

Physics with Forward LHC Neutrino Detectors

Invisibles 2021



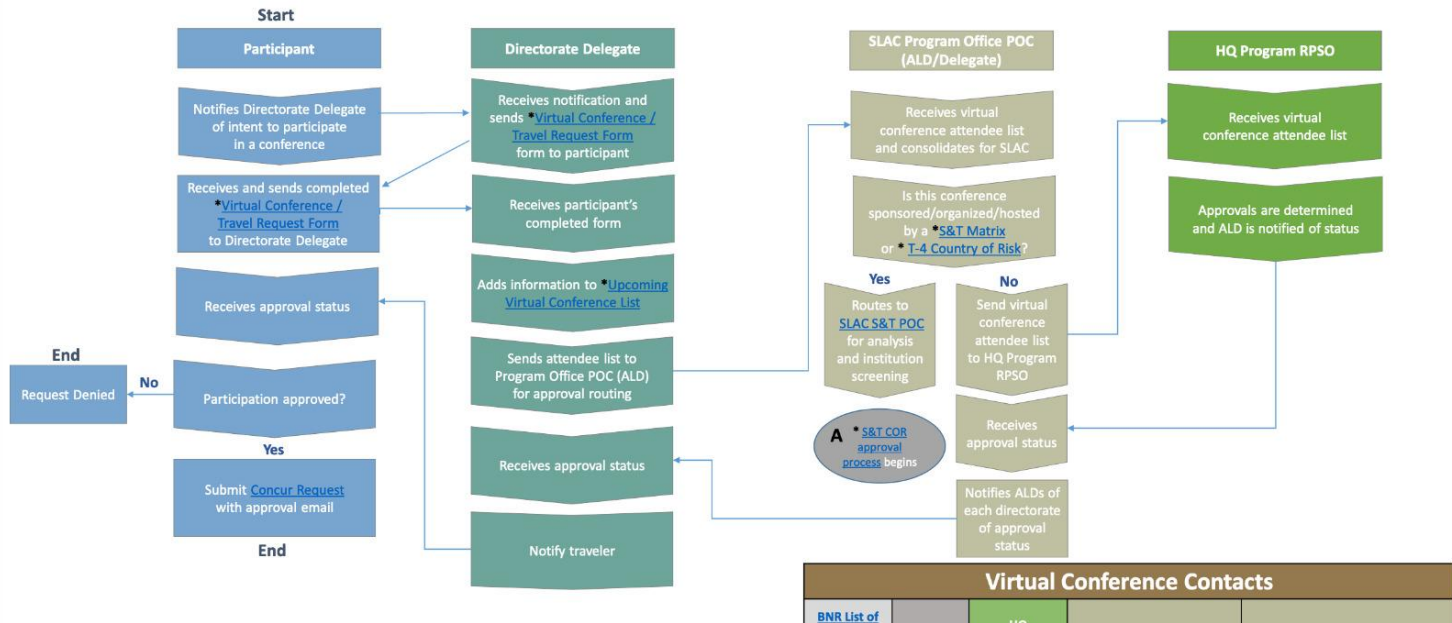
Felix Kling

SLAC NATIONAL
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Administration required to attend this conference

SLAC Foreign Virtual Conference Approval Process

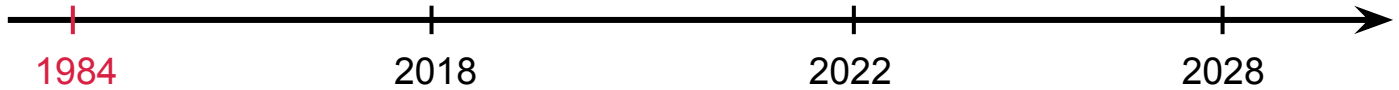
The Office of Science (SC) guidance below addresses steps that must be taken before laboratory employees participate in virtual scientific conferences, workshops, seminars, and similar activities that are sponsored/organized/hosted by foreign entities (e.g., foreign institutions or governments).



the entire process took about 2 months ...

Neutrinos at the LHC

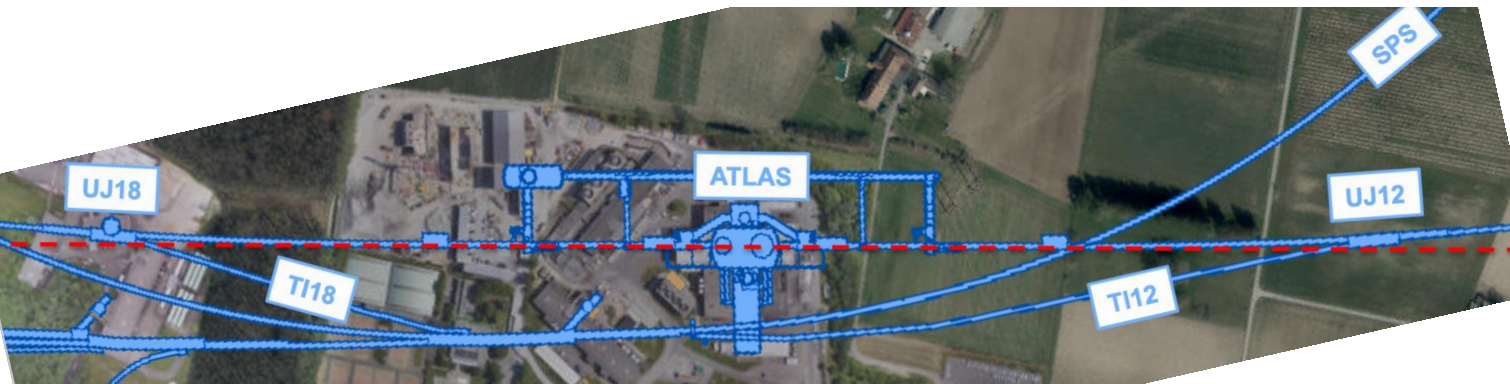
Neutrinos at the LHC



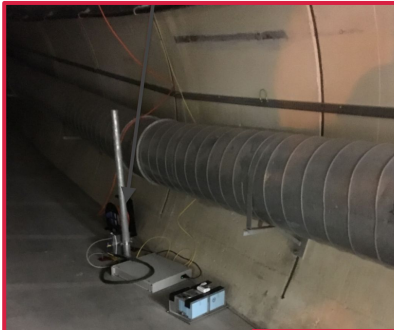
Neutrinos detected from many sources, but not from colliders.

But there is a huge flux of neutrinos in the forward direction, mainly from π , K and D meson decay. De Rujula et al. (1984)

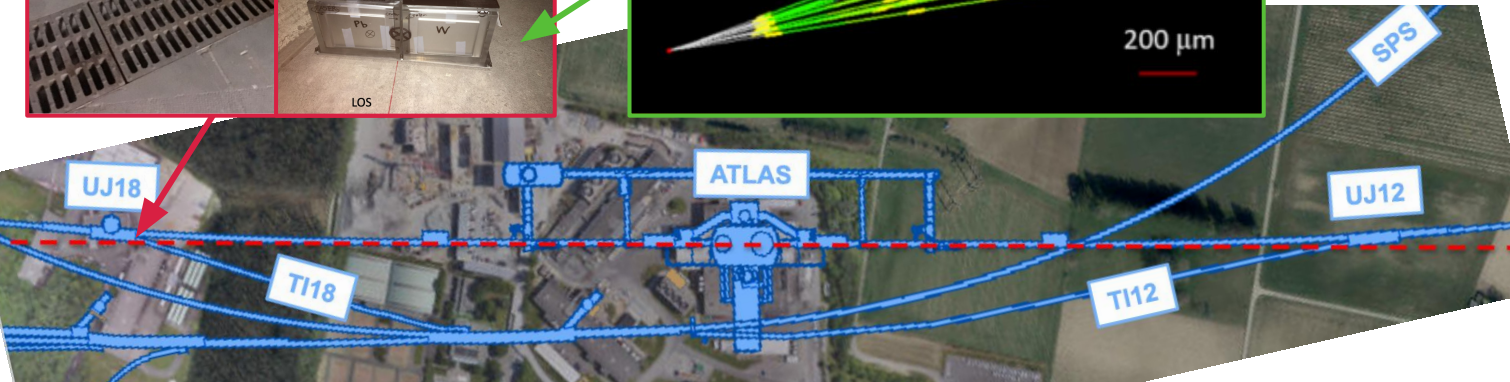
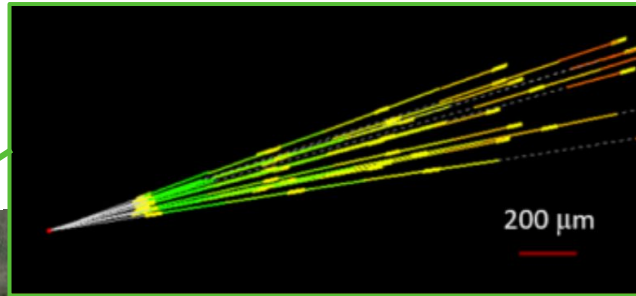
ATLAS provides an **intense** and **strongly collimated** beam of **TeV-energy** neutrinos along **beam collision axis**.



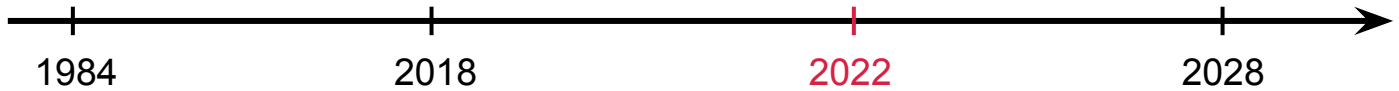
Neutrinos at the LHC



In 2018, the FASER collaboration placed ~30 kg **pilot emulsion detectors** in T118 for a few weeks. $O(10)$ neutrino interactions expected
First neutrino interaction candidates were recently reported.

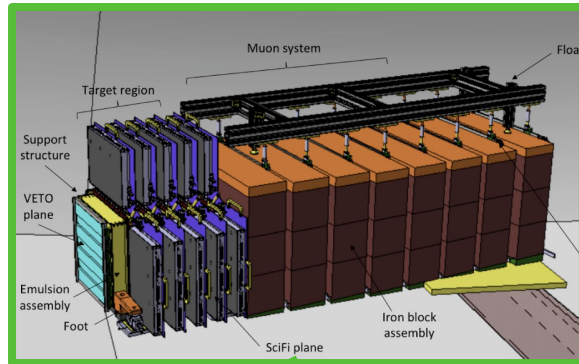


Neutrinos at the LHC

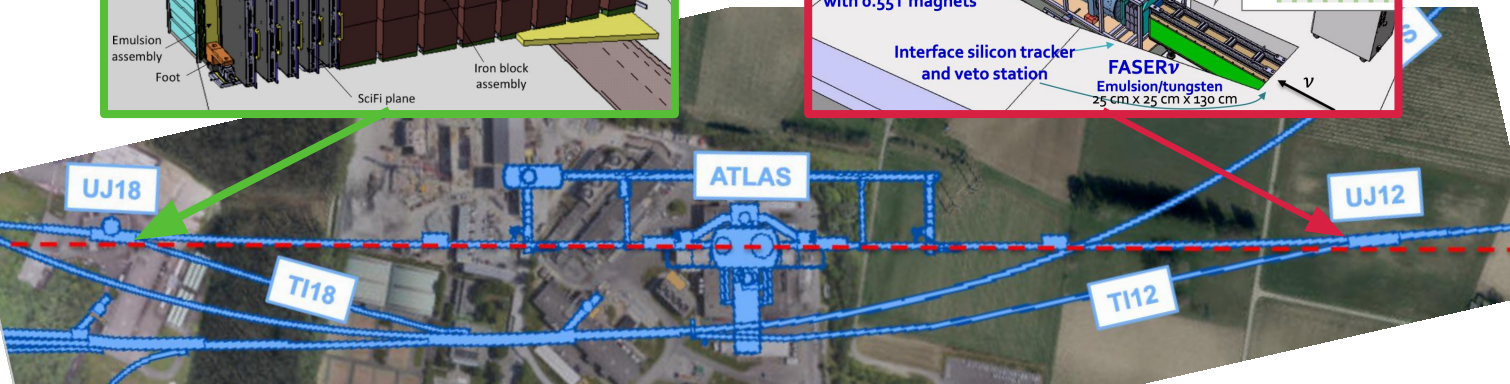
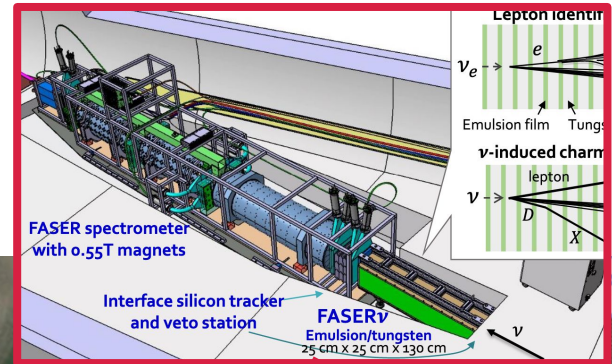


During Run 3 of the LHC, two new experiments will detect LHC neutrinos. FASERv: 1000 neutrinos, 10000 muon neutrinos and 10 tau neutrino CC interactions.

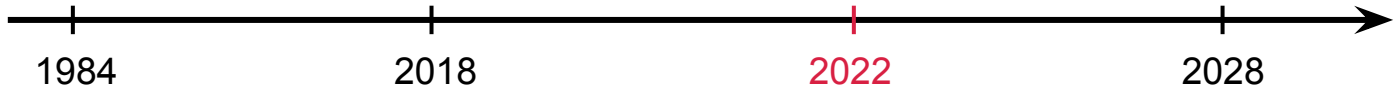
SND@LHC



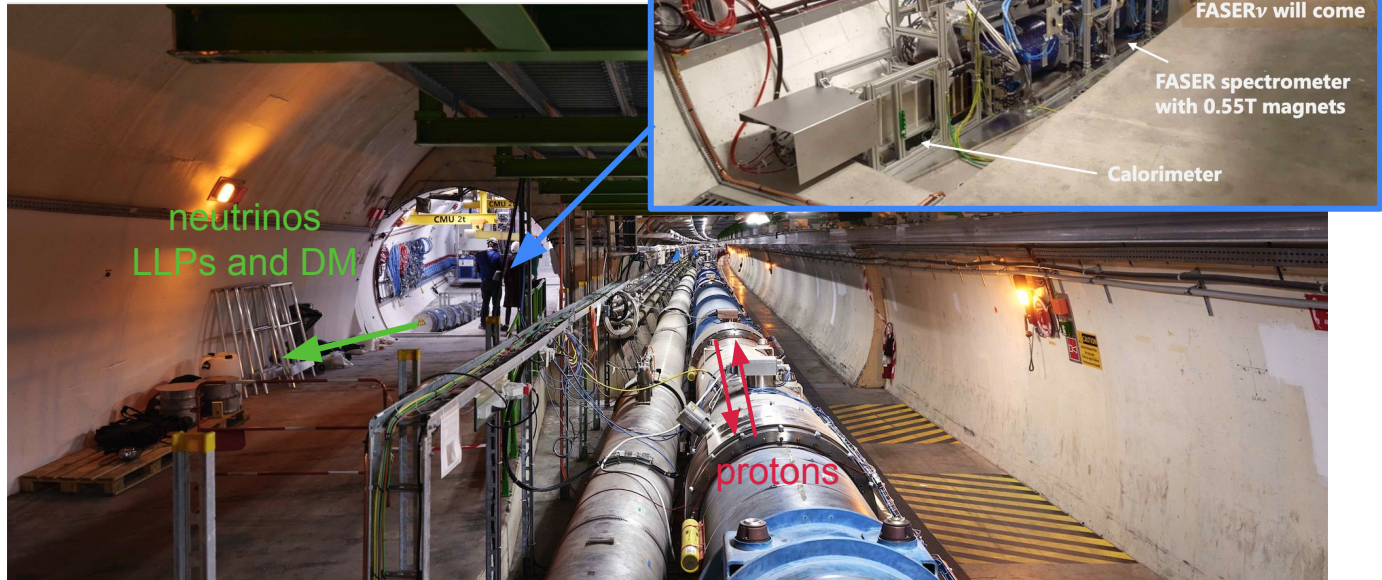
FASERnu



Neutrinos at the LHC



FASER detector was successfully installed into the TI12 tunnel in March 2021



John Osborne, Kincso Balazs, Jonathan Gall

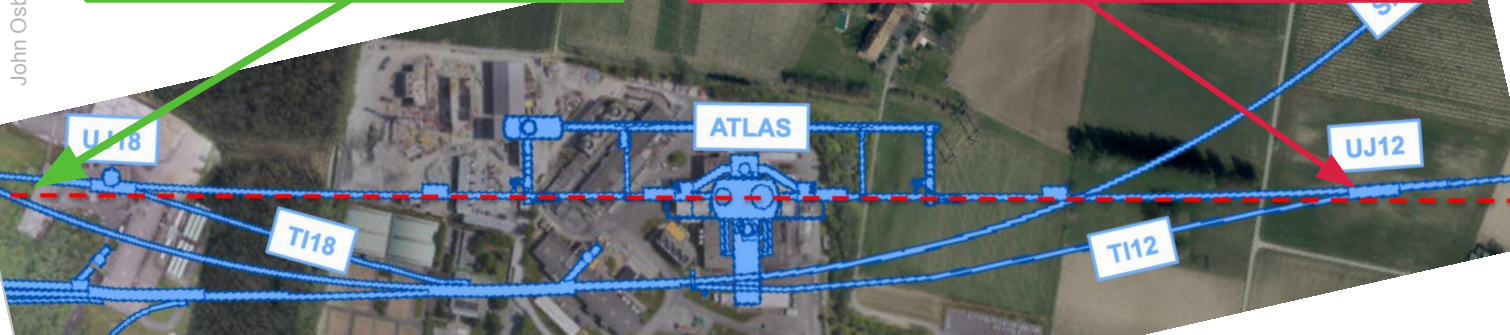
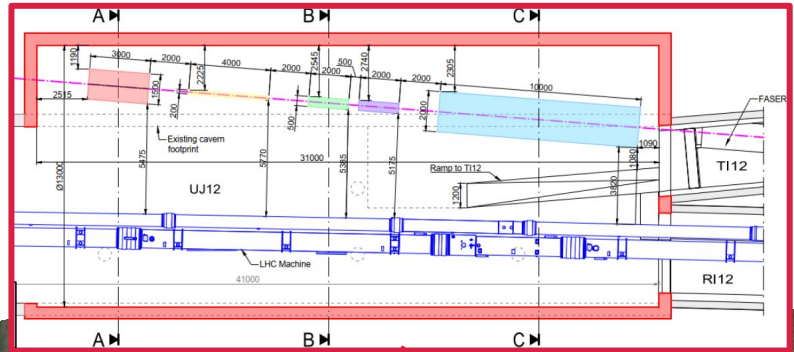


2018

2022

2028

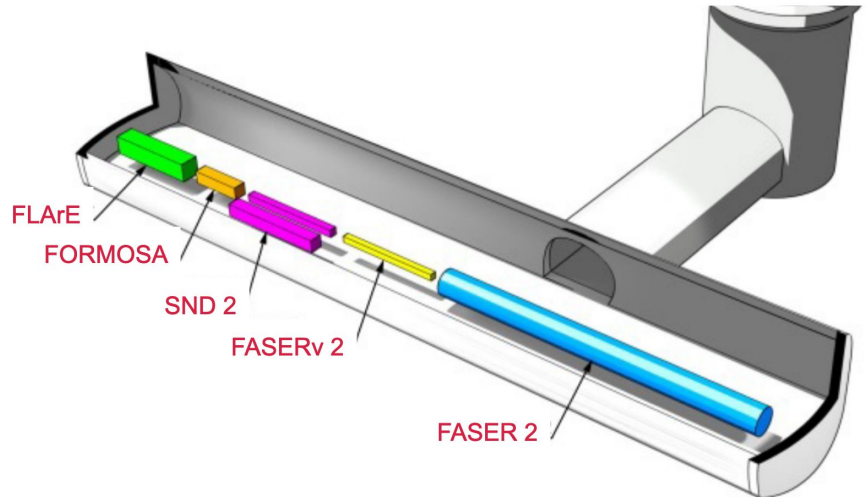
A 3D perspective view of the underground structure. A red shaft, labeled 'SHAFT ST1395418_01', is the central vertical feature. A red cavern, labeled 'CAVERN ST1395416_01', is located at the base of the shaft. A red access tunnel, labeled 'ACCESS TUNNEL ST1395417_01', connects the cavern to the shaft. Several grey pipes are shown: 'SPS' (Surface Pressure Supply) runs horizontally across the top; 'PM18' (Pressure Main) is a vertical pipe; 'R18' (Return Air) is a horizontal pipe; 'RT18' (Return Tunnel) is a horizontal pipe; 'T18' (Tunnel) is a horizontal pipe; 'UJ18' (Underground Junction) is a horizontal pipe; and 'R18' (Return Air) is a horizontal pipe.



Neutrinos at the LHC

A suite of experiments were proposed for the FPF.

We are currently writing a physics potential summary.
You are welcome to join!



Snowmass LOI: <https://zenodo.org/record/4059893>

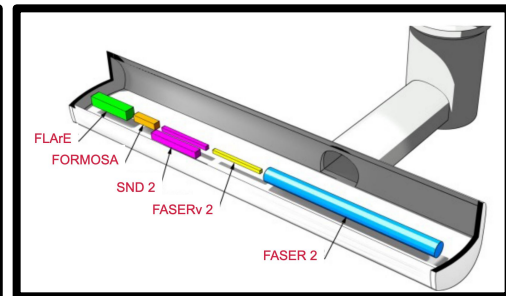
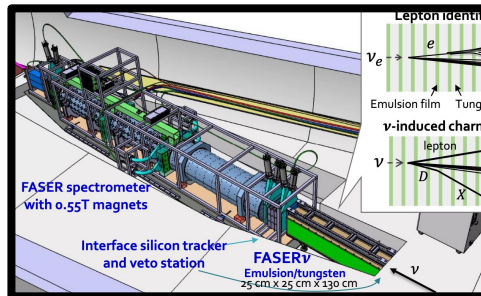
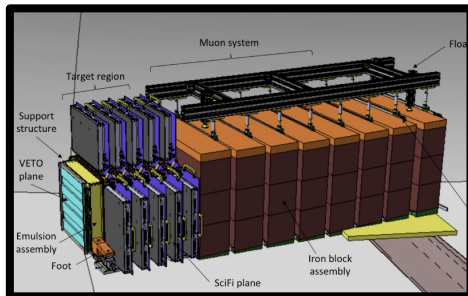
Kickoff Meeting: <https://indico.cern.ch/event/955956/>

2nd Workshop (last week): <https://indico.cern.ch/event/1022352>

Neutrinos at the LHC

Neutrino experiments at the LHC will greatly enhance the LHC's physics potential for **BSM physics searches**, **neutrino physics** and **QCD**.

In this talk, I will highlight some exciting examples.



Neutrino Physics

Neutrino Physics

The LHC neutrino beam is broad, with mean energies around 1 TeV, exceeding the energies of all other artificial neutrino sources.

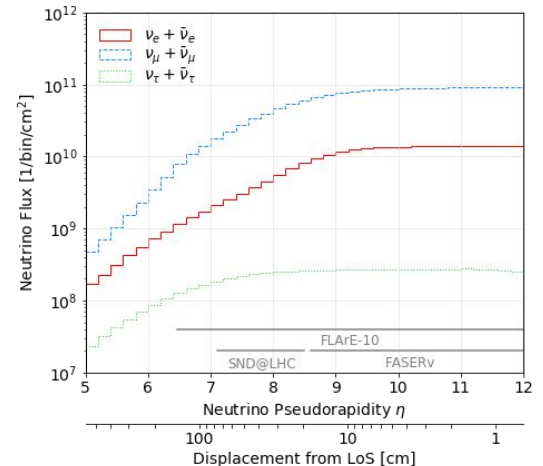
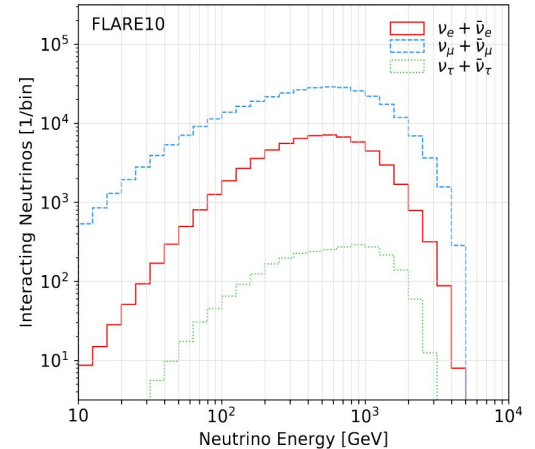
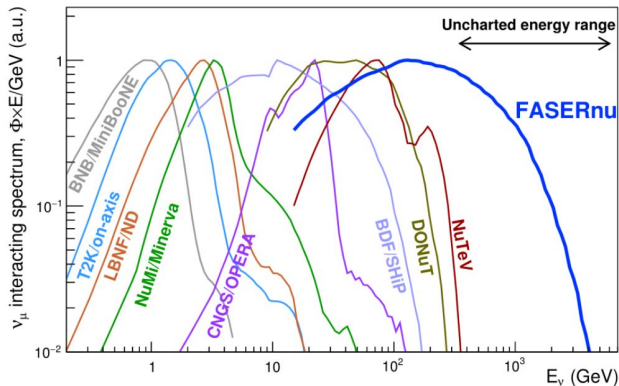
It originates from a variety of sources: pion, kaon, hyperon and charm decays.

FASERv (~1ton, 150/fb):

1k ν_e , 10k ν_μ , 10 ν_τ interactions

FPF (~10ton, 3000/fb):

100k ν_e , 1M ν_μ , 1k ν_τ interactions



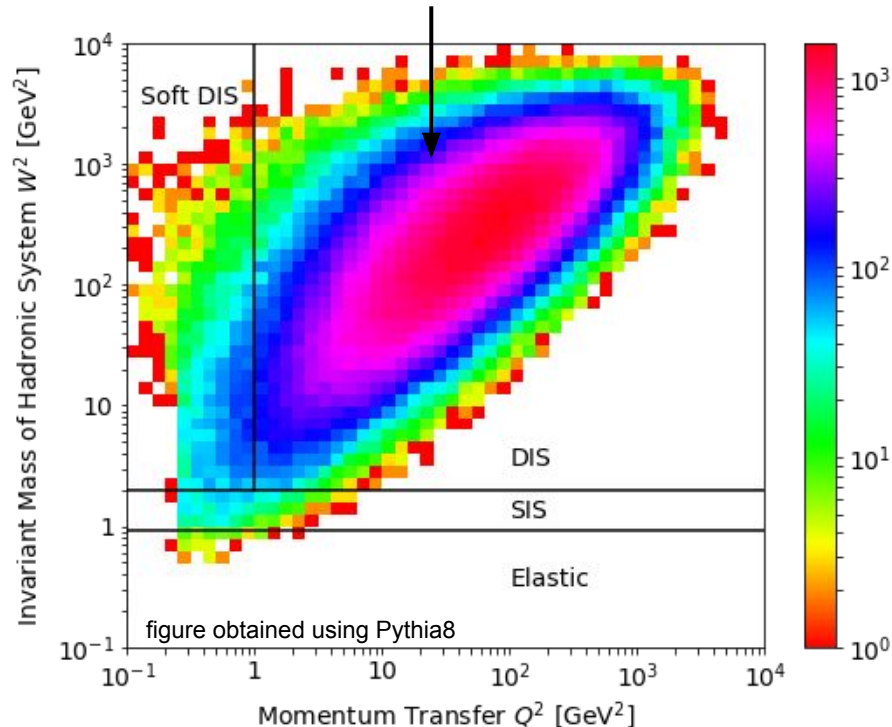
Neutrino Physics

Due to the high energy, most interactions are described by DIS: $\nu q \rightarrow l q'$

invariant mass of hadronic system \sim hadronic system particle multiplicity

DIS: $O(10)$ particles in final state

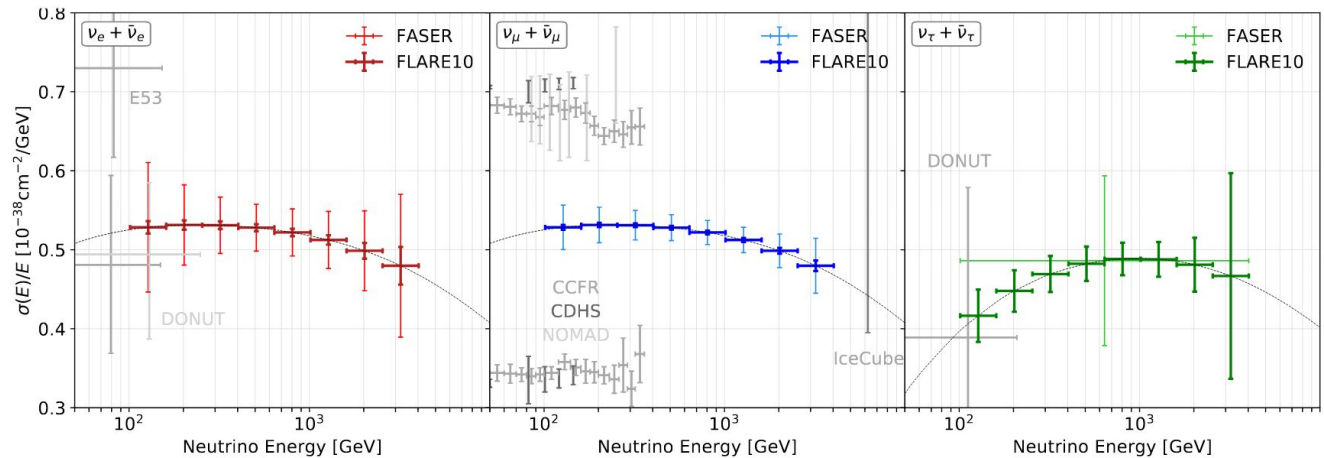
SIS: transition between resonance production to DIS



typical momentum transfer $|Q| \sim 10\text{GeV}$

Neutrino Physics

Using LHC neutrinos, one can measure **neutrino cross section** at unexplored TeV energies for all three flavors. Both CC and NC are possible.



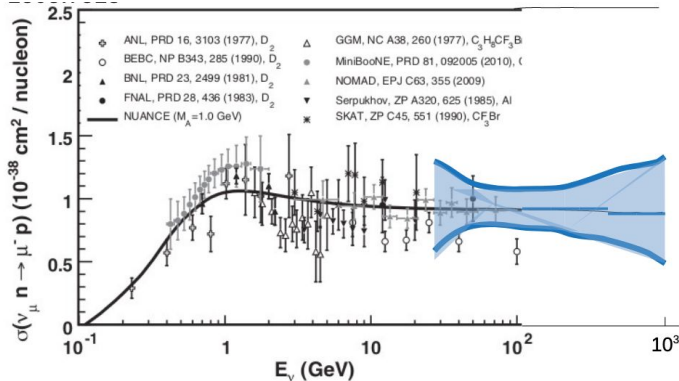
statistical uncertainty only

FASERv will detect ~ 10 **tau neutrino interactions**, which is similar to DONUT and OPERA. Thousands of tau neutrino events possible at HL-LHC, allowing for precision studies of tau neutrino properties.

Neutrino Physics

more than 1k quasi-elastic and resonant events expected

	CCQE				CCRES				NCEL	NCRES
	ν_e	ν_μ	$\bar{\nu}_e$	$\bar{\nu}_\mu$	ν_e	ν_μ	$\bar{\nu}_e$	$\bar{\nu}_\mu$	all	all
Event Rate	58	590	47	366	167	1673	184	1219	175	1206



Measurement of inclusive interaction cross section possible

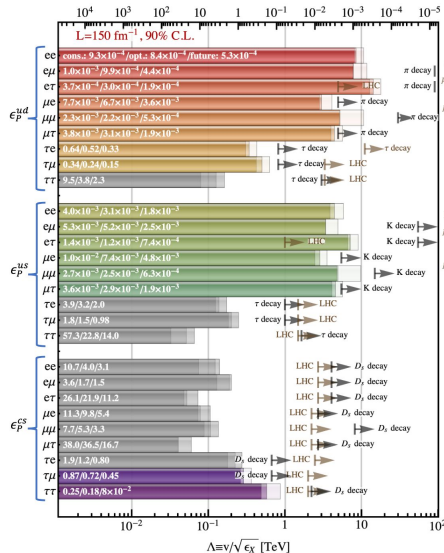
For high-energy neutrinos, typically >95% of E_ν goes into the outgoing lepton

Consistency check of the neutrino spectrum & cross section measurements

Neutrino Physics

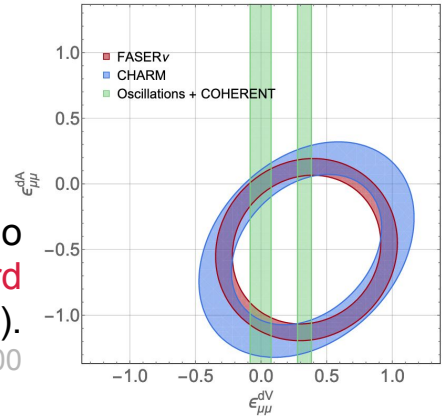
Interactions of LHC neutrino can also be used to constrain **SM EFT** coefficients

Falkowski, González-Alonso, Kopp, Soreq, Tabrizi 2015.12136

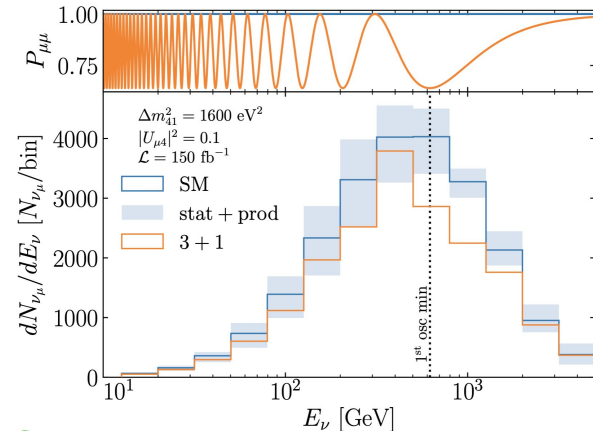


NC measurements could also constrain **neutrino non-standard interactions (NSI)**.

Abraham, Ismail, Kling 2012.10500



SM neutrino oscillations are expected to be negligible at FASERv. However, sterile neutrinos with mass $\sim 40 \text{ eV}$ can cause oscillations. FASERv could act as a short-baseline neutrino experiment.



QCD and Cosmic Rays

QCD and Cosmic Rays

Forward particle production is poorly constrained by other LHC experiments. FASERv's **neutrinos flux measurements** will provide novel complimentary constraints that can be used to validate/improve MC generators.

We need to **quantify** and reduce these uncertainties.

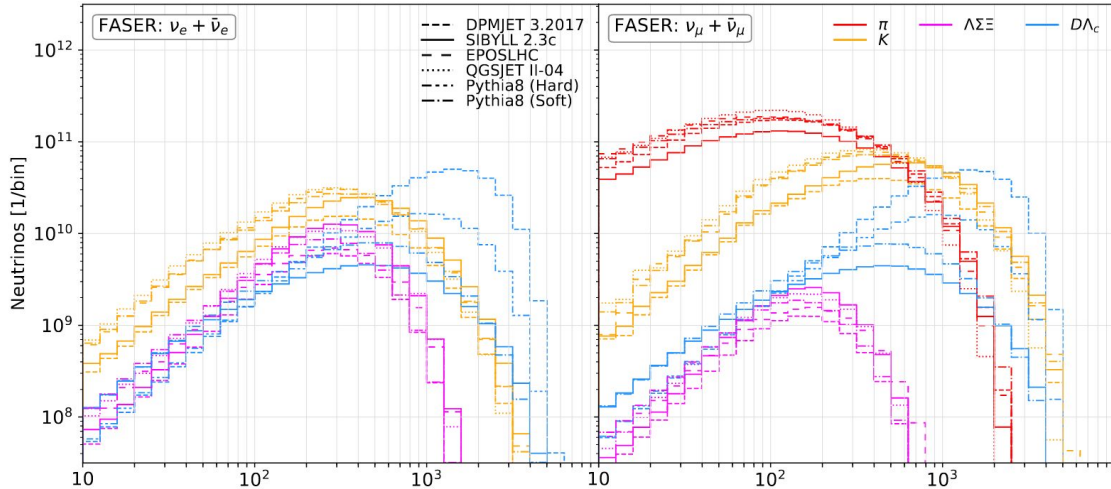
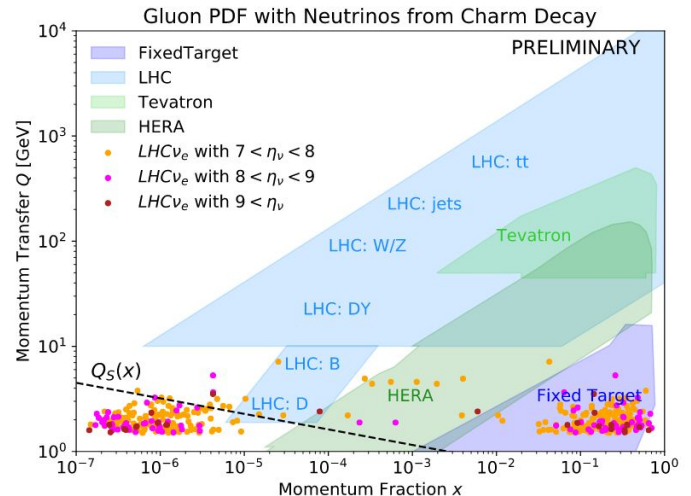
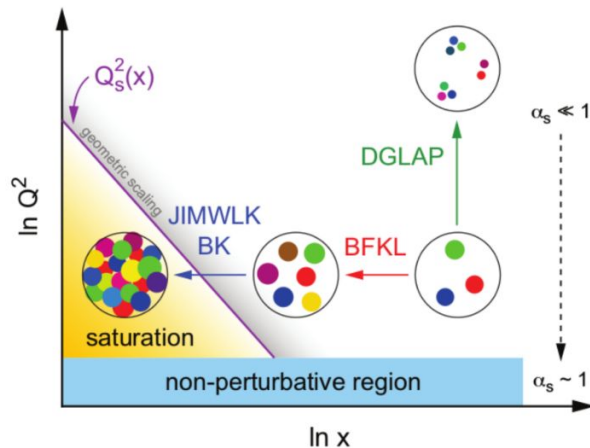


FIG. 5. Neutrino energy spectrum for electron neutrinos (left) and muon neutrinos (right) passing through FASER ν . The vertical axis shows the number of neutrinos per energy bin that go through the detector's cross sectional area for an integrated luminosity of 150 fb^{-1} . We separate the different production modes: pion decays (red), kaon decays (orange), hyperon decays (magenta) and charm decays (blue). The different line styles correspond to predictions obtained from different commonly used event generators.

QCD and Cosmic Rays

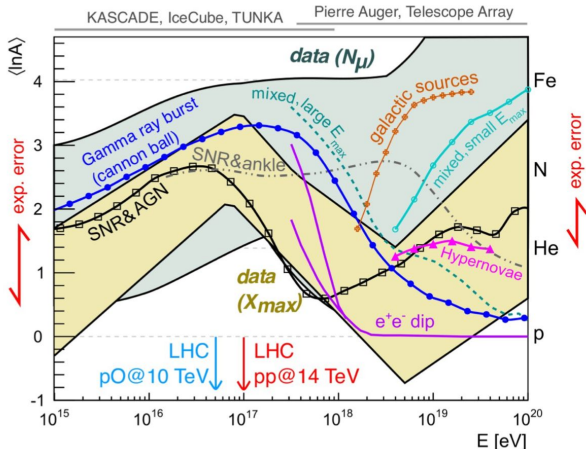
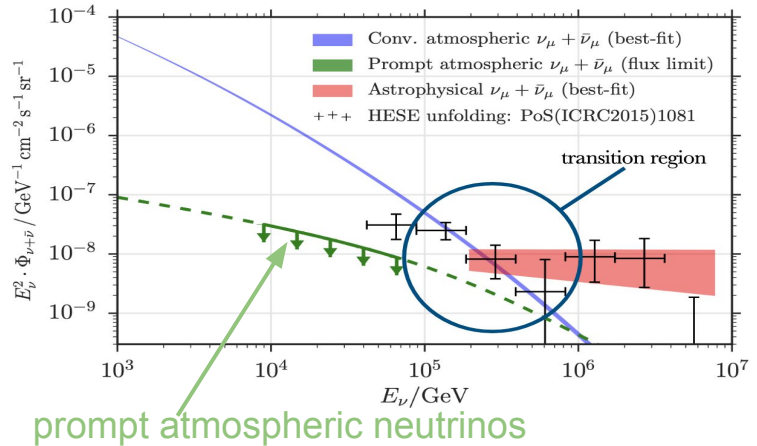
Electron neutrinos at high energy and tau neutrinos are mainly produced in charm decays: $g g \rightarrow c \bar{c}$, $c \rightarrow D$, $D \rightarrow K l \nu$

Neutrinos from charm decay could allow to test transition to **small-x factorization**, constrain **low-x gluon PDF**, probe **gluon saturation**, and probe **intrinsic charm**.



QCD and Cosmic Rays

Measuring forward charm production at the LHC would help to constrain the (currently very poorly constrained) **prompt atmospheric neutrino flux** at IceCube.



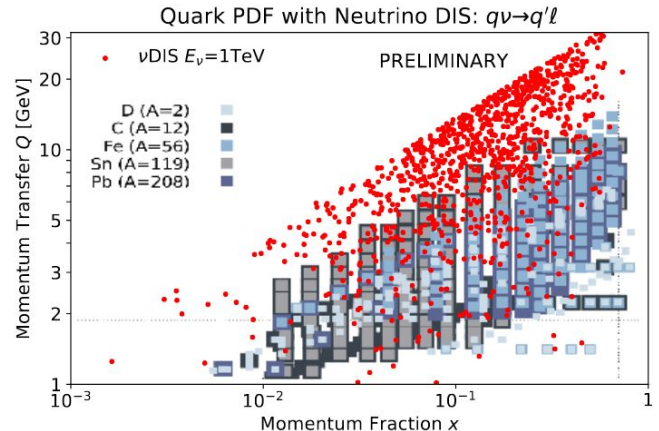
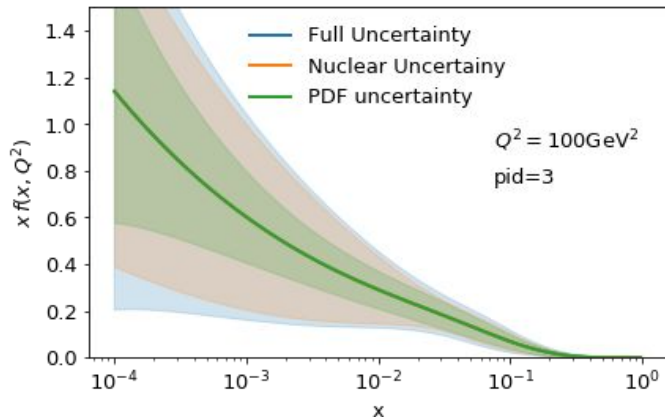
Cosmic Ray experiments have reported an excess in the number of muons over expectations computed using extrapolations of hadronic interaction models tuned to LHC data at the few σ level (**muon problem in CR physics**).

Measurements of forward hadron production (kaons) at the LHC are crucial to solve this issue.

QCD and Cosmic Rays

One can also use DIS neutrino scattering to probe (nuclear) PDFs: shadowing, anti-shadowing, EMC effect for different nuclear targets

In particular, charm associated neutrino events ($\nu s \rightarrow l c$) are sensitive to the poorly constrained **strange quark PDF**, and can help to resolve existing tension between different measurements.



BSM Physics

BSM Physics



Simple Model: Dark Matter charged under $U(1)_D$

$$\mathcal{L} \supset -\frac{\epsilon}{2} F^{\mu\nu} F'_{\mu\nu} - \frac{1}{2} m_{A'}^2 A'^2 - m_\chi^2 \chi^2 - ig_D A' \chi^2$$

coupling to SM via small
mixing with SM photon

massive gauge boson:
dark photon

dark photon
couples to DM

Phenomenology depends on masses:

$m_{A'} > 2m_\chi$: dark photon promptly decays in DM \rightarrow LHC produces DM beam

$m_{A'} < 2m_\chi$: dark photon can only decay to SM $\rightarrow A'$ is long-lived

$m_{A'} = 0$: dark matter becomes millicharged

BSM Physics - Dark Matter

DM Scattering in Neutrino Detector

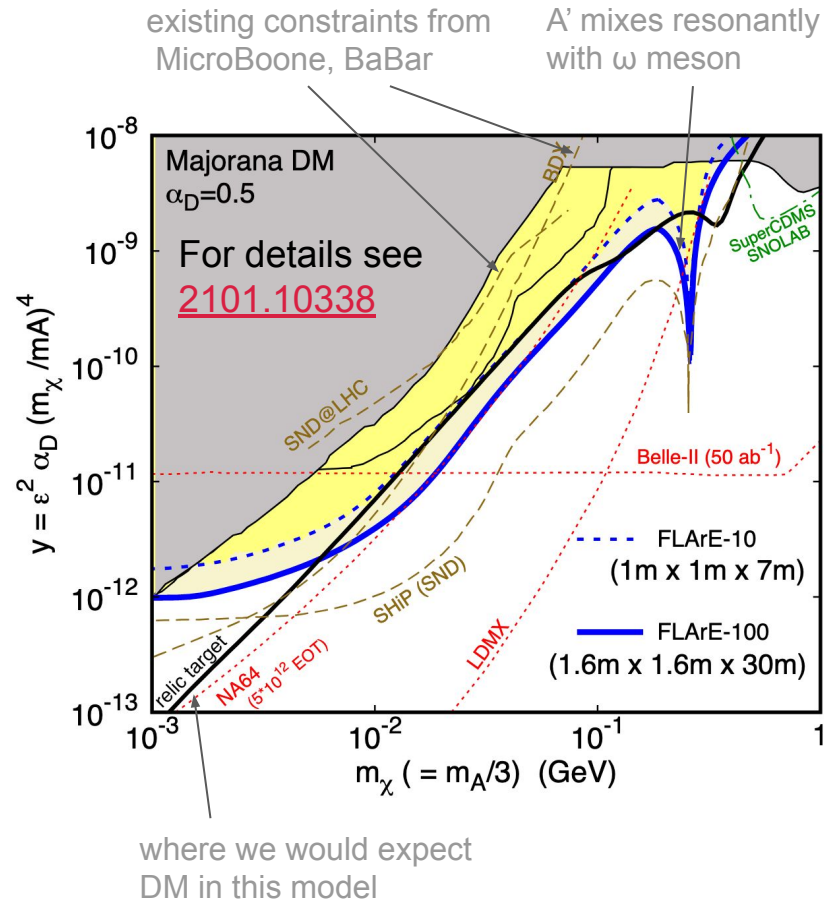
A huge number of high-energy mesons are produced in forward direction (hadronization of beam remnants)

A' produced via decays $\pi^0 \rightarrow A' \gamma$
or A' Bremsstrahlung $pp \rightarrow pp A'$

Prompt decay $A' \rightarrow XX$
produces DM beam

DM scatters on electrons: $X e \rightarrow X e$.
Typical electron energy $\sim 1\text{-}10$ GeV

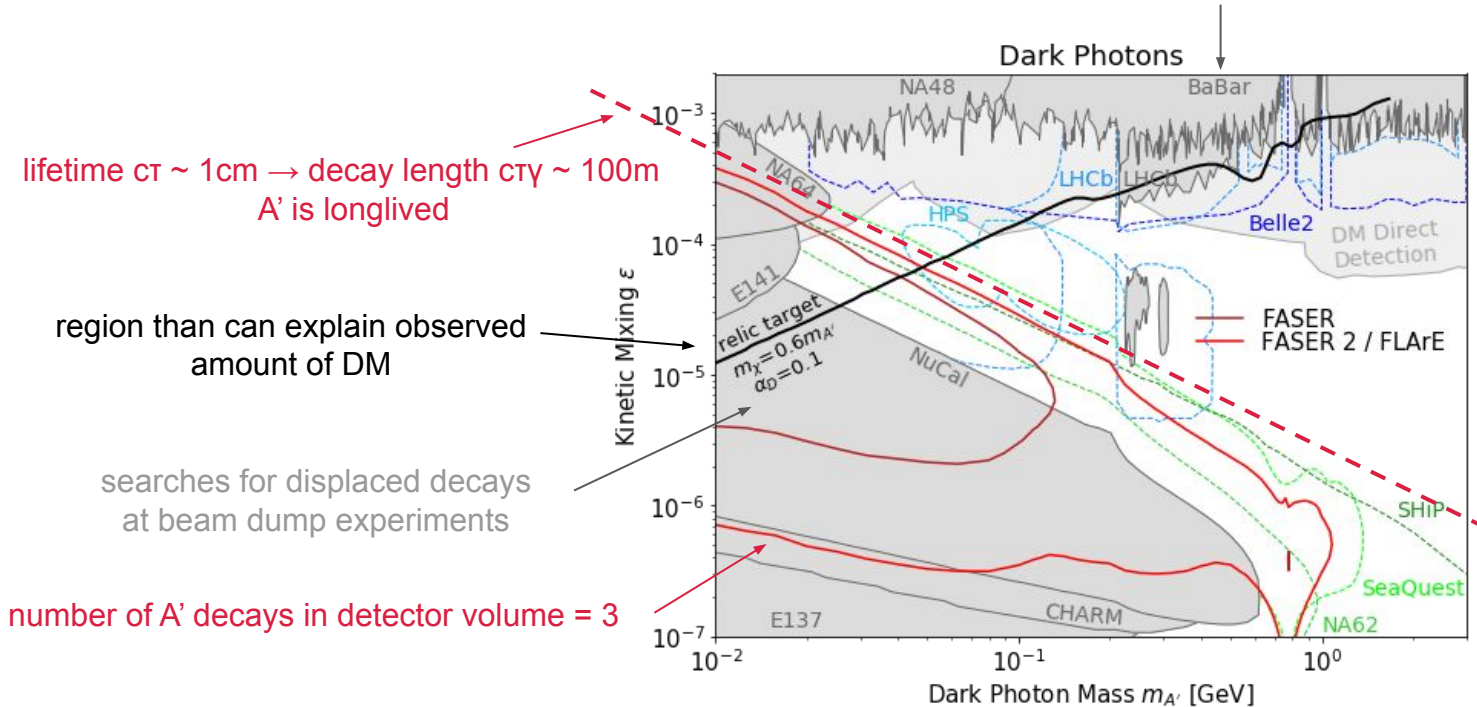
possible background: $\nu e \rightarrow \nu e$ with
typical electron energy $\sim 0.1\text{-}1$ TeV



BSM Physics - LLPs

If $m_{A'} < 2m_X$: A' decays to SM particles \rightarrow Long lived particle decays

searches for prompt di-electron resonance



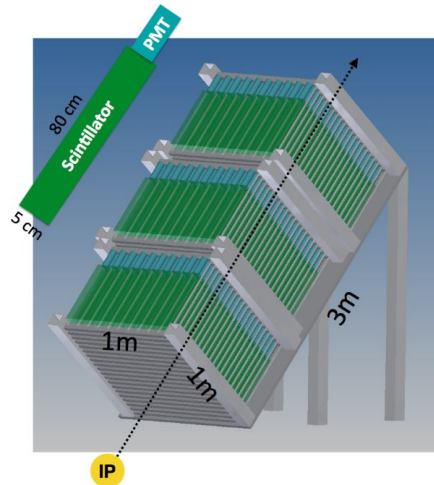
For details and many more models see [1811.12522](#).

BSM Physics: MCPs

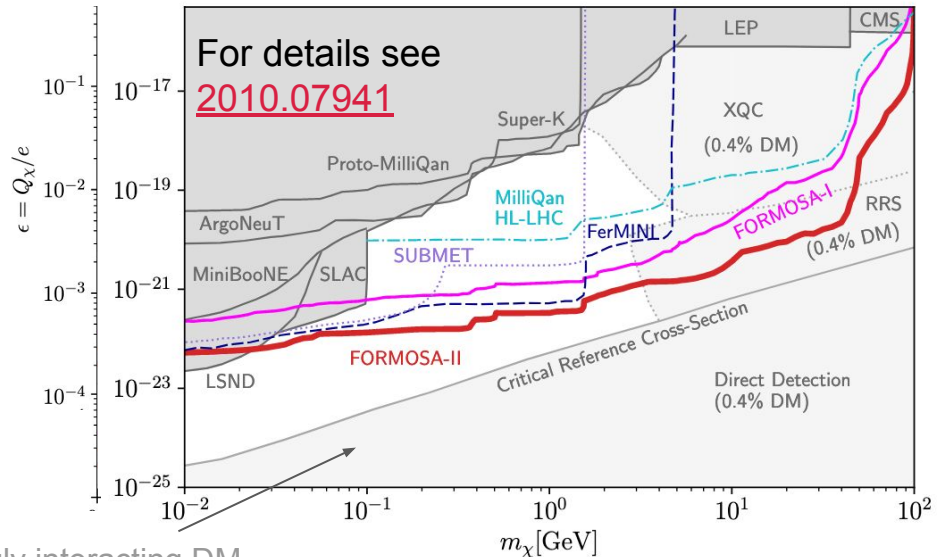
If $m_A'=0$: X is effectively **milli-charged** with $Q=\epsilon e \rightarrow$ search for minimum ionizing particle with very small dE/dx

MilliQan was proposed as dedicated LHC experiment to search for MCPs near CMS.
But it was noted that signal flux is ~ 100 times larger in forward direction.

LAr detector could in principle also look for MCPs (example: ArgoNeuT).



milliQan detector: [1607.04669](#)



MCPs are an example of strongly interacting DM.

Above DD bounds: DM absorbed in earth crust.

Popular model to explain EDGES anomaly.

Summary

With FASER and SND@LHC, two new experiment will soon start to perform neutrino measurements at the LHC.

They also paves the way for a forward search and neutrino program at the HL-LHC, opening up many many new opportunities for **neutrino physics**, **BSM physics searches** and **QCD measurements**, significantly extending the LHC's physics program.

We would like to invite the Invisibles 2021 community to help us explore and better understand the physics potential of this program.

We are currently writing a physics potential summary.
You are welcome to join!



contact me via felixk@slac.stanford.edu or
visit our workshop page <https://indico.cern.ch/event/1022352>