

UNIVERSITÄT GREIFSWALD Wissen lockt. Seit 1456



Mass determination of ^{99g,m}In and ^{98m}In INTC-P-553

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64th Meeting of the INTC

Shell evolution around ¹⁰⁰Sn

- Nuclear shell model predicts shell closures (magic numbers)
- Model calculations perform well for closed shells + few nucleons in valence space
- Vicinity of doubly magic N = Z = 50 ¹⁰⁰Sn ideal case for shell model studies
- Neutron deficient In isotopes as ¹⁰⁰Sn core with single p-hole and n or nholes
- → nn- and np-interaction studies



Shell evolution around ¹⁰⁰Sn



Shell evolution around ¹⁰⁰Sn



99 10 50 1# s 1/2^{-#} 3.1 s 9/2^{+#} 50 # s 1/2^{-#} 0.1 s 9/2^{+#} 0.1 s 9/2^{+#} B⁺ = 100% B⁺ = 100% B⁺ = 100% B⁺ = 200% B⁺ = 0.9(4)% 0.4 s 9/2^{+#}

Shell evolution in odd-even ⁹⁹In

- Precision measurement of ⁹⁹In (1/2⁻) isomeric state
 - Effective single proton separation energy along isotopic chain (type-I shell evolution)
 - Excitation energy for shell model benchmarking
- Type-II shell evolution
 - Configuration space dependendent shell evolution within same nucleus

→ Will provide essential experimental input for shell model calculations





Shell evolution in odd-odd ⁹⁸In



- Effective *np*-hole interaction and pairing studies in $g_{9/2}$ orbit $\rightarrow J = 0$ to J = 9 multiplet (ordering?)
- *np* matrix element directly equivalent to excitation spectrum!
- Indirect evaluation of ground state by measuring (9⁺) excited state

→ Benchmarking of shell model calculations

Amended from J. Park *et al.,* Phys. Rev. C **99**, 034313 (2019)





Wigner effect in self-conjugate ⁹⁸In

- Excess binding energy in N = Z nuclei through pn coupling
- Accessible through lowest T = 0 and T = 1 states
 - Lowest T = 0 state can become ground state due to neighboring shell closure!
- Anomaly in ⁵⁸Cu also expected for ⁹⁸In
- Binding energy difference described through interplay of Wigner + pairing + symmetry energies

A. O. Macchiavelli, Phys. Rev. C 61, 041303(R) (2000)
Y. Y. Cheng *et al.*, Phys. Rev. C 91, 024313 (2015)



→ Test of this prediction by pinning down J = 9 excitation energy



- MR-ToF: versatile and fast
 - Mass separation mode
 - Mass spectrometry mode
 - Resolving power up to 250k
- "Contaminants" can be used as calibrants





- Newly commissioned PI-ICR technique: high precision even with low yields
- ToF between species projected on phase difference
- Shortest storage time to date: ^{101m}In with 65ms





qB M. Mougeot et al., in preparation (2020) ω $\overline{m_{ion}}$ 12 preliminary $101mIn^{+}$ Y position (mm) Center Φ_{exc} -8¹⁰¹gIn -12-8 -4 8 X position (mm)



- ToF between species projected on phase difference
- Shortest storage time to date: ^{101m}In with 65ms



Summary



Backup: Resolving Powers

MR-ToF MS:



Penning trap PI-ICR:

$$R = \frac{m}{\Delta m} \approx \frac{\nu_{+}}{\Delta \nu_{+}} \approx \pi \frac{\nu_{+} t_{obs} r_{+}}{\Delta r_{+}}$$





Backup: Type-II shell evolution in odd-even In

• Type-II shell evolution

- Configuration space dependendent shell evolution within same nucleus
- Excitation energy calculations very much depend on choice of orbital model space





- Element creation and light curves during Type-I X-ray bursts along to proton trip line influenced by waiting points cause by slow beta⁺-decay
- Light curve

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- Remaining ashes make up crust of neutron stars above neutron core
- Models sensitive to photodisintrigration rate which depends on the proton-capture Q-value
- Various key elements influence rp-process significantly (type 1 and 2)
- ⁹⁹In and ⁹⁸In as type 2 elements are important



Backup: weak interaction studies

- Test of CVC hypothesis and through this determination of V_{ud} and test of unity of first row of CKM-matrix
- Q_{ec}-value of superallowed beta decay of ⁹⁸gIn contributes to the calculation of the global Ft-value
- Very challenging measurement of **R** and $T_{1/2}$ needed, alongside Q_{ec} -value. Measurement might be one of the first steps
- Will stimulate new theoretical approaches like *ab initio* calculations to determine theoretical corrections



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