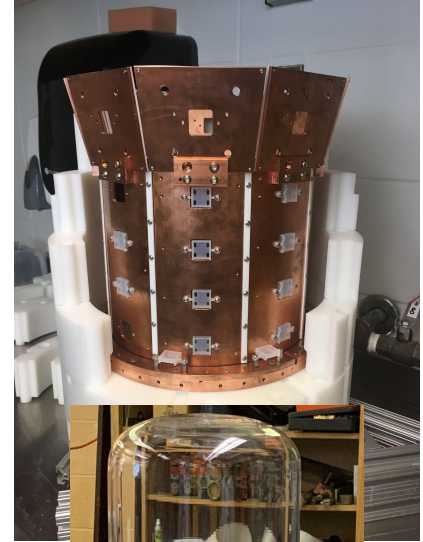


The Scintillating Bubble Chamber Experiment



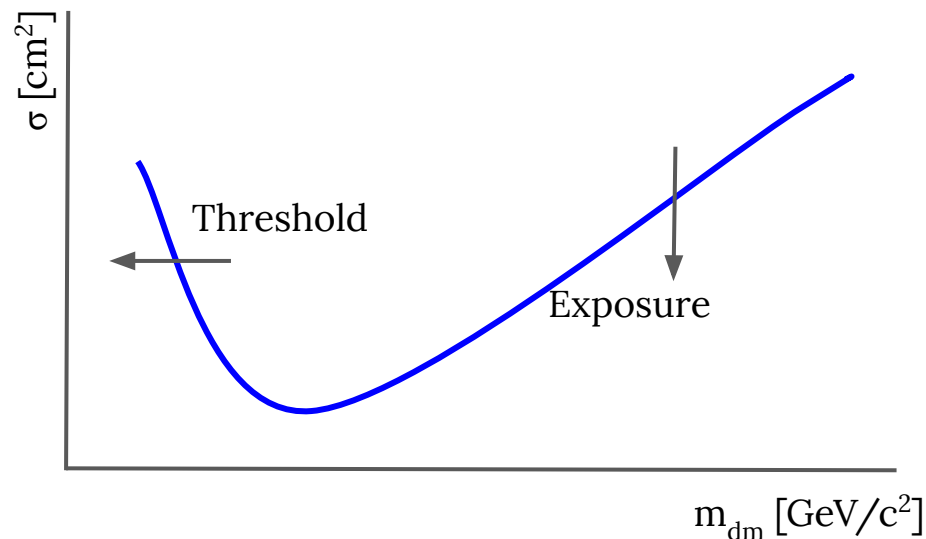
Arthur B. McDonald
Canadian Astroparticle Physics Research Institute

Ben Broerman
IDM 2022



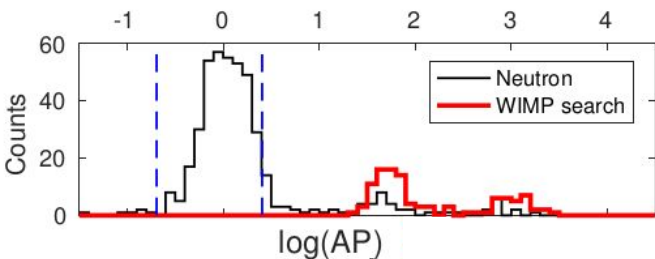
GeV-scale dark matter

- $m_{\text{DM}} \sim m_p$ is motivated if baryon asymmetry is related to DM abundance (e.g. ADM models)
- Difficult kinematics of low mass DM-nucleon scattering require low thresholds
 - Future success here still needs particle identification and scalability

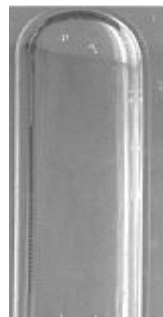
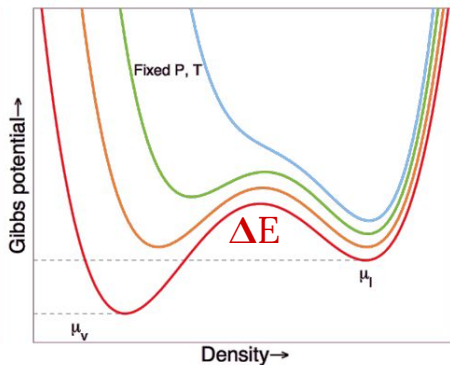


Bubble chambers

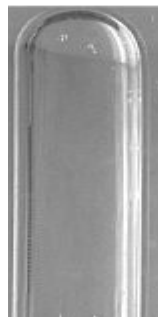
- Bubble chambers maintain target fluid in a superheated state
- High efficiency low n.r. threshold, β/γ insensitivity, n.r./ α discrimination:



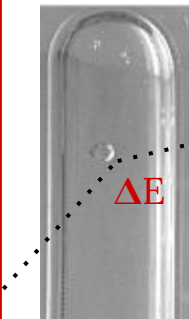
$\{x,y,z\}$ + acoustics



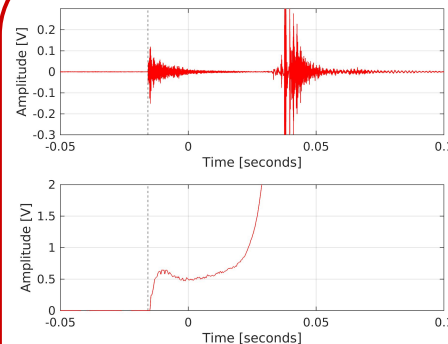
Pressure \downarrow , 2 minima form



Further P \downarrow ,
Metastable superheated



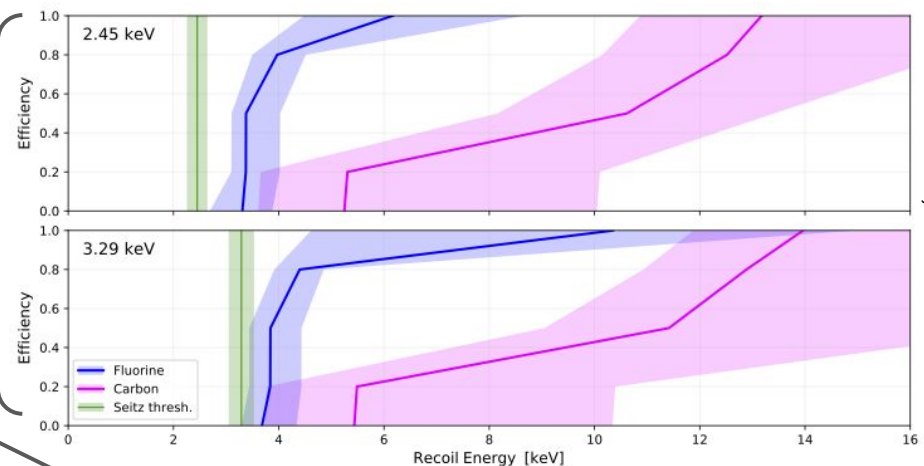
If ΔE deposited, local phase change



See E. Vázquez-Jáuregui's PICO talk

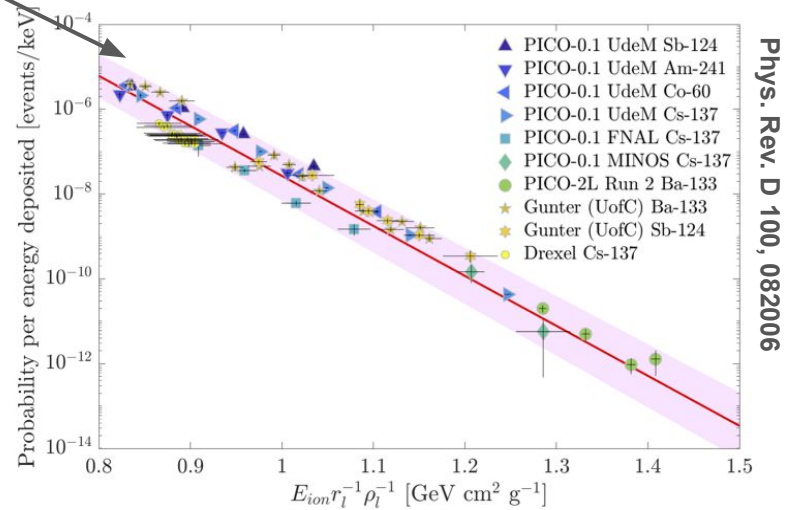
Bubble chambers

- Issues with freon B.C.'s
 - Threshold detectors: no energy information
 - β/γ rejection fails at low thresholds



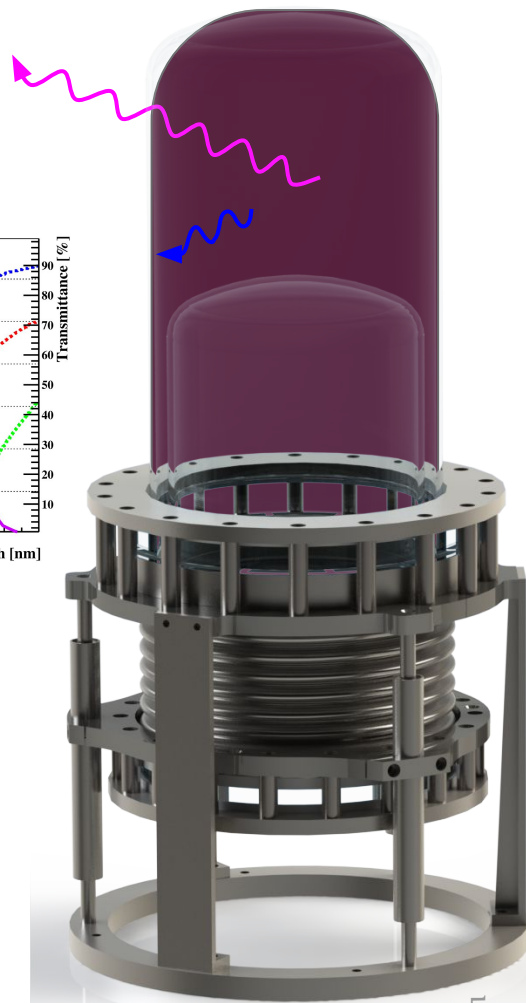
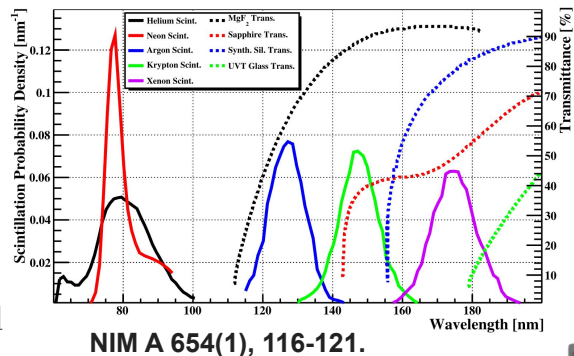
- Noble liquid B.C.'s
 - Scintillation provides energy information
 - Lack molecular vibrational modes to provide local heating (possible increase in β/γ suppression at low thresholds)

- **No bubbles in pure Xe B.C., but nucleation in Xe + ethylene (quenched scintillation)**

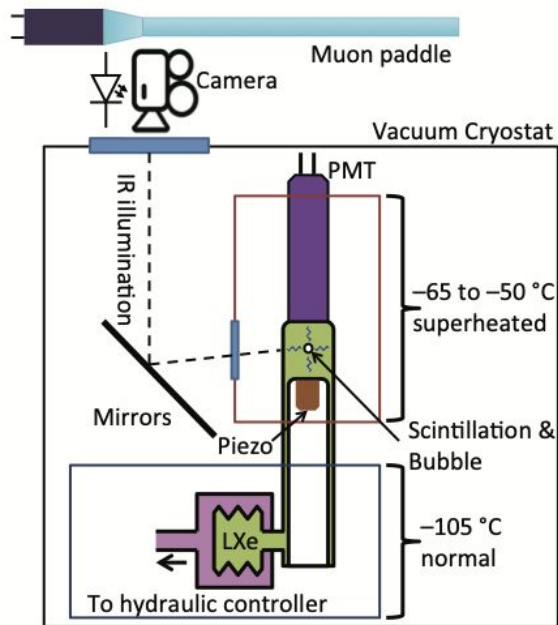


Liquid noble bubble chambers

- Argon 3x lighter than Xe
 - Ar for low mass SI, Xe-doping to get the light out
- Cryogenic temperatures
 - Everything that moves, must move cold
- Signal timing difference
 - Scintillation $\mathcal{O}(\text{ns})$
 - Bubble formation/acoustic emission $\mathcal{O}(\mu\text{s})$
 - Recompression $\mathcal{O}(\text{ms})$

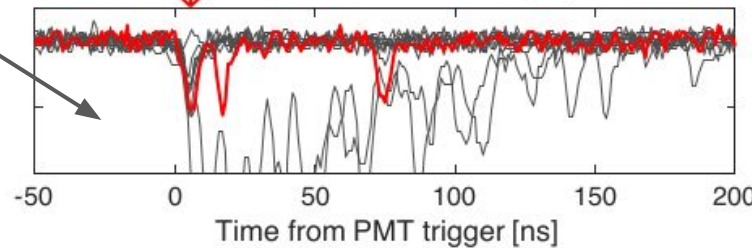
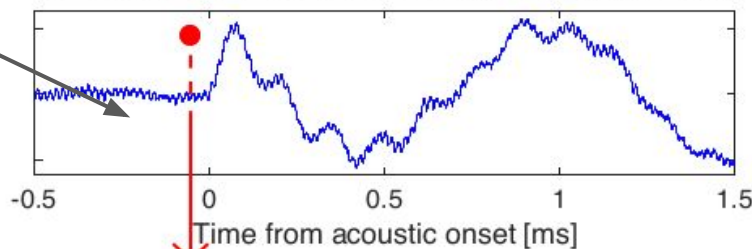
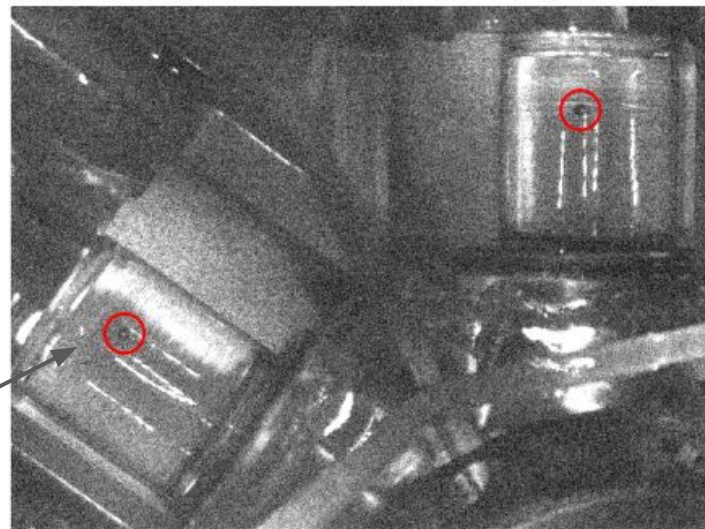


Prototype LXe bubble chamber



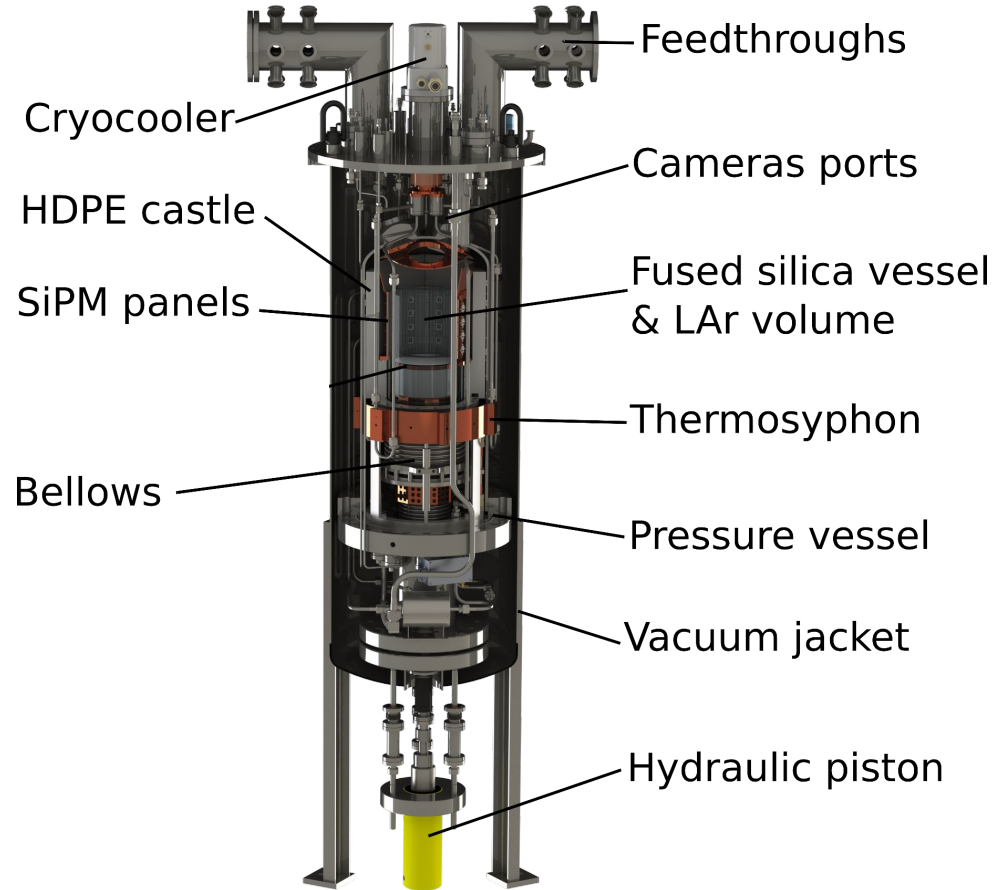
Simultaneous measurement of

- Bubble position (camera)
- Acoustic emission (piezo transducer)
- Scintillation (UV-grade PMT)



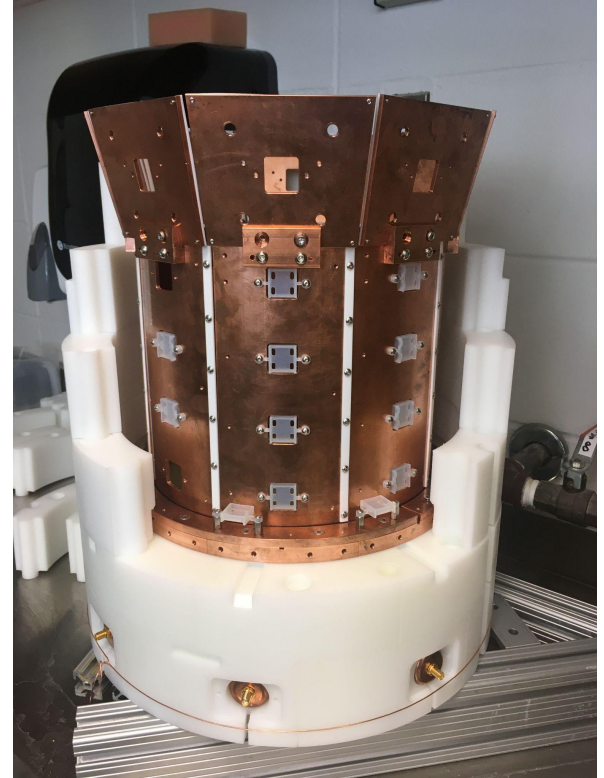
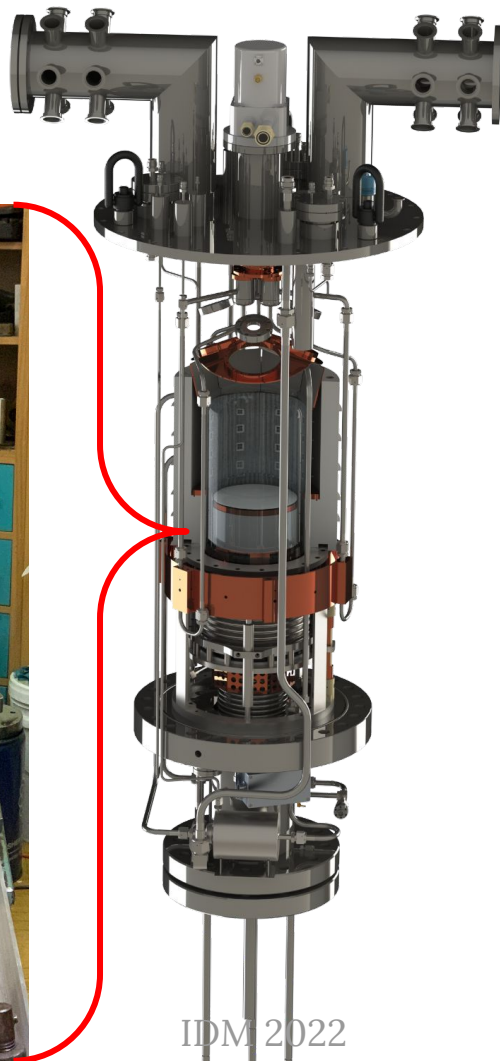
SBC detectors

- Two functionally-similar detectors operated at:
 - SBC-LAr10 @ Fermilab (engineering, calibration, CE ν NS)
 - SBC-SNOLAB @ SNOLAB (low bkg. dark matter)
- 10 kg LAr + Xe as a wavelength shifter targeting 100 eV n.r. threshold
- Cold region ~90 K, warm region ~130 K
- Expanded ~30 psi, compressed ~200 psi



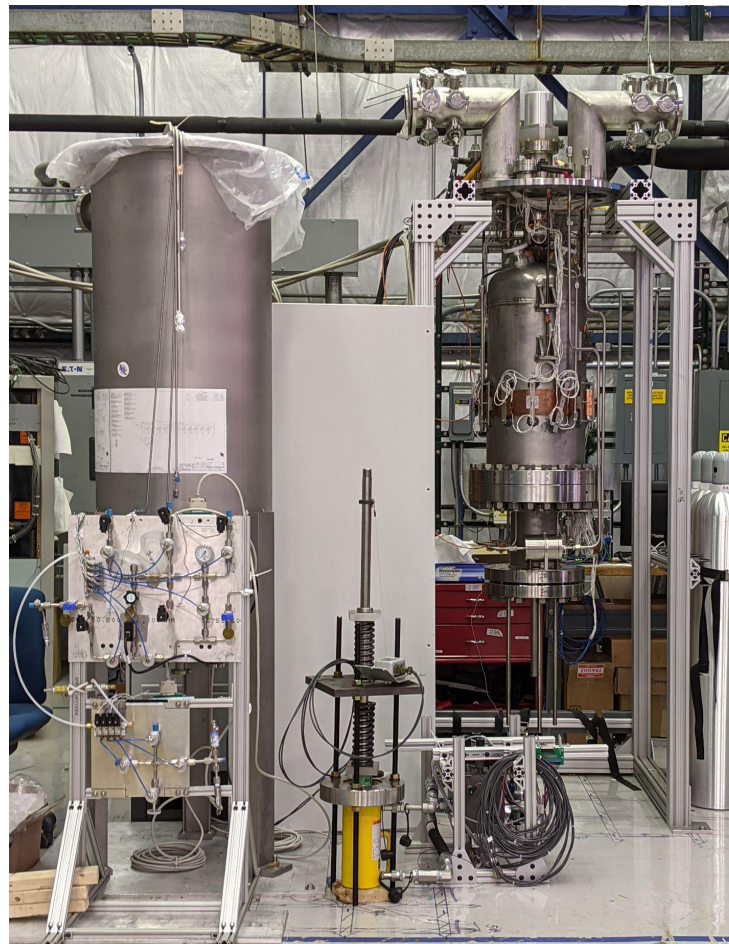
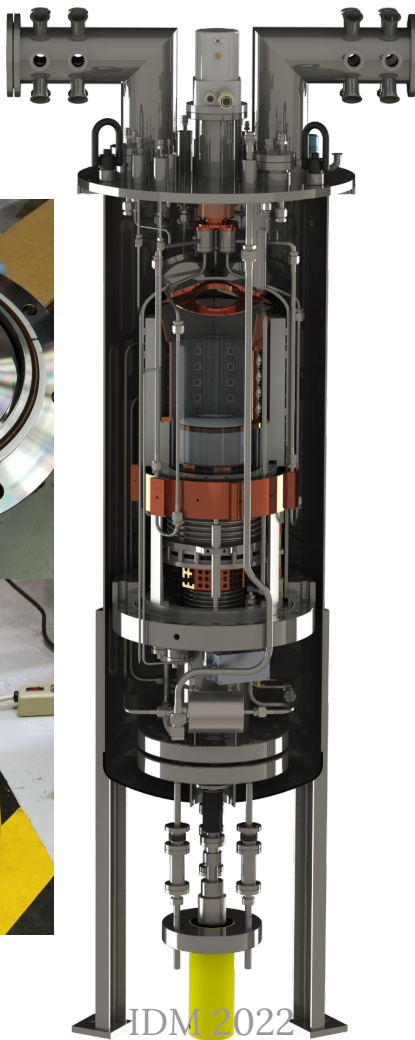
Inner vessel

Fused silica
vessels and
bellows

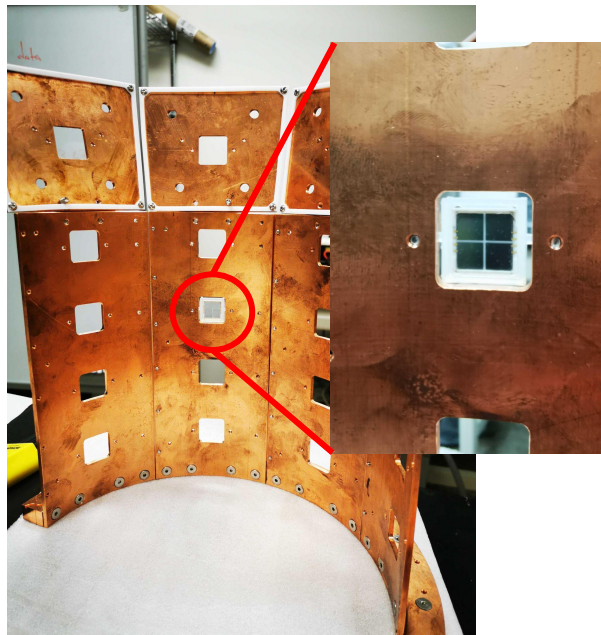


HDPE castle, Cu SiPM holders,
and piezo transducers

Outer vessel

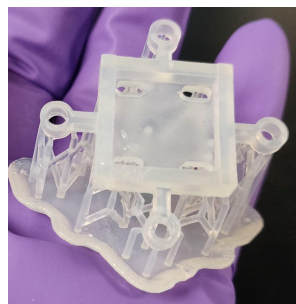


SiPMs, piezos, and cameras

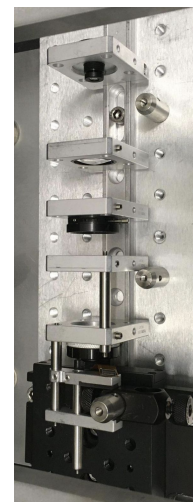
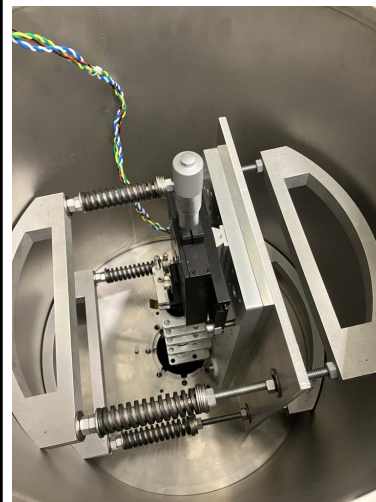


SNOLAB: FBK (25-30% QE)
Fermilab: Hamamatsu (20-25% QE)

Custom PZT piezo
transducers



Camera and lens setup



Backgrounds

α -induced

Surface Po, bulk ^{222}Rn decays
[purification/cleaning]

(α, n) in detector materials
& LCF_4 [material selection,
leaching]

β -induced

^{39}Ar [stray SiPM signals]

n -induced

Fast n [shielding/ LCF_4
veto, multiple scatters]

ν -induced

^8B CE ν NS [irreducible]

γ -induced

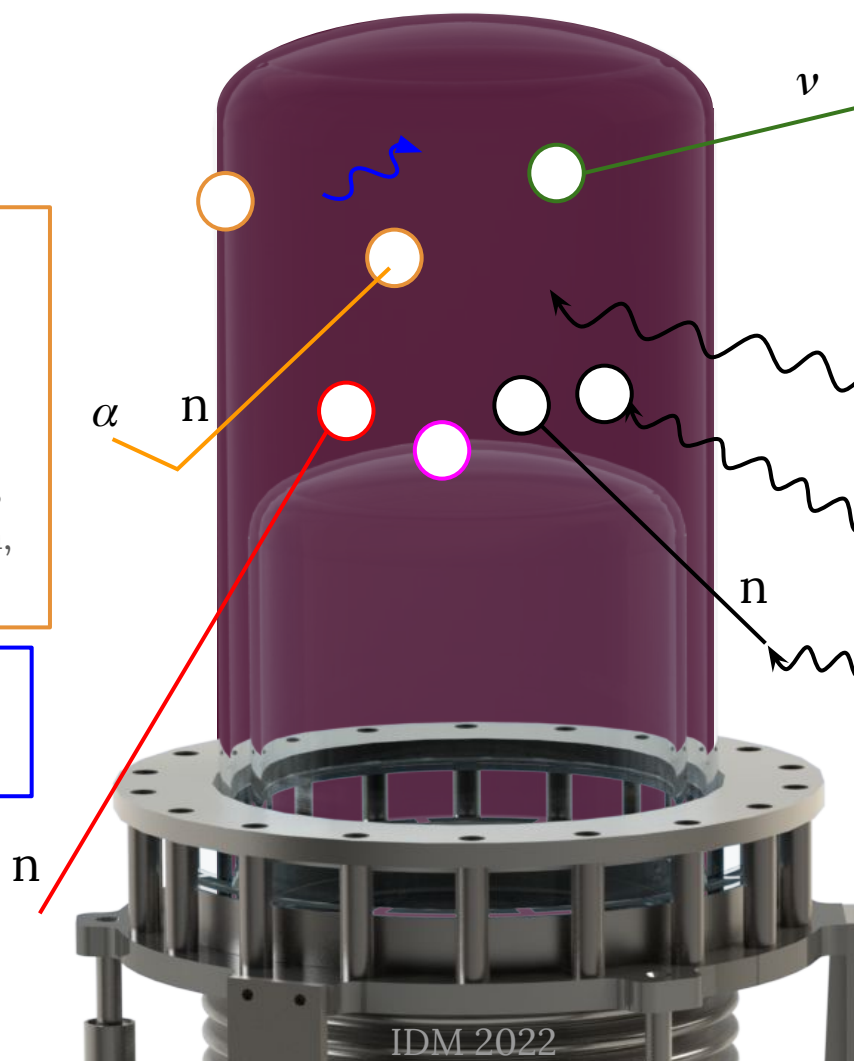
Compton [stray SiPM
signals, material
selection/shielding]

Thomson (^{40}K , ^{208}Tl)
[shielding]

(γ, n) reactions in water
tank [shielding]

Pathological

Surface roughness,
thermal stability



Initial calibration goals

1) Stable operation

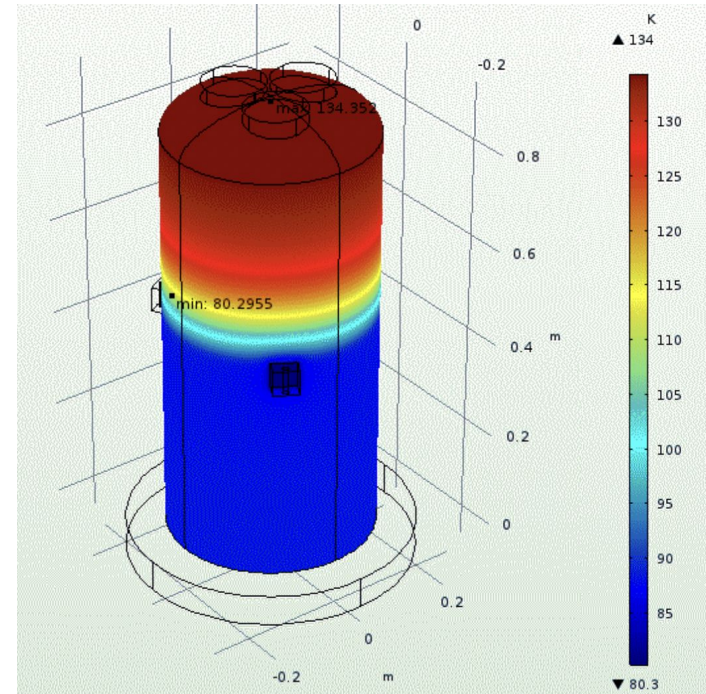
- Thermal model verification
- Homogeneous response across sensitive volume
- Manageable wall nucleation rate

2) Gamma calibration

- Confirming no e.r. nucleation at keV-scale
- Probe ultimate thermodynamic limit
- Investigate e.r. rejection below at/below 100 eV

3) Nuclear recoil calibration

- Photoneutron (keV)
- Thomson scattering (sub-keV)
- Tagged neutron capture (<100 eV)



PICO-60 $\mathcal{O}(0.02/\text{liveday}/\text{cm}^2)$

Initial calibration goals

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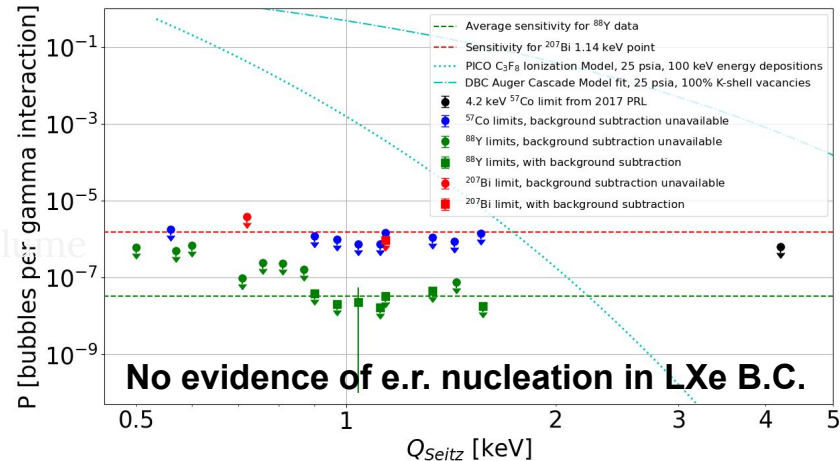
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- Tagged neutron capture (<100 eV)



Thermodynamic limit
(spontaneous nucleation)

Ar : ~40 eV

Xe: ~75 eV

Initial calibration goals

1) Stable operation

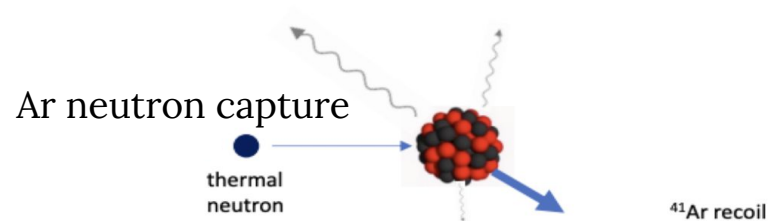
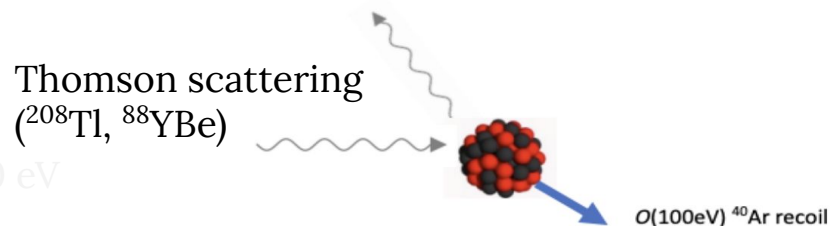
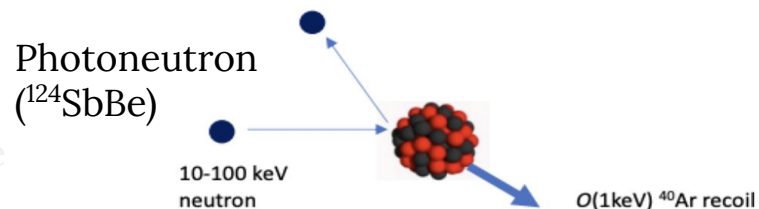
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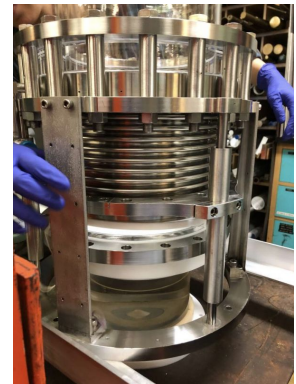
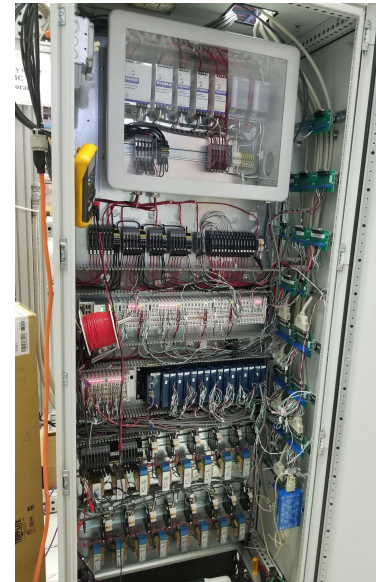
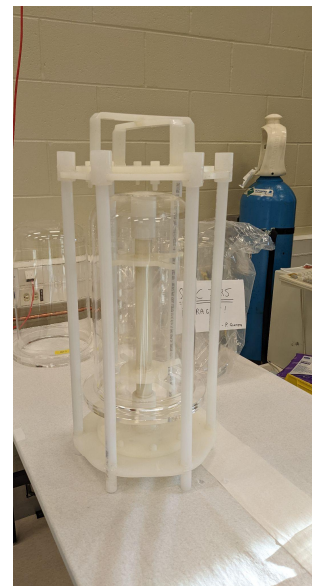
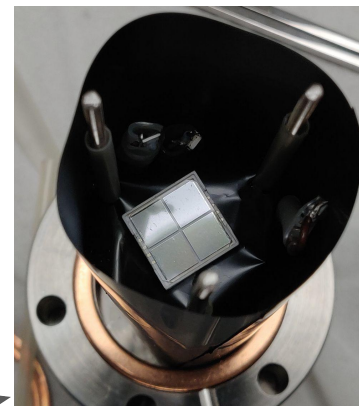
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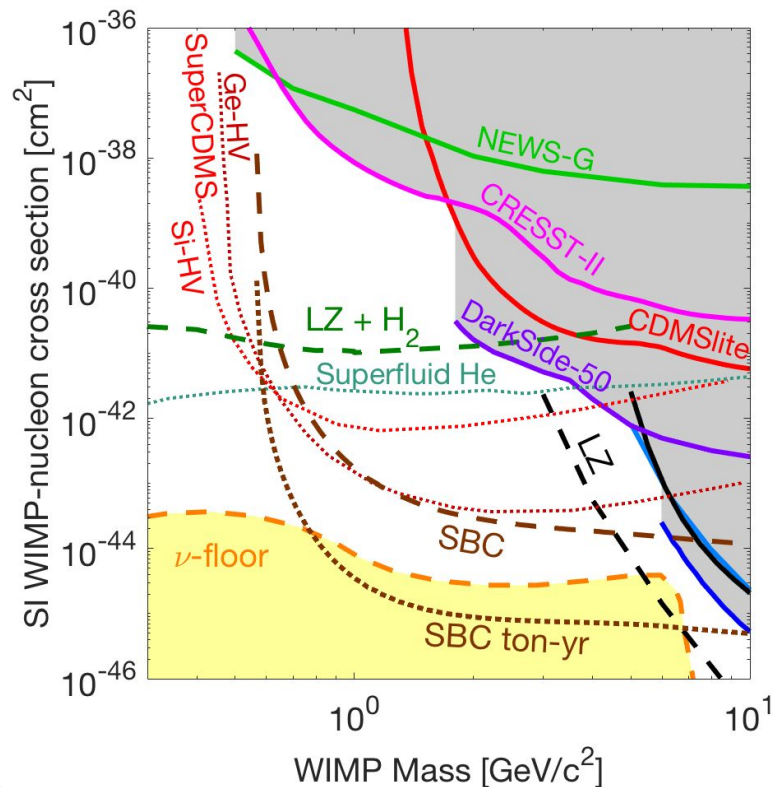


Current status

- Component testing
 - Cryogenic seal testing and inner vessel test assembly
 - SiPM cold tests/ DAQ development
 - Camera & optics development
 - CF_4 scintillation studies
 - Background studies
- Fermilab detector
 - Pressure vessel and vacuum jacket construction
 - Commissioning hydraulic system and slow controls
- SNOLAB detector
 - Space allocated underground
 - Procurement of long-lead items



Physics potential



- n.r. region of interest: 0.1 keV - 10 keV
- SNOLAB chamber, 10 kg-year can reach $10^{-43} \text{ cm}^2 @ 1 \text{ GeV}/c^2$
- If we achieve calibration goals and have long-term stable operation at SNOLAB, ton-year can reach boundary of the Ar ν -fog

For CE ν NS potential: PRD **103**, L091301 (2021)

Conclusion

- Noble liquid bubble chamber will probe GeV-scale dark matter
- Commissioning at Fermilab through 2022, SNOLAB construction in 2023
- SBC collaboration consists of institutions from Canada, USA, Mexico (new collaborators always welcome)
- Check out our Snowmass white paper (in preparation)



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**UNIVERSITY OF
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Fermilab



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