

# Lessons Learned from CCM120



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# Coherent CAPTAIN-Mills (CCM)



CAPTAIN = "Cryogenic Apparatus for Precision Tests of Argon Interactions with Neutrinos"



The University of Manchester



COLUMBIA UNIVERSITY  
IN THE CITY OF NEW YORK

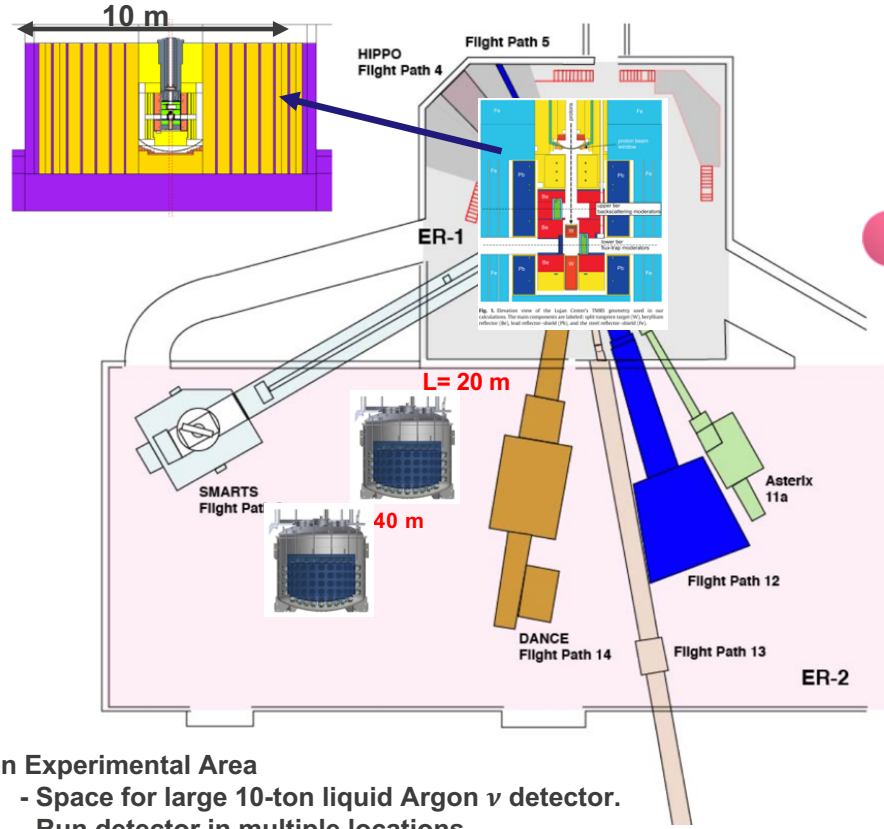


# LANSCCE-Lujan Facility 20 Hz

## 270 ns beam width, FWHM = 135 ns

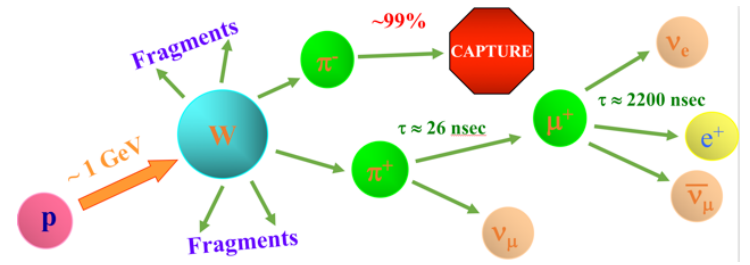
## 100 kW max

Nuclear Instruments and Methods in Physics Research A 594 (2008) 373–381  
 Nuclear Instruments and Methods in Physics Research A 632 (2011) 101–108

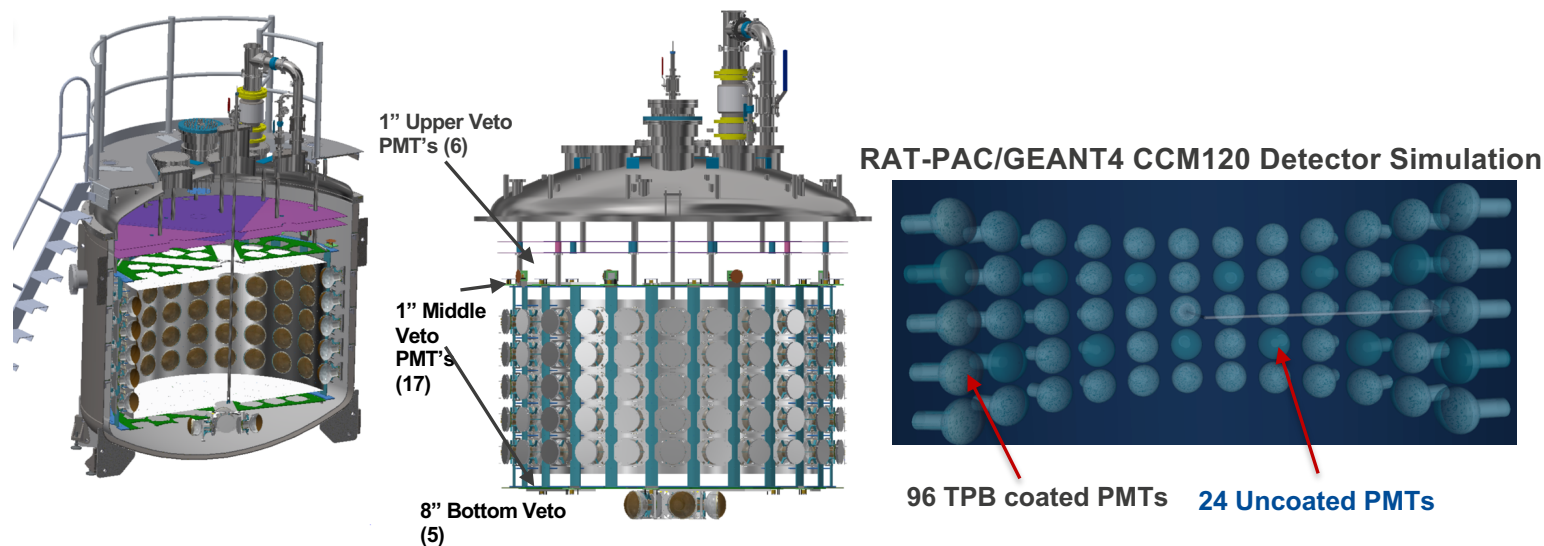


- Lujan Experimental Area**
- Space for large 10-ton liquid Argon  $\nu$  detector.
  - Run detector in multiple locations.
  - Room to deploy shielding, large overhead crane, power, etc

Intense source muon neutrinos: target MCNP simulation flux  $4.74 \times 10^5 \nu/cm^2/s$  at 20 m



# Integrated and Active Veto Regions for Background Rejections



- 7 tons LAr Fiducial volume, 3 tons LAr Veto (2-3 radiation lengths).
- Active Veto region crucial to rejecting cosmic rays and other external backgrounds.
- Detailed CCM200 RAT-PAC/GEANT4 simulation predicts 10-20 keV detection threshold.
- For CCM200 predict  $\sim 0.5$  PE/keVnr

# The CCM120 Detector

- LAr cold test entire SBND PDS system: 96 TPB coated + 24 uncoated PMT's, mounts, cables, feedthrus, HV, electronics, trigger, DAQ, calibration, simulations and data analysis.
- Built detector **August-Dec 2018** at LANSCE/Lujan center (100 kW neutron/stopped pion neutrino source)

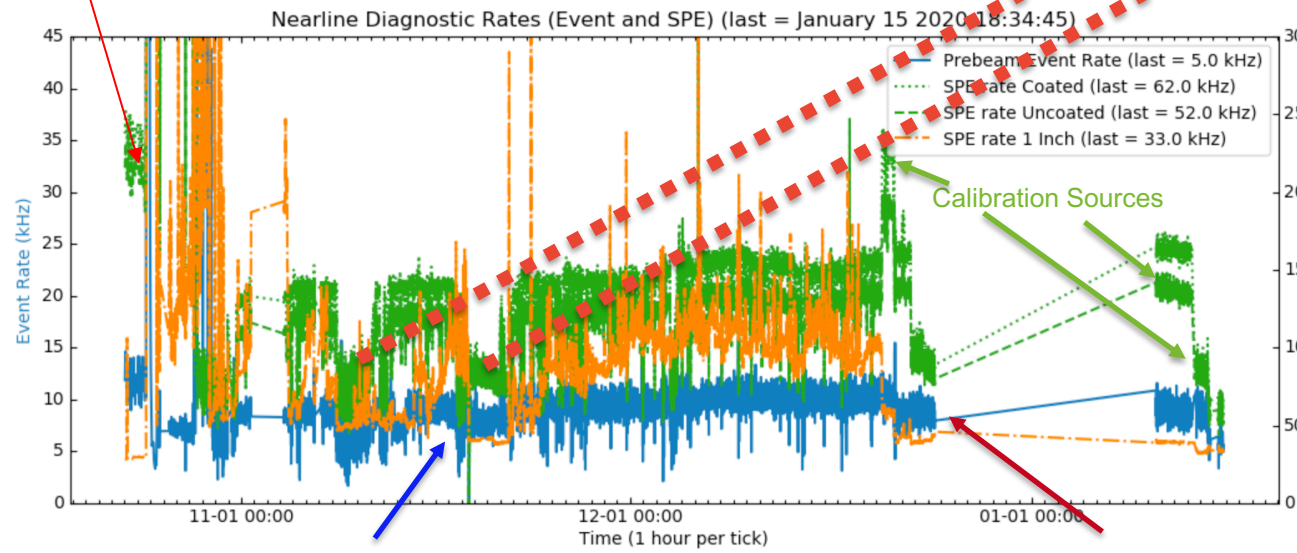
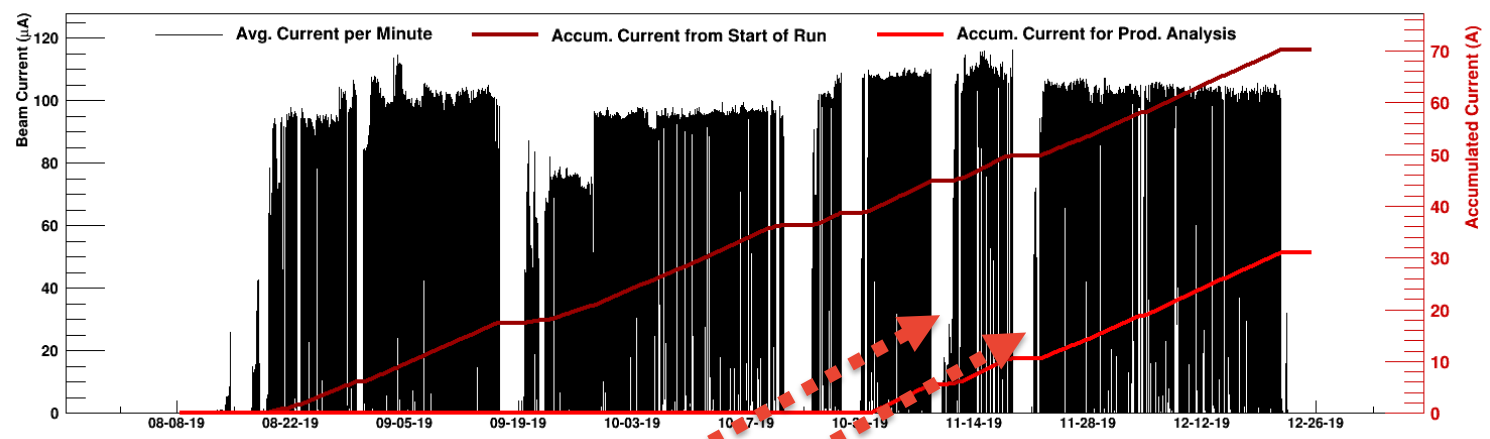


TPB coated PMTs    Uncoated PMTs

TPB coated reflector foils.  
Maximize light output to detect  
dark matter and neutrino events

# Run Cycle 2019

Noisy veto PMT found and turned off

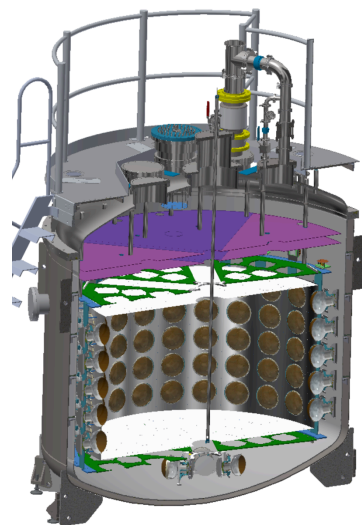


Collected  $17.89 \times 10^{20}$  POT in ~1.5 wall clock months

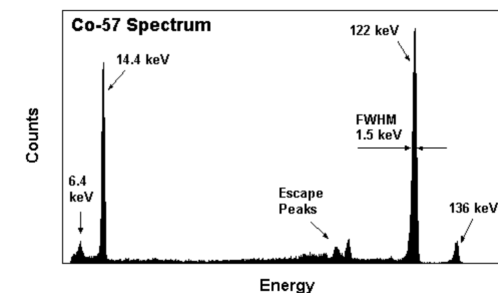
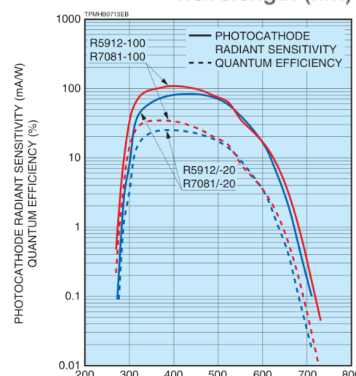
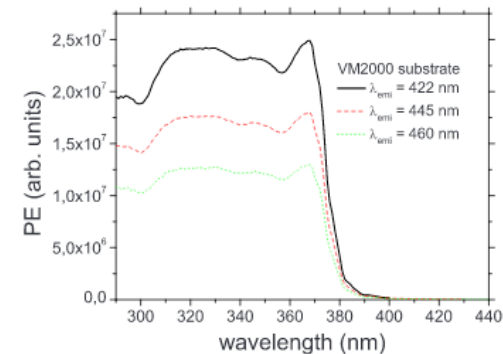
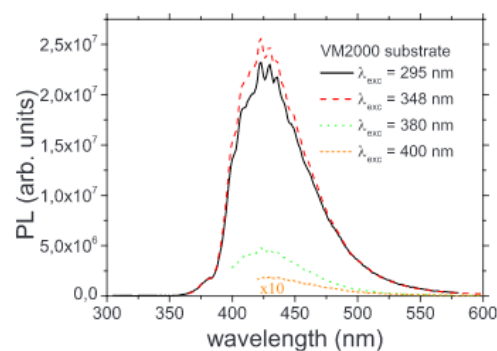
High threshold

Run Cycle ended

# Detector calibrations



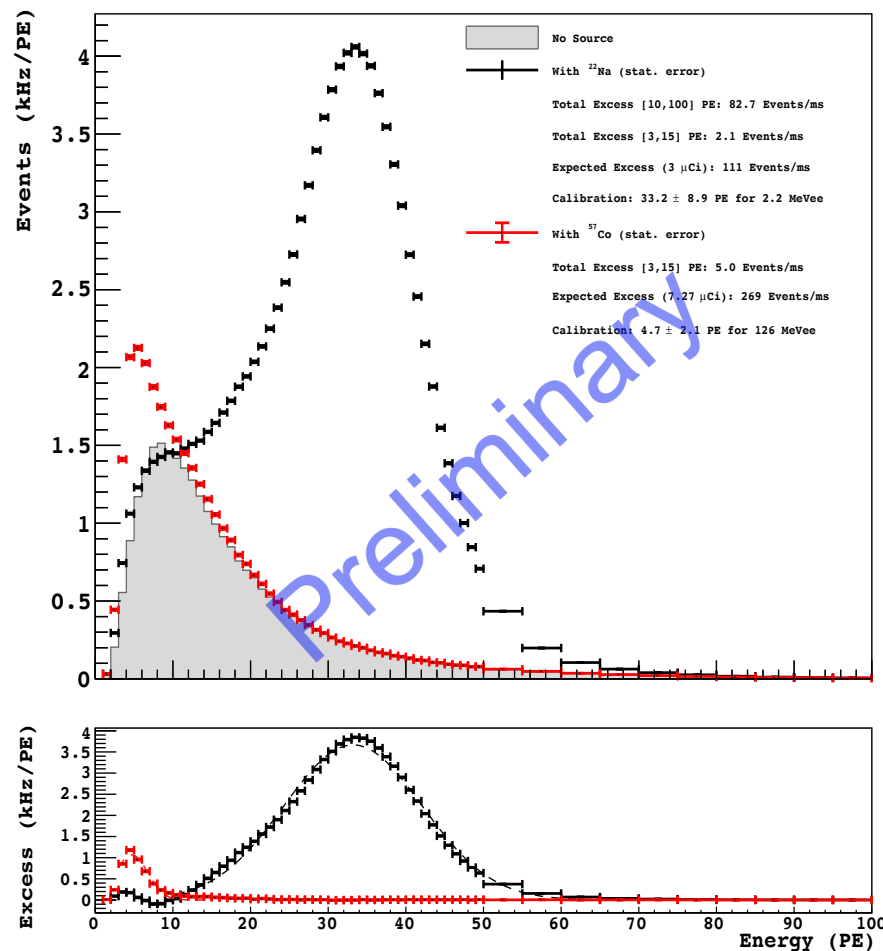
- 120 R5912 PMT's, wavelength shifting TPB foils
- LAr emission peaks at 128 nm, below the range where PMTs are sensitive



- Laser/Diffuser for 213/532 nm calibrations to test TPB response for foils and PMTs.
- LED calibrations for PMT gain/timing
- Co-57 source provide energy scale calibration 122 keV gamma-ray.
- Na-22 source provide energy scale calibration 2.2 MeV gamma-rays
- Radioactive sources provides position reconstruction calibration.

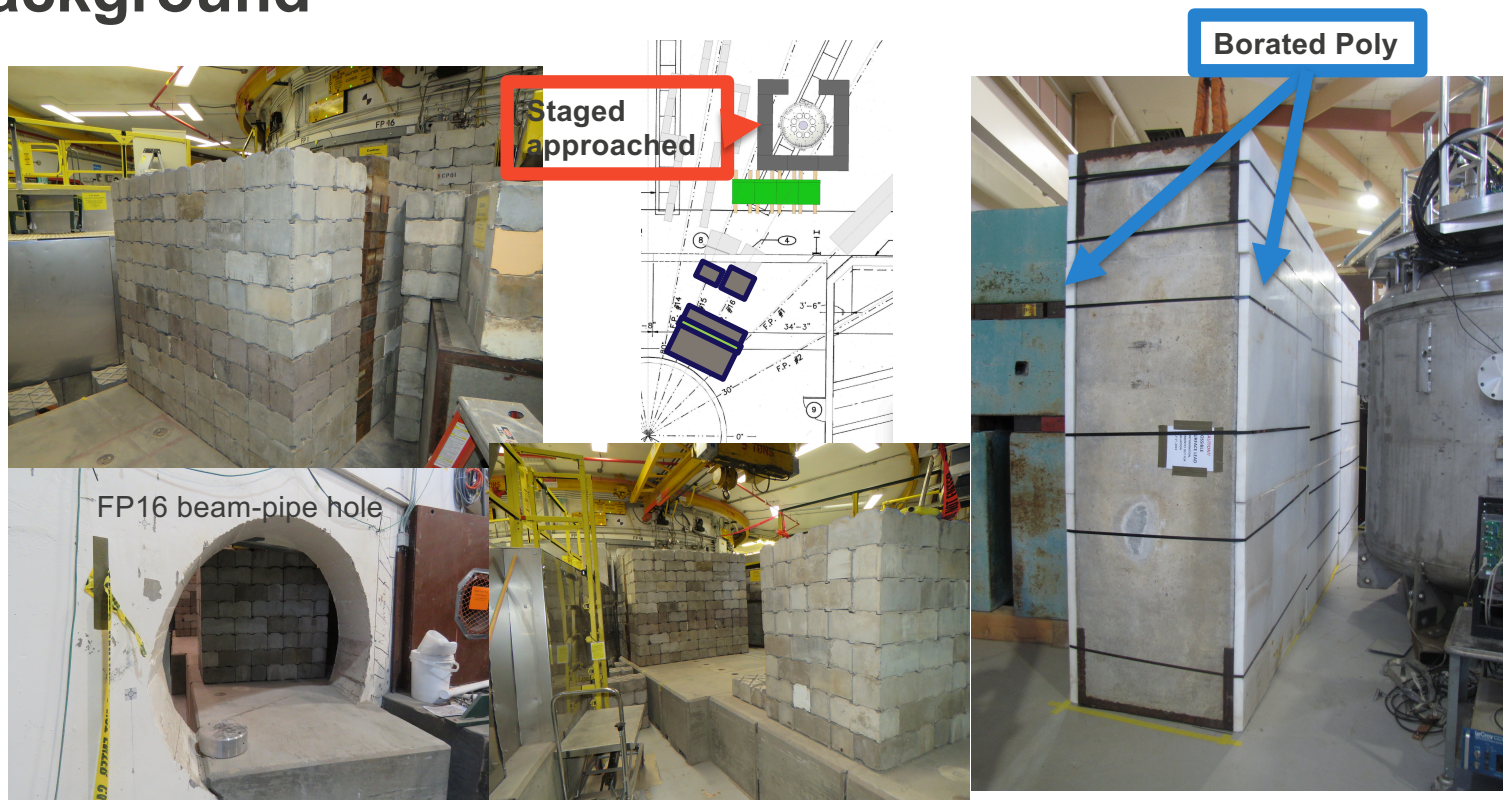
# Results from Calibration

- Impurities from not recirculating or filtering the argon led to low light levels O(ppm) O<sub>2</sub> reduced the 128 nm light attenuation length from O(10 m) to ~40 cm
- According to simulations the 4.7 PE peak for Co57 is an artifact of the event cuts, the real peak is 1.8 PE
- Na22  $33.2 \pm 8.9$  PE for 2.2 MeV
- Both Co57 and Na22 rates are within 25% of simulation prediction



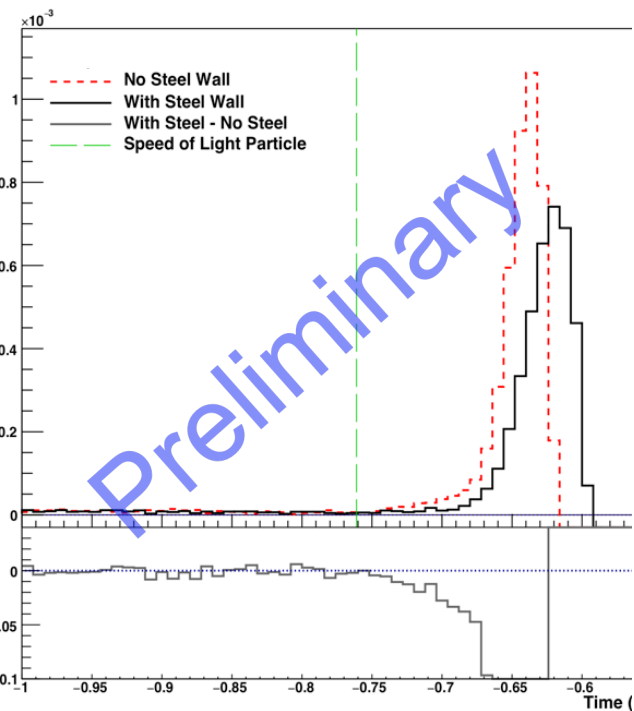


# Added Shielding to Reduce Neutron Related Background

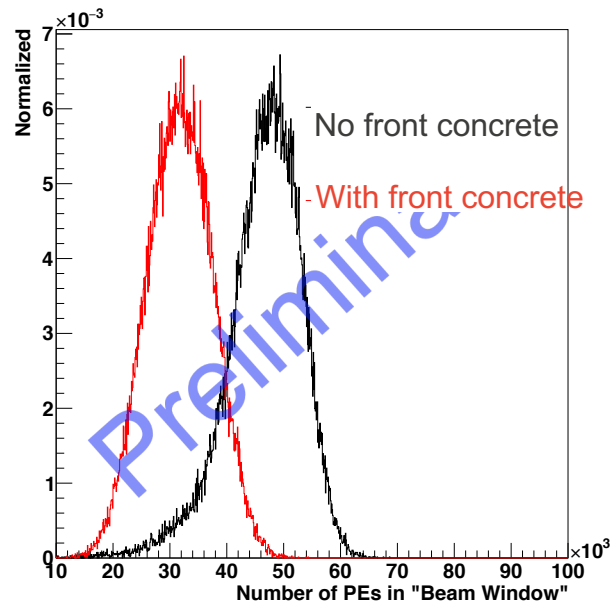


At the beginning of 2019 run we methodically and purposefully added and modified shielding to understand and reduce our background rates

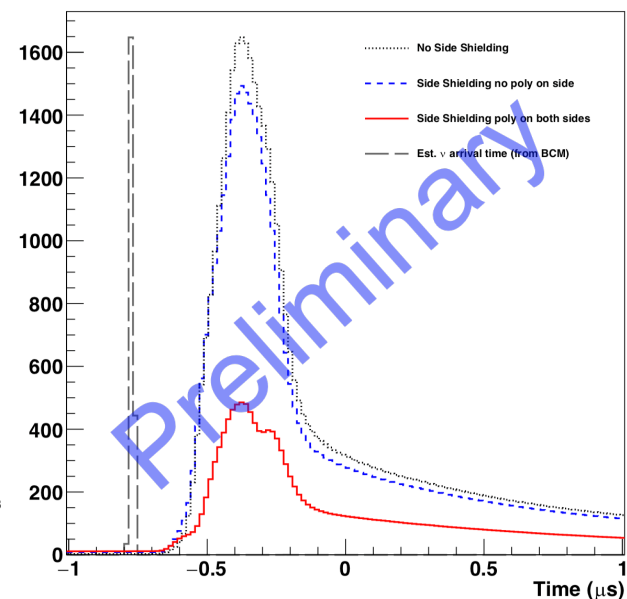
# Effects of shielding



**Steel shield: decreases beam-related rate and increases delay**



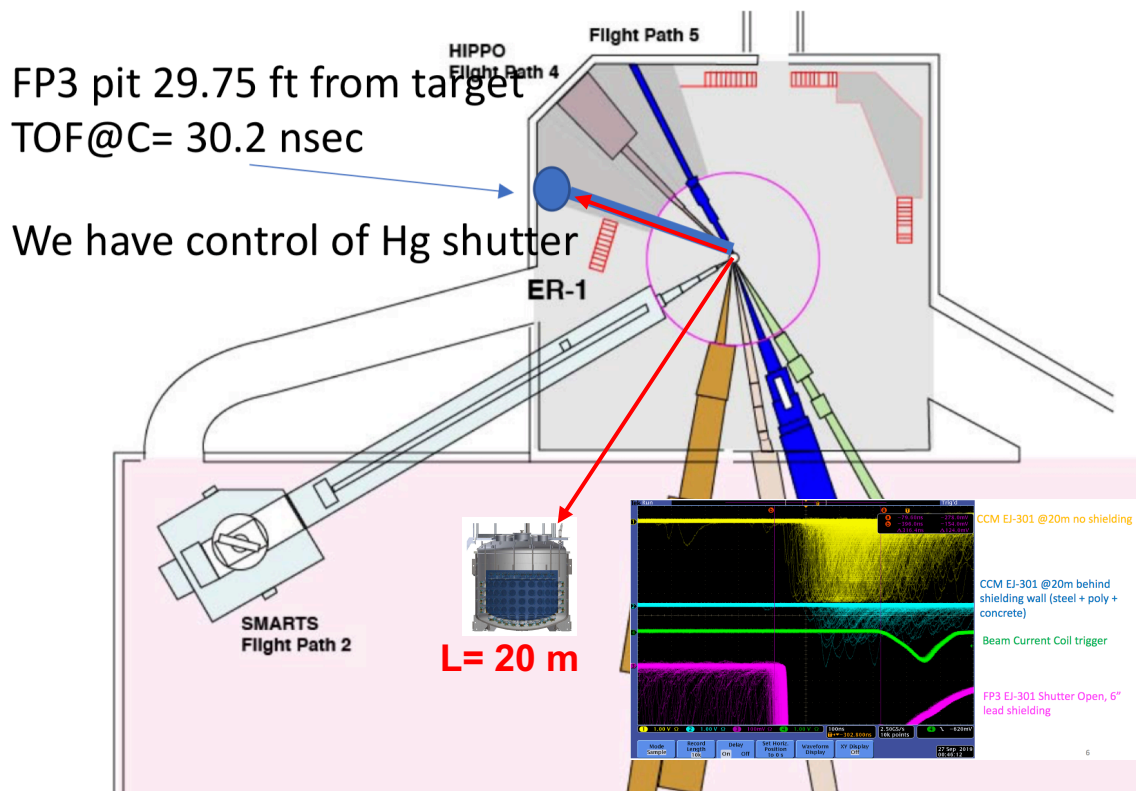
**Front concrete: decreases beam-related energy deposition in CCM**



**Side shielding: decreases beam-related rate**

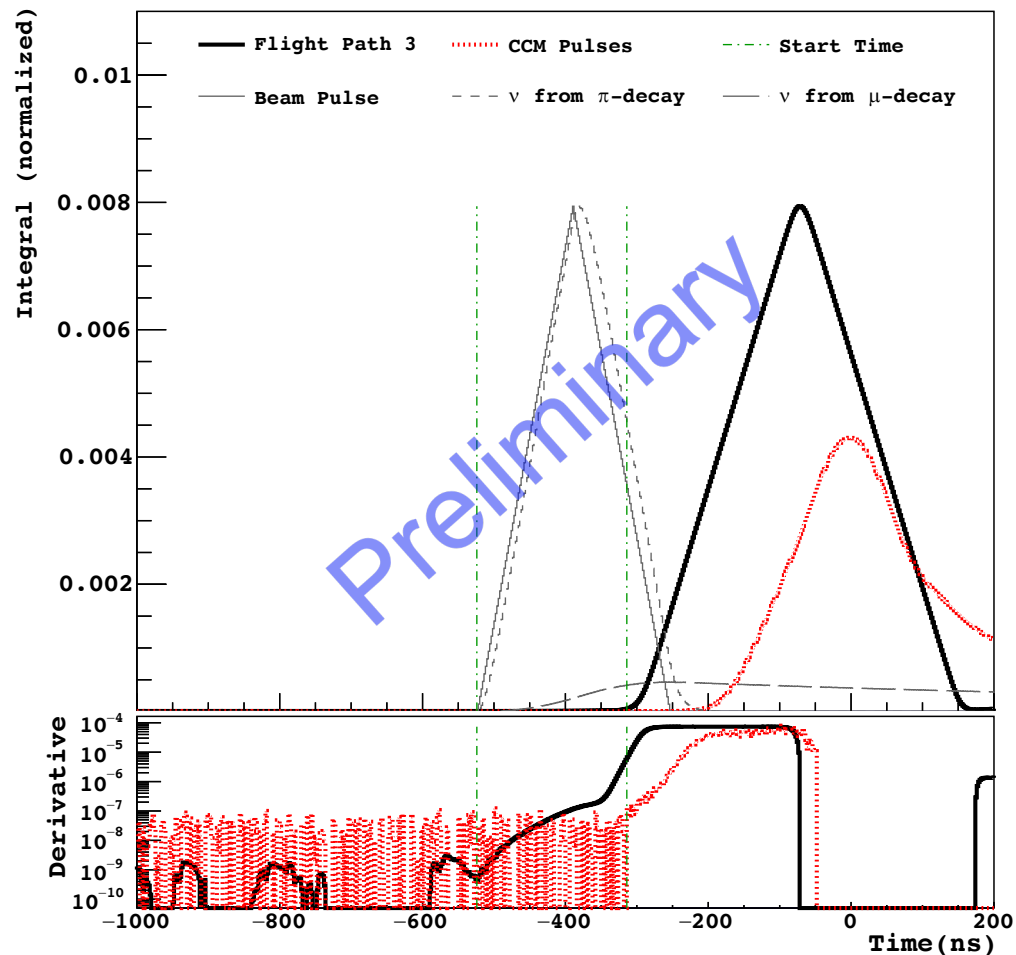
# EJ301 Detector Placed in Flight Path 3

- Liquid scintillator detector sitting on beam line with no shielding between it and the target to observe  $\gamma$ -flash
- Used to measure the time offset between the  $t_0$  and the CCM events time
- Not all time delays in the trigger signal could be measured



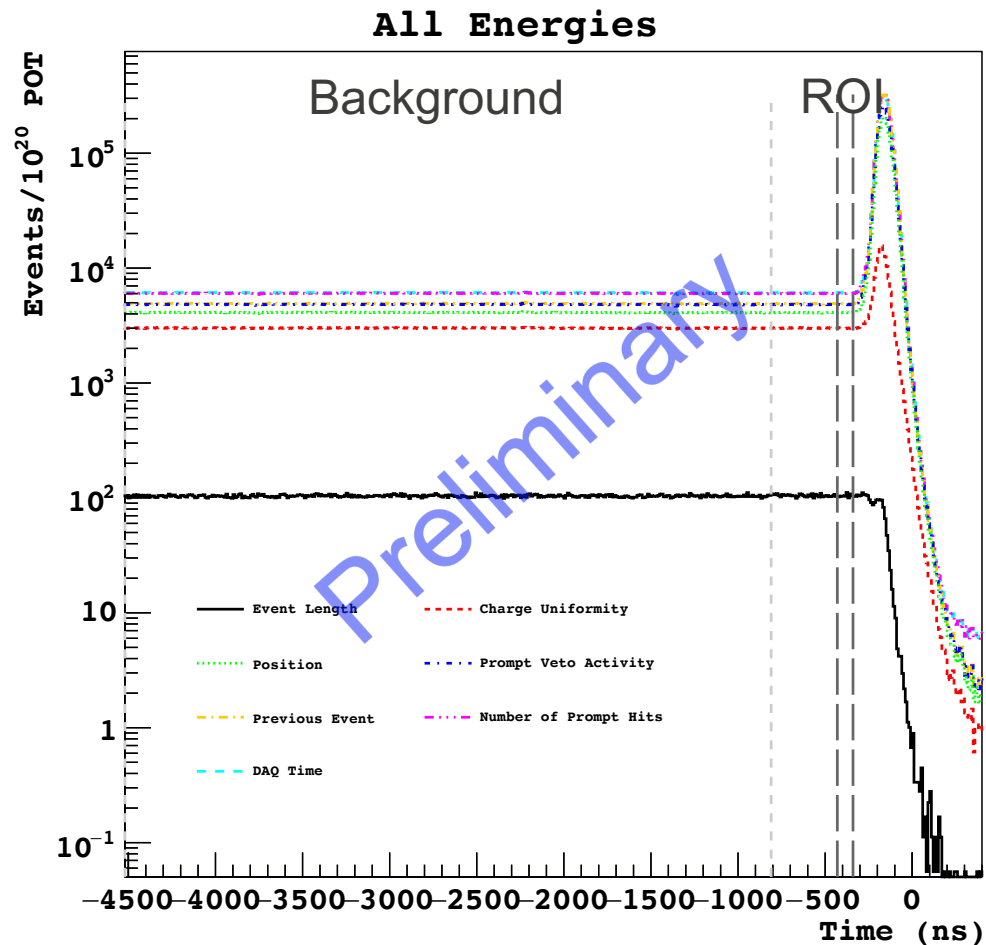
# Beam Related Background Free Region

- Based on the turn on of the FP3 detector, we expect speed of light particles from  $\pi^0$  decay to arrive 210 ns before events we seen in CCM
- Because of change in efficiencies of cuts near the CCM turn on the signal region will be 190 ns
- 190 ns consists of:
  - 80% of  $\pi^0$ -decay events
  - 74% of  $\pi^\pm$ -decay events
  - 4% of  $\mu^\pm$ -decay events that would fall within our DAQ window

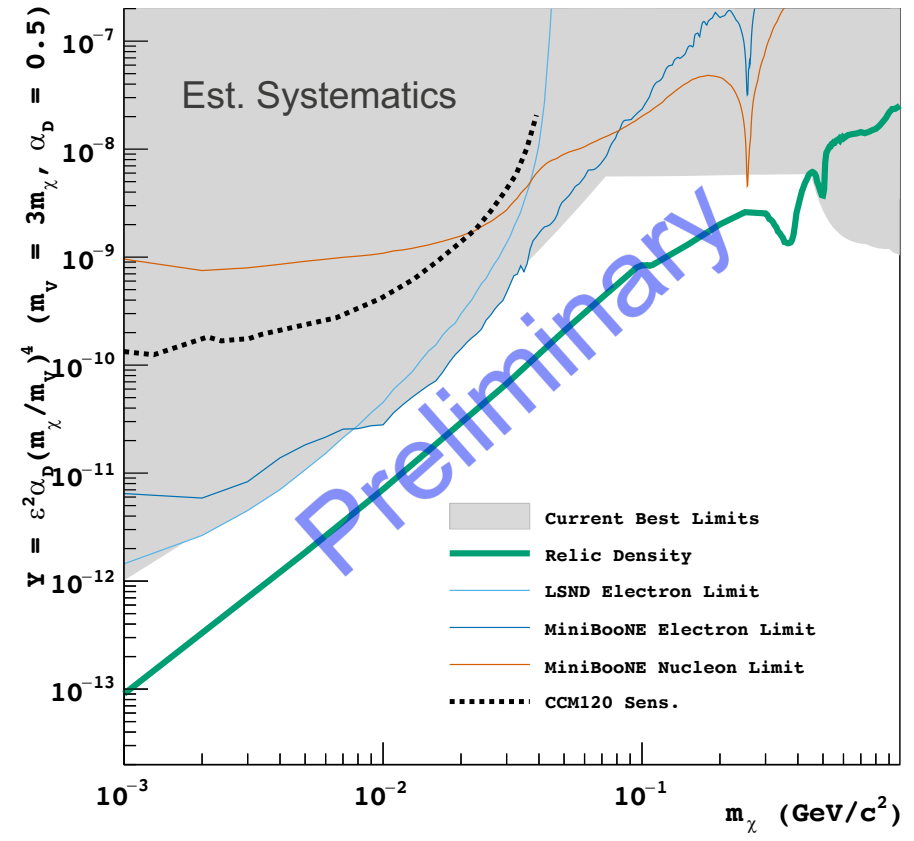
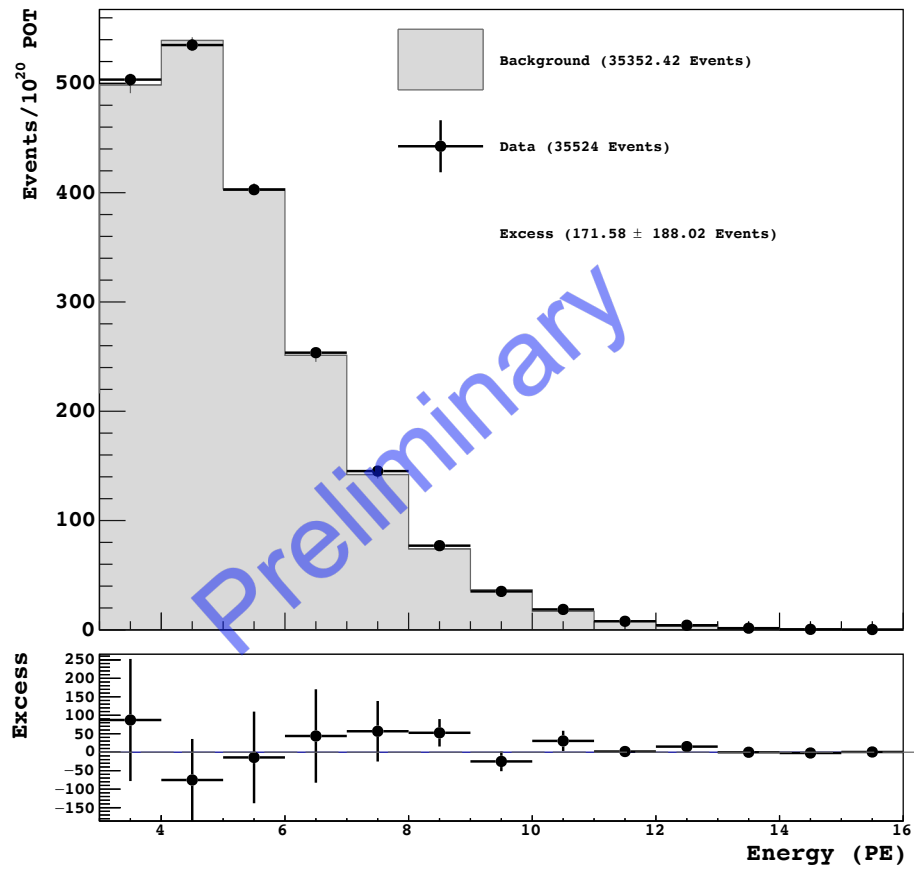


# Beam Related Background Free Region

- Prompt light only analysis
- Dynamic event lengths allow a poor-mans PID
  - Maximize dark matter over Ar39 puts the length cut at 44 ns
- Pre-beam is flat in time allowing a good prediction of what to expect in the ROI
- ROI is a beam-related background free region, so the prediction on the number of events is statistical only (systematics will be on DM signal)



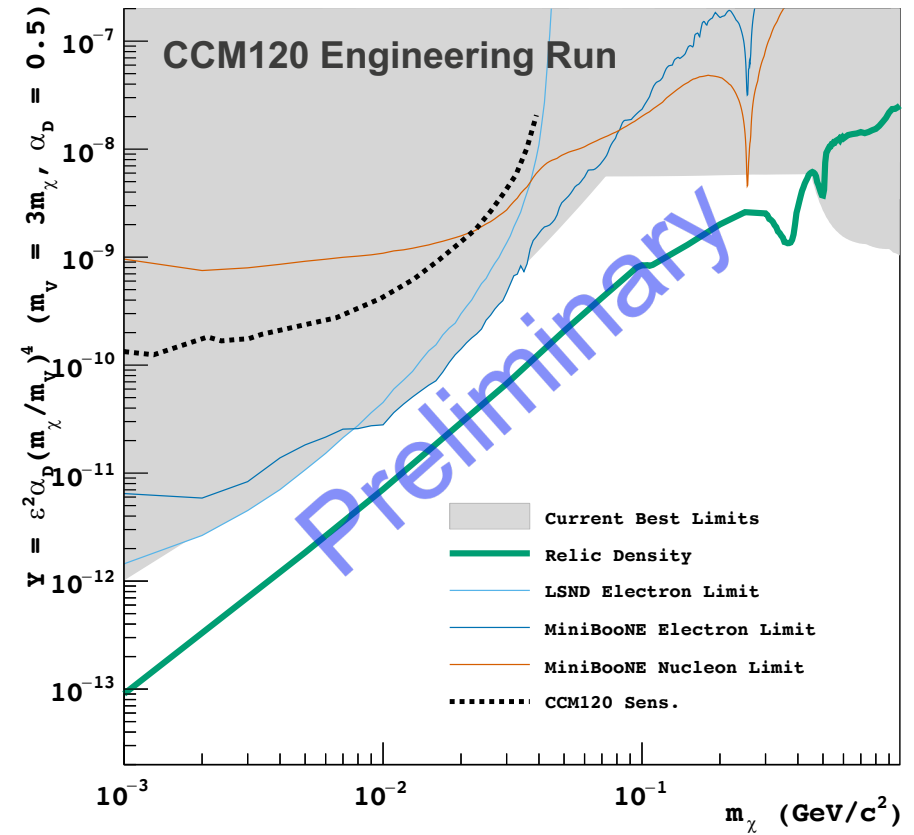
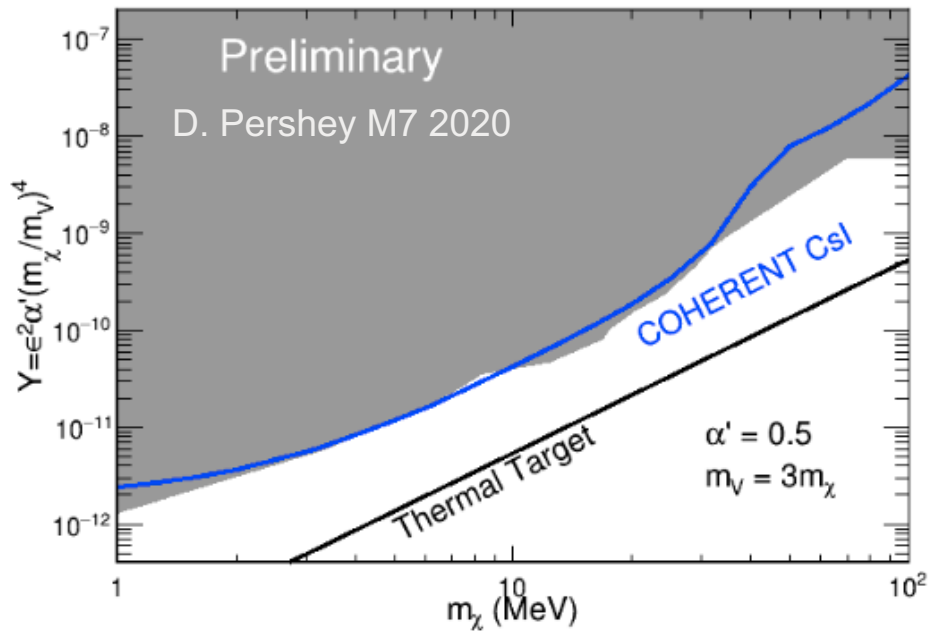
# Observed Events and Predicted Sensitivity



Preliminary

Preliminary

# Compared to COHERENT CsI



# Lessons Learned

- Measured  $t_0$ : Have ~200 ns beam-related neutron background free region
- Beam-unrelated background is 3x higher than expected Ar39 rate
- O(ppm) contamination of O<sub>2</sub> and H<sub>2</sub>O reduced the attenuation length of 128 nm light from O(10m) to ~40 cm
- Even with the bad argon, and the high background rate, and only 1.5 months of data taking, we predict we will set a strong nucleon only dark matter search



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# Upgrades to the Coherent CAPTAIN Mills experiment for the upcoming CCM200 run

Eric Renner<sup>1</sup>

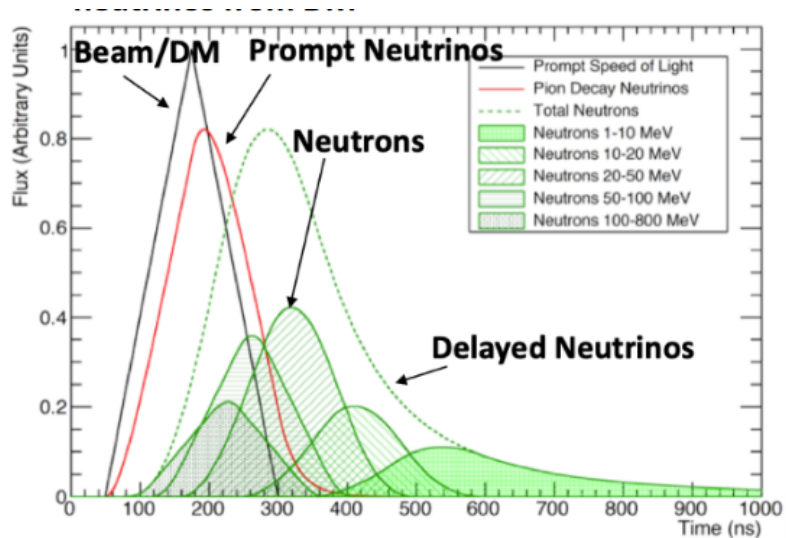


<sup>1</sup>Los Alamos National Laboratory  
*Physics Division*  
*P-25: Subatomic Physics*

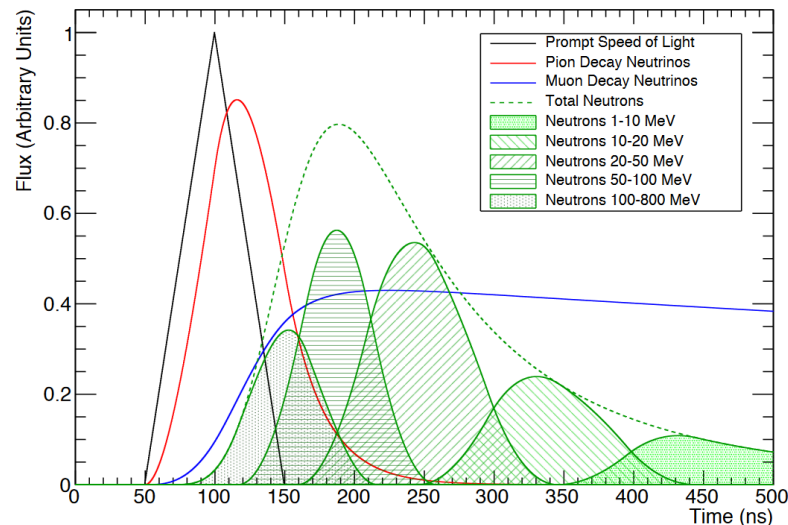
# Anticipated upgrades for the upcoming run...

- Double the number of main and veto PMTs
- Decrease the width of the LANSCE proton beam pulse
- Filter O<sub>2</sub> and H<sub>2</sub>O contaminants that absorb light from scattering events
- Eliminate the LAr boil off to maintain consistent purity levels inside the cryostat
- Shield the detector from unwanted background events (gamma rays, stray neutrons, etc.)

# Improved beam tuning to decrease beam width increases the sensitivity of CCM to prompt neutrinos



275ns Beam Width



100ns Beam Width

- With 100ns beam width, our region of interest is outside the measured neutron wave ( $E = 20\text{-}50\text{MeV}$ )

\*Figures from T.J. Schaub – Searching for Sterile Neutrinos and Accelerator Produced Dark Matter with the Coherent CAPTAIN-Mills (CCM) Detector at the Los Alamos Neutron Science Center

# Contaminants in bulk fluid from the gas plant

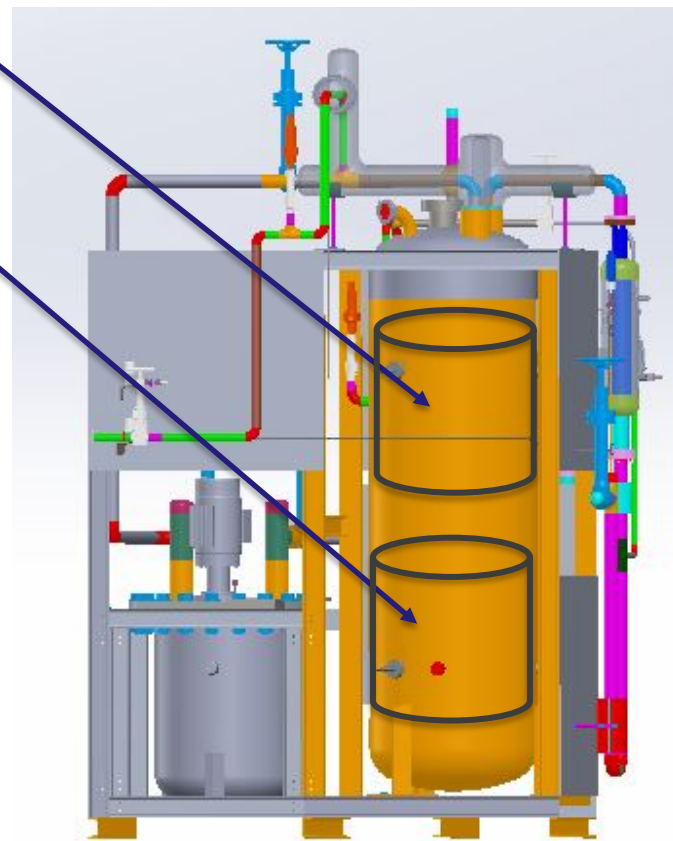
- O<sub>2</sub> = 1.95ppm
  - Absorb scintillation light.
- H<sub>2</sub>O = 0.01ppm
  - Absorb scintillation light
  - We will evacuate the cryostat and additional piping to minimize outgassing
- N<sub>2</sub> = 2.50ppm
  - Effects triplet light output. Can handle up to ~10ppm
- Verified by simulations tuning to our calibration runs

## Analysis from Matheson gas company

Weight: [lbs]	Tare :	<u>32020</u>
	Gross :	<u>74120</u>
	Net :	<u>42100</u>
	Gallons :	<u>3619.1</u>
	SCF :	<u>407149</u>
Analysis	Pretest :	<u>2.23</u> PPM
	Post Fill Assay :	<u>99.999</u> %
	Oxygen :	<u>1.95</u> PPM
	N <sub>2</sub> :	<u>2.50</u> PPM
	CO :	<u>N/A</u> PPM
	H <sub>2</sub> O :	<u>0.01</u> PPM
	THC :	<u>N/A</u> PPM
	CO <sub>2</sub> :	<u>N/A</u> PPM
	Hydrogen :	<u>N/A</u> PPM
	Odor :	<u>None</u>

# Filtration skid to remove O<sub>2</sub> and H<sub>2</sub>O contamination

- 4A molecular sieve material to remove water contamination
- Cu Alumina to remove oxygen contamination
- Proven MicroBooNE filtration skid design can achieve concentration in the single digit ppb range



# O<sub>2</sub> Contamination causes light absorption in the cryostat

- Goal to achieve absorption length of greater than 250cm
- Current absorption length is ~40cm

Concentration of O <sub>2</sub> (ppb)	Absorption Length in LAr (128nm Light)
0 (pure LAr)	2,000 cm
2	1,700 cm
20	1,000 cm
100 (*)	250 cm
200	180 cm
2000 (from plant)	20 cm
20000	2 cm

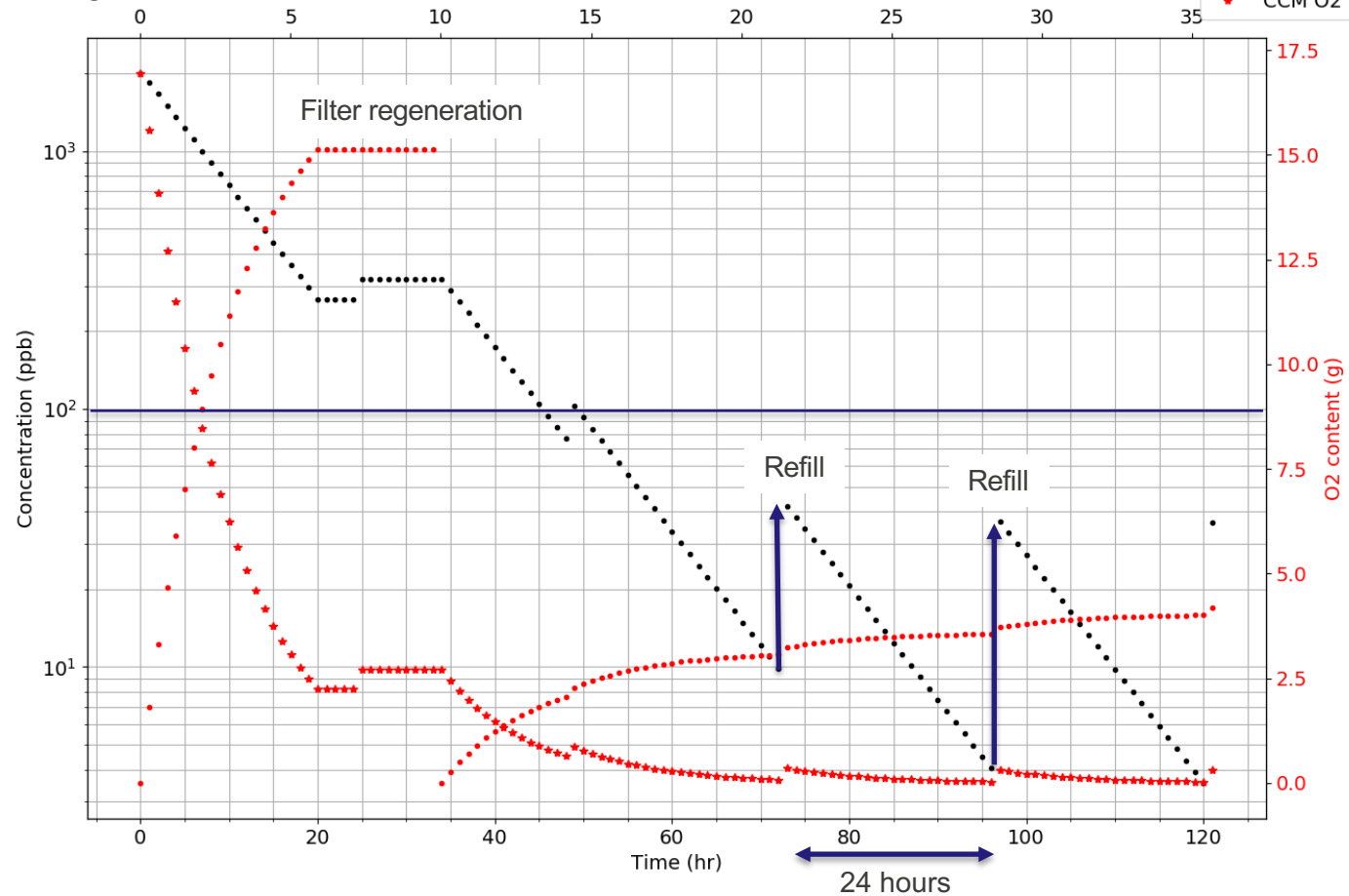
Table generated by E. Dunton

# Model of filtration skid performance

— Goal Concentration  
Absorption length > 250cm

Volume exchanges (1 VE = ~3.4 hr)

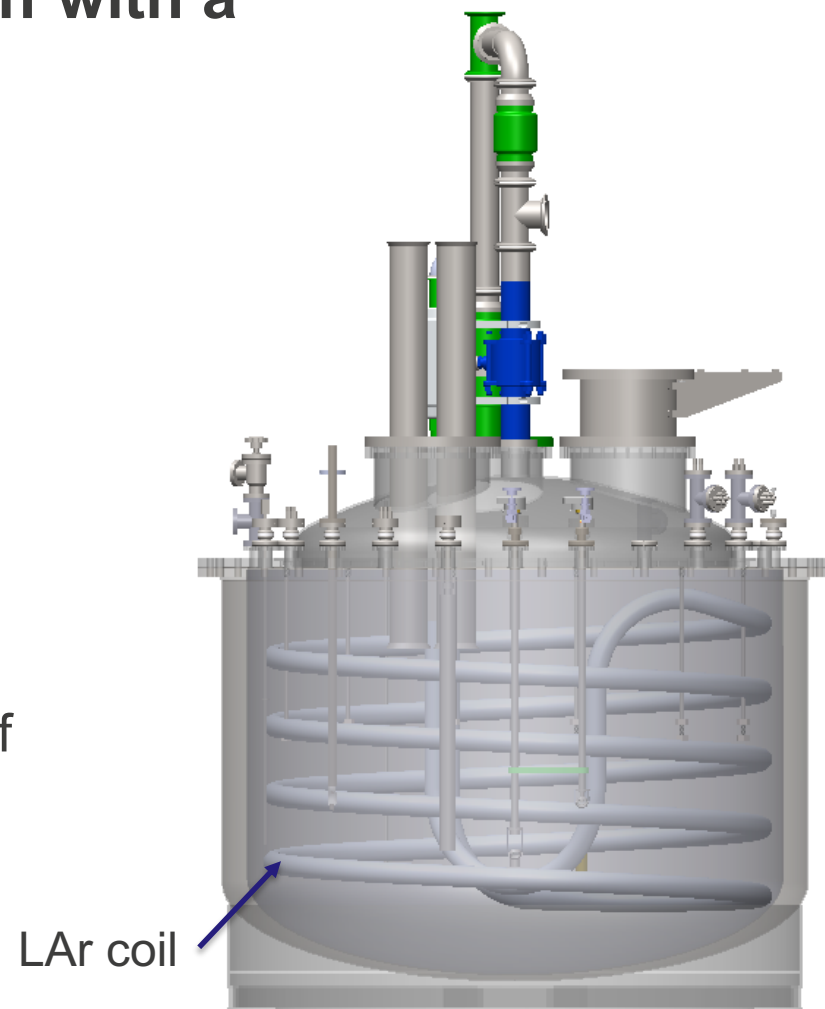
- CCM O2 Conc
- Filter O2 Cont
- ★ CCM O2 Cont



- Refill with 220L of “dirty” Ar every 24 hours leads to a periodic concentration curve (35ppb mag.)

# Eliminate 220L per day boil off with a LN bath heat exchanger

- 60' LAr line submerged in a bath of liquid nitrogen
- Maximum Cooling Power is  $1.6kW$ .
- Can control the head pressure and total submerged length of pipe to match the required cooling power of  $1kW$



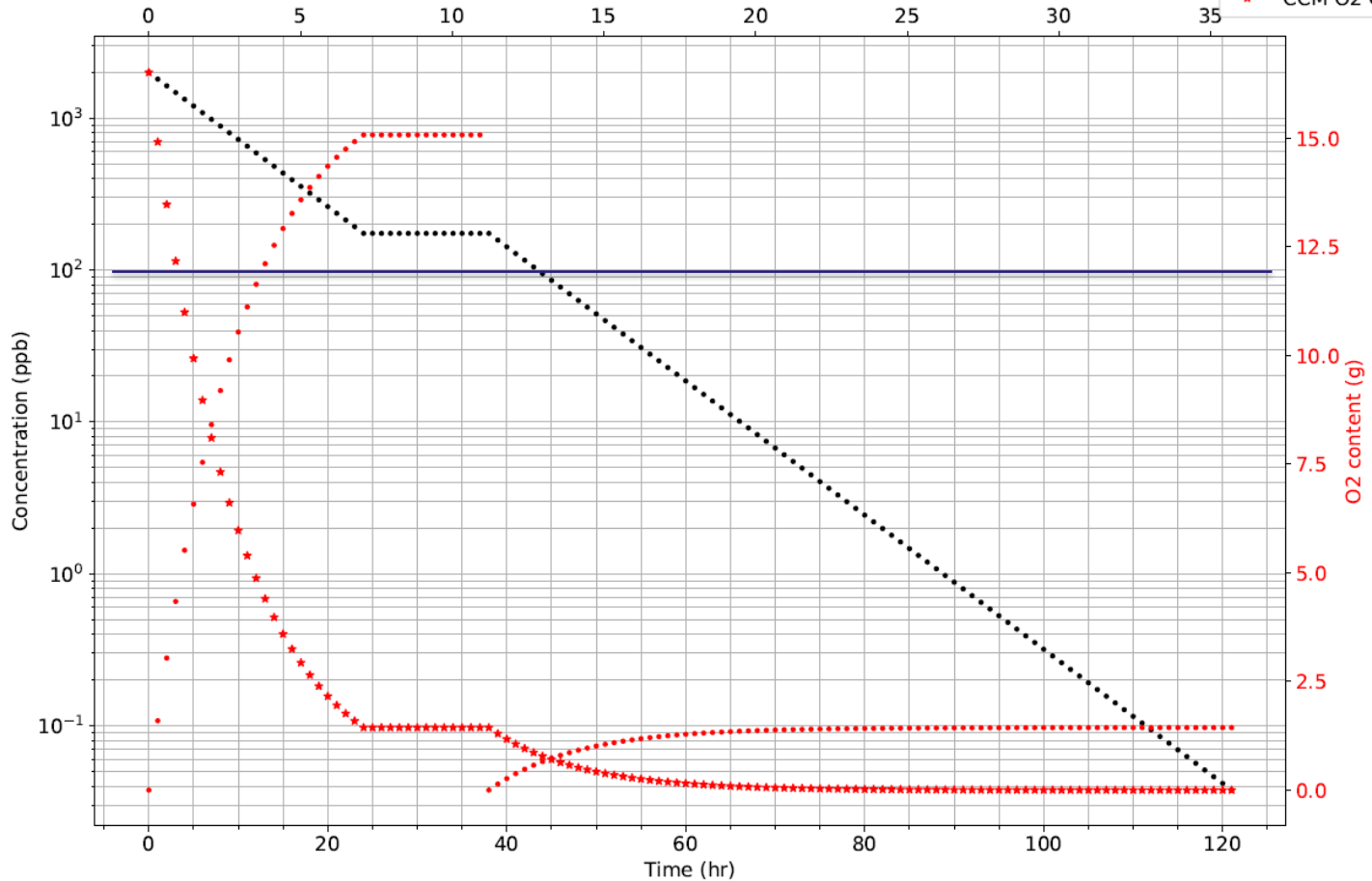


# Model of filtration skid performance

Goal Concentration  
Absorption length > 250cm

Volume exchanges (1 VE = ~3.4 hr)

- CCM O2 Conc
- Filter O2 Cont
- \* CCM O2 Cont



- Maintain lower purity in the detector without boil off

# Shield the detector from background events.

- Current configuration, leads to gamma rays, sky shine.

- Concrete Blocks

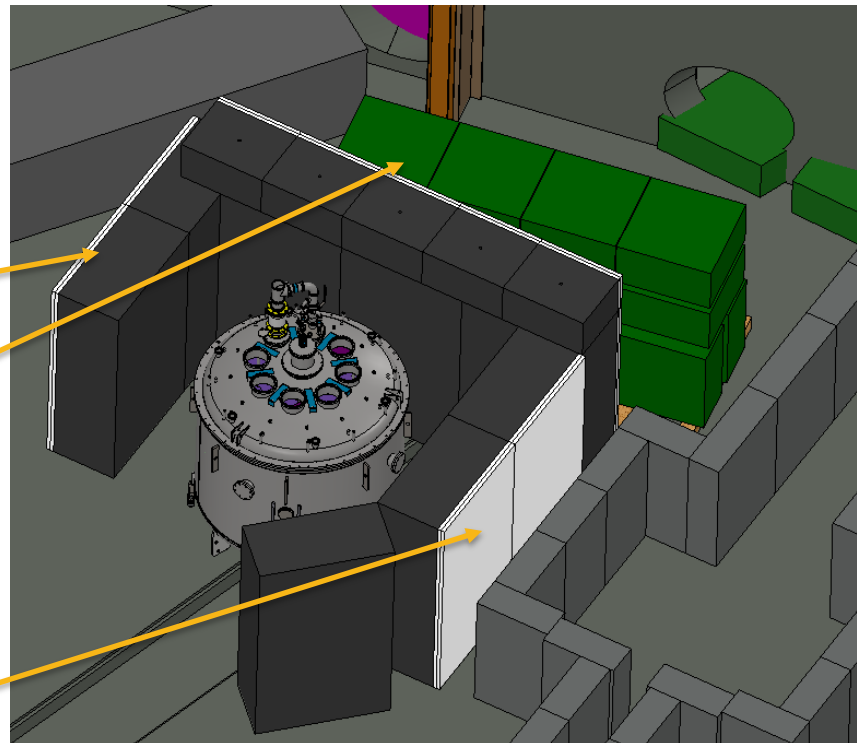
- 36" thick

- Steel Blocks

- 52" thick

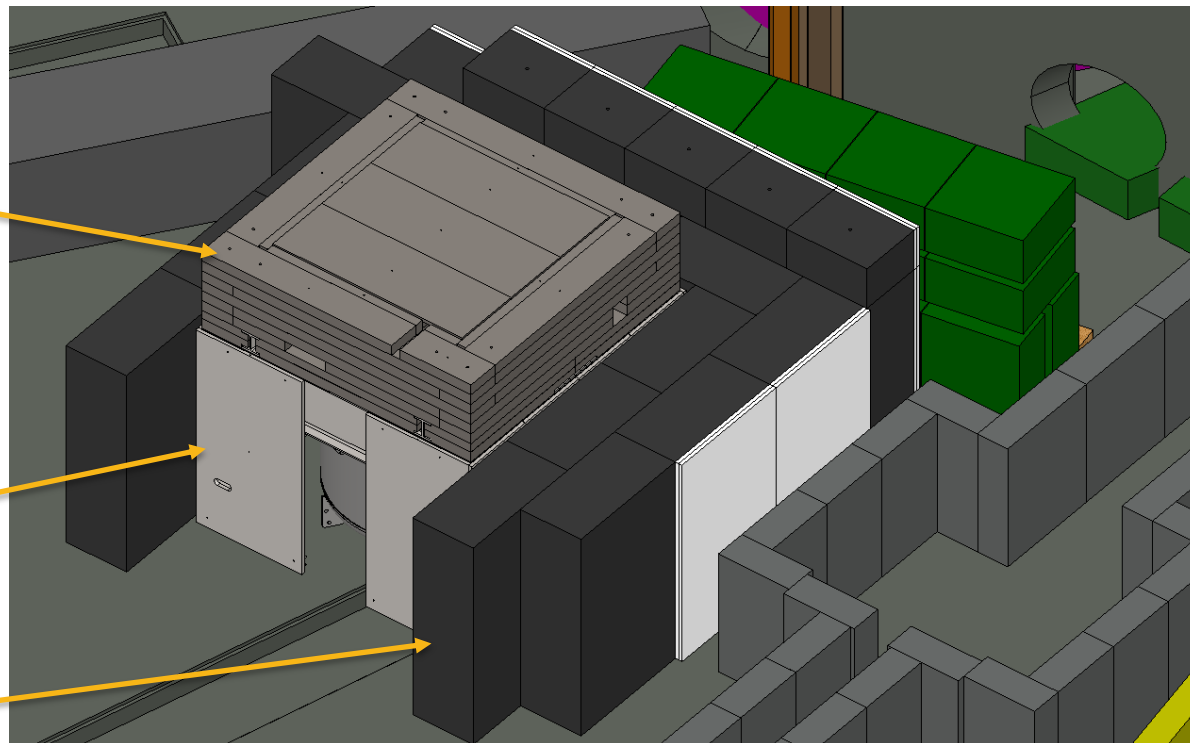
- Borated poly

- 2" thick



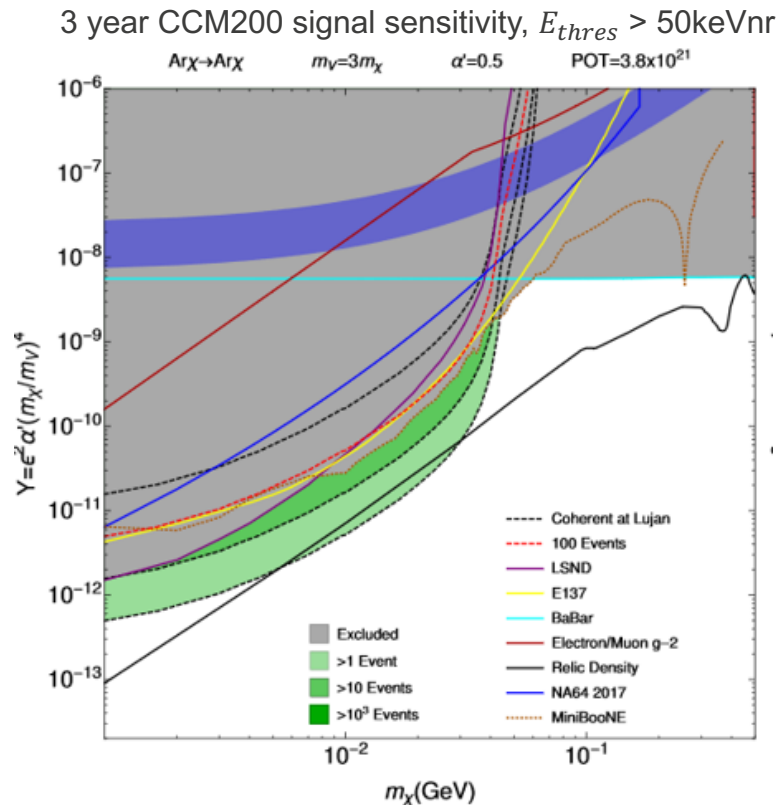
# New shielding designed to address out of time beam induced background

- Steel to be installed on top of the detector to shield from fast neutrons
- 2" steel walls can capture gamma rays created inside the concrete blocks.
- Additional Concrete



# CCM200 will be an impressive experiment and make massive improvements to CCM120

- Decreased width of beam pulse at LANSCE will increase sensitivity in the region of interest.
- The filtration skid will reduce O2 and H2O contaminants that absorb light from scattering events and increase sensitivity of the detector.
- The heat exchanger will eliminate LAr boil off to maintain consistent purity levels inside the cryostat.
- Shielding installed around the detector will decrease unwanted background events (gamma rays, stray neutrons, etc.).
- CCM R&D lessons learned provide critical design input to FNAL dedicated stopped pion source facility. See J. Zettlemoyer talk on Thursday



This work was funded by LANL LDRD and HEP DMNI