

New Physics in a Hidden Sector

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IPP Townhall July 15, 2020

The Standard Model Works!

$$\mathcal{L}_{SM} = -\frac{1}{4} G_{\mu\nu} G^{\mu\nu} - \frac{1}{4} W_{\mu\nu} W^{\mu\nu} - \frac{1}{4} B_{\mu\nu} B^{\mu\nu}$$

$$+ |D_{\mu} H|^2 + \mu^2 |H|^2 - \lambda |H|^4$$

$$+ \bar{u}(i\not{\partial} - m_u)u + \bar{d}(i\not{\partial} - m_d)d + \bar{e}(i\not{\partial} - m_e)e + \bar{\nu} i\not{\partial} \nu$$

but there
are issues!

$$+ \frac{i g}{\sqrt{2}} (\bar{u} \quad \bar{c} \quad \bar{t}) \begin{pmatrix} V_{ud} & V_{us} & V_{ub} \\ V_{cd} & V_{cs} & V_{cb} \\ V_{td} & V_{ts} & V_{tb} \end{pmatrix} W^{+} (1 - \gamma^5) \begin{pmatrix} u \\ c \\ t \end{pmatrix}$$

$$+ (\lesssim 10^{-10}) \frac{g_s^2}{32\pi^2} G_{\mu\nu} G^{\mu\nu} + \text{many other terms} \dots$$

The Standard Model Works!

$$\mathcal{L}_{SM} = -\frac{1}{4} G_{\mu\nu} G^{\mu\nu} - \frac{1}{4} W_{\mu\nu} W^{\mu\nu} - \frac{1}{4} B_{\mu\nu} B^{\mu\nu} \\ + |D_{\mu} H|^2 + \mu^2 |H|^2 - \lambda |H|^4$$

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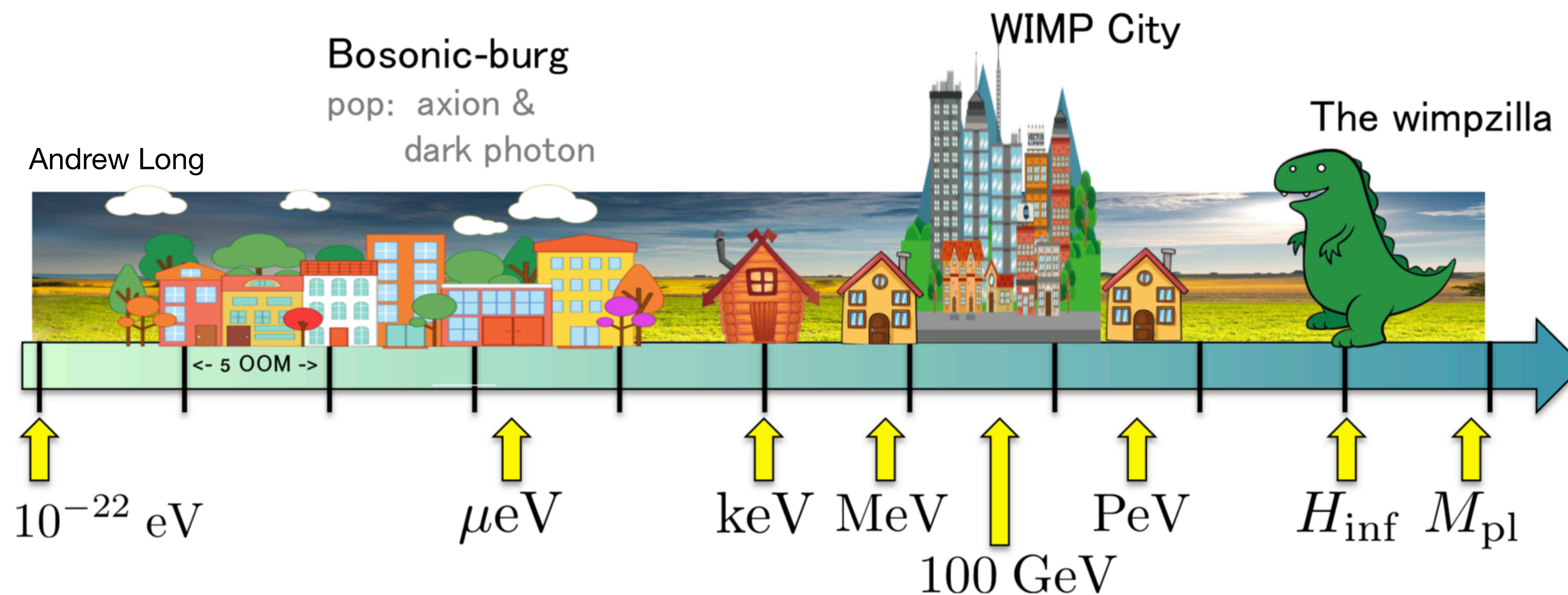
$$+ (\lesssim 10^{-10}) \frac{g_s^2}{32\pi^2} G_{\mu\nu} G^{\mu\nu} + \text{many other terms} \dots$$

Such as:
No
Dark
Matter

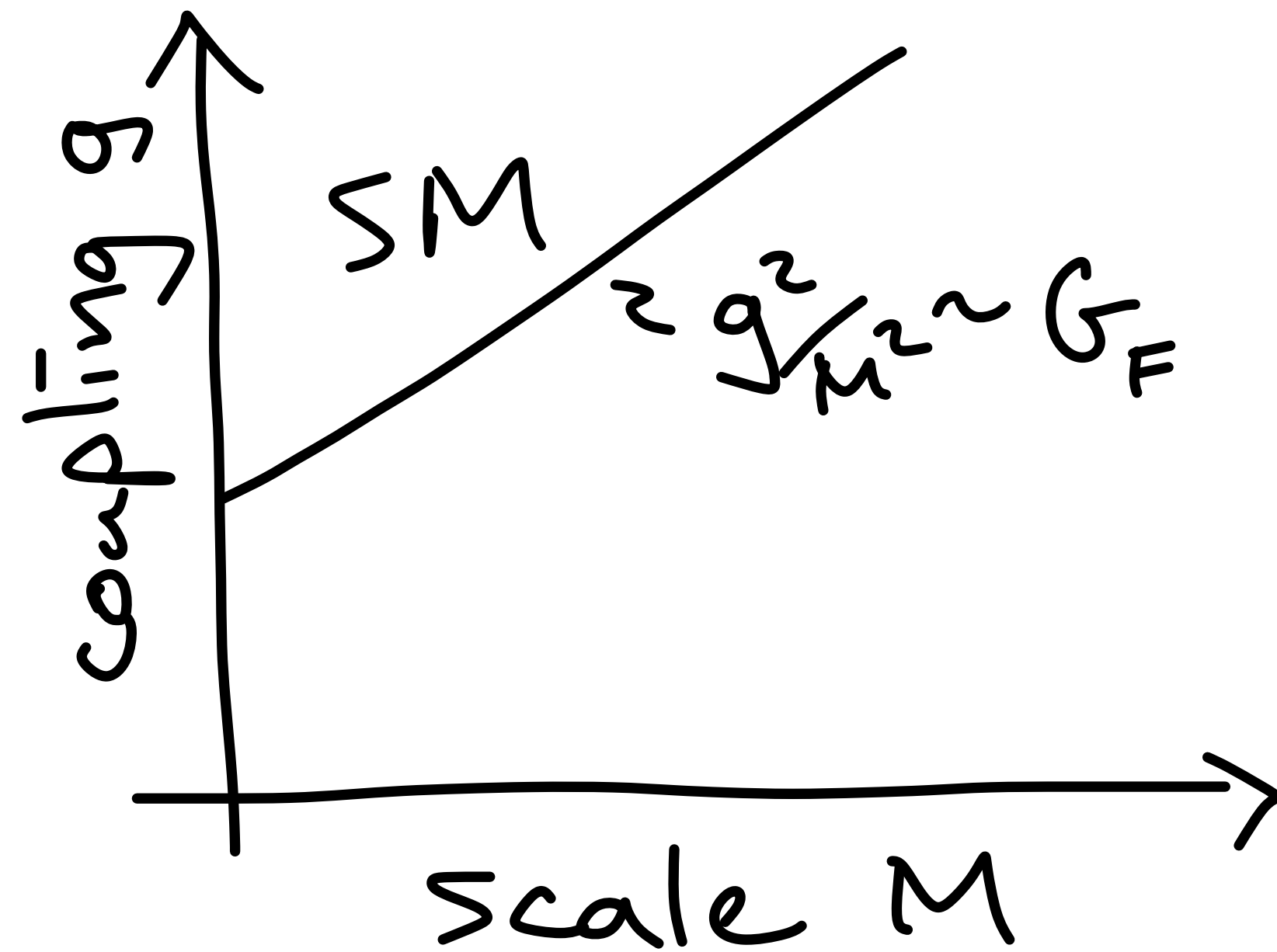
So what do we know about DM?

- cold, i.e. nonrelativistic (massive)
- doesn't interact (nongravitationally) too much

What's its mass?

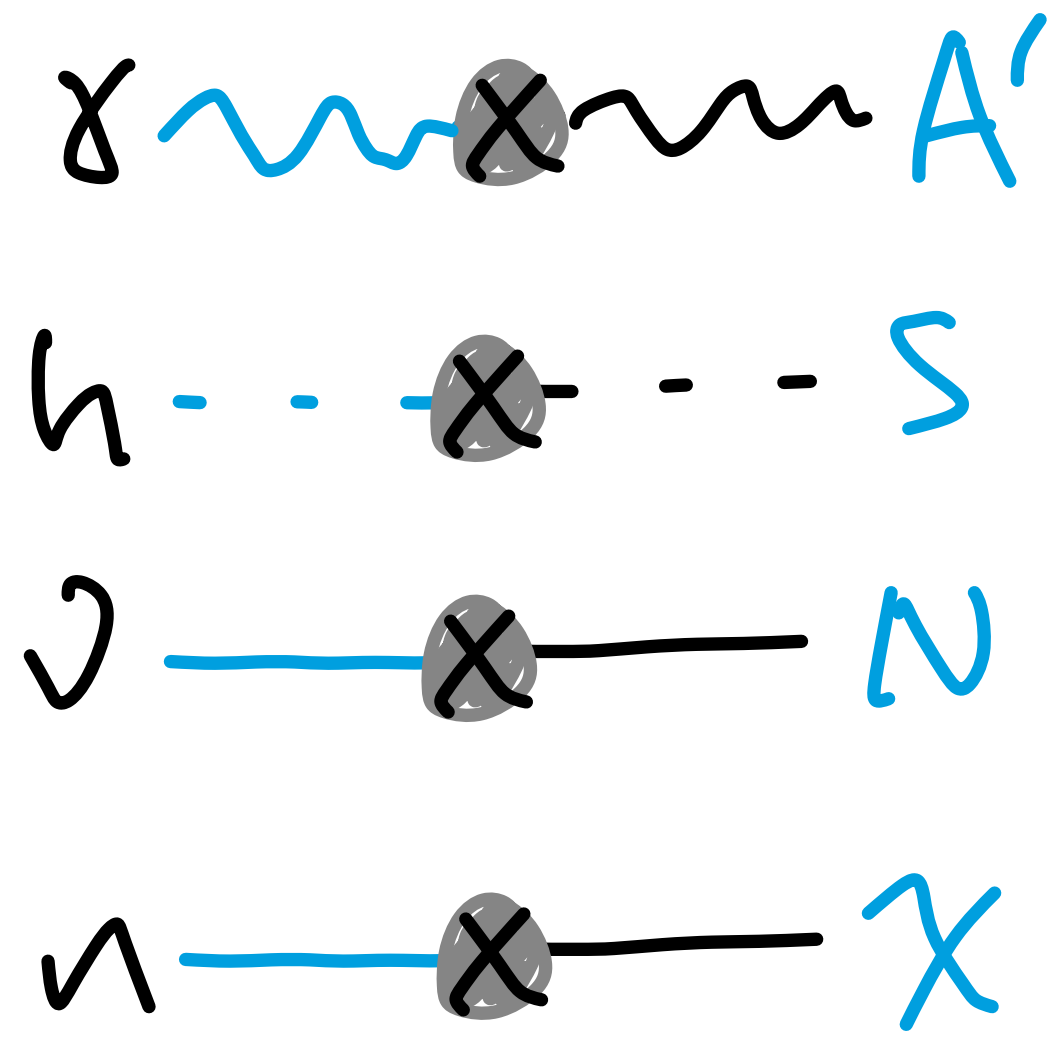
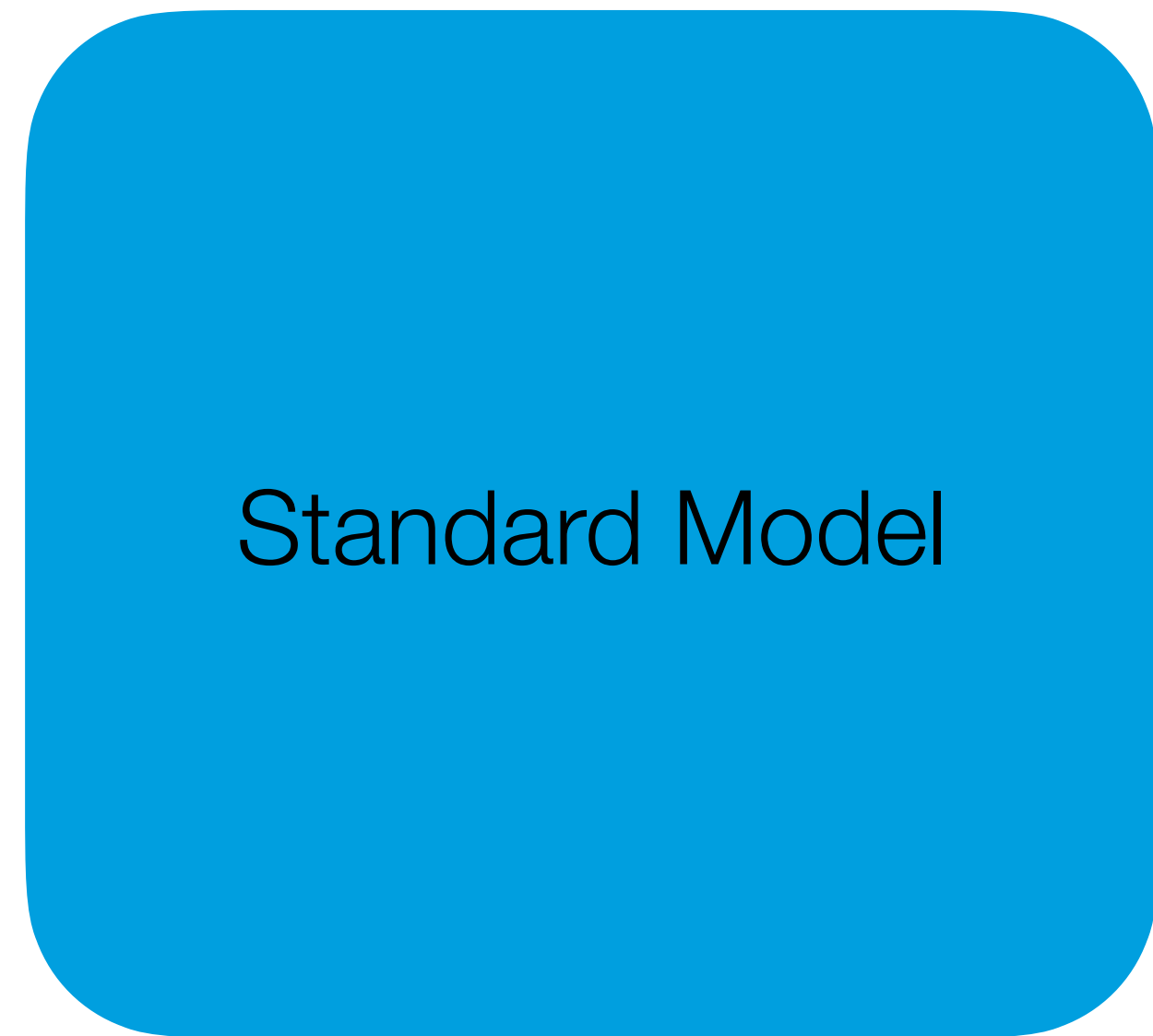


Life on the (SM) frontier



How can we have new physics at low scales without messing up SM??

Portals, or how to (not too feebly) couple light new physics to the SM



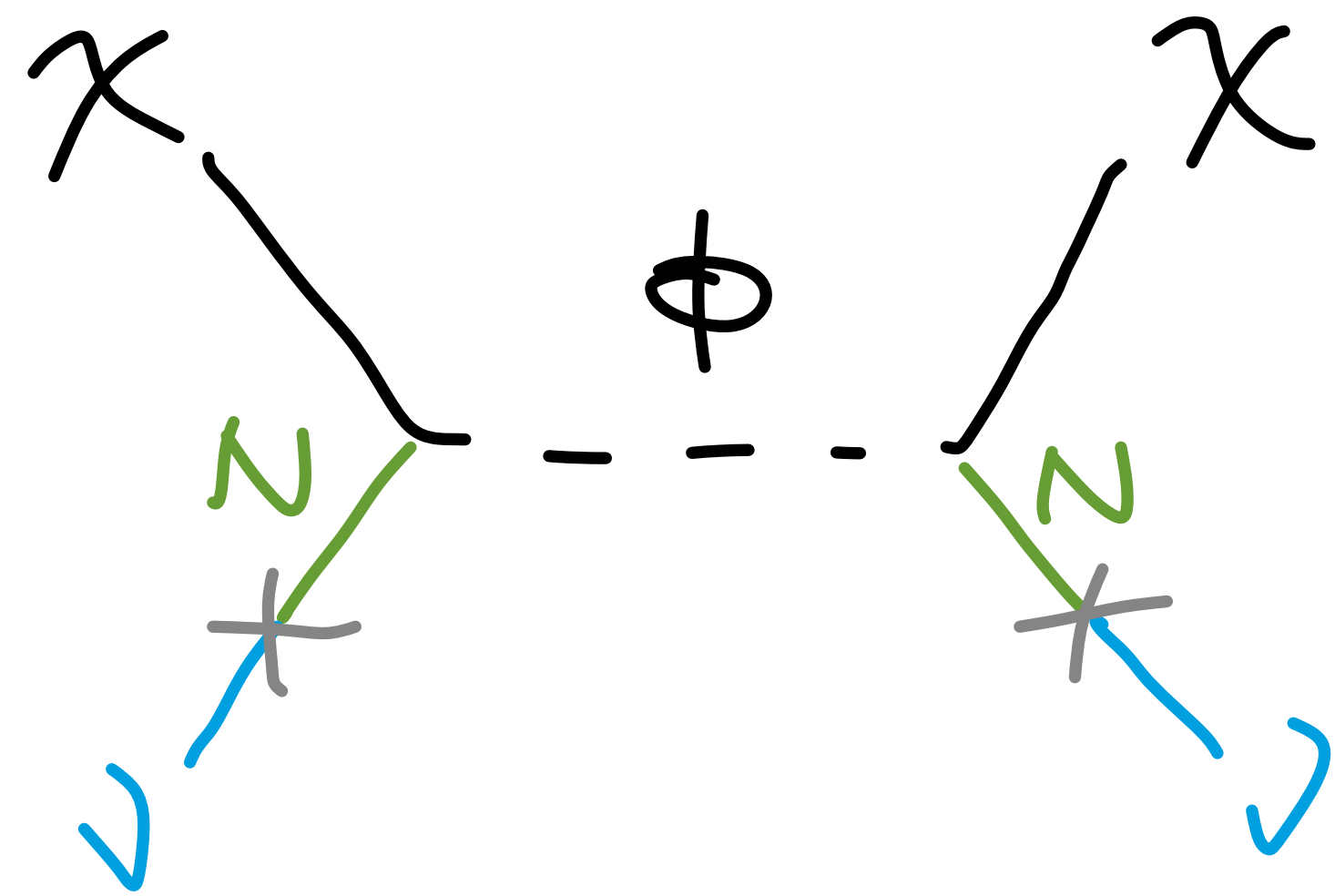
} will discuss these
(see talk by D. Curtin for others)

∇ Portal

$$\mathcal{L} \supset -y \bar{L} H^* N - g \bar{N} \phi \chi \rightarrow -m_0 \bar{\nu} N - g \bar{N} \phi \chi$$

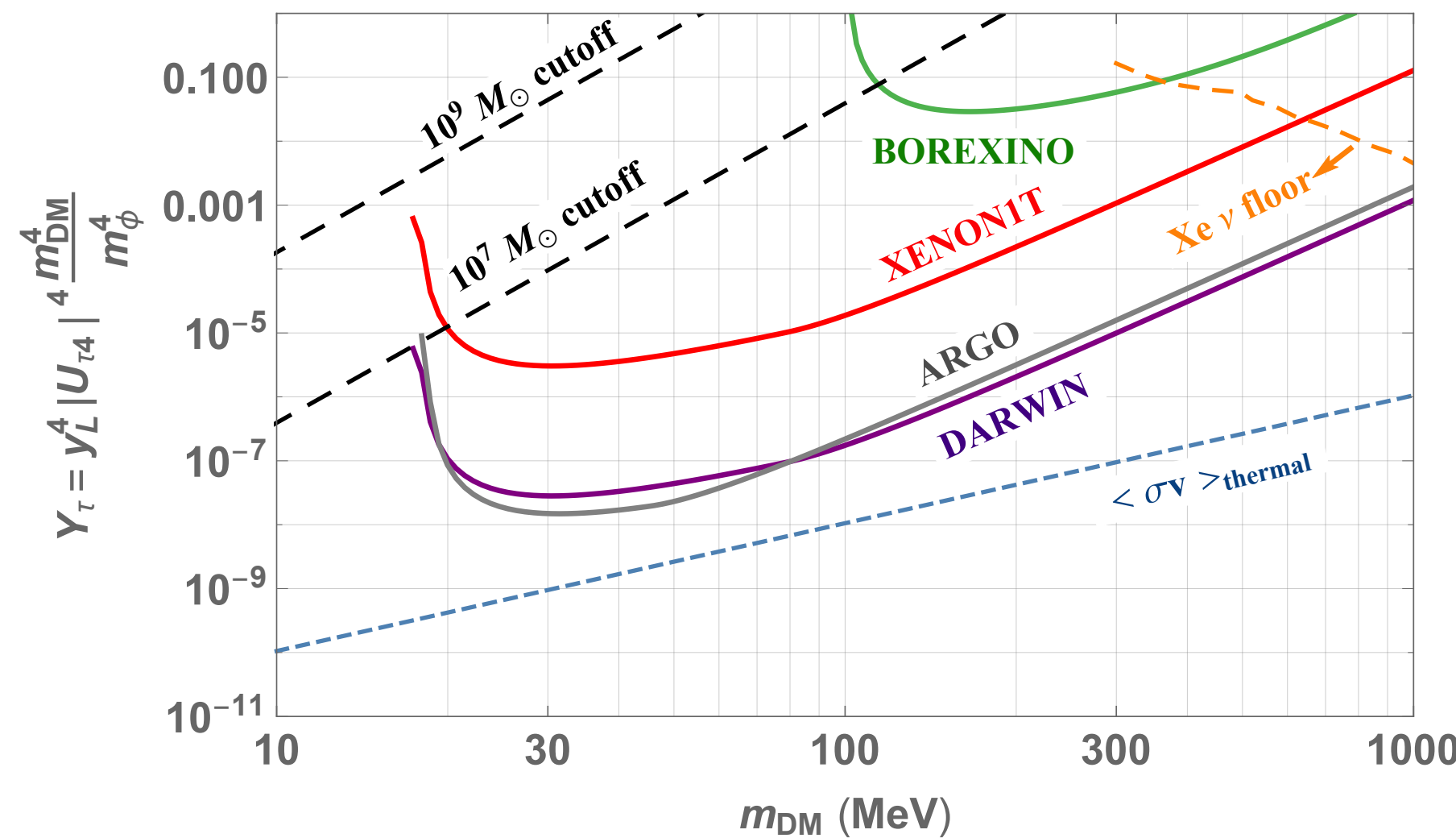
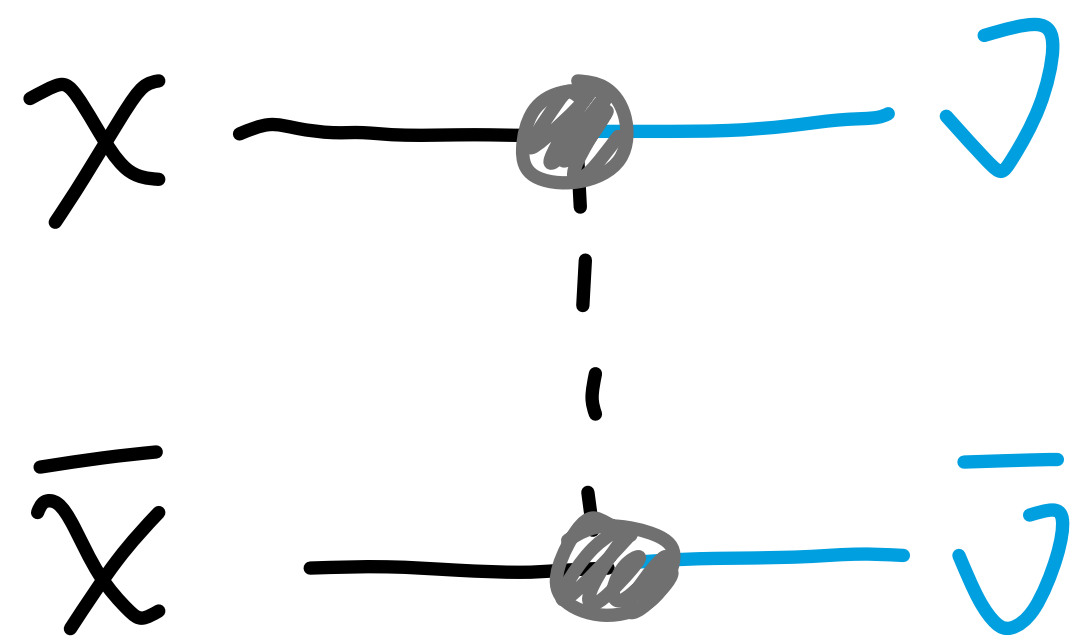
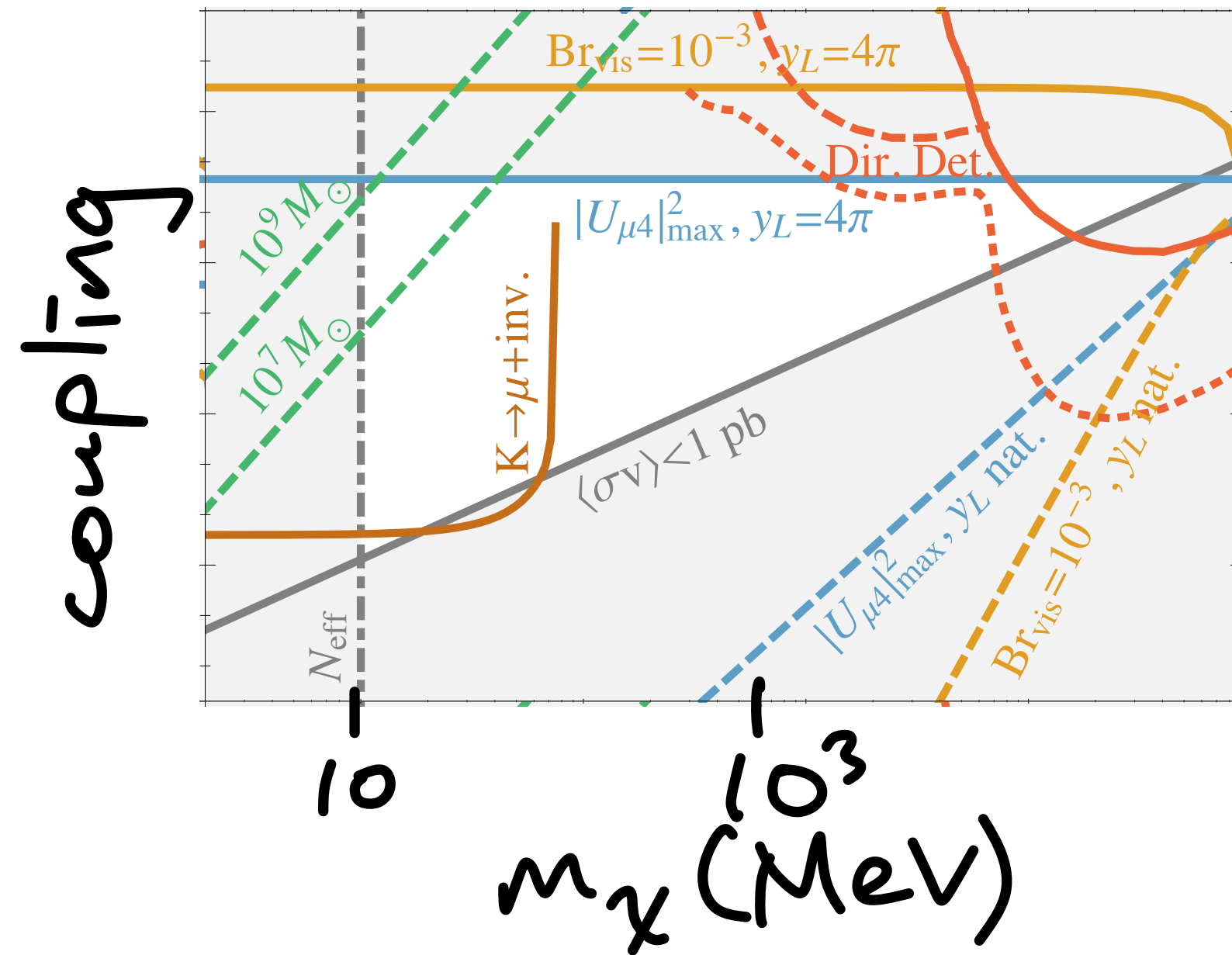
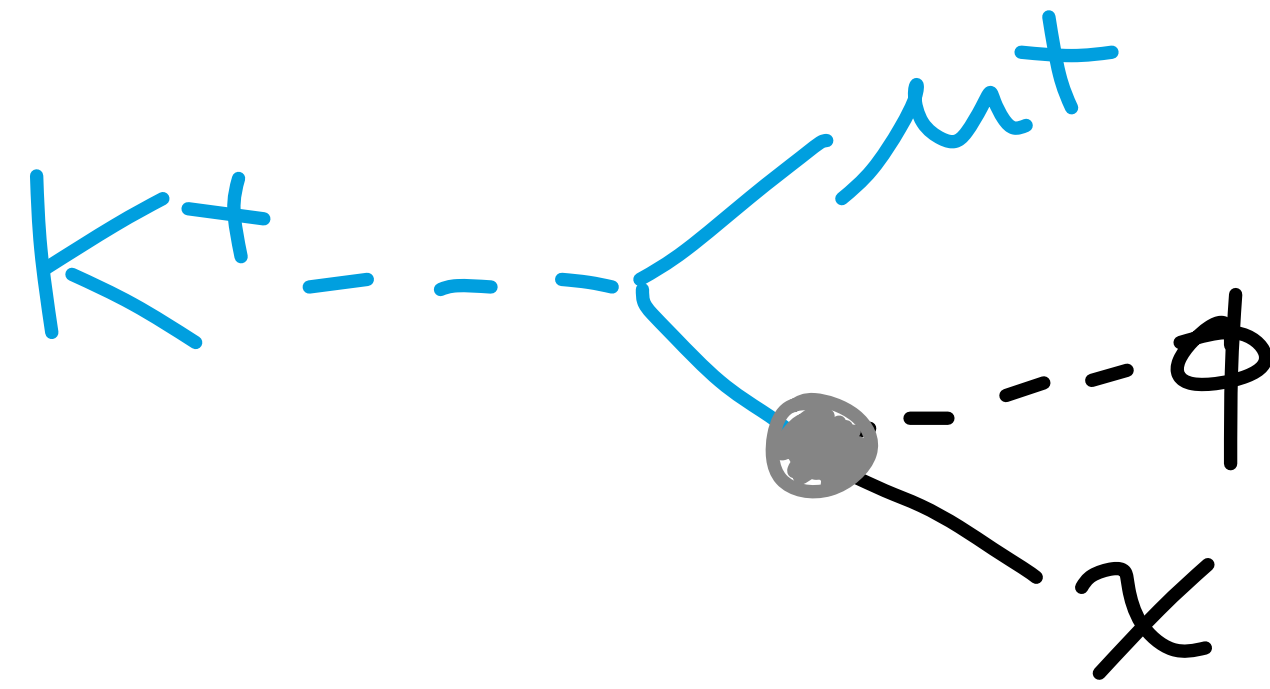
mediator "dark sector"

In early universe DM scatters on ν 's



\Rightarrow Affects structure formation

✓ Portal-Probes



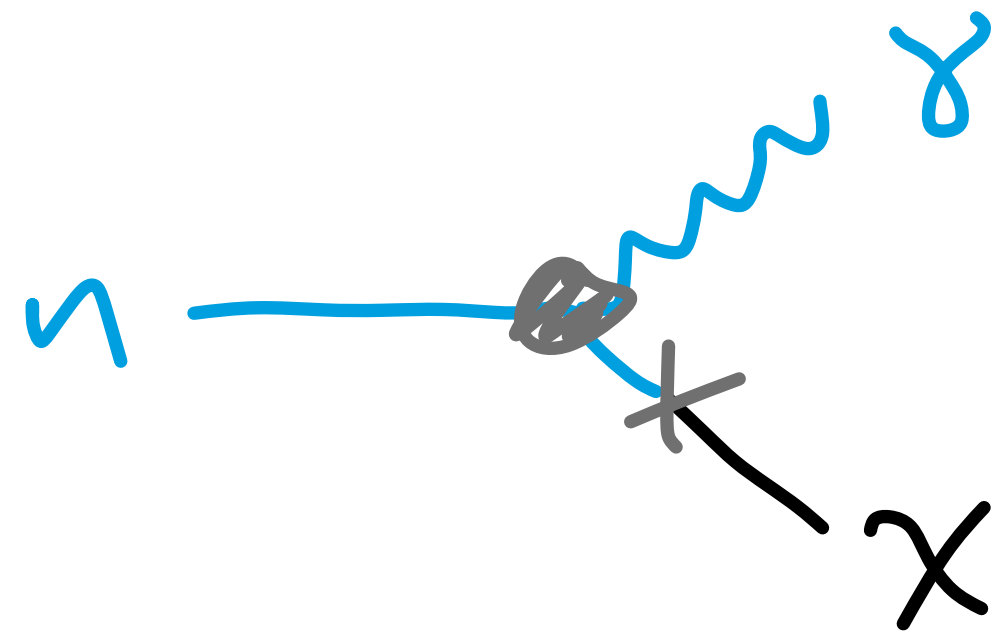
Batell, Han, DM,
ShamsEstaghi
1709.07001
(see talk by
D. Bryman on NABZ)

DM & N. Raj
1812.05103

n Portal

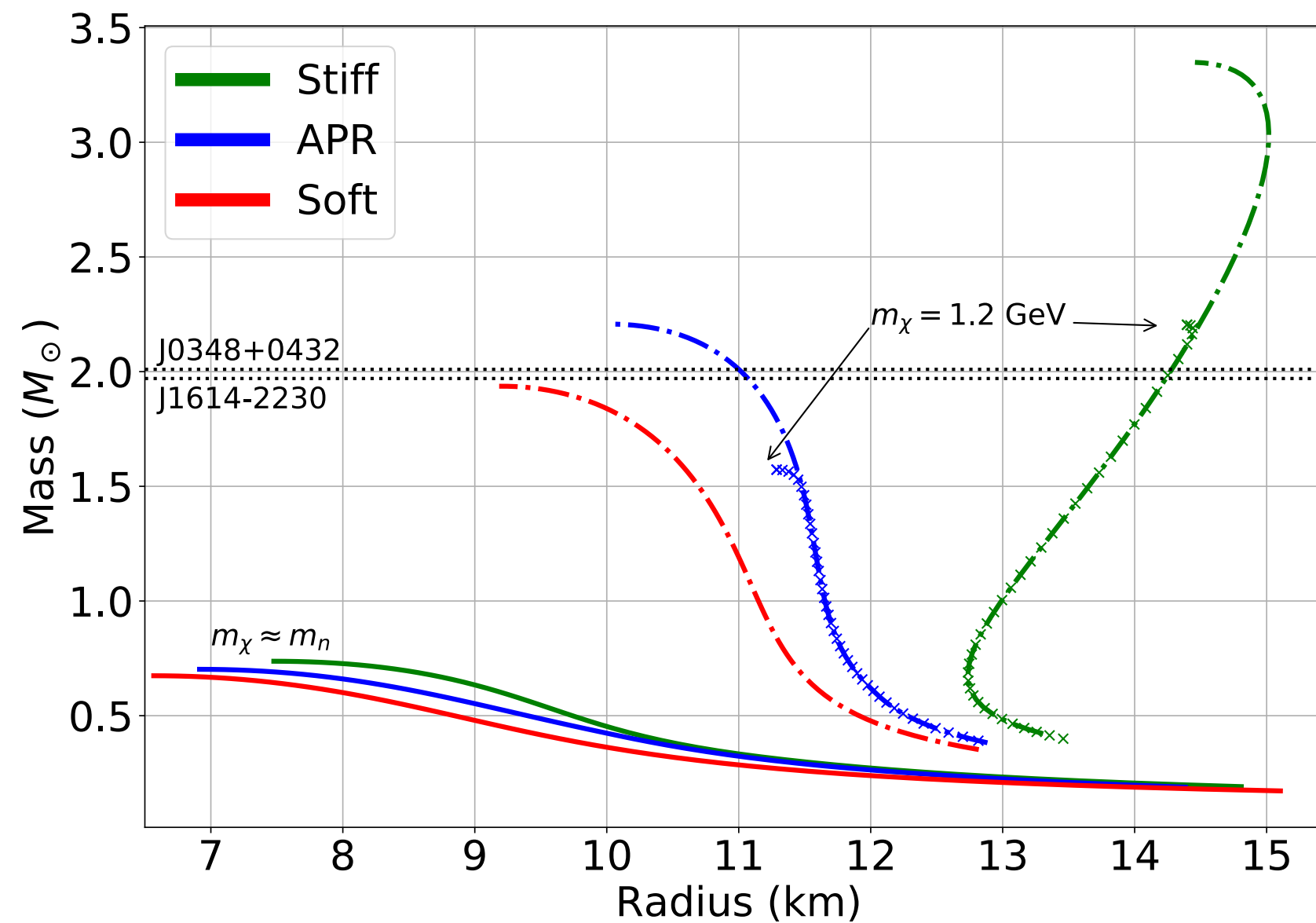
$$\mathcal{L} \supset \frac{1}{\Lambda^2} \bar{u} \bar{d} \bar{d} \chi \rightarrow \delta \bar{u} \chi$$

Leads to exotic n decay

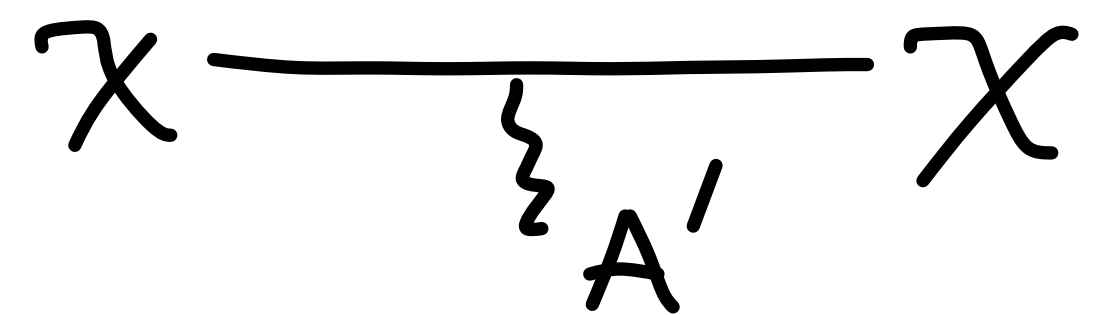


Br ~ 1% can explain n lifetime anomaly

neutron stars important

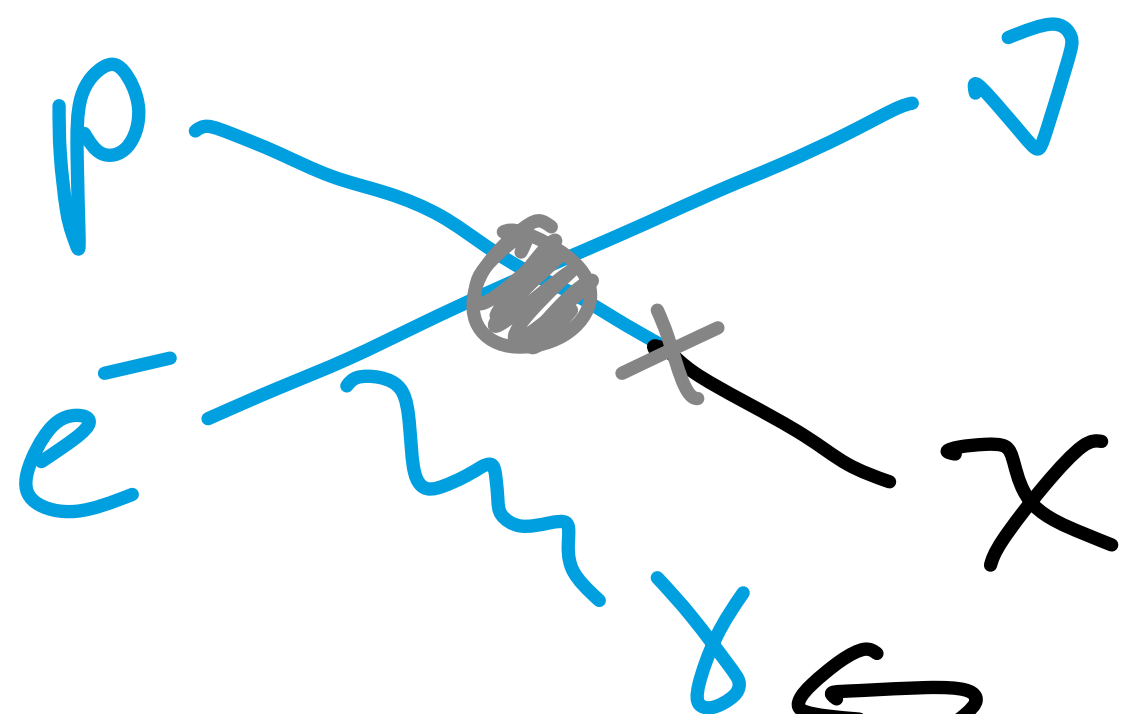


DM, Nelson, Reddy, Zhou
1802.08244
(& others)
See Cline & Cornell

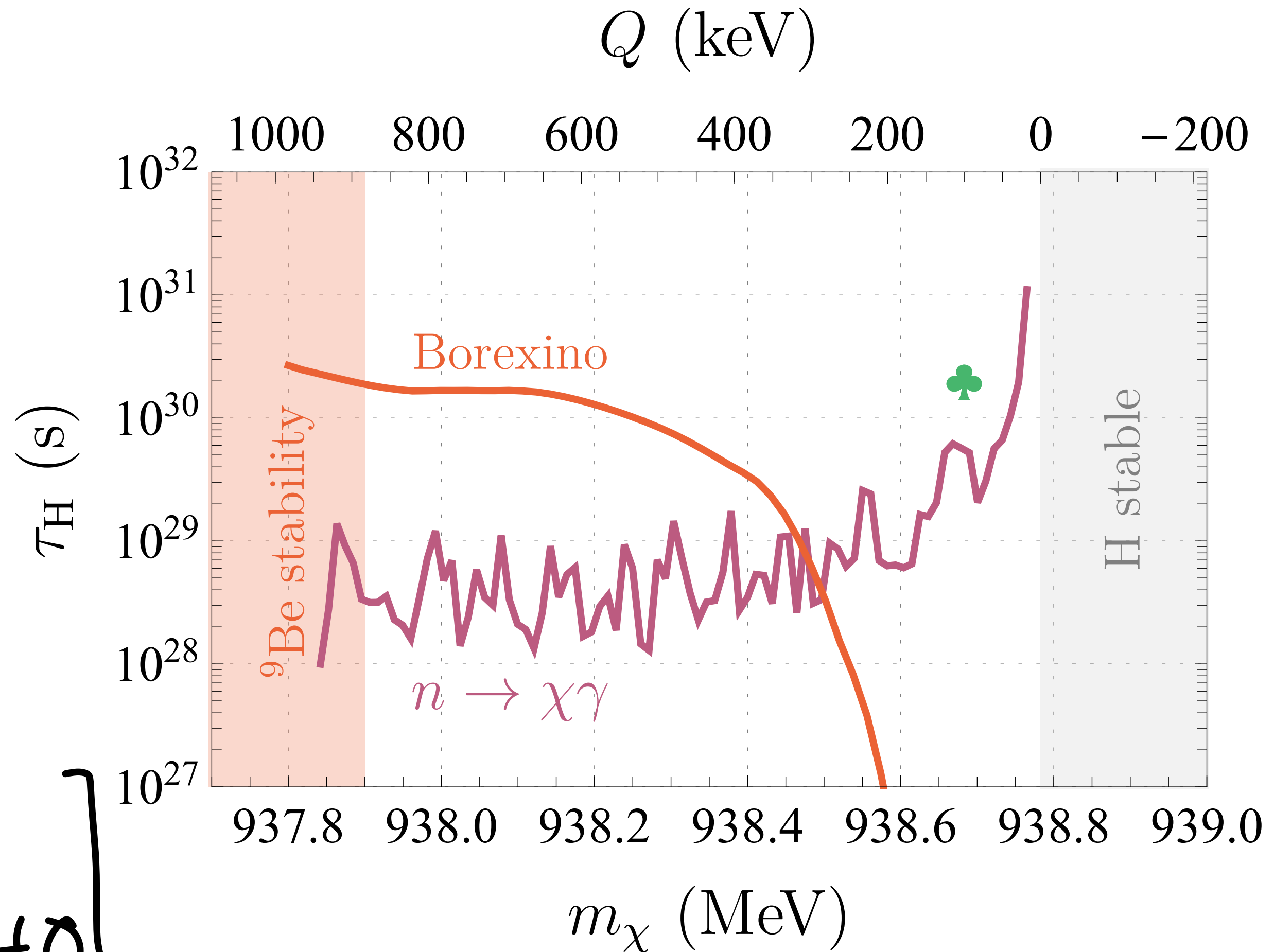


n Portal - Further Probes

Atomic hydrogen unstable:



look for this



[DM, Pospelov 2003.02270
 DM, Pospelov, Raj 2006.15140]

Wrap Up

SM is successful but has deficiencies

↳ DM, Baryon Asymmetry, Inflation, ...

We don't have any "no lose theorems"

⇒ need to have an open mind

"Energy frontier" is an obvious place to look

But we need a diversity of exp'ts

⇒ understand range of new physics
they are sensitive to!

A Lesson from History?



UTLICEPP-82-04
Jun. 1982

"KAMIOKANDE"

- The KAMIOKA Nucleon Decay Experiment -

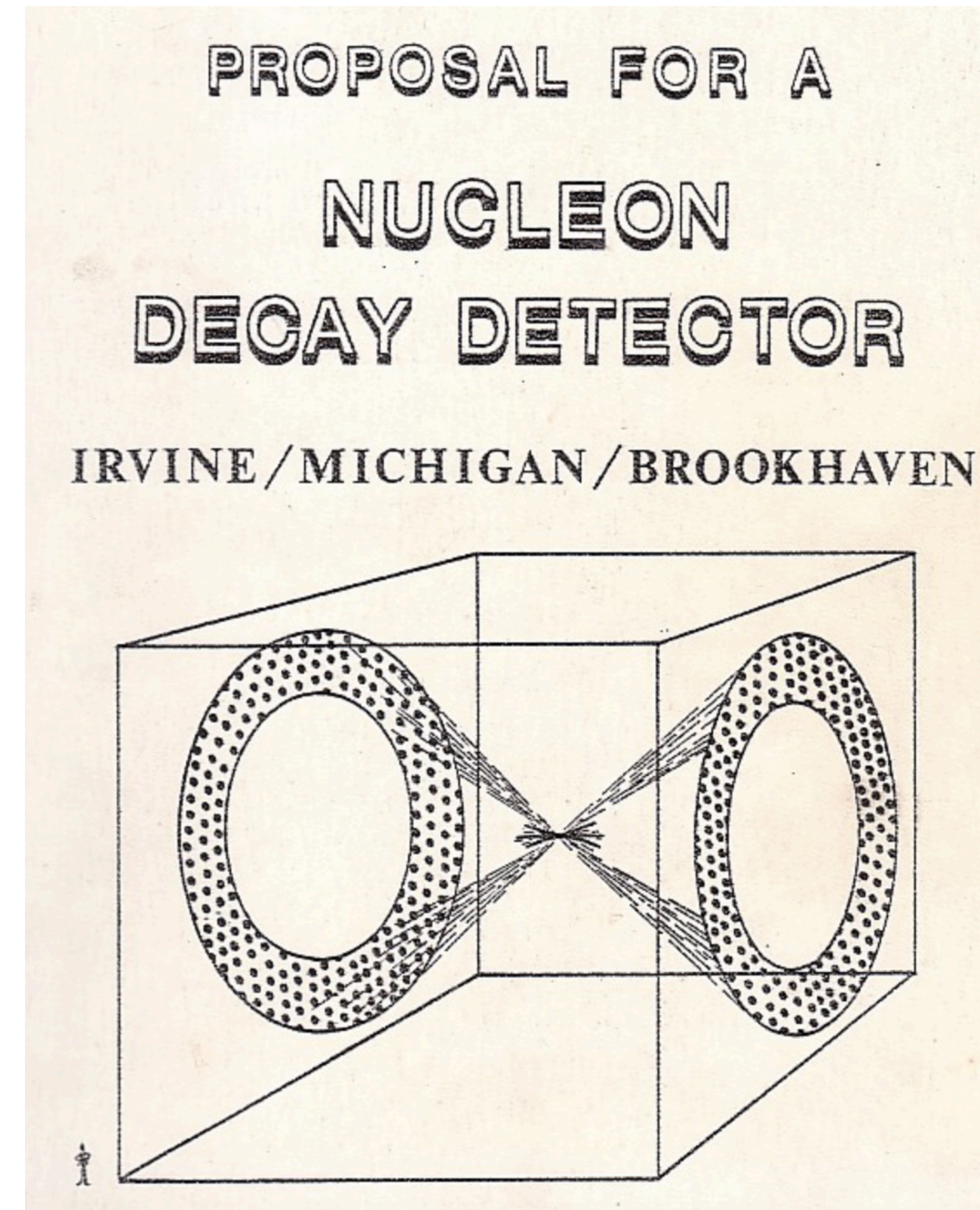
H. Ikeda, A. Nishimura, H. Sugawara, and K. Takahashi
KEK, National Laboratory for High Energy Physics
Oho-machi, Tsukuba-gun, Ibaraki-ken 305, Japan

K. Arisaka, T. Fujii, T. Kajita, K. Kawagoe, M. Kobayashi,
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A Lesson from History?

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高工研四

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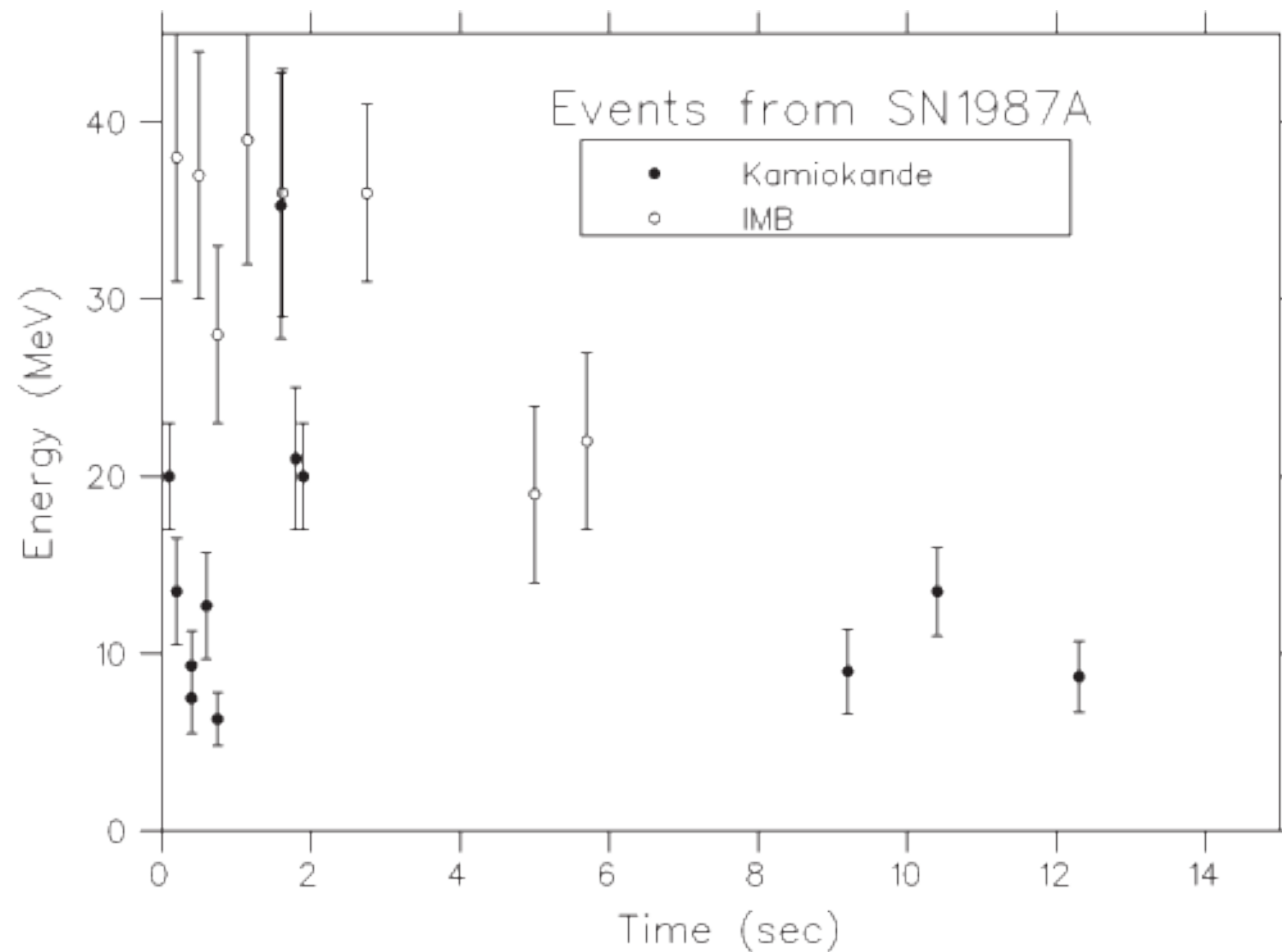
K.A

M.Ko

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