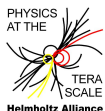


The Alignment of the CMS Silicon Tracker

Gero Flucke

(on behalf of the CMS Collaboration)

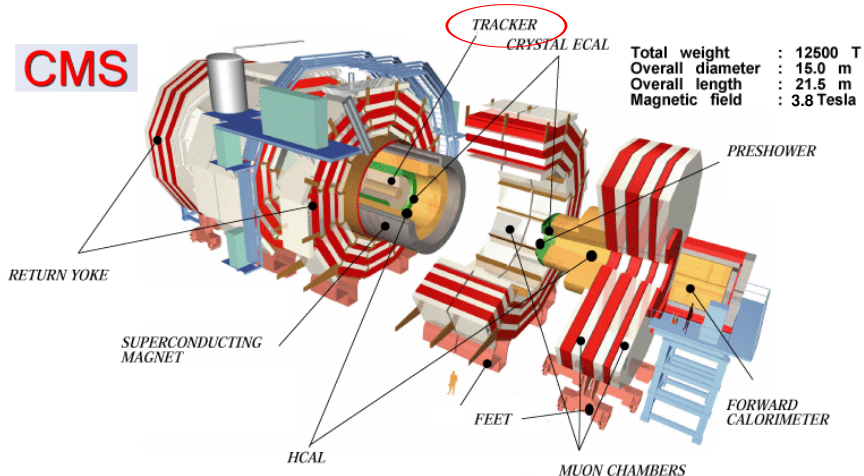


14th International Workshop on
Advanced Computing and Analysis Techniques
in Physics Research
September 5-9, 2011
Brunel University, Uxbridge, London

Outline

- The CMS silicon tracker
- Alignment method and tools
- Results
- Summary

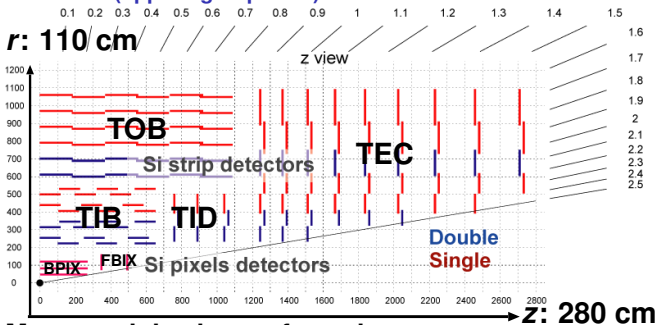
The CMS Detector at the LHC



- Large Solenoid: $B = 3.8$ T.
- All Silicon Inner Tracker.

The CMS Tracker: All Silicon

rz-view (upper right quarter)



Single Hit Resolution

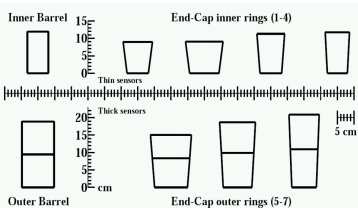
Pixel:

up to $\sigma = 9 \mu\text{m}$

Strip:

$\sigma \approx 23 - 60 \mu\text{m}$

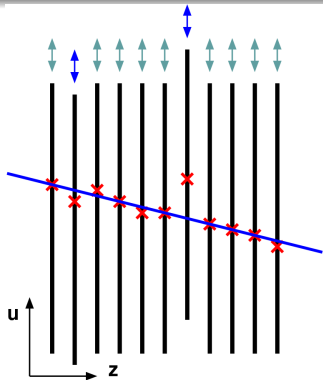
Many module shapes for strips



- 1440 silicon pixel modules
- 15148 silicon strip modules (24244 sensors)
- Strips generally measure $r\phi$ direction
- Some radii: additional modules rotated by 100 mrad

Alignment challenge: 200 k parameters

Track Based Alignment: Principle



Simple Example

- 11 parallel planes measuring 1D
 - **displaced** in measurement direction
 - fit 10^4 straight tracks: $u = F_{\mathbf{a}}(z) = a_1 + a_2 \cdot z$
 - residual $r_i = m_i - F_{\hat{\mathbf{a}}}$ at plane i :
shift of plane i leads to $\langle r_i \rangle \neq 0$
 - cannot simply shift plane by $-\langle r_i \rangle$:
depends on shifts of other planes
- ⇒ tracks correlate alignment parameters

Proper Treatment: Global Fit Approach (e.g. Least Squares)

- **Simultaneous fit of all parameters**: shifts, track parameters!
- Minimise **sum of squares of residuals**, $\chi^2(\mathbf{a}) = \sum_k \left(\frac{m_k - F_{\mathbf{a}}}{\sigma_k} \right)^2$.
- $\mathbf{a} = (\mathbf{a}^{global}, \mathbf{a}_1^{local}, \dots, \mathbf{a}_n^{local})^T$
 - global: alignment parameters,
 - local: track parameters.

Global Fit Approach

- Linearising track model and minimisation requiring $\frac{d\chi^2(a)}{da} = 0$:
⇒ Normal equations of least squares $\mathbf{C} \mathbf{a} = \mathbf{b}$.
- Local parameters appear in part of the data only:
⇒ Block structure in \mathbf{C} , use matrix algebra to reduce problem:

$$\mathbf{C}' \mathbf{a}^{global} = \mathbf{b}'.$$

- Matrix \mathbf{C}' , vector \mathbf{b}' summing up contributions from all tracks.
- Solving $\mathbf{C}' \mathbf{a}^{global} = \mathbf{b}'$ provides alignment solution in one step.
⇒ All correlations from tracks taken care of.
- Need clever algorithms for $> 100\,000$ global parameters:
⇒ **Millepede II** and **General Broken Lines Track Refit**.

Millepede II: Experiment Independent Global Fit Tool

(originally by V. Blobel, further developed by C. Kleinwort)

Task

- Setting up and Solving Matrix Equation

$$\mathbf{C}' \mathbf{a}^{global} = \mathbf{b}' ,$$

- from millions of tracks (containing outlier hits),
- \mathbf{C}' is $n \times n$ matrix:
 - here $n \approx 200\,000$,
 - typically sparse.

⇒ **Very demanding for memory and CPU.**

Input from Experiment

- Linearised track fit information:
 - residuals with uncertainties,
 - derivatives $\frac{\partial F}{\partial a^{local}}$ and $\frac{\partial F}{\partial a^{global}}$,
- Global parameter constraints: $\sum d_i a_i^{global} = e$.

Features: Computing Aspects

- More powerful successor of Millepede I.
- Stand alone Fortran program.
- Reading (zipped) binary input from Fortran or C(++).
- Optimised for speed:
 - iterative MINRES to solve $\mathbf{C}' \mathbf{a}^{global} = \mathbf{b}'$,
 - CPU intense parts parallelised using OpenMP®,
 - local fit detects bordered band matrices (\Rightarrow Broken Line Fit),
 \Rightarrow reading data from disc and memory access remaining bottlenecks.
- Optimised for memory space:
 - symmetric \mathbf{C}' would need 160 GB in double precision,
 - reduction due to sparsity
 - compression by bit packed addressing of continuous non-zero blocks,
 - and by single precision for elements summing up from few tracks.

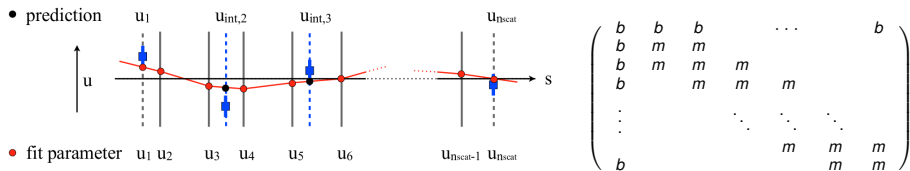
Track Fit

- Charged particle in magnetic field: need 5 helix parameters.
 - Traversing material: multiple scattering effects.
(relevant for “heavy” tracking detectors)
 - Usually treated by progressive track fit: Kalman filter.
 - Millepede II needs global fit:
 - ⇒ 2 scattering angles per thin scatterer,
 - ⇒ $5 + 2n_{scat}$ explicit track parameters.
 - Reaching > 50 parameters for cosmic tracks in CMS tracker.
- ⇒ Danger of CPU consuming single track fits when building matrix equation $\mathbf{C} \mathbf{a}^{global} = \mathbf{b}$.

Way out:

- General Broken Lines Track Refit

General Broken Lines Track Refit



Concept: Define Track Parameters with Local Meaning

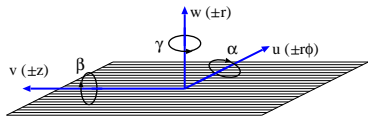
- Reparametrise: $\mathbf{a}^{local} = (\Delta q/p, \mathbf{u}_1, \dots, \mathbf{u}_{n_{scat}})$.
- \mathbf{u}_i : 2D offsets in local system at each scatterer.
- Predictions \mathbf{u}_{int} for measurements: interpolating between scatterers.
- Kink angles from triplets of adjacent scatterers.

⇒ Local fit $\mathbf{A} \cdot \mathbf{x} = \mathbf{b}$:

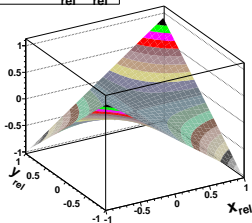
- bordered band matrix, band width $m \leq 5$, border size $b = 1$.
- Fast solution by root free Cholesky decomposition:
 - Effort to calculate \mathbf{x} : $\sim n_{par} \cdot (m + b)^2$, \mathbf{A}^{-1} : $\sim n_{par}^2 \cdot (m + b)$
- Equivalent to standard CMS Kalman filter track fit.

Input Data

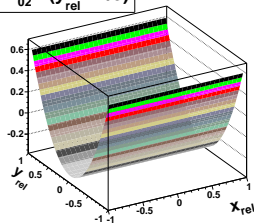
- Loosely selected isolated muons: 15 million.
- Muon pairs from $Z \rightarrow \mu^+ \mu^-$ decays: 375 thousand pairs.
- Low momentum tracks: 3 million.
- Cosmic tracks (e.g. recorded in between LHC fills): 3.6 million.



$$w_{11} * (x_{rel} * y_{rel})$$



$$w_{02} * (y_{rel}^2 - 1/3)$$



Alignment Algorithm and Parameters

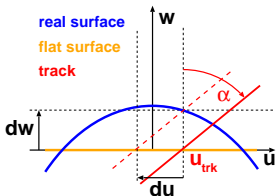
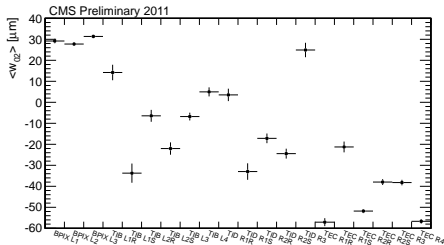
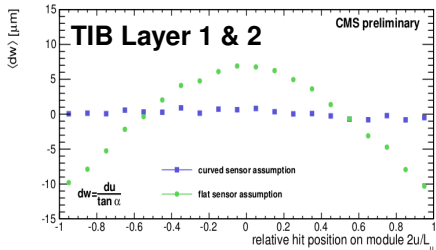
- Millepede II algorithm with $\sim 200\,000$ free alignment parameters.
- 8 (9) parameters per strip (pixel) sensor:
 - 5/6 rigid body like parameters (one insensitive for strips),
 - 3 bow parameters.
- Time dependent rigid body parameters for larger pixel structures:
 - 9 time periods in common fit,
 - ⇒ moving structures, modules constant within.
- $Z \rightarrow \mu^+ \mu^-$ combined object, adding Z mass “measurement” (⇒).

Millepede II at Work

- 246 zipped binary files (\sum 46.5 GB), read 13 times.
- 22.6 million local fit objects,
 - bordered band matrix structure: $\max(\text{border}) = 9$, $\max(\text{width}) = 4$.
- MINRES iterating 3 times (tightening outlier rejection) to solve $\mathbf{C}' \mathbf{a}^{global} = \mathbf{b}'$.
- 200 614 fit parameters (including 138 Lagrange multipliers).
- Matrix with 31% non-zero off-diagonal entries, compression ratio 40%,
 \Rightarrow fits well into 32 GB memory.
- Total CPU 44.5 h, **Wall 9:50 h** using 8 threads on Intel® Xeon® L5520, 2.27 GHz.

\Rightarrow Very efficient usage of resources with fast turnaround for analysis!

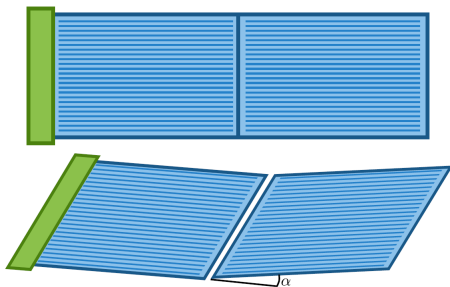
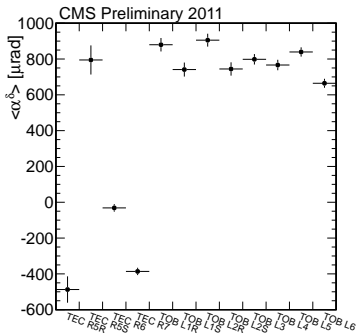
Silicon Sensors are Bowed, not Flat



- Residual du , track slope $\tan \alpha$:
map residual \perp sensor, $dw = du / \tan \alpha$.
- ⇒ See **bowed sensor surface**.
- Alignment of bows **accounts for that**.
- Sensor type/mounting specific bow strengths.

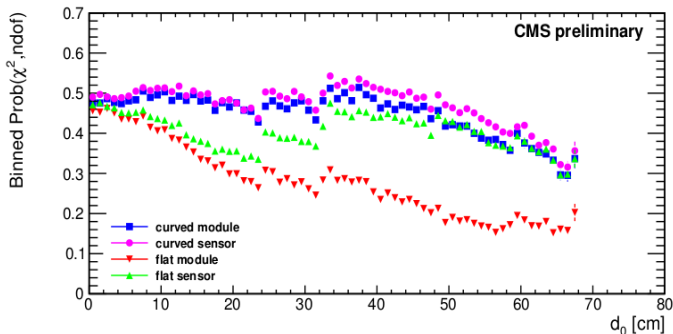
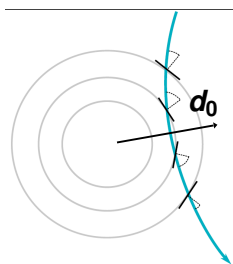
- High relevance in BPIX layer 1: $\langle w_{02} \rangle = 30 \mu\text{m}$, large track slope:
⇒ systematic residuals $> 100 \mu\text{m}$ at sensor edge.

Kink within Two Sensor Modules



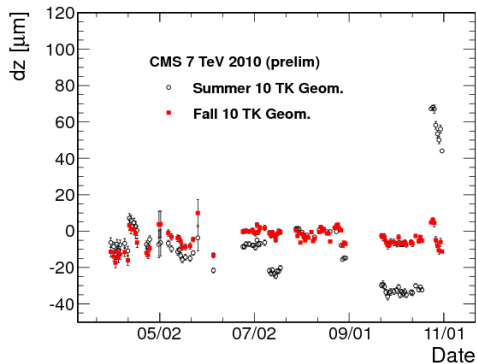
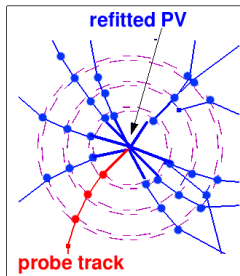
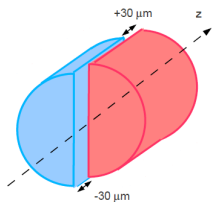
- Long modules in outer strip tracker (TOB and outer TEC): two daisy chained sensors.
- Sensors rotated against each other: kink angle.
- Typical kink is $2\alpha^\delta = 1.6$ mrad.
- Larger effect than sensor bow.

Sensor Bow Treatment Improves Cosmic Tracking



- Cosmic tracks mainly come from above.
- Increasing d_0 increases average track angle from sensor normal, \Rightarrow increasing sensitivity to deviation from flat sensors.
- Average goodness of fit vs d_0 demonstrates improvements from flat modules via flat sensors, bowed modules to bowed sensors.
- Remaining structure related to radii of layers: material.

Pixel Movements and Monitoring

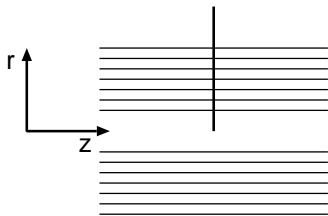


- Pixel half barrels move wrt each other in time.
- Monitoring separation along z using unbiased vertex-track residuals.
- Time dependence of pixel structure alignment: // properly caring of separation as function of time.
- B-tagging insensitive to remaining $10 \mu\text{m}$ effect.

Problem of Track Based Alignment: Weak Modes

Minimising residuals can be **insensitive** to certain **global distortions**.

- Potential bias on track parameters.
- **Dependent on data** fed into matrix.



Example: Telescoping

Shift in z growing linear with radius r

- Magnetic field $\mathbf{B} \parallel z$:
tracks are straight lines in rz
- This distortion does not change that!
- \Rightarrow Bias in η

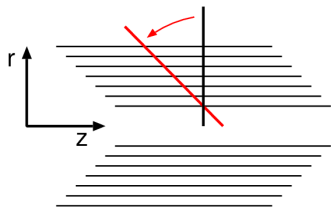
Solution:

- Adding cosmic tracks.
- Telescope effect bends track:
not allowed by track model.

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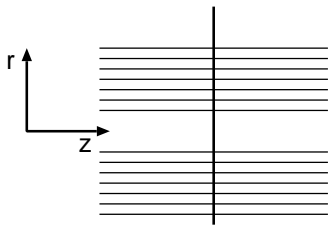
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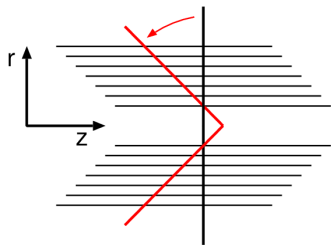
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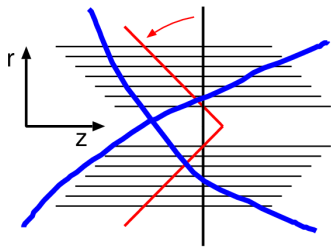
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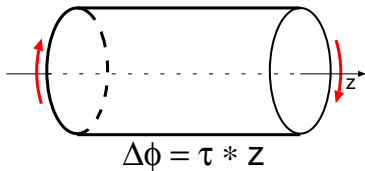
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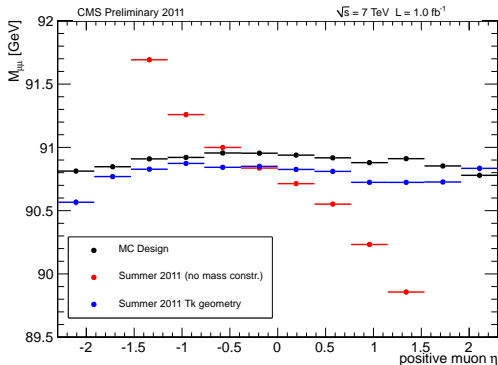
Solution:

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Weak Mode “Twist” and the Z Mass



Twist distortion is weak mode even using cosmics.



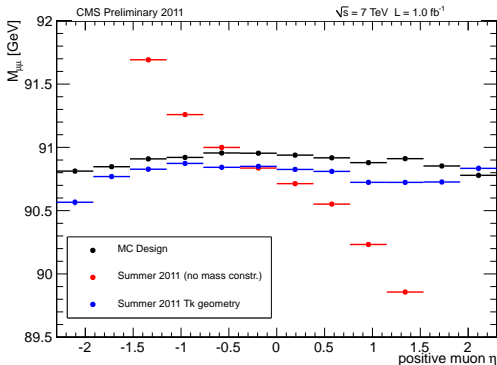
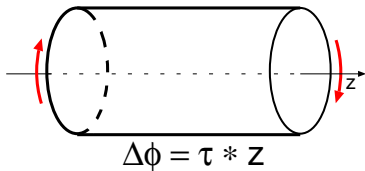
- Results in curvature changes, biasing measured p_t of positive or negative tracks oppositely.

⇒ Reconstructed Z mass depends on muon charge and η : red curve.

Way out:

- Use Z mass information in alignment fit.

Controlling the Twist



- Re-parametrise muons from Z decay as common local fit object:
 - from 2×5 independent parameters to common set of 9.
 - Add Z mass as virtual measurement, RMS as uncertainty.
- ⇒ Keeps twist under control: **blue curve**, almost as in simulation with design conditions (**black curve**).

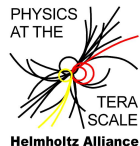
Summary

- Large CMS silicon tracker: a challenge for alignment.
- Track based alignment with ~200 k parameters:
 - determining sensor bows and kinks,
 - following time dependent movements of large pixel structures,
 - controlling momentum changing weak modes using the Z mass.
- Main working horse: Millepede II with General Broken Lines,
⇒ global fit approach in < 10 h Wall time.

CMS Silicon Tracker Alignment

Serving physics analysis with high precision for discoveries.

Supporting:



Millepede II

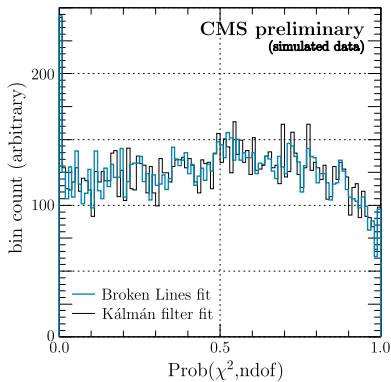
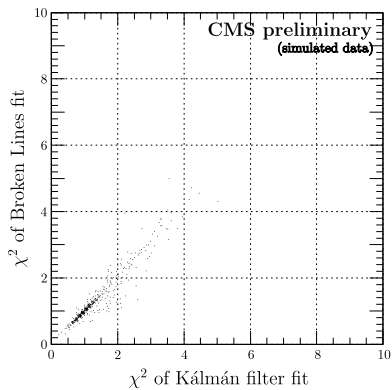
See Wiki https://www.wiki.terascale.de/index.php/Millepede_II.

General Broken Lines Track Refit

See  Wiki

<https://www.wiki.terascale.de/index.php/GeneralBrokenLines>.

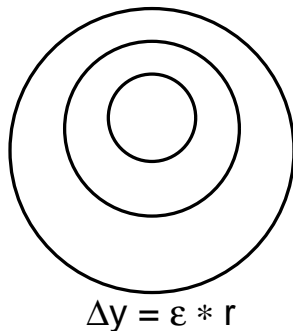
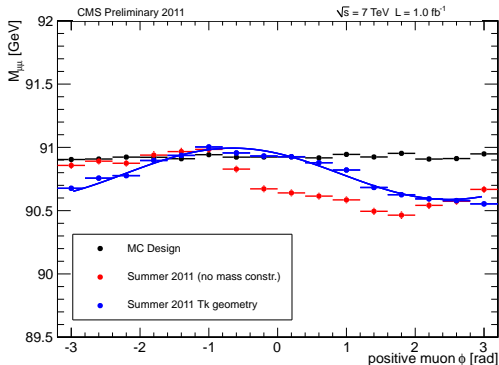
General Broken Lines: Equivalence with Kalman Fit



Test on Simulated Muon Tracks

- Good correlation of χ^2/ndf values.
- No difference in probability distributions.
- No momentum/track angle dependence seen (not shown).

Open Issue: ϕ Dependence of Reconstructed Z Mass



- Charge and ϕ dependent bias remains despite usage of Z mass.
- But sine wave fit reveals small amplitude of $\sim 0.2 \text{ GeV}$.
- Simulation: sine wave can be modeled with sagitta-like distortion.
 - Increasing Z statistics?
- Studies ongoing:
 - Using straight cosmics from B -field off data?
 - Small biases in track model?