ON THE PECULIAR BEHAVIOUR OF NEUTRONS VIA NANO-STRUCTURES

Maaza@tlabs.ac.za Maazam@unisa.ac.za



OUTLINE

1-U2ACN2: 2-HISTORICAL BACKGROUND: FROM CHADWICK TO

RAUCH

3-NEUTRON OPTICS & WAVE-PARTICLE DUALITY
4-NEUTRON TUNNELLING & NEUTRON LIFETIME
5-NEUTRON TRANSMISSION THROUGH ¹⁰B !!!!

4-NEUTRON OPTICS & FORESIGHT: ESS



MULTISKILLED Human Capital DEVELOPMENT-



MULTISKILLED Human Capital DEVELOPMENT-



MULTISKILLED Human Capital DEVELOPMENT-







Contents lists available at ScienceDirect

Applied Surface Science



journal homepage: www.elsevier.com/locate/apsusc

Structural and optical properties of nano-structured tungsten-doped ZnO thin films grown by pulsed laser deposition

B.D. Ngom ^{a,b,c,*}, T. Mpahane ^c, N. Manyala ^d, O. Nemraoui ^c, U. Buttner ^e, J.B. Kana ^f, A.Y. Fasasi ^g, M. Maaza ^{a,c}, A.C. Beye ^{a,b}

^a The African Laser Centre, CSIR campus, P.O. Box 395, Pretoria, South Africa

^b Groupes de Laboratoires de physique des Solides et Sciences des Matériaux, Faculté des sciences et Techniques Université Cheikh Anta Diop de Dakar (UCAD), B.P. 25114 Dakar-Fann Dakar, Senegal

^cNANO-Sciences Laboratories, Materials Research Group, iThemba LABS, National Research Foundation, South Africa 🛸

^d Department of Physics and Electronics National University of Lesotho, Lesotho.

^e Engineering Department, University of Stellenbosch, South Africa

^fDepartment of Physique University of Yaoundé 1, Cameroon____

^g Centre for Energy Research and Development, Obafemi Awolowo University, Ile-Ife, Osun State, Nigeria –

2-HISTORICAL BACKGROUND: FROM CHADWICK VIA FERMI TO RAUCH

Postulated 1920: Rutherford
Confirmed 1923: Chadwick
Sensitive to the 4 forces
Cornerstone for the Standard model

Ideal QM tool / Wave-particle duality
Total Reflection 1946
Interferometry 1973
Size: r_n~1F [area~ 10⁻²⁵ cm²=0.1 "barn"]

2-HISTORICAL BACKGROUND: GLUON NATURE

Sensitive to the 4 forces Cornerstone of the Standard model Neutron decay



Standard model – structure of the universe

Particle physicists believe that matter – everything created in the universe by the Big Bang about 14 billion years ago – is made up of 12 types of fundamental particle and six force carriers. These building blocks, which cannot be broken down any further, make up the *Standard Model* theory



2-HISTORICAL BACKGROUND: GLUON NATURE

Sensitive to 4 forcesCornerstone of the Standard model

□Neutron decay



S.P.-2015, Pakistan 26-29/08/2015

2-HISTORICAL BACKGROUND: GLUON NATURE Cornerstone for the Standard model Lifetime: T_n=885.7 +/- 0.8 s

PARTICLE PHYSICS

Physicists close in or neutron puzzle

Researchers are narrowing down their measurements of how long the subatomic particle survives on its own.

BY ALEXANDRA WITZE IN DENVER, COLORADO

Physicists are drawing nearer to answering a long-standing mystery of the Universe: how long a neutron lives.

Neutrons are electrically neutral particles that usually combine with protons to make up atomic nuclei. Some neutrons are not bound up in atoms; these free-floating neutrons decay radioactively into other particles in minutes.

But physicists can't agree on precisely how long it takes a neutron to die. Using one laboratory approach, they measure the average neutron lifetime as 14 minutes 39 seconds. Using a different approach, they get 8 seconds longer. "We don't know why they're different," says Shannon Hoogerheide, a physicist at the National Institute of Standards and Technology (NIST) in Gaithersburg, Maryland. "We really need to understand and eliminate this discrepancy." She and other scientists debated new ways to solve the problem this month at a meeting of the American Physical Society in Denver, Colorado.

Pinpointing the lifetime of a neutron is important for understanding how much hydrogen, helium and other light elements formed in the first few minutes after the Universe was born 13.8 billion years ago. Scientists also think that pinning down the neutron's lifetime would help to constrain measurements of other subatomic particles.

One way of clocking the neutron's lifespan is to put some of the particles in a bottle and count how many are left after a period of time. This 'bottle' method has been tried at several laboratories, including the Los Alamos National Laboratory in New Mexico¹ and the Institut Laue–Langevin in Grenoble, France. On average, they come up with a neutron lifetime of 14 minutes 39 seconds.

The other way is to feed neutrons into a detector that counts the protons created as the neutrons decay. This 'beam' method has been used at NIST and at the Japan Proton Accelerator Research Complex in Tokai. The Japanese work has just begun, but the NIST team reported in 2013 that its neutrons live eight seconds longer, on average, than those in the bottle method².

That's a big problem, because the beam and bottle measurements don't overlap, even when their margins of error are taken into account. So physicists have been looking for ways to explain why neutrons might be disappearing from bottles faster than from beams.

One possibility is that one of the two methods is doing something wrong. In that case, researchers might want to combine beam



2-HISTORICAL BACKGROUND: WAVE PARTICLE DUALITY

□Ideal Quantum Mechanics tool $\lambda = h/mv$



2-HISTORICAL GROUND: WAVE NATURE WAVE PARTICLE DUALITY

□Total Reflection "Fermi, 1946", □Polarization "Gukasov, 1954", Thin Film Interference "Maier-Leibnitz, 1962", Prism Deflection "Landkammer/ Korpiun, 1966", □Neutron Interferometry "Rauch, 1973", **Supermirrors** "Mezei, 1978", □Neutron Tunneling "Refl. mode, Steyerl, 1981", □Sagnac Effect "Werner, 1985", Aharonov-Bohm Effect "Collela, 1993", Neutron Tunneling "Ref./Trans "Maaza & al, 1996", □Neutron Goos-Hanschen"de Haan & al, 2010". **Zeeman Tunneling** "Ref. "Maaza & al, 1997"

2-HISTORICAL GROUND: WAVE NATURE WAVE PARTICLE DUALITY Neutron Interferometry "H. Rauch, 1973",



NEUTRON OPTICS: THERMAL NEUTRONS

.Neutron energy Energy range .0.0-0.025 eV .0.025 eV .0.025-0.4 eV .0.4-0.5 eV .0.5-1 eV .1-10 eV .10-300 eV .300 eV-1 MeV .1-20 MeV .> 20 MeV

Cold neutrons Thermal neutrons Epithermal neutrons Cadmium neutrons Epi-Cadmium neutrons Slow neutrons Resonance neutrons Intermediate neutrons **Fast neutrons Ultrafast neutrons**

RESEARCH REACTORS: THERMAL NEUTRON .Controlled Fission + Thermalization + Moderation



3-NEUTRON OPTICS :REFRACTIVE INDEX WavePacket & Schrodinger Equation.

The de Broglie wave-particle duality [15] associates a wavevector, $\mathbf{k} = m\mathbf{v}/\hbar$, with a neutron propagating through a medium, where \mathbf{v} is the neutron velocity, m its mass, and $\hbar = 1.05457 \times 10^{-34}$ Js is the reduced Planck constant. The wavefunction can be represented as a 3D wavepacket (in Dirac notation),

$$|\Psi(t)\rangle = \int d\mathbf{k} \ \mu_{\mathbf{k}} e^{-i\omega_{\mathbf{k}}t} |\mathbf{k}\rangle, \tag{1.1}$$

with $\omega_{\mathbf{k}} = E/\hbar$, where E is the energy, $\mu_{\mathbf{k}}$ is the probability amplitude. The state $|\mathbf{k}\rangle$ is a plane wave component with position representation $\psi_{\mathbf{k}} = \langle r | \mathbf{k} \rangle = e^{i\mathbf{k}\cdot\mathbf{r}}$, with wavevector $\mathbf{k} = k_x \hat{e}_x + k_y \hat{e}_y + k_z \hat{e}_z$ and $\mathbf{r} = x \hat{e}_x + y \hat{e}_y + z \hat{e}_z$.



The propagation of a neutron through a medium is governed by the matter-wave Schrödinger equation. In a time independent potential, under steady state conditions, the Schrödinger equation in the position representation is,

$$\Big[-\frac{\hbar^2}{2m}\nabla^2 + V(\mathbf{r})\Big]\Psi(\mathbf{r}) = E\Psi(\mathbf{r}),$$

$$V(\mathbf{r}) = \sum_{j} \frac{2\pi\hbar^2}{m} b_j \delta(\mathbf{r} - \mathbf{r}_j),$$

where $V(\mathbf{r})$ is the potential energy of the particle, and E is the total energy of the p The equation can be written in a form that is similar to the Helmholtz equation in coptics as,

$$\nabla^2 \Psi(\mathbf{r}) + K(\mathbf{r})^2 \Psi(\mathbf{r}) = 0,$$

with $K(\mathbf{r}) = \sqrt{2m[E - V(\mathbf{r})]}/\hbar$ is the medium dependent wavevector. $V(\mathbf{r})$ c various forms including electromagnetic, gravitational, and nuclear. Of interest to us is the nuclear and magnetic interaction which will be explored in the next section in detail. One common way to characterize them is via the refractive index, defined as [2, 3],

$$n \equiv \frac{K(\mathbf{r})}{k} = \sqrt{\frac{2m[E - V(\mathbf{r})]}{2mE}} \simeq 1 - \frac{\overline{V}}{2E} , \qquad V(\mathbf{r}) = \sum_{j} \frac{2\pi\hbar^2}{m} b_j \delta(\mathbf{r} - \mathbf{r}_j),$$

where $k = \sqrt{2mE\hbar^{-2}}$ is the wavevector in free space. The approximation in Eq. (1.7) is valid for thermal neutrons as the potential $V(\mathbf{r})$ can be expressed in a form equal to the optical potential, \overline{V} , which for most materials is of the order $10^{-5}eV$.

$$V_O = \frac{2\pi\hbar^2}{m}N\overline{b}, \text{ where } N\overline{b} = \sum_j b_j\delta(\mathbf{r} - \mathbf{r}_j).$$
(1.9)

Nb is the local mean scattering length density, which represents the response of the overall system as multiple copies of a single atom of the same kind. Generally, \overline{b} is complex with typical values in the femtometre scale. Moreover, this value leads to a scattering cross-section defined as $\overline{\sigma}_s = 4\pi |\overline{b}|^2$ and an absorption cross-section defined as $\overline{\sigma}_a = 4\pi \text{Im}[b]^2 k^{-1}$, where $\text{Im}[\overline{b}]$ is the imaginary component.

When a neutron beam is shined on an absorbing target of effective thickness D, the intensity of transmitted neutron is related to the incident intensity, I_0 , by [19],

$$I = I_0 e^{-\sigma_a N D}. \tag{1.10}$$

Materials with high $\bar{\sigma}_a$ including lead, cadmium, gadolinium, are commonly used as neutron absorbers. For most materials, $\text{Re}[\overline{b}] > 0$, with the few exceptions of ²H, ⁴⁸Ti, and ⁶²Ni

The general form of the complex refractive index from a spin-independent scattering material is [3, 2],

$$n \equiv 1 - \frac{\lambda^2 N}{2\pi} \sqrt{\overline{b}^2 - \left(\frac{\overline{\sigma}_r}{2\lambda}\right)^2} + i \frac{\overline{\sigma}_r N \lambda}{4\pi},\tag{1.11}$$

$$V_O = \frac{2\pi\hbar^2}{m} N\overline{b}, \text{ where } N\overline{b} = \sum_j b_j \delta(\mathbf{r} - \mathbf{r}_j). \tag{1.9}$$

Nb is the local mean scattering length density, which represents the response of the overall system as multiple copies of a single atom of the same kind. Generally, \overline{b} is complex with typical values in the femtometre scale. Moreover, this value leads to a scattering cross-section defined as $\overline{\sigma}_s = 4\pi |\overline{b}|^2$ and an absorption cross-section defined as $\overline{\sigma}_a = 4\pi \text{Im}[b]^2 k^{-1}$, where $\text{Im}[\overline{b}]$ is the imaginary component.

When a neutron beam is shined on an absorbing target of effective thickness D, the intensity of transmitted neutron is related to the incident intensity, I_0 , by [19],

$$I = I_0 e^{-\sigma_a N D}. \tag{1.10}$$

Materials with high $\bar{\sigma}_a$ including lead, cadmium, gadolinium, are commonly used as neutron absorbers. For most materials, $\text{Re}[\overline{b}] > 0$, with the few exceptions of ²H, ⁴⁸Ti, and ⁶²Ni

The general form of the complex refractive index from a spin-independent scattering material is [3, 2],

$$n \equiv 1 - \frac{\lambda^2 N}{2\pi} \sqrt{\overline{b}^2 - \left(\frac{\overline{\sigma}_r}{2\lambda}\right)^2} + i \frac{\overline{\sigma}_r N \lambda}{4\pi},\tag{1.11}$$

bNi= +10.31 b⁵⁸Ni= +14.41

bTi= - 3.438 b⁴⁸Ti= - 6.08

bV= -0.3824 bMn= -3.73

bDy= -+16.92 **T**_aDy= -+994.13



Table 1. Neutron scattering lengths and cross sections of the elements and their isotopes.

Column	Symbol	Unit	Quantity
1			element
2	Ζ		atomic number
3	Α		mass number
4	ľ(p)		spin (parity) of the nuclear ground state
5	C	%	natural abundance (For radioisotopes the half-life is given instead.)
6	b	fm	bound coherent scattering length
7	b,	fm	bound incoherent scattering length
8	s,	barn ¹	bound coherent scattering cross section
9	S,	barn	bound incoherent scattering cross section
10	s,	barn	total bound scattering cross section
11	s,	barn	absorption cross section for 2200 m/s neutrons ²

VARLEY F. SEARS

AECL Research, Chalk River Laboratories Chalk River, Ontario, Canada K0J 1J0

(1) (2)	1 b E =	am = 1 25.30	00 fm² meV, <i>k</i> = 3.4	494 Å~1, / = 1.79	8 A					
_	Z	A	<i>I</i> (π)	с	b	b _i	σ	σ,	σ,	σ,
н	1				-3.7390(11)		1,7568(10)	80,26(6)	82.02(6)	0.3326(7)
		1	1/2(+)	99.985	-3.7406(11)	25.274(9)	1.7583(10)	80,27(6)	82.03(6)	0.3326(7)
		2	1(+)	0.015	6.671(4)	4.04(3)	5.592(7)	2.05(3)	7.64(3)	0.000519(7)
		3	1/2(+)	(12.32 a)	4.792(27)	-1.04(17)	2.89(3)	0.14(4)	3.03(5)	0
He	2				3.26(3)		1.34(2)	0	1.34(2)	0.00747(1)
		3	1/2(+)	0.00014	5.74(7)	-2.5(6)	4.42(10)	1.6(4)	6.0(4)	5333.(7.)
		4	0(+)	99.99986	3.26(3)	+2.566(3)/ 0	1.34(2)	0	1.34(2)	0
Li	3				-1.90(2)		0.454(10)	0.92(3)	1.37(3)	70.5(3)
		6	1(+)	7.5	2.00(11)	-1.89(10) +0.26(1)/	0.51(5)	0.46(5)	0.97(7)	940.(4.)
		7	3/2(-)	92.5	-2.22(2)	-2.49(5)	0.619(11)	0.78(3)	1.40(3)	0.0454(3)
Be	4	9	3/2()	100	7.79(1)	0.12(3)	7.63(2)	0.0018(9)	7.63(2)	0.0076(8)
в	5				5.30(4) 0.213(2)/		3.54(5)	1.70(12)	5.24(11)	767.(8.)
		10	3(+)	20.0	-0.1(3) -1.066(3)i	-4.7(3) +1.231(3) <i>i</i>	0.144(8)	3.0(4)	3.1(4)	3835.(9.)
		11	3/2(-)	80.0	6.65(4)	-1.3(2)	5.56(7)	0.21(7)	5.77(10)	0.0055(33)
с	6				6.6460(12)		5.550(2)	0.001(4)	5.551(3)	0.00350(7)
		12	0(+)	98.90	6.6511(16)	0	5.559(3)	0	5.559(3)	0.00353(7)
		13	1/2(-)	1.10	6.19(9)	-0.52(9)	4.81(14)	0.034(11)	4.84(14)	0.00137(4)

	Z	A	/(π)	c	<i>b</i> _c	b _i	σ,	σ	σ,	σ,
	22	46 47 48	0(+) 5/2(–) 0(+)	8.2 7.4 73.8	-3.438(2) 4.93(6) 3.63(12) -6.08(2)	0 3.5(2) 0	1.485(2) 3.05(7) 1.66(11) 4.65(3)	2.87(3) 0 1.5(2) 0	4.35(3) 3.05(7) 3.2(2) 4.65(3)	6.09(13) 0.59(18) 1.7(2) 7.84(25)
		49 50	7/2(-) 0(+)	5.4 5.2	1.04(5) 6.18(8)	5.1(2) 0	0.14(1) 4.80(12)	3.3(3) 0	3.4(3) 4.80(12)	2.2(3) 0.179(3)
V :	23	50 51	6(+) 7/2(-) 5/2(-)	0.250 99.750 100	-0.3824(12) 7.6(6) -0.402(2) -3.73(2)	6.35(4) 1 79(4)	0.01838(12) 7.3(1.1) 0.0203(2) 1.75(2)	5.08(6) 0.5(5) E 5.07(6) 0.40(2)	5.10(6) 7.8(1.0) 5.09(6) 2.15(3)	5.08(4) 60.(40.) 4.9(1) 13.3(2)
Fe 2	26	55 56 57	0(+) 0(+) 1/2(-)	5.8 91.7	9.45(2) 4.2(1) 9.94(3) 2.3(1)	0	11.22(5) 2.2(1) 12.42(7)	0.40(2) 0.40(11) 0 0 0.3(3) E	11.62(10) 2.2(1) 12.42(7) 1.0(3)	2.56(3) 2.25(18) 2.59(14) 2.48(30)
		58	0(+)	0.3	15.(7.)	0	28.(26.)	0	28.(26.)	1.28(5)
Ni	78	58 60 61 62	0(+) 0(+) 3/2(-) 0(+)	68.27 26.10 1.13 3.59	10.3(1) 14.4(1) 2.8(1) 7.60(6) -8.7(2)	0 0 ±3.9(3) 0	13.3(3) 26.1(4) 0.99(7) 7.26(11) 9.5(4)	5.2(4) 0 0 1.9(3) 0	18.5(3) 26.1(4) 0.99(7) 9.2(3) 9.5(4)	4.49(16) 4.6(3) 2.9(2) 2.5(8) 14.5(3)
		64	0(+)	0.91	-0.37(7)	õ	0.017(7)	õ	0.017(7)	1.52(3)

3-NEUTRON OPTICS : TOTAL REFLECTION



3-NEUTRON OPTICS :TOTAL REFLECTIONQuantum Mechanics governed phenomena Maximize the reactor capabilities via n-guides



4-NEUTRON TUNNELING: NEUTRON LIFETIME

□ 200 Å Ni-1000 Å ^{nat}V-200Å Ni /0.1 mmSilicon □ $D_r = 200$ Å Ni □ $D_s = 1000$ Å ^{nat}V □ Θ i = 0.5 deg □ Fabry-Perot structure (FPResonance Equation).



$\tanh(2\pi x_{\pm}D_r) = \tan(2\Phi_{\pm})/\tan(2\pi D_s/\Lambda)$

4-NEUTRON TUNNELING:NEUTRON LIFETIME 200 Å Ni-1000 Å ^{nat}V-200Å Ni /0.1 mmSilicon D_r =200 Å Ni □D_s=1000 Å ^{nat}V □Θi =0.5 deg



4-NEUTRON TUNNELING:NEUTRON LIFETIME □Resonance modes: 5 modes □D_r = 200 Å Ni □D_s=1000 Å ^{nat}V □Θi =0.5 deg









4-NEUTRON TUNNELING:NEUTRON LIFETIME 200 Å Ni-1000 Å ^{nat}V-200Å Ni /0.1 mmSilicon D_r =200 Å Ni □D_S=1000 Å ^{nat}V □Θi =0.5 deg



Resonance Order	λ_r (Å)	$\Delta\lambda_r$ (Å)	$k_r(10^{-3}\text{\AA}^{-1})$	$\Delta k_r (10^{-4} \text{\AA}^{-1})$	τ, (ms)
$\frac{1}{2}_{\frac{3}{4}} \tau \approx m/hk\Delta k$	17.38 12.99 10.56 9.02	1.09 0.85 0.73 0.61	5.05 6.75 8.31 9.73	3.17 4.42 5.74 6.58	0.99 0.53 0.33 0.25
5	1.01	0.01	11.15	0.04	0.10

 Reactors are limited by heat removed from core
 Pulsed sources have not reached yet that limit

High flux cold sources



 $au pprox m/hk \Delta k$

□ Cornerstone for the Standard model □ Lifetime: T_n=885.7 +/- 0.8 s

PARTICLE PHYSICS

Physicists close in on neutron puzzle

Researchers are narrowing down their measurements of how long the subatomic particle survives on its own.

BY ALEXANDRA WITZE IN DENVER, COLORADO

Physicists are drawing nearer to answering a long-standing mystery of the Universe: how long a neutron lives.

Neutrons are electrically neutral particles that usually combine with protons to make up atomic nuclei. Some neutrons are not bound up in atoms; these free-floating neutrons decay radioactively into other particles in minutes.

But physicists can't agree on precisely how long it takes a neutron to die. Using one laboratory approach, they measure the average neutron lifetime as 14 minutes 39 seconds. Using a different approach, they get 8 seconds longer. "We don't know why they're different," says Shannon Hoogerheide, a physicist at the National Institute of Standards and Technology (NIST) in Gaithersburg, Maryland. "We really need to understand and eliminate this discrepancy." She and other scientists debated new ways to solve the problem this month at a meeting of the American Physical Society in Denver, Colorado.

Pinpointing the lifetime of a neutron is important for understanding how much hydrogen, helium and other light elements formed in the first few minutes after the Universe was born 13.8 billion years ago. Scientists also think that pinning down the neutron's lifetime would help to constrain measurements of other subatomic particles.

One way of clocking the neutron's lifespan is to put some of the particles in a bottle and count how many are left after a period of time. This 'bottle' method has been tried at several laboratories, including the Los Alamos National Laboratory in New Mexico¹ and the Institut Laue–Langevin in Grenoble, France. On average, they come up with a neutron lifetime of 14 minutes 39 seconds.

The other way is to feed neutrons into a detector that counts the protons created as the neutrons decay. This 'beam' method has been used at NIST and at the Japan Proton Accelerator Research Complex in Tokai. The Japanese work has just begun, but the NIST team reported in 2013 that its neutrons live eight seconds longer, on average, than those in the bottle method².

That's a big problem, because the beam and bottle measurements don't overlap, even when their margins of error are taken into account. So physicists have been looking for ways to explain why neutrons might be disappearing from bottles faster than from beams.

One possibility is that one of the two methods is doing something wrong. In that case, researchers might want to combine beam



ZEEMAN-BEAM POLARIZATION



ISOTOPICAL EFFECT $Q_{r^{\pm}} = (l\pi - \varphi_{\pm})/D_s$



Maaza et al, MRS Advances2019

4-LITERATURE: NEUTRON LIFETIME

- C.G. Shull, Phys. Rev. 179 (1969) 752.
- [2] H. Maier-Leibnitz, T. Springer, Z. Phy. 167 (1962) 386.
- [3] F.J. Landkammer, Z. Phys. 189 (1966) 113.
- [4] H. Kurz, H. Rauch, Z. Phys. 220 (1969) 419.
- [5] H. Rauch, W. Treimer, U. Bonse, Phys. Lett. A 47 (1974) 369.
- [6] S.A. Werner, R. Collela, A.W. Overhauser, C.F. Eagen, Phys. Rev. Lett. 35 (1975) 1053.
- [7] B.P. Schoenborn, D.L.D. Caspar, O.F. Kammerer, J. Appl. Cryst. 7 (1974) 508-510.
- [8] C.F. Majkrzack, Physica (Utrecht) 136B (1986) 69–74.
- [9] F. Mezei, P.A. Dagleisch, Commun. Phys. 2 (1977) 41.
- [10] J.B. Hayter, H.A. Mook, J. Appl. Cryst. 22 (1989) 35-41.
- [11] V.F. Sears, Neutron Optics, Oxford University Press, 1989.
- [12] S.A. Werner, A.G. Klein, in: K. Sköld, D.L. Price (Eds.), Neutron Scattering, Academic Press, New York, 1986, pp. 303–318.
- [13] E. Fermi, W.H. Zinn, Phys. Rev. 70 (1946) 103.
- [14] E. Fermi, L. Marshall, Phys. Rev. 71 (1947) 666.
- [15] H. Maier-Leibnitz, T. Springer, Z. Phys. 176 (1962) 386-391.
- [16] V.F. Sears, Neutron News 3 (1992) 3.
- [17] G.P. Felcher, R.T. Kampwirth, K.E. Gray, R. Felici, Phys. Rev. Lett. 52 (1984) 1539.
- [18] V.K. Ignatovich, The Physics of Ultra Cold Neutrons, Oxford-Clarendon Press, 1990.
- [19] A. Steyerl, K.A. Steinhauser, H. Sheckenhofer, S.S. Malik, Phys. Rev. Lett. 44 (1980) 1306.
- [20] A. Steyerl, T. Ebisawa, K.A. Steinhauser, M. Utsuro, Z. Phys. B41 (1981) 283–290.
- [21] A. Steyerl, W. Drexel, S.S. Malik, E. Gutsmeidl, Physica. B 151 (1988) 36–43.
- [22] M. Maaza, L.P. Chernenko, D.A. Korneev, B. Pardo, C. Sella, F. Bridou, Phys. Lett. A 218 (1996) 312-318.
- [23] M. Maaza, B. Pardo, S. Malik, C. Sella, L.P. Chernenko, D.A. Korneev, Phys. Lett. A 195 (1994) 1-8.
- [24] M. Maaza, B. Pardo, F. Bridou, Nucl. Instr. Meths. Phys. Res. A326 (1993) 531-537.
- [25] M. Maaza, B. Pardo, Phys. Lett. A 181 (1993) 276–282.
- [26] M. Maaza, B. pardo, J.P. Chauvineau, A. Raynal, F. Bridou, A. Menelle, Phys. Lett. A 223 (1996) 145.
- [27] M. Maaza, B. pardo, J.P. Chauvineau, F. Bridou, A. Menelle, Phys. Lett. A 235 (1997) 19.
- [28] M. Maaza, B. Pardo, F. Bridou, J. App. Cryst. 26 (1993) 327.
- [29] M. Maaza, B. Pardo, Opt. Commun. 142 (1997) 84.
- [30] V.K. Ignatovich, F. Radu, Phys. Rev. B 64 (2001) 13. 205408.
- 1211 E Padu U Zabel Neutron News 17 (2) (2006) 20-22

4-LITERATURE: NEUTRON LIFETIME

- S.A. Werner, R.Collela, A.W.Overhauser, C.F.Eagen, Phys.Rev.Lett.35 (1975) 1053. [2] S.A. Werner, A.G. Klein, in Neutron Scattering, Eds K. Skold, D.L. Price, Methods of Experimental Physics, Vol.23 (Academic Press, New York, (1986) p. 303 [3] C.G. Shull, Phys. Rev. 179 (1969) 752. [4] H. Rauch, Neutron Matter Quantum Optics, Foundations Physics (2012), 42, p.760-777 [5] H. Maier-Leibnitz, T. Springer, Z. Phys. 167 (1962) 386. [6] E. Fermi, W.H. Zinn, Phys. Rev. 70 (1946) 103; [7] F.J. Landkammer, Z. Phys. 189 (1966) 113. [8] V.F Sears, Oxford, Can. J. Phys. 56, (1978) 1261 [9] H. Kurz, H. Rauch, Z. Phys. 220 (1969) 419 [10] H. Rauch, W. Treimer, U. Bonse, Phys. Lett. A 47 (1974) [11] H. Rauch, G. Badurek, A. Zeilinger, W. Bauspiess, U.Bonse, Phys. Rev. D 14 (1977) 1177 [12] H. Maier-Leibnitz, T. Springer, J. Nucl. Energy 17 (1963) 217 [13] A. Steyerl, T. Ebisawa, K.A. Steinhauser, Z. Utsuro, Z. Phys. B 41 (1981) 283 [14] H.Rauch, S.A. Werner, Neutron Interferometry. Clarendon, Oxford (2000) [15] M. Maaza, B. Pardo, S. Malik, C. Sella, L.P. Chernenko, D.A. Korneev, Phys. Lett. A 218 (1996) 312-318 [16] M. Maaza, B. Pardo, F. Bridou, Nucl. Instr. Meths. Phys. Res. A326 (1993) 531-537 [17] M. Maaza, B. Pardo, Phys. Lett. A 181 (1993) 276–282 [18] M. Maaza, B. Pardo, J.P. Chauvineau, A. Raynal, F. Bridou, A. Menelle, Phys. Lett. A 223 (1996) 145 [19] M. Maaza, B. Pardo, J.P. Chauvineau, F. Bridou, A. Menelle, Phys. Lett. A 235 (1997) 19 [20] M. Maaza, B. Pardo, F. Bridou, J. App. Cryst. 26 (1993) 327 [21] M. Maaza, B. Pardo, Opt. Commun. 142 (1997) 84 [22] P. Croce, B. Pardo, Nouv. Rev. Opt. Appl. 1 (4) (1970) 229 [23] K.A. Steinhauser, A. Steyerl, H. Schekenhofer, S.S. Malik, Phys. Rev. Lett. 44 (1980) 1306 [24] Y.P. Feng, C.F. Majkrzack, SK. Sinha, D.G. Wiesler, H.Zhang and H.W. Deman. Phys. Rev. B 49 (1994) 10814. [25] S.K. Sinha. M. Tolan, G. Vacca, Z. Li. M. Rafailovich, J.Sokolov, H. Lorenz, J.P. Kotthaus. Y.P. Feng, Cl. Guibel and D. Abernathy, Proc. MRS Symp. On Molecules in confined geometries, eds. J.M. Drake, Y. Klopfter and R. Koppelman (1995).
- [26] J.J. Norton, E.J. Kramer, R.A.L. Jones, F.S. Bates, H.R. Brown, G.P. Felcher, R. Kleb, J. Phys. II 4 (1994) 367–376.

4-LITERATURE: NEUTRON LIFETIME







Most important neutron absorbers: ¹⁰B as ¹⁰B₄C in control rods, or Boric acid as a coolant water additive in PWRs.
 Other important neutron absorbers that are used in nuclear reactors Xe, Cd, Hf, Gd, Co, Sm,, Dy, all of which usually consist of mixtures of various isotopes—some of which are excellent neutron-absorbers. These also occur in combinations such as Mo₂B₅, HfB₂, TiB₂,

⊛_i≈1 Deg



Absorbing slab

Isotope	absorption cross- section - [barns]	yield as fission product and by product decay [%]	half-life
Xe-135	2,700,000	6.4	9.1 hours
Gd-157	250,000	0.01	stable
Gd-155	61,000	0.08	stable
Sm-149	42,000	1.1	stable
Cd-113	20,000	0.01	nearly stable
Sm-151	10,000	0.6	90 years
Eu-151	7,700	0.6	nearly stable
B-10	3,800	< 0.0001	stable

R=0%, T=0% A=100% Neutron capture

		Thermal	cross sectio	n (barn) 🚺	- East cr	ds. <u>pection</u> (b	H.		Reflectivity (R) = I_{μ}/I_{μ}
		Scattering	Capture	Fission	Scattering	Capture	Fist		l,
	<u>¹H</u>	20	0.2	-	4	0.00004	-	Specular	
Moderator	² H	4	0.0003	-	3	0.000007	-	$\theta_i = \theta_r$	1
	¹² C	5	0.002	-	2	0.00001	-		
	¹⁹⁷ Au	8.2	98.7	-	4	0.08	-		
	⁹⁰ Zr	5	0.006	-	5	0.006	-		
Structural	⁵⁶ Fe	10	2	-	20	0.003	-		
materials,	⁵² Cr	3	0.5	-	3	0.002	-		
Uners	⁵⁹ Co	6	37.2	-	4	0.006	-		RAC:
	⁵⁸ Ni	20	3	-	3	0.008	-		D4C.
	<u>16</u> O	4	0.0001	-	3	0.0000003	-		
	¹⁰ B	2	200	-	2	0.4	-		nat B:
Absorber	¹¹³ Cd	100	30,000	-	4	0.05	-		20% 10R
Aboorbor	¹³⁵ Xe	400,000	2,000,000	-	5	0.0008	-		
	115ln	2	100	-	4	0.02	-		80% ''B
Fuel	²³⁵ U	10	99	583 ^[5]	4	0.09	1		Neutron capture
	238U	9	2	0.00002	5	0.07	0.3	3	
	²³⁹ Pu	8	269	748	5	0.05	2		

Neutron 0.2 barn



¹⁰B 200 barns



D_{Ti}= 12 nm D_{B4C}= 2 nm (a)9 bilayers (b)25 (c)50

Содородные болого с напалания Пара оженено областия. Спатка народство сла продека призак ожененое областия А25-53.

Преимурарство водородных бомб не реалкрустоя без достаточного скатие легиото крарного горочаго (а 10-20 да). Знартие объязных водначатых выдаста на достаточно для осущоствления такого сжатия. Многочисланные нацелянны, проводнамосо а КС-11 в 1951-35 гг. не правали к реализмо пробламы содания мощных, акономичных и транспортабляных водородных болаб, объязнатия в.В.

При габаритах порядка РДС-6с такие водородные бомбы оказываются дажа не более окономичными, чем атомные бомбы на принципа целиюй ровкции.

Для получения высоких о использовать совершенно нов обжатие», т.е. сжатие водород гии ворыее эспоногательной а

В 1951, 1952 и 1954 гг. в Сі питанні акрародової банб, в то была настипна болба модиность 1954 г. была испитана транот ностью 14 мля тонк. Столь ани нацевну мненяю, с большой казыствуют о применення атон

Раснат мощности опытис проведенный на электронной лемии прикладной математики 1.9 млн тома



Zeldovich-Vinogradov equation (X-Rays)

$$\Gamma = \frac{\Delta I_{\rm H}}{\Delta I_{\rm H} + \Delta I_{\rm L}} = \frac{\Delta I_{\rm H}}{d} \qquad (4.7)$$
$$\tan(\pi \Gamma_{\rm opt}) = \pi \left[\Gamma_{\rm opt} + \frac{\beta_{\rm L}}{\beta_{\rm H} - \beta_{\rm L}} \right] \qquad (4.8)$$

(Vinogradov and Zeldovich, 1977) (also see Borrmann, 1941)



- Sharp interfaces needed for scattering
- Thin high-Z layer to minimize absorption
- Low-Z layer best as a "spacer"



Highly Absorbing layer: B4C

Weakly Absorbing Transparent layer:

D_{TI}= 12 nm D_{B4C}= 2 nm (a)9 bilayers (b)25 (c)50

tiah-

Zeldovich-Vinogradov equation (X-Rays)

⊛_i≈1 Deg



D_{eff}≈D_{B4C}/sinΘ

D_{eff}≈200nm≈10⁵F



Highly Absorbing layer: <mark>B4C</mark>

Weakly Absorbing Transparent layer:

Tĭ

D_{TI}= 12 nm D_{B4C}= 2 nm (a)9 bilayers (b)25 (c)50

1st order Bragg Peak Periodicity interference R=12.4 %



Kiessig Fringes: Air-ML & ML-Substrate interference



Highly Absorbing layer: B4C

Weakly Absorbing Transparen layer:

D_{TI}= 12 nm D_{B4C}= 2 nm (a)9 bilayers (b)25 (c)50

5-NEUT TUNNELING:TRANS BORON GATES 1st order Bragg Peak: R=12.4% (a),60.6%(b), 73.5%(c)





Highly Absorbing layer: B4C

Weakly Absorbing Transparen layer: Ti

D_{TI}= 12 nm D_{B4C}= 2 nm (a)9 bilayers (b)25 (c)50

5-NEUT TUNNELING:TRANS BORON GATES



Monochromation and apodization with $Ti-B_4C$ multilayers in neutron optics[†]

M. Maaza^{a, b, *}, A. Menelle^a, J.P. Chauvineau^c, B. Pardo^c, A. Raynal^c, F. Bridou^c, C. Sella^d, T. Megademini^c

 ^a Laboratoire Léon Brillouin, Commissariat à l'Energie Atomique-Centre National de la Recherche Scientifique, Bat. 563, Centre d'Etudes Nucléaires de Saclay 91191, Gif-sur- Yvette, France
 ^b Atominstitut der Österreichischen Universitäten, Schüttelstrasse 115, A-1020 Wien, Austria
 ^c Institut d'Optique Théorique et Appliquée, Université de Paris-Sud, Centre d'Orsay, Bat. 503, 91405 Orsay, France
 ^a Laboratoire de Physique des Matériaux, Centre National de la Recherche Scientifique, 1 place Aristide Briand, 92195 Meudon-Bellevue, France

^e Départment de Physique, Université des Sciences et Techniques de Masuku, B.P. 595, Franceville, Gabon



- B.P. Schoenborn, C.L.D. Caspar and O.F. Kammerer, J. Appl. Crystallogr. 7 (1974) 508-510.
- [2] C.F. Majkrzack, Appl. Optics 23 (1984) 3524–3528.
- [3] F. Mezei and P. A. Dagleish, Commun. Phys. 2 (1977) 41-43.
- [4] O. Schaerpf, Physica B 156-157 (1989) 631-638.
- [5] J.B. Hayter and H.A. Mook, J. Appl. Crystallogr. 22 (1989) 35-41.
- [6] T. Ebisawa, T. Akiyoshi, N. Achiwa, S. Yamada and S. Okamoto, Annu. Rep. Res. Reactor Inst. Kyoto Univ 14 (1981) 10-18.
- [7] H. Kiessig, Ann. Phys. 10 (1931) 715–723.
- [8] M. Mâaza, Z. Jiang, F. Samuel, B. Farnoux and B. Vidal, J. Appl. Crystallogr. 25 (1992) 789-796.
- [9] A.M. Afanas'ev and Yu. Kagan, Zh. Eksp. Teor. Fiz. 48 (1965) 327-335.
- [10] G. Bormann, Z. Phys. 42 (1941) 157-162.
- [11] G. Bormann, Z. Phys. 127 (1950) 297-305.
- [12] M. Laue, Acta. Crystallogr. 2 (1949) 106-113.
- [13] H. Rauch and D. Tuppinger, Z. Phys. A 322 (1985) 427-432.
- [15] M. Mâaza, V. Mikerov, B. Pardo, J.P. Chauvineau, A. Menelle, J. Raynal, F. Bridou and T. Megademini, J. Appl. Crystallogr., submitted.
- [16] A. Menelle, M. Mâaza, B. Pardo, J.P. Chauvineau, J. Raynal, F. Bridou and T. Megademini, J. Appl. Crystallogr., submitted.
- [17] D. Schwahn, A. Miksôvsky, H. Rauch, E. Seidl and G. Zugarek, Nucl. Instr. Meth. A 239 (1985) 229-234.
- [18] J.P. Chauvineau, Rev. Phys. Appl. 23 (1988) 1645–1652.
- [19] B. Pardo, J.M. Andre and T. Megademini, Rev. Phys. Appl. 23 (1988) 1579-1597.
- [20] F. Abelès, Ann. Phys. (Paris) 5 (1950) 777-795.
- [21] T. Megademini, Thesis of Paris XI University (1990).
- [22] A.V. Vinogradov and B.Y. Zeldovich, Appl. Opt. 16 (1977) 89–103.

THANK YOU, NDIALIVUHA CHUKRAN, MERCI, DANKE



S.P.-2015, Pakistan 26-29/08/2015

Acknowledgments

- UNESCO & IAEA.
- Abdus Salam International Centre for Theoretical Physics, Trieste-Italy.
- Department of Science & Technology of South Africa, Pretoria-South Africa.
- Ministry of Foreign Affairs, Roma-Italy.
- Science & Technology Directorate, French Embassy, Pretoria-South Africa.
- ELETTRA Synchrotron Facility, Trieste-Italy
- African Laser Centre, Pretoria-South Africa.
- African Union-Science & Technology Commission, Addis Ababa-Ethiopia.
- National Research Foundation of South Africa, Pretoria-South Africa.
- Academy of Sciences for the Developing World, Trieste-Italy.
- Organization of Women in Science for the Developing World, Trieste-Italy.
- International Centre for Science & Technology-UNIDO, Trieste-Italy.
- Centre National pour la Recherche Scientifique, Paris-France.
- The EU-FP7 ICPCNANONET, Brussels-Belgium.
- The National Institute for Materials Sciences NIMS, Tsukuba-Japan.
- Nelson Mandela African University of Science & Technology, Abuja-Nigeria.
- I' Oreal-UNESCO Foundation, Paris-France.
- University of South Africa.
- Islamic Academy of Sciences, Amman-Jordan.
- iThemba LABS, Western Cape-South Africa.







S.P.-2015, Pakistan 26-29/08/2015