



THE CYGNO/INITIUM EXPERIMENT

G. Dho* for CYGNO coll:

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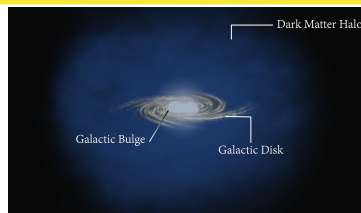


Part of this project has been funded by the European Union's Horizon
2020 research and innovation programme under the ERC Consolidator
Grant Agreement No 818744



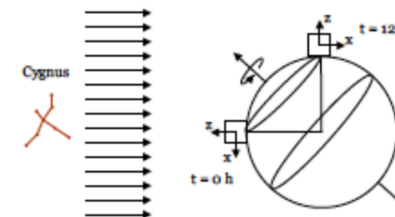
DIRECT DIRECTIONAL DARK MATTER SEARCH

- Dark Matter (DM) is considered a well established paradigm of our universe and our Galaxy is believed to reside in a halo of Dark Matter (DM), possibly WIMP-like particles.

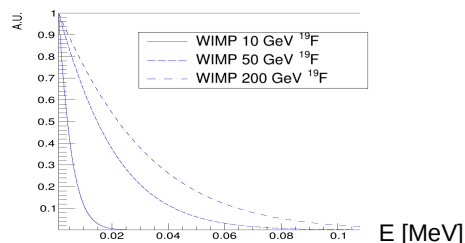


Credit: L. Jaramillo and O. Macias Virginia Tech

- The motion of the Earth together with the Sun produces an apparent wind of DM particles.

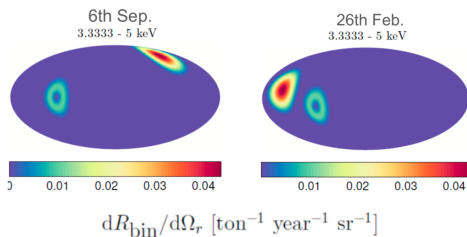
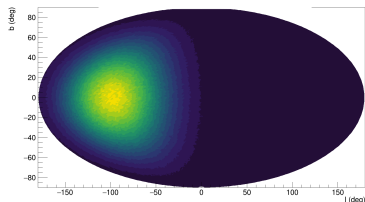


- DM of about 1–10 GeV/c² can induce nuclear recoils of light elements of few keV.



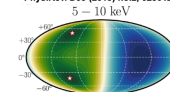
- The angular distribution of the recoils will have a very peculiar structure that no background can mimic and can be used as a powerful tool for **positive discovery, rejection of neutrino floor and DM astronomy**

F recoil due to WIMP

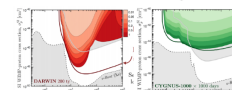


Capability to probe DM nature

15) The Gaia Sausage gives rise to peaks off-center from Cygnus



Distribution for 5-10 keV Fluorine recoils with a 100 GeV WIMP
 Halo model + SHM + Sausage



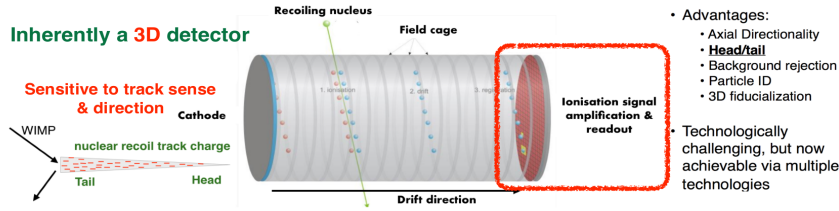
Phys.Rev. D98 (2018) no.10, 103006

The CYGNO project aims at building a large scale directional detector whose main goal will be the direct detection of DM

DETECTOR WORKING PRINCIPLE

- CYGNO features:

Time Projection Chamber (TPC)



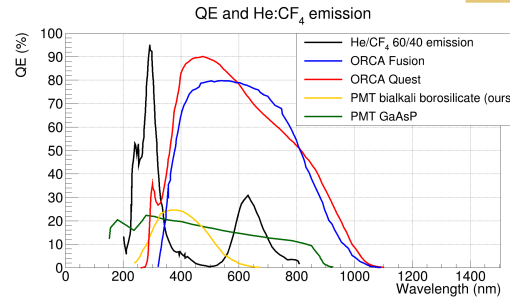
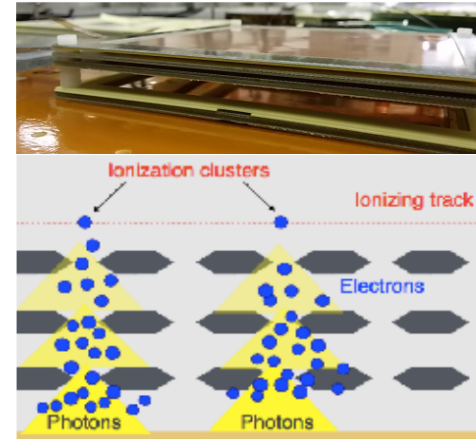
- Advantages:
 - Axial Directionality
 - Head/tail
 - Background rejection
 - Particle ID
 - 3D fiducialization
- Technologically challenging, but now achievable via multiple technologies

- Intrinsic 3D capable detector
- Operated at room temperature and atmospheric pressure
- He:CF₄ 60/40 gas mixture for Spin-dependent and independent sensitivity at low WIMP masses
- 3D fiducialization in drift direction through diffusion fit
- O(keV) energy threshold and background rejection

with GEM amplification optically readout

sCMOS camera

- Single photon sensitivity
- High granularity (2304x2304 pixels)
- Nice match with CF₄ emission

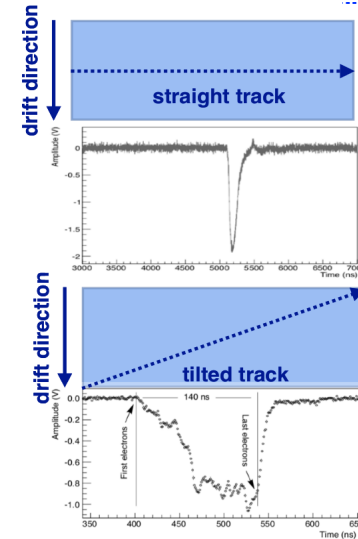


Energy
x-y coordinate



Full 3D
reconstruction

PMT



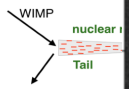
Energy
z coordinate

DETECTOR WORKING PRINCIPLE

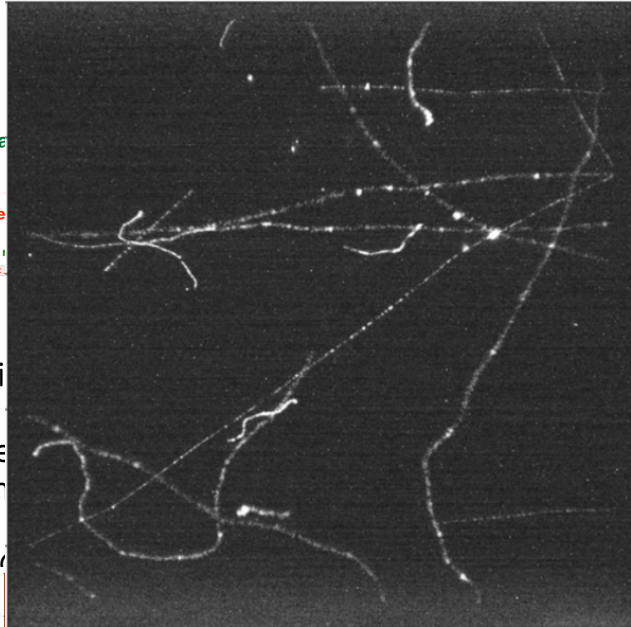
- CYGNO features:

Inherently a

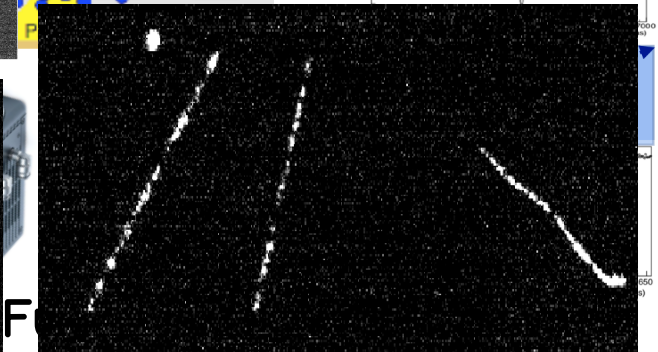
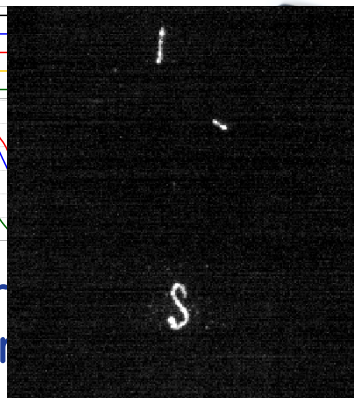
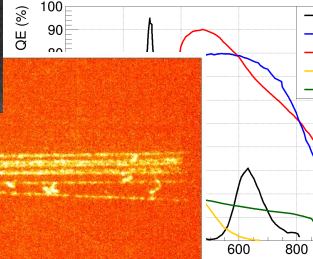
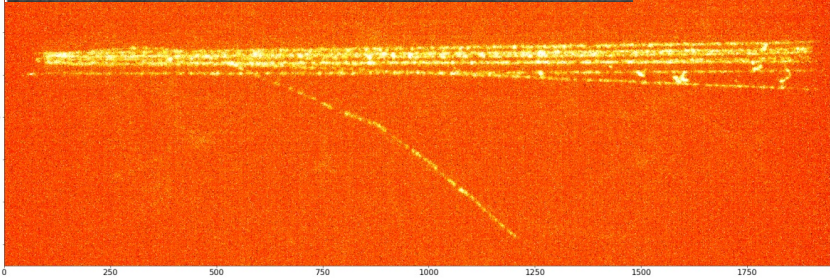
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Energy
z coordinate

CYGNO TIMELINE

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NEAR
FUTURE

FUTURE

PHASE 0: R&D and prototypes

2015/16
ROMA1

2017/18
LNF

2019/22
LNF/LNGS

PHASE 1: O(1) m³ Demonstrator

2023/26
LNF/LNGS

PHASE 2: 30 m³ Experiment

2026..
LNGS

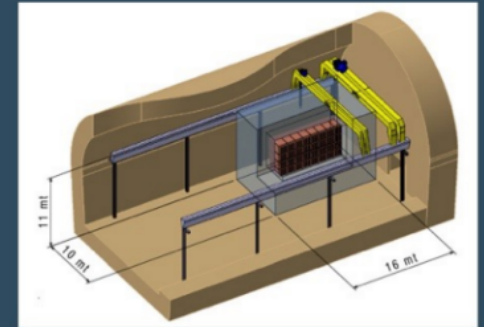
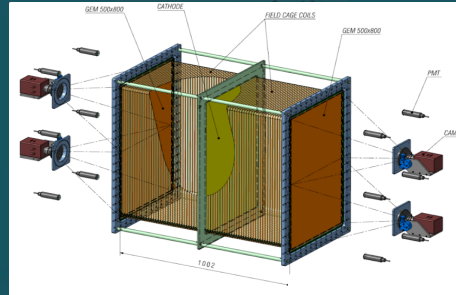
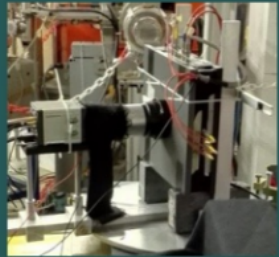
ORANGE

LEMON

LIME

CYGNO_04

CYGNO_30



- 1 cm drift

- 3D printing
- 20 cm drift

- 50 cm drift
- underground tests
- MC validation

- background
- materials test, gas purification
- scalability

- Physics research

FUNDED

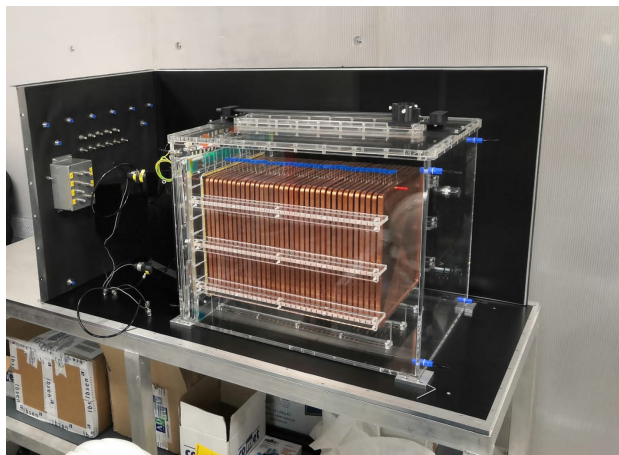
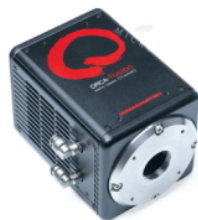
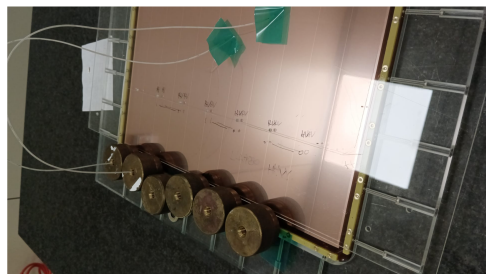
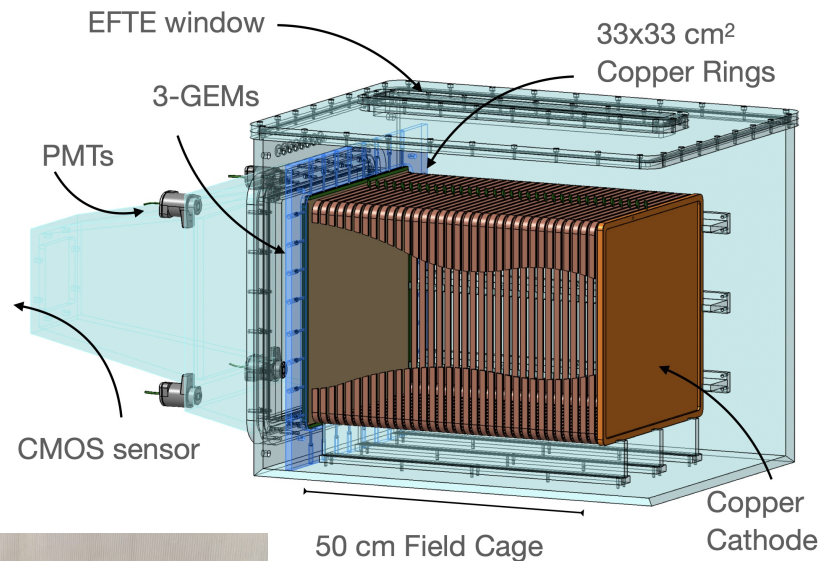
Instruments 6 (2022) 1, 6
JINST 15 (2020) P08018
JINST 15 (2020) P10001
NIM A 999 (2021) 165209

JINST 15 (2020) 12, T12003
Measur.Sci.Tech. 32 (2021) 2, 025902
2019 JINST 14 P07011

IDM 2022, G. Dho

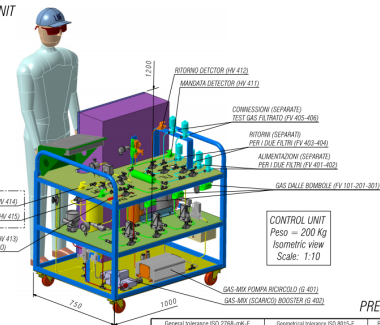
PHASE-0: LIME

- 50 l detector with 50 cm drift (copper ring field cage)
- 1 sCMOS and 4 PMTs (Hamamatsu R7378)
- 3 GEMs 50 μm thick of 33x33 cm^2 readout area
- He:CF₄ at atmospheric pressure (900 mbar once at LNGS)
- Complex gas system furnished with recirculation system



GAS SYSTEM - CONTROL UNIT

- GAS MIXING
- FLUSHING CONTROL
- PRESSURE REGULATION
- EXHAUST GAS PUMPING



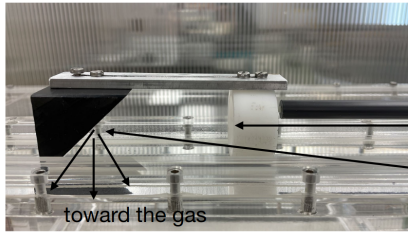
PRELIMINARY

General Information	2023-08-04	Operating Information	00-0010-4	Responsibility	00-1002
INFN	AS	AS	AS	AS	AS
Fisica Nucleare	AS	AS	AS	AS	AS
Research Division - SEM	AS	AS	AS	AS	AS
200 Kg	AS	AS	AS	AS	AS
1:10	AS	AS	AS	AS	AS
202	AS	AS	AS	AS	AS

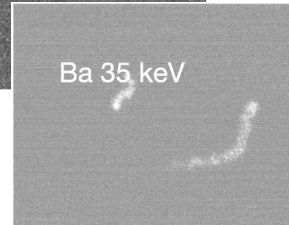
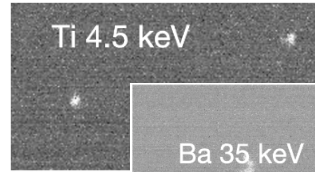
July 18th 2022

PHASE-0: LIME PERFORMANCES

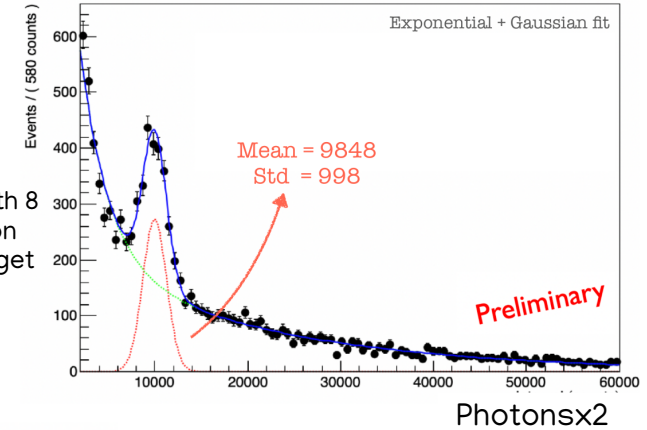
- Tests were performed overground to test the energy response on ER



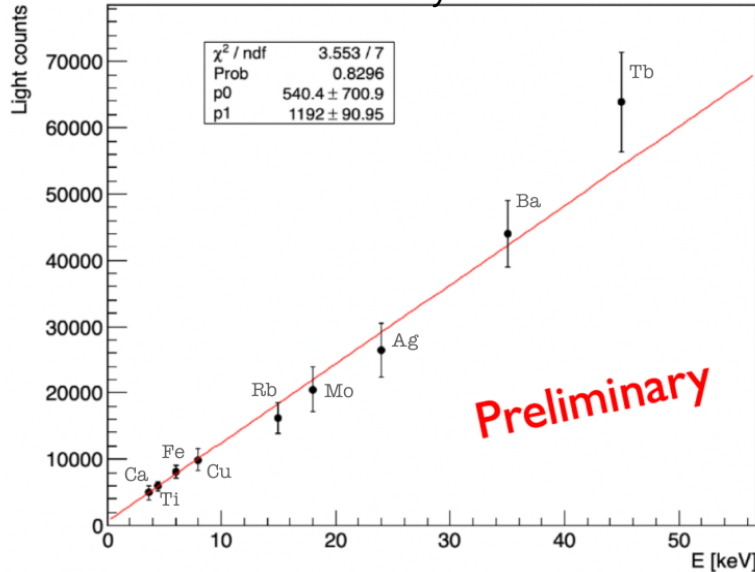
Target of different materials to be irradiated by ^{55}Fe or Am source



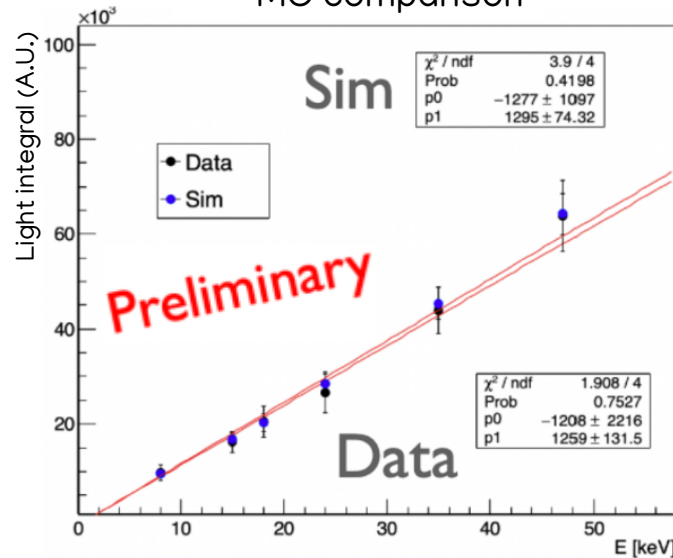
Example with 8 keV emission from Cu target



Linearity



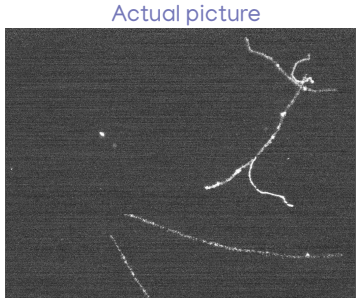
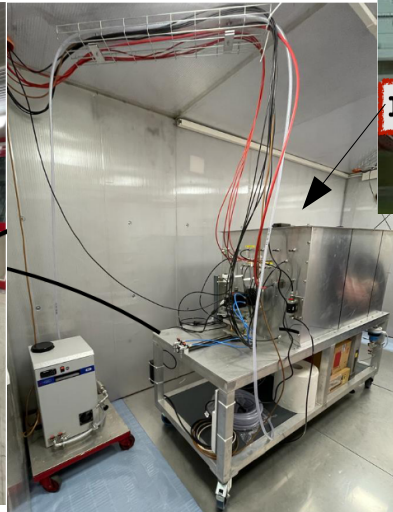
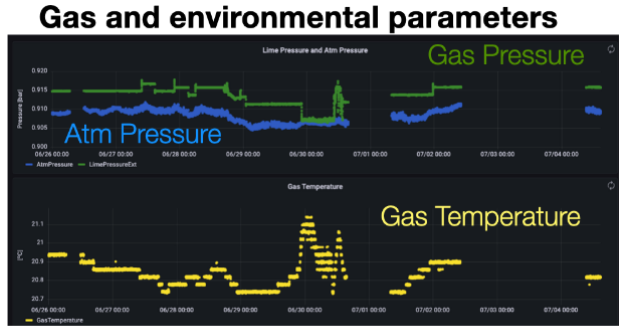
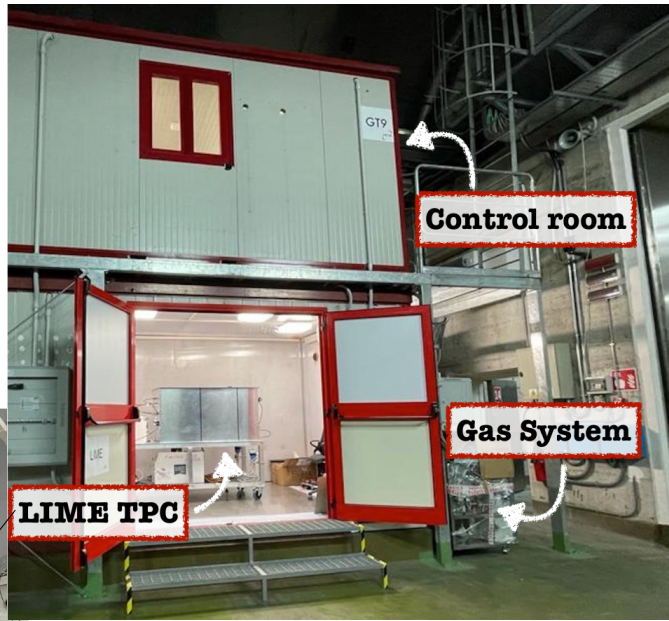
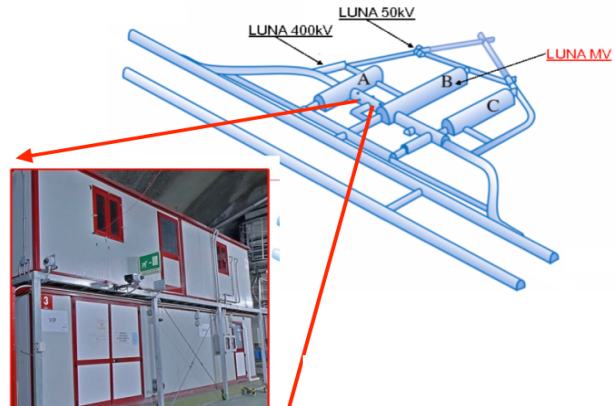
MC comparison



- Good linearity and track reconstruction from few up to 45 keV
- Good energy resolution around 14% in this energy range
- Good matching with simulation

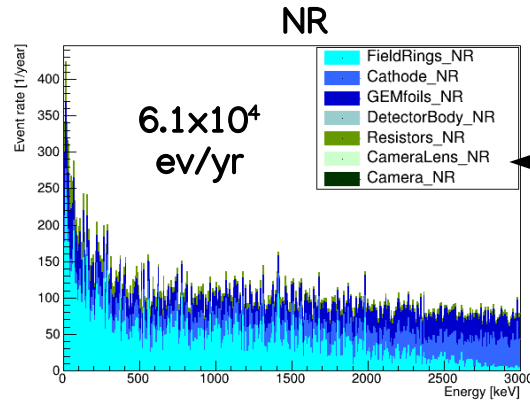
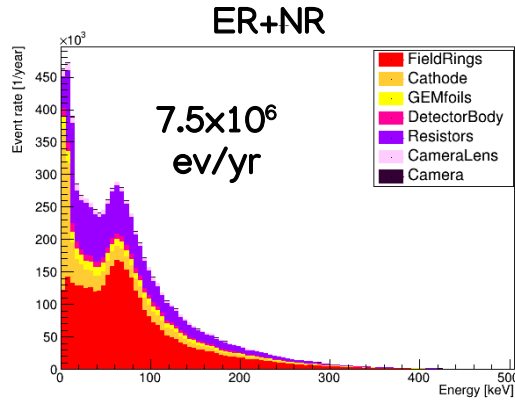
PHASE-0: LIME @ LNGS

- LIME is working underground with continuous data taking with an automatic system to remotely control it



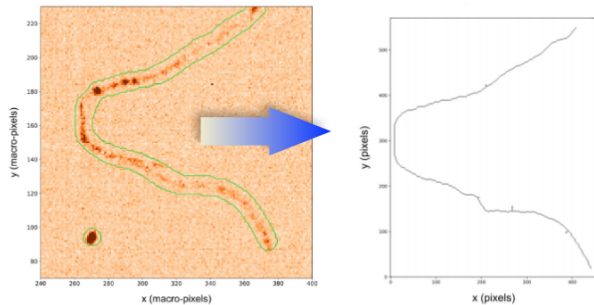
PHASE-0: LIME BACKGROUND

- LIME background simulation performed with GEANT4 after measuring @ LNGS the radioactivity of the materials

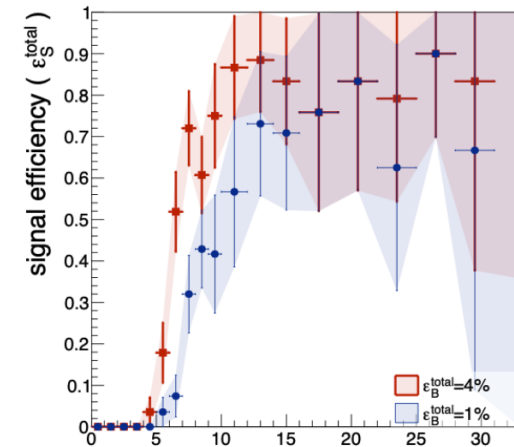


- Shielding configuration under optimization (10 cm Cu + 40 cm water, returns 2×10^5 ev/yr)
- Internal background dominates after shielding
- Geometrical cuts can suppress 96% of total recoils (99,9% of NR)
- Expected 2.4×10^5 ev/yr after cuts below 20 keV (6 ev/yr NR)

- Background rejection capabilities were tested with LEMOn prototype



- Algorithm based on DBSCAN recognizes the tracks and allows topological studies
- A simple cut on the photon density per pixel of the sCMOS pictures allows to reject 96% of background at 6 keV keeping about 40% signal efficiency
- Machine learning techniques exploiting more topological information are under development



PHASE-0: LIME MEASUREMENTS

- Foreseen measurements with LIME @ LNGS

NO SHIELD CONFIGURATION

- x Characterization of the detector with ^{55}Fe and $^{241}\text{AmBe}$ sources
- x **External background studies**, to cross-check simulation

2 months

6 CM COPPER

- x **External background studies**, to cross-check simulations
- x ^{55}Fe calibration

1 month

10 CM COPPER

- x **External background studies**, to cross-check simulations
- x **Measurement of underground neutron flux.**
Expected **200** events above **20 keV** in 4 months

4 months

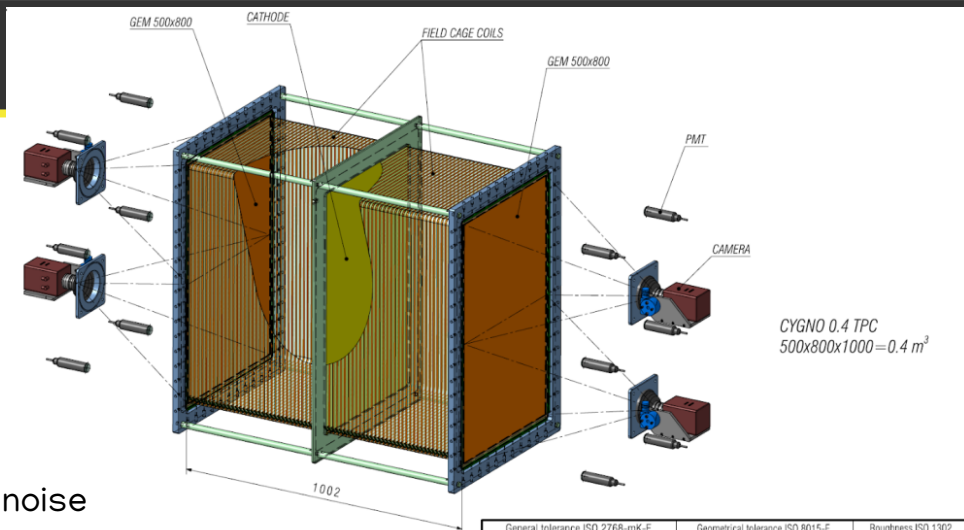
10 CM COPPER AND 40 WATER

- x **Internal background studies** for final MC validation, when internal and external background are expected to have the same intensity

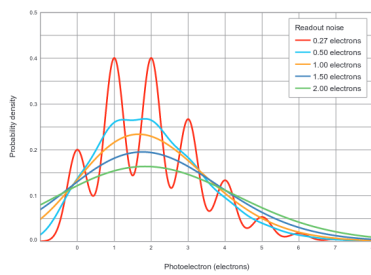
10 months

PHASE-1: CYGNO_04

- **Structure:** TPC in back-to-back configuration, 50 cm drift per side and 0,4 m³ total volume
- **Amplification:** Triple thin GEM stack of 50x 80 cm² per side
- **Readout:** Optical with 2 sCMOS (Hamamatsu ORCA Quest) and 6 PMTs per side
- **Vessel:** Very low radioactive PMMA
- **Shielding:** 10 cm Cu +100 cm water with polyethylene base
- **Purpose:**
 - Prove the scalability of the technology to large volumes using **two cameras per side** (better than LIME)
 - Test the purification gas system



Extremely low noise

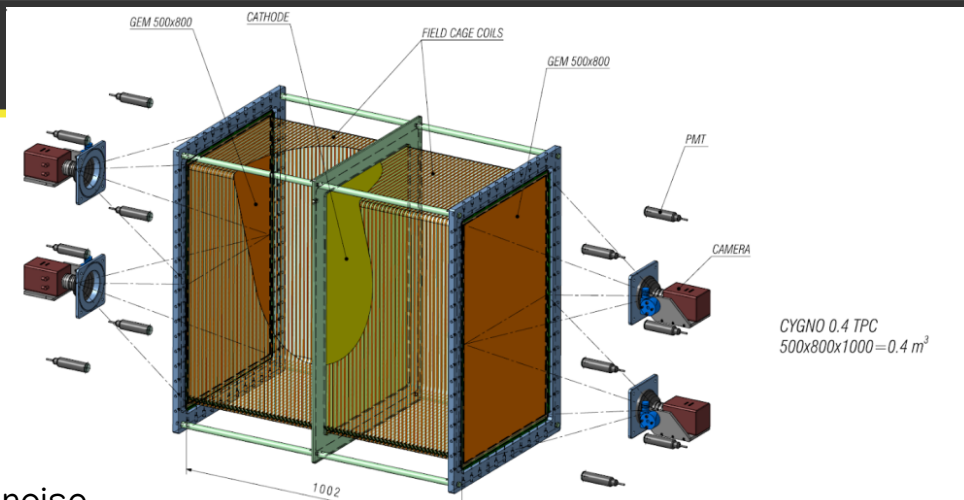
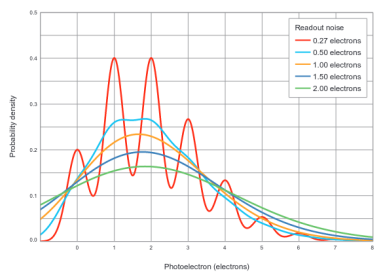


General tolerance ISO 2768-mK-E		Geometrical tolerance ISO 8015-E		Roughness ISO 1302		
NATIONAL INSTITUTE FOR NUCLEAR PHYSICS FRASCATI NATIONAL LAB RESEARCH DIVISION - SEM		SEM RESOLUTION	SIZE A3 	DATE: _____ NAME: _____ DATE: _____ NAME: _____ DATE: 11/06/2022 DATE: _____ SCALE 1:8 1/3	DATE: _____ NAME: _____ DATE: _____ NAME: C.Capocchia CHECKED: _____ APPROVED: _____	CY4-01-P

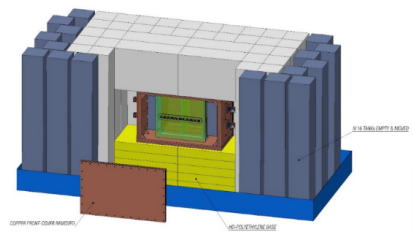
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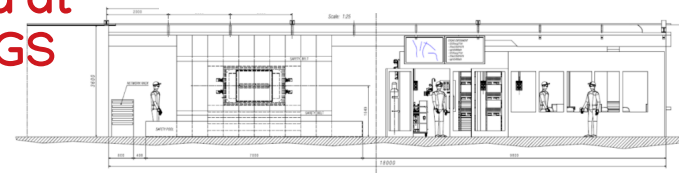
Extremely low noise



General tolerance ISO 2768-mK-E		Geometrical tolerance ISO 8015-E		Roughness ISO 1302	
NATIONAL INSTITUTE FOR NUCLEAR PHYSICS FRASCATI NATIONAL LAB RESEARCH DIVISION - SEM		SIZE A3 PROJECTION		DATE: _____ NAME: _____ DATE: _____ NAME: _____ DATE: 11/06/2022 NAME: C. Capocchia DATE: _____ NAME: _____ DATE: _____ NAME: _____	CHECKED APPROVED CY4-01-P
CYGNO EXPERIMENT CYGNO 0.4 DETECTOR TPC COMPONENTS SCHEME					

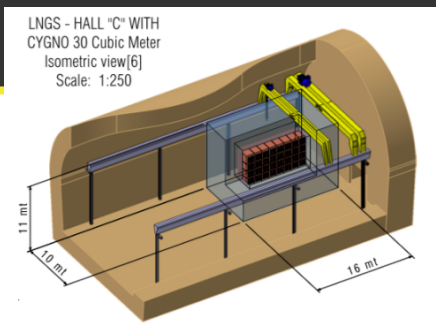


TDR submitted to be hosted at Hall F @ LNGS



PHASE-2: CYGNO_30

- O(30) m³ detector for the DM physics case



- Expected SI and SD 90% CL are obtained by simulation and statistical analysis based on Bayesian technique

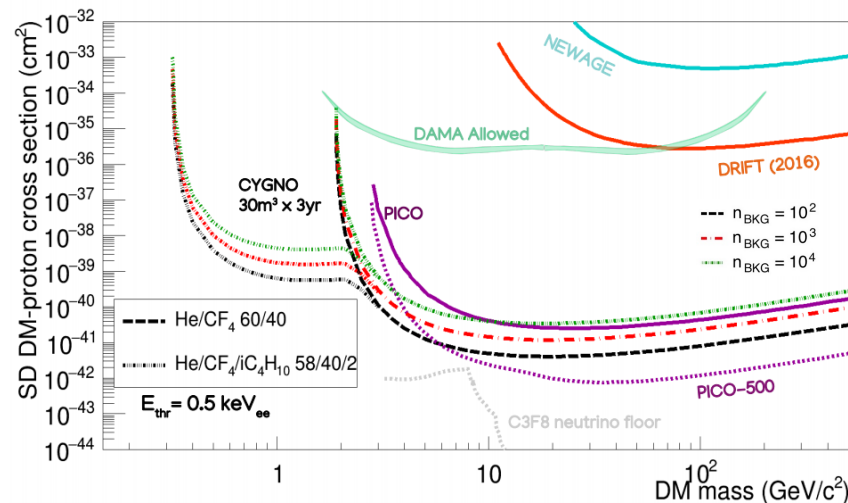
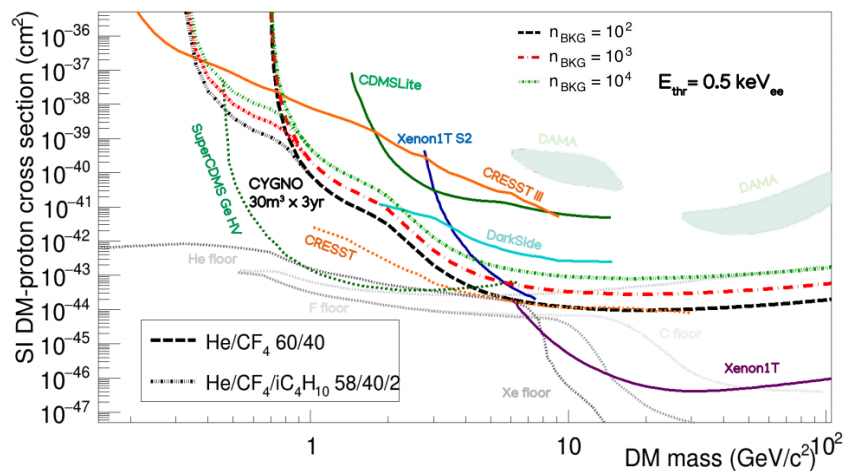
Energy threshold 0.5 keV

Observable 3D full directional angular distribution

Background model different number of events in a flat isotropic distribution

Quenching factor obtained by SRIM simulation

Angular resolution 30x30 degree² flat



ONGOING R&D

- Some R&D are still ongoing to improve the future performances

- Internal background reduction

- × Building low radioactivity camera sensor
- × Optimize field cage materials
- × Lenses made of fused silica



Suprasil
PMMA
POLYCARB

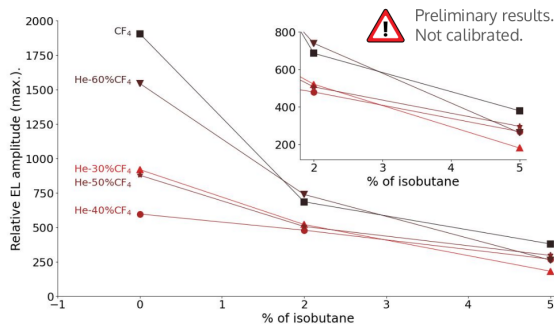
Feasibility study for low radioactivity lens

Ready to be produced

- Hydrogen addition in gas mixture

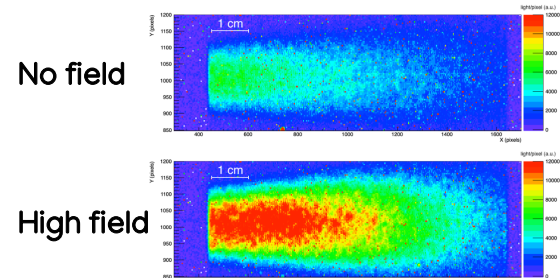
- × Introducing hydrogen rich gas to improve the sensitivity at low mass without losing light yield

	Minimum detectable DM mass for 0.5 keV _{ee} energy threshold	Minimum detectable DM mass for 1 keV _{ee} energy threshold
H	300 MeV/c ²	500 MeV/c ²
He	700 MeV/c ²	1 GeV/c ²
C	1.4 GeV/c ²	1.9 GeV/c ²
F	1.9 GeV/c ²	2.5 GeV/c ²

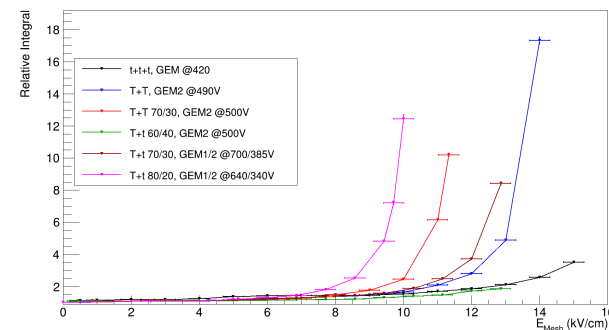


- Electroluminescence to enhance light

- × Apply strong electric fields below the GEM to produce light without additional charge

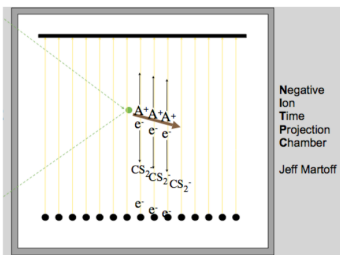


Baracchini, E. et al. JINST 15 08 (2020)



ONGOING R&D: NID OPERATION

- In synergy with CYGNO, the INITIUM project aims at utilizing negative ion operation to improve track performances



– J. Martoff et al., NIM A 440 355

- Adding SF₆ as high electronegative gas, electrons get attached to the dopant close to the ionization position o(100) um
- Anions drift acting as the image carrier but with a longitudinal and transverse diffusion reduced to the thermal limit

+ SF₆ for negative ion drift



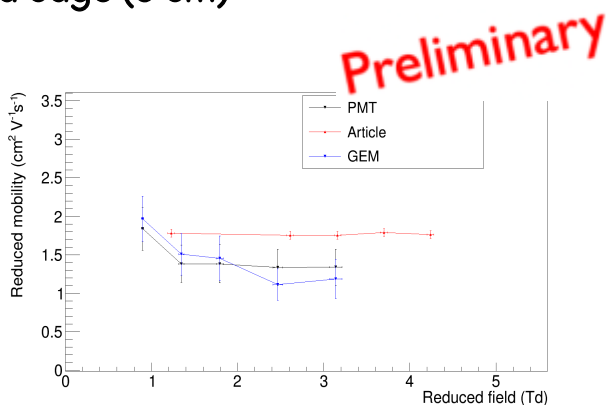
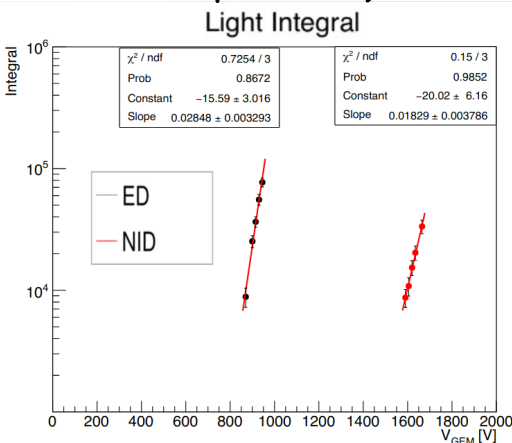
Part of this project has been funded by the European Union's Horizon 2020 research and innovation programme under the ERC Consolidator Grant Agreement No 818744

- Using a small prototype it was possible to measure negative ion drift at nearly atmospheric pressure with optical readout using an ²⁴¹Am alpha source

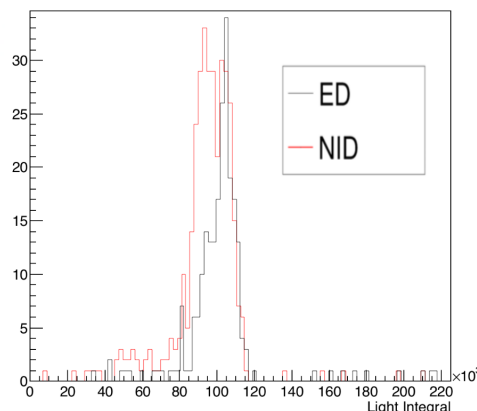
880 mbar (mountain atm pressure) short field cage (5 cm)

He:CF₄:SF₆ 59:39.4:1.6

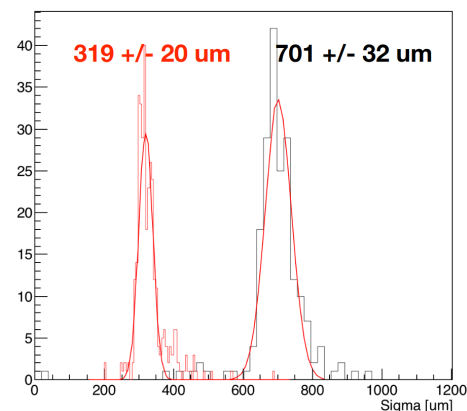
650 mbar long field cage (15 cm)



Light Integral @ 12.5 cm drift distance

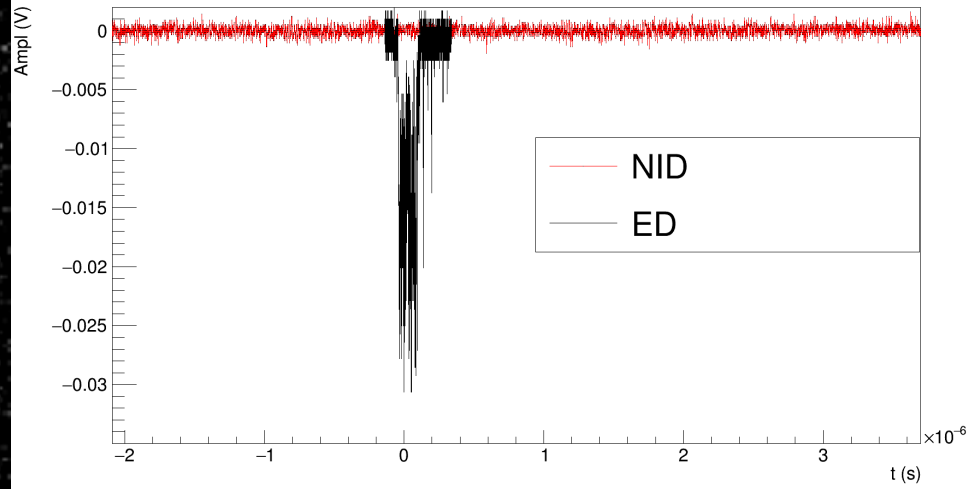
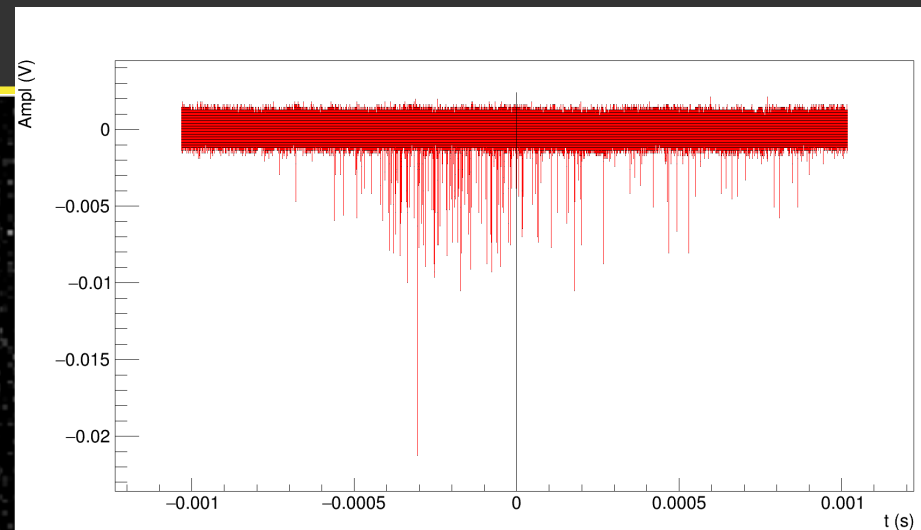
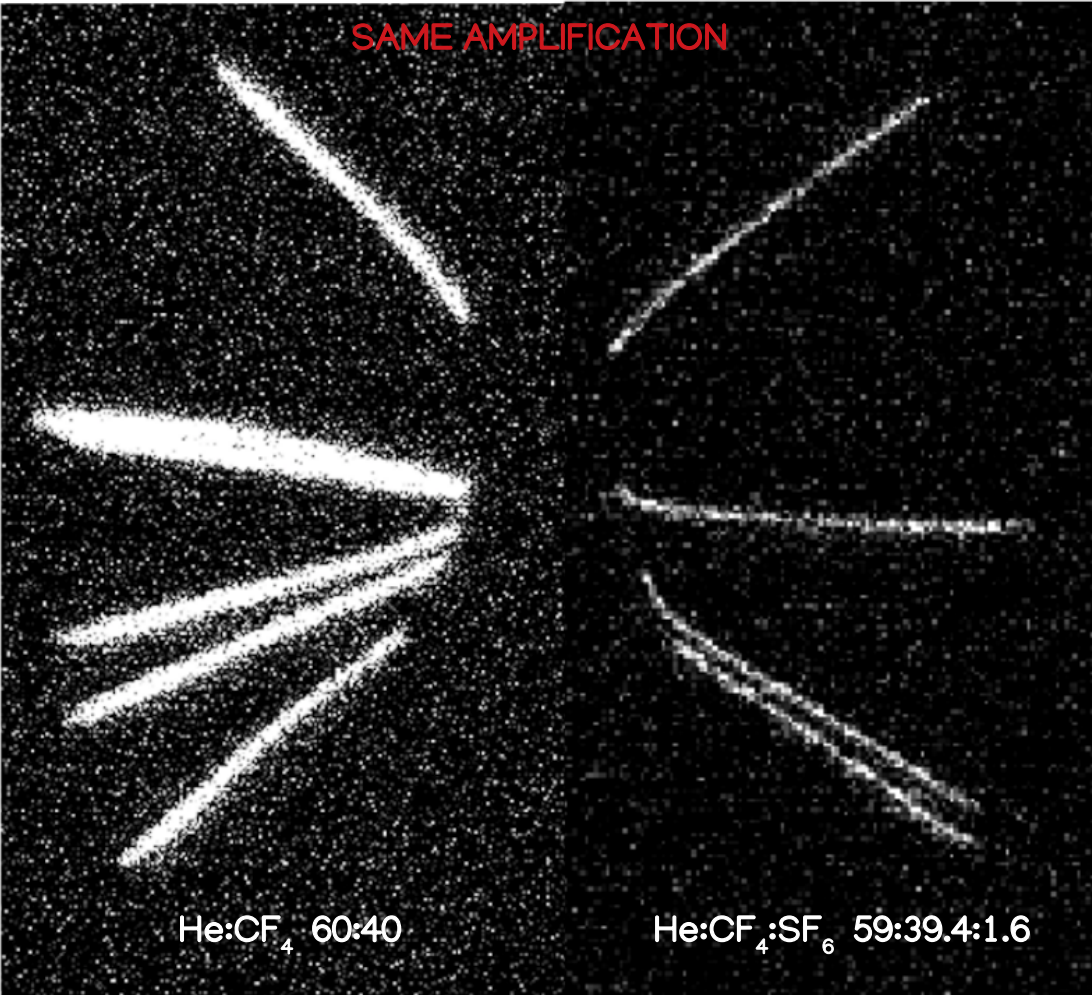


Sigma @ 12.5 cm drift distance



ONGOING R&D: NID OPERATION

SAME AMPLIFICATION

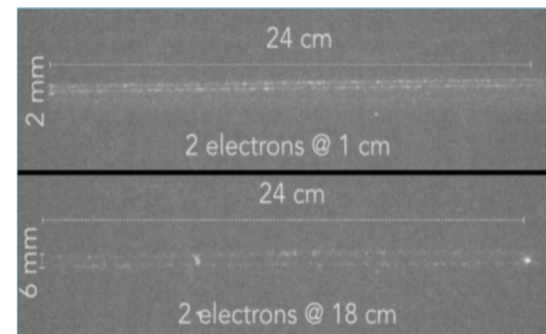
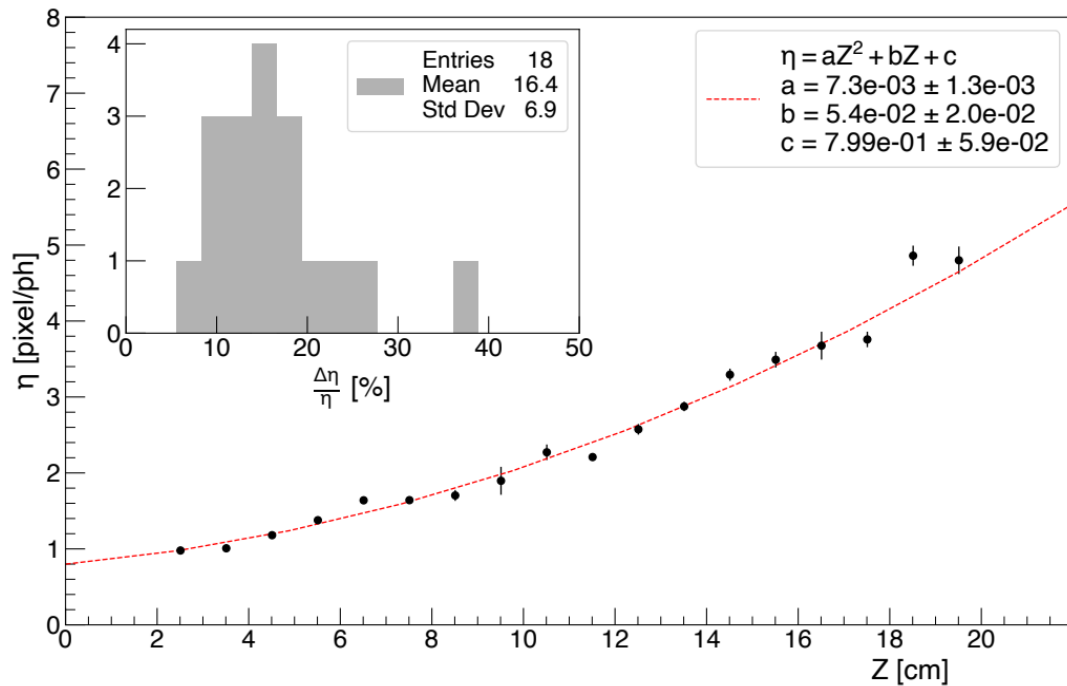


CONCLUSIONS

- The CYGNO collaboration is moving toward the realization of large scale Dark Matter detector for $O(\text{GeV})$ WIMP-like particle based on a gaseous Time Projection Chamber.
- The R&D efforts that showed the $O(\text{keV})$ energy threshold and background rejection, the fiducialization, all lead to the realization of LIME, the largest prototype built with 50 l volume. Now it is taking data in the underground laboratories at LNGS to show the performances in a realistic environment towards the construction of phase 1.
- The future steps toward the construction of the large scale detector are defined with two separate phases. Phase 1 is under development to prove scalability, it is already funded and the TDR has been submitted for it to be hosted at LNGS. Phase 2 is under development and the sensitivity studies show interesting prospects.
- R&D is being carried out to improve the performances of the detector from the radioactivity of the materials to the gas mixture optimization
- The study on negative ion doped gas mixtures is showing relevant and promising results

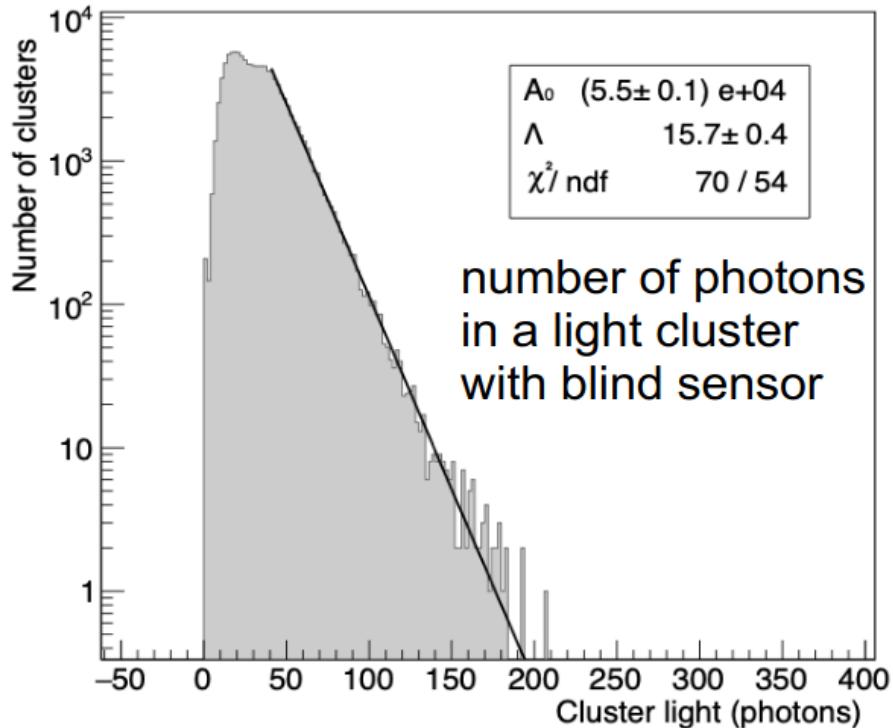
BACKUP

Z (DRIFT DIRECTION) FIDUCIALIZATION



- Electron transverse diffusion can be exploited to infer the track Z coord.
- Track transverse light profile measured to have gaussian shape which enlarges linearly with Z
- Under study the method using a PMT

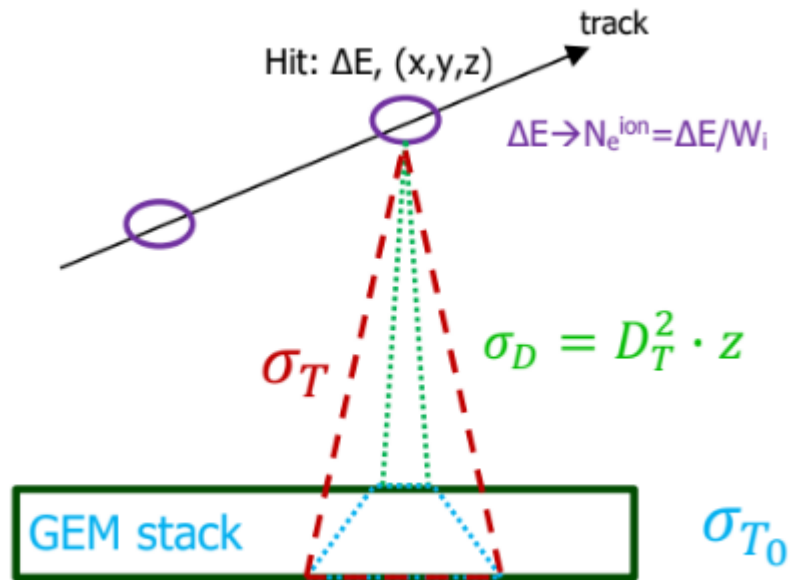
ENERGY THRESHOLD



Clusters found due to electronic noise:

- Clusters containing >400 photons are less than 10 per year
- LIME setup allows to see about 650 photons per keV
- Energy threshold assumed 1 keV_{ee}
- Positive that 0.5 keV threshold could be achieved with better algorithms

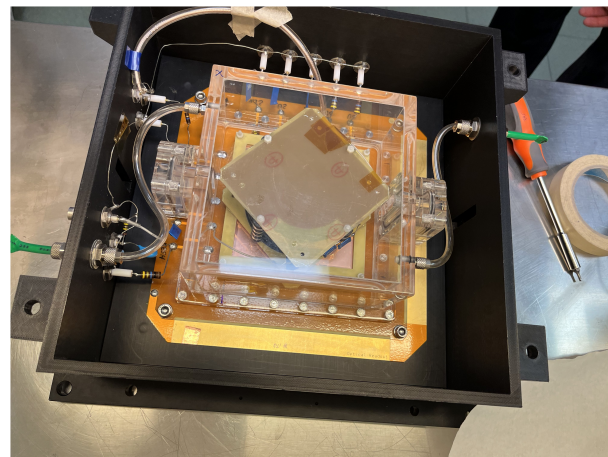
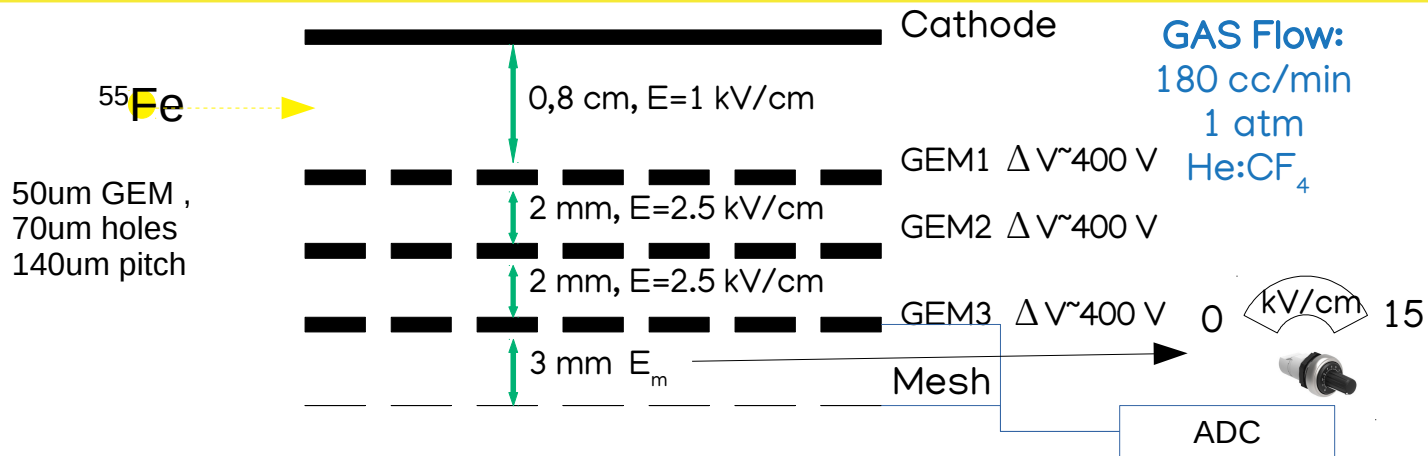
SIMULATION



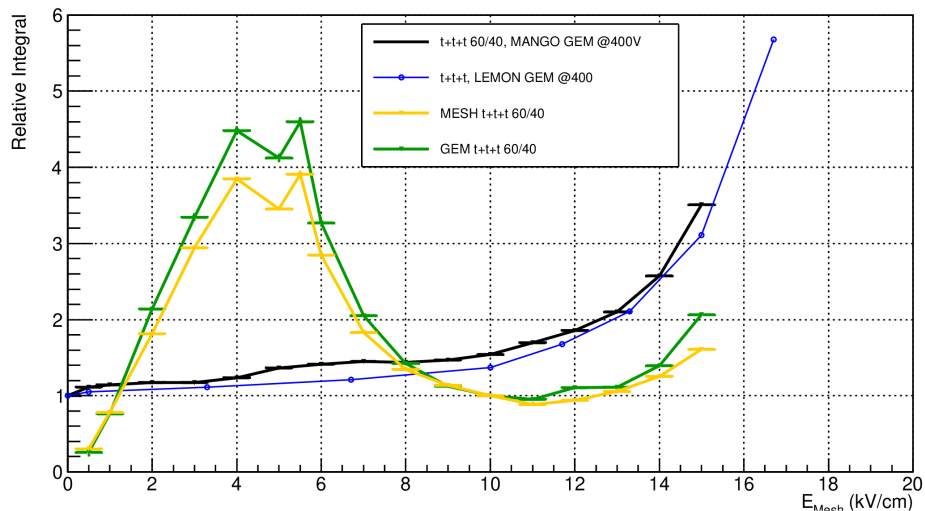
The simulation produces pictures equivalent to the format of the data. It includes:

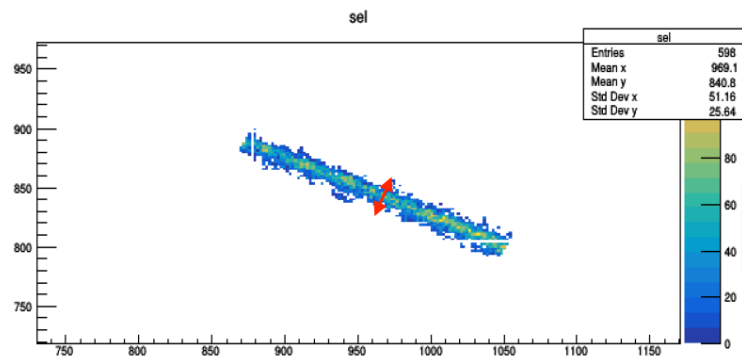
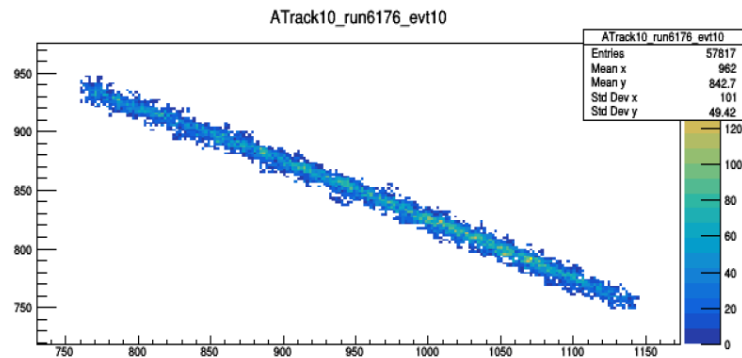
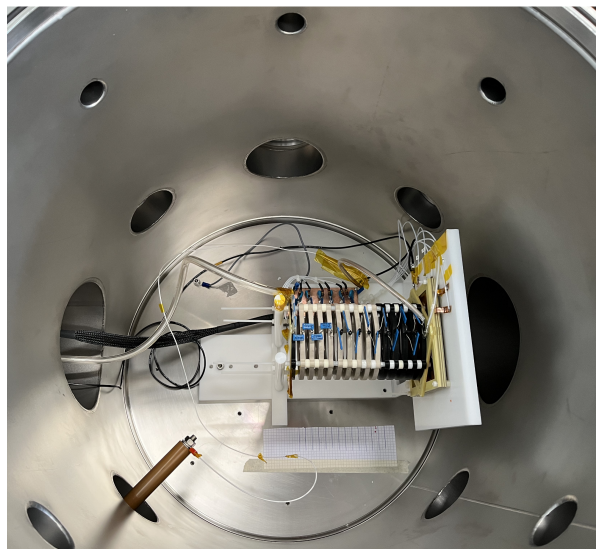
- GEANT4 or SRIM to represent energy deposit on gas
- Primary charge production
- Parametrized diffusion in the gas and through the GEMs
- Parametrized GEM amplification process and saturation effect on third GEM
- Photon production and collection efficiency of the detector
- Electronic noise of the sensor

ELECTROLUMINESCENCE



Process	Threshold (eV)	Energy loss (eV)
Direct vibrational excitation v_4	0.078	0.078
v_3	0.159	0.159
Indirect vibrational excitation	4.0	0.4
Electron attachment	4.3	4.3
Electronic excitation (dissociation into neutral fragments) [†]	12.5 (10)	12.5 (10)
Dissociative ionization [†]	15.9	15.9





Tracks selection:

- track length > 8 mm
- track slimness < 0.3
- # of peaks in the transverse profile == 1 (select single tracks)
- Chi2/nDOF of transverse fit profile < 5 (remove additional multiple tracks)

