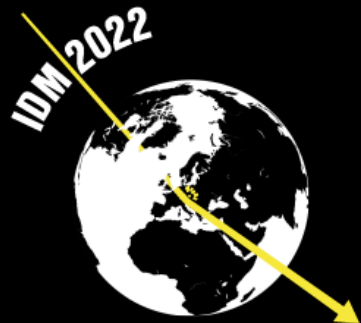
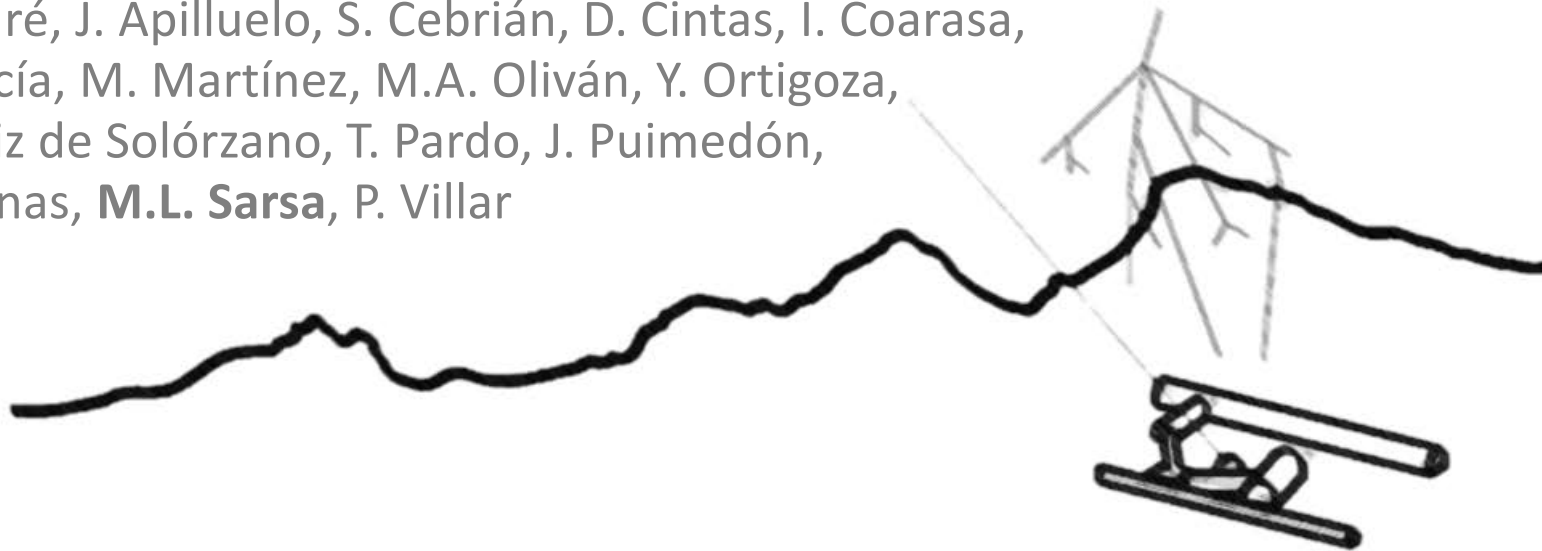


ANAIS-112 annual modulation results and prospects to test DAMA/LIBRA beyond 3σ

J. Amaré, J. Apilluelo, S. Cebrián, D. Cintas, I. Coarasa, E. García, M. Martínez, M.A. Oliván, Y. Ortigoza, A. Ortiz de Solórzano, T. Pardo, J. Puimedón, A. Salinas, M.L. Sarsa, P. Villar

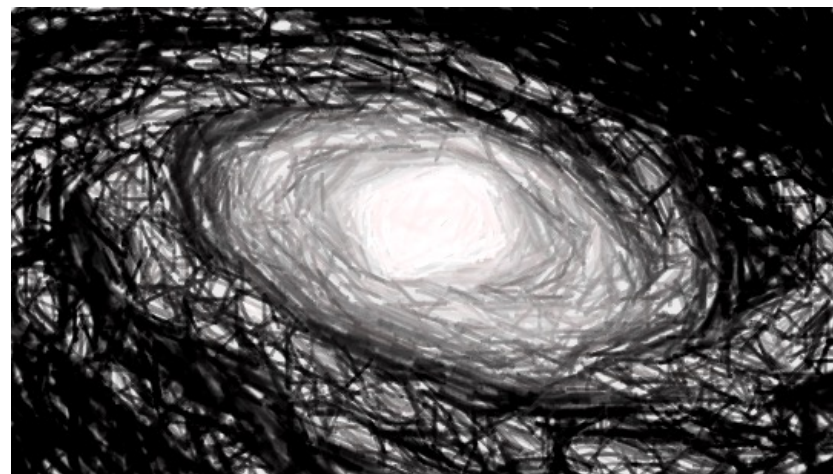
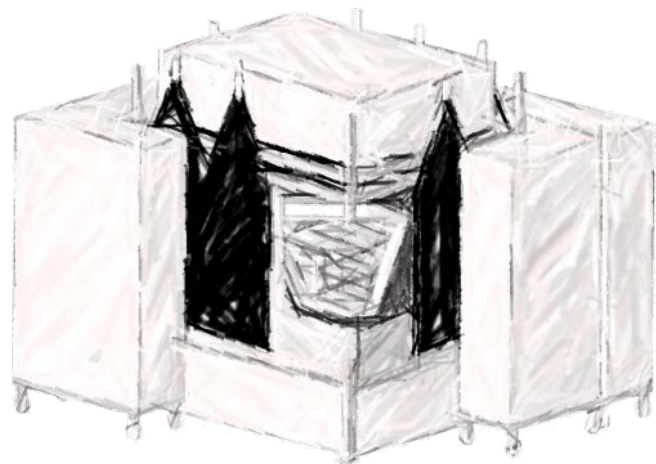


14th International Conference on Identification of Dark Matter

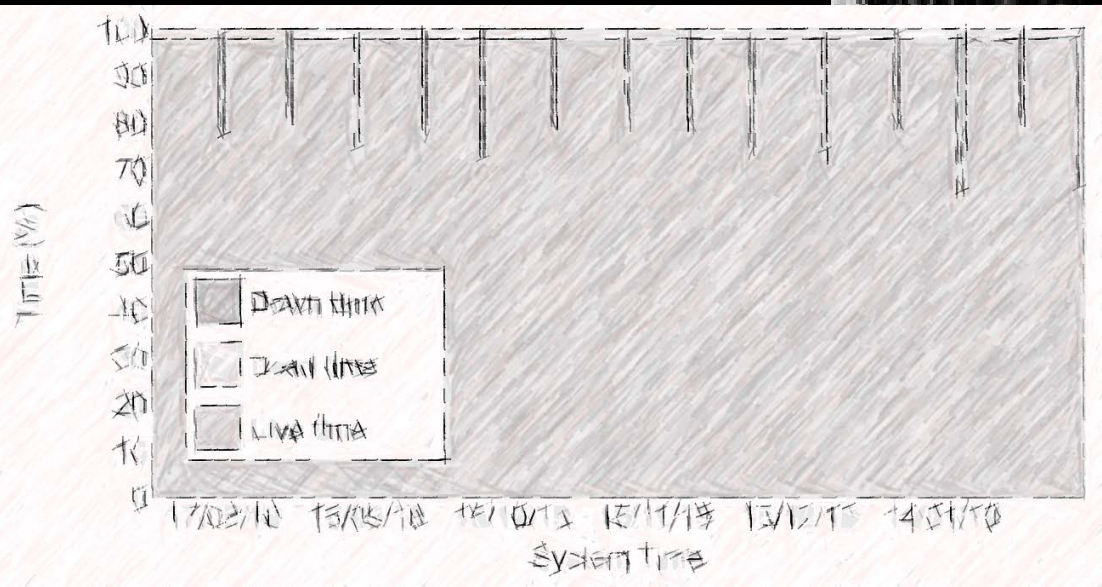
18-22 July 2022
Vienna, Austria



Centro de Astropartículas y Física de Altas Energías
Universidad Zaragoza



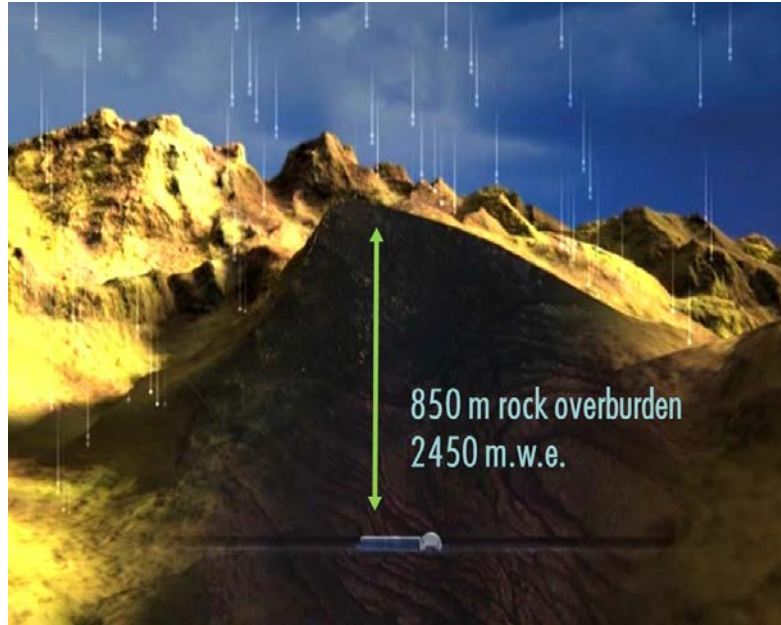
- **ANAIS-112 experimental set-up and performance**
- **Neutron calibration program of ANAIS-112 detectors**
- **New ANAIS-112 event selection protocol using Machine Learning techniques**
- **ANAIS-112 three years annual modulation re-analysis and update of ANAIS-112 sensitivity prospects**
- **Summary and Outlook**



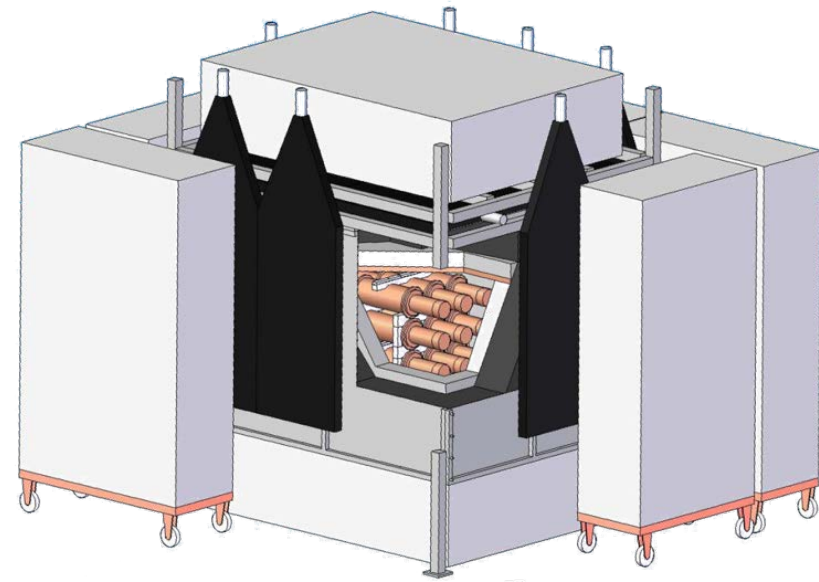
ANAIS-112 experimental set-up and performance



Annual modulation with NaI Scintillators

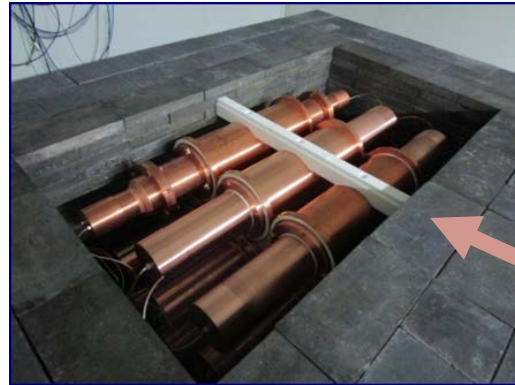


- Confirmation of DAMA-LIBRA modulation signal -> **same target and technique / different** experimental approach / **different** environmental conditions affecting **systematics**
- **At Canfranc Underground Laboratory, SPAIN** taking data since **August 2017**
- 3x3 matrix of 12.5 kg cylindrical modules = **112.5 kg of active mass** grown @ Alpha Spectra, Inc.
- HE PMTs coupled at LSC clean room





Annual modulation with NaI Scintillators



RELEVANT EXPERIMENTAL FEATURES

Mylar windows built-in allowing for calibration at LE with ^{109}Cd sources

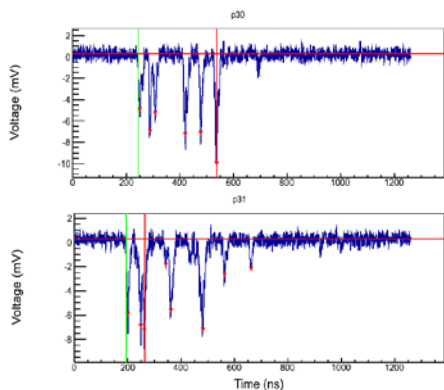
Excellent light collection in all the nine modules, at the level of 15 p.e./keV

Active muon veto system (16 plastic scintillators)

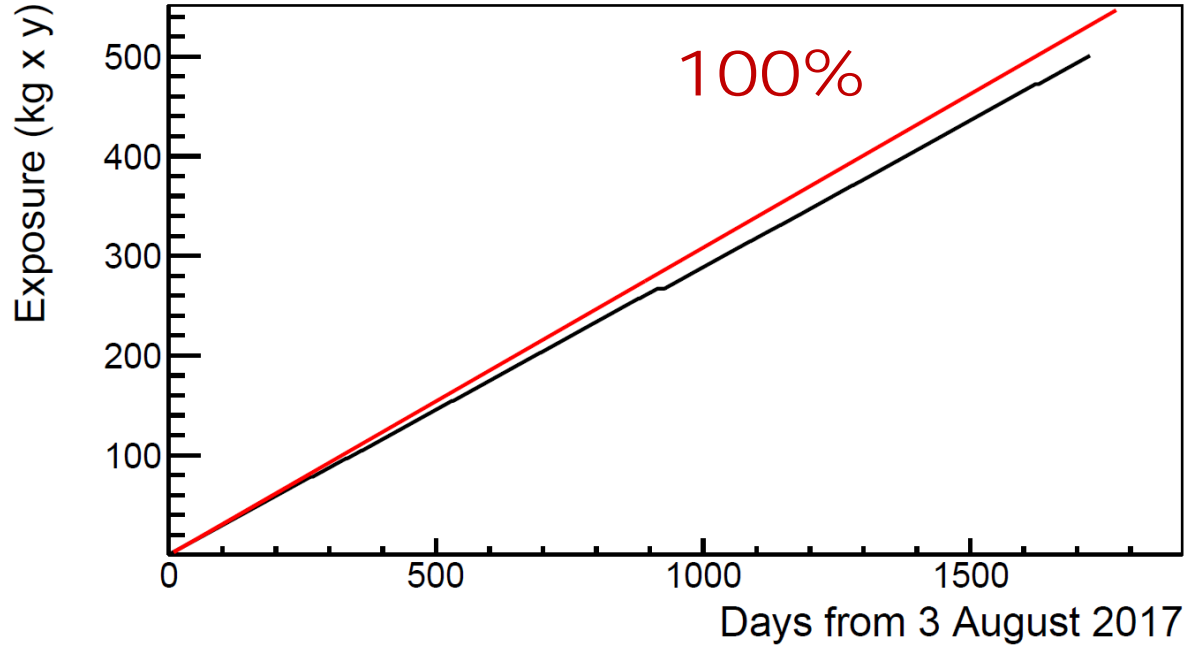
Individual PMT signals digitized and fully processed

(14 bits / 2 GS/s), trigger at p.e. level for each PMT + Logical AND coincidence in 200ns window

Robust / low noise / tested with previous prototypes



The ANAIS-112 Dark Matter Run started on August 3, 2017



524.44 kg x y @ June 7, 2022

ANAIS-112 accumulated exposure

Five-year exposure will be completed by August 2022 with about 95% of live time

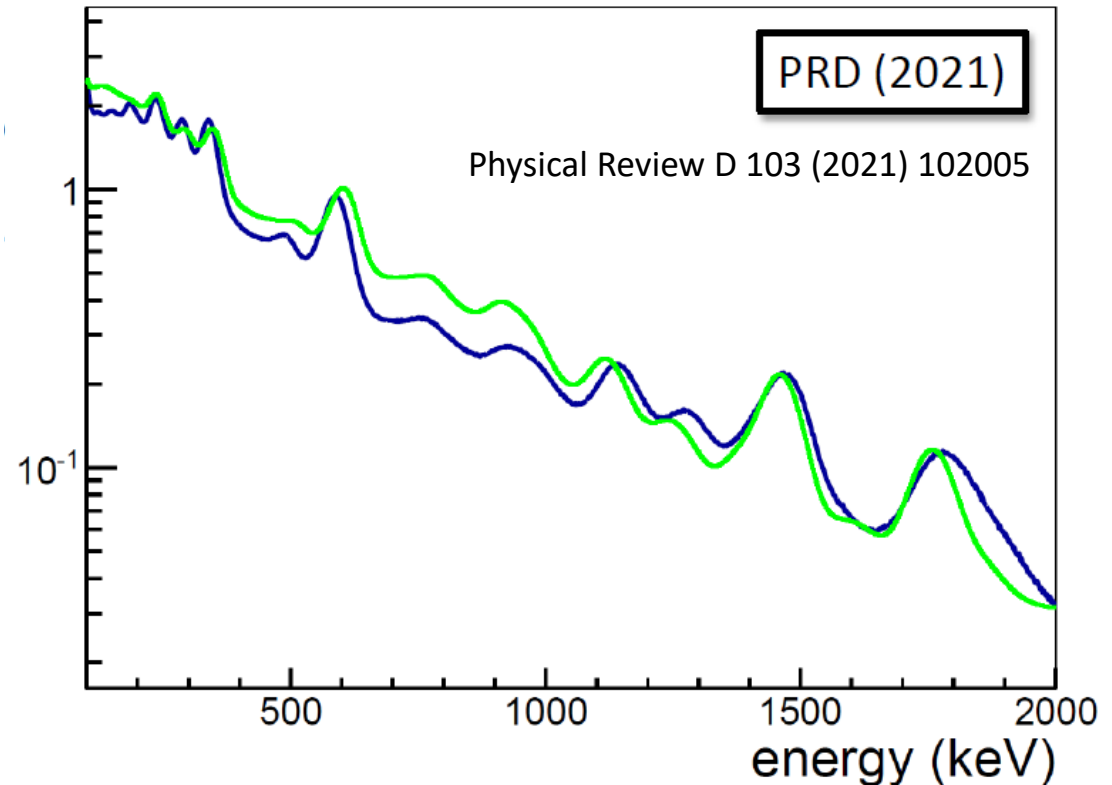
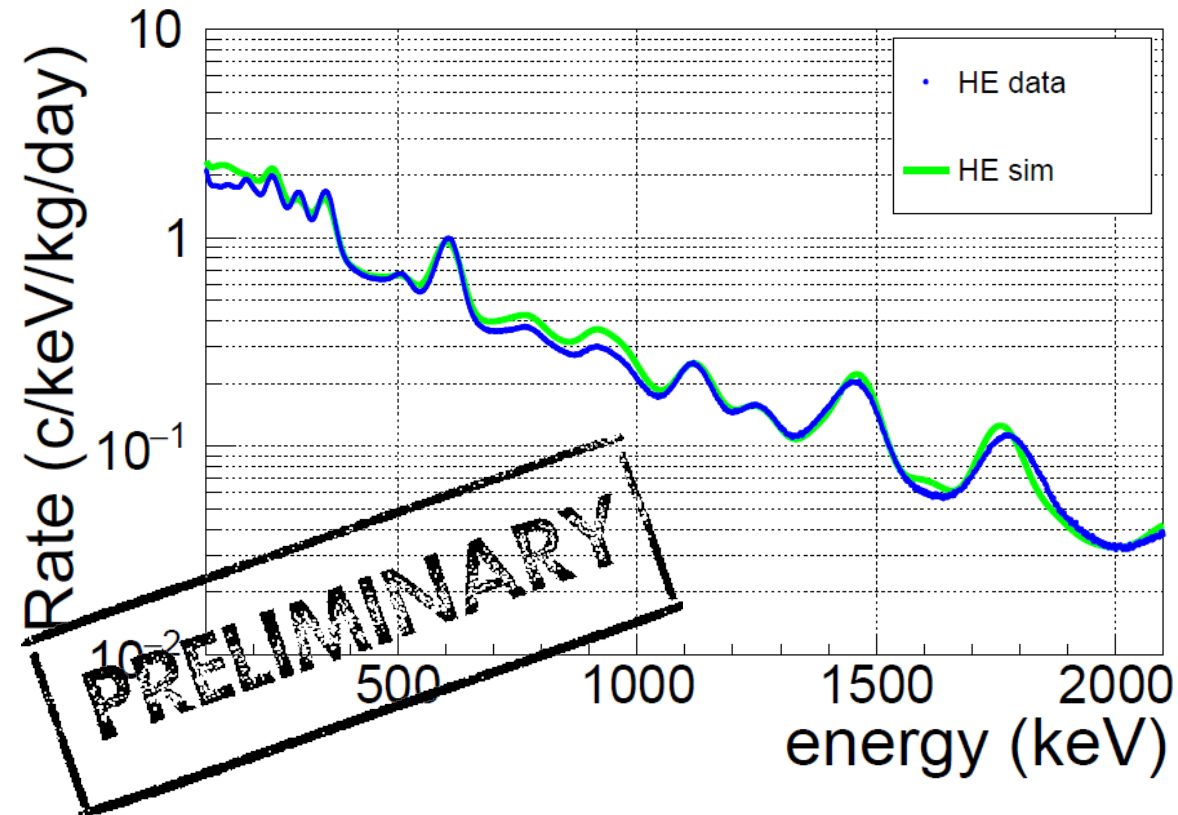


Time period	Accumulated Live Time	Live Time (%)	Down Time (%)	Dead Time (%)
08/03/2017-07/31/2018	341.722 days	94.5	2.6	2.9
08/01/2018-08/28/2019	374.302 days	95.6	2.4	2.1
08/29/2019-08/13/2020	333.791 days	95.2	2.5	2.3
08/13/2020-08/18/2021	354.667 days	95.9	1.7	2.4
08/19/2021-06/07/2022	297.022 days	95.0	3.2	1.9
TOTAL LIVE	1701.504 days			
TOTAL	1766.25 days			

Calibration @ HE (above 100 keV)

Improvement of the estimate of energy above 100 keV by better linearizing the pulse area – QDC relation for each module (using Chebyshev' polynomials +pol1 +pol2 instead of a logistic function) and calibrating quadratically with background lines

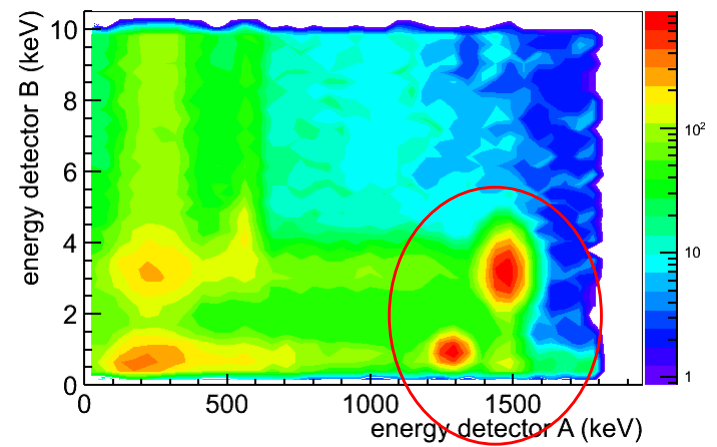
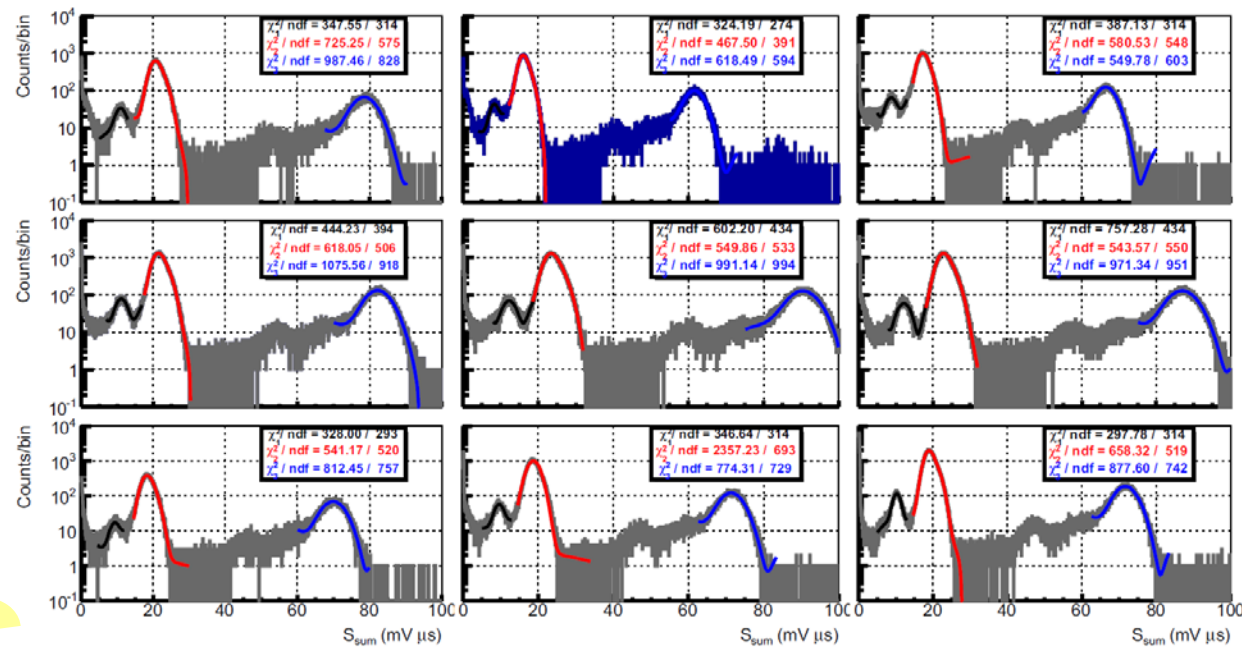
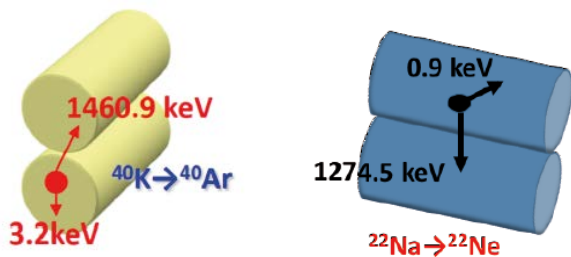
IMPROVEMENT IN MATCHING THE MC SIMULATED BACKGROUND



Calibrating the ROI with high accuracy

Two calibration steps:

- periodical external calibration using ^{109}Cd (88.0, 22.6 and 11.9 keV) to correct gain drifts every two weeks
- ^{40}K and ^{22}Na internal contamination background lines (3.2 and 0.9 keV) identified by coincidences with HE gammas every 1.5 months



3.2 and 0.87 keV

ROI calibrated with 22.6, 11.9, 3.2 and 0.9 keV

Calibrating the ROI with high accuracy

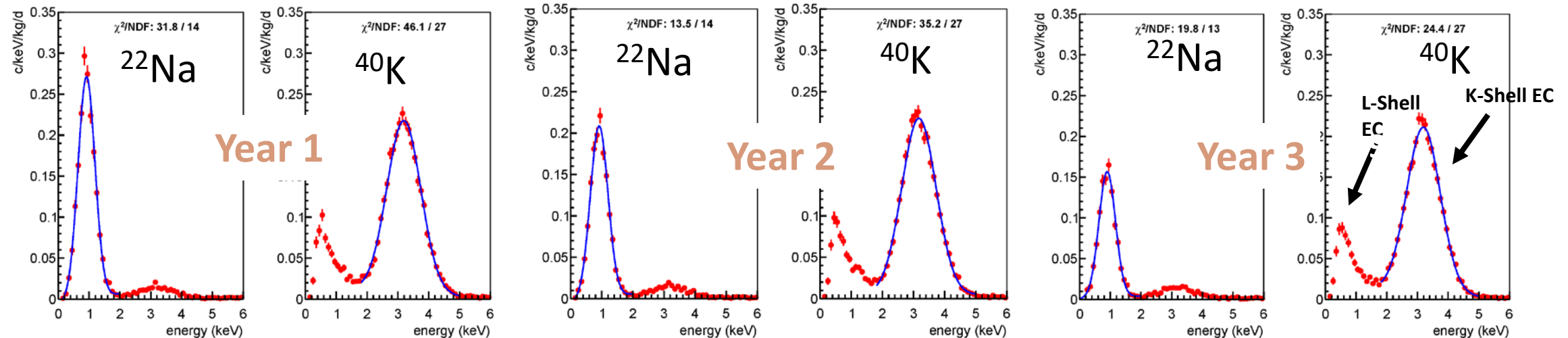
Two calibration steps:

- periodical external calibration using ^{109}Cd (88.0, 22.6 and 11.9 keV) to correct gain drifts every two weeks
- ^{40}K and ^{22}Na internal contamination background lines (3.2 and 0.9 keV) identified by coincidences with HE gammas every 1.5 months

ROBUST CALIBRATION AT LOW ENERGY

checked by ^{40}K and ^{22}Na peaks mean values and non-degradation of energy resolution in the analysed period

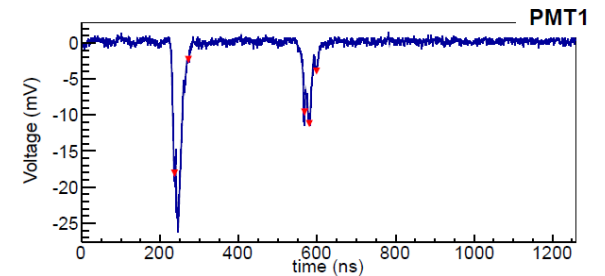
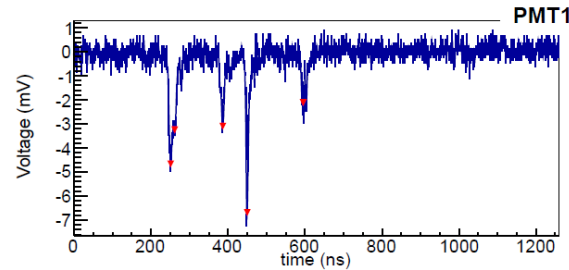
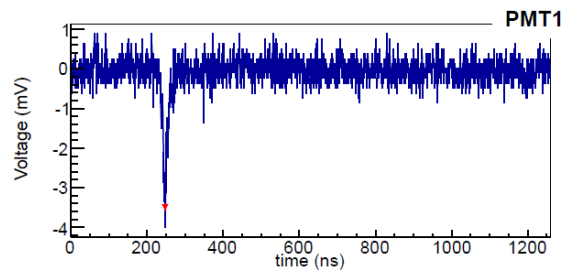
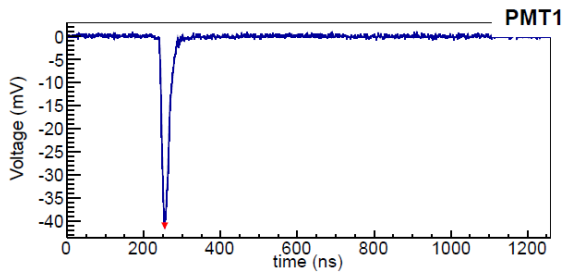
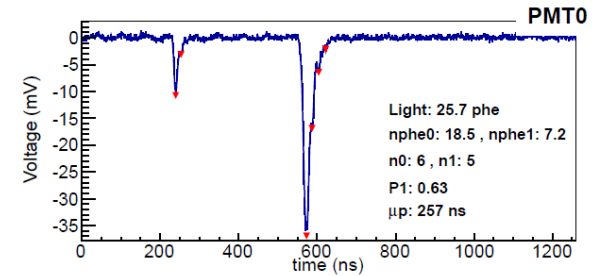
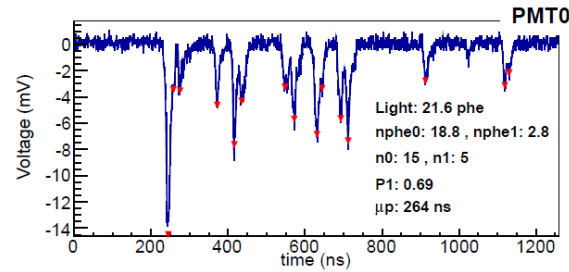
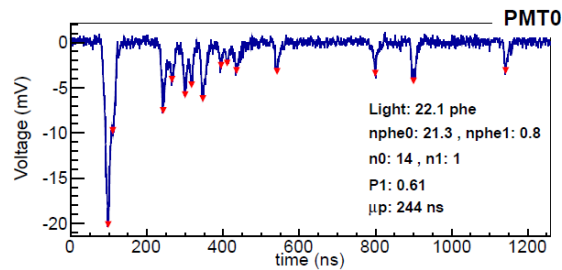
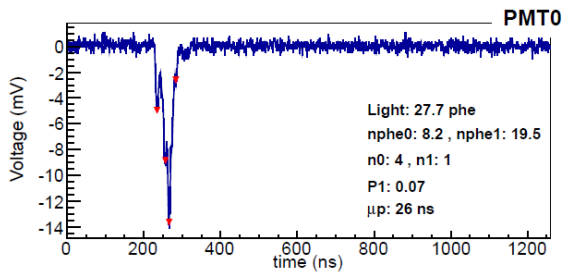
Clear demonstration of triggering below 1 keV



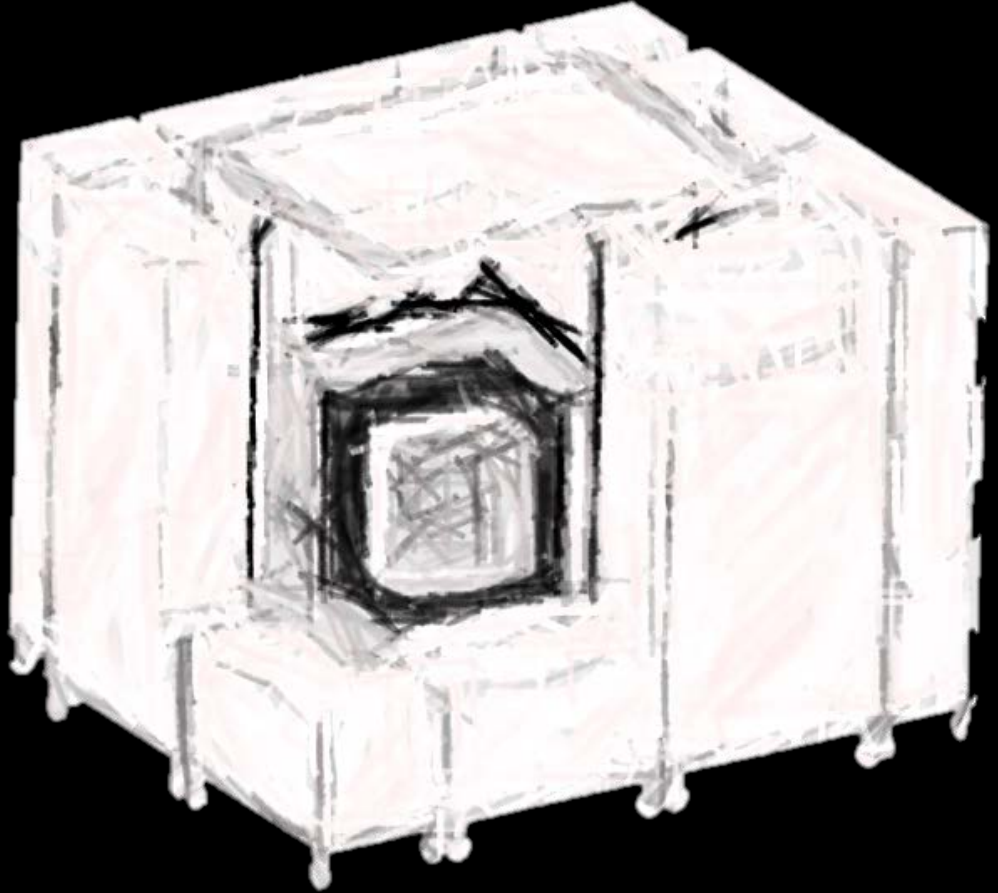
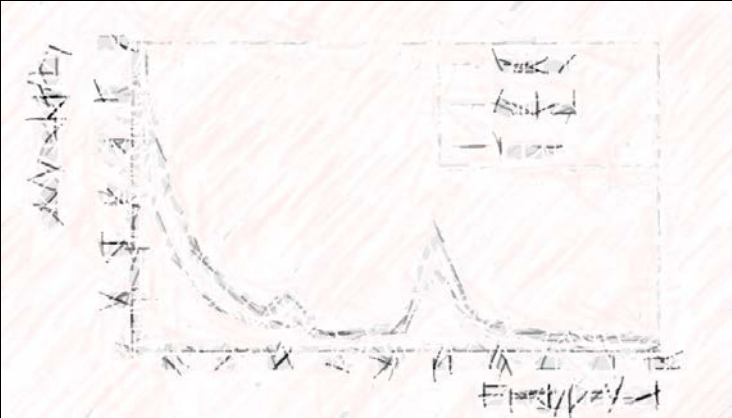
In 2018 a BLANK module was commissioned for taking data in similar conditions to ANAIS-112 modules similar size, housing and PMTs, without NaI(Tl) crystal same DAQ, independent shielding close to ANAIS-112



ANAIS-112 BLANK module

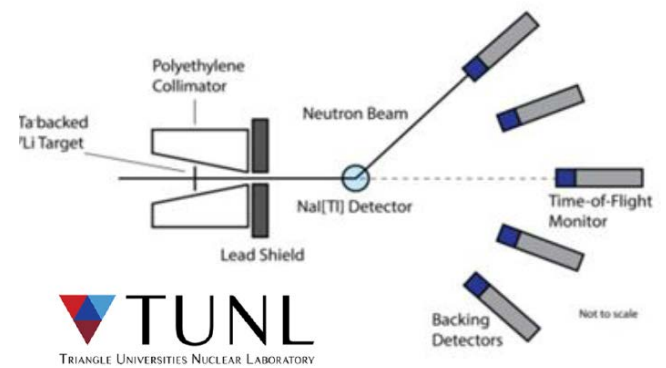


ANAIIS-112 neutron calibration program



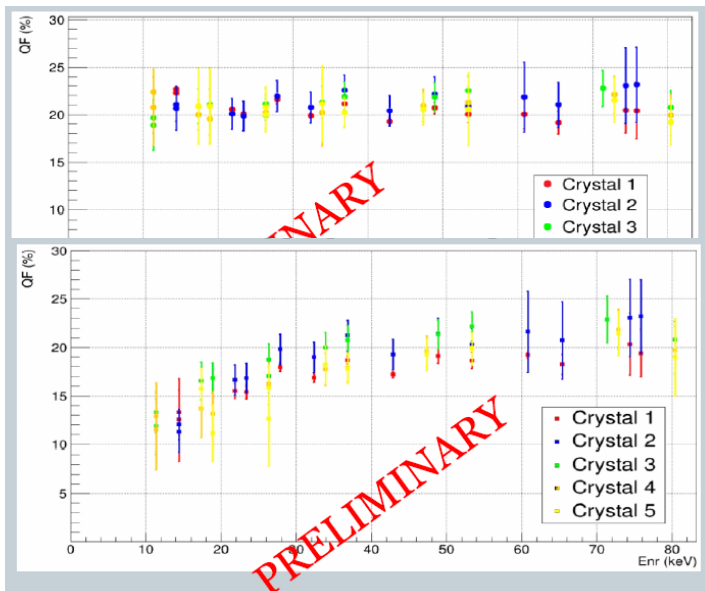
UNCERTAINTIES IN SCINTILLATION QUENCHING FACTORS ARE A RELEVANT SYSTEMATIC EFFECT for the comparison between DAMA/LIBRA and ANAIS

Na - QF derived using different methods and different calibration procedures do not agree. Dependences on the crystal properties can not be discarded



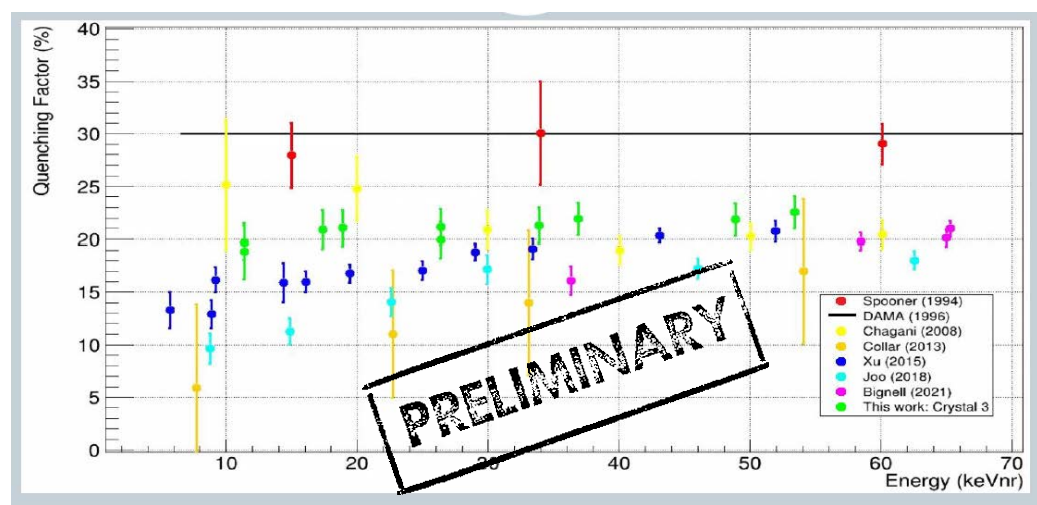
Five small AS crystals
 Collaboration between Yale, Duke and Zaragoza
Measurements in 2018 @TUNL

Our best (preliminary) estimate for Na-QF in NaI(Tl)-AS crystals is 20%



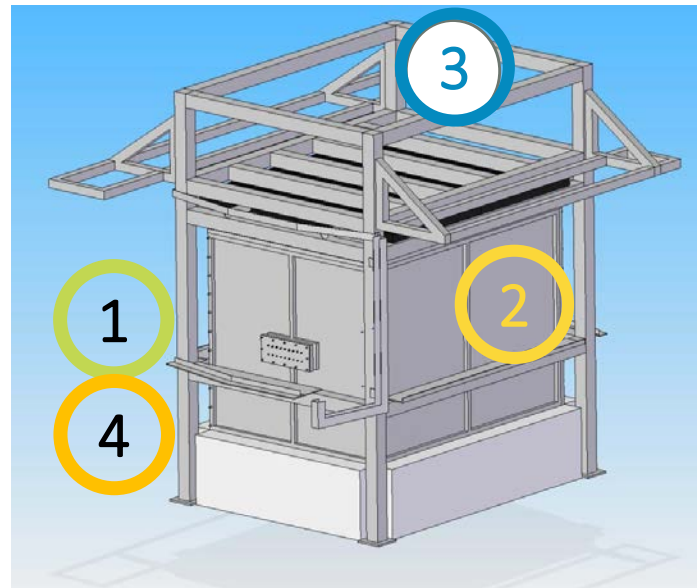
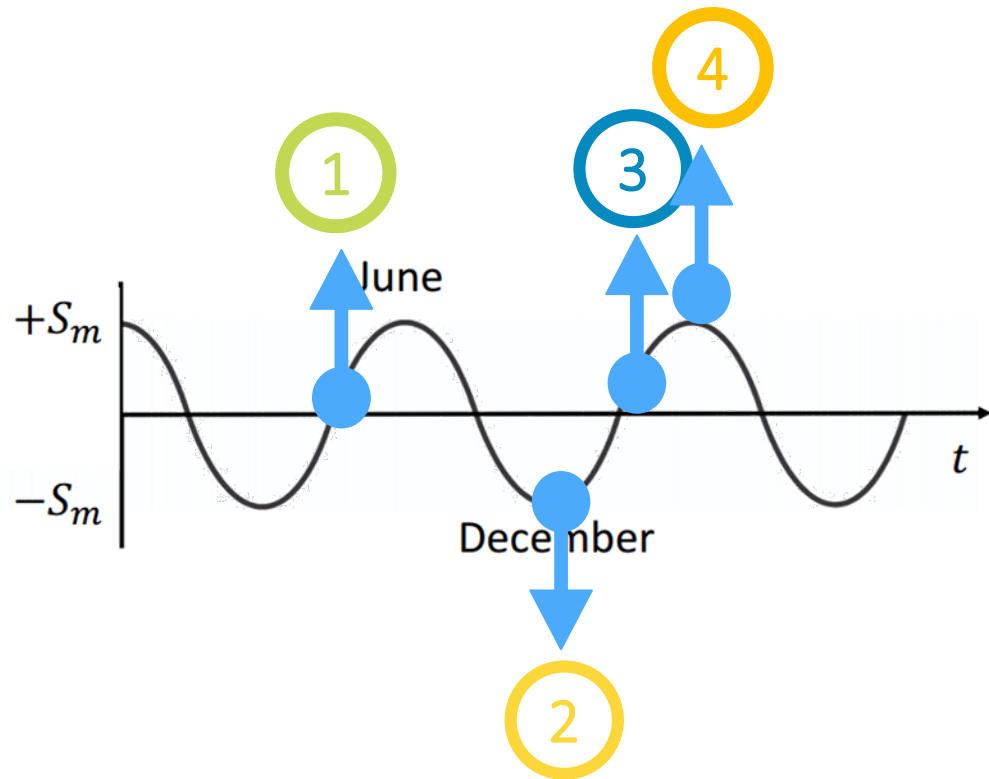
Linearizing response in energy using ¹³³Ba data

Assuming proportional response using 57.6 keV energy deposition



“onsite” neutron calibrations with the full ANAIS-112 set-up

Four calibration runs since April 2021 using ^{252}Cf neutron source at different positions in the ANAIS-112 set-up

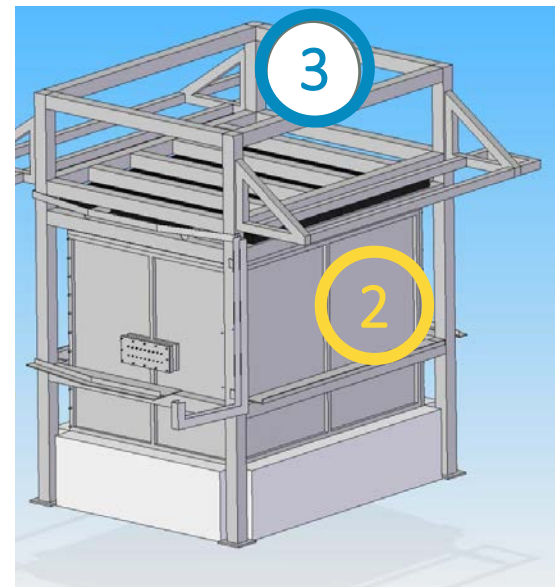
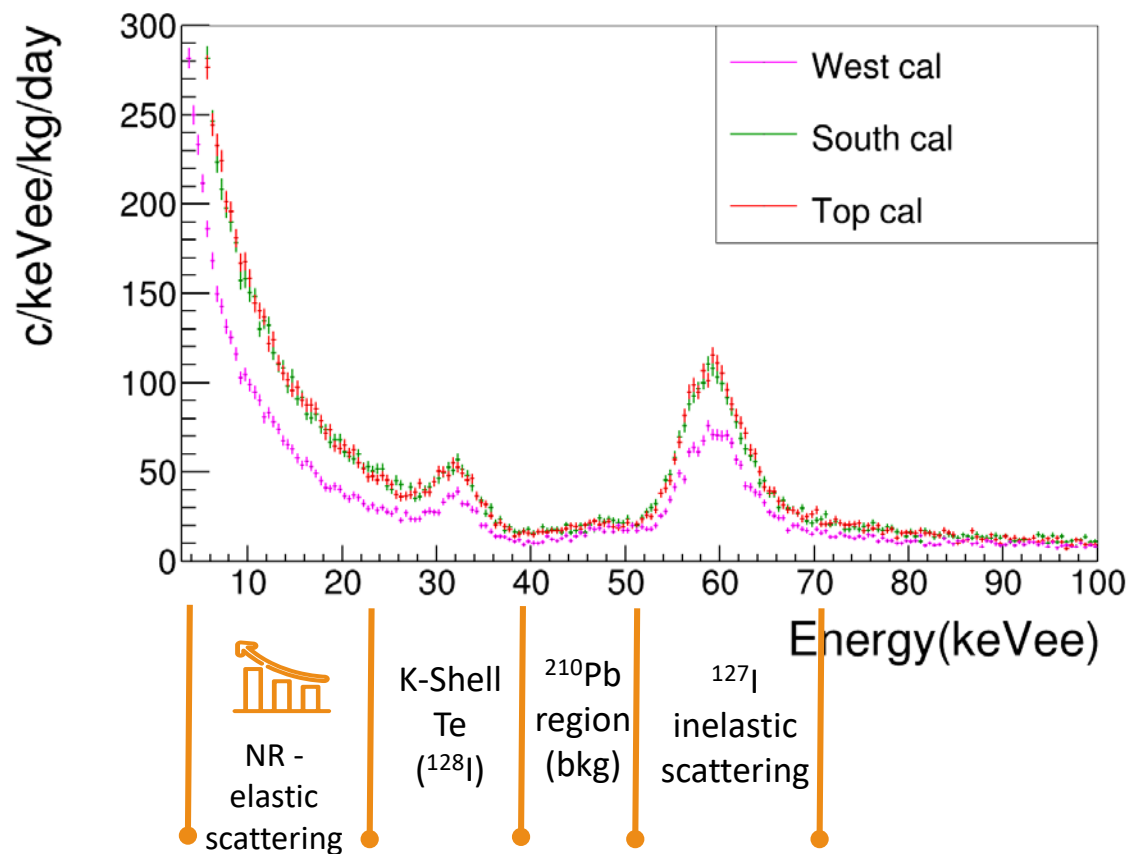


Neutrons produce many “bulk scintillation events” at very low energy

- We have a large population of events for training our new ML analysis and checking efficiency stability
- We expect to derive some information on QF by comparing simulations and measurements

“onsite” neutron calibrations with the full ANAIS-112 set-up

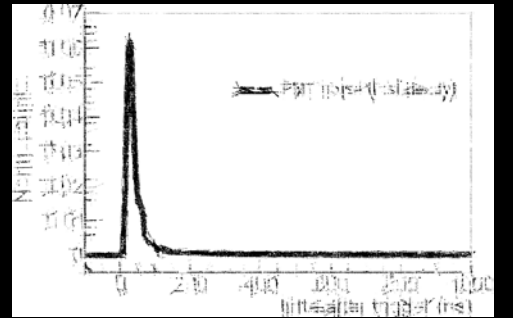
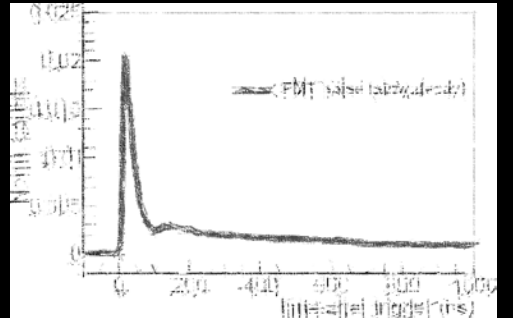
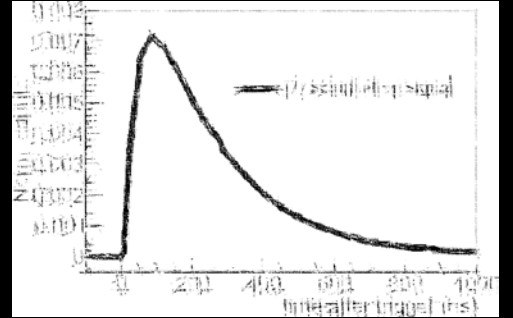
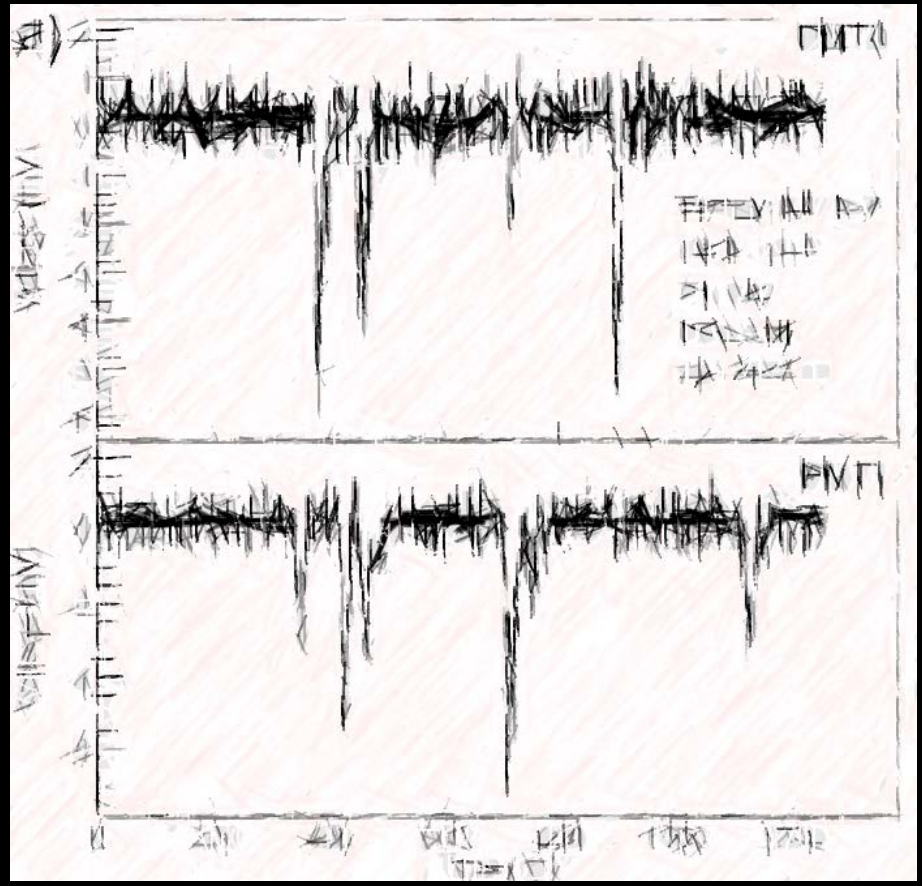
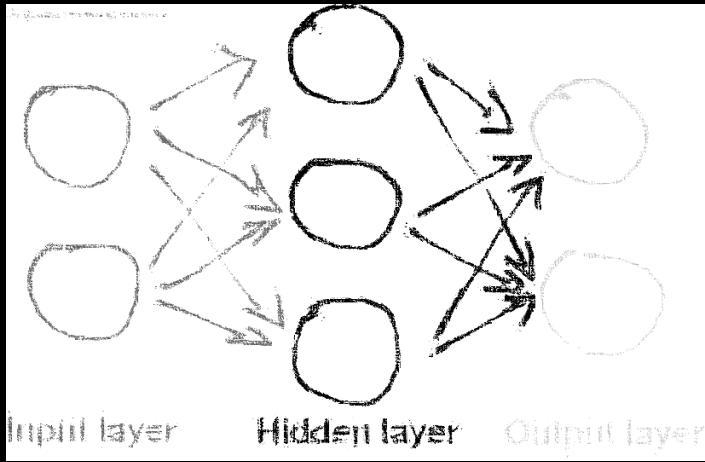
Four calibration runs since April 2021 using ^{252}Cf neutron source at different positions in the ANAIS-112 set-up



Neutrons produce many “bulk scintillation events” at very low energy

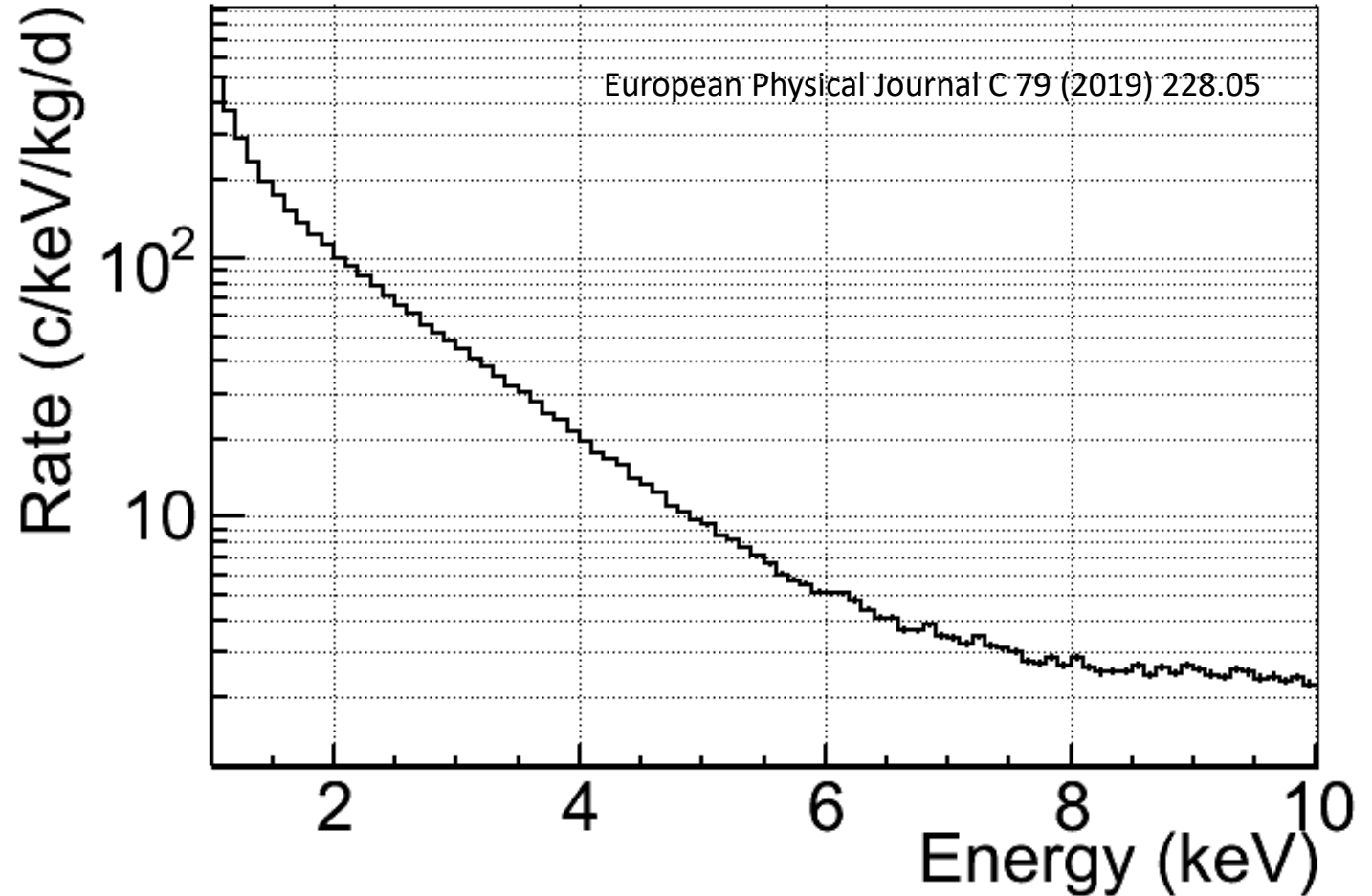
- We have a large population of events for training our new ML analysis and checking efficiency stability
- We expect to derive some information on QF by comparing simulations and measurements

New ANAIS-112 event selection protocol using Machine Learning techniques



ANAIS-112 event selection for the three-year analysis (PRD 2021)

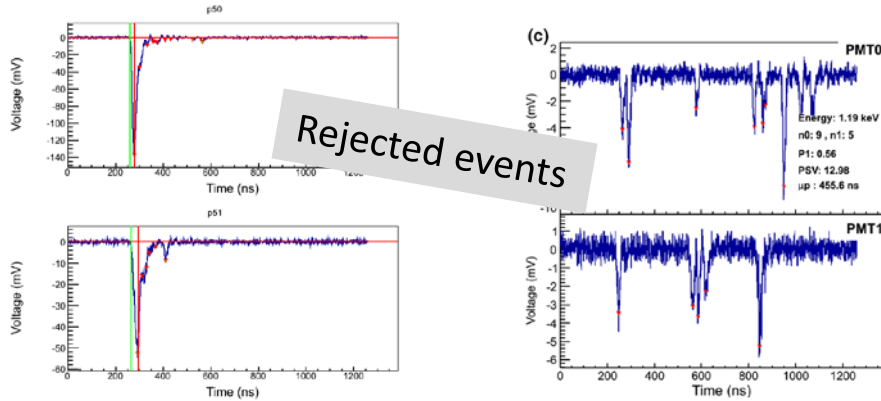
Raw data below 10 keV are dominated by non-bulk scintillation events



10 % data unblinded background @ ROI (random distribution)

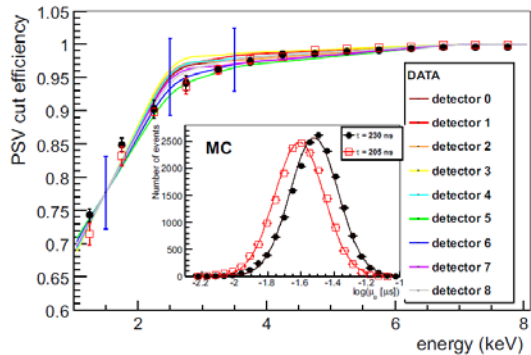
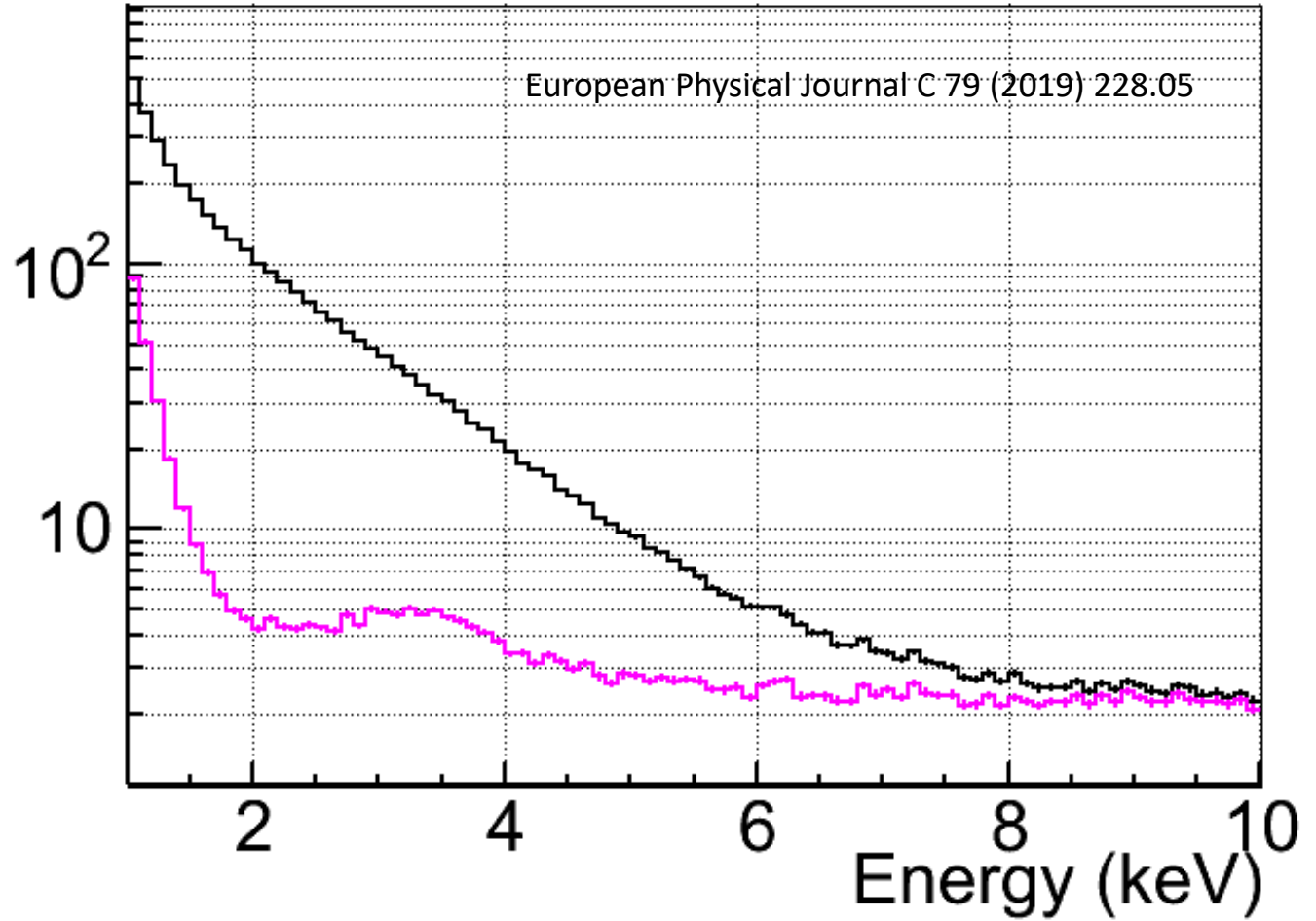
PULSE SHAPE cut to properly select events with pulse shapes from NaI(Tl) scintillation

$$P_1 = \frac{\int_{100\text{ ns}}^{600\text{ ns}} A(t)dt}{\int_0^{600\text{ ns}} A(t)dt} \quad \mu_p = \frac{\sum A_p t_p}{\sum A_p}$$

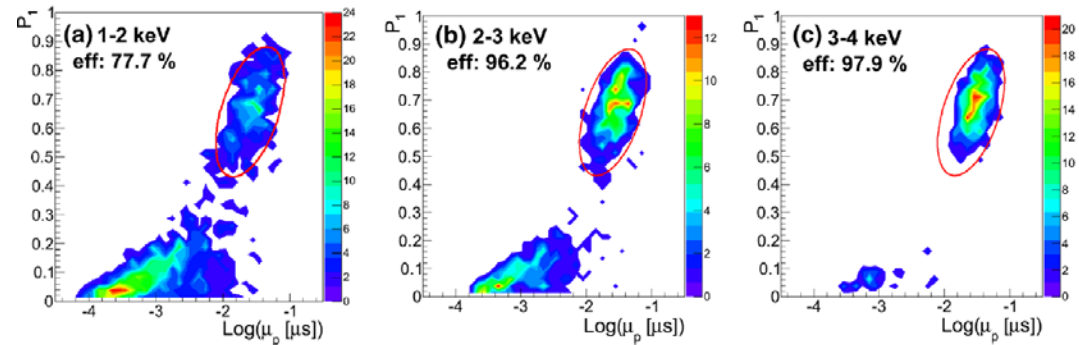


Rejected events

Rate (c/keV/kg/d)

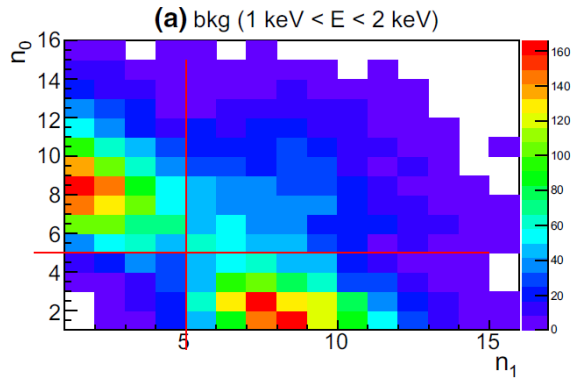


← Efficiency determined from ²²Na and ⁴⁰K events

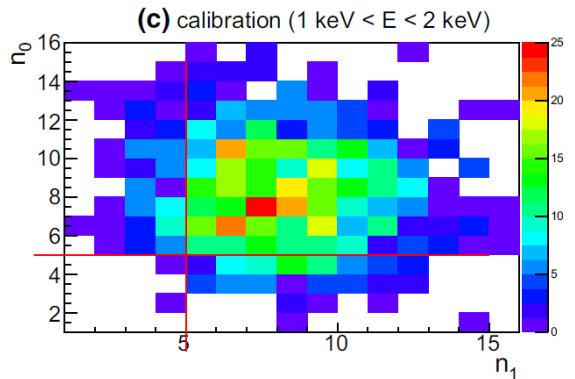


Asymmetric events rejection (E<2keV):

Requirement on number of peaks >4 at each PMT

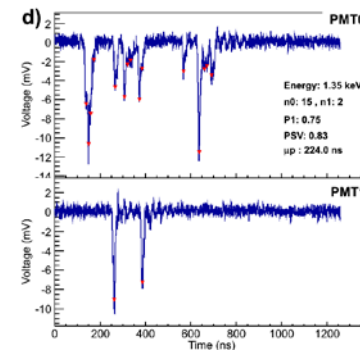
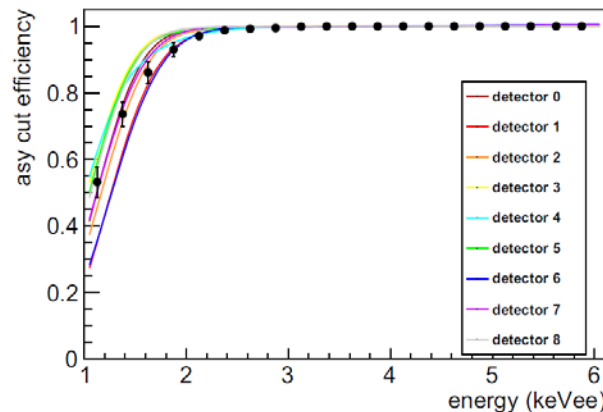
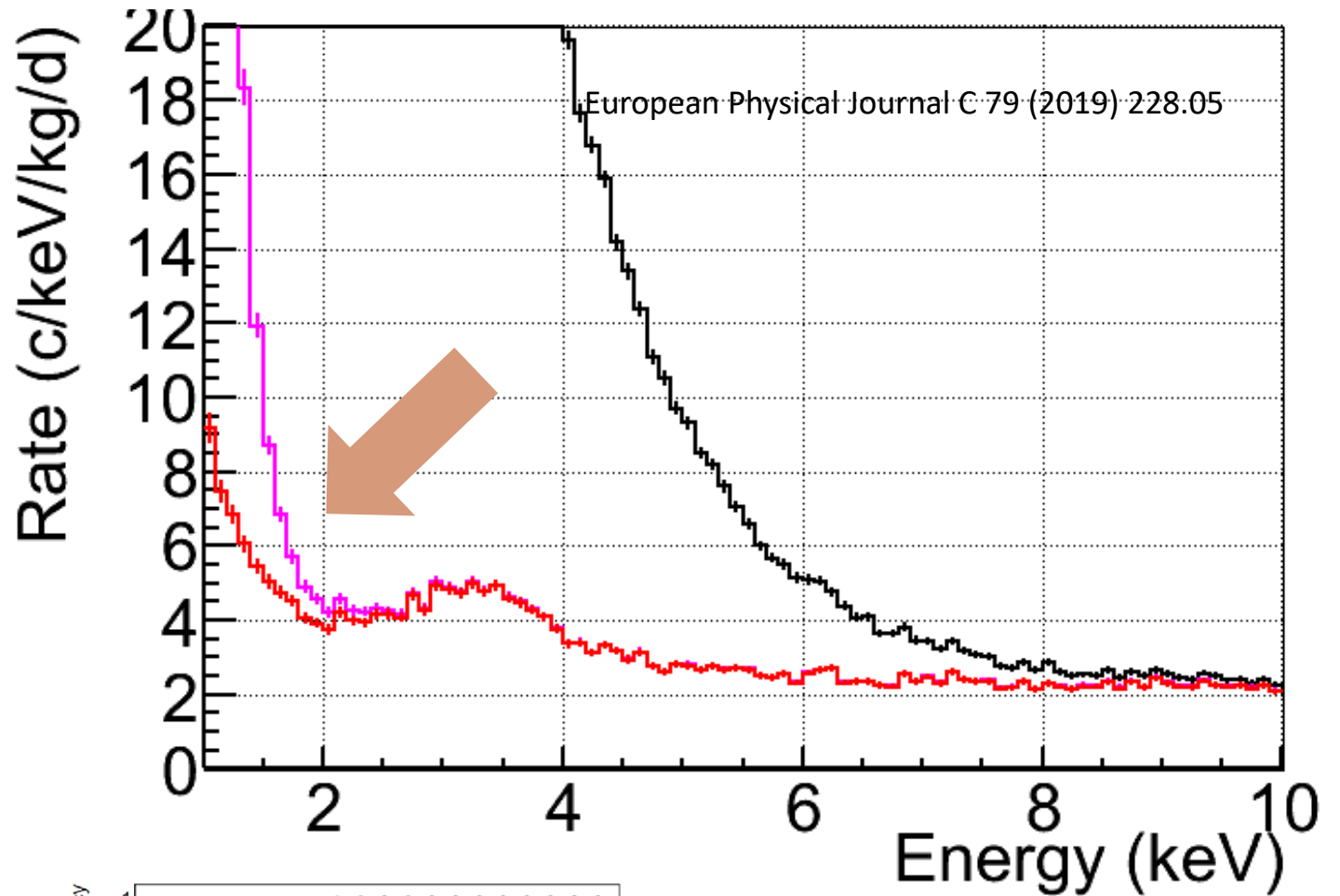


BKG



CAL

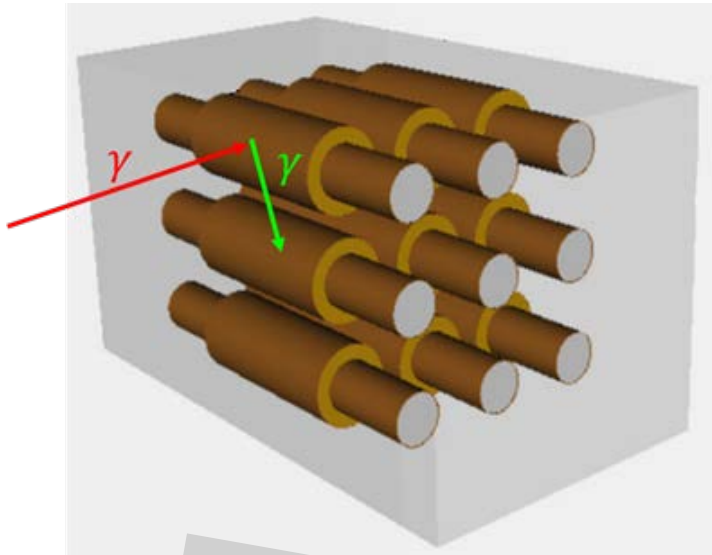
Efficiency determined from Cd calibration evts



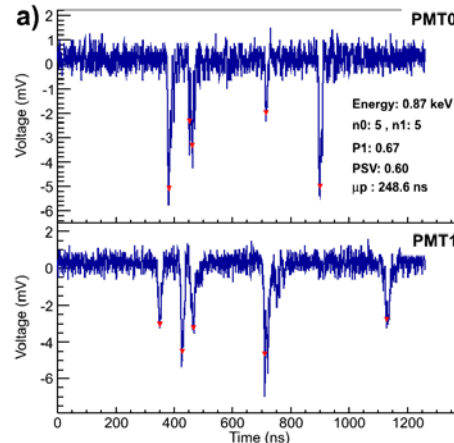
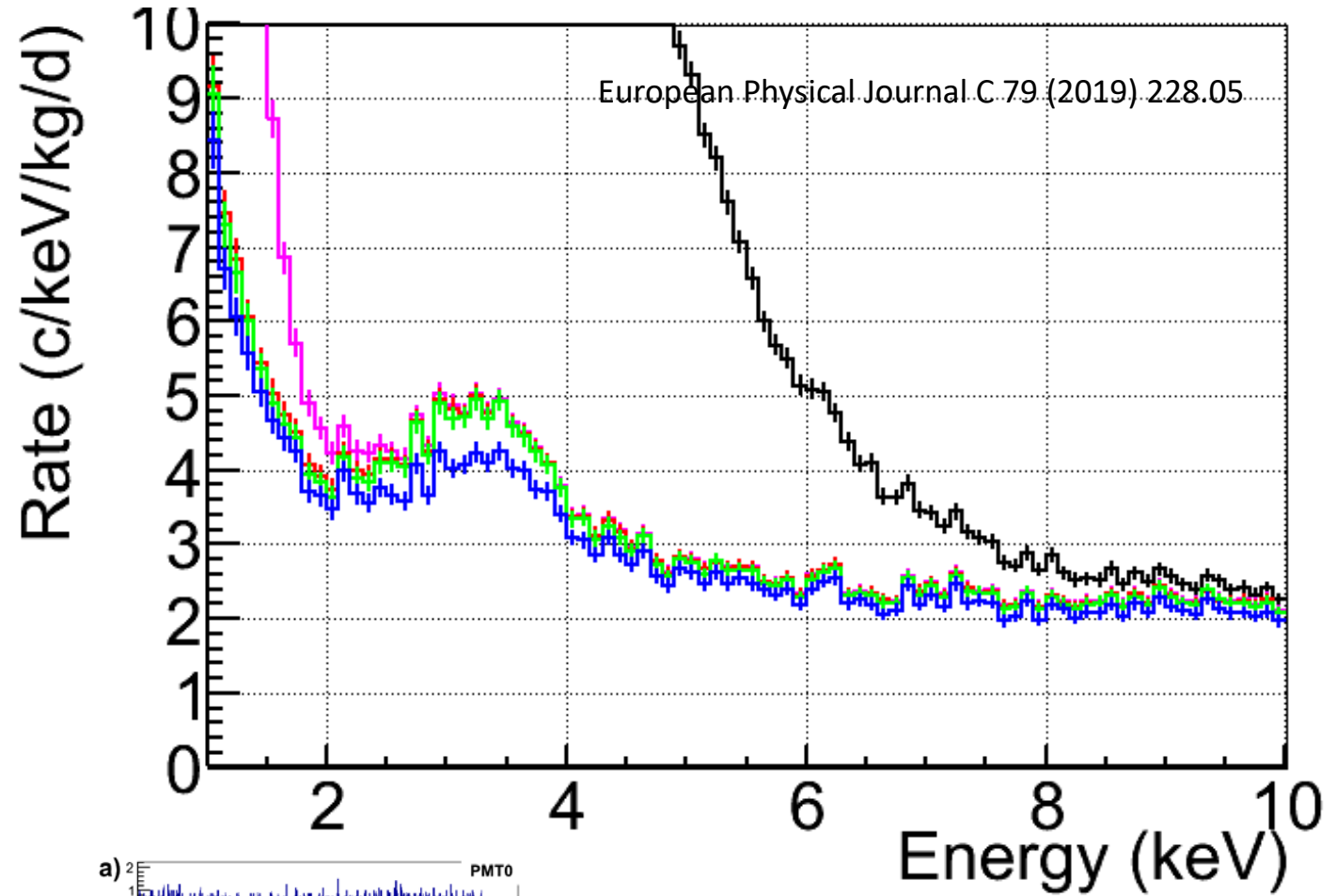
Rejected event

Events arriving more than 1 second after a muon

Single Hit events



Rejected events having multiplicity > 1

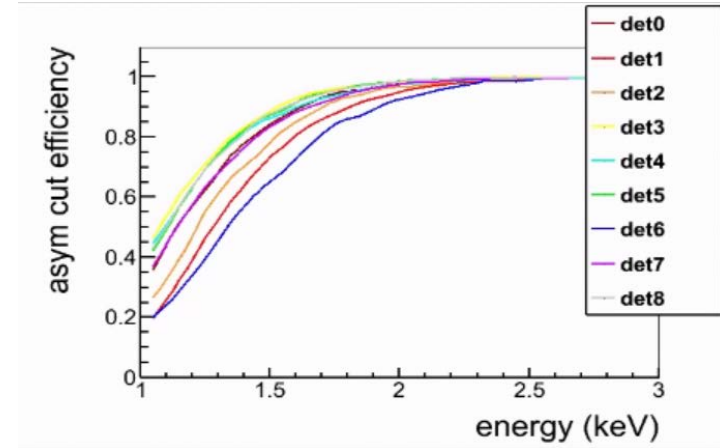
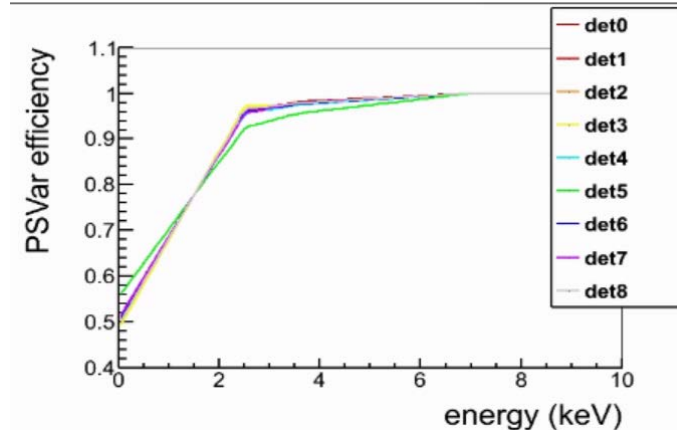
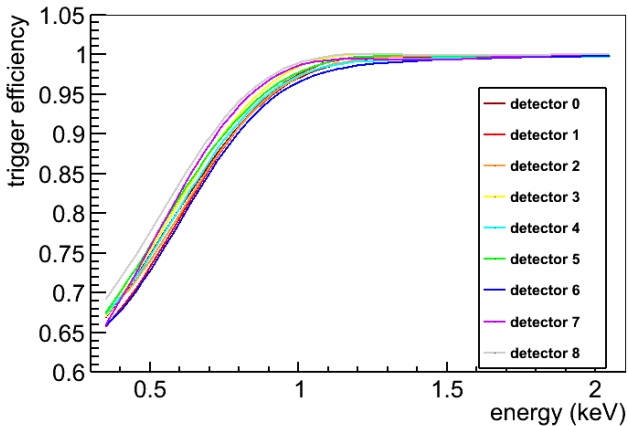
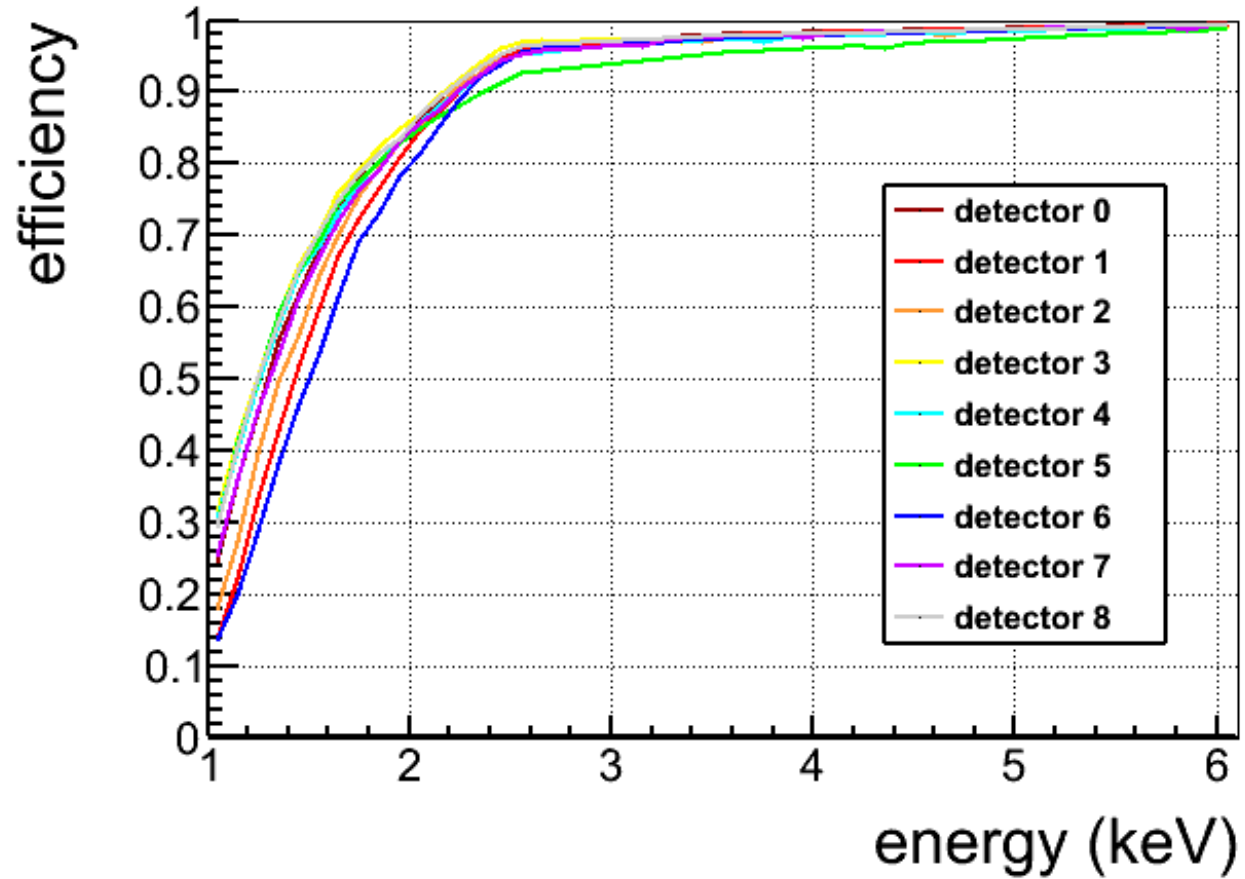


Accepted event

1 keV_{ee}

**Efficiencies used in PRD21 analysis
are calculated with the full
statistics detector by detector
-> 3 years**

$$\epsilon(E, d) = \epsilon_{trg}(E, d) \times \epsilon_{PSA}(E, d) \times \epsilon_{asy}(E, d),$$



Boosted Decision Tree (BDT) has been trained for events' filtering

NEW ANALYSIS

Following preliminary results in

I. Coarasa *et al* 2021 *J. Phys.: Conf. Ser.* 2156 012036

-Multivariate analysis

-More powerful than standard filtering by combining several variables into one parameter

$$P_1 = \frac{\int_{t=100\text{ns}}^{t=600\text{ns}} A(t) dt}{\int_{t=0\text{ns}}^{t=600\text{ns}} A(t) dt}$$

$$P_2 = \frac{\int_{t=0\text{ns}}^{t=50\text{ns}} A(t) dt}{\int_{t=0\text{ns}}^{t=600\text{ns}} A(t) dt}$$

$$\text{CAP}_x = \frac{\int_0^{t=x} A(t) dt}{\int_{t=0\text{ns}}^{t=t_{\text{max}}} A(t) dt}$$

$$\mu_p = \frac{\sum_i A_i t_i}{\sum_i A_i}$$

$$A_{\text{asymple}} = \frac{n_{\text{pre},0} - n_{\text{pre},1}}{n_{\text{pre},0} + n_{\text{pre},1}}$$

Good events

- Neutron calibration events from 1 to 2 keV ($P_1 > 0.35$)

Noise events

- Blank module evts (10-28 phe)
80% $P_1 > 0.35$ and 20% $P_1 < 0.35$

PRELIMINARY

-equilibrated training populations (>30000 evts) and independent from bkg data

-70% used for training
30% used for testing

-CUT defined for every energy bin and detector

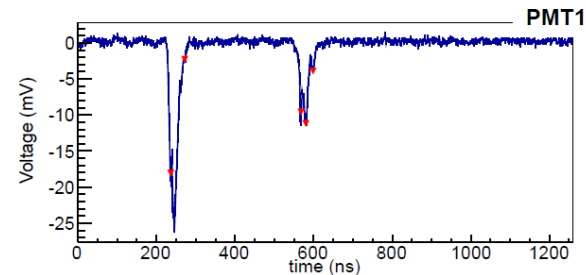
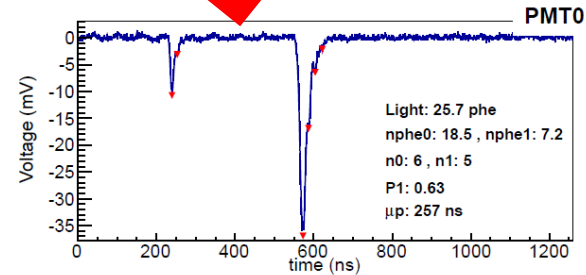
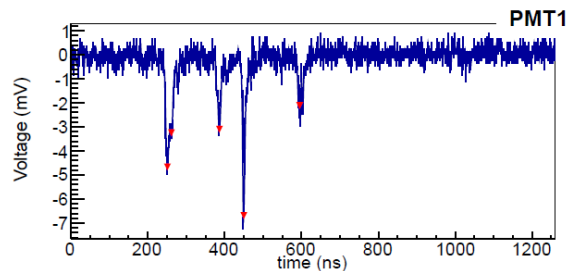
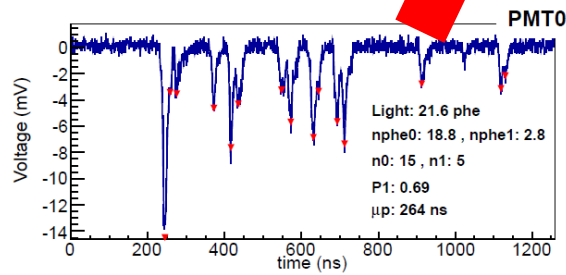
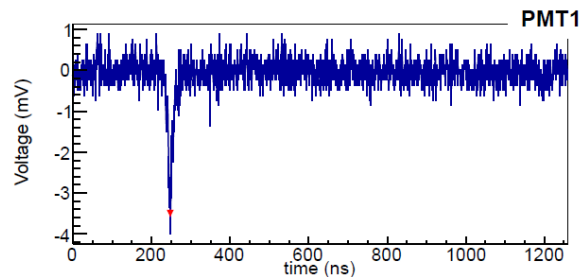
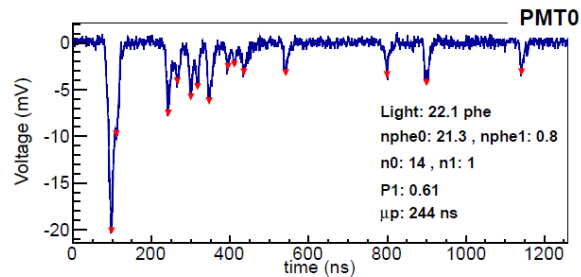
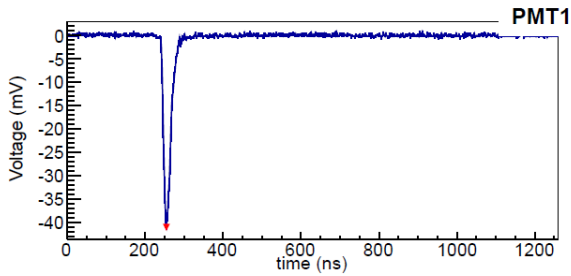
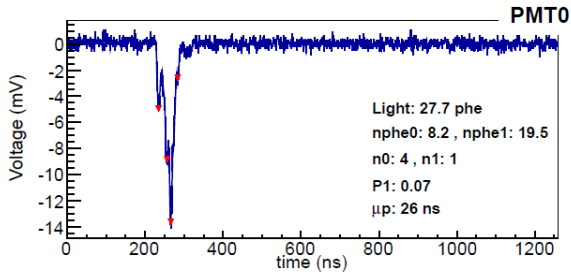
Boosted Decision Tree (BDT) has been trained for events' filtering

Examples of pulses from blank module selected for the BDT training

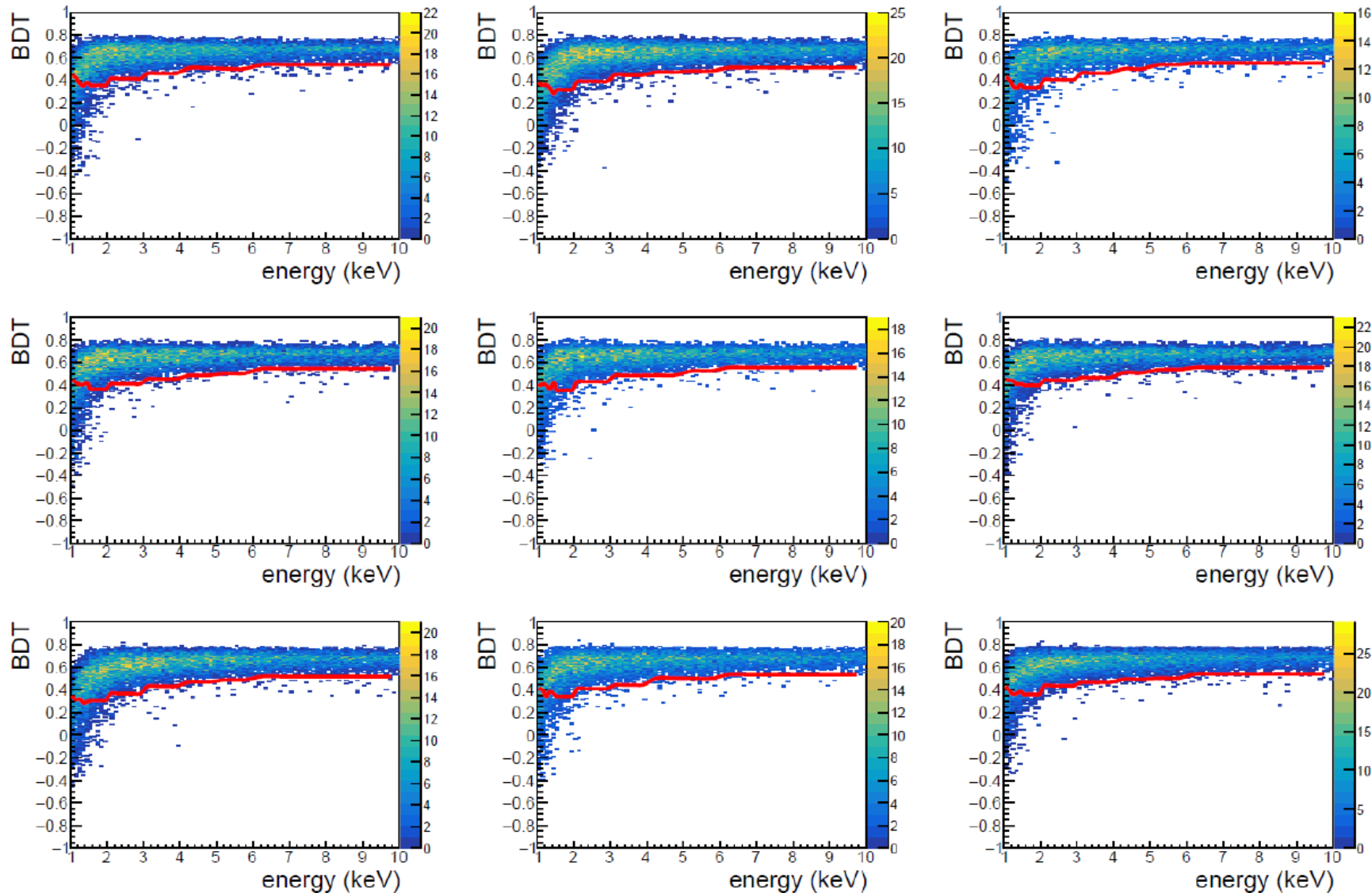
20% $P1 < 0.35$

These two passed the previous ANAIS-112 cuts

80%
 $P1 > 0.35$

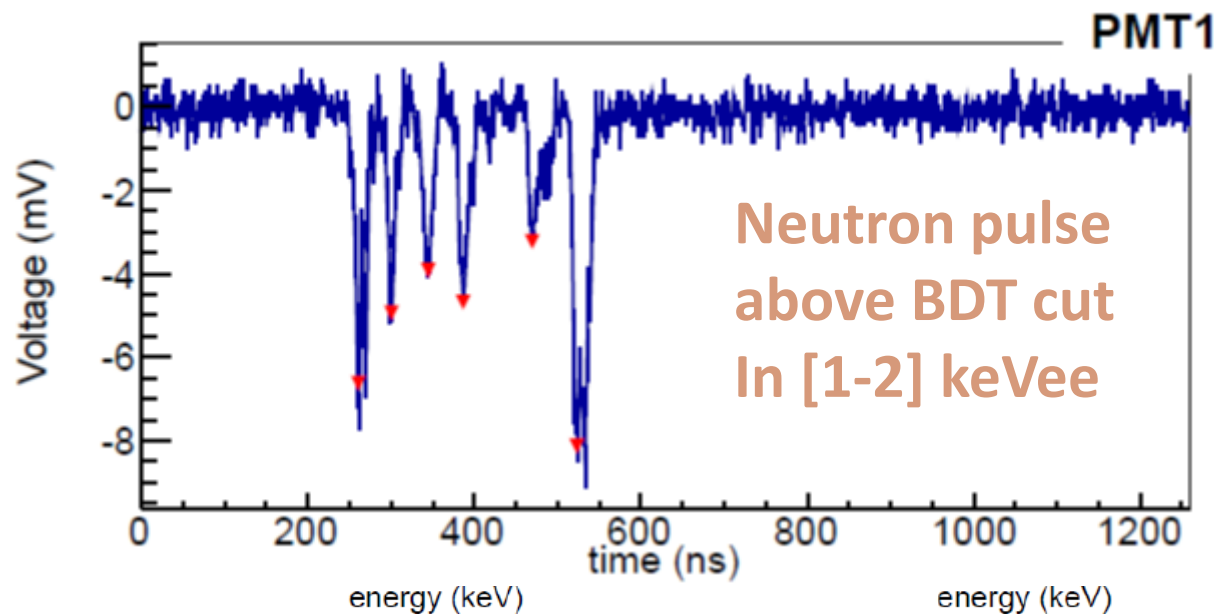
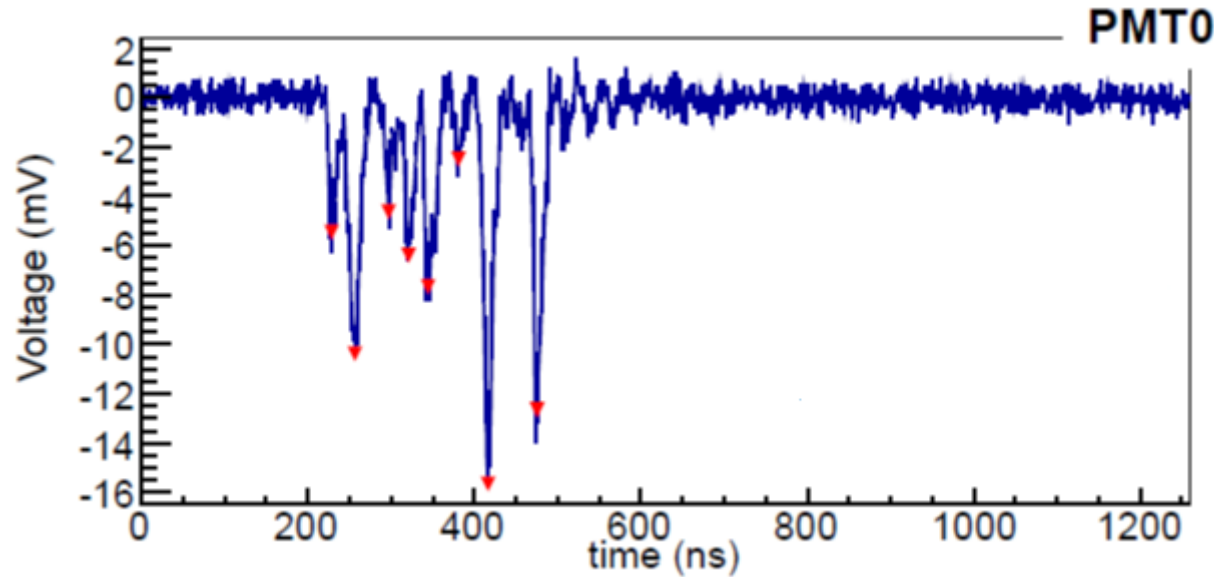
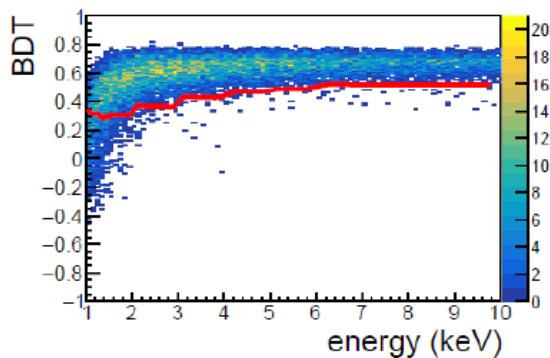
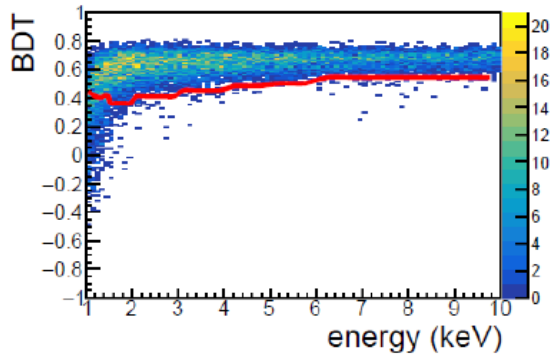
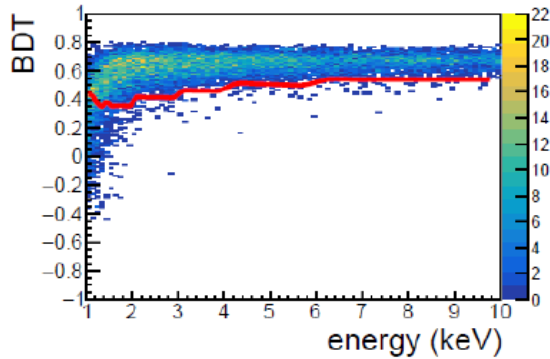


Boosted Decision Tree (BDT) has been trained for events' filtering



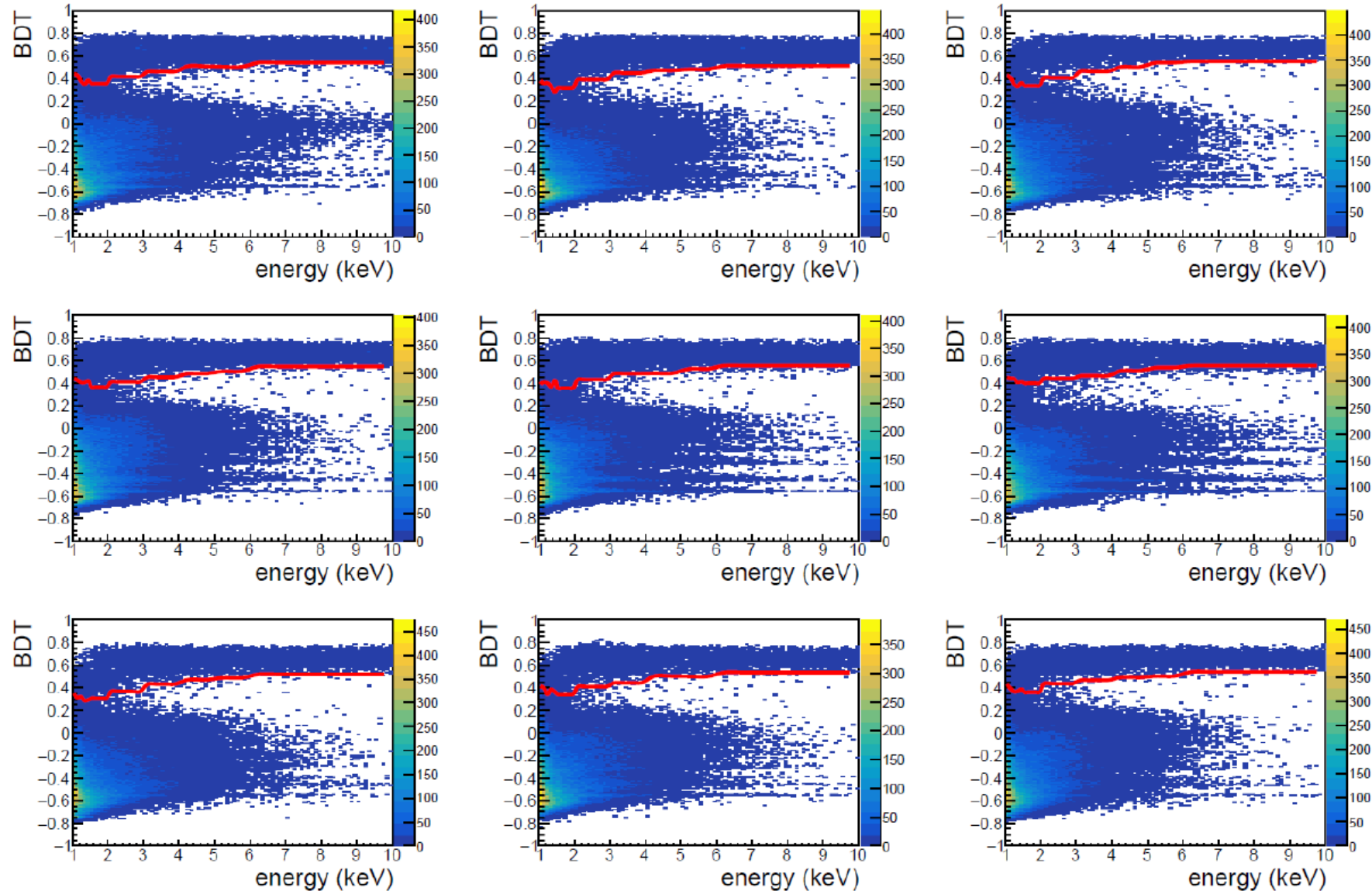
CUT on the BDT parameter defined for every energy bin using the neutron population - coincident events

Boosted Decision Tree (BDT) has been trained for events' filtering



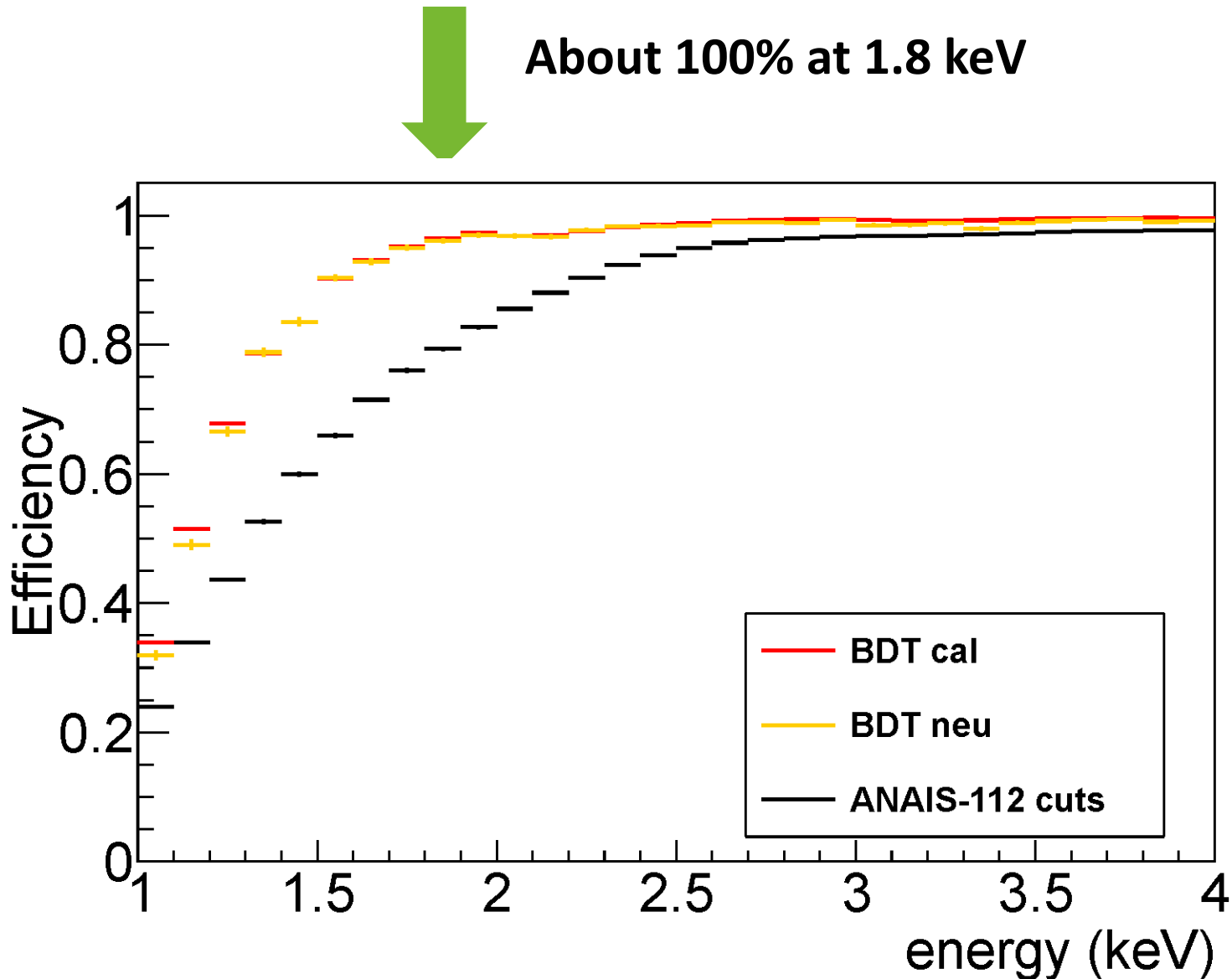
BDT parameter defined for every energy bin
removing the neutron population - coincident events

Boosted Decision Tree (BDT) has been trained for events' filtering



CUT on BDT parameter applied to the background (10% events of the first year)

Boosted Decision Tree (BDT) has been trained for events' filtering



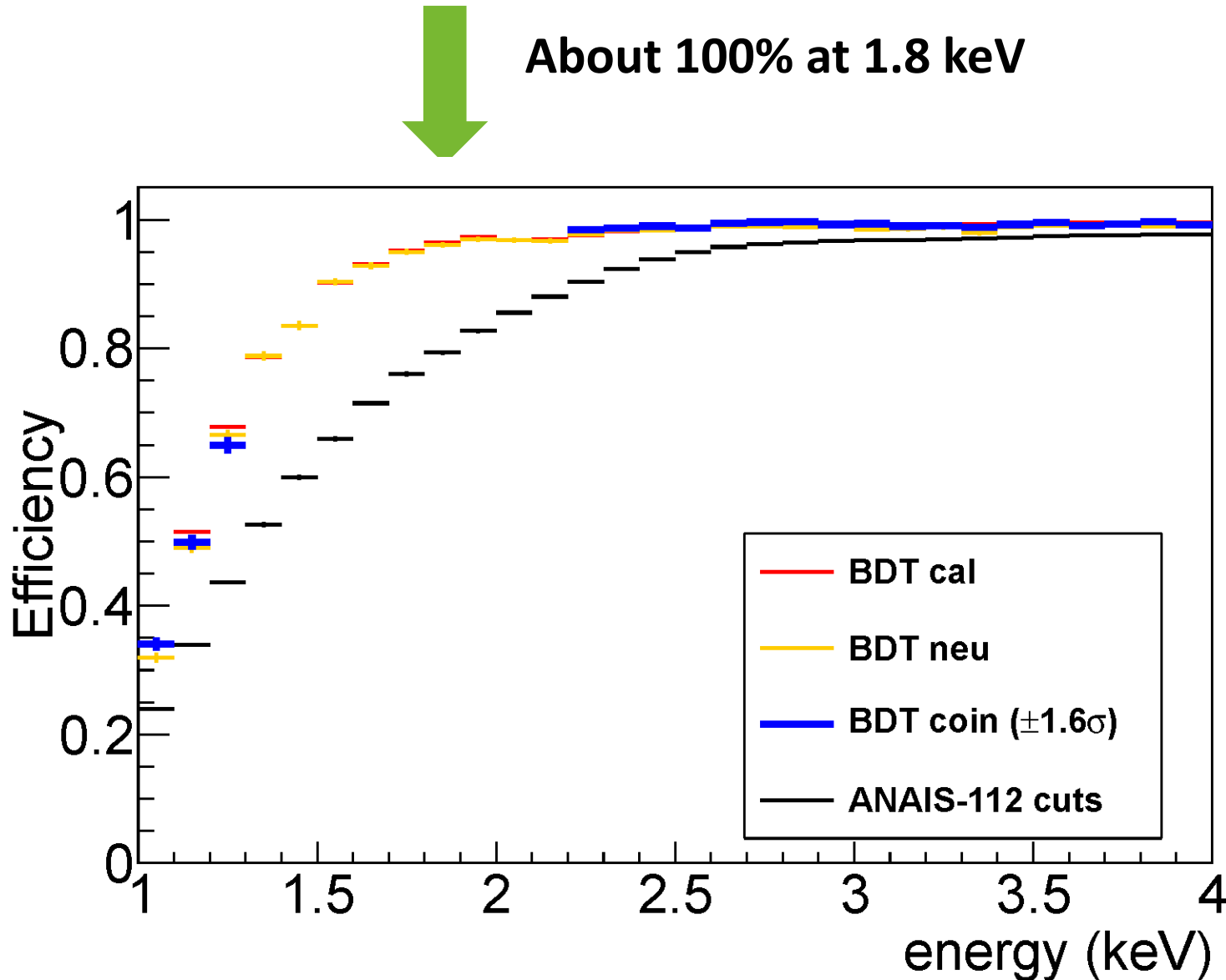
Remarkable improvement in efficiency in the ROI

Efficiencies determined with different populations of scintillation events:

- neutrons
- ^{109}Cd events

FULLY COMPATIBLE

Boosted Decision Tree (BDT) has been trained for events' filtering



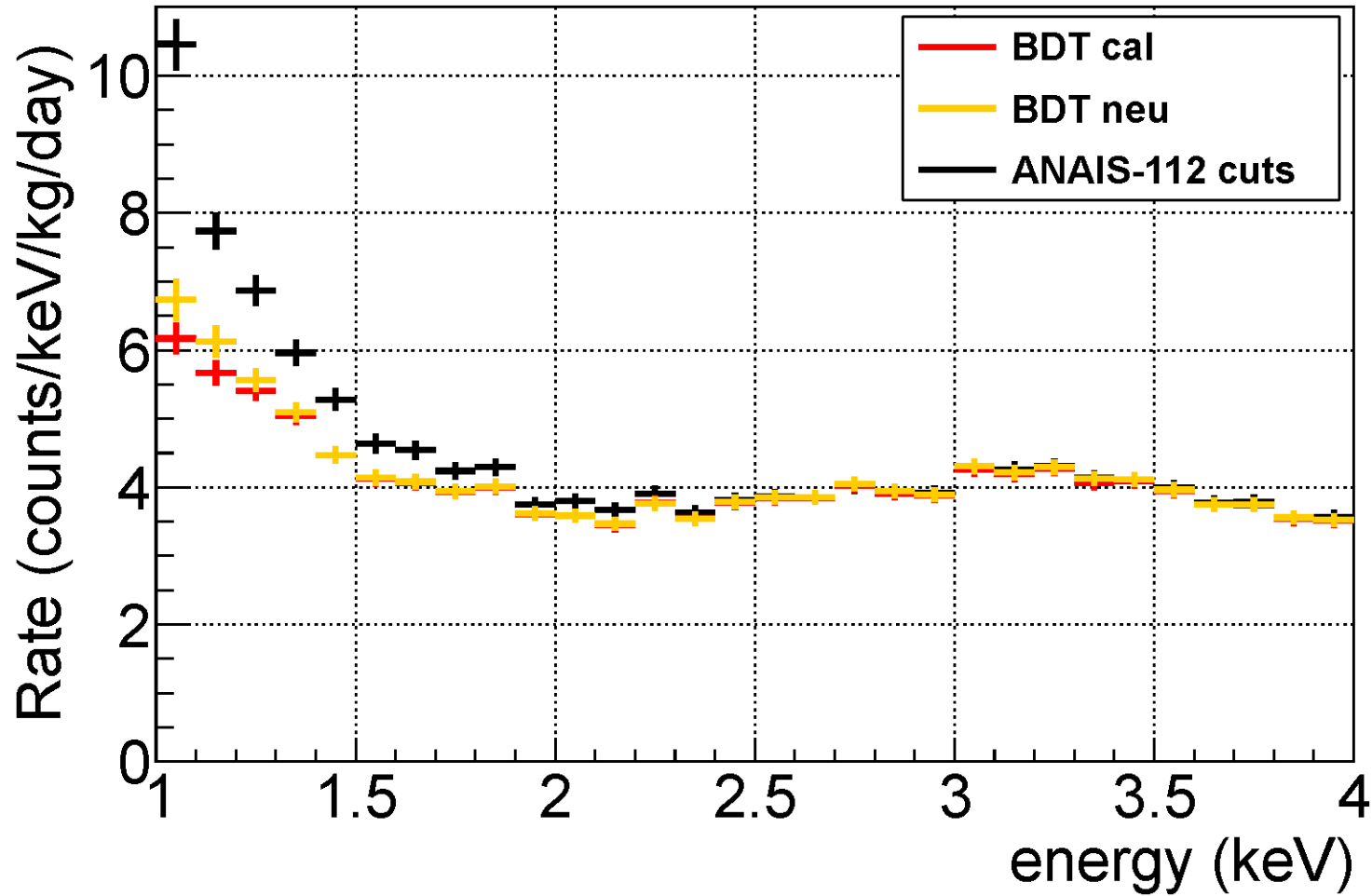
Remarkable improvement in efficiency in the ROI

Efficiencies determined with different populations of scintillation events:

- neutrons
- ^{109}Cd events
- ^{40}K and ^{22}Na coincident evts

FULLY COMPATIBLE

Boosted Decision Tree (BDT) has been trained for events' filtering



Relevant background reduction

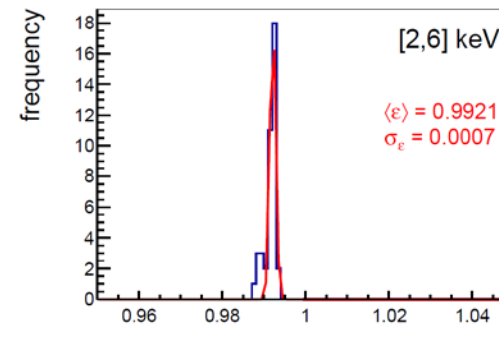
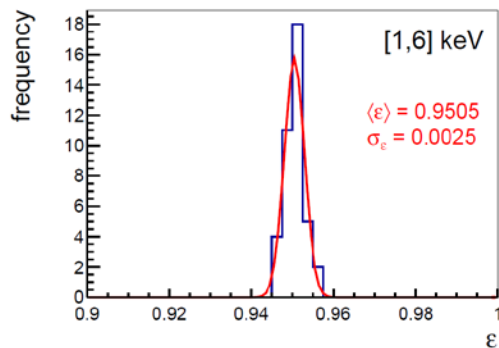
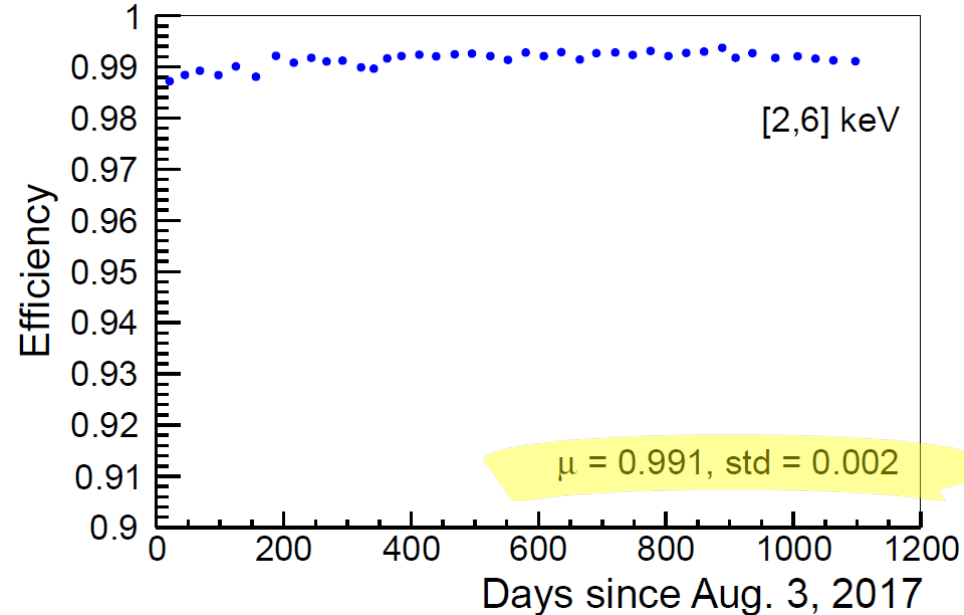
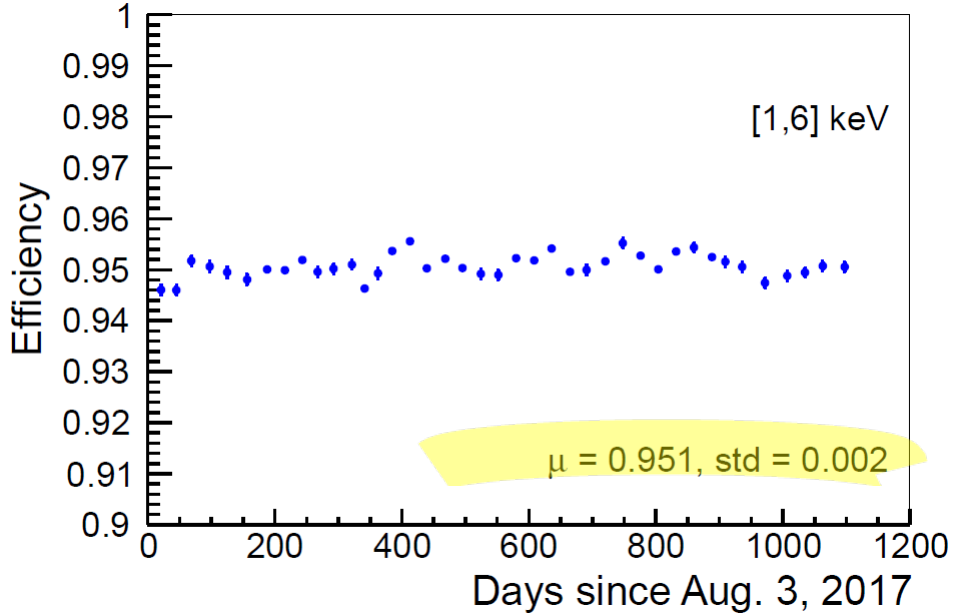
17.2% in [1-2] keV



We are working on the reduction of our analysis threshold profiting from the improvement in efficiency achieved by the BDT filtering

Efficiency stability and associated systematic uncertainty

We are working on determining the possible variation in time of the BDT's efficiencies
 Using ^{109}Cd data for the first three years with all detectors averaged



0.5%

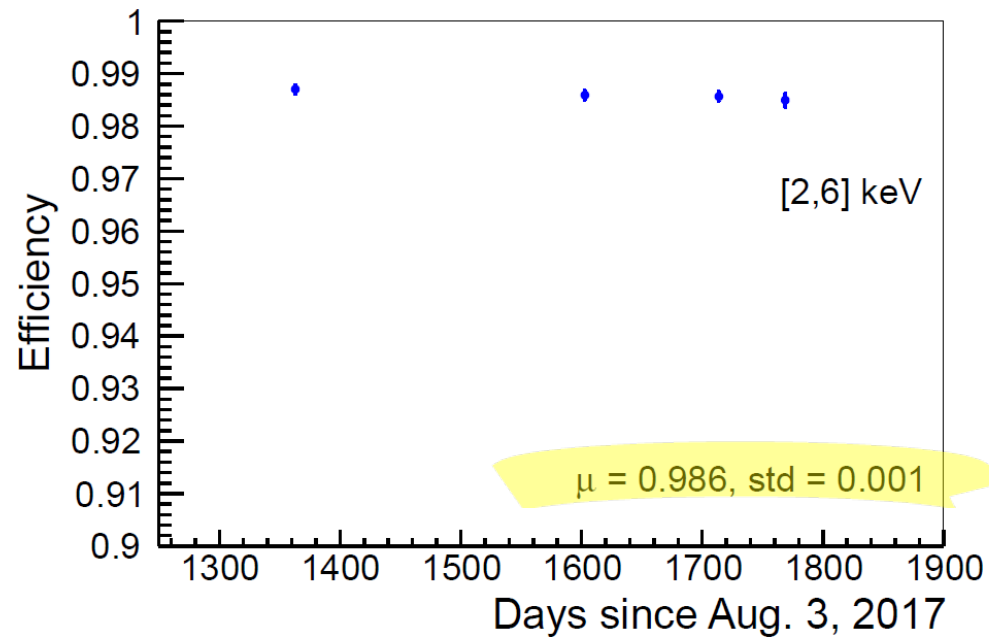
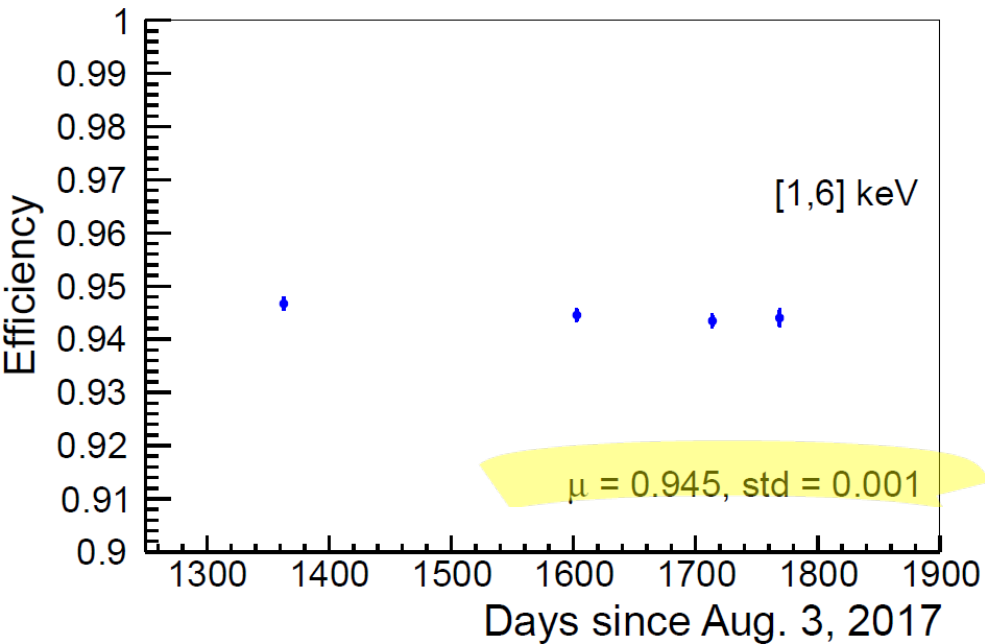
0.33%

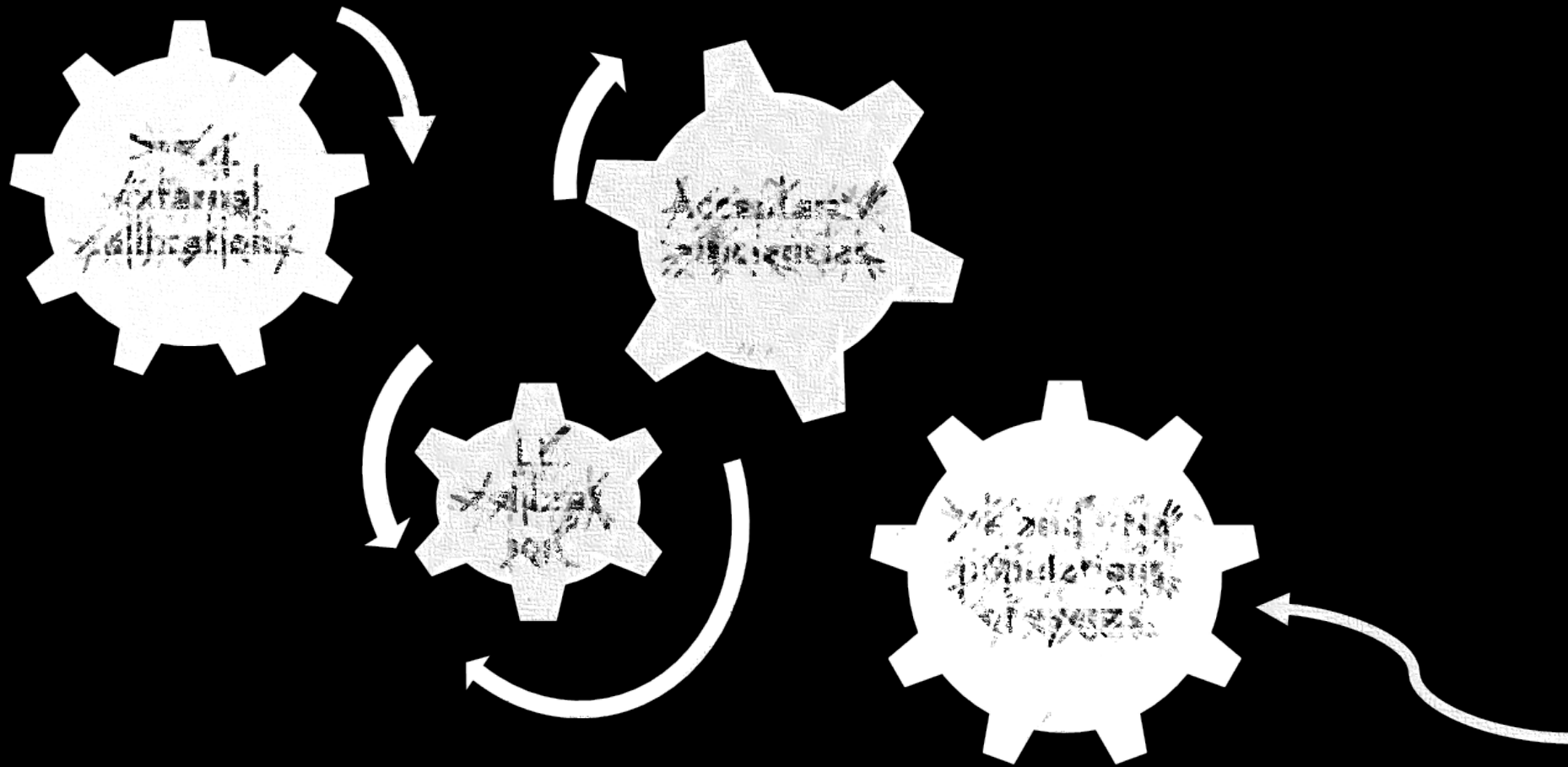


Systematic uncertainty is taken as half the difference between maximum and minimum values

Efficiency stability and associated systematic uncertainty

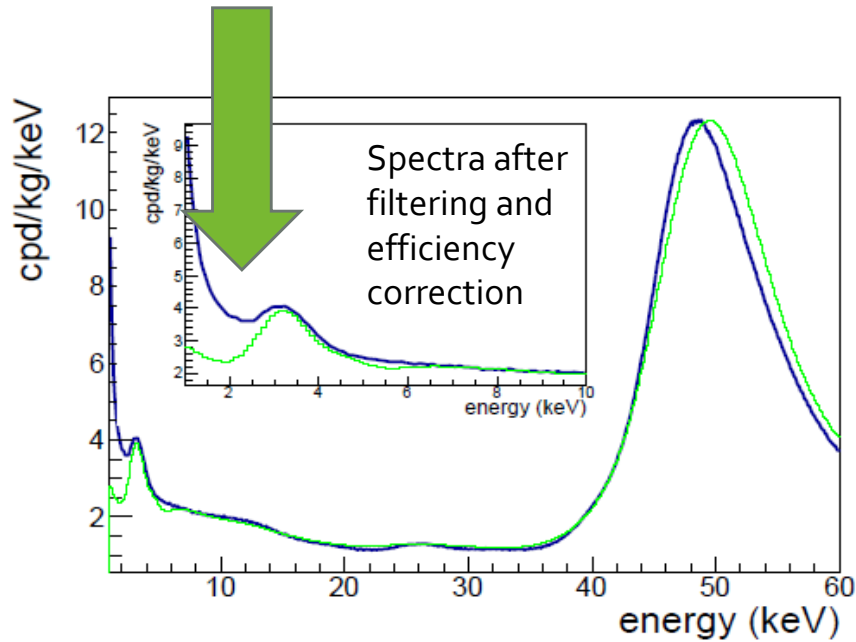
We are working on determining the possible variation in time of the BDT's efficiencies
Using neutron data (4th-5th years)





**ANAIS-112 annual modulation analysis
strategy – DATA BLINDED**

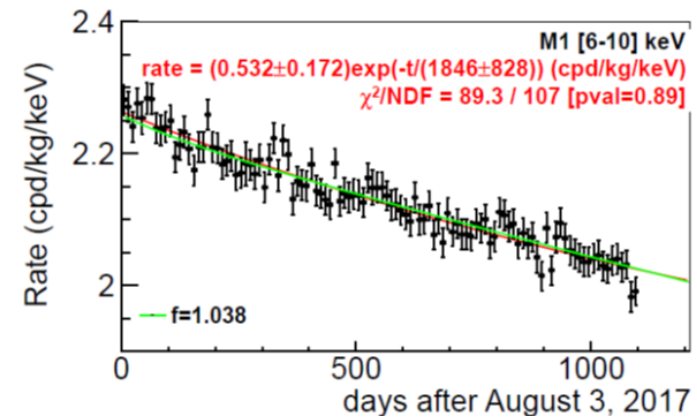
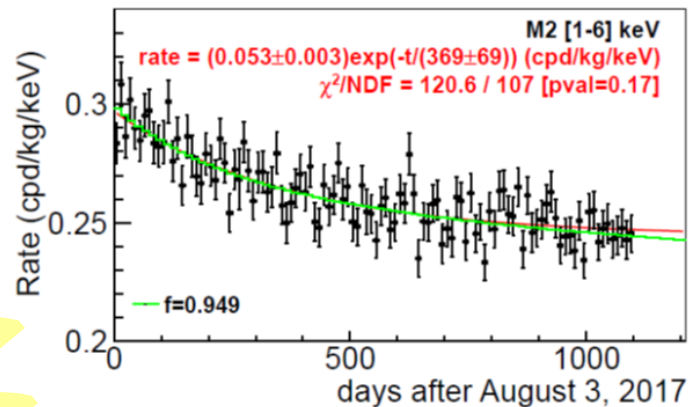
Robust background model



- ROI background dominated by ^{210}Pb , ^{40}K and ^3H -> higher than DAMA/LIBRA
- Good agreement in all energy regions, but underestimate in 1-2 keV region
- Background model established before unblinding data

It predicts time evolution of the background detector by detector

It reproduces satisfactorily time evolution outside the ROI and in the ROI for multiple-hit events



313.95 kg x y (three-year exposure): Annual modulation search

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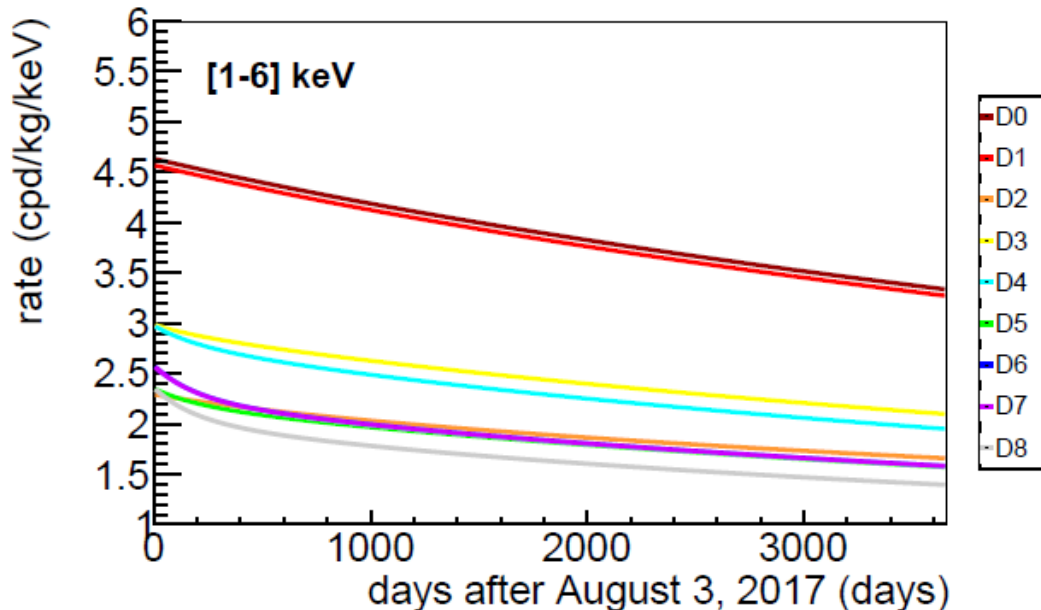
Background probability distribution function drawn from ANAIS-112 background model for every detector to account for possible systematic effects related with the different backgrounds and efficiencies of the different modules

MODEL INDEPENDENT ANALYSIS

Minimizing:

$$\chi^2 = \sum_{i,d} \frac{(n_{i,d} - \mu_{i,d})^2}{\sigma_{i,d}^2}$$

$$\mu_{i,d} = [R_{0,d}(1 + f_d \phi_{bkg,d}^{MC}(t_i)) + S_m \cos(\omega(t_i - t_0))] M_d \Delta E \Delta t,$$



$n_{i,d}$, $\sigma_{i,d}$ are number of events (and Poisson uncertainty) in 10 days bins corrected by live time and efficiency for detector d

ω and t_0 are fixed while S_m is set free (modulation hypothesis) or fixed to 0 (non-modulation hypothesis)

ANAIS-112 three year results – annual modulation analysis (PRD2021)

Null hyp $\chi^2/\text{n.d.f.}$: 1075.81/972 [$p_{\text{val}} = 0.011$]

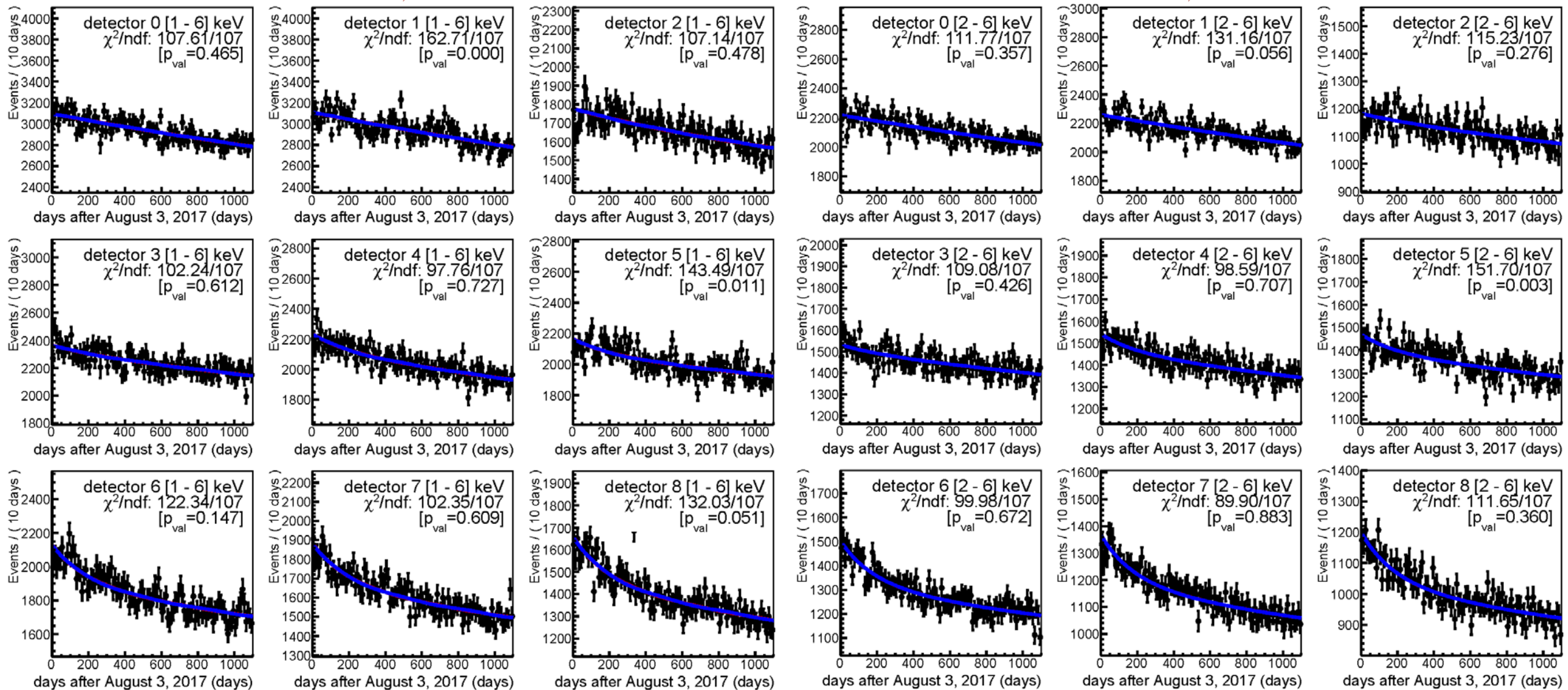
Mod hyp $\chi^2/\text{n.d.f.}$: 1075.15/971 [$p_{\text{val}} = 0.011$]

$$S_m = (-0.0034 \pm 0.0042) \text{ (cpd/kg/keV)}$$

Null hyp $\chi^2/\text{n.d.f.}$: 1018.19/972 [$p_{\text{val}} = 0.148$]

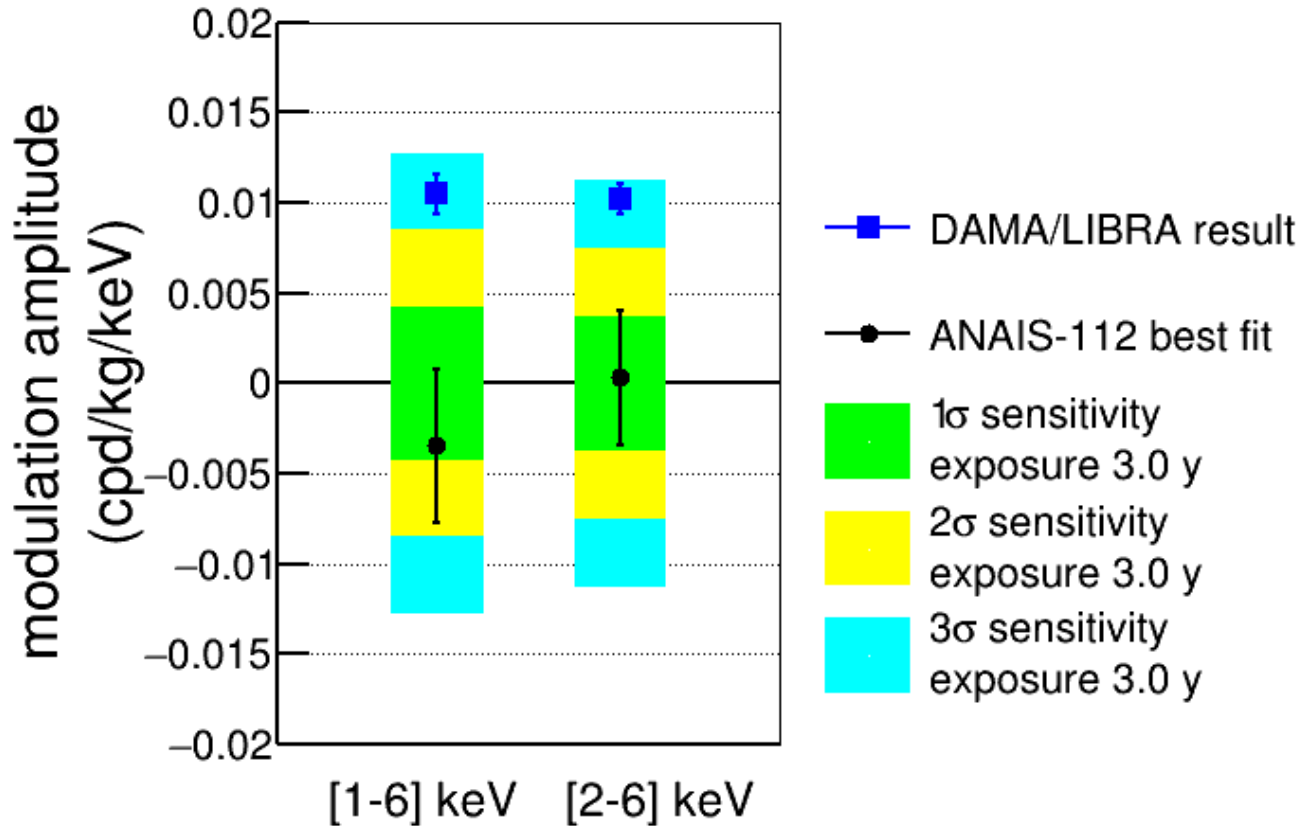
Mod hyp $\chi^2/\text{n.d.f.}$: 1018.18/971 [$p_{\text{val}} = 0.143$]

$$S_m = (0.0003 \pm 0.0037) \text{ (cpd/kg/keV)}$$



ANAIS-112 three year results – annual modulation analysis (PRD2021)

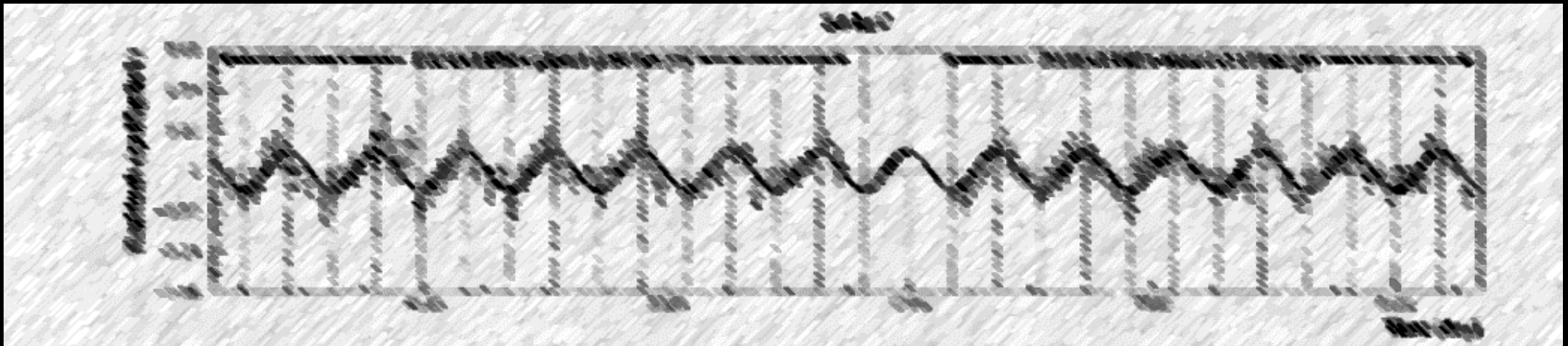
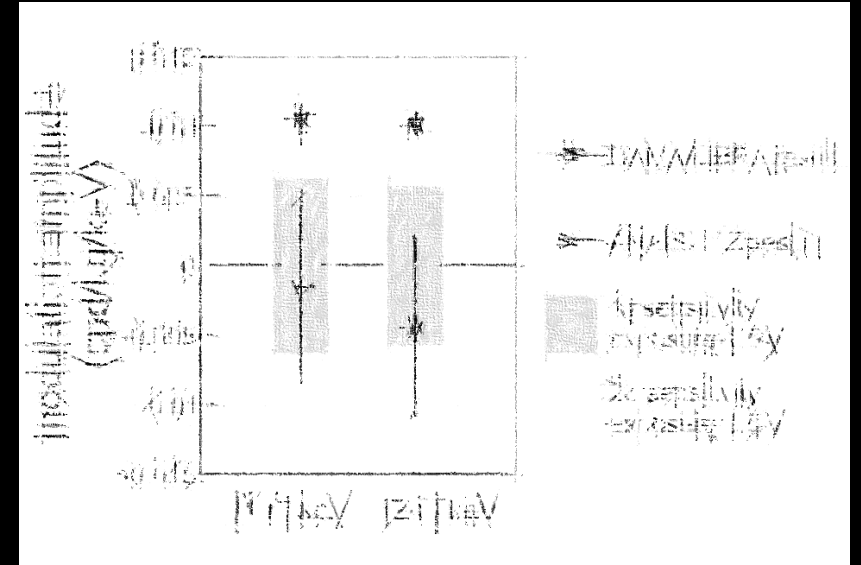
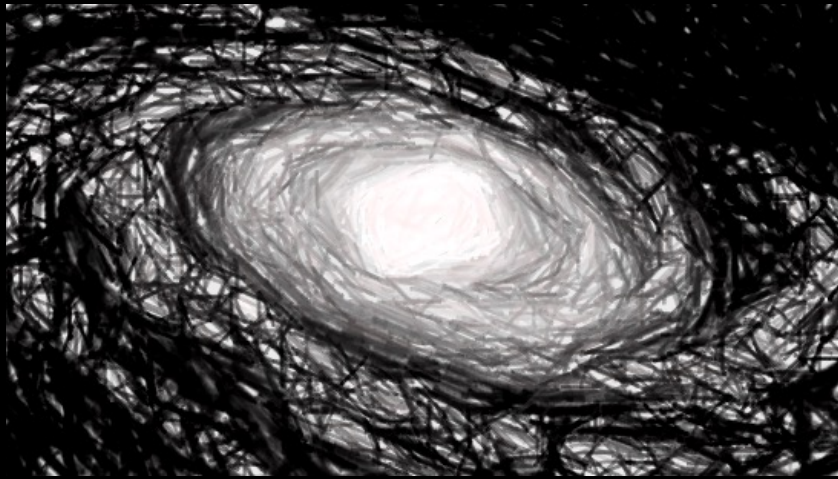
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- Best fits are incompatible with DAMA/LIBRA result at 3.3 and 2.6 σ in [1-6] and [2-6] keV energy regions
- Sensitivity is at 2.5 and 2.7 σ in [1-6] and [2-6] keV energy regions



ANAIS-112 Three Years Annual Modulation re-analysis



ANAIS-112 three year results – annual modulation re-analysis using BDTs

Null hyp χ^2/ndf : 1010.88/972 [$p_{\text{val}}=0.188$]

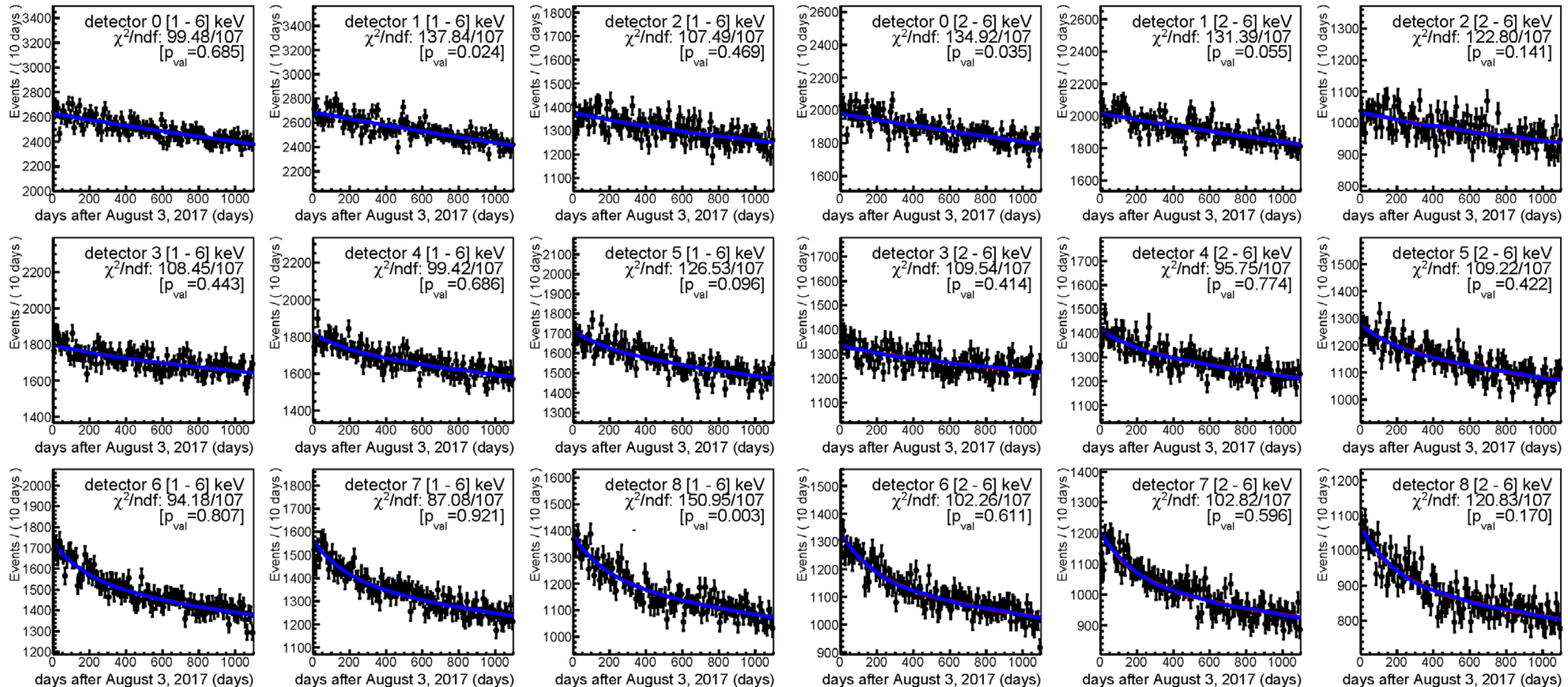
Mod hyp χ^2/ndf : 1010.62/971 [$p_{\text{val}}=0.183$]

$$S_m = (-0.0016 \pm 0.0032) \text{ (cpd/kg/keV)}$$

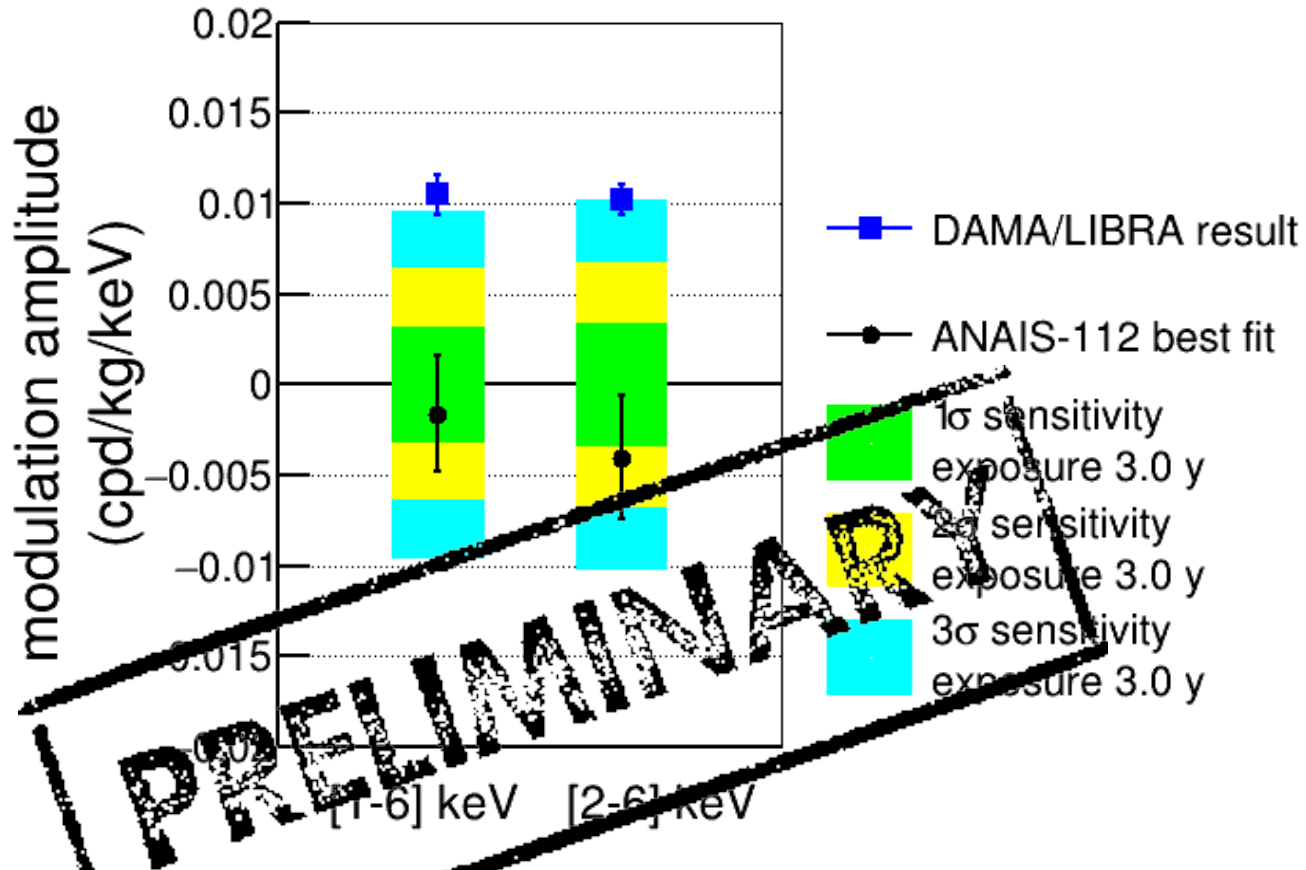
Null hyp χ^2/ndf : 1029.87/972 [$p_{\text{val}}=0.096$]

Mod hyp χ^2/ndf : 1028.50/971 [$p_{\text{val}}=0.098$]

$$S_m = (-0.0040 \pm 0.0034) \text{ (cpd/kg/keV)}$$

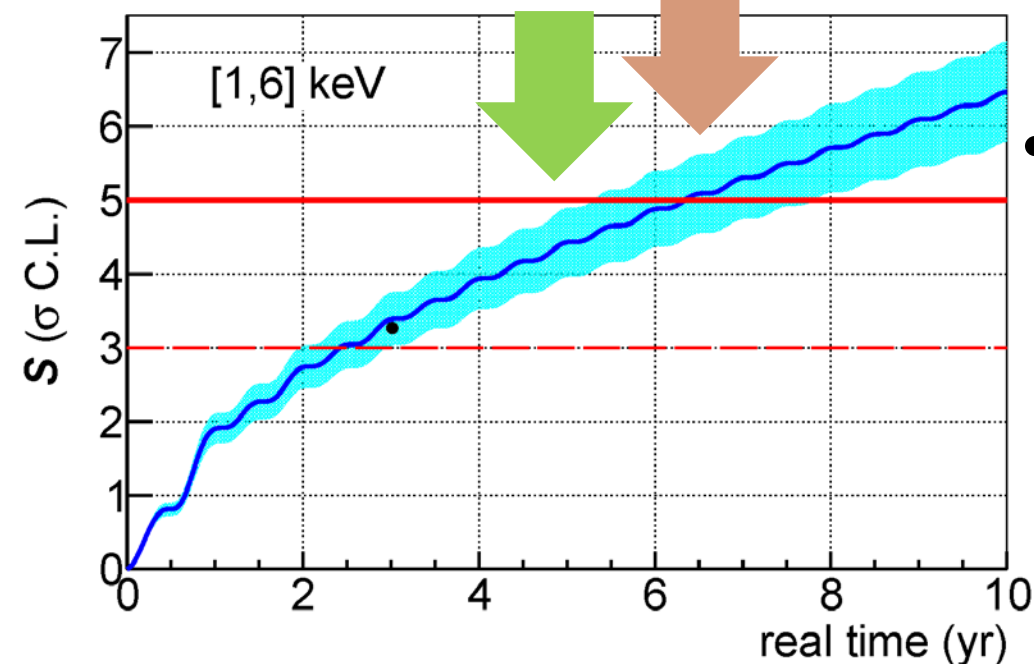
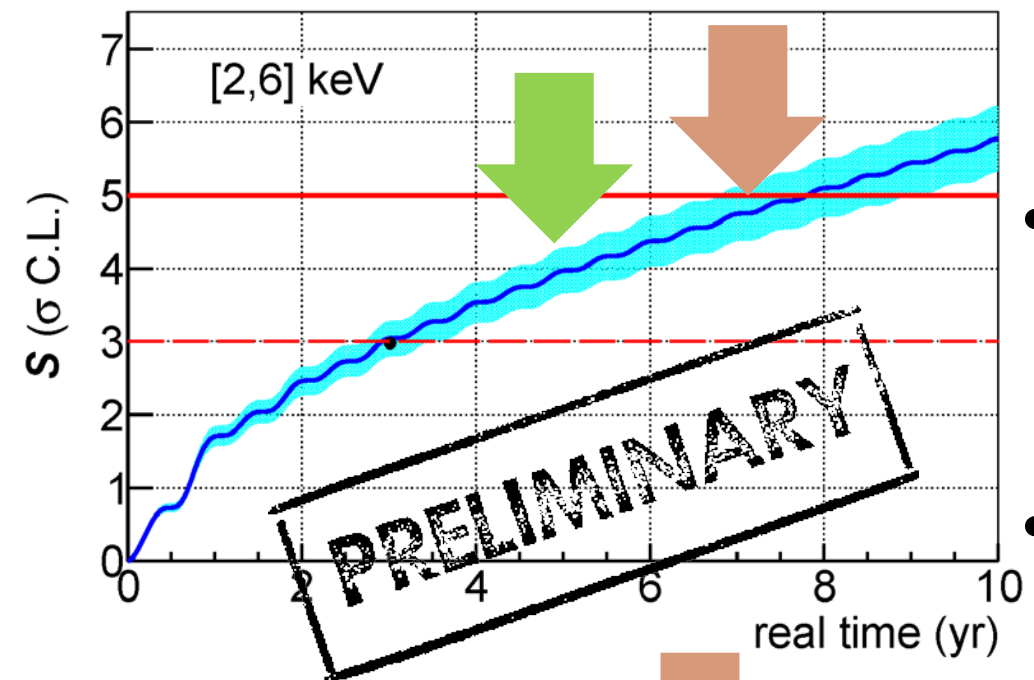


ANAIS-112 three year results – annual modulation re-analysis using BDTs



- Best fits are incompatible with DAMA/LIBRA result at 3.8 and 4.2 σ in [1-6] and [2-6] keV energy regions
- Sensitivity is at 3.3 and 3.0 σ in [1-6] and [2-6] keV energy regions



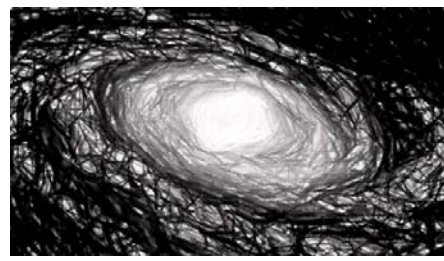
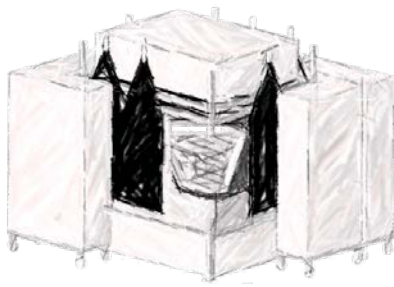


- Our “a priori” sensitivity estimates have been adapted to the new BDT – filtering: IMPROVED background and efficiencies improve our prospects
- Including systematic uncertainty in the efficiency: 0.4% systematic uncertainty quadratically added to the statistical in each energy bin and detector
- We should be at 5σ from DAMA/LIBRA result with two more years of data taking in [1-6] keV region without further analysis improvement

Statistical significance of our result is determined by the standard deviation of the modulation amplitude distribution, $\sigma(S_m)$

Summary and Outlook





ANAIS-112 data taking is progressing smoothly – **almost 5 years !**
Improved analysis will be applied by the end of the year to the full exposure

Neutron calibration “onsite” and complementary QF measurements are relevant for understanding systematics -> more coordinated work from the community would be required

Monitoring “onsite” possible modulations of neutron flux, muon flux, radon content in the air, etc.

Three years annual modulation re-analysis almost ready to publish (NO MODULATION OBSERVED WITH **3σ SENSITIVITY**)

Testing at 5 sigma is at hand with at least 2 more years of data-taking

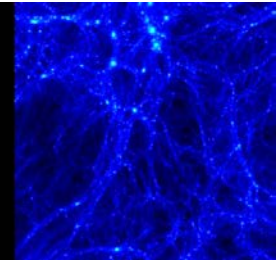
Planning to make our data public at The Dark Matter Data Center (ORIGINS cluster) soon

Thank you for your attention



MultiDark

Multimessenger Approach
for Dark Matter Detection



LSC

Laboratorio Subterráneo de Canfranc



Centro de Astropartículas y
Física de Altas Energías
Universidad Zaragoza