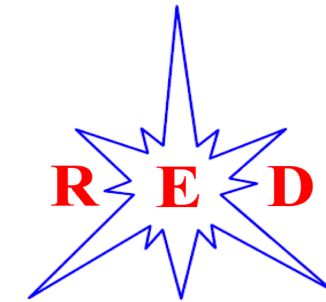




# Status of the RED-100 experiment

Rudik D. G.

NRNU MEPhI



RUSSIAN EMISSION DETECTOR

Magnificent CEvNS 2020

Cyberspace

16-20.11.2020



# RED-100 collaboration



**RUSSIAN EMISSION DETECTOR**

**D.Yu. Akimov<sup>a</sup>, I.S. Aleksandrov<sup>a,g</sup>, V.A. Belov<sup>a,b</sup>, A.I. Bolozdynya<sup>a</sup>, Yu.V. Efremenko<sup>f</sup>,  
A.V. Etenko<sup>a,d</sup>, A.V. Galavanov<sup>a,c</sup>, D.V. Gouss<sup>a</sup>, Yu.V. Gusakov<sup>a,c</sup>, Dj.Ed. Kdib<sup>a</sup>, A.V. Khromov<sup>a</sup>,  
A.M. Konovalov<sup>a,b</sup>, V.N. Kornoukhov<sup>a,e</sup>, A.G. Kovalenko<sup>a,b</sup>, A.A. Kozlov<sup>a</sup>, E.S. Kozlova<sup>a,b</sup>,  
A.V. Kumpan<sup>a</sup>, A.V. Lukyashin<sup>a,h</sup>, Yu.A. Melikyan<sup>a,g</sup>, V.V. Moramzin<sup>a</sup>, A.V. Pinchuk<sup>a</sup>,  
O.E. Razuvaeva<sup>a,b</sup>, D.G. Rudik<sup>a</sup>, A.V. Shakirov<sup>a</sup>, G.E. Simakov<sup>a,b</sup>, V.V. Sosnovtsev<sup>a</sup>, A.A. Vasin<sup>a</sup>**

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Moscow, 115409, Russia

<sup>b</sup> Institute for Theoretical and Experimental Physics named by A.I. Alikhanov of National Research  
Center “Kurchatov Institute”, Moscow, 117218, Russia

<sup>c</sup> Joint Institute for Nuclear Research, Dubna, Moscow region, 141980, Russia

<sup>d</sup> National Research Center “Kurchatov Institute”, Moscow, 123098, Russia

<sup>e</sup> Institute for Nuclear Research, Moscow, 117312, Russia

<sup>f</sup> University of Tennessee, Knoxville, TN 37996-1200, USA

<sup>g</sup> National Research Tomsk Polytechnic University, Tomsk, 634050, Russia

<sup>h</sup> Moscow State University, Moscow, 119991, Russia

Our goal is to detect and study CEvNS @ close  
vicinity of reactor core with RED-100 detector

Supported by



НАУКА  
И ИННОВАЦИИ  
РОСАТОМ\*



Russian  
Science  
Foundation



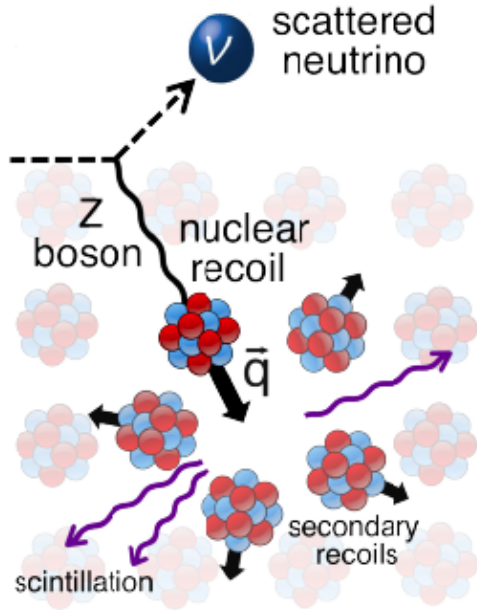
RUSSIAN  
FOUNDATION  
FOR BASIC  
RESEARCH



\* Science and innovations Rosatom



# Coherent Elastic Neutrino Nucleus Scattering (CEvNS)

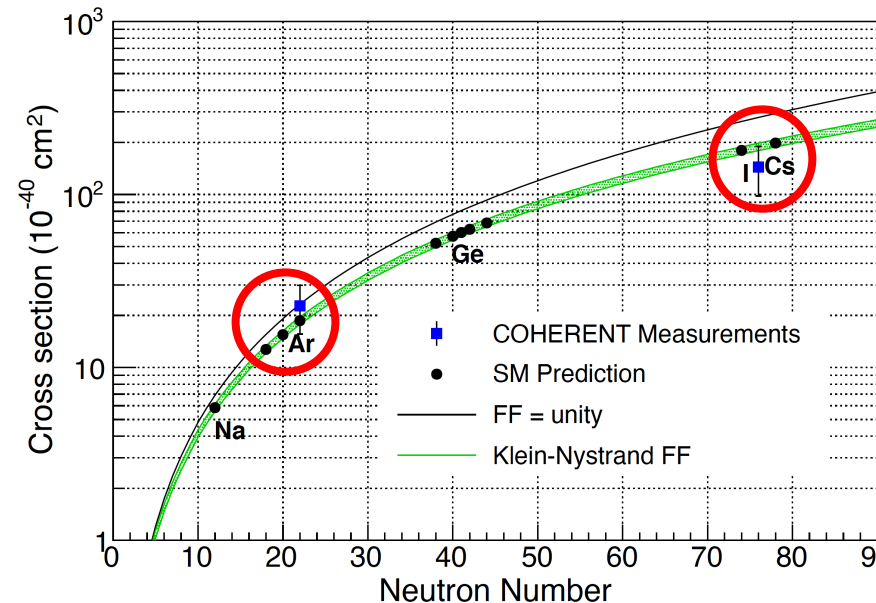
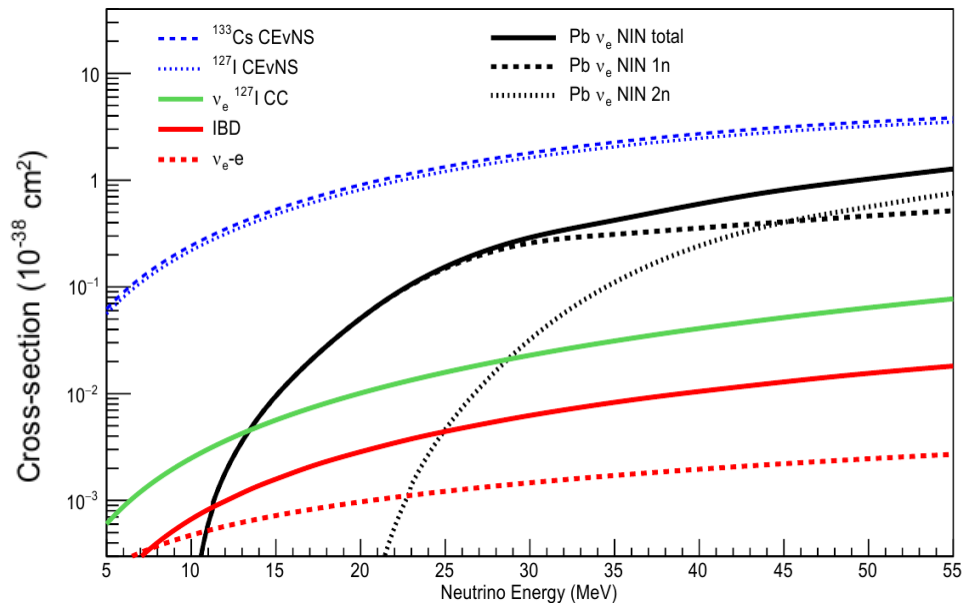


Predicted more than 40 years ago within Standard Model (SM)

- ❖ *D.Z. Freedman, Coherent effects of a weak neutral current, Phys. Rev. D 9 (1974) 1389*
- ❖ *Kopeliovich V B, Frankfurt L L JETP Lett. 19 145 (1974); Pis'ma Zh. Eksp. Teor. Fiz. 19 236 (1974)*

$$\frac{d\sigma}{d\Omega} = \frac{G^2}{4\pi^2} k^2 (1 + \cos \theta) \frac{(N - (1 - 4 \sin^2 \theta_W) Z)^2}{4} F^2(Q^2) \propto N^2$$

where  $G$  – Fermi constant,  $Z$  – number of protons,  $N$  – number of neutrons,  $F(Q^2)$  – nuclear form factor,  $Q$  – momentum transfer,  $k$  – neutrino energy,  $\theta_W$  – Weinberg angle

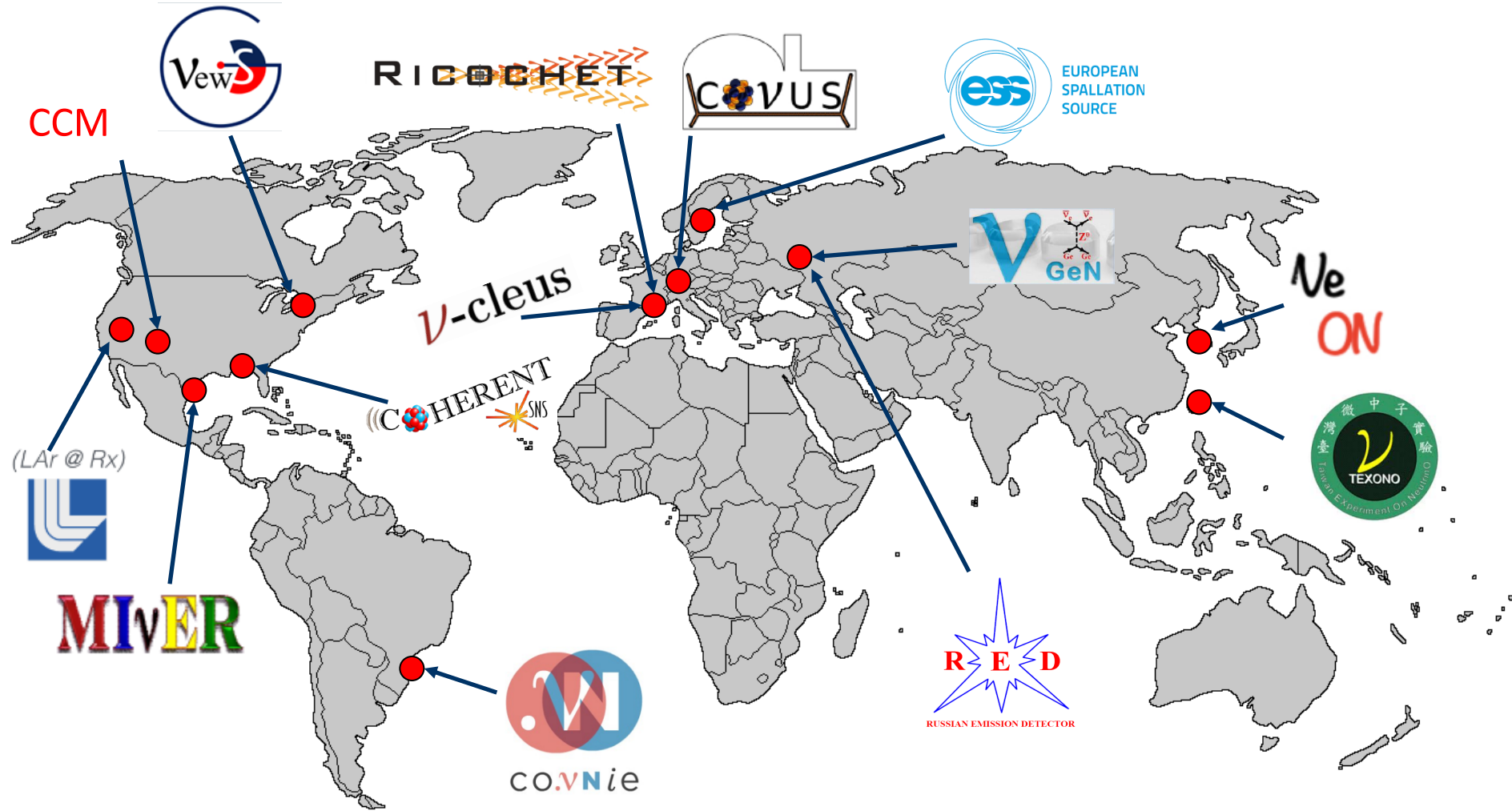


First observations:

**Cs & I** - Experimental point by COHERENT: *Science* Vol. 357 (2017) 1123

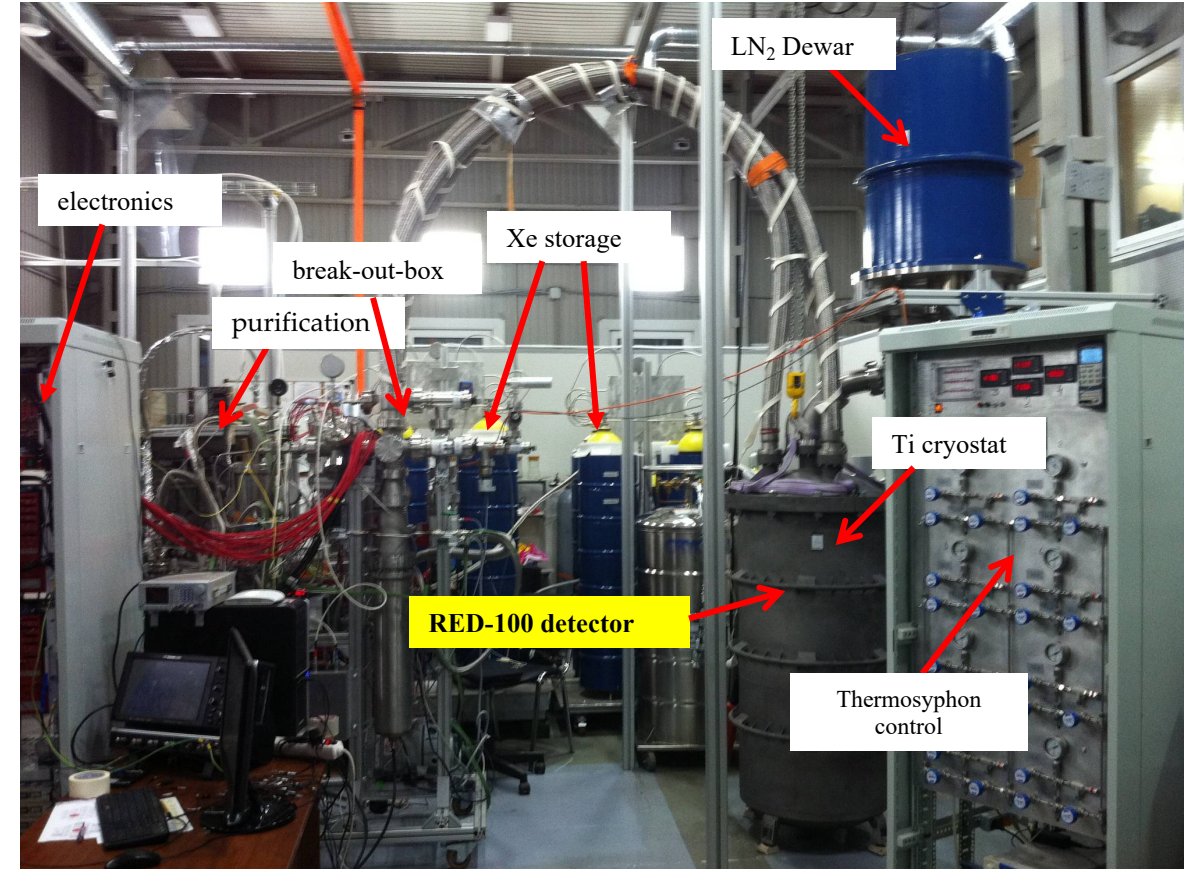
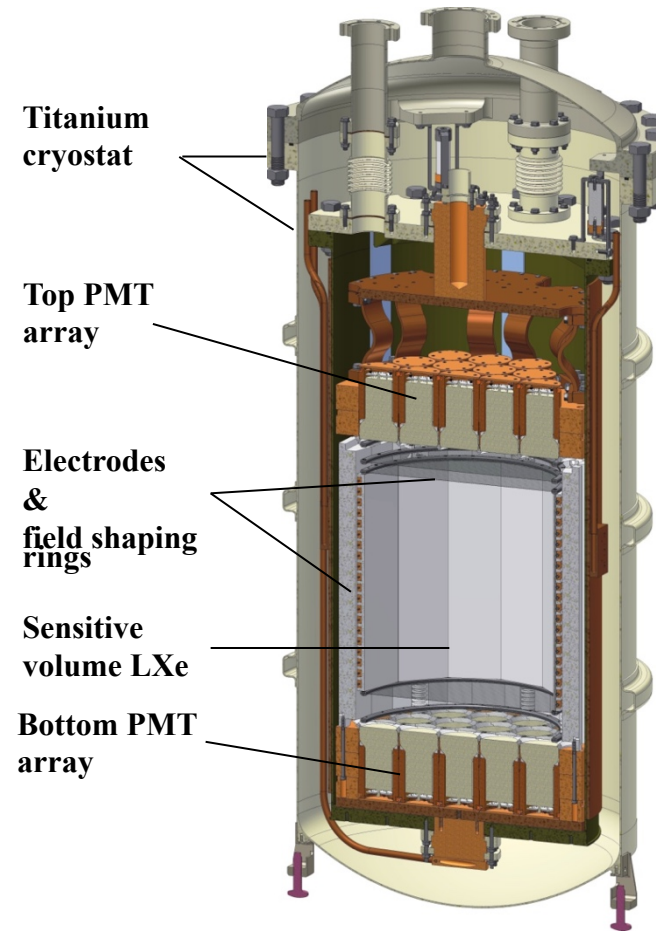
**LAr** - Experimental point by COHERENT: *arXiv: 2003.10630* (2020)

# CEvNS around the World



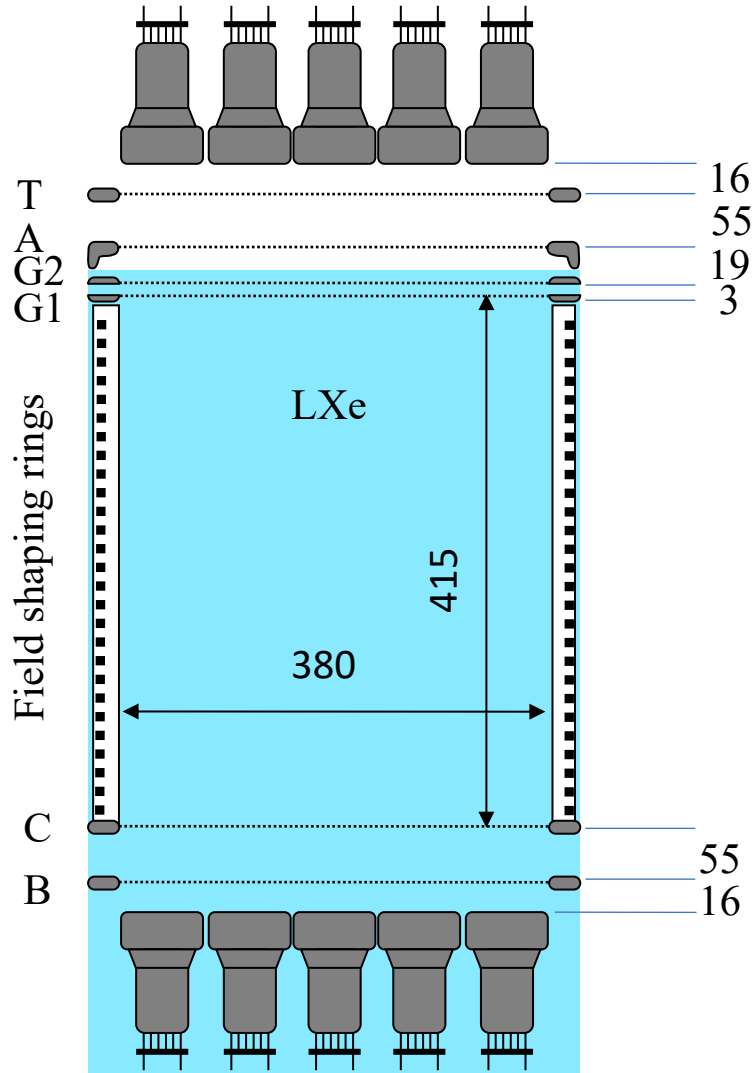
# RED-100

- Two-phase noble gas emission detector
- Contains ~200 kg of LXe (~ 100 kg in FV)
- 38 PMTs  
Hamamatsu R11410-20 (19 in each PMT array)
- Thermosyphon-based cooling system (LN<sub>2</sub>)



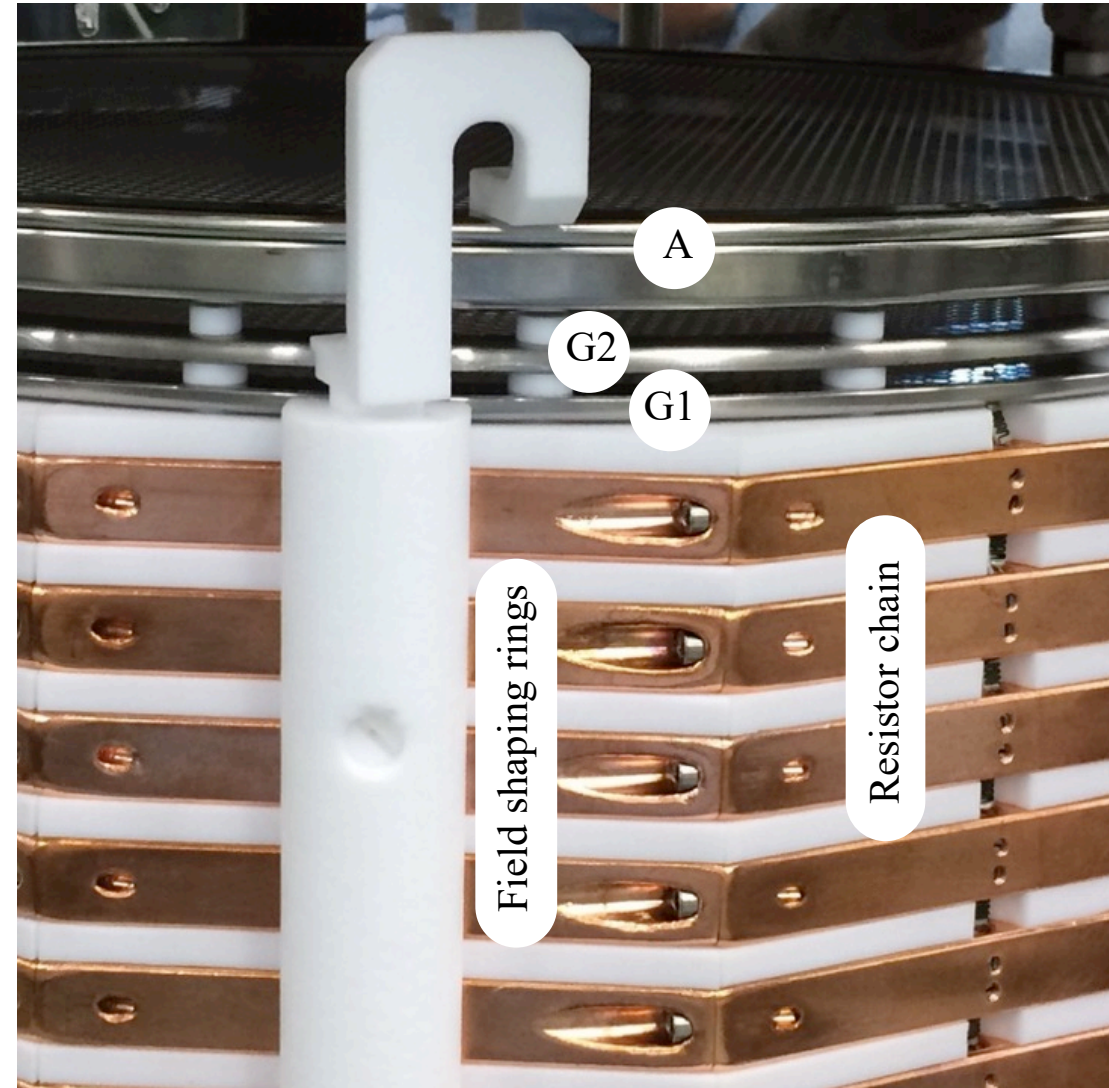


# RED-100: schematic layout of grids and PMTs



Sizes of the drift volume and distances between grids are in **mm**.

T and B – top and bottom grounded grids,  
 A – anode grid,  
 G1 – electron shutter grid,  
 G2 – extraction grid,  
 C – cathode grid



# Two-phase emission detector technique

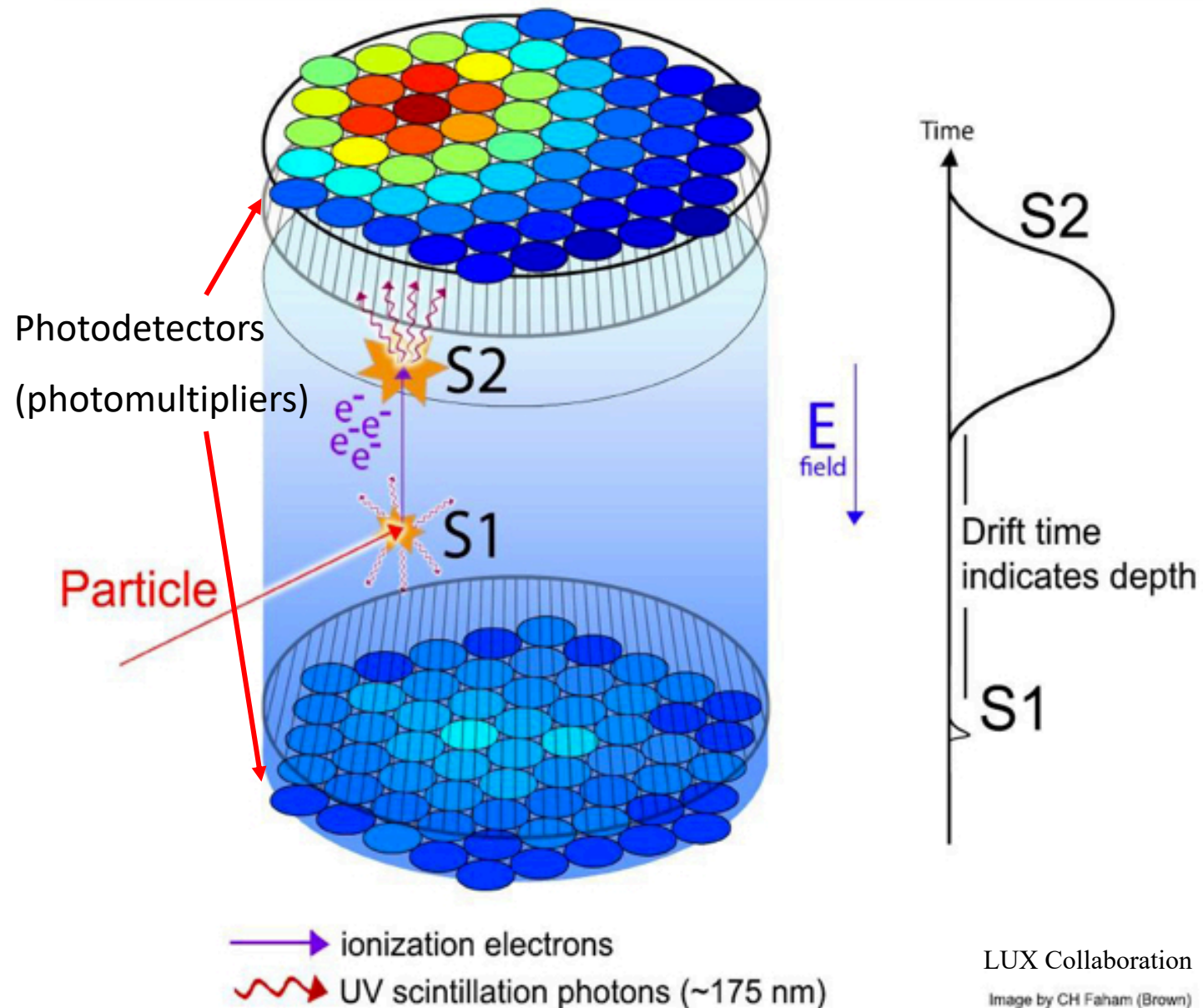
Very suitable for CEvNS study. It combines the advantages of gas detectors: the possibility of proportional or EL amplification, XYZ positioning, and the possibility to have the large mass!

This method was proposed by Russian scientists in MEPhI in 1970:

B.A. Dolgoshein, V.N. Lebedenko, B.U. Rodionov, JETF Letters (in Russian), 1970, v. 11, p. 513

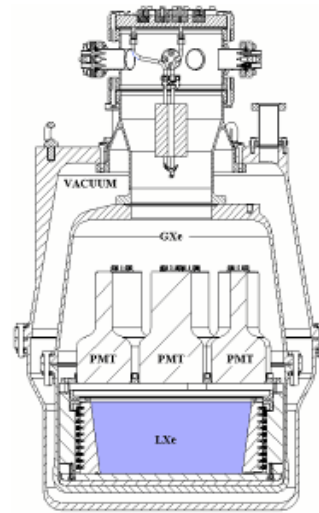
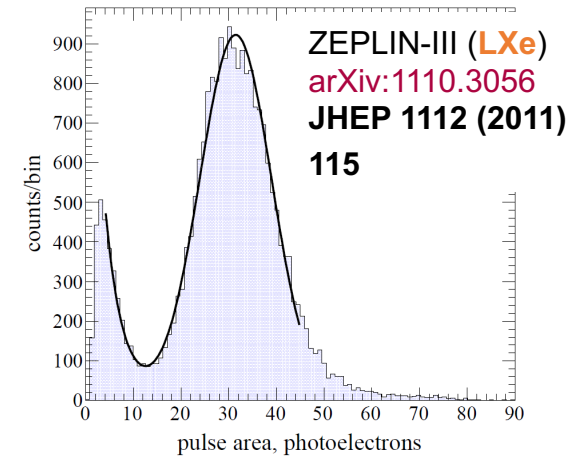
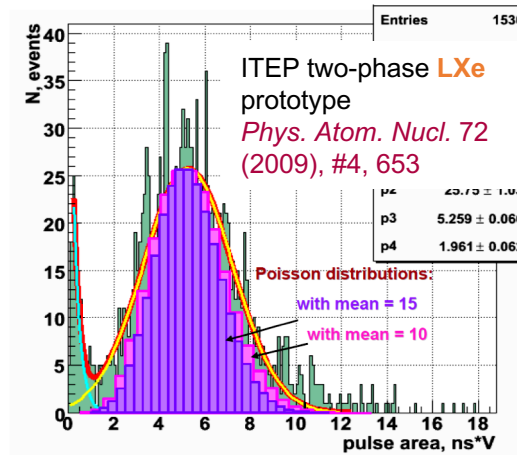
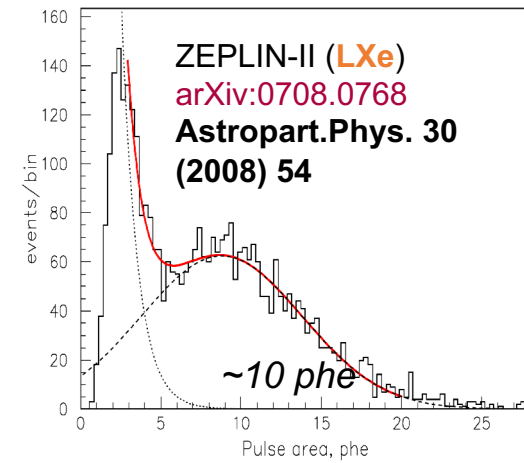
Two-phase emission detector with PMT matrices for rare events study:

Bolozdynya A. I., Egorov O. K., Rodinov B. U., Miroschnichenko V. P. (1995). Emission detectors. IEEE Trans. Nucl. Sci. 42:565-569

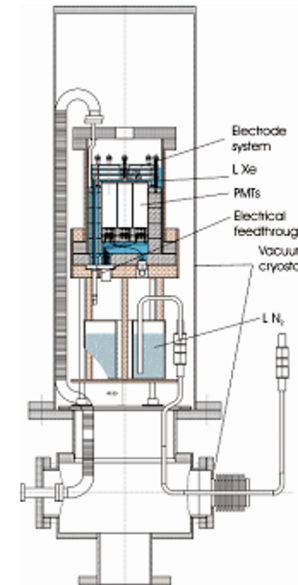


# Single Electron (SE) detection

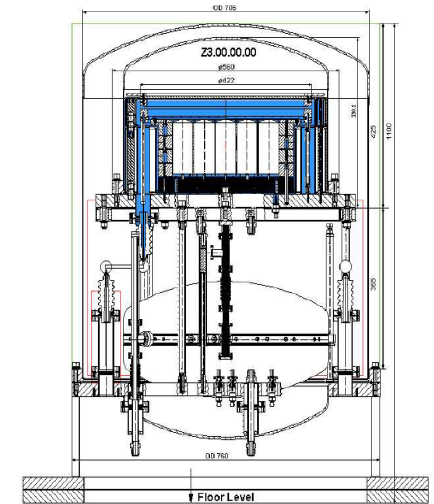
- Capability to detect single ionization electrons (SE) was demonstrated
- Projects for CEvNS with LXe two-phase detectors appeared



Proposals on CEvNS detection:



ITEP&INR LXe: JINST 4 (2009) P06010 [arXiv:0903.4821]



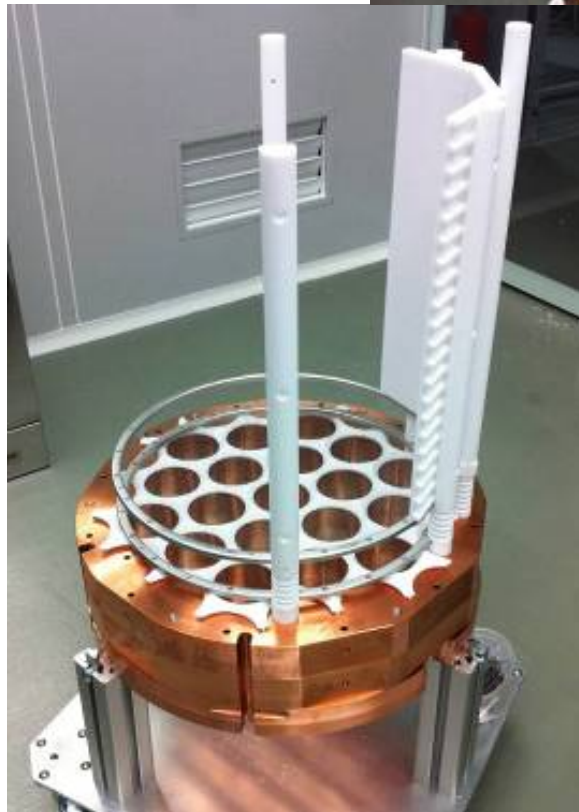
ZEPLIN-III Collaboration LXe: JHEP 1112 (2011) 115 [arXiv:1110.3056]



# RED-100 assembling

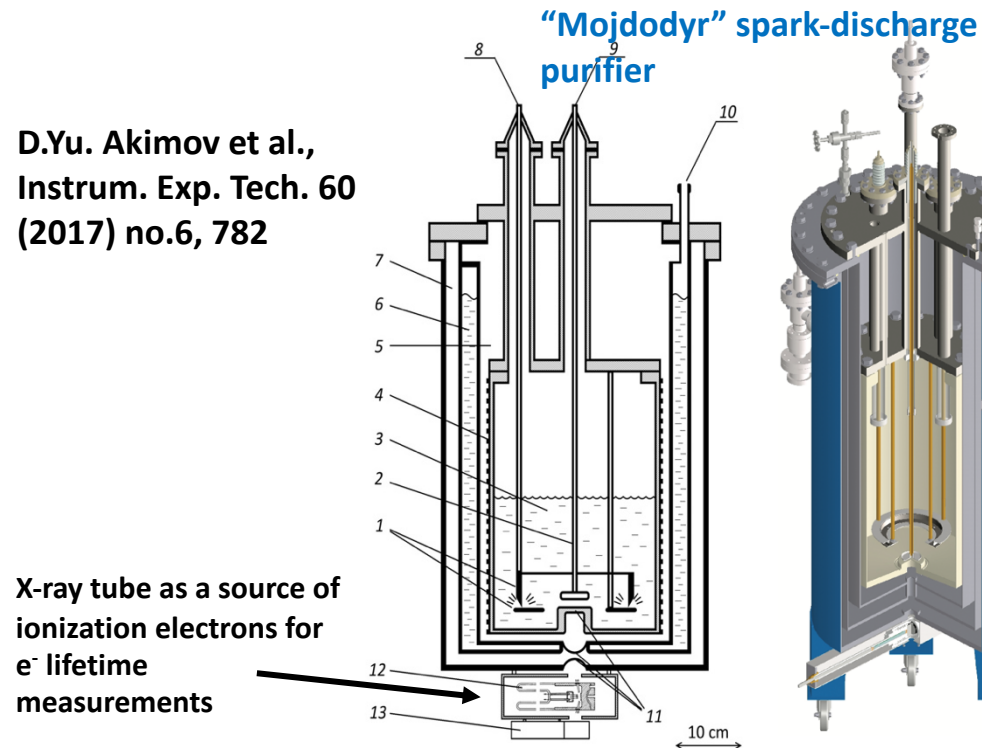
- RED-100 was assembled and tested in the MEPhI laboratory

Akimov D. Yu., et al. JINST 15.02 (2020): P02020.



# RED-100 performance: LXe purity

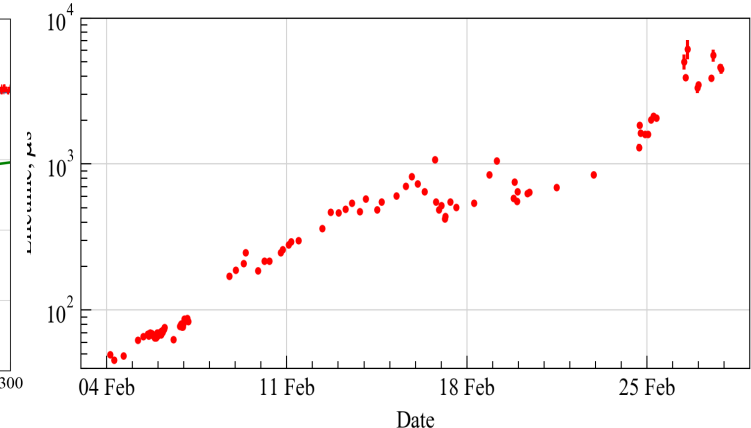
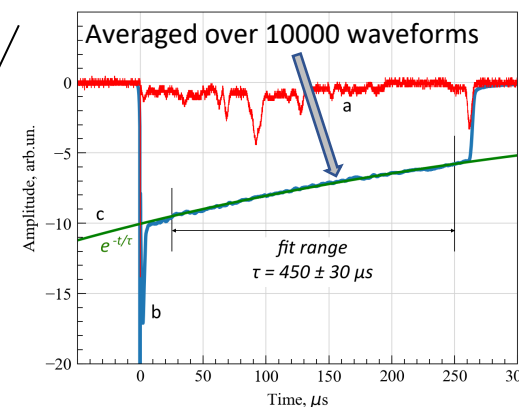
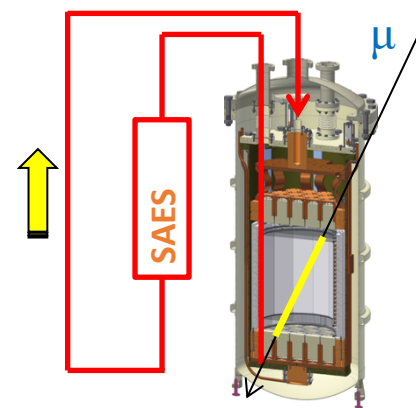
- Electronegative impurities catch the ionization electrons
- Purification in two stages
  - 1<sup>st</sup>: spark discharge technique with “Mojdodyr”
  - 2<sup>nd</sup>: continues circulation of Xe through RED-100 and SAES
- Electron lifetime of several milliseconds was achieved



Xenon was contaminated by highly-electronegative impurities presumably due to the use of a special fluorine-containing high-molecular-weight lubricant in gas centrifuges.

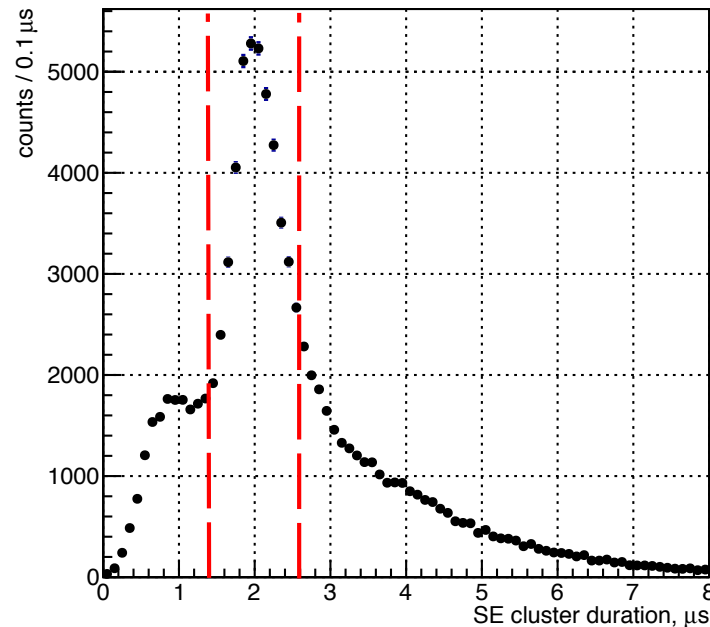
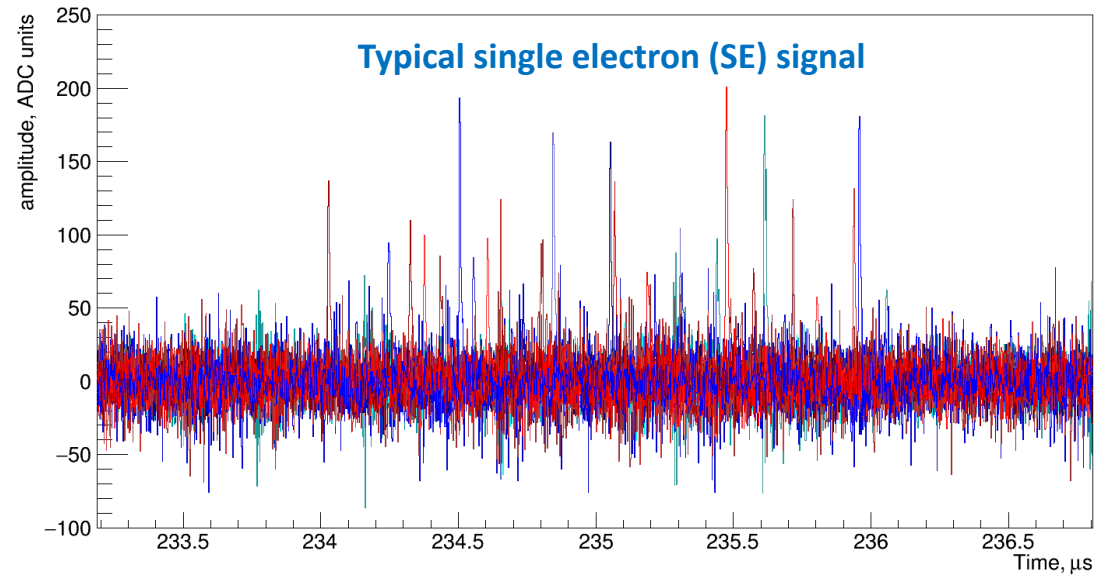
After purification, the achieved lifetime  $\geq 50 \mu\text{s}$  for  $\sim 200 \text{ kg}$  of LXe

Electron lifetime was measured by cosmic muons passed through the detector:

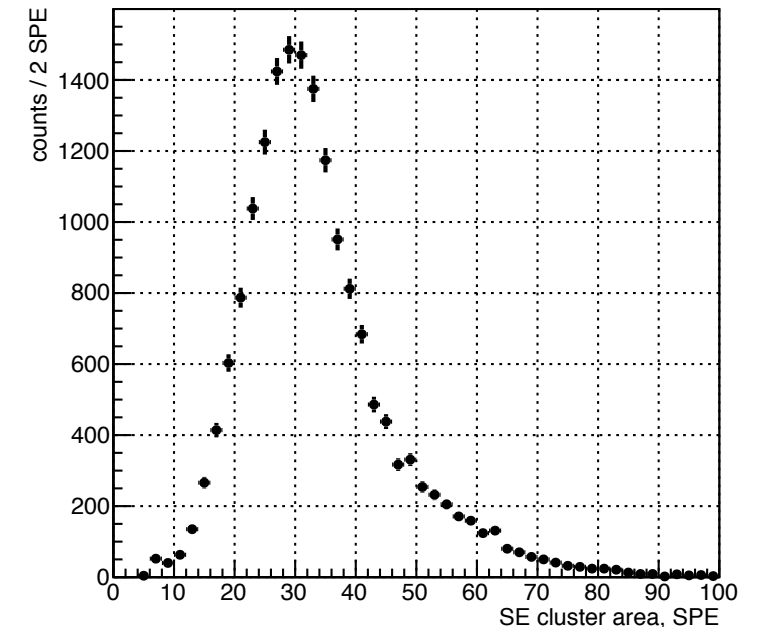


# RED-100 performance: SE

- SE is a cluster of individuals SPEs (single photo electrons)
- Typical duration  $\sim 2 \mu\text{s}$
- $\sim 30$  SPE/SE for RED-100



Distribution of SE duration



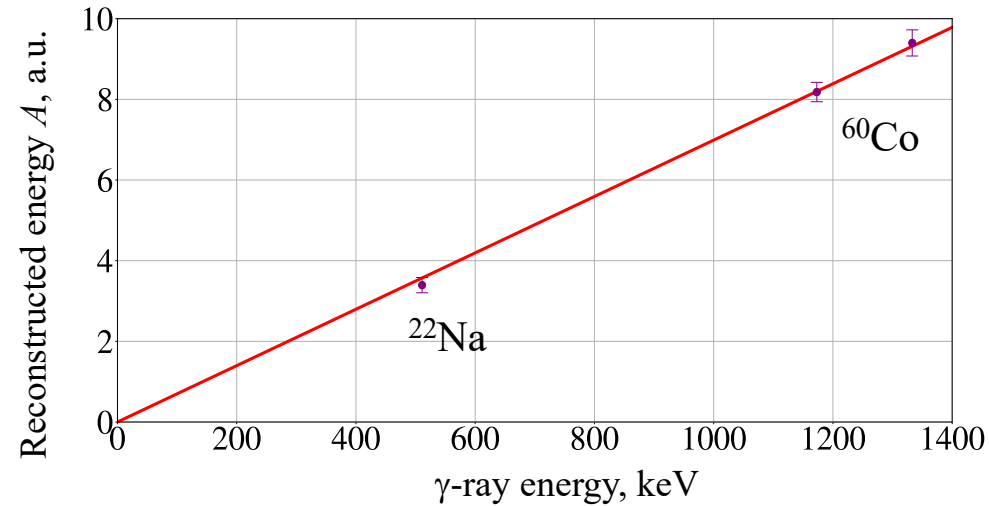
Distribution of SE area



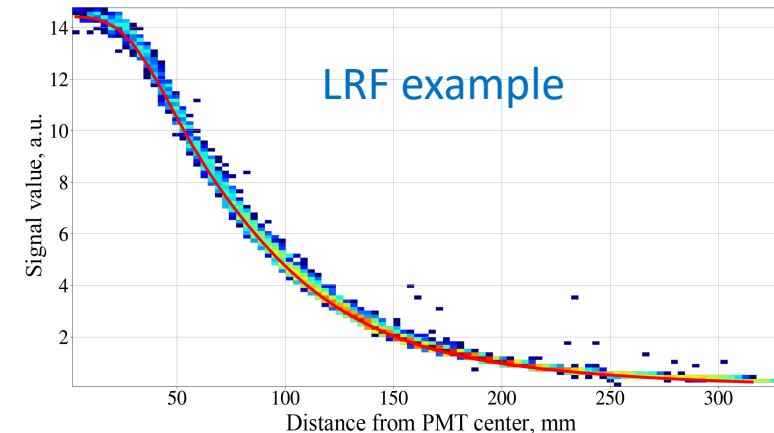
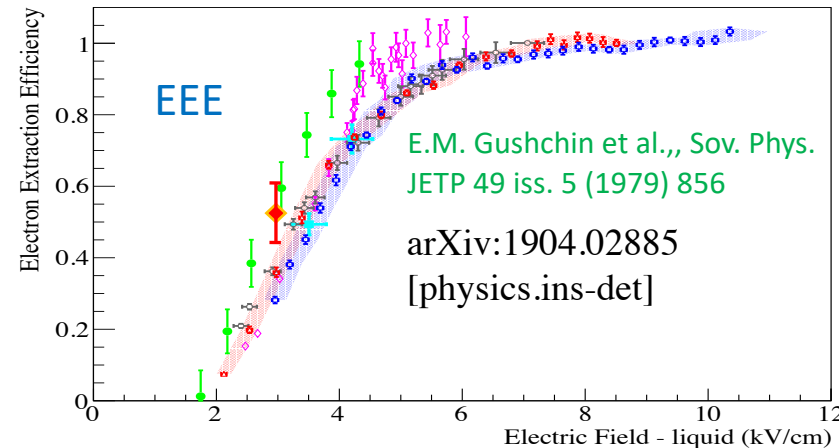
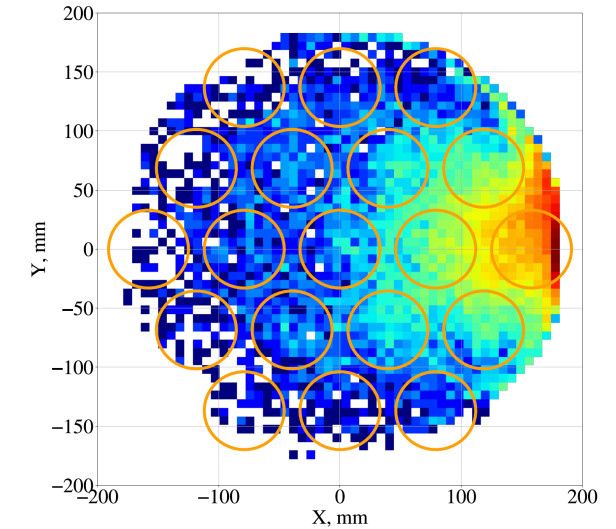
# RED-100 performance: gamma calibration

- Gamma calibration was done
- Position reconstruction tested
- LRF obtained for the top PMT plane
- Electron extraction efficiency (EEE)
  - S2-based only
  - $N_{SE} = {}^{22}\text{Na}$  peak position/SE area
  - $N_E$  – from NEST @  $E_{dr} = 0.217$  kV/cm
  - $N_E^*$  – corrected for electron lifetime
  - **EEE** =  $N_{SE}/N_E^* = 0.54 \pm 0.08$   
@  $E_{extr} = 3.0 \pm 0.1$  kV/cm

Gamma calibration



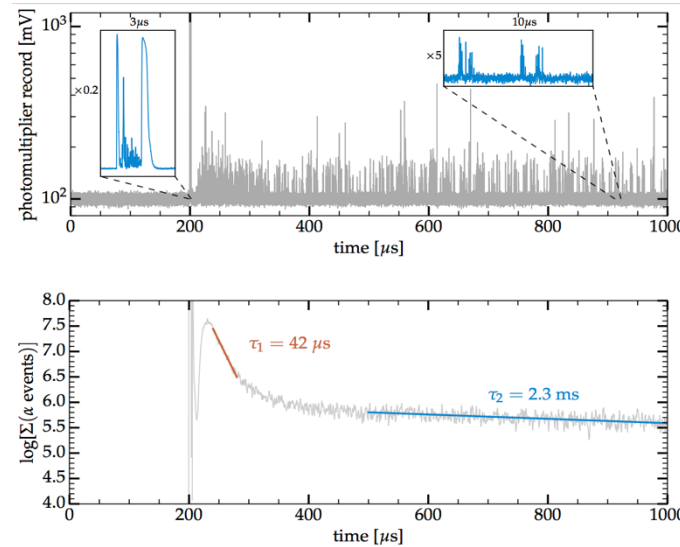
Position reconstruction for  ${}^{22}\text{Na}$



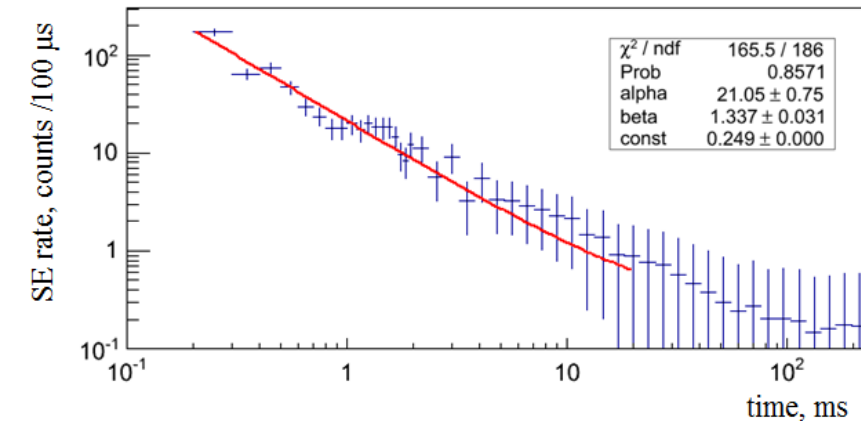
# RED-100 performance: “spontaneous” SE

- An increasing of SE rate after energy deposition in liquid noble gas detector was observed by several groups
- Two components:
  1. short, but more intense, caused by emission of the electrons trapped at LXe surface
  2. long, but less intense; unknown mechanism, decreases with time as purity increase; possibly, catching and releasing electrons by impurities (correlation with purity (of LAr) was also observed in DS50)
- Electron shutter in RED-100
  - To minimize 1<sup>st</sup> component
  - Muon is a trigger
  - SE rate was reduced by factor of about 3
  - Still high SE rate of the second component (250 kHz) in the lab
  - Expecting reduction at the site of KNPP in a factor of about 5

P. Sorensen, K. Kamdin  
JINST 13 (2018) no.02, P02032

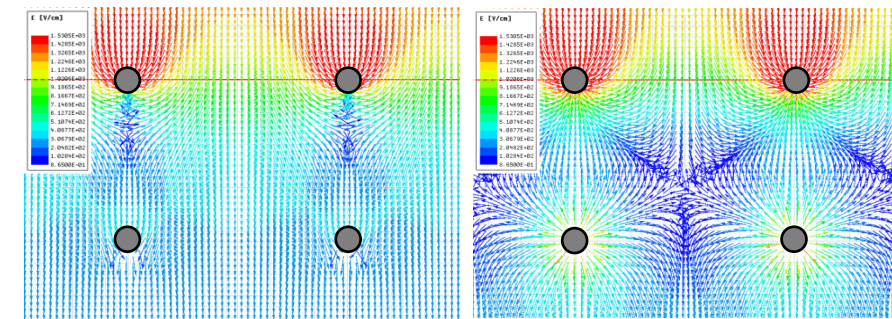
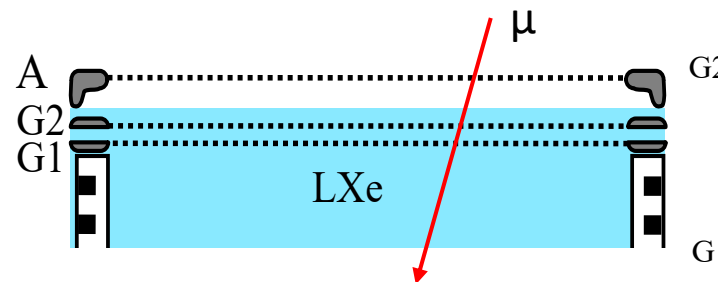


JINST 11 (2016) no.03, C03007



Observed in ZEPLIN-III: JHEP 1112 (2011) 115, [arXiv:1110.3056](https://arxiv.org/abs/1110.3056) [physics.ins-det]

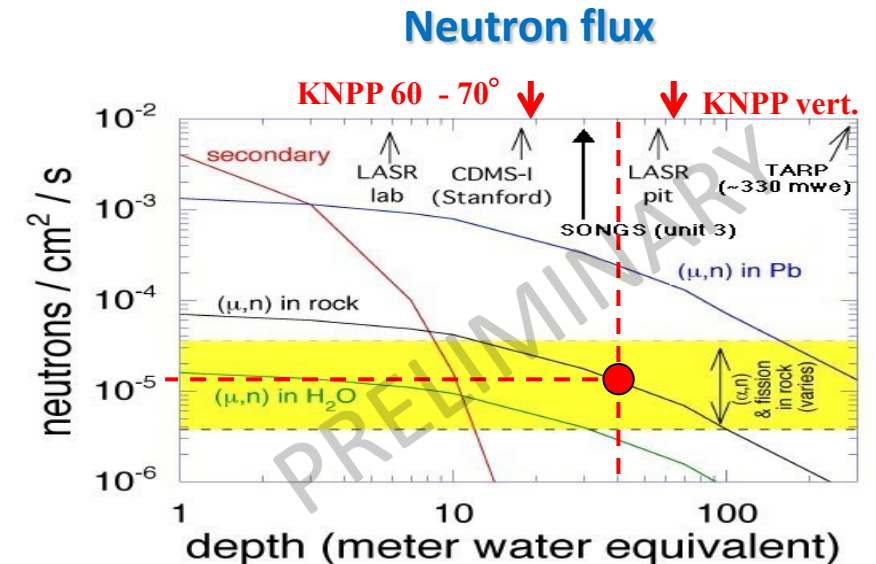
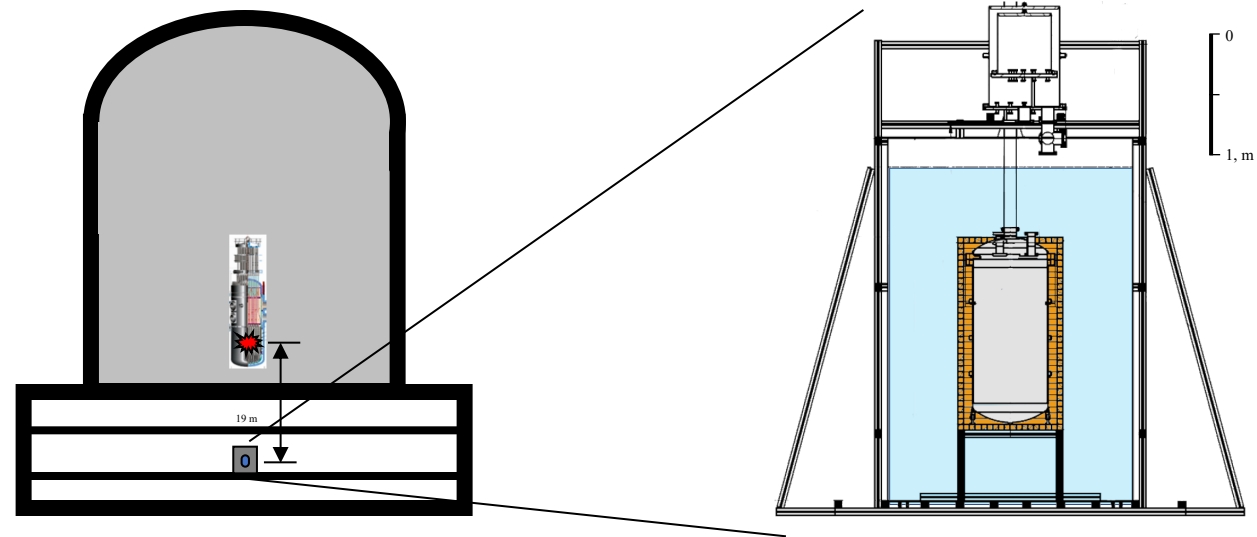
D.Yu. Akimov et al., Two-phase emission low-background detector (in Russian), Utility model patent RU 184222 U1, 2018



open:  $U_{G2} - U_{G1} = 50 \text{ V}$  closed:  $U_{G2} - U_{G1} = -250 \text{ V}$

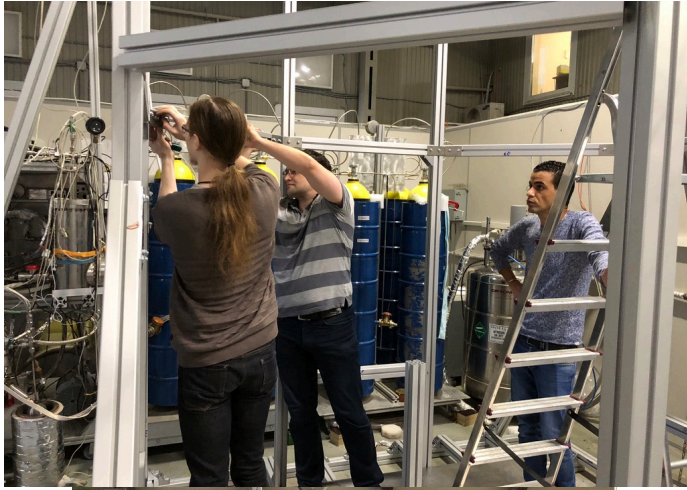
# RED-100 at KNPP

- KNPP – Kalinin Nuclear Power Plant
- 19 m from the reactor core
- Antineutrino flux  $\sim 1.35 \cdot 10^{13} \text{ cm}^{-2}\text{s}^{-1}$
- $\sim 65 \text{ m.w.e.}$  in vertical direction
- Passive shielding:
  - 5 cm Cu
  - $\sim 60 \text{ cm H}_2\text{O}$



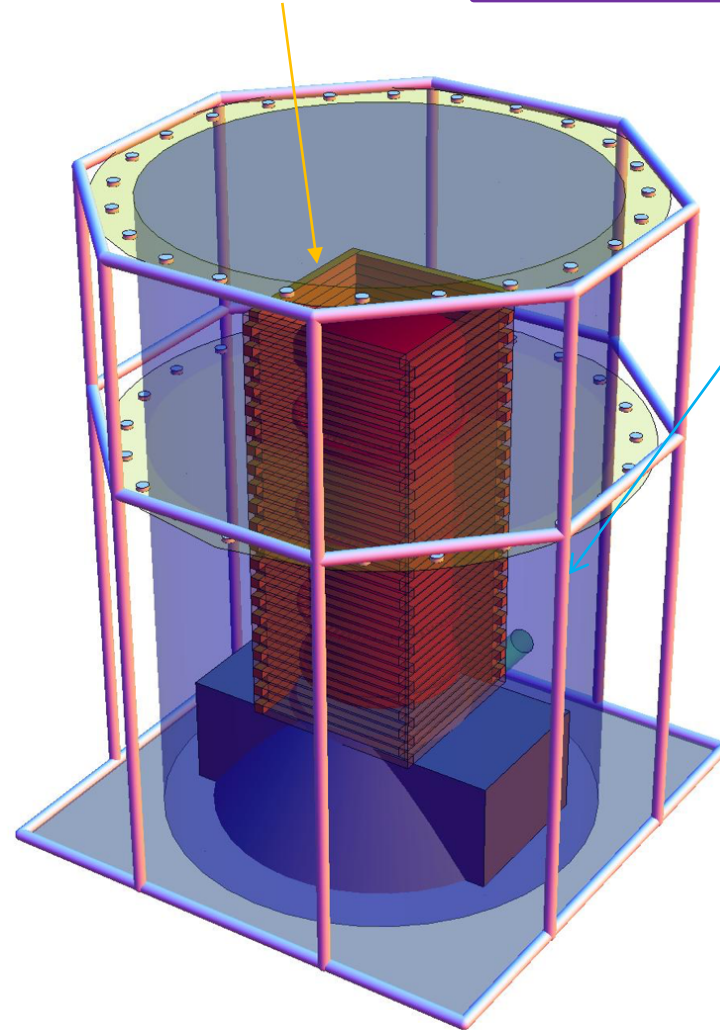


# Passive shielding

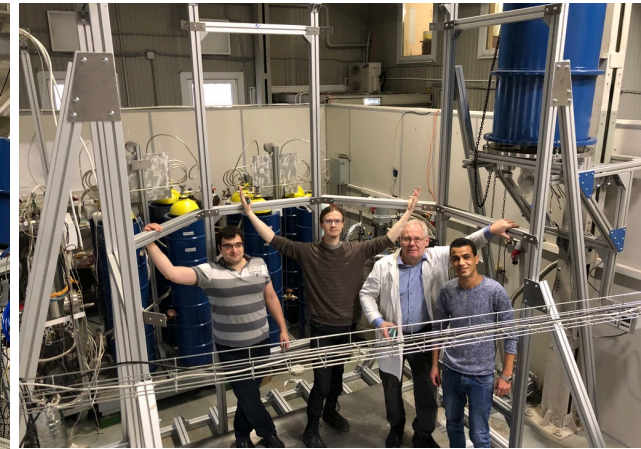
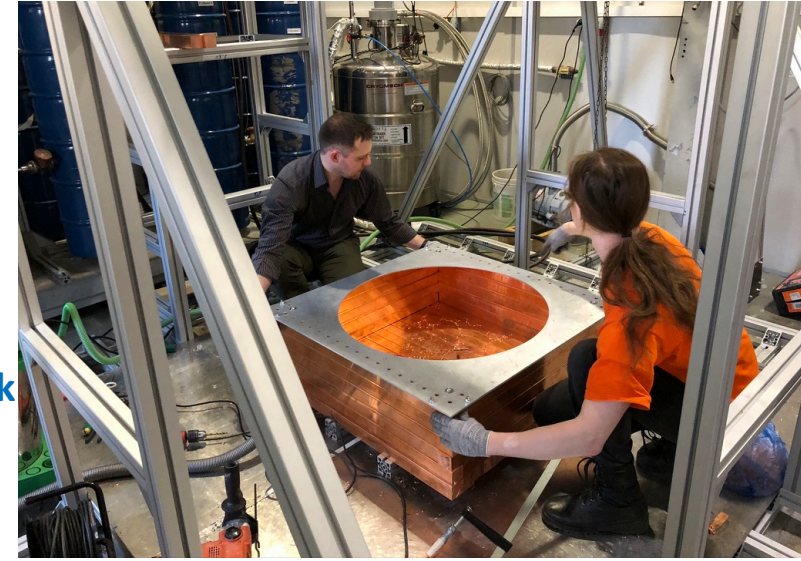


Copper shielding

Assembled and tested  
in MEPHl



Water tank

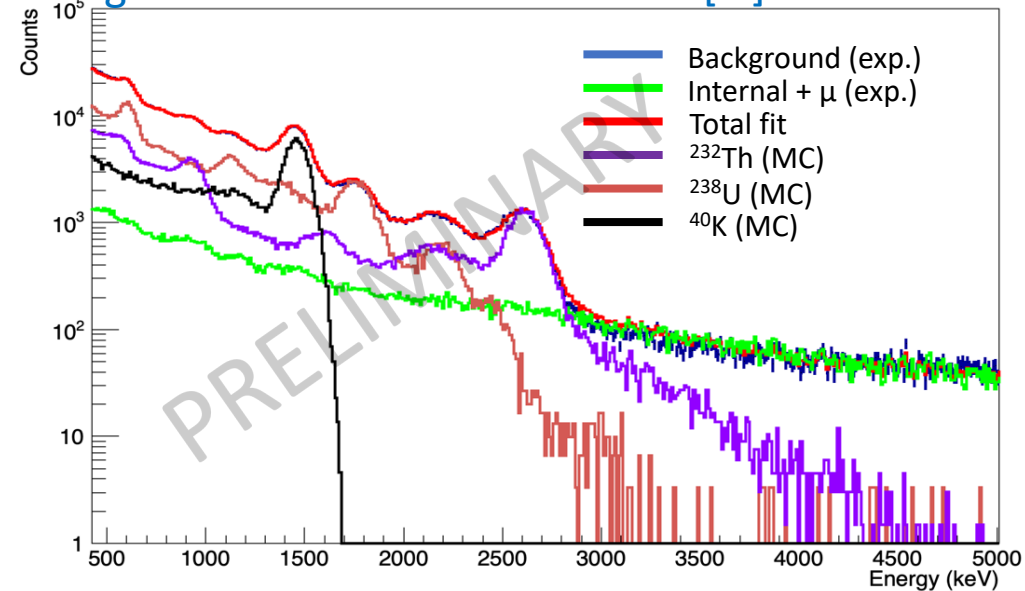




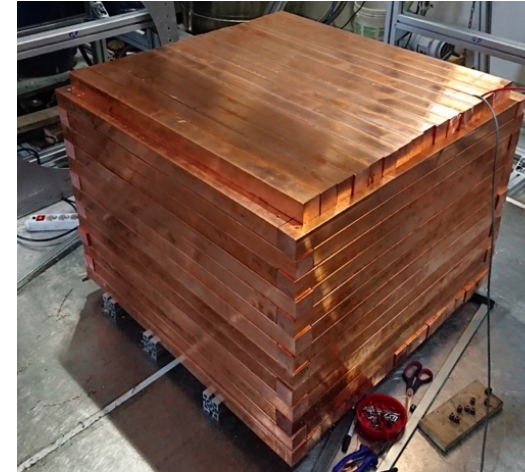
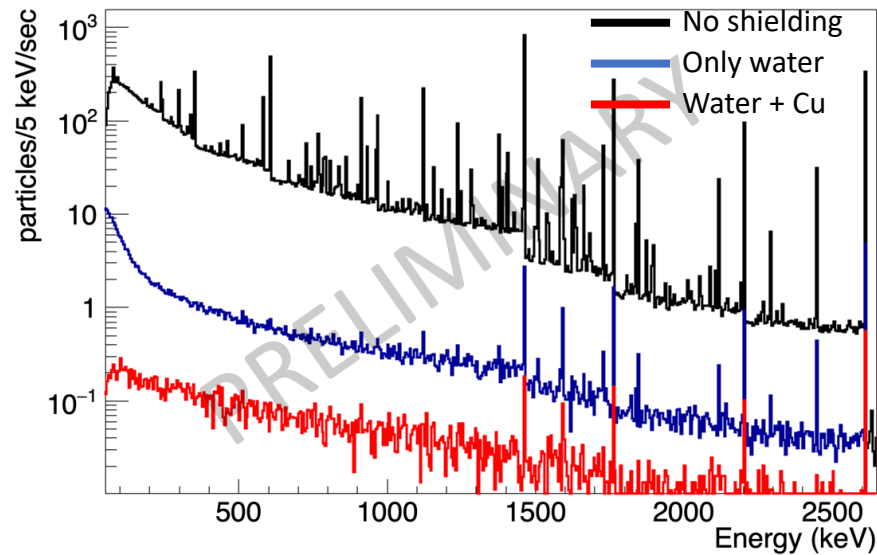
# Passive shielding

- Restrictions in space and  $\text{kg/m}^2$  at KNPP
- 5 cm Cu and  $\sim 60$  cm water
- Reduction of gamma background in a factor of  $\sim 700$  according to tests and MC
- MC for neutron background reduction is on going
- Paper is submitted

Background measurement with NaI[Tl] and fit with MC



MC background at the external cryostat surface



# Estimation of CEvNS count rate at KNPP

- Main background → accidental coincidence of several spontaneous SE (multi-electron event = ME)
- But:
  - CEvNS events are point-like events
  - Background is mostly NOT point-like
- Naive point-like cut was tested, and it works!
- Several ML approaches, CNN and DL networks were tested → further improvement (PRELIMINARY)

ME value in electrons	Estimated ME background at KNPP, events/160kg/day			Expected CEvNS count rate at KNPP, events/160kg/day		
	no cut	Naive point-like	DL point-like	no cut	Naive point-like	DL point-like
2	$5.3 \cdot 10^7$	$1.8 \cdot 10^7$	$0.8 \cdot 10^7$	465	283	317
3	$4.4 \cdot 10^5$	$0.9 \cdot 10^5$	$1.7 \cdot 10^4$	129	79	100
4	$2.7 \cdot 10^3$	348	34.2	35.5	21.7	29.3
5	13.7	1.1	0.05	10.6	6.4	9.1
6	$5.7 \cdot 10^{-2}$	$3.0 \cdot 10^{-3}$	$6.3 \cdot 10^{-5}$	1.9	1.2	1.7

**We can detect CEvNS with threshold of  $\geq 4$  SE**



Taken into account:

- Recent data on ionization yield in LXe for NR
- $EEE = NSE / N * E = 0.54 \pm 0.08$
- Factor of 5 reduction of muon rate  $\Rightarrow$  50 kHz spontaneous SE rate
- Poisson flow of spontaneous SE



# Timeline



- Despite COVID-19 situation we are in time!
- 2020
  - October: water tank was tested
  - November: RED-100 is moved to KNPP!
- 2021
  - Winter: deployment; background measurements and analysis
  - Spring: start of data taking
- 2022
  - Data analysis
- Prolongation of experiment (?)

# Conclusion

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- RED-100 was assembled and tested @ MEPHI
  - Excellent electron lifetime of several milliseconds
  - Electron extraction efficiency =  $0.54 \pm 0.08$  @  $3.0 \pm 0.1$  kV/cm
  - SE gain is about 30 SPE/SE
  - The electron shutter was tested: the spontaneous SE rate suppressed but still high
- The results of the first lab test are available: Akimov D. Yu., et al. JINST 15.02 (2020): P02020
- Estimations based on our tests show the possibility to detect CEvNS at KNPP with a threshold  $\geq 4$  SE (progress is ongoing)
- RED-100 already at KNPP!

Thank you for your attention!

# Backup

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# CEvNS around the World

Gaseous spherical proportional counters



Composite of Zn- and Ge-based bolometric detectors



Germanium detectors

CaWO<sub>4</sub> and Al<sub>2</sub>O<sub>3</sub> bolometric detectors

(LAr @ Rx)



LAr detectors



(CCM)

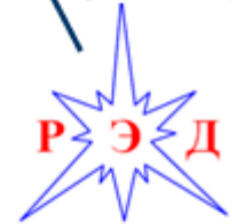


Super-CDMS-style Ge and Si detectors

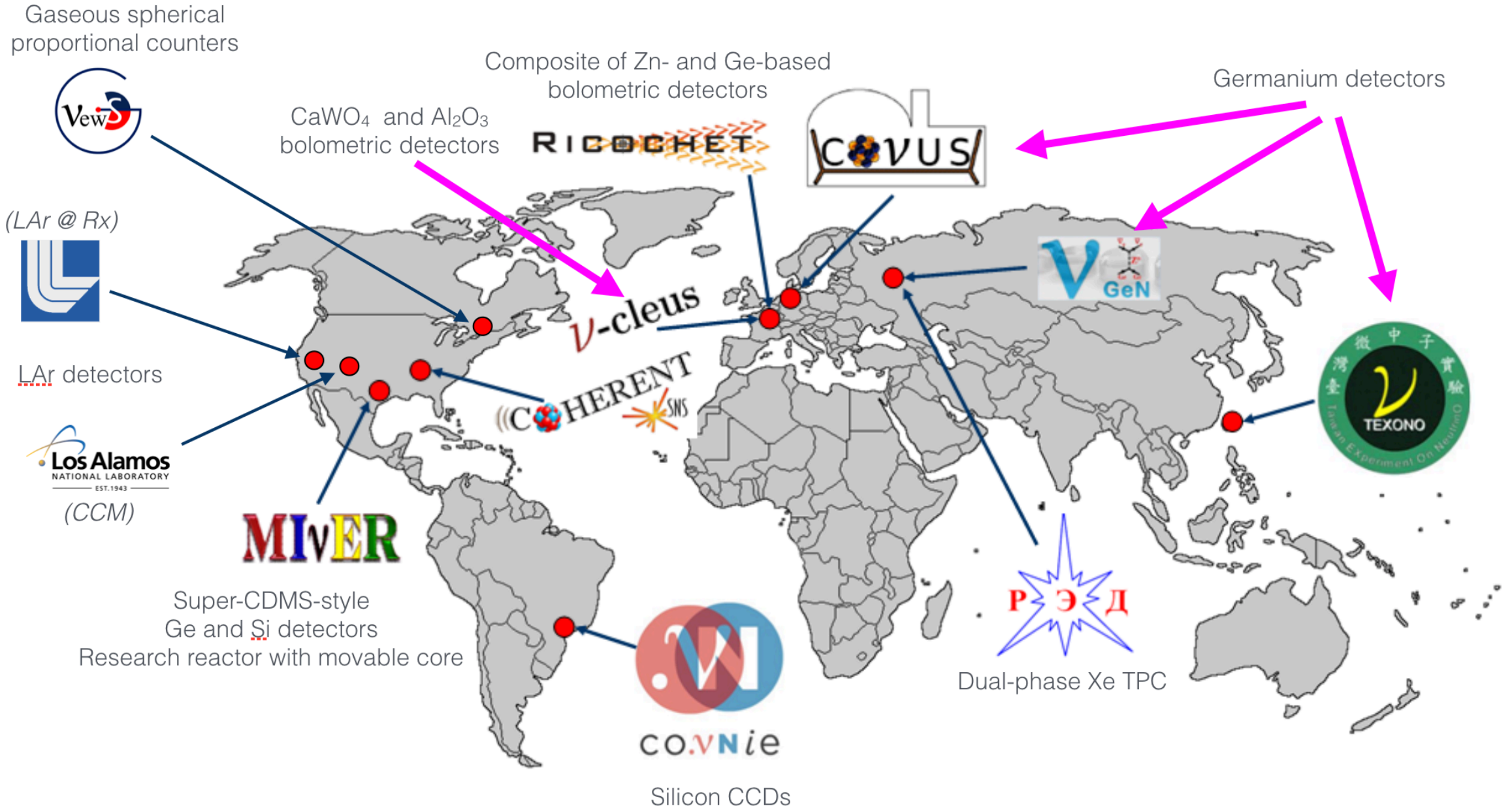
Research reactor with movable core



Silicon CCDs

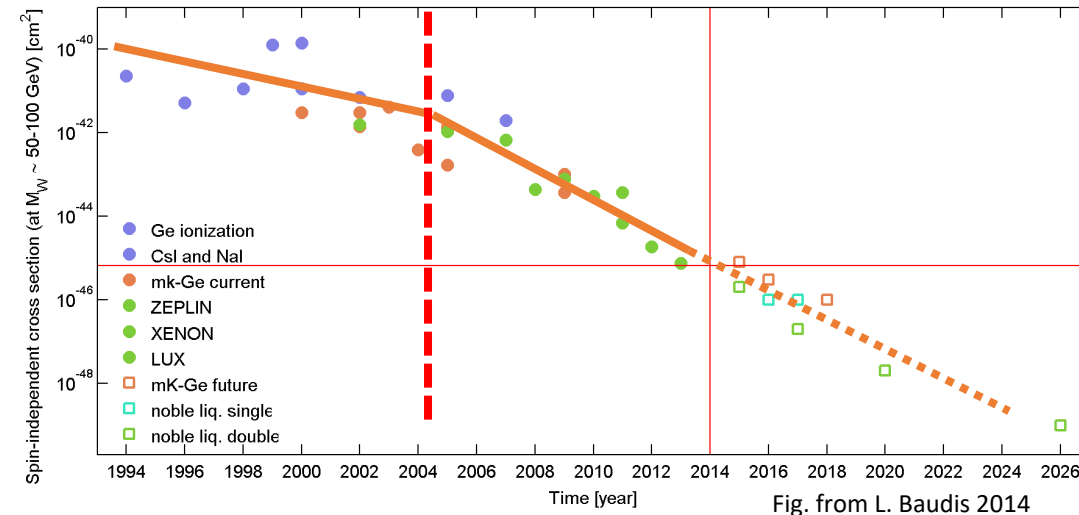


Dual-phase Xe TPC

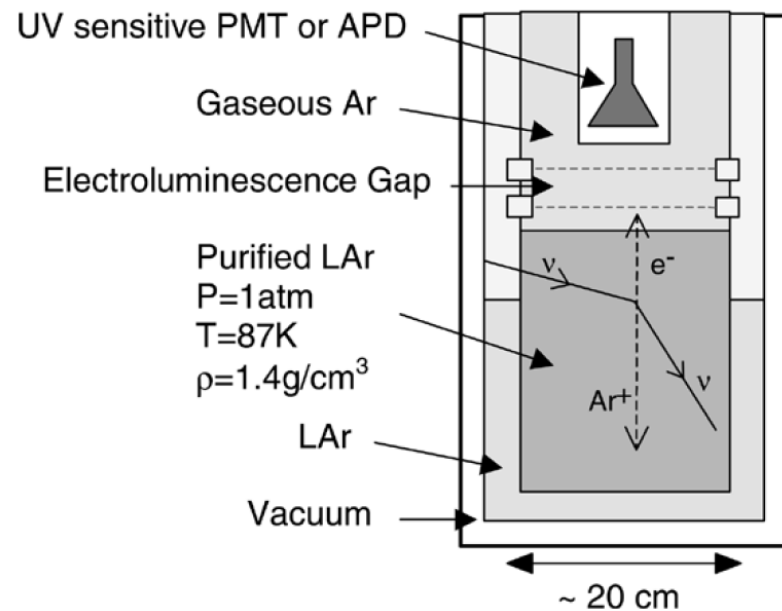


# Noble gas detectors and CEvNS

In Dark Matter search experiments, the progress of setting limits has increased significantly when liquid noble gas detectors (two-phase) started operation



1<sup>st</sup> proposal (in 2004); LAr detector



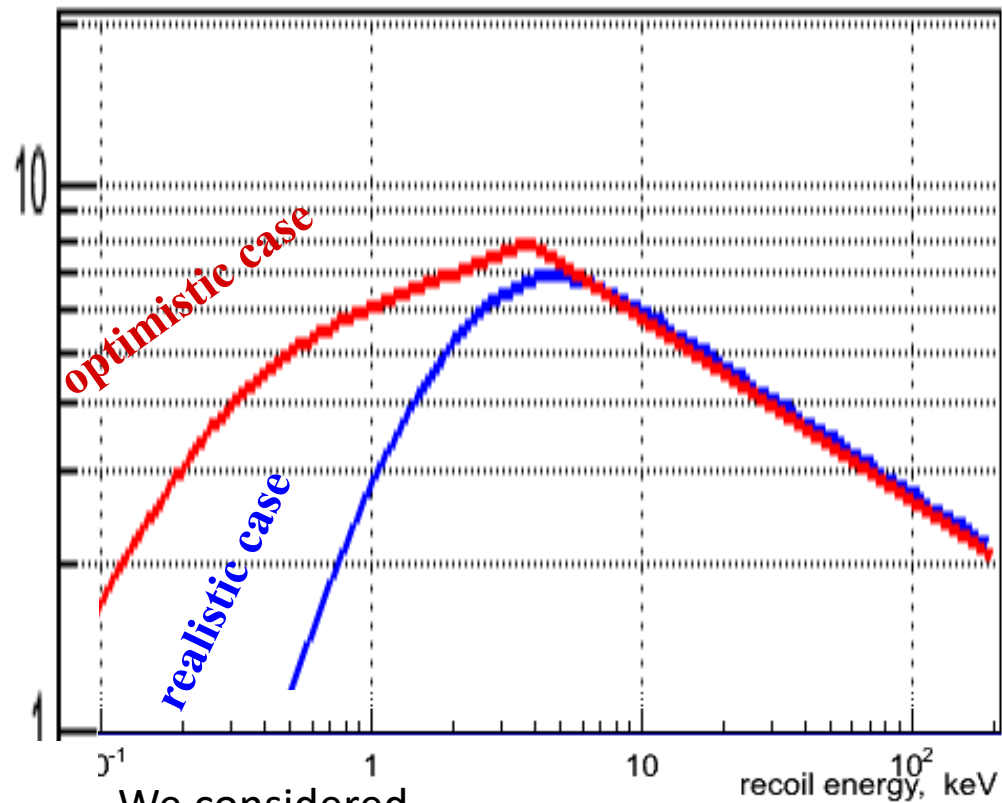
C. Hagmann and A. Bernstein,  
**Two-Phase Emission Detector for Measuring  
 Coherent Neutrino-Nucleus Scattering**  
 IEEE Trans.Nucl.Sci. 51 (2004) 2151

From Akimov D. talk @ INSTR20

# Ionization yield for sub-keV nuclear recoils

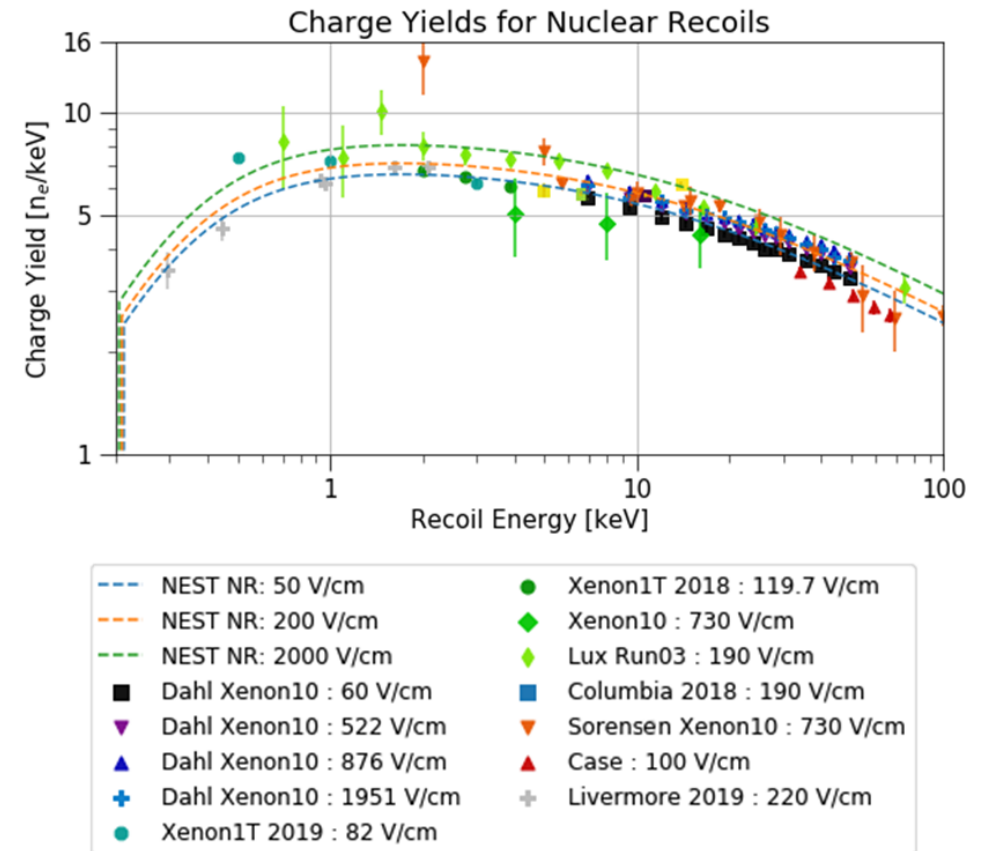
7 years ago

There were no data  $< 4 \text{ keV}_{nr}$



We considered  
“optimistic” and “realistic” scenarios

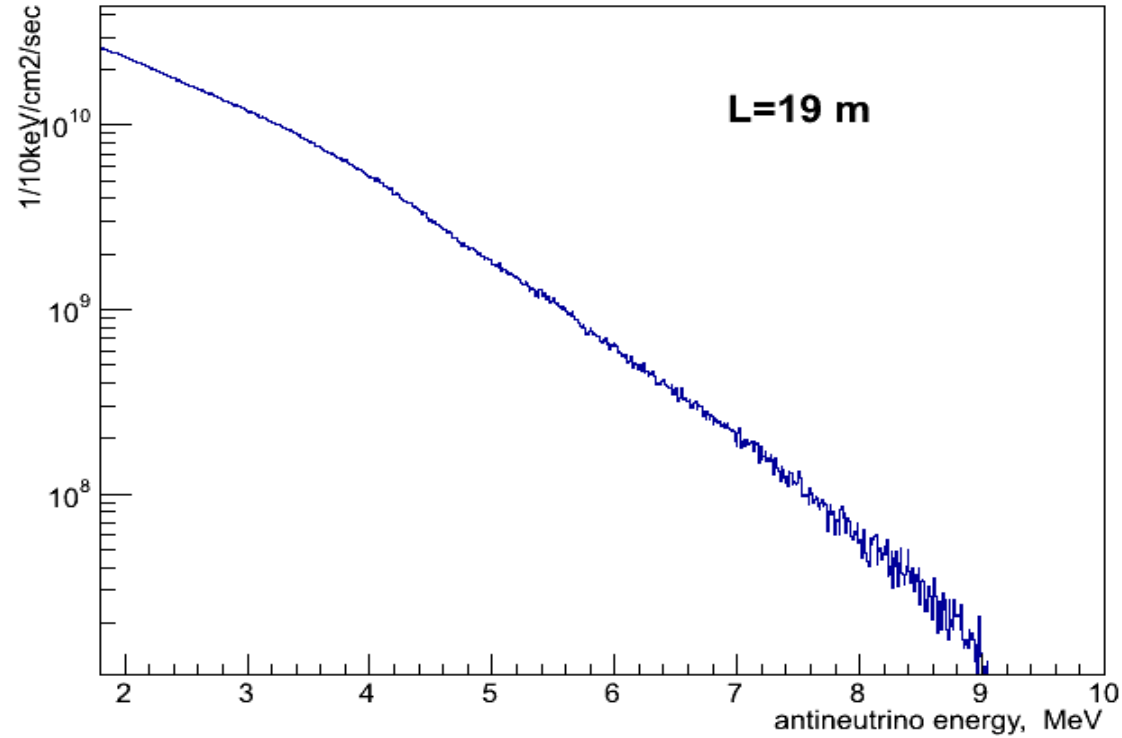
Now



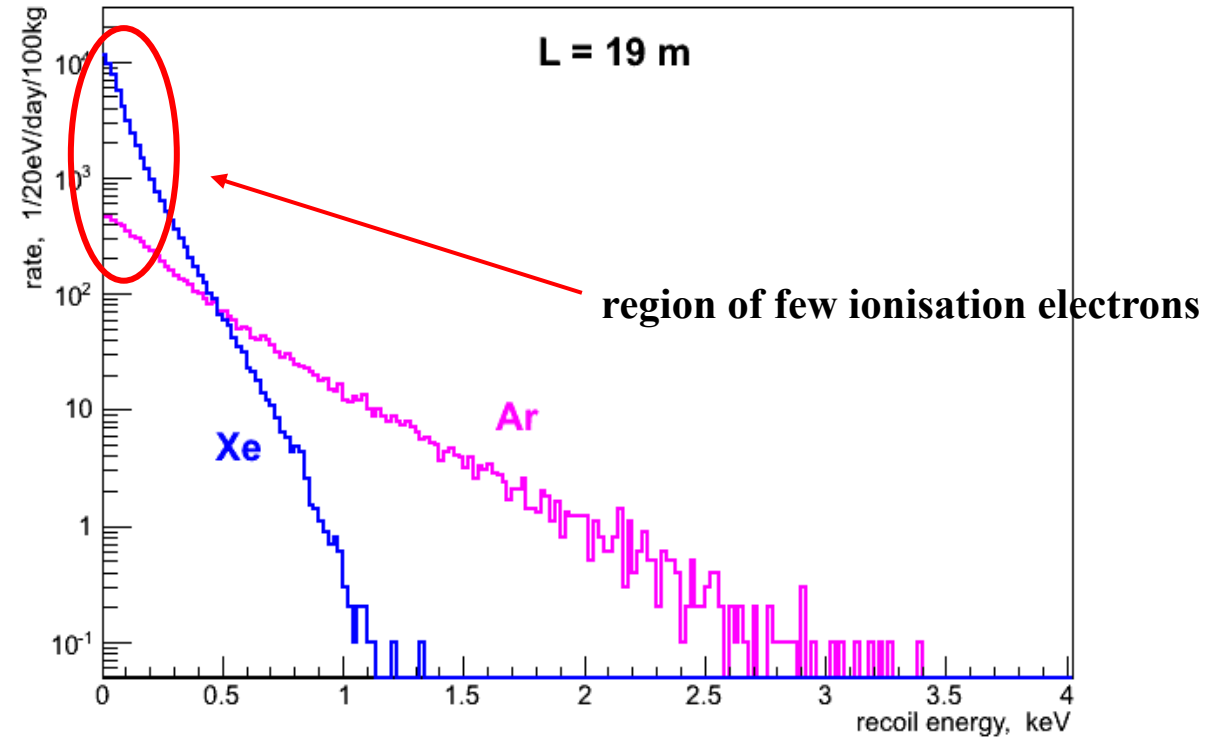


# Energy spectrum

$\bar{\nu}_e$  energy spectrum from nuclear reactor



Xe and Ar nuclear recoil spectra



**This is very challenging task, but feasible!**

# Additional improvements to improve CEvNS/bckg

**1 To increase EEE by increasing extraction (G2-A) electric field  $\Rightarrow$  CEvNS signal  $\hat{\uparrow}$ , however SE rate  $\hat{\uparrow}$ , but not significantly**

*For this purpose, additional Teflon isolator is installed between G2 and A*

**2 To introduce smart blocking for the muon events: the higher muon deposited energy, the longer blocking time of the shutter (up to several hundred ms)**

**3 To study the influence of LXe purity on the rate of spontaneous SE events**

**4 To improve algorithm of point-like events selection**



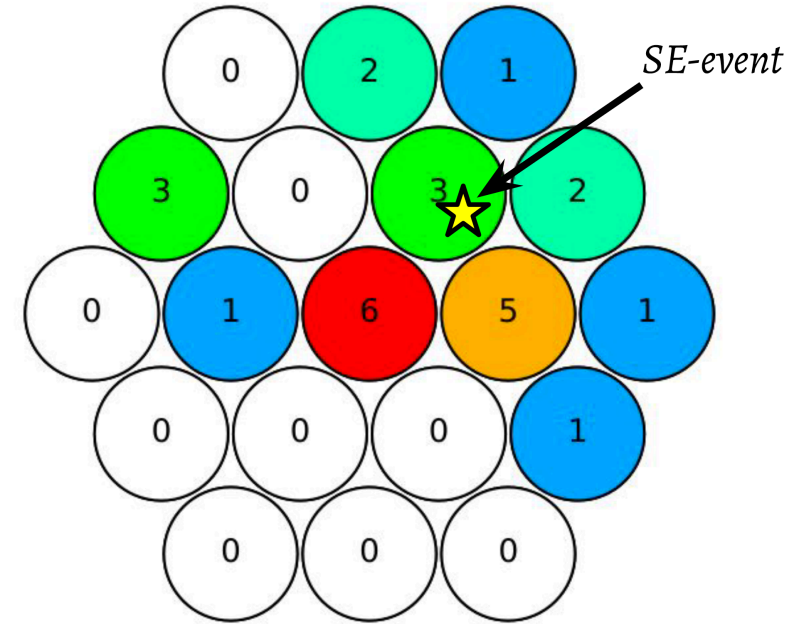
# Simulating of optical photons distribution

The optical simulation of RED-100 is done via ANTS-2 software.  
Number of photons in electroluminescence from single electron was estimated with NEST code (v.2.0.1)

The simulated event is a set of numbers of photons detected by each PMT in top array.  
Simulated events: 1SE, 2SE, ..., 6SE

**Background events:** 1+1, 1+1+1, 2+1, 2+2, 1+3... (sum $\leq$ 6)  
in different points uniformly scattered on the XY plane.

**CEvNS events:** 2,3,4,5 or 6 SE in one point



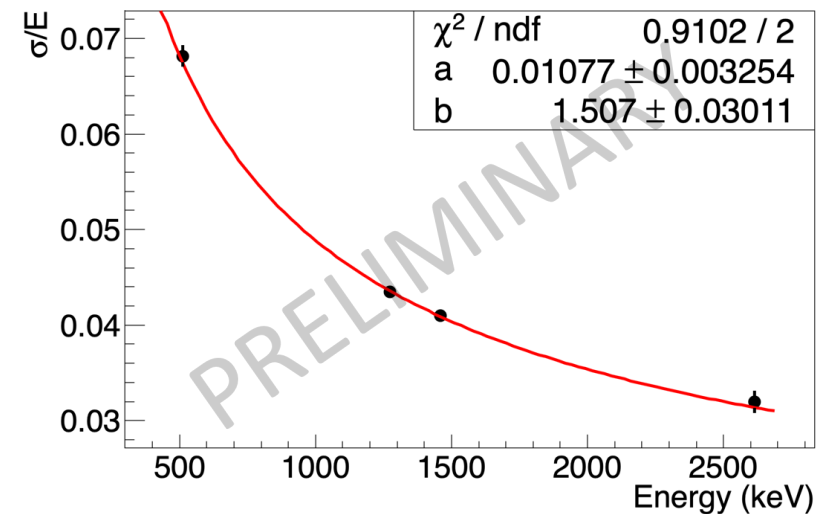
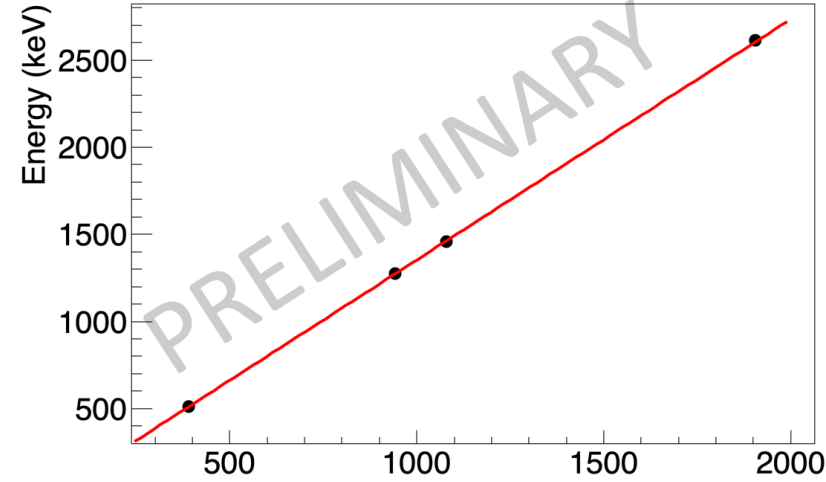
Example of simulated event (1 SE)  
Numbers in circles are the detected photons by PMTs in top plane.

*A. Morozov, et al. "ANTS2 package: simulation and experimental data processing for Anger camera type detectors." Journal of Instrumentation 11.04 (2016): P04022.*



# Background measurements with NaI(Tl)

- Calibrated NaI(Tl) detector was used for the background study
- Thick Cu+Pb shielding was used to get internal + muons background
  - 15 cm of Pb from the bottom
  - At least 15 cm Cu from each side
  - 5 cm Pb belt
- Publication is under preparation



# Gamma background @ KNPP measured by DANSS colab.

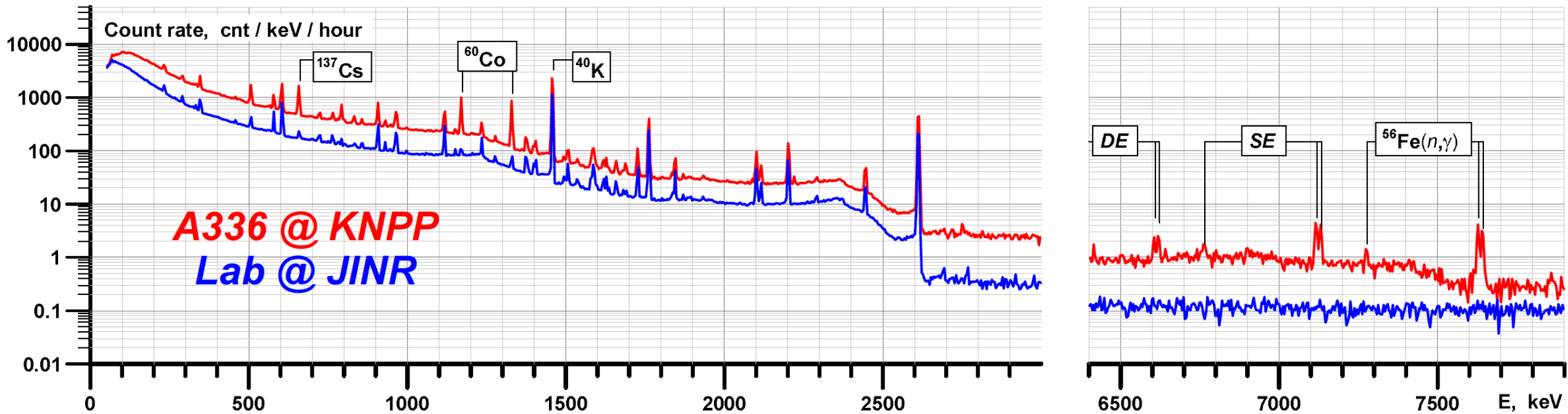


Figure 2: Gamma-background measured with 1.3 kg HPGe detector at the DANSS position in the KNPP room A336 and in the JINR laboratory.

arXiv:1606.02896v3