# The $SU(2)_D$ lepton portals for $(g-2)_\mu$ , $M_{\rm W}$ and dark matter

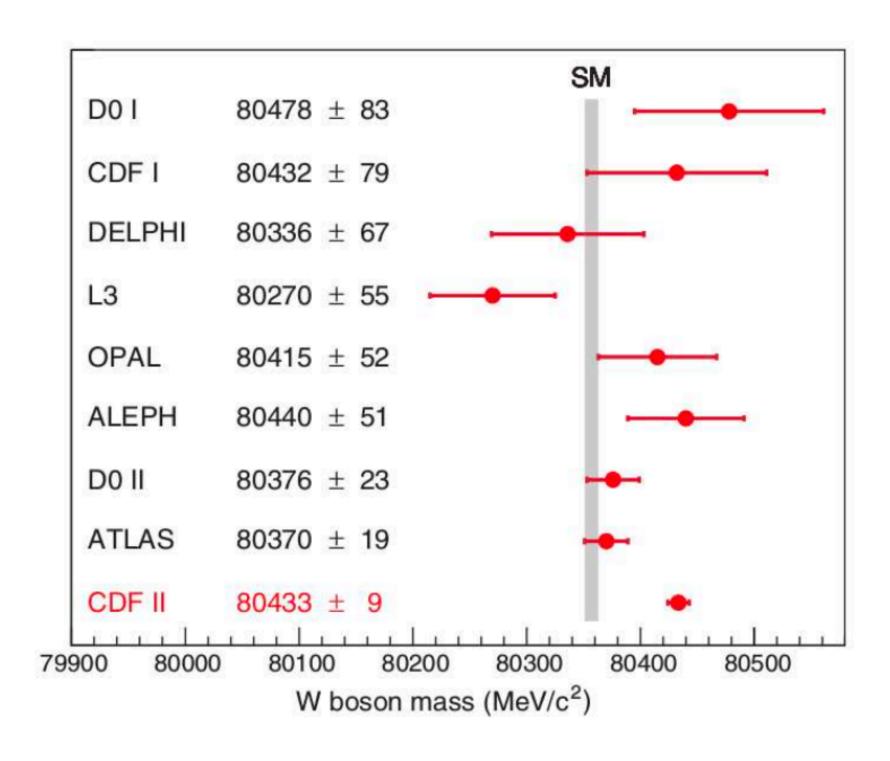
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Based on 2205.04016 - [PRD 106 (2022) 1, 015008]



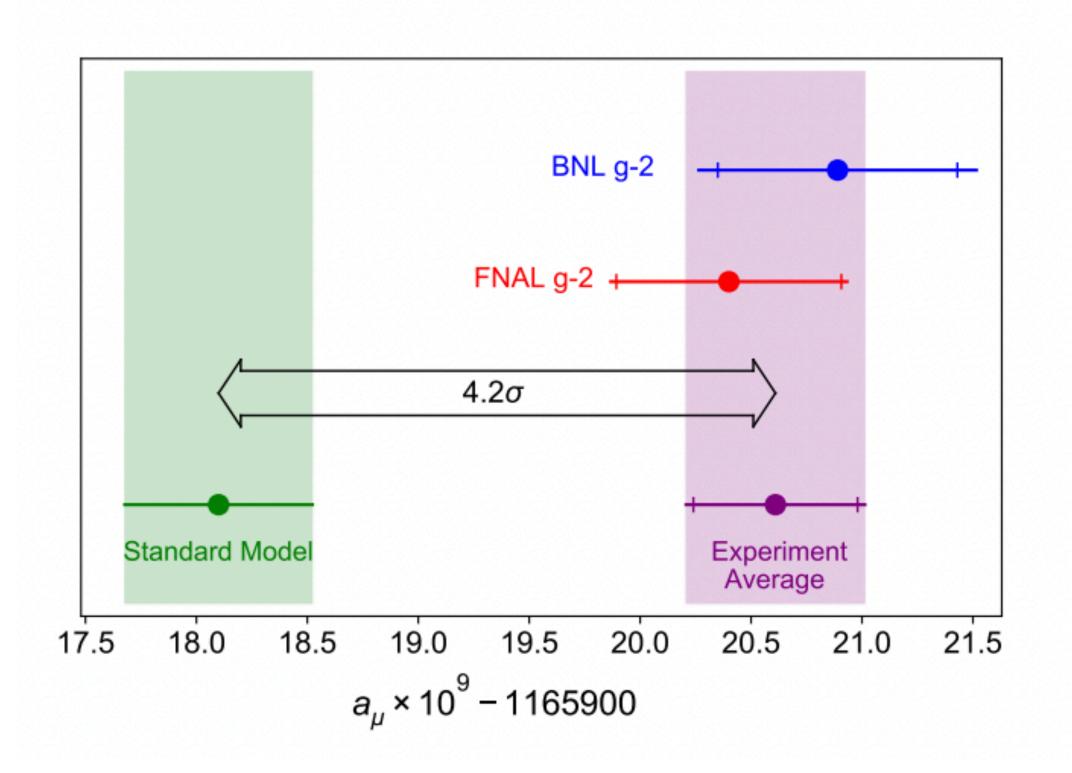
# $M_{ m W}$ and $(g-2)_{\mu}$ measurements



$$M_{\rm W}^{CDFII} = 80.4335 \text{ GeV} \pm 9.4 \text{ MeV}$$

$$M_{\rm W}^{SM} = 80.357 \,\,{\rm GeV} \pm 6 \,\,{\rm MeV}$$

 $7\sigma$  deviation



$$\Delta a_{\mu} = a_{\mu}^{\text{exp}} - a_{\mu}^{\text{SM}} = 251(59) \times 10^{-11}$$

 $4.2\sigma$  deviation

Both suggest new physics!
Can we explain both simultaneously?
Can we explain DM too?

## What kind of new physics?

We consider the SM + an extra local  $SU(2)_D$  symmetry.

The Z2 parity, originates from a combination of the dark isospin symmetry and a global  $U(1)_{G}$  symmetry in the Higgs sector

#### Seesaw lepton masses.

$$\mathcal{L}_{L,\text{mass}} = -M_E \bar{E}' E' - \left[ (\bar{e}_L, \bar{E}_L) \mathcal{M}_L \begin{pmatrix} e_R \\ E_R \end{pmatrix} + \text{h.c.} \right]$$

• After diagonalization: 
$$\begin{pmatrix} e_L \\ E_L \end{pmatrix} = U_L \begin{pmatrix} l_{1L} \\ l_{2L} \end{pmatrix}$$
,  $\begin{pmatrix} e_R \\ E_R \end{pmatrix} = U_R \begin{pmatrix} l_{1R} \\ l_{2R} \end{pmatrix}$ 

$$\mathcal{L}_{L,\text{mass}} = -m_{l_1}\bar{l}_1l_1 - m_{l_2}\bar{l}_2l_2 - M_E\bar{E}'E'$$

$$m_{l_1} \approx m_0 - \frac{m_R m_L}{M_E}$$
 $m_{l_2} \approx (M_E^2 + m_L^2 + m_R^2)^{1/2}$ 

$$\mathcal{M}_L = \begin{pmatrix} m_0 & m_L \\ m_R & M_E \end{pmatrix}$$

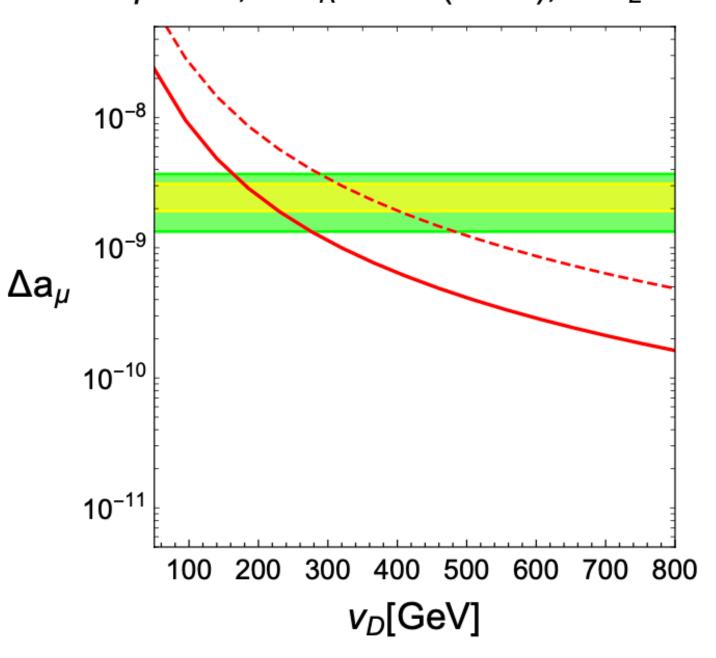
 $m_0$ : bare lepton mass  $m_R, m_L$ : mixing masses

• Lepton masses are naturally small since they are a result of a simultaneous symmetry breaking of  $SU(2)_D$  and the EW gauge symmetry,  $m_L \neq 0$  and  $m_R \neq 0$ .

## Contributions to the $(g-2)_{\mu}$ and $M_W$

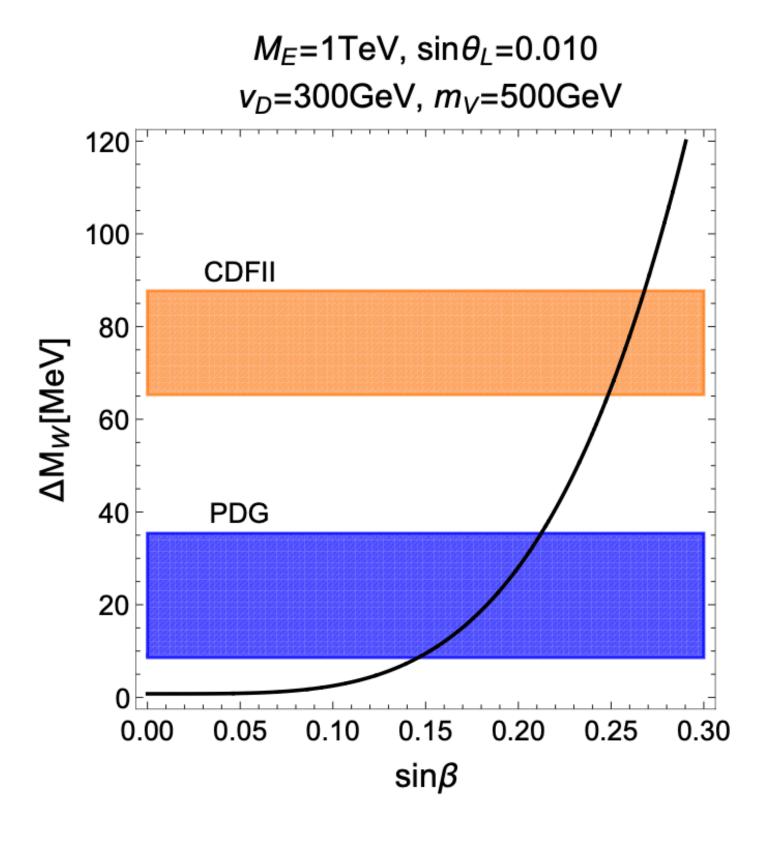
 $M_E$ =1TeV,  $m_V$ =500GeV  $m_{\phi}$ =800GeV  $\sin \beta$ =0.25,  $\sin \theta_R$ =0.011(0.033),  $\sin \theta_L$ =0.010

For the favored correction,  $m_{\phi}$  close to the TeV scale



$$\Delta a_{\mu}^{V,E} \simeq \left\{ \begin{array}{l} \frac{g_D^2 M_E m_{\mu}}{16\pi^2 m_{V^0}^2} \left(c_V^2 - c_A^2\right) + \frac{g_D^2 M_E m_{\mu}}{32\pi^2 m_{V^0}^2} \left(\hat{c}_V^2 - \hat{c}_A^2\right), & M_E \gg m_{V^0}, \\ \frac{g_D^2 M_E m_{\mu}}{4\pi^2 m_{V^0}^2} \left(c_V^2 - c_A^2\right) + \frac{g_D^2 M_E m_{\mu}}{8\pi^2 m_{V^0}^2} \left(\hat{c}_V^2 - \hat{c}_A^2\right), & m_{\mu} \ll M_E \ll m_{V^0}. \end{array} \right.$$

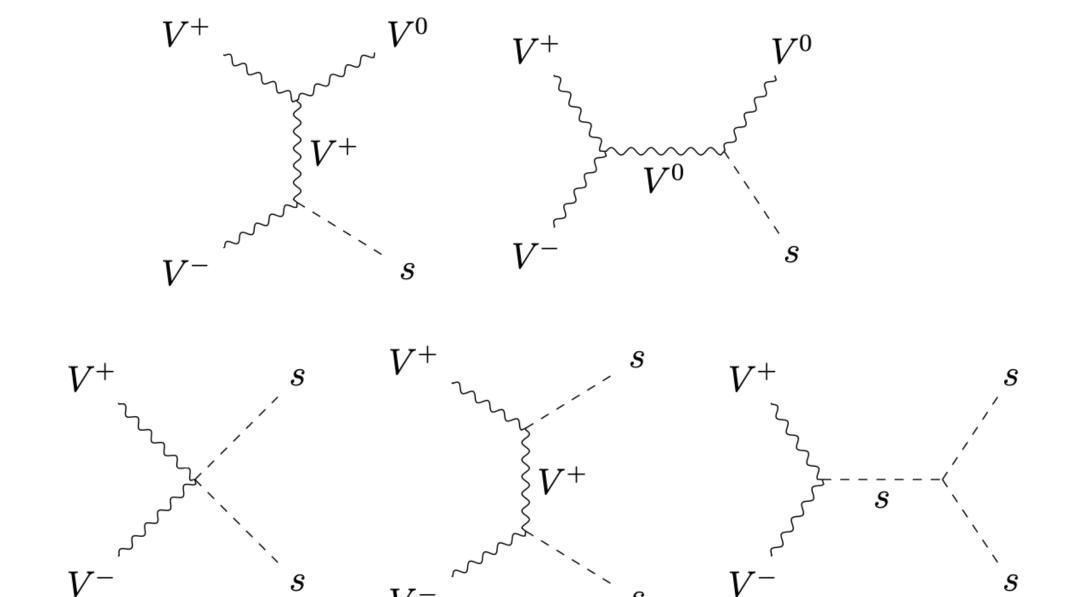
$$\Delta a_{\mu}^{h,E} \simeq \frac{m_{\mu}^2}{24\pi^2 m_{h_i}^2} \bigg[ |v_i^E|^2 + |a_i^E|^2 + \frac{3M_E}{m_{\mu}} (|v_i^E|^2 - |a_i^E|^2) \Big( \ln \Big( \frac{m_{h_i}^2}{M_E^2} \Big) - \frac{3}{2} \Big) \bigg]$$



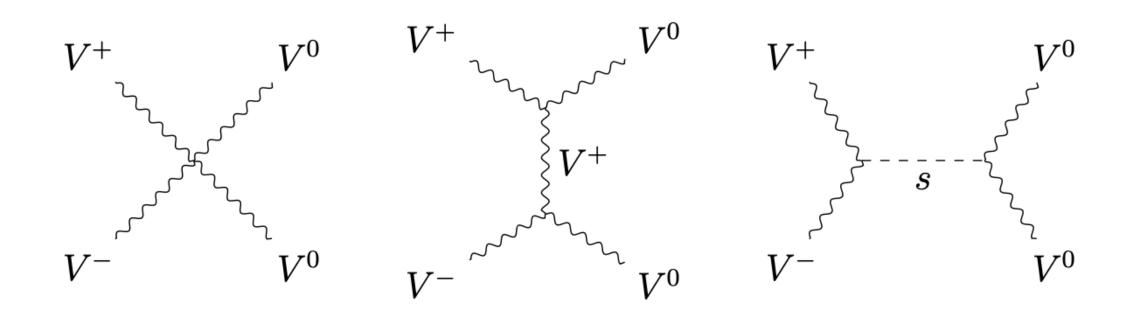
$$\Delta 
ho_H \simeq \left\{ egin{array}{ll} rac{s_W^2 g_D^2}{g_Y^2} rac{M_Z^2}{m_{V^0}^2} \sin^4 eta, & m_{V^0} \gg M_Z, \ & -rac{s_W^2 g_D^2}{g_Y^2} \sin^4 eta, & m_{V^0} \ll M_Z. \end{array} 
ight.$$

#### Dark Matter

$$V^+V^- \rightarrow V^0s, ss$$



$$V^+V^- \to V^0V^0$$



Because of the Z-V mass mixing,  $m_{V^0}$  is slightly larger than  $m_{V^\pm}$ 

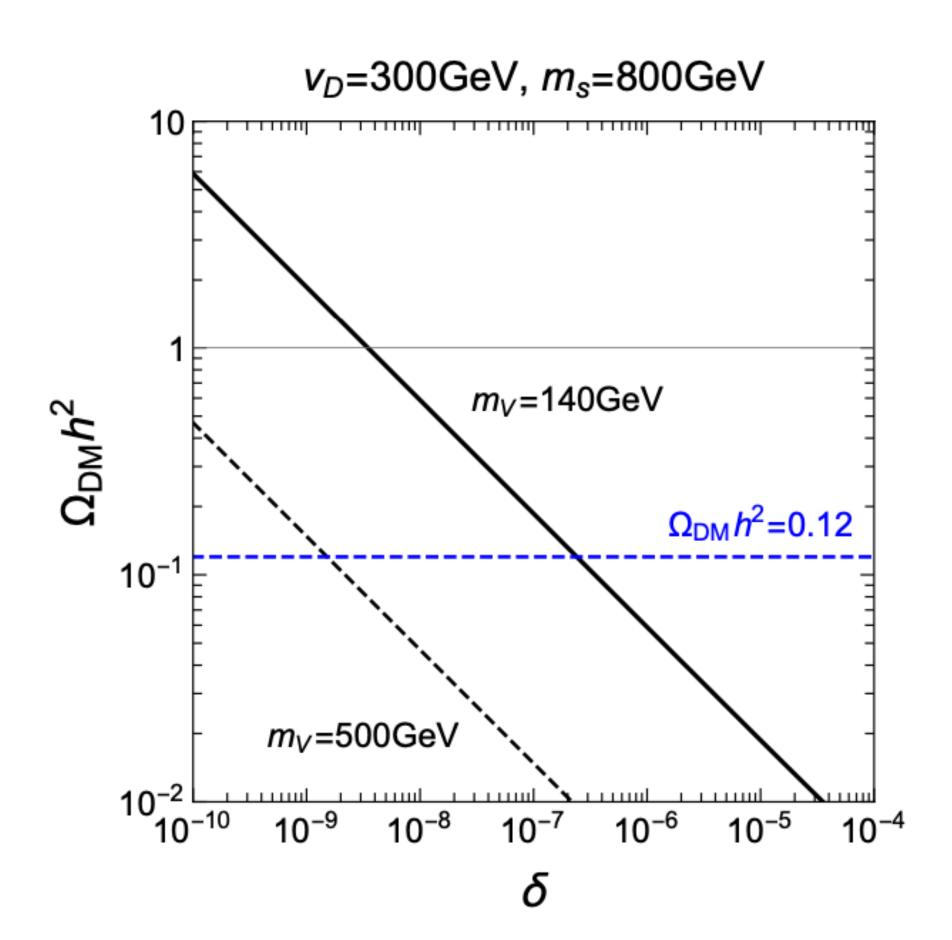
Closed for heavy s (as favored for XENON1T)

The channel is allowed due to a non-zero DM velocity at F.O.

#### Dark Matter

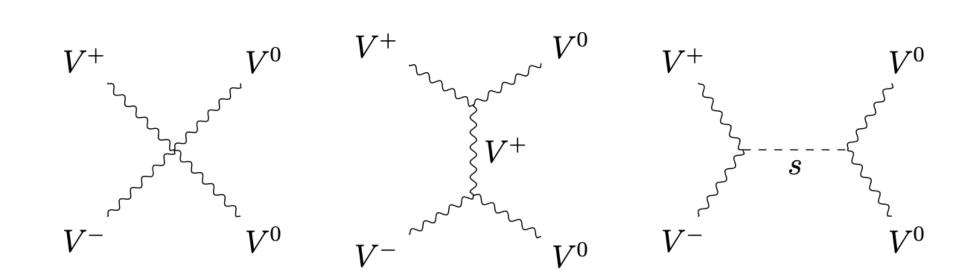
$$\Omega_{\rm DM} h^2 = 0.2745 \left(\frac{Y_{\rm DM}}{10^{-11}}\right) \left(\frac{m_{V^+}}{100 \,\text{GeV}}\right)$$

$$\Omega_{\rm DM} h^2 = 0.2745 \left(\frac{Y_{\rm DM}}{10^{-11}}\right) \left(\frac{m_{V^+}}{100\,{\rm GeV}}\right) \qquad \delta \simeq 2.2 \times 10^{-5} \left(\frac{\Delta \rho_H}{1.3 \times 10^{-3}}\right) \left(\frac{500\,{\rm GeV}}{m_{V^0}}\right)^2$$



- The relic abundance condition is insensitive to  $m_s$  and mixing angles
- Crucially dependent on the mass splitting  $\delta \equiv m_{V^0}/m_{V^+} - 1$
- For a fixed  $v_D$ , a larger  $SU(2)_D$  dark coupling (larger mass) leads to a larger annihilation crosssection so the relic density decreases.

#### Forbidden channel



Closed for 
$$v_{rel} \lesssim \sqrt{8\delta} \approx 220 \, \mathrm{km/s}$$

Does not lead to observable signatures for  $\delta \gtrsim 6 \times 10^{-7}$ 

$$V^{+}V^{-} \to hh$$

$$V^{+}V^{-} \to V^{0}Z$$

$$V^{+}V^{-} \to SMSM$$

#### **Subdominant channels**

- Suppressed by small mixing angles
- They may lead to signals in CMB or cosmic rays.

#### Direct detection

$$V^{\pm}q \rightarrow V^{\pm}q$$

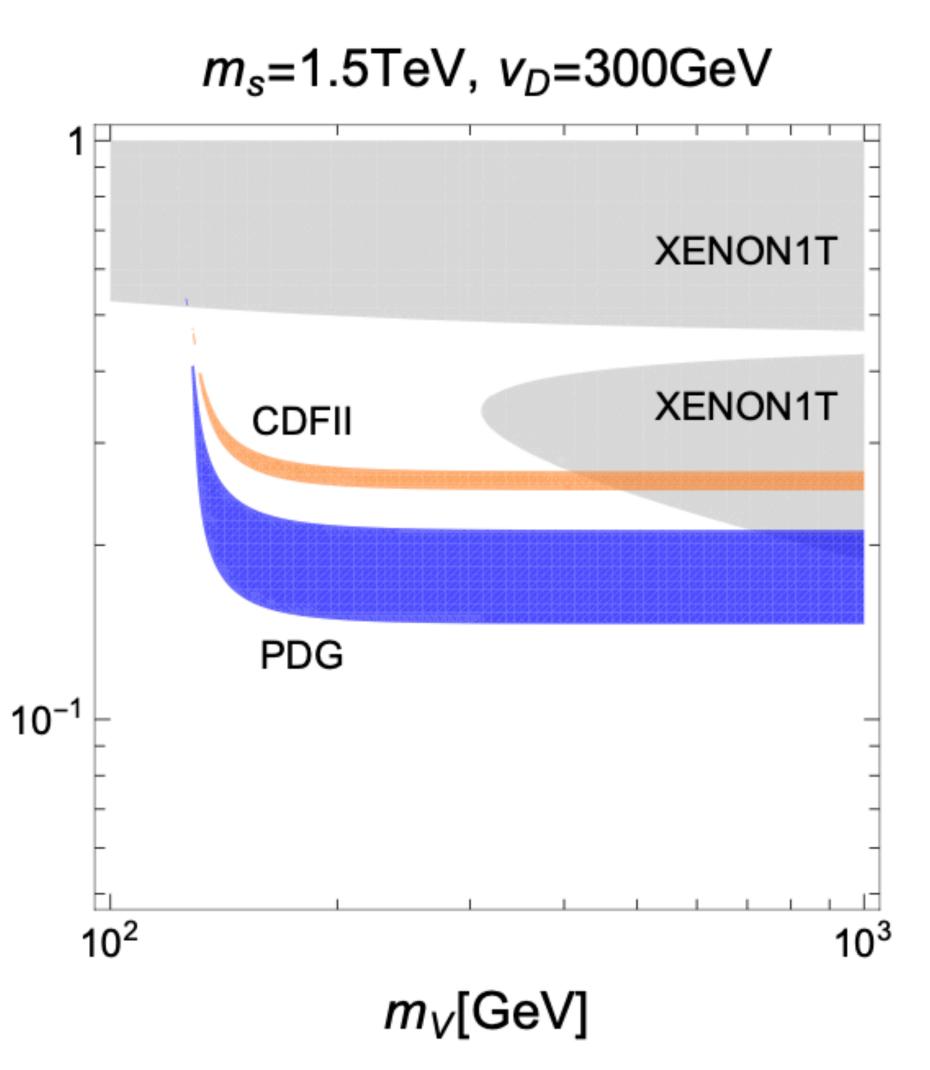
- Possible through SM Higgs and singlet scalar exchanges.
- It is subdominant but can be constrained by the direct detection.
- Spin-independent elastic scattering:

$$\mathcal{L}_{V^{\pm}-q} = \lambda_{\text{eff}} m_q V_{\mu}^{+} V^{-\mu} \bar{q} q$$

$$\lambda_{\text{eff}} = \frac{\sqrt{2}}{2v} v_D g_D^2 \sin \theta_h \cos \theta_h \left( \frac{1}{m_s^2} - \frac{1}{m_h^2} \right) - \frac{1}{2} g_D^2 \sin^2 \beta \left( \frac{\sin^2 \theta_h}{m_s^2} + \frac{\cos^2 \theta_h}{m_h^2} \right)$$

Alignment limit 
$$\sin \theta_h = -\frac{v}{\sqrt{2}v_D} \sin^2 \beta$$
,

$$m_s \gg m_h$$



### Summary

- We extended the SM with an extra  $SU(2)_D$  gauge symmetry.
- The vector-like leptons and  $SU(2)_D$  gauge bosons contribute to the muon g 2.
- The mass mixing between the Z boson and the dark  $V^0$  contributes to the W boson mass.
- A combination of the  $U(1)_G$  in the Higgs sector and the dark isospin leads to a Z2 parity allowing for stable candidates for DM.
- The forbidden annihilation channel explains the correct relic density.
- Direct detection bounds can be satisfied in the alignment limit of the mixing between the SM Higgs and the singlet scalar of  $SU(2)_D$ .