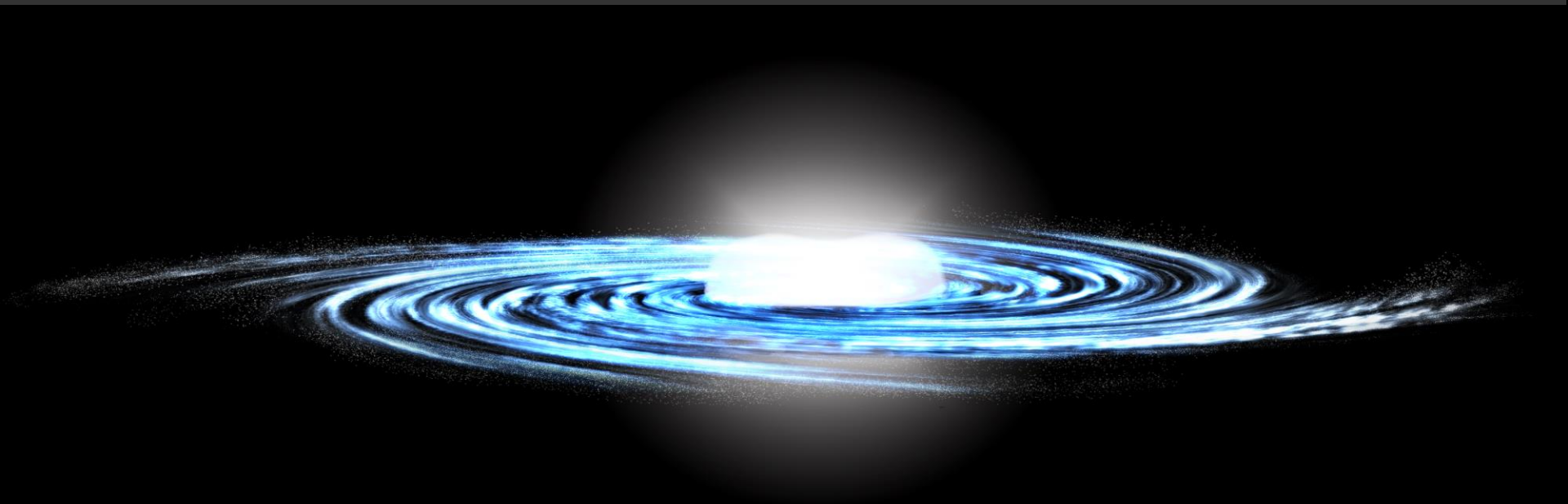
Assessing the Impact of Hydrogen Absorption on the Characteristics of the Galactic Center Excess



Oscar Macias

with M. Pohl (DESY), P. Coleman (UofC), C. Gordon (UofC)

ApJ 929 (2022) 2, 136 [arXiv:2203.11626]

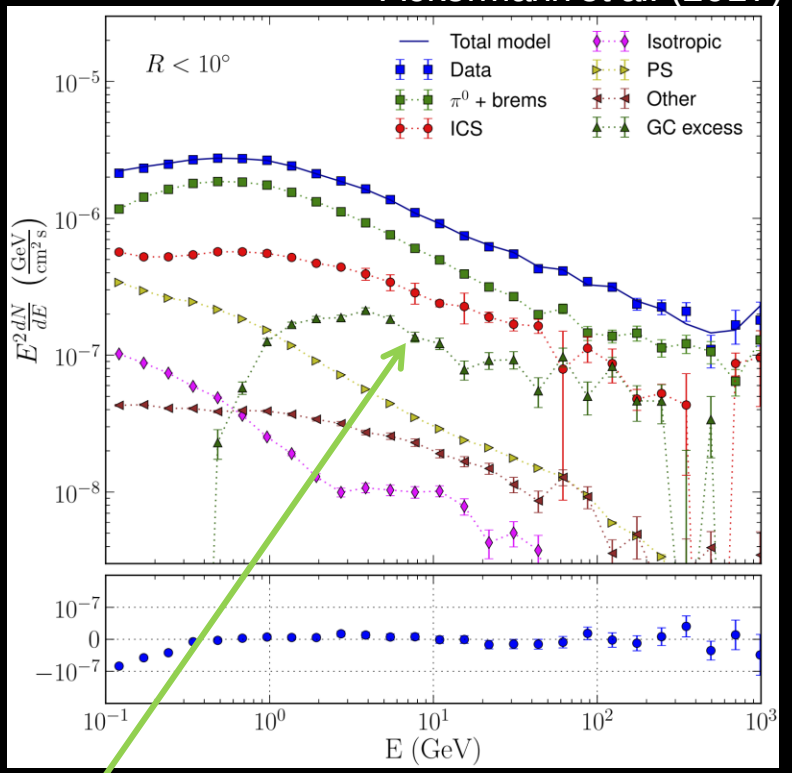
IDM 2022,
Technical University Vienna

UNIVERSITY OF AMSTERDAM

GRAPPA
GRavitation AstroParticle Physics Amsterdam

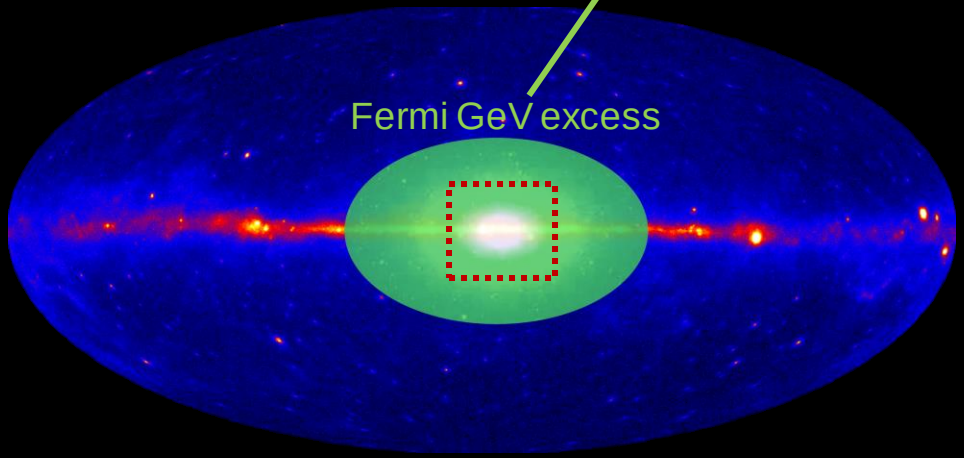
The Fermi GeV Excess

Ackermann et al. (2017)



At the Galactic Center

Goodenough & Hooper (2009)
Vitale & Morselli (2009)
Hooper & Goodenough (2011)
Hooper & Linden (2011)
Boyarsky et al (2011)
Abazajian & Kaplinghat (2012)
Gordon & Macias (2013)
Macias & Gordon (2014)
Abazajian et al (2014, 2015)
Calore et al (2014)
Daylan et al (2014)
Selig et al (2015)
Huang et al (2015)
Gaggero et al (2015)
Carlson et al (2015, 2016)
de Boer et al (2016)
Yang & Aharonian (2016)
Fermi Coll. (2016)
Horiuchi et al (2016)
Linden et al (2016)
Ackermann et al (2017)
Macias et al (2018)
Bartels et al (2018)
Balaji et al (2018)
Zhong et al (2019)
Macias et al (2019)
Chang et al (2020)
Buschmann et al (2020)
Leane & Slatyer (2020)
Abazajian et al (2020)
List et L (2020)
Di Mauro (2020)
Burns et al (2020)
Cholis et al (2022)
Pohl, Macias+(2022)
…



At mid-lat

Hooper & Slatyer (2013)
Huang et al (2013)
Zhou et al (2014)
Daylan et al (2014)
Calore et al (2014)
…

The Fermi GeV Excess

Ackermann et al. (2017)

At the Galactic Center

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…

At mid-lat

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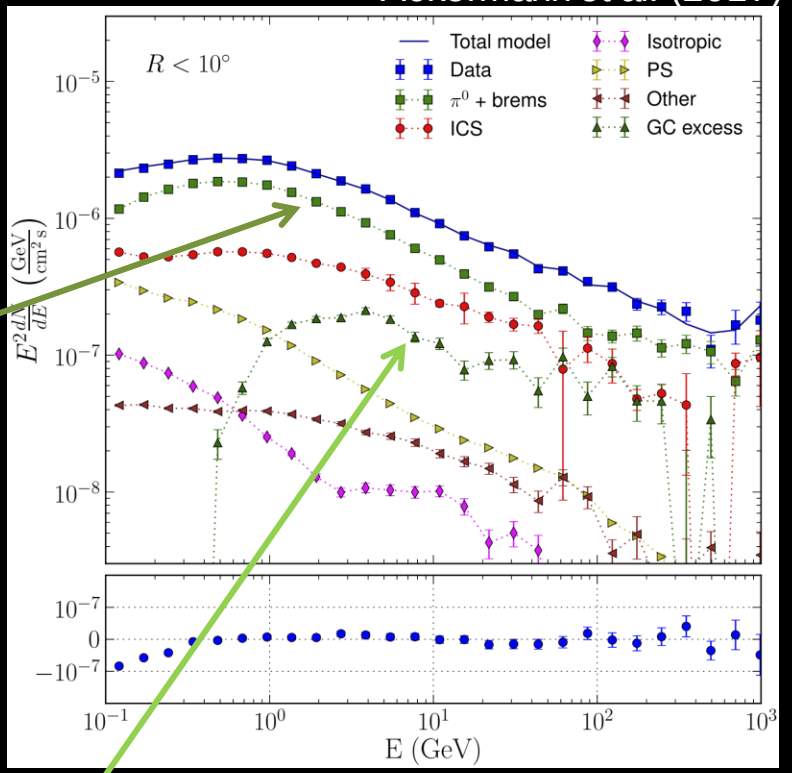
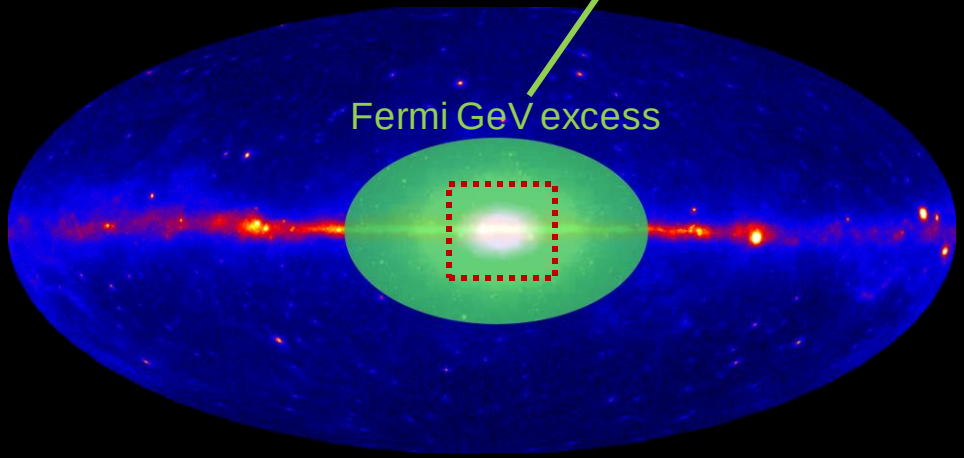
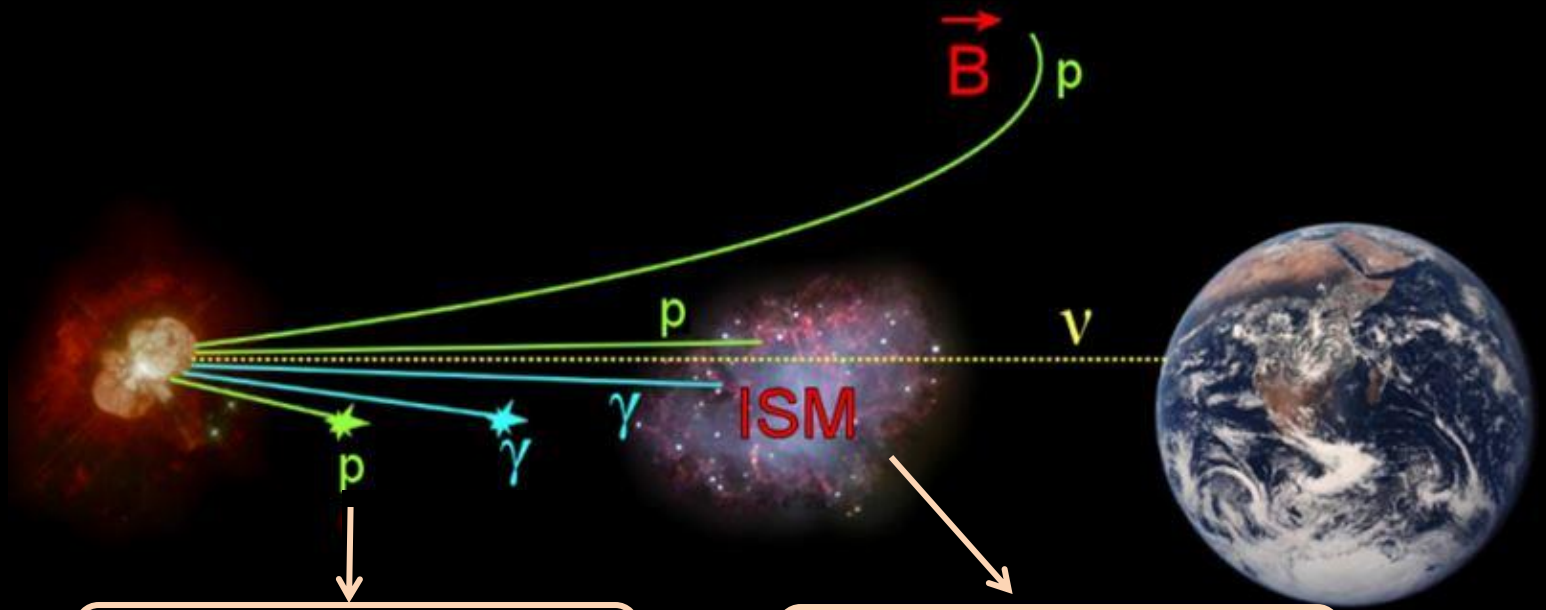
Huang et al (2013)
Zhou et al (2014)
Daylan et al (2014)
Calore et al (2014)
…

Gas-correlated gamma rays dominate in this sky region





Ingredients of the Gas-correlated emission

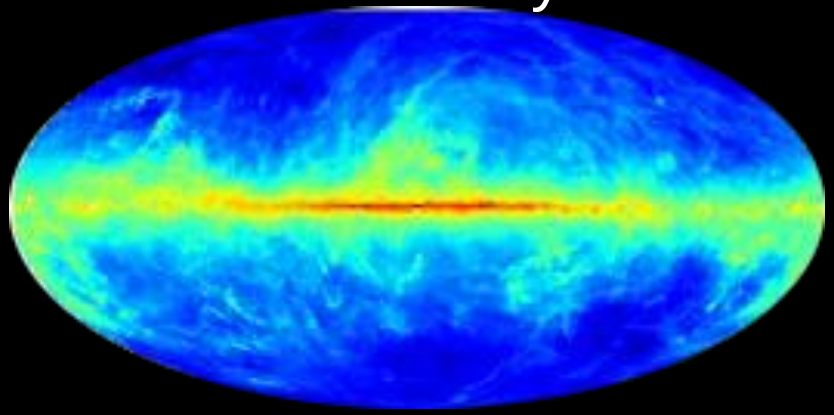


Cosmic Ray Densities

Interstellar Gas

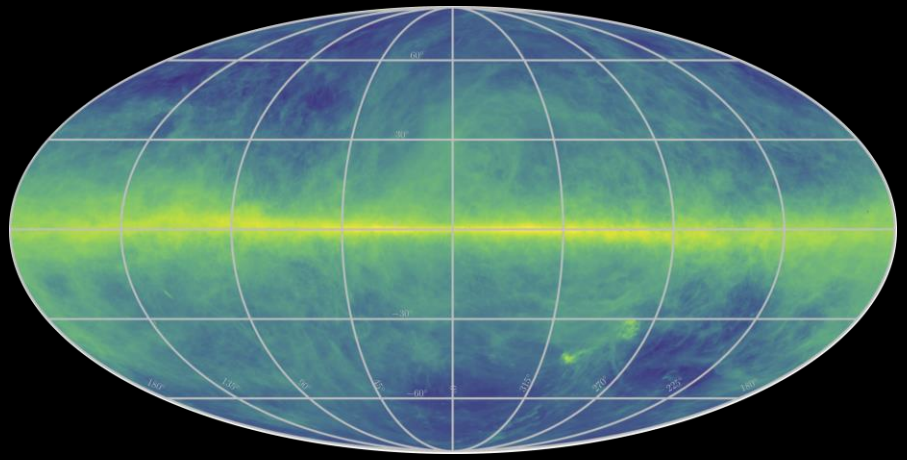
See Mattia's and Ilias' talks!

Gamma rays



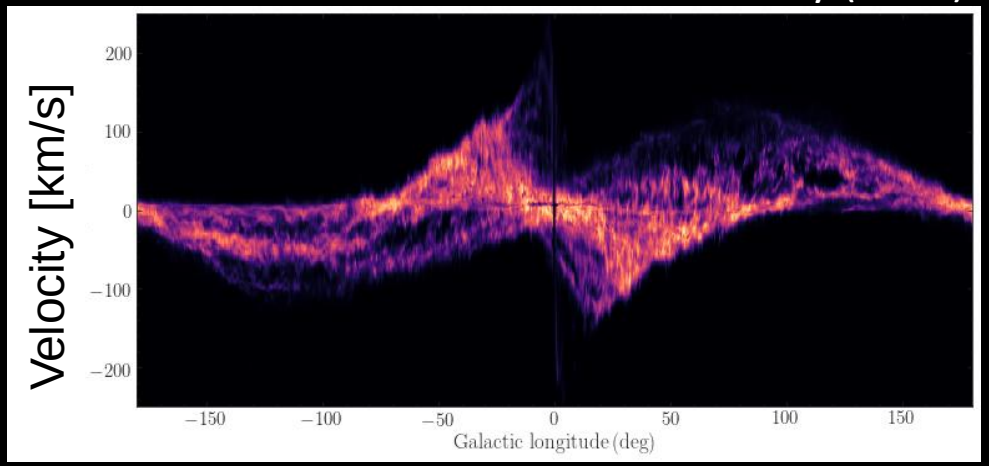
Interstellar Hydrogen Measurements

HI4PI survey, Ben Bekhti et al. (2016)



Total hydrogen column density

J. Bovy (2021)



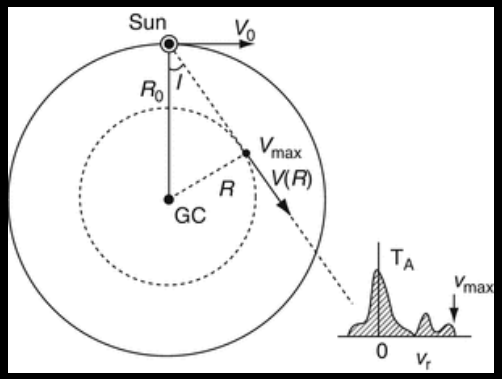
Line-of-sight **HI velocity**

➔ What is needed is the spatial distribution of hydrogen along the line of sight

Constructing the new hydrogen maps

Previous approach

- Assumes circular motion of hydrogen gas (interpolation method)



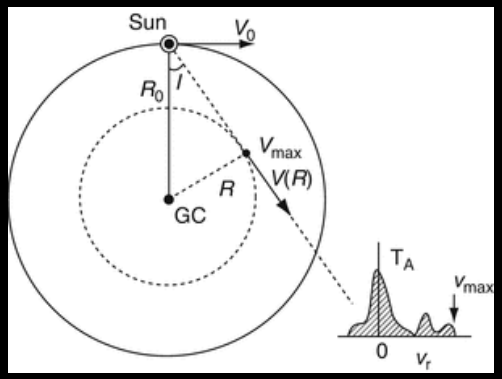
Circular orbits of gas

Constructing the new hydrogen maps

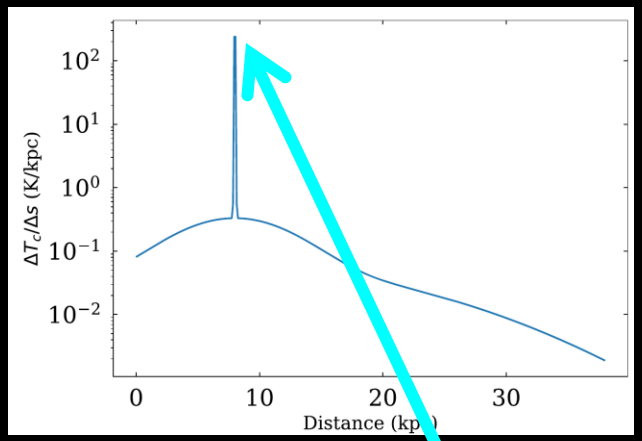
Previous approach

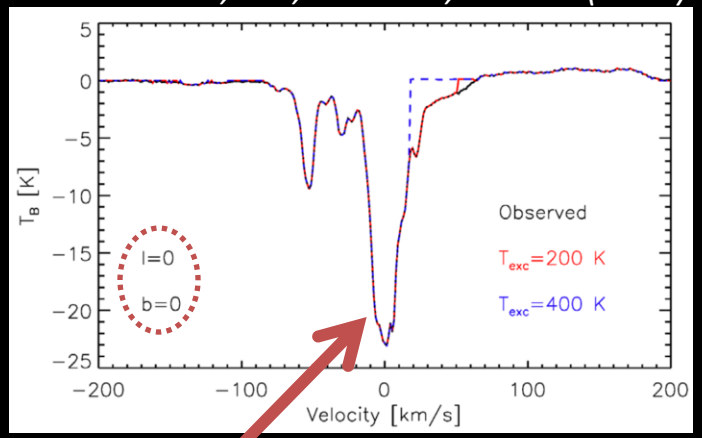
- Assumes circular motion of hydrogen gas (interpolation method)
- Disregards continuum emission and absorption.

Pohl, OM, Coleman, Gordon (2022)



Circular orbits of gas





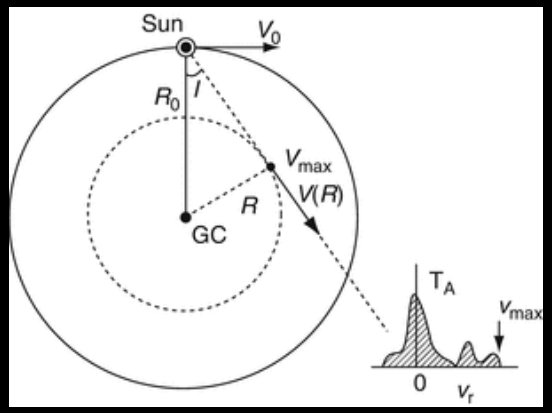
Strong **continuum** and **absorption** features @ GC

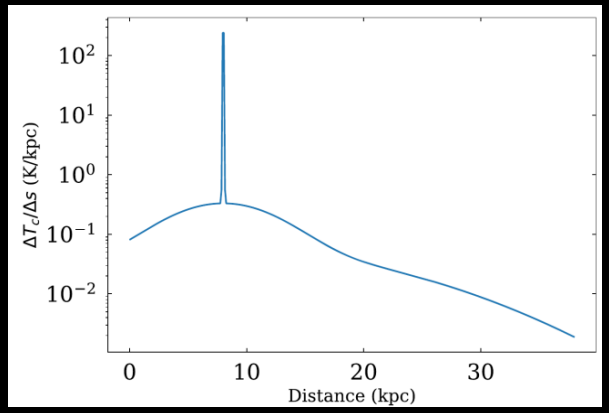
Constructing new hydrogen maps

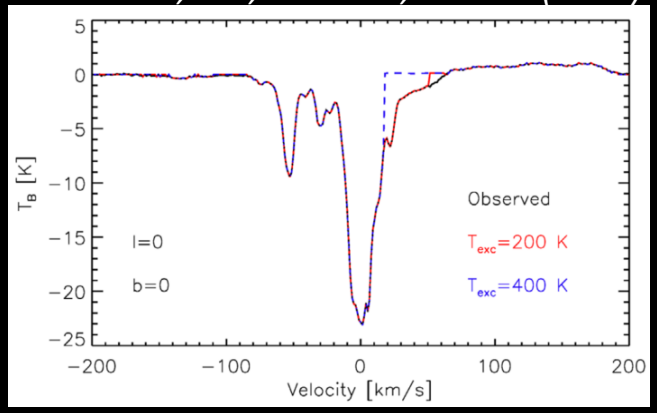
Previous approach

- Assumes circular motion of hydrogen gas (interpolation method)
- Disregards continuum emission and absorption.

Pohl, OM, Coleman, Gordon (2022)

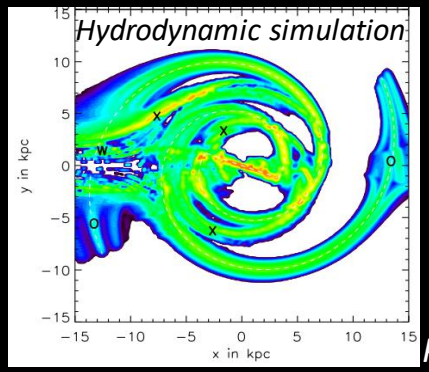






New approach

- Gas-flow model from hydrodynamic simulations
- Solve the radiative transfer equation



Pohl+ (2009)

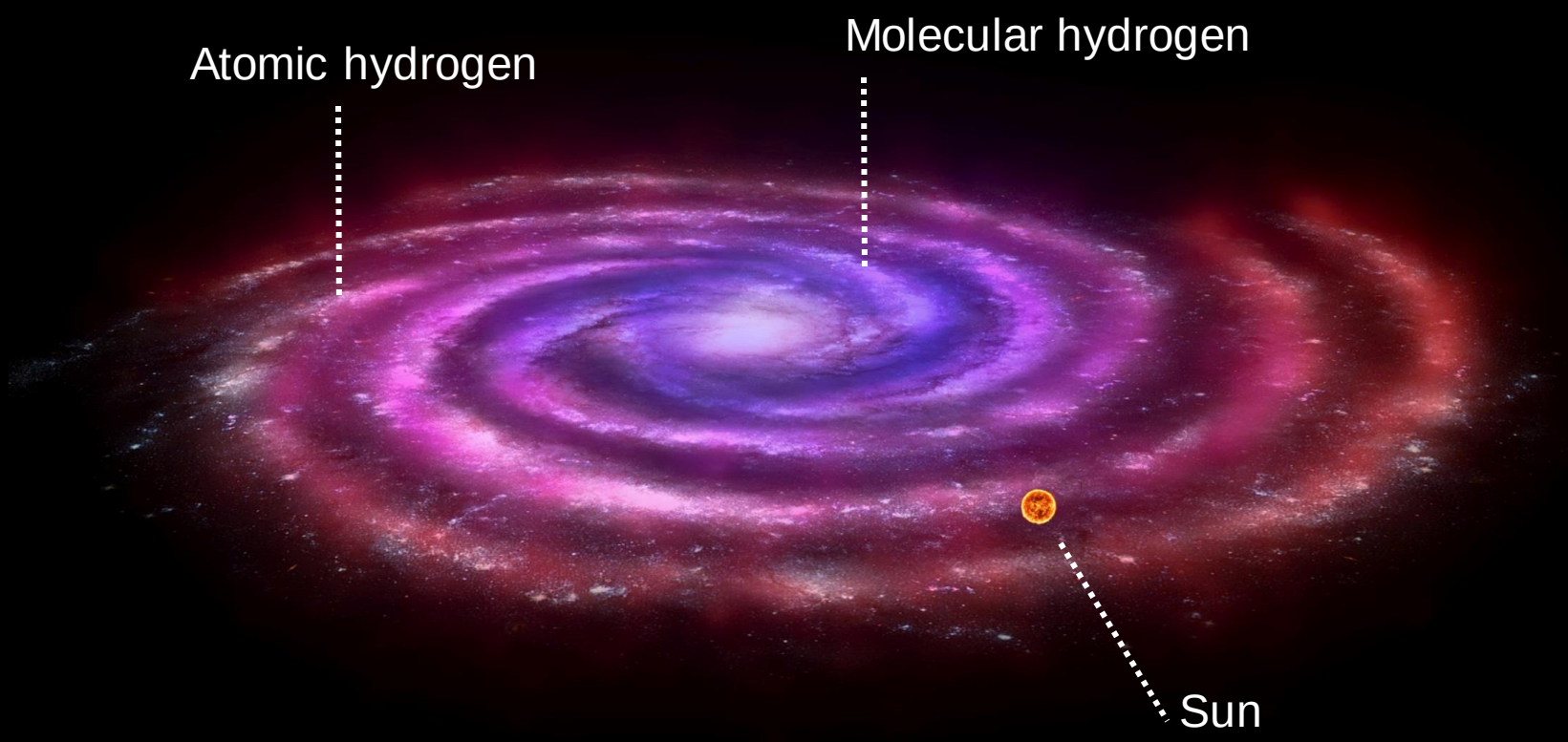
Line emission

$$\frac{dI}{ds} = j_c + j_l - \alpha_l I$$

Continuum emission

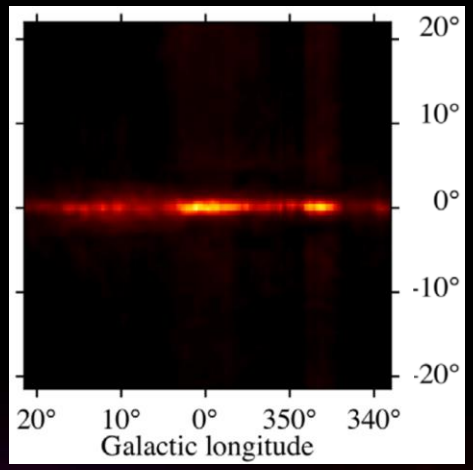
Absorption

Modeling Strategy: gas maps divided in rings



Modeling Strategy: gas maps divided in rings

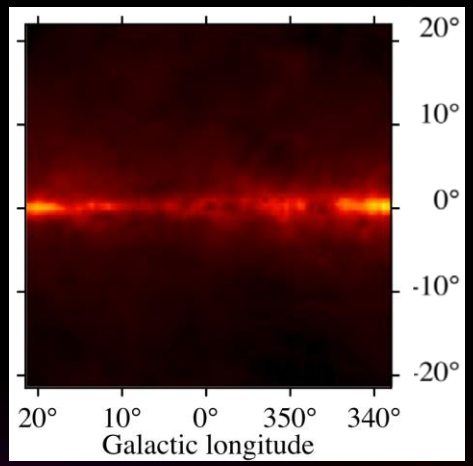
Modeling Strategy: gas maps divided in rings



0 – 3.5 kpc,
Hydrogen gas

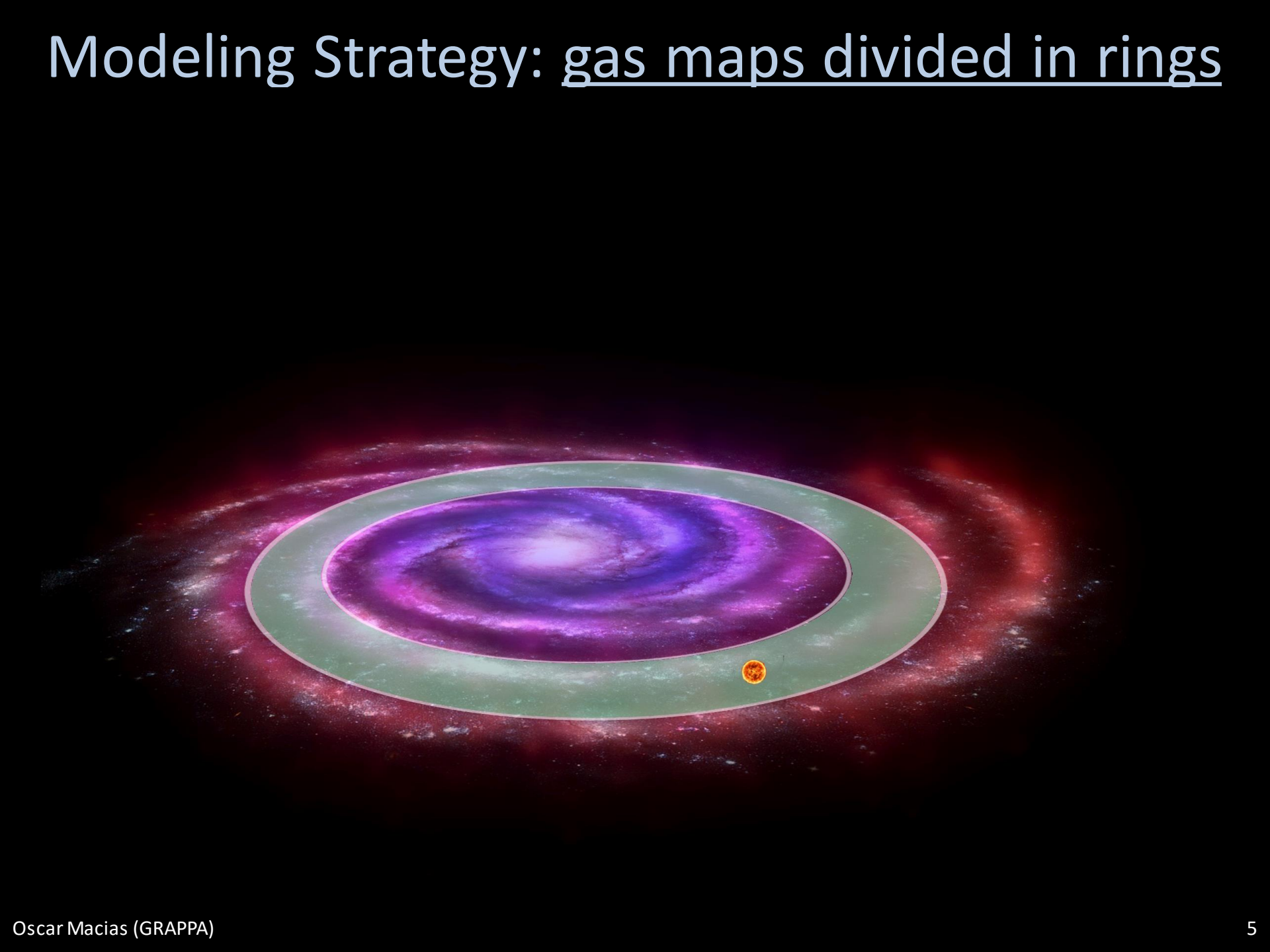
Modeling Strategy: gas maps divided in rings

Modeling Strategy: gas maps divided in rings

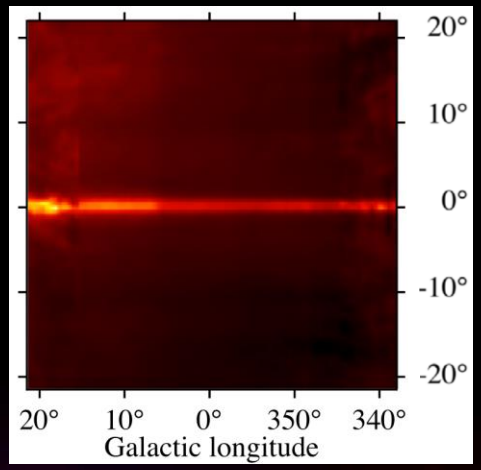


3.5 - 8.0 kpc,
Hydrogen gas

Modeling Strategy: gas maps divided in rings

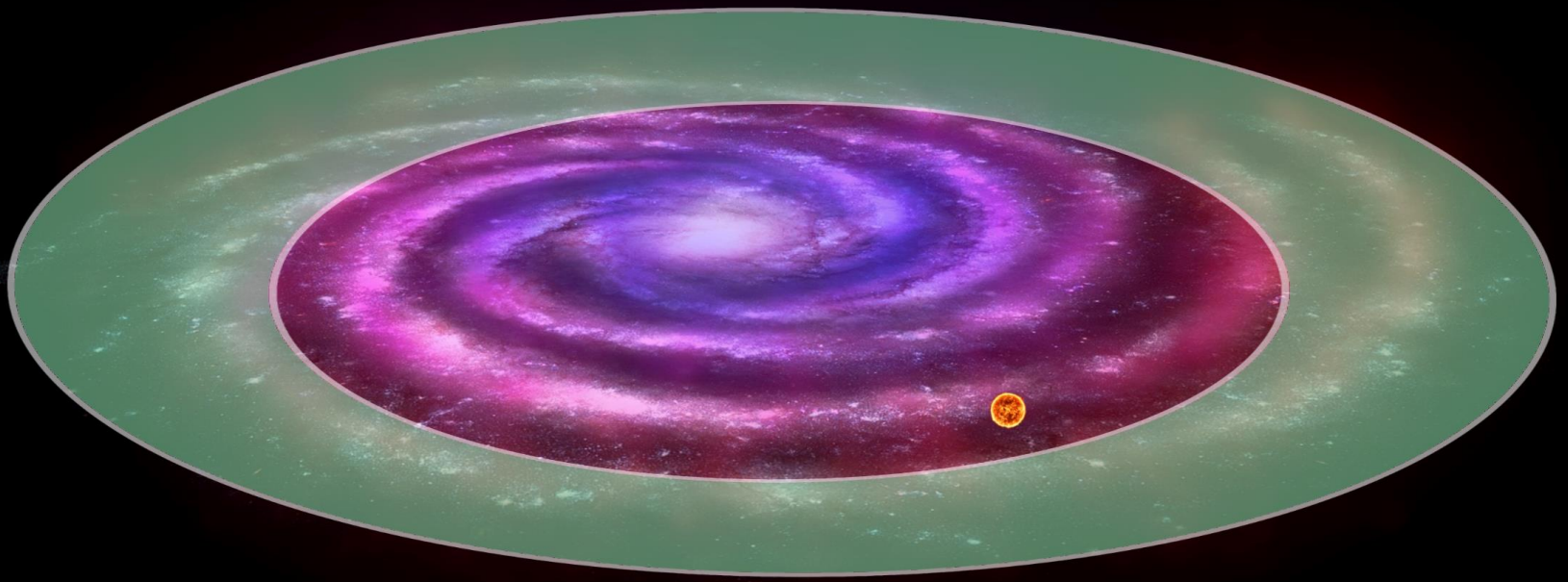


Modeling Strategy: gas maps divided in rings

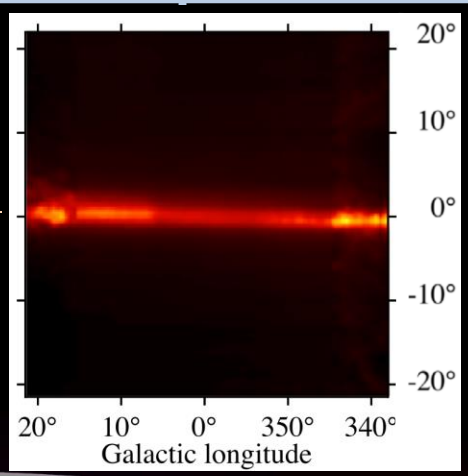


8 - 10 kpc,
Hydrogen gas

Modeling Strategy: gas maps divided in rings



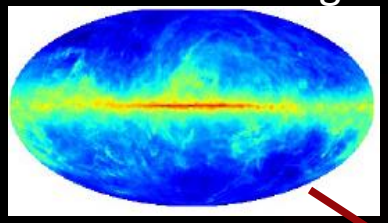
Modeling Strategy: gas maps divided in rings



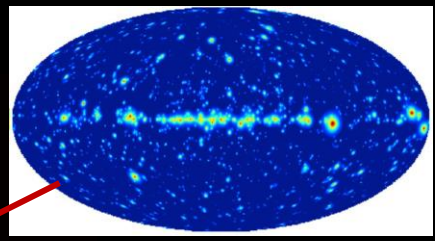
10 - 50 kpc,
Hydrogen gas

Modeling strategy: template fitting

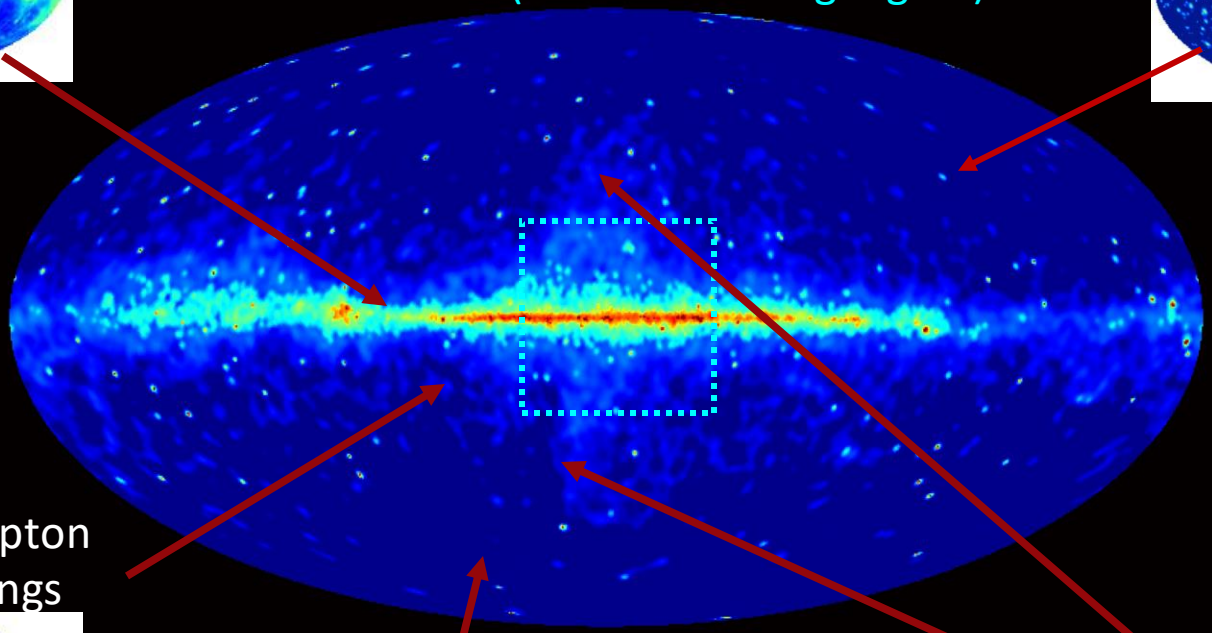
New hydrogen maps divided in 4 rings



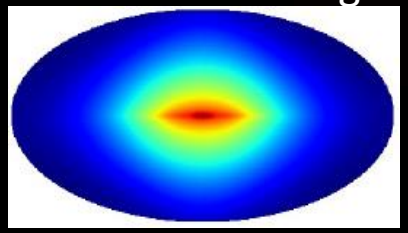
4FGL Catalog



Fermi-LAT data (inner 40x40 deg region)

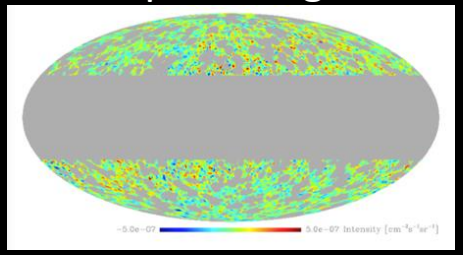


3D Inverse Compton divided in 6 rings

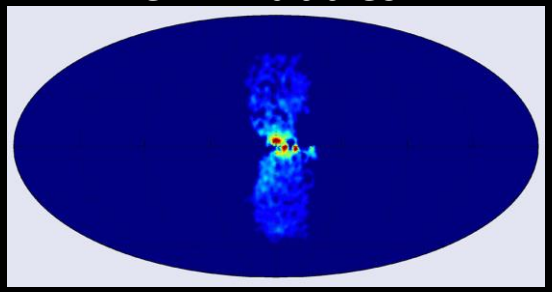


Porter et al. (2017)

Isotropic background

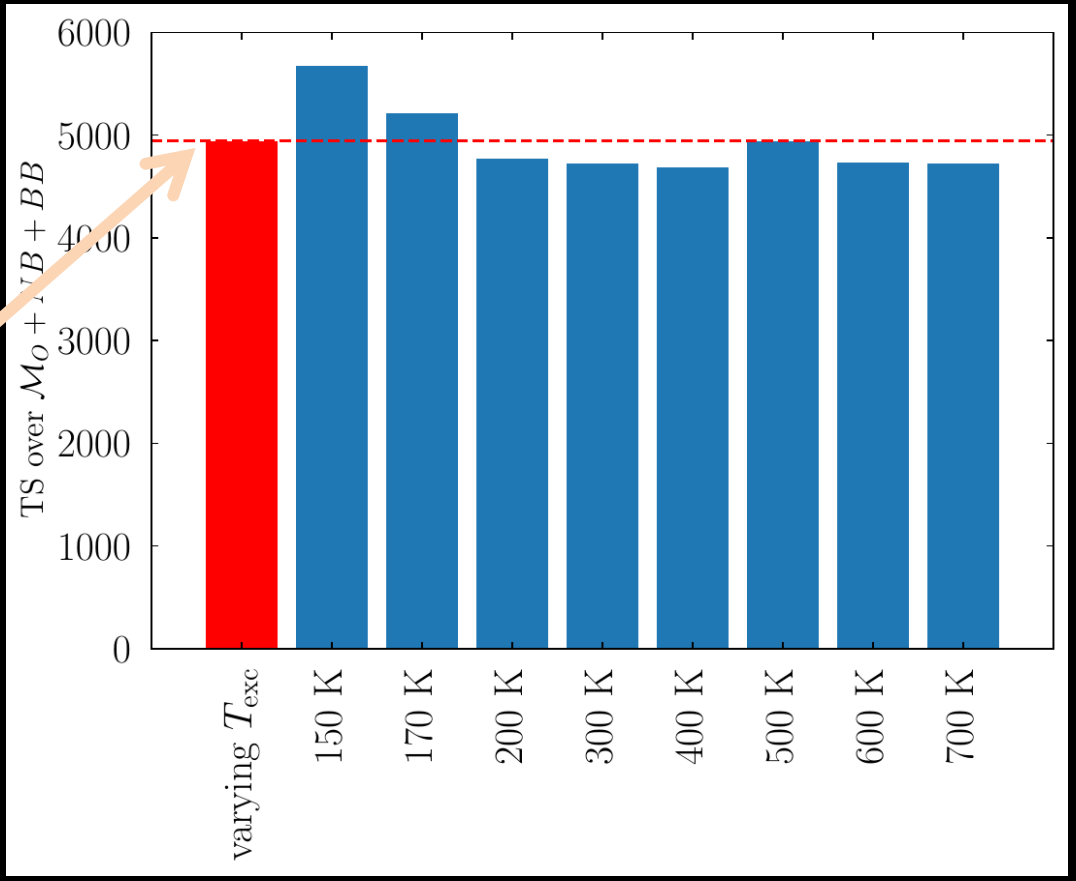


Fermi Bubbles



New background model **much better**

Significant improvement over our previous hydrodynamic maps



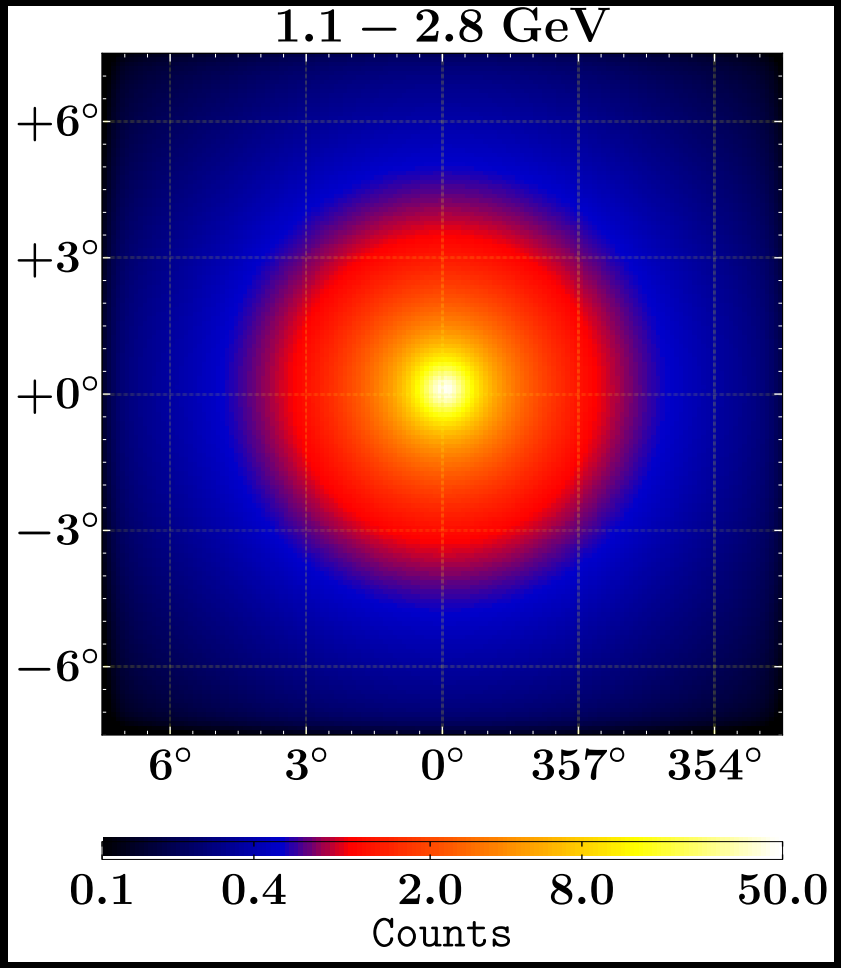
Test Statistic over our old background model

Pohl, OM, Coleman, Gordon (2022)

→ New hydrogen maps are publicly available at **https://doi.org/10.5281/zenodo.6276721**

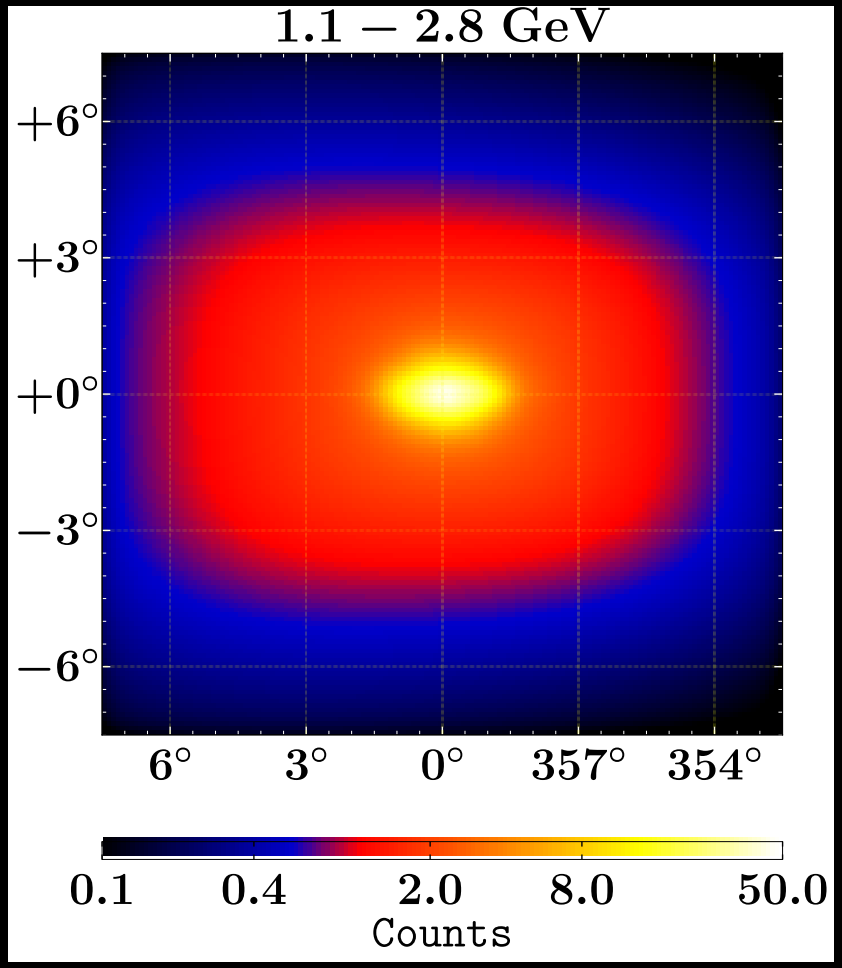
Spherical symmetry vs bulge shape

Navarro-Frenk-White profile



vs

Stellar Distribution



→ We want to test which spatial template is preferred by the data

Statistical Results

Base model includes: new gas maps, 3D IC, Fermi bubbles, 4FGL sources, isotropic

Baseline Model	Additional Source	ΔTS	Significance
Base	Cored ellips.	0.0	0.0σ
Base	Cored	0.1	0.0σ
Base	BB	282.2	15.3σ
Base	NFW ellips.	647.2	24.2σ
Base	NFW	807.1	27.3σ
Base	NB	1728.9	40.8σ
Base+NB	Cored ellips.	0.1	0.0σ
Base+NB	Cored	0.7	0.0σ
Base+NB	NFW ellips.	1.0	0.0σ
Base+NB	NFW	3.4	0.2σ
Base+NB	BB	261.0	14.7σ
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Base+NB+BB	Cored ellips.	0.4	0.0σ
Base+NB+BB	Cored	0.7	0.0σ
Base+NB+BB	NFW	2.6	0.1σ

Pohl, OM, Coleman, Gordon (2022)

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Pohl, OM, Coleman, Gordon (2022)

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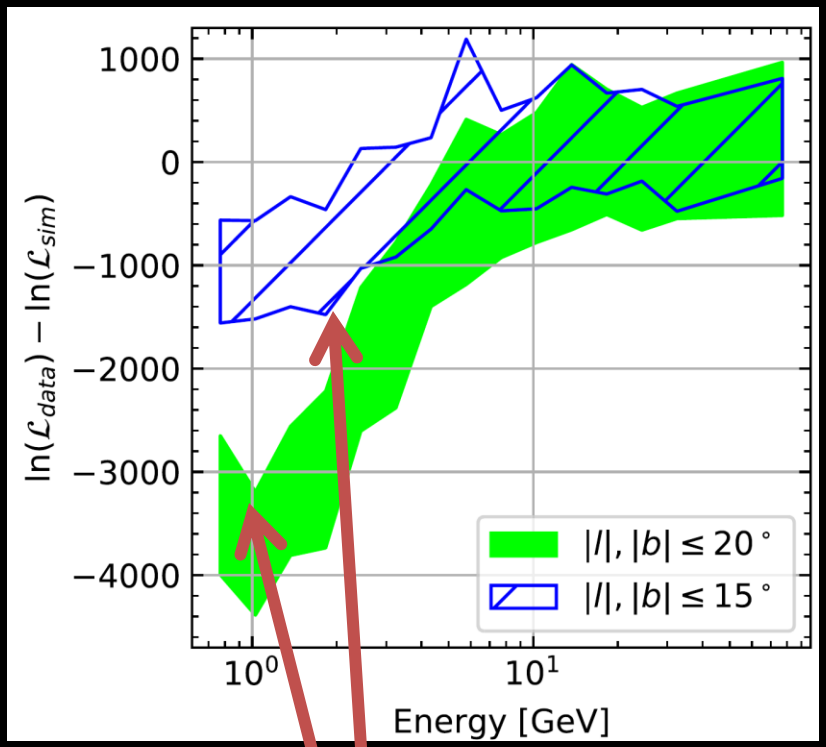
See also Macias et al. Nat. Astr. (2018) and Bartels et al. Nat.Astr. (2018)

Goodness of the fit

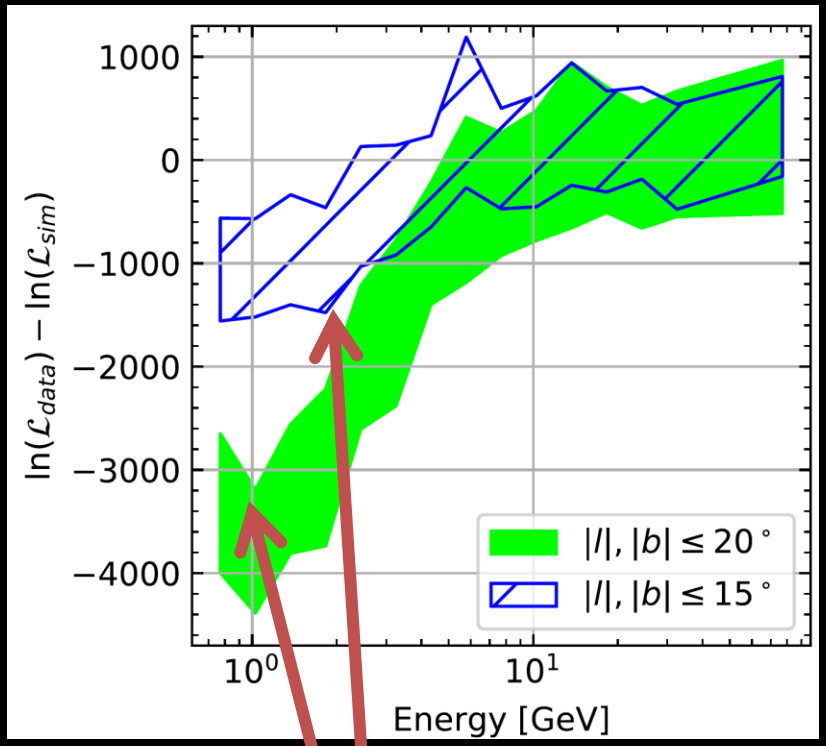
Pohl, OM, Coleman, Gordon (2022)



Comparisons with Monte Carlo Expectations (point sources fitted)

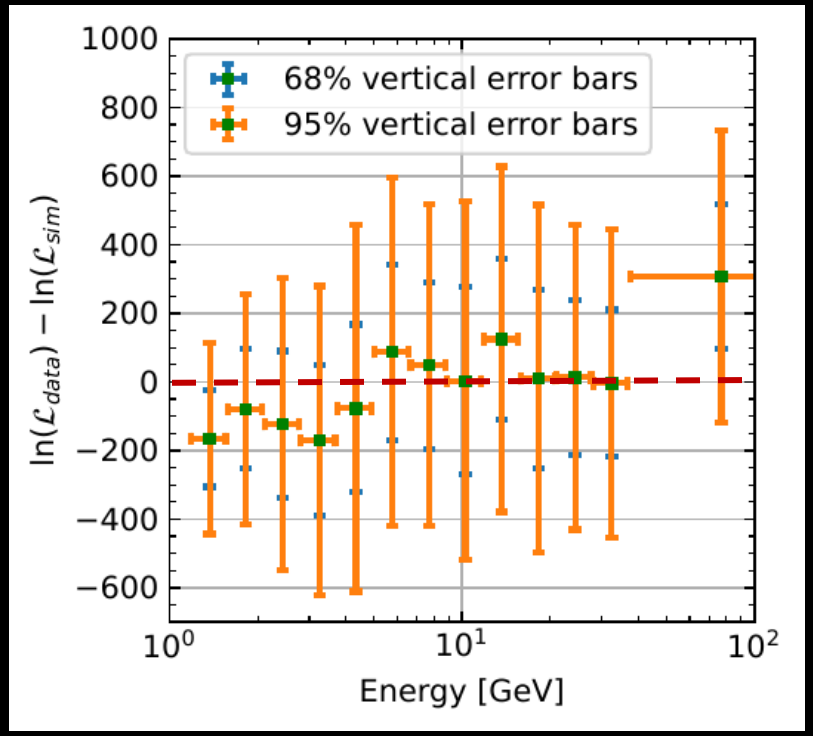
Goodness of the fit

Pohl, OM, Coleman, Gordon (2022)



Comparisons with Monte Carlo Expectations (point sources fitted)

Gordon, OM, Pohl (In preparation)



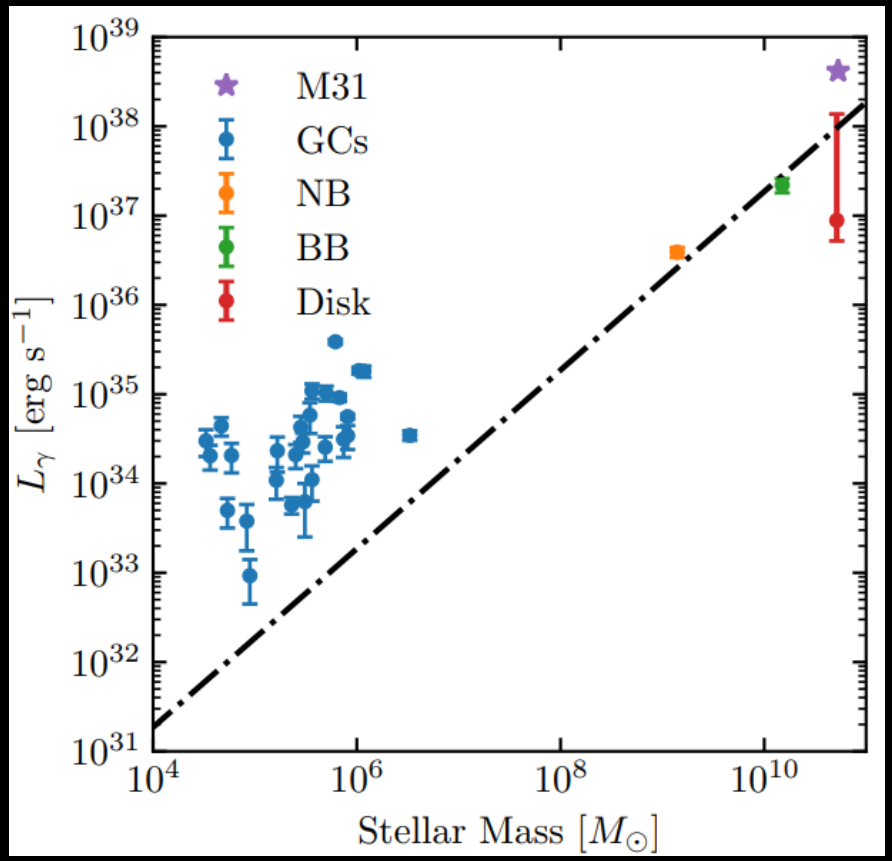
Log-likelihood results agree with MC expectations when point sources are masked

Comparison of regions

Gamma ray to mass ratios
Shows environmental dependence of MSP gamma rays

- Boxy bulge, nuclear bulge, Milky Way disk (from MSPs) consistent with each other
- Globular clusters higher by factor ~10-40, explained by large dynamical channel
- M31 also higher, consistent with its higher encounter rate.

Song, OM, Horiuchi, Crocker, Nataf (2021)



See also Bartels et al (2018), Eckner et al (2018)

Slide credit: S. Horiuchi

Conclusions

We have released new hydrogen gas maps which account for continuum emission and absorption.

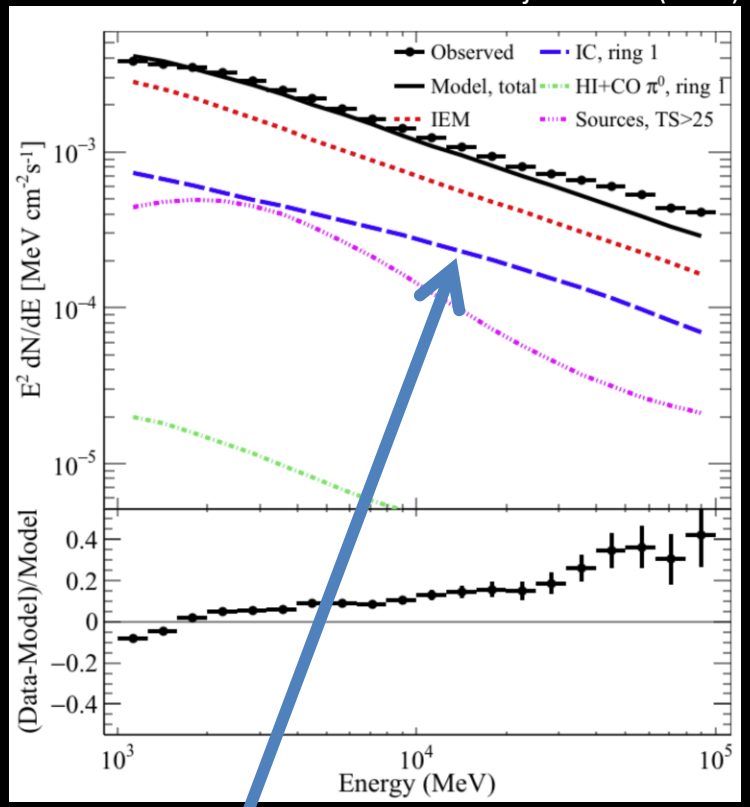
The new gas maps provide a much better fit to Fermi-LAT observations of the Galactic center.

We confirm previous results that the stellar bulge hypothesis is preferred to the dark matter explanation for the GCE.

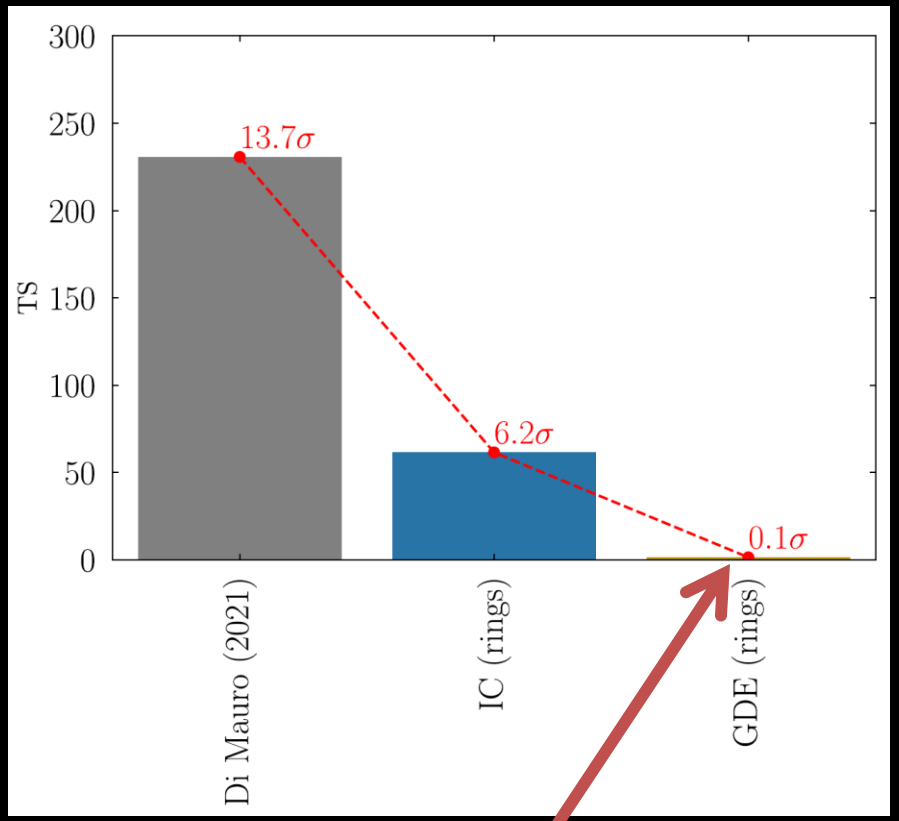
BACK-UP SLIDES

Flexibility of the Galactic diffuse emission

Ajello et al. (2015)



Pohl, OM, Coleman, Gordon (2022)



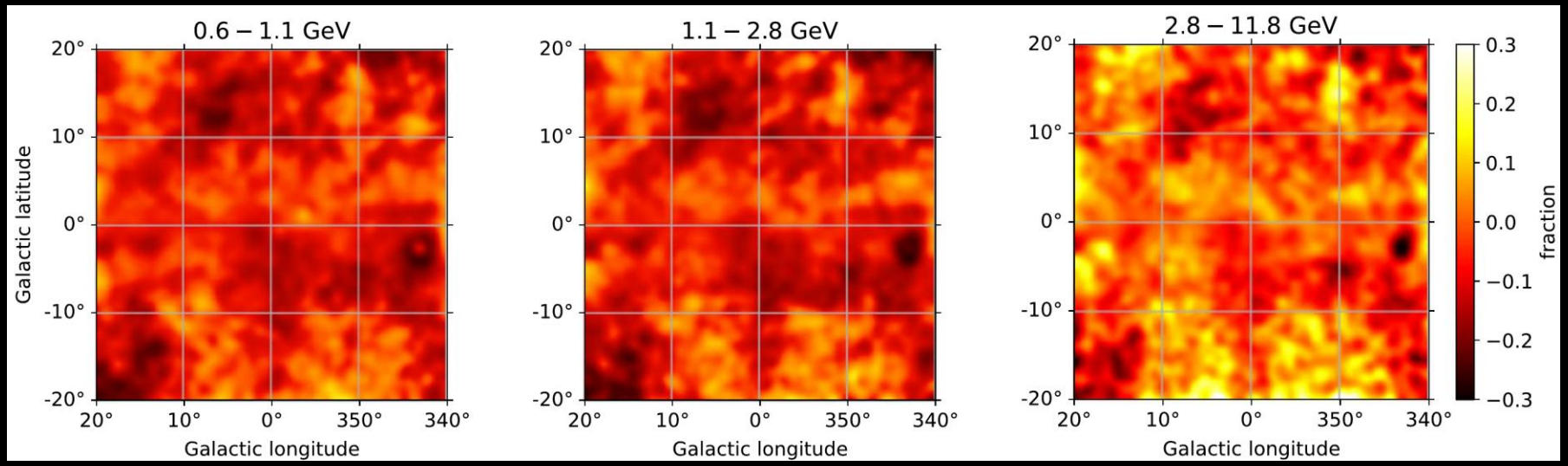
Enhanced IC emission in the inner Galaxy

The evidence for dark matter goes away when the GDE models are divided in rings

Residual Images

Residuals are small, except in a few localized regions

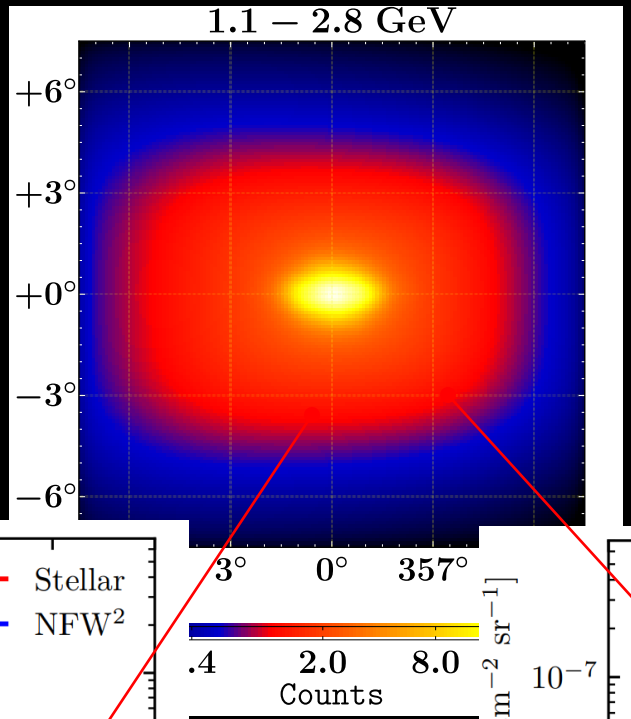
Pohl, OM, Coleman, Gordon (2022)



Some correlated residuals remain, possibly due to mismodeling of point sources and/or the Fermi bubbles.

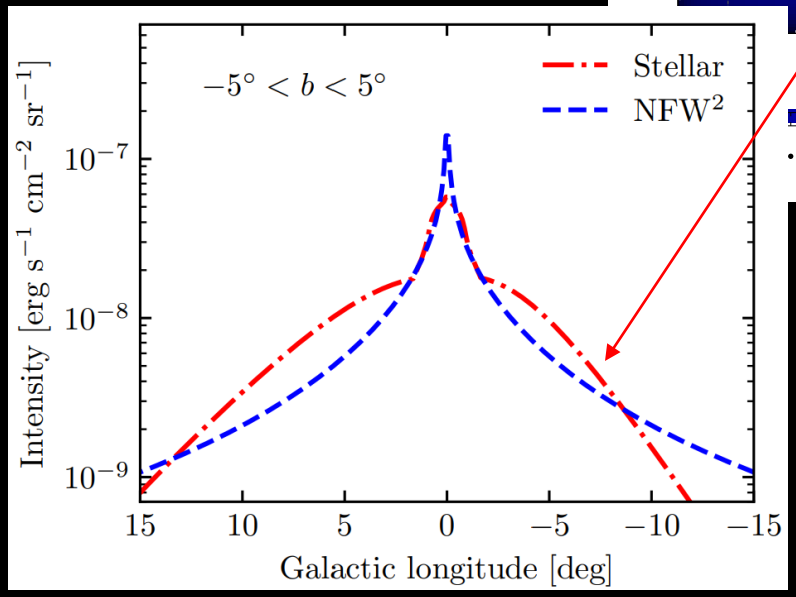
How are the two distinguished?

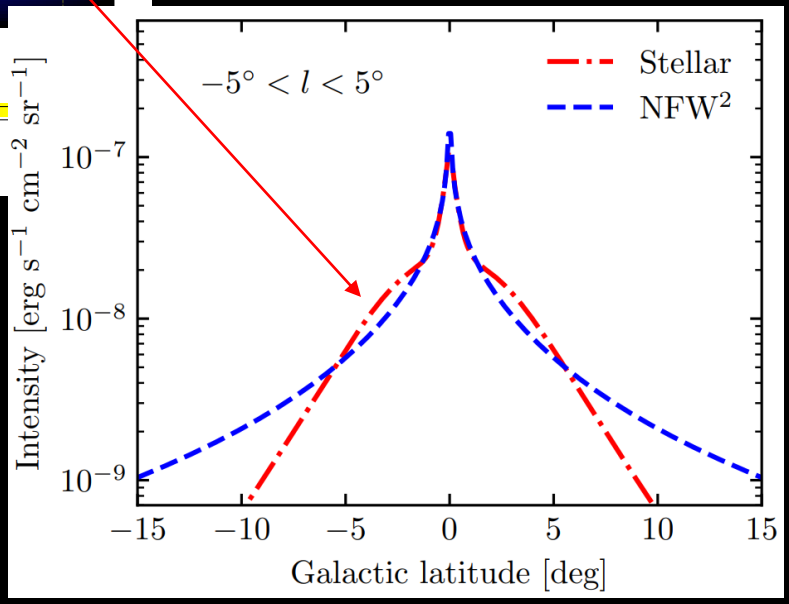
Q: why didn't previous dark matter model deformation studies discover a deviation from sphericity?



A: Tilted bar seen from the side has unique radius-dependent longitudinal shape

Song, OM, Horiuchi (2019)





nature astronomy

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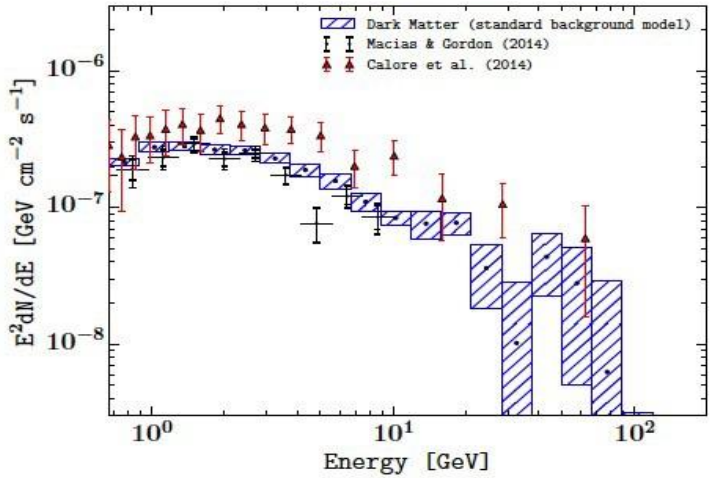
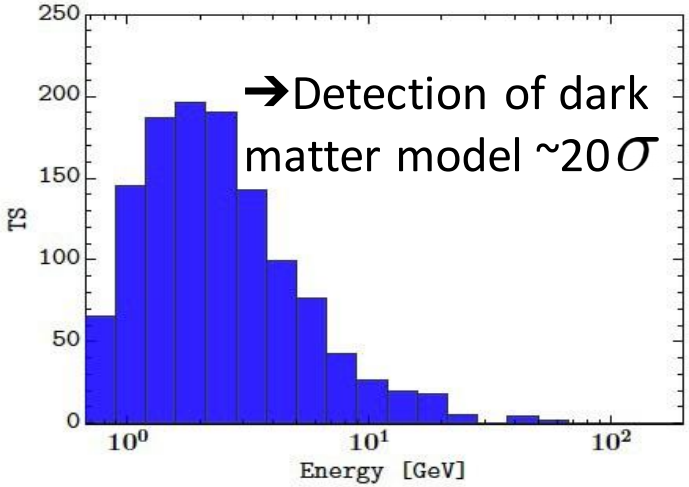
https://doi.org/10.1038/s41550-018-0414-3

Galactic bulge preferred over dark matter for the Galactic centre gamma-ray excess

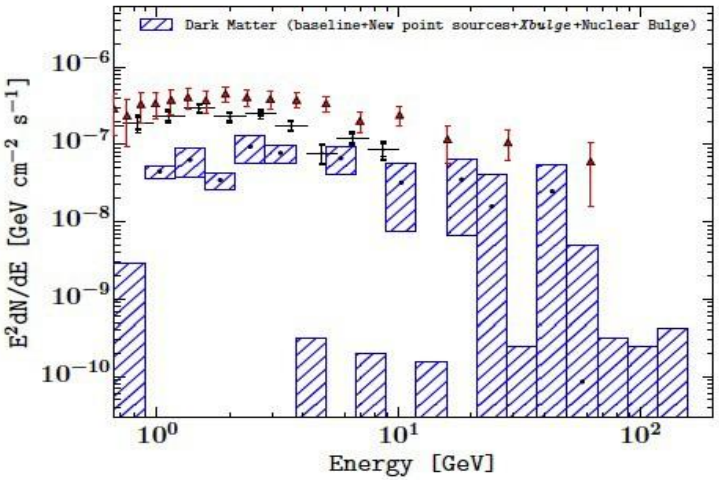
Macias+(2018)

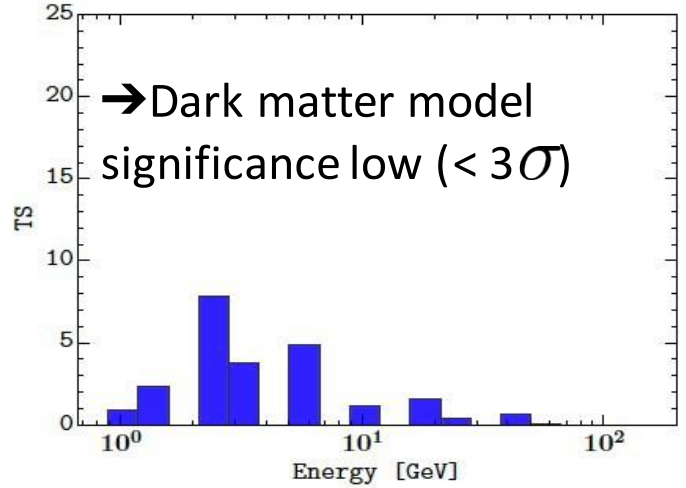




➔Detection of dark matter model ~20σ

↖ **WITHOUT bulge** (representative of previous studies)





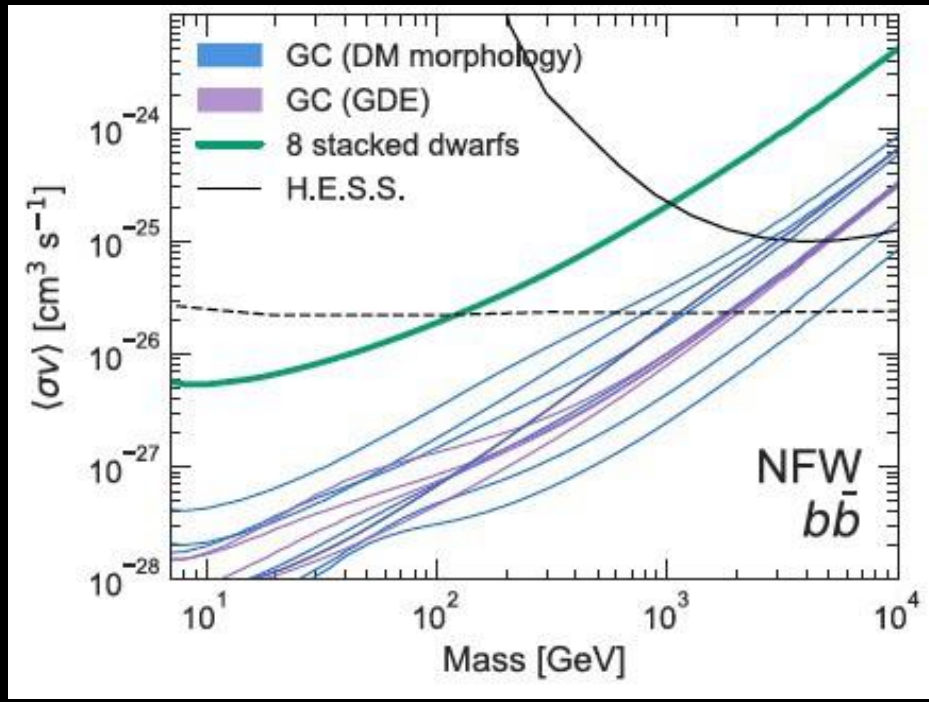
➔Dark matter model significance low (< 3σ)

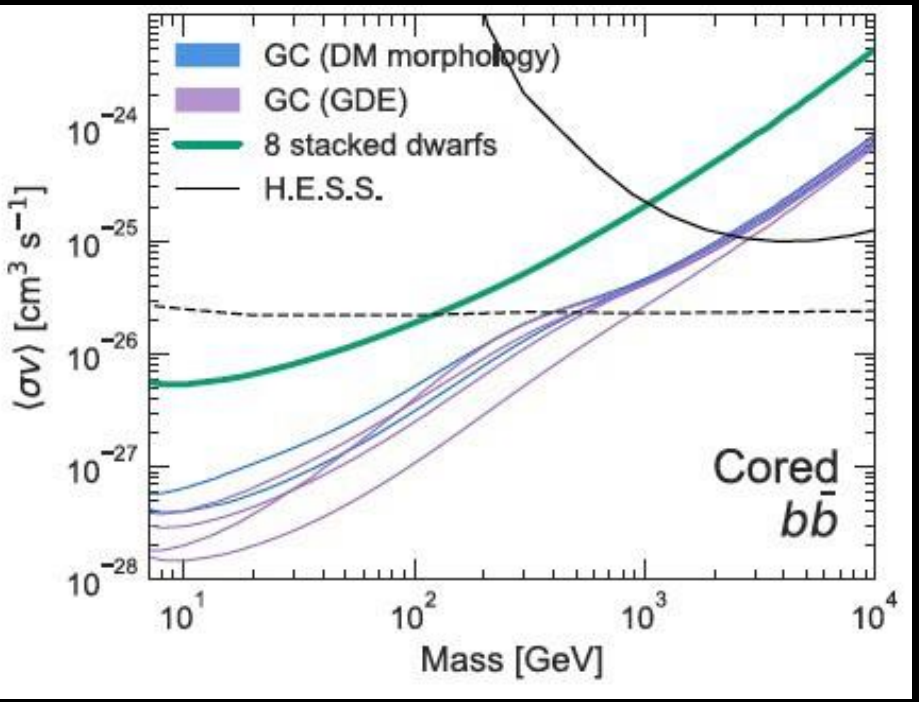
↖ **WITH bulge**
Including our bulge model the data no longer needs a dark matter component

Improved sensitivity to dark matter

We addressed a major systematic, which allowed us to realize the potential of the Galactic Center to constrain dark matter

Abazajian, Horiuchi, Kaplinghat, Keeley, OM (2020)





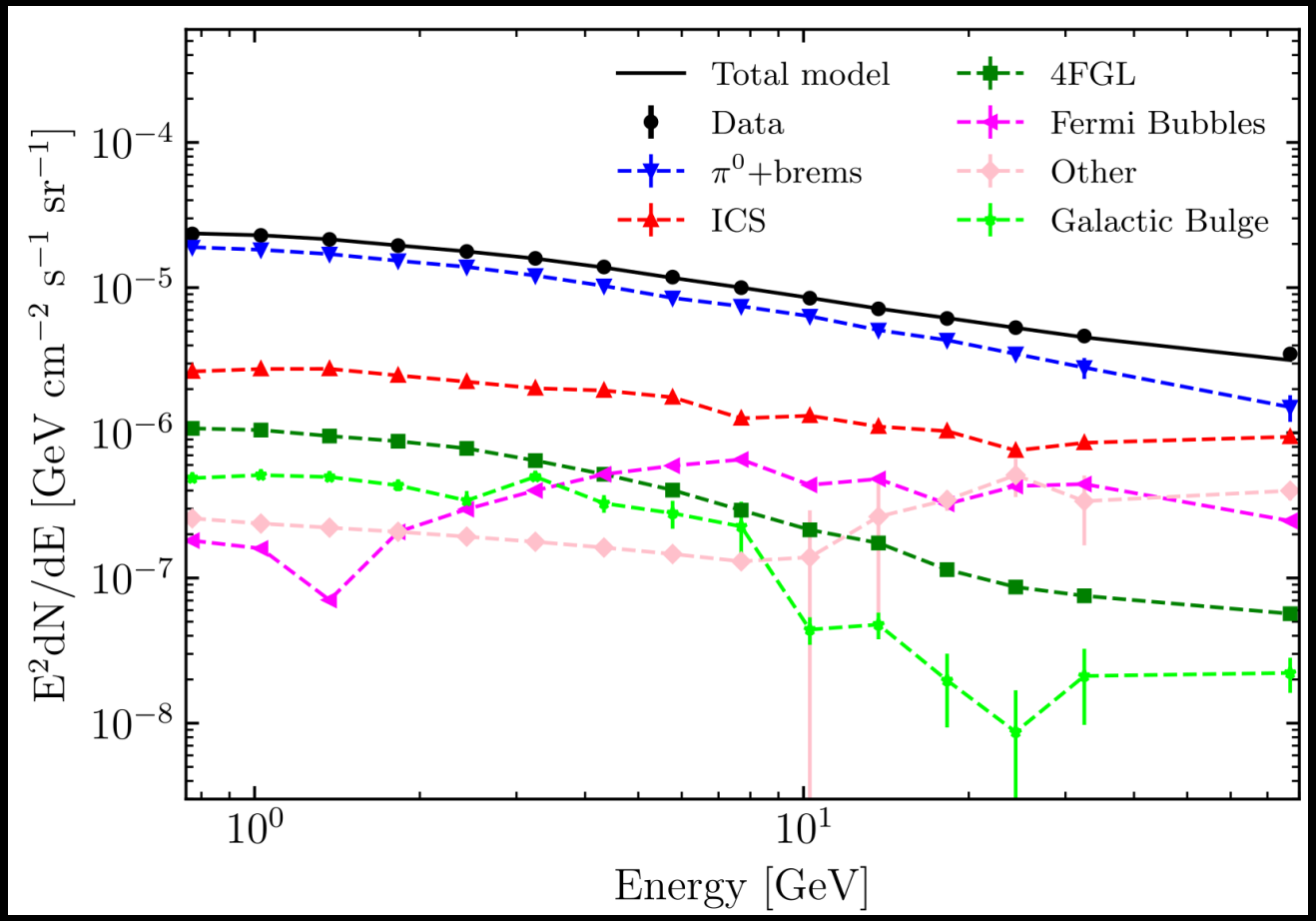
- Impacts of NFW slope [0.5,1.5] & sphericity
- Impacts of background modeling

- Impacts of core (1 kpc) & sphericity
- Impacts of background modeling

→ Tests thermal dark matter out to ~500 GeV

Spectrum of the Fermi GeV excess

Pohl, OM, Coleman, Gordon (2022)



→ **Spectrum of the stellar bulge template**

Checks for systematics

Gas maps: using the gas maps used by the Fermi Diffuse models yield the same conclusions

Macias et al. (2018)

Base	Source	$\log(\mathcal{L}_{\rm Base})$	$\log(\mathcal{L}_{\rm Base+Source})$	$\rm TS_{Source}$	σ	Number of source parameters
baseline+NB+Boxy	NFW	-172005.9	-171999.0	13.8	1.4	19
baseline+NFW	NB+Boxy	-172167.9	-171999.0	337.8	18.3	2×19
baseline*	NFW	-173565.0	-172929.2	1272	34.6	19
baseline*+NFW	NB+Boxy	-172929.2	-172533.0	792.4	28.2	2×19
baseline*+NB+Boxy	NFW	-172547.4	-172533.0	28.8	3.0	19

Point sources: using none or the 2FIG point source catalog yield the same conclusions

baseline	2FIG	-172461.4	-170710.5	3501	37.3	81×19
baseline+2FIG	Boxy	-170710.5	-170536.3	348.4	18.7	19
baseline+2FIG	NFW	-170710.5	-170484.6	452	19.9	19
baseline+2FIG	NB	-170710.5	-170470.5	480	20.6	19
baseline+2FIG+NB	NFW	-170470.5	-170387.8	165	11.1	19
baseline+2FIG+NB	Boxy	-170470.5	-170317.2	306.6	17.5	19
baseline+2FIG+NB+Boxy	NFW	-170317.2	-170313.5	7.4	0.5	19

Galactic plane mask: using a $|b| < 1$ deg mask yields the same conclusions

baseline	NFW	-430824.6	-430696.9	255	14.4	19
baseline	Boxy	-430824.6	-430626.1	397	18.5	19
baseline	NP	-430824.6	-430189.9	1269	35.6	22×19
baseline+NP	NFW	-430189.9	-430097.0	186	12.0	19
baseline+NP	Boxy	-430189.9	-430035.8	308	16.1	19
baseline+NP+Boxy	NFW	-430035.8	-430026.3	19	2.0	19

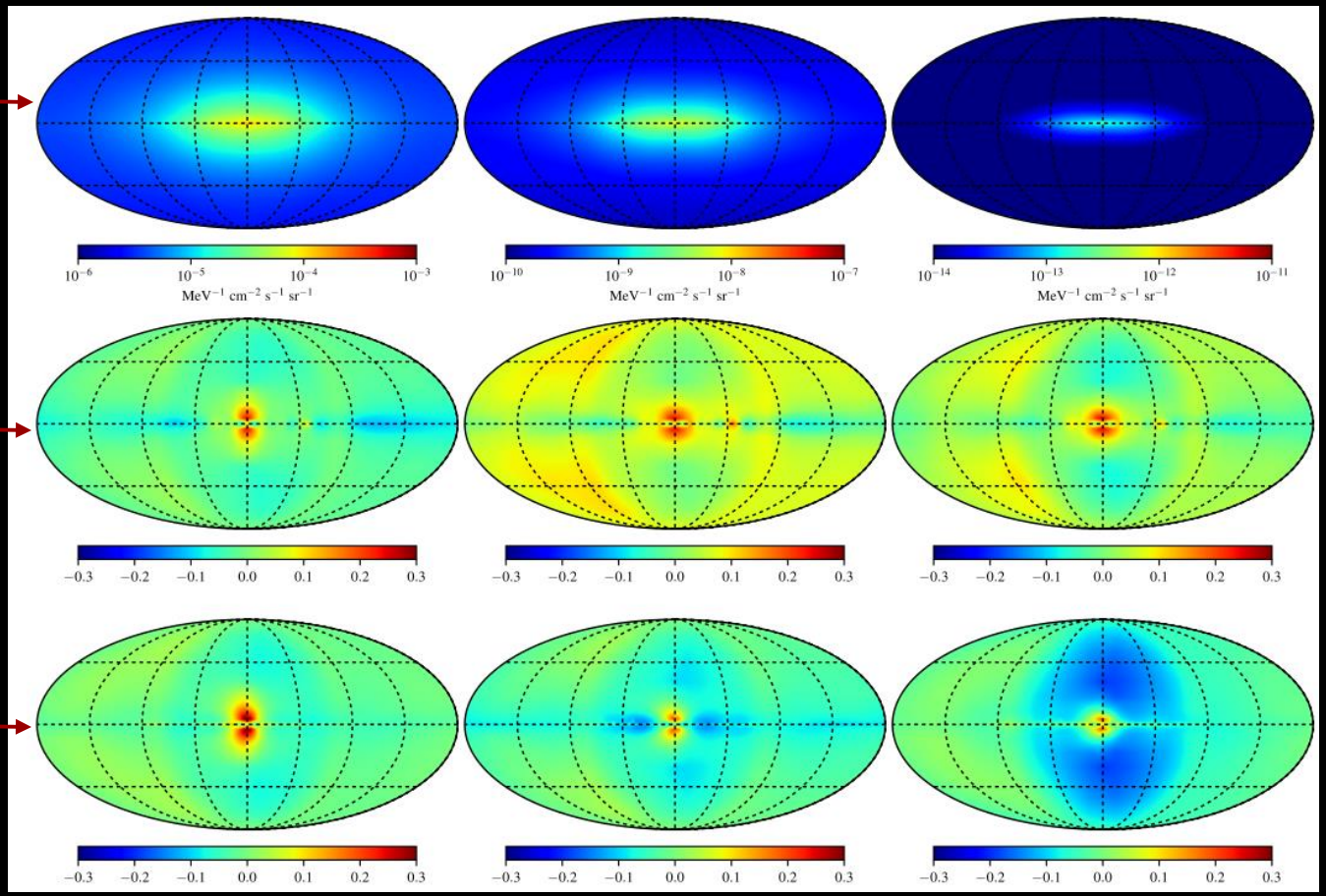
New 3D Inverse Compton Maps

10.6 MeV | 1.2 GeV | 79 GeV

Standard 2D IC

New 3D IC:
Fractional residual for
R12 Model

New 3D IC:
Fractional residual for
F98 Model

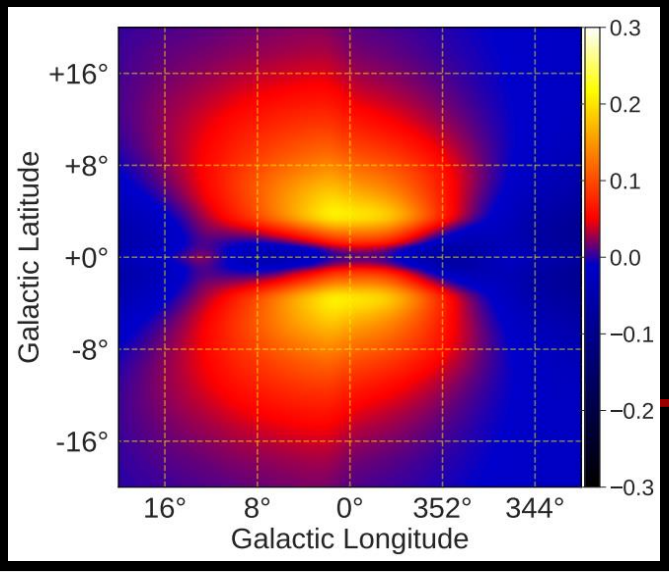


Porter et al. ApJ. (2017)

New 3D Inverse Compton maps

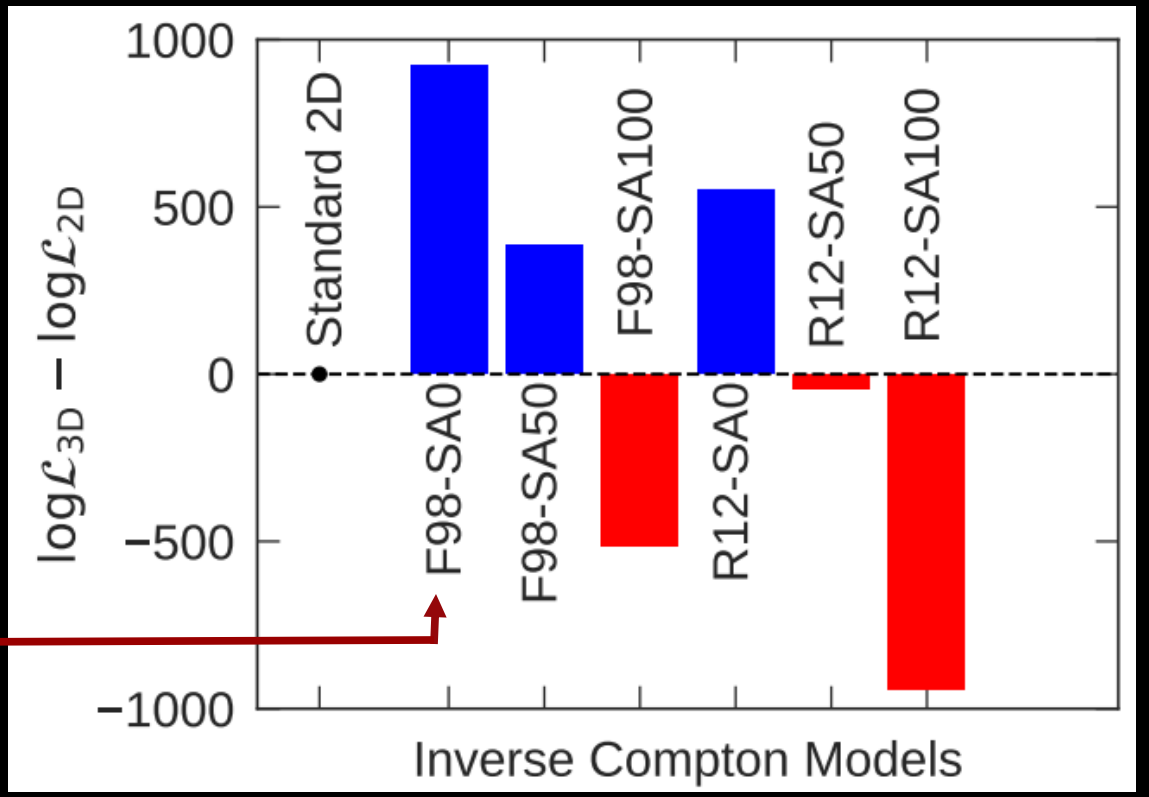
The data highly prefers the new 3D IC map *F98-SA0* over the Standard 2D IC map.

Macias et al. (2019)



Fractional residual for **F98** Model in the Galactic Center

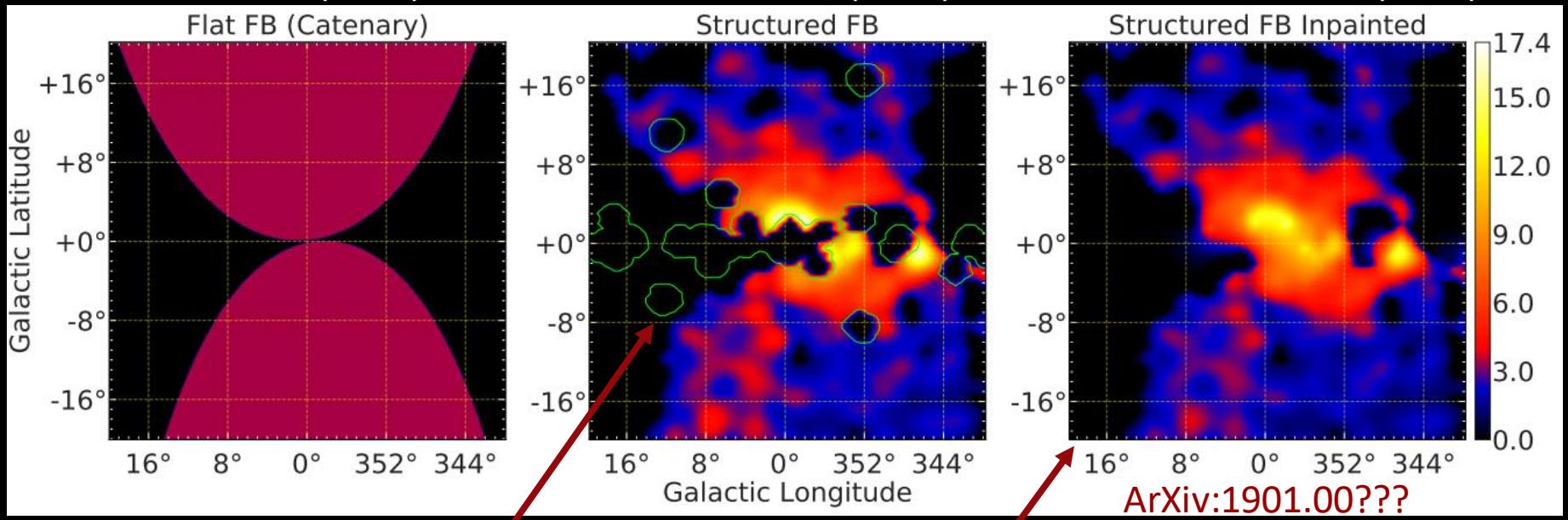
Macias et al. JCAP (2019)



Low-latitude Fermi Bubbles map

Used an inpainting method to correct artifacts in available maps

Acero et al. (2016) | Ackermann et al. (2017) | Macias et al. (2019)



Artifacts due to point source mask

Correction after inpainting method

Low-latitude Fermi Bubbles map

Inpainted Fermi bubbles map significantly improves the fit

Acero et al. (2016) Ackermann et al. (2017) Macias et al. (2019)

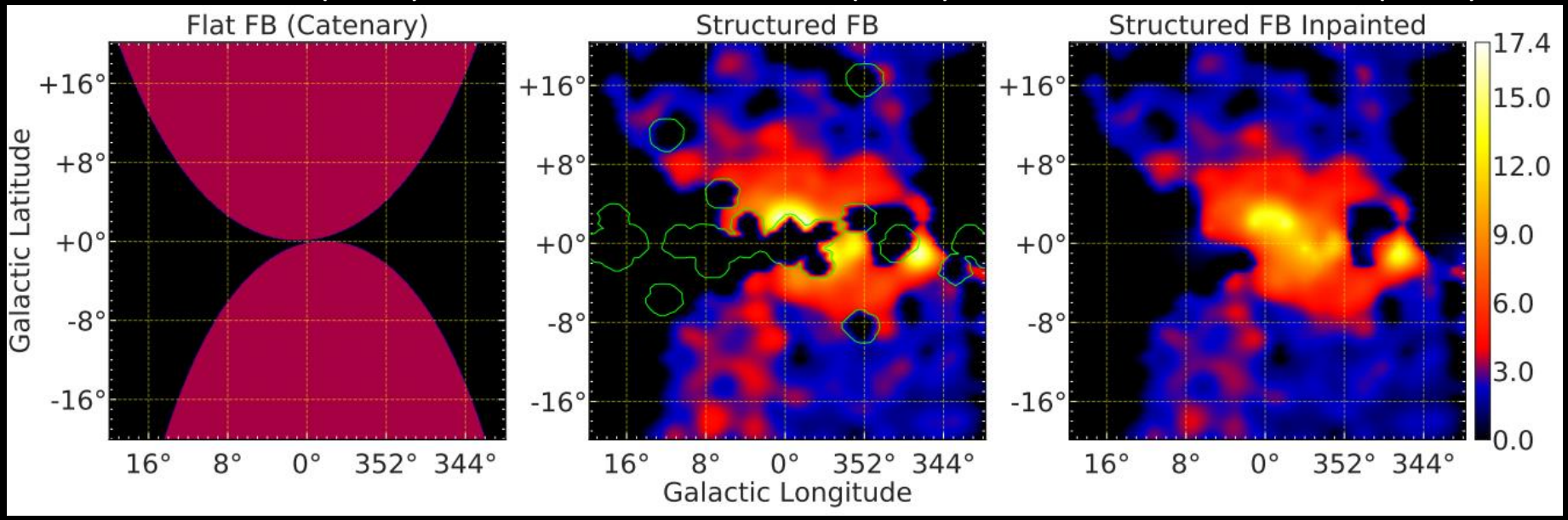


TABLE II. **Summary of the likelihood analyses for 3 alternative *Fermi* bubbles maps.**

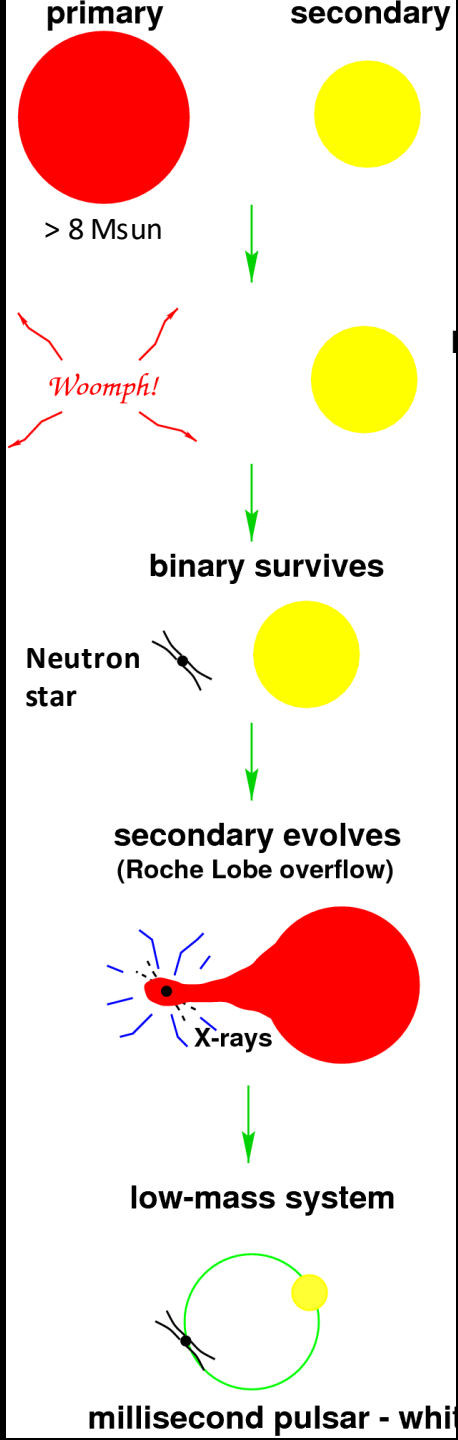
Base	Source	$\log(\mathcal{L}_{\rm Base})$	$\log(\mathcal{L}_{\rm Base+Source})$	$\mathrm{TS}_{\rm Source}$	Number of source parameters	Reference for FB template
baseline	Catenary	-2486188.1	-2486753.1	1130	15	[42]
baseline	Structured FB	-2486188.1	-2487322.3	2268	15	[43]
baseline	Structured FB (Inpainted)	-2486188.1	-2487802.1	3228	15	adapted from [43]

Macias et al. (2019)

Pulsar Recycling Scenario

Recycling of old neutron stars:

- ○ Initially a normal NS, but B-field decays or is buried in the superconducting core (resurface must be stopped)
- ✗ High natal kick velocities might wash out the stellar-bulge vs GCE spatial correlation [e.g., Boodram&Heinke (2022)].
- ✗ Expected a higher number of observed LMXBs in the GC [Haggard+(2017)].



nature astronomy

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https://doi.org/10.1038/s41550-022-01658-3

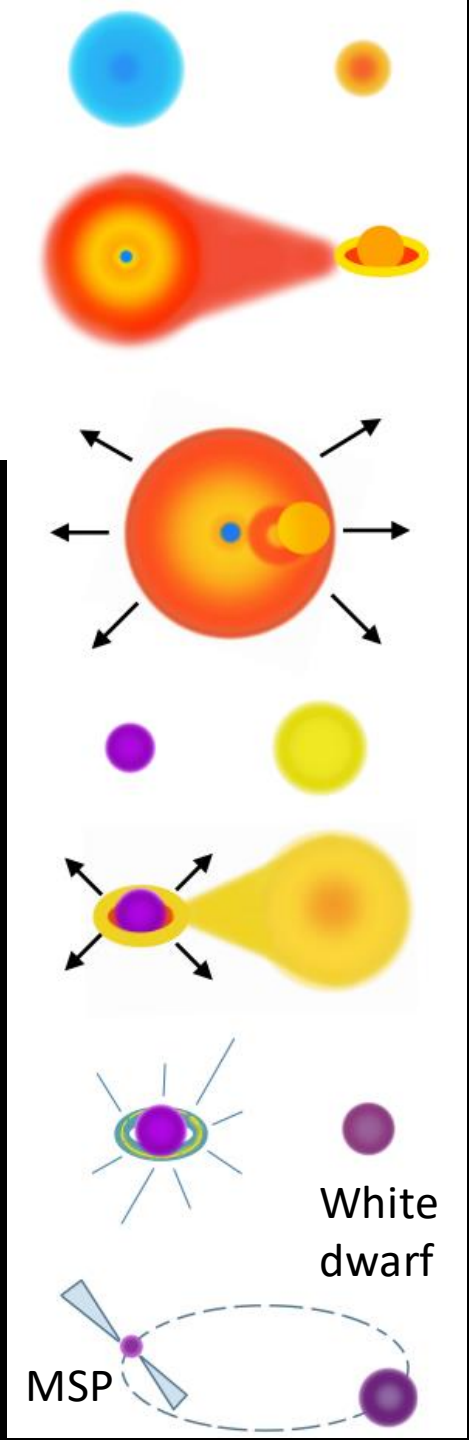
Check for updates

Millisecond pulsars from accretion-induced collapse as the origin of the Galactic Centre gamma-ray excess signal

Gautam, Crocker, Ferrario, Ruiter, Ploegg, Gordon, OM (2022)

Accretion-Induced Collapse of White dwarfs:

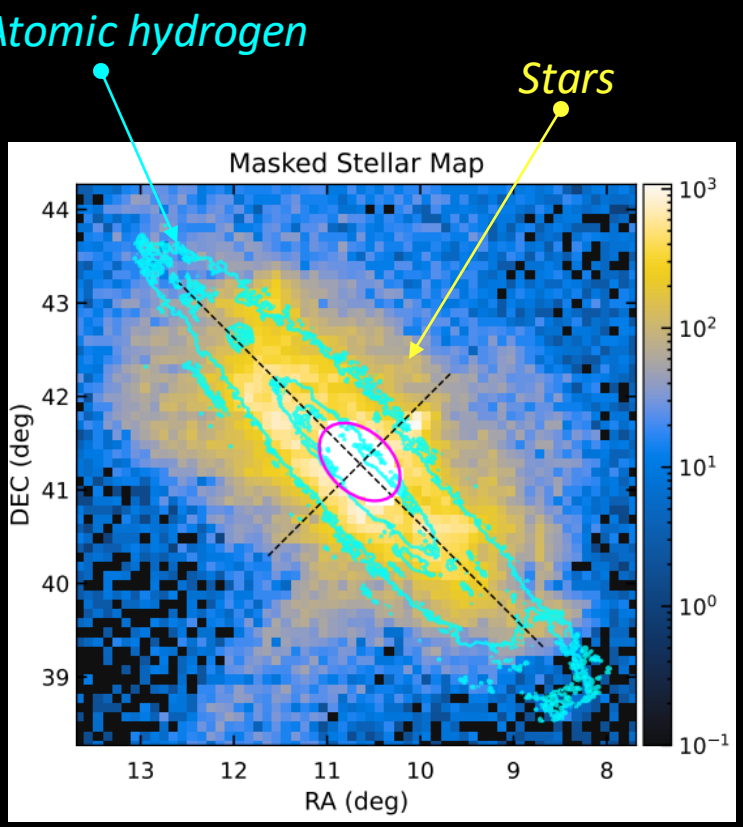
- Magnetic flux conservation yields the correct B-field.
- Very low or in-existent pulsar natal kicks (consistent with spatial correlation of GCE and stellar bulge).
- Does not generate LMXBs at any stage of the MSPs evolution.



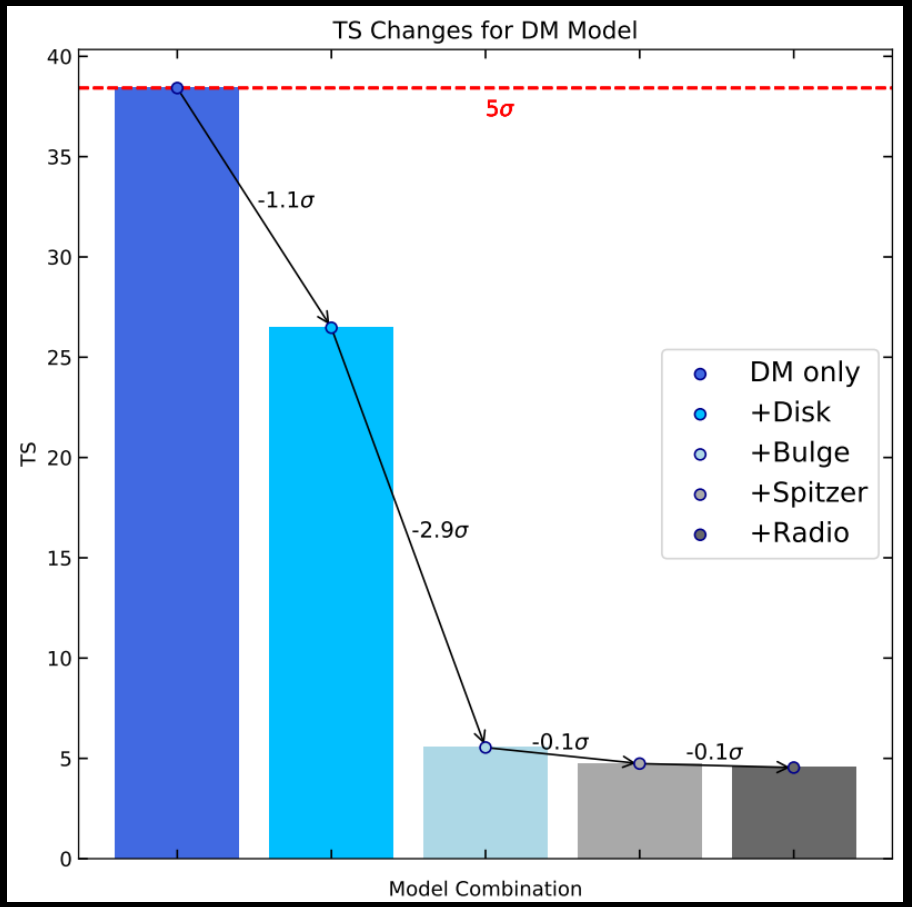
The Andromeda GeV excess

Gamma ray evidence for dark matter annihilations?

No evidence found once the stellar disk+bulge are included in the fit



Zimmer, OM, Ando, Horiuchi, Crocker (2022)

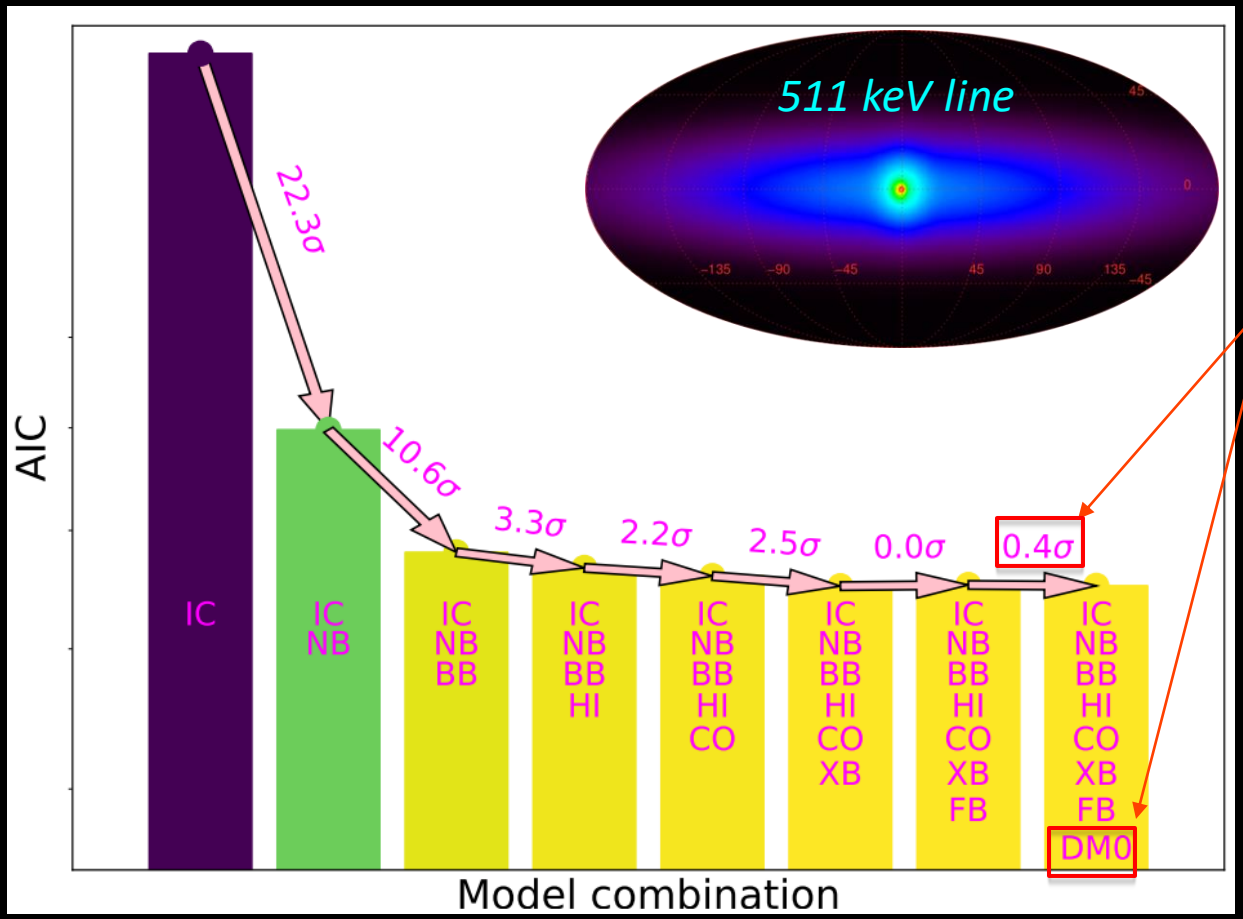


Zimmer, OM, Ando, Horiuchi, Crocker (2022)

See also Armand & Calore (2021), Di Mauro (2019)

Positron Annihilation excess @ GC

Positron annihilation evidence for dark matter annihilations?
No evidence found once the stellar disk+bulge are included in the fit



WITH bulge
Including our bulge model the data no longer needs a dark matter component

Siegert, Crocker, OM, Calore, Horiuchi, Song (2021)