

# R-Parity Violating $0\nu\beta\beta$ Decay with Light Neutralinos

Patrick D. Bolton<sup>1</sup>, Frank F. Deppisch<sup>1</sup> and P. S. Bhupal Dev<sup>2</sup>

<sup>1</sup> University College London, Gower Street, London, UK, WC1E 6BT <sup>2</sup> Washington University, St. Louis, MO 63130, USA



## 1 Introduction

In general, one can write down renormalisable terms in the supersymmetric (SUSY) superpotential that violate either lepton number ( $L$ ) or baryon number ( $B$ ), or so-called **R-parity** ( $R_p = (-1)^{3B+L+2S}$ , where  $S$  is the particle spin),

$$W_{R_p} = \underbrace{\lambda LLE^c}_{\Delta L=1} + \underbrace{\lambda' LQD^c}_{\Delta L=1} + \underbrace{\lambda'' UC^c D^c}_{\Delta B=1},$$

where  $L$ ,  $E^c$ ,  $Q$ ,  $U^c$  and  $D^c$  are chiral superfields containing the SM fermions and their scalar SUSY partners.

It has long been known that the  $\Delta L = 1$   $\lambda'$  term contributes to neutrinoless double beta ( $0\nu\beta\beta$ ) decay with the exchange of either a **neutralino** ( $\tilde{\chi}_i^0$ ) or **gluino** ( $\tilde{g}$ ) [1]. In the Minimal SUSY Standard Model (MSSM), the neutralinos are admixtures of the neutral fermionic SUSY partners of the  $U(1)_Y$  and  $SU(2)_L$  gauge bosons and Higgs doublets. The gluinos are the color-charged fermionic SUSY partners of the  $SU(3)_c$  gluons.

It is usually assumed that the lightest neutralino is **heavier** than the average momentum exchange of  $0\nu\beta\beta$  decay, i.e.  $m_{\tilde{\chi}_1^0} \gg p_F \sim 100$  MeV. However, indirect collider bounds on the lightest neutralino can be evaded so that it can be very light or even massless [2]. The  $0\nu\beta\beta$  decay process can therefore display **short-range** ( $m_{\tilde{\chi}_1^0} \gg p_F$ ) or **long-range** ( $m_{\tilde{\chi}_1^0} \ll p_F$ ) behaviour.

## 2 RPV Contribution to $0\nu\beta\beta$ Decay

For a heavy neutralino, the presence of the R-parity violating (RPV) coupling  $\lambda'_{111}$  results in the **dimension-9** Lagrangian

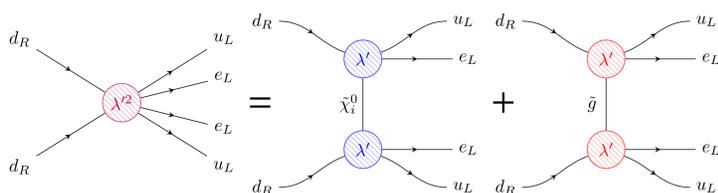
$$\mathcal{L}_9 = \frac{G_F^2 \cos^2 \theta_C}{2m_p} \left( \epsilon_1^{RRL} J_R J_R + \epsilon_2^{RRL} J_R^{\mu\nu} J_{R\mu\nu} \right) j_L + \text{h.c.},$$

where  $j_L = \bar{e}(1 + \gamma_5)e^c$ ,  $J_R = \bar{u}(1 + \gamma_5)d$  and  $J_R^{\mu\nu} = \bar{u}\sigma^{\mu\nu}(1 + \gamma_5)d$  for  $\sigma_{\mu\nu} = \frac{i}{2}[\gamma_\mu, \gamma_\nu]$ , i.e. **scalar** and **tensor** quark currents [1].

► The scalar and tensor coefficients can be decomposed as

$$\epsilon_1^{RRL} = \eta_{\tilde{\chi}} + \eta_{\tilde{\chi}\tilde{e}} + \eta_{\tilde{\chi}\tilde{f}} + \eta_{\tilde{g}} + \eta'_{\tilde{g}}, \quad \epsilon_2^{RRL} = -\frac{1}{4}(\eta_{\tilde{\chi}} + \eta_{\tilde{g}}),$$

where the  $(\eta_{\tilde{\chi}}, \eta_{\tilde{\chi}\tilde{e}}, \eta_{\tilde{\chi}\tilde{f}})$  and  $(\eta_{\tilde{g}}, \eta'_{\tilde{g}})$  factors encode the contributions from neutralino and gluino exchange diagrams respectively, i.e.



► The  $0\nu\beta\beta$  decay half-life is then given by

$$(T_{1/2}^{0\nu})^{-1} = G_{0\nu} \left| \epsilon_1^{RRL} M_1^{RR} + \epsilon_2^{RRL} M_2^{RR} \right|^2,$$

where  $G_{0\nu}$  is a phase space factor and  $M_1^{RR}$ ,  $M_2^{RR}$  are nuclear matrix elements.

► **QCD corrections** induce an RGE running and mixing of the coefficients  $\epsilon_1^{RRL}$  and  $\epsilon_2^{RRL}$  from the scale of new physics  $\Lambda_{NP} \sim 1$  TeV to the QCD scale  $\Lambda_{QCD} \sim 1$  GeV [3].

## 3 Collider Constraints and a Light Neutralino

• The factors  $\eta_{\tilde{\chi}}, \eta_{\tilde{\chi}\tilde{e}}, \eta_{\tilde{\chi}\tilde{f}}, \eta_{\tilde{g}}$  and  $\eta'_{\tilde{g}}$  depend on the selectron ( $\tilde{e}_L$ ), up squark ( $\tilde{u}_L$ ), down squark ( $\tilde{d}_R$ ), neutralino and gluino masses.

► Collider experiments **ATLAS** and **CMS** have excluded portions of the SUSY partner mass parameter space via searches for e.g.,  $\tilde{e}_L \rightarrow e\tilde{\chi}_1^0$ ,  $\tilde{q} \rightarrow q\tilde{\chi}_1^0$  and  $\tilde{g} \rightarrow q\bar{q}\tilde{\chi}_1^0$  (missing  $E_T$ ). Direct constraints on  $\lambda'_{111}$  can be made from measurements of  $V_{ud}$  and  $R_\pi$ .

• To describe the **long-range** behaviour of a light neutralino, we replace

$$\eta_{\tilde{\chi}}, \eta_{\tilde{\chi}\tilde{e}}, \eta_{\tilde{\chi}\tilde{f}} \propto \frac{\lambda'_{111}}{m_{\tilde{\chi}_1^0}} \rightarrow \frac{\lambda'_{111} m_{\tilde{\chi}_1^0}}{\langle \mathbf{p}^2 \rangle + m_{\tilde{\chi}_1^0}^2}$$

where  $\langle \mathbf{p}^2 \rangle = 150$  MeV. A similar **interpolating** function can be used for a sterile neutrino  $N$  with mass  $m_N$  above or below  $p_F$ .

## 4 Bounds on SUSY Parameter Space from $0\nu\beta\beta$ Decay and Other Experiments

• Lower limit on the  $0\nu\beta\beta$  half-life  $(T_{1/2}^{0\nu})_{\text{exp}}$  can be related to the RPV prediction. The resulting inequality can be rearranged for  $\lambda'_{111}$ ,

$$(T_{1/2}^{0\nu})_{\text{exp}} < T_{1/2}^{0\nu} \Rightarrow \lambda'_{111} < F(m_{\tilde{\chi}_1^0}, m_{\tilde{e}_L}, m_{\tilde{u}_L}, m_{\tilde{d}_R}, m_{\tilde{g}}, (T_{1/2}^{0\nu})_{\text{exp}})$$

i.e. an upper limit on  $\lambda'_{111}$  as a function of  $(T_{1/2}^{0\nu})_{\text{exp}}$  and the SUSY masses.

► **Parameter Scan**: Perform scan of SUSY masses allowed by collider experiments, with absolute lower limit from phenomenological MSSM analyses  $\Rightarrow$  Find corresponding upper limit on  $\lambda'_{111}$ .

► Best  $0\nu\beta\beta$  limits for  $m_{\tilde{e}_L} = 700$  GeV (lower limit from ATLAS for  $m_{\tilde{\chi}_1^0} = 0$  GeV) and  $m_{\tilde{e}_L} = 90$  GeV (OPAL lower limit for  $m_{\tilde{\chi}_1^0} > m_{\tilde{e}_L}$ ).

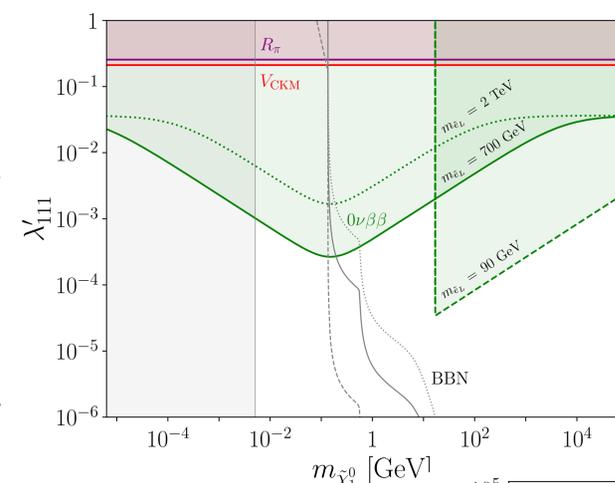
• Big Bang Nucleosynthesis (BBN): Naive limits by requiring the lifetime of the lightest neutralino be less than 1 second. Also depends on **gravitino** mass.

•  $(g-2)_\mu$  anomaly has persisted at Fermilab g-2 experiment [6].

► For a neutralino and degenerate selectron and smuon, obtain

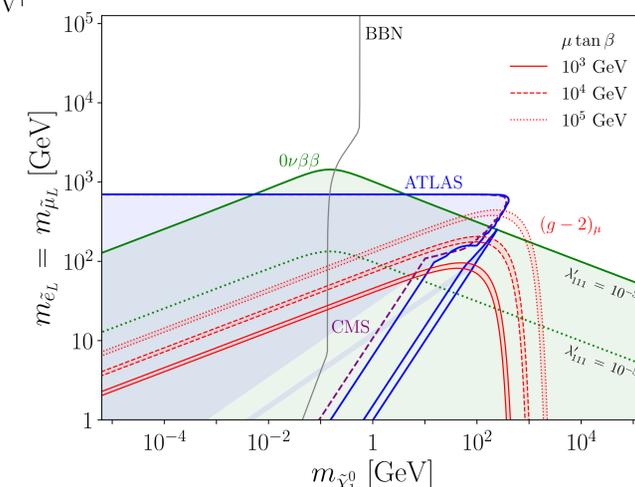
$$\Delta a_\mu^{\tilde{\chi}_1^0} \approx \frac{\alpha_Y m_\mu^2 \mu \tan \beta}{6\pi m_{\tilde{\chi}_1^0}^3} G\left(\frac{m_{\tilde{e}_L}^2}{m_{\tilde{\chi}_1^0}^2}\right),$$

where  $G(x)$  is a loop function [5]. Defines a **favoured region** in  $(m_{\tilde{\chi}_1^0}, m_{\tilde{e}_L})$  plane depending on MSSM parameters  $\mu$  and  $\tan \beta$ .



**Left**: Excluded regions in  $(m_{\tilde{\chi}_1^0}, \lambda'_{111})$  plane from BBN,  $0\nu\beta\beta$  decay (shown for different selectron masses  $m_{\tilde{e}_L}$ ),  $V_{ud}$  and  $R_\pi$ .

**Right**: Excluded regions in  $(m_{\tilde{\chi}_1^0}, m_{\tilde{e}_L})$  plane from BBN,  $0\nu\beta\beta$  decay (shown for different  $\lambda'_{111}$  values) and colliders compared to the  $(g-2)_\mu$  favoured region.



## 5 Conclusions

To conclude, we have performed a detailed study of the **RPV** contribution to  $0\nu\beta\beta$  decay with an arbitrary mass lightest **neutralino**  $\tilde{\chi}_1^0$ . Previous studies have only considered neutralinos more massive than the average momentum exchange of  $0\nu\beta\beta$  decay,  $p_F \sim 100$  MeV. We made use of an **interpolating** function to encapsulate the change from short-range ( $m_{\tilde{\chi}_1^0} \gg p_F$ ) to long-range ( $m_{\tilde{\chi}_1^0} \ll p_F$ ) behaviour depending on the  $\tilde{\chi}_1^0$  mass.

In deriving constraints on the RPV coupling  $\lambda'_{111}$  as a function of  $m_{\tilde{\chi}_1^0}$ , we performed a scan over the relevant SUSY masses allowed by collider constraints. We compared the  $0\nu\beta\beta$  decay constraints to naive limits from BBN and the region favoured by the  $(g-2)_\mu$  anomaly.

## References

- [1] M. Hirsch, H. V. Klapdor-Kleingrothaus and S. G. Kovalenko, Phys. Rev. D, 53:1329-1348 (1996).
- [2] H. K. Dreiner, S. Heinemeyer, O. Kittel, U. Langenfeld, A. M. Weber and G. Weiglein, Eur. Phys. J. C, 62:547-572 (2009).
- [3] M. González, M. Hirsch and S. Kovalenko, Phys. Rev. D, 93(1):013017 (2016).
- [4] G. Aad et al., Eur. Phys. J. C, 80(2):123 (2020).
- [5] T. Moroi, Phys. Rev. D 56, 4424 (1997).
- [6] B. Abi et al., Phys. Rev. Lett., 126(14):141801 (2021).