

# Dark matter bound states

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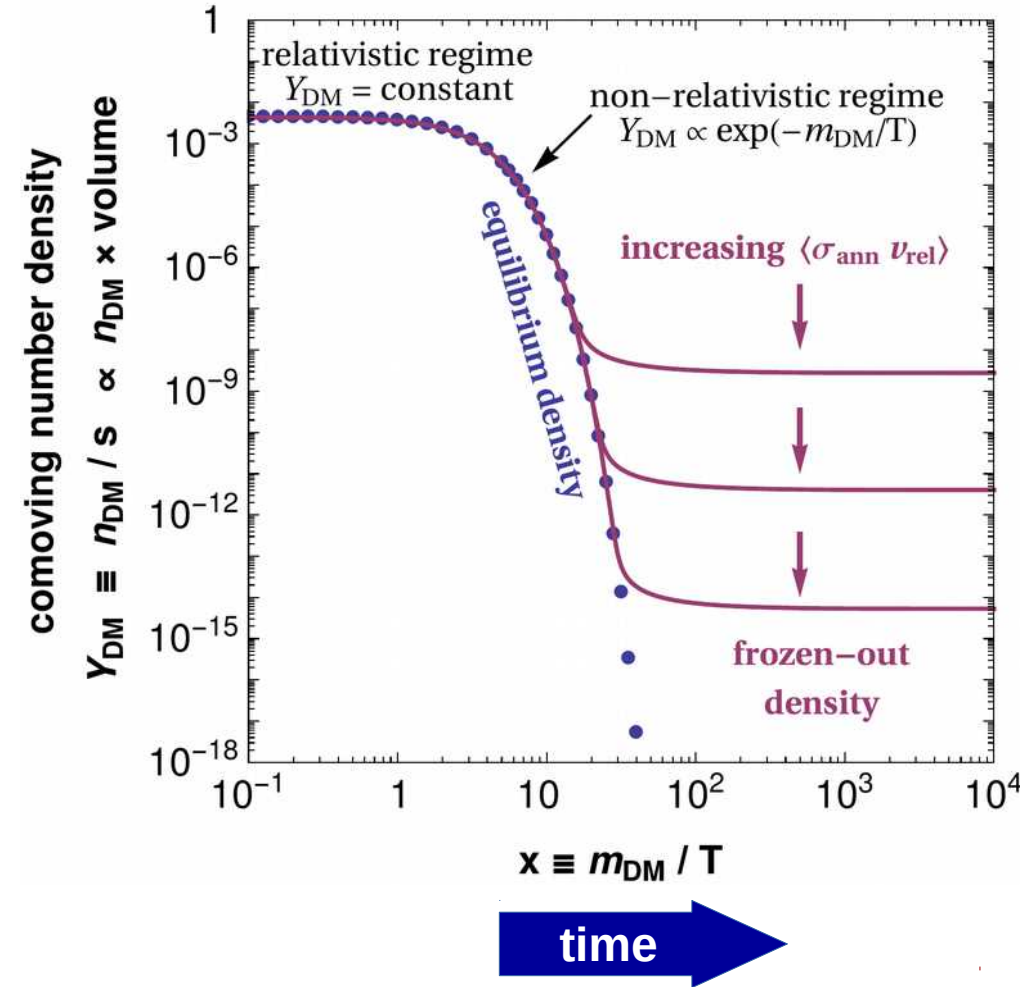
Invisibles meeting, 02 June 2021

# Dark matter production in the early universe

Why do we care?

Production depends on  
the couplings of DM to other particles,  
which are the very probes of the DM properties

# Dark matter production via thermal freeze-out



$$T > m_{\text{DM}}$$

DM kept in chemical & kinetic equilibrium with the plasma, via



$$n_{\text{DM}} \sim T^3 \quad \text{or} \quad Y_{\text{DM}} = \text{constant}$$

$$T < m_{\text{DM}}$$

$Y_{\text{DM}} \propto \exp(-m_{\text{DM}}/T)$ , while still in equilibrium

$$T < m_{\text{DM}} / 25$$

Density too small, annihilations stall  
⇒ **Freeze-out!**

$$\Omega \simeq 0.26 \times \left( \frac{1 \text{ pb} \cdot c}{\sigma_{\text{ann}} v_{\text{rel}}} \right)$$

1 pb ~  $\sigma_{\text{Weak}}$   
WIMP miracle!

# WIMPs and variations

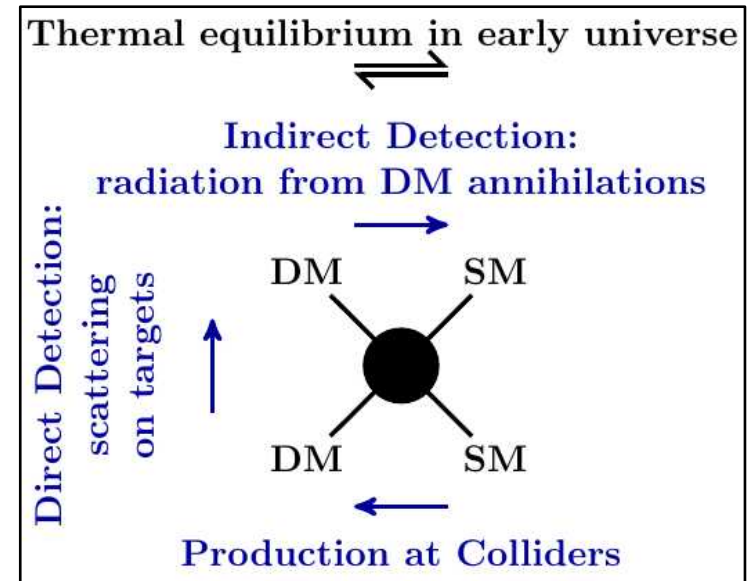
Weakly coupled to SM  
via  $W^\pm, Z, H$   
e.g. LSP in SUSY

or

weakly coupled to SM  
via non-SM interactions,  
e.g.  $\delta L = \frac{\bar{X} \gamma^\mu X \bar{q} \gamma_\mu q}{\Lambda^2}$

or

weakly coupled to light dark-sector  
particles that couple (feebly) to SM,  
e.g. DM coupled to dark photon  
kinetically mixed with Hypercharge



Significant  
constraints.

No discovery  
so far.

# What now?

## Diversify dark matter searches

- **Heavier DM**

Particles with  $m \gtrsim \text{TeV}$  coupled to SM via the Weak or other interactions not constrained by collider experiments

→ existing and upcoming telescopes observing multi-TeV sky with increasing sensitivity, e.g. HESS, IceCube, CTA, Antares

- **Lighter DM**

Particles with  $m \lesssim \text{few GeV}$ , possibly coupled to SM via a portal interaction, not constrained by older direct detection experiments

→ development of new generation of direct detection experiments

- Simple thermal-relic WIMP models live in the (multi-)TeV scale.
- Thermal-relic DM can be as heavy as  $\text{few} \times 100 \text{ TeV}$ .

**How heavy can thermal-relic DM be, and what are the underlying dynamics of heavy ( $\gtrsim \text{TeV}$ ) thermal-relic DM?**

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# Long-range interactions

If dark matter is very heavy, then:

$$\lambda_B \sim \frac{1}{\mu v_{\text{rel}}}, \quad \frac{1}{\mu \alpha} \lesssim \frac{1}{m_{\text{mediator}}} \sim \text{interaction range}$$

$\mu$ : reduced mass ( $m_{\text{DM}}/2$ )

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## Relevant for various models

- Self-interacting DM
- DM explanations of astrophysical anomalies, e.g. galactic positrons, IceCube PeV neutrinos
- WIMP DM with  $m_{\text{DM}} > \text{few TeV}$ . [Hisano et al. 2002]
- WIMP DM with  $m_{\text{DM}} < \text{TeV}$ ,  
in scenarios of DM co-annihilation with coloured partners.



# Implications of long-range interactions

## Sommerfeld effect

distortion of scattering-state wavefunctions  
⇒ affects all cross-sections, incl annihilation

- Freeze-out ⇒ changes correlation of parameters (mass – couplings)
- Indirect detection signals

## Bound states

- **Unstable bound states**  
⇒ extra annihilation channel
  - **Freeze-out**
  - Indirect detection
  - Novel low-energy indirect detection signals
- **Stable bound states (particularly important for asymmetric DM)**
  - Novel low-energy indirect detection signals
  - Affect DM self-interactions (screening)
  - Inelastic scattering in direct detection experiments (?)

## Outline

### Bound states and density of thermal relic DM

An iceberg floating in a blue ocean under a blue sky with light clouds. The tip of the iceberg is above the water, and the much larger part is submerged. The text 'Sommerfeld effect' is written in red on the tip, and 'Bound states' is written in red on the submerged part.

Sommerfeld  
effect

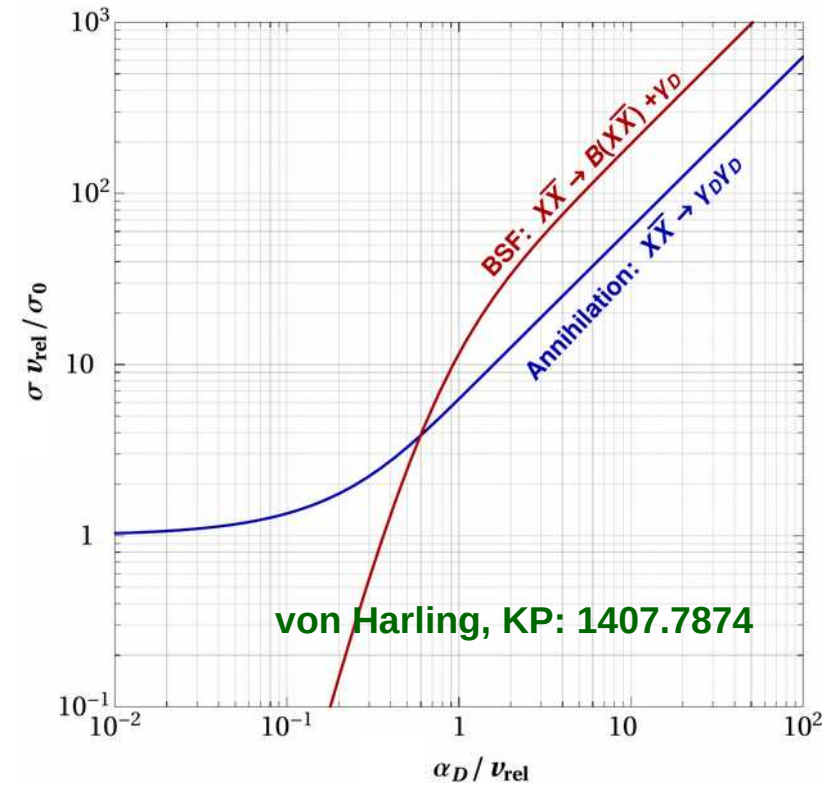
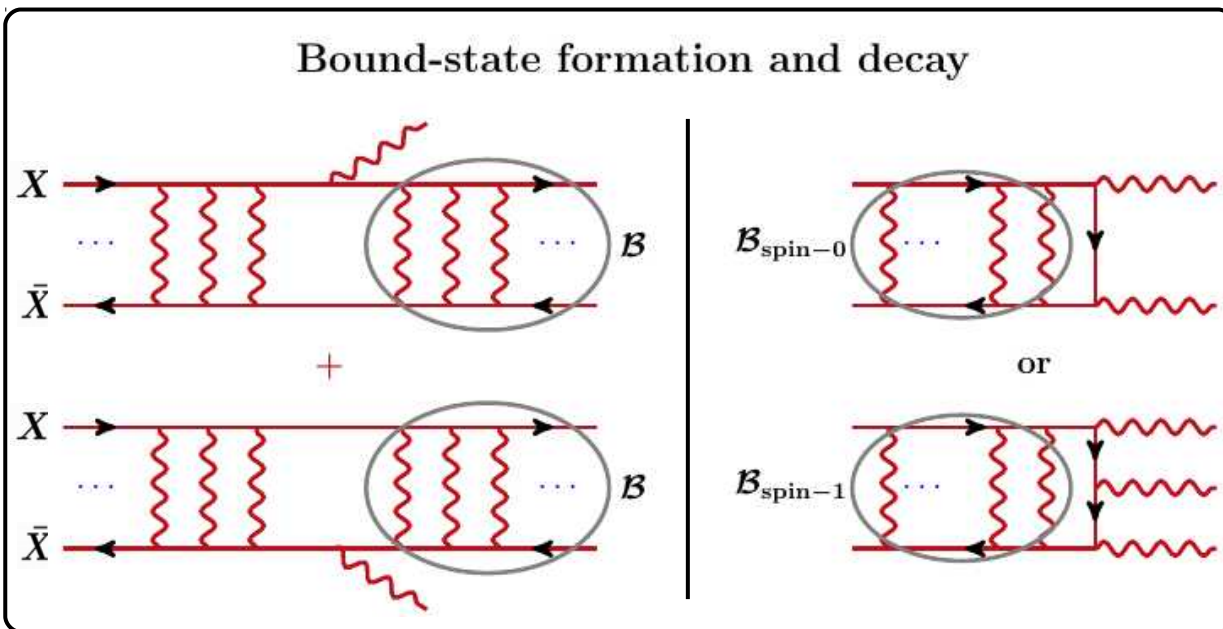
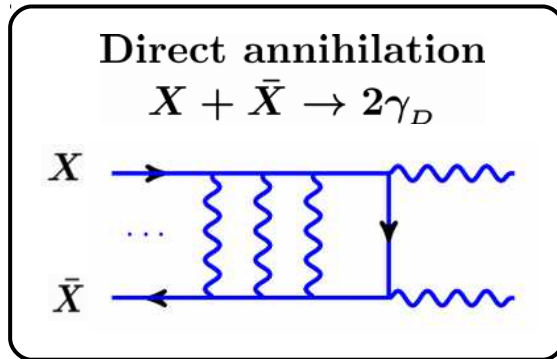
Bound  
states

- ◆ Dark U(1) sector
- ◆ Unitarity limit and long-range interactions
- ◆ Neutralino-squark coannihilation scenarios
- ◆ The Higgs-doublet as a light mediator

# Dark U(1) sector

# Thermal freeze-out with long-range interactions

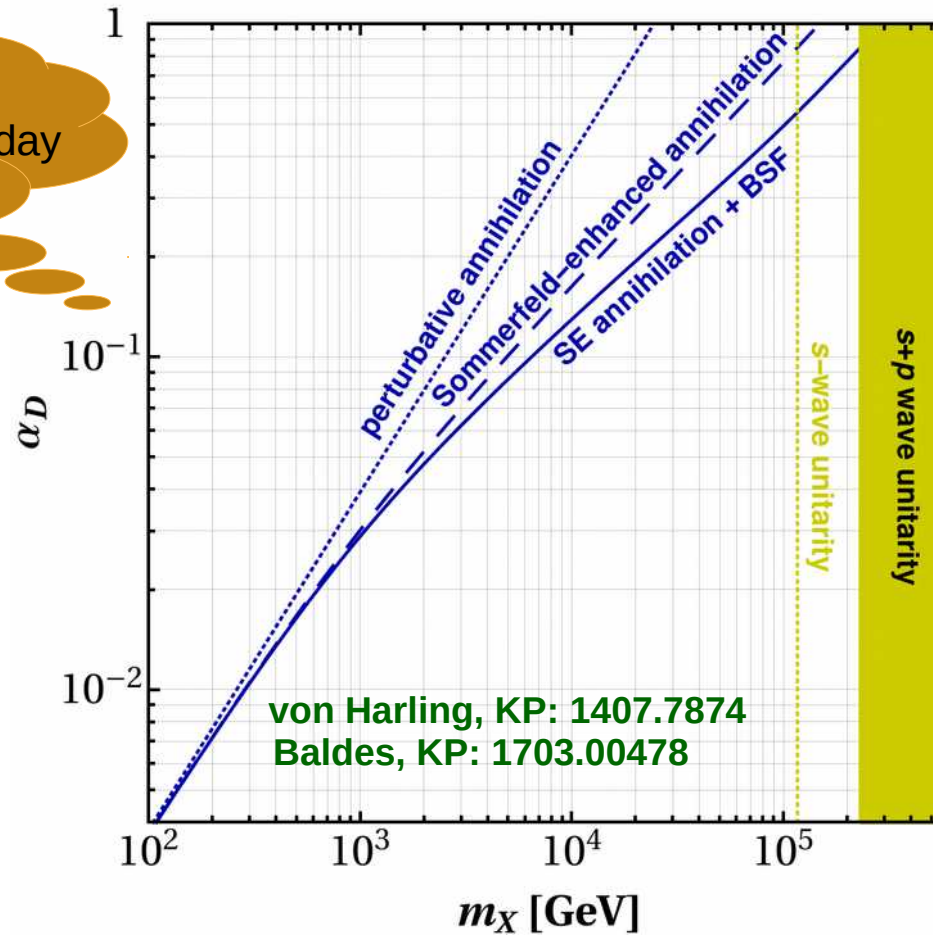
Dark U(1) model: Dirac DM  $X, \bar{X}$  coupled to  $\gamma_D$



# Thermal freeze-out with long-range interactions

Dark U(1) model: Dirac DM  $X, \bar{X}$  coupled to  $\gamma_D$

Important because it determines DM interactions today (direct, indirect detection)

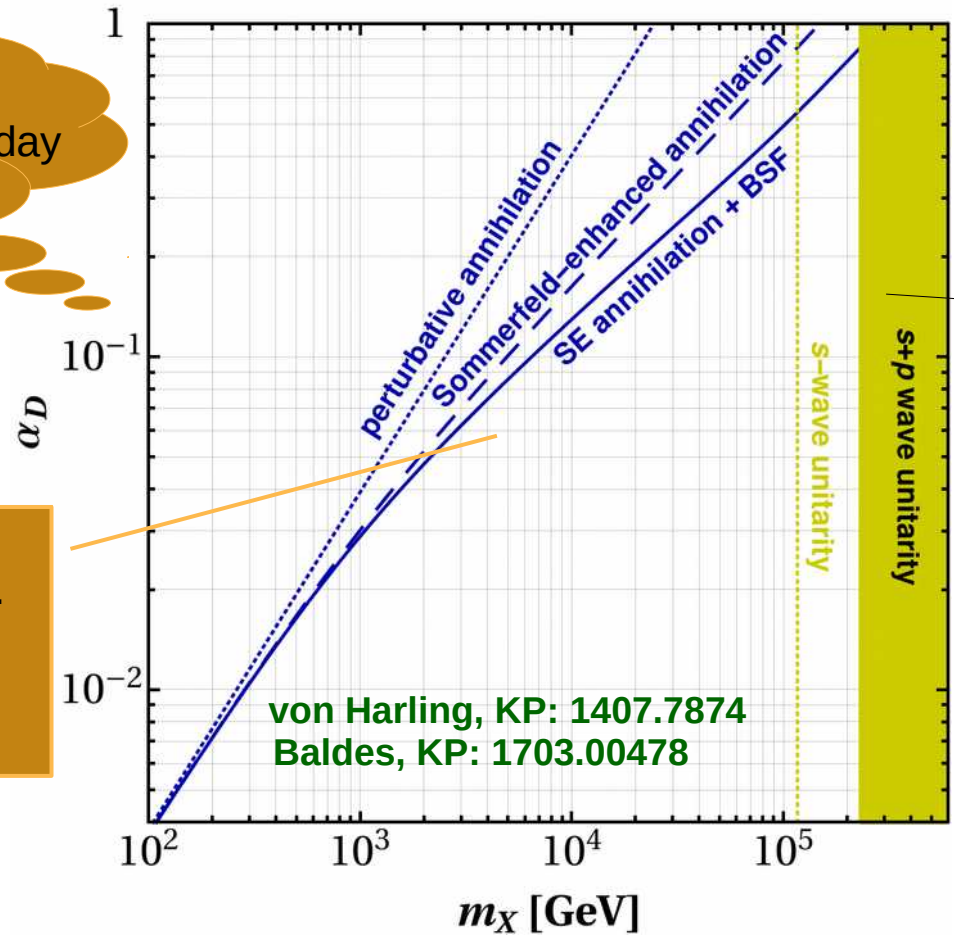


# Thermal freeze-out with long-range interactions

Dark U(1) model: Dirac DM  $X, \bar{X}$  coupled to  $\gamma_D$

Important because it determines DM interactions today (direct, indirect detection)

Long-range effects indeed become at  $m_{DM} \gtrsim$  few TeV.  
Verifies expectation from unitarity arguments!



Dominant annihilation mode: **s-wave**.  
Dominant BSF mode: **p-wave**  
Same order!  
Higher partial waves Important / dominant in multi-TeV regime.  
**DM may be even heavier!**

# Unitarity limit and long-range interactions

# Partial-wave unitarity limit

$$\sigma_{\text{inel}}^{(\ell)} v_{\text{rel}} \leq \sigma_{\text{uni}}^{(\ell)} v_{\text{rel}} = \frac{4\pi(2\ell + 1)}{M_{\text{DM}}^2 v_{\text{rel}}}$$

Implies upper bound on the mass of thermal-relic DM

Griest, Kamionkowski (1990)

$$\sigma_{\text{ann}} v_{\text{rel}} \simeq 2.2 \times 10^{-26} \text{ cm}^3/\text{s} \leq \frac{4\pi}{M_{\text{DM}}^2 v_{\text{rel}}}$$

$$\langle v_{\text{rel}}^2 \rangle^{1/2} = (6T/M_{\text{DM}})^{1/2} \xrightarrow[M_{\text{DM}}/T \approx 25]{\text{freeze-out}} 0.49$$

$$\Rightarrow M_{\text{uni}} \simeq \begin{cases} 117 \text{ TeV,} & \text{self-conjugate DM} \\ 83 \text{ TeV,} & \text{non-self-conjugate DM} \end{cases}$$

- Assumes contact-type interactions,  $\sigma v_{\text{rel}} = \text{constant}$
- Considers only s-wave annihilation



# Partial-wave unitarity limit

$$\sigma_{\text{inel}}^{(\ell)} v_{\text{rel}} \leq \sigma_{\text{uni}}^{(\ell)} v_{\text{rel}} = \frac{4\pi(2\ell + 1)}{M_{\text{DM}}^2 v_{\text{rel}}}$$

Parametric dependence implies that  $\sigma_{\text{uni}}$  can be approached / attained only by long-range interactions.

Long-range interactions imply bound states, which may form by higher partial waves

- Upper limit on mass of thermal relic DM higher
- **In viable thermal scenarios, expect long-range behavior at  $m_{\text{DM}} \gtrsim \text{few TeV!}$**

Baldes, KP: 1703.00478

# Neutralino-squark co-annihilation scenarios

# Neutralino in SUSY models

## Squark-neutralino co-annihilation scenarios

- Degenerate spectrum  $\rightarrow$  soft jets  $\rightarrow$  evade LHC constraints
- Large stop-Higgs coupling reproduces measured Higgs mass and brings the lightest stop close in mass with the LSP

$\Rightarrow$  DM density determined by “effective” Boltzmann equation

$$n_{\text{tot}} = n_{\text{LSP}} + n_{\text{NLSP}}$$

$$\sigma_{\text{ann}}^{\text{eff}} = \left[ n_{\text{LSP}}^2 \sigma_{\text{ann}}^{\text{LSP}} + n_{\text{NLSP}}^2 \sigma_{\text{ann}}^{\text{NLSP}} + n_{\text{LSP}} n_{\text{NLSP}} \sigma_{\text{ann}}^{\text{LSP-NLSP}} \right] / n_{\text{tot}}^2$$

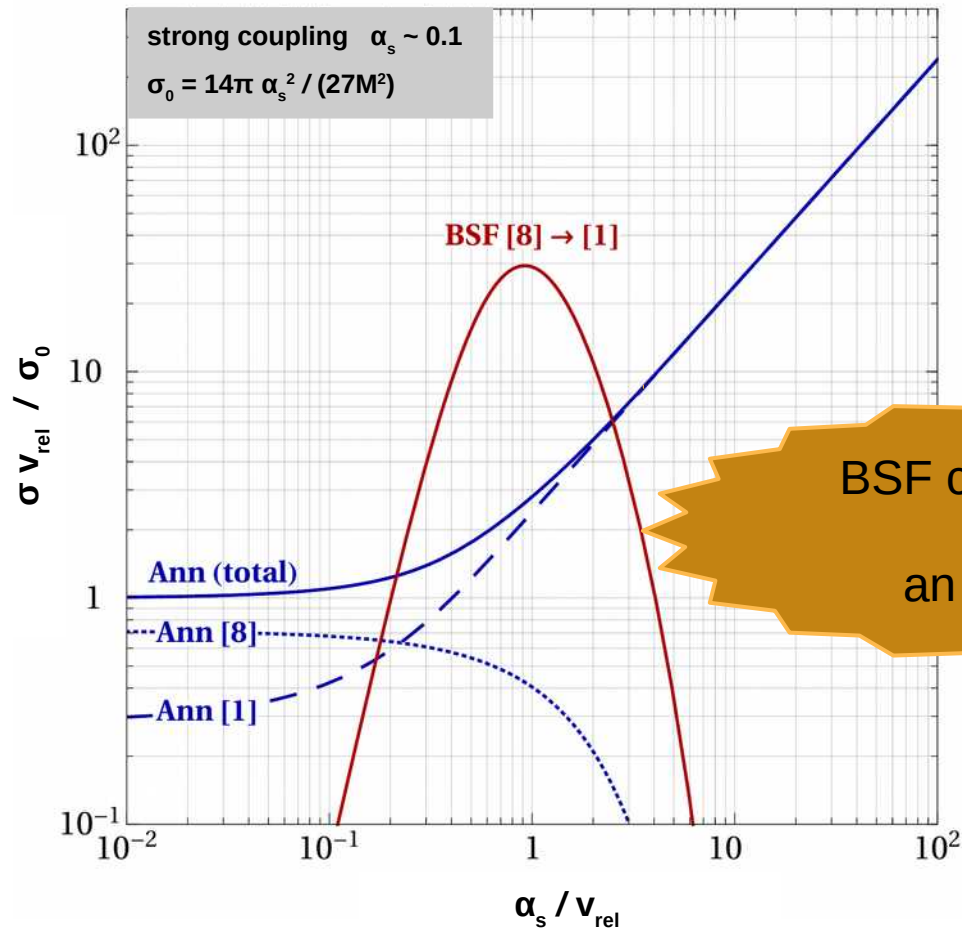
Scenario probed in colliders.  
 Important to compute DM density accurately!  
 $\rightarrow$  QCD corrections

# DM coannihilation with scalar colour triplet MSSM-inspired toy model

$$\begin{aligned} \mathcal{L} \supset & \frac{1}{2} \bar{\chi}^c i \not{\partial} \chi - \frac{1}{2} m_\chi \bar{\chi}^c \chi \\ & + \left[ (\partial_\mu + i g_s G_\mu^a T^a) X \right]^\dagger \left[ (\partial^\mu + i g_s G^{a,\mu} T^a) X \right] - m_X^2 |X|^2 \\ & + (\chi \leftrightarrow X, X^\dagger) \text{ interactions in chemical equilibrium during freeze-out} \end{aligned}$$

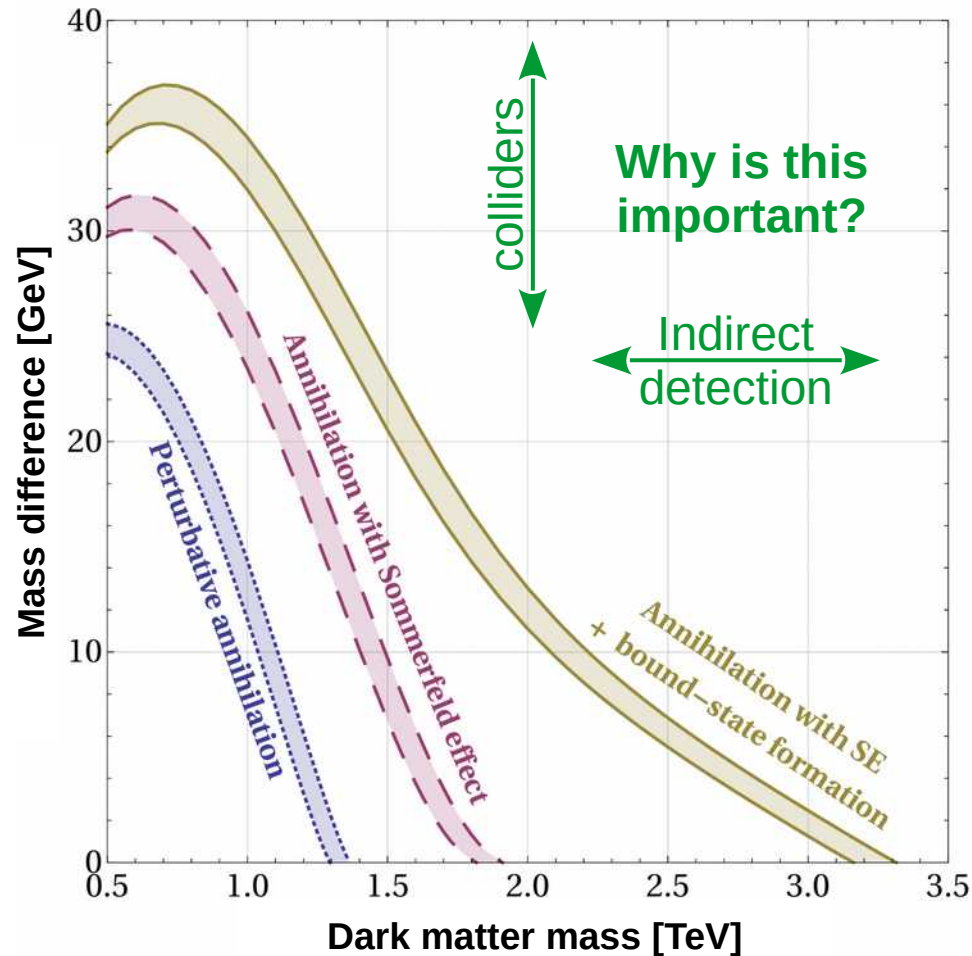
# DM coannihilation with scalar colour triplet MSSM-inspired toy model

## Bound-state formation vs Annihilation



BSF can exceed Annihilation  
by more than  
an order of magnitude!

# DM coannihilation with scalar colour triplet MSSM-inspired toy model



Effect on relic density:  
much much larger than  
obs uncertainty in  $\Omega_{\text{DM}}$

# The Higgs doublet as a light mediator

# The Higgs as a light mediator

- Sommerfeld enhancement of direct annihilation
- Binding of bound states

Harz, KP: 1711.03552

Harz, KP: 1901.10030

- Formation of bound states via Higgs (*doublet*) emission ?

Capture via emission of neutral scalar suppressed,  
due to selection rules: quadruple transitions

KP, Postma, Wiechers: 1505.00109  
An, Wise, Zhang: 1606.02305  
KP, Postma, de Vries: 1611.01394

Capture via emission of charged scalar [or its Goldstone mode]  
very very rapid: monopole transitions !

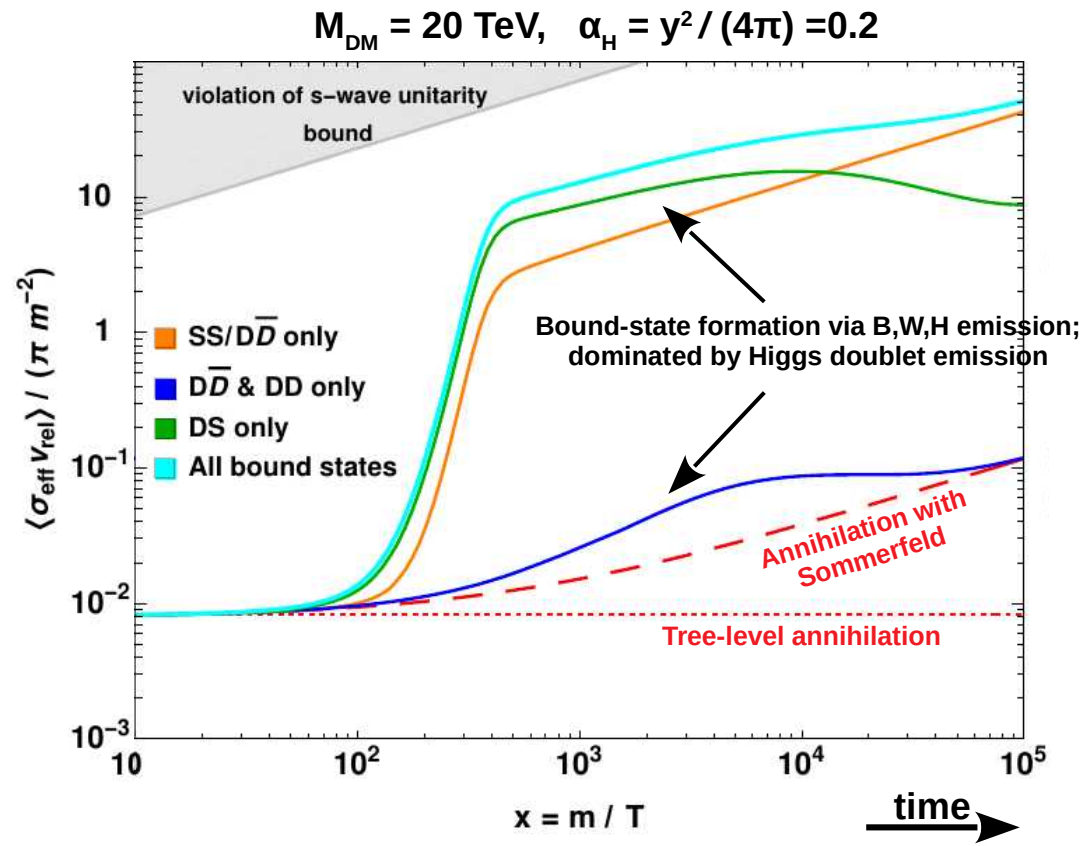
Ko, Matsui, Tang: 1910.04311  
Oncala, KP: 1911.02605  
Oncala, KP: 2101.08666/7

Sudden change in effective Hamiltonian precipitates transitions.  
Akin to atomic transitions precipitated by  $\beta$  decay of nucleus.



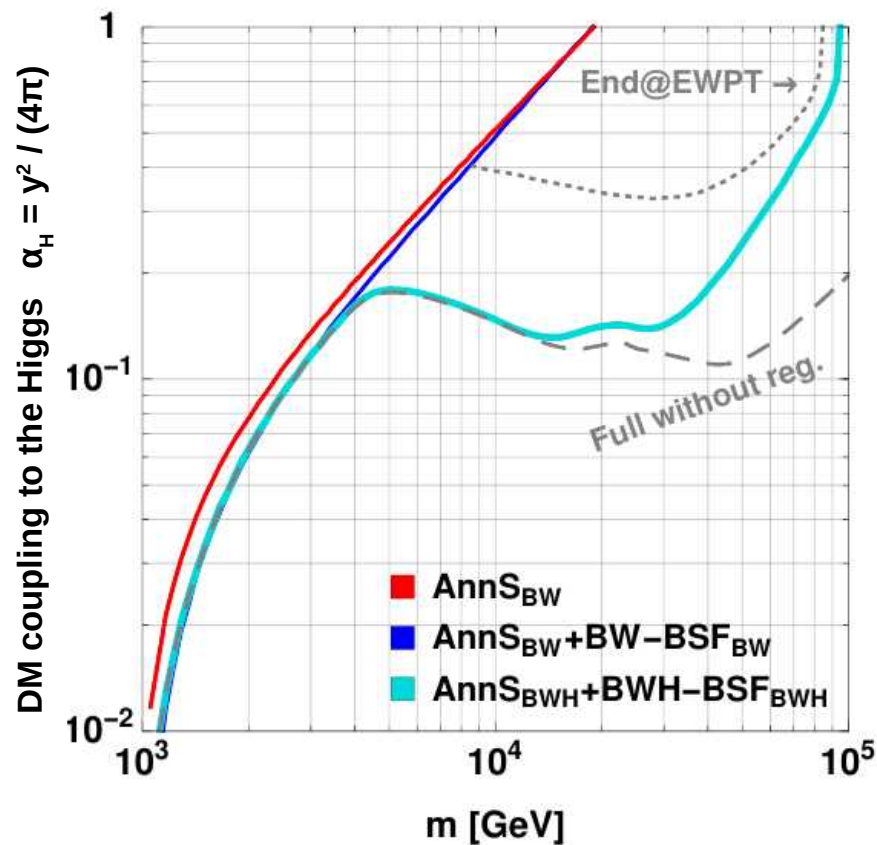
# Renormalisable Higgs-portal WIMP models

Mass-degenerate **Singlet-Doublet** coupled to the Higgs:  $L \supset -y \bar{D} H S$   
 $D$  &  $S$  co-annihilate; freeze-out begins before the EWPT if  $M_{DM} > 5\text{TeV}$



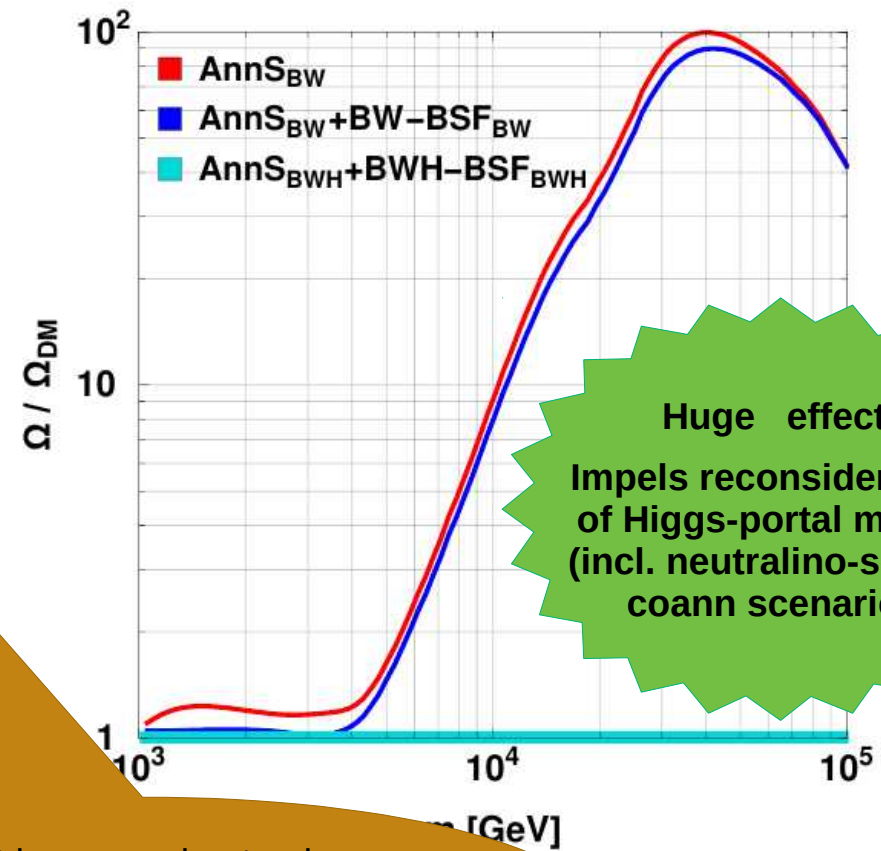
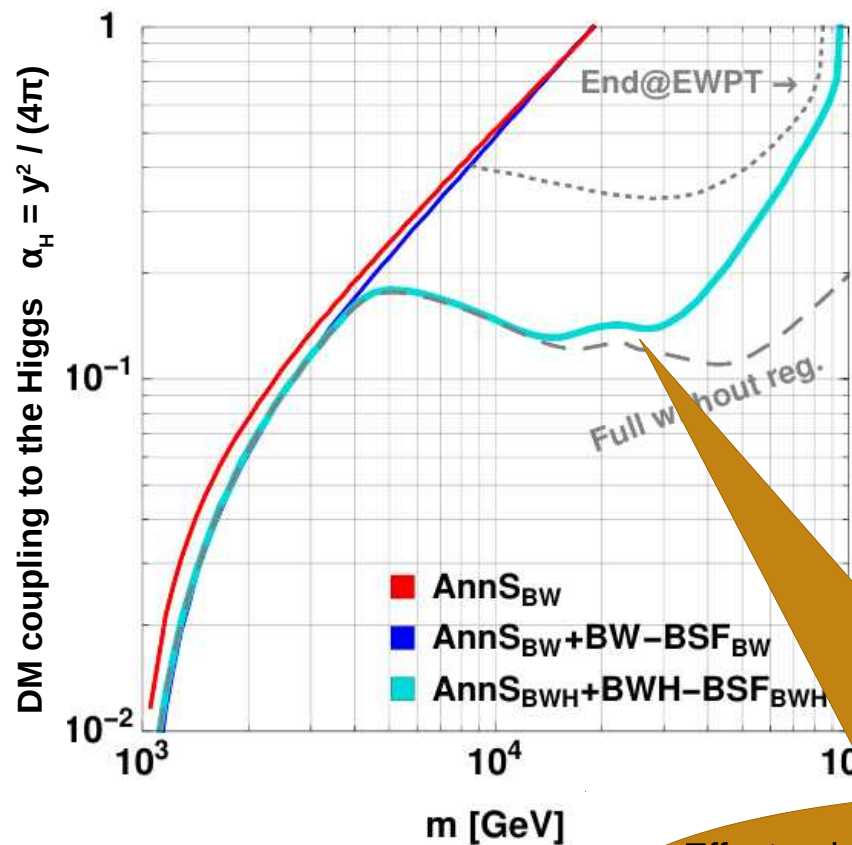
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**Huge effect!**  
 Impels reconsideration  
 of Higgs-portal models  
 (incl. neutralino-squark  
 coann scenarios)

Effect only at large  $\alpha_H$ , due to phase-space  
 suppression in Higgs emission.

For higher multiplets, important effect  
 expected also at lower  $\alpha_H$ .

# Conclusion

- **Bound states impel complete reconsideration of thermal decoupling at / above the TeV scale.**

Unitarity limit can be approached / realised only by attractive long-range interactions  $\Rightarrow$  bound states play very important role! [Baldes, KP: 1703.00478](#)

- **Important experimental implications:**
  - **DM heavier than anticipated:** multi-TeV probes very important.
  - **Indirect detection**
    - ♦ Enhanced rates due to BSF
    - ♦ Novel signals: low-energy radiation emitted in BSF
    - ♦ Indirect detection of asymmetric DM
  - **Colliders:** improved detection prospects due increased mass gap in coannihilation scenarios