

# FORMOSA & Forward-DUNE - Looking Forward to Millicharged Dark Sectors and New Neutrino Physics

### Yu-Dai Tsai, Fermilab / U Chicago

Saeid Foroughi, Felix Kling, Yu-Dai Tsai, arXiv:2010.07941

"Tau & Heavy Neutrino Electric Dipole Moment," Foroughi, Kling, Tsai, (see link!)

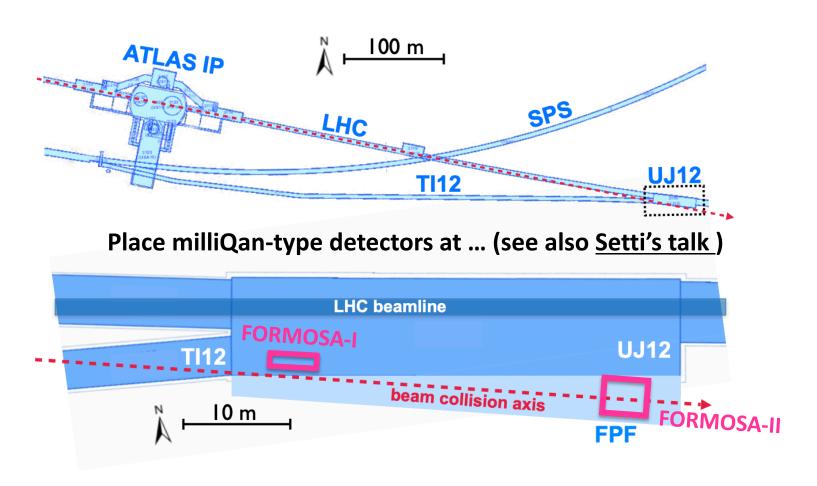
Forward proto-DUNE, Kling, Tsai (+ Feng, Foroughi, Cavanna), one slide

Email: ytsai@fnal.gov; arXiv: https://arxiv.org/a/tsai y 1.html

# When Energy meets Intensity

- The High-Intensity Energy Frontier: FPF
- General LLP: FASER
- Specified (MCP): FORMOSA
- Neutrino: FASERv + Forward-DUNE
- Filling low-mass / high-mass gap (dark sector)
- low-energy / high-energy gap (neutrino)

### FORMOSA: FORward MicrOcharge SeArch



Formosa means "beautiful" in Portuguese and is the ancient name of Taiwan

### Outline

- Intro of millicharged particles (MCP)
- FORMOSA and its sensitivity reach
- Probing the strongly-interacting millicharged dark matter (SIMDM) window
- New idea: Forward (proto-)DUNE

### Millicharged Particle & Dark Matter

# Finding Minicharge

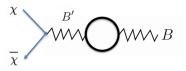
- Is electric charge quantized and why? A long-standing question!
- SM U(1) allows arbitrarily small (any real number) charges. Why don't we see them? Motivates Dirac quantization, Grand Unified Theory (GUT), to explain such quantization (anomaly cancellations fix the SM  $U(1)_V$  charge assignments)
- MCP (not confined) is predicted by Superstring theories:
   Wen, Witten, Nucl. Phys. B 261 (1985) 651-677, Youtube: <a href="link">link</a>
   Link to string compactification and quantum gravity (Shiu, Soler, Ye, PRL '13)
- Conservatively, testing if e/3 is the minimal charge
- MCP could have natural link to dark sector (through dark photon, etc)
- Could account for (sometimes fractional) dark matter (DM) abundance;
   Used for the cooling of gas temperature to explain the EDGES anomaly
   [EDGES collab., Nature, (2018); Barkana, Nature, (2018)]. Refs in backup slides.
- Review of Accelerator Probes of MCP: TODAY at 4 pm Central Time! [link]

### MCP Model

A particle fractionally (or irrationally) charged under SM U(1)

hypercharge, 
$$\mathcal{L}_{ ext{MCP}} = i ar{\chi} ( \partial \!\!\!/ - i \epsilon' e \!\!\!/ \!\!\!/ + M_{ ext{MCP}} ) \chi$$

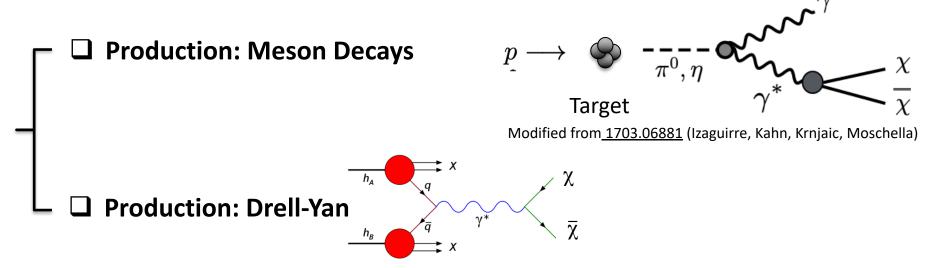
- $\epsilon'$  can in principle be arbitrarily small.
- Can just consider these Lagrangian terms by themselves (no extra mediator, i.e., dark photon). Completely legal!
   Naively violating the empirical charge quantization.
- We are simply search for MCP!
   Minimal assumptions = most robust constraints/probes.
- This could come from vector portal Kinetic Mixing (Holdom, '85)
  - a nice origin to the above terms
  - help give rise to dark sectors: a dark photon model
  - easily compatible with **Grand Unification Theory**
  - see backup slides



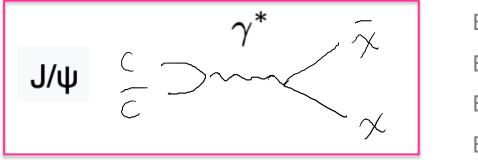
# **Important Notes!**

- Our search is simply a search for particles (**fermion**  $\chi$ ) with  $\{$ mass, electric charge $\} = \{m_{\chi}, \epsilon e\}$
- Minimal theoretical inputs/parameters
   (need high-energy/intensity machines to probe in MeV –
   GeV+ mass regime)
- Dark photon not necessary for MCP particles (Neff, Self-Int.
   could provide extra constraints)
- Similar bound/sensitivity applies to scalar MCPs

### **Important Production Channels of MCP**



☐ Heavy (vector) mesons are important for high-mass mCP's in high-energy beams



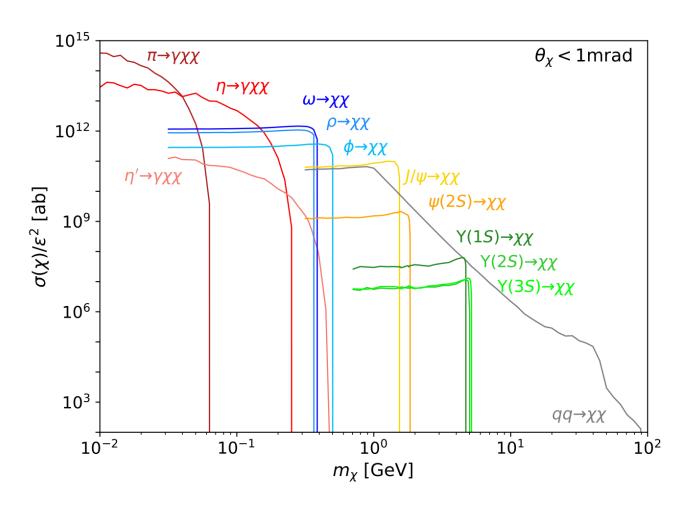
$$BR(\pi^0 \rightarrow 2v) = 0.99$$

$$BR(\pi^0 \rightarrow \gamma e^- e^+) = 0.01$$

$$BR(\pi^0 \rightarrow e^- e^+) = 6 * 10^{-6}$$

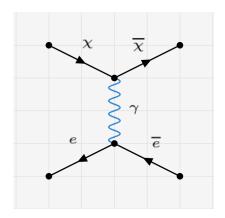
$$BR(J/\psi \to e^-e^+) = 0.06$$

### **MCP Production/Flux**



Saeid Foroughi, Felix Kling, Yu-Dai Tsai, <u>arXiv:2010.07941</u> **Flux much more enhanced comparing to the transverse region** 

### MCP Detection: Ionization or Hard Scattering



- MCP's & light-mediator scattering: favors scattering with small energy transfer
- Different from Neutrino (exchange heavy boson), MCP exchange photon
- studying ionization / scintillation or measuring low-energy electron recoils
- Ionization (eV-level): ~ very low-energy scattering: (triple / quadruple scintillation)
   MilliQan: arXiv:1410.6816, Haas, Hill, Izaguirre, Yavin
- "Hard" (MeV-level) electron elastic scattering: (single / double hit)
   see, e.g., Magill, Plestid, Pospelov, Tsai, 1806.03310 (MCP in neutrino Experiments)

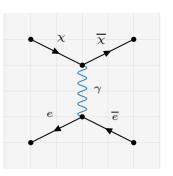
### MCP Detection: Electron Scattering

- $Q^2$  is the squared 4-momentum transfer.
- lab frame:  $Q^2 = 2m_e (E_e m_e)$ ,  $E_e m_e$  is the electron recoil energy.
- Expressed in **recoil energy threshold**,  ${E_e}^{(min)}$ , we have

$$\sigma_{e\chi} \simeq 2.6 \times 10^{-25} \text{cm}^2 \times \epsilon^2 \times \frac{1 \text{ MeV}}{E_e^{(\text{min})} - m_e}.$$

Sensitivity greatly enhanced by accurately measuring low energy

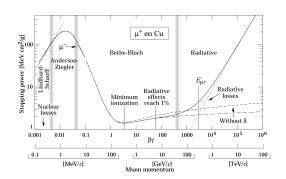
electron recoils for MCP's & light-mediator scattering



### MCP Detection: Ionization

- Want very low momentum transfer: ionization and scintillation signature
- Signature proportional to dE/dx of the MCP, referred to as energy loss/stopping power
- Can be approximated with the Bethe-Bloch Formula (various modified versions and detailed considerations.)

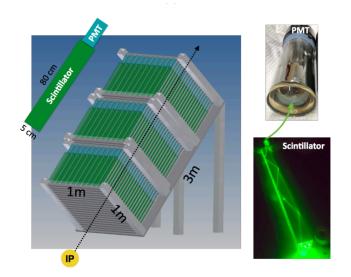
$$\left\langle -\frac{dE}{dx} \right\rangle \propto \epsilon^2.$$

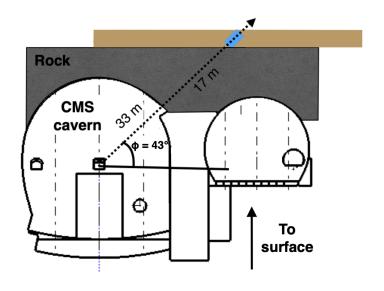


intentionally make the plot small so we don't get into too much details of this. <a href="http://pdg.lbl.gov/2020/reviews">http://pdg.lbl.gov/2020/reviews</a>

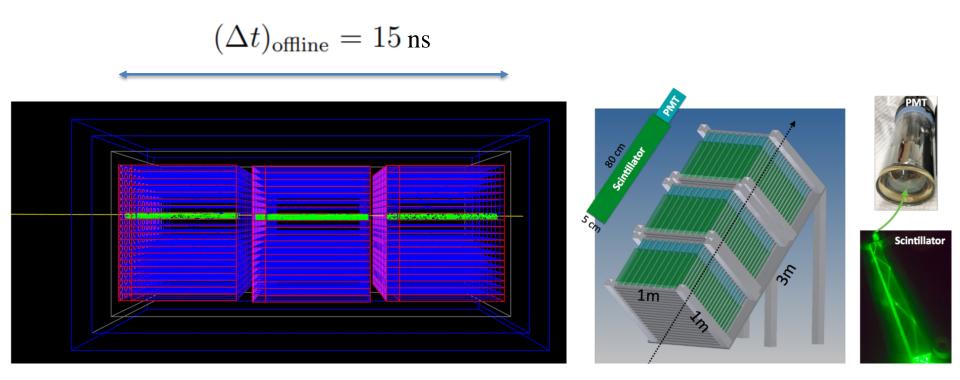
# MilliQan

- MilliQan: arXiv:1410.6816, Haas, Hill, Izaguirre, Yavin
- Also will show an update from the milliQan-prototype!
- From milliQan Collaboration:
- LOI: <u>1607.04669</u>
- First result update:
   2005.06518
   (will get into this more)





# MCP Detector Concept



1607.04669, (MilliQan Collaboration)

Ball, Brooke, Campagnari, <u>De Roeck</u>, Francis, Gastal, Golf, Goldstein, **Haas**, **Hill**, **Izaguirre**, Kaplan, Magill, Marsh, Miller, Prins, Shakeshaft, Stuart, Swiatlowski, **Yavin** 

### Photoelectrons (PE) from Scintillation

 The averaged number of photoelectron (PE) seen by the detector from single MCP is:

$$N_{PE} \propto \left\langle -\frac{dE}{dx} \right\rangle \times l_{scint}, \ \left\langle -\frac{dE}{dx} \right\rangle \propto \epsilon^2.$$

 $\langle dE/dx \rangle$  is the "mass stopping power" (PDG 2018)

One can use modified **Bethe-Bloch Formula** to get an approximation

- $N_{PE} \sim \epsilon^2 \times 10^6 \text{ for 1}$  meter plastic scintillation bar
- $\epsilon \sim 10^{-3}$  roughly gives one PE



### **Signature: Triple Coincidence**

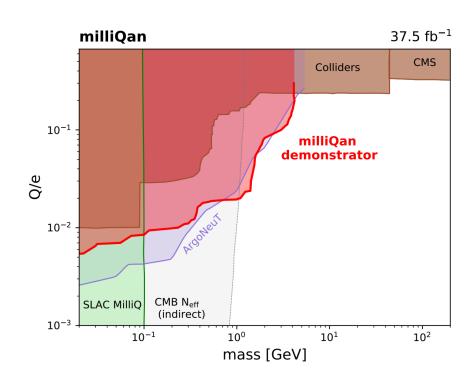
 Based on Poisson distribution, zero event in each bar correspond to

 $P_0 = e^{-N_{PE}}$ , so the probability of seeing triple incident of one or more photoelectrons is:

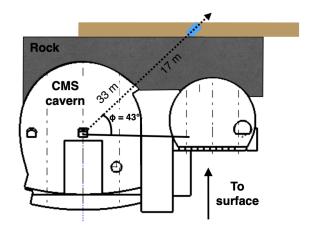
$$P = \left(1 - e^{-N_{PE}}\right)^3$$

•  $N_{x,detector} = N_x$  (going through detector) x P.

### Proto-milliQan Update!

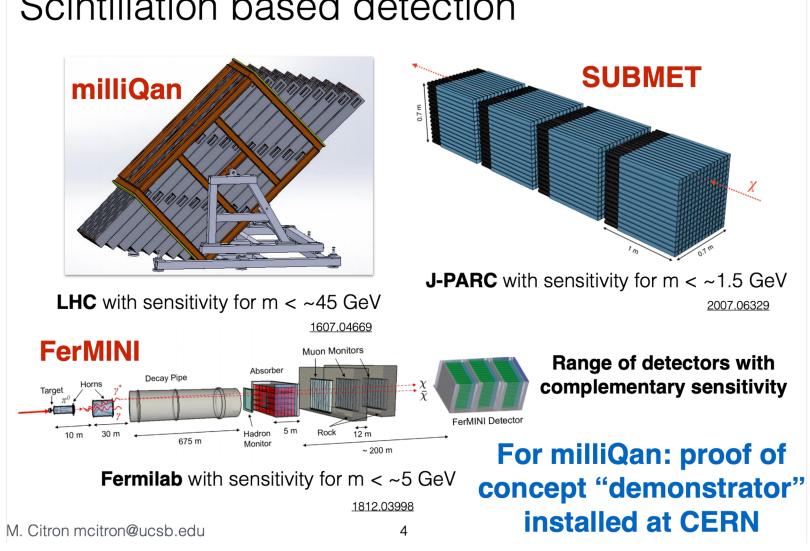


MilliQan prototype: updates on the results <a href="https://arxiv.org/pdf/2005.06518.pdf">https://arxiv.org/pdf/2005.06518.pdf</a> (milliQan Collaboration)



- ~1% of the full milliQan (total of 18 bars)
- LHC '18 of 37.5 fb-1 at a center-of-mass energy of 13 TeV.
- A prototype scintillator-based detector is deployed to conduct the first search at a hadron collider
- 20 and 4700 MeV is excluded at 95% confidence level for charges between 0.006e and 0.3e

### Scintillation based detection



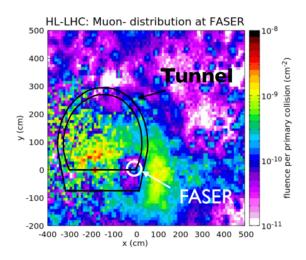
Directly from **Matthew Citron**'s talk at NF3 kickoff meeting: link

### A New Challenge in the Forward Region

#### **Stronger Beam related background:**

- New challenge arises due to large flux of **High-Energy Muons** (and secondary particles) from the beam collisions
- **FLUKA simulation** performed by CERN STI group estimated muon flux <  $\sim$  1 Hz/cm2 resulting in roughly  $\sim$  one muon every **100**  $\mu s$
- Feasible task: implementing an online-veto of large-PE events

But this is not the full story ...

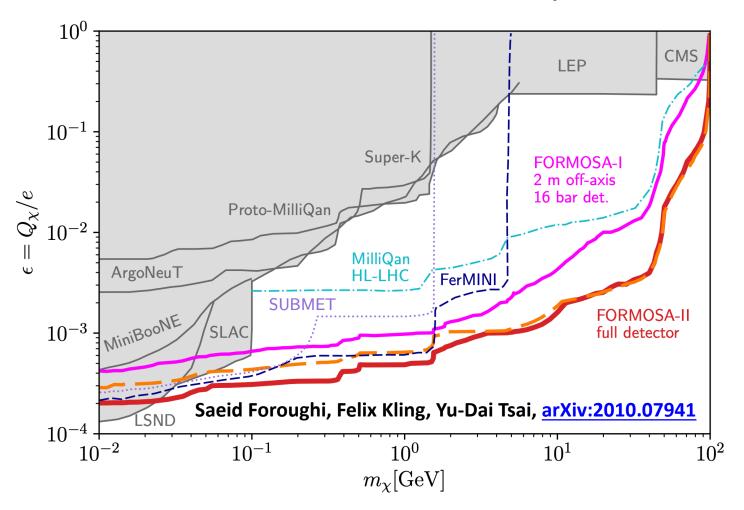


- lacktriangled Afterpulses: small pulses occurring with a delay time of  $\delta t \lesssim 10~\mu s$  after the initial pulse
- FASER Collaboration, arXiv:1812.09139

- ❖ Remove the afterpulse background by vetoing ~10% of the data
- Better PMTs with reduced afterpulse duration can also improve the live-time efficiencies

Thank Saeid, Matthew, and Felix for help preparing this slide.

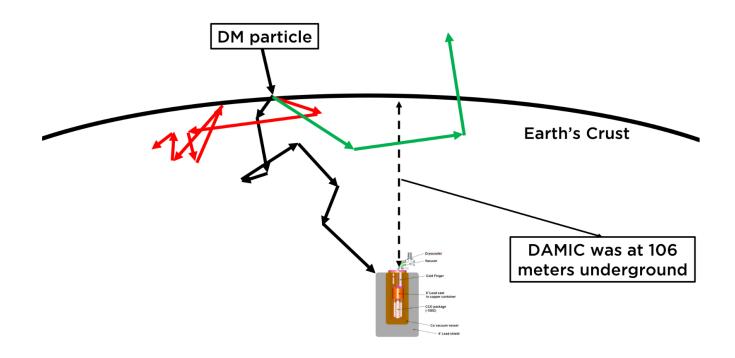
### **FORMOSA Sensitivity**



FORMOSA-I:  $\sim$  0.2 m  $\times$  0.2 m  $\times$  4 m; 4 layers of 16 scintillator bars @UJ12/TI12 tunnel. (can use the milliQan demonstrator:  $\sim$  5 cm  $\times$  5 cm  $\times$  3 m; 3 layers of 6 scintillator bars ) FORMOSA-II:  $\sim$  1 m  $\times$  1 m  $\times$  4 m; 4 layers of 400 scintillator bars @ FPF.

# Strongly Interacting Dark Matter

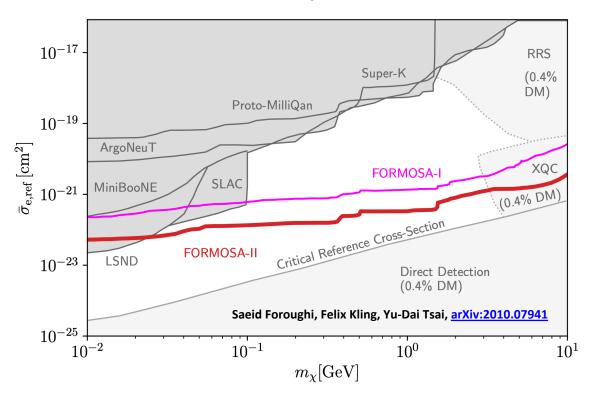
DM-SM Interaction too strong that attenuation stop the particles from reach the direct detection detector



**DMATIS (Dark Matter ATtenuation Importance Sampling),** Mahdawi & Farrar '17 Also, see Mahdawi & Farrar, arXiv:1804.03073

### Probe of Millicharged Dark Matter

$$\bar{\sigma}_e \simeq \frac{16\pi\alpha^2\epsilon^2\mu_{\chi e}^2}{q_{ref}^2}, \ q_{ref} = \alpha m_e$$



- Here we plot the critical reference cross-section see <u>1905.06348</u> (Emken, Essig, Kouvaris, Sholapurkar)
- Cosmic-Ray Production/Super-K Detection <u>2002.11732</u>
- FORMOSA can help close the Millicharged SIDM window!

### **Looking Ahead**

### **Action plan:**

- Run 3: placing milliQan demonstrator:
  - ~ 5 cm × 5 cm × 3 m; 3 layers of 6 scintillator bars @UJ12/TI12 tunnel.
- FORMOSA-I:
  - $\sim 0.2 \text{ m} \times 0.2 \text{ m} \times 4 \text{ m}$ ; 4 layers of 16 scintillator bars @UJ12/TI12 tunnel.
- FORMOSA-II:
  - $\sim 1$  m  $\times 1$  m  $\times 4$  m; 4 layers of 400 scintillator bars @ FPF.

### Other new ideas? Neutrino physics?

### Forward Proto-DUNE (Forward-DUNE!)

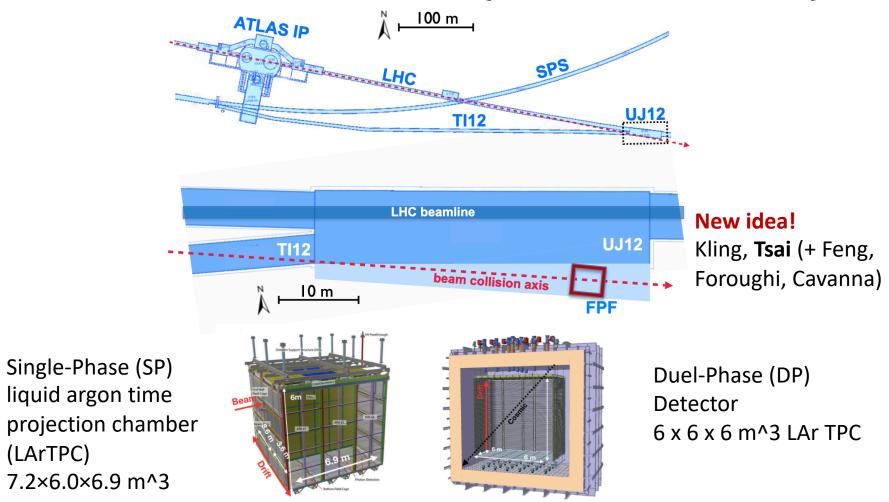


Figure 3: Left: draft of ProtoDUNE-SP [2]. Right: draft of ProtoDUNE-DP [3]

DUNE Collaboration (arXiv:1706.07081 + arXiv:1409.4405) Updates, see, e.g. arXiv:1910.10115 & arXiv:2007.06722

### Thank you!

Special thanks to my collaborators and the organizers.

Thank Albert for the invitation!

Para uma formosa aventura!

### Review: Accelerator Probes of MCP

- HEP/Astro Results Forum talk, "Accelerator Probes on Millicharged Particles and Dark Matter".
- HEP/Astro Results Forum website: <a href="https://sites.google.com/site/lhcresultsforumtalks/">https://sites.google.com/site/lhcresultsforumtalks/</a>
- Link to subscribe to this mailing list:
   <a href="https://utlists.utexas.edu/sympa/subscribe/lhcresultsforum">https://utlists.utexas.edu/sympa/subscribe/lhcresultsforum</a>

The relevant literatures are growing fast, please let us know if we miss mentioning your important works.

# Backup Slides and Information

### More on MCP/DM & 21-cm Cosmology

Some more reference of Millicharged DM (mDM) and constraints.

See, e.g.,

McDermott, Yu, Zurek, 1011.2907;

EDGES explanation to  $\sim 100$  GeV.

Muñoz, Dvorkin, Loeb, 1802.10094, 1804.01092;

Berlin, Hooper, Krnjaic, McDermott, 1803.02804;

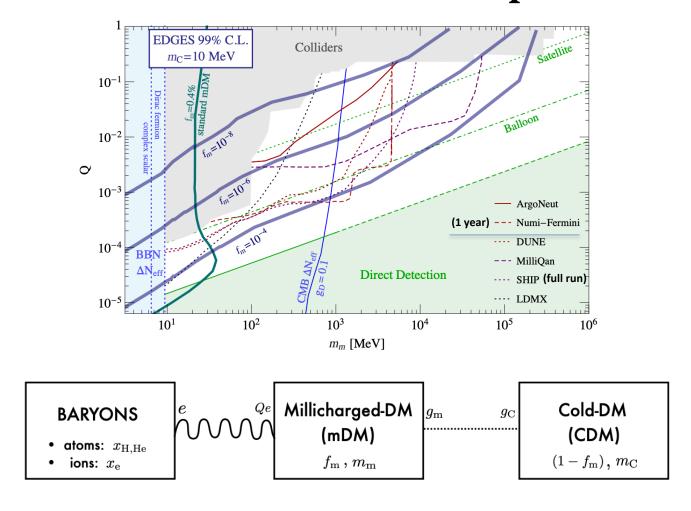
Kovetz, Poulin, Gluscevic, Boddy, Barkana, Kamionkowski, 1807.11482;

Liu, Outmezguine, Redigolo, Volansky, 1908.06986:

"Reviving Millicharged Dark Matter for 21-cm Cosmology,"

Introduces a long-range force between a subdominant mDM and the dominant cold dark matter (CDM) components. Leads to efficient cooling of baryons in the early universe. Extend the range of viable mDM masses for

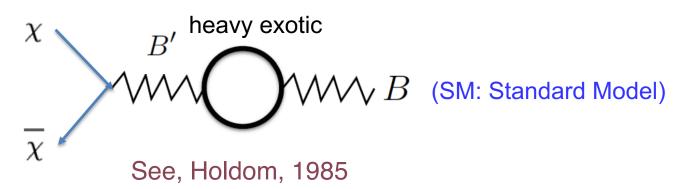
# MDM as an EDGES Explaner



Liu, Outmezguine, Redigolo, Volansky, '19 For more refs, see backup slides

# Kinetic Mixing and MCP Phase

 Coupled to new dark fermion χ



$$\mathcal{L} = \mathcal{L}_{SM} - \frac{1}{4}B'_{\mu\nu}B'^{\mu\nu} - \frac{\kappa}{2}B'_{\mu\nu}B^{\mu\nu} + i\bar{\chi}(\partial \!\!\!/ + ie' \!\!\!\!/ B' + iM_{MCP})\chi$$

- New fermion χ charged under new gauge boson B'.
- Millicharged particle (MCP) can be a low-energy consequence
  of massless dark photon (a new U(1) gauge boson) coupled to
  a new fermion (become MCP in a convenient basis.)
- Important Notes: Our search is simply a search for particles (fermion  $\chi$ ) with {mass, electric charge} =  $\{m_{\chi}, \epsilon e\}$

### Accelerator Probes of MCP: Reference

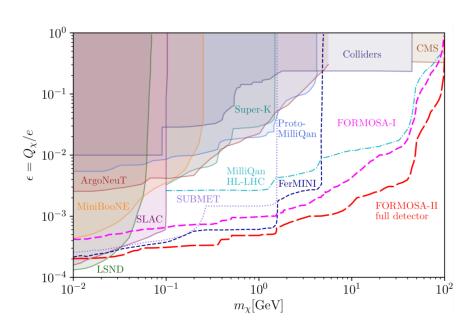
#### Works from other people and collaborations:

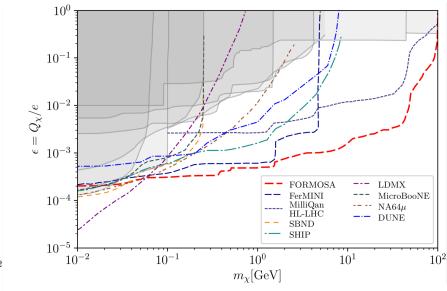
- SLAC mQ (hep-ex/9804008)
- Early collider bounds (LEP+) (hep-ph/0001179)
- Invisible decays of positronium (hep-ex/0609059)
- LHC CMS (1210.2311)
- milliQan (1410.6816 + 1607.04669 + 2005.06518)
- LDMX MCP probe (1807.01730)
- NA64 MCP Probe (1810.06856)
- LArTPC's (1902.03246) / ArgoNeuT (1911.07996)
- MAPP (conf. proceeding, LHCP 2019)
- Electron colliders (1909.06847)
- SUBMET (2007.06329)
- ...

#### Our works:

- MCPs in neutrino experiments (LSND, MiniBooNE, DUNE) (1806.03310)
- FerMINI (1812.03998), <u>Fermilab today</u>
- Cosmic-Ray production and detection in Super-K (2002.11732)
- Forward MCP Search (coming ~ next week)
- My review talk: <a href="https://indico.cern.ch/event/910753/contributions/3831616/">https://indico.cern.ch/event/910753/contributions/3831616/</a>

# Compilation





- Our search is simply a search for particles (fermion  $\chi$ ) with {mass, electric charge} =  $\{m_{\chi}, \epsilon e\}$
- The result for scalar MCP is similar
- $\epsilon \equiv Q_{\chi}/e$ .