

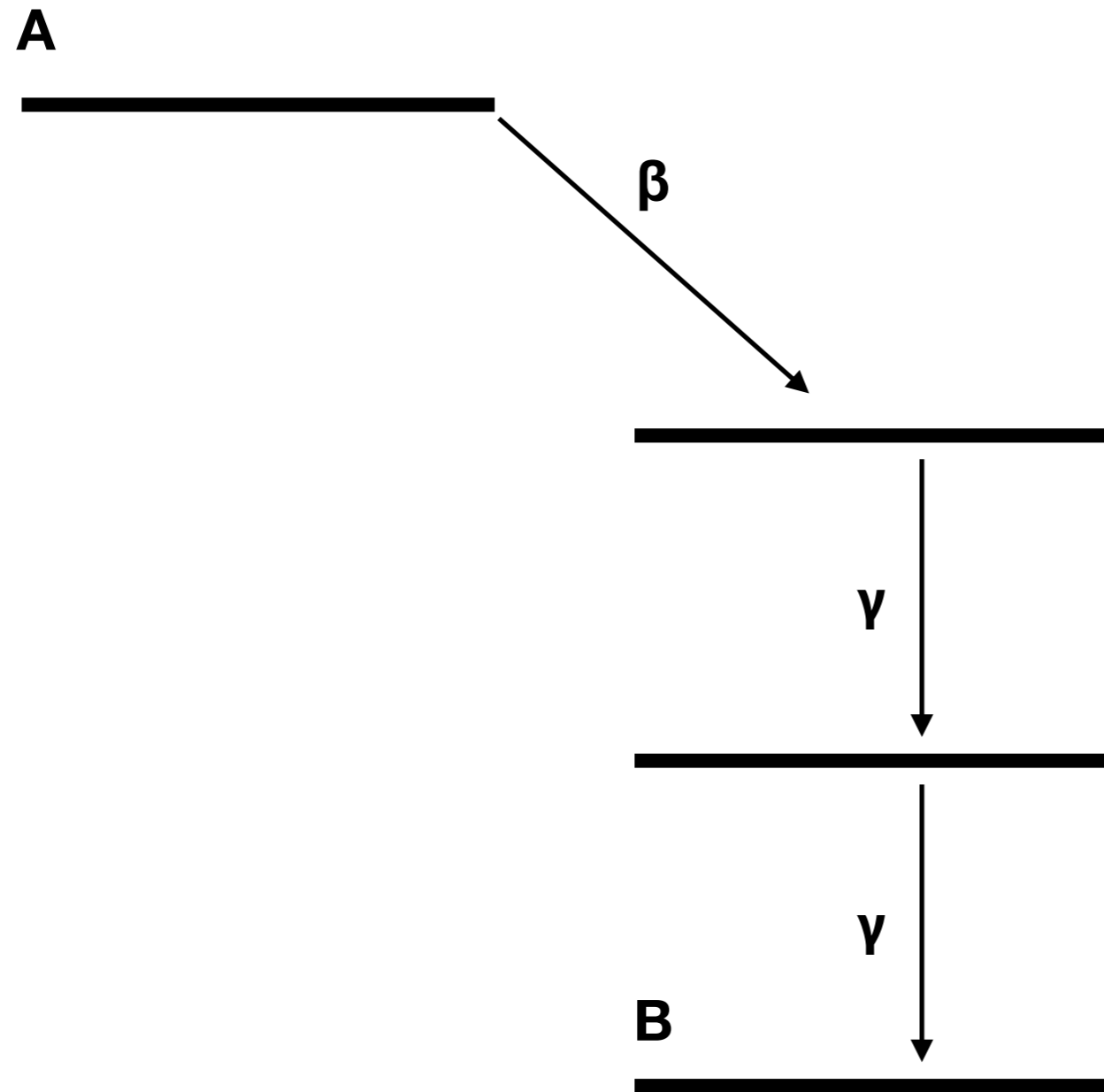
Gamma Nuclear Decays Hiding from Investigators Experiment (GANDHI)

Surjeet Rajendran
with

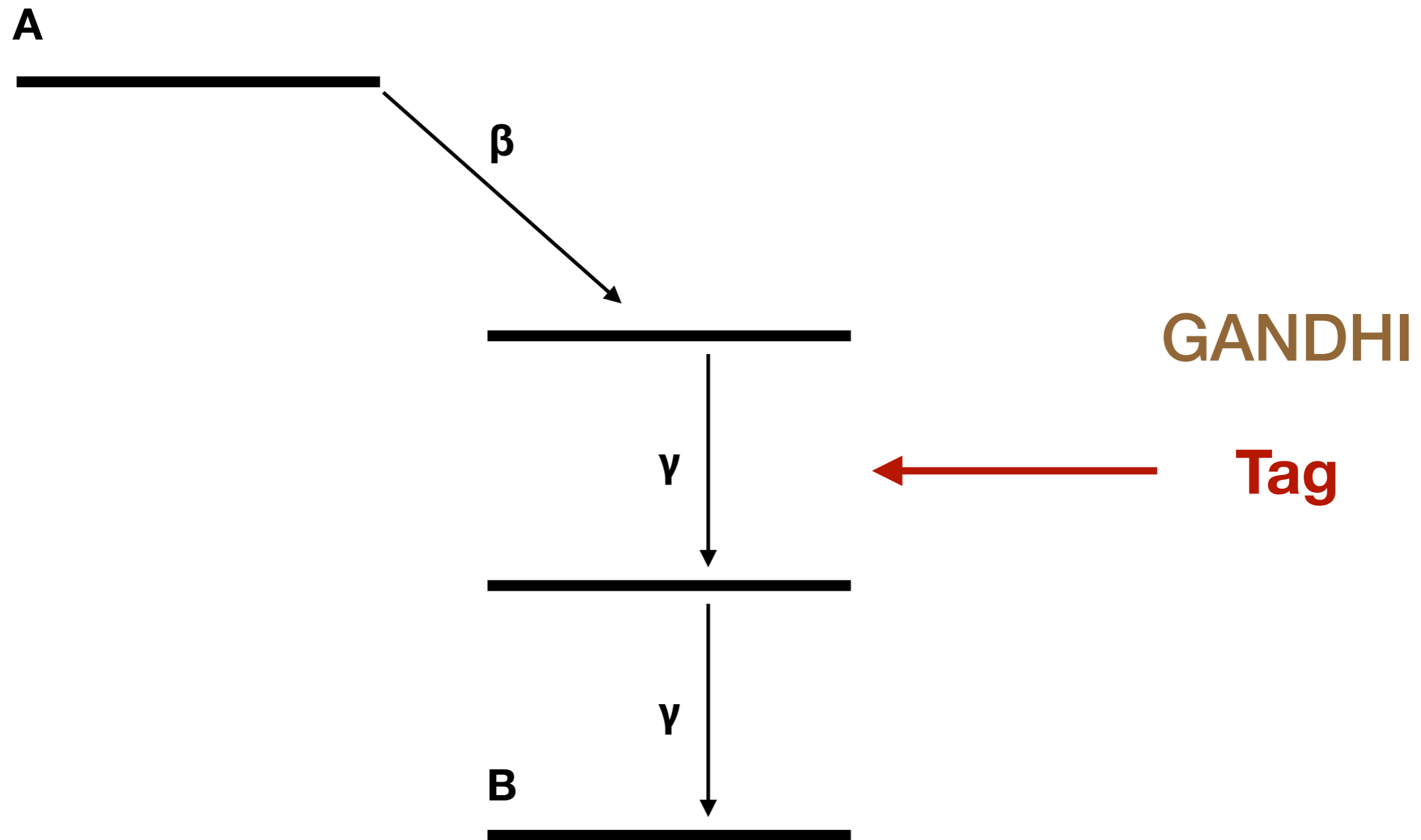
Giovanni Benato, Alexey Drobizhev and Hari Ramani

Proof of Concept:
Rupak Mahapatra (TAMU)

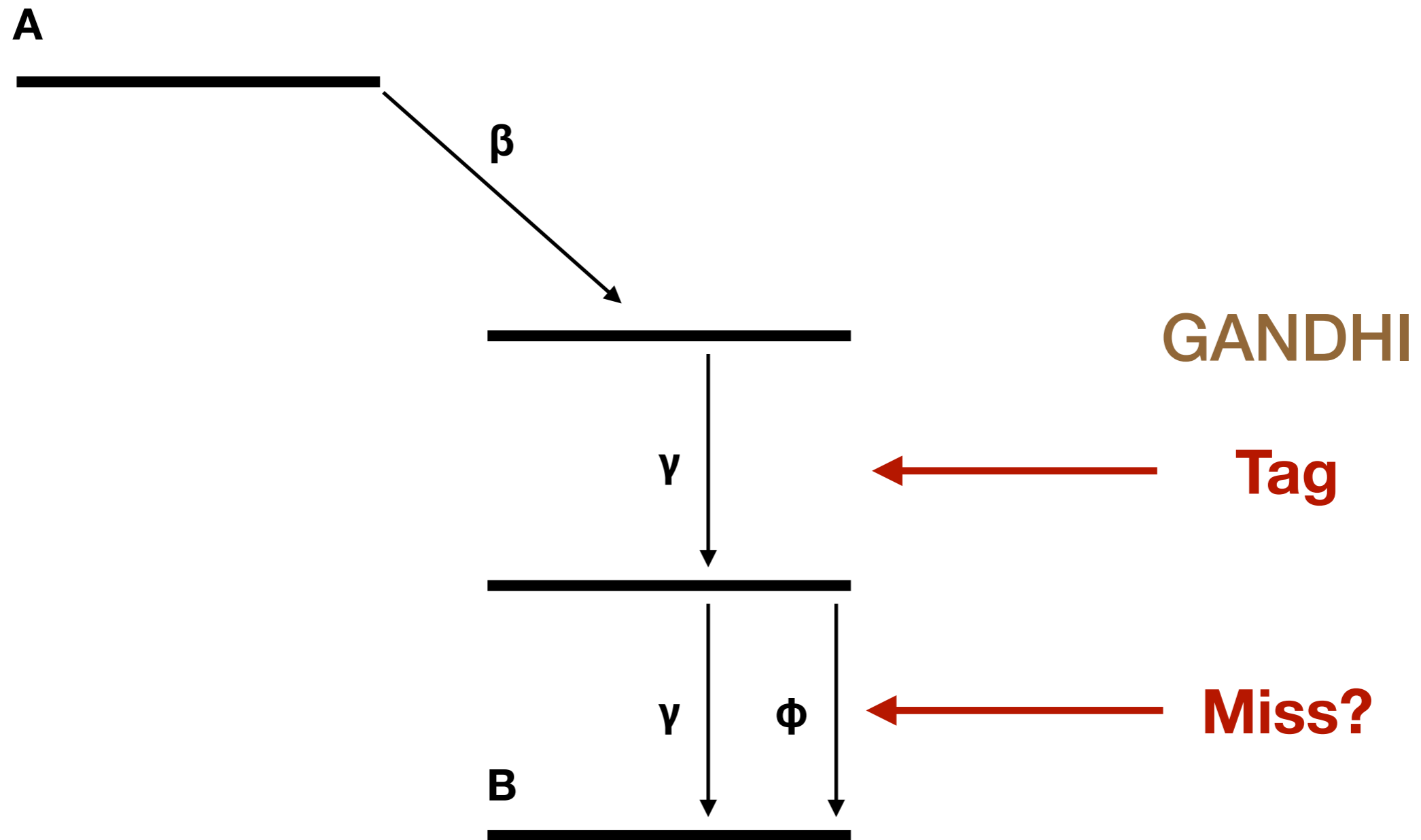
Missing Energy in Gamma Cascades



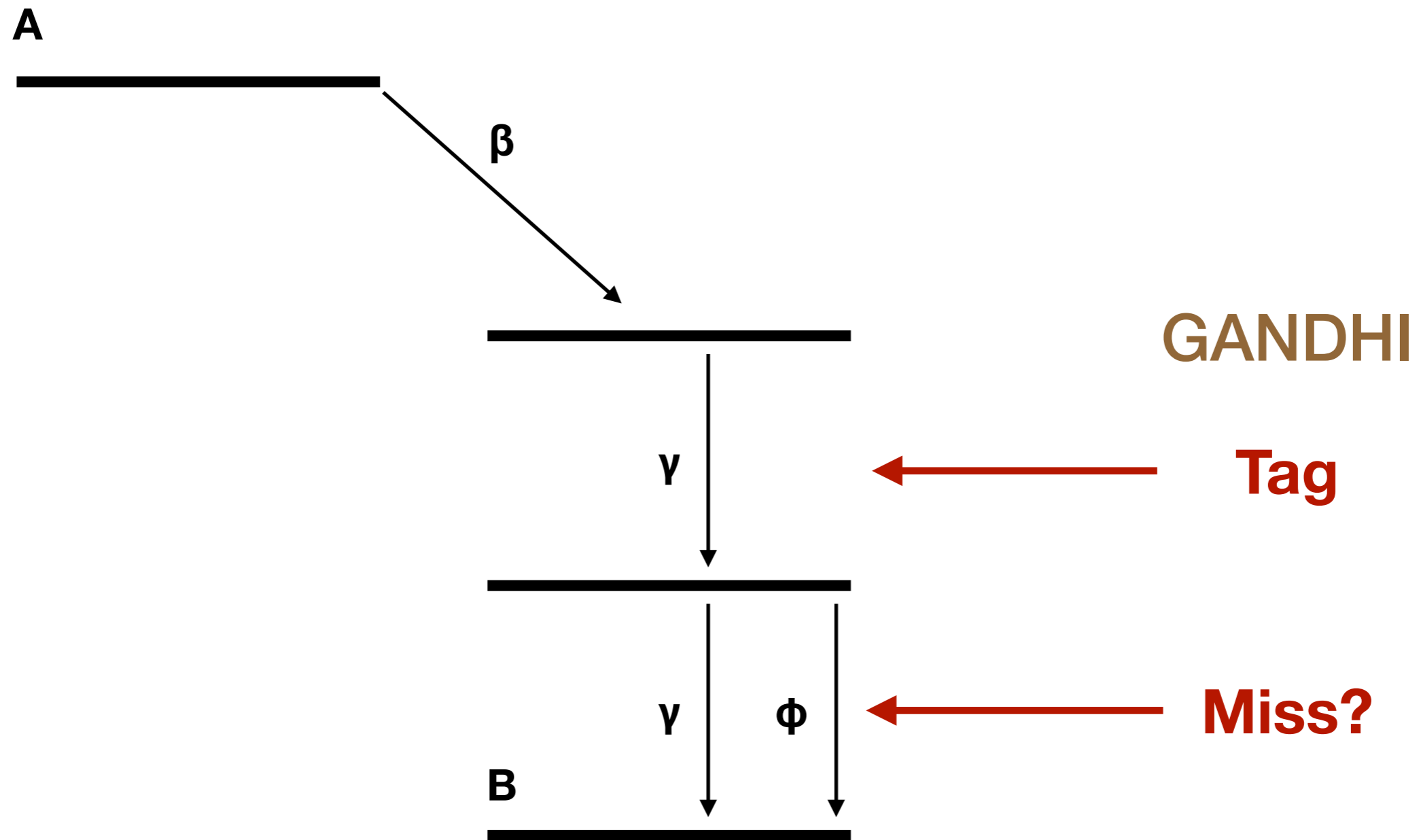
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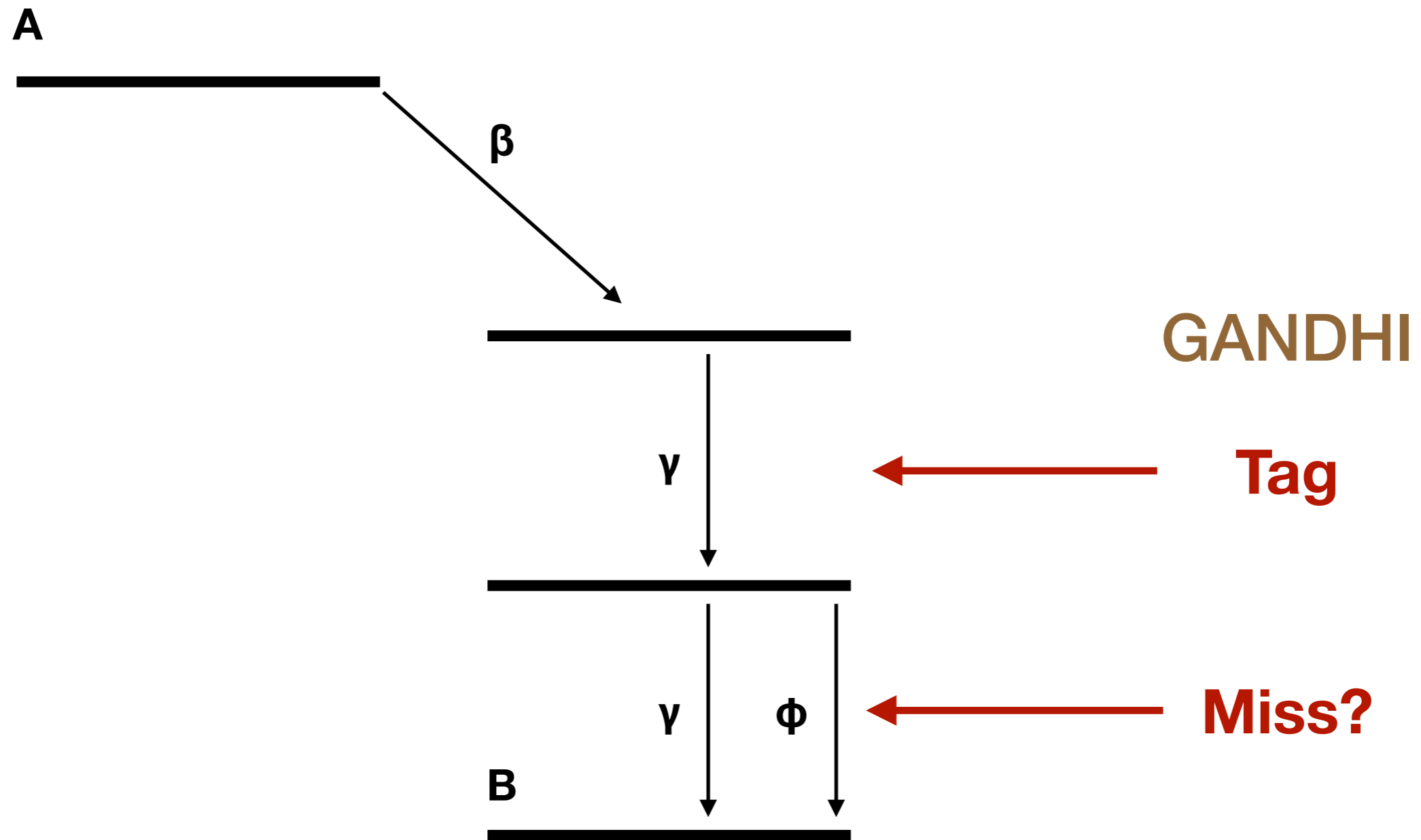


Missing Energy in Gamma Cascades



Aim: Single Event for Discovery

Missing Energy in Gamma Cascades



Aim: Single Event for Discovery

How well can we do?

Baryonically coupled ϕ , mass $< \sim$ MeV

Outline

1. Nuclei
2. Setup
3. Theory/Reach
4. Conclusions

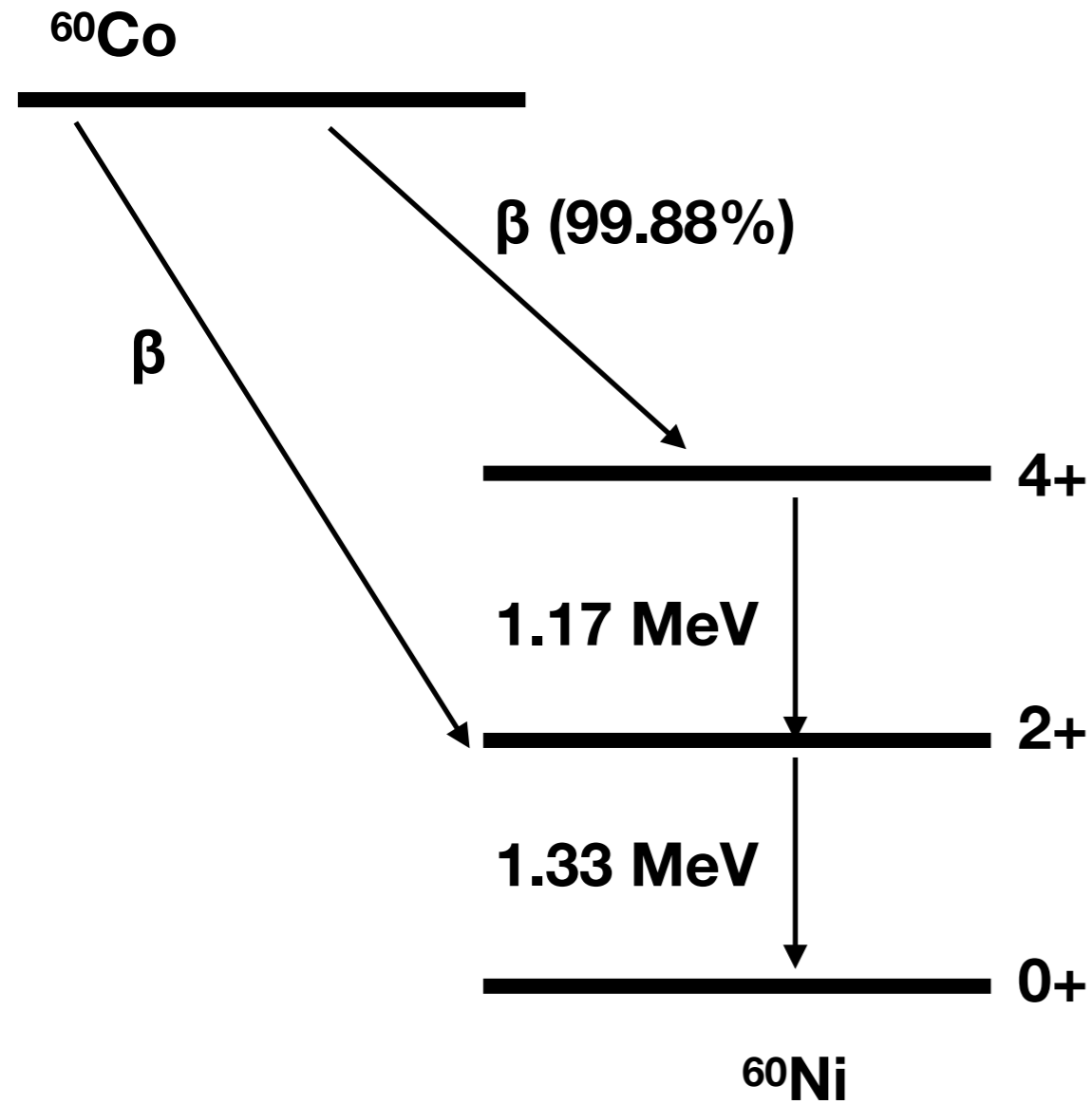
Nuclei

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Lifetime, Cascade Efficiency, Availability

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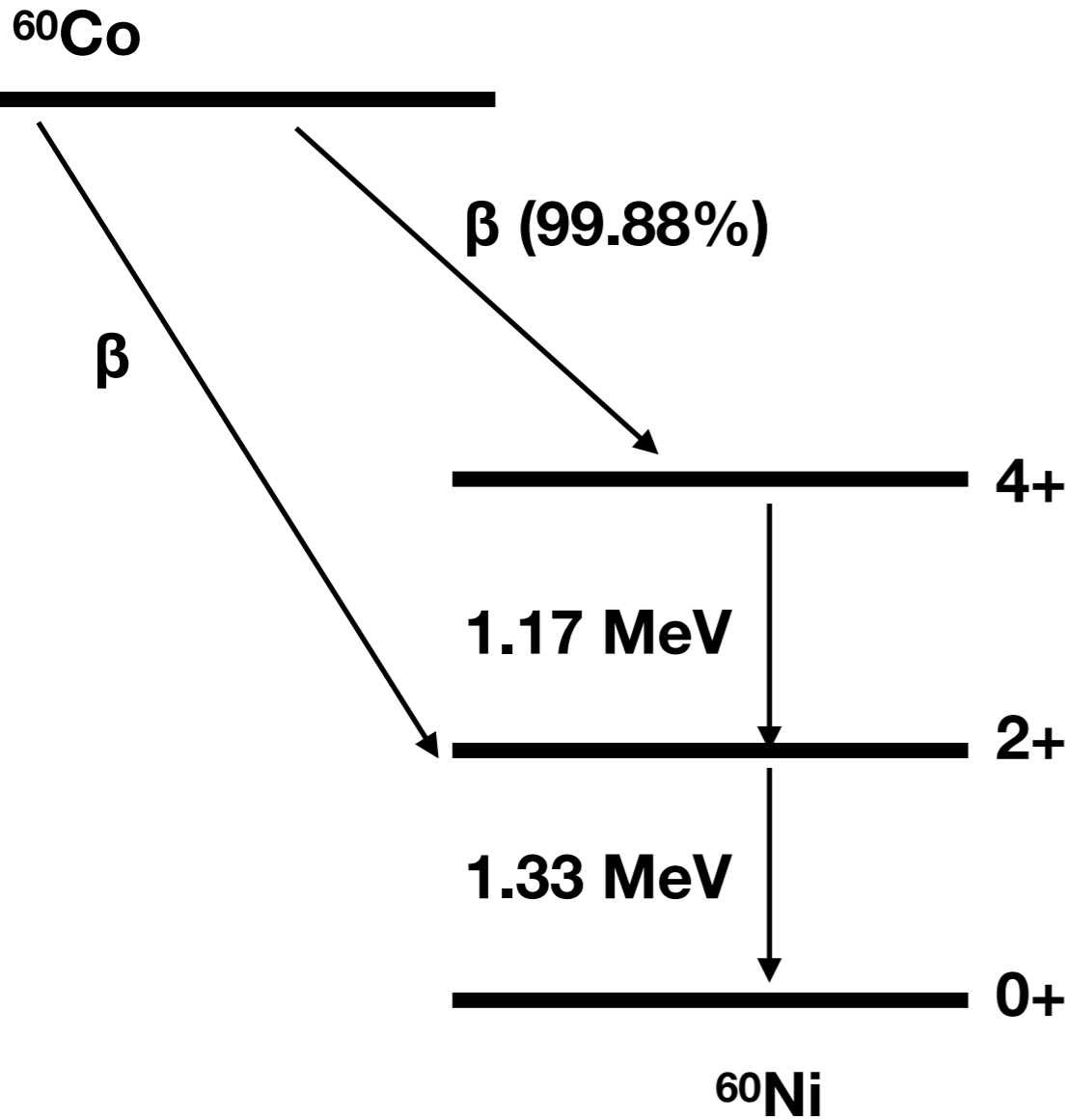


$t_{1/2} \sim 5$ years

Similar energy Gammas

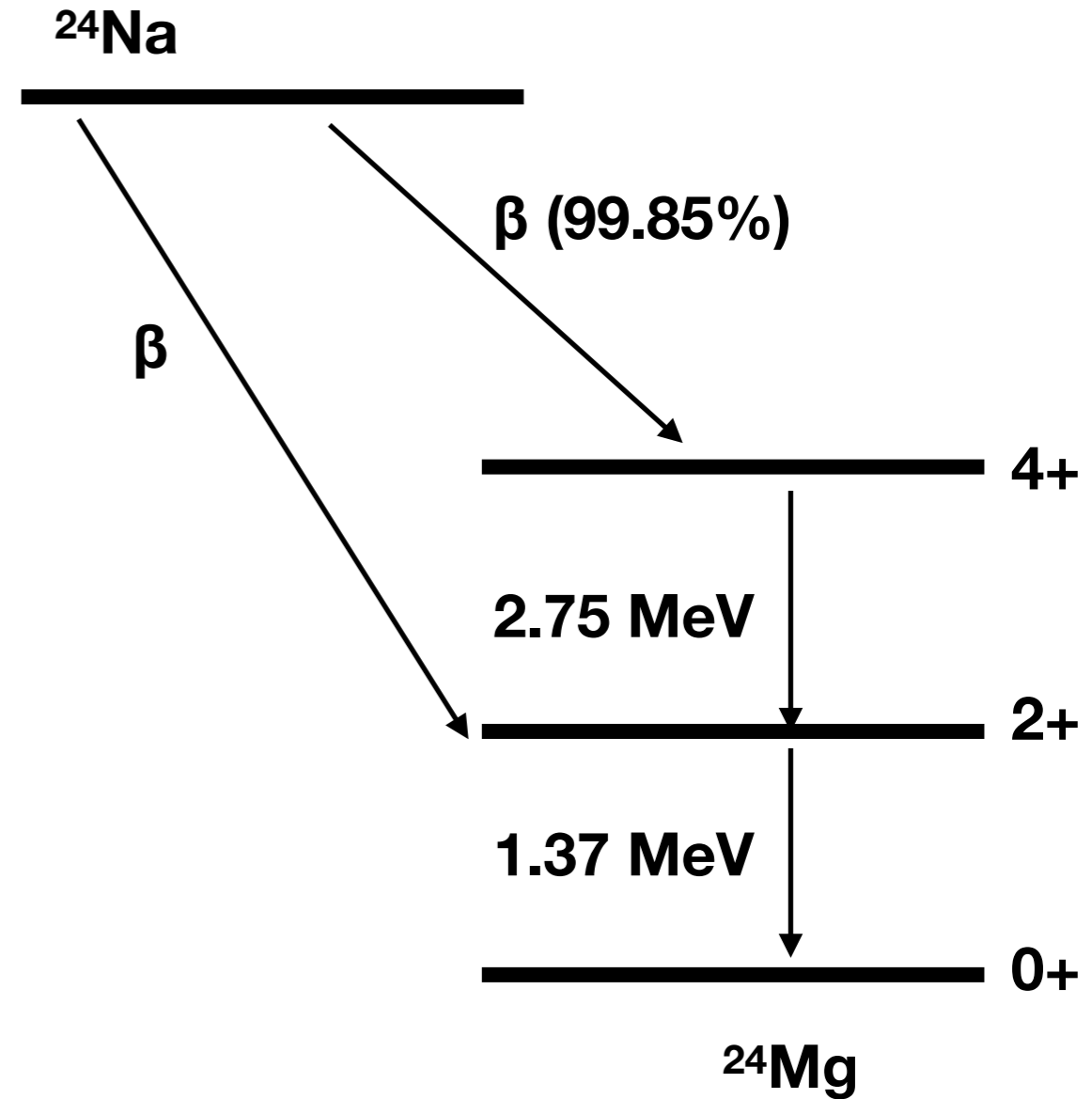
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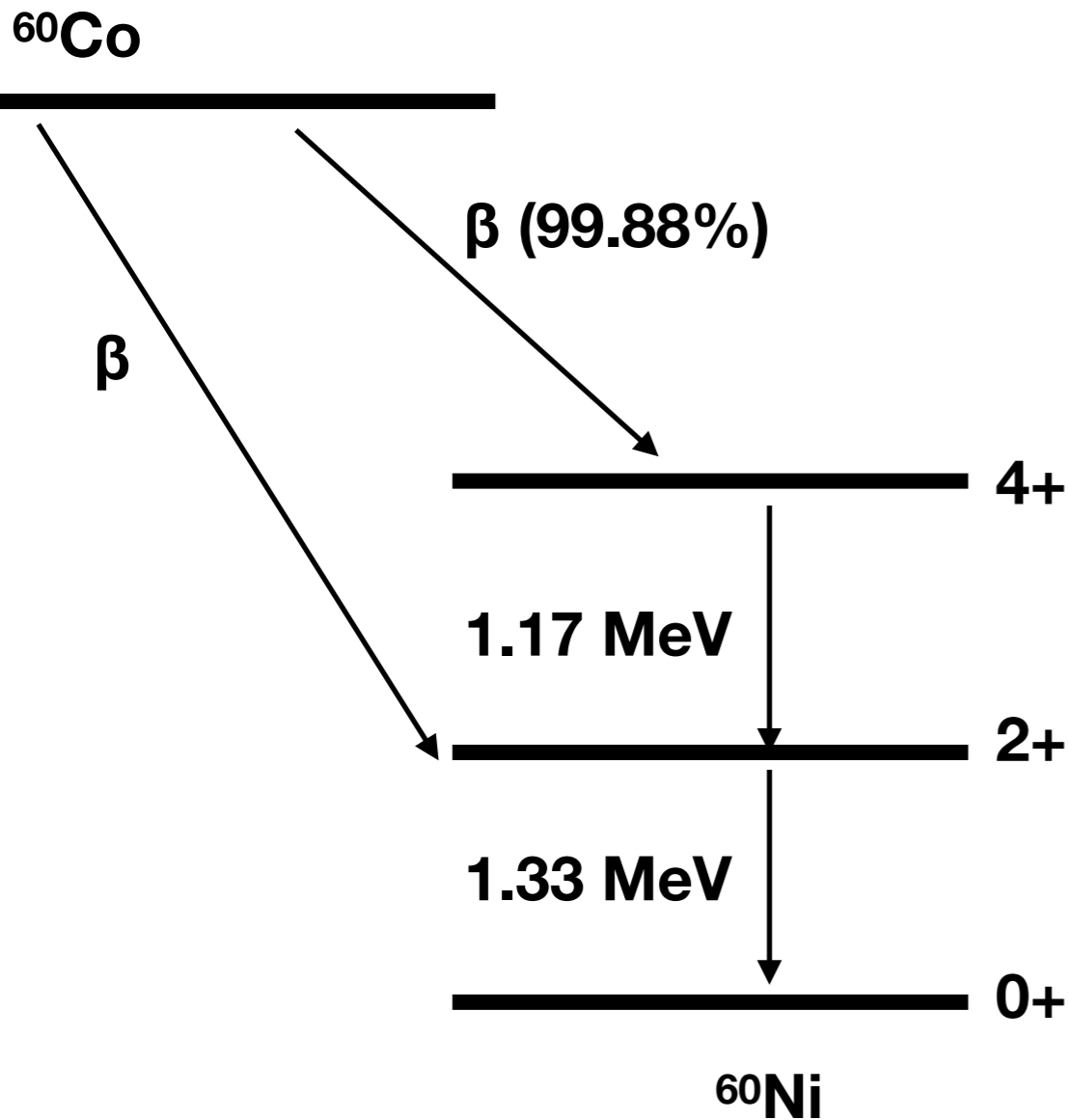


$t_{1/2} \sim 15$ hr

Medical Isotope

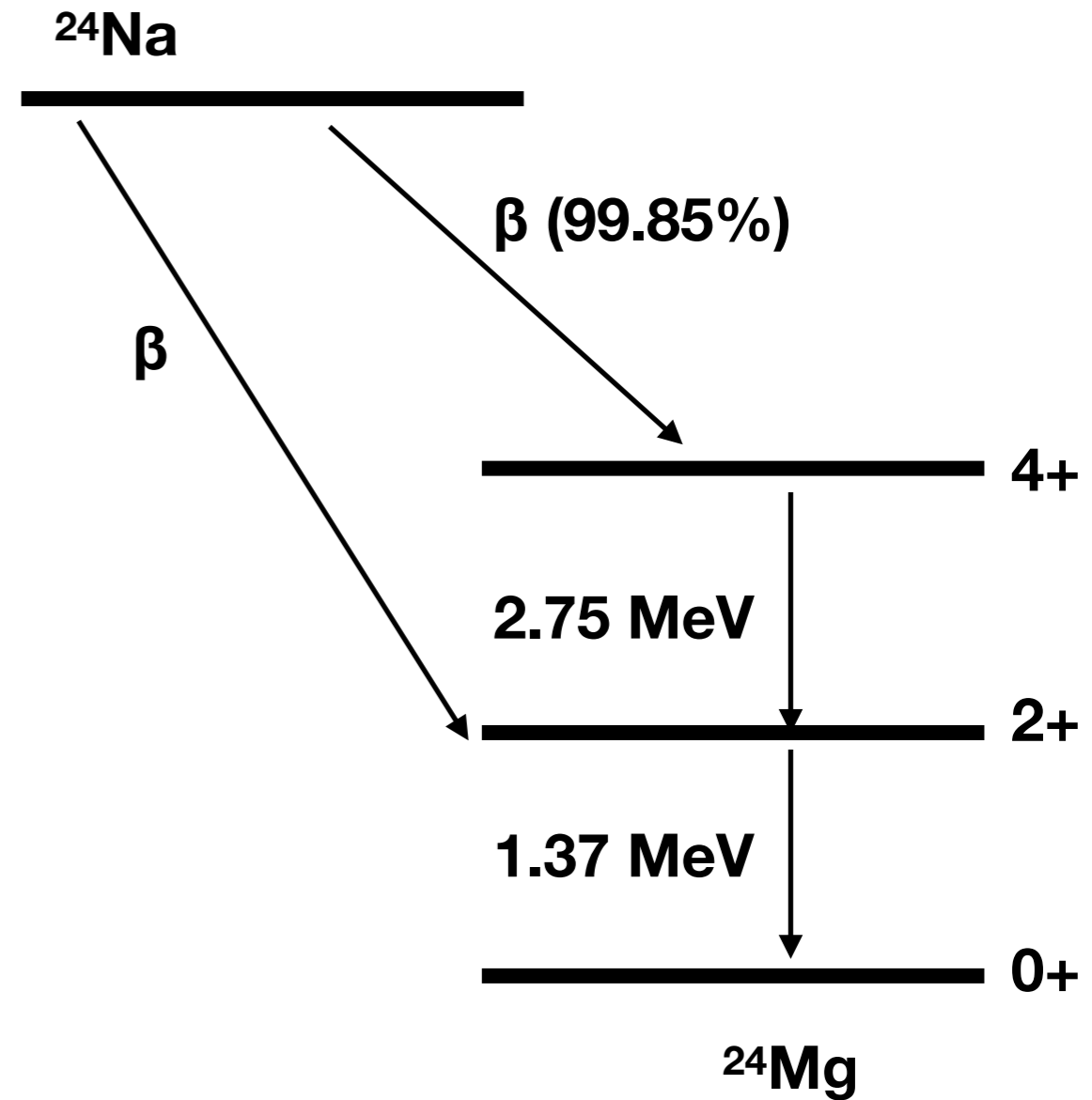
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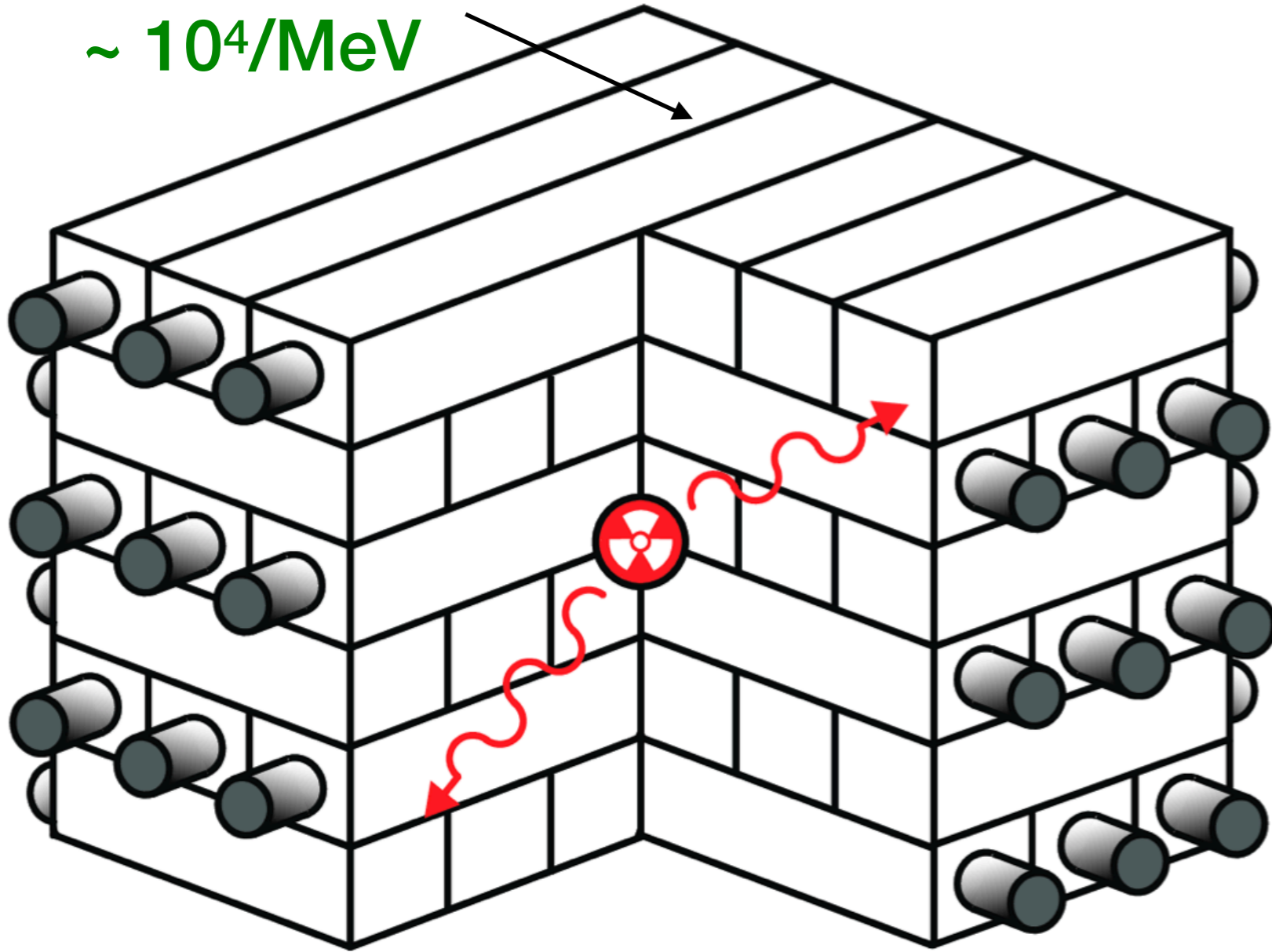
Medical Isotope

Parity of States \rightarrow scalars and vectors

Setup

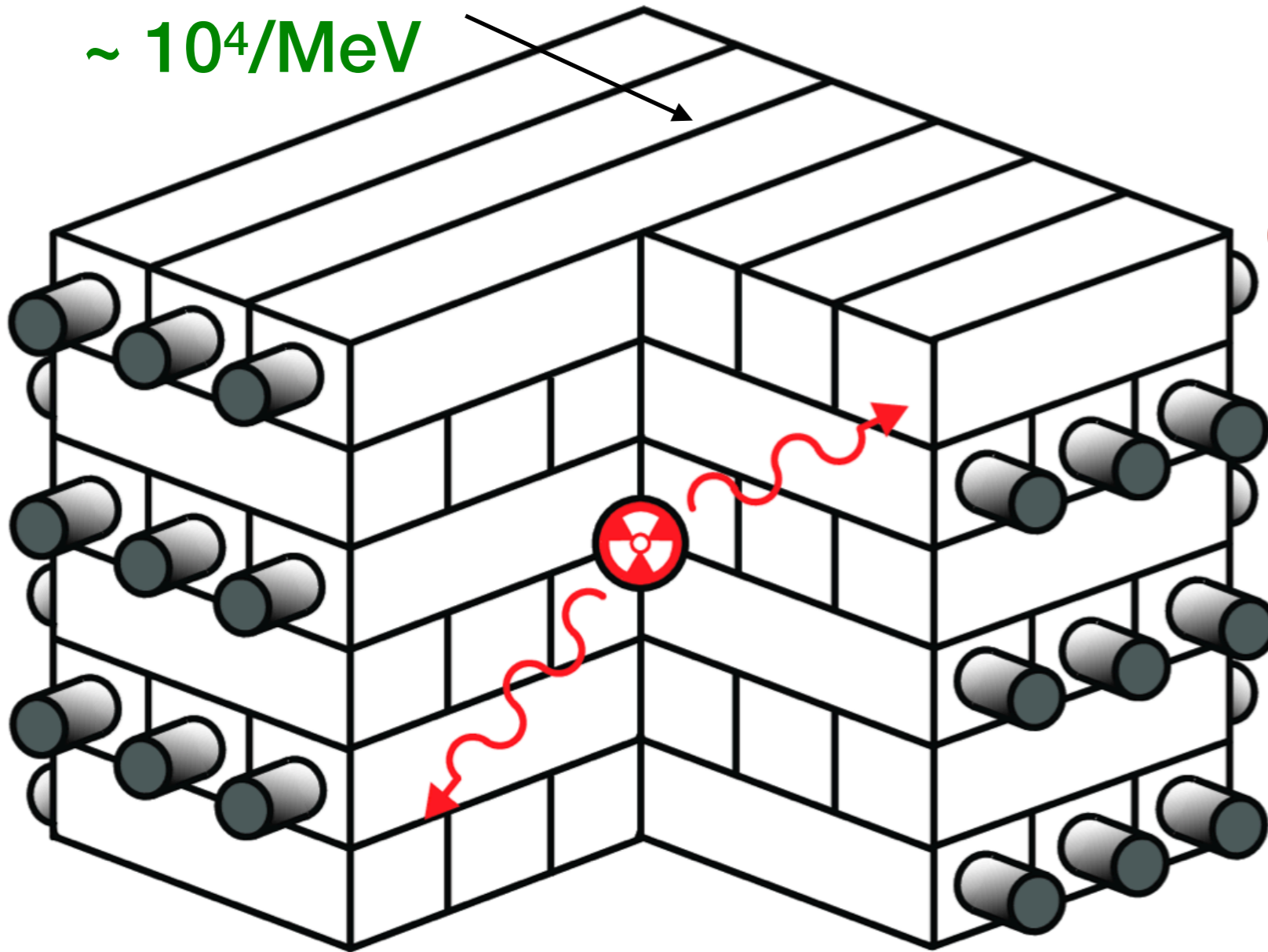
Setup

Scintillator
 $\sim 10^4/\text{MeV}$



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Initial Goal: 10^{-11}
Eventual Goal: 10^{-14}

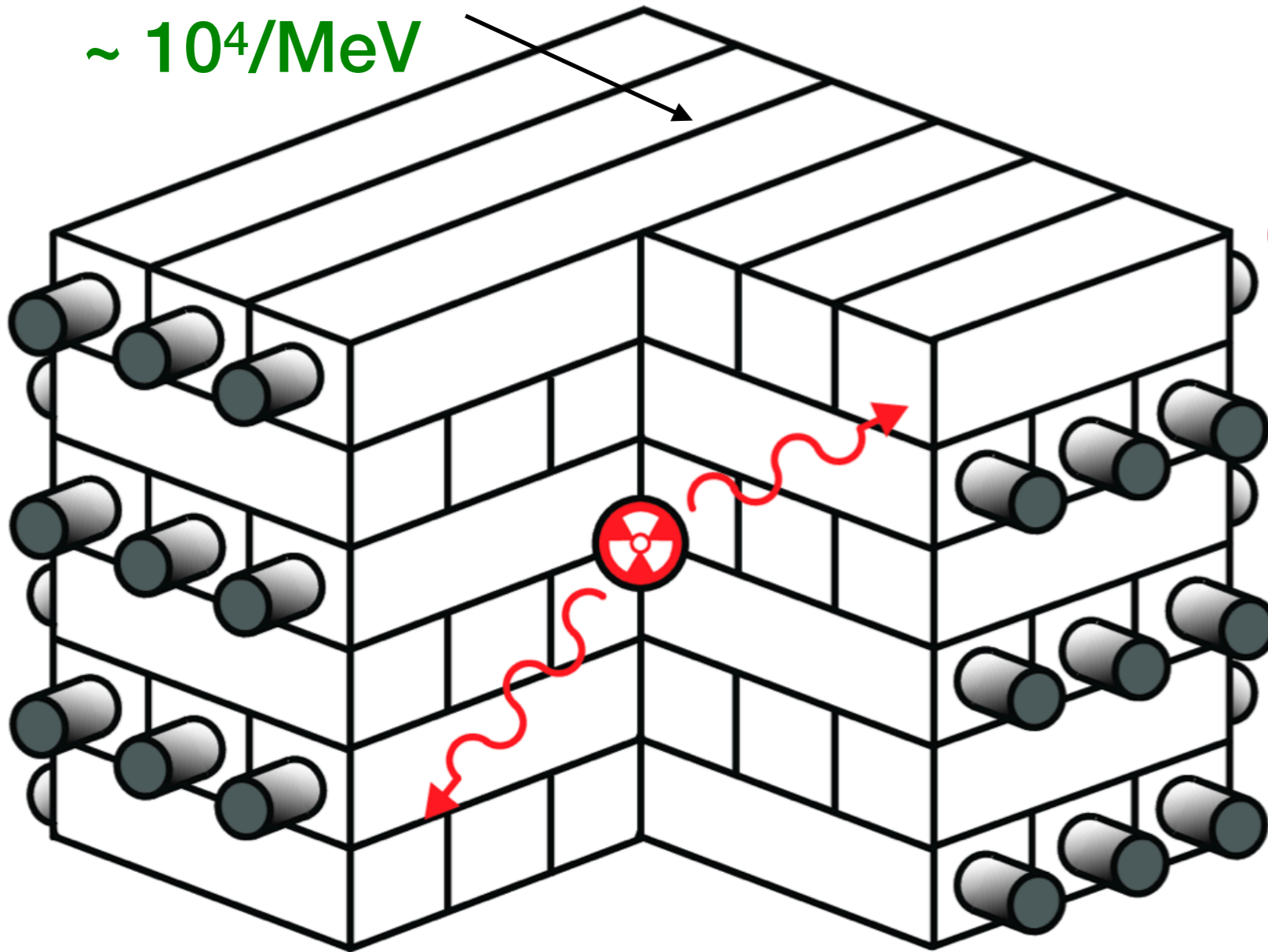
Observe Individual Event
No pile up

High Event Rate
Fast Scintillator

Plastics or Crystals
 $\sim \text{ns}$ response

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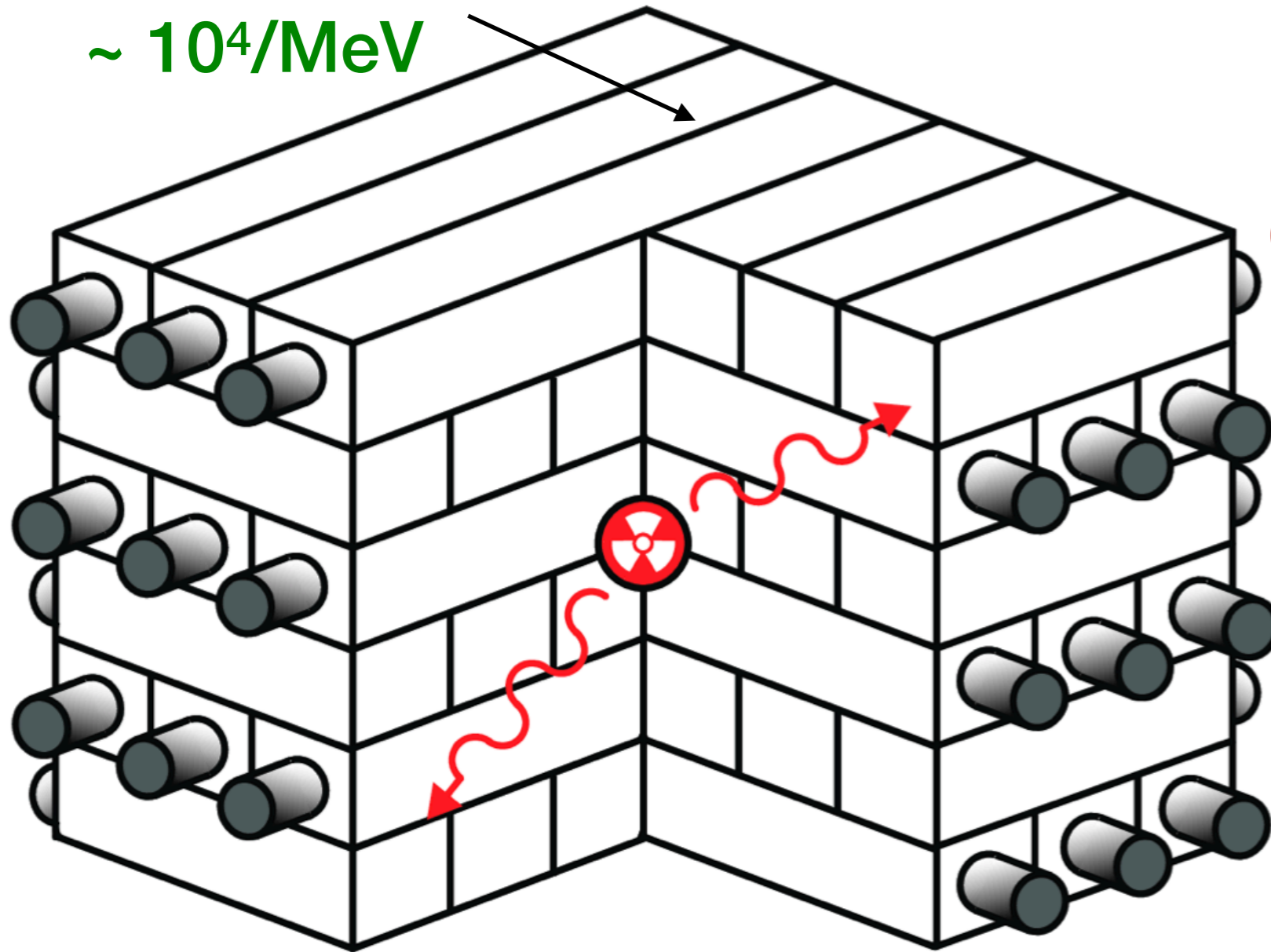
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~ 30 radiation lengths

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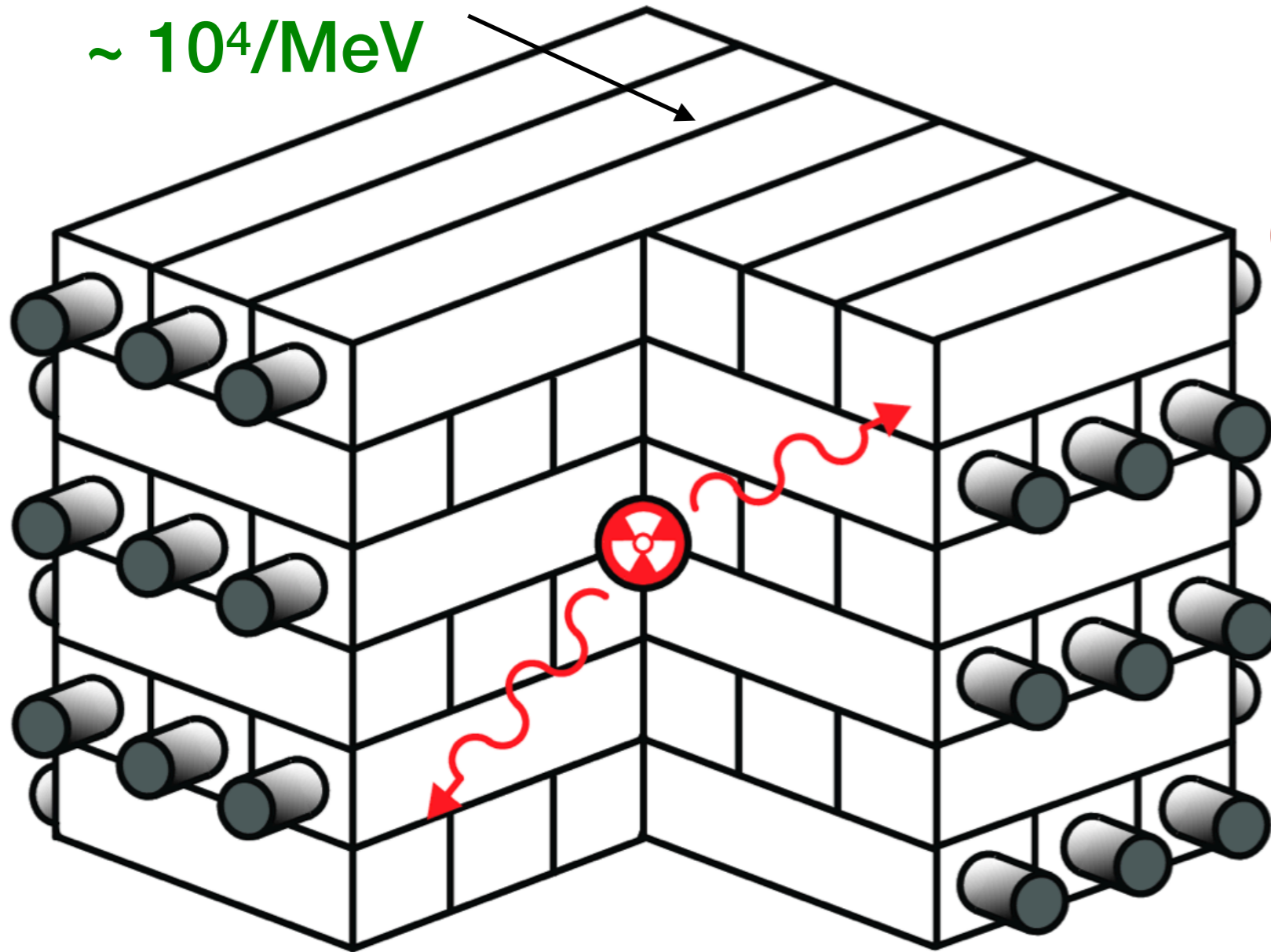
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Plastics: ~ 10 m, cheap, make large modules

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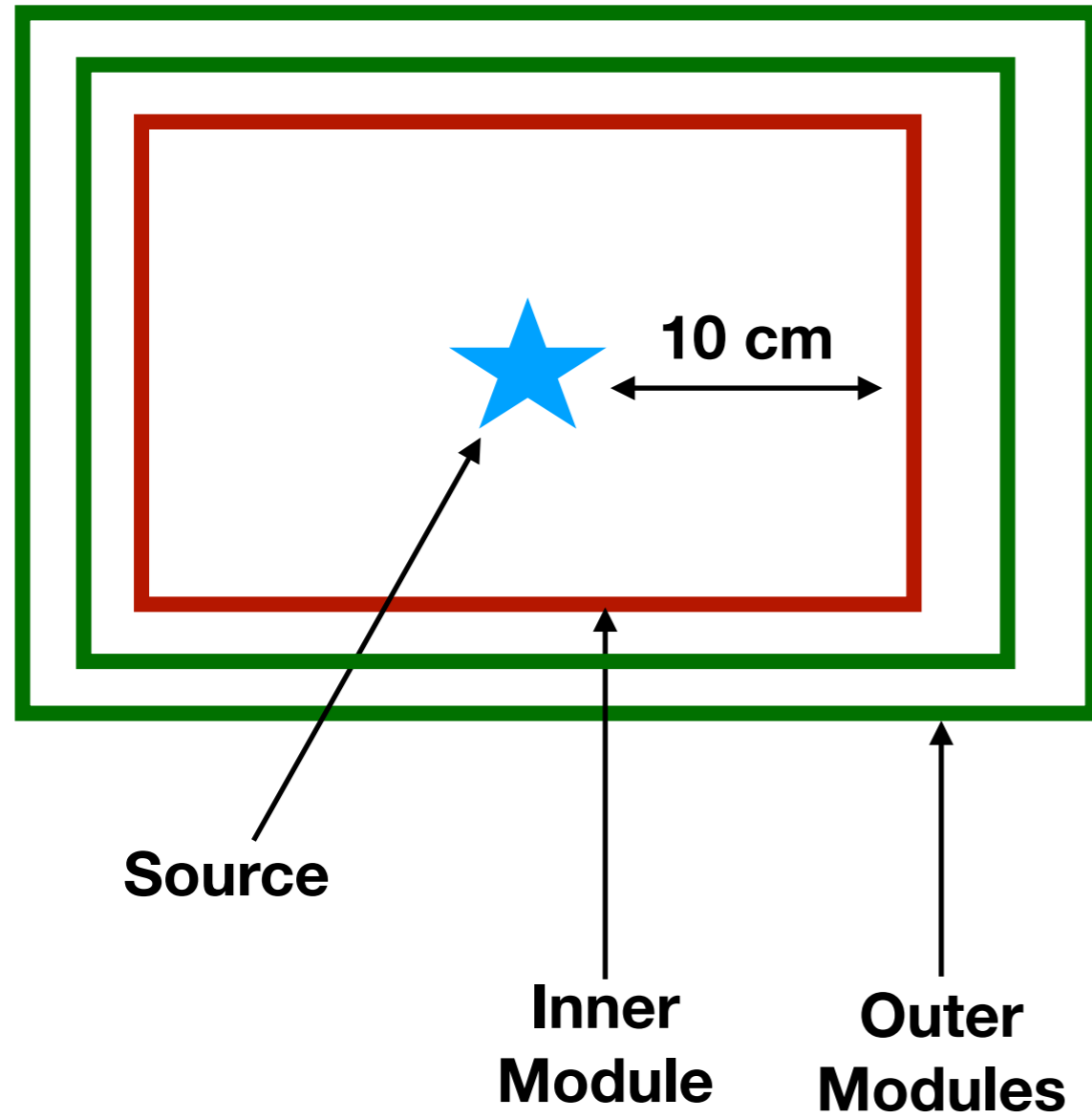
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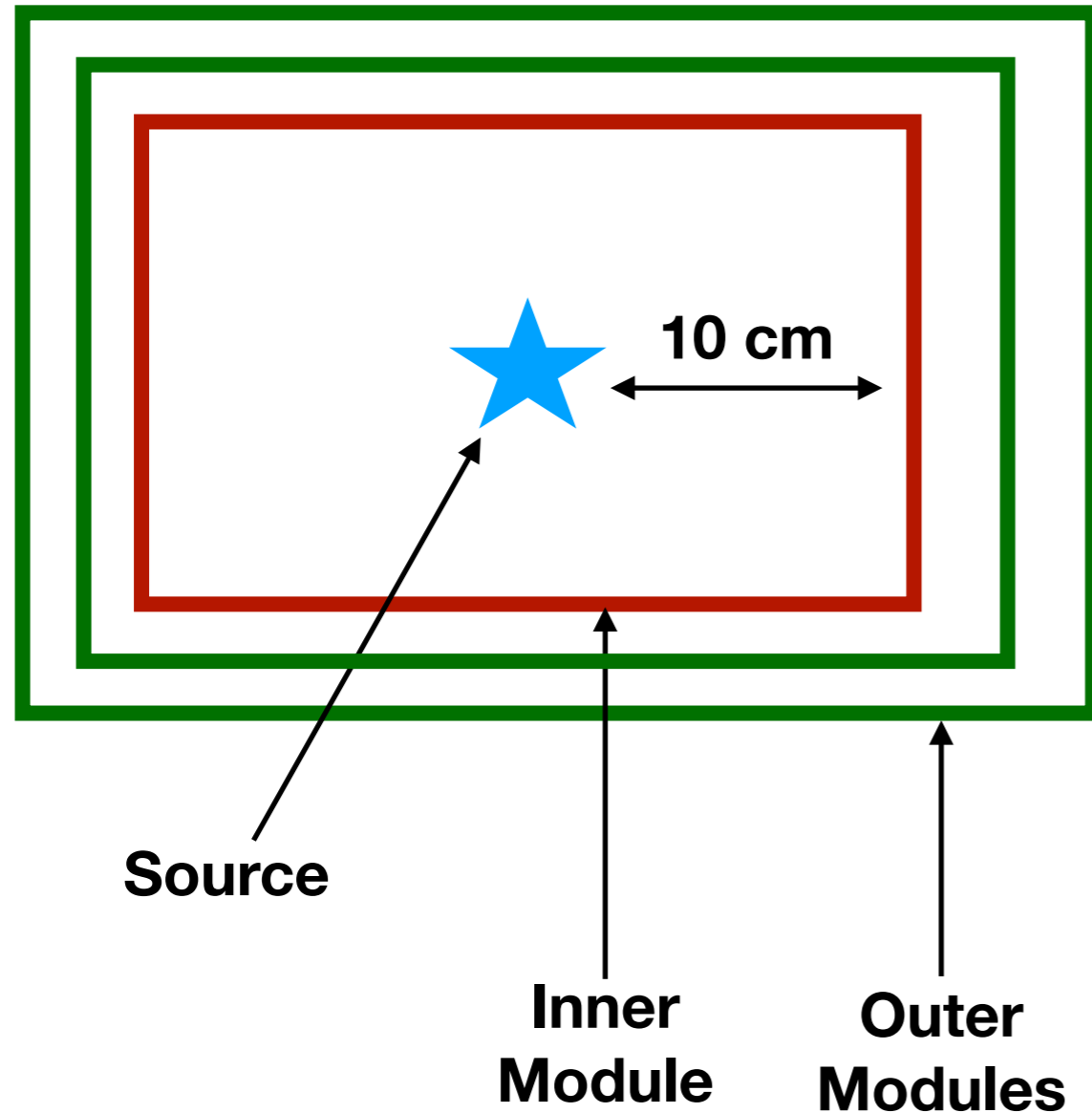
Plastics: ~ 10 m, cheap, make large modules

Crystals: ~ 2 m, harder to grow. CMS E-cal

Protocol



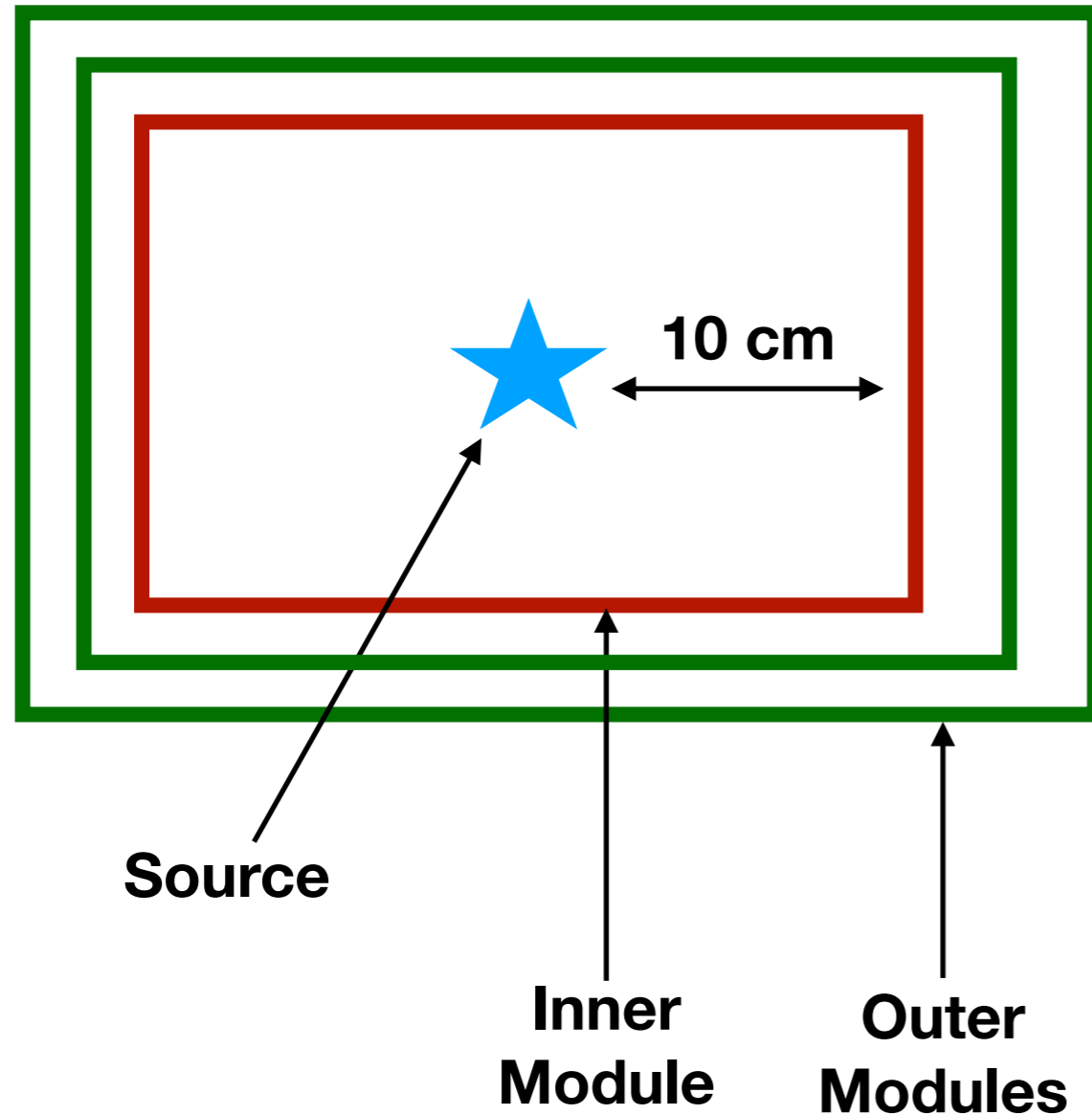
Protocol



Signal

1. Observe β activity consistent with initial decay
2. Within \sim ns, observe 1st γ in inner module
3. In that \sim ns, no other energy in detector

Protocol



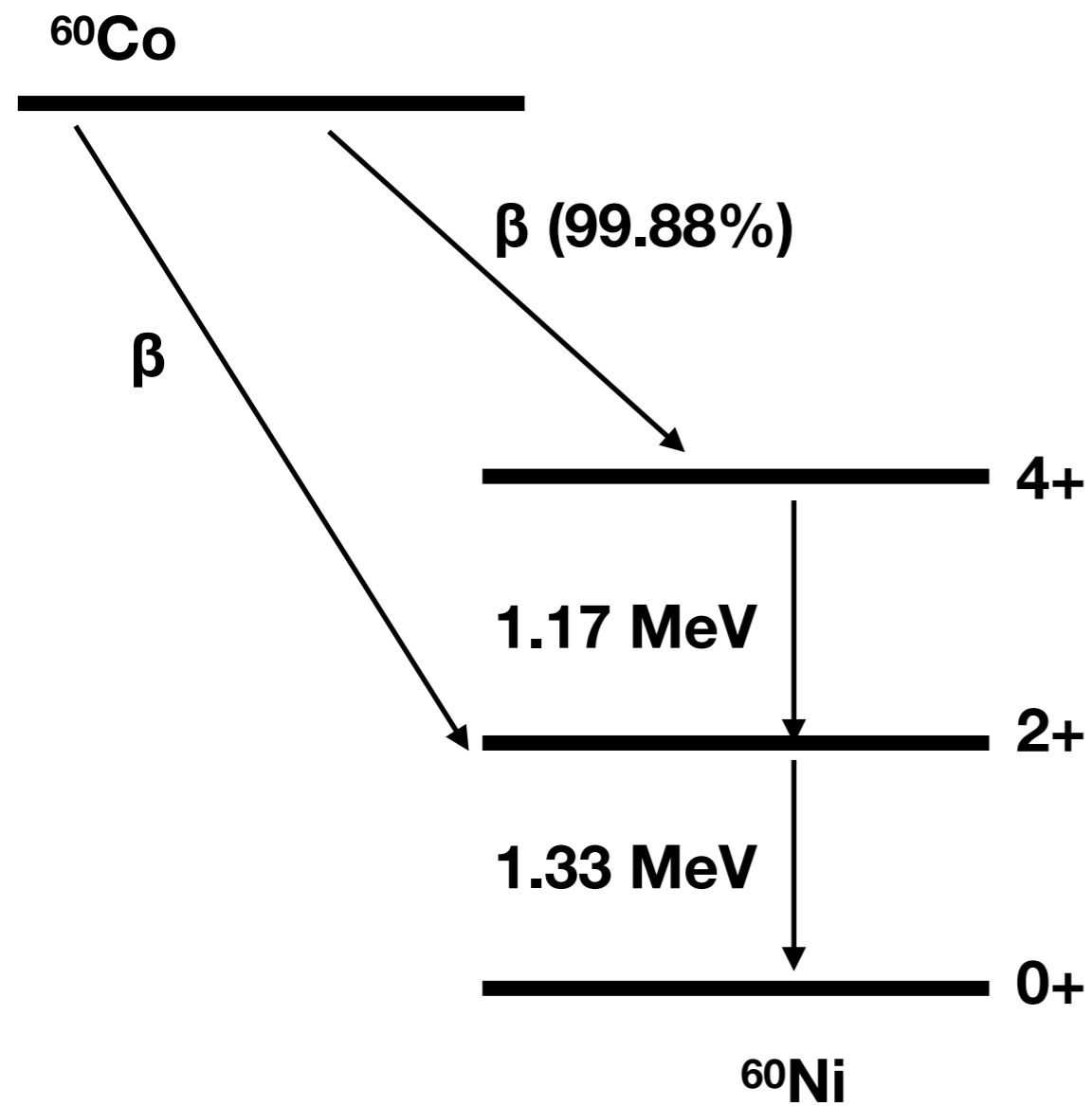
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Backgrounds?

Intrinsic Background for ^{60}Co

Can 2nd γ fake 1st?



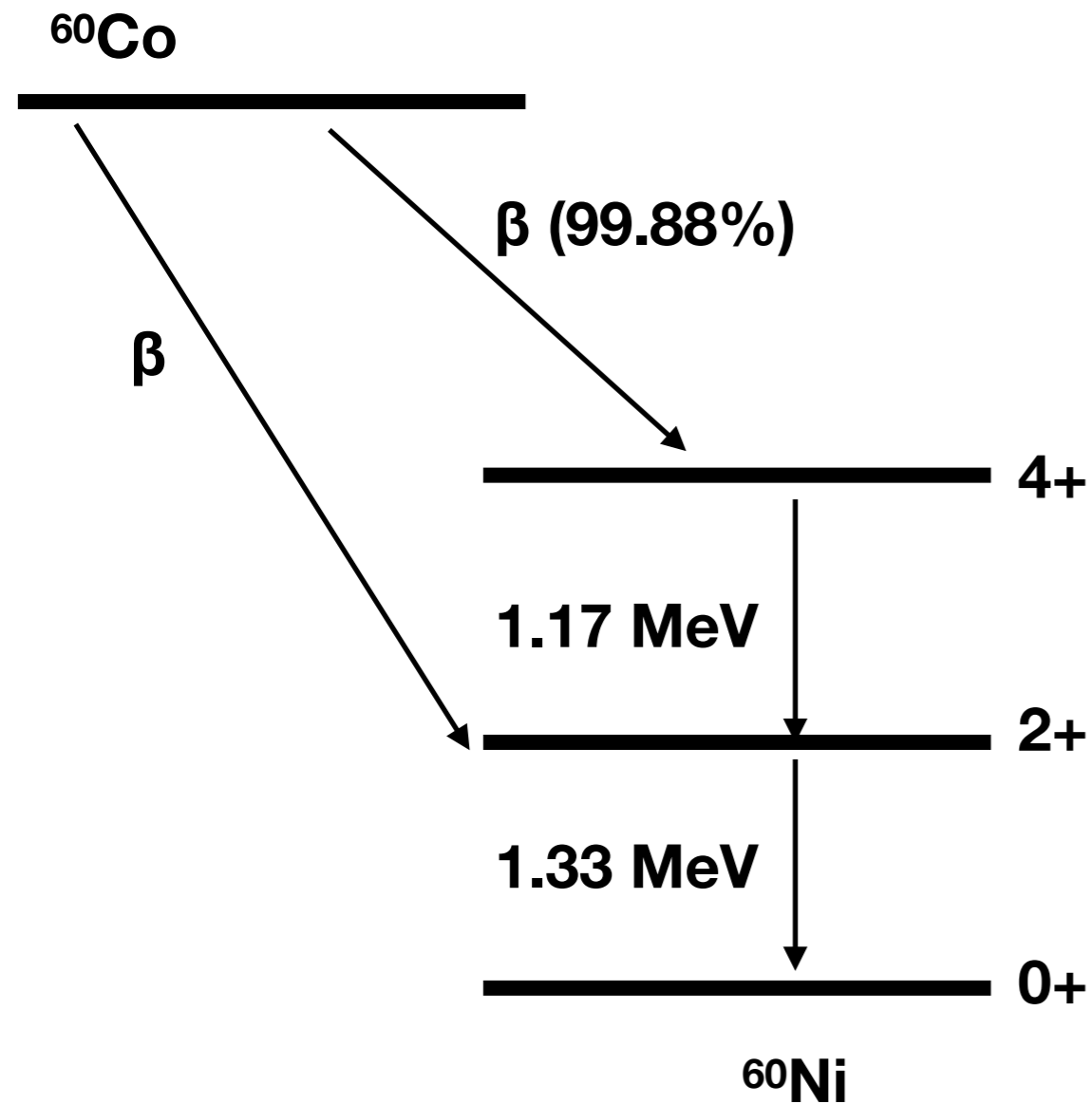
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Energy Resolution

Produce both. Confuse 1.33 MeV γ for 1.17 MeV γ

Requiring single γ only eliminates background



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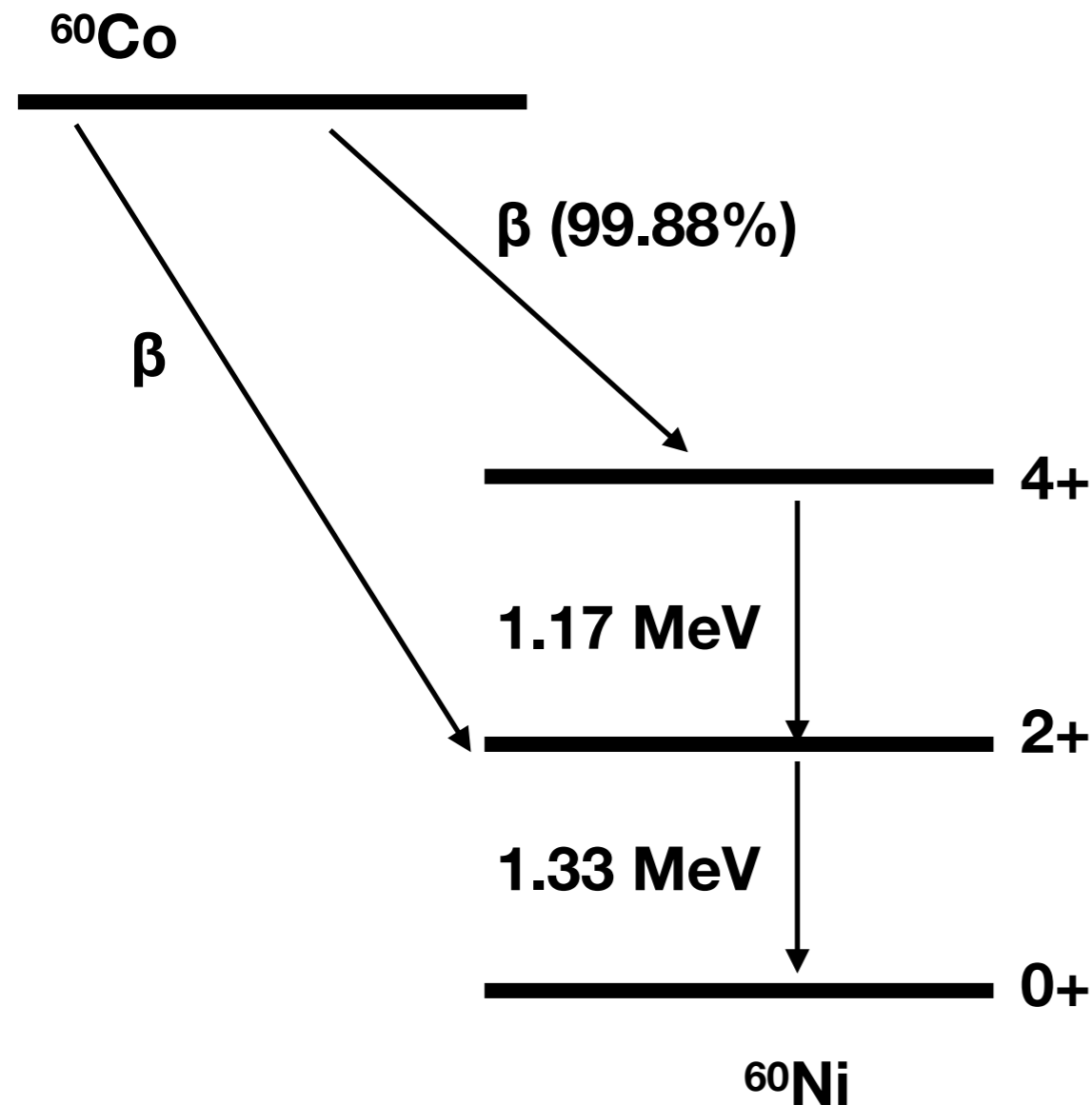
Requiring single γ only eliminates background

Soft β to 2+ and Soft Compton γ

Populate 2+ @ 10^{-3} .

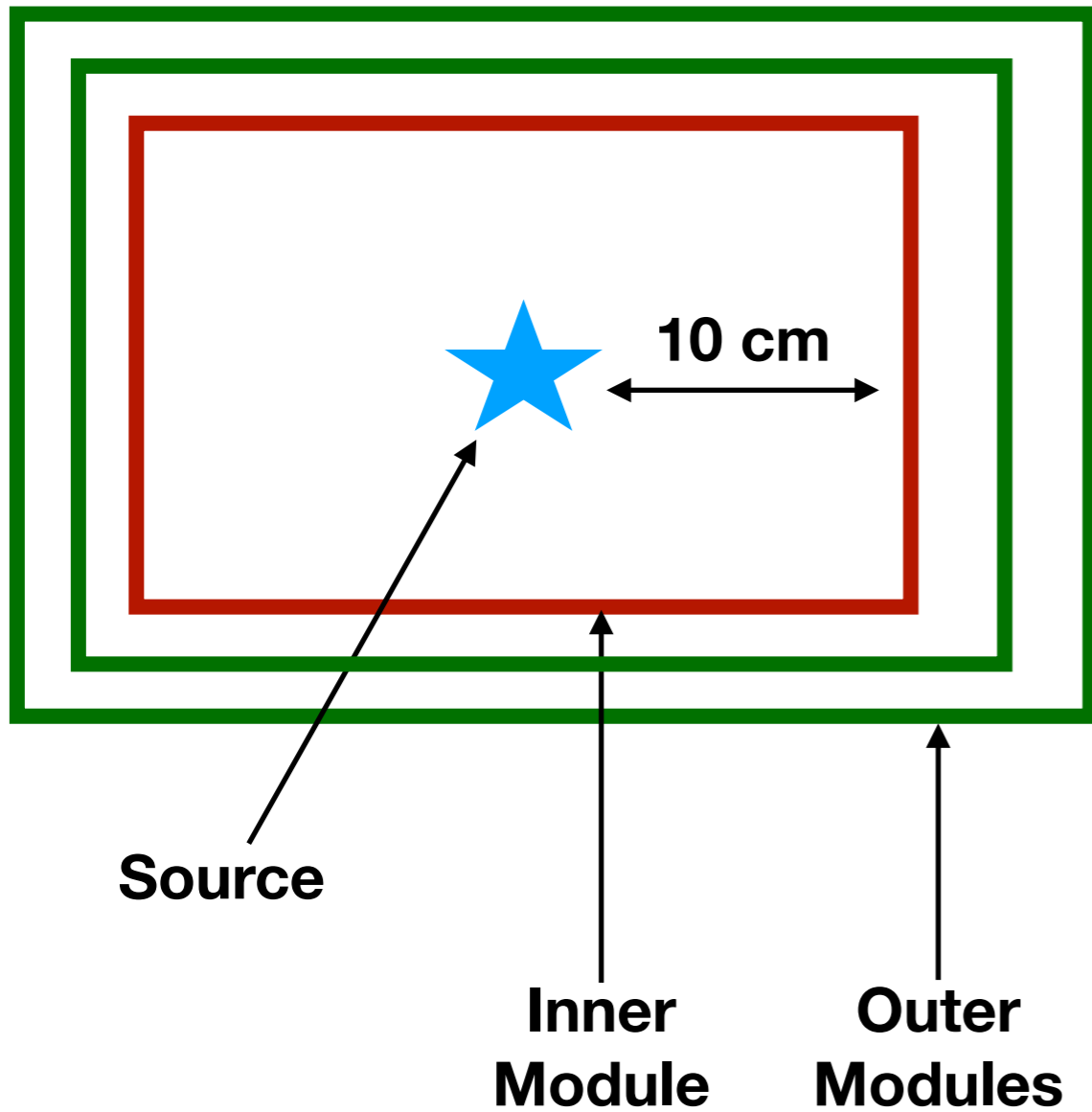
Soft β + Soft 1.33 MeV = β to 4+ and 1.17 γ ?

Soft β + Energy Resolution of 1.33 MeV?



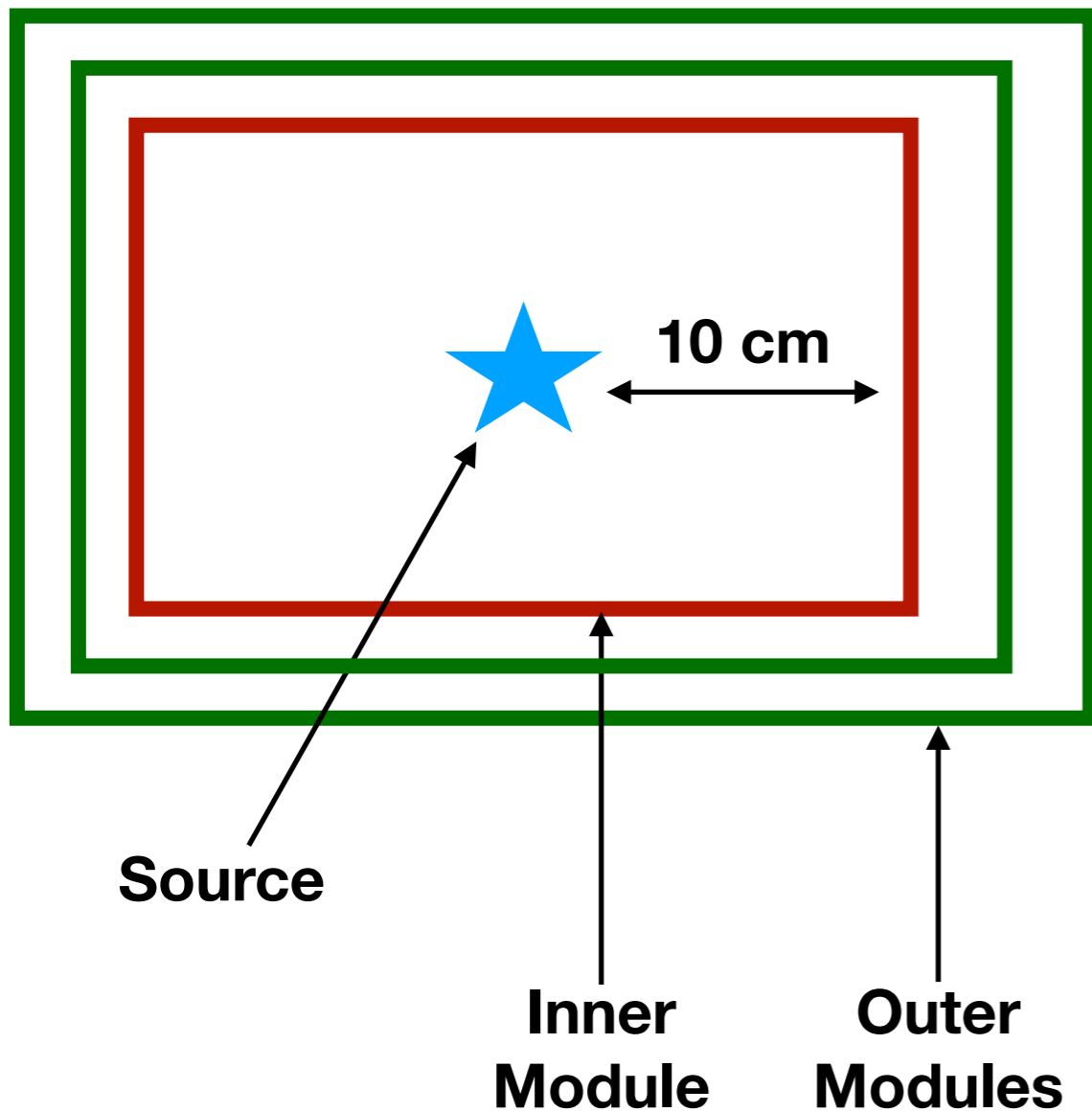
Geometry

Soft β to 2+ and Soft Compton γ



Geometry

Soft β to 2+ and Soft Compton γ



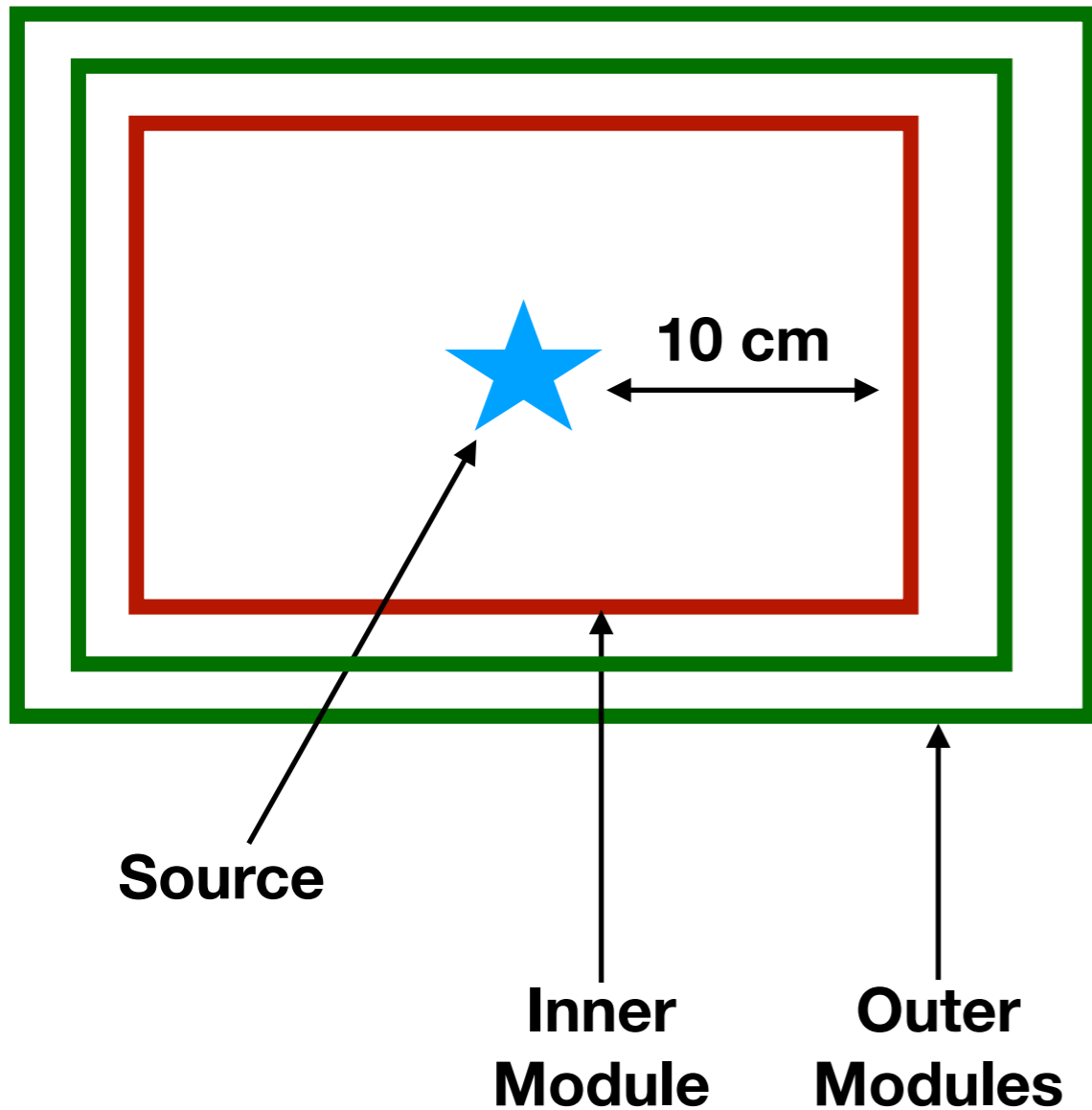
Geometry separates β & γ .

Confusion only if both hit same scintillator (\sim cm)

Simulated reach $\sim 10^{-11}$

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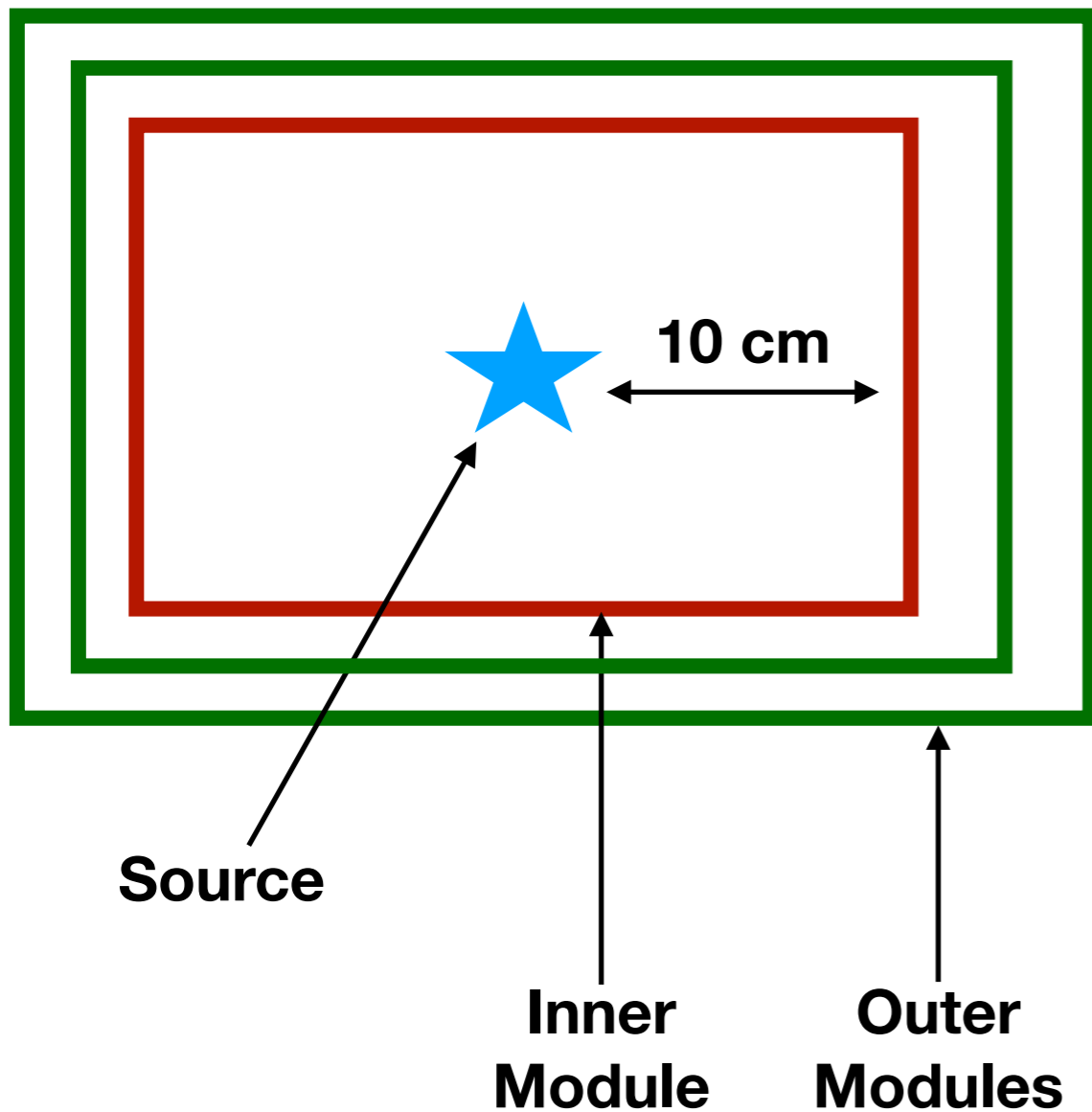
Possible Elimination?

Separate source from inner module.

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Absent in ^{24}Na where $E_1 \gg E_2$

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Soft β to $2+$ and mis-measured energy

Measure energy from light yield (LY)

Light yield set by quantum efficiency of photodetector (Q)

Plastic Scintillators: LY \sim 10000/MeV

PMT: Q \sim 0.25

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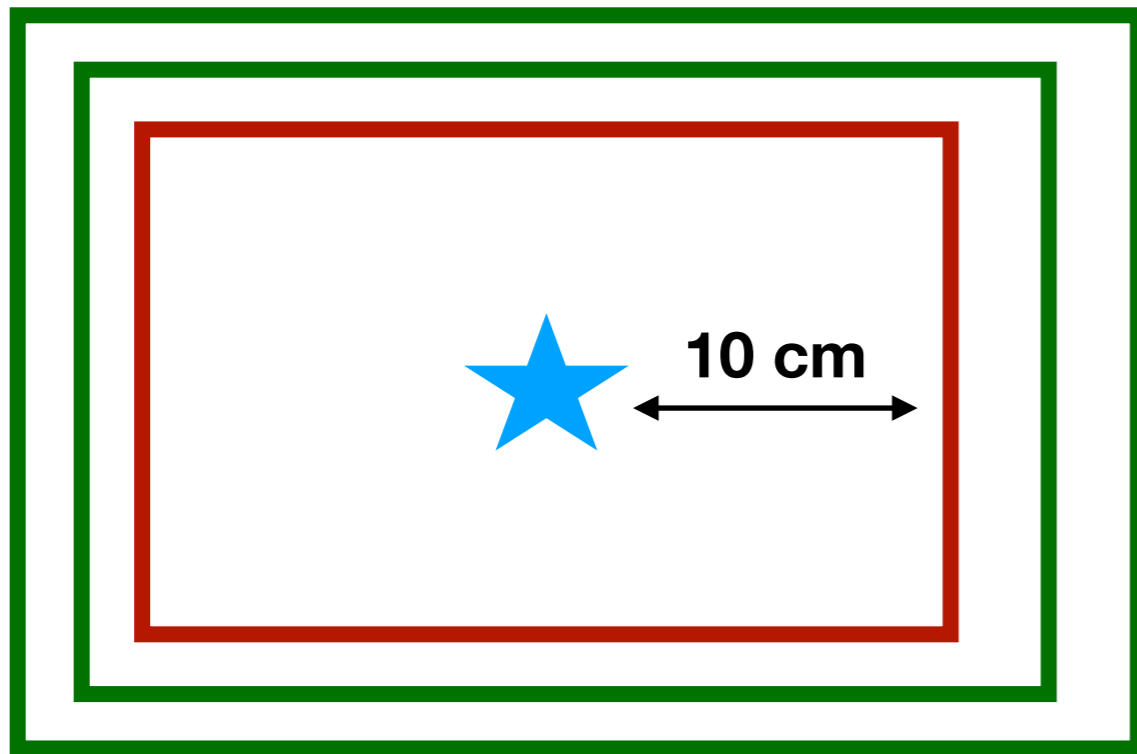
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Other Backgrounds



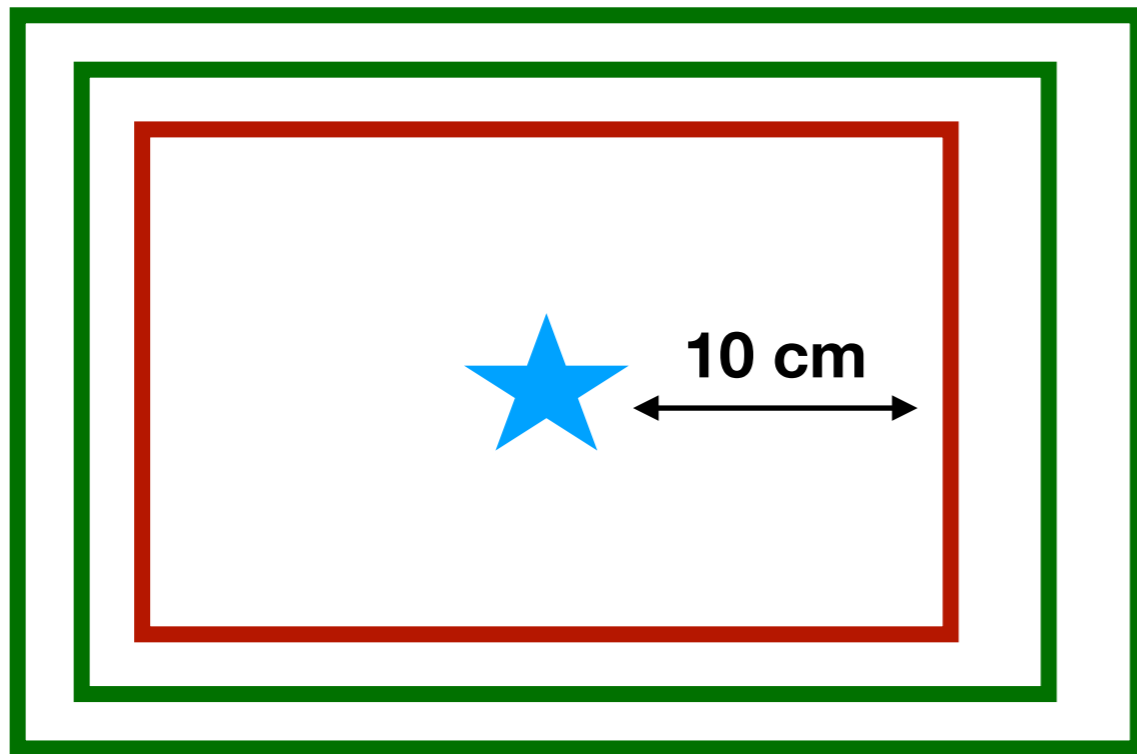
Detector Dead Volumes?

Well calibrated inner modules

Radiation Damage $< 10^4$ Grays

Further limit through separation

Other Backgrounds



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Radioactive Contaminants

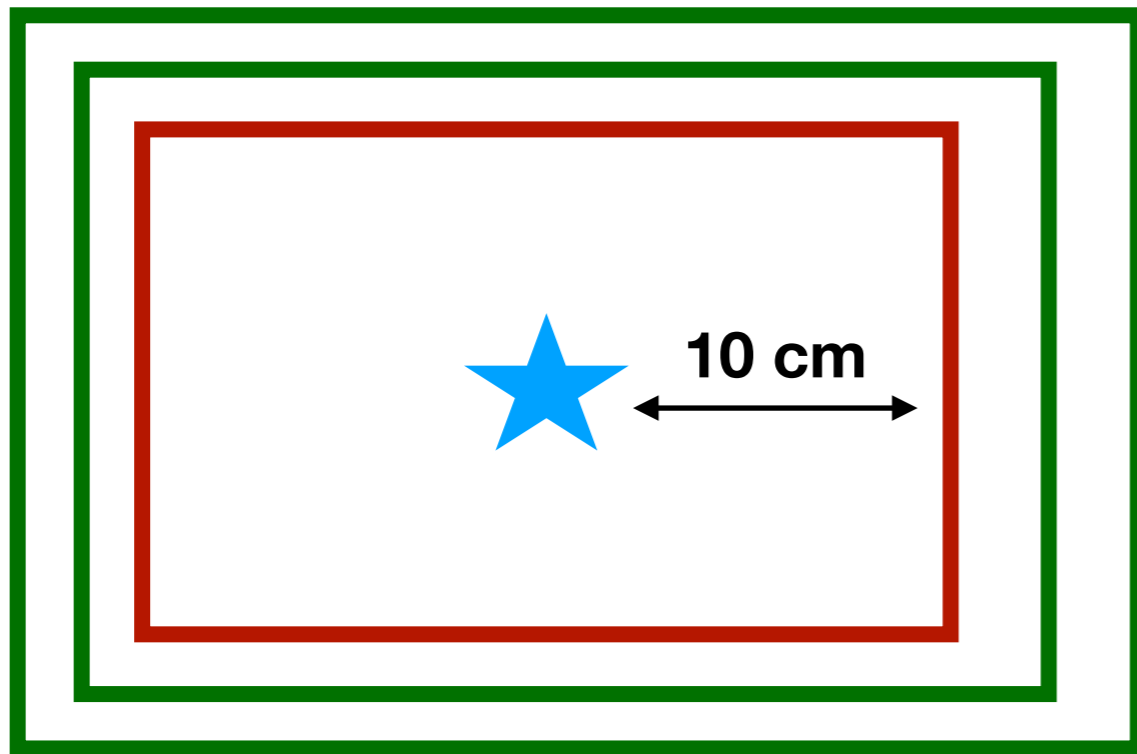
Long lived β at right energy?

None for ^{24}Na .

^{40}K for ^{60}Co - mBq/gm in some plastics.

Demand well separated β and γ in central module, ns timing

Triggers



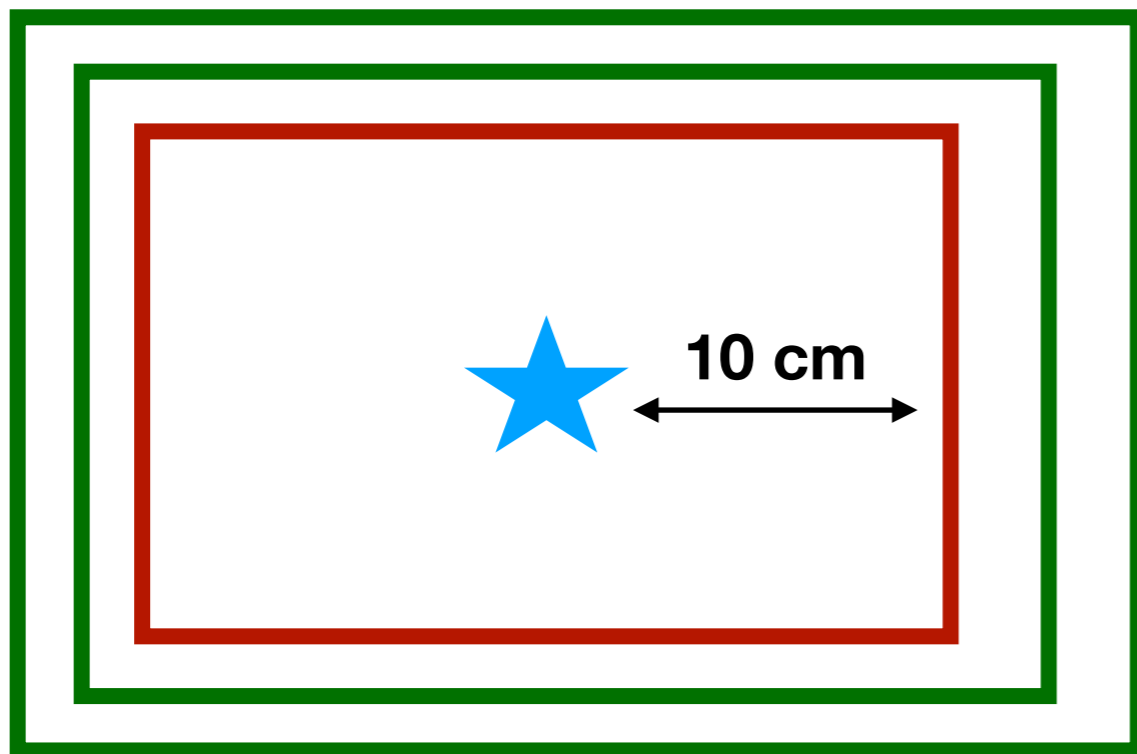
Cosmic Rays

Veto event with energy outside
inner module

Require well separated β and γ in
inner modules within \sim ns

Many radiation lengths separate
inner module from environment

Triggers



Trigger

@ 10^{-11} , not as hard as LHC

@ 10^{-14} , comparable to LDMX

Cosmic Rays

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Theory/Reach

Model

$$\mathcal{L} \supset g_p \phi \bar{\Psi}_p \Psi_p + \mu^2 \phi^2$$

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Need Branching fraction in E2 transitions.

Similar to γ transitions

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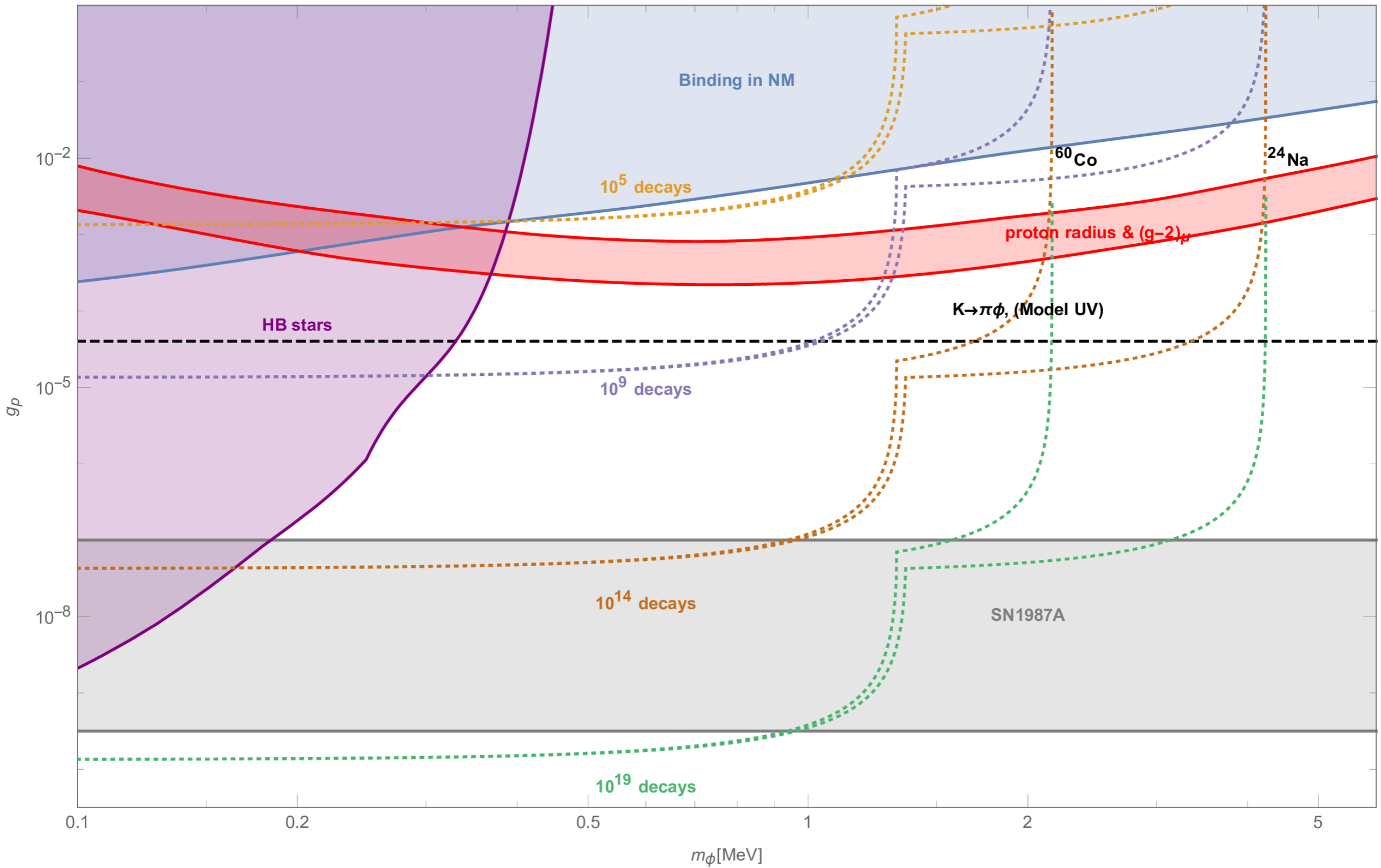
$$\frac{\Gamma_{\phi}}{\Gamma_{\gamma}} \sim \frac{g_p^2}{e^2}$$

Poor constraints on baryonic forces > 100 keV

Relevant for light dark matter experiments

Potentially cause Type 2 Supernova

Constraints



Conclusions

Probe Past Supernova? ($> 10^{12}/s$)

Not limited by availability of source. Complex Handling!

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Avoid pile up?

Resolve individual events - hard to get good energy
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Gamma Cascades in forbidden channels? Enhanced branching fraction for scalars?

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Axions: M1 transitions - $^{65}\text{Cu} \rightarrow ^{65}\text{Ni}$?