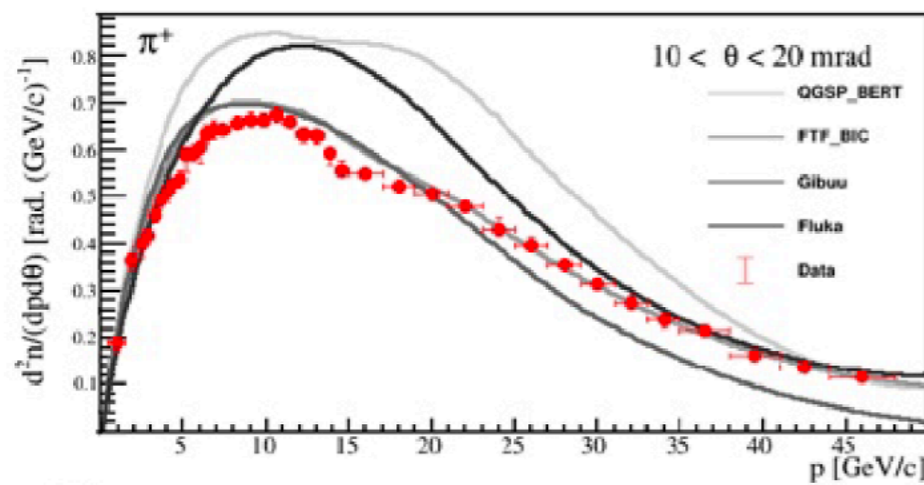
Beam model



Beam model is obtained from a full GEANT simulation of the particle transport reweighed by the NA61 results

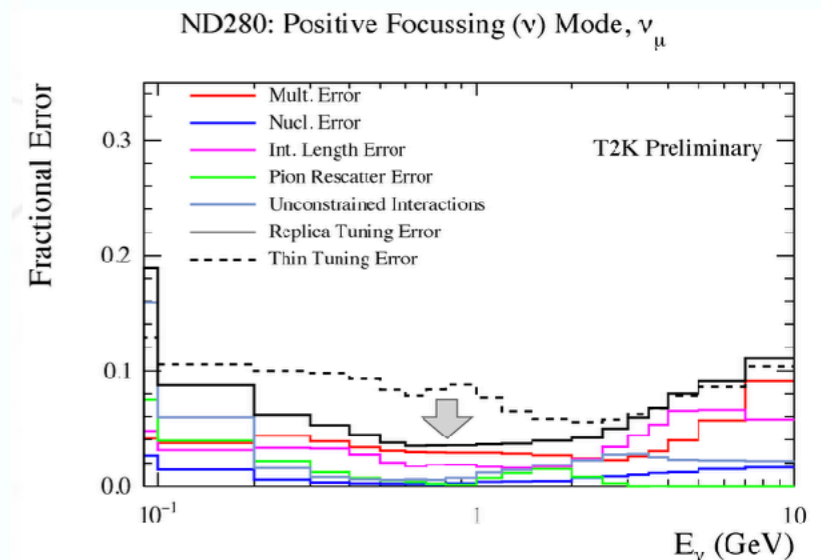
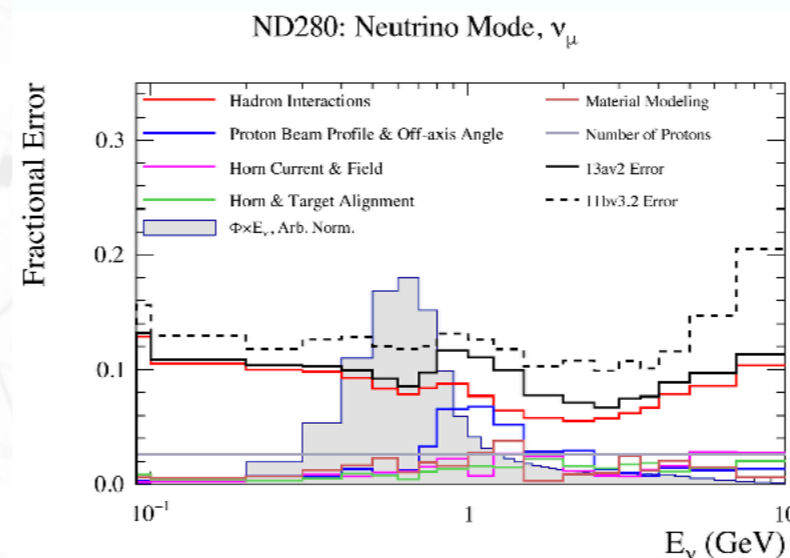
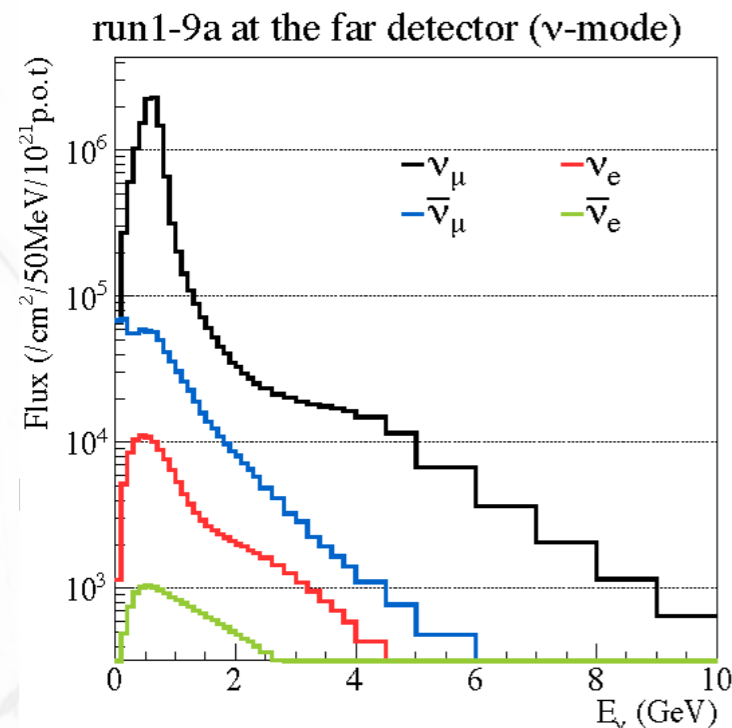
Input

Beam monitors



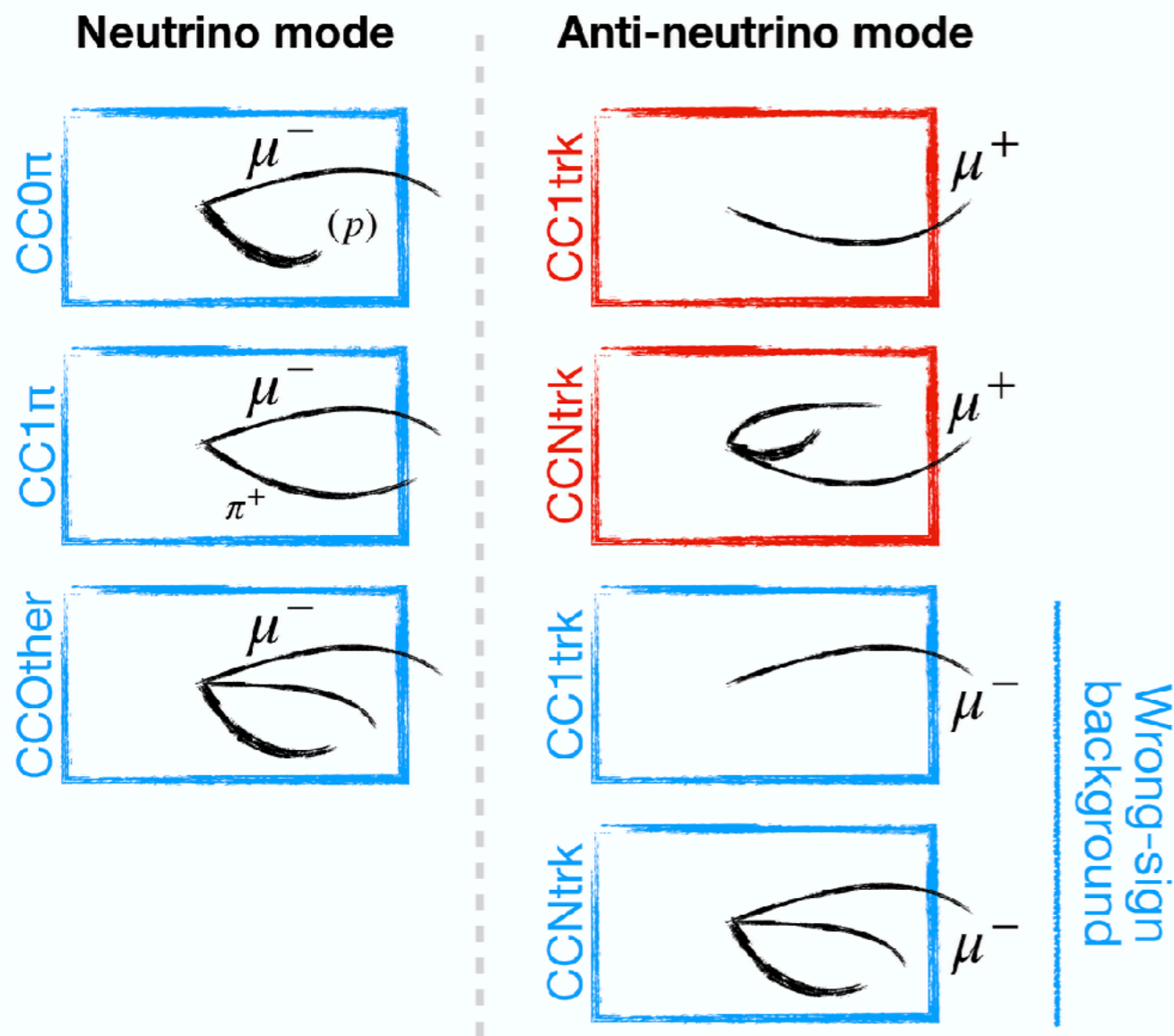
GEANT 3

Output



Including error covariance matrix

Near detector data

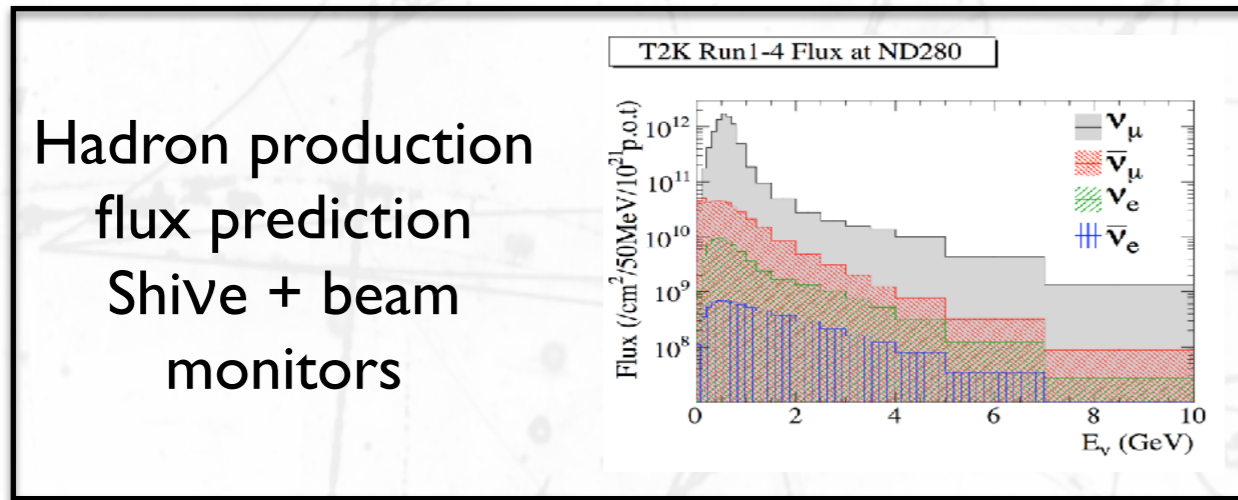
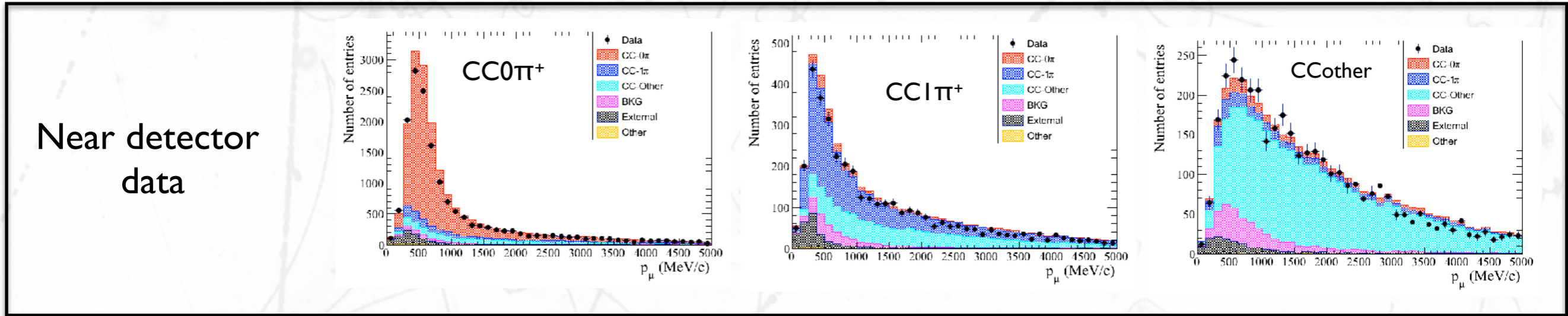


Each sample for two targets FGD1,2 ~ (C,O)

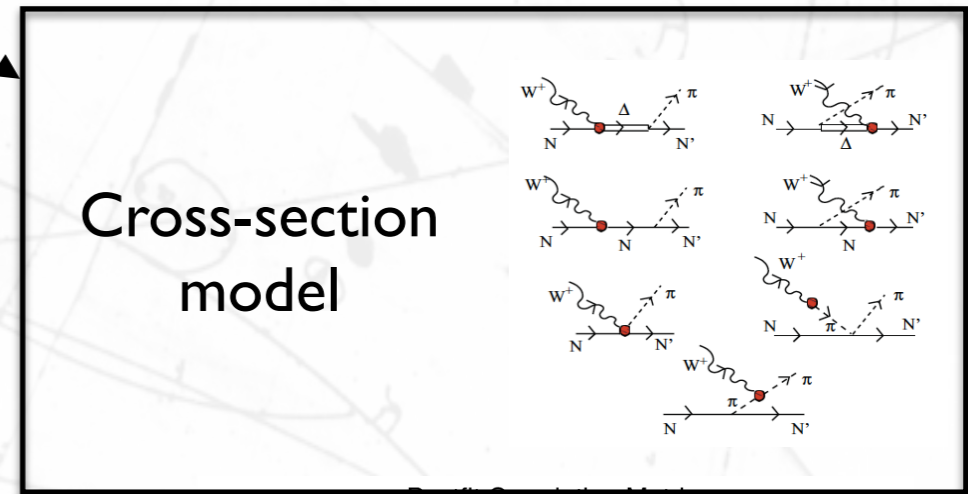
CCOther are CC events with multiple π 's, π^0 or π^- candidates

- Analysis of muon kinematics: $(p_\mu \cos \theta_\mu)$
 - E_ν obtained from interaction model in generators.
- Three samples of neutrinos (two for antineutrinos): enriched in
 - CCQE \rightarrow CC0 π
 - CCRes \rightarrow CC1 π
 - CC-DIS \rightarrow CCOther
- the different samples also have different E_ν dependencies.
- Wrong Sign background:
 - neutrinos in anti-neutrino mode.
- Neutrino interaction in Oxygen from FGD2 data.

Conceptually

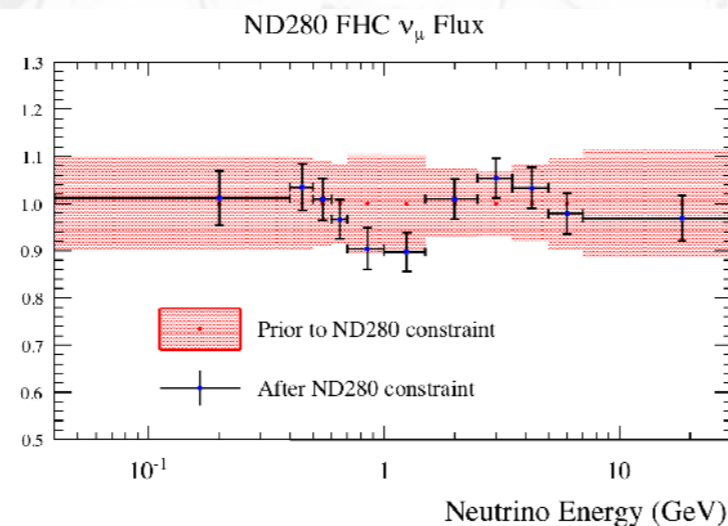


feed back

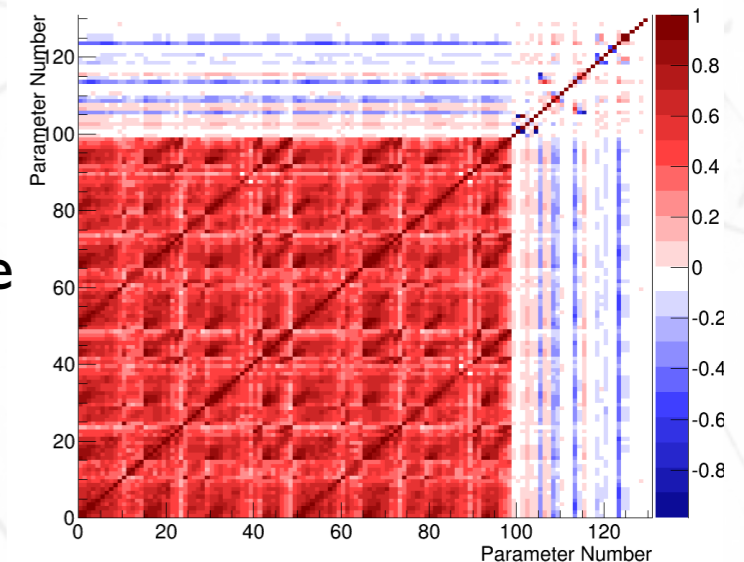


Postfit Correlation Matrix

Corrected flux and cross-section model



& error covariance matrix



Oscillation fits



$\nu_{\mu} \rightarrow \nu_e$ and $\bar{\nu}_{\mu} \rightarrow \bar{\nu}_e$ combined analysis within the 3v oscillation paradigm (PMNS).

Other oscillation parameters from 2018 PDG values.

Binned likelihood comparing data to MC predictions.

Bins of reconstructed energy from lepton kinematics assuming CCQE two body interactions.

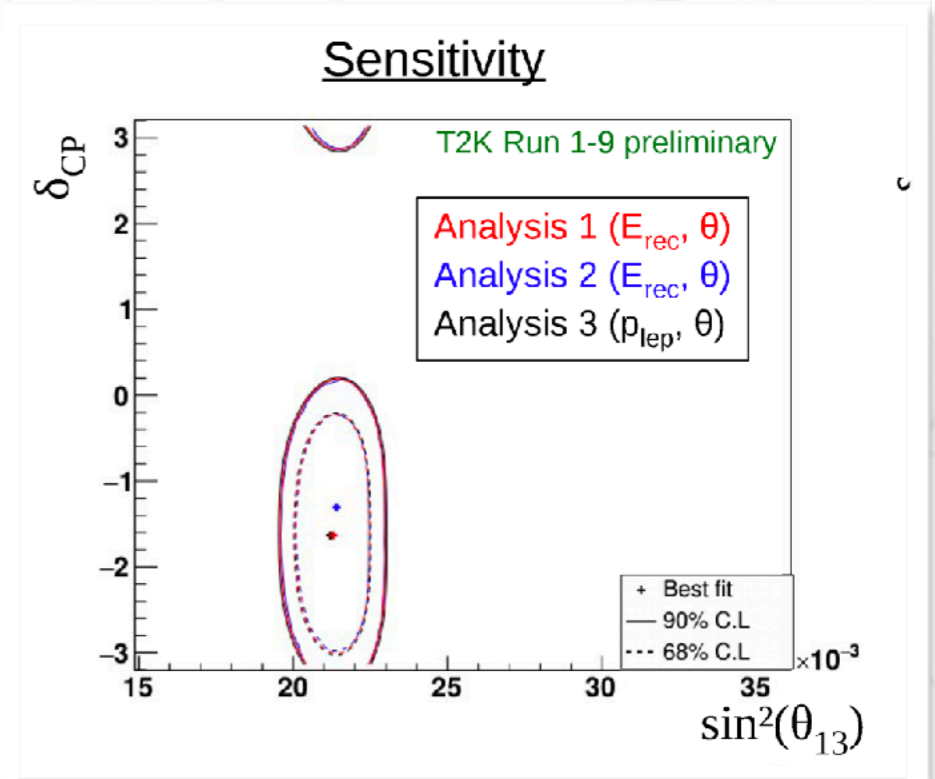
ν_e sample also bins in θ_e

Bayesian Markov Chain MonteCarlo and 2 frequentist approach.

Frequentists confidence intervals (grid search) agree with the Bayesian factors and credible intervals.

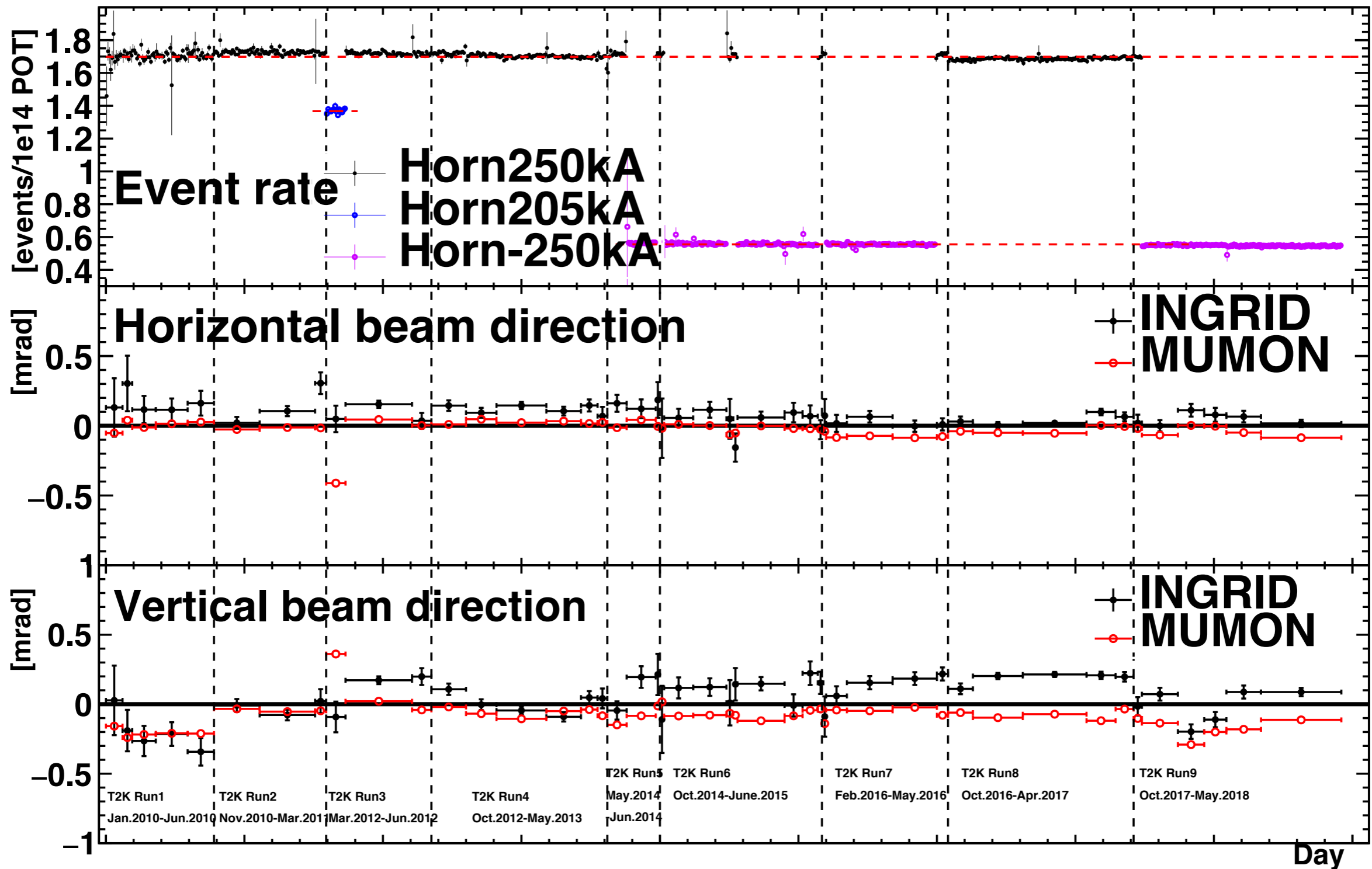
$$-2 \ln \lambda(\overline{\delta}_{CP}; \mathbf{a}) = 2 \sum_{i=1}^N \left[n_i^{obs} \ln \left(\frac{n_i^{obs}}{n_i^{exp}} \right) + n_i^{exp} - n_i^{obs} \right] + (\mathbf{a} - \mathbf{a}_0)^T \mathbf{C}^{-1} (\mathbf{a} - \mathbf{a}_0)$$

$$E_{rec} = \frac{ME_{\mu} - m_{\mu}^2/2}{M - E_{\mu} + |\vec{p}_{\mu}| \cos \theta_{\mu}}$$



T2K results

Beam stability

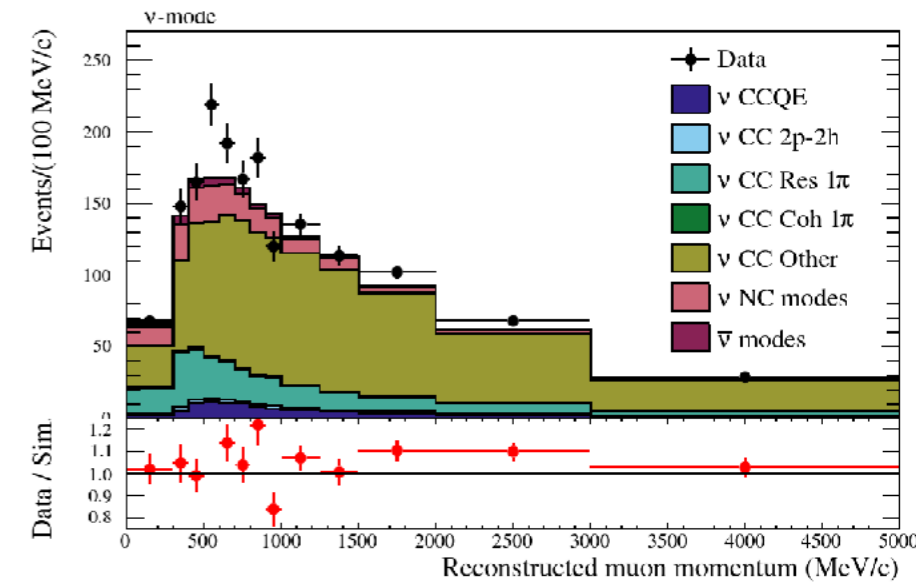
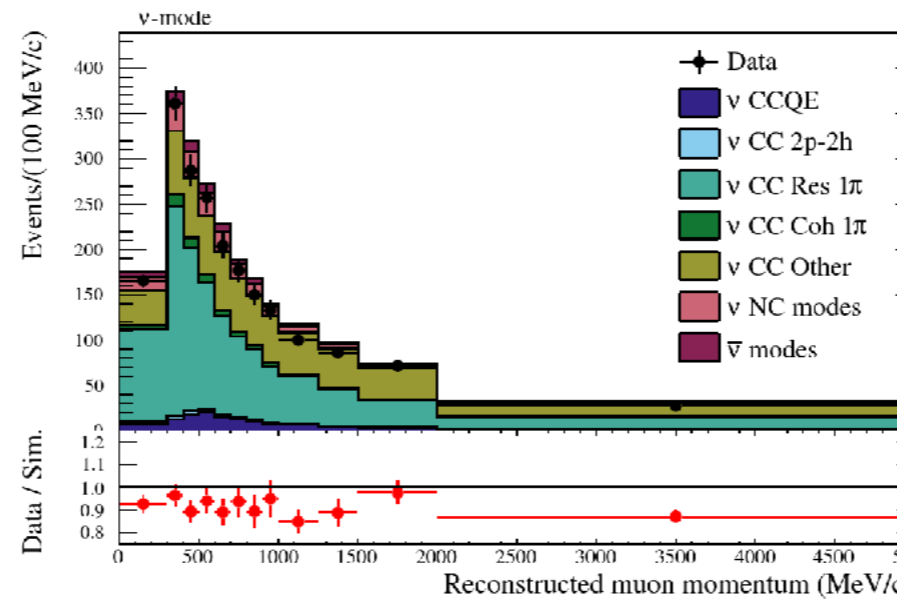
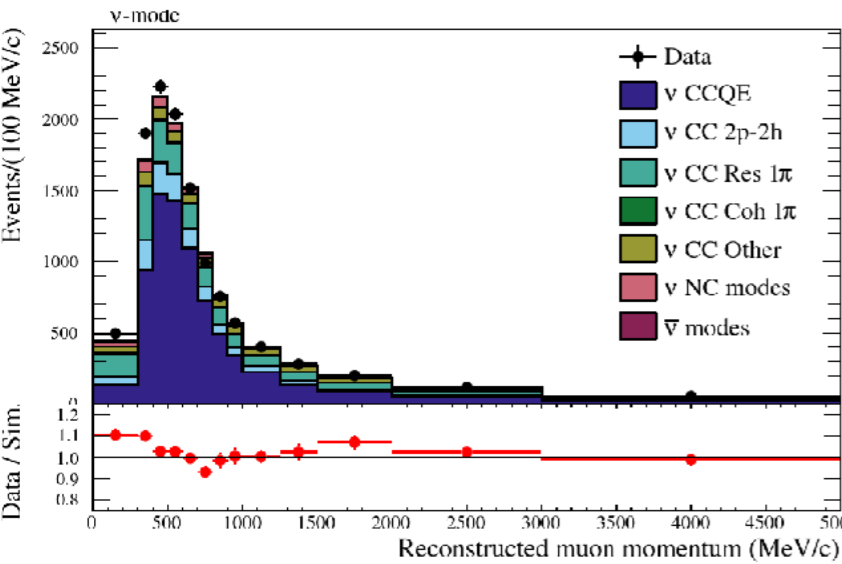


ND input samples

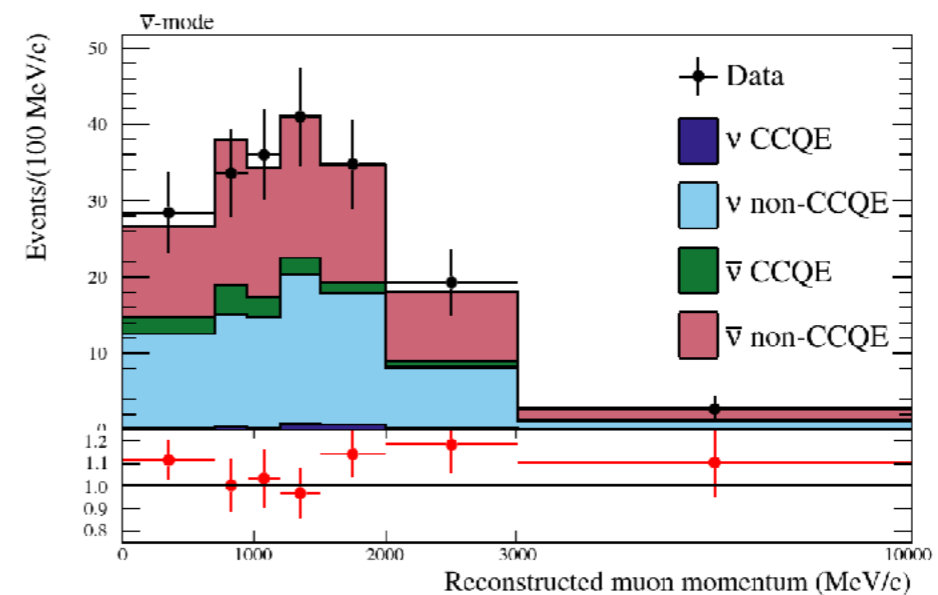
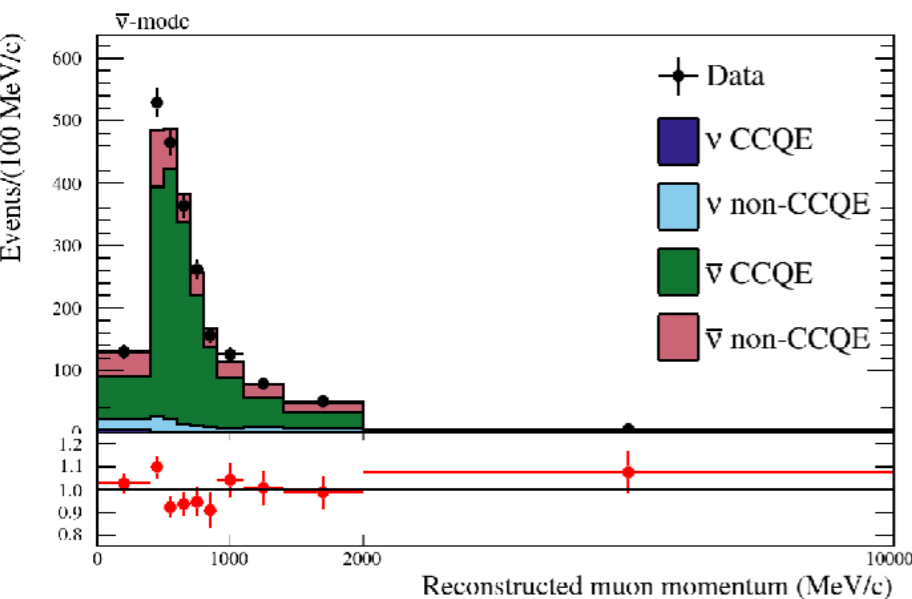


We use 14 different sample in $(p_\mu, \cos \theta_\mu)$

Forward Horn Current



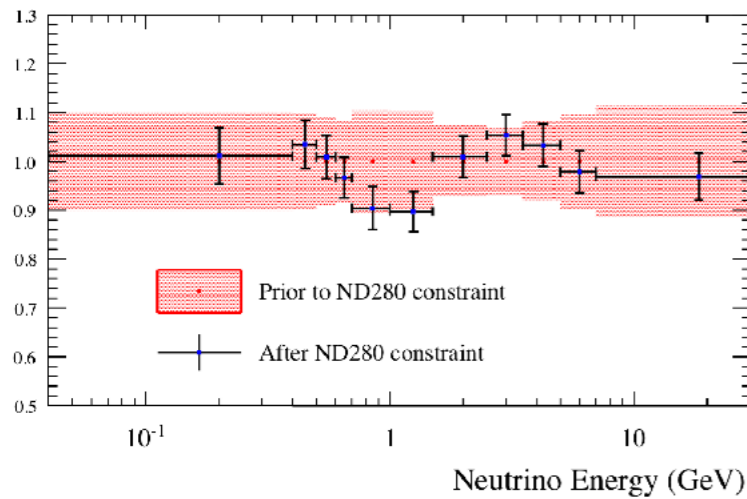
Reversed Horn Current



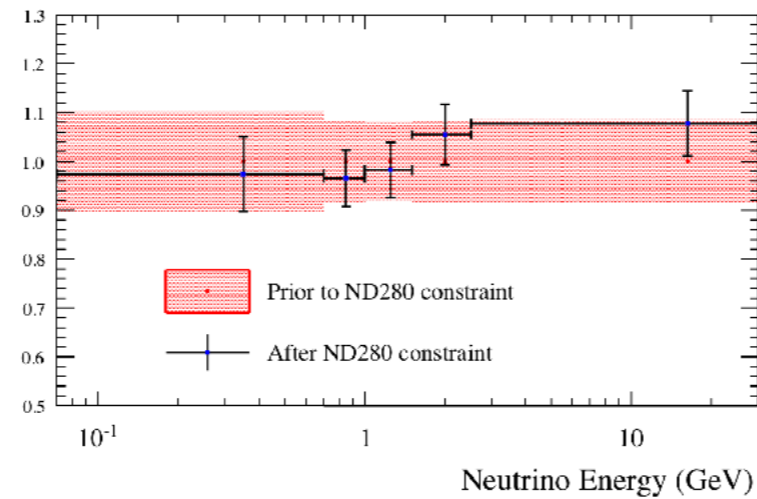
Examples of ND fits



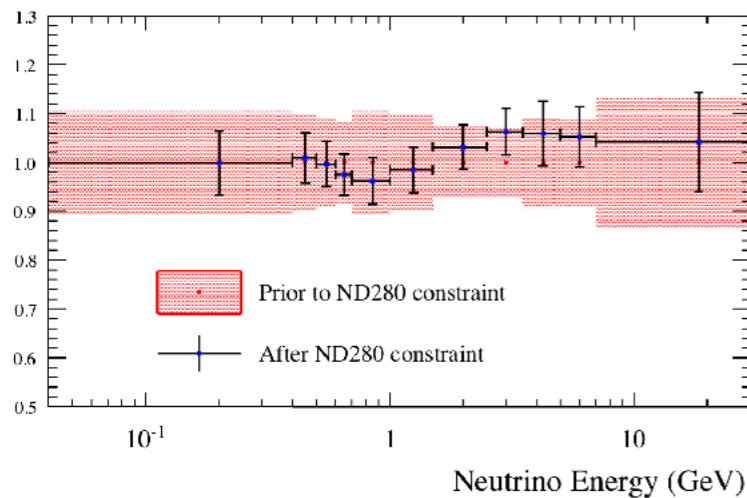
ND280 FHC ν_μ Flux



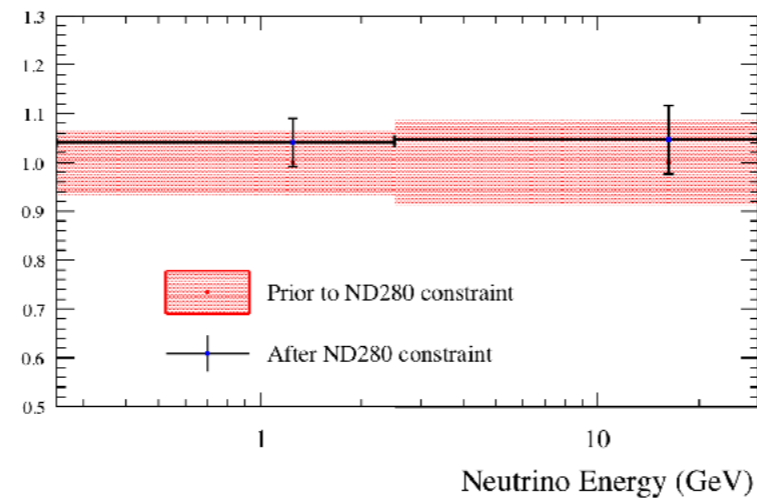
ND280 FHC $\bar{\nu}_\mu$ Flux



ND280 RHC $\bar{\nu}_\mu$ Flux

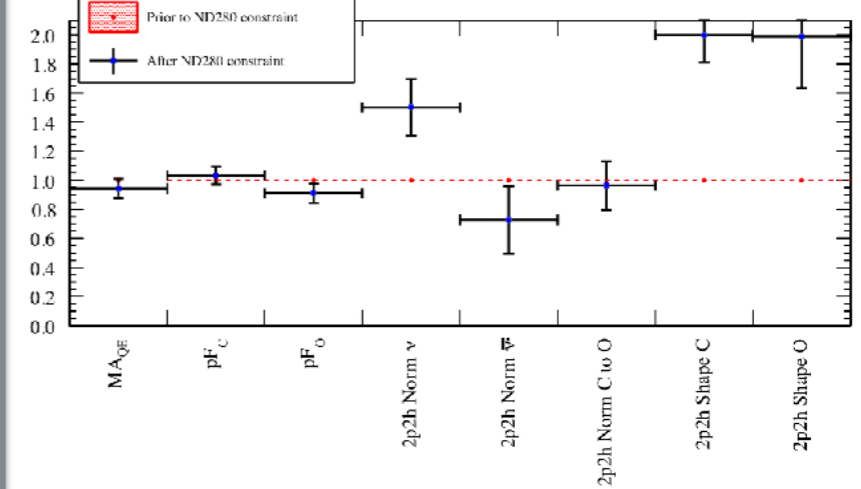


ND280 RHC ν_e Flux

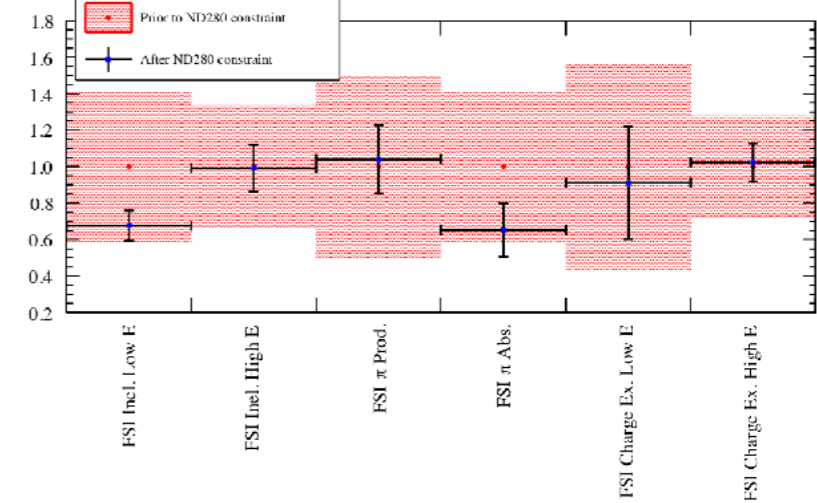


Flux parameters

CC0 π parameters



FSI parameters

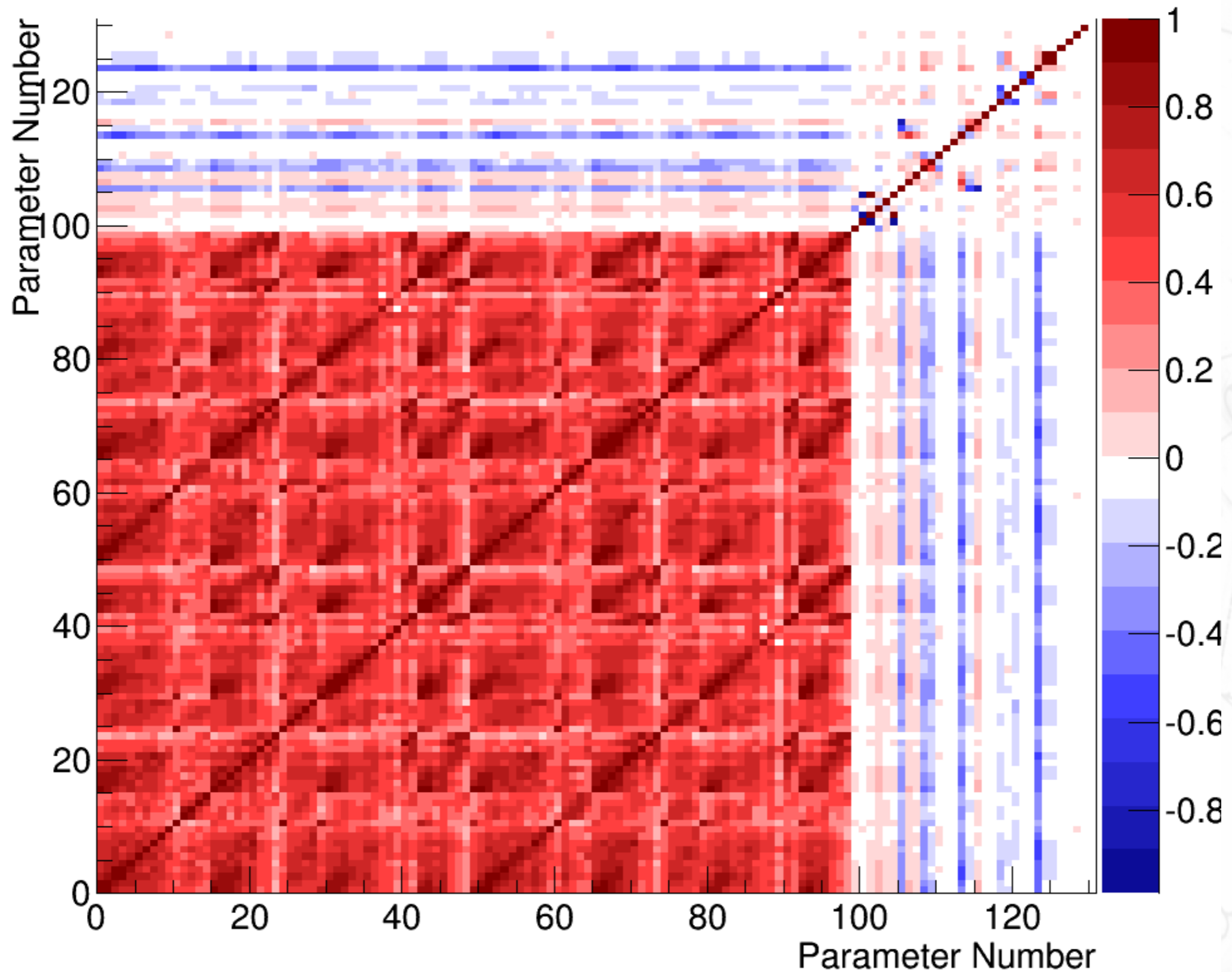


X-section parameters

ND fits



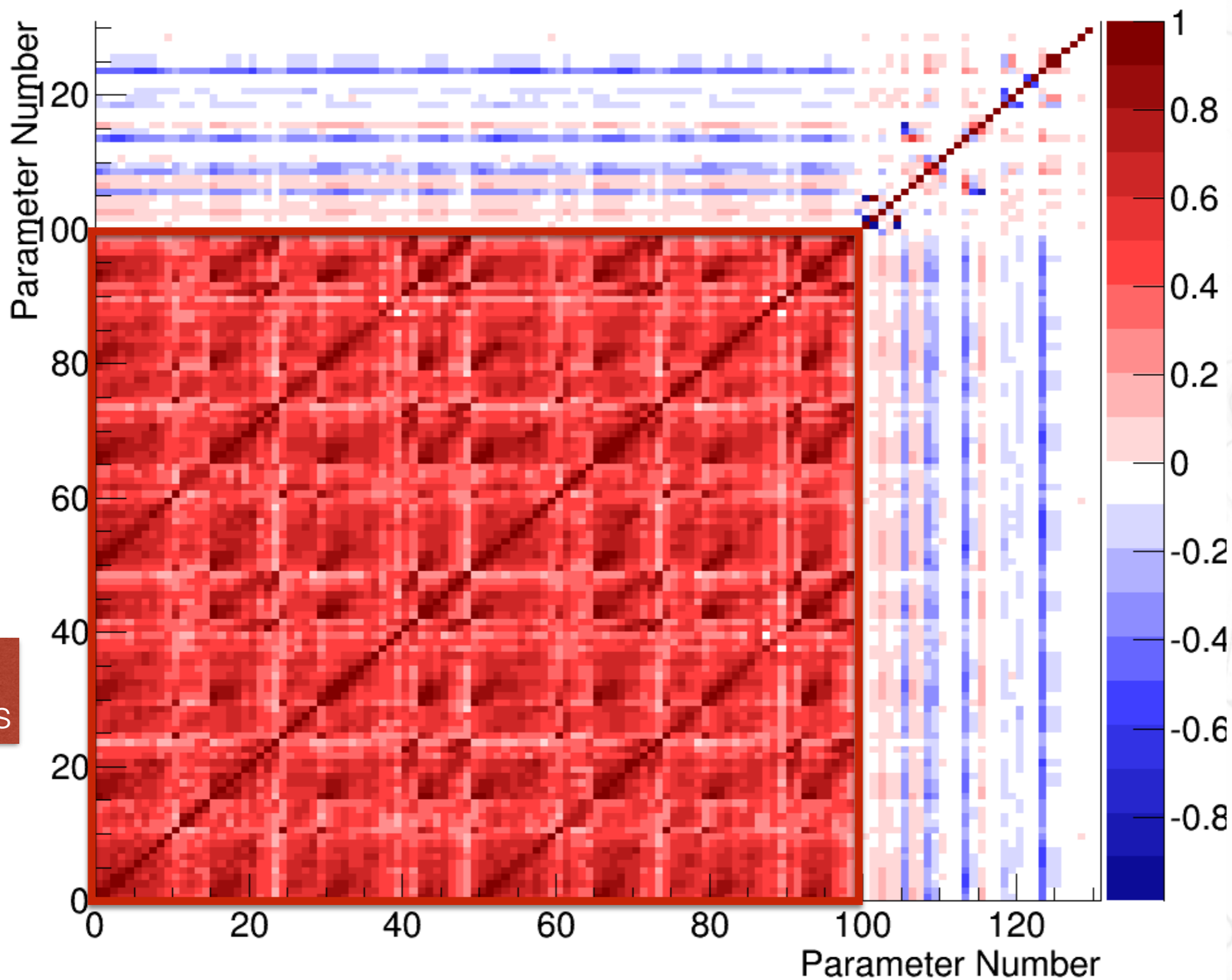
Postfit Correlation Matrix



ND fits



Postfit Correlation Matrix

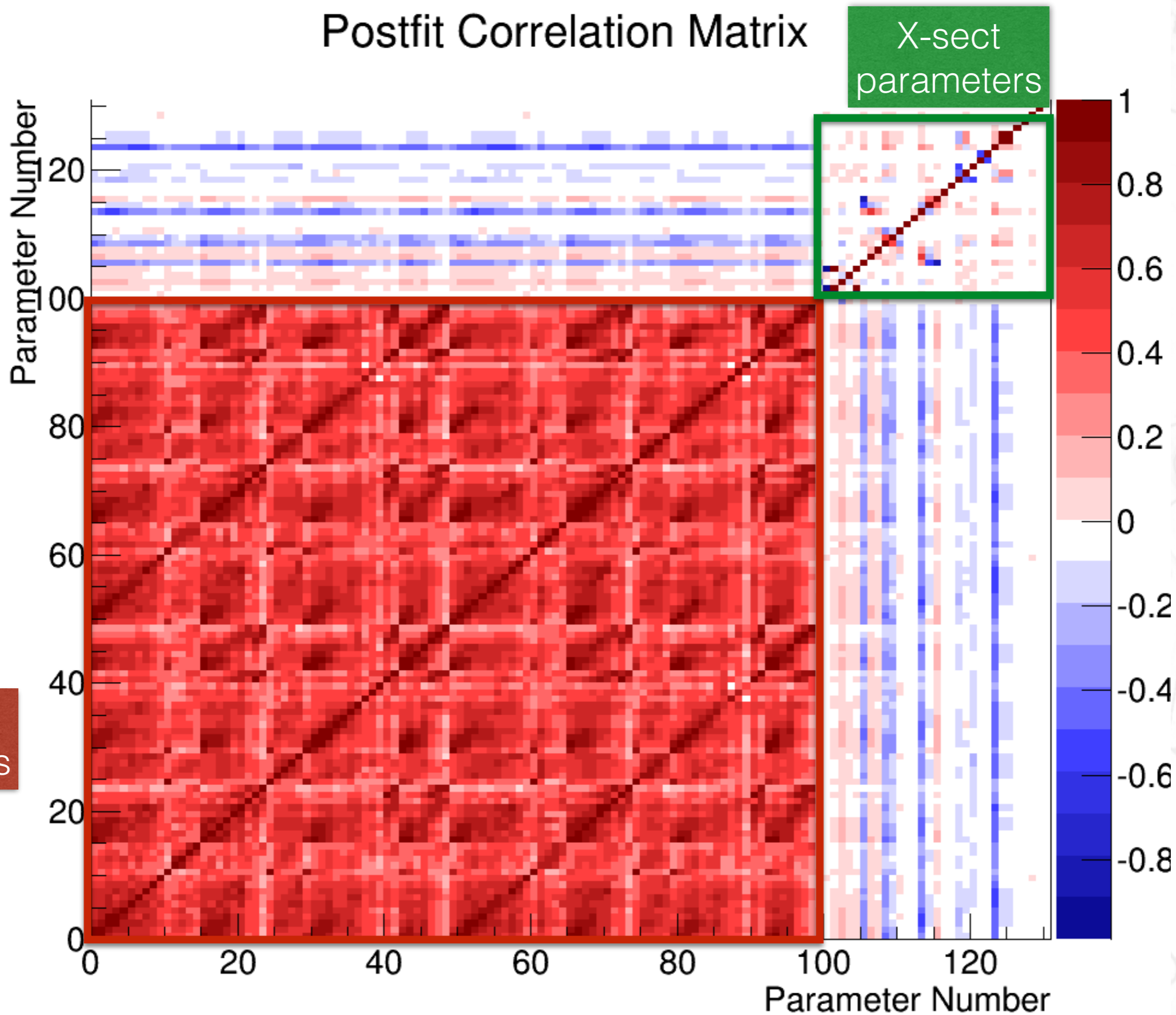


Flux
parameters

ND fits



Postfit Correlation Matrix



Flux parameters

X-sect parameters

SK data **T2K**

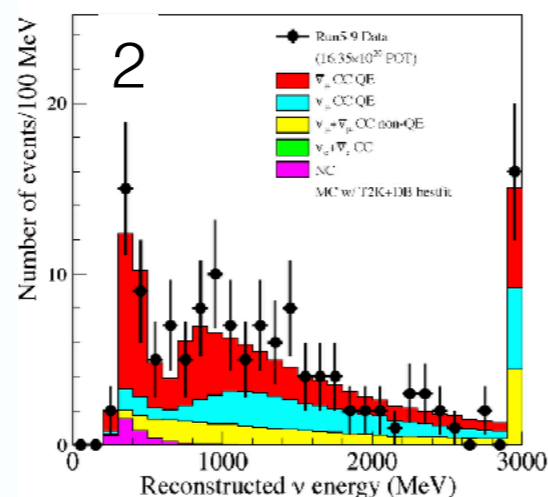
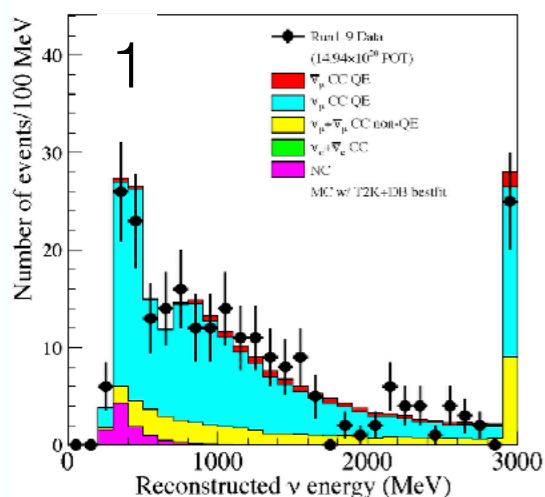
Neutrino mode

Anti-neutrino mode

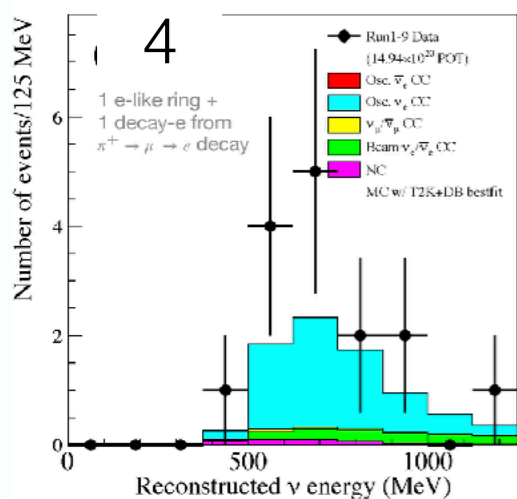
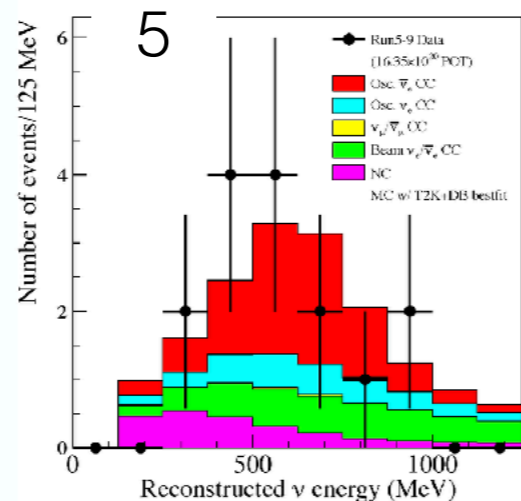
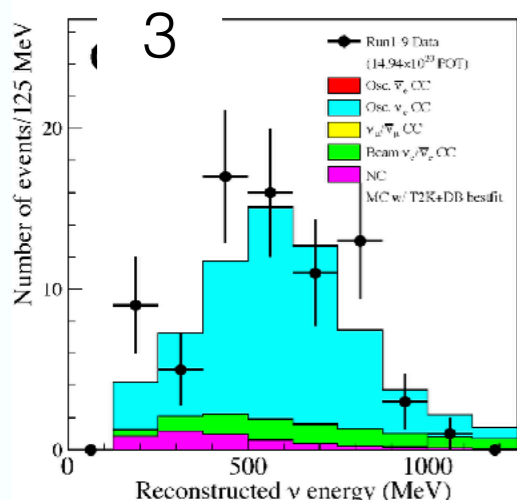
5 SK samples

1. muon candidate, neutrino mode.
2. muon candidate, antineutrino model
3. electron candidates, neutrino mode.
4. electron candidate with a charged pion (Michel electron) neutrino mode.
5. electron candidate, antineutrino mode

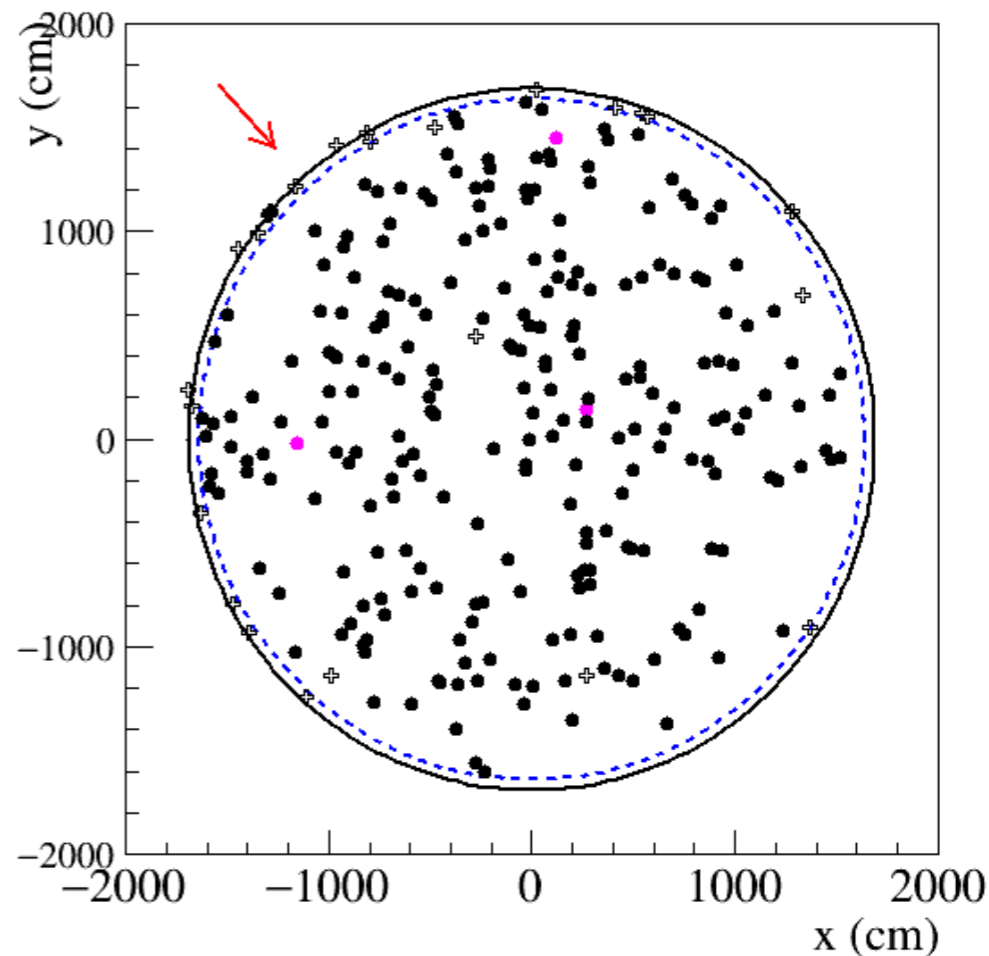
μ-like ring



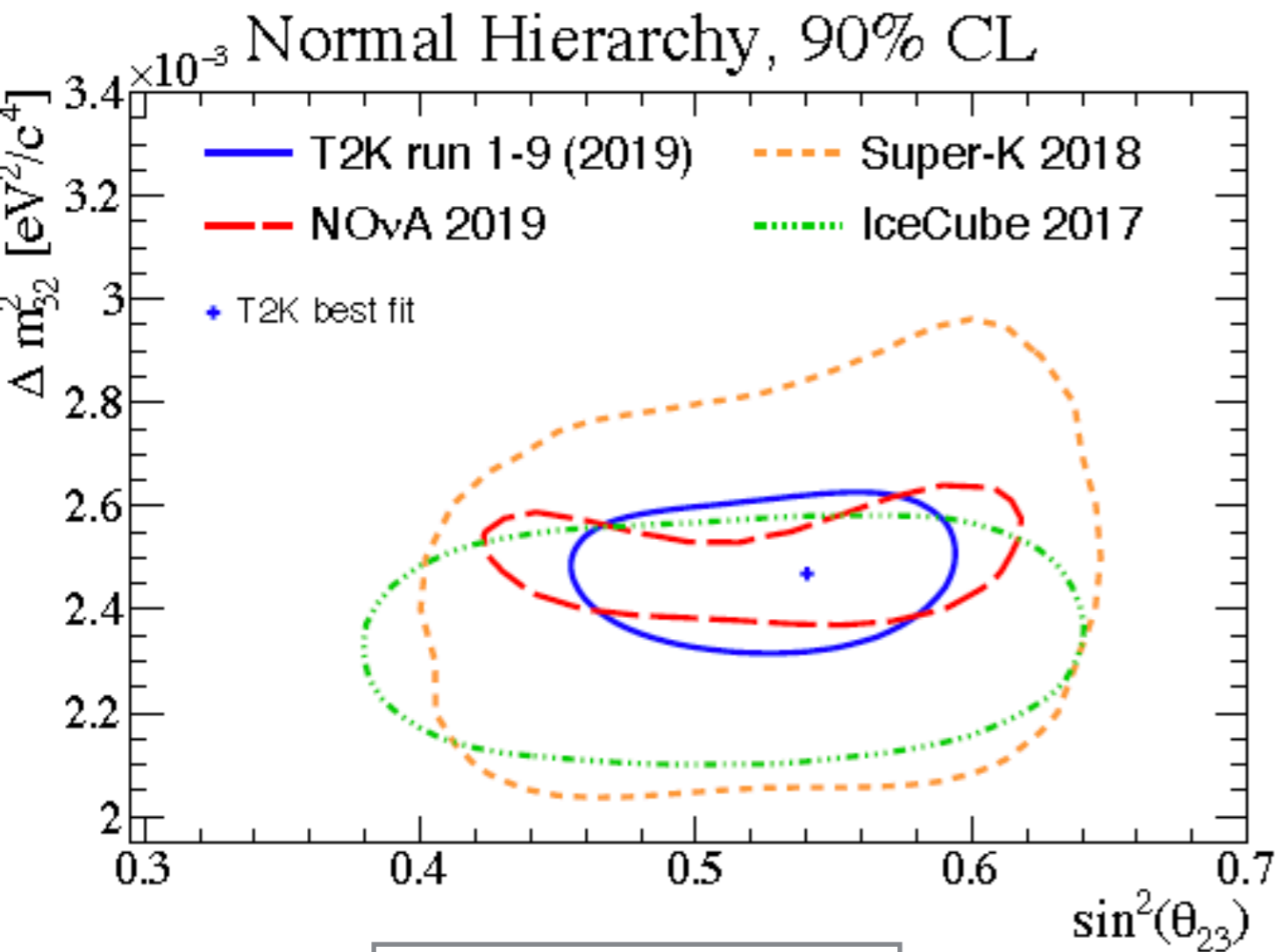
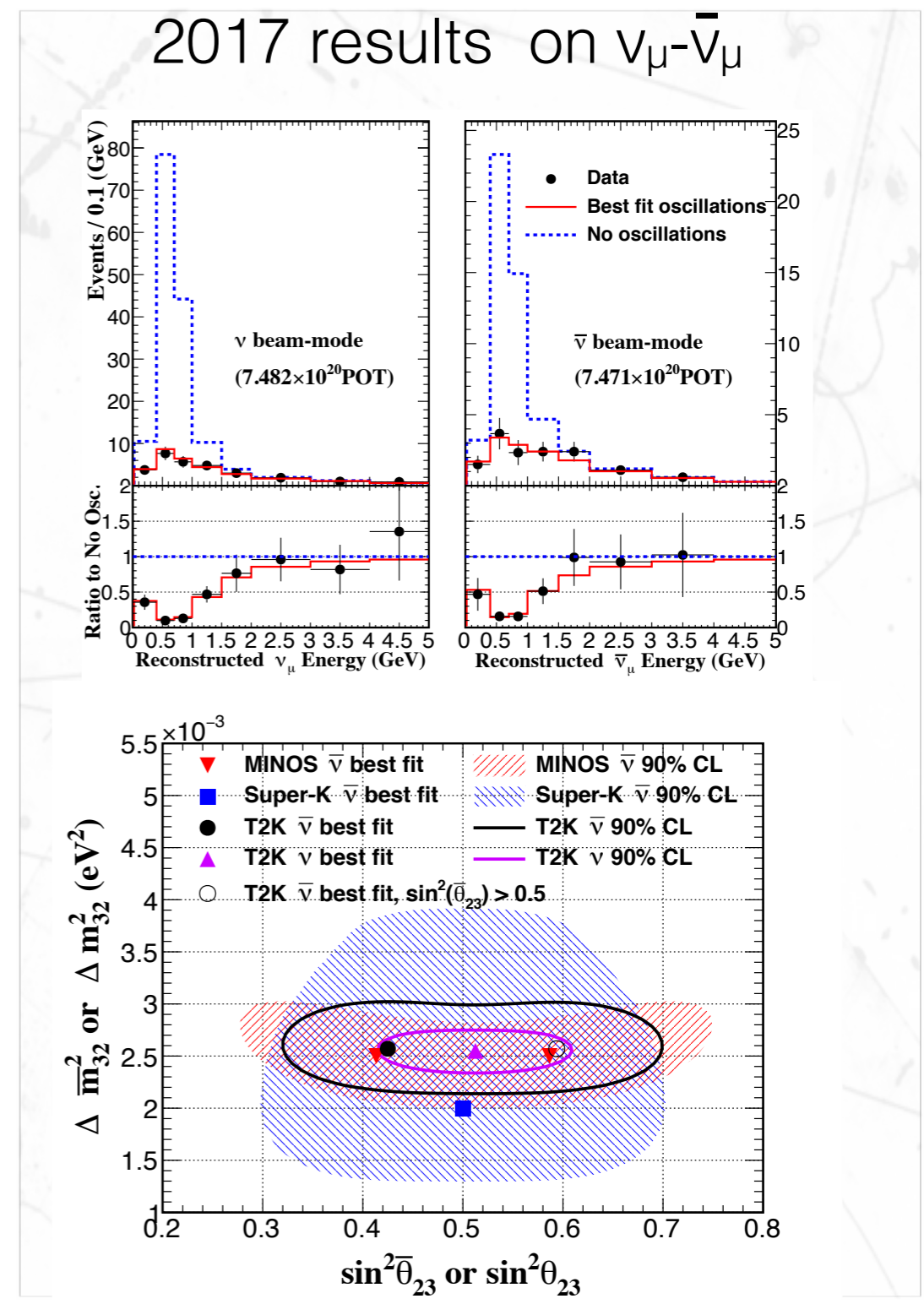
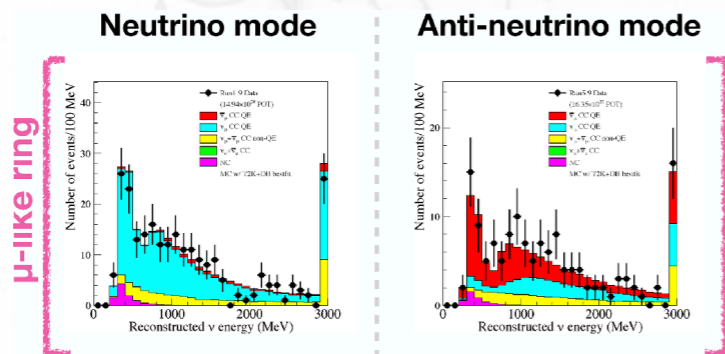
e-like ring



No CC1π sample in anti-neutrino mode because π^- produced in $\bar{\nu}$ interaction are mostly absorbed before decay.



$\bar{\nu}_\mu$ disappearance



World leading result

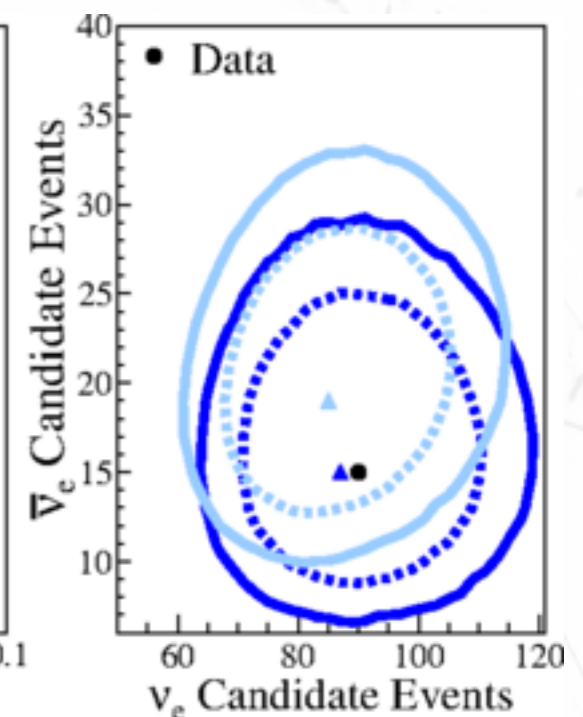
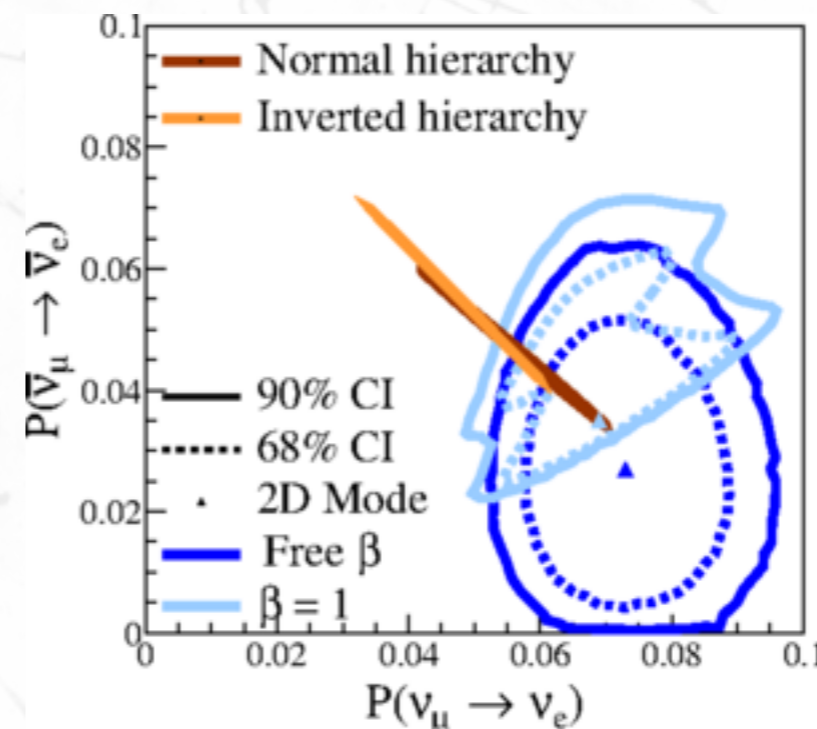
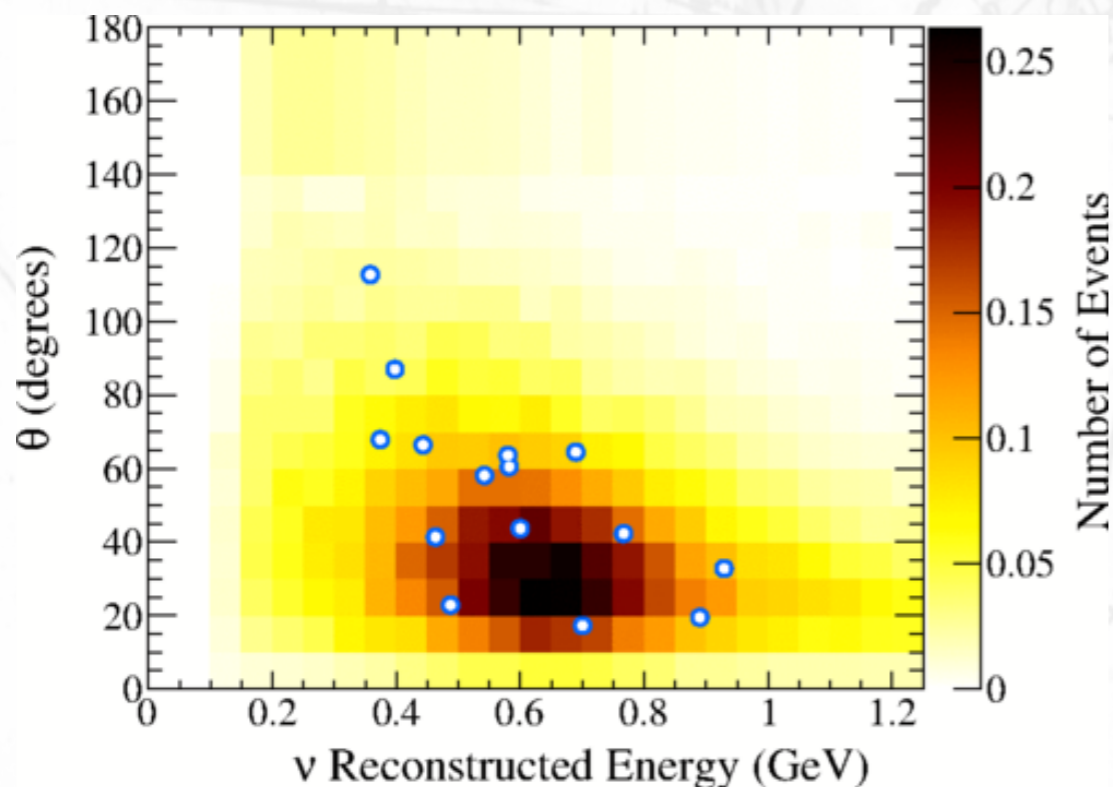
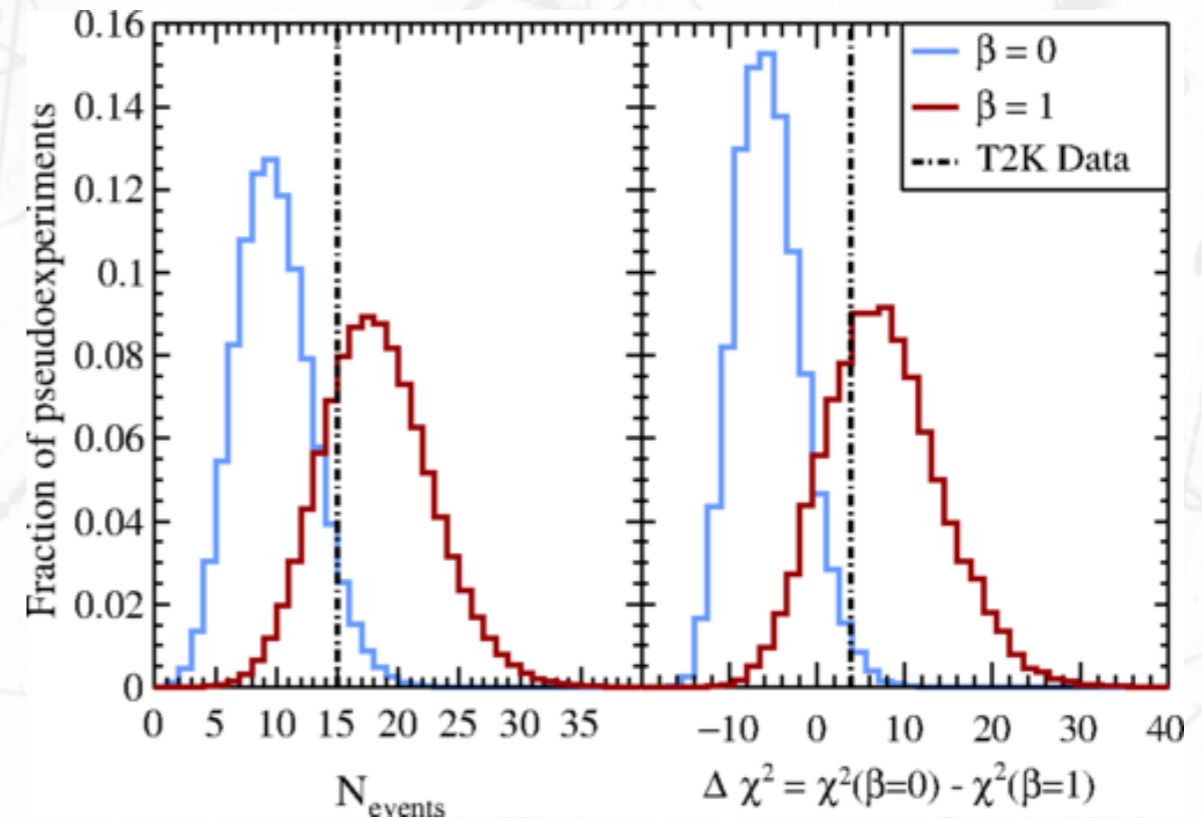
$\bar{\nu}_e$ appearance



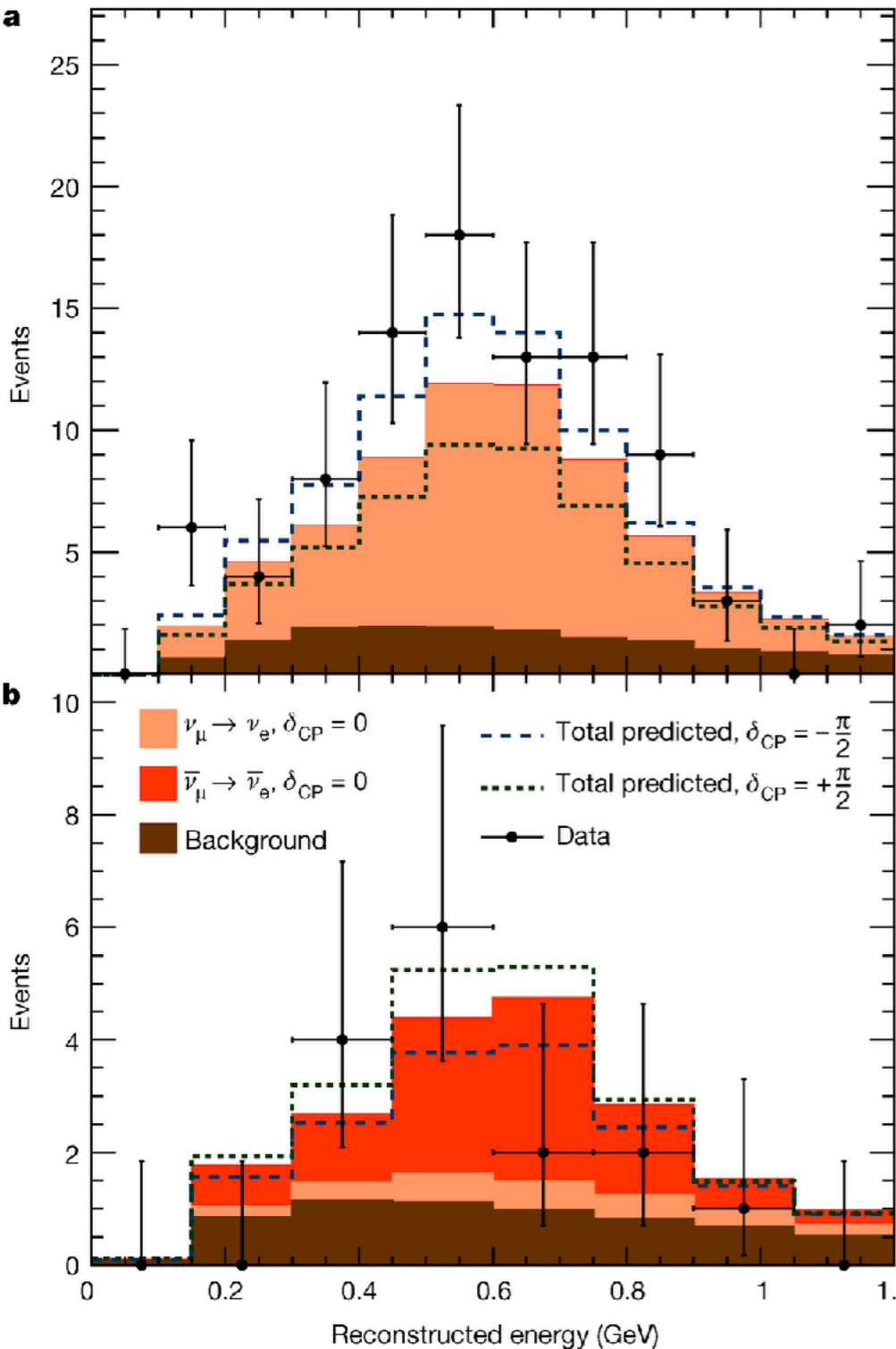
- High background from ν_e .
- strategy is to parametrise the oscillation and measure β

$$P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e) = \beta \times P_{\text{PMNS}}(\bar{\nu}_\mu \rightarrow \bar{\nu}_e).$$

- No $\bar{\nu}_e$ appearance disfavoured to 2.4σ



CP violation phase



c

	1e0de ν -mode	1e0de $\bar{\nu}$ -mode	1e1de ν -mode
$\nu_{\mu} \rightarrow \nu_e$	59.0	3.0	5.4
$\bar{\nu}_{\mu} \rightarrow \bar{\nu}_e$	0.4	7.5	0.0
Background	13.8	6.4	1.5
Total predicted	73.2	16.9	6.9
Systematic uncertainty	8.8%	7.1%	18.4%
Data	75	15	15

$\nu_e/\bar{\nu}_e$ Systematic Uncertainty

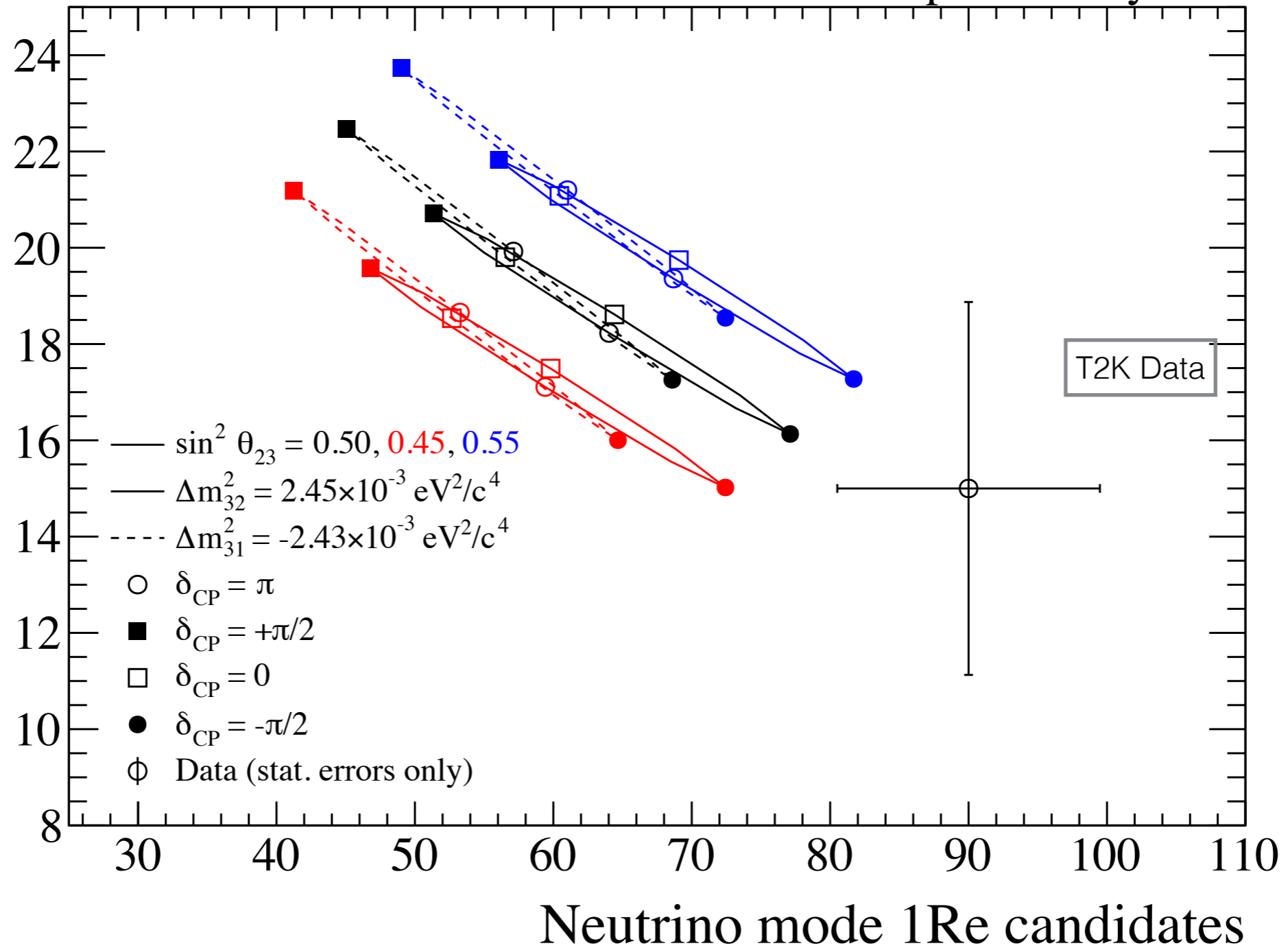
Type of Uncertainty	$\nu_e/\bar{\nu}_e$ Candidate Relative Uncertainty (%)
Super-K Detector Model	1.5 %
Pion Final State Interaction and Rescattering Model	1.6
Neutrino Production and Interaction Model Constrained by ND280 Data	2.7
Electron Neutrino and Antineutrino Interaction Model	3.0
Nucleon Removal Energy in Interaction Model	3.7
Modeling of Neutral Current Interactions with Single γ Production	1.5
Modeling of Other Neutral Current Interactions	0.2
Total Systematic Uncertainty	6.0 %

CP violation phase



T2K Run 1-9 preliminary

Antineutrino mode 1Re candidates



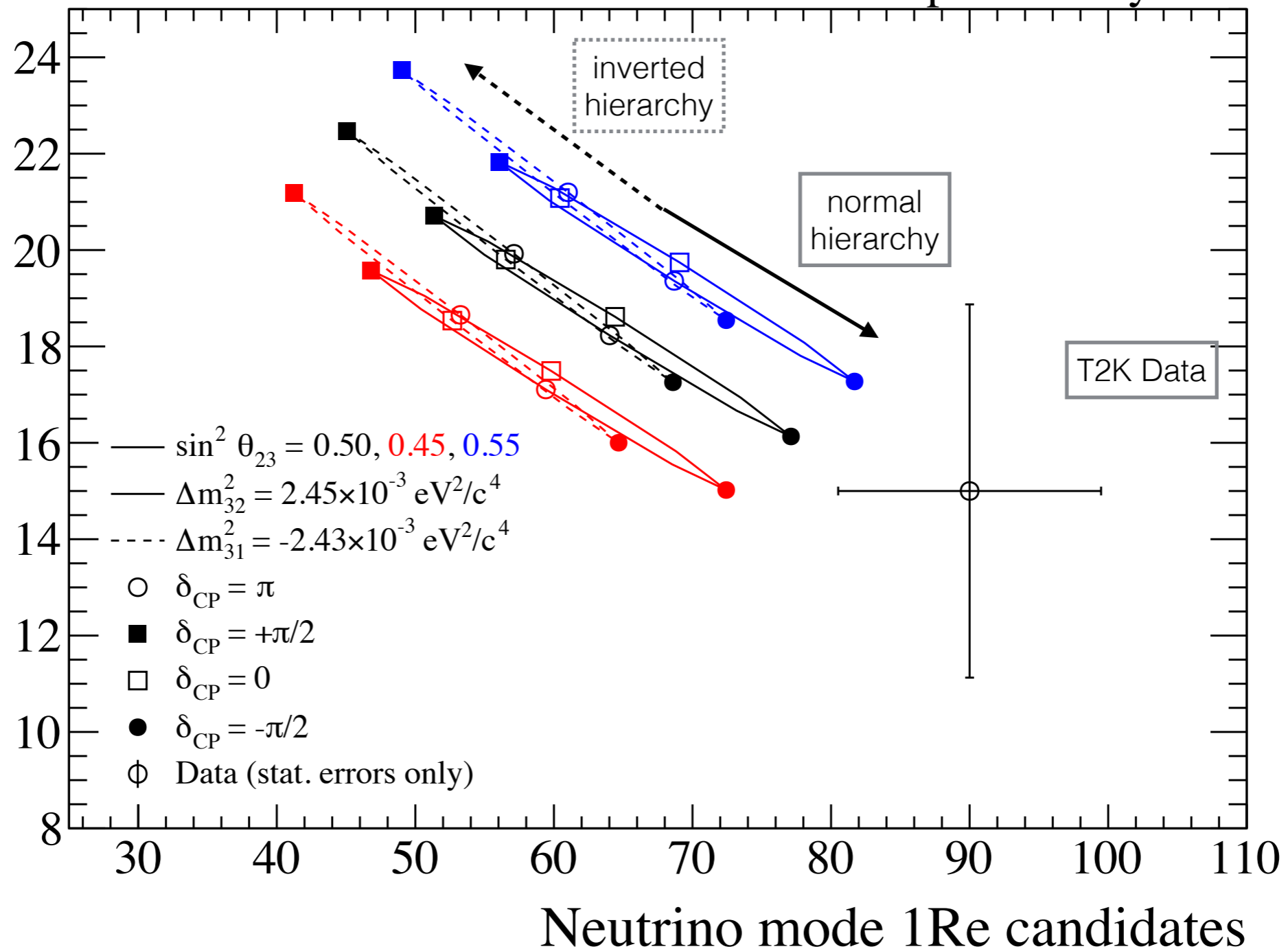
ν energy dependency is not reflected in this plot

CP violation phase



T2K Run 1-9 preliminary

Antineutrino mode 1Re candidates



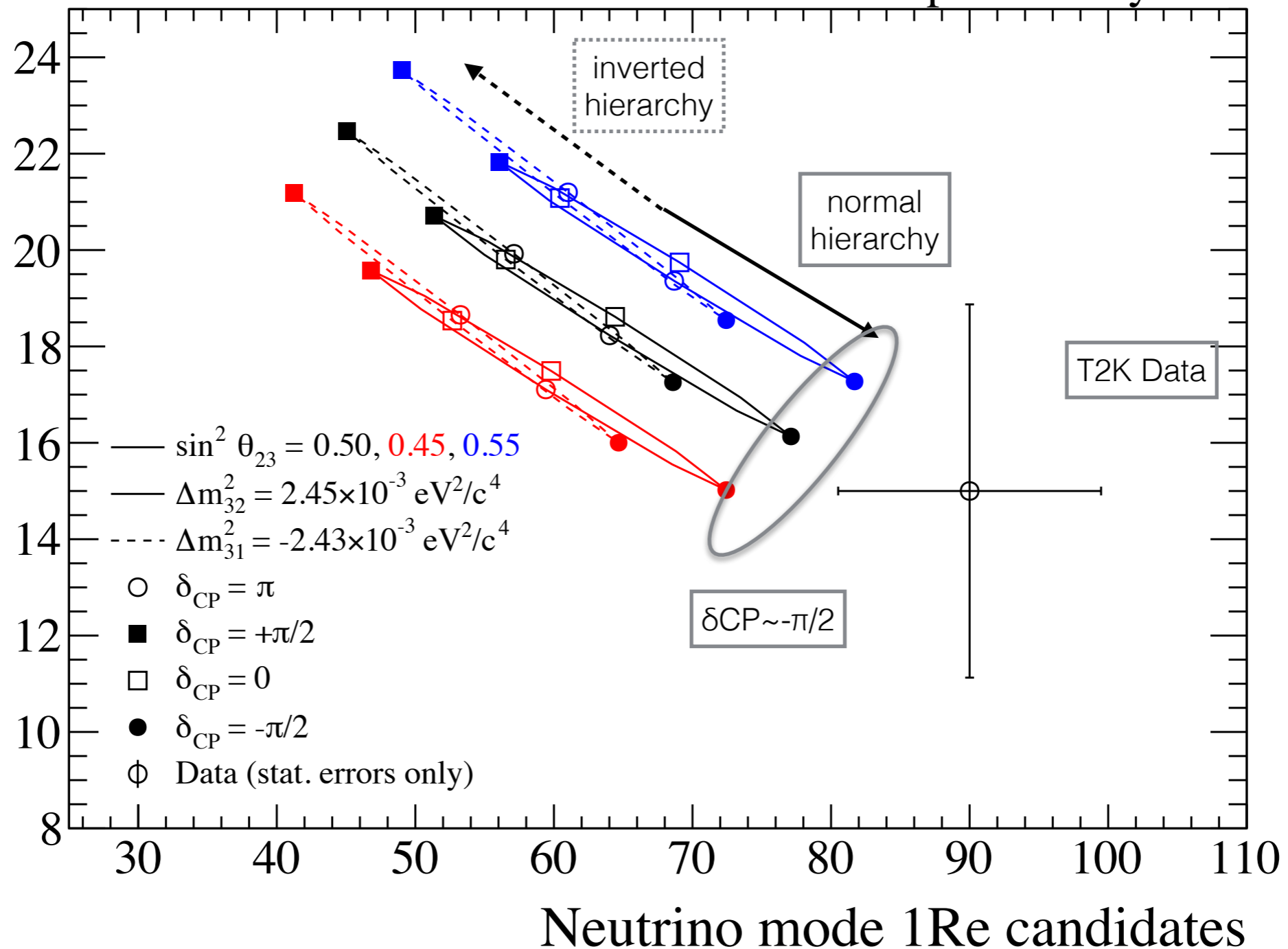
ν energy dependency is not reflected in this plot

CP violation phase



T2K Run 1-9 preliminary

Antineutrino mode 1Re candidates



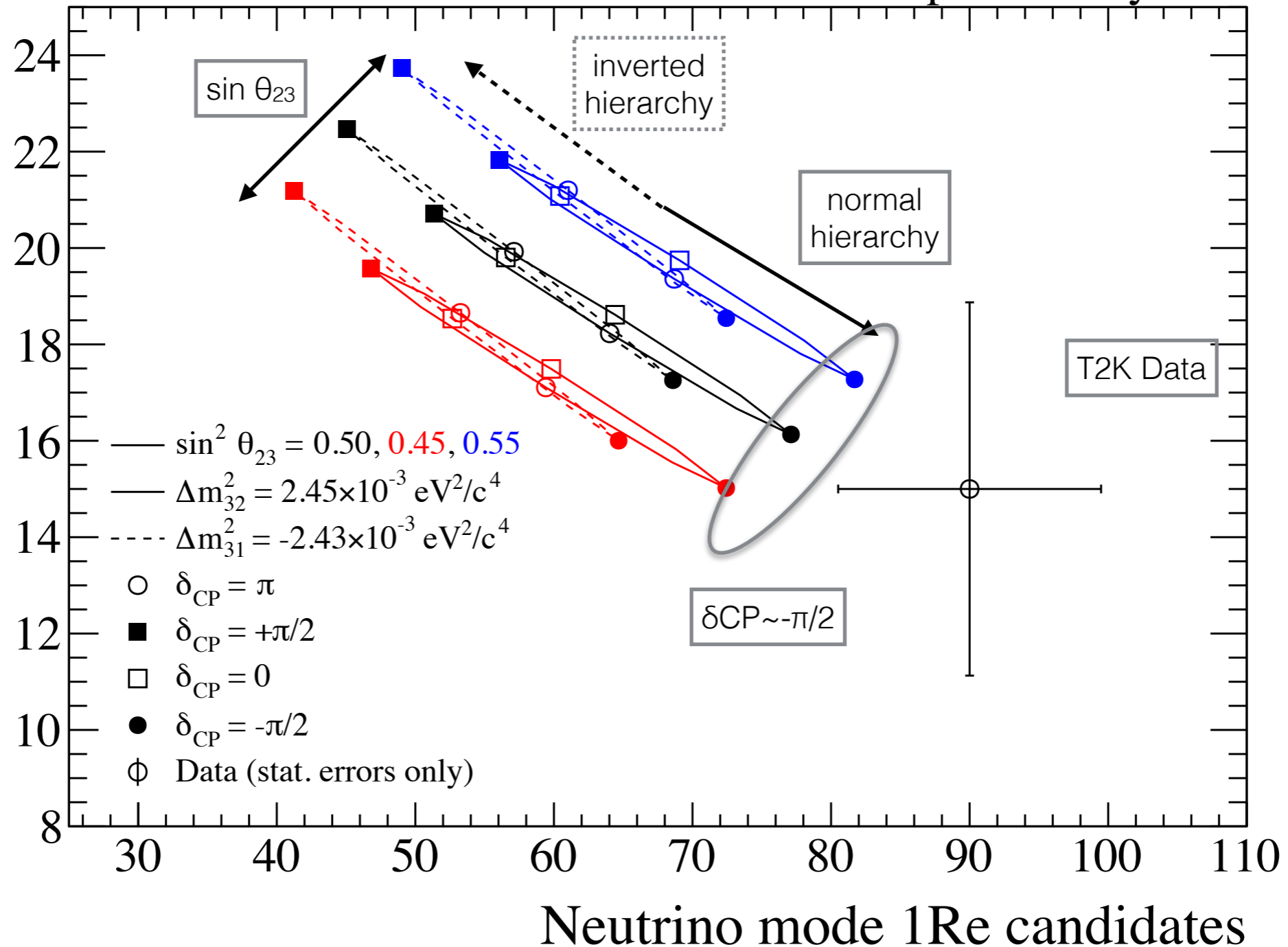
ν energy dependency is not reflected in this plot

CP violation phase



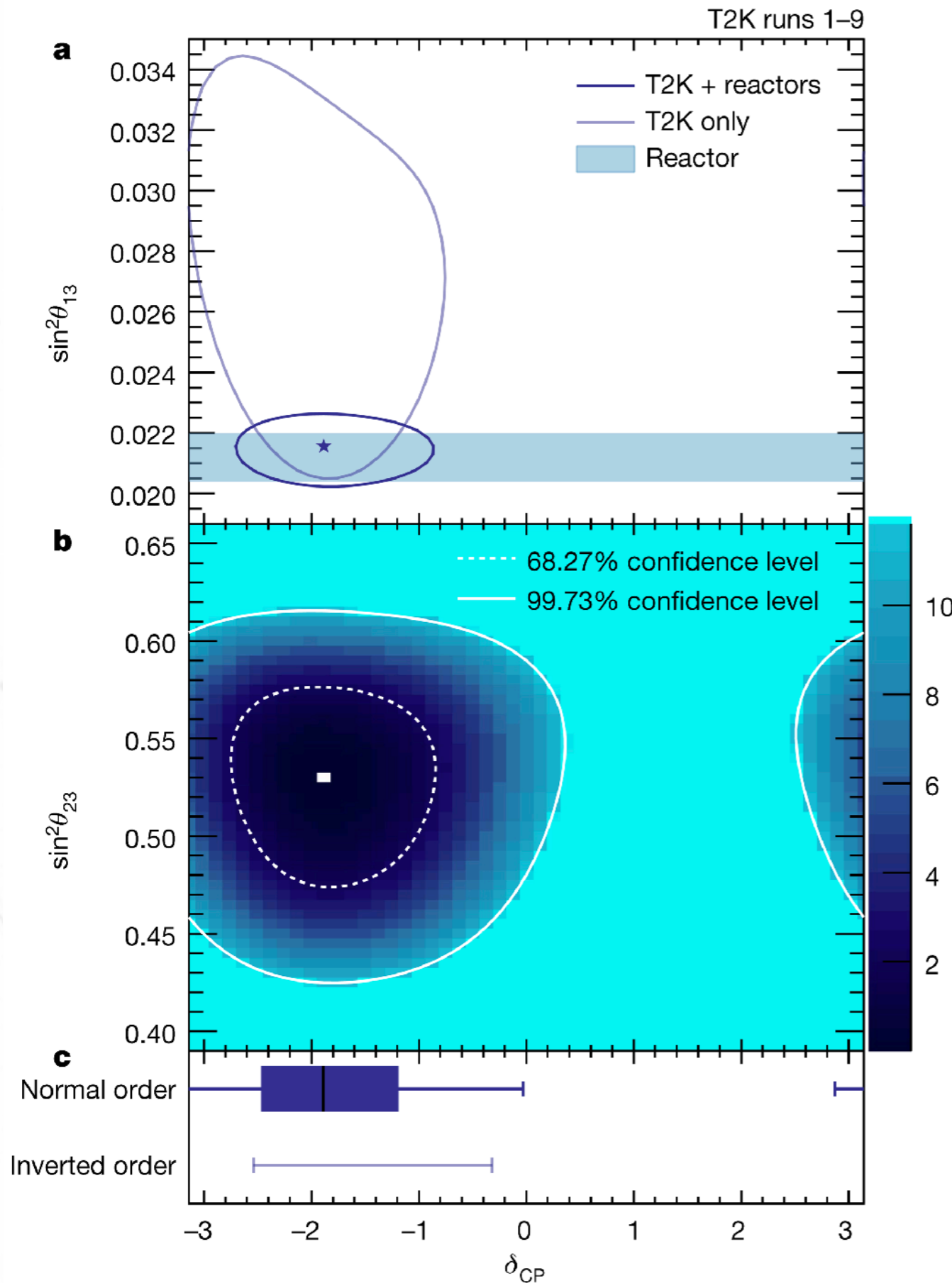
T2K Run 1-9 preliminary

Antineutrino mode 1Re candidates



v energy dependency is not reflected in this plot

δ_{CP} measurement



Fit uses the value of θ_{13} from reactor experiments

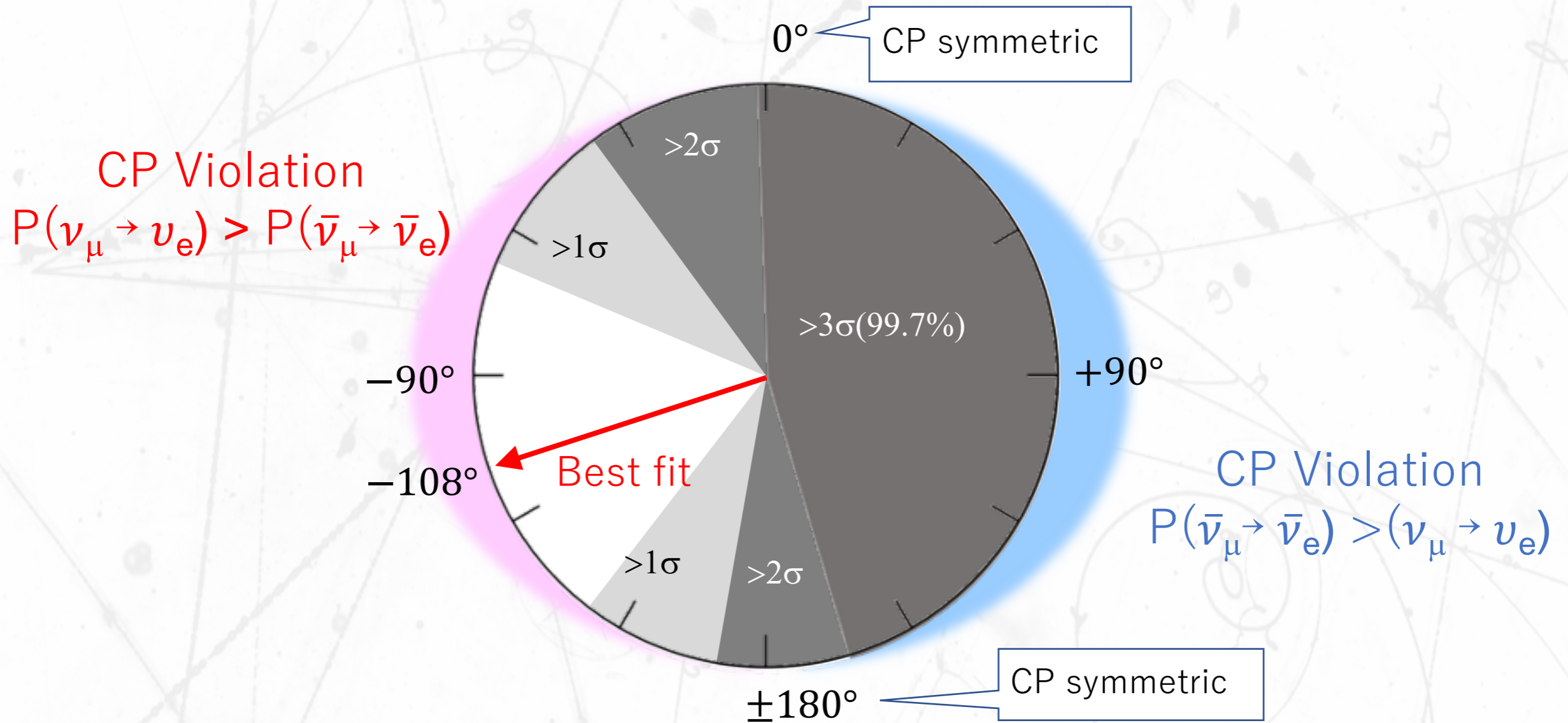
Data also prefers Normal Hierarchy with a posterior probability of 89%

CP violation phase



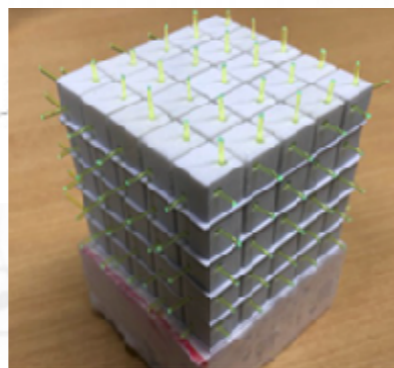
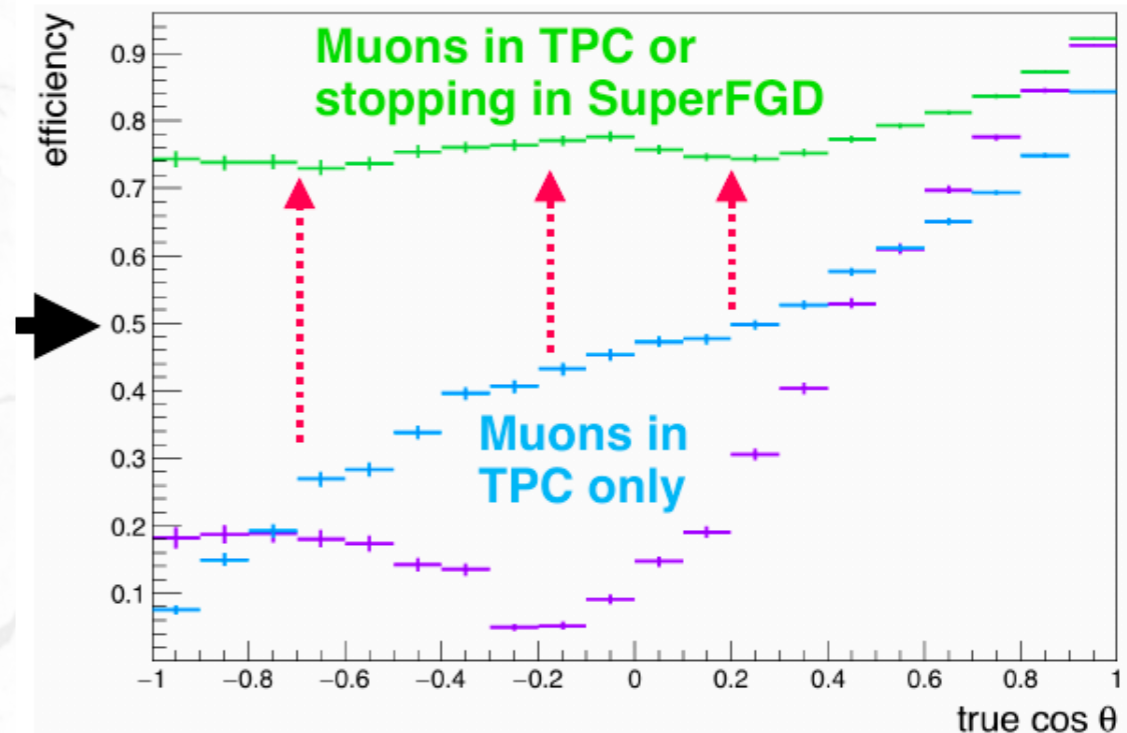
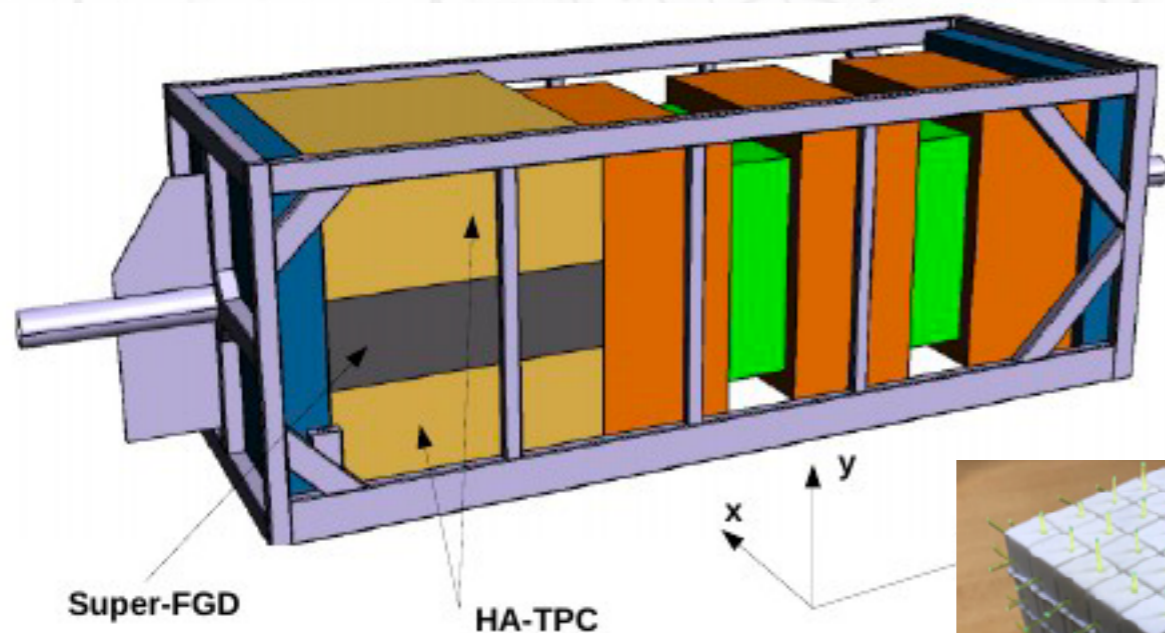
T2K result excludes most of the $\delta_{CP} > 0$ values @ 99.7% CL

CP phase

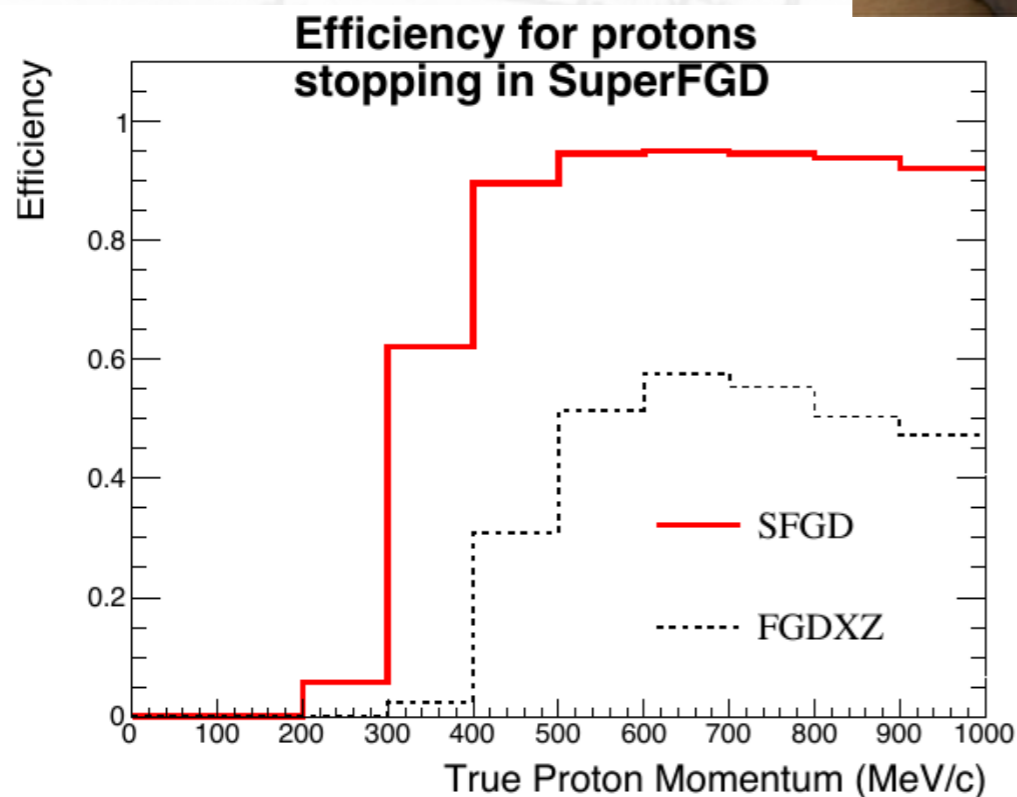


Next steps and beyond

ND280 upgrade



ND280 upgrade goals

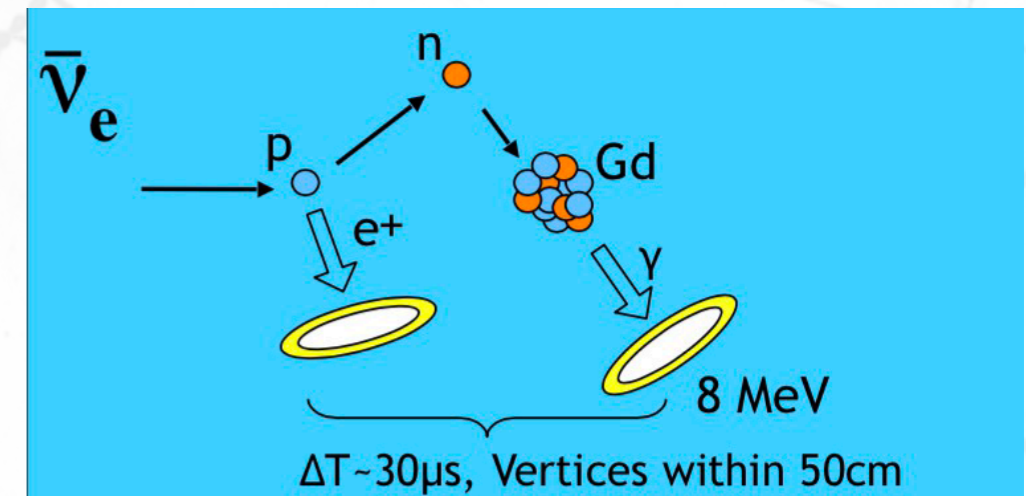


- quasi-3D imaging.
 - Improved target tracking.
 - Improved proton detection threshold.
 - neutron detection capabilities
- Improved high angle acceptance:
 - High Angle TPC's.
- x 2 in statistics for equal p.o.t.
- Time of Flight for background reduction.

SK Gadolinium



- SK Gadolinium project
 - enhance neutron detection
 - improve low-energy $\bar{\nu}_e$ detection (non-T2K goal).
 - may provide wrong-sign background constraint in $\bar{\nu}_e$
 - more data samples.
- Leak repairs to SK tank finished in 2019.
- Load $\text{Gd}_2(\text{SO}_4)_3$ in stages up to 0.2%.
- Loading to start in 2020.

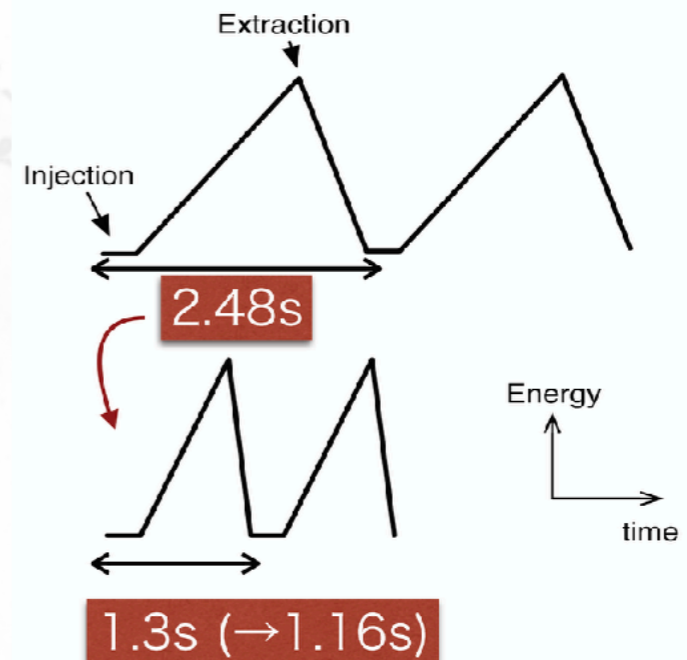
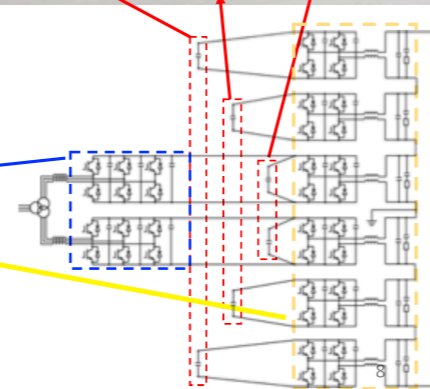


Beam upgrade



- A new power supply was designed with capacitor banks for the cycle of 1.3 s.
- The power supply for the BM3 family was constructed and installed at D4.
- It has been tested with the BM3 family.

Capacitor Banks for BM3



$$f_{\text{rep}}=0.4 \text{ Hz} \oplus \text{PPP} = 2.7 \times 10^{14} \oplus 30 \text{ GeV} = 515 \text{ kW}$$

515 kW stable operation in 2019

MR Power Supply approved JFY 2020

$$f_{\text{rep}}=0.77 \text{ Hz} \oplus \text{PPP} = 2.2 \times 10^{14} \oplus 30 \text{ GeV} = 810 \text{ kW}$$

exp. >800 kW by 2023

RF upgrade and Machine development

$$f_{\text{rep}}=0.86 \text{ Hz} \oplus \text{PPP} = 3.2 \times 10^{14} \oplus 30 \text{ GeV} = 1.3 \text{ MW}$$

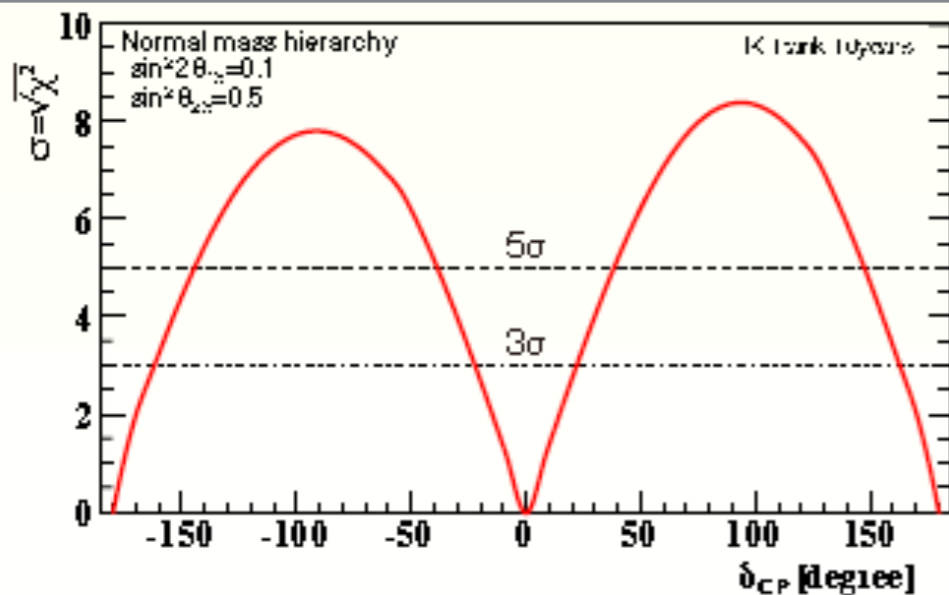
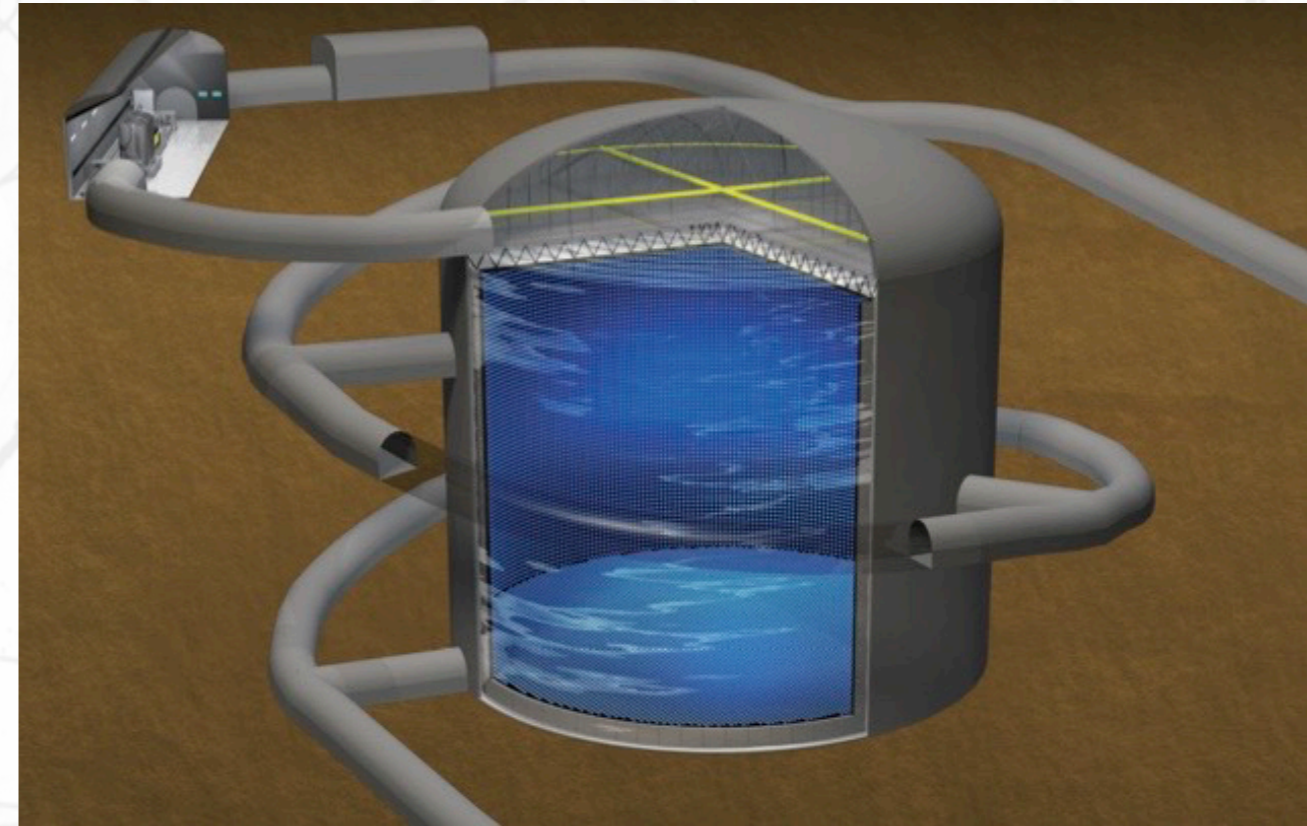
HyperKamiookande



Approved early 2020

- 1000-2000 $\nu_e + \bar{\nu}_e$ events.
 - 115 in T2K
- $> 5\sigma$ discovery of CP violation.
- Precise measurement of θ_{23}

Same neutrino spectrum as T2K.
 Same ND280 as in T2K
 T2K results in x-sect and oscillation
 fosters future HK results.



	SK	HK
Site depth	Mozumi (1000m)	Tochibora (650m)
# PMT	11,129	40,000
Photo-coverage	40 %	40% (x2 QE)
Mass	50 ktons	260 ktons
Fiducial mass	22.5 ktons	188 ktons

Conclusions & step forward



- Long Base Line technology is very **mature**.
 - research of years show the requirements for a **precision measurement using this technology**.
 - This includes **hadron production experiments, nuclear theory, beam monitor technology and advanced statistical methods**.
- Closing the **measurement of the PMNS matrix**.
 - Atmospheric angle close to maximal.
 - Rejected $\delta_{CP} > 0$ with 99.7% C.L.
 - Mild preference for normal hierarchy.
- T2K improving **statistical and systematic errors** control:
 - **Upgrade of ND280** for x-section
 - Use of **NA61/SHINE T2K replica-target** measurements will reduce significantly the flux errors.
- And, this is just the beginning:
 - **SK-Gd** may improve the results for antineutrinos.
 - T2K measurements paves the road for the approved **HyperKamiokande!**

T2K Collaboration thanks NA61 and CERN for their invaluable contribution to the success of the experiment.

nature



THE MIRROR CRACK'D

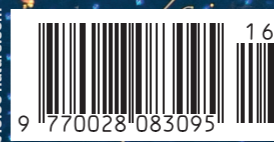
An indication of matter–antimatter
symmetry violation in neutrinos

Coronavirus
The models driving
the global response
to the pandemic

Hot source
Remnants of
primordial nitrogen
in Earth's mantle

Origin of a species
Revised age for Broken
Hill skull adds twist to
human evolution

Vol. 580, No. 7803
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Supporting slides

Combined analysis

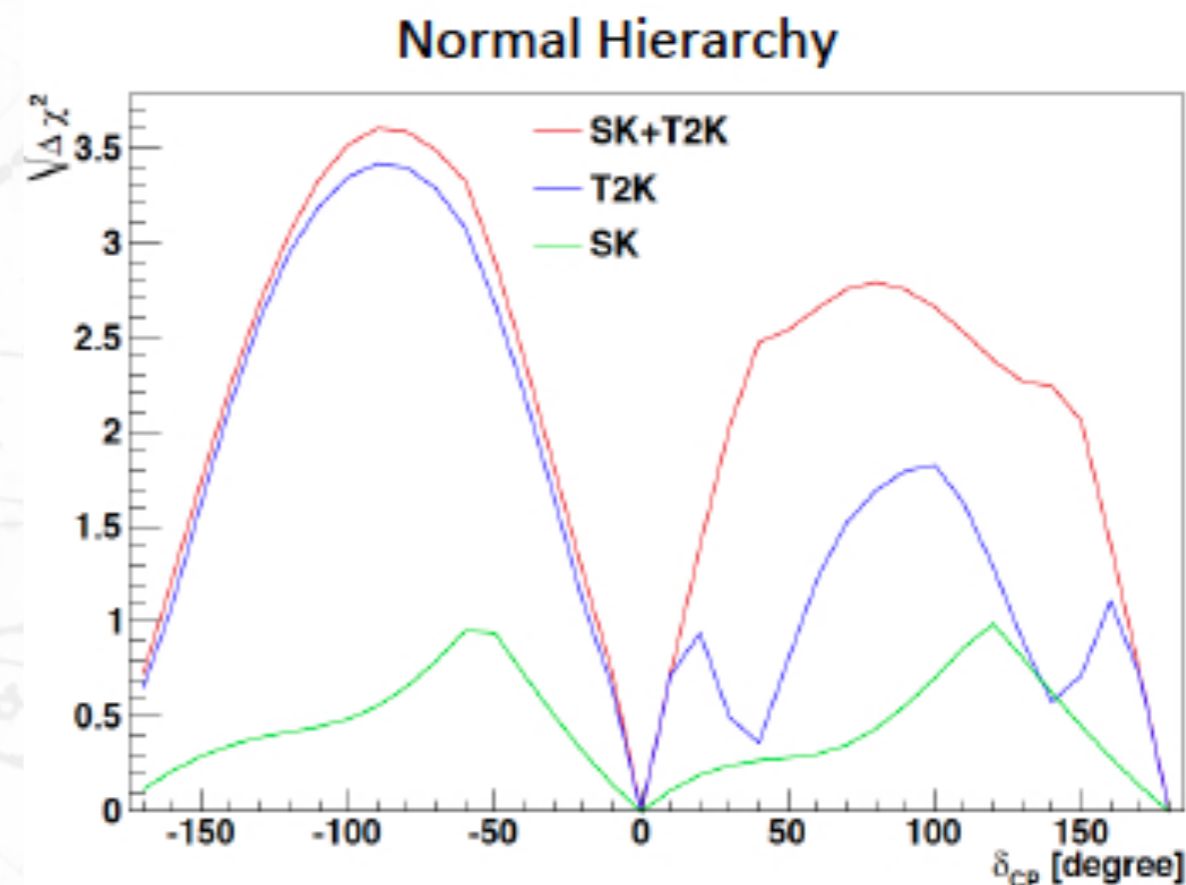
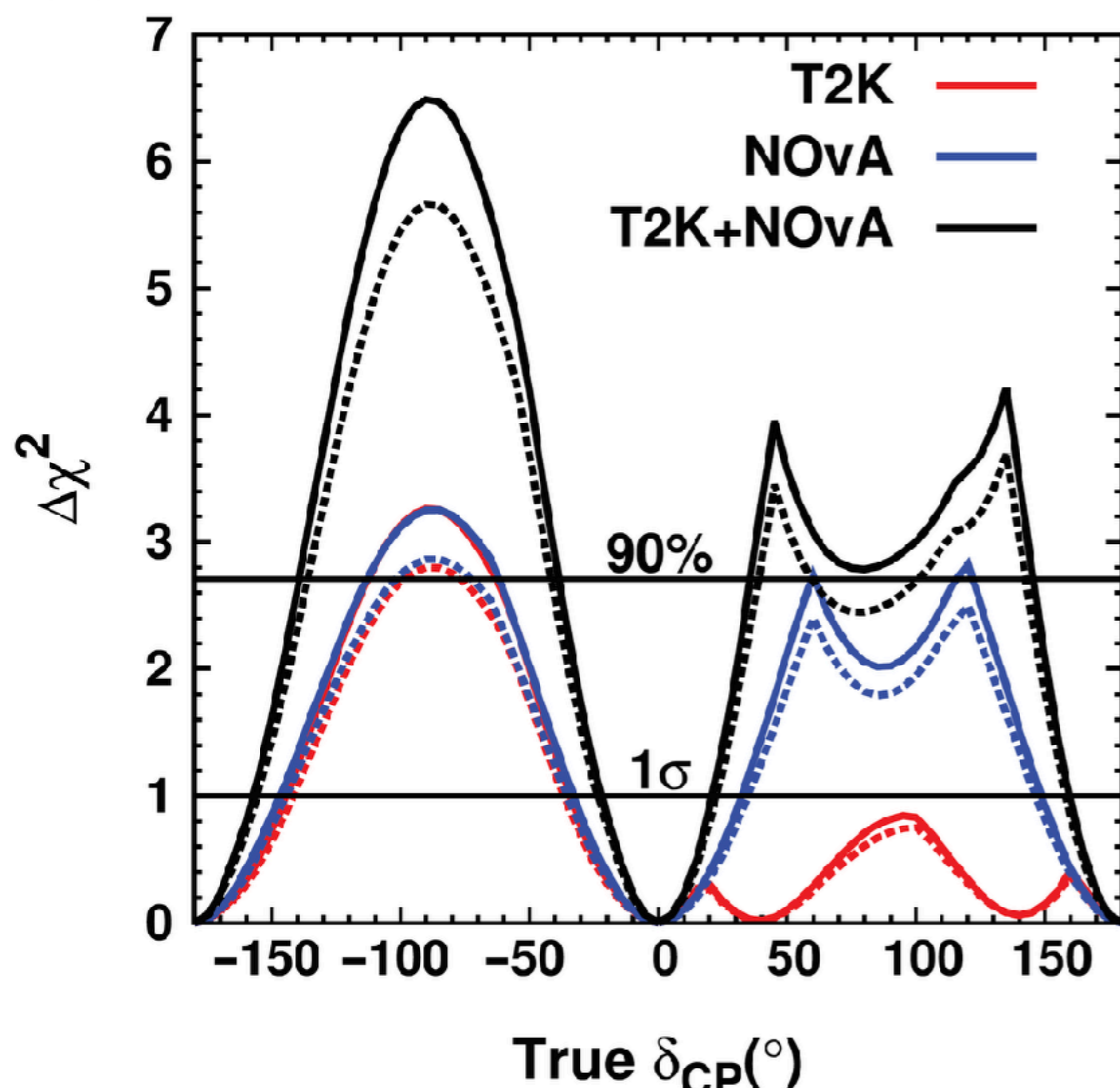


Sensitivity
T2K+Nova

*Progress of Theoretical
and Experimental Physics,*
Volume 2015, Issue 4,
April 2015, 043C01

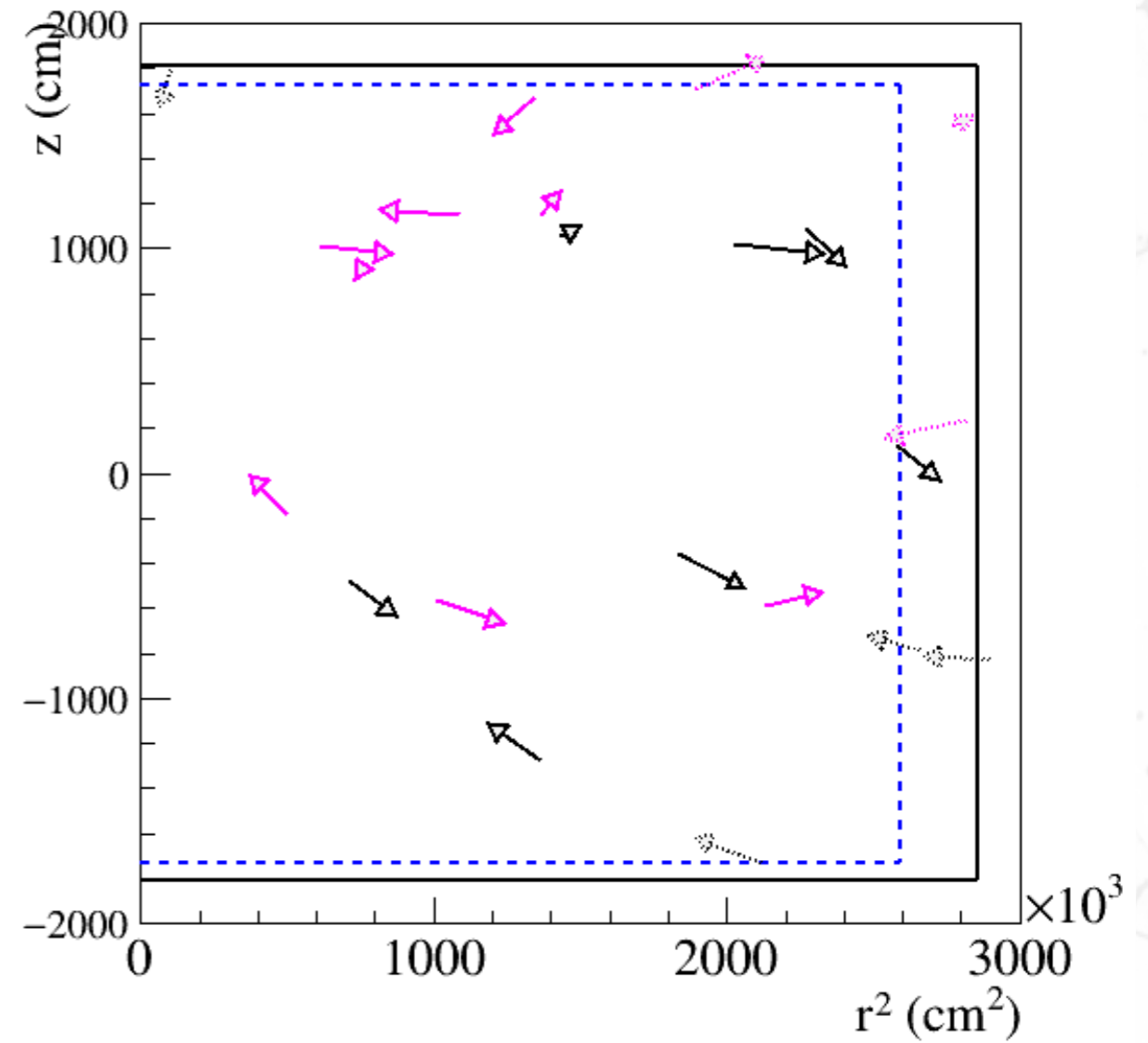
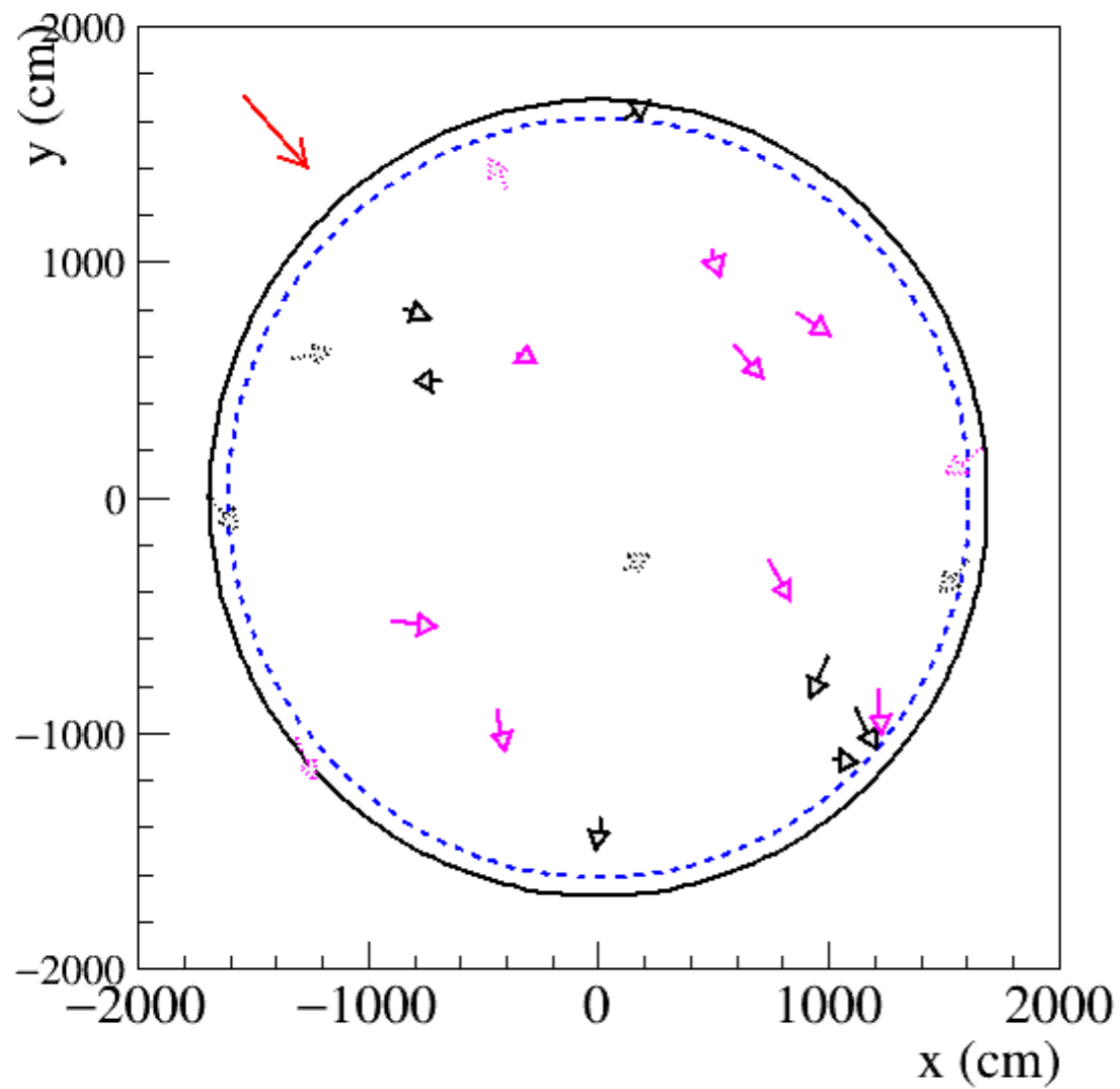
Sensitivity T2K+SK

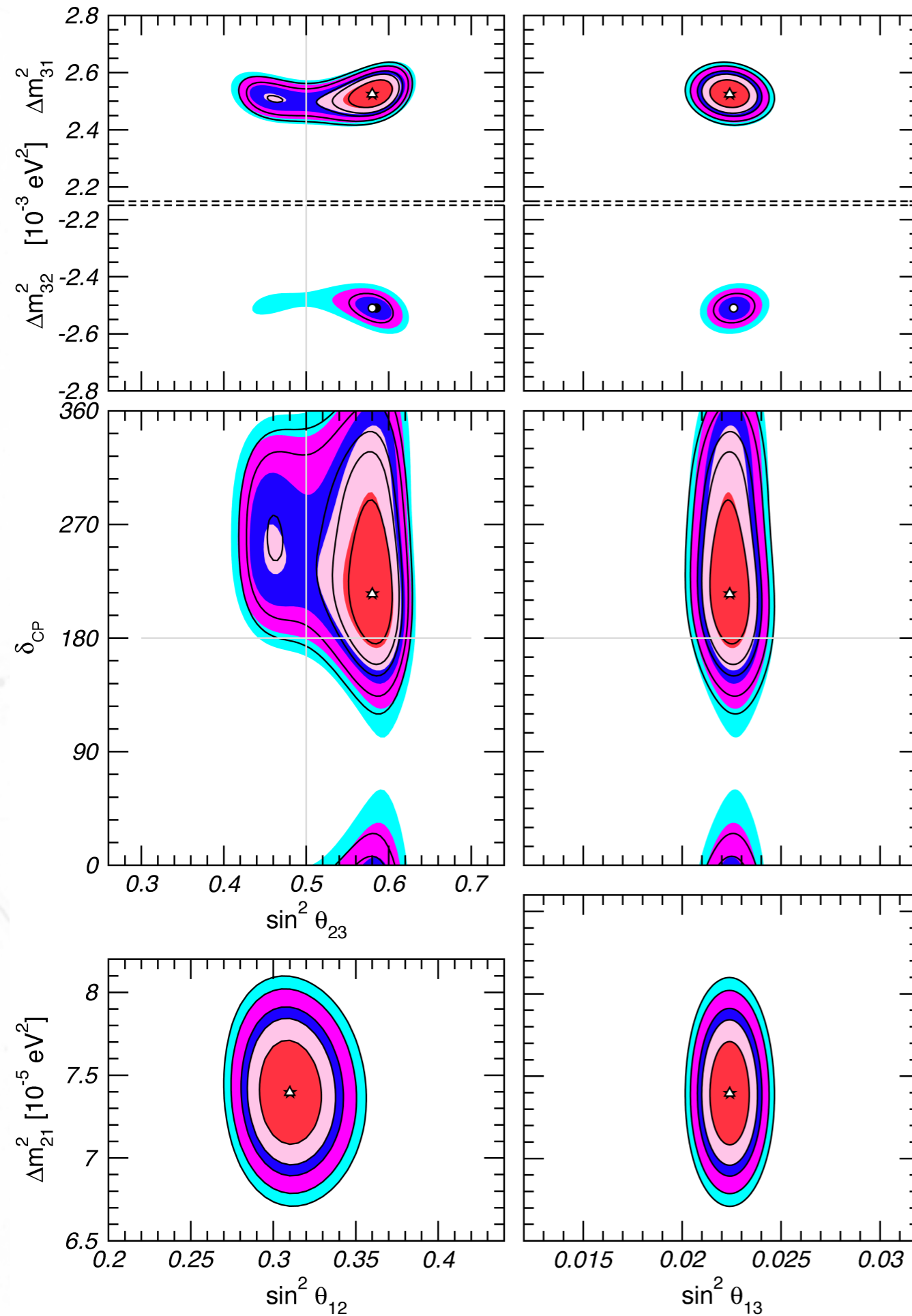
(b)



T2K formed working groups with Nova and SK to provide combined oscillation analyses

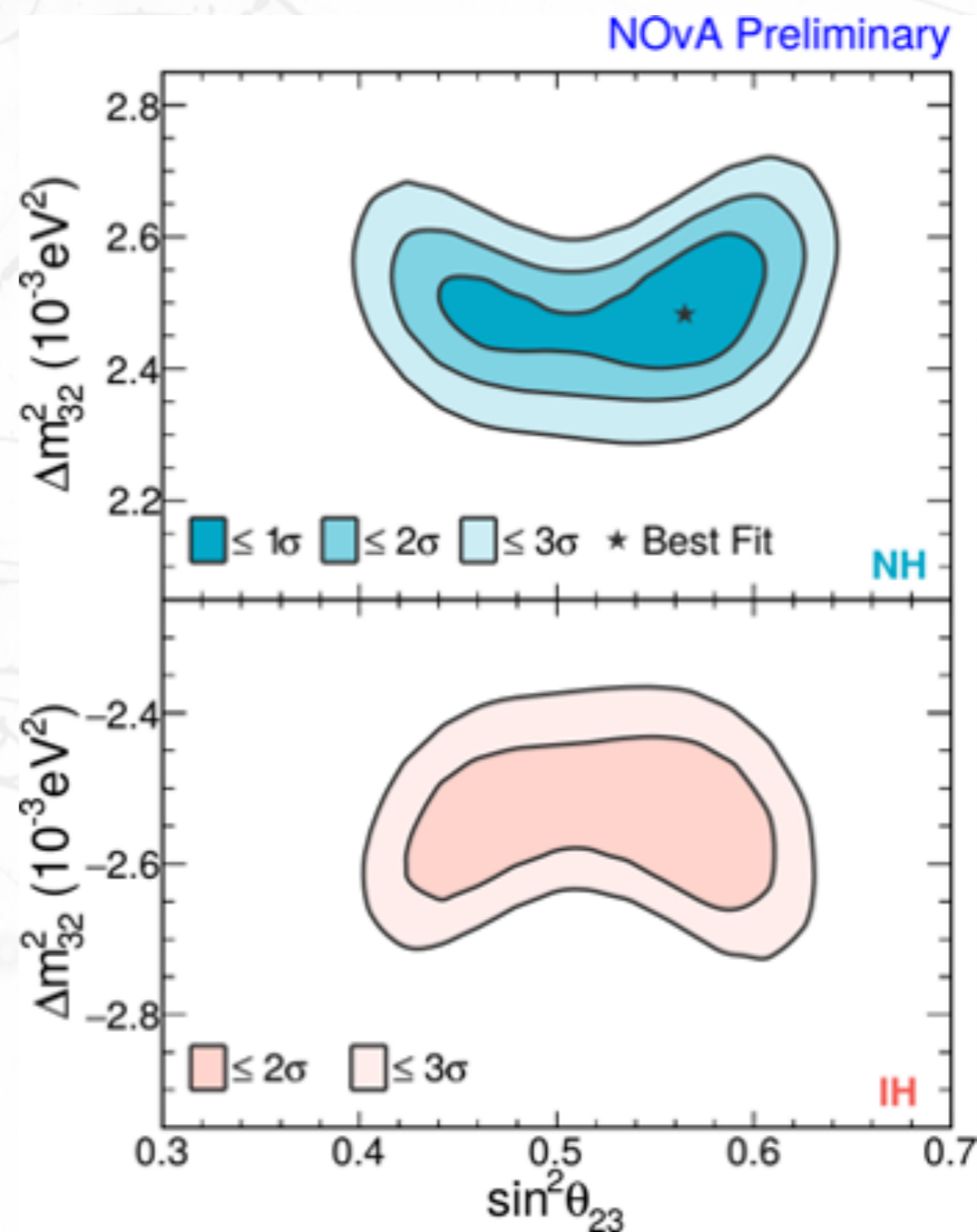
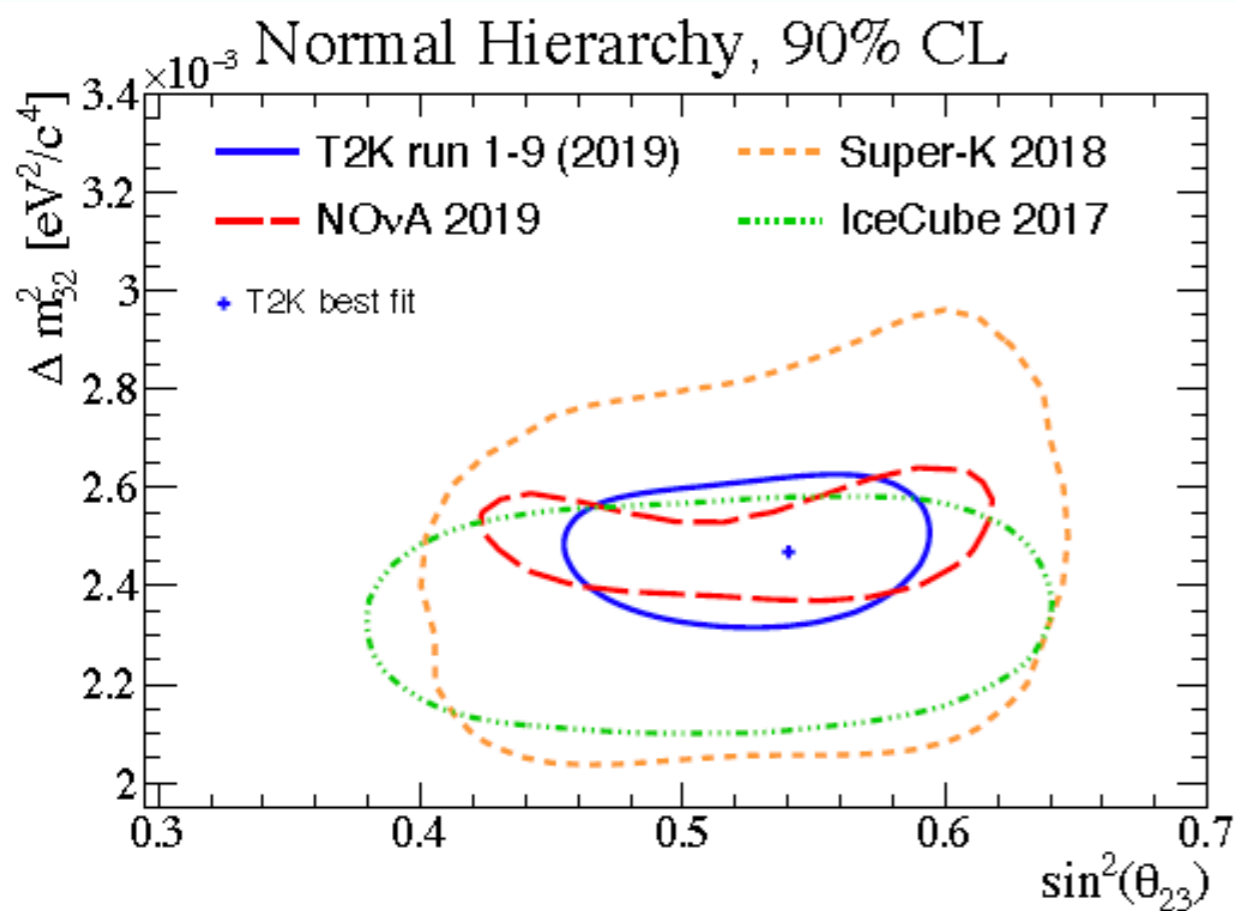
$\bar{\nu}_e$ vertex distribution





Global Fits

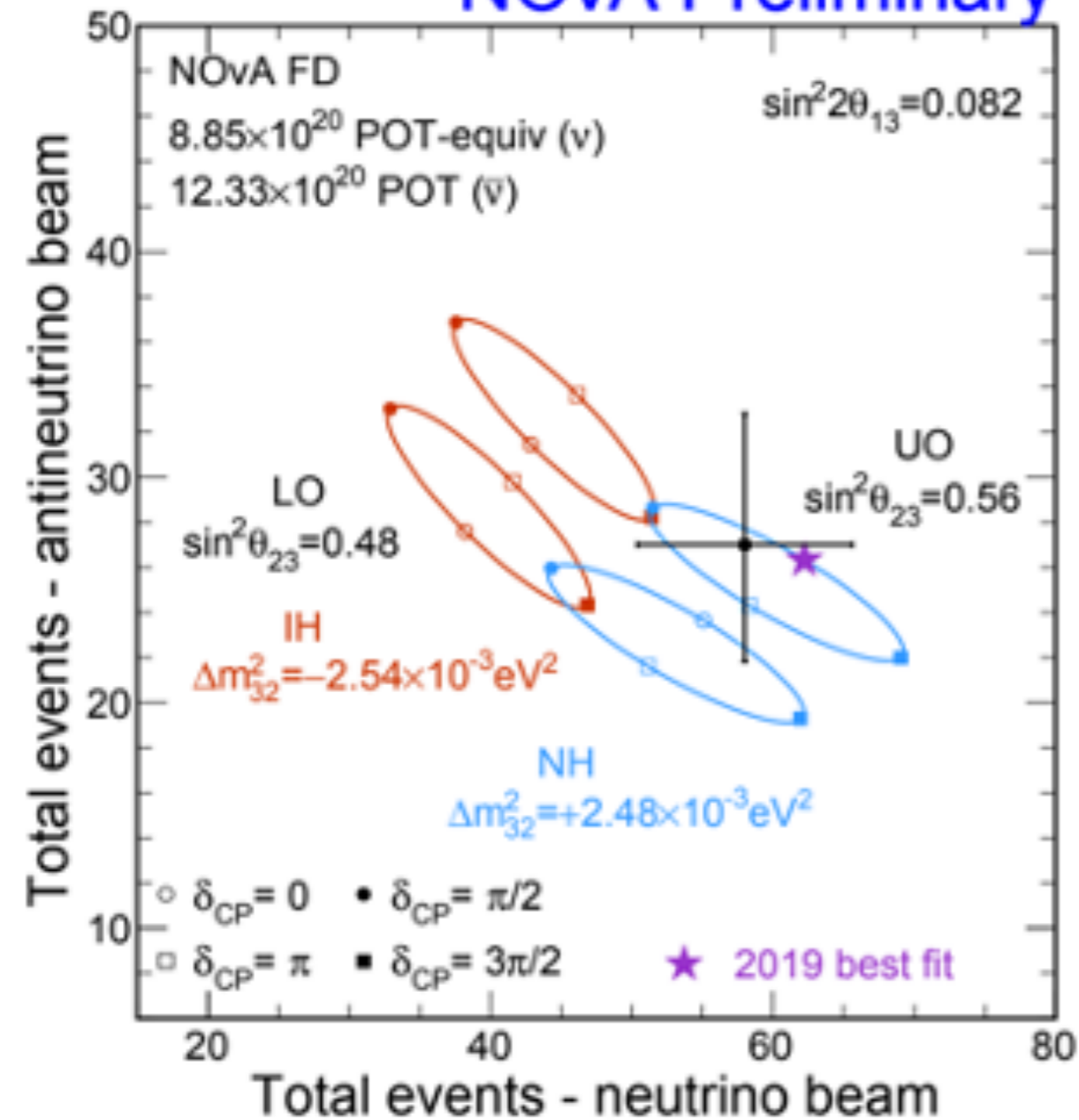
Nova vs T2K results



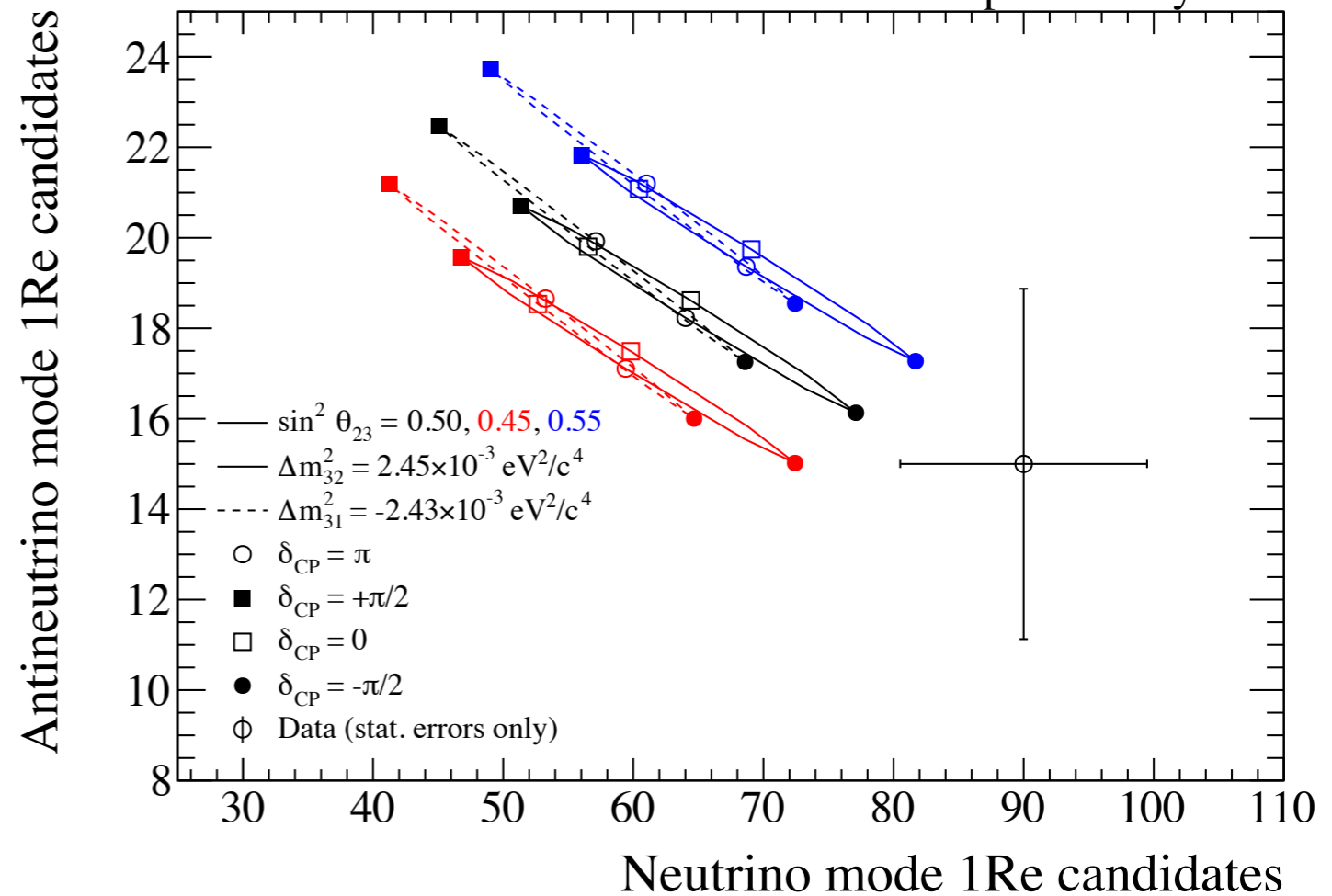
Nova results



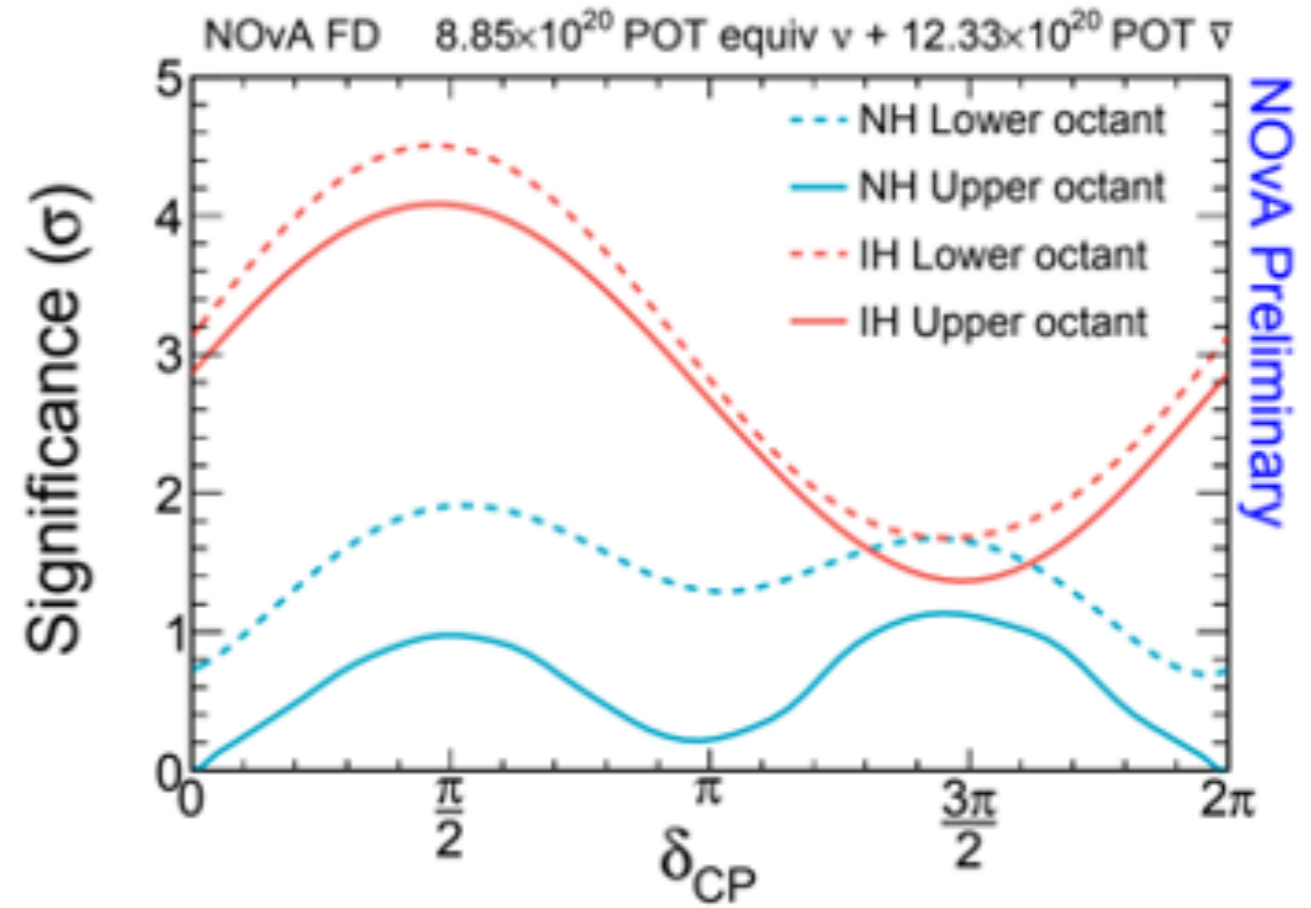
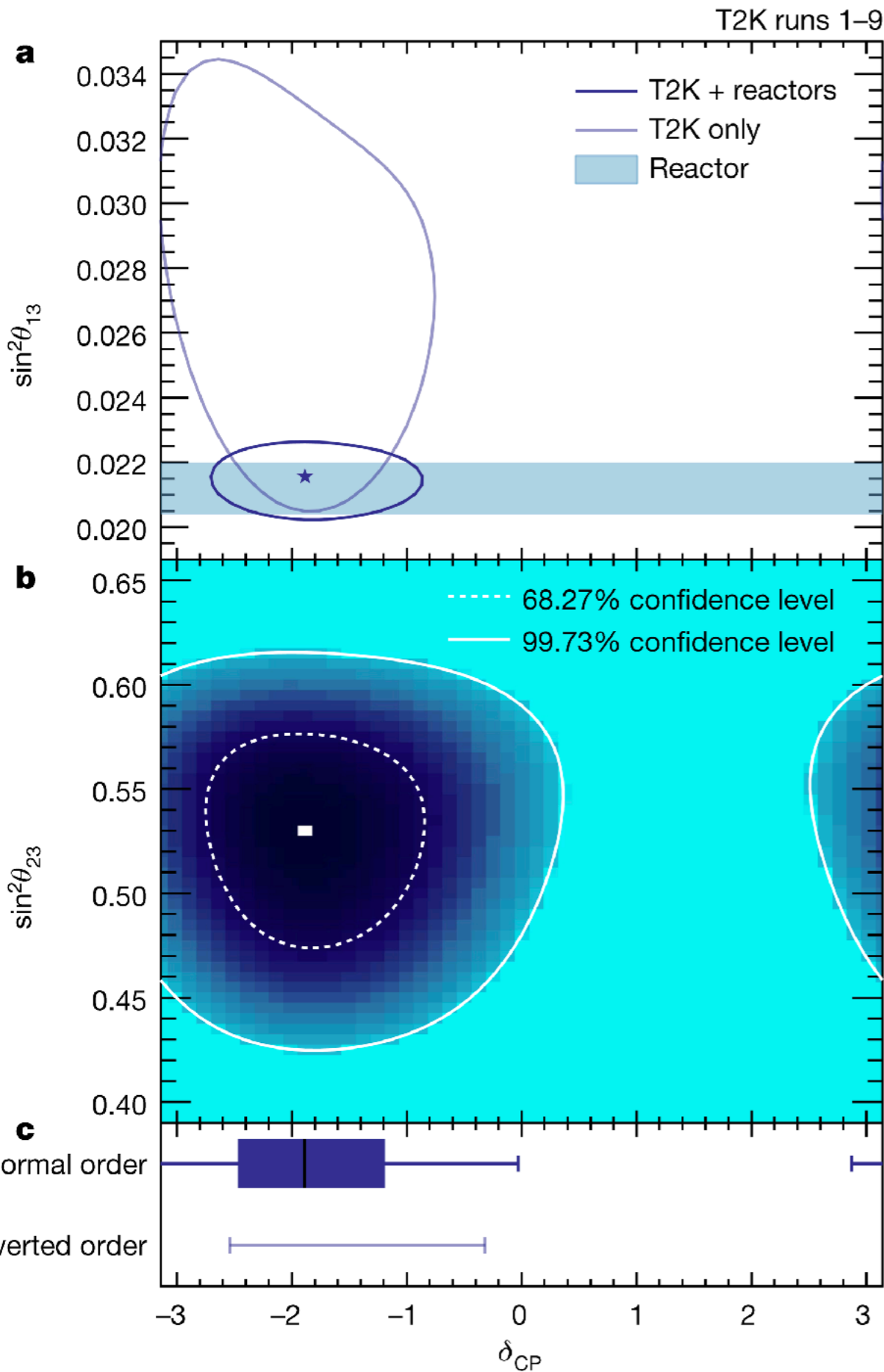
NOvA Preliminary



T2K Run 1-9 preliminary



T2K vs Nova



Approx oscillation formulae



Appearance

$$P(\nu_\mu \rightarrow \nu_e) = 4c_{13}^2 s_{13}^2 s_{23}^2 \sin^2 \Delta_{31} \times \left(1 \pm \frac{2a}{\Delta m_{31}^2} (1 - s_{13}^2) \right)$$

Leading term

$$+ 8c_{13}^2 s_{12} s_{13} s_{23} (c_{12} c_{23} \cos \delta - s_{12} s_{13} s_{23}) \cos \Delta_{32} \sin \Delta_{31} \sin \Delta_{21}$$

CP Conserving

ν vs. $\bar{\nu}$
sign
change

$$\mp 8c_{13}^2 s_{13}^2 s_{23}^2 \cos \Delta_{32} \sin \Delta_{31} \frac{aL}{4E} (1 - 2s_{13}^2)$$

Matter effect

$$\mp 8c_{13}^2 c_{12} c_{23} s_{12} s_{13} s_{23} \sin \delta \sin \Delta_{32} \sin \Delta_{31} \sin \Delta_{21}$$

CP Violating

$$+ 4s_{12}^2 c_{13}^2 (c_{12} c_{23} + s_{12}^2 s_{13}^2 s_{23}^2 - 2c_{12} c_{23} s_{12} s_{13} s_{23} \cos \delta) \sin^2 \Delta_{21}$$

Solar term

$$c_{ij} = \cos \theta_{ij}, \quad s_{ij} = \sin \theta_{ij}, \quad \Delta_{ij} = \Delta m_{ij}^2 \frac{L}{4E_\nu}, \quad a = 2\sqrt{2} G_F n_e E$$

Disappearance

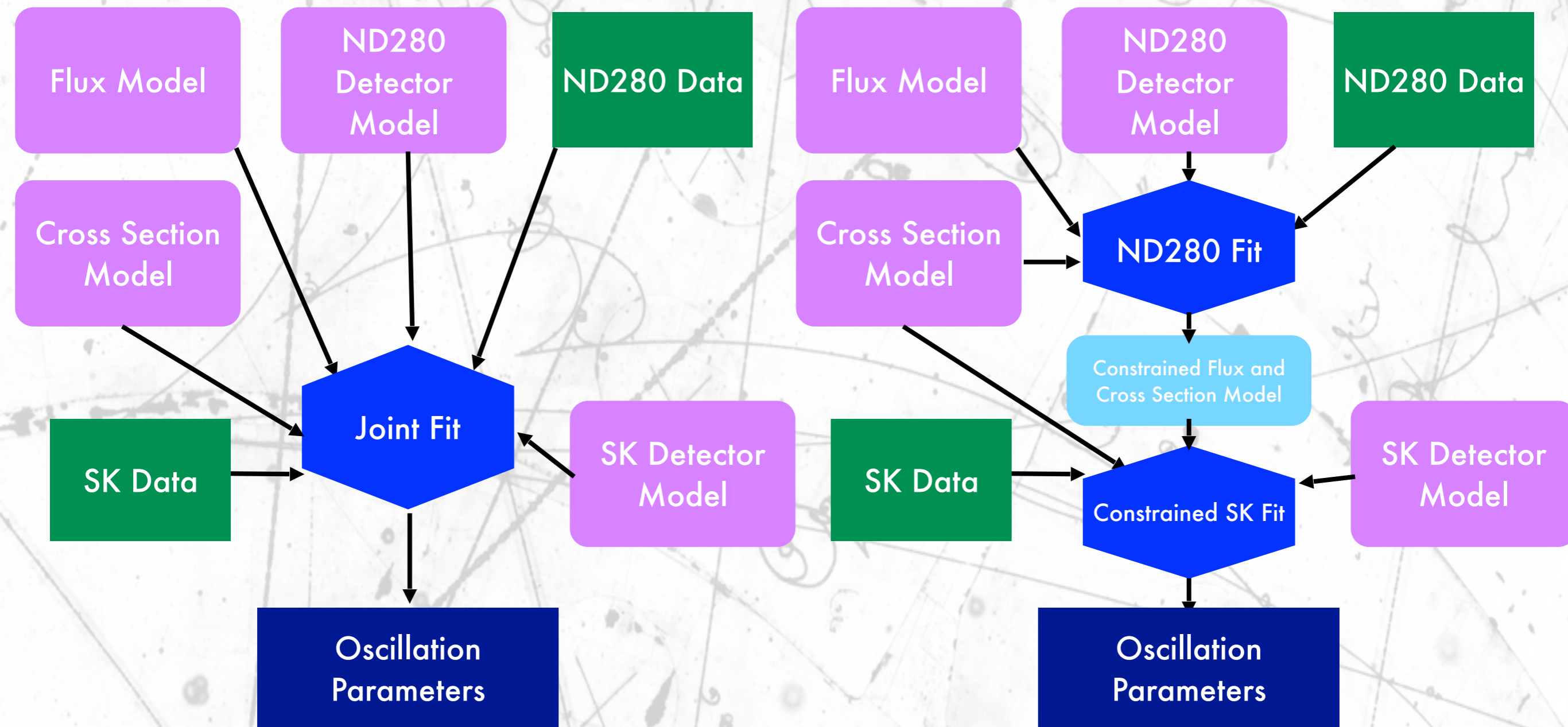
$$P(\nu_\mu \rightarrow \nu_\mu) \approx 1 - \left(\cos^4 \theta_{13} \cdot \sin^2 2\theta_{23} + \sin^2 2\theta_{13} \cdot \sin^2 \theta_{23} \right) \cdot \sin^2 \frac{\Delta m_{32}^2 \cdot L}{4E_\nu}$$

Statistical methods



	Analysis 1	Analysis 2	Analysis 3
Kinematic variables for 1Re sample at SK	Erec- θ	p_e - θ	Erec- θ
Likelihood	Binned Poisson Likelihood Ratio	Binned Poisson Likelihood Ratio	Binned Poisson Likelihood Ratio
Likelihood Optimization	Markov Chain Monte Carlo	Gradient descent and grid scan	Gradient descent and grid scan
Contours/limits produced	Bayesian Credible Intervals	Frequentist Confidence Intervals with Feldman-Cousins (credible intervals supplemental)	Frequentist Confidence Intervals with Feldman-Cousins
Mass Hierarchy Analysis	Bayes factor from fraction of MCMC points in each hierarchy	Bayes factor from likelihood integration	Frequentist p-value from generated PDF
Near Detector Information	Simultaneous joint fit	Constraint Matrix	Constraint Matrix
Systematics Handling	Simultaneous fit then marginalization	Marginalization during fit	Marginalization during fit

Statistical methods



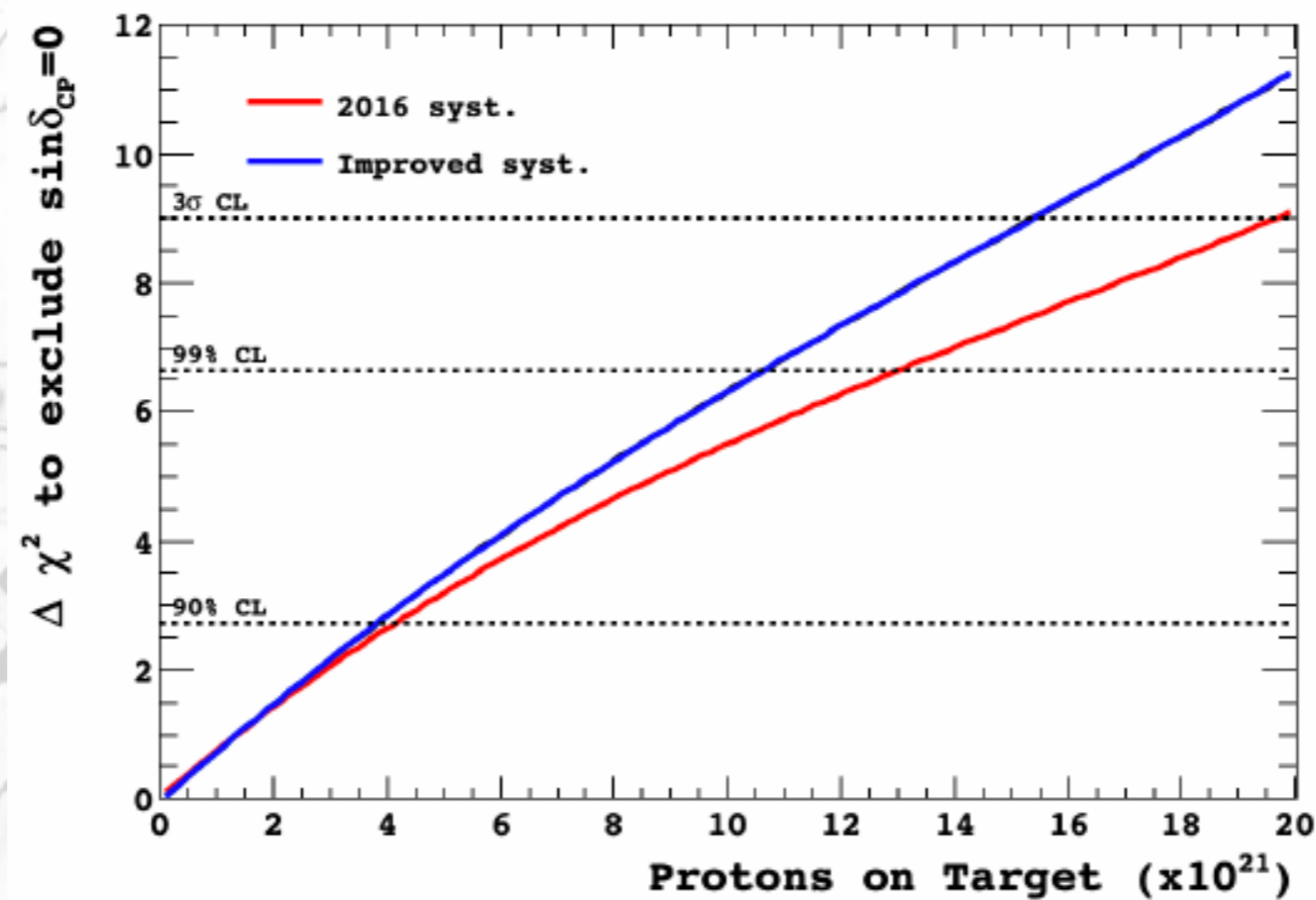
Joint Fit

Separate ND and SK Fits

T2K: impact of ND



- Why is this so important?:
 - Main systematics are from the **beam modelling** (improved with hadro-production experiments) and **x-section modelling**.

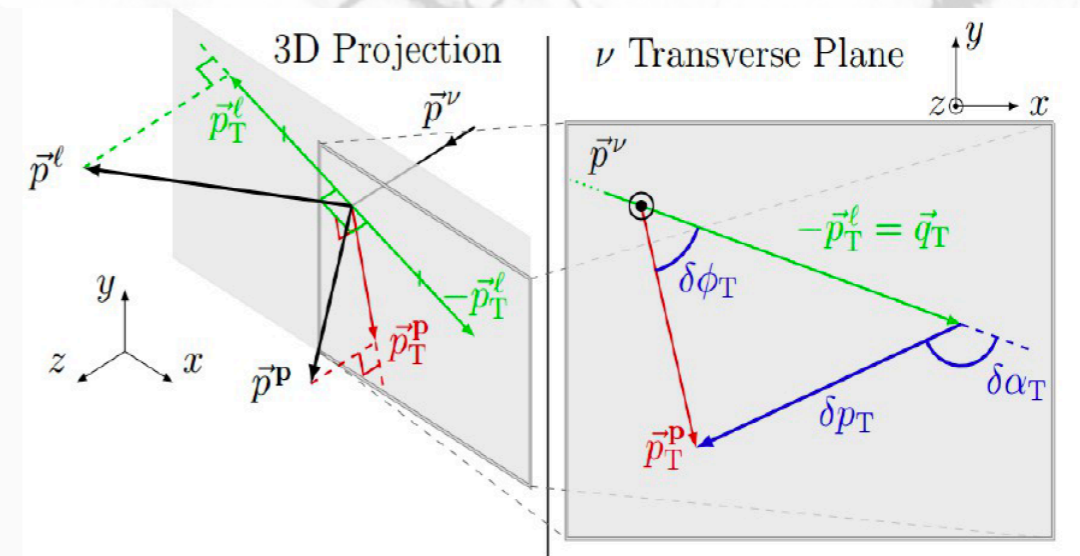


- T2K has still **significant x-section** systematic errors.
- T2K measurements are important for **HK, Dune, Nova** and **atmospheric neutrino oscillations**.

ND280 upgrade



How to use this proton information:
“Single Transverse Variables” and beyond!
 → measurements of Fermi momentum, binding energy, 2p2h...

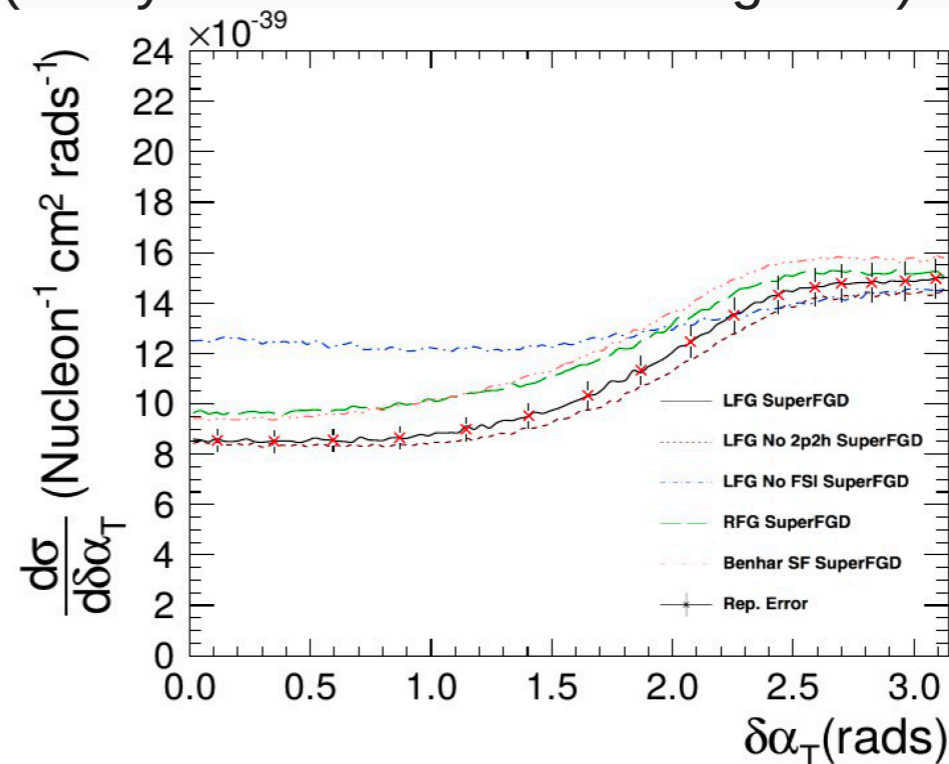
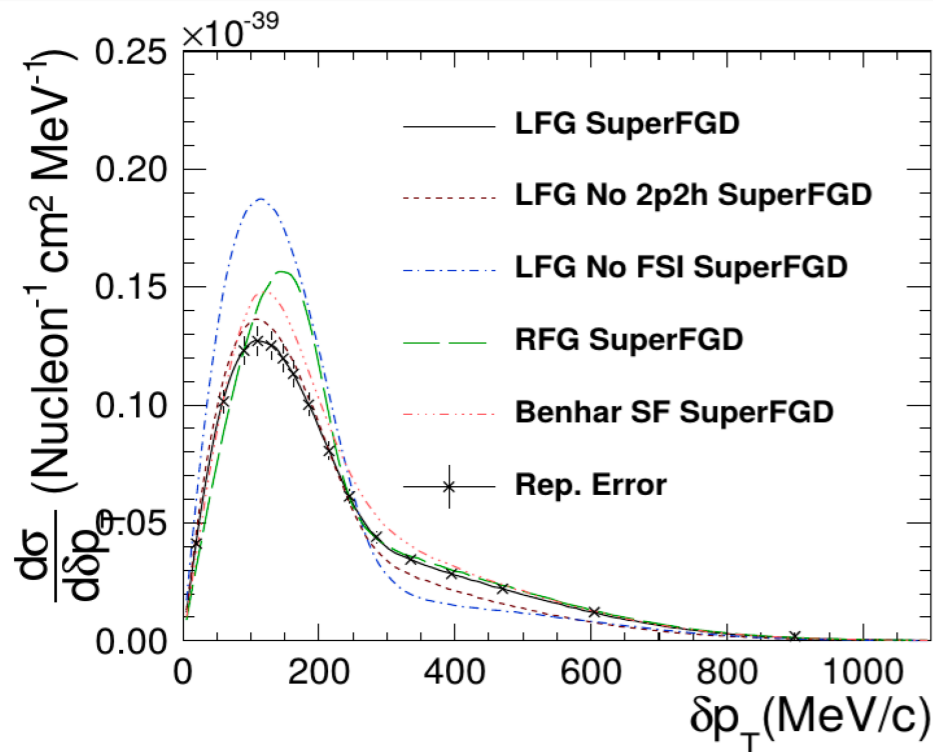


δp_T is a direct measurement of **Fermi momentum: shape measurement <10% precision in each bin with 8×10^{21} POT**

$\delta \alpha_T$ shape is highly sensitive to proton FSI

→ allows to constrain it to $\sim 1\%$: **not anymore an issue to use protons in the ND fit for the oscillation analysis!**

(today 30% from e-scattering data)

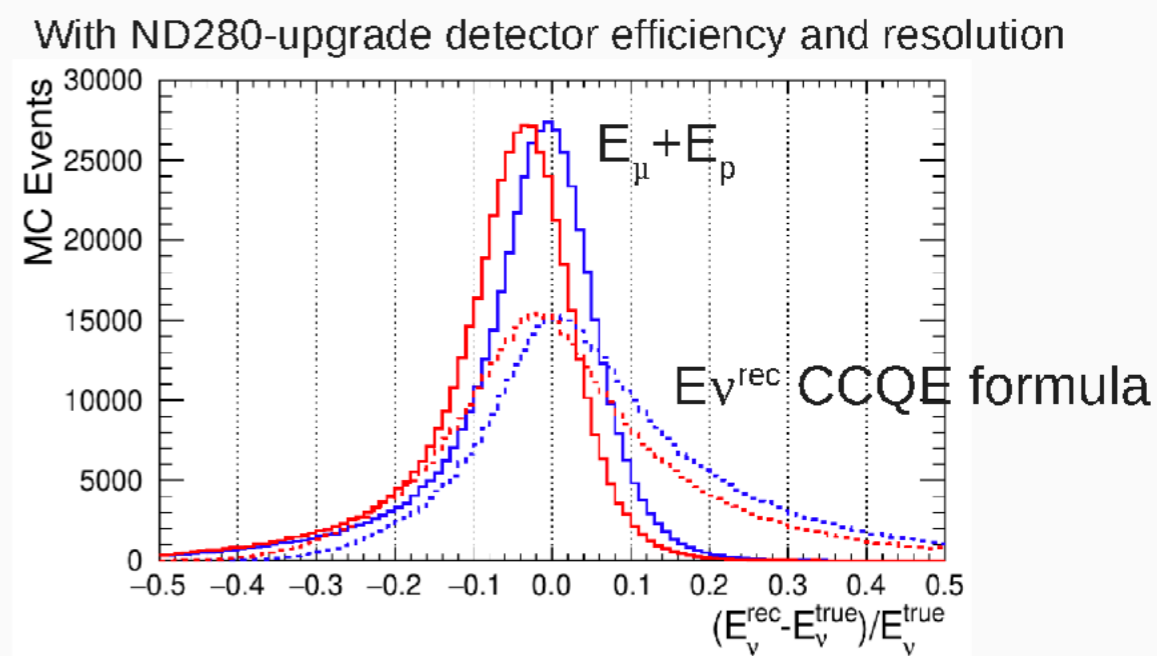
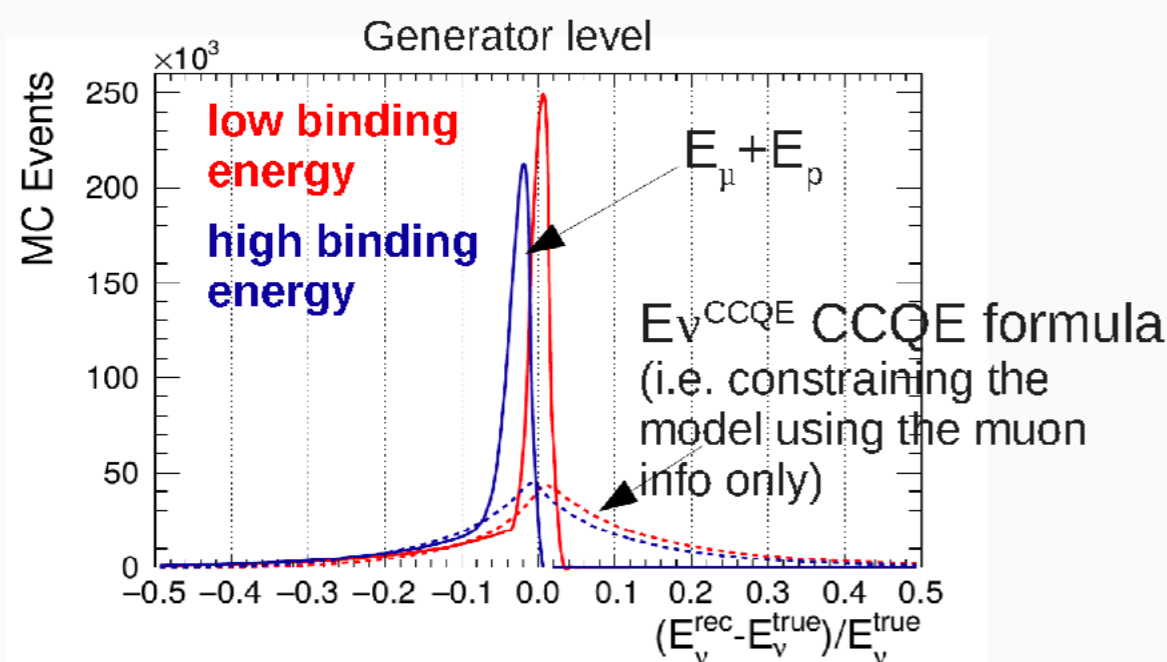


ND280 upgrade



Another variable: total energy

- The E_{ν}^{rec} CCQE formula does not include information on the outgoing proton \rightarrow $E_{\mu}+E_p$ is a much better estimator of the true neutrino energy



Smearing of E_{ν}^{rec} is dominated by Fermi momentum,
smearing of $E_{\mu}+E_p$ is dominated by flux (and detector effects)

\rightarrow **$E_{\mu}+E_p$ is a much more robust estimator of true E_{ν} and of binding energy**

- This is just the appetizer! We are starting investigating possible other variables and combinations \rightarrow a lot of new sensitivity

A good example of the 'iterative' process: new detector + *DATA* \rightarrow new ways of doing analysis / looking at our systematics \rightarrow improvements of oscillation analysis!

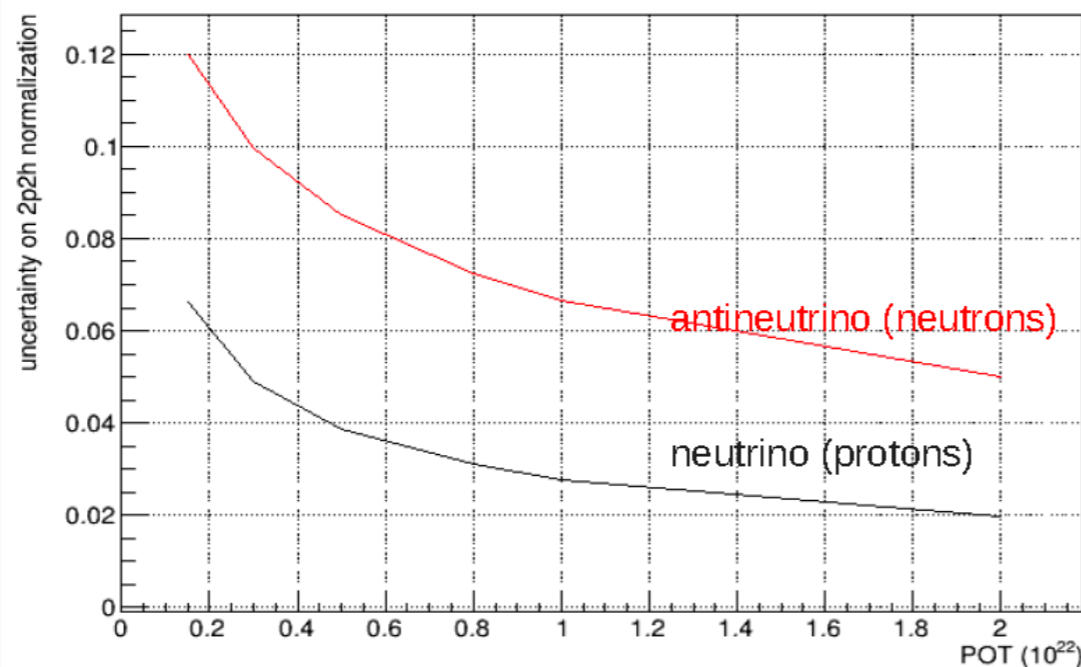
ND280 upgrade



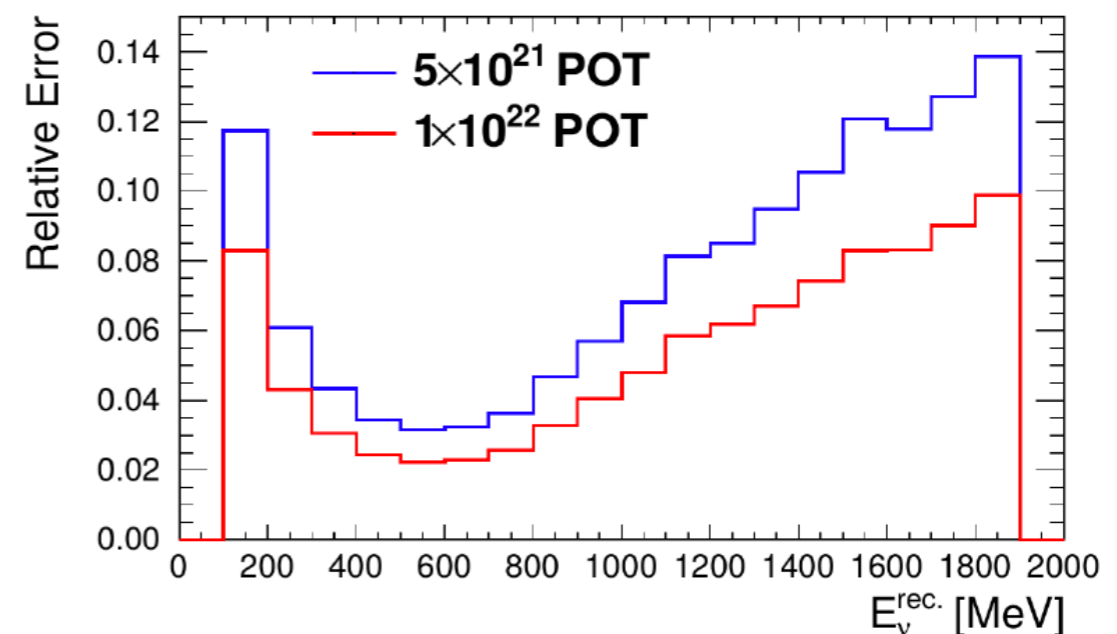
SuperFGD: neutrons

- The superFGD can detect neutrons with ~60% efficiency
- If the path is long enough (>50 cm) also neutron energy is measured with resolution 15-30% (to be calibrated with neutron test-beam at LosAlamos)
- The background can be rejected by reducing the fiducial volume (no reliable simulation available yet)

The same analyses shown for protons can be repeated for neutrons.
Example of fitting single transverse variables



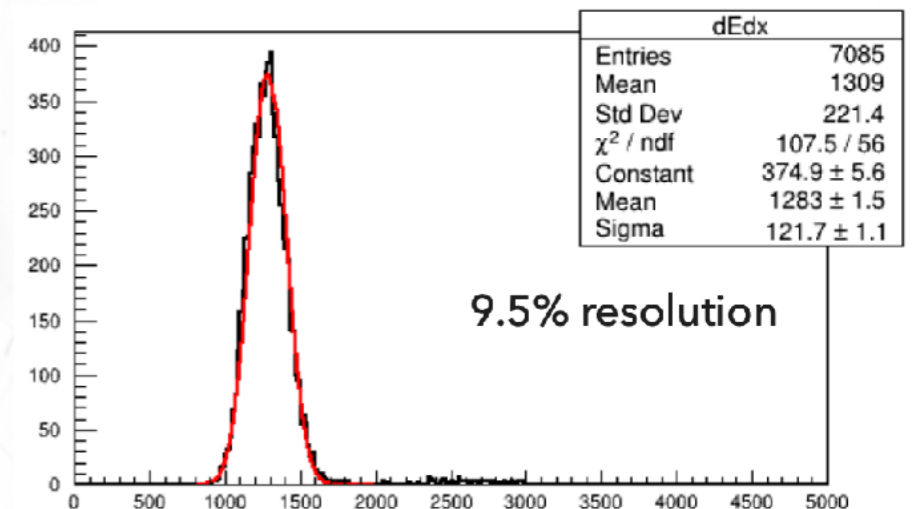
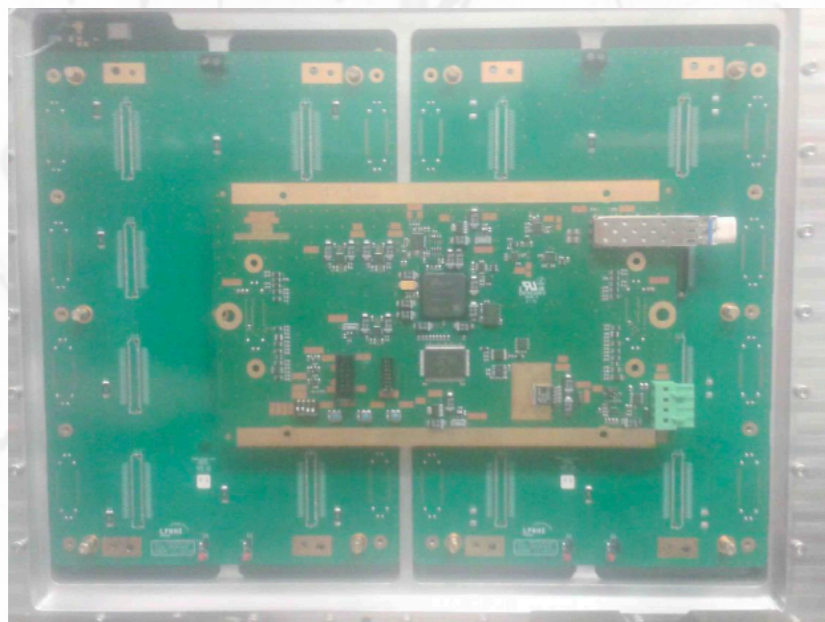
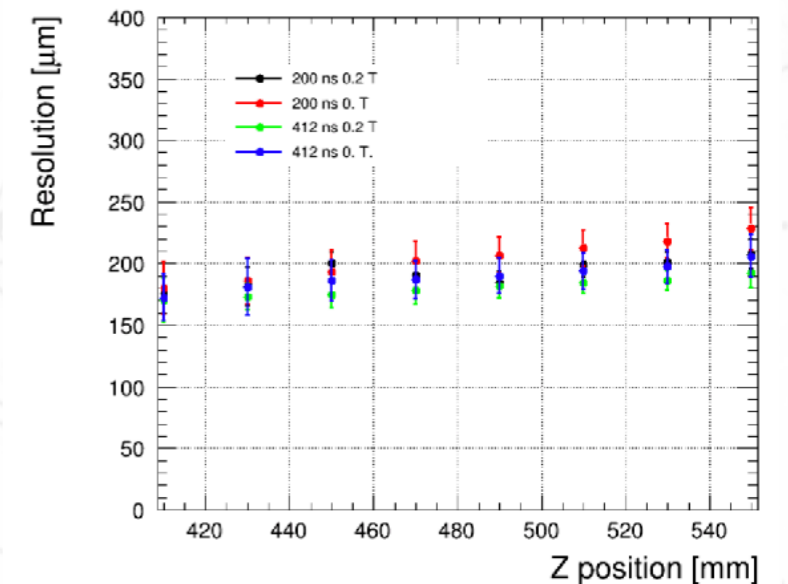
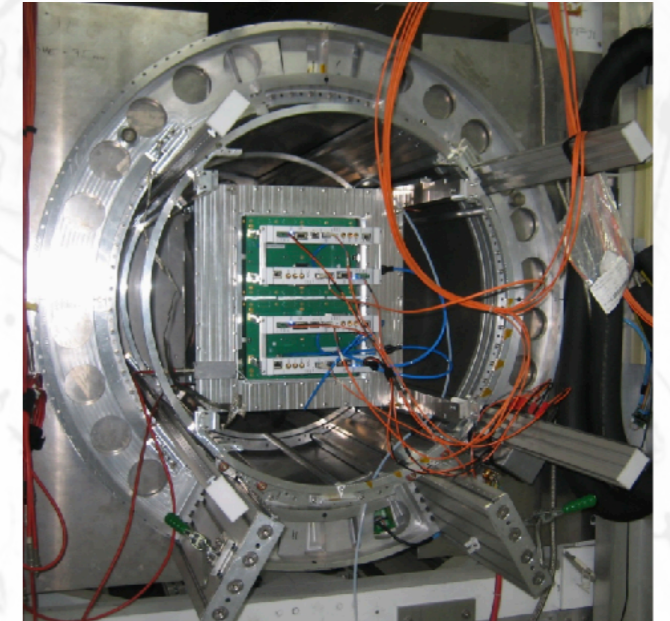
A lot of interesting physics with neutron tagging (e.g. DSNB, increased atm MH sensitivity)
→ e.g. neutron multiplicity as a function of neutrino energy



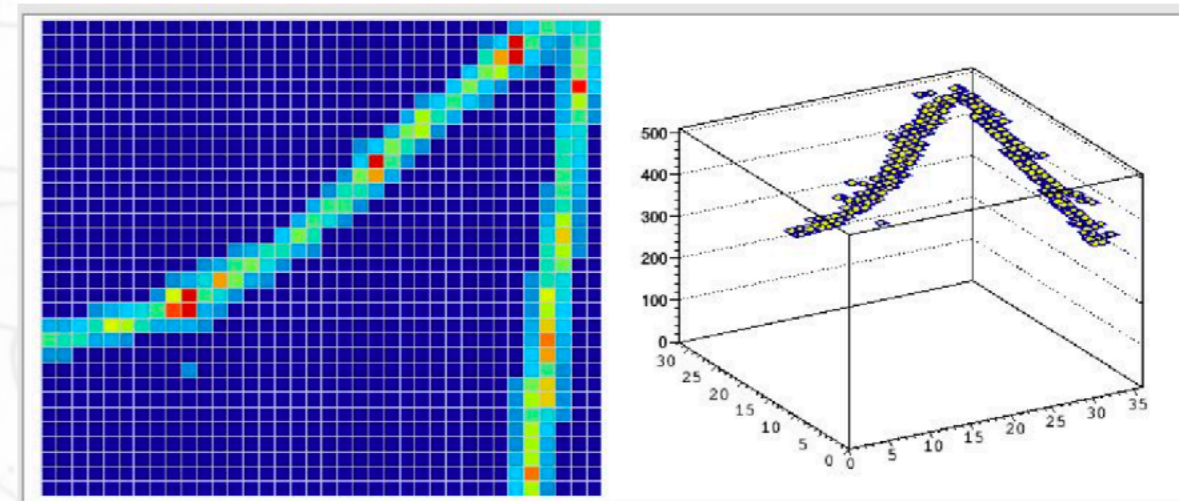
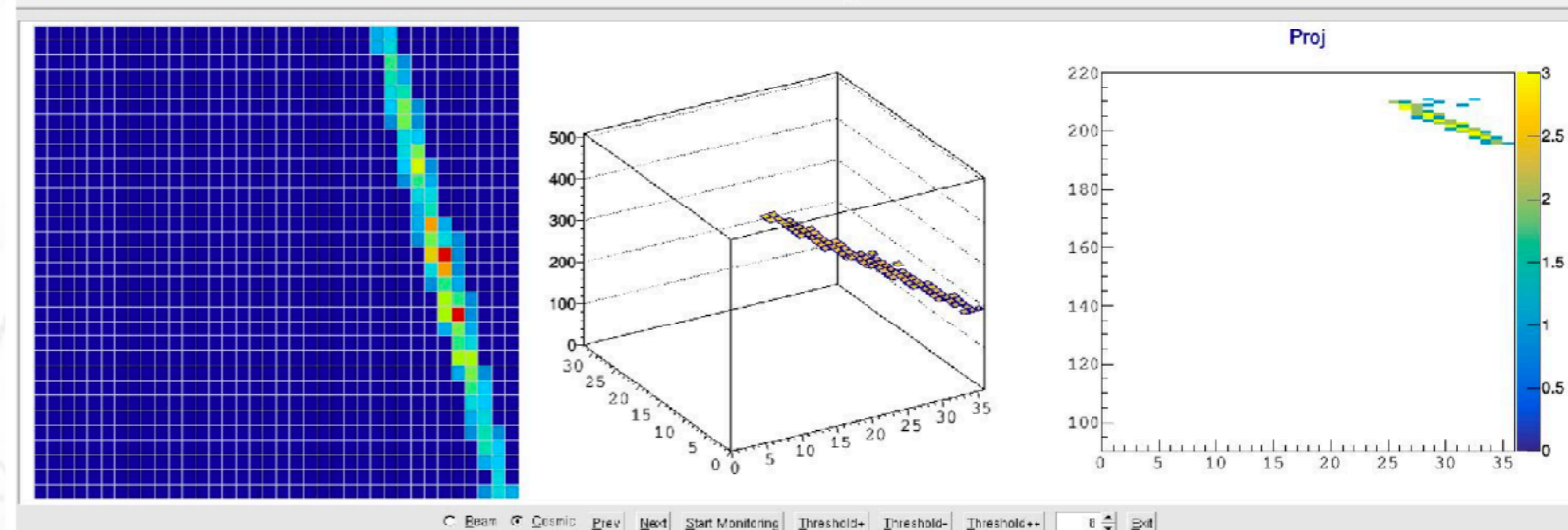
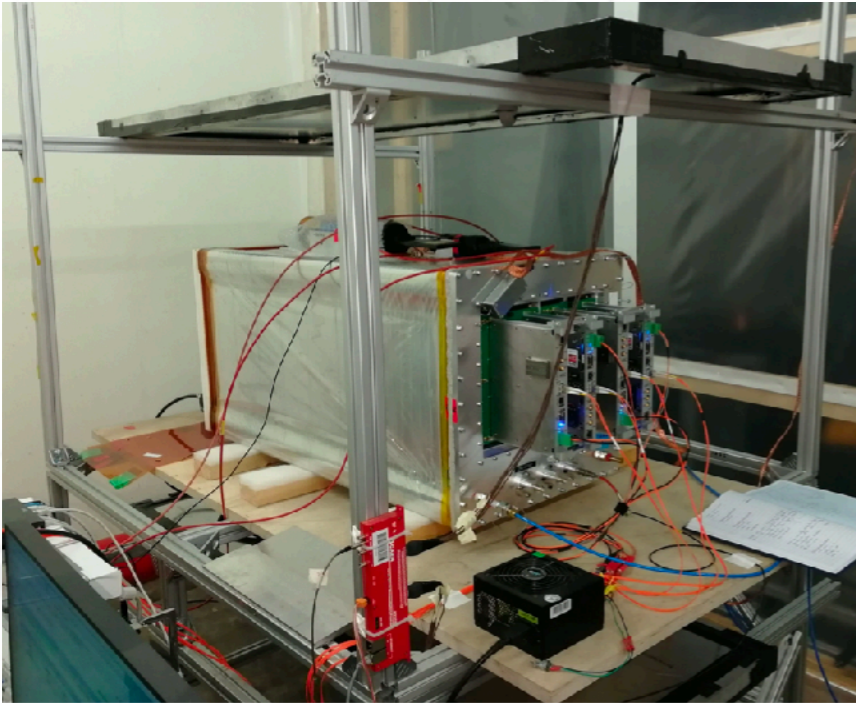
HA-TPC

@DESY

- Test Beam with Resistive MicroMegas at DESY in June 2019
 - 4 GeV electrons, analysis on-going
 - Excellent spatial resolution ($\sim 200 \mu\text{m}$ for horizontal tracks) and dE/dx resolution
- HA-TPC electronics:
 - First Front-End mezzanine (FEM) prototype has been tested
 - First Front-End-Card (FEC) will be delivered in Jan 2020



HA-TPC prototype @ CERN



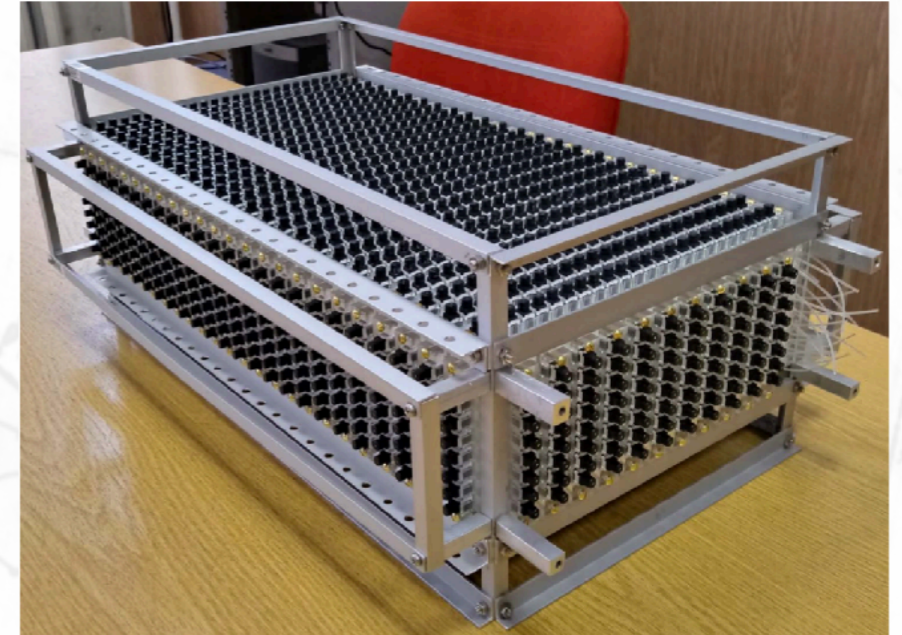
- Stable operation at 18 kV
- 2nd prototype to improve gas leak rate expected in Feb 2020
- First TPC field cage expected in June
 - External review committee for the TPC field cage design has been formed → expect 1st meeting in January

Super-FGD

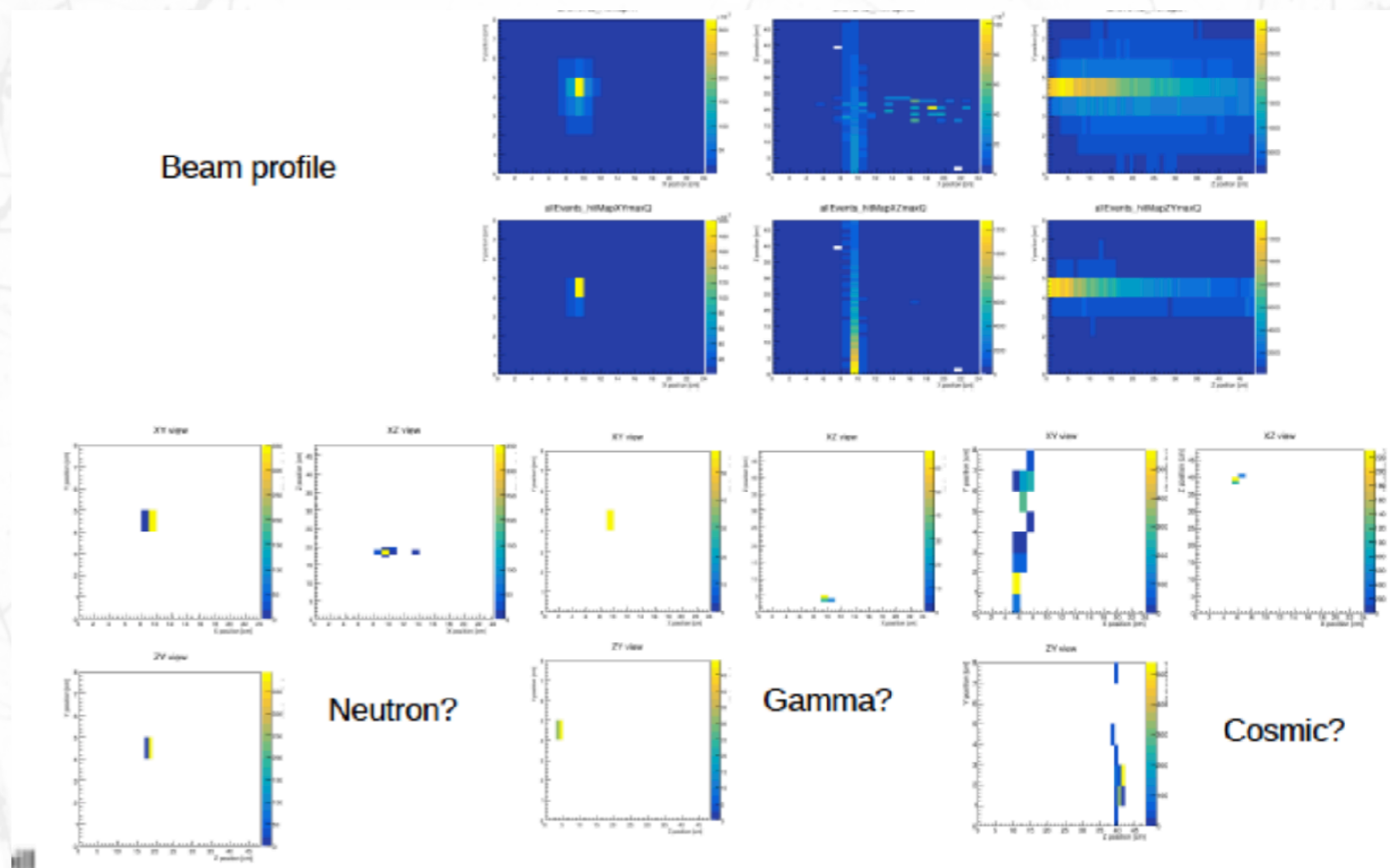
- Assembly with Fishing Line at INR →
 - 27 full size (192 x 184 cubes) x-y layers assembled
 - 56 z layers (15 x 192 cubes) → corresponding to the full height of the Super-FGD
- All cubes will be produced by Jan 2021
- Review to discuss feasibility of assembly method organized by T2K → Fishing-Line method has been chosen as primary option for the assembly
- Design of the Super-FGD electronics is ongoing → all CITIROC chips have been bought



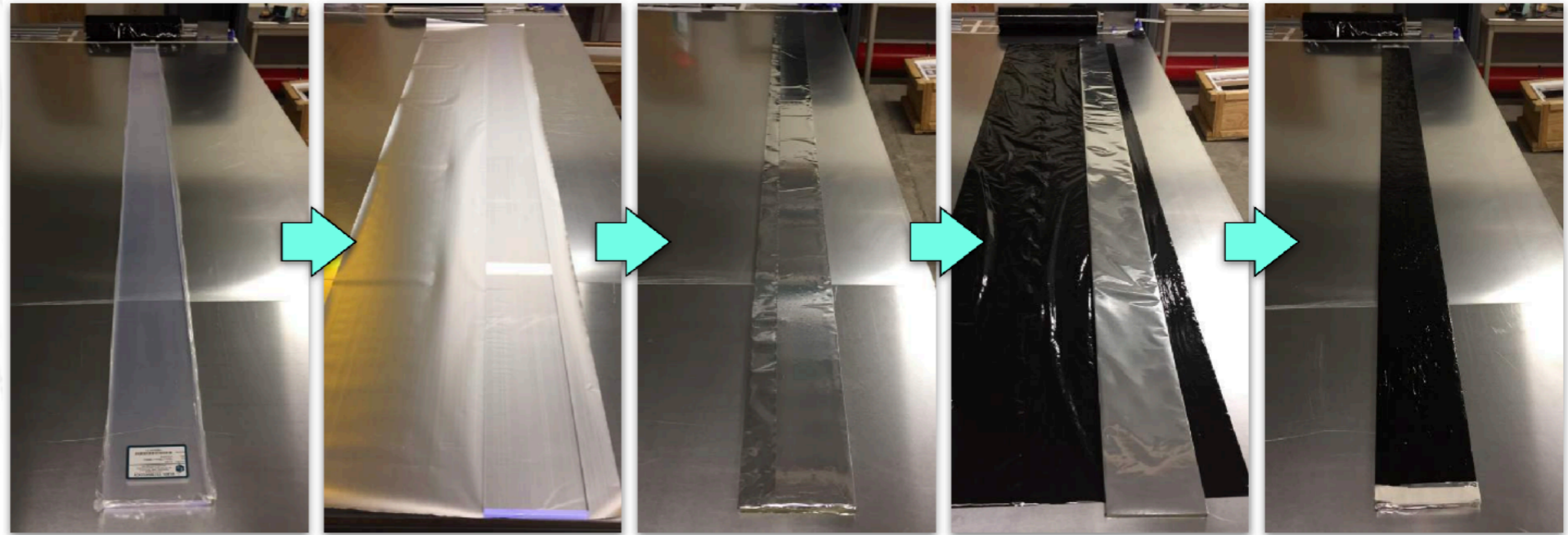
Super-FGD prototype neutron tests at LANL



- Data taking in December
- Neutron beam profile clearly visible
- Analysis of the data is on-going



Time Of Flight



3 layers of Al foil + 6 layers of black stretch film

- Start assembling scintillator bars
- Most of the components already received
- First ToF module assembled

- Mock-up basket



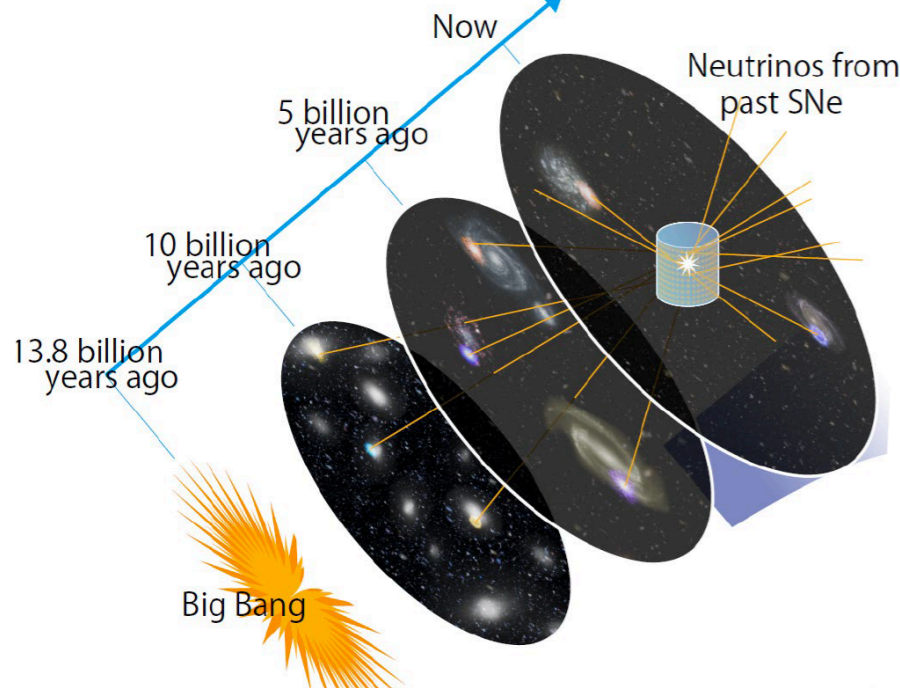
- The 6 ToF modules will be installed into the “mini basket” that has been delivered to CERN
- Mock-up of the upstream part of the real ND280 basket
- It will be used to test integration of the different sub-detectors

SK-Gd

Diffused Supernova Neutrino Backgrounds

- Neutrinos produced from the past SN bursts and diffused in the current universe.
 - ~ a few SN explosions every second $\rightarrow O(10^{18})$ SNe so far in this universe
 - Can study history of SN bursts with neutrinos

$$\frac{dF_\nu}{dE_\nu} = c \int_0^{z_{\max}} R_{\text{SN}}(z) \frac{dN_\nu(E'_\nu)}{dE'_\nu} (1+z) \frac{dt}{dz} dz$$

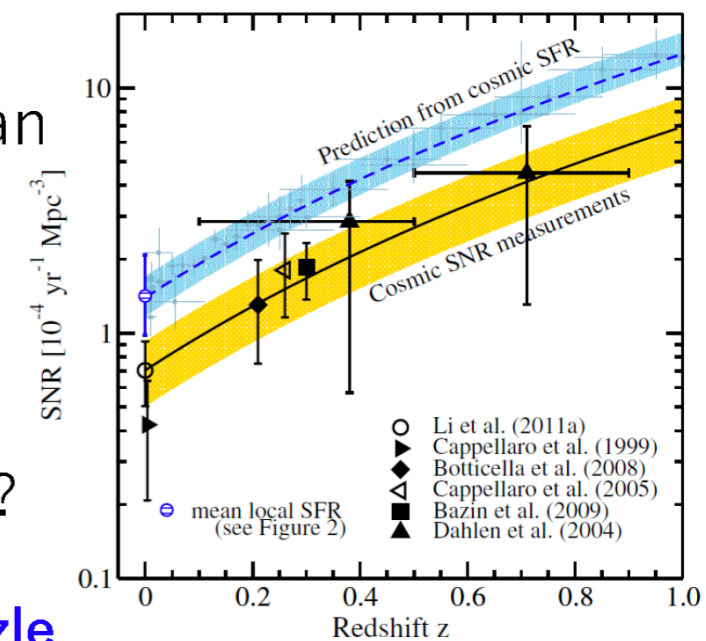


SN rate problem

- Observed SN burst rate lower than prediction from cosmic star formation rate
 - Invisible dim supernova?
 - Black-hole formation?
 - Something blocking optical light?

DSNB signal could resolve the puzzle

H. Horiuchi et al, *Astrophys. J.* **738**, 154 (2011)



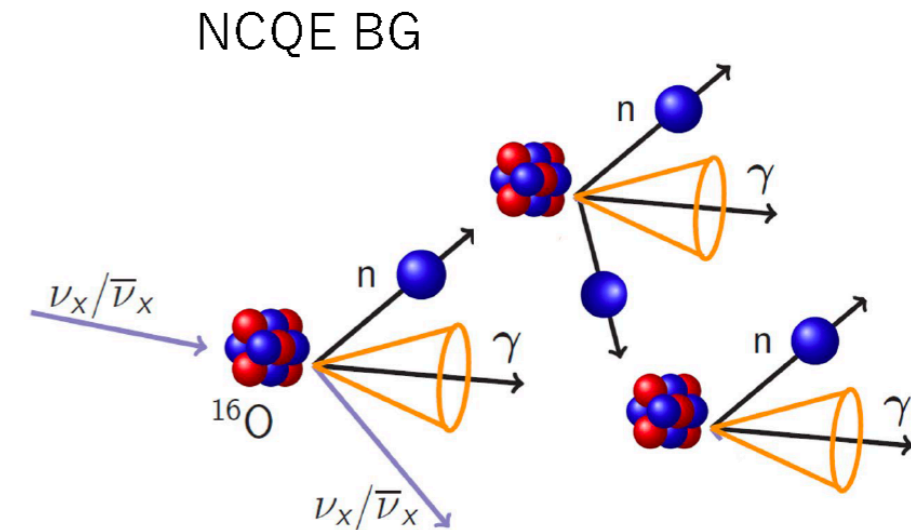
SK-Gd

NCQE measurements with T2K

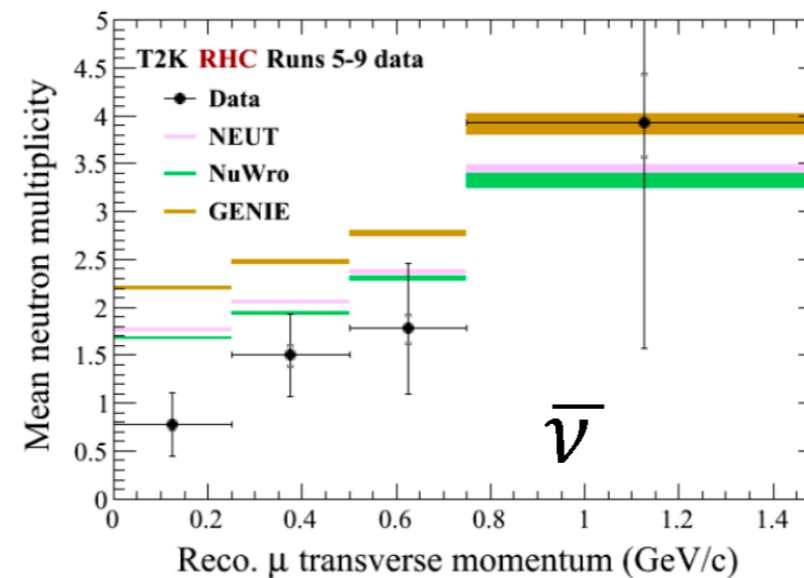
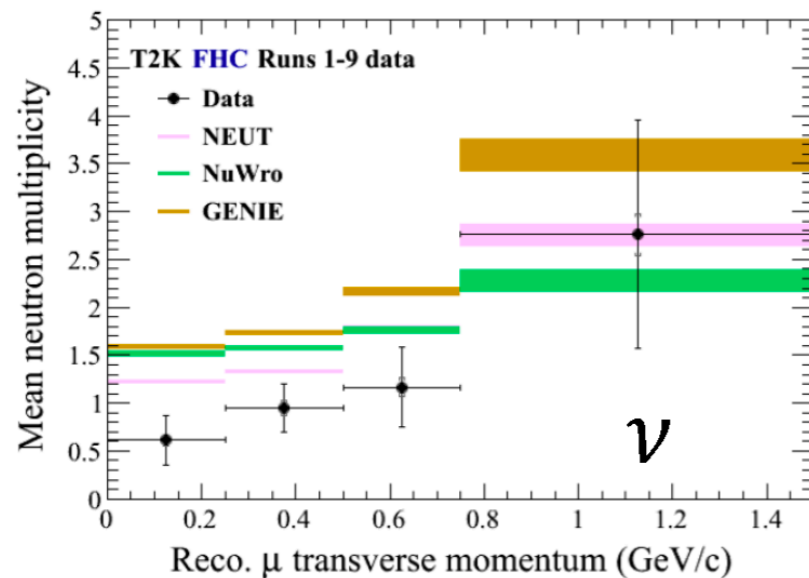
- Another important data: Neutron multiplicity

Multiple neutrons produced through hadronic final-state interactions (FSI) in nuclei, and secondary interactions (SI) in the detector medium

➡ Key to reduce NCQE BG

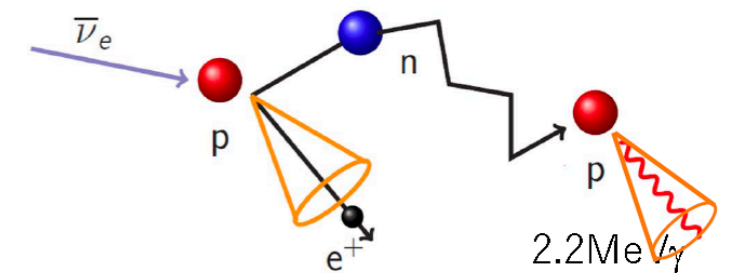


Measured mean neutron multiplicity (CC) and MC predictions



Large discrepancy causes ~44% systematic error for NCQE BG estimation

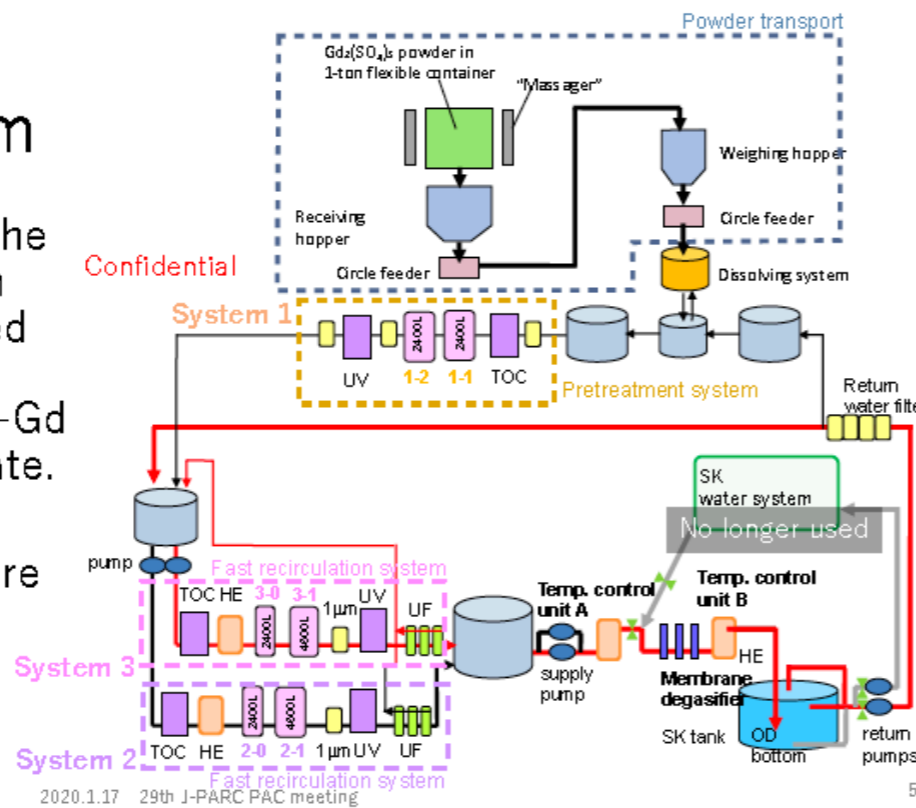
DSNB



SK-Gd

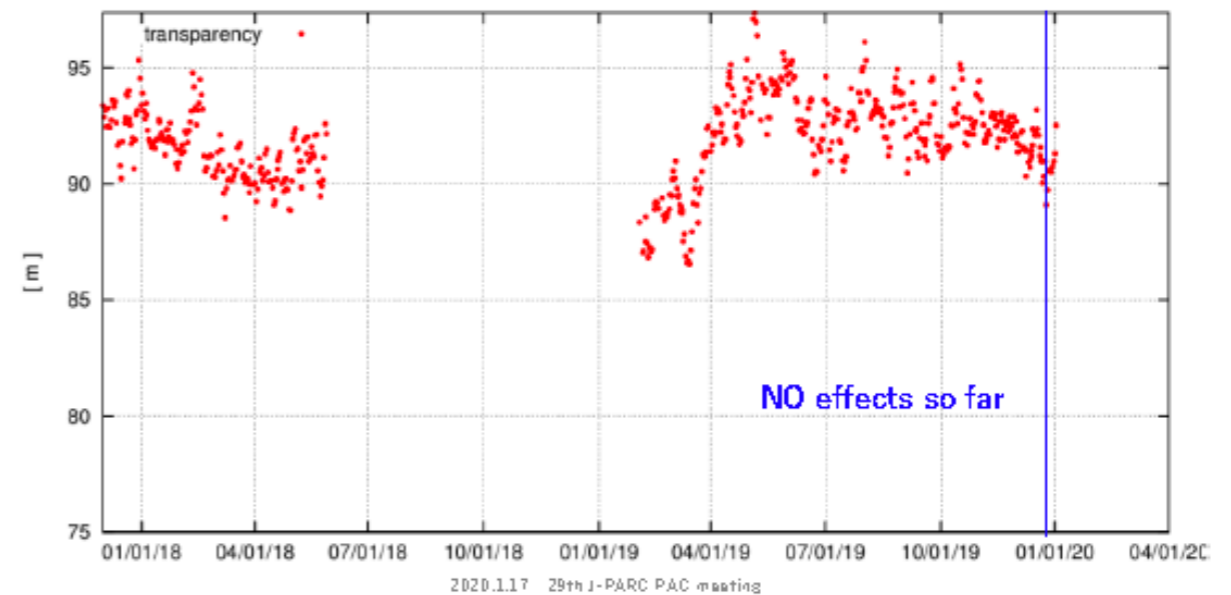
Migration of the water system

- Since Dec. 24, 2019, the SK pure water system has been disconnected from the recirculation loop and the new SK-Gd is in use with 60t/h rate.
- Specially developed resins for $Gd_2(SO_4)_3$ are under final test.
- Full 120t/h power is under preparation



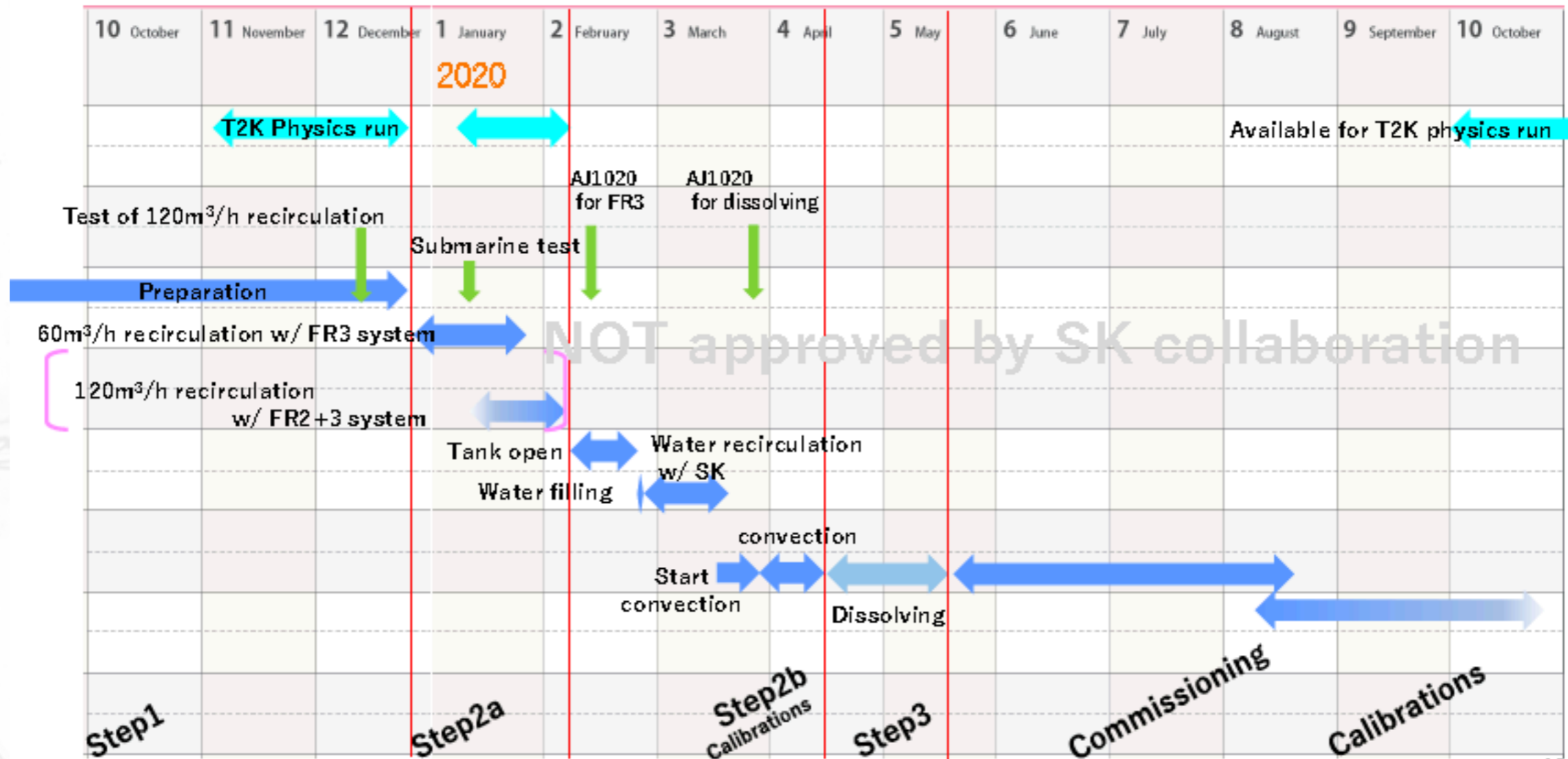
The water transparency

- After 2 weeks of SK-Gd water system operation



SK-Gd

T1 Schedule w/ 2.2 m water draining ver. 2019.12.27



First attempt to dissolve Gd salt stopped because of COVID-19