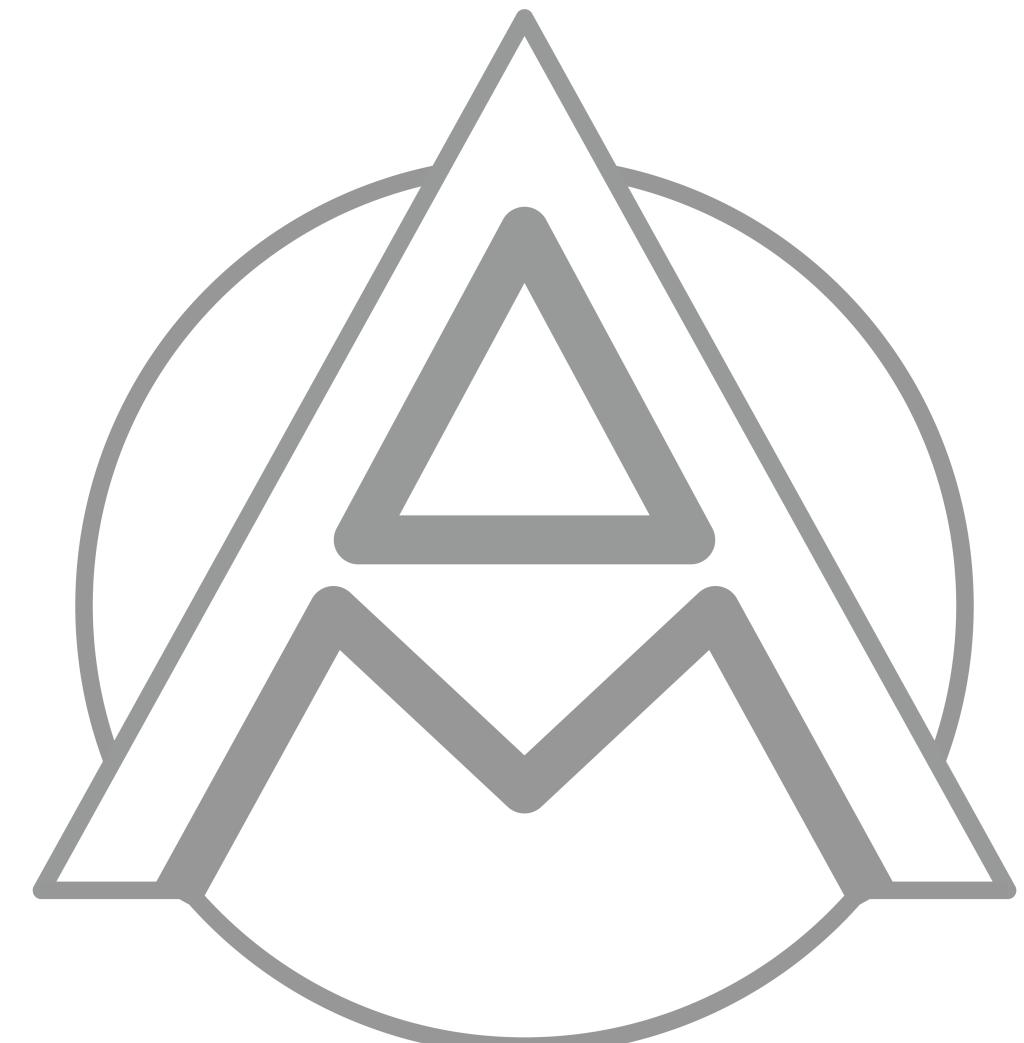


# Revitalising axion-longitudinal photon conversion

arXiv:2005.00078, arXiv:2006.10415

Ciaran O'Hare, Andrea Caputo, Alex Millar, Edoardo Vitagliano



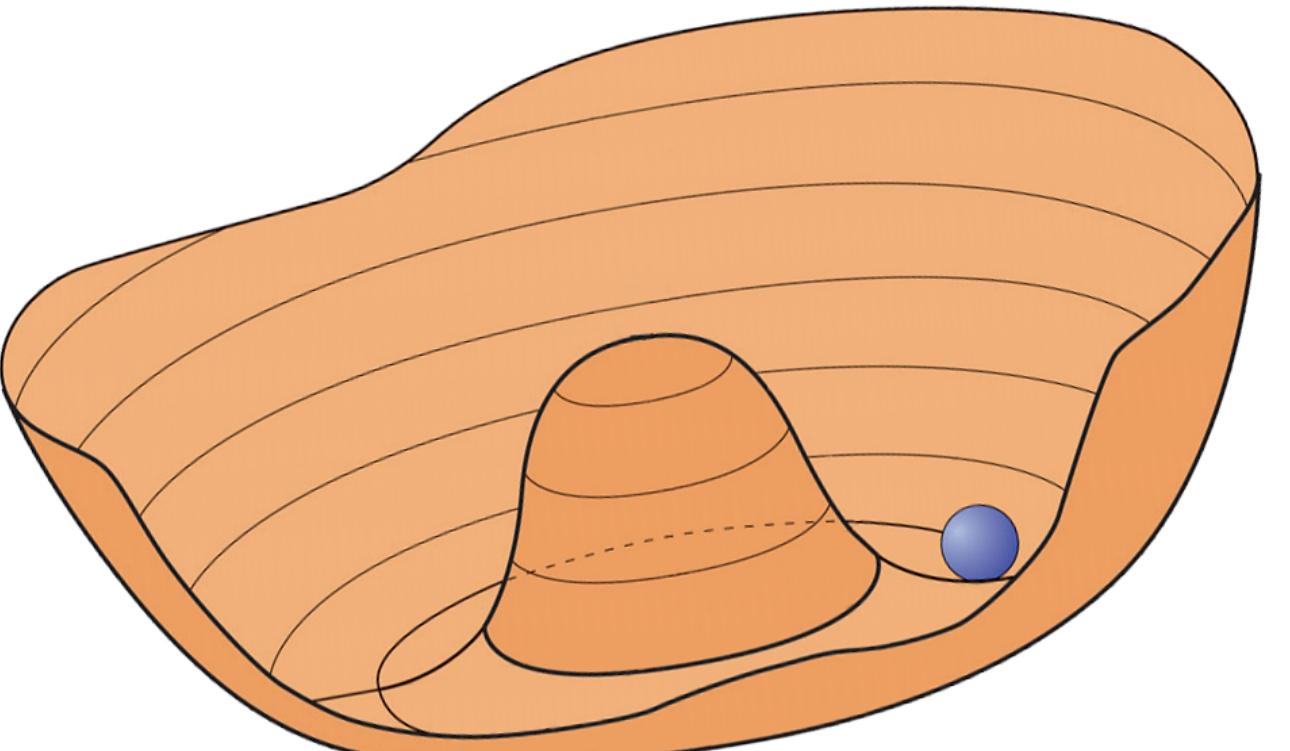
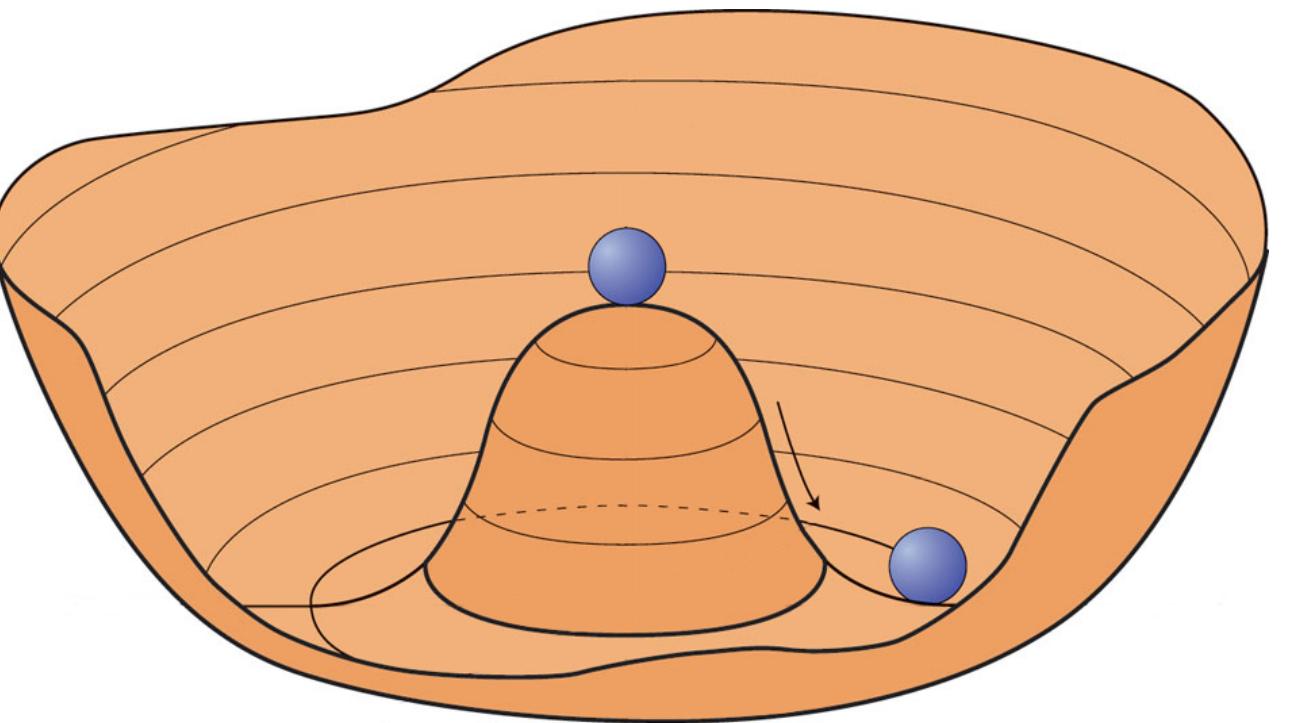
# Axions

Strong CP problem: SM has a CP violating angle  $\theta \frac{g^2}{32\pi^2} G\tilde{G}$

Experimental limit:  $\theta \lesssim 5 \times 10^{-9}^\circ$

Axion Solution:

$$\theta \rightarrow a(\mathbf{x}, t)$$



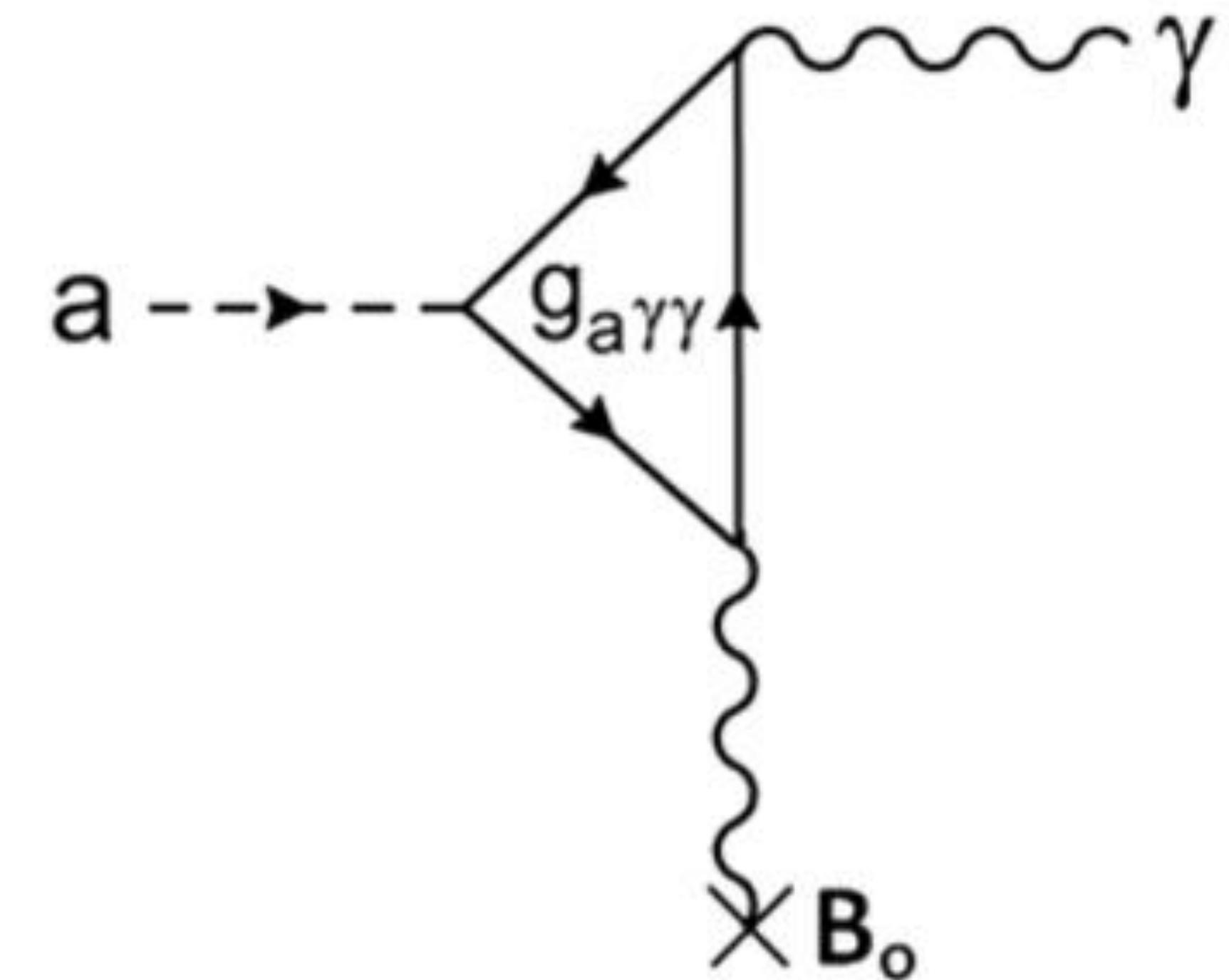
# Axions

Also provides a good DM candidate

Axions and ALPs interact with photons

$$\mathcal{L} = -\frac{1}{4}F_{\mu\nu}F^{\mu\nu} - J^\mu A_\mu + \frac{1}{2}\partial_\mu a\partial^\mu a - \frac{1}{2}m_a^2 a^2 - \boxed{\frac{g_{a\gamma}}{4}F_{\mu\nu}\tilde{F}^{\mu\nu}a},$$

Many searches use axion-transverse  
photon conversion in B-field

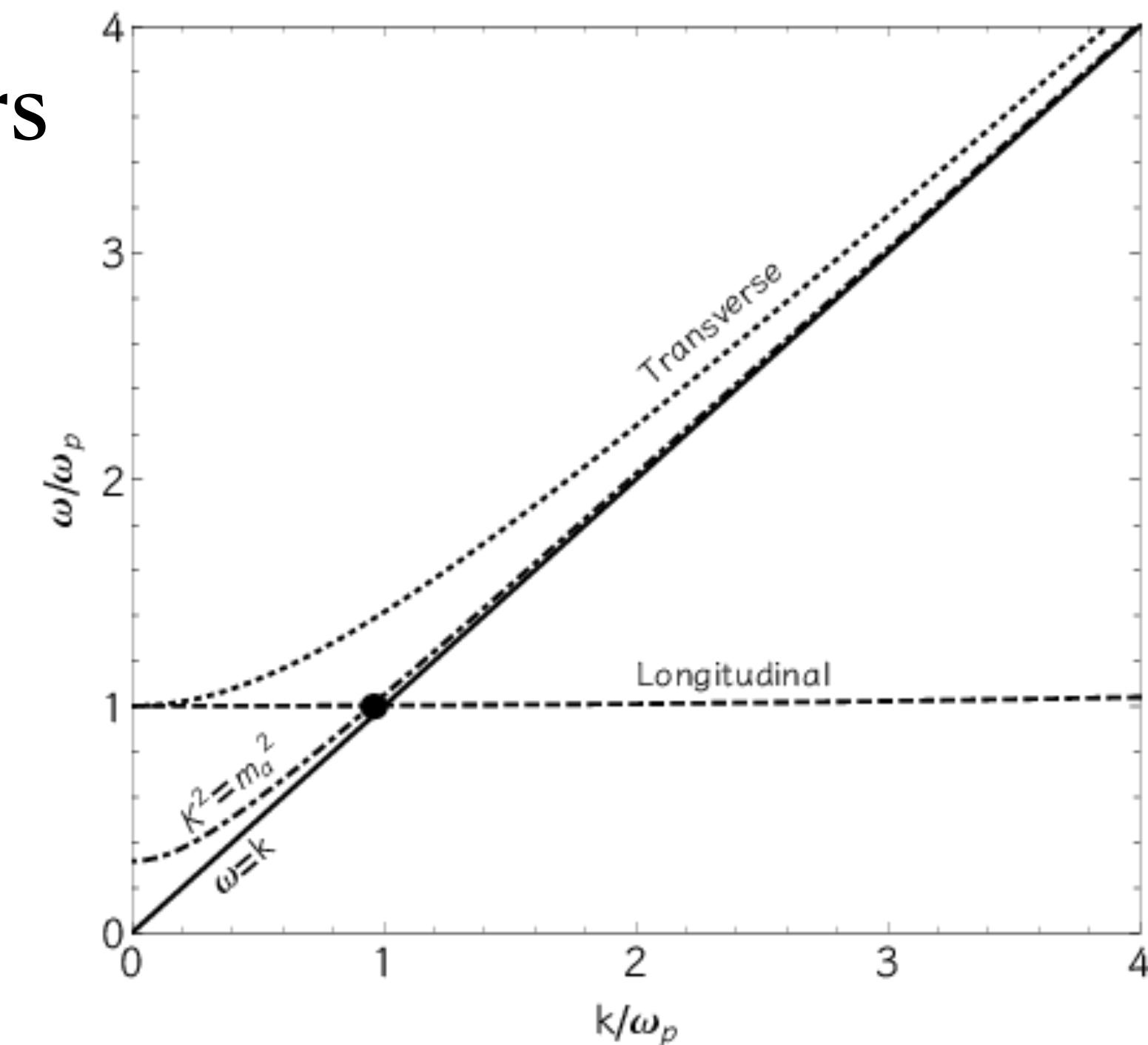


# Axion-L plasmon conversion

Axions can also resonantly mix with longitudinal plasmons in a magnetic field

Not entirely new, but overlooked for many years  
(9803486, 0410006, 0810.4380, 1807.06828)

Needs: a plasma and B-field

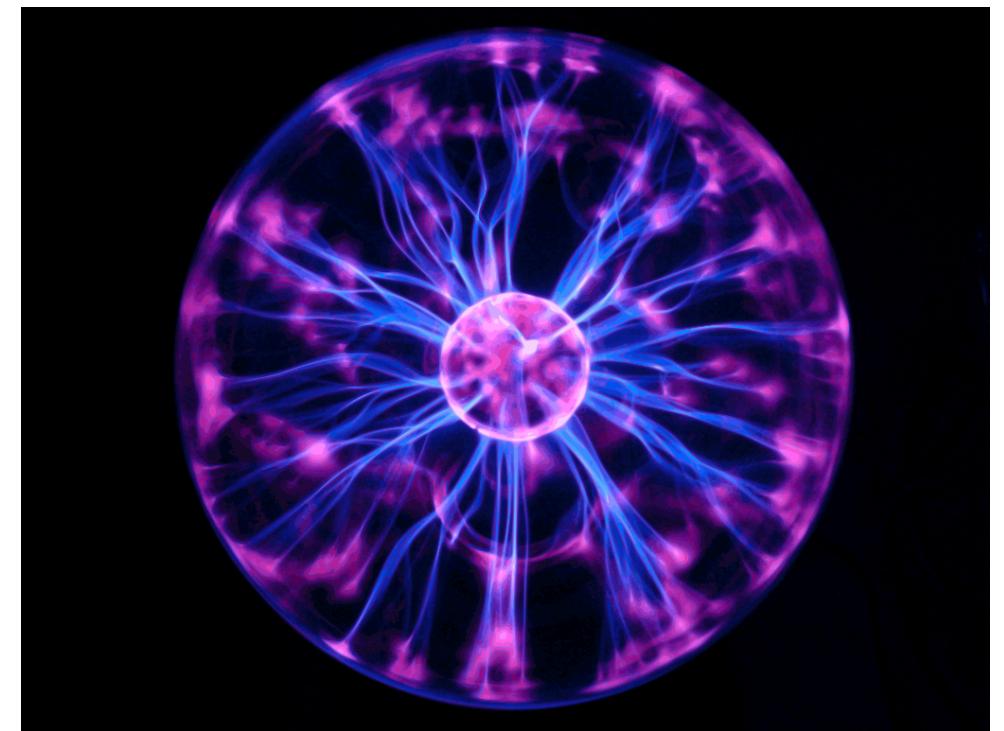


# Conversion Environments

**Photon** → **Axion**

High T, High B

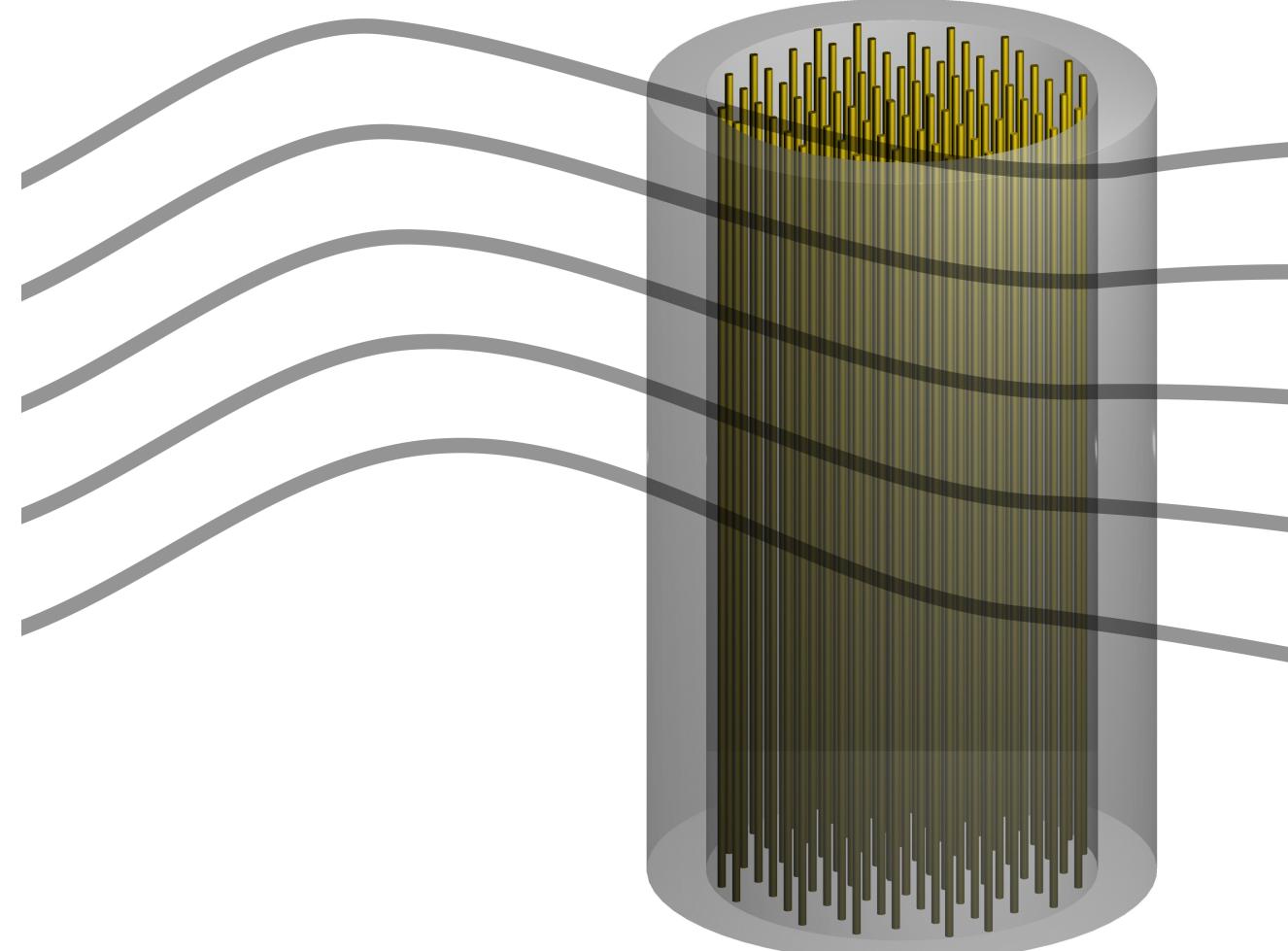
The Sun, WD, NS, plasma shining through wall (1710.01906, 1901.05910)



**Axion** → **Photon**

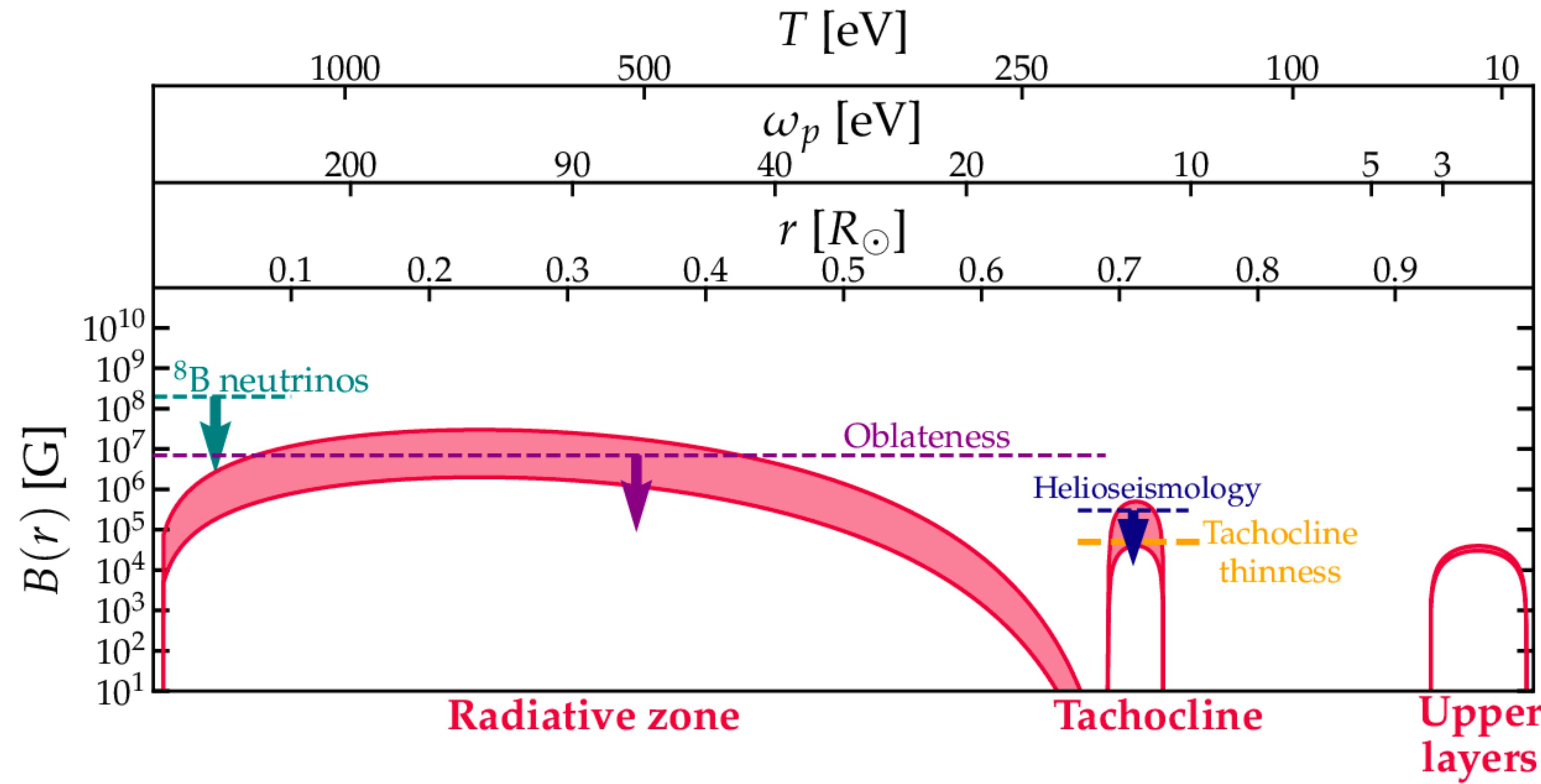
Low T, High B, Lowish  $\omega_p$

NS, plasma haloscopes (1904.11872, 2006.06836)...



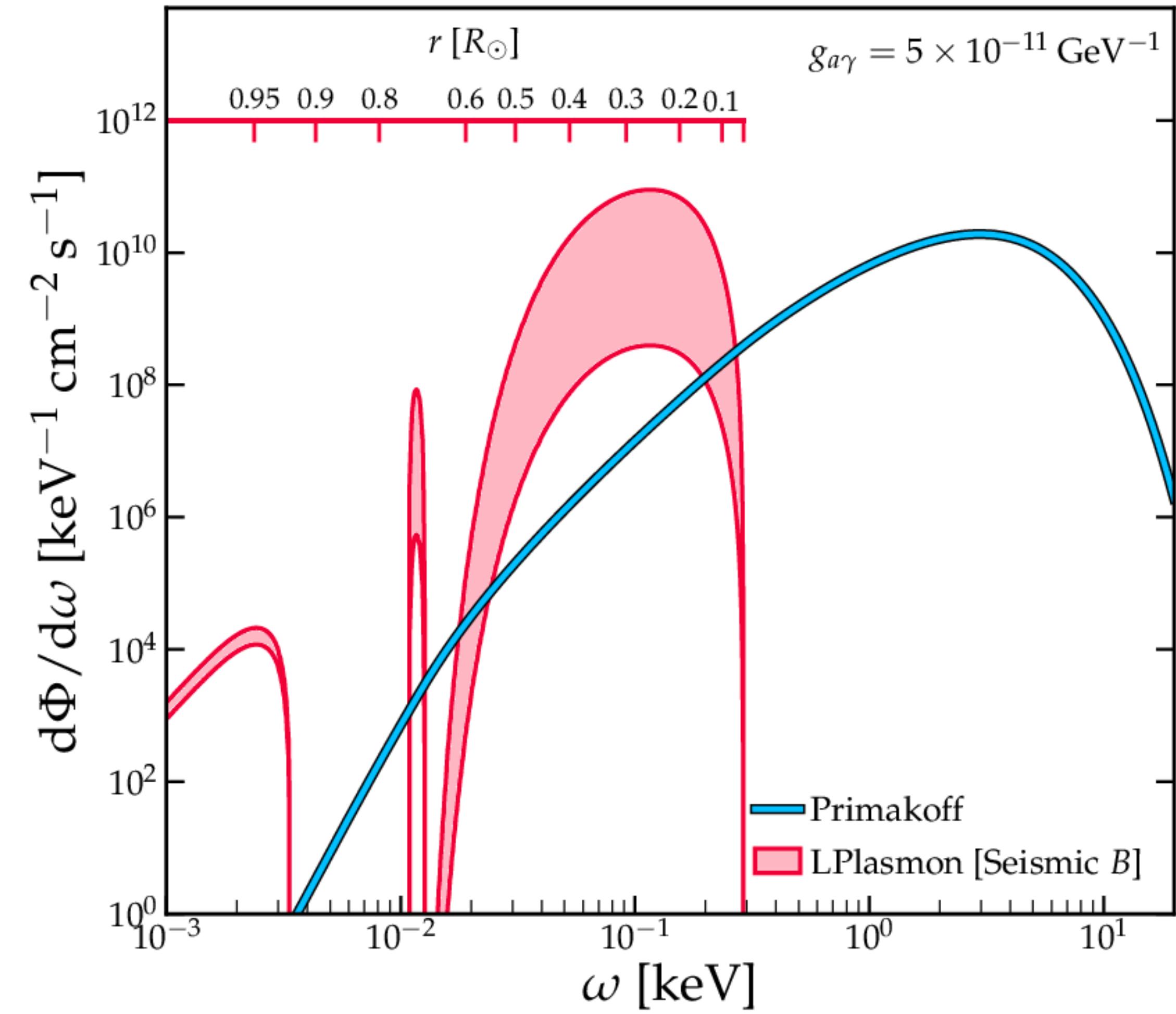
# Solar magnetic fields

- Very hard to access the inner regions of the Sun: weak bounds



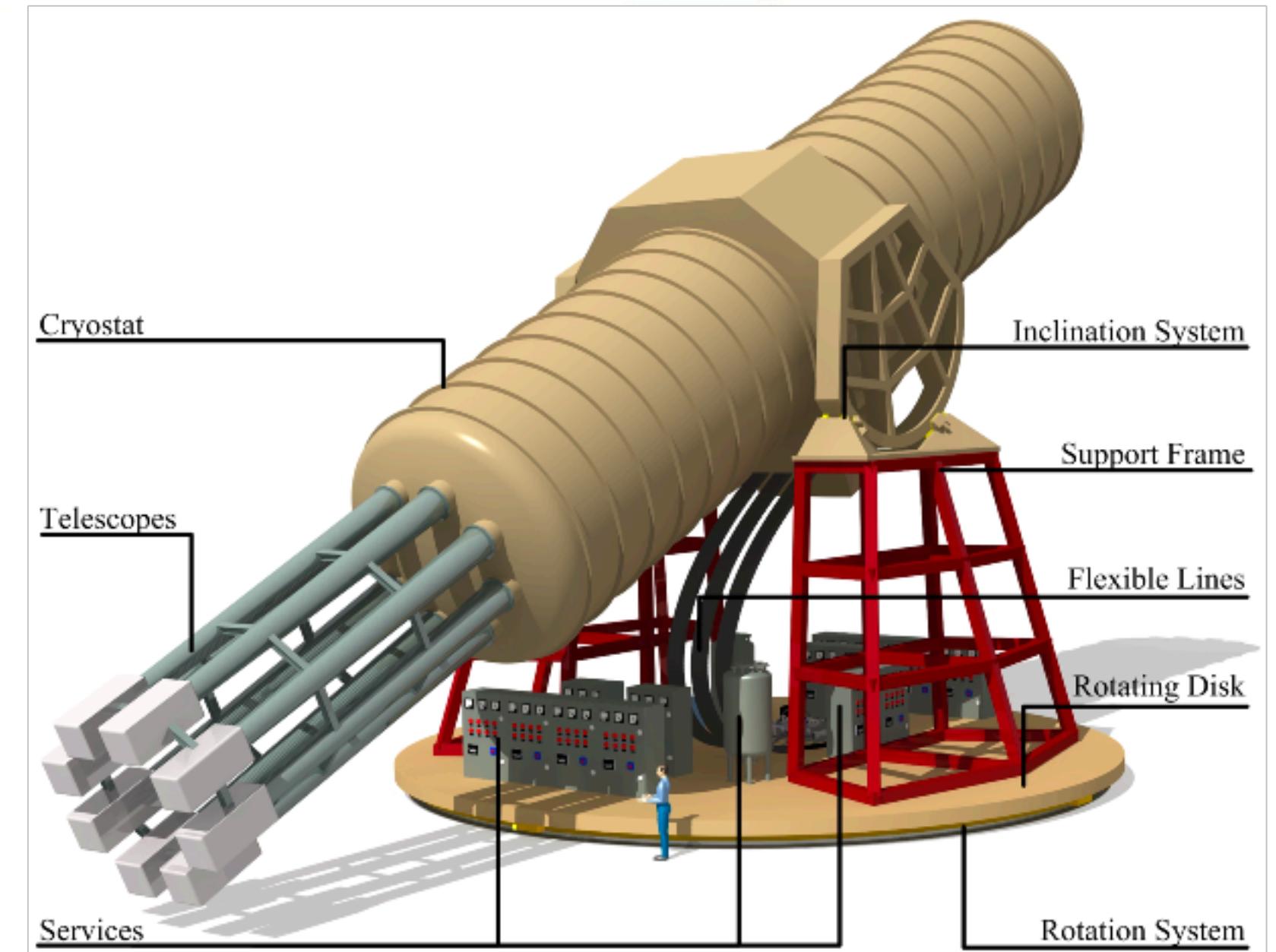
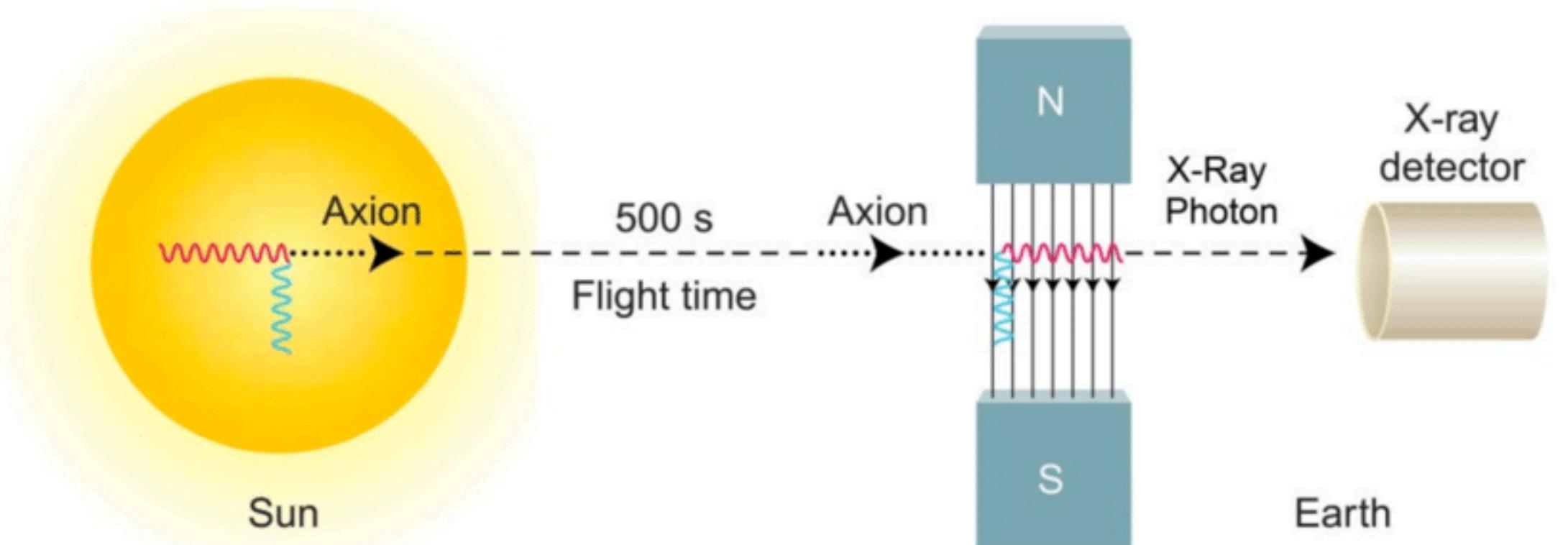
# Axion-L plasmon conversion in the sun

- Axions are produced at the plasma frequency at a given solar radius
- Flux proportional to the square of the B-field
- Ideal for magnetometry (requires at least one axion...)



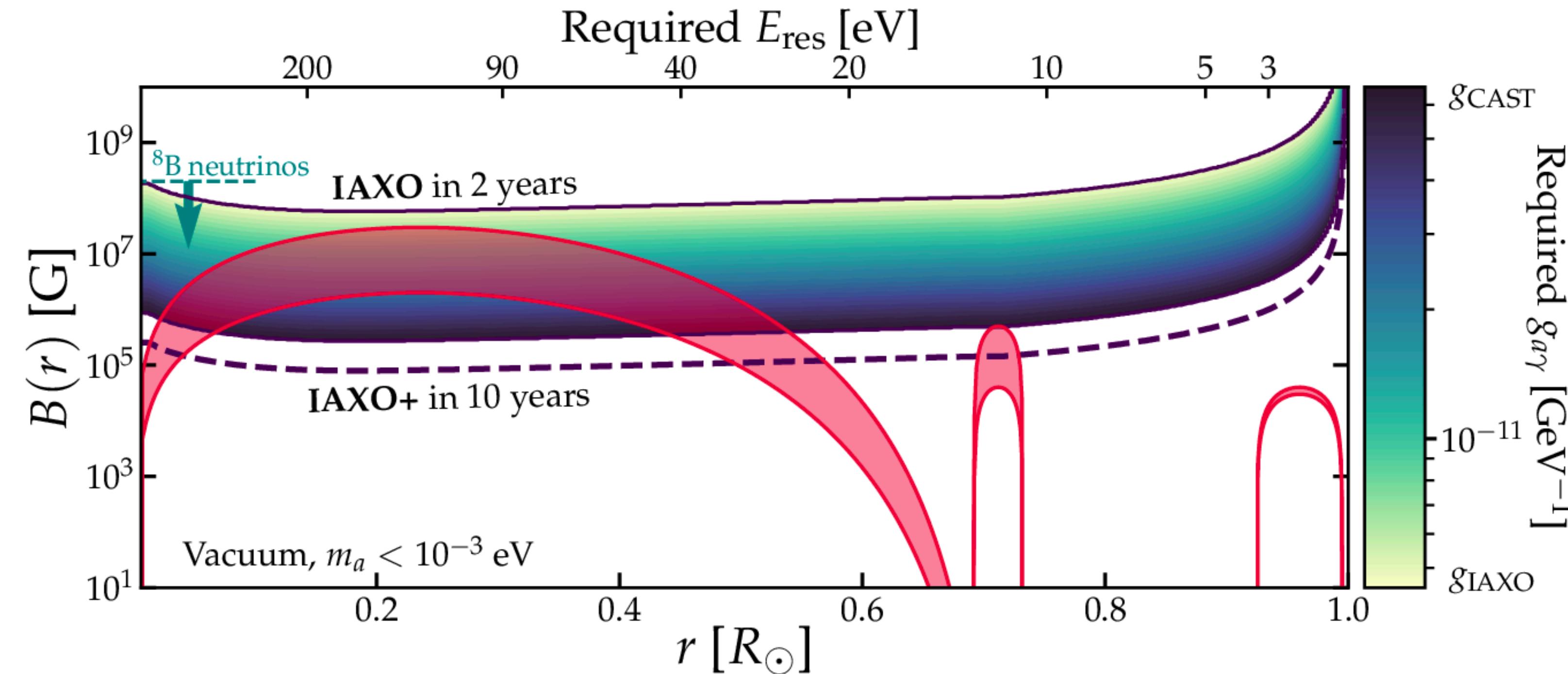
# IAXO: next generation helioscope

- Converts axions into photons with a B-field
- Previous experiment used an old LHC magnet
- IAXO a proposed (much, much bigger) axion observatory
- Could you use it to map the solar B-field?



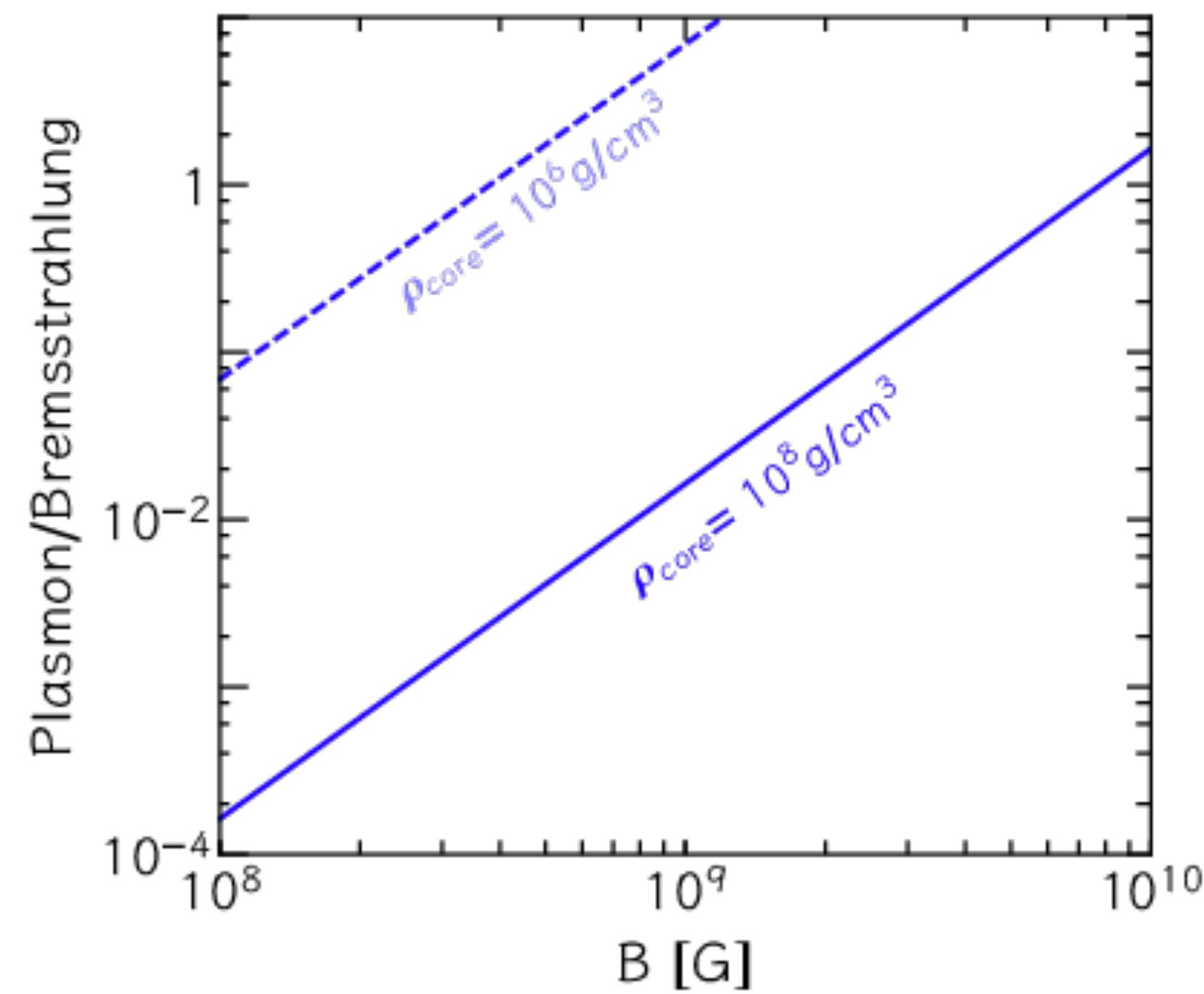
# IAXO as a magnetometer

- LP flux is in the low energy end of the spectrum, with less total energy
- IAXO's sensitivity depends on the energy resolution, which is not yet determined, but between 200 and 2 eV are being considered

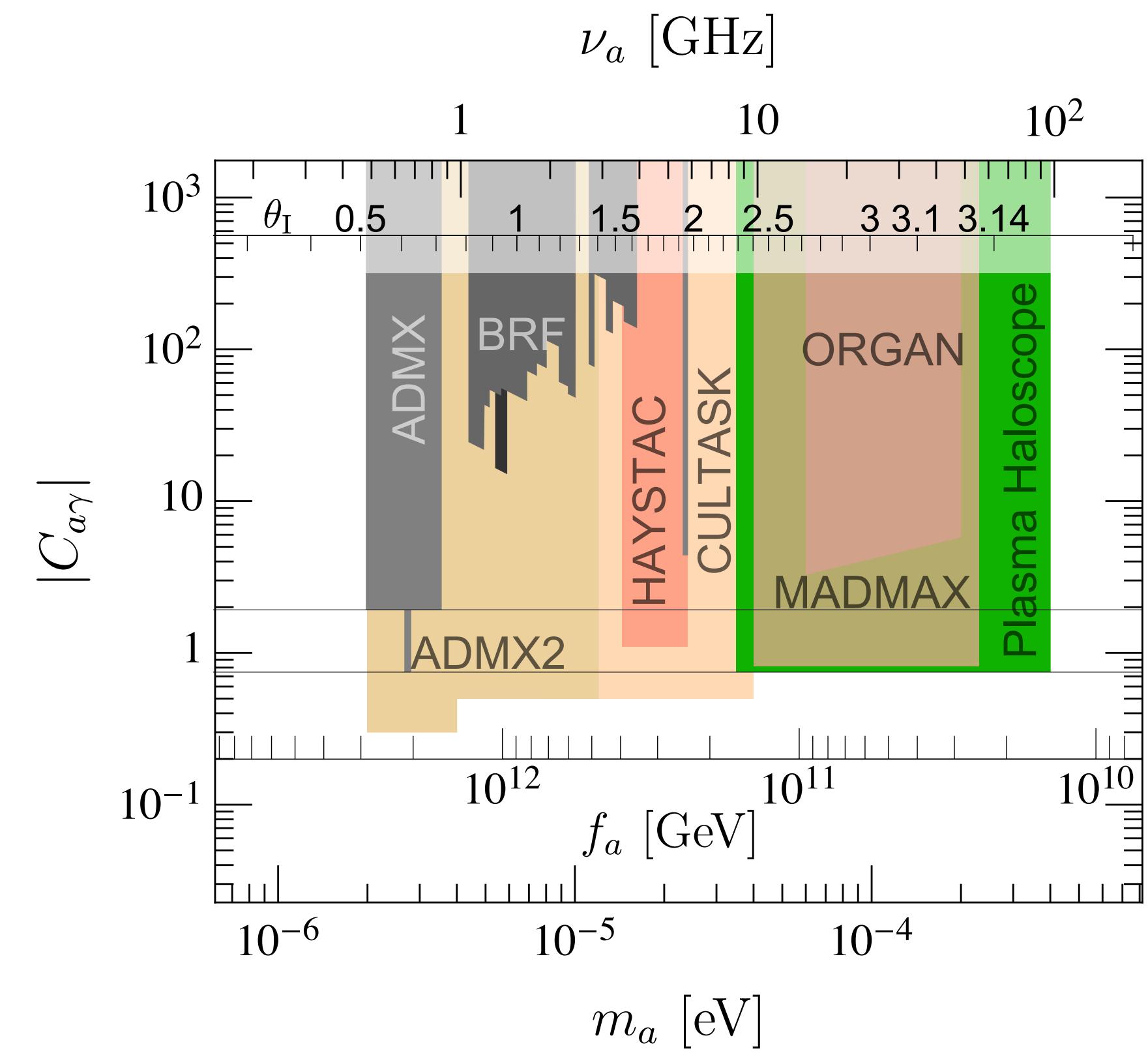


# Other environments

WDs can give X-ray signals



Tuneable plasmas can search for DM



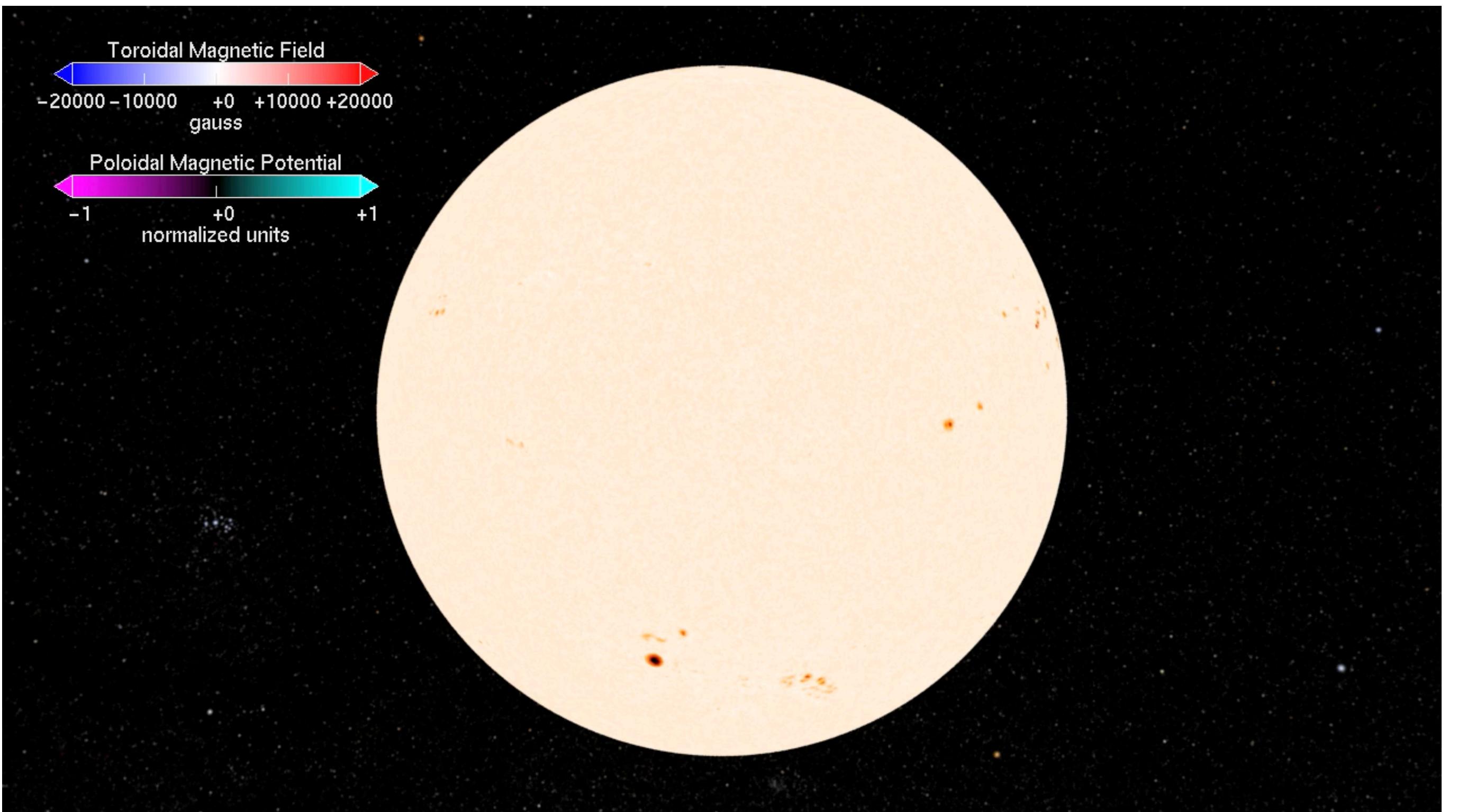
# Conclusions

- Axion-L plasmon conversion very relevant for astrophysical and lab environments
- Axions could provide the best probe yet of the solar magnetic field, particularly in the inner regions
- Relevant also to the new plasma haloscope proposal (arXiv:1904.11872)
- Part of the new AxionDM research environment at SU

# Solar magnetic fields

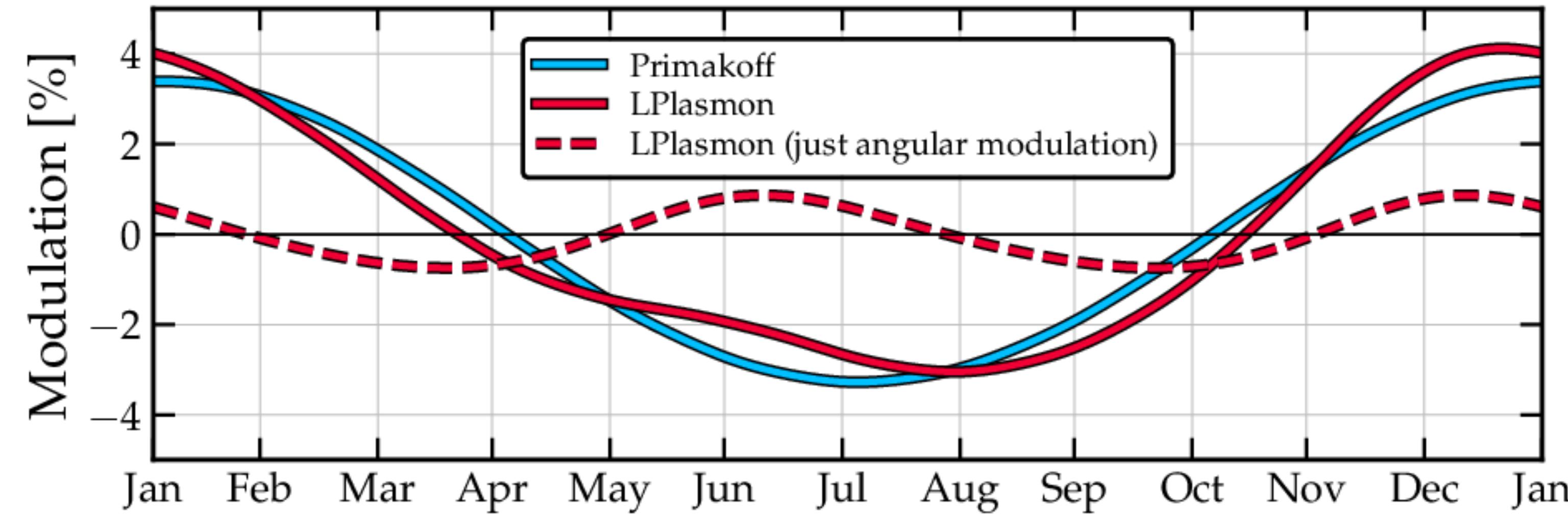
The solar B-field is mostly toroidal:

Axions preferentially emitted in the equatorial plane of the sun (luckily, where the Earth sits)



# Modulation

- Annual and biannual modulations of the signal may allow the orientation of the B-field to be determined
- Annual modulation: Earth-Sun distance varies
- Biannual modulation: Earth moves in and out of the equatorial plane



# IAXO as a magnetometer

