# Mono Photon Searches at 3 TeV With Polarised Beams

#### Outline

- Introduction
- Background cross-section calculation and event simulation.
- Background cross-sections and Syst errors

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- Background dN/dEγ distributions
- 95%CL cross-section calculation method
- 95%CL cross-section plots and results.

June®0Summary and outlook ising/IN2P3/LAPP

# Introduction

For the CLICdp Yellow report and for the additional contributions prepared for Granada, the exclusion limits for Dark Matter models were computed using 95% upper limits, computed using backgrounds without polarisation. Right handed polarised e<sup>-</sup> beams reduce the main SM background, v v y. The goal of this study is to estimate the increase in sensitivy taking into account the systematic errors.

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# **Background Cross-sections 3TeV**

Background cross-section values calculated for:  $10^{\circ} < \theta y < 170^{\circ}$  and Pty/ $\sqrt{s} > 0.02$ , without and with ebeam polarisation.

- Cross-section calculation and event generation done using Whizard with beam spectrum, isr function and n(1...3) matrix element photons.
- Events with Isr y's overlapping ME y's were rejected (Note: CLICdp-2020-004, Filip, Pawel ...).
- Fast simulation used to take into acount y energy resolution and efficiency and  $e^+ e^- \gamma$  veto efficiency. <sup>June, 2020</sup> JeV fast simulation based on extrapolation of full simulation.<sup>3</sup>

# **Background Cross-sections 3TeV**

	Polarisation		
	No	Pe-:-80%	Pe-:+80%
Process	σ[fb]	σ[fb]	σ[fb]
$e^{+} e^{-} \rightarrow v \overline{v} \gamma + v \overline{v} \gamma \gamma + v \overline{v} \gamma \gamma \gamma \gamma (\gamma)$	1058	1880	235
$e^{+} e^{-} \rightarrow e^{+} e^{-} \gamma + e^{+} e^{-} \gamma \gamma + e^{+} e^{-} \gamma \gamma \gamma (\gamma)$	1925	1960	1890

Background cross-section values for:  $10^{\circ} < \theta \gamma < 170^{\circ}$  and Pty/ $\sqrt{s} > 0.02$ .  $\sigma(\nu \nu \gamma)$  for Pe-:+80%,  $\sigma$  reduced by a factor 4.5 w.r.t. no polarisation and a factor 8, w.r.t. Pe-:-80% June 2020 J-J Blaising/IN2P3/LAPP

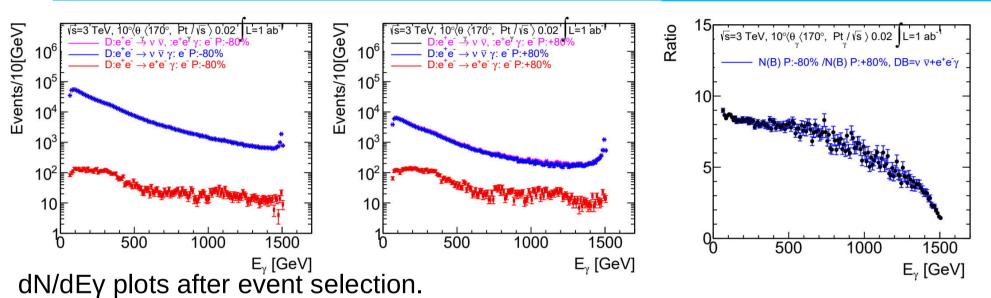
#### Systematic errors

Systematic error	Value
Event selection $v \overline{v} \gamma$	0.002
Event selection $e^+ e^- \gamma$	0.01 *
Luminosity	0.002
Polarisation	0.0025 **

The values are those used in the ILC paper arXiv:2001.03011v1. \* ILC does not assign a systematic error on the e<sup>+</sup> e<sup>-</sup> y veto. \*\*The sign of polarisation can be reversed at each bunch train, => experimental uncertainties will be strongly correlated.

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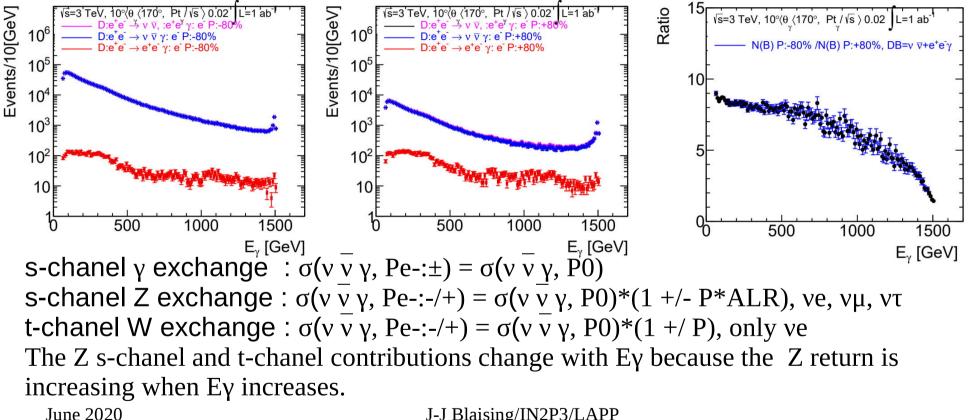
# Background dN/dE<sub>Y</sub> JL=1ab<sup>-1</sup>



Left : dN/dEy (Pe-:- 80%)  $v \bar{v} \gamma$ , e<sup>+</sup> e<sup>-</sup>  $\gamma$  events and sum: N(e<sup>+</sup> e<sup>-</sup>  $\gamma$ )/N( $v \bar{v} \gamma$ )=0.005 Midle: dN/dEy (Pe-:+80%)  $v \bar{v} \gamma$ , e<sup>+</sup> e<sup>-</sup>  $\gamma$  events and sum: N(e<sup>+</sup> e<sup>-</sup>  $\gamma$ )/N( $v \bar{v} \gamma$ )=0.04 Right :dN/dEy(Bsum) (Pe-:-80%) / dN/dEy B(sum) (Pe-:+80%) The shape of dR/dE $\gamma$  is due to the  $v \bar{v} \gamma$  s and t-channel contributions changing with  $E \bar{\gamma}^{020}$  J-J Blaising/IN2P3/LAPP

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# Background dN/dEy (L=1ab<sup>-1</sup>



# 95% Upper Limit Calculation

Definitions (Reminder)

To compute the 95% CL upper limit the likelihood ratio test statistic p=p(s+b)/p(b) (F1) is used:

Nobs=number of observed events, b=number of background events b+s = number of background + signal events.

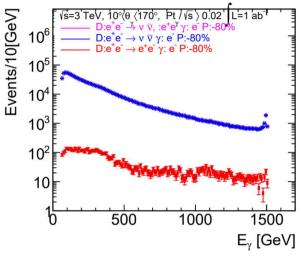
p=probability, CL=Confidence Level, CL=1-p

A 1 sided 95% CL upper limit corresponds to  $p \ge 0.025$  and  $z \le 2$ 

$$p = \frac{\sum_{n=0}^{Nobs} (b+s)^n \frac{e^{-(b+s)}}{n!}}{\sum_{n=0}^{Nobs} (b)^n \frac{e^{-b}}{n!}}$$
(F1)

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## **σ(95%) for** Pe-:-80%

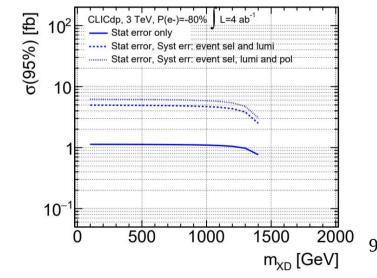


-dN/dEy for  $v \bar{v} v$  and  $e^+ e^- v$  events and sum for Pe-:-80%. For a DM mass mXD, Eymax=  $\sqrt{s/2-2*mXD^2}/\sqrt{s}$ . The background b is computed for 50 GeV < Ey < Eymax using F1. The s excluded at 95%CL is computed using F2.  $\sigma(95\%) = s/Lumi. \sigma(95\%)$  is computed for 200 GeV < mXD < 1400 GeV without and with systematic errors.

With systematic errors :  $\sigma(95\%)=6$  fb.

$$b = \int_{E_{\gamma_{min}}}^{E_{\gamma_{max}}} dN / dE_{\gamma}$$
 (F1)

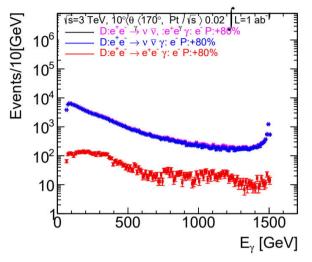
$$\frac{\sum_{n=0}^{Nobs} (b+s)^n \frac{e^{-(b+s)}}{n!}}{\sum_{n=0}^{Nobs} (b)^n \frac{e^{-b}}{n!}} \ge 0.025 \text{ (F2)}$$



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## σ(95%) for Pe-:+80%

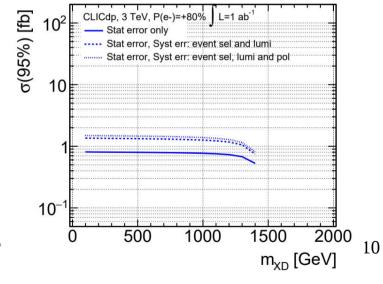


-dN/dEy for  $v \bar{v} v$  and  $e^+ e^- v$  events and sum for Pe-:-80%. For a DM mass mXD, Eymax=  $\sqrt{s/2-2*mXD^2}/\sqrt{s}$ . The background b is computed for 50 GeV < Ey < Eymax using F1. The s excluded at 95%CL is computed using F2.  $\sigma(95\%) = s/Lumi. \sigma(95\%)$  is computed for 200 GeV < mXD < 1400 GeV without and with systematic errors.

With systematic errors :  $\sigma(95\%)=1.4$  fb.



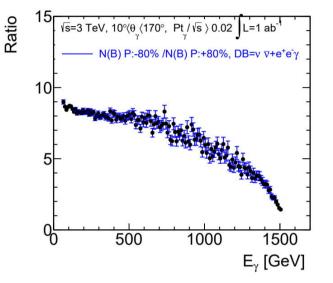
$$\frac{\sum_{n=0}^{Nobs} (b+s)^n \frac{e^{-(b+s)}}{n!}}{\sum_{n=0}^{Nobs} (b)^n \frac{e^{-b}}{n!}} \ge 0.025 \text{ (F2)}$$



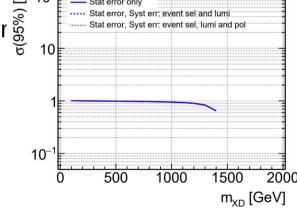
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## σ(95%) for Pe-:-80%/Pe-:+80%



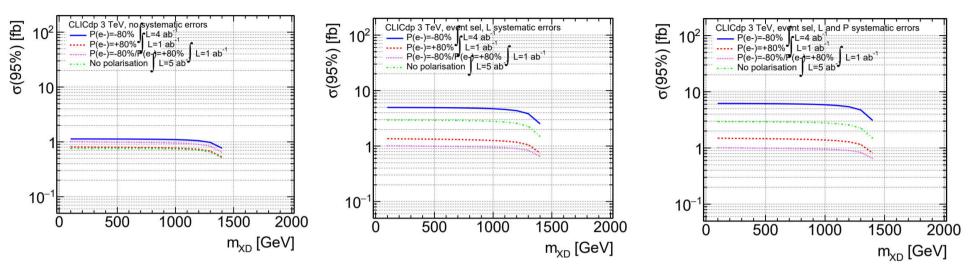
Plot: dRB/dEy: RB=NB( Pe-:-80%)/NB(Pe-:+80%) (B only) RB and  $\Delta$ RB computed for 50 GeV < Ey < Eymax In presence of signal: RBPS=(NB:-80%+NS)/(NB:+80%+NS) Compute NS to get RBPS=RB-2\* $\Delta$ RB.  $\sigma$ (95%,RB)=NS/Lumi Compute  $\sigma$ (95%,RB) for 200 GeV < mXD < 1400 GeV without and with systematic errors. With systematic errors:  $\sigma(95\%, RB)=1$  fb. It increases by 2% w.r.t No Syst Err 3 TeV, P(e-)=-80% and P(e-)=+80% Stat error only Stat error, Syst err: event sel and lum



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#### 95% CL cross-section values



95% upper limit  $\sigma$ , for different polarisation conditions: L, R, L/R, No. Left : without systematic errors Middle: with systematic errors: event selection and luminosity Right : with systematic errors: event selection, luminosity and polarisation W.r.t no polarisation, P(e-)=+80% is decrasing the  $\sigma$ (95%) by 2 and RP by 3. June 2020

# Summary

Right handed e<sup>-</sup> polarised beams reduce the main SM background, v v  $\gamma$ , increasing significantly the sensitivy. Taking into account the systematic errors:  $\sigma(95\%, \text{Pe-:-}80\%, \text{L}=1ab^{-1}) = 1.4$  fb for MXD=1 TeV  $=\sigma(95\%, \text{Pe-:}0, \text{L}=5ab^{-1})/2$  $\sigma(95\%, \text{RP}, \text{L}=1ab^{-1}) = 0.95$  fb for MXD=1 TeV  $=\sigma(95\%, \text{Pe-:}0, \text{L}=5ab^{-1})/3$ 

# Outlook

Using axial, axial-vector and scalor simplified DM models, study:

- Mass determination and cross-section measurement.
- Chiral properties: couplings identification

# **Additional slides**

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### **Radiative Neutrinos Diagrams**

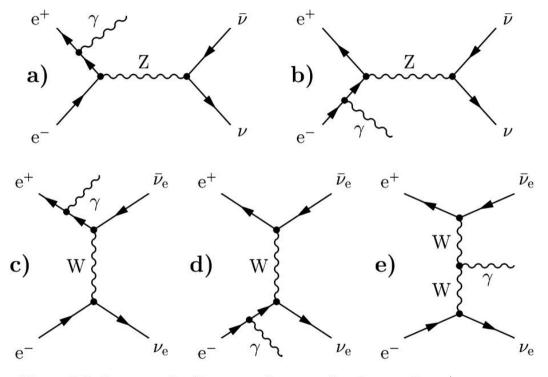
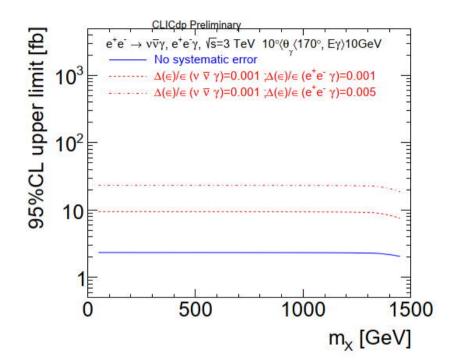


Figure 2.6: Lowest-order Feynman diagrams for the reaction  $e^+e^- \rightarrow \nu \bar{\nu} \gamma$ .

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# **2019** σ(95%CL)



σ(95%CL) values used by Ulrike to compute the DM exclusion limits for Granada.

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