

$$\mathcal{L}_{EW} = i\bar{\psi}\gamma^\mu D_\mu\psi - \frac{1}{4}\mathbf{W}^{\mu\nu} \cdot \mathbf{W}_{\mu\nu} - \frac{1}{4}B^{\mu\nu} B_{\mu\nu} + \mathcal{L}_\Phi$$

$$\mathcal{L}_\Phi = |D_\mu\Phi|^2 - \mu^2|\Phi|^2 - \lambda|\Phi|^4 + \mathcal{L}_{Yukawa}$$

$$\mathcal{L}_{QCD} = i\bar{\psi}_a\gamma^\mu D_\mu^{ab}\psi_b - \frac{1}{4}G_a^{\mu\nu} G_{\mu\nu}^a$$

Beyond Standard Model Physics

At the LHC

Piyabut Burikham

The SM

- Quantum gauge theory of 3 “fundamental” interactions, *gravity excluded*.
- Gauge group: $SU(3)$ $SU(2)$ $U(1)$
- Unitary group of [*color weak hypercharge*]
- Verified only up to 100 GeV energy (Tevatron)
- But dimensional analysis suggests quantum gravity scale $\sim 10^{18}$ GeV
- **New Physics** inevitable! (but somewhat remote)

$$\mathcal{L} = \mathcal{L}_{EW} + \mathcal{L}_{QCD} \quad \left. \vphantom{\mathcal{L}} \right\} \text{Before Symmetry breaking}$$

$$\mathcal{L}_{EW} = i\bar{\psi}\gamma^\mu D_\mu\psi - \frac{1}{4}\mathbf{W}^{\mu\nu} \cdot \mathbf{W}_{\mu\nu} - \frac{1}{4}B^{\mu\nu}B_{\mu\nu} + \mathcal{L}_\Phi$$

$$\mathcal{L}_\Phi = |D_\mu\Phi|^2 - \mu^2|\Phi|^2 - \lambda|\Phi|^4 + \mathcal{L}_{Yukawa}$$

$$\mathcal{L}_{QCD} = i\bar{\psi}_a\gamma^\mu D_\mu^{ab}\psi_b - \frac{1}{4}G_a^{\mu\nu}G_{\mu\nu}^a$$

$$\mathbf{W}_{\mu\nu} = \partial_\mu\mathbf{W}_\nu - \partial_\nu\mathbf{W}_\mu - g\mathbf{W}_\mu \times \mathbf{W}_\nu$$

$$G_{\mu\nu}^a = \partial_\mu G_\nu^a - \partial_\nu G_\mu^a - igf^{abc}G_{\mu b}G_{\nu c}$$

Gauge fields: B
(hypercharge), W
(weak), G (gluon)

Cov.Derivative:
Dynamics&gauge
interactions, generators
T, Ts

$$D_\mu = \partial_\mu + ig\mathbf{W}_\mu \cdot \mathbf{T} + ig'\frac{1}{2}B_\mu Y$$

$$D_\mu^{ab} = \partial_\mu\delta^{ab} + ig_s\mathbf{G}_\mu \cdot \mathbf{T}_s^{ab}$$

Representation matters!!

- Fermions in fundamental representations
- Gauge bosons in adjoint representations
- **SM Higgs** in **fund**amental representation
- Representation = a way fields group together forming single multi-component field
- Can we have **ad**joint fermions(SUSY), scalars(SUSY, extraD) in nature? → New Physics

Particle contents

$$\psi_L^j = \begin{pmatrix} \nu_L^j(u_L^j) \\ \ell_L^j(d_L^j) \end{pmatrix}$$

$$\psi_R^j = \nu_R^j(u_R^j) \text{ and } \ell_R^j(d_R^j)$$

$$\Phi = \begin{pmatrix} \phi^+ \\ \phi^0 \end{pmatrix}$$

$$= \exp\left(\frac{i\zeta(x) \cdot \mathbf{T}}{v\sqrt{2}}\right) \begin{pmatrix} 0 \\ (v + H(x))/\sqrt{2} \end{pmatrix}$$

$$\mathcal{L}_{Yukawa} = -\frac{\sqrt{2}}{v} \left[M_{ij}^d \overline{\psi_R^{di}} \Phi^\dagger \psi_L^j + M_{ij}^u \overline{\psi_R^{ui}} (i\tau_2 \Phi^*)^\dagger \psi_L^j \right] + \text{h.c.}$$

- 3 generations of fermions (all detected) **Why??**
- Scalar doublet, with Yukawa couplings
- Self-interacting potential selects **vacuum**

EW Gauge charges

- Left-handed fermions form weak SU(2) doublets.
- Right-handed fermions are weak singlets.
- Left and Right have different gauge charges T, Y
- *But same Q (electric charge)*
- $Q = T_3 + Y/2$
- **Observe:** no right-handed neutrino \rightarrow **extended** SM includes a number of ν_R (first BSM physics at Kamiokande, “*neutrino has masses!*”)

	T	T_3	$\frac{1}{2}Y$	Q
ν_{eL}	$\frac{1}{2}$	$\frac{1}{2}$	$-\frac{1}{2}$	0
e_L	$\frac{1}{2}$	$-\frac{1}{2}$	$-\frac{1}{2}$	-1
u_L	$\frac{1}{2}$	$\frac{1}{2}$	$\frac{1}{6}$	$\frac{2}{3}$
d_L	$\frac{1}{2}$	$-\frac{1}{2}$	$\frac{1}{6}$	$-\frac{1}{3}$
e_R	0	0	-1	-1
u_R	0	0	$\frac{2}{3}$	$\frac{2}{3}$
d_R	0	0	$-\frac{1}{3}$	$-\frac{1}{3}$

	T	T_3	$\frac{1}{2}Y$	Q
ϕ^+	$\frac{1}{2}$	$\frac{1}{2}$	$\frac{1}{2}$	1
ϕ^0	$\frac{1}{2}$	$-\frac{1}{2}$	$\frac{1}{2}$	0

EW breaking: $SU(2)_L \times U(1)_Y \rightarrow U(1)_{em}$

$$\mathcal{L}_{Yukawa} = -\frac{\sqrt{2}}{v} \left[M_{ij}^d \overline{\psi_R^{di}} \Phi^\dagger \psi_L^j + M_{ij}^u \overline{\psi_R^{ui}} (i\tau_2 \Phi^*)^\dagger \psi_L^j \right] + \text{h.c.}$$

- Scalar doublet gains vev, breaks $SU(2)_L \times U(1)_Y$ spontaneously (i.e. Vacuum breaks symmetry, action still preserving sym) through Yukawa coupling. $V = 246$ GeV.

$$\mathcal{L}_{EW} = \mathcal{L}_{KE} - \frac{m}{v} H \bar{\psi} \psi + \mathcal{L}_{em} + \mathcal{L}_{neutral} + \mathcal{L}_{charged} + \mathcal{L}_{Higgs}$$

After sym breaking, $SU(3)_c$ remains, EW action acquires mass terms.

SM Higgs mechanism(Breaking& Mixings)

$$\mathcal{L}_{em} = -eQ\bar{\psi}\gamma^\mu\psi A_\mu$$

$$V = \frac{\mu^2}{2}(v + H)^2 + \frac{\lambda}{4}(v + H)^4$$

$$\mathcal{L}_{neutral} = -\frac{g}{\cos\theta_W}\bar{\psi}_\alpha\gamma^\mu g_\alpha\psi_\alpha Z_\mu$$

Higgs self-potential

$$\mathcal{L}_{charged} = -\frac{g}{\sqrt{2}}\bar{\psi}_L^i\gamma^\mu(T^+W_\mu^+ + T^-W_\mu^-)\psi_L^i$$

$$\mathcal{L}_{Higgs} = \frac{1}{4}g^2W^+W^-(v + H)^2 + \frac{1}{8}\frac{g^2}{\cos^2\theta_W}ZZ(v + H)^2 - V$$

where $g_\alpha = T_3 - Q\sin^2\theta_W, -Q\sin^2\theta_W$ for $\alpha = L, R$, and $\tan\theta_W = g'/g$. θ_W is the Weinberg angle, representing the mixing between B and W^3 to form the photon $A = B\cos\theta_W + W^3\sin\theta_W$, and the $Z = -B\sin\theta_W + W^3\cos\theta_W$. $W^\pm = (W^1 \mp iW^2)/\sqrt{2}$ are the charged gauge bosons from the mixing of the original $W^{1,2}$.

Gauge boson masses

$$1/g^2 + 1/g'^2 = 1/e^2 \quad \left\{ \begin{array}{l} \text{Coupling mixing} \end{array} \right.$$

- $m_W = gv/2$, $m_Z = m_W/\cos\theta_W$
Z is heavier. $\left\{ \begin{array}{l} \rho = M_W^2/M_Z^2 \cos^2 \theta_W \\ \text{Rho parameter} = 1 \\ \text{in SM Higgs} \end{array} \right.$

If the sym breaking is NOT SM Higgs mechanism,
Rho does NOT have to be 1. \rightarrow New Physics

Fermion mixing

- weak eigenstates = mixing of mass eigenstates

$$u^i = U_{iq}q \text{ for } q = u, c, t$$

$$d^i = D_{iq'}q' \text{ for } q' = d, s, b$$

- Only *charged* current feels the mixing \rightarrow effectively **CKM**(Cabibbo-Kobayashi-Maskawa) matrix applies to (d,s,b) . For $V = U_L^\dagger D_L$

$$\mathcal{L}_{\text{charged}} = -\frac{g}{\sqrt{2}} \left(\bar{q}_L \gamma^\mu W_\mu^+ V_{qq'} q'_L + \bar{q}'_L V_{q'q}^* \gamma^\mu W_\mu^- q_L \right)$$

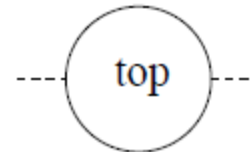
- For neutrinos, called **MNS**(Maki-Nakagawa-Sakata) matrix.

What about Higgs mass?

- At tree level, $m_H \propto \mu$ but this means nothing since for the scalar field, quantum corrections are *enormous*. For **cut**off scale Λ ,

$$-\frac{3}{8\pi^2}\lambda_t^2\Lambda^2$$

from the top



$$\frac{1}{16\pi^2}g^2\Lambda^2$$

from the gauge boson



$$\frac{1}{16\pi^2}\lambda^2\Lambda^2$$

from the Higgs



Cutoff scale 10^{16-18} GeV

- Quantum corrections are enormous. Higgs mass cannot be < 1 TeV unless fine tuning occurs. \rightarrow **Fine** tuning problem in SM Higgs.
- Several **BSM** models to address fine tuning problem.
- Among many: Little Higgs models, low-scale SUSY models such as MSSM (Minimal SUSY SM), Technicolour models, Extra-D models
- The most boring possibility = SM Higgs with fine tuning!!

Why needing $m_H < 1 \text{ TeV}??$

Any $2 \rightarrow 2$ scattering amplitude can be expanded using partial waves:

$$A(s, t) = 16\pi \sum_{j=M}^{\infty} (2j+1) a_j(s) d_{mm'}^j(\cos \theta)$$

Wigner functions $d_{mm'}^j(\cos \theta)$

- $J = 0, 1$ part of σ ($ZZ, W^+W^- \rightarrow W^+W^-, ZZ$) proportional to **2nd power** of $(E_{\text{cm}}/m_W) \rightarrow$ Will **violate unitarity** around $E \approx m_W$.
- Higgs exchange in t -channel will cancel this contribution iff $m_H < 1 \text{ TeV}!!$

NP inevitable around 1 TeV 😊!!

- Most boring scenario: SM Higgs with fine tuning of the mass so that $m_H < 1$ TeV. 😞
- OR Other NPs coupling to W, Z show up around 1 TeV.

Lists: (as well as combinations)

- Non-SM Higgs models (Little Higgs)
- Higgsless models
- Extra-D models (KK, TeV-braneworld)
- Composite models (technicolor, preon)
- SUSY models (top-down, bottom-up)

Motivations for NP models

- In addition to unitarity argument that **NP must** show up around 1 TeV, *hierarchy problem* or *fine tuning problem* is also a motivation.
- Large mass gap between Planck scale (or GUT scale) and EW breaking scale, **10^{18} GeV** and **100 GeV**. **Nothing** in between? Really?
- A scalar such as Higgs receives quantum corrections to its **mass** proportional to cutoff scale square $\Lambda^2 \rightarrow$ if Λ huge, *fine tuning* is required for $m_H < 1$ TeV.

EW precision observables (any NPs need to pass.)

- ρ parameter: $\rho = M_W^2 / M_Z^2 \cos^2 \theta_W$
LEP2 results found rho **very** close to 1.
- 1-loop Higgs contributions to $m_{W,Z}$ *constrain* SM Higgs mass.

$$\hat{\rho} \approx 1 - \frac{11G_F M_Z^2 \sin^2 \theta_W}{24\sqrt{2}\pi^2} \ln \frac{m_h^2}{M_Z^2}$$

Global fits leading to $\rho_0 = 1.0008^{+0.0017}_{-0.0007}$,
 $114.4 \text{ GeV} \leq M_H \leq 427 \text{ GeV}$,

Parameter	Value	SM value
m_t [GeV]	176.1 ± 7.4	176.9 ± 4.0
M_W [GeV]	80.454 ± 0.059	80.390 ± 0.018
M_Z [GeV]	91.1876 ± 0.0021	91.1874 ± 0.0021
Γ_Z [GeV]	2.4952 ± 0.0023	2.4972 ± 0.0012
$\Gamma(\text{had})$ [GeV]	1.7444 ± 0.0020	1.7435 ± 0.0011
$\Gamma(\text{inv})$ [MeV]	499.0 ± 1.5	501.81 ± 0.13
$\Gamma(\ell^+\ell^-)$ [MeV]	83.984 ± 0.086	84.024 ± 0.025

EW precision measurements

- Some of EW values any NPs cannot violate.
- Strongest constraints usually come from Z-pole precision measurements.

*More updated values in PDG.

Extra-D models

- Roughly 3 categories: ADD, RS, Braneworld
- ADD(Antoniadis-Arkani Hamed-Dimopoulos-Dvali): Large compactified flat extra-D
- RS(Randall-Sundrum): infinite curved extra-D (Anti-de Sitter space)
- Braneworld (Witten-Horava-Antoniadis-Dvali): we live on the worldvolume of Dbranes, **only gravity** can probe extra-D!!

ADD scenario

- Extra-D compactified in a **torus** (flat) \rightarrow KK (Kaluza-Klein) modes with nth mode mass: $m_{KK} = nh/2\pi R$.

$$S_E = \frac{\bar{M}_D^{2+\delta}}{2} \int d^4x d^\delta y \sqrt{-\det g} \mathcal{R}(g)$$

$$\bar{M}_{Pl}^2 = \bar{M}_D^{2+\delta} V_\delta = \bar{M}_D^{2+\delta} (2\pi R)^\delta$$

$$\bar{M}_{Pl} = M_{Pl}/\sqrt{8\pi} = 2.4 \times 10^{18} \text{ GeV}$$

$1/R \sim 0.4 \text{ meV} - 7 \text{ MeV}$ ($\delta=2 - 6$) for $M_D = 1 \text{ TeV}$



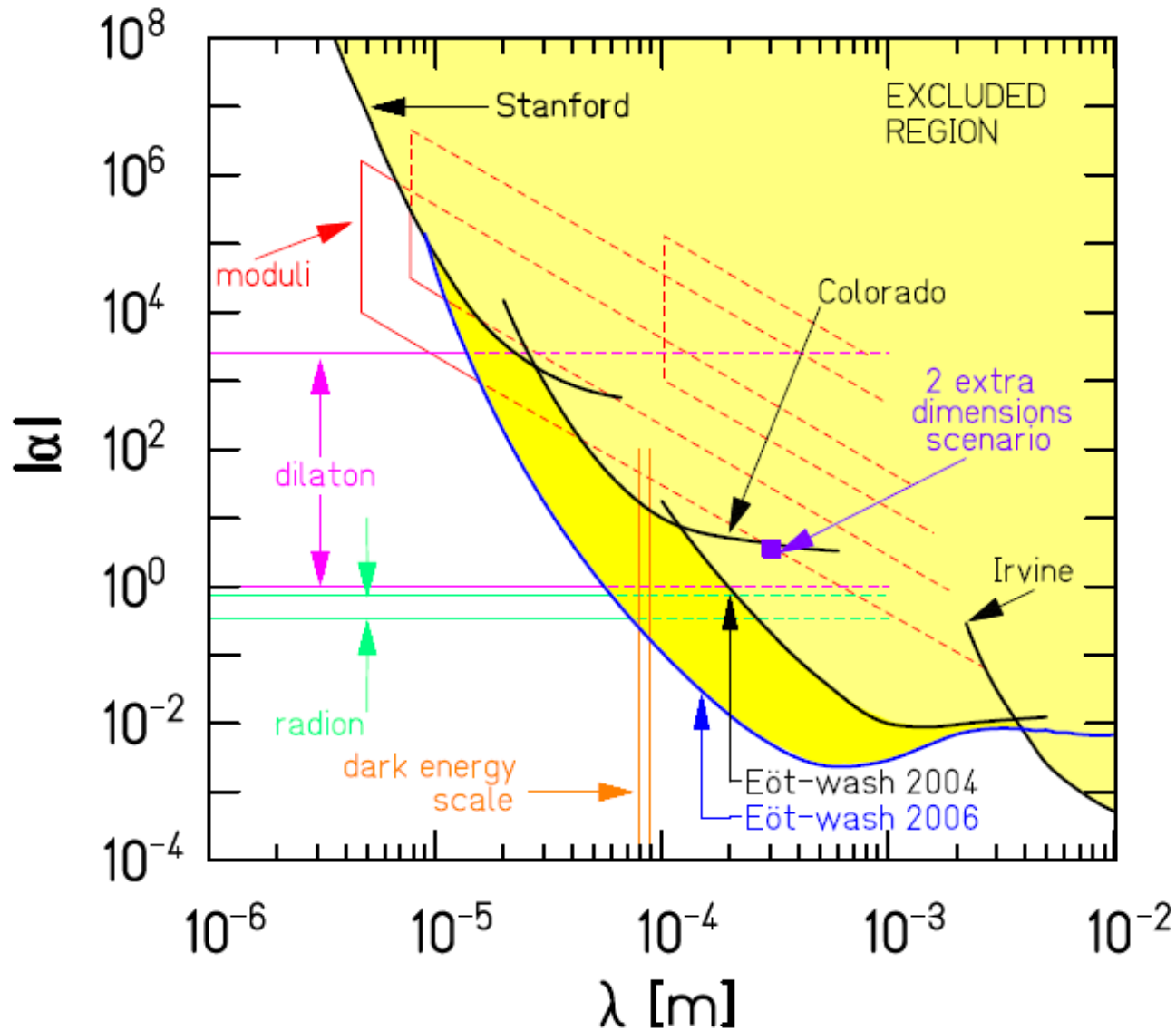
Small KK levels

- Sufficiently large $R \rightarrow$ quite small quantum gravity scale $M_D \rightarrow$ *low scale quantum gravity!!*
- But how BIG can it be?? Most stringent *universal* constraints from table-top experiments, *e.g.* Eotwash
- Parametrized as deviation from inverse-square law: **KK-graviton** in extra-D generates **Yukawa** potential

$$V(r) = -G_N \frac{m_1 m_2}{r} [1 + \alpha \exp(-r/\lambda)]$$

$$\alpha = 8 \delta/3 \quad \blacksquare \quad R < 37 (44) \mu\text{m} \text{ at } 95\% \text{ CL for } \delta = 2 \quad (1)$$

$$R^{-1} = M_D \left(M_D / \overline{M}_{\text{Pl}} \right)^{2/\delta} \quad \text{Weaker bound on } M_D \text{ for } \delta > 2.$$



* **Stronger** bound comes from Supernovae cooling via radiation of **extra-D** d.o.f. such as **KK-gravitons**

- Bounds from KK-gauge, radion, dilaton: $M_D > 3.6$ TeV for $\delta = 2$.
- Bounds from KK-gravitons from supernovae cooling: $M_D > 14$ (1.6) TeV for $\delta = 2$ (3).
- Stronger bounds from luminosities of pulsar hit by KK-gravitons: $M_D > 750$ (35) TeV for $\delta = 2$ (3).

ADD **open** questions:

- **R**adius stabilization, what mechanism fixing the radius of extra-D? Why this value?
- Still don't know **how to quantize** gravity, worse when quantum gravity scale is this small!!

Collider signals

$$\mathcal{L}_{\text{int}} = \pm \frac{4\pi}{\Lambda_T^4} \mathcal{T}, \quad \mathcal{T} = \frac{1}{2} \left(T_{\mu\nu} T^{\mu\nu} - \frac{1}{\delta + 2} T^\mu{}_\mu T^\nu{}_\nu \right) \quad \left. \vphantom{\mathcal{L}_{\text{int}}} \right\} \text{Dim.8 operator}$$

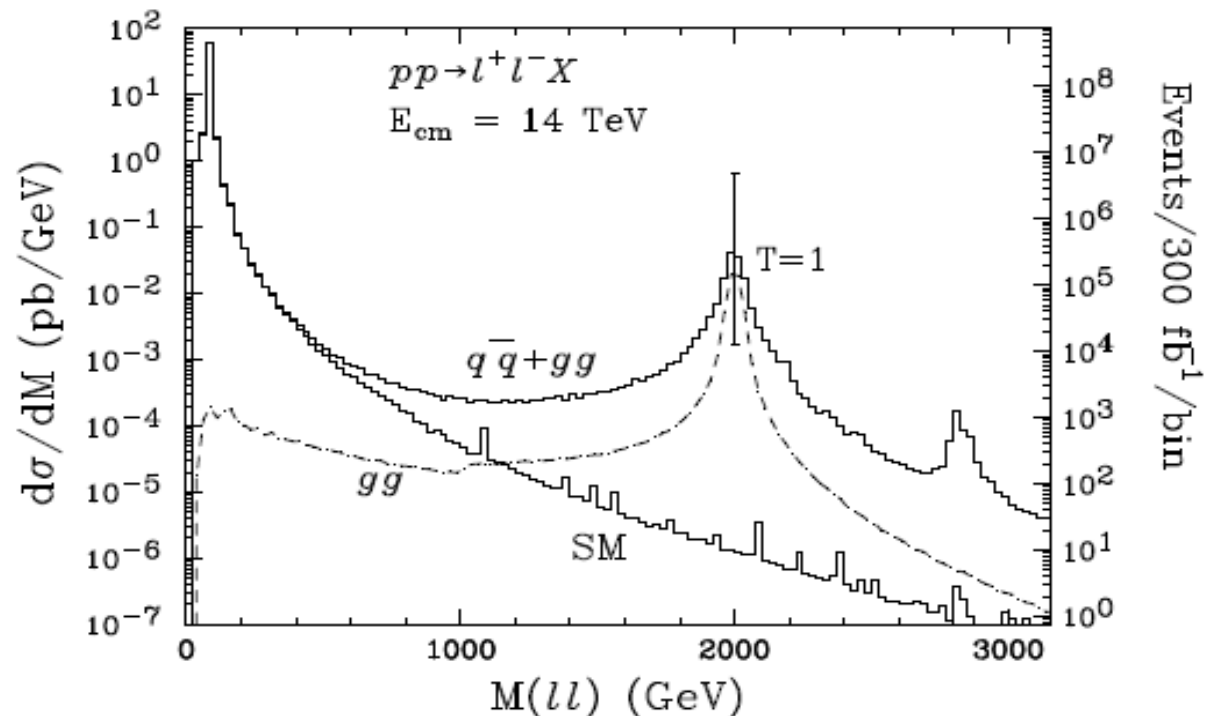
$$\mathcal{L}_{\text{int}} = \pm \frac{4\pi}{\Lambda_\Upsilon^2} \Upsilon, \quad \Upsilon = \frac{1}{2} \left(\sum_{f=q,\ell} \bar{f} \gamma_\mu \gamma_5 f \right)^2 \quad \left. \vphantom{\mathcal{L}_{\text{int}}} \right\} \text{Dim.6 operator}$$

- Effective interactions induced by graviton exchanges, **tree**-level(dim.8) and **loop**(dim.6).
- Current lower bound on scales $\approx 1\text{-}10$ TeV. Still **visible** at LHC if exists!
- Best channels: **lepton pair**, **diphoton** production

TeV-string signals

- If quantum gravity scale is as low as TeVs and if the correct QG theory is string-like, LHC signals are enormous!!
- SR (string resonances)

or stringy
excitations could
enhance SM
scatterings
(P. Burikham et. al.).



RS scenario

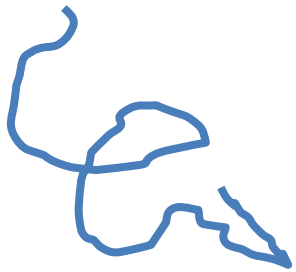
- Curved or warped extra-D, 5-D space, the 5th compactified on half-circle. S^1/Z_2

$$ds^2 = \exp(-2k|y|) \eta_{\mu\nu} dx^\mu dx^\nu - dy^2$$

- 2 branes at opposite ends or *fixed pts*, with negative and positive tension. Negative-tension brane = **IR** brane ($y=\pi R$) where SM particles **localized**, positive-tension brane = **UV** brane ($y=0$).
- Bulk cosmological constant **fine-tuned** to exactly cancel apparent 3-D cosmological constant.

$$ds^2 = \exp(-2k|y|) \eta_{\mu\nu} dx^\mu dx^\nu - dy^2$$

- Spacetime *not* factorized, metric on 4-D exponentiated by 5th coordinate.
- Gravitation redshift by factor $1/g_{00}^{1/2}$, energies from UV-brane viewed by IR-world redshifted by this factor.



$$m_{IR} = m_{UV} \exp(-\pi k R)$$



Large hierarchy generated!

- For $kR \approx 12$, $M_{UV} = \text{Planck mass } (10^{18} \text{ GeV})$, $M_{IR} = \text{EW mass } (100 \text{ GeV})$ can be generated.
- **Radius** stabilization via Goldberger-Wise mechanism ('99).

- Nth mode KK has mass:

$$m_n = \frac{x_n}{x_1} m_1$$

- x_n is nth zero of Bessel function. m_1 is mass parameter.

- *Interaction:*

$$\mathcal{L} = -\frac{T^{\mu\nu}}{M_{\text{Pl}}} h_{\mu\nu}^{(0)} - \frac{T^{\mu\nu}}{\Lambda_\pi} \sum_{n=1}^{\infty} h_{\mu\nu}^{(n)}$$

0-th mode

KK modes

Collider Signals

$$\mathcal{L}_{\text{int}} = \pm \frac{4\pi}{\Lambda_\pi^4} \mathcal{T}, \quad \mathcal{T} = \frac{1}{2} \left(T_{\mu\nu} T^{\mu\nu} - \frac{1}{\delta + 2} T^\mu{}_\mu T^\nu{}_\nu \right) \quad \left. \vphantom{\mathcal{T}} \right\} \text{Dim.8 operator}$$

- Similar to ADD int. except there is NO dim.6 operator. \rightarrow cannot tell which model for certain if found.

$$\Lambda_\pi = \bar{M}_{\text{Pl}} \exp(-\pi k R)$$

- Again, best channels are dilepton, diphoton.
- Current bounds from D0 and CDF: $\Lambda_\pi > 4.3(2.6)$ TeV for $m_1 = 500(700)$ GeV.
- LHC can probe upto 10 TeV for m_n, Λ_π .


Radion

- Size of extra-D determined by radion, radion **stabilization** is crucial. For a radion r :

$$\mathcal{L} = -rT/\Lambda_\varphi \quad \Lambda_\varphi = \sqrt{24}\Lambda_\pi$$

- Trace anomaly from SM makes $gg \rightarrow r$ large, main channel to be searched at LHC.
- Radion can mix with Higgs through scalar-curvature int. For 4-D induced metric:

$$S_{mix} = -\xi \int d^4x \sqrt{-\det g_{ind}} R(g_{ind}) H^+ H$$

 **mixing parameter**

- Higgs-radion mixing \rightarrow search for radion is the same as Higgs.
- Radion stabilization requires radion mass **less** than KK-gravitons.

RS **open** questions:

- Why 5-D? Gauge and gravity in 5-D are **non**-renormalizable.
- GUT? How to quantize gravity? **String** theory at higher scales?
- Other questions remain, cosmological constant, baryogenesis, **DM** is proposed to be **lightest** KK-mode.

SM in flat extra-D

- Massive KKs in RGE (Renorm Group Eqn.) → GUT in extra-D at low scales, as low as TeVs!!
- In contrast to ADD, extra-D must be **smaller**, around TeV^{-1} since we do NOT observe KK SM particles below 1 TeV!
- Typical model: extra-D S/Z_2
- *Varieties*: Gauge bosons in the bulk, fermions&Higgs in the bulk, ALL in bulk (UED).

Gauge bosons in extra-D

- KK masses of gauge bosons: $M_n^2 = M_0^2 + \frac{n^2}{R^2}$
- KK bosons coupling: $(g_n = \sqrt{2}g)$

Shift of observables by factor V $V = 2 \sum_n \left(\frac{g_n^2}{g^2} \right) \frac{M_Z^2 R^2}{n^2} \sim \frac{2}{3} \pi^2 M_Z^2 R^2$

- Constraints from Precision EW data from Tevatron, HERA, LEP2 $\rightarrow 1/R > 6.8$ TeV.
- LHC at 100 fb^{-1} can probe to $1/R \sim 16$ TeV.

- Fermions at different location in extra-D overlap with Higgs wave function differently → observed **mass hierarchy**.
- **Universal** Extra-D(UED) → All SM particles live in bulk with **KK parity** (discrete sym Z_2 of KK number).
- Conservation of momentum in extra-D → KK-parity conservation → KK particles as DM candidate
- LHC can probe up to $1/R \sim$ **1.5** TeV.

GUT in extra-D

- Scherk-Schwarz mechanism to breaks symmetries such as GUT and SUSY.

$$+ : \cos \frac{n y}{R}$$

$$- : \sin \frac{n y}{R}$$

- Under S/Z_2 ($y \rightarrow -y$): only even states have 0^{th} mode. **Odd** states missing at the **fixed pts** $y=0, \pi R$ where our low E world lives.

- Could have $N=2$ SUSY in bulk and $N=1$ for 0^{th} modes at fixed pt if imposing A, λ even and Φ, ψ odd.
- Choosing diff boundary conditions for diff fields \rightarrow orbifolding $\rightarrow Z_2$ orbifold.

GUT in extra-D

- Orbifolding: can project out certain unwanted states by choosing odd b.c. so their 0^{th} modes won't show up at the **fixed pt world**.
- Can fudge while proton wont decay according to GUT.
- → Can make unification better at lower scale and yet proton decay is not too fast.

SM in warped extra-D

- In RS scenario, only Higgs need to live on the IR-brane to solve the fine tuning problem of the Higgs mass. \leftarrow corrections are redshifted.
- But if SM also lives in warped bulk, interesting things happen.
- In 5-D AdS space (AdS/CFT):
 1. motion along 5th D = RG flow of 4D theory
 2. local sym in 5D = global sym in 4D
- RS model \leftrightarrow walking technicolor model!!

SM in warped

- Can explain mass hierarchy if locate fermions on diff position along 5th coordinate. Higgs on IR-brane overlap differently with each fermion → diff. masses.
- EW precision observables especially ρ *parameter* excluded basic model → need to enlarge gauge group of EW in bulk to $SU(2)_L$ $SU(2)_R$ $U(1)$ to impose custodial symmetry preserving value of ρ .

Higgsless models

- No Higgs!! Breaks EW using orbifolding. Lightest KK gauge bosons identified with W, Z with masses $\sim 1/R$.
- Scatterings of SM without Higgs need something to unitarize the amplitude at $E \sim 4\pi m_W / g \sim 1$ TeV \rightarrow heavy KK gauge bosons do the job postponing this to 10 TeV!
- Above 10 TeV, strong dynamics take over, bound states expected to form.
- Require warped extra-D, custodial sym, still cannot predict top-quark mass.

SUSY models

- In a sense, this is extra-D models with Grassmann extra dimensions!
- Poincare sym. \rightarrow max'al extension to contain SUSY generators, $Q \sim \sqrt{P}$. \leftarrow *theoretical beauty*
- *Motivations (apart from beauty)*: fermionic loop contributes the same as bosonic loop but with opposite sign \rightarrow natural loop cancellations!
- Loop cancellation is promising for many purposes.

Unwanted Loops

- Quantum gravity suffers *loop* complications. Each order of loops is worse than the previous. → **unrenormalizable**.
- Loops induce **anomalies** (= breaking of classical sym by quantum effects).
- Pheno level: loop corrections to scalar mass proportional to Λ^2 → *fine tuning problem*.
- *SUSY* ensures loop cancellation at 1-loop order. ← *not only beautiful but also useful!*

SUSY algebra

$$Q|\text{Boson}\rangle = |\text{Fermion}\rangle, \quad Q|\text{Fermion}\rangle = |\text{Boson}\rangle$$

$$\{Q, Q^\dagger\} = P^\mu,$$

$$\{Q, Q\} = \{Q^\dagger, Q^\dagger\} = 0,$$

$$[P^\mu, Q] = [P^\mu, Q^\dagger] = 0,$$

- Q transforms boson to fermion and vice versa.
- P is a vector with spin 1 \rightarrow Q acts as spin $\frac{1}{2}$.
- *1-particle* states = irreducible rep. of SUSY algebra called *supermultiplets*

- Spin-statistic thm \rightarrow #bosonic d.o.f. = #fermionic d.o.f. for *each* supermultiplet \rightarrow Every particle must have its superpartner in SUSY theory.
- *chiral* multiplet = 2-component Weyl fermion + 2 real scalars (or 1 complex scalar)
- *vector* multiplet = 2-helicity gauge boson + 2-component Weyl fermion
- **partners** of Weyl in chiral mulp = *sfermions*, partners of gauge bosons in vector mulp = *gauginos*. Same **representations** in each mulp.

- If including gravity such as sugra: massless graviton (2) + massless gravitino (2), graviton has spin 2, gravitino has spin 3/2.
- For $N = \#$ of SUSY generator Q 's. **Extended** SUSY with $N > 1$ in 4D **not** allow chiral fermions \rightarrow unrealistic. But extra-D models $N > 1$ are realistic if chiral fermions can be obtained, *e.g.* at fixed pts.
- **Anomaly** free Higgs mulp = at least 2 chiral mulp for Higgs, one for up-type and one for down-type fermion.

MSSM (minimal supersymmetric SM)

Names		spin 0	spin 1/2	$SU(3)_C, SU(2)_L, U(1)_Y$
squarks, quarks ($\times 3$ families)	Q	$(\tilde{u}_L \ \tilde{d}_L)$	$(u_L \ d_L)$	$(\mathbf{3}, \mathbf{2}, \frac{1}{6})$
	\bar{u}	\tilde{u}_R^*	u_R^\dagger	$(\bar{\mathbf{3}}, \mathbf{1}, -\frac{2}{3})$
	\bar{d}	\tilde{d}_R^*	d_R^\dagger	$(\bar{\mathbf{3}}, \mathbf{1}, \frac{1}{3})$
sleptons, leptons ($\times 3$ families)	L	$(\tilde{\nu} \ \tilde{e}_L)$	$(\nu \ e_L)$	$(\mathbf{1}, \mathbf{2}, -\frac{1}{2})$
	\bar{e}	\tilde{e}_R^*	e_R^\dagger	$(\mathbf{1}, \mathbf{1}, 1)$
Higgs, higgsinos	H_u	$(H_u^+ \ H_u^0)$	$(\tilde{H}_u^+ \ \tilde{H}_u^0)$	$(\mathbf{1}, \mathbf{2}, +\frac{1}{2})$
	H_d	$(H_d^0 \ H_d^-)$	$(\tilde{H}_d^0 \ \tilde{H}_d^-)$	$(\mathbf{1}, \mathbf{2}, -\frac{1}{2})$

- Chiral supermultiplet in MSSM, superpartners must have the same gauge charges, reps.
- 2 Higgs multiplets.

MSSM

Names	spin 1/2	spin 1	$SU(3)_C, SU(2)_L, U(1)_Y$
gluino, gluon	\tilde{g}	g	(8, 1, 0)
winos, W bosons	$\tilde{W}^\pm \tilde{W}^0$	$W^\pm W^0$	(1, 3, 0)
bino, B boson	\tilde{B}^0	B^0	(1, 1, 0)

- Gauge mulps in MSSM. Gluino, wino, bino
- After EW breaking: photino, wino, zino
- Apparently, SUSY must be broken since we don't see partners with exactly the same masses around. \rightarrow probably spontaneously.

MSSM

- To break SUSY

$$\mathcal{L} = \mathcal{L}_{\text{SUSY}} + \mathcal{L}_{\text{soft}}$$

$$\begin{aligned} \mathcal{L}_{\text{soft}}^{\text{MSSM}} = & -\frac{1}{2} \left(M_3 \tilde{g} \tilde{g} + M_2 \tilde{W} \tilde{W} + M_1 \tilde{B} \tilde{B} + \text{c.c.} \right) \\ & - \left(\tilde{u} a_u \tilde{Q} H_u - \tilde{d} a_d \tilde{Q} H_d - \tilde{e} a_e \tilde{L} H_d + \text{c.c.} \right) \\ & - \tilde{Q}^\dagger m_Q^2 \tilde{Q} - \tilde{L}^\dagger m_L^2 \tilde{L} - \tilde{u} m_{\tilde{u}}^2 \tilde{u}^\dagger - \tilde{d} m_{\tilde{d}}^2 \tilde{d}^\dagger - \tilde{e} m_{\tilde{e}}^2 \tilde{e}^\dagger \\ & - m_{H_u}^2 H_u^* H_u - m_{H_d}^2 H_d^* H_d - (b H_u H_d + \text{c.c.}) . \end{aligned}$$

- Soft terms: mass terms, positive mass dimension coupling terms which **violate** SUSY.

Breaking MSSM

- Even when SUSY is broken, masses of partners are different. If $m_{\text{soft}} = \text{largest scale in soft terms}$, then the correction to Higgs mass is

$$\Delta m_H^2 = m_{\text{soft}}^2 \left[\frac{\lambda}{16\pi^2} \ln(\Lambda_{\text{UV}}/m_{\text{soft}}) + \dots \right]$$

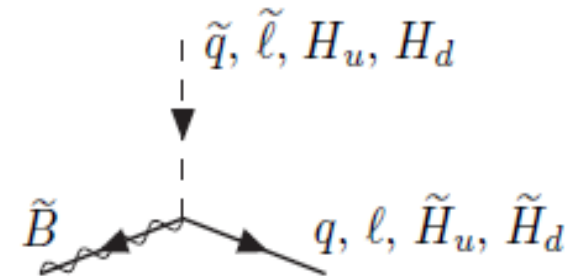
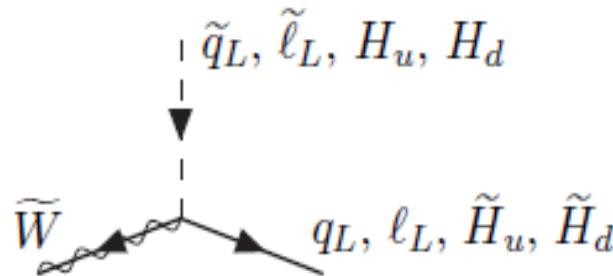
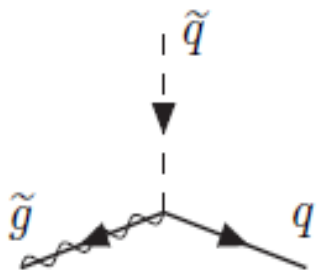
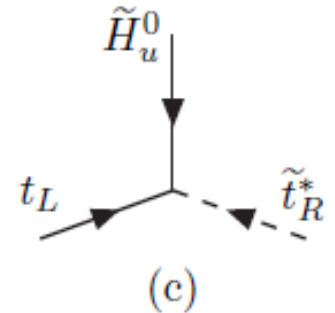
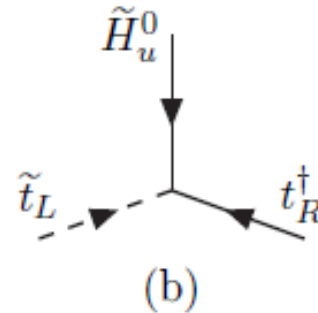
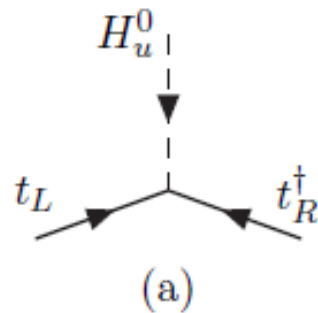
- Only Log divergence with respect to cutoff Λ , much better than quadratic divergence in non-SUSY models.
- However, partners masses cannot be too huge to solve fine tuning problem.

- For $\Lambda = M_{Pl}$, m_{soft} should be less than 1 TeV to solve the fine tuning problem.
- Thus if correct, LHC should discover the superpartners.
- Any reasons **why** only superpartners are not light enough to have already been observed?
- scalars such as sfermions, Higgs can have *gauge-inv.* mass terms $m^2|\phi|^2$ of order of m_{soft}
- Higgsinos, gauginos are in real representation
→ can also have *gauge-inv.* mass terms of order of m_{soft} as well.

Interactions in MSSM

- Apply SUSY transformation to SM vertices \rightarrow MSSM vertices! (Higgs sector considered separately)
- Caution: SUSY particles must appear in **pairs** (R -parity conservation).

e.g. top Yukawa, gauge couplings



Interactions in MSSM

- Some ints. Not determined by gauge int. of SM \rightarrow Higgs sector. All dimensionful couplings depend on μ :

$$- \mathcal{L}_{\text{higgsino mass}} = \mu(\tilde{H}_u^+ \tilde{H}_d^- - \tilde{H}_u^0 \tilde{H}_d^0) + \text{c.c.}$$

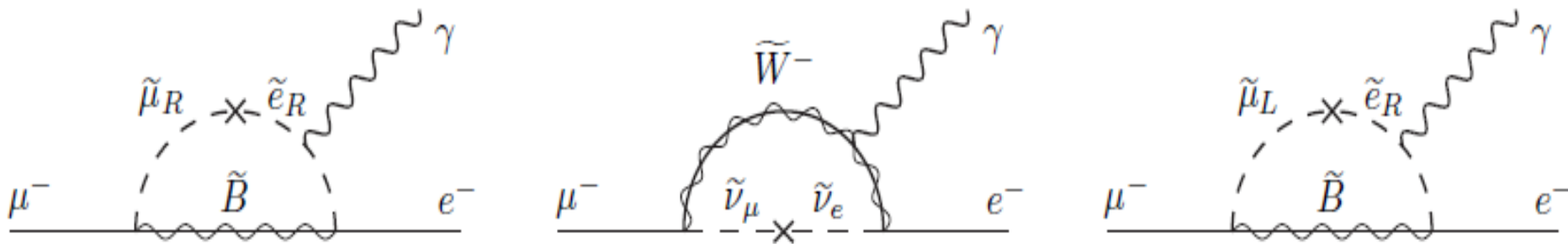
$$- \mathcal{L}_{\text{supersymmetric Higgs mass}} = |\mu|^2 (|H_u^0|^2 + |H_u^+|^2 + |H_d^0|^2 + |H_d^-|^2)$$

$$\mathcal{L}_{\text{supersymmetric (scalar)}^3} = \mu^* (\tilde{u} y_u \tilde{u} H_d^{0*} + \tilde{d} y_d \tilde{d} H_u^{0*} + \tilde{e} y_e \tilde{e} H_u^{0*})$$

- μ should be about 100 GeV to get right EW scale, but why not the M_{Pl} or M_{GUT} ?? \rightarrow μ problem.

Constraints on MSSM (105 new parameters)

- FCNC (Flavor Changing Neutral Current): right-handed slepton can mix, inducing FC processes:

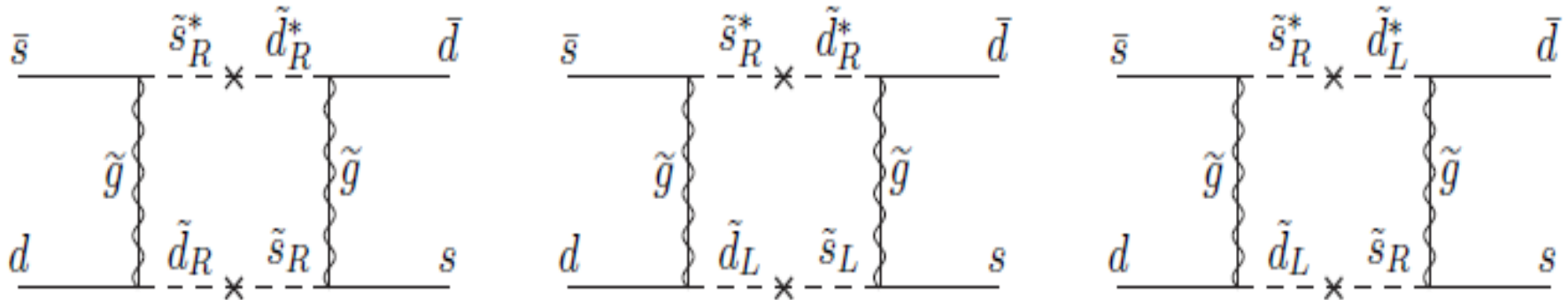


- This process occurs in SM by mixing of μ & e -neutrinos (SUSY transforms the 2nd diagram).
- $\text{Br}(\mu \rightarrow e\gamma)_{\text{exp}} < 1.2 \times 10^{-11} \rightarrow$ suppressed slepton mixing.

Constraints on MSSM

- Also strong constraints on squark mixing from

Kaon mixing: $K^0 \leftrightarrow \bar{K}^0$



- Therefore, MSSM usually assumes NO mixing of squarks, sleptons. \leftarrow *flavor-universal* SUSY breaking!

Constraints from SM

- Soft-SUSY breaking *universality*:

In family space at SUSY breaking scale,

$$m_{\mathbf{Q}}^2 = m_{\mathbf{Q}}^2 \mathbf{1}, \quad m_{\mathbf{U}}^2 = m_{\mathbf{U}}^2 \mathbf{1}, \quad m_{\mathbf{D}}^2 = m_{\mathbf{D}}^2 \mathbf{1}, \quad m_{\mathbf{L}}^2 = m_{\mathbf{L}}^2 \mathbf{1}, \quad m_{\mathbf{E}}^2 = m_{\mathbf{E}}^2 \mathbf{1}$$

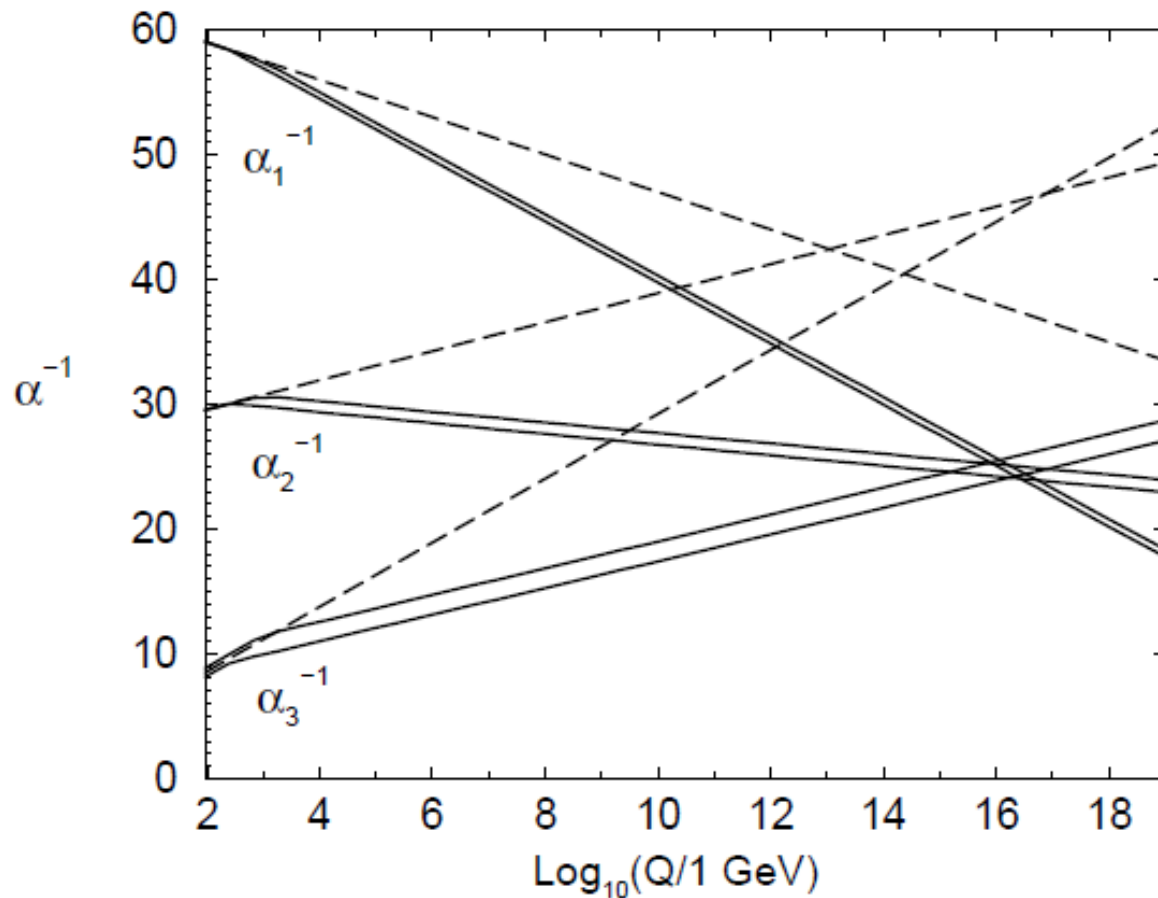
$$\mathbf{a}_{\mathbf{u}} = A_{u0} \mathbf{y}_{\mathbf{u}}, \quad \mathbf{a}_{\mathbf{d}} = A_{d0} \mathbf{y}_{\mathbf{d}}, \quad \mathbf{a}_{\mathbf{e}} = A_{e0} \mathbf{y}_{\mathbf{e}},$$

$$\arg(M_1), \arg(M_2), \arg(M_3), \arg(A_{u0}), \arg(A_{d0}), \arg(A_{e0}) = 0 \text{ or } \pi$$

So that only CKM phases break CP.

- This flavor-blind *ad hoc* conditions require theoretical explanations in the top-down approach.

gaugecouplings
unified better in
MSSM



$$\frac{d}{dt} \alpha_a^{-1} = -\frac{b_a}{2\pi}$$

$$t = \ln(Q/Q_0)$$

*Stephen Martin:
The Supersymmetry
Primer

$$(b_1, b_2, b_3) = \begin{cases} (41/10, -19/6, -7) & \text{Standard Model} \\ (33/5, 1, -3) & \text{MSSM} \end{cases}$$

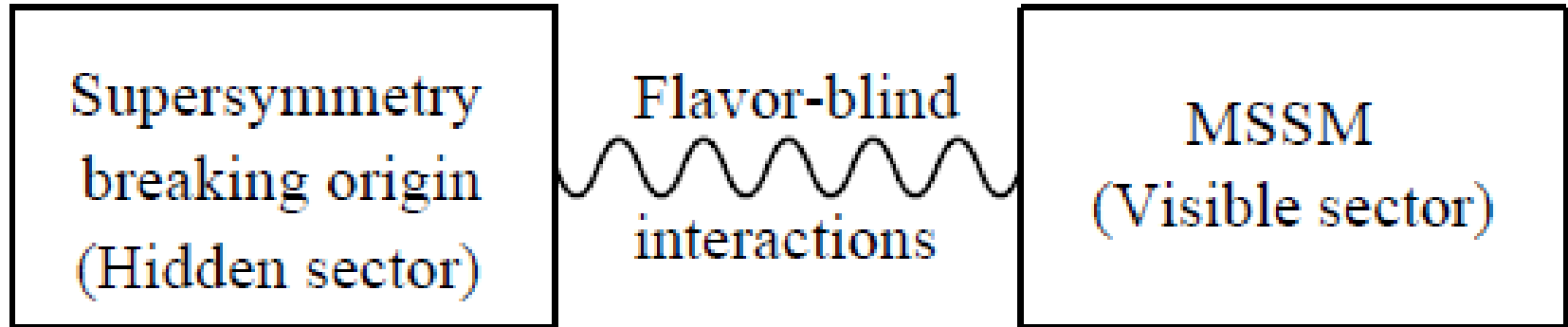
LSP as **DM** candidate

- MSSM has discrete sym called *R-sym*:

$$R = (-1)^{3(B-L)+2S}$$

- positive for SM particles, negative for superpartners.
- superpartners always created in pairs, LSP can**not** into pure SMs.
- LSP with mass few hundred GeVs can serve as cold DM candidate!!

mSUGRA



- Hidden sector breaks SUSY, then gravity mediates the breaking to MSSM sector, resulting in flavor-**blind** or flavor **universal** breaking.
- Can explain conditions mentioned earlier.

mSUGRA

- In generation space at the SUSY breaking scale:

$$M_3 = M_2 = M_1 = m_{1/2},$$

$$m_Q^2 = m_U^2 = m_D^2 = m_L^2 = m_E^2 = m_0^2, \quad m_{H_u}^2 = m_{H_d}^2 = m_0^2,$$

$$a_u = A_0 y_u, \quad a_d = A_0 y_d, \quad a_e = A_0 y_e,$$

$$b = B_0 \mu,$$

- Remaining parameters :

$$m_{1/2}, m_0^2, A_0, B_0, \text{ and } \mu$$

mSUGRA

- Remaining parameters after **EW breaking** :
for $v_u, v_d = \text{vev of } H_u, H_d,$

$$m_0, A_0, m_{1/2}, \tan \beta, \text{ and } \text{sgn}(\mu)$$

$$\tan \beta = v_u/v_d \quad v_d^2 + v_u^2 = 4m_W^2/g^2 \simeq (246 \text{ GeV})^2$$

$$\sin(2\beta) = \frac{2b}{m_{H_u}^2 + m_{H_d}^2 + 2|\mu|^2},$$

Masses at EW scale

$$m_Z^2 = \frac{|m_{H_d}^2 - m_{H_u}^2|}{\sqrt{1 - \sin^2(2\beta)}} - m_{H_u}^2 - m_{H_d}^2 - 2|\mu|^2$$

- All** mass spectrum determined by SM parameters plus these additional 5.

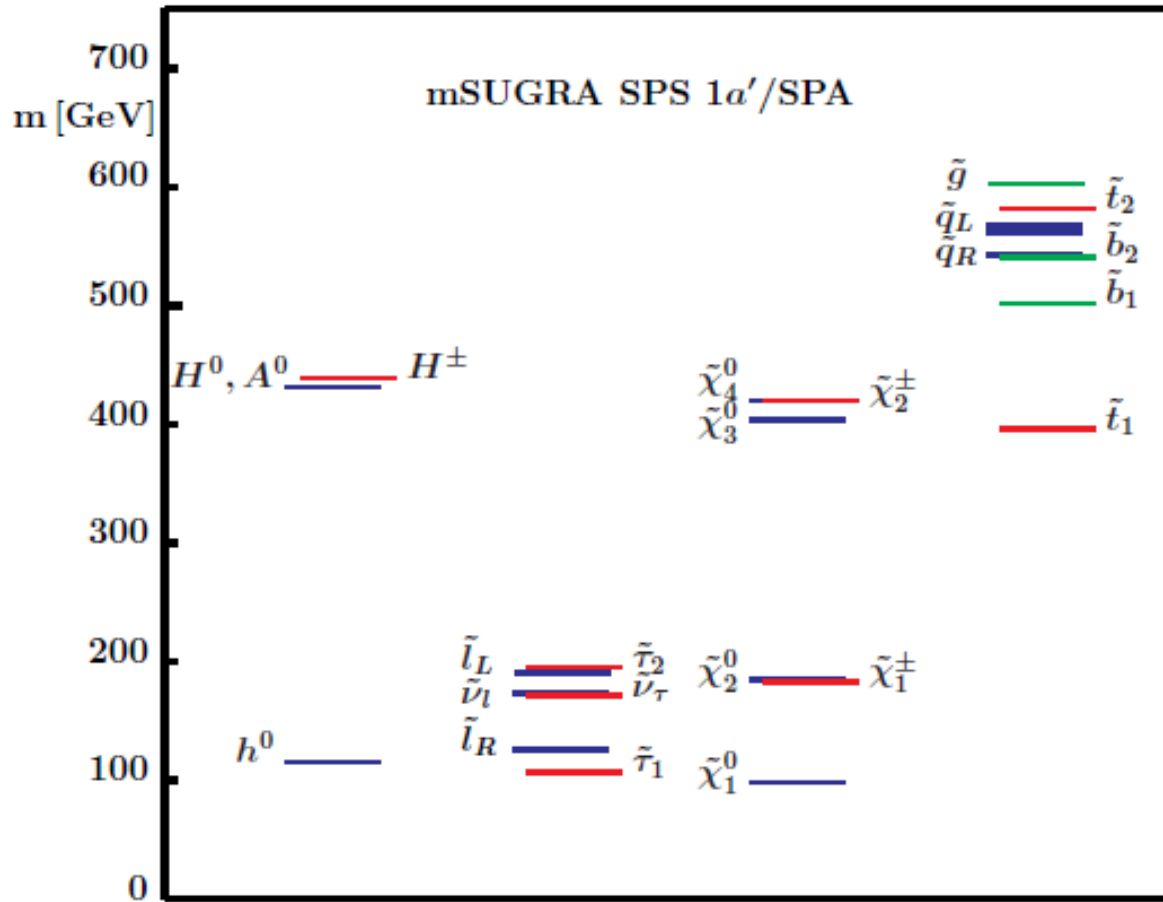
mSUGRA

$$y_b/y_t = (m_b/m_t) \tan \beta$$

- For natural Yukawa coupling, $y_b \sim y_t \rightarrow$ large $\tan\beta$ is preferred. \leftarrow But upper bound from EW breaking conditions exists.
- *Reference points* or benchmark points: SPA=SUSY Parameter Analysis, points in parameter space with realistic values consistent with LSP as CDM and EW precision.

mSUGRA mass spectrum at SPS 1a'/SPA reference point (SUSY scale 1 TeV)

$M_{1/2} =$	250 GeV	$\text{sign}(\mu) =$	+1
$M_0 =$	70 GeV	$\tan \beta(\tilde{M}) =$	10
$A_0 =$	-300 GeV		



- Squarks heavier than sleptons and gauginos
- LSP=lightest neutralino
- Higgs light enough to be found at LHC.

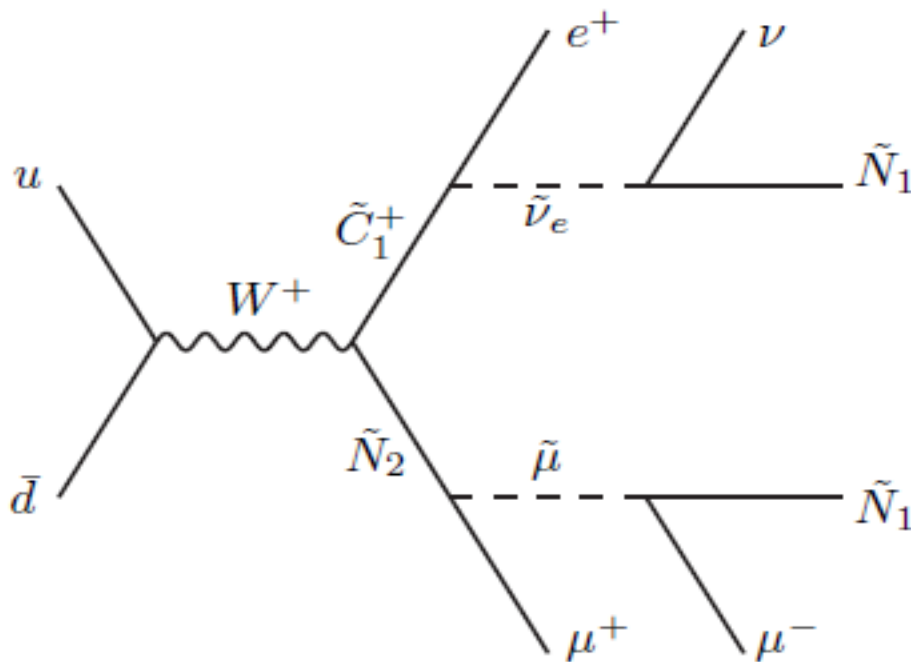
http://arxiv.org/PS_cache/hep-ph/pdf/0511/0511344v2.pdf
for details.

Experimental signatures

- **Constraints** from $\mu \rightarrow e\gamma$, $b \rightarrow s\gamma$, neutral meson mixing, electric dipole moment for e , n , anomalous magnetic moment of muon.
- **Collider** signatures:
- Superpartners produced in pairs then decay to SM particles and invisible LSPs (Lightest superpartners). *e.g.* at hadron colliders

$$\begin{array}{lll} u\bar{d} \rightarrow \tilde{C}_i^+ \tilde{N}_j, & d\bar{u} \rightarrow \tilde{C}_i^- \tilde{N}_j, & gg \rightarrow \tilde{g}\tilde{g}, \quad \tilde{q}_i \tilde{q}_j^*, \\ u\bar{d} \rightarrow \tilde{\ell}_L^+ \tilde{\nu}_\ell & d\bar{u} \rightarrow \tilde{\ell}_L^- \tilde{\nu}_\ell^*, & gq \rightarrow \tilde{g}\tilde{q}_i, \end{array}$$

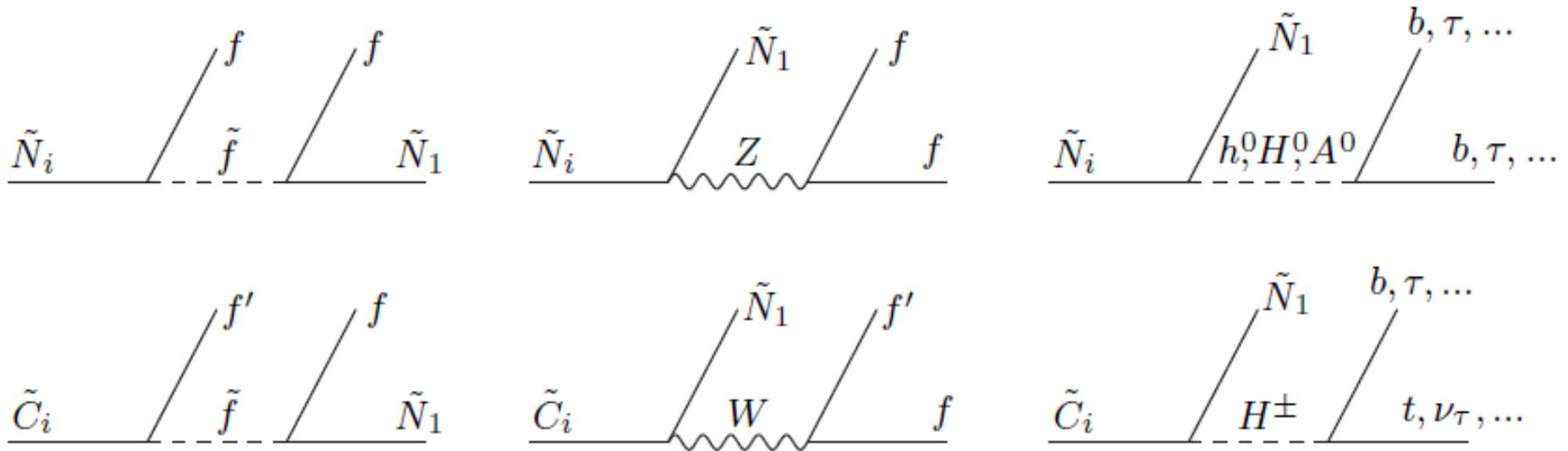
Collider signatures



- Example of charginos, neutralino production. They will finally decay into LSP, assumed to be the lightest neutralino here.
- Looking for $\mu^+\mu^-e^+ + \cancel{E}_T$ with no jets. (*trilepton* signals)

Collider signatures

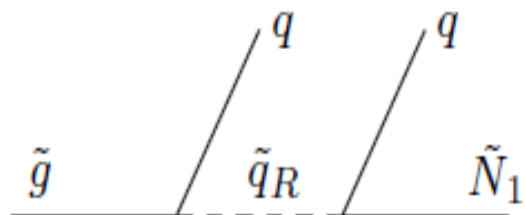
- Many decay channels: chargino, neutralino



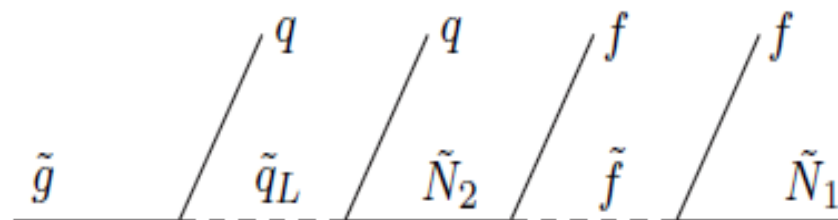
- slepton decays:

$$\tilde{l} \rightarrow l\tilde{N}_i, \quad \tilde{l} \rightarrow \nu\tilde{C}_i, \quad \tilde{\nu} \rightarrow \nu\tilde{N}_i, \quad \tilde{\nu} \rightarrow l\tilde{C}_i$$

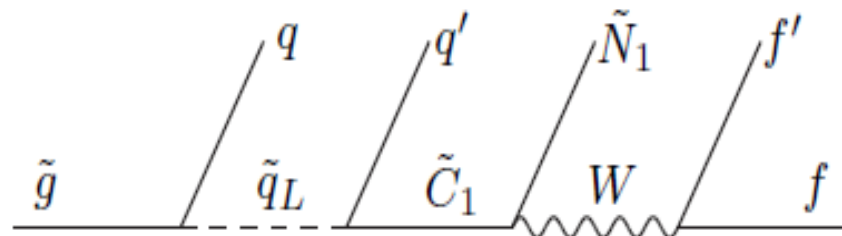
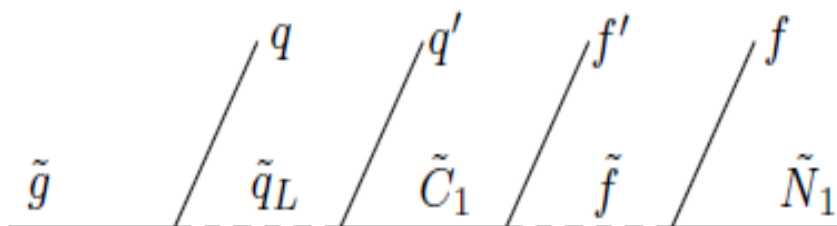
- Gluino decays:



(a)



(b)



- Lightest Higgs possible to be discovered at LHC:

Very large
Gluon fusion

$$gg \rightarrow h^0$$



$$h^0 \rightarrow \gamma\gamma$$

$$h^0 \rightarrow ZZ^{(*)} \rightarrow e^+e^-e'^+e'^-$$

$$h^0 \rightarrow WW^{(*)} \rightarrow e^+\nu e'^-\bar{\nu}$$

MSSM Higgs at LHC

- Other channels: weak bosons fusion, W, Z radiated from quarks (forward jets):

$$W^+W^- \rightarrow h^0 \text{ and } ZZ \rightarrow \bar{h}^0$$

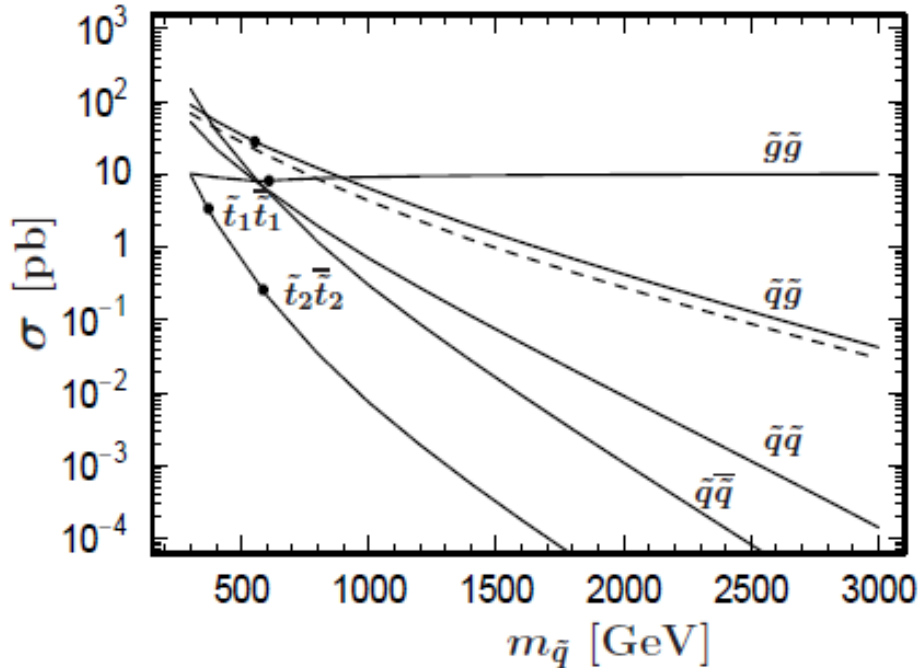
$$pp \rightarrow h^0 jj \rightarrow \tau^+\tau^- jj \text{ or } WW^{(*)} jj$$

- Also because of large Yukawa coupling:
Higgs radiated from a top

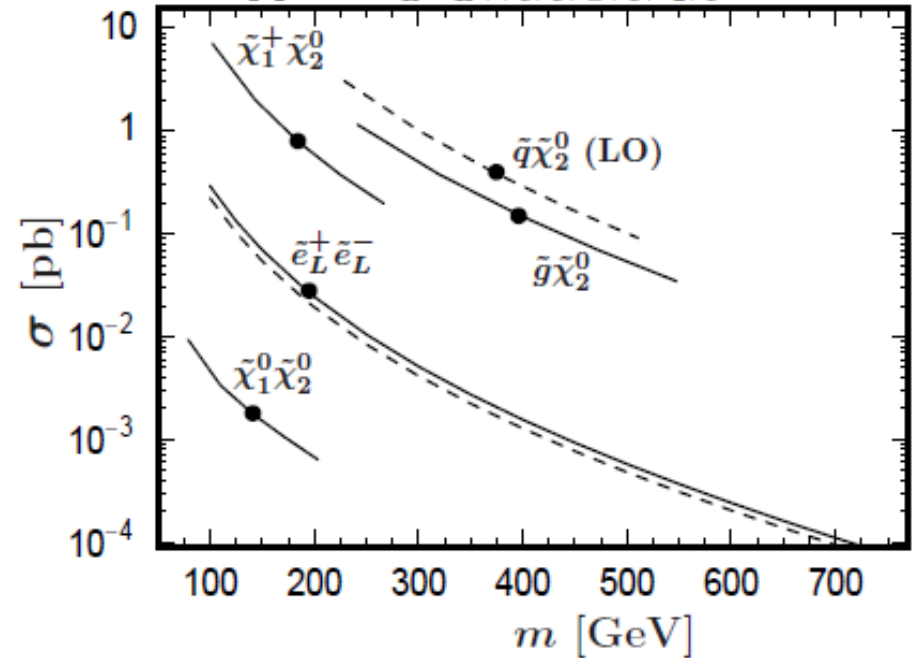
$$pp \rightarrow t\bar{t}h^0 \rightarrow t\bar{t}b\bar{b}.$$

Some hypotheticalal plots at LHC

$pp \rightarrow \tilde{q}\tilde{q}, \tilde{q}\tilde{\bar{q}}, \tilde{q}\tilde{g}, \tilde{g}\tilde{g}, \tilde{t}_i\tilde{\bar{t}}_i + X$



$pp \rightarrow \tilde{e}_L^+\tilde{e}_L^-, \tilde{\chi}\tilde{\chi}, \tilde{g}\tilde{\chi}, \tilde{q}\tilde{\chi} + X$



http://arxiv.org/PS_cache/hep-ph/pdf/0511/0511344v2.pdf

for details.

NP lists

More and more and much more.

- Little Higgs, Fat Higgs, SUSY in extra-D
- Strong dynamics: technicolor models, preons
- Flavor gauge theories, horizontal symmetries
- NMSSM, mmSUGRA (gravitino variations), GMSB, AMSB → more parameters!
- String models, low-scale string pheno (e.g. P. Burikham et. al. 😊) TeV-scale BH, string balls etc.
- LHC Pb-Pb collision → Quark-Gluon Plasma*

References

- PDG (Particle Data Group) and references therein, <http://pdg.lbl.gov/>
- SPIRES, <http://www.slac.stanford.edu/spires/> to search for articles and references.