

DARK MATTER AT COSMIC DAWN

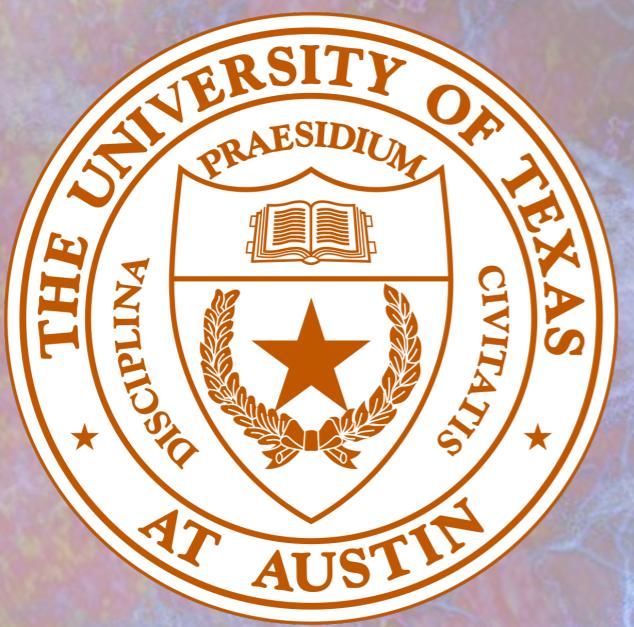
JULIAN B. MUÑOZ

CENTER FOR

ASTROPHYSICS

HARVARD & SMITHSONIAN

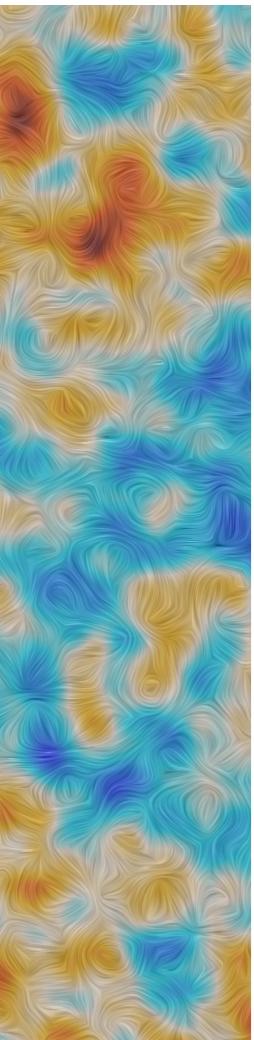
w/ Blas, Bohr, Cyr-Racine, Dvorkin, HERA Coll.,
Loeb, **Sabti**, Vogelsberger, Zavala



How does cosmology measure things?

Image: ESA

CMB



$$z \approx 10^3$$

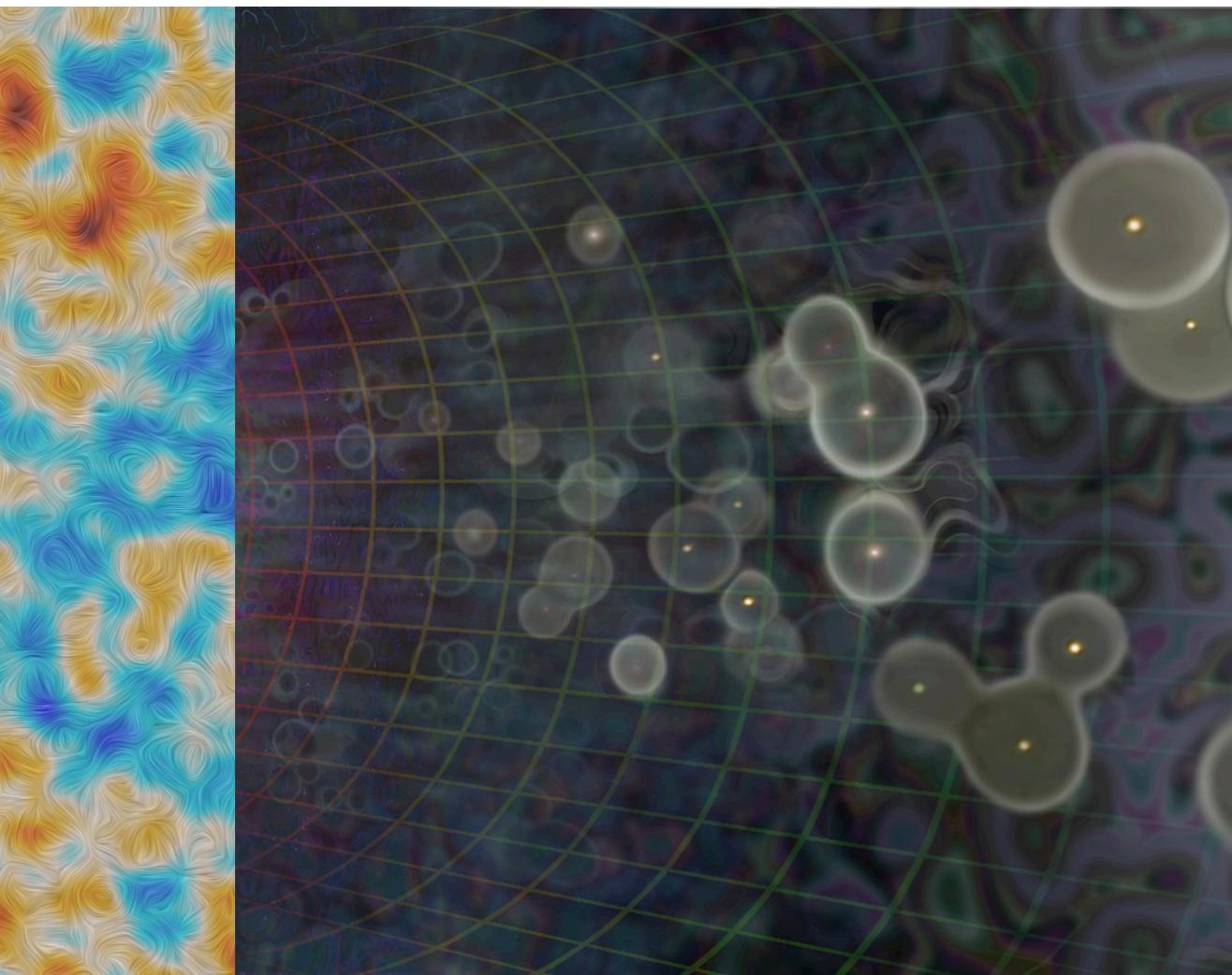
$\xrightarrow{4 \cdot 10^5}$ cosmic time [yr]

How does cosmology measure things?

Image: NASA/
CXC/M.WEISS

CMB

Cosmic Dawn



$z \approx 10^3$

$z \approx 30$

$4 \cdot 10^5$

10^8

cosmic time [yr]

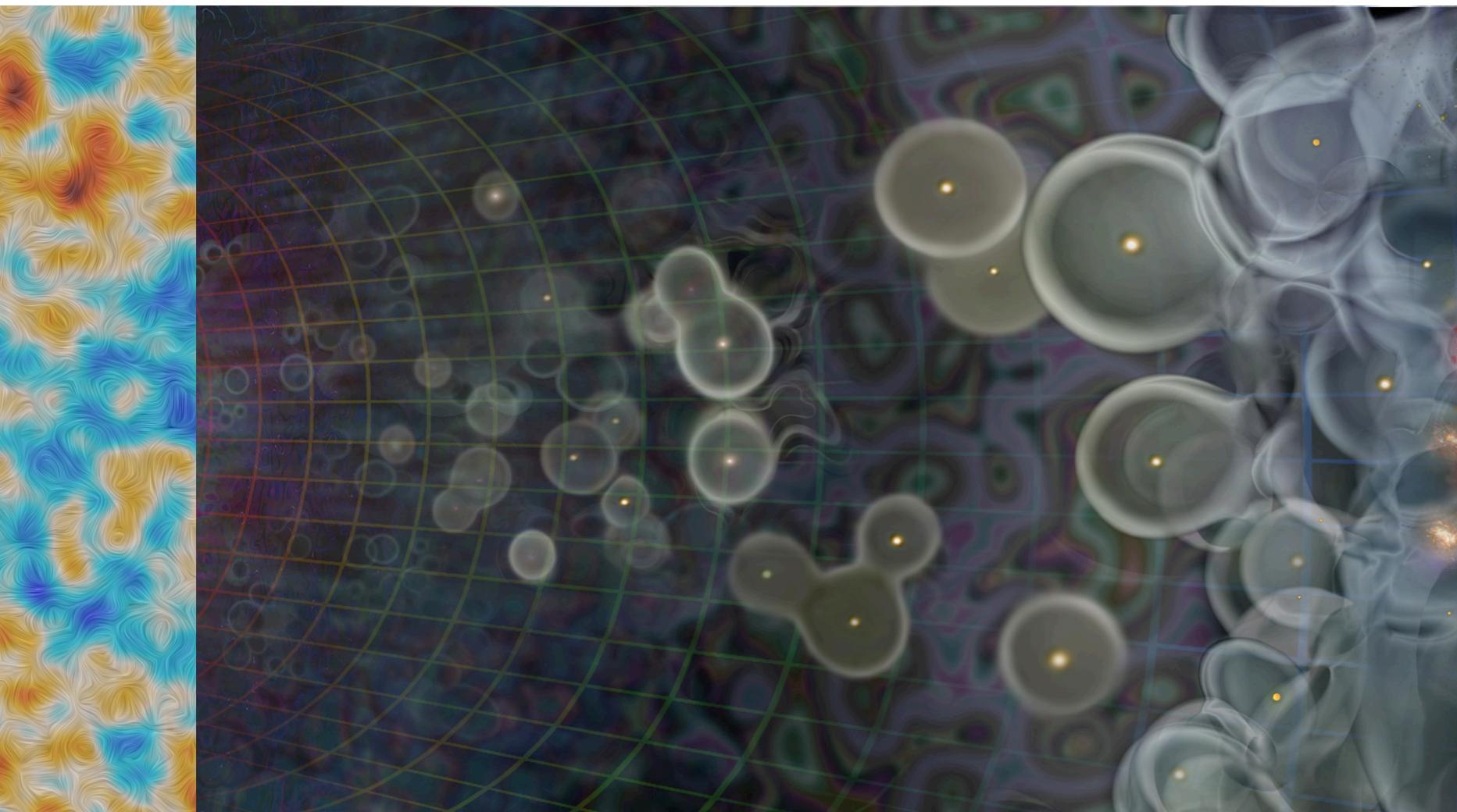
How does cosmology measure things?

Image: NASA/
CXC/M.WEISS

CMB

Cosmic Dawn

Reionization



$z \approx 10^3$

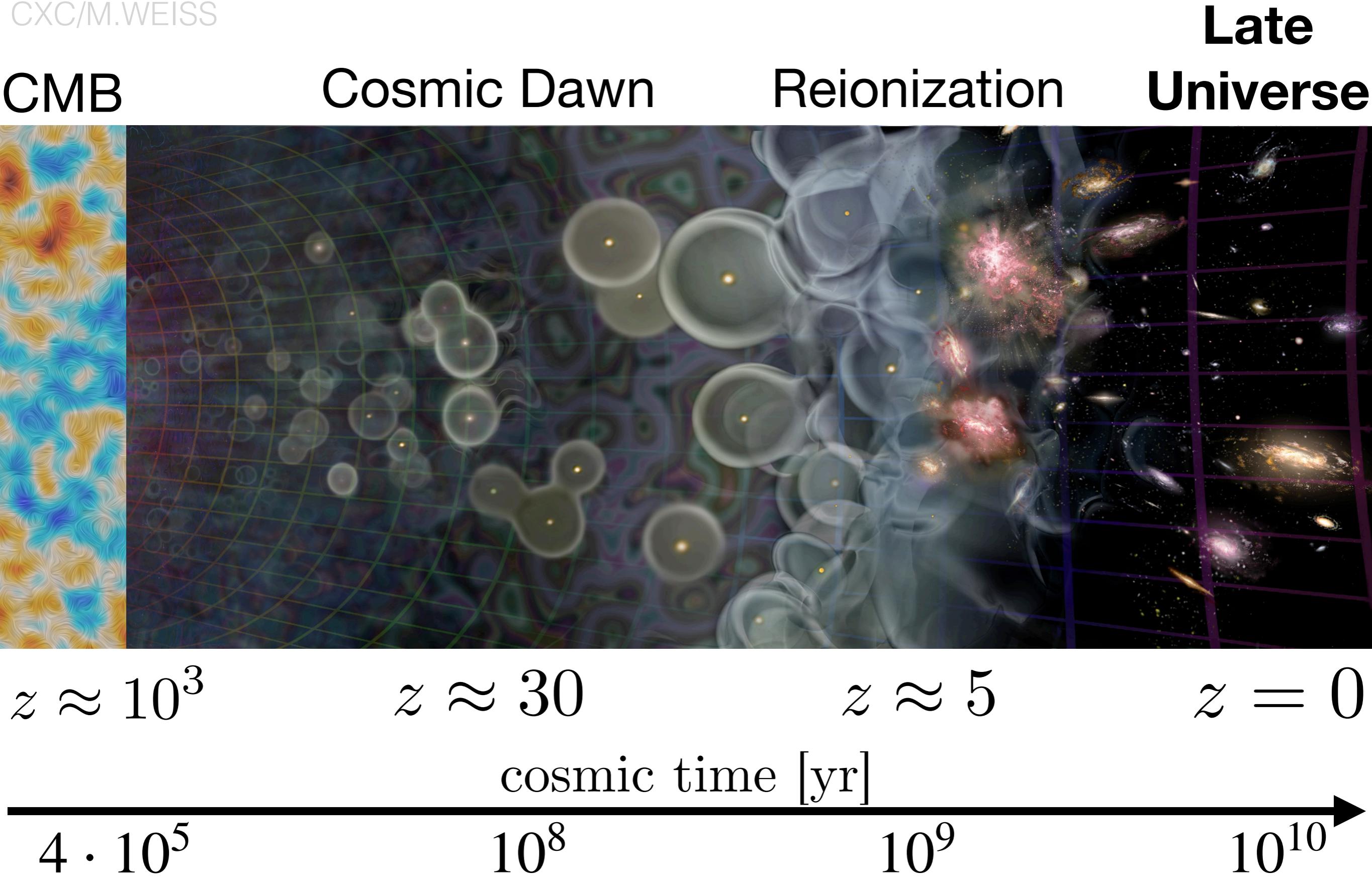
$z \approx 30$

$z \approx 5$

$\xrightarrow{4 \cdot 10^5 \quad 10^8 \quad 10^9}$ cosmic time [yr]

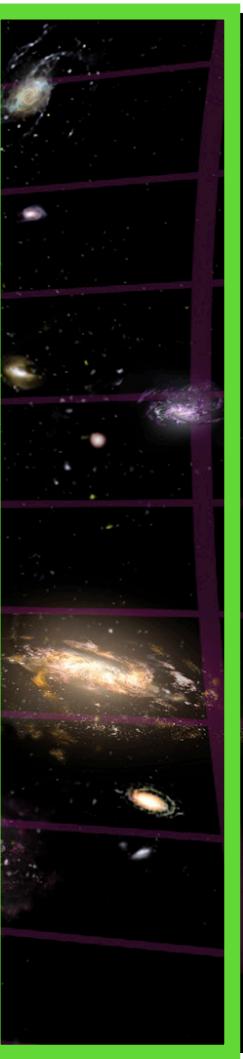
How does cosmology measure things?

Image: NASA/
CXC/M.WEISS



Late
Universe

CMB



$z \approx 10^3$

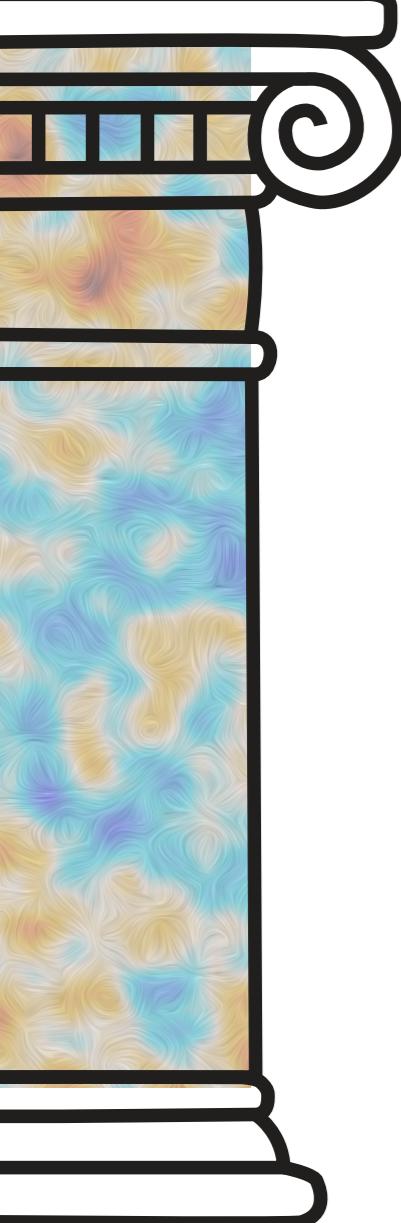
cosmic time [yr]

$4 \cdot 10^5$

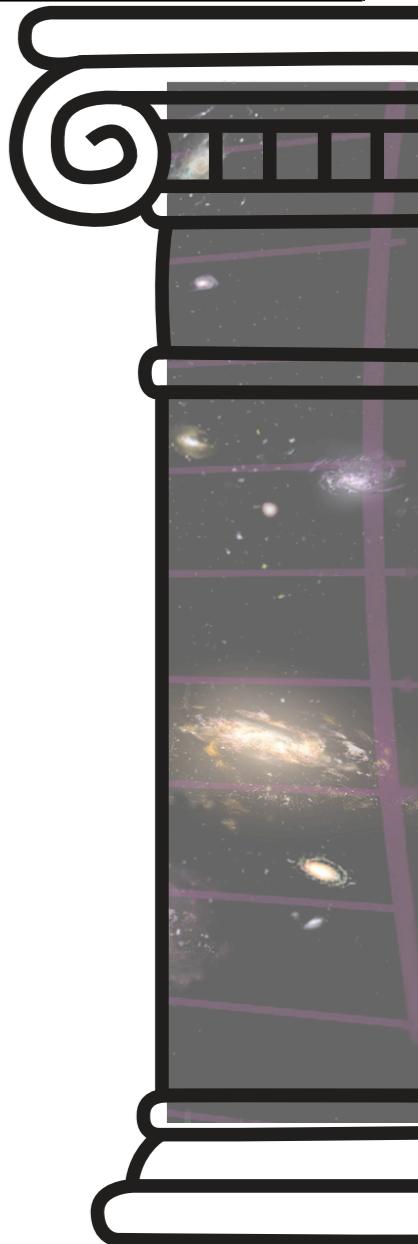
$z = 0$

10^{10}

Cosmology

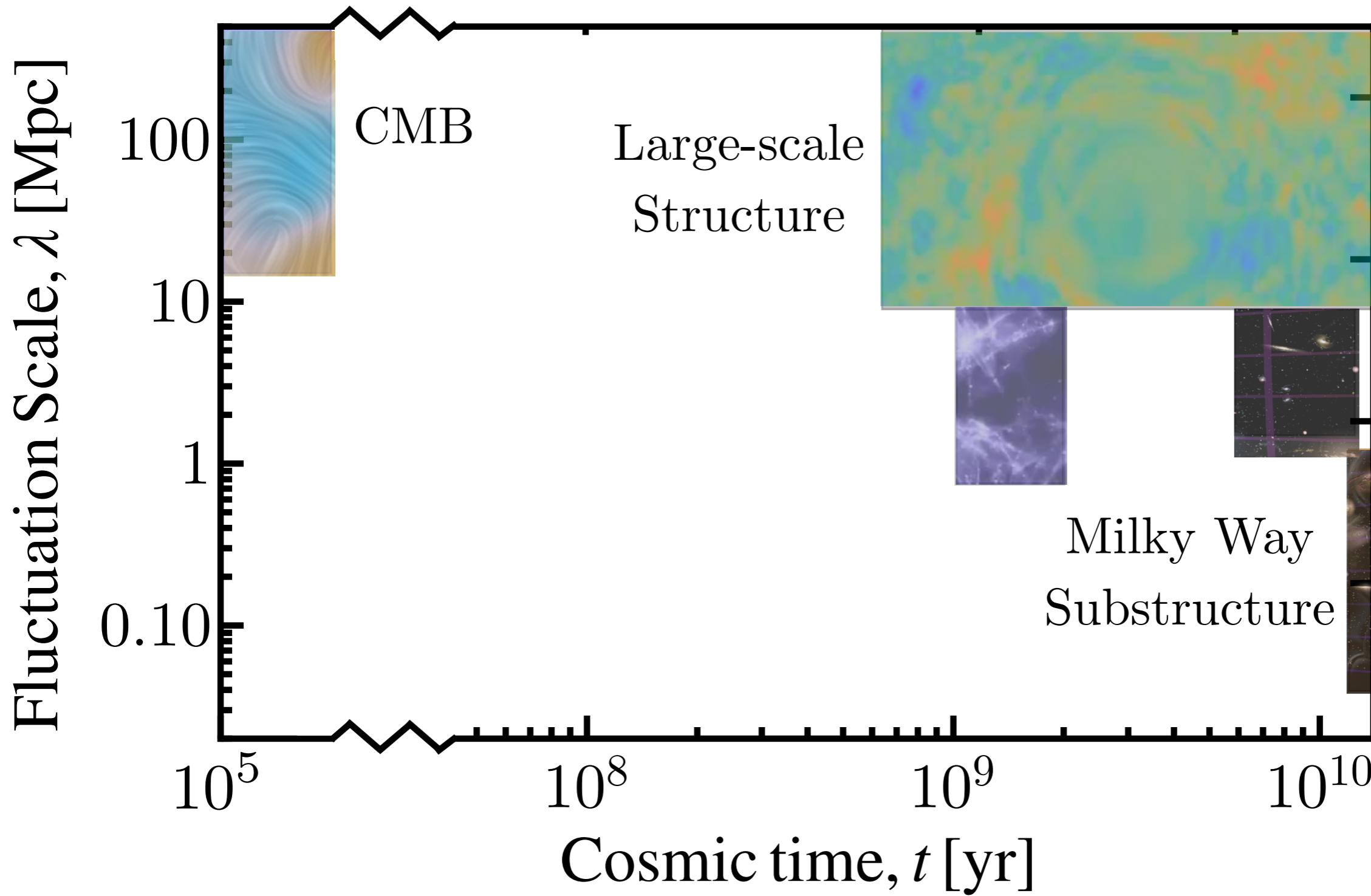


CMB

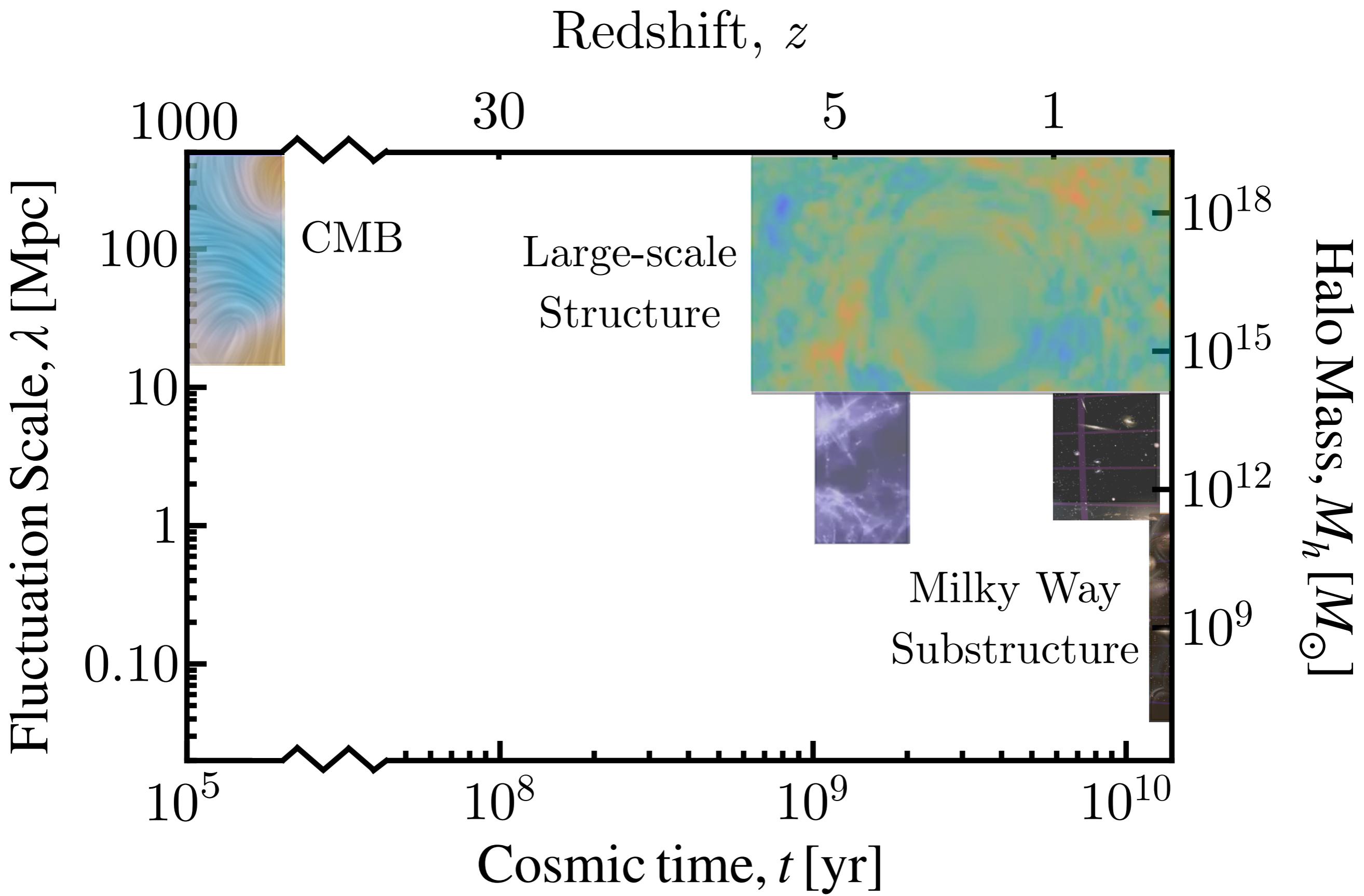


Local
Universe

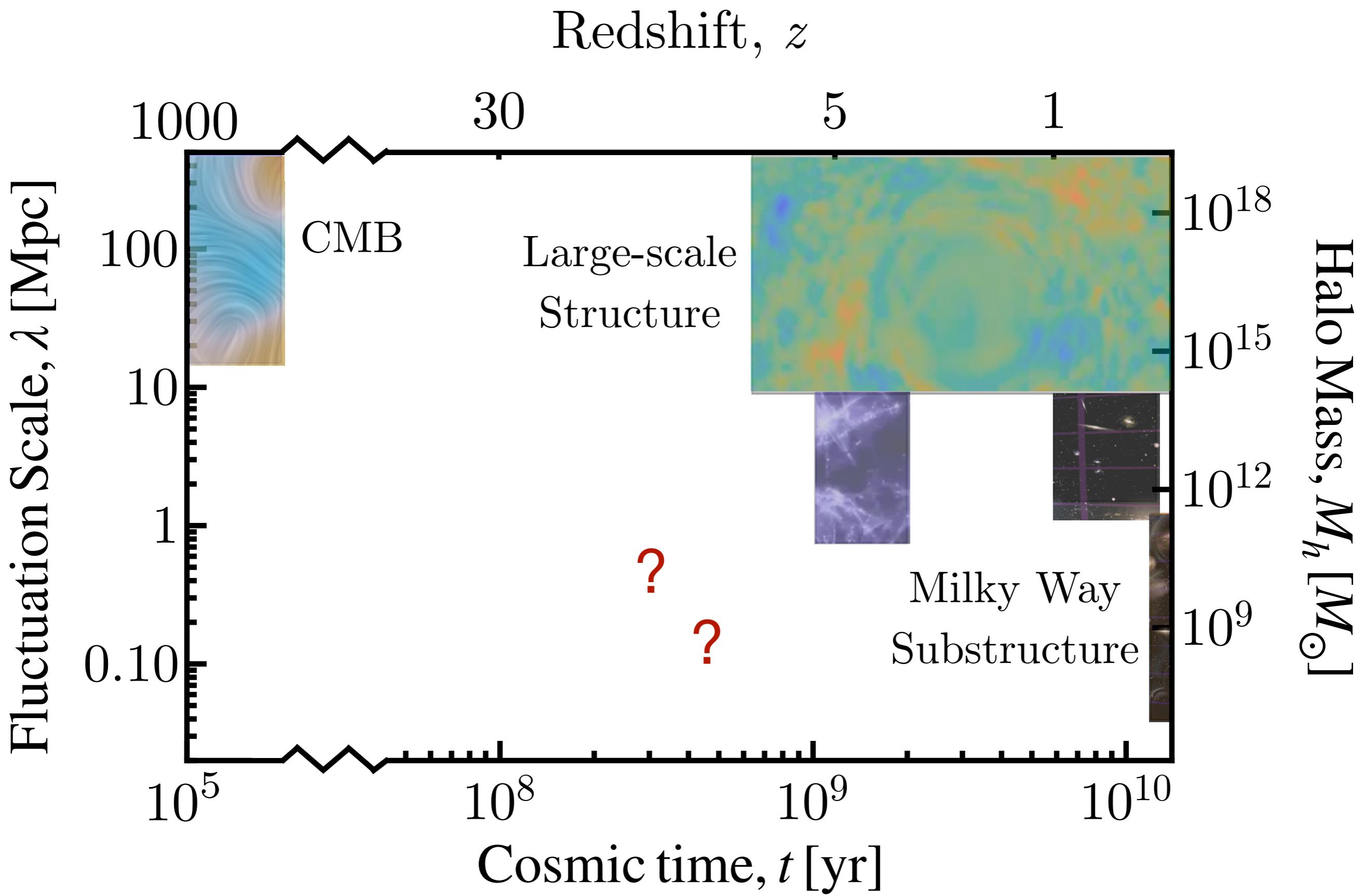
DM is cold ...



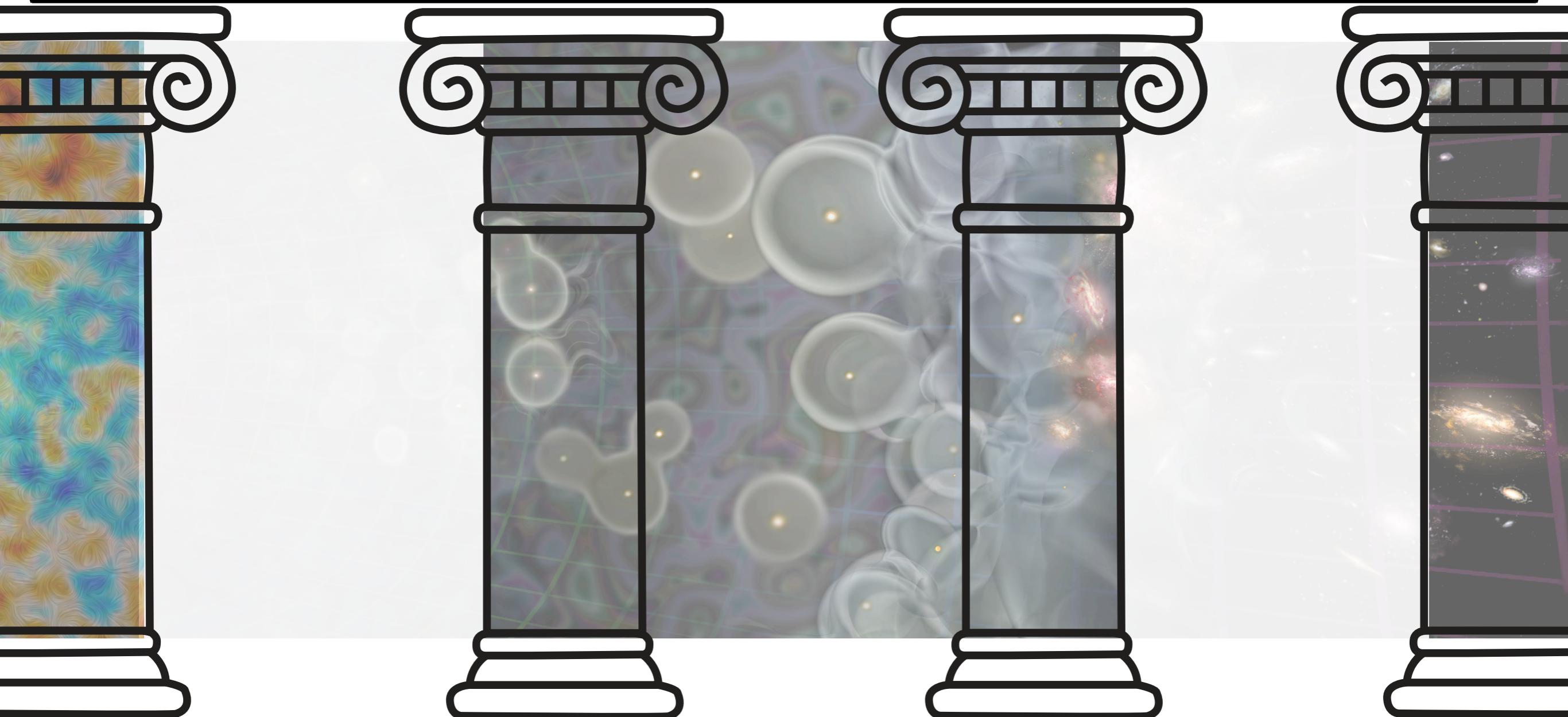
DM is cold ...



DM is cold as far as we can tell



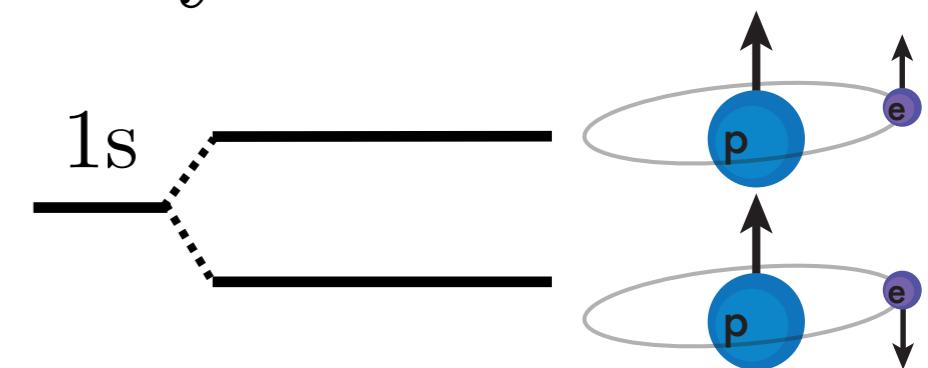
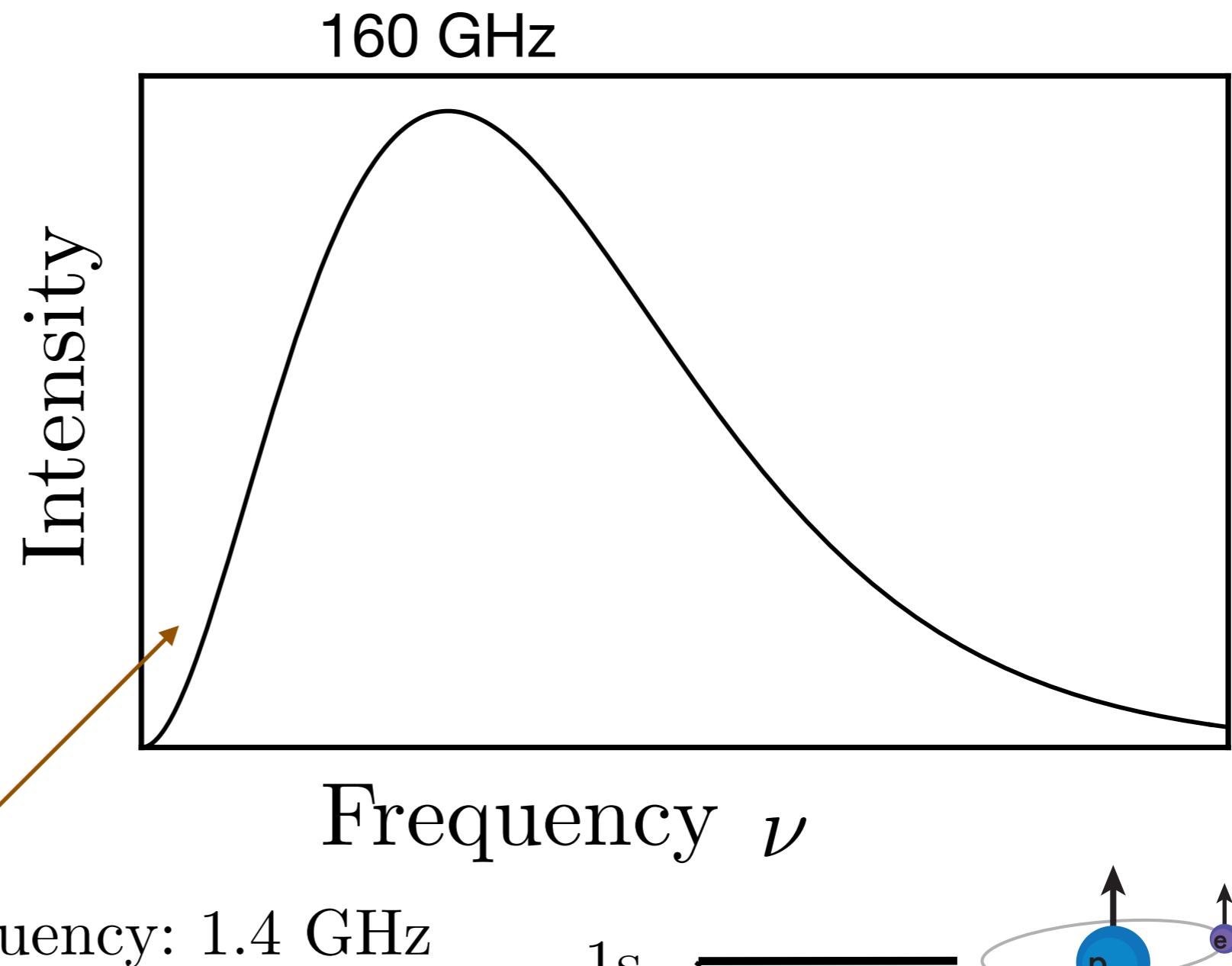
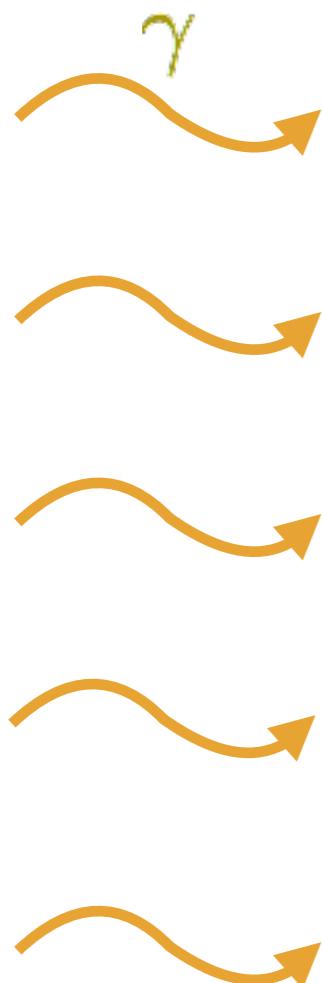
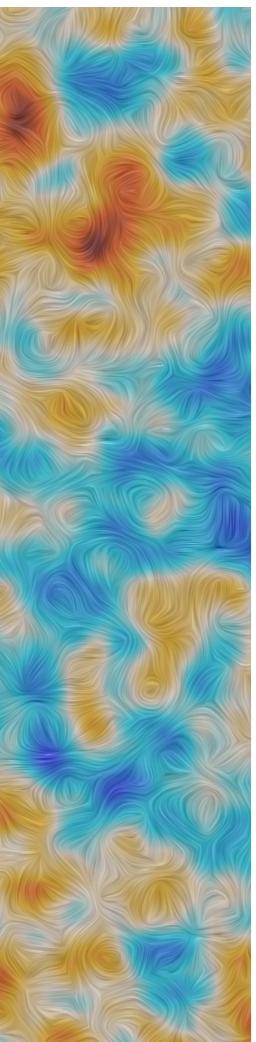
New pillars of cosmology



Cosmic Dawn and Reionization

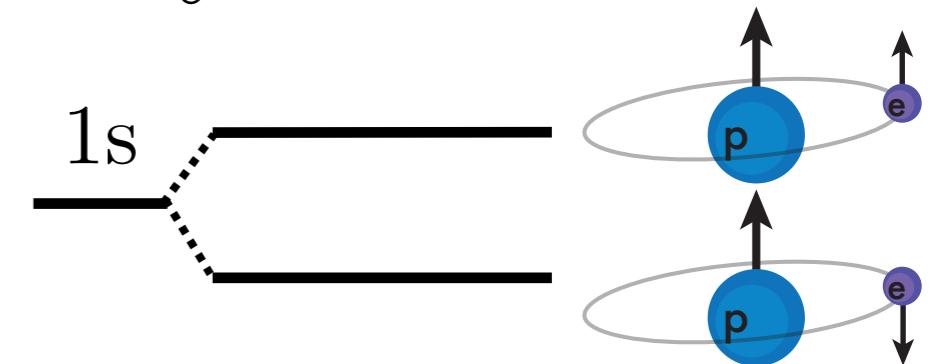
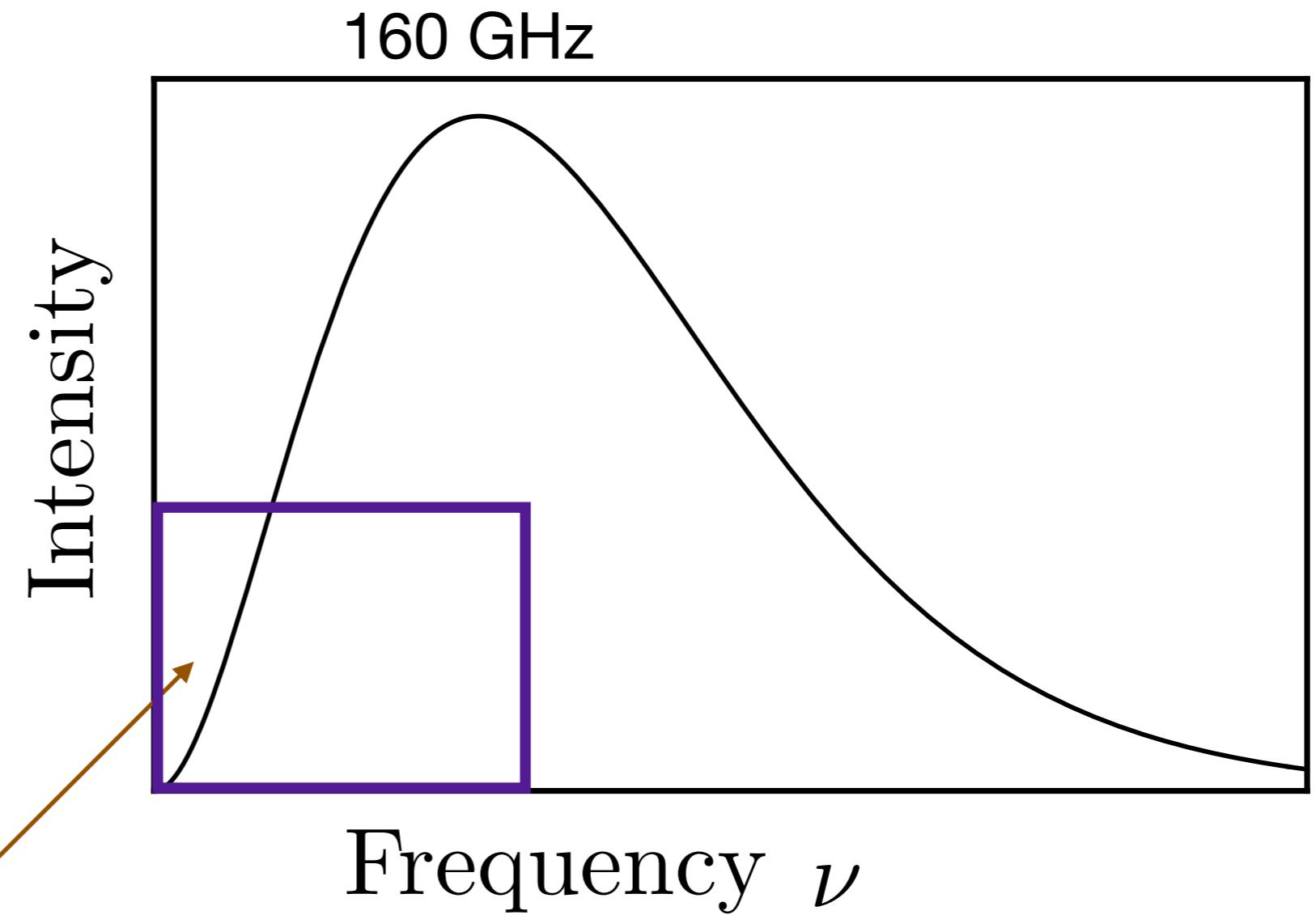
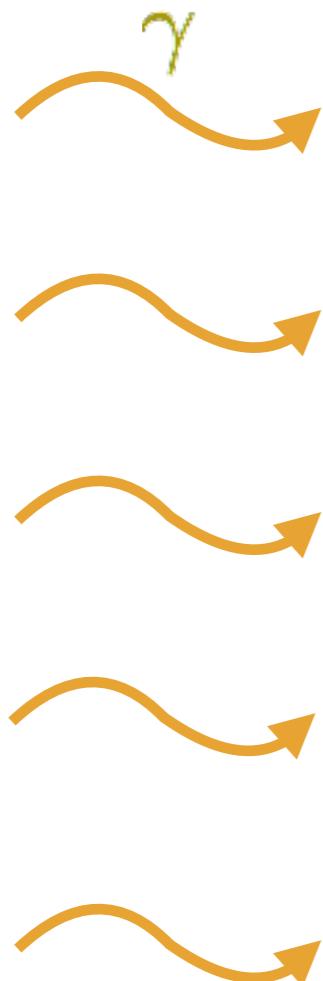
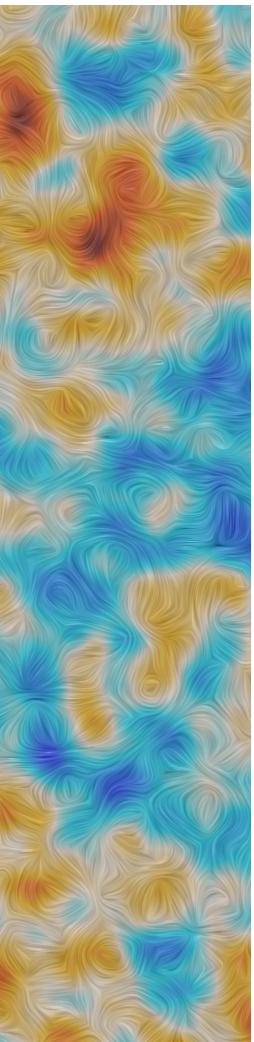
21-cm cosmology super-summarized

CMB



21-cm cosmology super-summarized

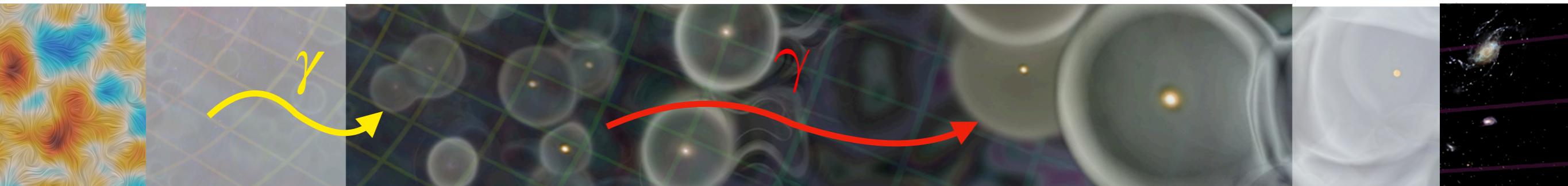
CMB



$z = 27$

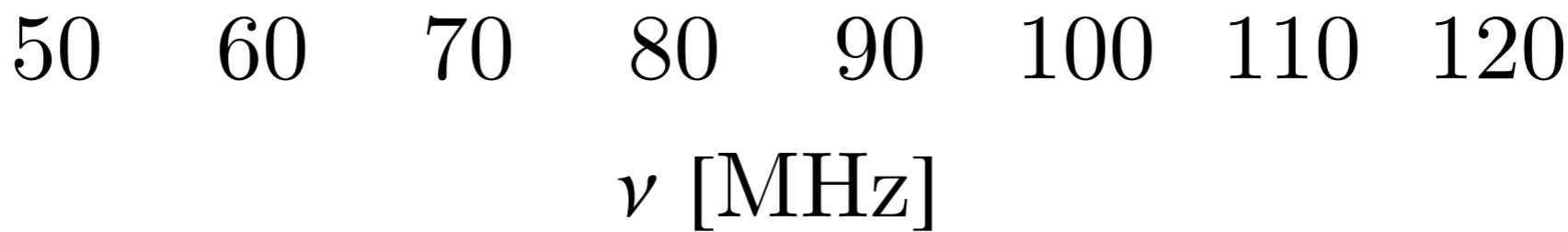
cosmic time
→

$z = 11$



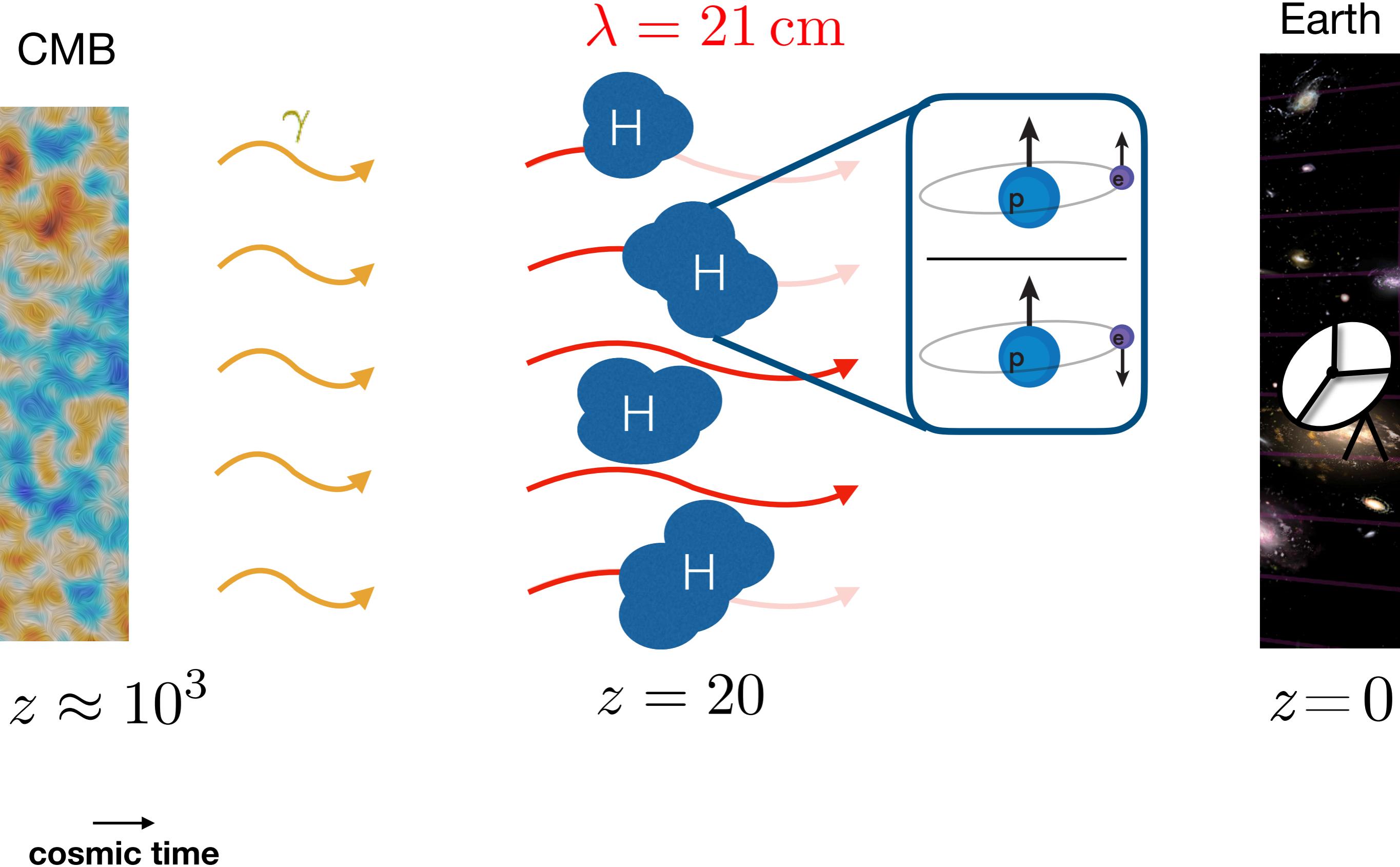
Intensity [unnormalized]

$$\nu = \frac{1.4 \text{ GHz}}{1 + z}$$



E.g., Pritchard & Loeb 2012

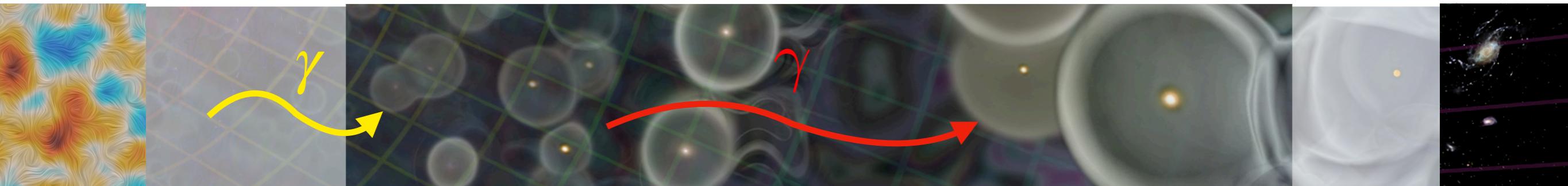
The basics of 21-cm: Absorption



$z = 27$

cosmic time

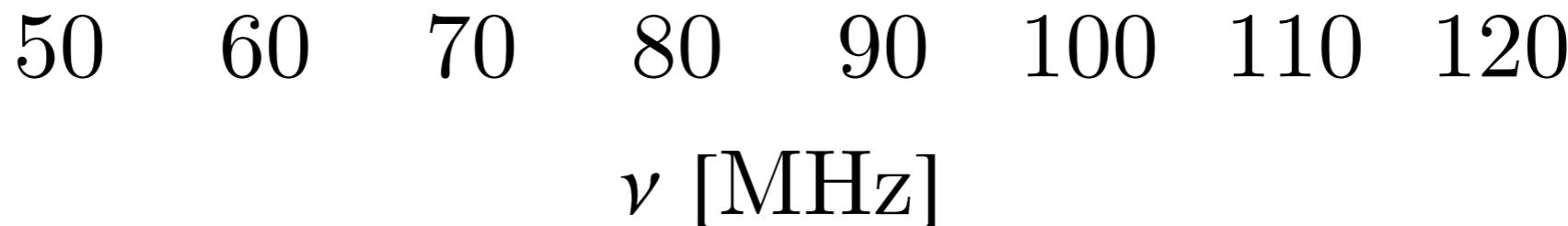
$z = 11$



Intensity [unnormalized]

$$\nu = \frac{1.4 \text{ GHz}}{1 + z}$$

Absorption

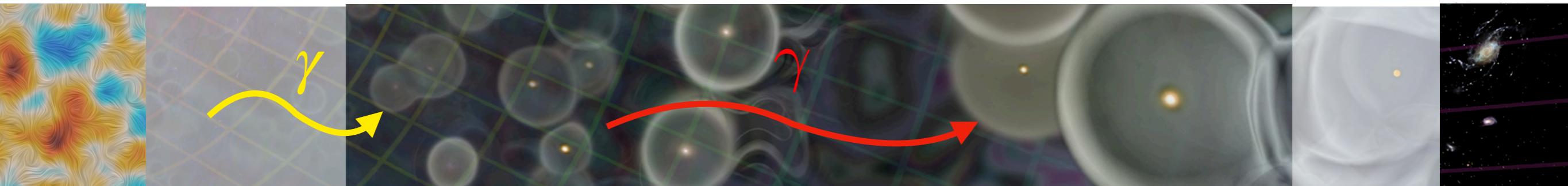


E.g., Pritchard & Loeb 2012

$z = 27$

cosmic time

$z = 11$

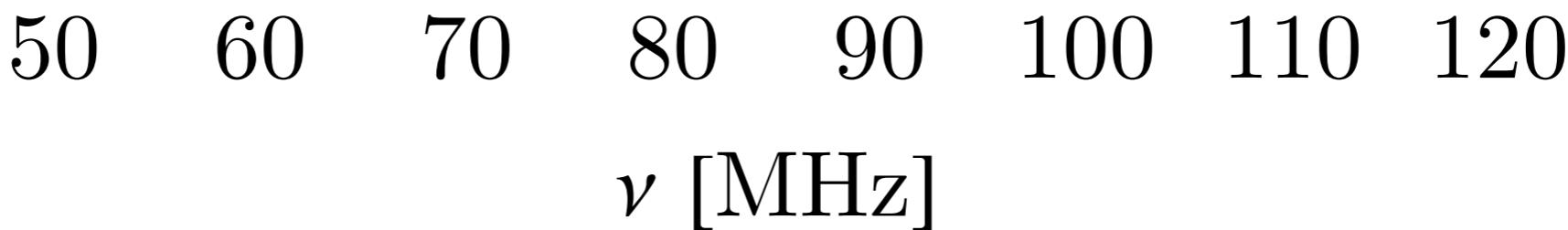


Intensity [unnormalized]

$$\nu = \frac{1.4 \text{ GHz}}{1 + z}$$

Emission

Absorption



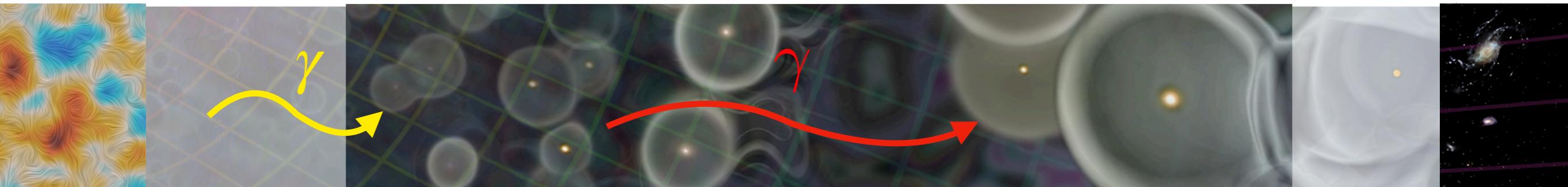
E.g., Pritchard & Loeb 2012



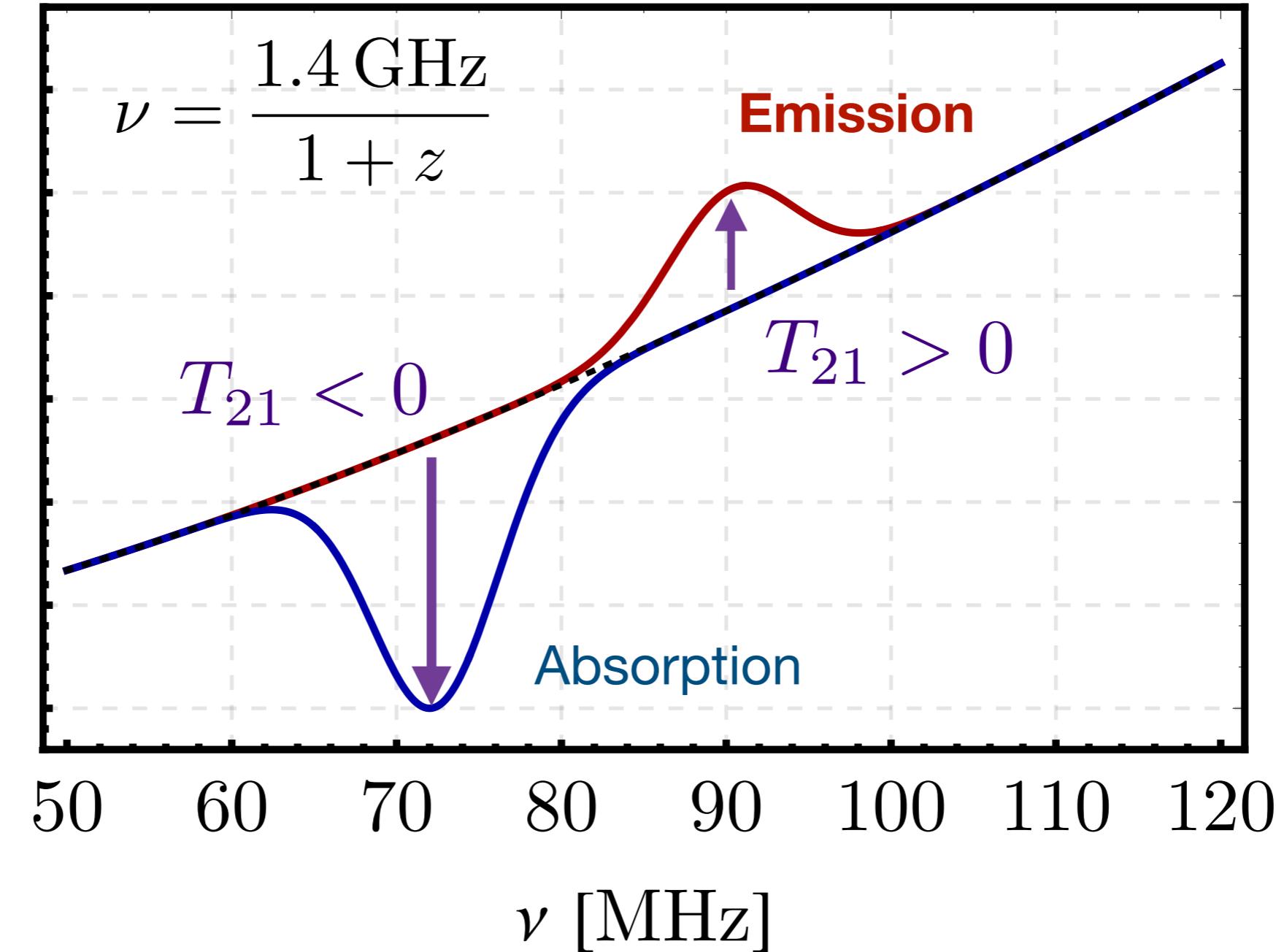
$z = 27$

cosmic time

$z = 11$

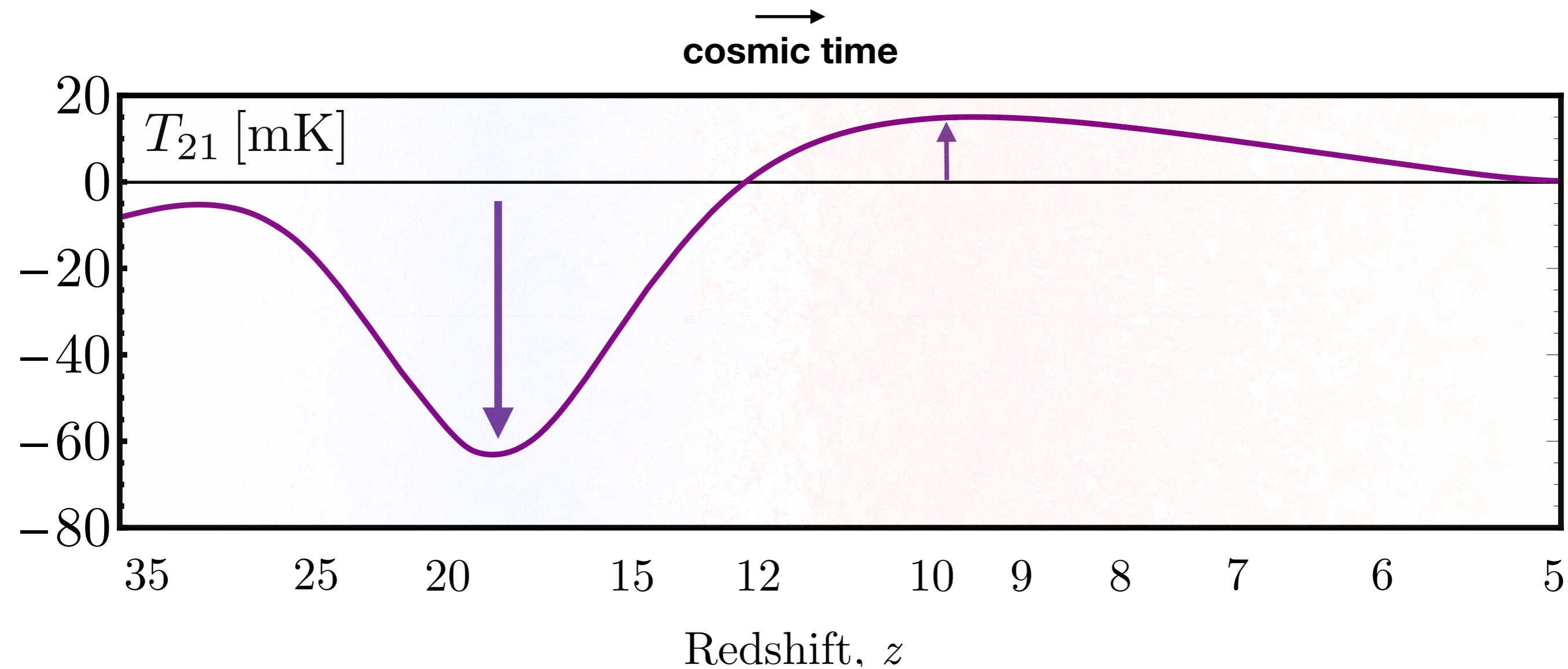


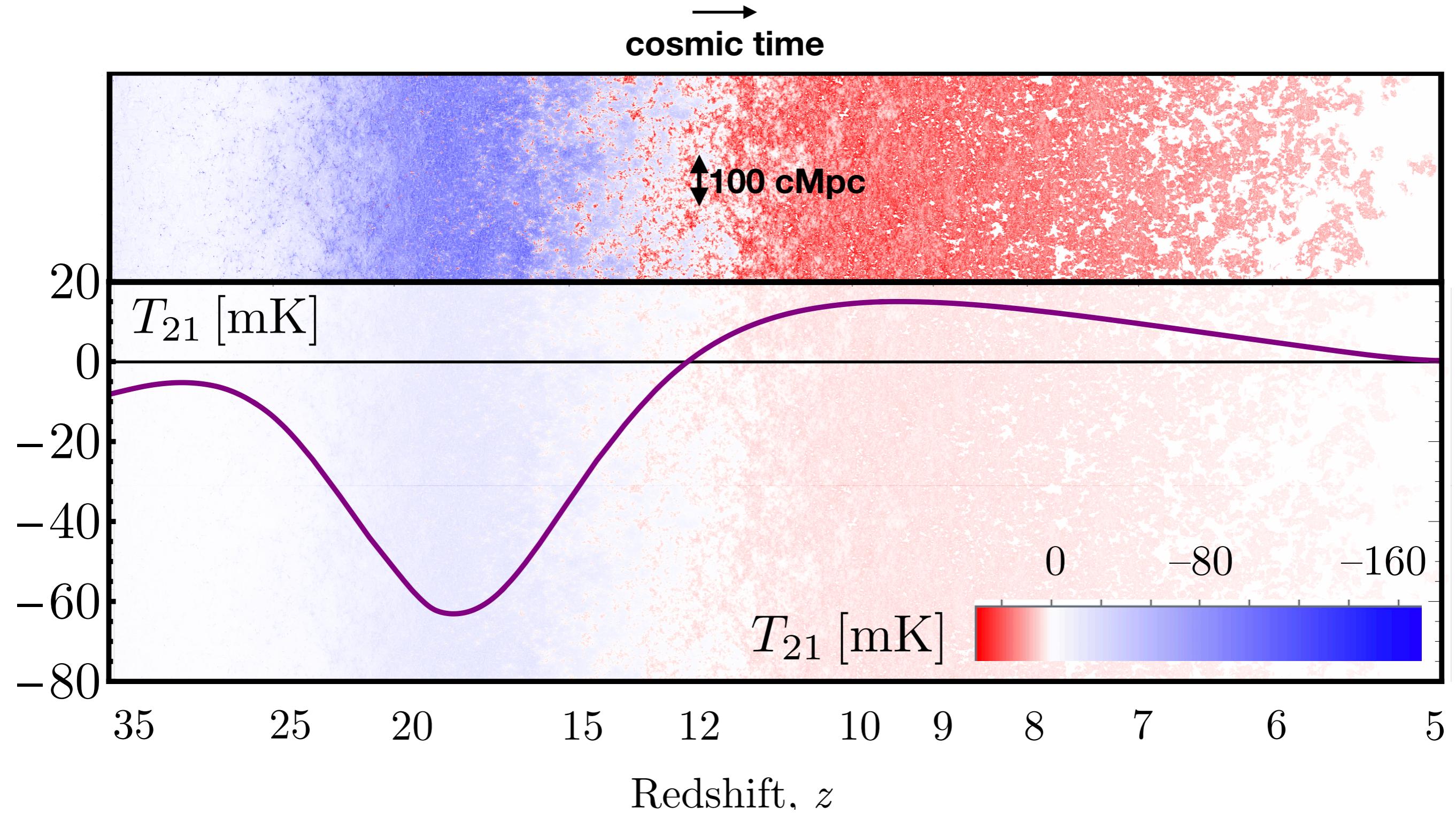
Intensity [unnormalized]

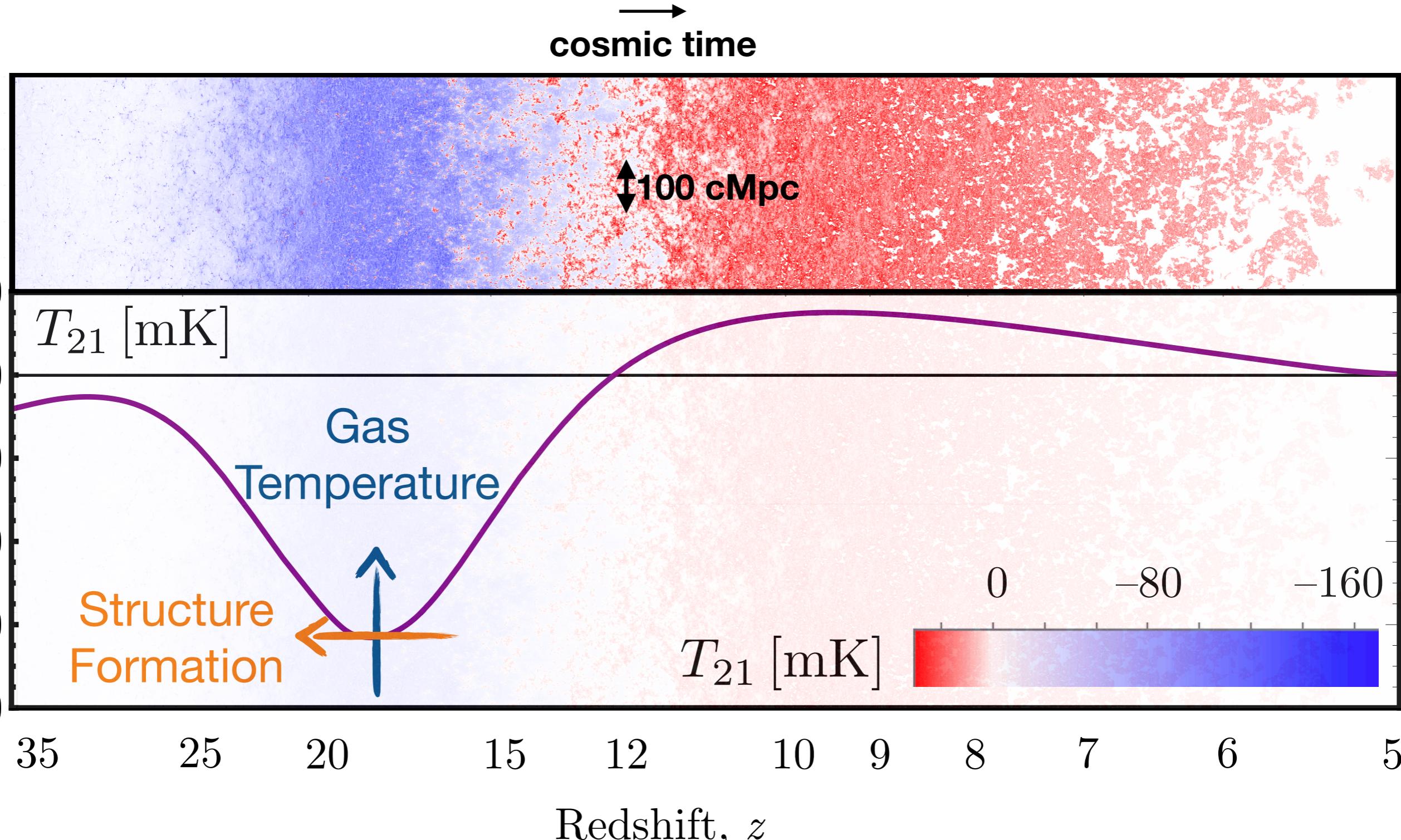


E.g., Pritchard & Loeb 2012

What would we see?

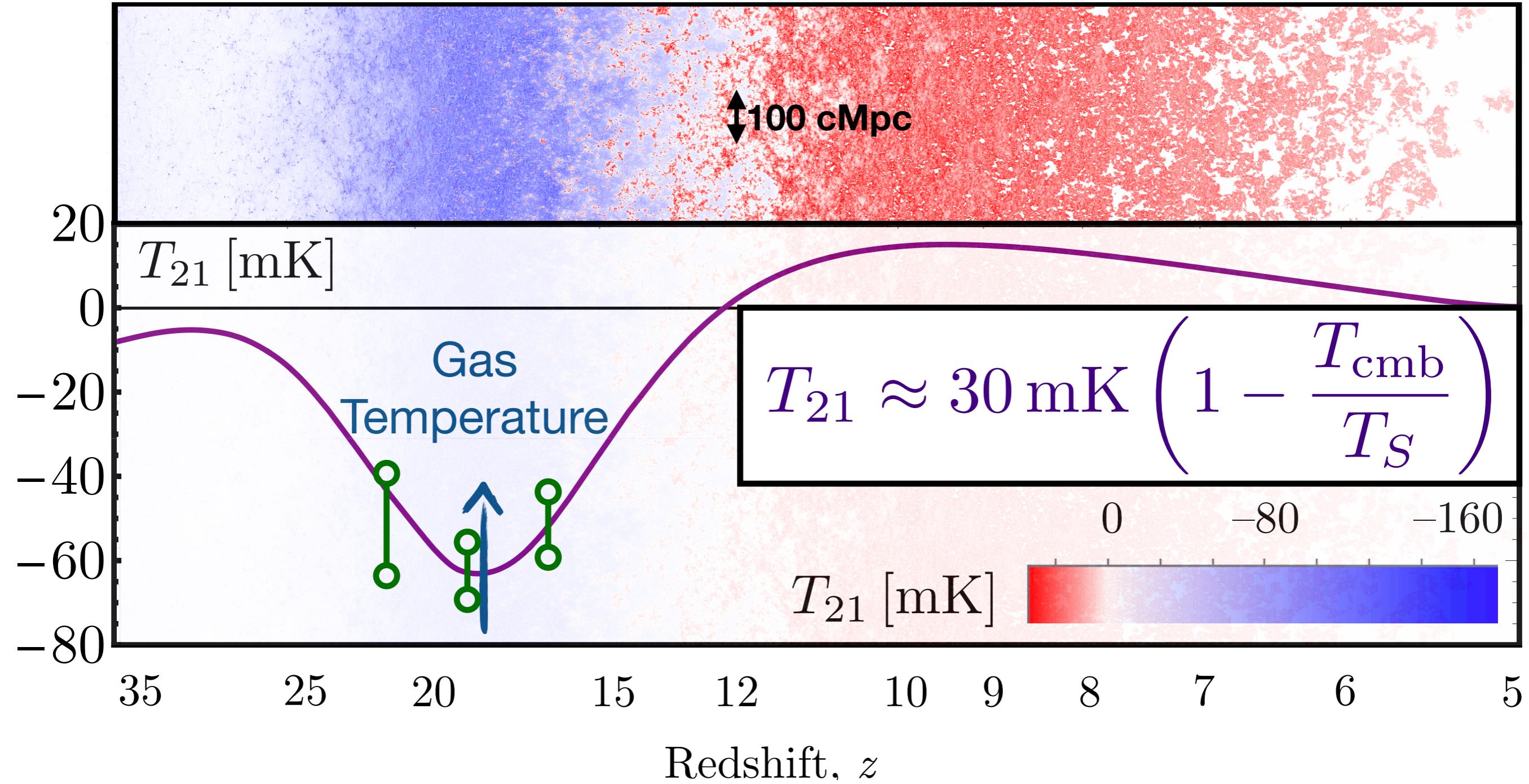






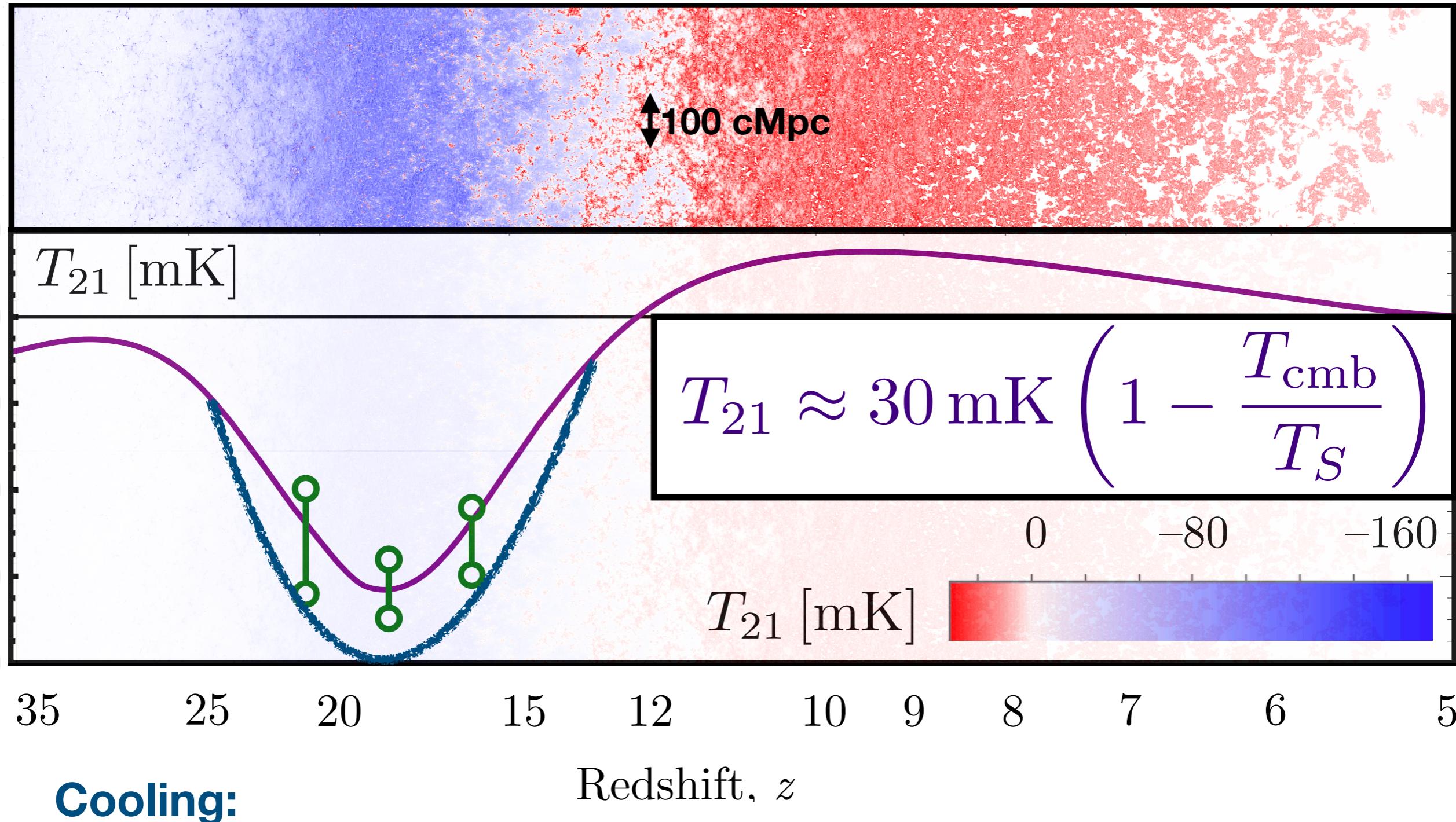
21-cm as a thermometer at high z

→
cosmic time



21-cm as a thermometer at high z

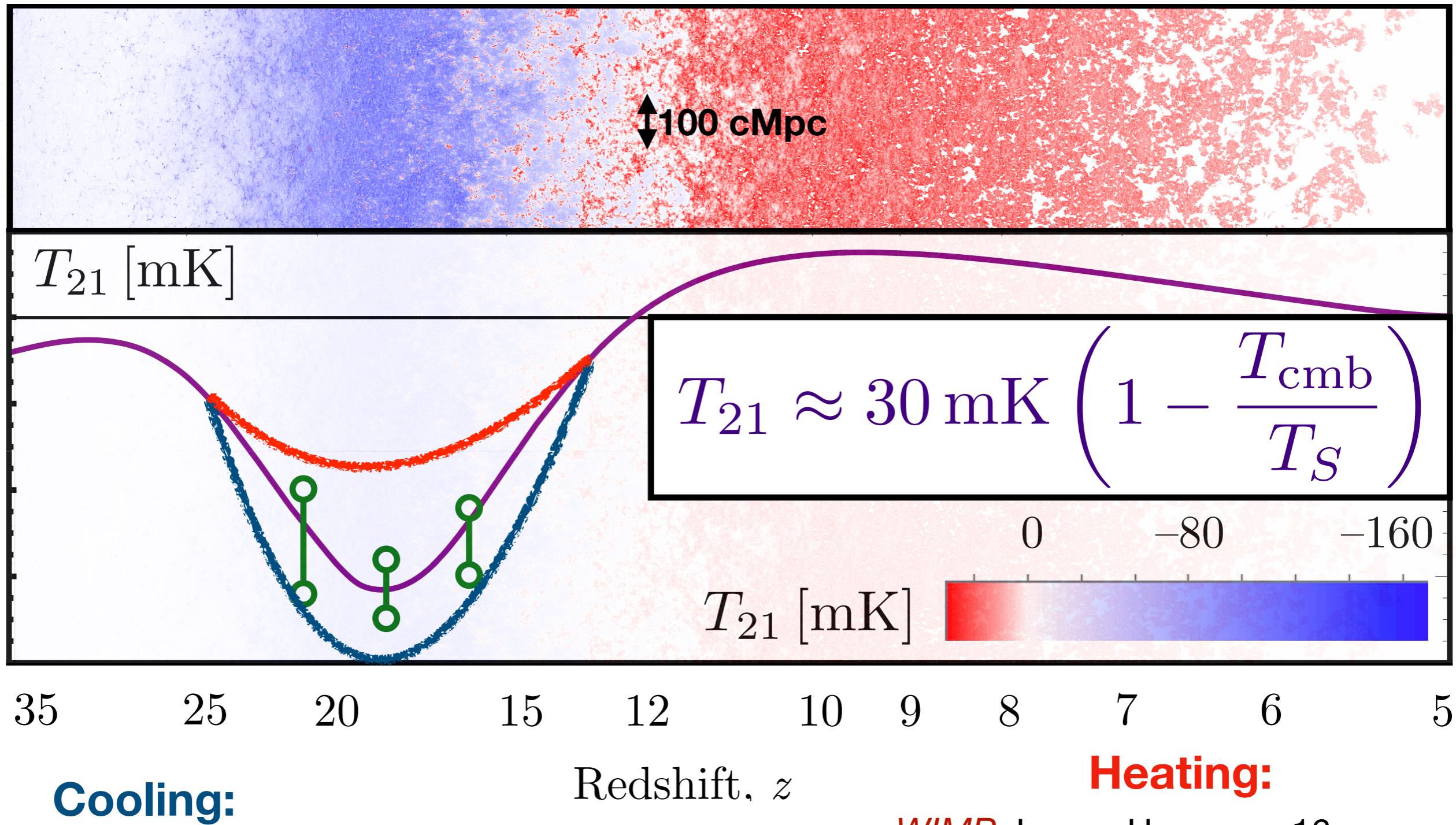
→
cosmic time



5th force: Tashiro+, **JBM**, Ali-Haïmoud & Kovetz 15, Barkana 18
mQ: **JBM** & Loeb, Barkana+, Berlin+ 18, Liu+ 19...
Radio: Ewall-Wice+, Fraser+, Pospelov+ 18, Caputo+21...

21-cm as a thermometer at high z

→
cosmic time



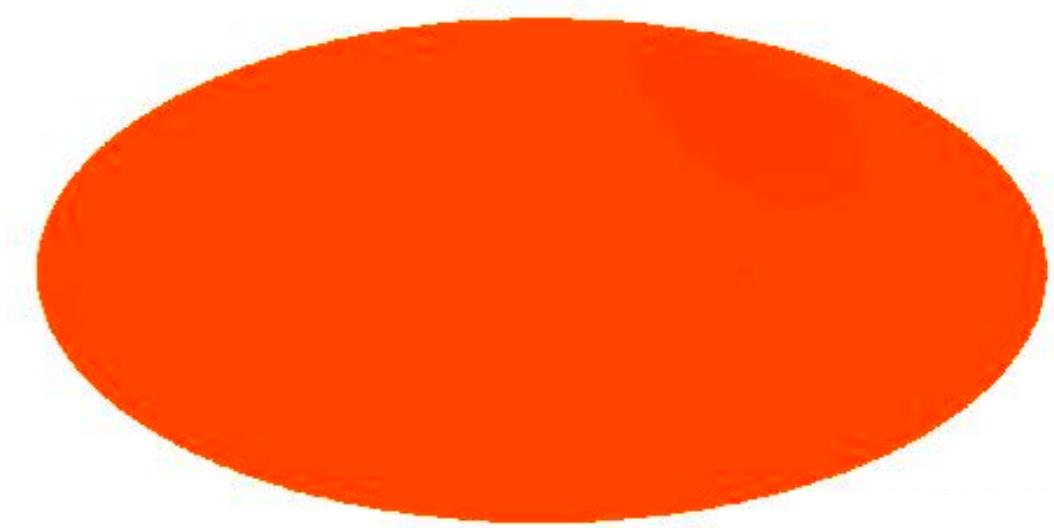
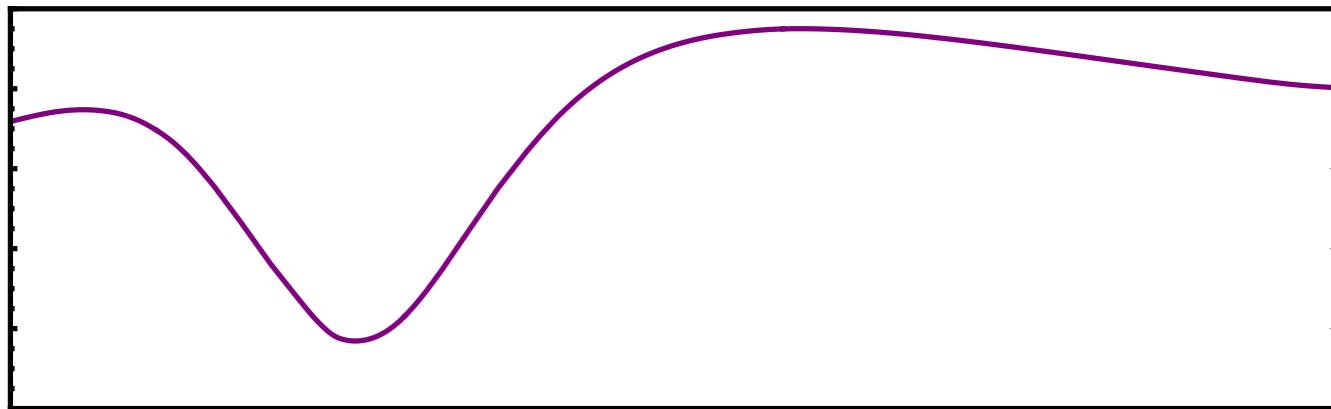
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 Radio: Ewall-Wice+, Fraser+, Pospelov+ 18, Caputo+21...

WIMP: Lopez-Honorez+ 16
 Liu & Slatyer, D'Amico+ 18
PBHs: Clark+ 18,
A' DM: Kovetz, Cholis, Kaplan 18

21-cm Global Signal
(EDGES, SARAS, PRIZM, MIST,...)

=

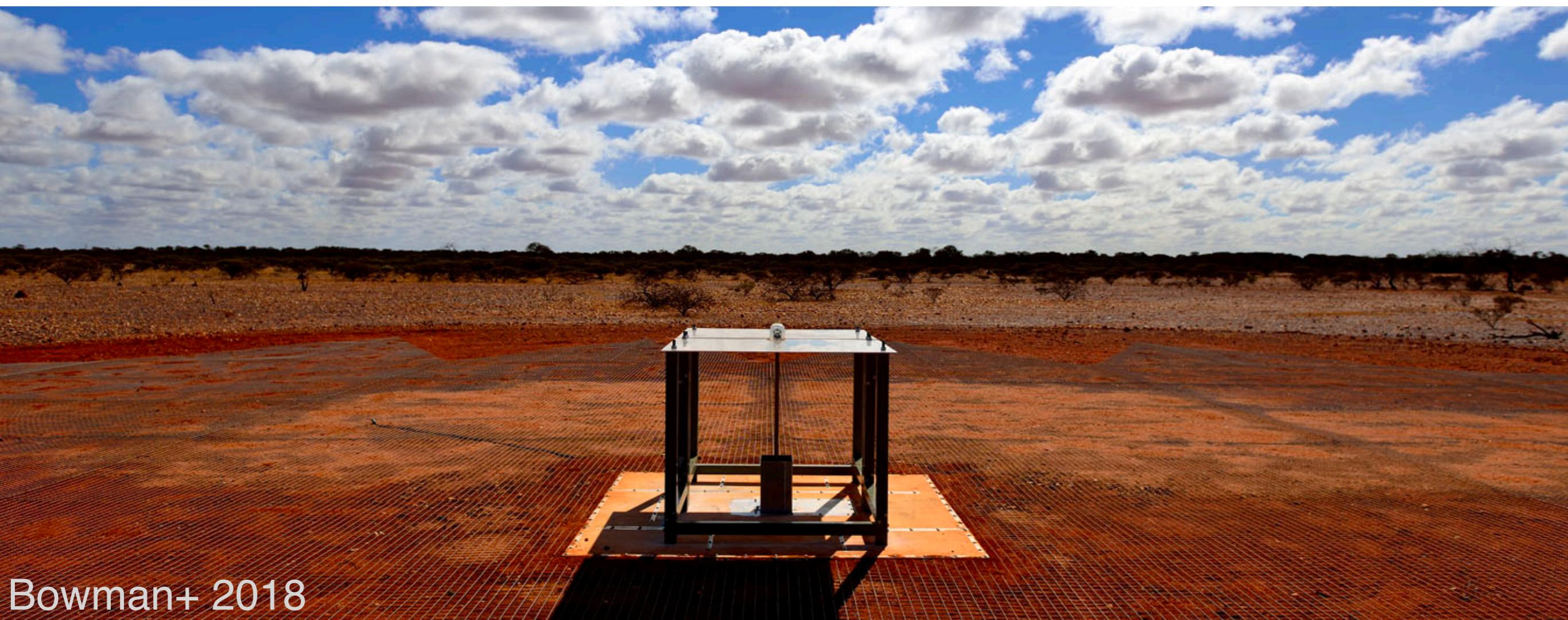
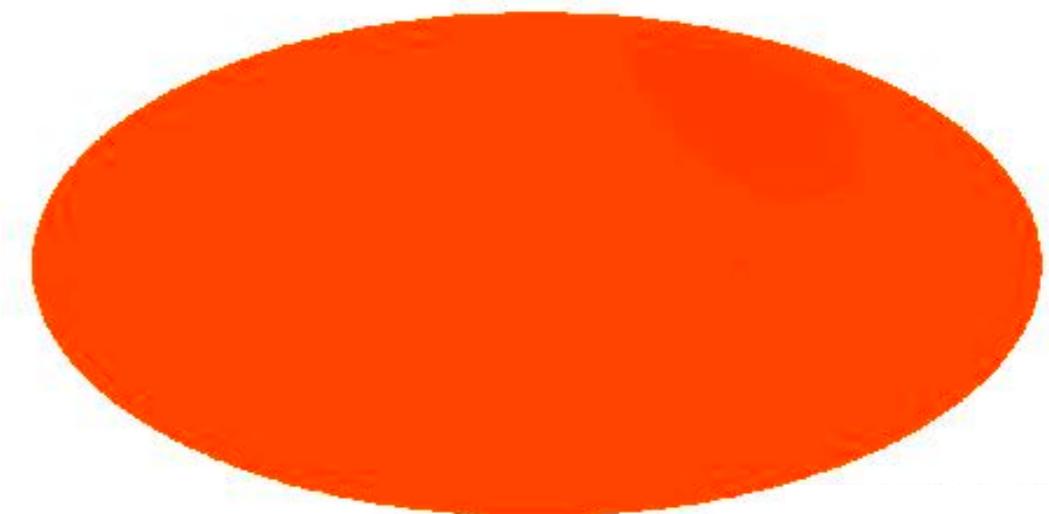
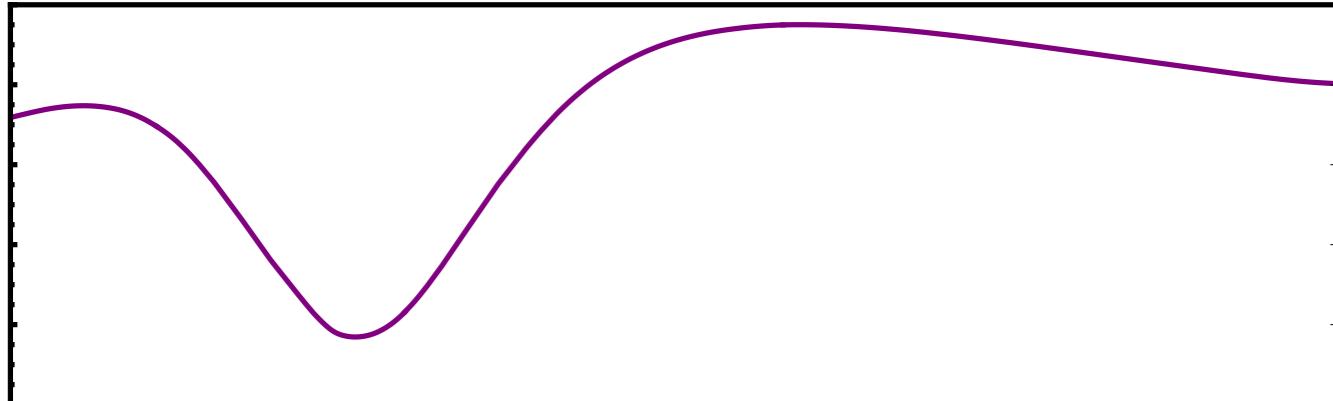
CMB Monopole



21-cm Global Signal
(EDGES, SARAS, PRIZM, MIST,...)

=

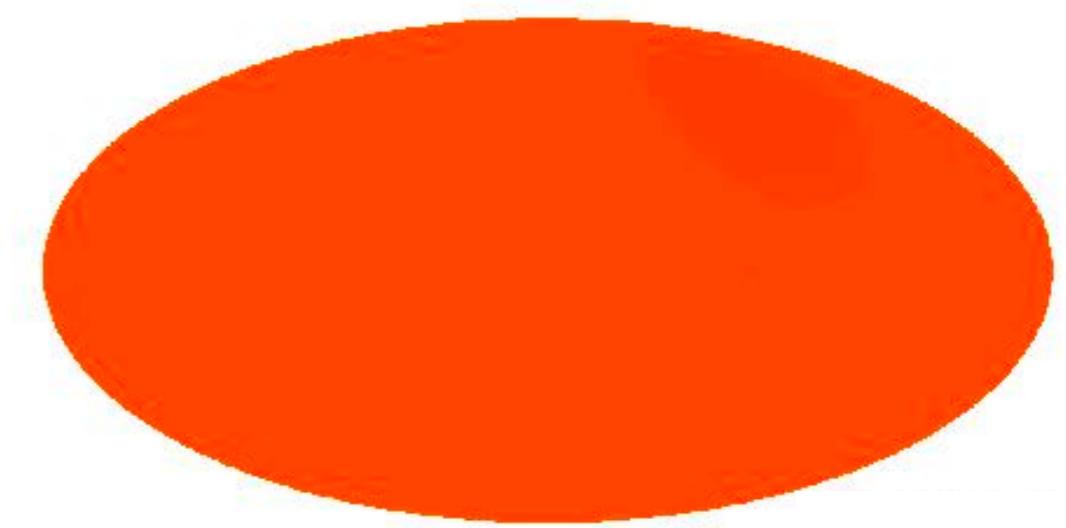
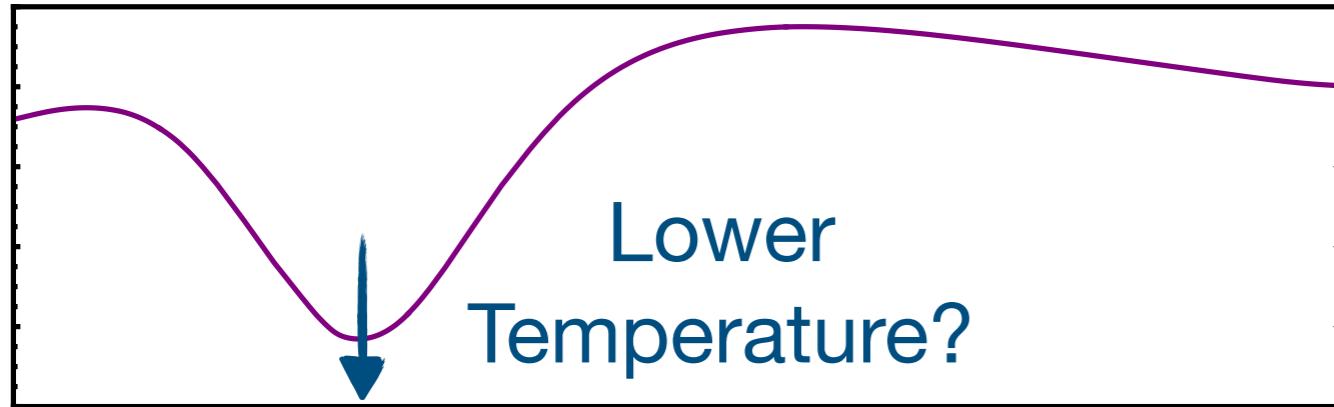
CMB Monopole



21-cm Global Signal
(EDGES, SARAS, PRIZM, MIST,...)

=

CMB Monopole



First detection? However, **foregrounds**....

EDGES
(Bowman+ 2018)

Hills+ 2018

Systematics?

SARAS 3
(Singh+ 2021)

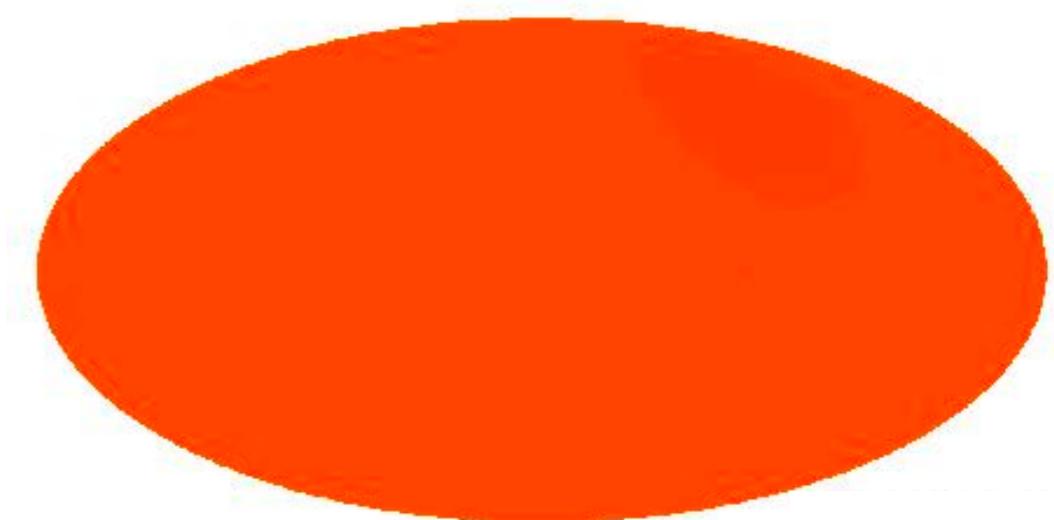
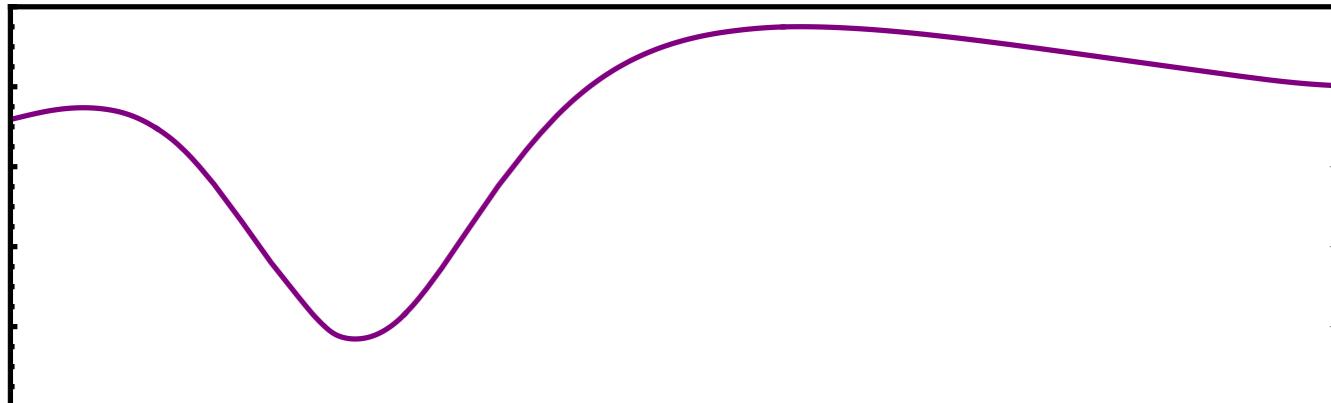
Dark-matter interactions?

mQ: JBM & Loeb 18, Barkana+, Berlin+ 18, Liu+ 19...

21-cm Global Signal
(EDGES, SARAS, PRIZM, MIST,...)

=

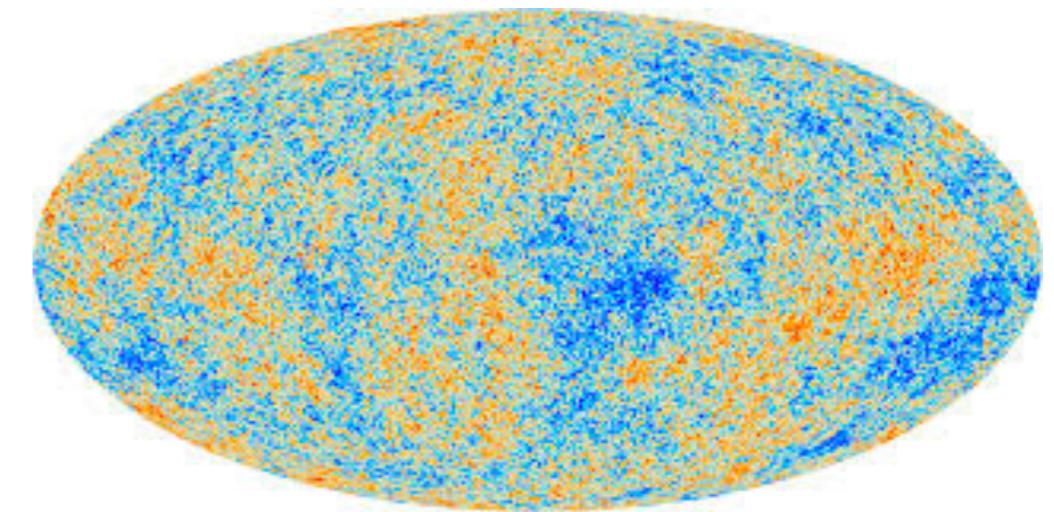
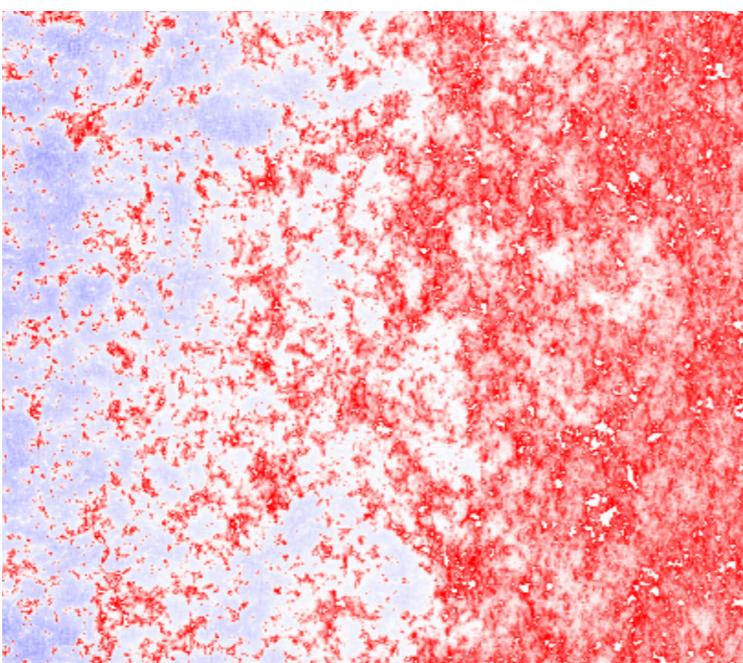
CMB Monopole



21-cm Fluctuations
(**HERA**, MWA, LOFAR, PAPER, SKA,...)

=

CMB Anisotropies



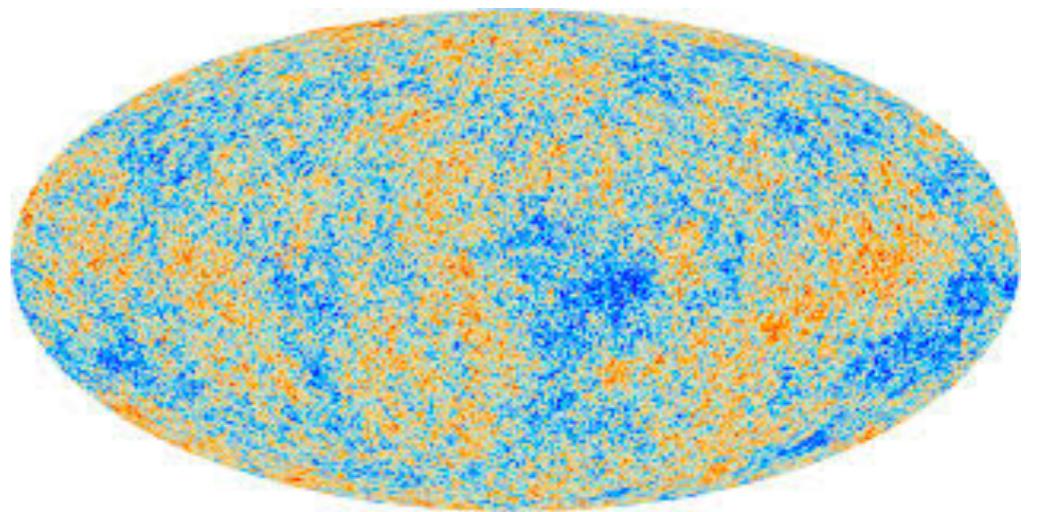
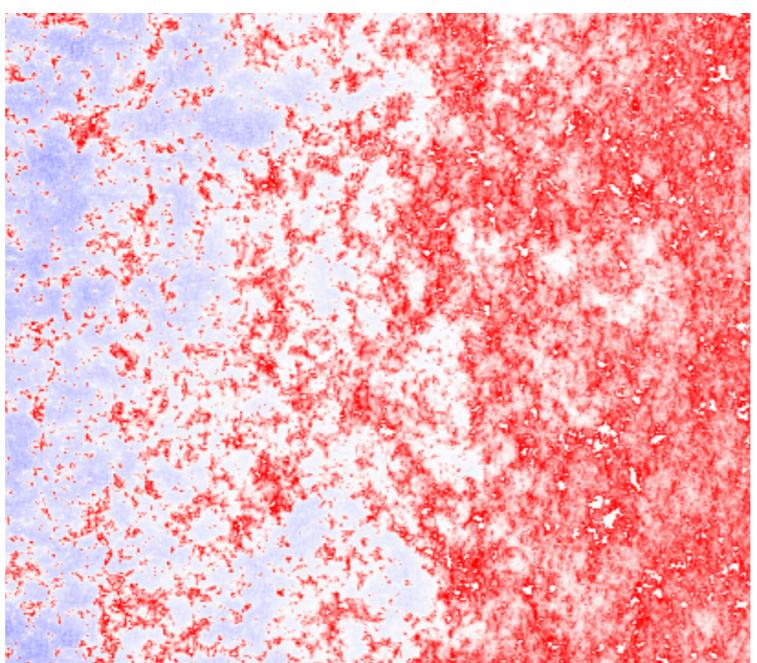


21-cm Fluctuations
(**HERA**, MWA, LOFAR, PAPER, SKA,...)

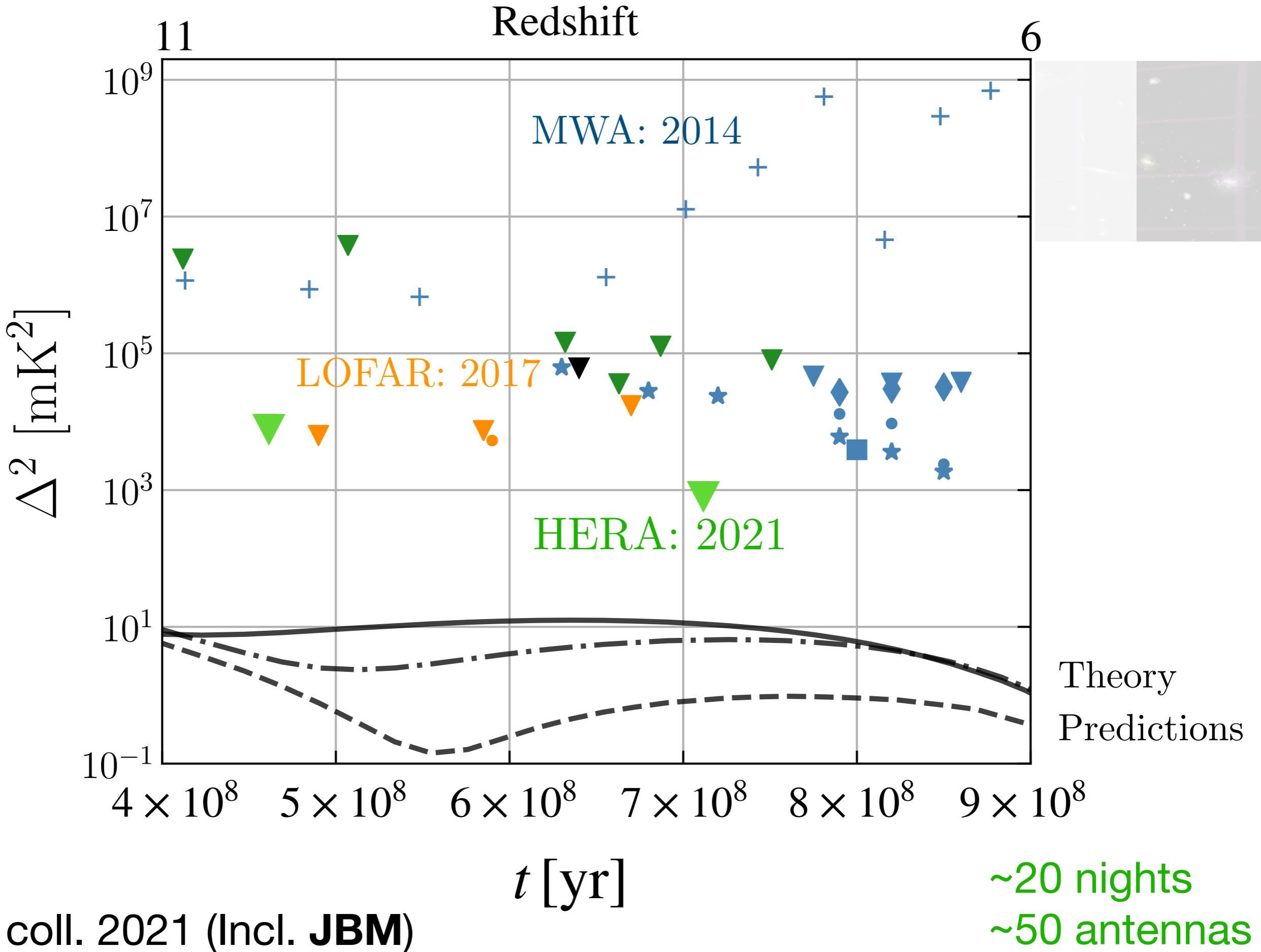
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CMB Anisotropies

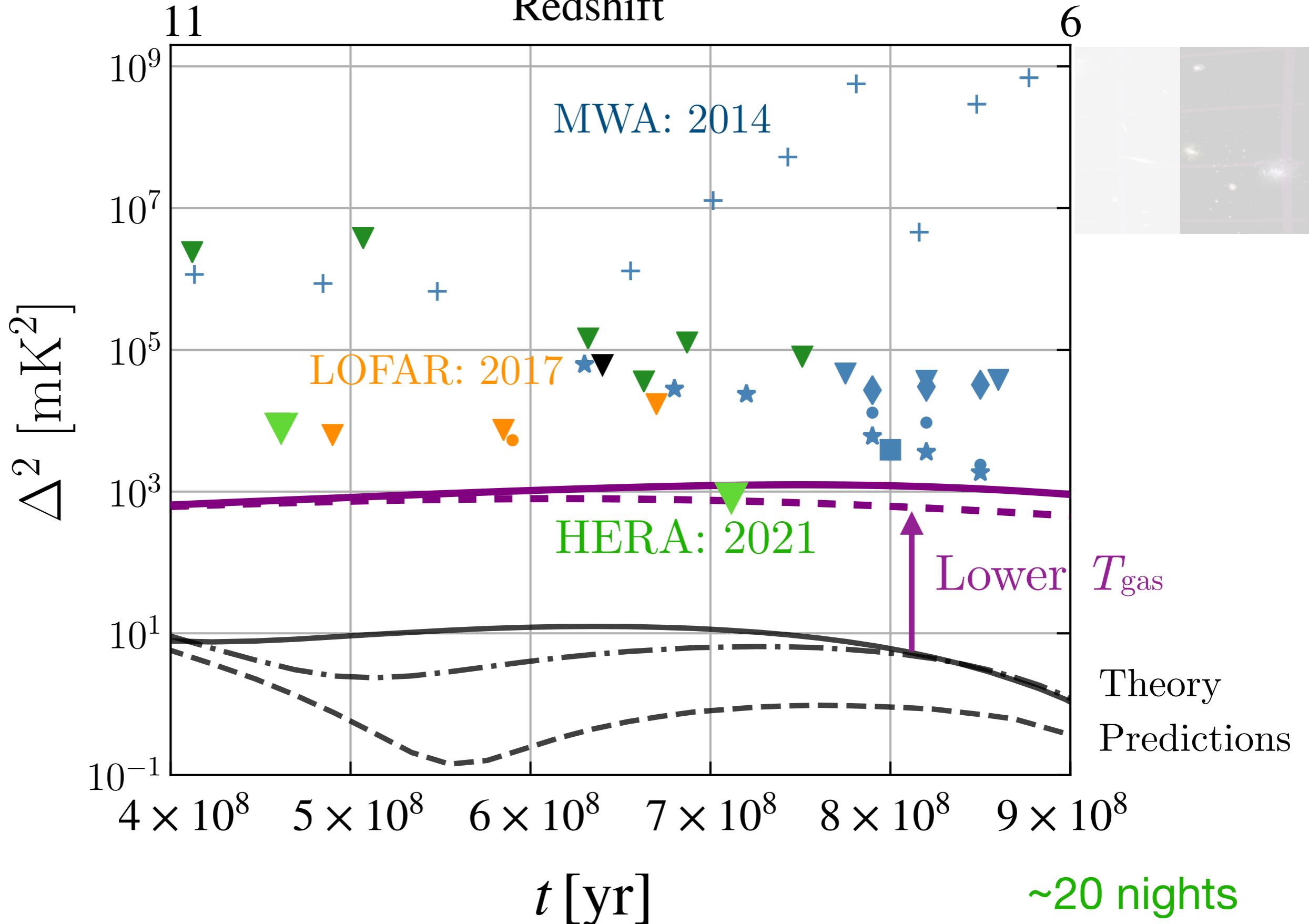


The status of 21-cm data



HERA vs EDGES

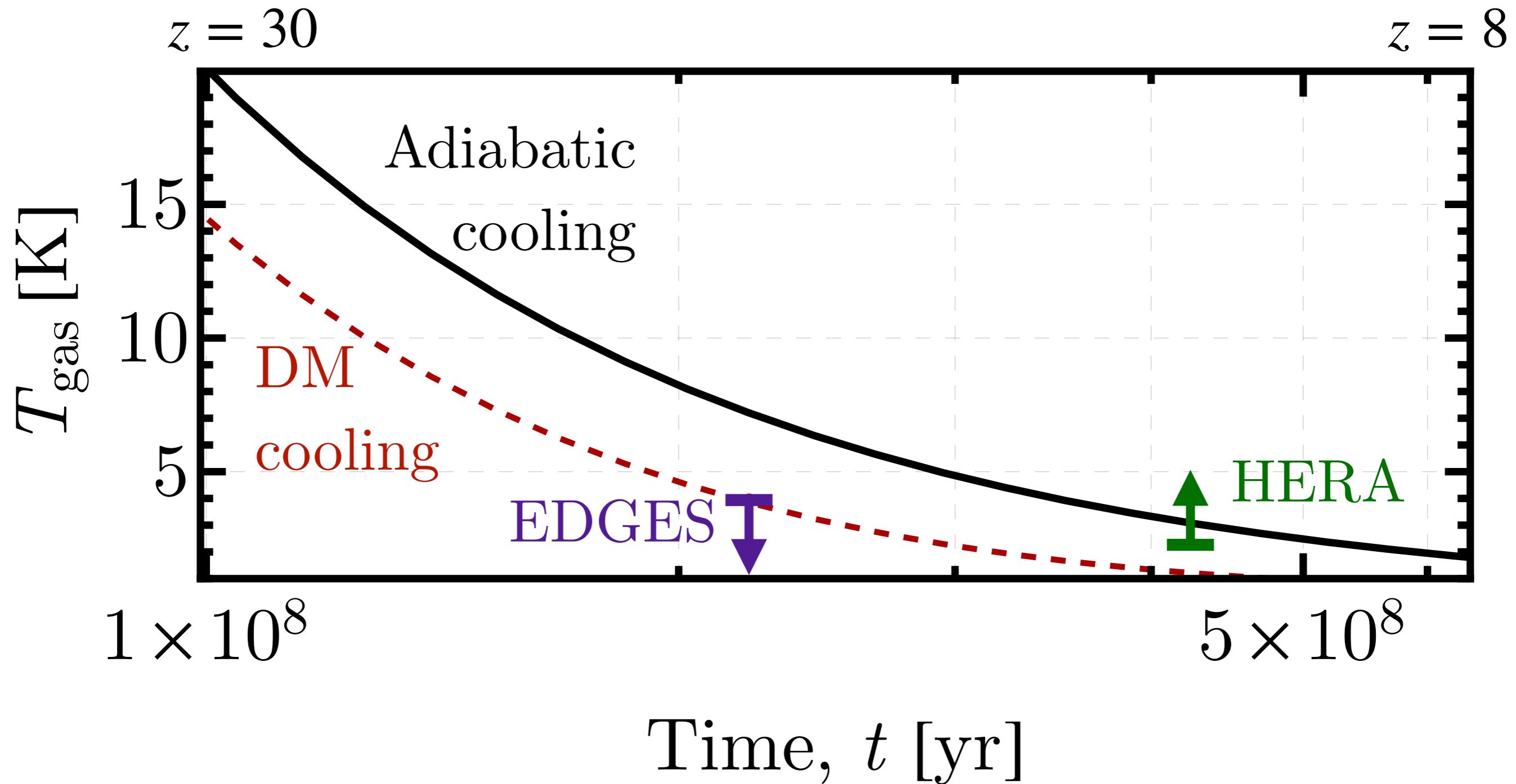
Redshift



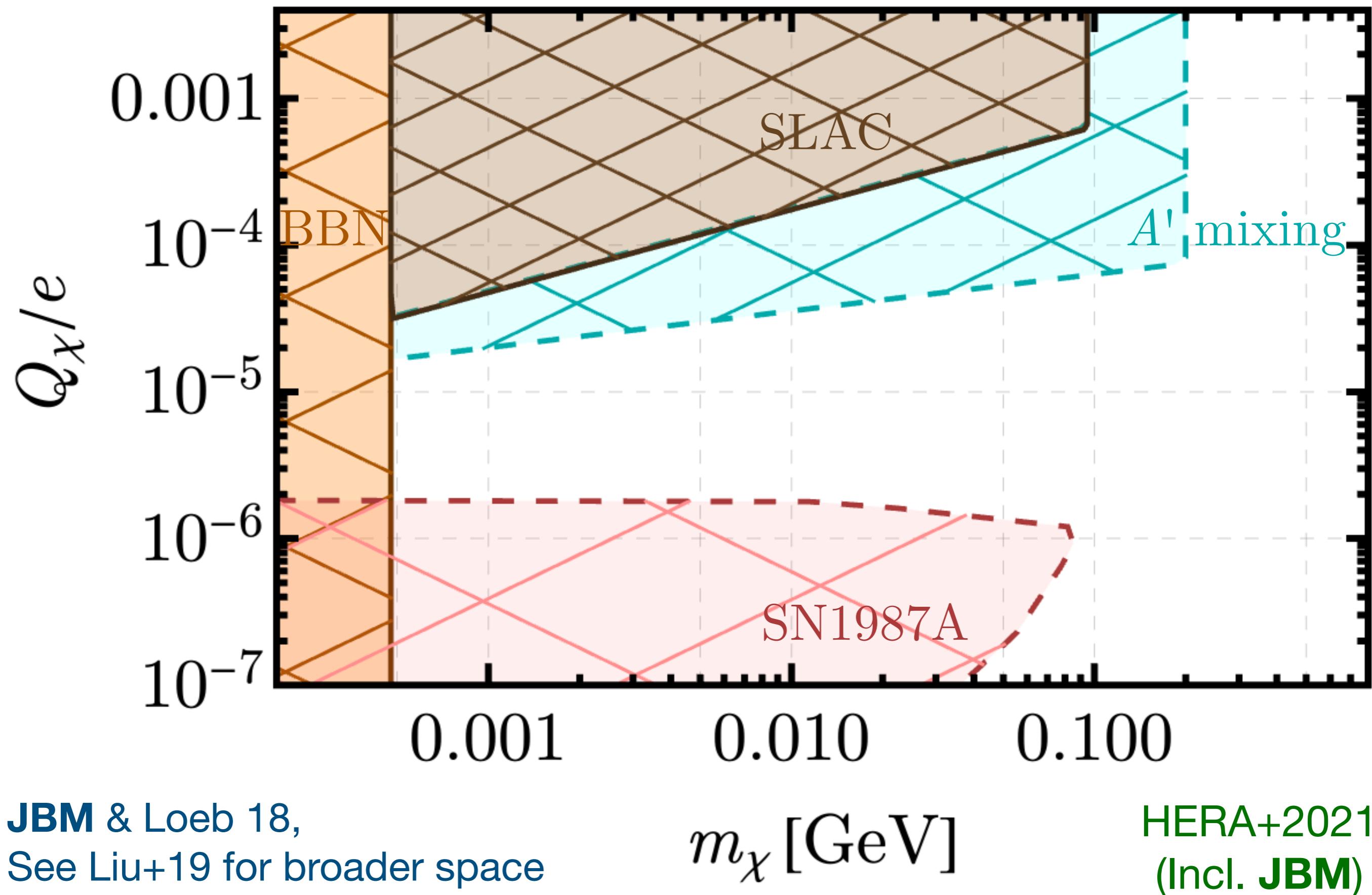
HERA coll. 2021 (Incl. JBM)

~20 nights
~50 antennas

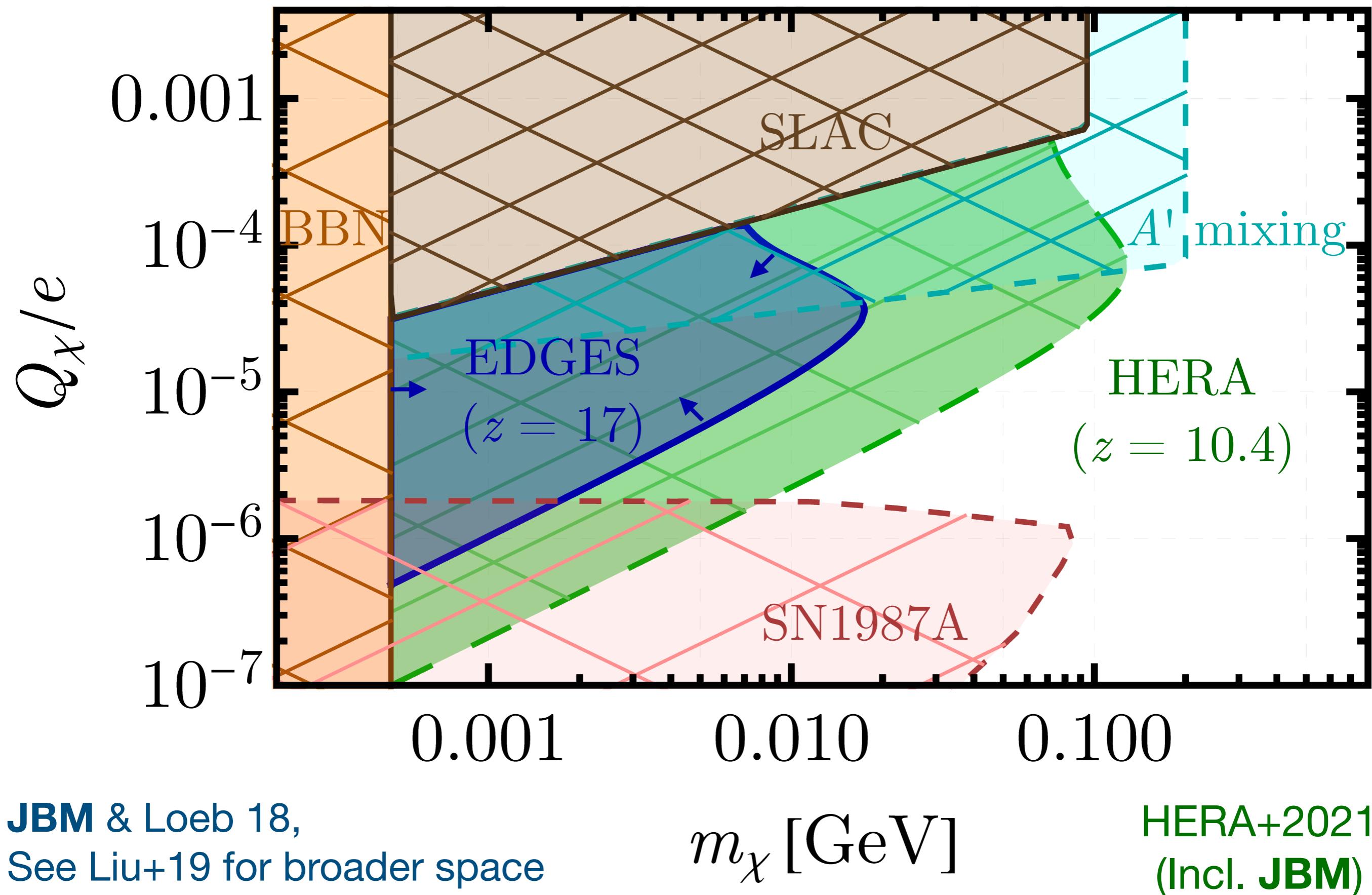
HERA vs EDGES



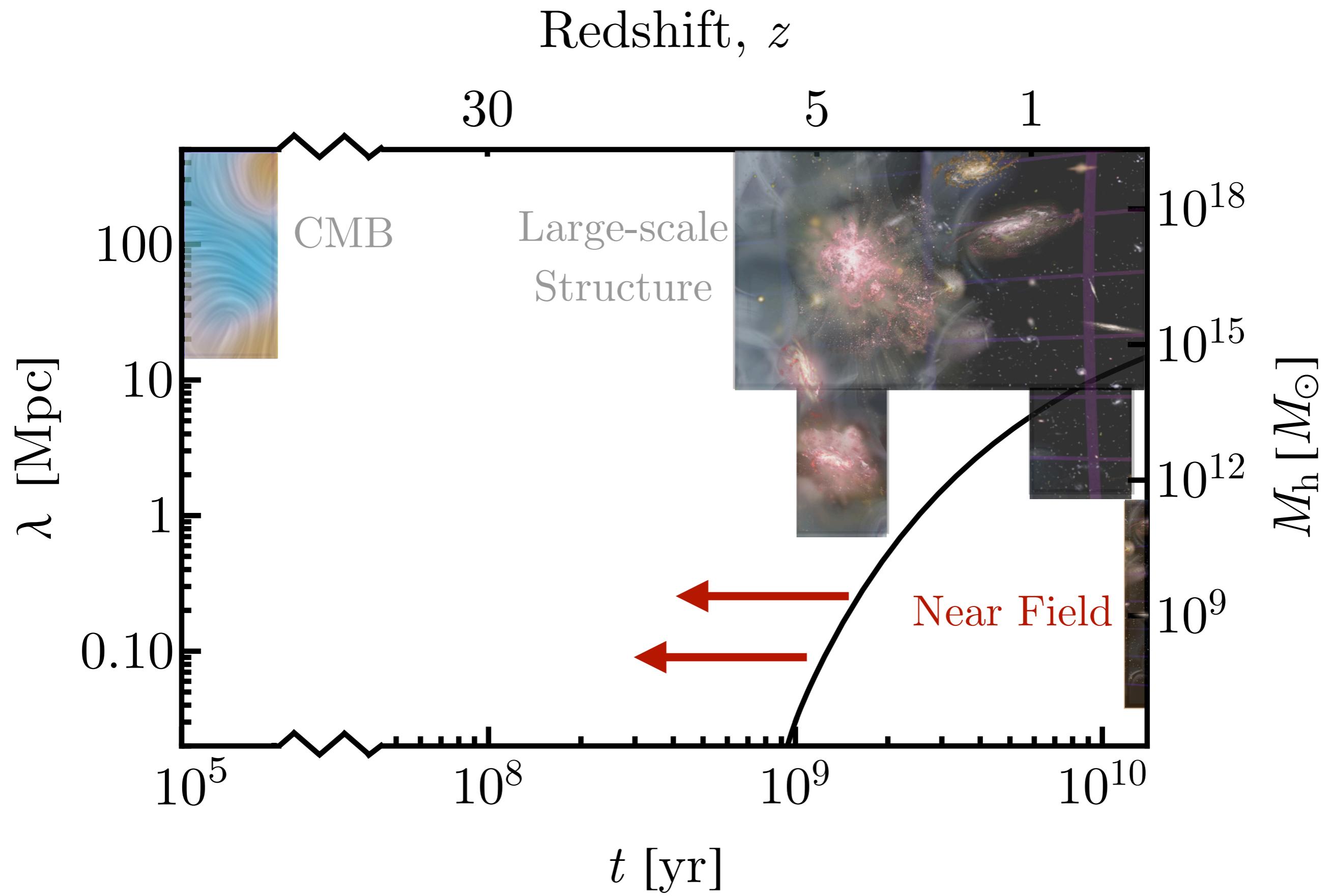
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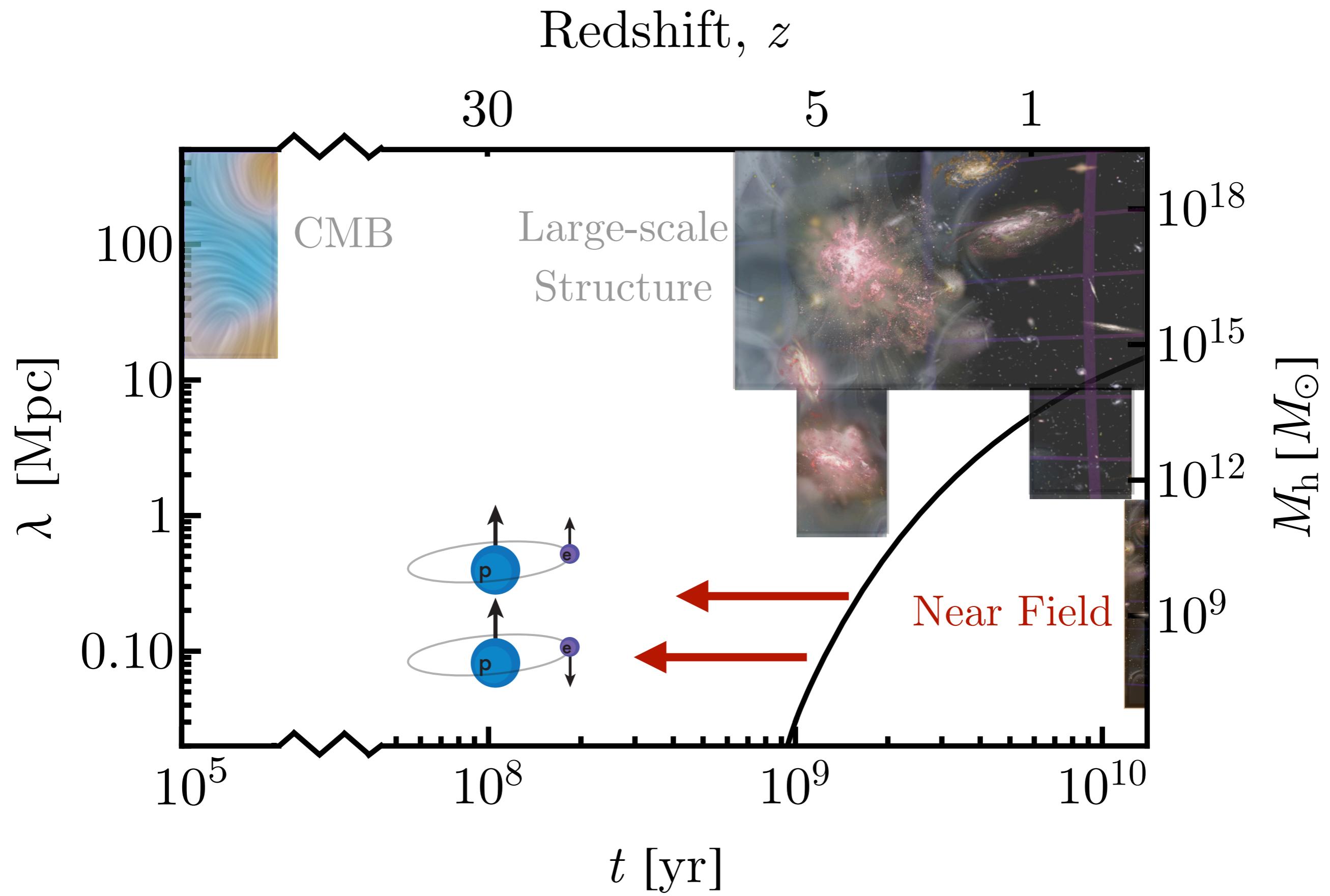
HERA vs EDGES



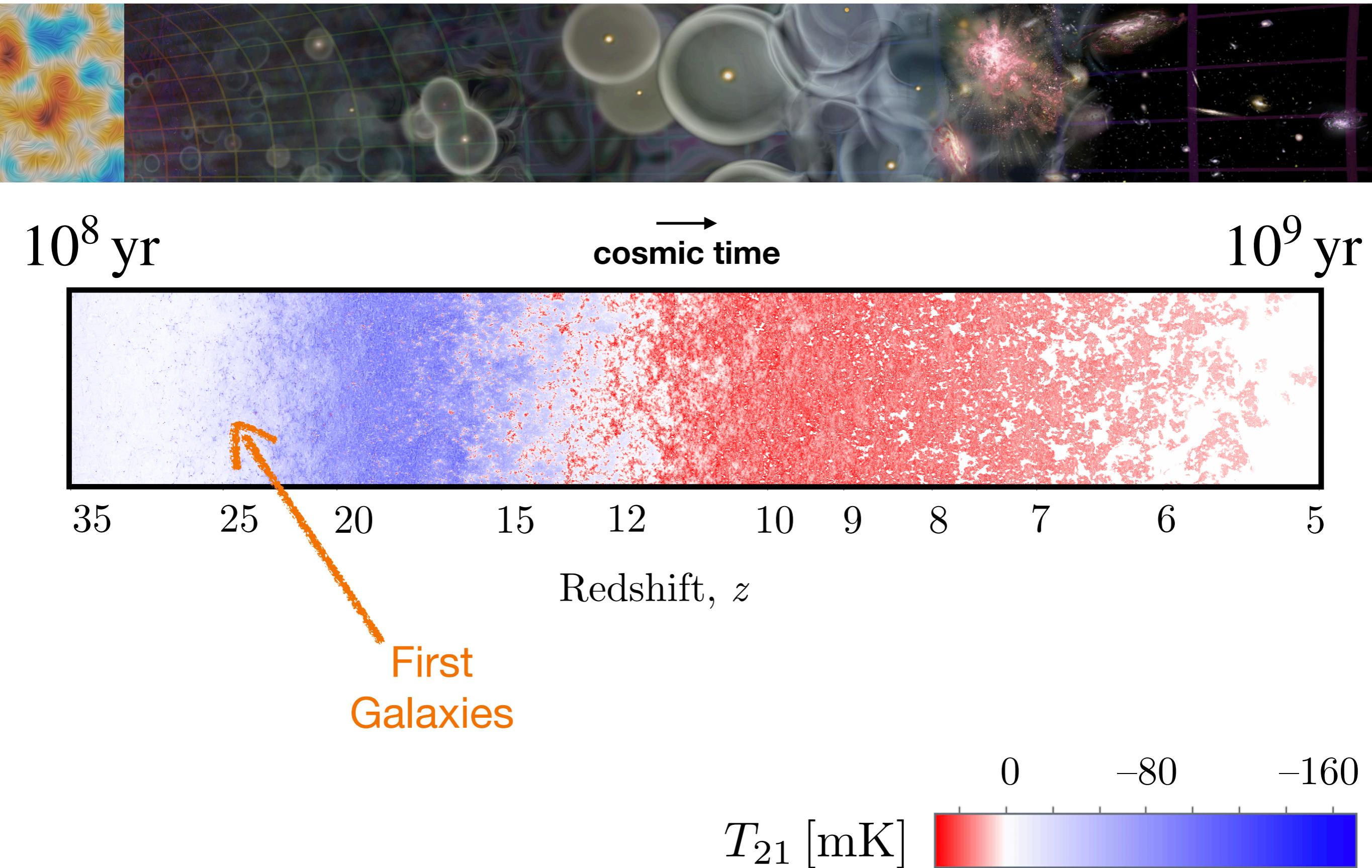
Probing the uncharted



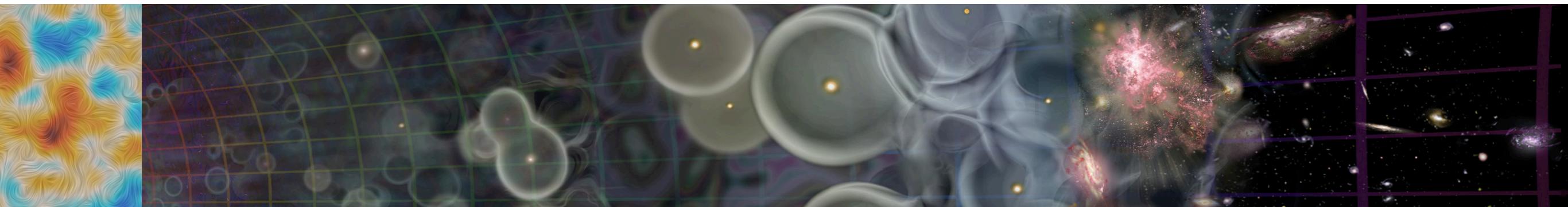
Probing the uncharted



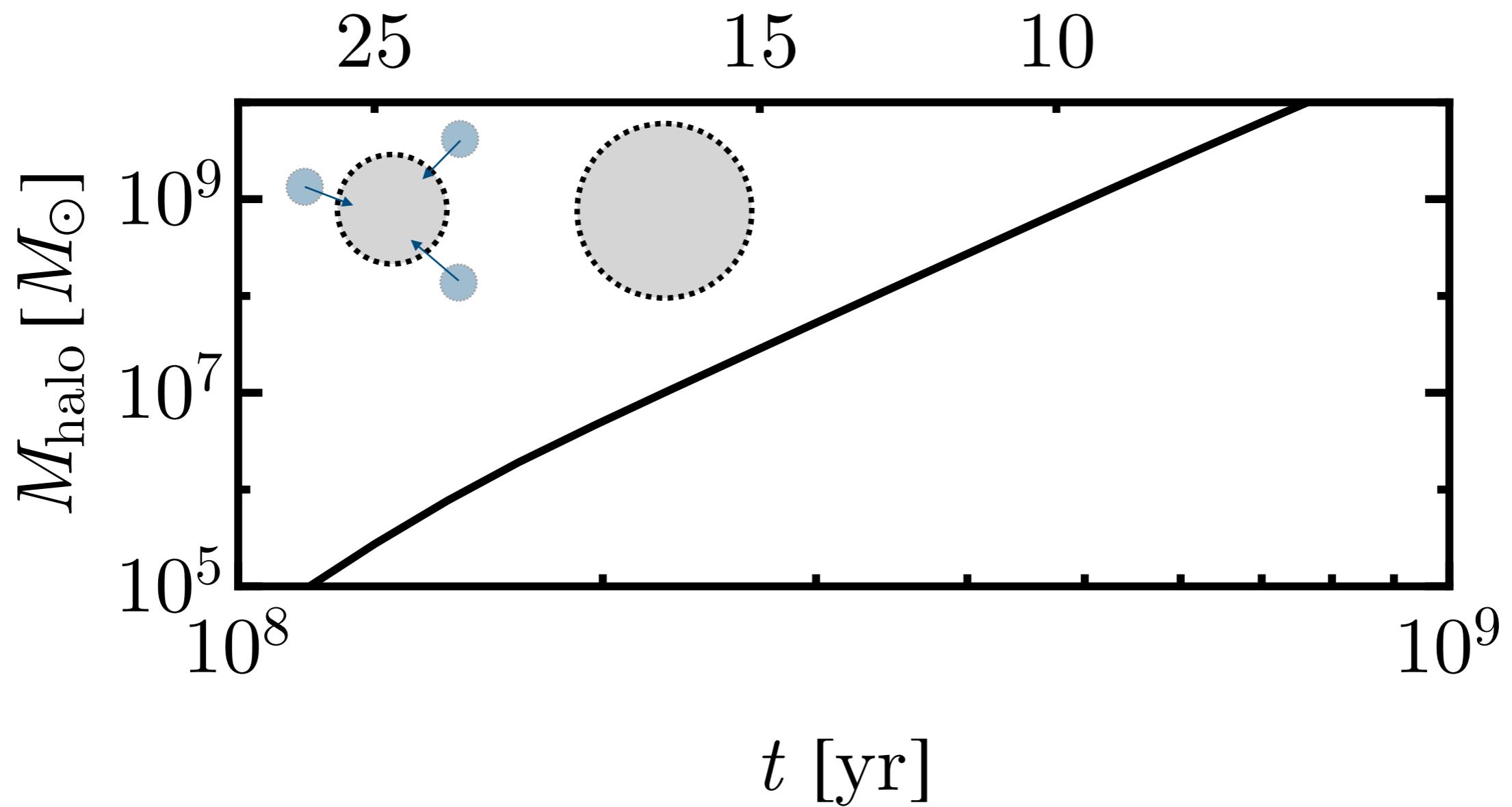
How does 21-cm help us with DM?



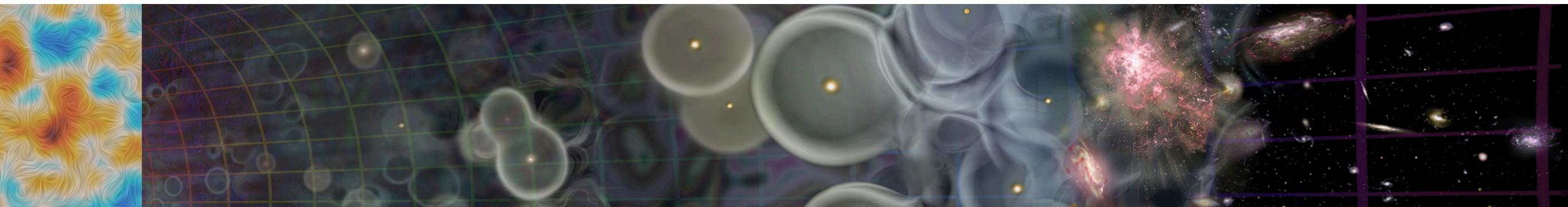
The first galaxies and DM



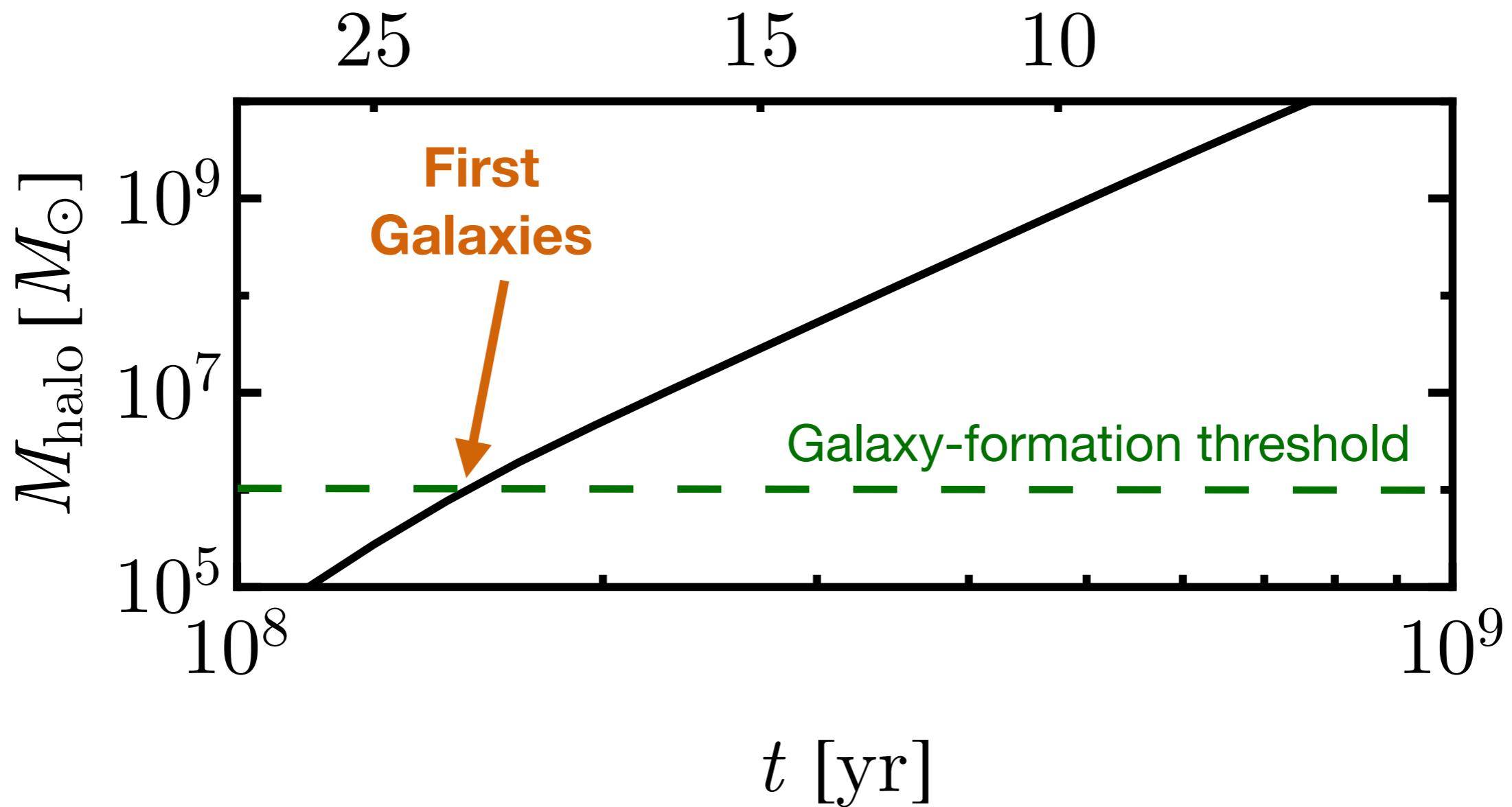
Redshift, z



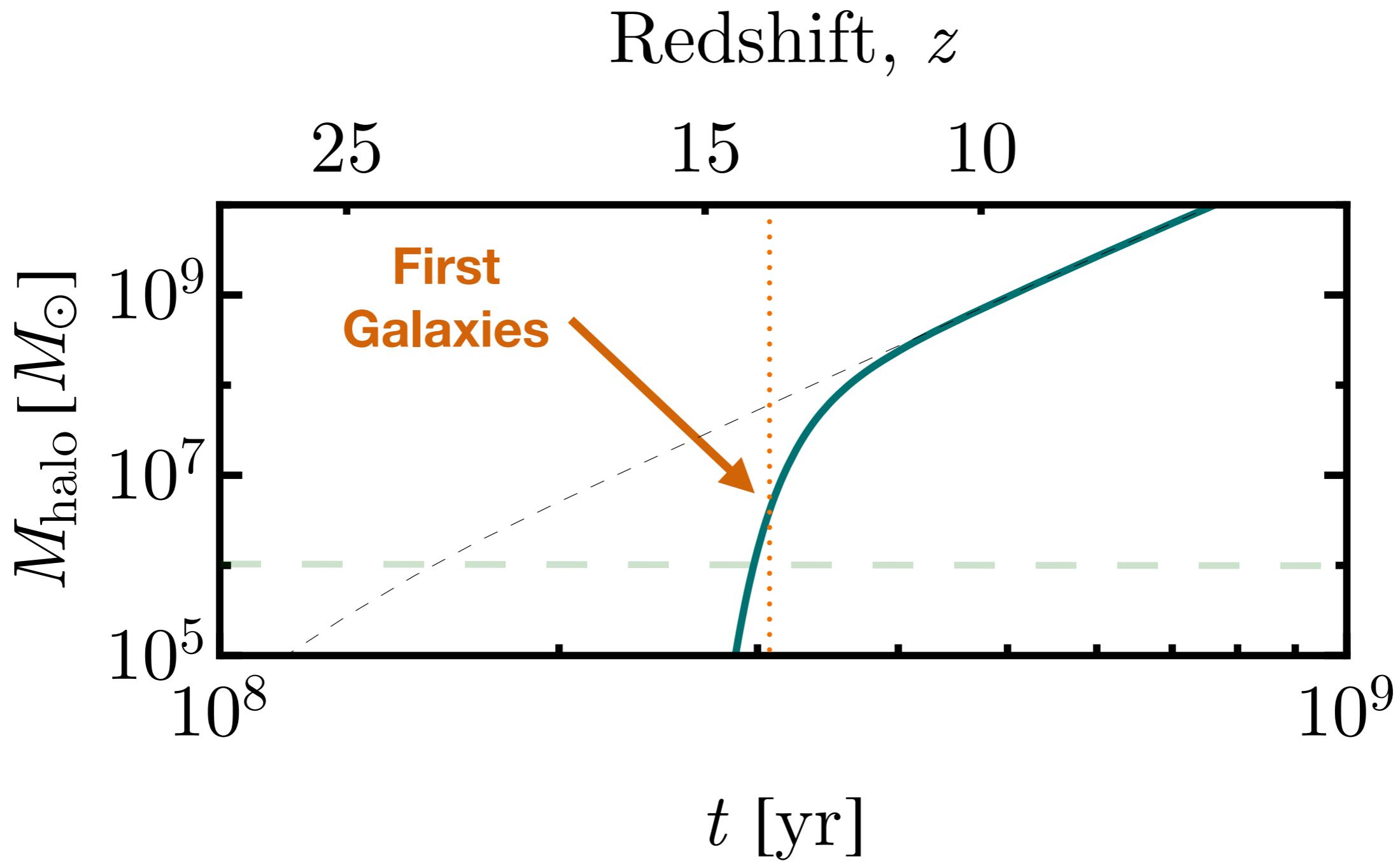
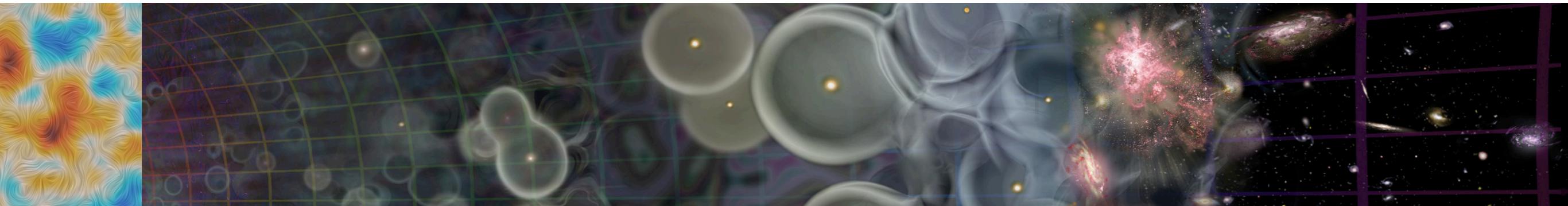
The first galaxies and DM



Redshift, z

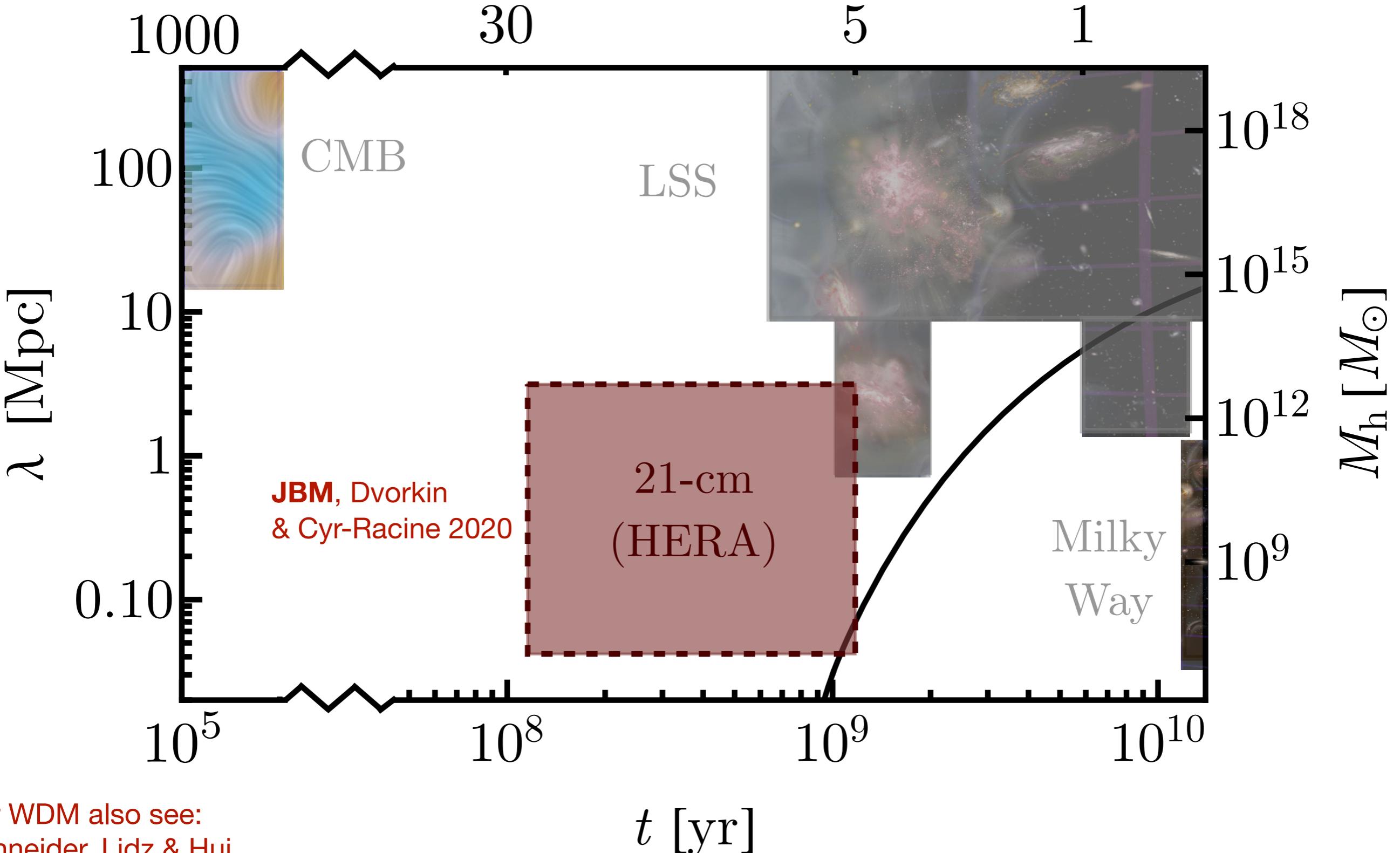


How does 21-cm help us with DM?



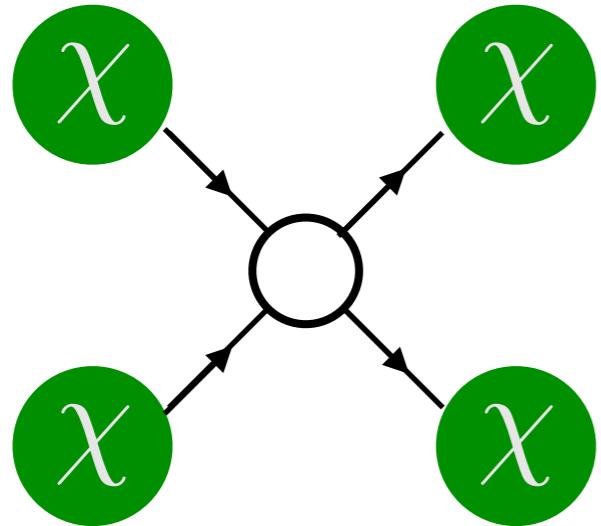
Probing the uncharted

Redshift, z

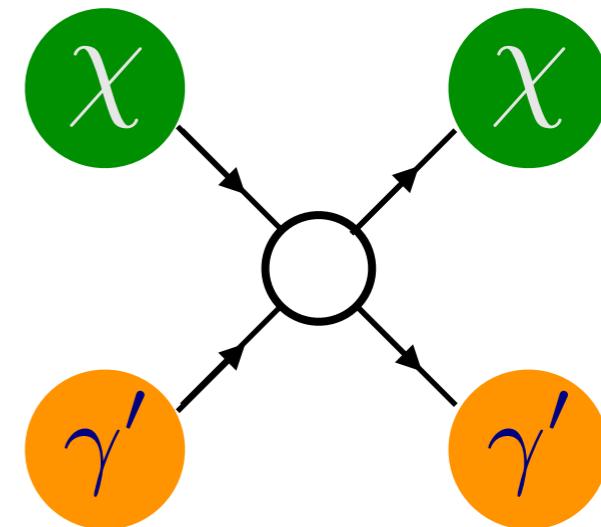


For WDM also see:
Schneider, Lidz & Hui,
Boyarsky+, Yoshiura+ 2018

Effective Theory of Structure Formation: ETHOS



DM-DM:
Halo profiles, etc.



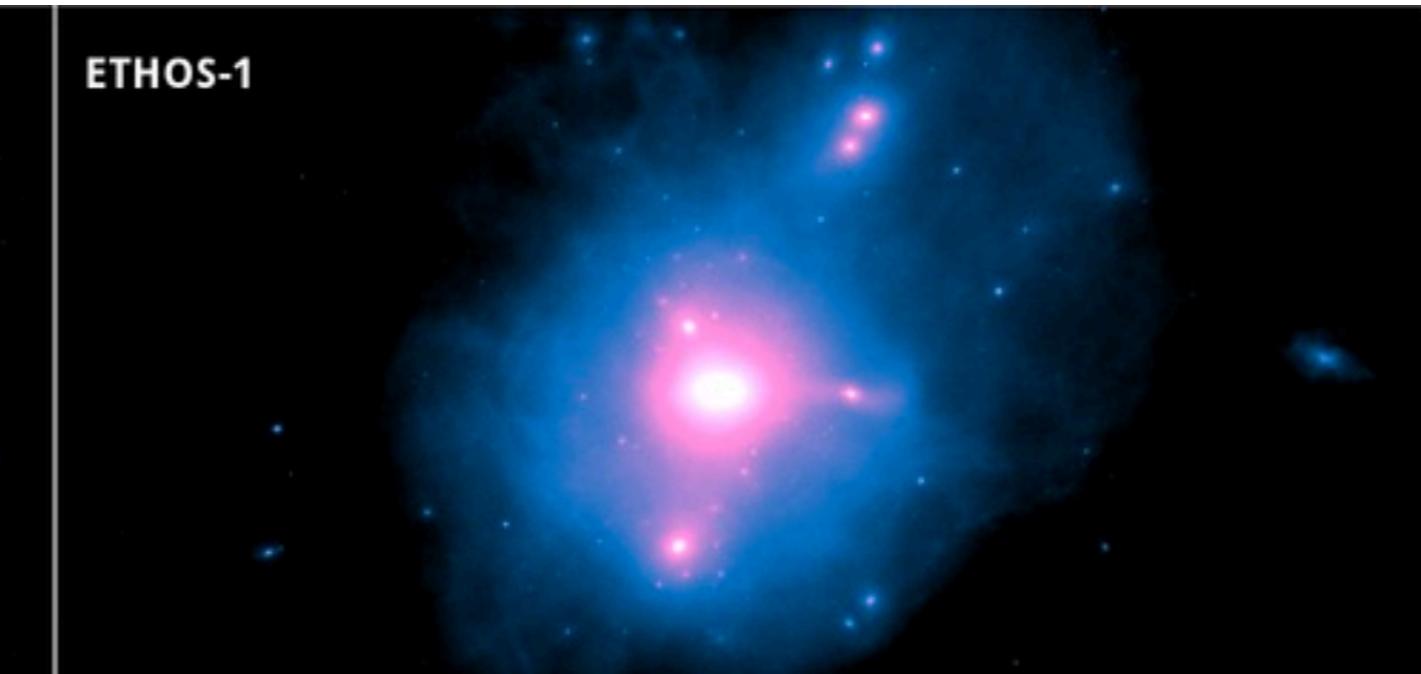
DM-DR:
Power Spectrum

Vogelsberger+ 2016, Cyr-Racine+2016, JBM+ 2021

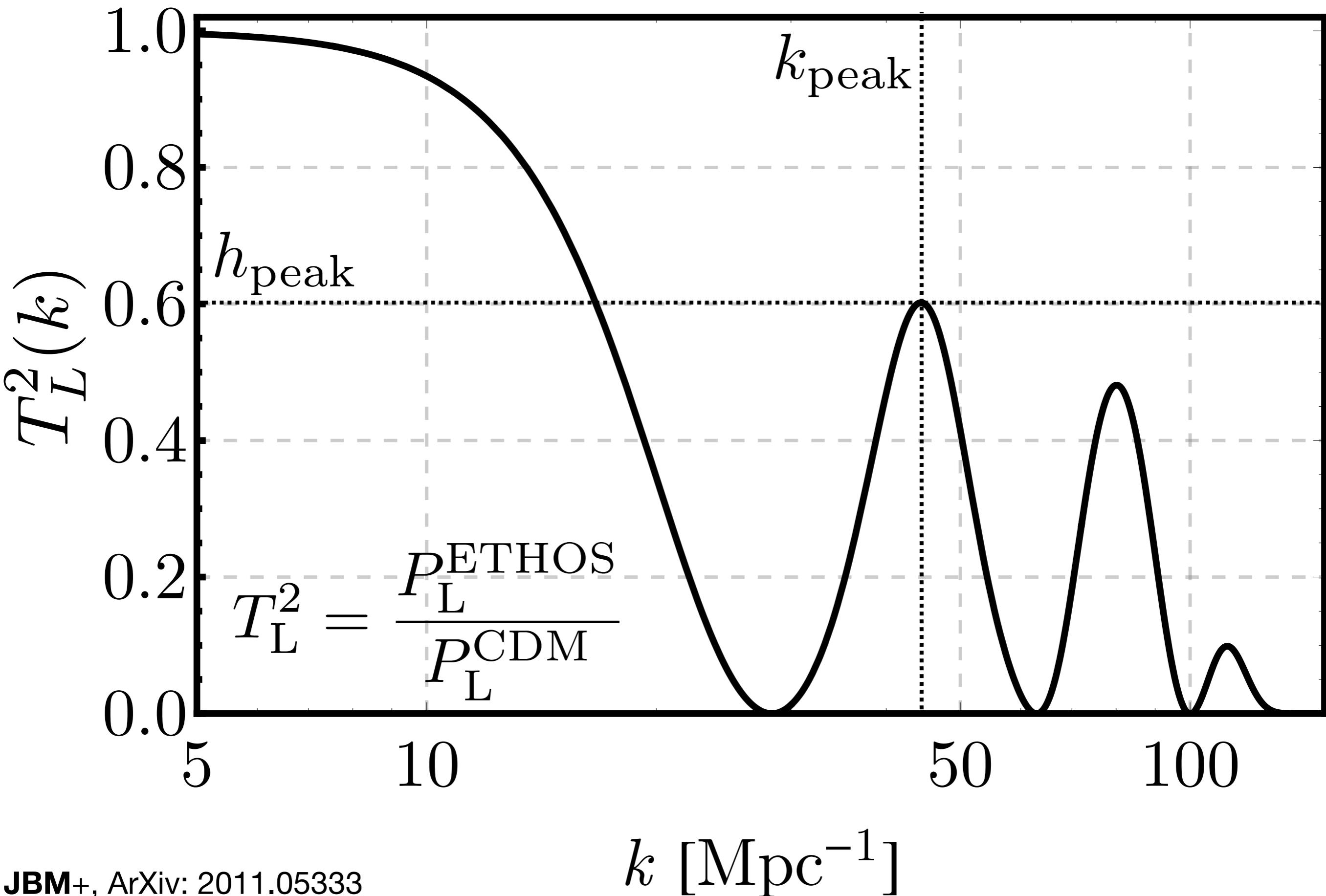
CDM



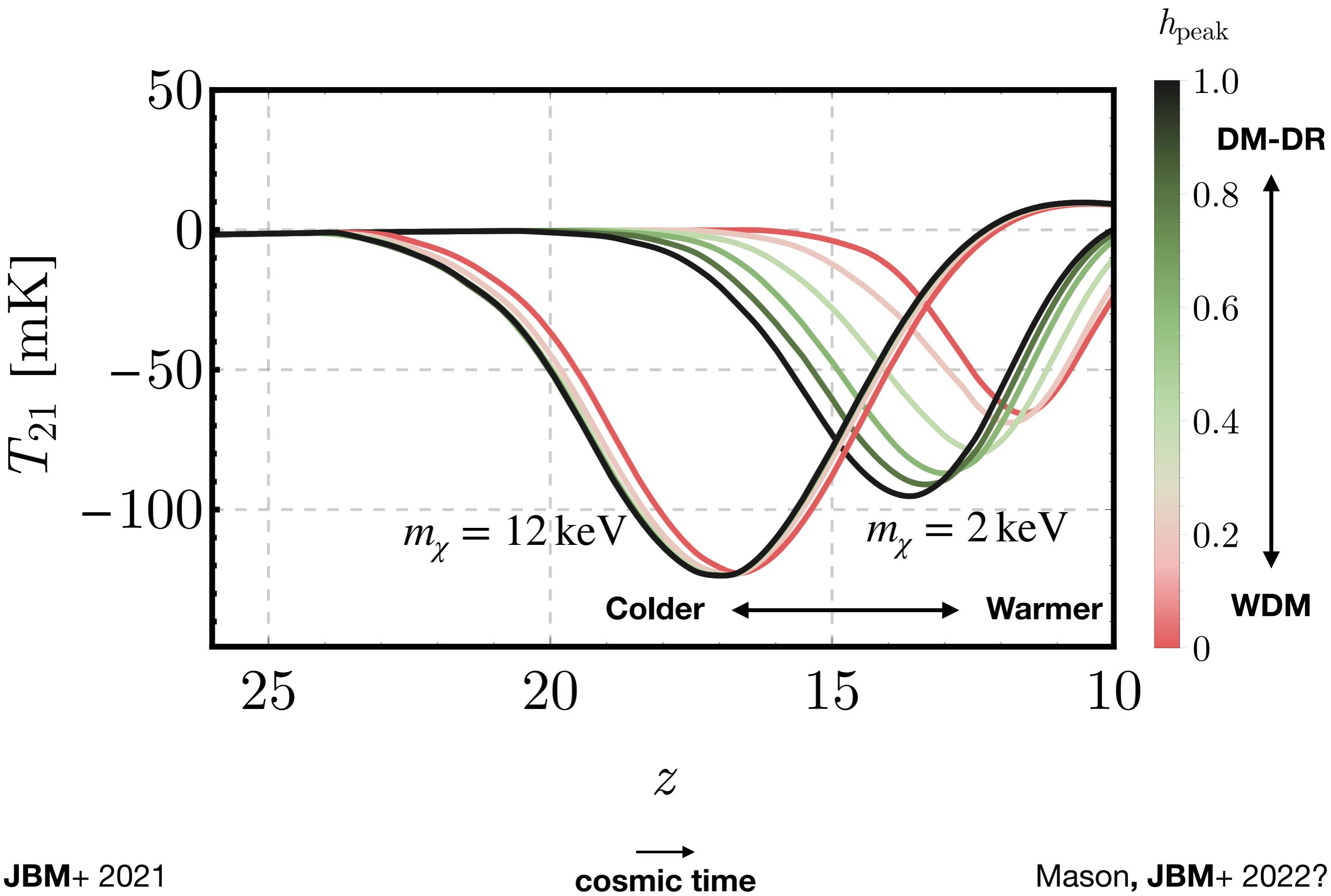
ETHOS-1



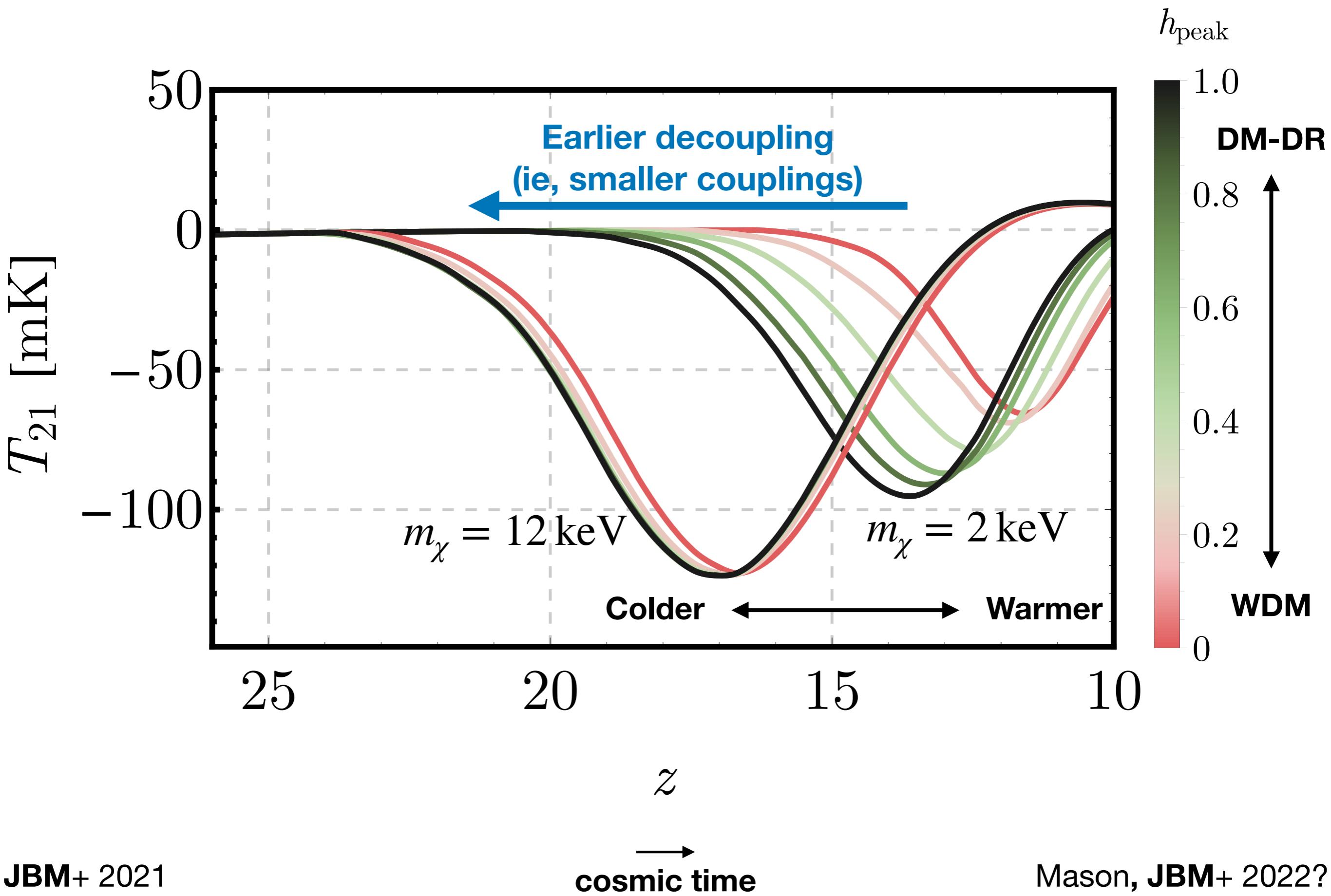
A two-parameter model



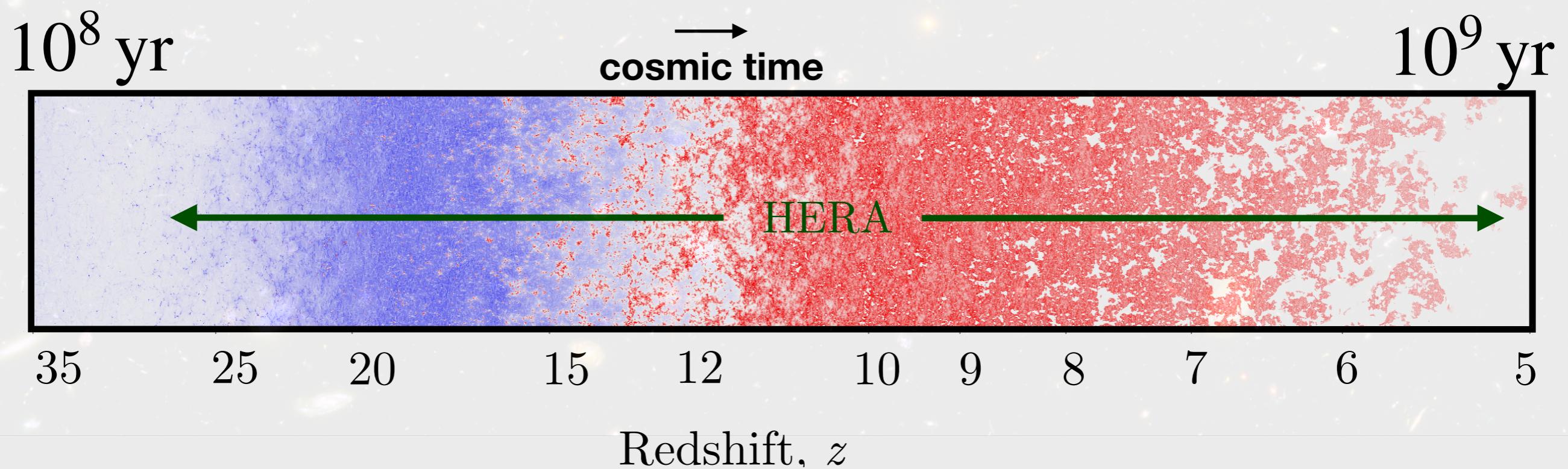
The 21-cm global signal in ETHOS



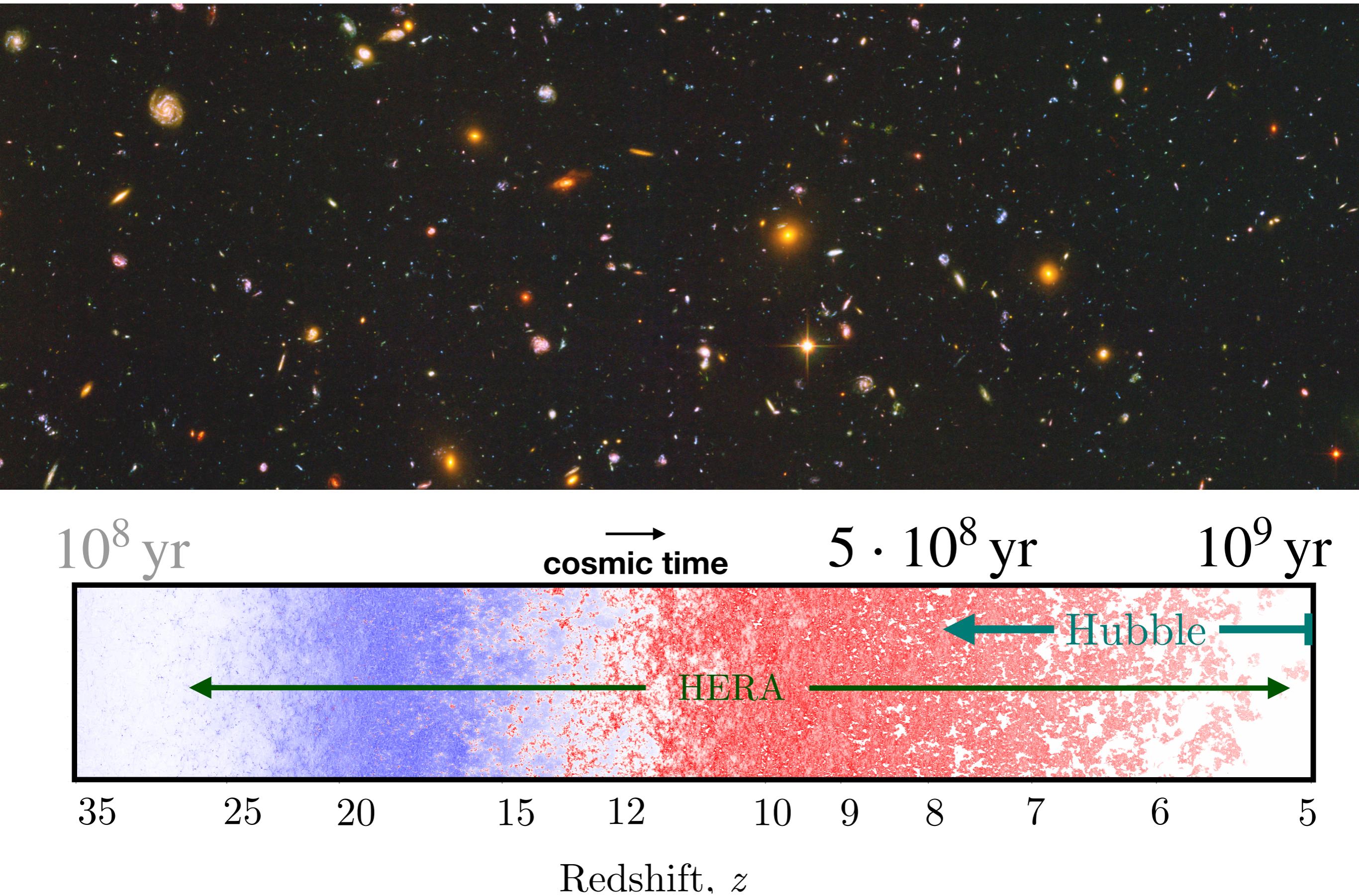
The 21-cm global signal in ETHOS



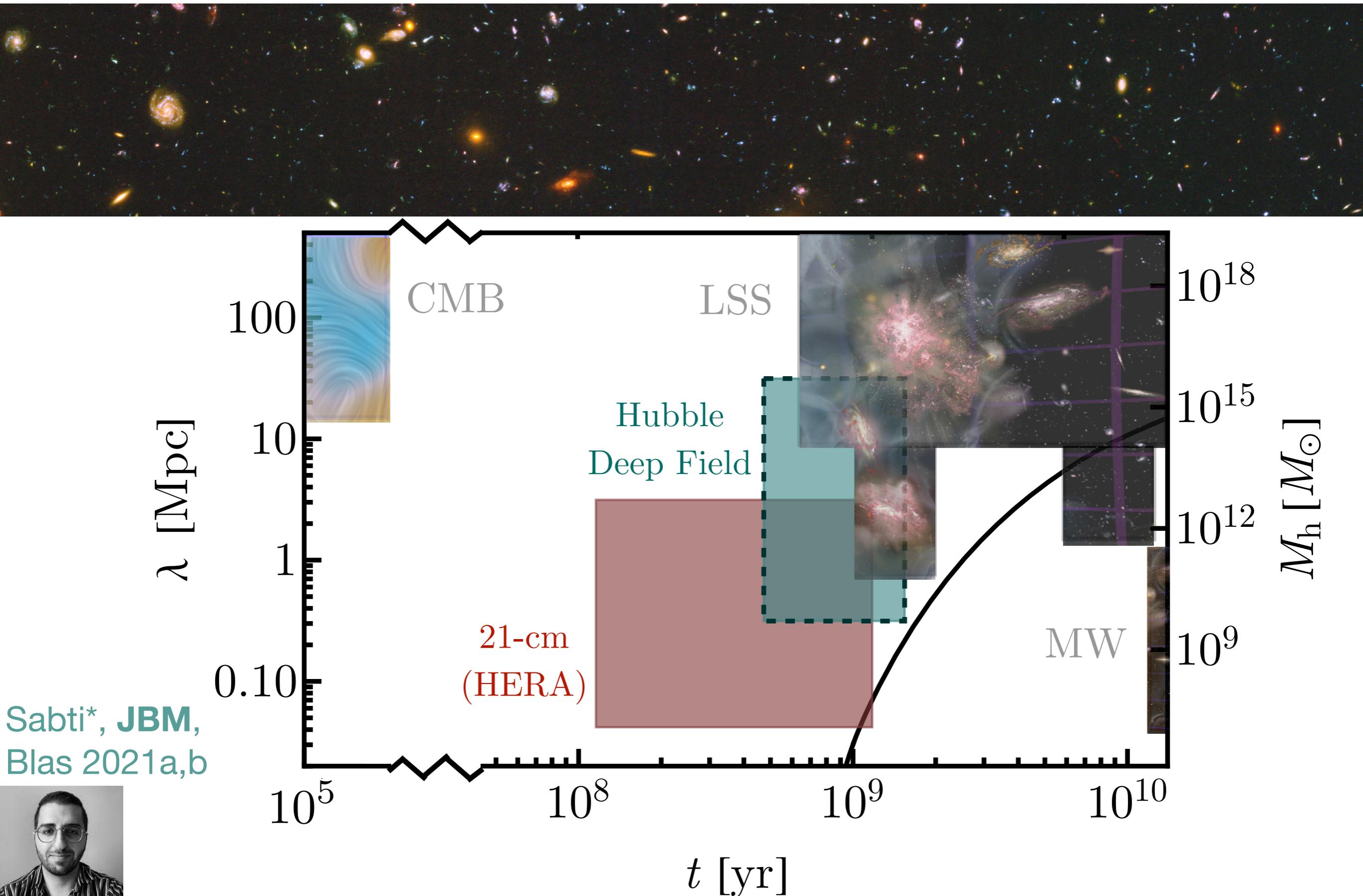
What can we say now?



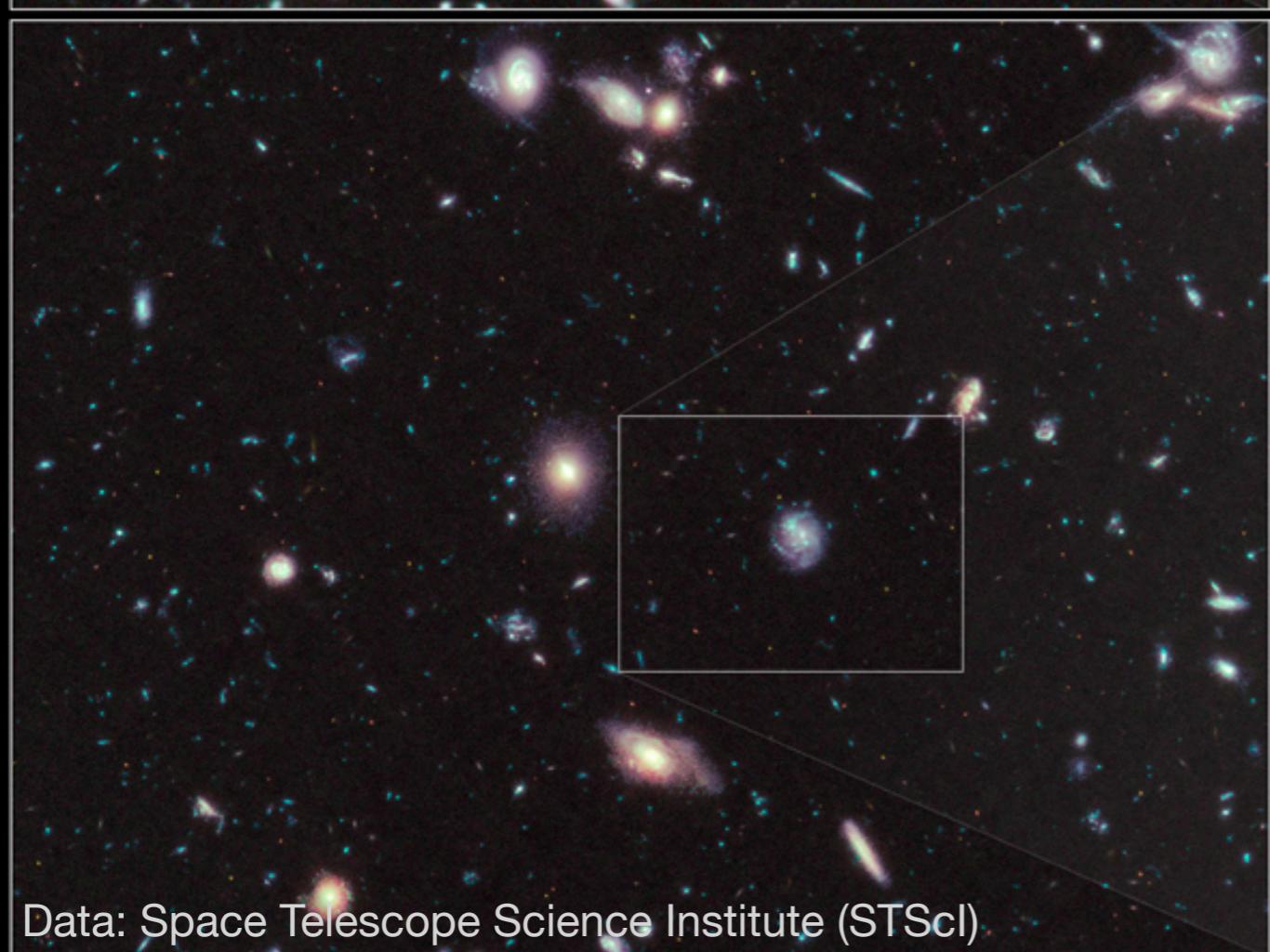
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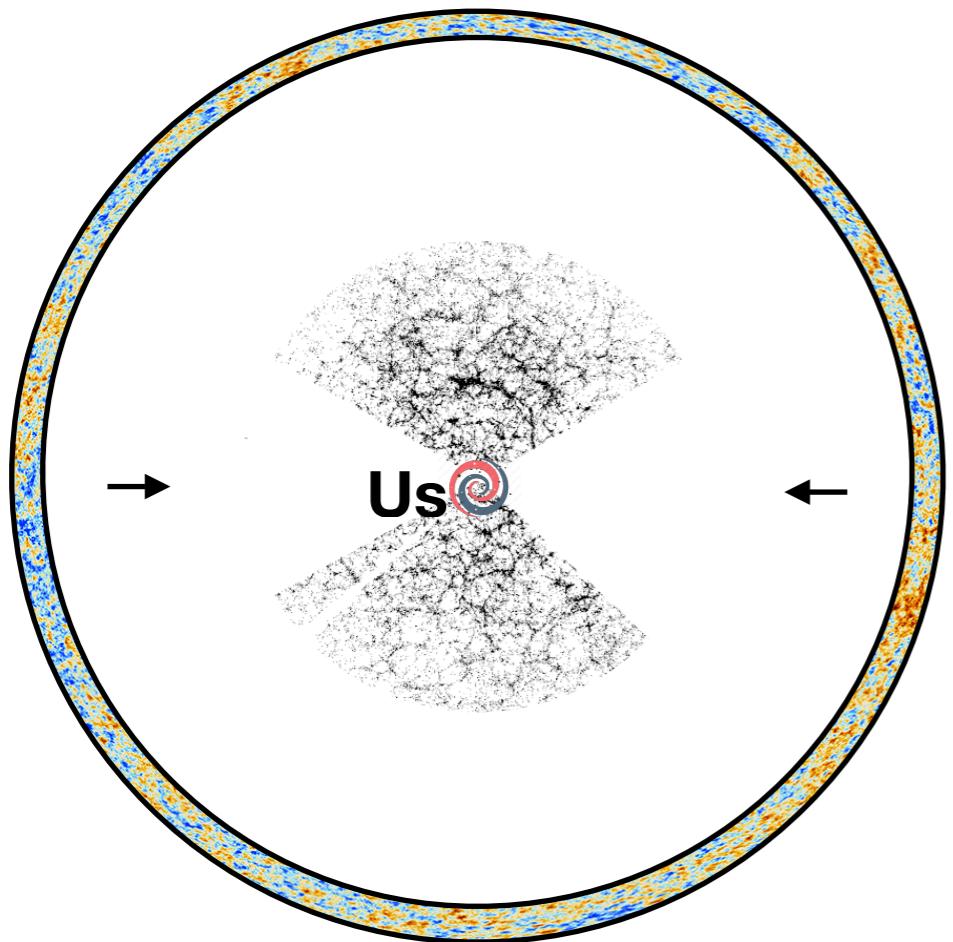
What can we say now?



What can we say in ~a year?



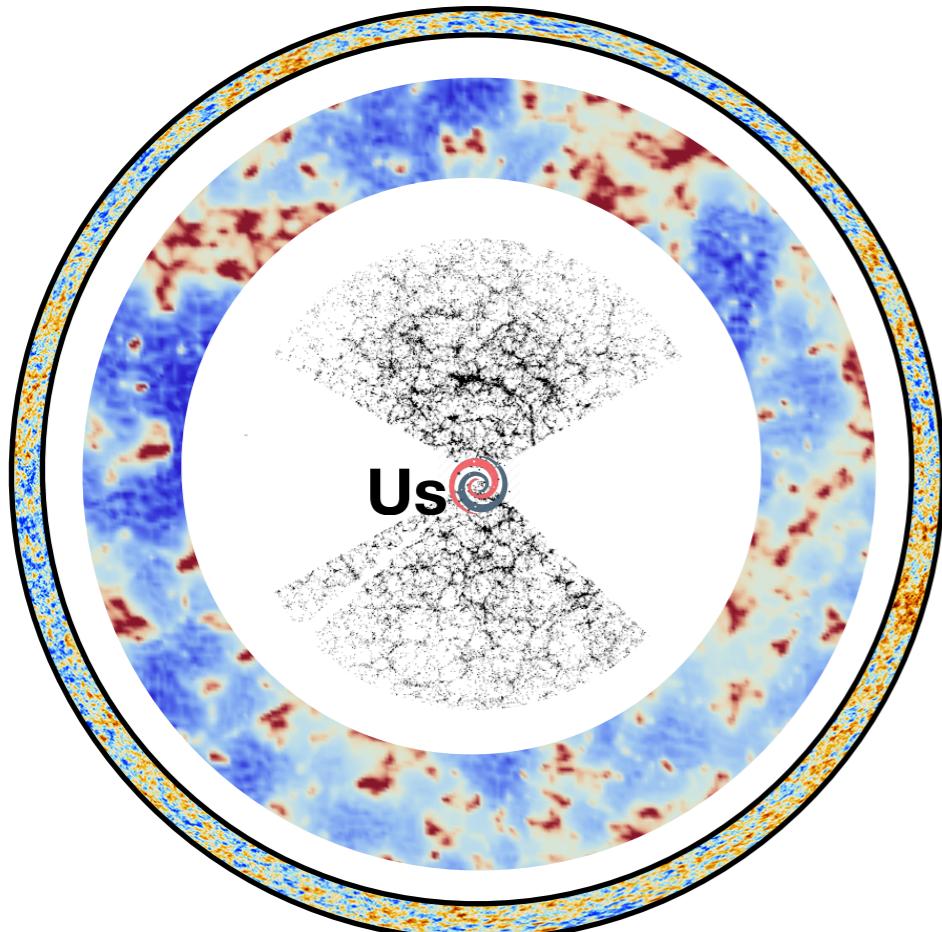
Summary



-We have not observed most of the Universe (yet)

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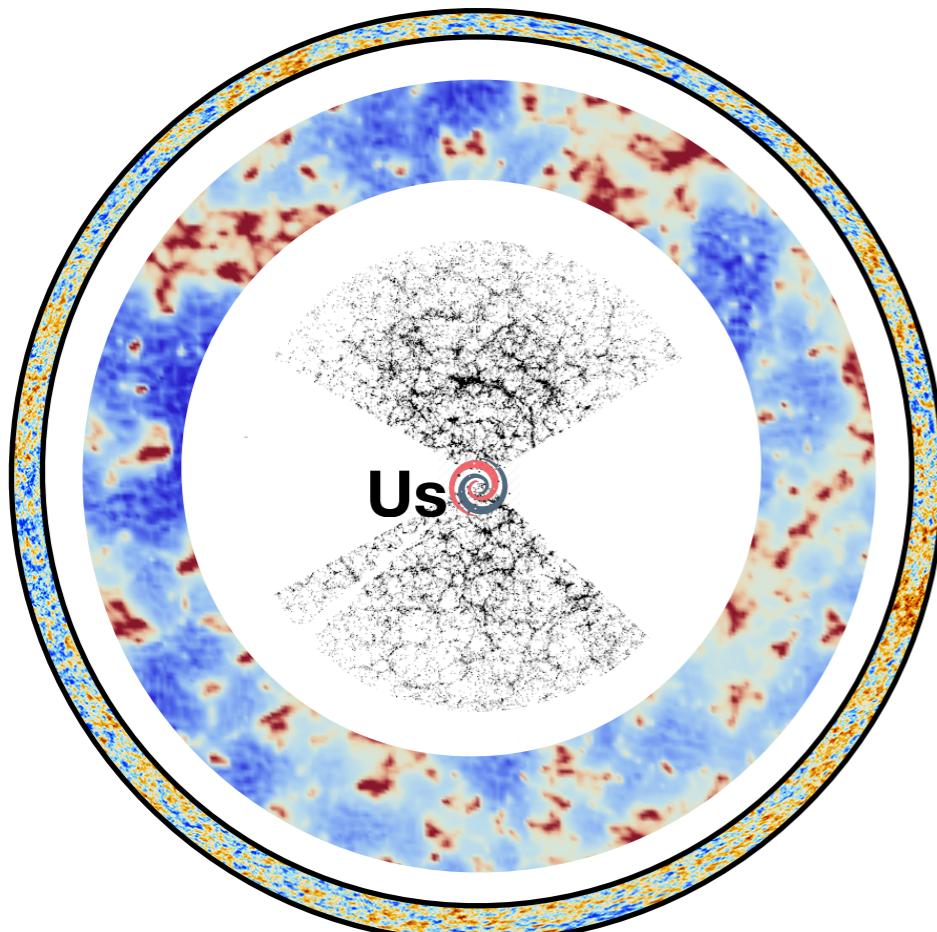
Data: SDSS/Planck



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-New lamppost for Dark Matter!

