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CENTER FOR
UNDERGROUND PHYSICS

A search for ${}^7\text{Li}$ solar axions with Li_2MoO_4 detectors in the AMoRE experiment

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INTRODUCTION

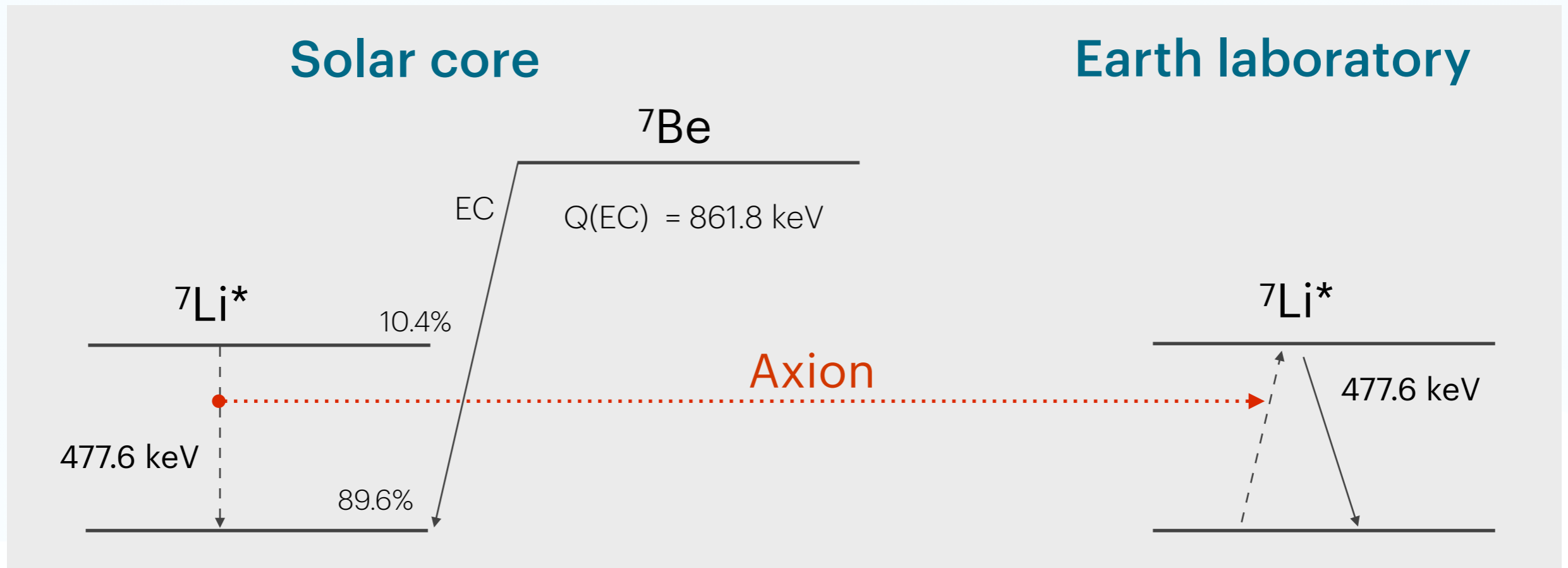
- Axion is a hypothetical particle considered as one of the candidates for the dark matter particles.
- Monoenergetic solar axions emitted in magnetic transitions by an excited state of ${}^7\text{Li}$ in the Sun could excite the same nuclei on the Earth to the corresponding level resonantly [1].
- Detection of subsequent gamma rays would be a sign of solar axion existence.
- This method is free from the uncertainty of the axion-photon coupling because both the emission and absorption occur not via the axion-photon coupling but via the axion-nucleon coupling [2].
- Latest result of axion mass limit of 8.6 keV at 90% C.L. [3].

[1] Shigetaka Moriyama, *Phys. Rev. Lett.* **75**, 3222 (1995).

[2] M.Krcmar et al., *Phys. Rev.* **D64**, 115016 (2001).

[3] P.Belli et al., *Phys. Lett.* **B711**, 41 (2012).

SEARCH FOR SOLAR AXIONS USING ${}^7\text{Li}$



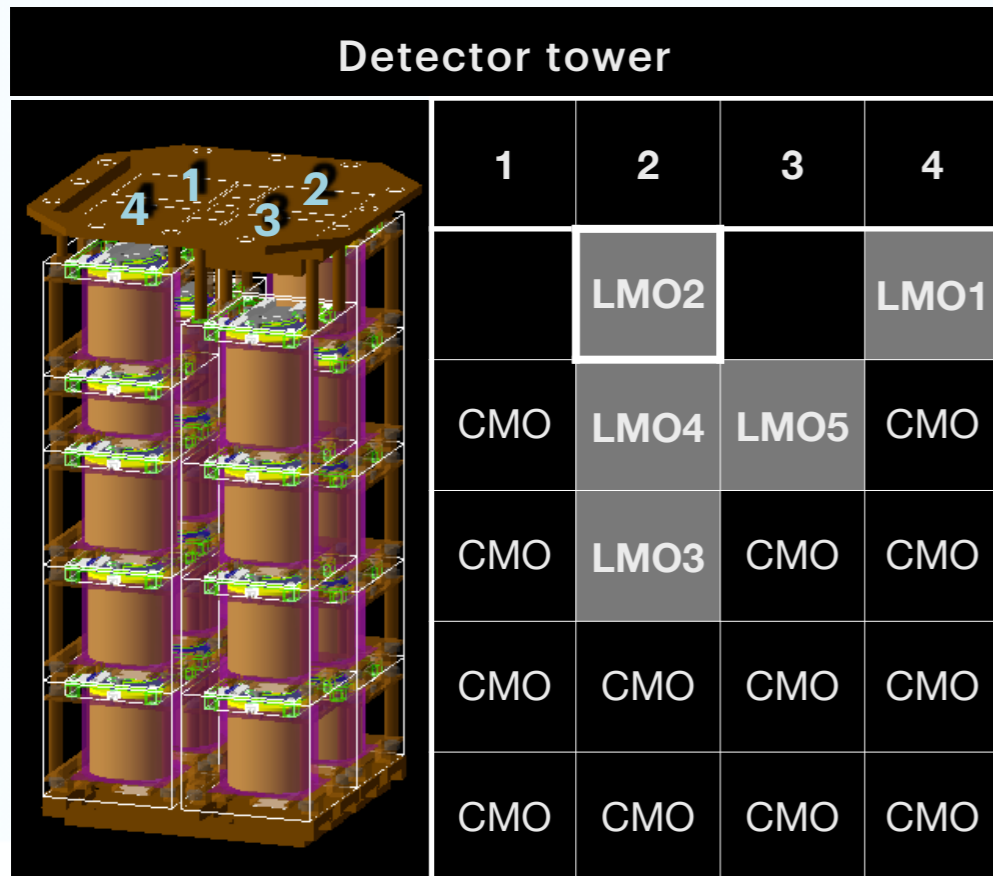
- Why ${}^7\text{Li}$?: populated in the pp chain by the electron capture decay of ${}^7\text{Be}$.
- The total number of resonant axion absorption process in ${}^7\text{Li}$:

$$R = N_7 \times t \times 1.74 \times 10^{-45} \times \left(\frac{m_a}{1\text{eV}}\right)^4$$

Derbin, JETP Letters, 81, 365 (2005).

(N_7 : the number of ${}^7\text{Li}$ in the target, t : measurement time in second)

AMORE-I DETECTOR



Area of γ peak at 477.6 keV

$$m_a = 1.55 \times 10^{11} \times \left(\frac{S}{\epsilon N_7 t} \right)^{\frac{1}{4}} eV$$

Detection efficiency

Measurement time

of ^7Li in the target

- Target: 1.6 kg of Li_2MoO_4 (LMO) crystals
- # of ^7Li in target: 1.03×10^{25} ^7Li

Full peak detection efficiency (for 5 LMOs: 13.5%)

LMO1 (13.92%)

LMO2 (13.95%)

LMO3 (15.65%)

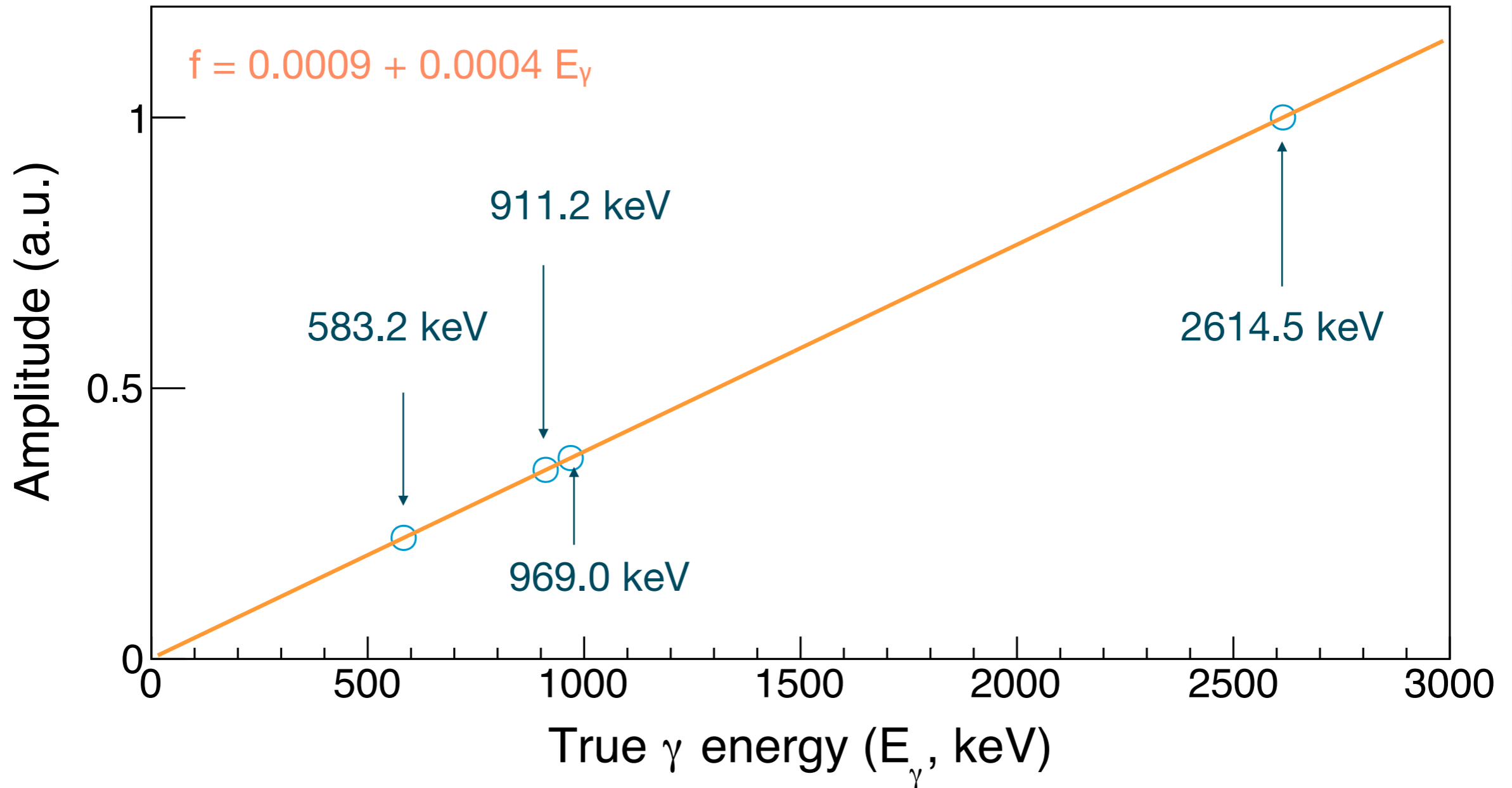
LMO4 (15.33%)

LMO5 (15.66%)

	0.13		12.26			12.38		0.13			0.03		0.02			0.52		0.05			0.27		0.27
0.35	0.06	0.33	0.50	0.30	0.52	0.35	0.03	0.02	0.12	0.51	0.11	0.02	0.41	12.39	0.44	0.06	0.15	0.35	13.36	0.19	0.06	0.09	0.21
0.07	0.02	0.09	0.09	0.07	0.04	0.08	0.39	12.43	0.53	0.17	0.27	0.52	0.33	0.12	0.06	0.03	0.02	0.02	0.01	0.04	0.06	0.09	0.21
0.01	0.00	0.01	0.00	0.02	0.00	0.01	0.02	0.00	0.01	0.01	0.25	0.52	0.28	0.09	0.07	0.02	0.06	0.03	0.02	0.01	0.04	0.02	0.01
0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.07	0.04	0.07	0.02	0.07	0.04	0.07	0.02	0.02	0.00	0.01	0.01	0.01	0.00	0.01	0.01	0.01

Near detector

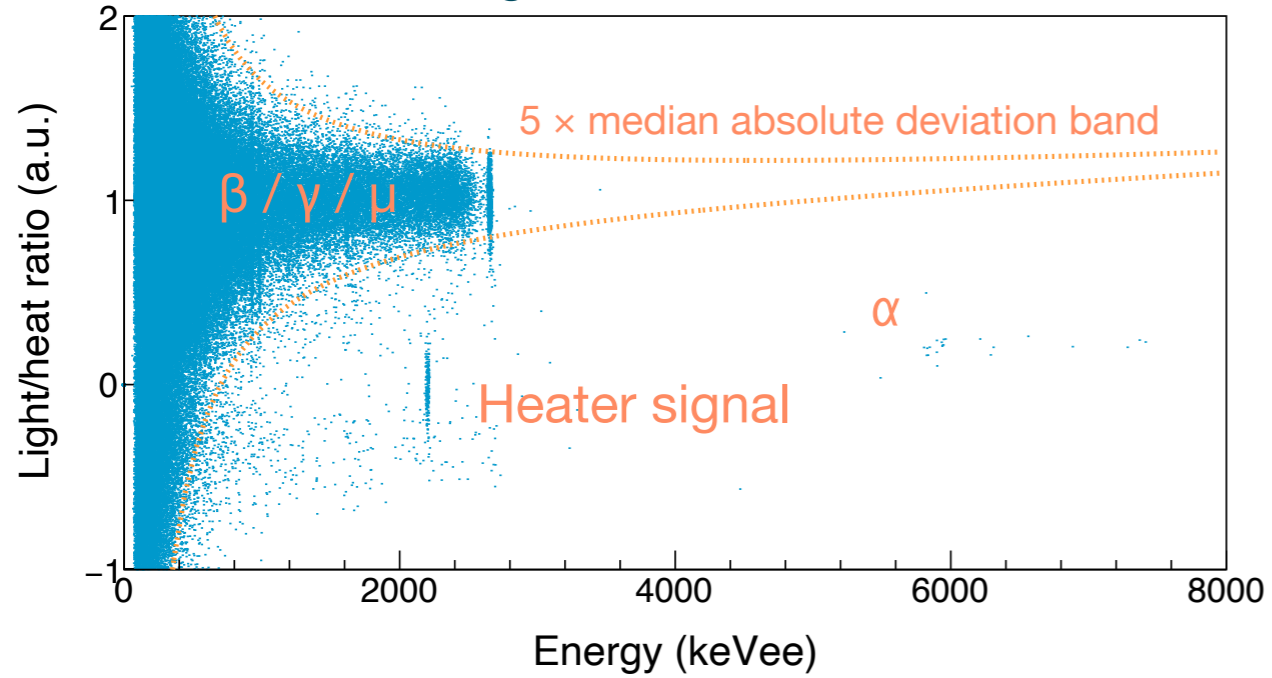
ENERGY CALIBRATION



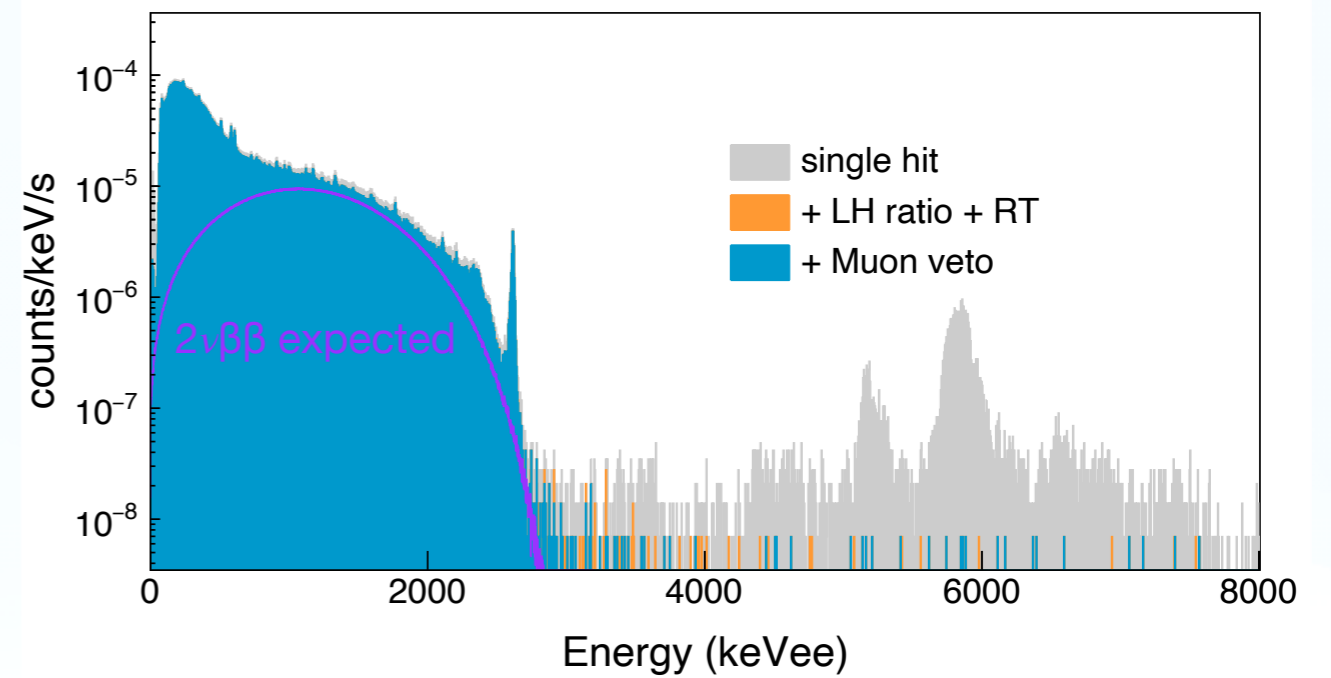
- Calibration source: ^{232}Th -rich welding rods

EVENT SELECTION

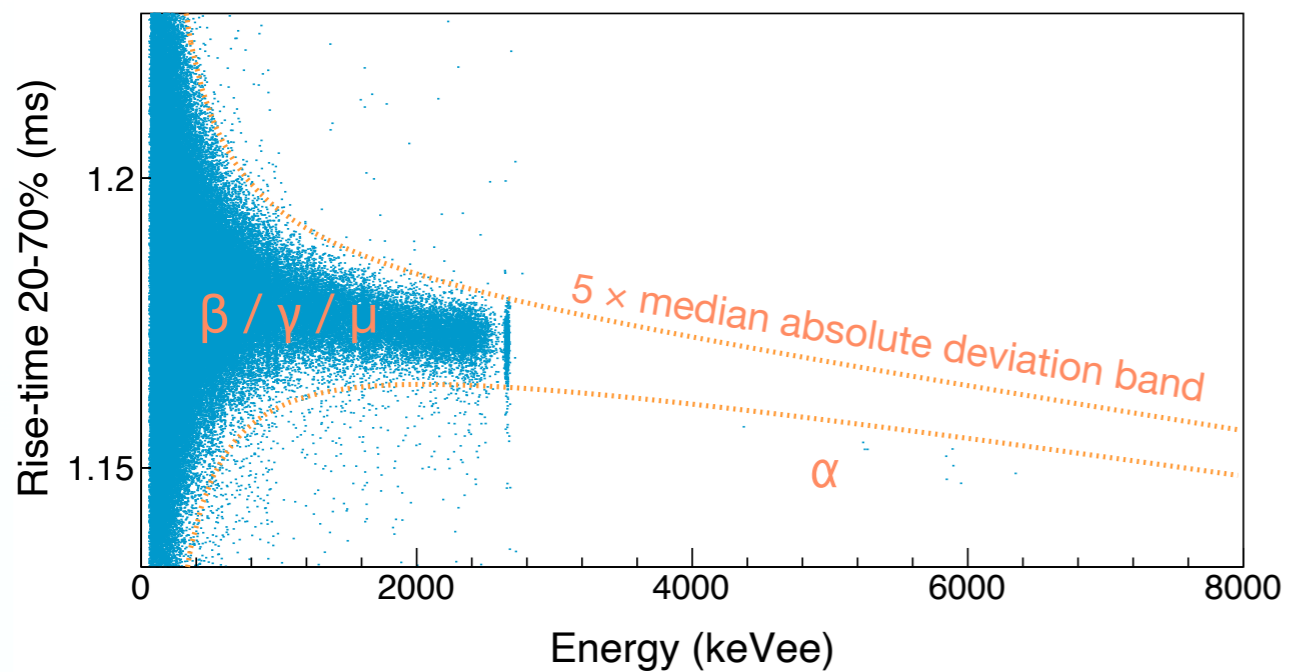
Light-heat ratio



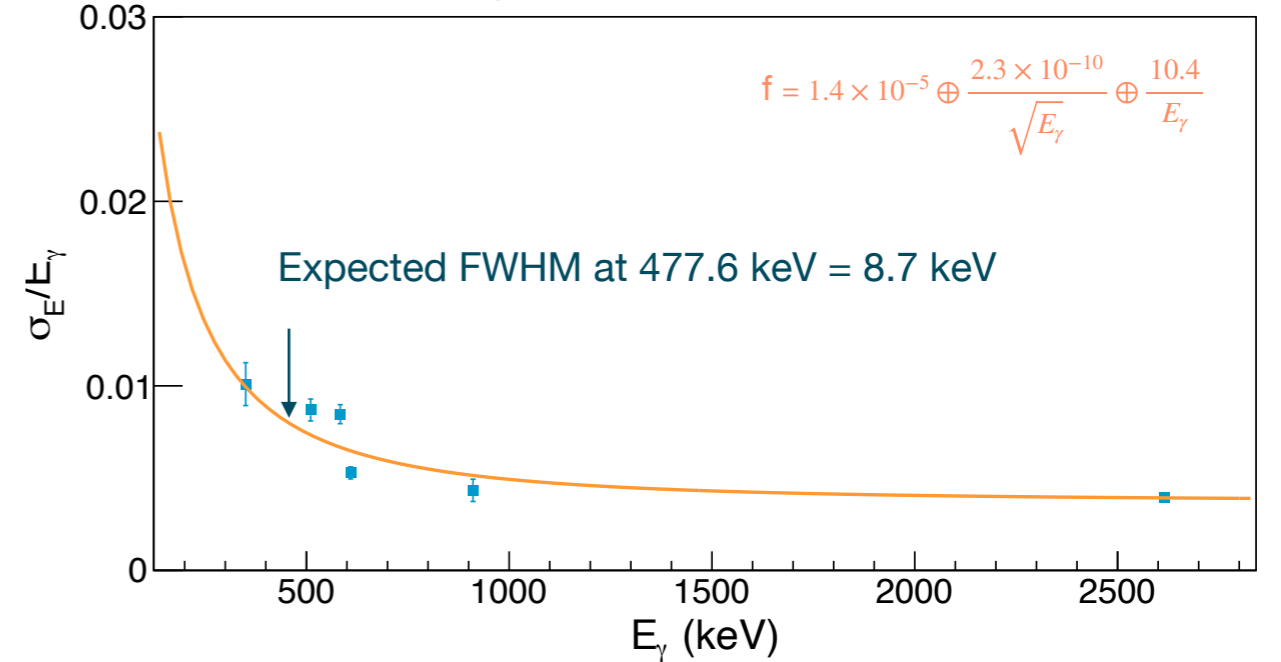
Background distribution



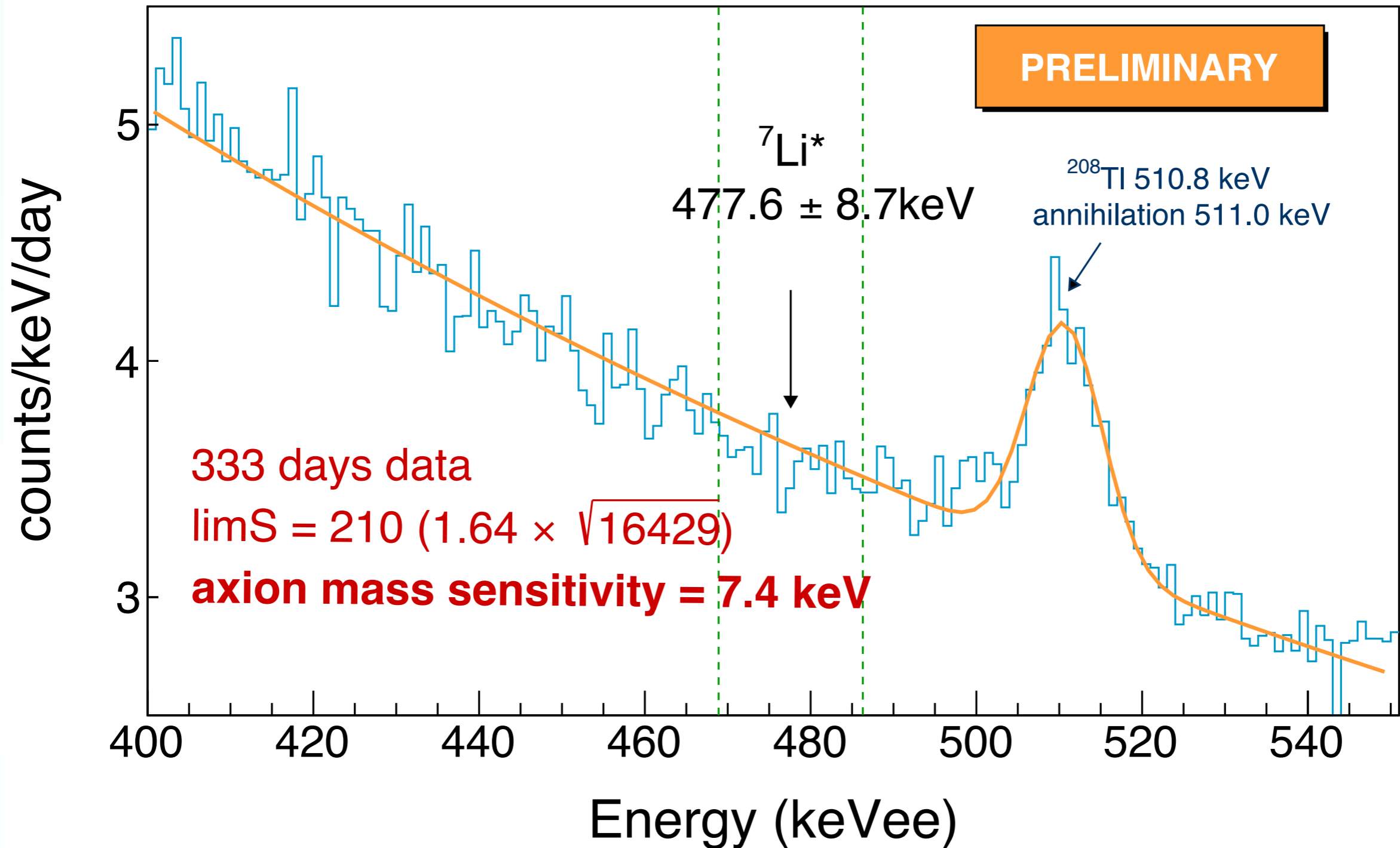
Rise-time 20-70%



Energy resolution function



PRELIMINARY AXION MASS SENSITIVITY FROM AMORE-I



SUMMARY AND FUTURE PLAN

- Studied ${}^7\text{Li}$ solar axions using the AMoRE-I data with five Li_2MoO_4 detectors and 1.46 kg·year.
- The axion mass sensitivity is 7.4 keV with $\text{lim}S = 210$ using the square root estimation.

P.Belli et al., Phys. Lett. B711, 41 (2012).

- Include the multi hit events in the analysis.
- Build a background model in region of interest.
- Analyze the detectors separately to take into the effects of different energy resolution.
- ~ 420 detectors (~178 kg Li_2MoO_4) will be used in AMoRE-II (Expected sensitivity ~2 keV with 1.28×10^{27} ${}^7\text{Li}$).