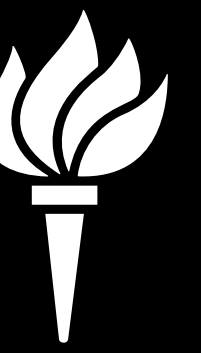




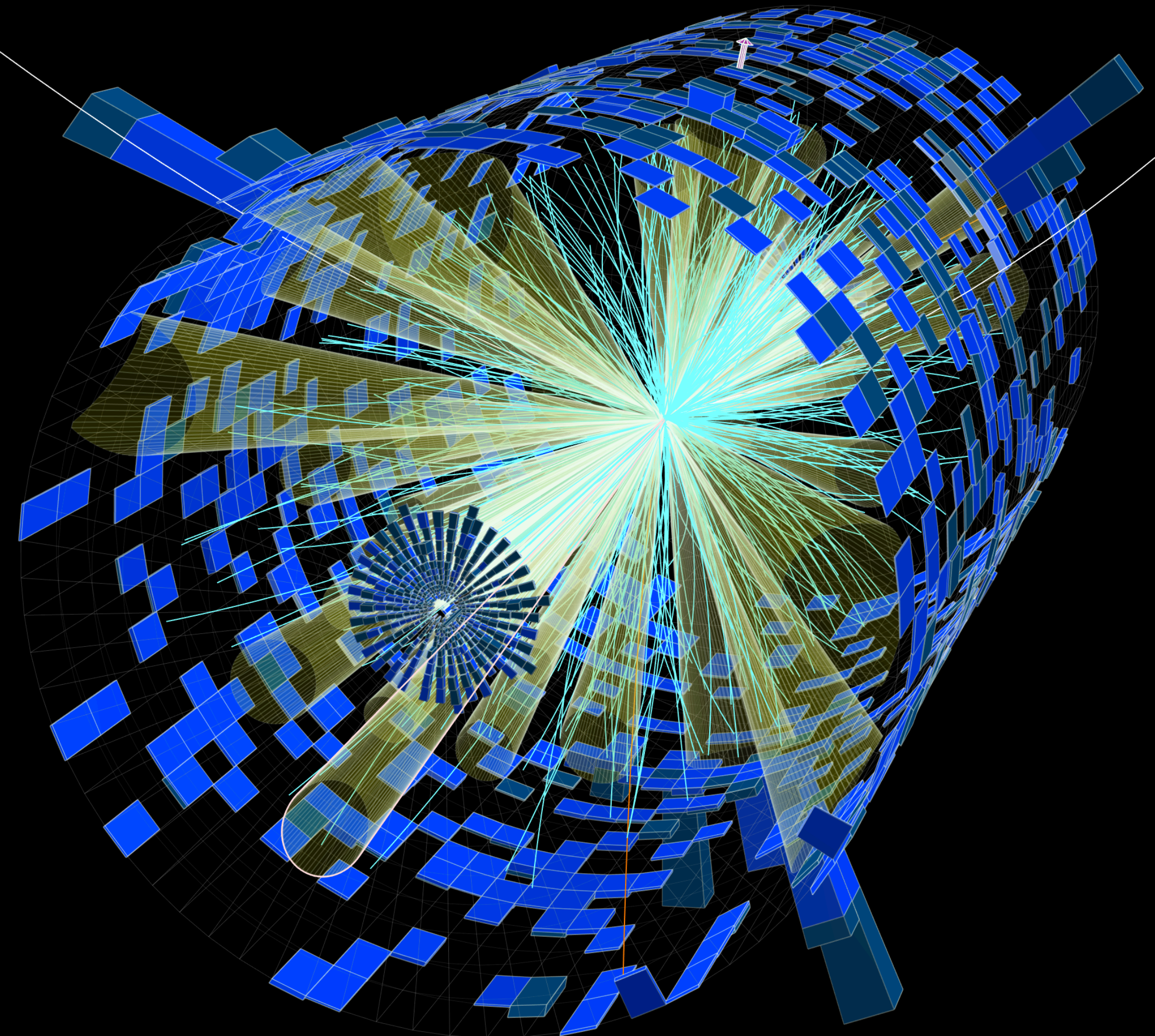
NYU CENTER FOR  
DATA SCIENCE

CENTER FOR  
COSMOLOGY AND  
PARTICLE PHYSICS



# CONSTRAINING EFTs & DM

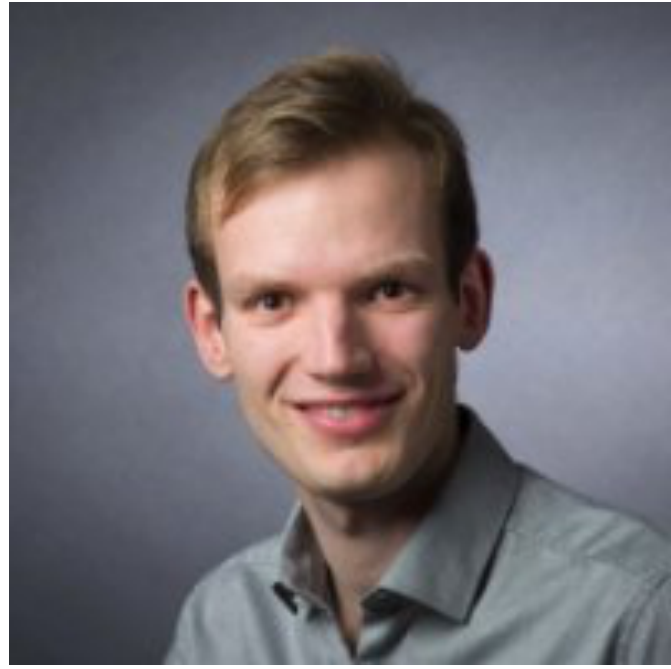
WITH SIMULATION-BASED INFERENCE



@KyleCranmer  
New York University  
Department of Physics  
Center for Data Science  
CILVR Lab



# Acknowledgements



Johann Brehmer



Gilles Louppe



Juan Pavez



Markus Stoye



Felix Kling



Irina Espejo



Sinclert Perez



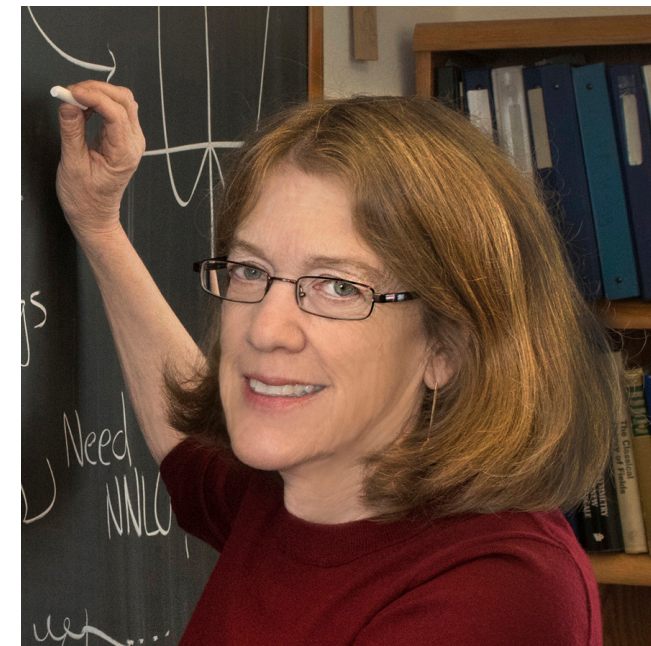
Sid Mishra-Sharma



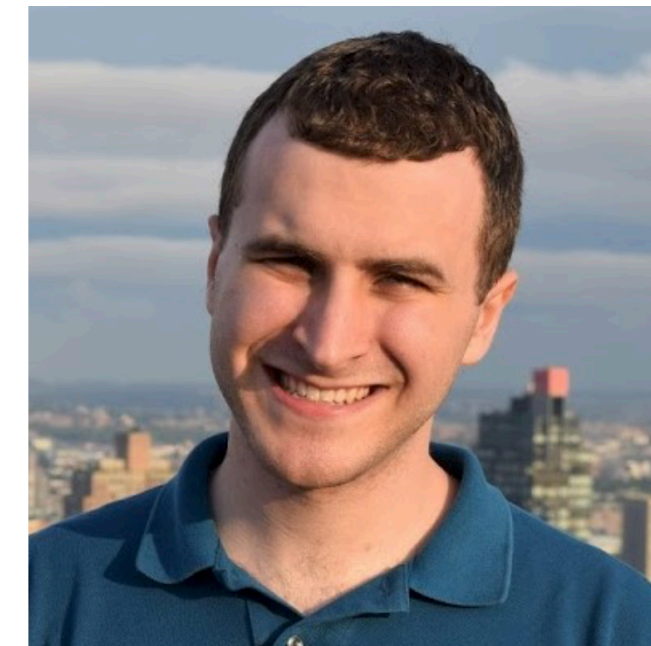
Joeri Hermans



Tilman Plehn



Sally Dawson



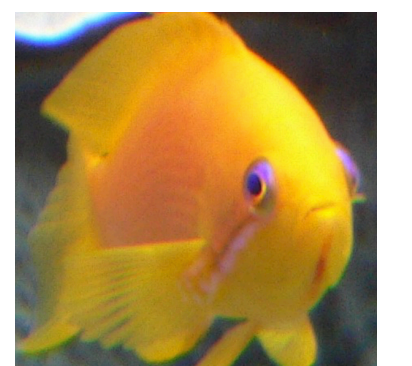
Sam Homiller



Zubair Bhatti



The SCAIFIN Project  
[scailfin.github.io](https://scailfin.github.io)





“I don't want your pity, I want change”

- Letetra Widman (Jacob Blake's sister)



**BLACK LIVES MATTER**



Many of these slides are borrowed from this talk by Johann Brehmer  
(who in turn has adapted many of my slides... so it's really hybrid 😊 )



# Constraining effective field theories with machine learning

Johann Brehmer  
New York University

ACAT 2019, Saas-Fee



# Motivation

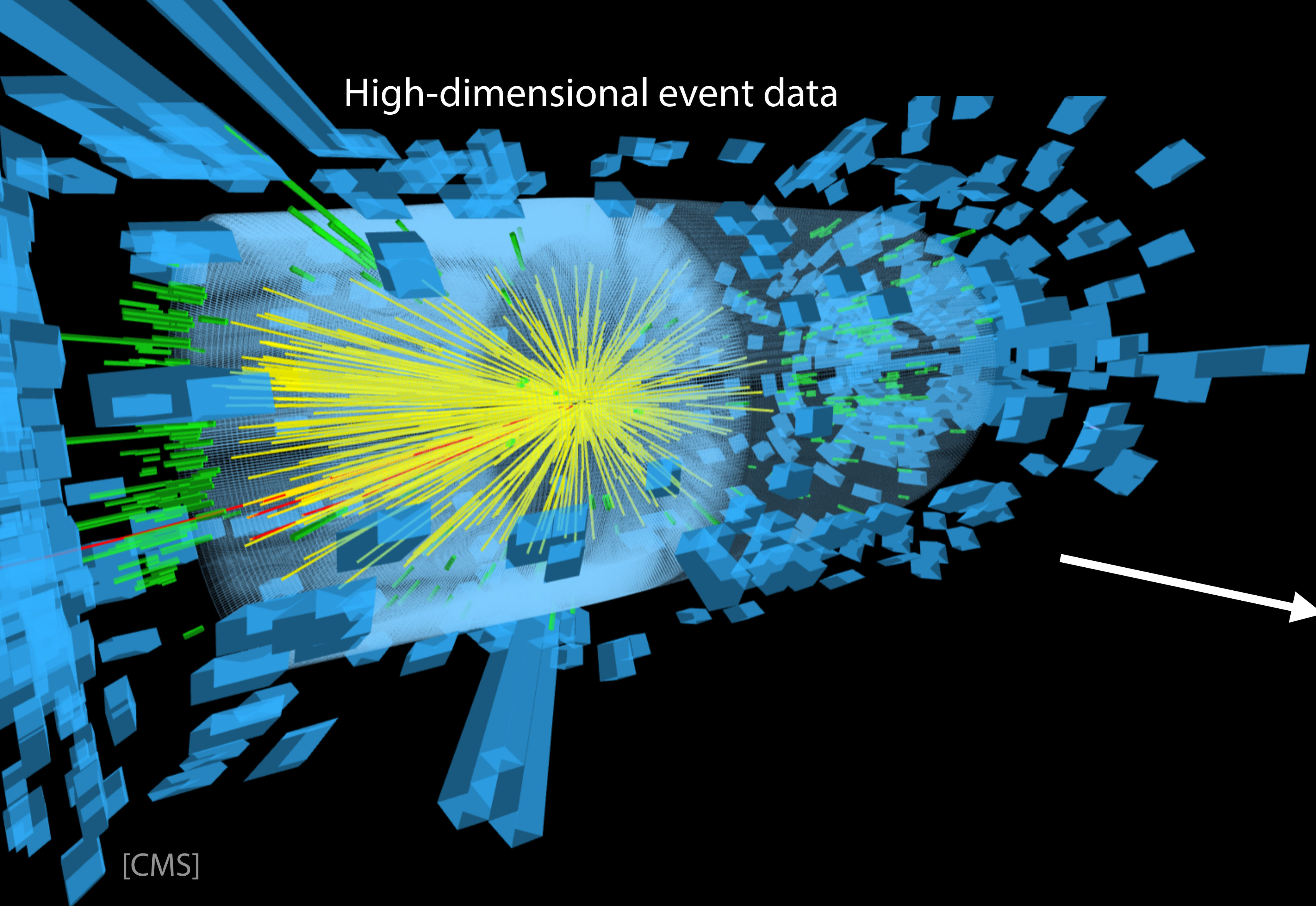
Obviously we would like to directly produce new particles at the LHC, and a robust search program is underway

But we can also study new physics indirectly by probing the subtle effects it would have on Standard Model processes.

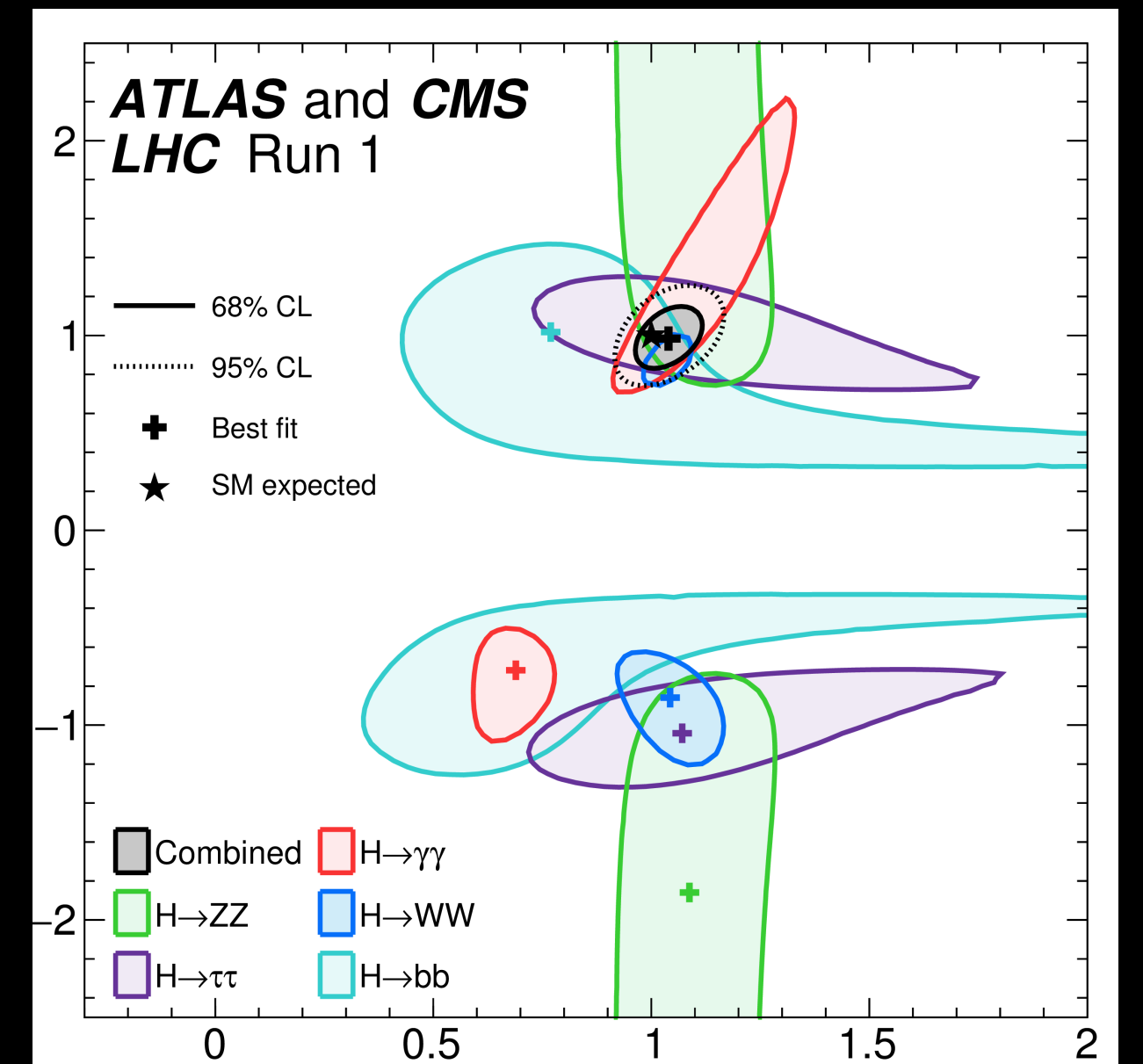
- These may well be the legacy measurements of the LHC
- Effective Field Theory is a systematic way of parametrizing those effects
- Effect size may be small, so we want to get the most out of the LHC data



High-dimensional event data



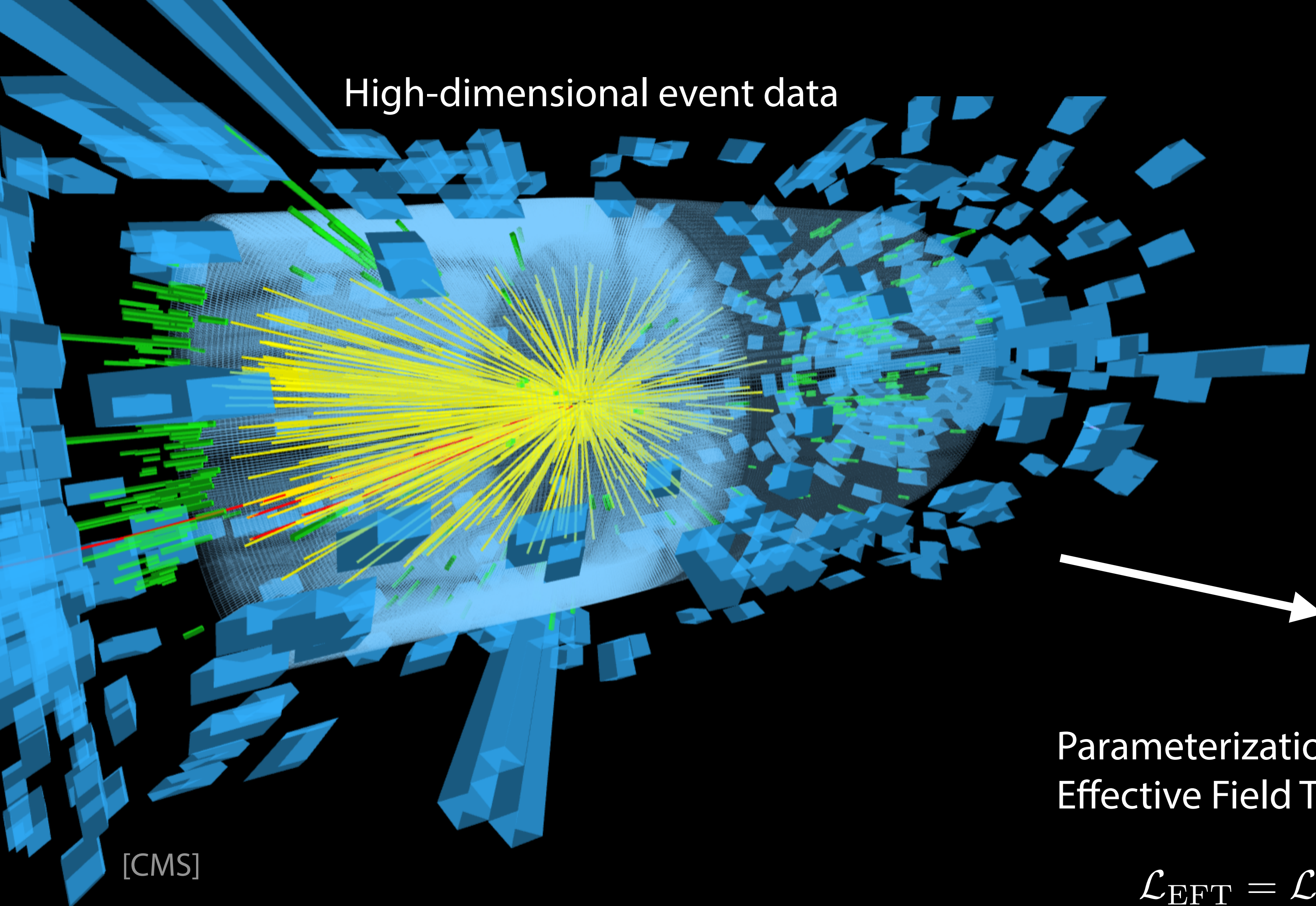
[CMS]



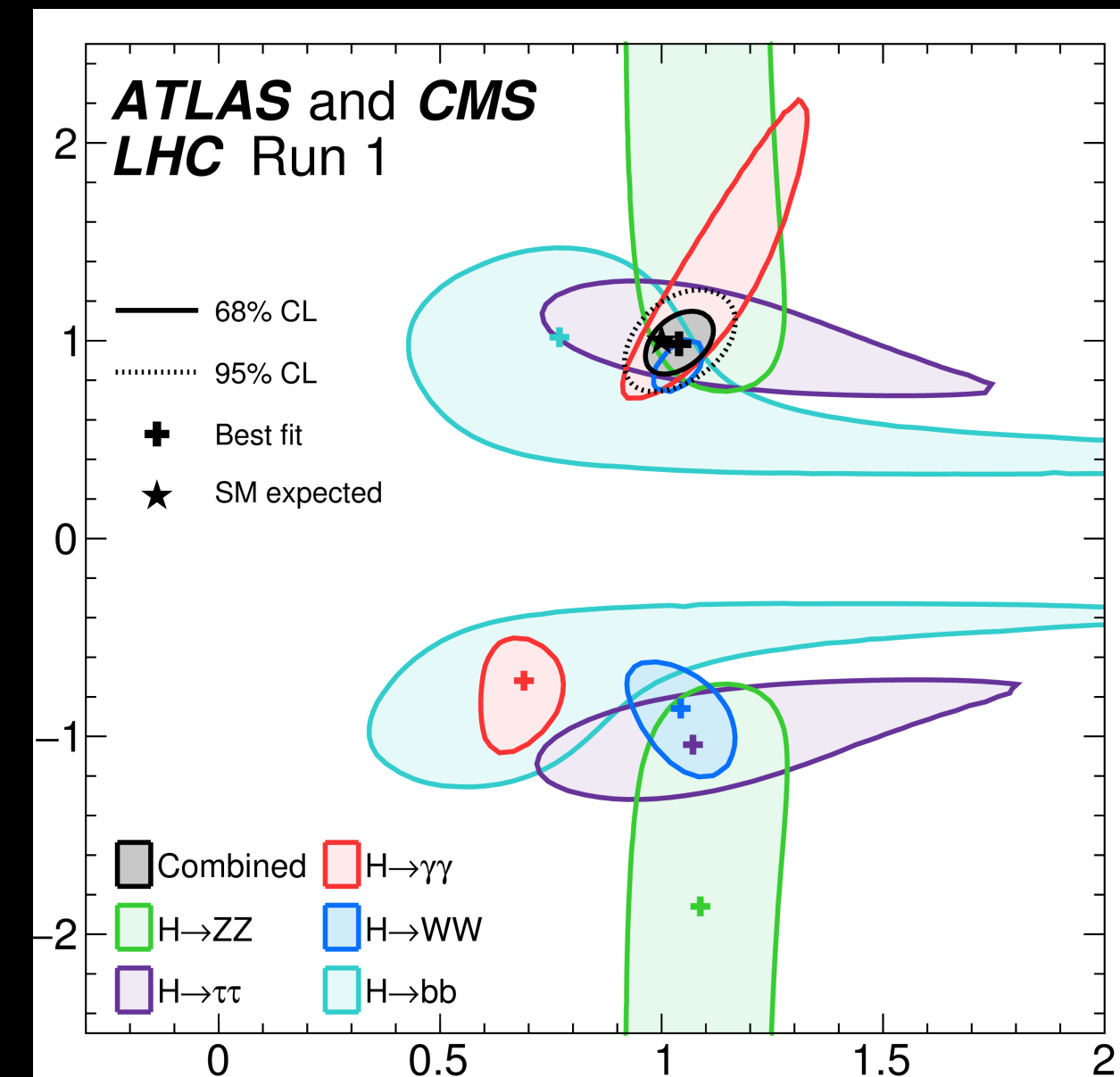
Precision constraints on  
new physics



High-dimensional event data



[CMS]



[ATLAS, CMS 1606.02266]

Precision constraints on new physics

Parameterization e.g. in Effective Field Theory:

systematic expansion of new physics around Standard Model

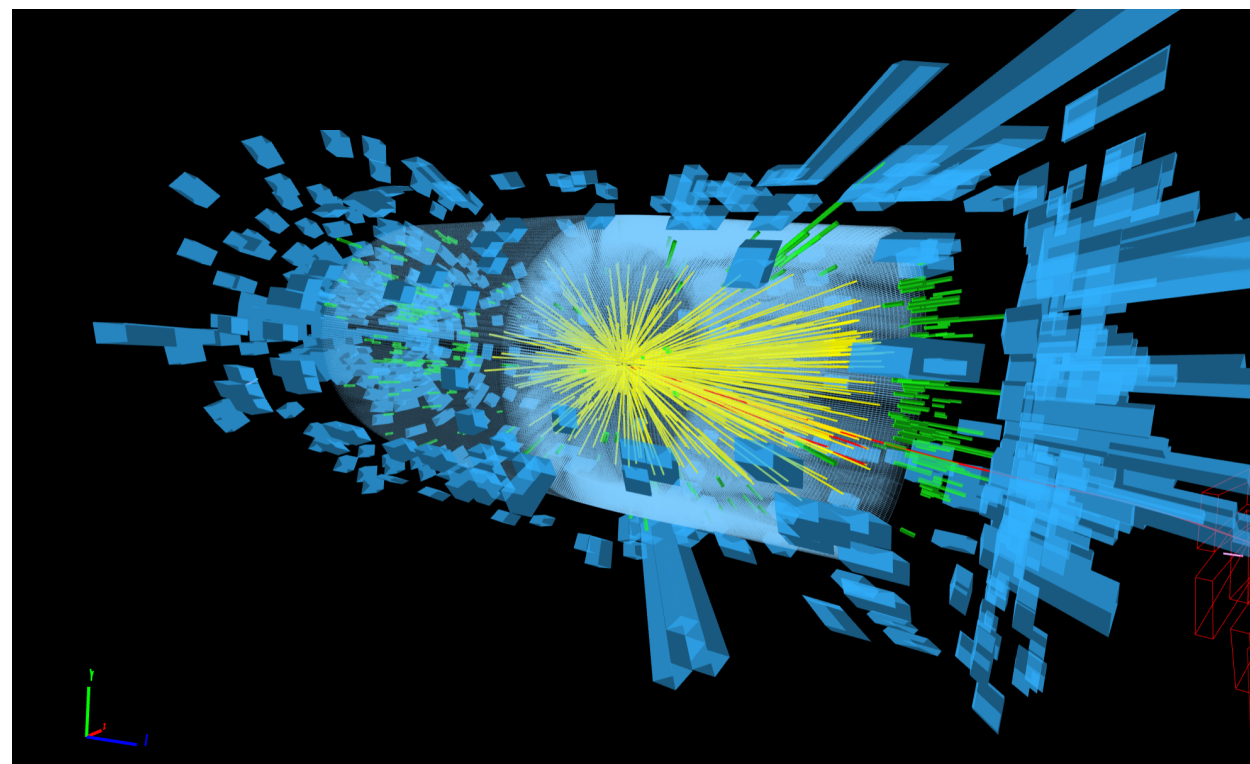
$$\mathcal{L}_{\text{EFT}} = \mathcal{L}_{\text{SM}} + \sum_i \frac{f_i}{\Lambda^2} \mathcal{O}_i + \dots$$

10s to 100s "universal" parameters to measure



# The likelihood is a key object

Let  $\theta$  denote the coefficients of higher dimensional operators in the Lagrangian,  $x$  be high-dimensional data associated to an event, and  $p(x | \theta) = \frac{1}{\sigma(\theta)} \frac{d\sigma}{d\theta}$  be the distribution for the data



High-dimensional event data  $x$

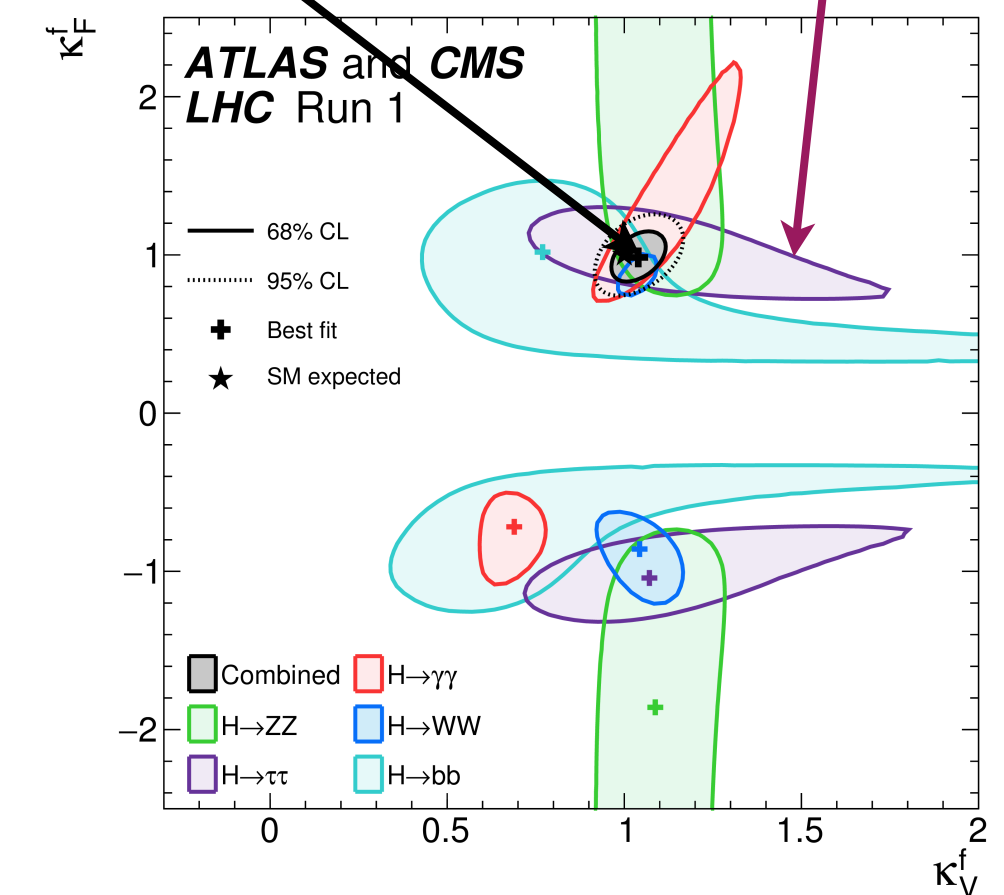


Likelihood function  
 $p(x|\theta)$



Maximum-likelihood estimator

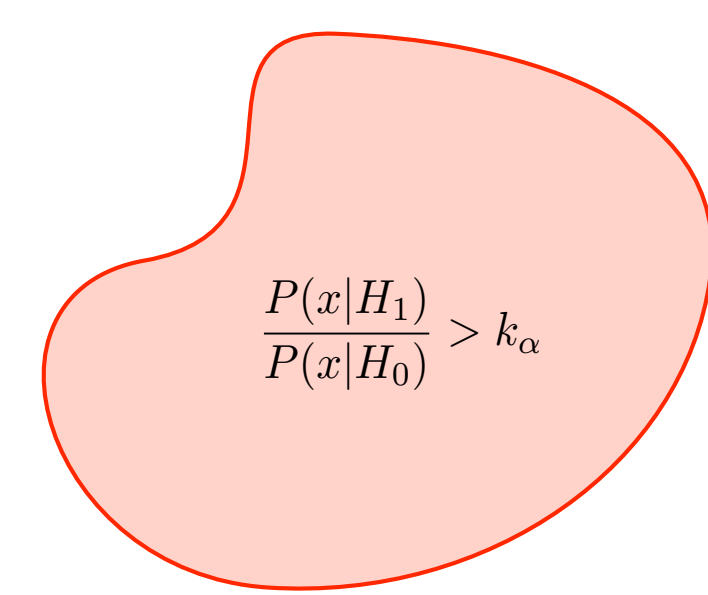
Confidence limits based on likelihood ratio tests



Constraints on parameters  $\theta$



# Recap on Likelihood Ratios


$$\frac{P(x|H_1)}{P(x|H_0)} > k_\alpha$$

For signal vs. background searches:

- **Neyman-Pearson Lemma**: optimal hypothesis test given by **likelihood ratio** (basis of Higgs search)
- Likelihood ratio  $\frac{p(x|\theta_0)}{p(x|\theta_1)}$  also used for exclusion contours

For estimates of parameters  $\hat{\theta}$

- **Cramér-Rao bound** states  $\text{cov}[\hat{\theta}|\theta_0]_{ij} \geq I_{ij}^{-1}(\theta_0)$  where  $I_{ij}$  is the **Fisher-information matrix** (Hessian of log-likelihood)
- Motivates **Information Geometry** as a phenomenological tool
- Maximum-likelihood (asymptotically) saturates the bound

Note:  $\nabla_\theta \log p(x|\theta)$  acts like a likelihood ratio locally



# Cramér-Rao Bound

The minimum variance bound on an unbiased estimator is given by the Cramér-Rao bound:

$$\text{cov}[\hat{\theta}|\theta_0]_{ij} \geq I_{ij}^{-1}(\theta_0)$$

Expected error of best-fit parameter

Inverse of Fisher information

Fisher information matrix (is also a Riemannian metric! )

$$I_{ij}[\theta] = -\mathbb{E} \left[ \frac{\partial^2 \log p(x|\theta)}{\partial \theta_i \partial \theta_j} \middle| \theta \right]$$

Maximum Likelihood Estimators *asymptotically* reach this bound

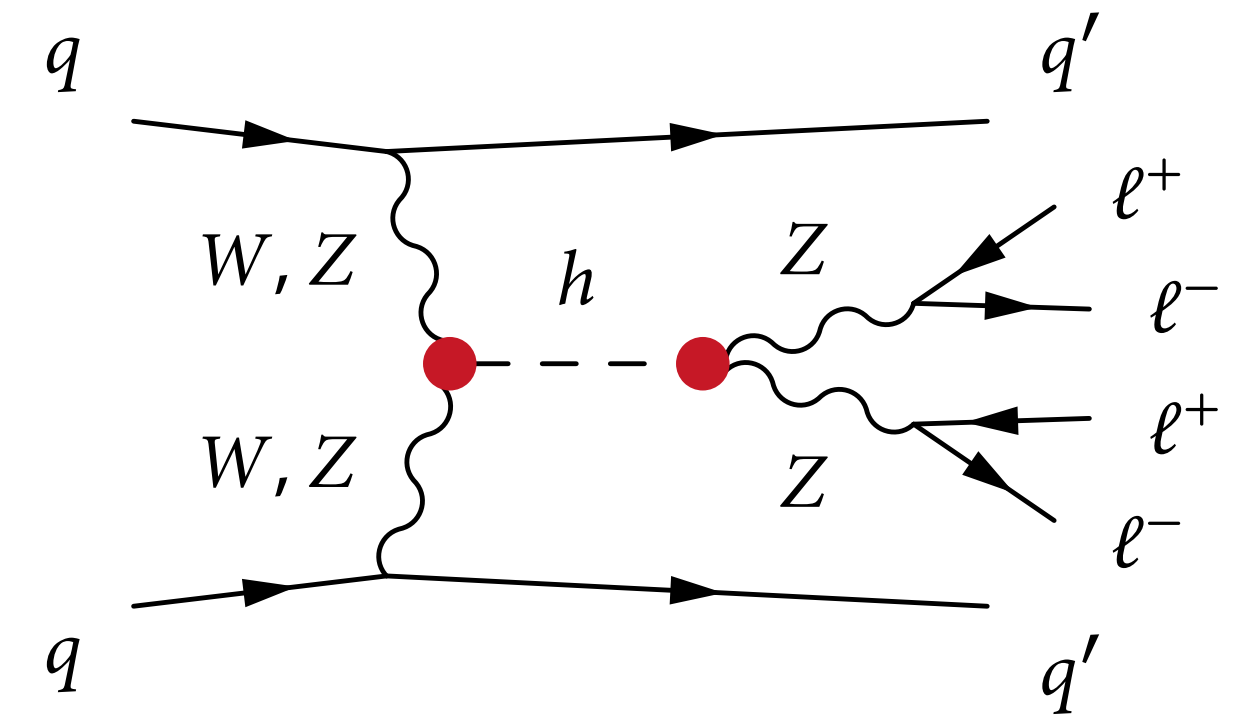


# Challenge for EFT

Let  $\theta$  denote the coefficients of higher dimensional operators in the Lagrangian,  $x$  be high-dimensional data associated to an event, and  $p(x | \theta) = \frac{1}{\sigma(\theta)} \frac{d\sigma}{d\theta}$  be the distribution for the data

- we want to compare any two points in EFT parameter space

- evaluate the **likelihood ratio**  $r(x|\theta_0, \theta_1) \equiv \frac{p(x|\theta_0)}{p(x|\theta_1)}$



Difficulty is that one changes the parameters of the EFT, the distributions  $p(x|\theta)$  change due to interference.

- It would be very computationally expensive (infeasible) to generate samples for every value of  $\theta$  and estimate  $p(x | \theta)$  with histograms. Small changes mean we need a lot of MC events!
- Ideally we could directly estimate the **score**  $t(x|\theta_0) \equiv \nabla_{\theta} \log p(x|\theta) \Big|_{\theta_0}$



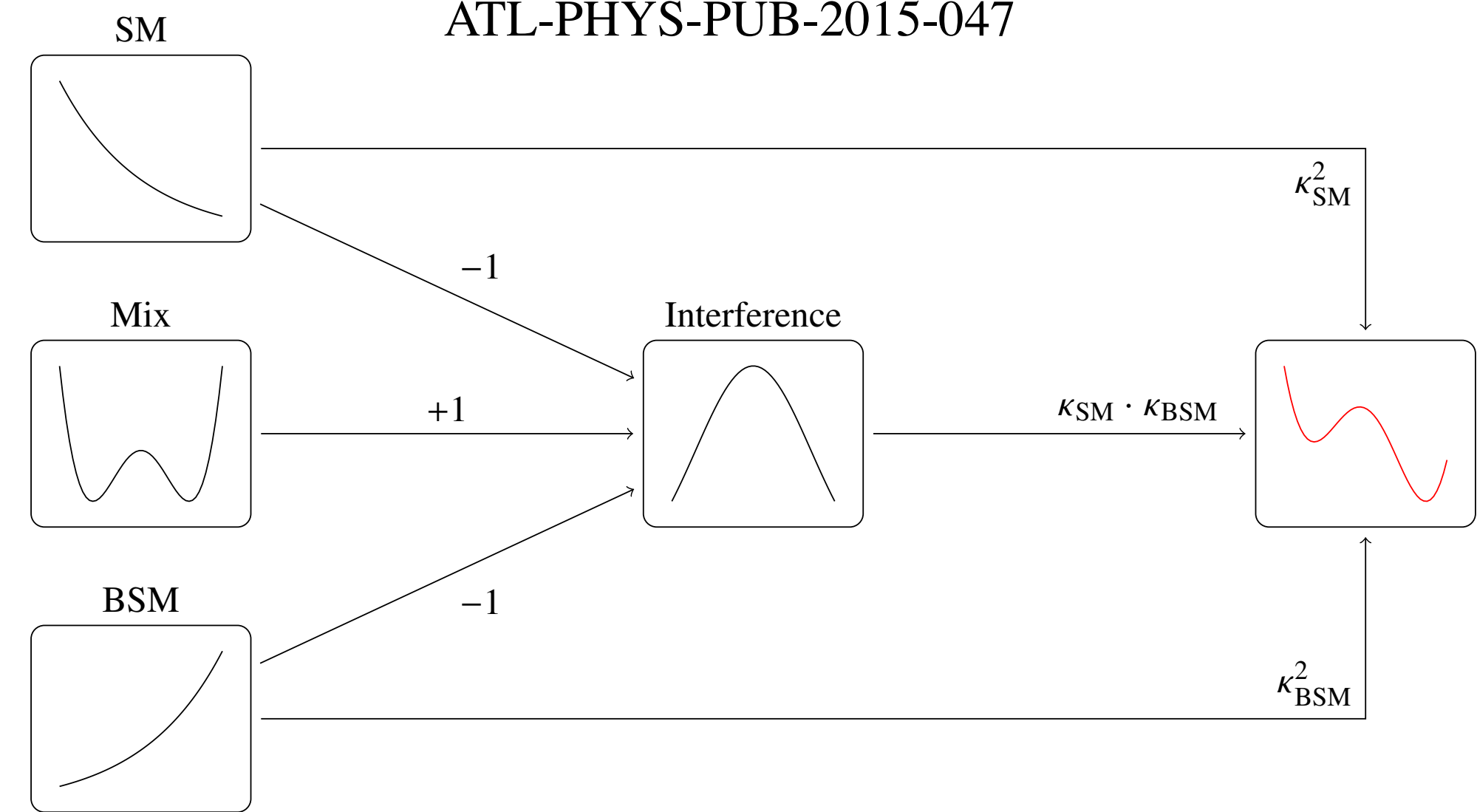
# EFT Embedded in a vector space

Difficulty is that one changes the parameters of the EFT, the distributions  $p(x|\theta)$  change due to interference.

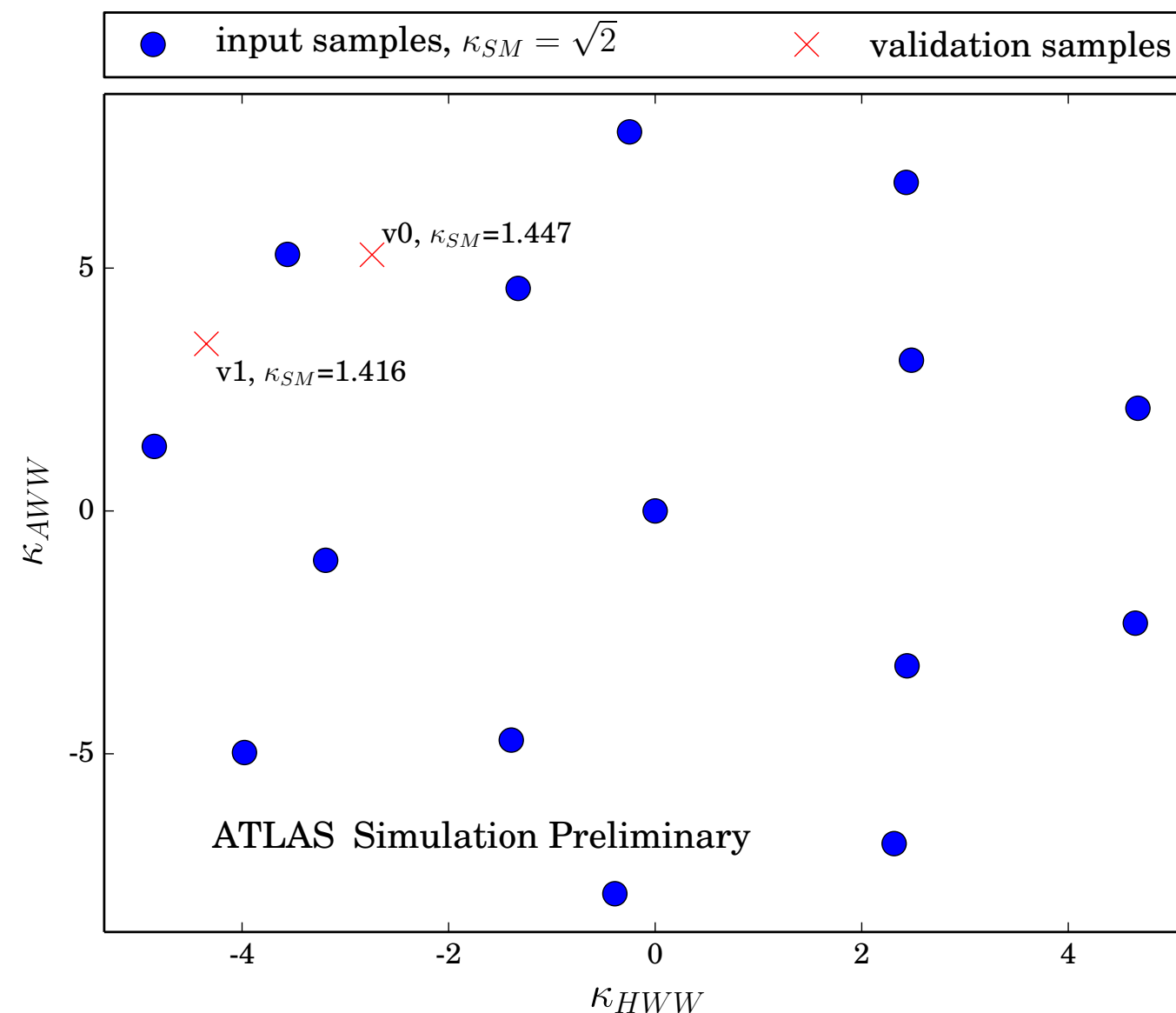
**But there is a trick:**

Simple example:

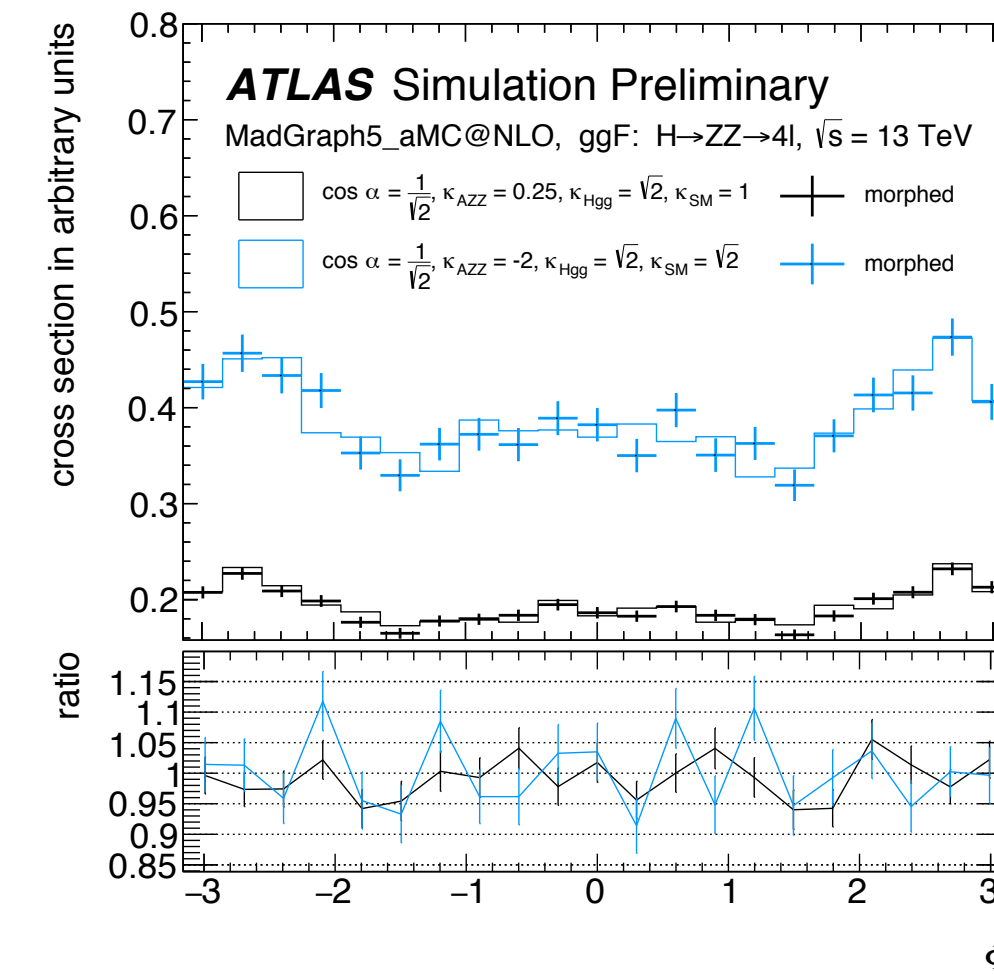
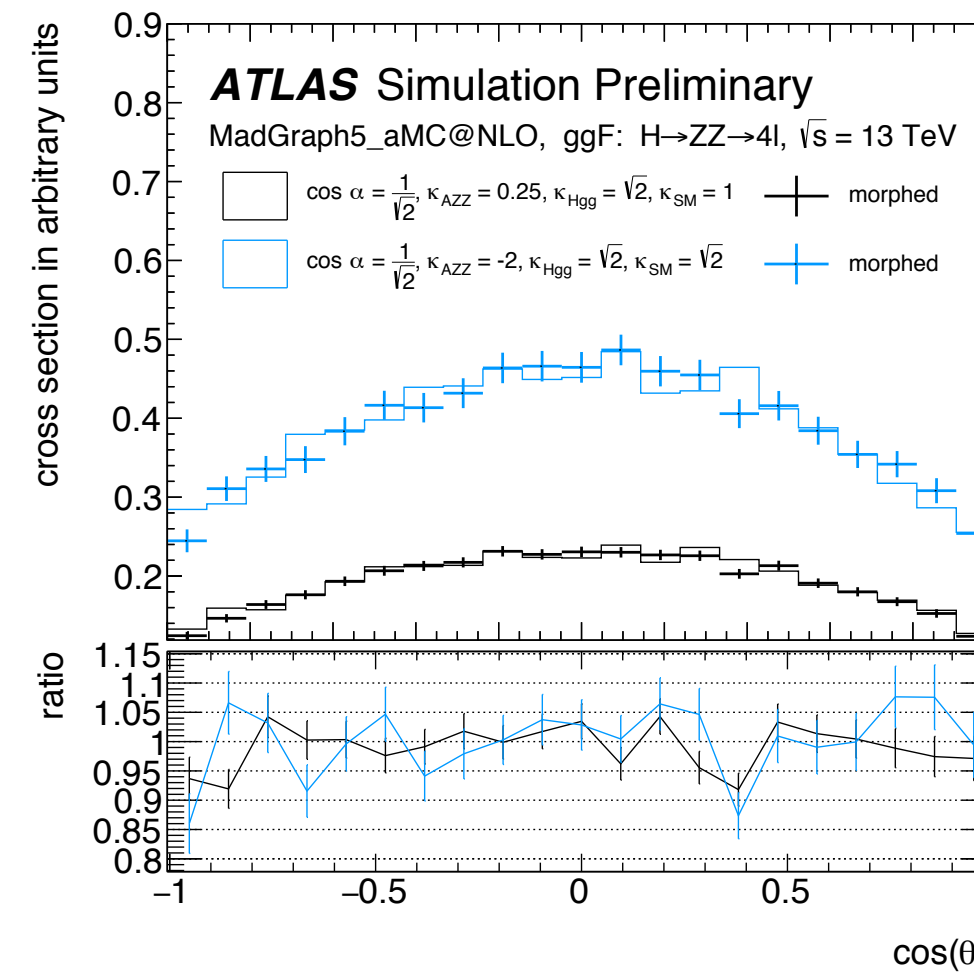
$$|g_1 M_{SM} + g_2 M_{BSM}|^2 = g_1^2 |M_{SM}|^2 + 2g_1 g_2 \text{Re} [M_{SM}^* M_{BSM}] + g_2^2 |M_{BSM}|^2$$



3-d vector space, distribution for any point in this space is linear mixture of distribution for 3 basis samples!



(real examples need more basis samples)





# EFT Decomposition

$$d\sigma \propto \left| \left( \mathcal{M}_{\text{SM}}^p + \sum_i \frac{f_i}{\Lambda^2} \mathcal{M}_i^p \right) \left( \mathcal{M}_{\text{SM}}^d + \sum_j \frac{f_j}{\Lambda^2} \mathcal{M}_j^d \right) \right|^2$$

Express EFT as a mixture:

$$p(x|\theta) = \sum_c w_c(\theta) p_c(x)$$

$w_c(\theta)$  are polynomials

$\nabla_{\theta} \log p(x|\theta)$  is now possible!

Process	Number of components for $n$ operators					$\Sigma$
	$\mathcal{O}(\Lambda^0)$	$\mathcal{O}(\Lambda^{-2})$	$\mathcal{O}(\Lambda^{-4})$	$\mathcal{O}(\Lambda^{-6})$	$\mathcal{O}(\Lambda^{-8})$	
$hV$ / WBF production	1	$n$	$\frac{n(n+1)}{2}$			$\frac{(n+1)(n+2)}{2}$
$h \rightarrow VV$ decay	1	$n$	$\frac{n(n+1)}{2}$			$\frac{(n+1)(n+2)}{2}$
Production + decay	1	$n$	$\frac{n(n+1)}{2}$	$\binom{n+2}{3}$	$\binom{n+3}{4}$	$\binom{n+4}{4}$

Table 1: Number of components  $c$  as given in Eq. (6) for different processes, sorted by their suppression by the EFT cutoff scale  $\Lambda$ .

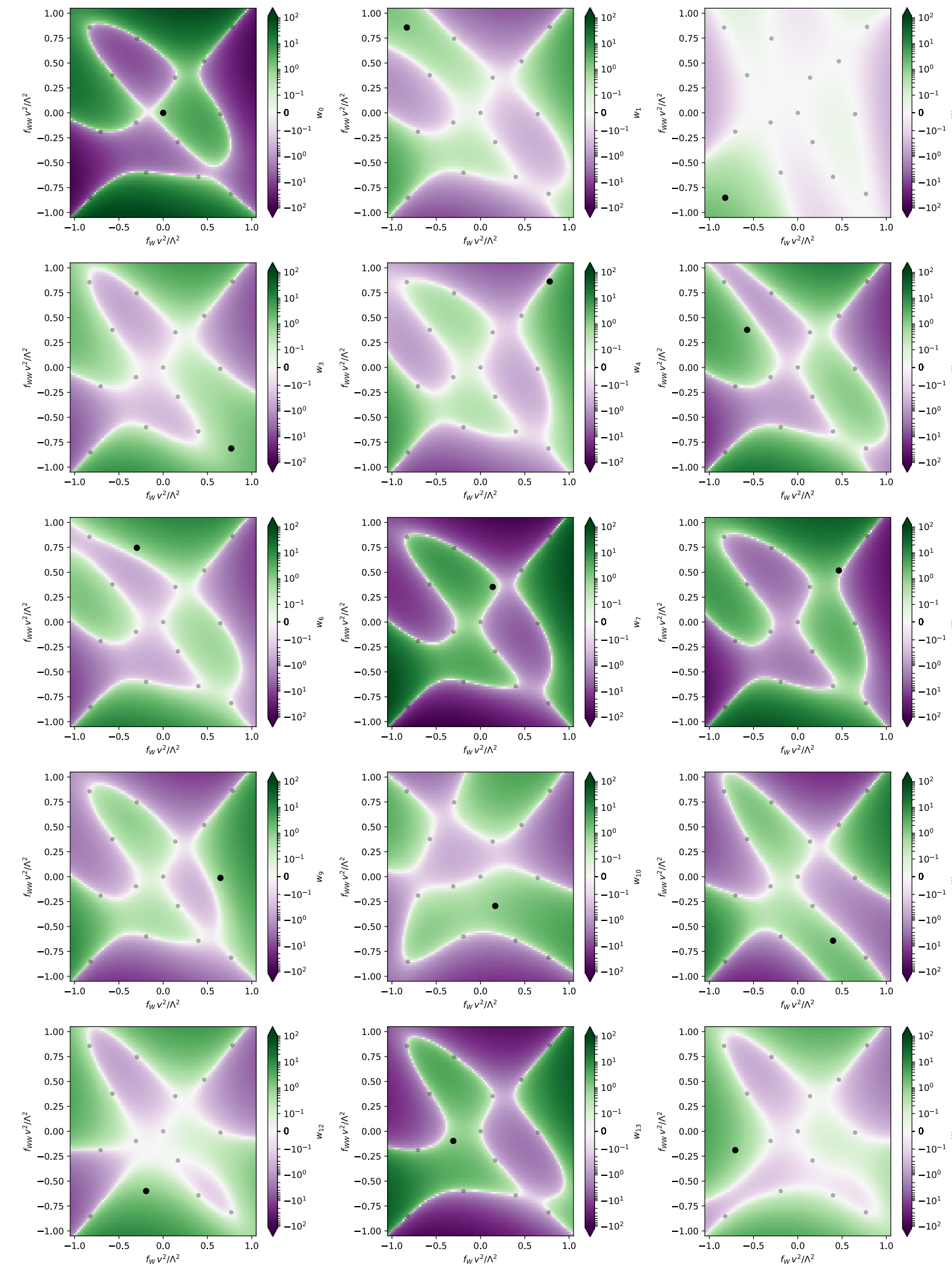


Figure 13: Morphing weights  $w_i(\theta)$  for basis points distributed over the full relevant parameter space.

For 2 BSM operators affecting VBF Higgs production and decay, we need a 15-D vector space

For 5 BSM operators we need 126-D vector space



# Information Geometry

- ▶ Theory language: dimension-6 operators of SM EFT,  $\mathcal{L} \supset \sum_i \frac{f_i}{\Lambda^2} \mathcal{O}_i$   
[W. Buchmuller, D. Wyler 85; K. Hagiwara, S. Ishihara, S. R. Szalapski, D. Zeppenfeld 93;  
B. Grzadkowski, M. Iskrzynski, M. Misiak, J. Rosiek 1008.4884; ...]

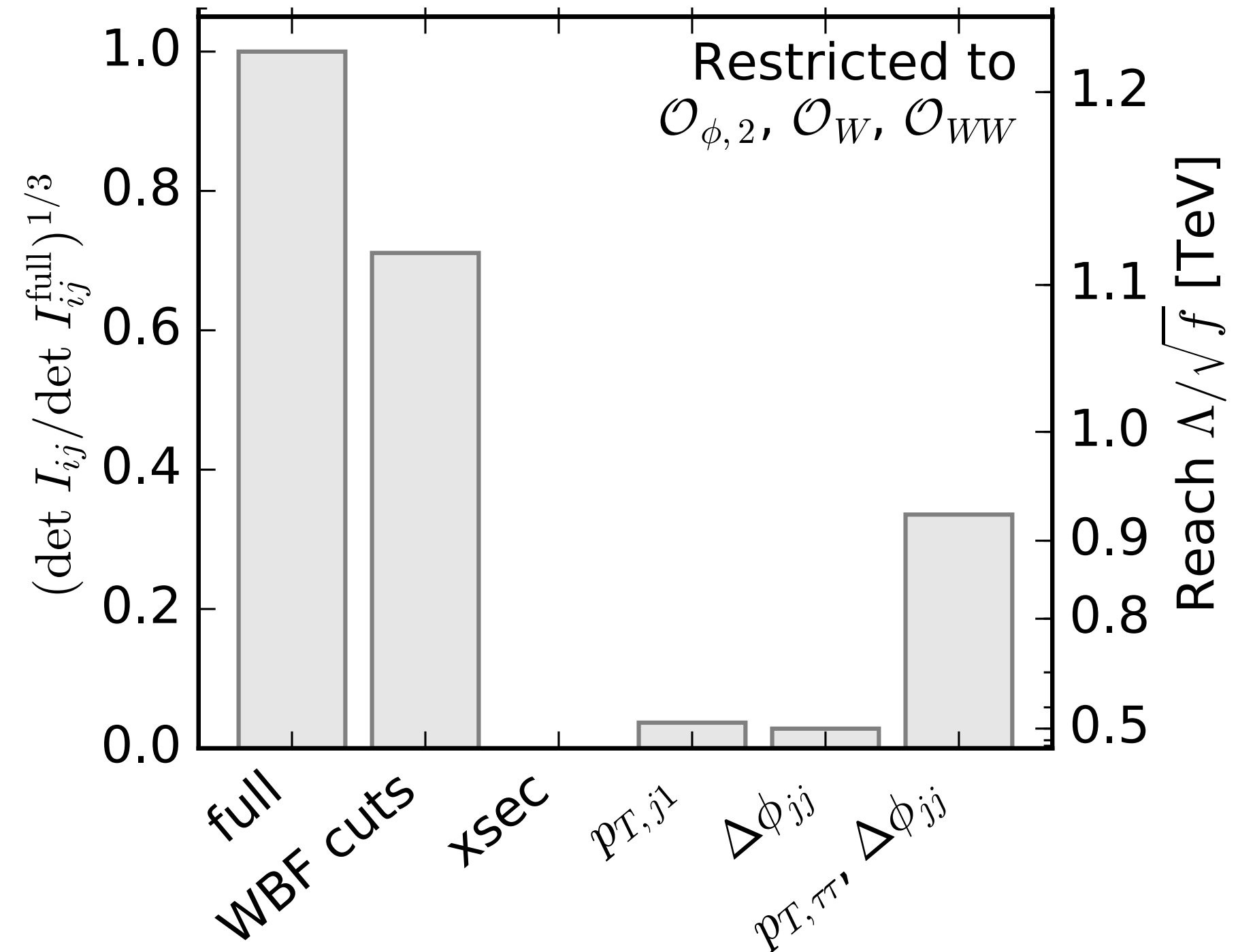
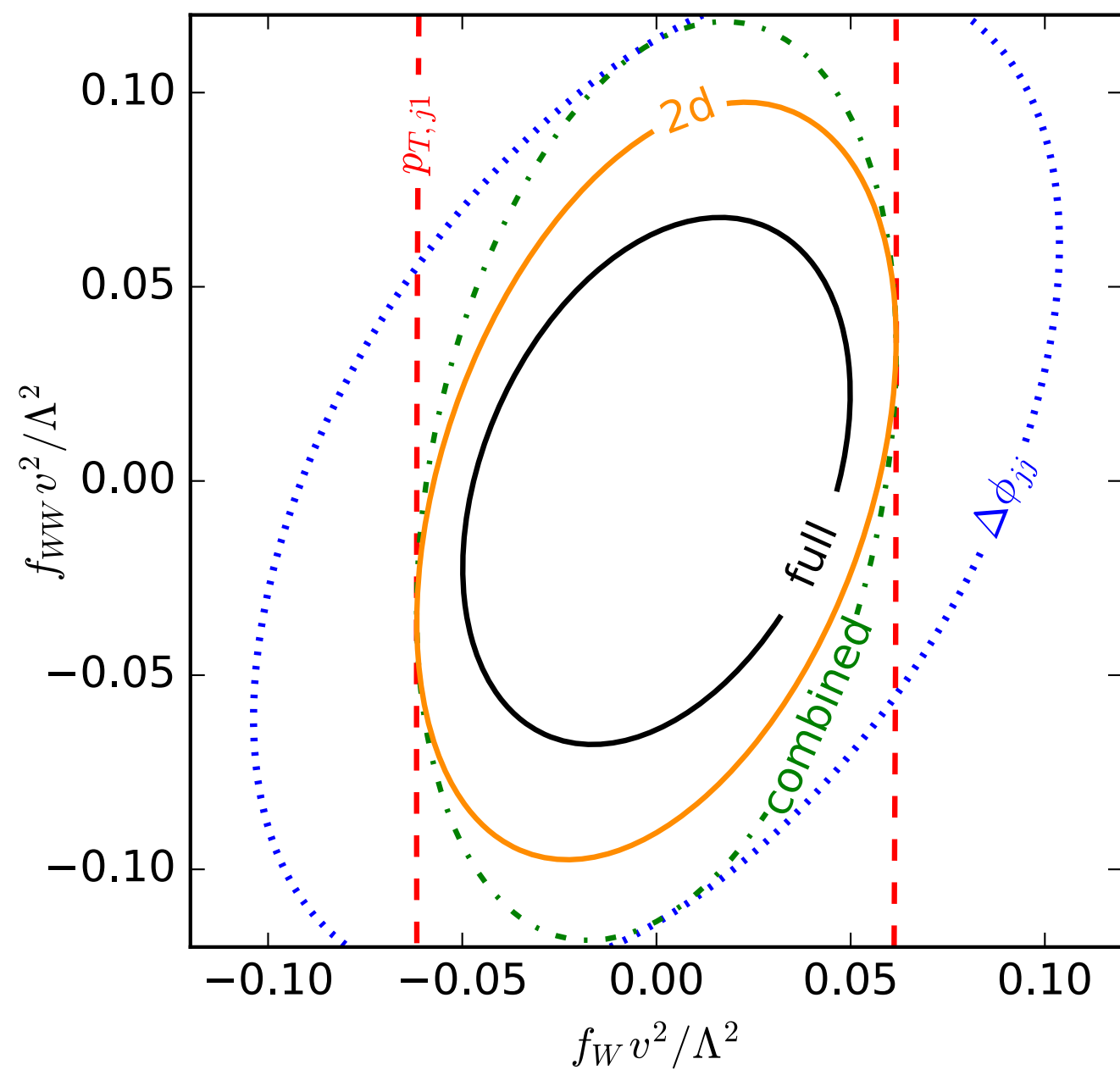
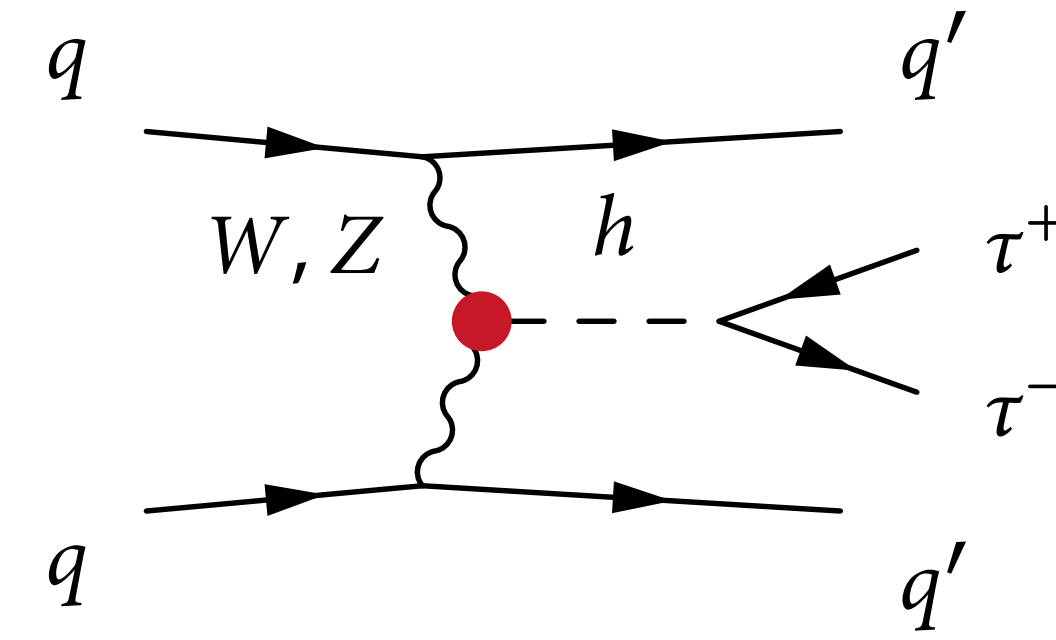
- ▶ Total rate:  $\mathcal{O}_{\phi,2} = \frac{1}{2} \partial^\mu (\phi^\dagger \phi) \partial_\mu (\phi^\dagger \phi)$

- ▶ New kinematic structures:

$$\mathcal{O}_B = i \frac{g}{2} (D^\mu \phi^\dagger) (D^\nu \phi) B_{\mu\nu} \quad \mathcal{O}_W = i \frac{g}{2} (D^\mu \phi)^\dagger \sigma^k (D^\nu \phi) W_{\mu\nu}^k$$

$$\mathcal{O}_{BB} = -\frac{g'^2}{4} (\phi^\dagger \phi) B_{\mu\nu} B^{\mu\nu} \quad \mathcal{O}_{WW} = -\frac{g^2}{4} (\phi^\dagger \phi) W_{\mu\nu}^k W^{\mu\nu k}$$

- ▶ CP violation:  $\mathcal{O}_{W\tilde{W}} = -\frac{g^2}{4} (\phi^\dagger \phi) W_{\mu\nu}^k \tilde{W}^{\mu\nu k}$
- ▶ Others strongly constrained by EWPD or redundant





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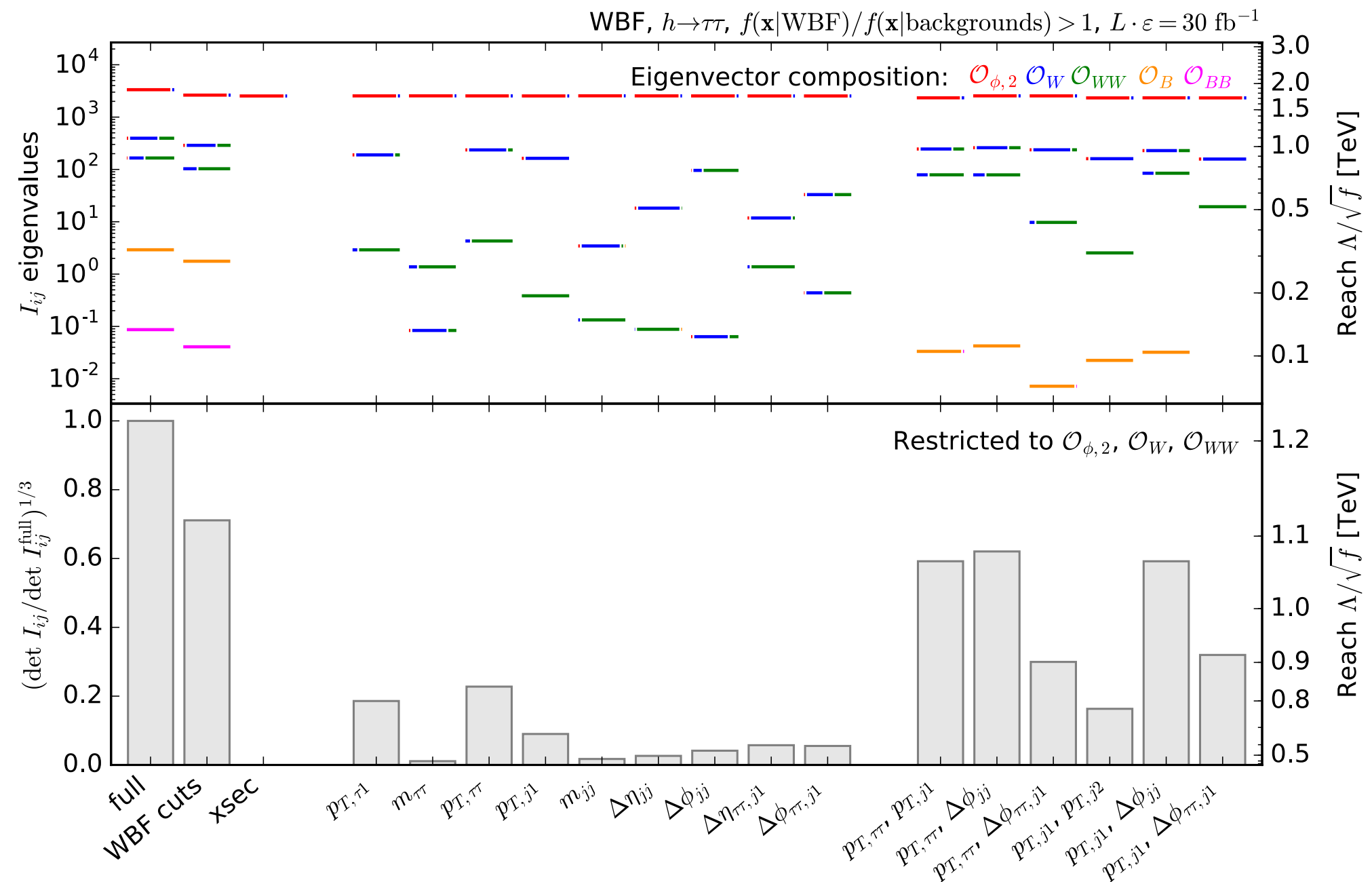
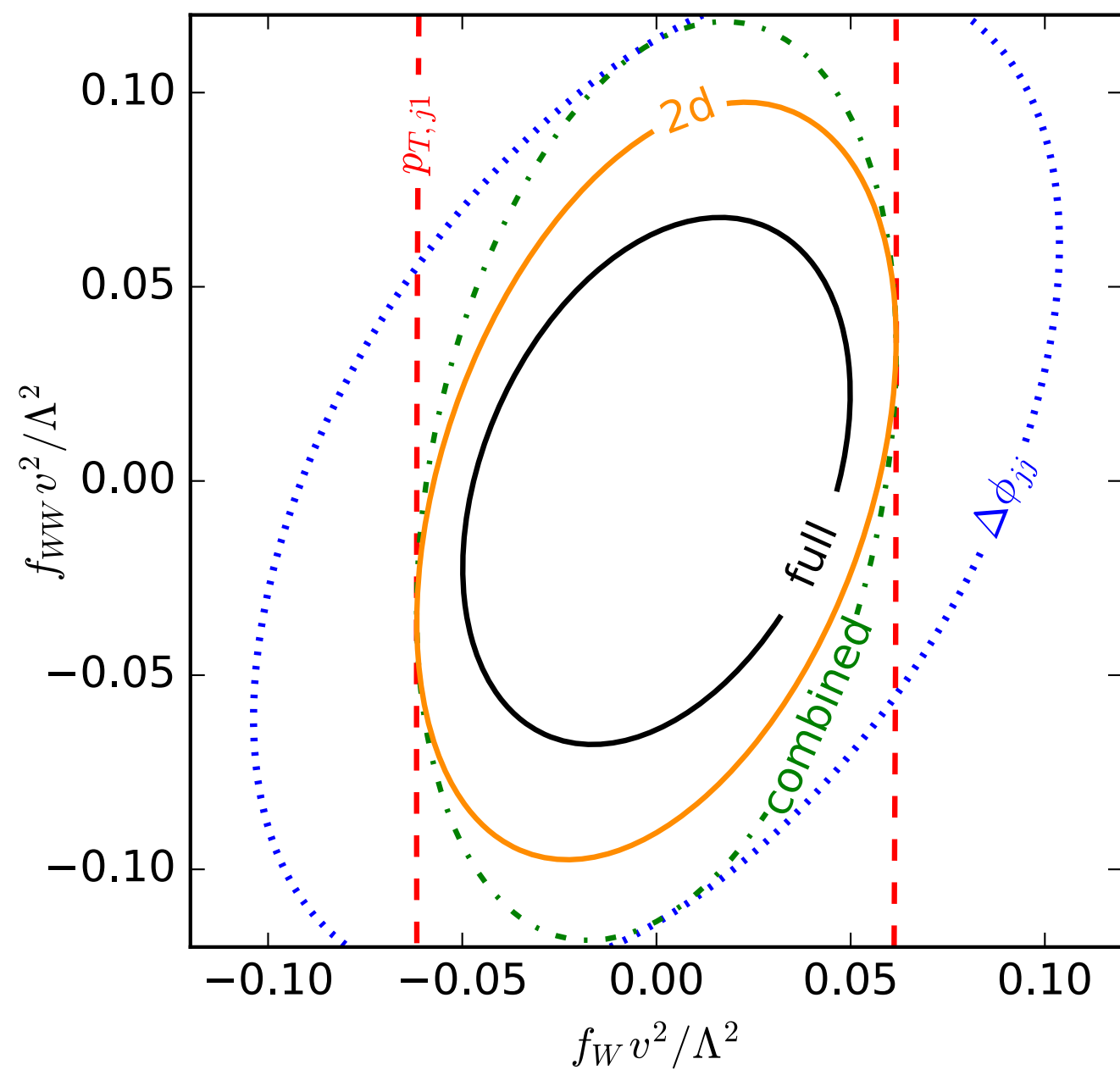
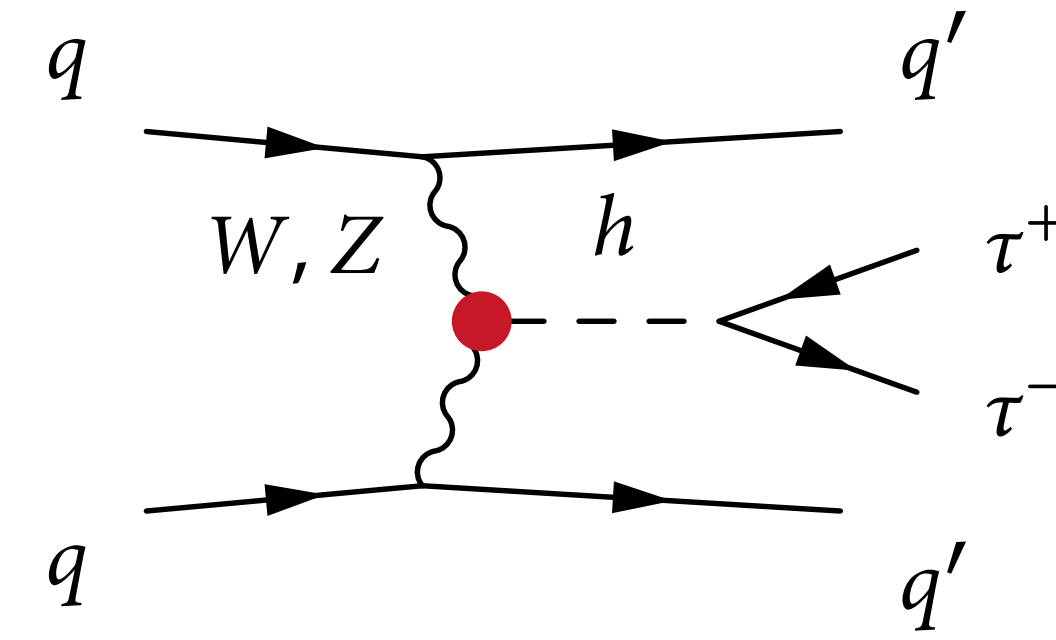
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## Information geometry provides a very powerful tool for phenomenology of EFT

- formal bounds on how well parameters can be measured
- exploit fully differential cross-section
- Eg: a global fit (13 parameters) & can profile/marginalize parameters you aren't interested in (eg. CP violating vs. CP conserving)

For an effective Higgs-gauge Lagrangian truncated at mass dimension six,

$$\mathcal{L} = \mathcal{L}_{\text{SM}} + \frac{f_i}{\Lambda^2} \mathcal{O}_i \quad (9)$$

our  $CP$ -even reference scenario consists of the renormalizable Standard Model Lagrangian combined with the five  $CP$ -even dimension-six operators in the HISZ basis [6, 7, 35],

$$\begin{aligned} \mathcal{O}_B &= i\frac{g}{2} (D^\mu \phi^\dagger)(D^\nu \phi) B_{\mu\nu} & \mathcal{O}_W &= i\frac{g}{2} (D^\mu \phi)^\dagger \sigma^k (D^\nu \phi) W_{\mu\nu}^k \\ \mathcal{O}_{BB} &= -\frac{g^2}{4} (\phi^\dagger \phi) B_{\mu\nu} B^{\mu\nu} & \mathcal{O}_{WW} &= -\frac{g^2}{4} (\phi^\dagger \phi) W_{\mu\nu}^k W^{\mu\nu k} \\ \mathcal{O}_{\phi,2} &= \frac{1}{2} \partial^\mu (\phi^\dagger \phi) \partial_\mu (\phi^\dagger \phi) . \end{aligned} \quad (10)$$

At the same mass dimension,  $CP$ -odd couplings are described by operators

$$\begin{aligned} \mathcal{O}_{B\tilde{B}} &= -\frac{g^2}{4} (\phi^\dagger \phi) \tilde{B}_{\mu\nu} B^{\mu\nu} \equiv -\frac{g^2}{4} (\phi^\dagger \phi) \epsilon_{\mu\nu\rho\sigma} B^{\rho\sigma} B^{\mu\nu} \\ \mathcal{O}_{W\tilde{W}} &= -\frac{g^2}{4} (\phi^\dagger \phi) \tilde{W}_{\mu\nu}^k W^{\mu\nu k} \equiv -\frac{g^2}{4} (\phi^\dagger \phi) \epsilon_{\mu\nu\rho\sigma} W^{\rho\sigma k} W^{\mu\nu k} . \end{aligned} \quad (11)$$

With the Levi-Civita tensor, these operators break down as  $C$ -conserving and  $P$ -violating.

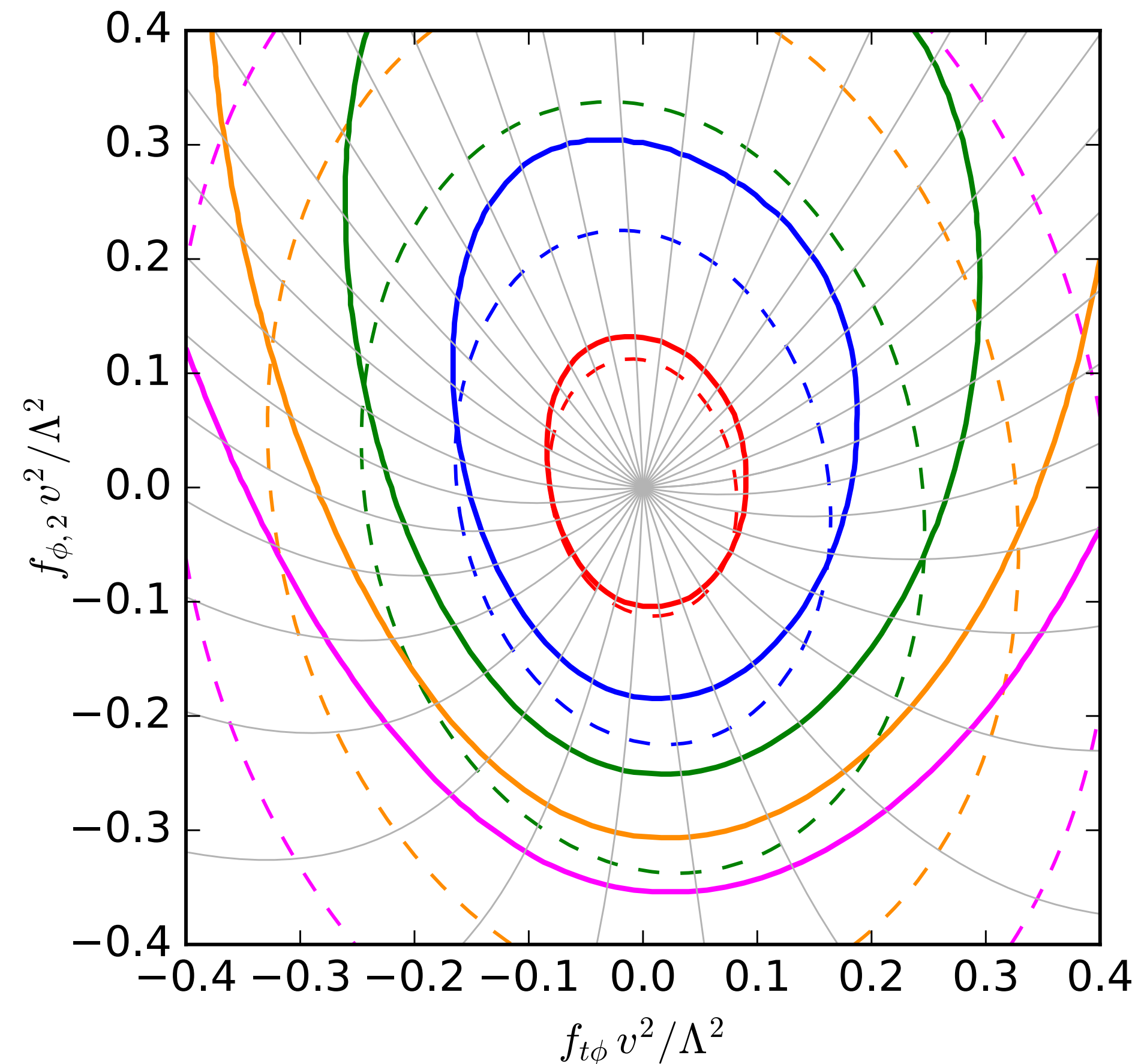
$$I_{ij} = \begin{pmatrix} f_{\phi,2} & f_W & f_B & f_{WW} & f_{BB} & f_{W\tilde{W}} & f_{B\tilde{B}} & \text{Im } f_W & \text{Im } f_B & \text{Im } f_{WW} & \text{Im } f_{BB} & \text{Im } f_{W\tilde{W}} & \text{Im } f_{B\tilde{B}} & f_{\phi,2} \\ -968 & 715 & 35 & -191 & -3 & 1 & 0 & 0 & 0 & 0 & 0 & -55 & -1 & f_W \\ -50 & 35 & 6 & -9 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & -2 & 0 & f_B \\ 54 & -191 & -9 & 321 & 3 & -1 & 0 & 0 & 0 & 1 & 0 & 72 & 1 & f_{WW} \\ 2 & -3 & 0 & 3 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 1 & 0 & f_{BB} \\ -7 & 1 & 0 & -1 & 0 & 359 & 4 & 41 & 1 & -81 & -1 & -1 & 0 & f_{W\tilde{W}} \\ 0 & 0 & 0 & 0 & 0 & 4 & 0 & 0 & 0 & -1 & 0 & 0 & 0 & f_{B\tilde{B}} \\ -1 & 0 & 0 & 0 & 0 & 41 & 0 & 6 & 0 & -12 & 0 & 0 & 0 & \text{Im } f_W \\ 0 & 0 & 0 & 0 & 0 & 1 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & \text{Im } f_B \\ 2 & 0 & 0 & 1 & 0 & -81 & -1 & -12 & 0 & 23 & 0 & 0 & 0 & \text{Im } f_{WW} \\ 0 & 0 & 0 & 0 & 0 & -1 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & \text{Im } f_{BB} \\ 36 & -55 & -2 & 72 & 1 & -1 & 0 & 0 & 0 & 0 & 0 & 21 & 0 & \text{Im } f_{W\tilde{W}} \\ 0 & -1 & 0 & 1 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & \text{Im } f_{B\tilde{B}} \end{pmatrix}, \quad (30)$$



# The Information Geometry of an EFT

• [JB, K. Cranmer, F. Kling, T. Plehn 1612.05261;  
JB, F. Kling, T. Plehn, T. Tait 1712.02350]

- With a metric, it now makes sense to talk about how far two points in theory space are from one another. And we can go beyond the ellipse and calculate geodesics!



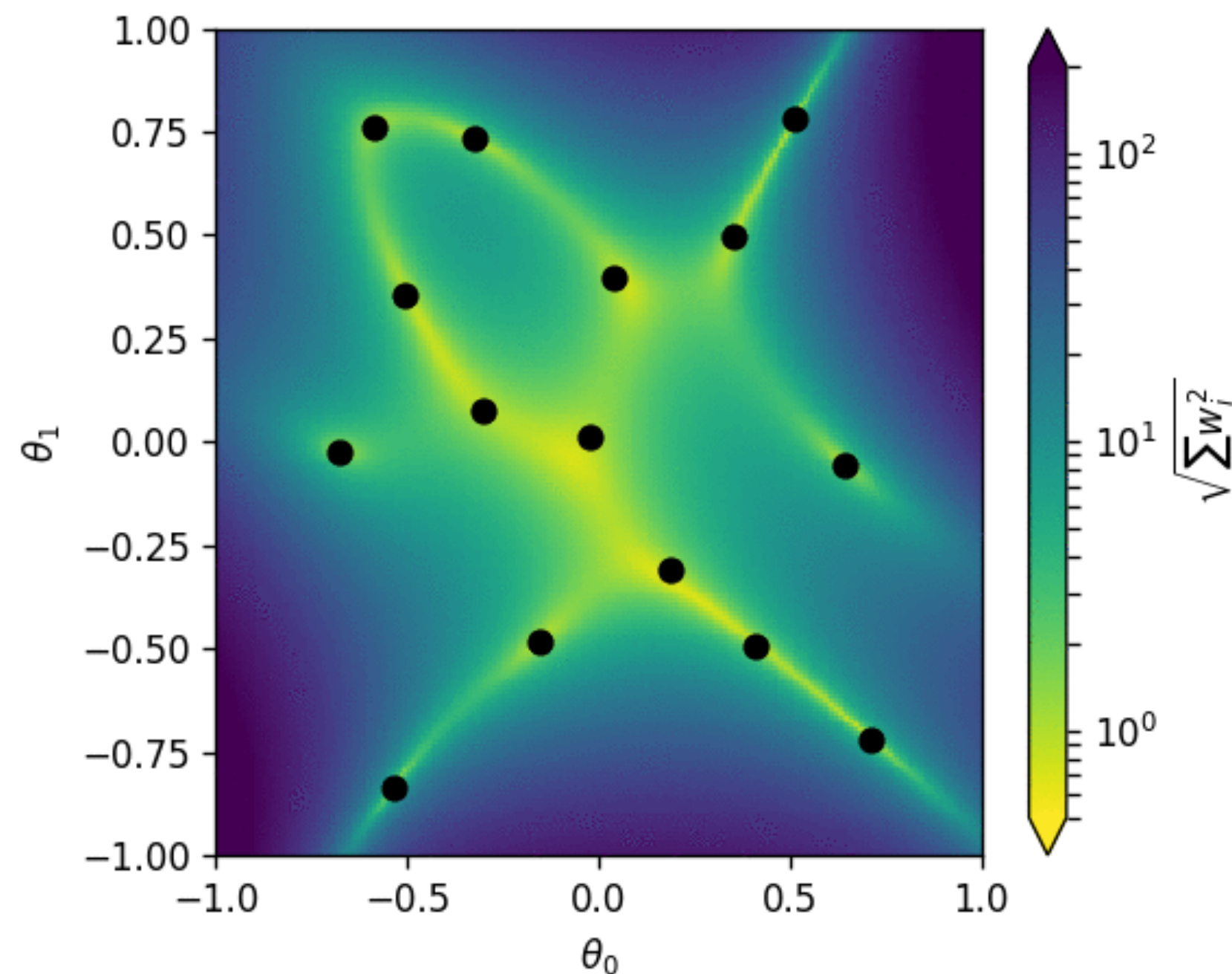
- **Parametrization-independent, geometric description** of sensitivity based on Fisher information



# MadMiner: Hands-on Tutorial

MadMiner automates the morphing for an arbitrary EFT and has utilities to to exploit information geometry to understand the phenomenology of EFT at a specific collider

- formal bounds on how well parameters can be measured
- exploit fully differential cross-section
- useful for Snowmass!



<https://cranmer.github.io/madminer-tutorial/>

**MadMiner Tutorial**

Introduction

**MadMiner tutorial**

This is a tutorial on [MadMiner](#) developed by Johann Brehmer, Felix Kling, Irina Espejo, and Kyle Cranmer. It is built using [Jupyter Book](#).

**Introduction to MadMiner**

Particle physics processes are usually modelled with complex Monte-Carlo simulations of the hard process, parton shower, and detector interactions. These simulators typically do not admit a tractable likelihood function: given a (potentially high-dimensional) set of observables, it is usually not possible to calculate the probability of these observables for some model parameters. Particle physicists usually tackle this problem of “likelihood-free inference” by hand-picking a few “good” observables or summary statistics and filling histograms of them. But this conventional approach discards the information in all other observables and often does not scale well to high-dimensional problems.

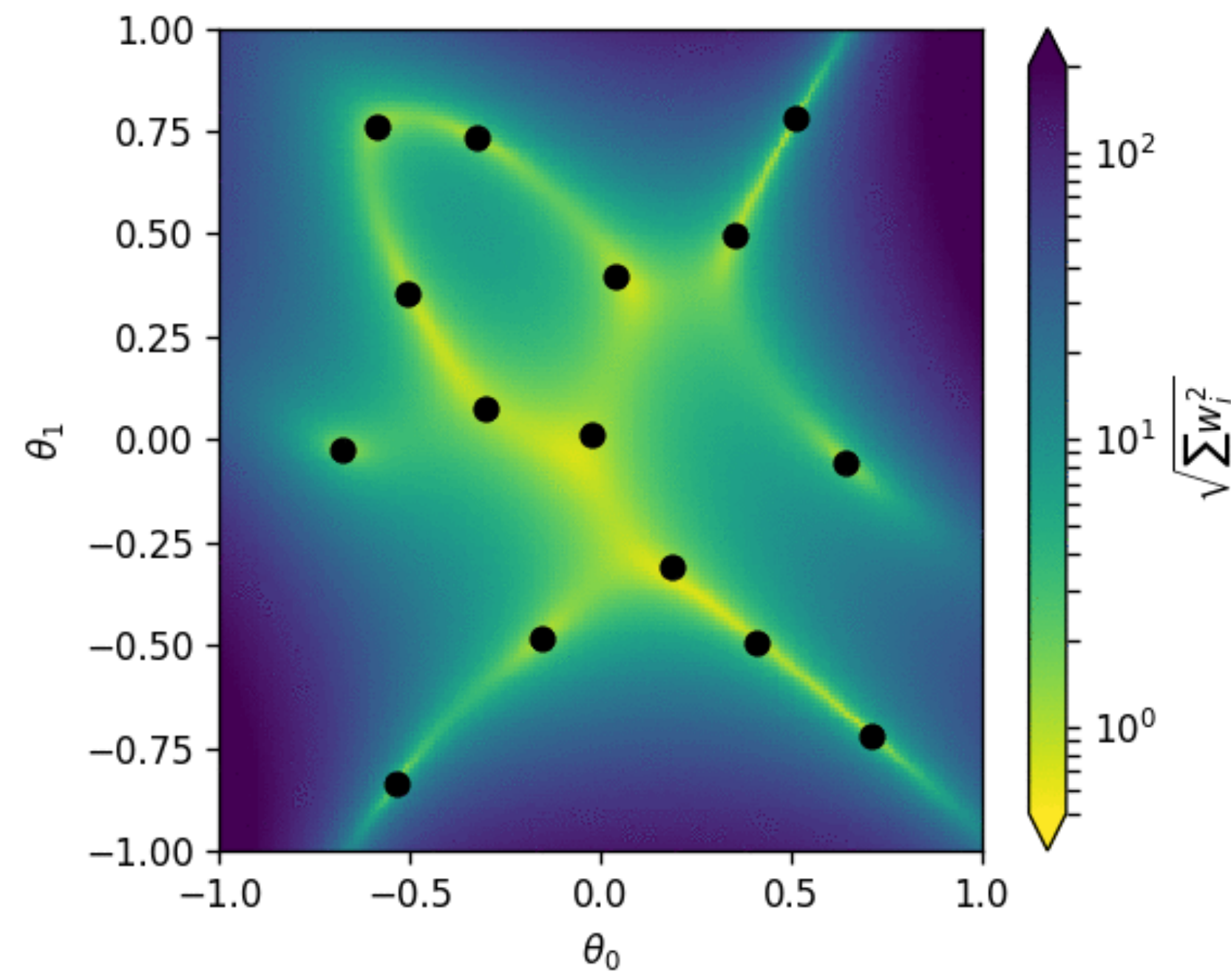
In the three publications “[Constraining Effective Field Theories With Machine Learning](#)”, “[A Guide to Constraining Effective Field Theories With Machine Learning](#)”, and “[Mining gold from implicit models to improve likelihood-free inference](#)”, a new approach has been developed. In a nut shell, additional information is extracted from the simulations that is closely related to the matrix elements that determine the hard process. This “augmented data” can be used to train neural networks to efficiently approximate arbitrary likelihood ratios. We playfully call this process “mining gold” from the simulator, since this information may be hard to get, but turns out to be very valuable for inference.



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Now for some bad news....



Particle physics processes do not have a tractable likelihood function.

# Modeling particle physics processes

Theory  
parameters  
 $\theta$





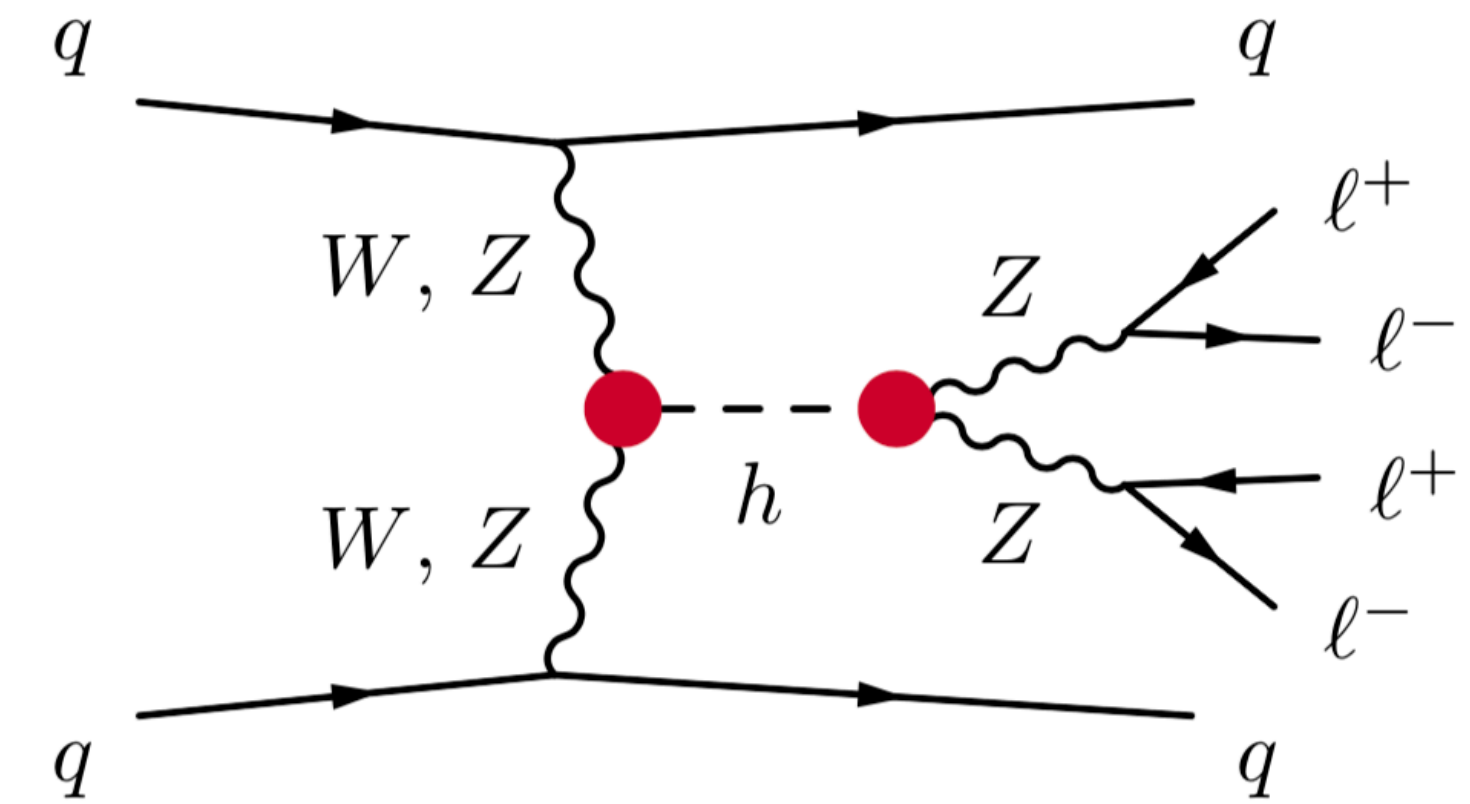
# Modeling particle physics processes

Latent variables

Parton-level  
momenta

Theory  
parameters

$z_p$  ←  $\theta$



← Evolution

# Modeling particle physics processes

Latent variables

Shower  
splittings

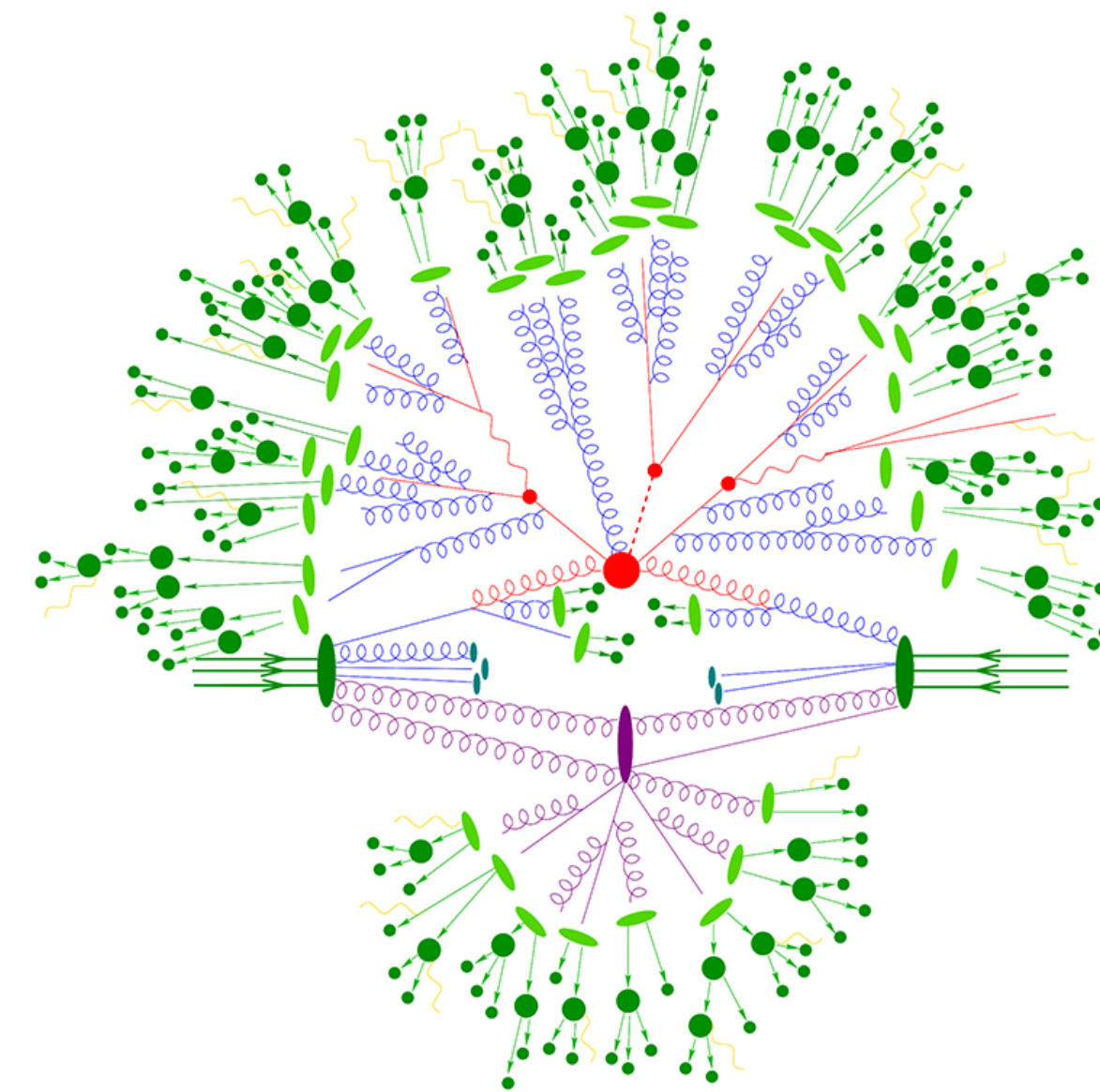
Parton-level  
momenta

Theory  
parameters

$z_s$

$z_p$

$\theta$



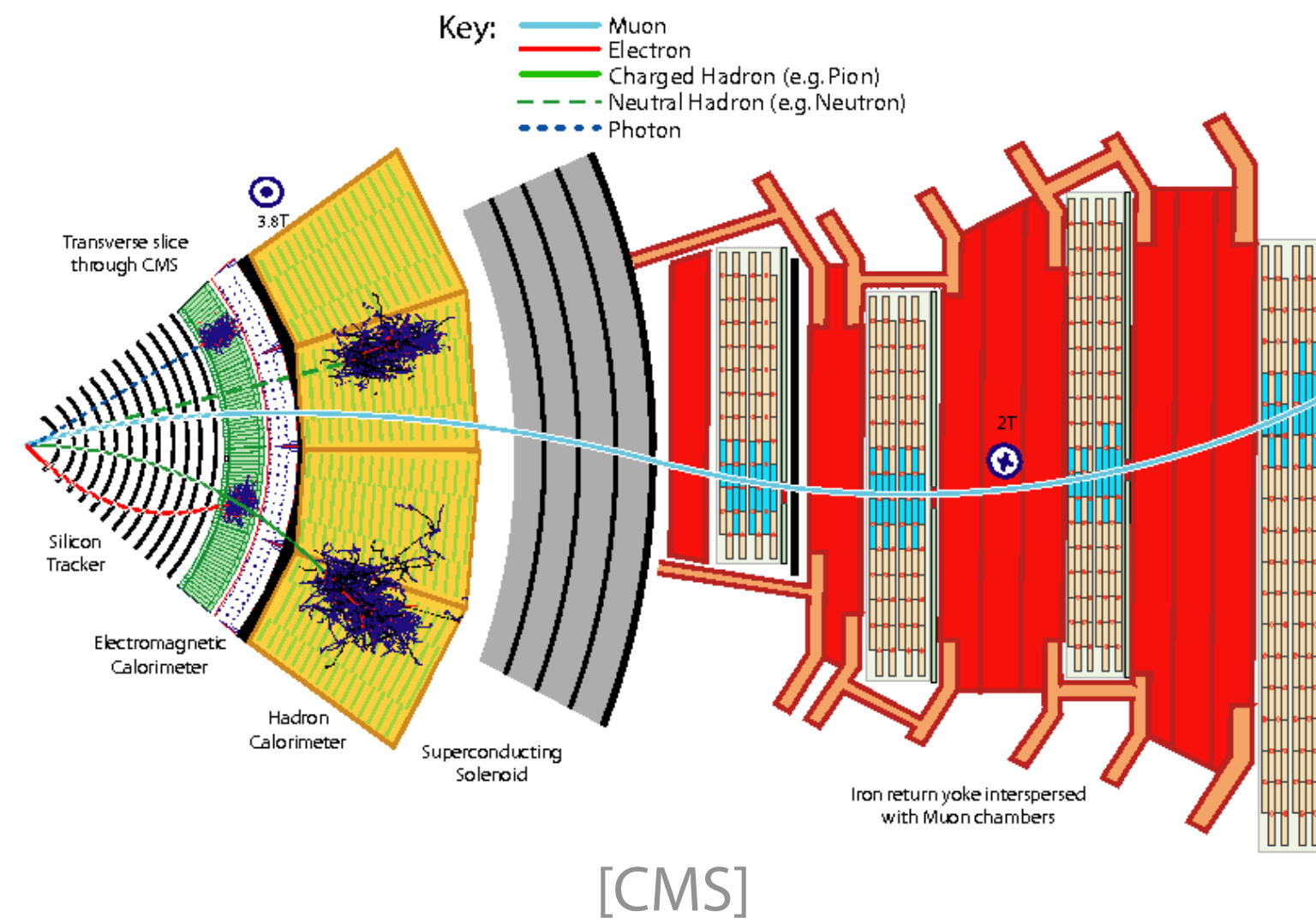
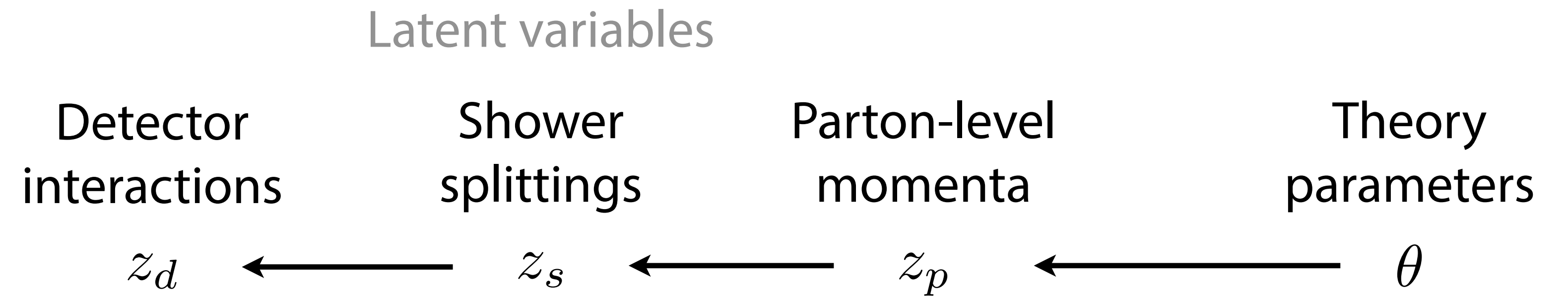
[F. Krauss]



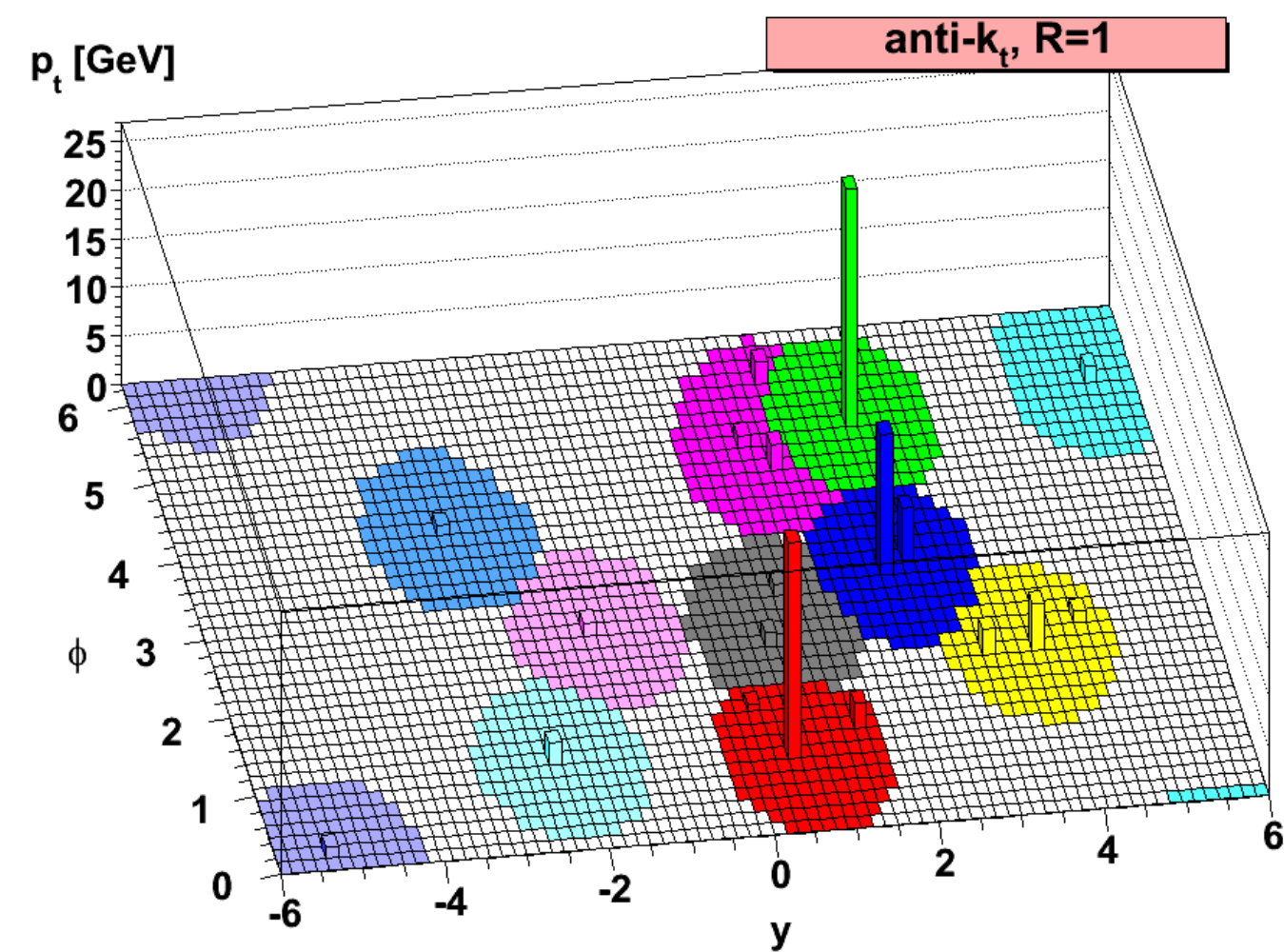
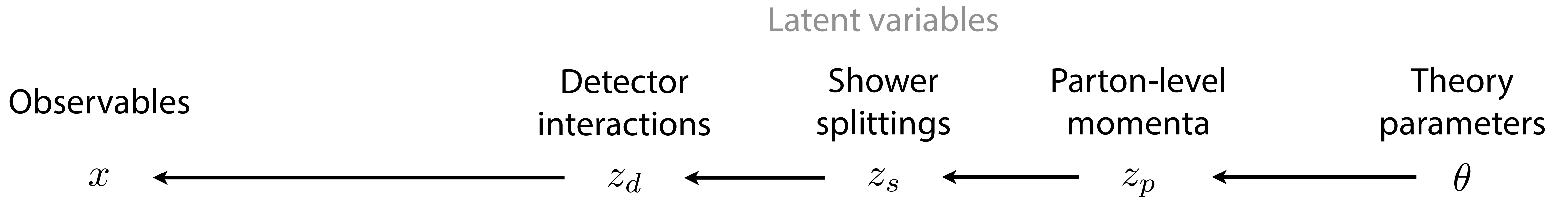
Evolution



# Modeling particle physics processes



# Modeling particle physics processes

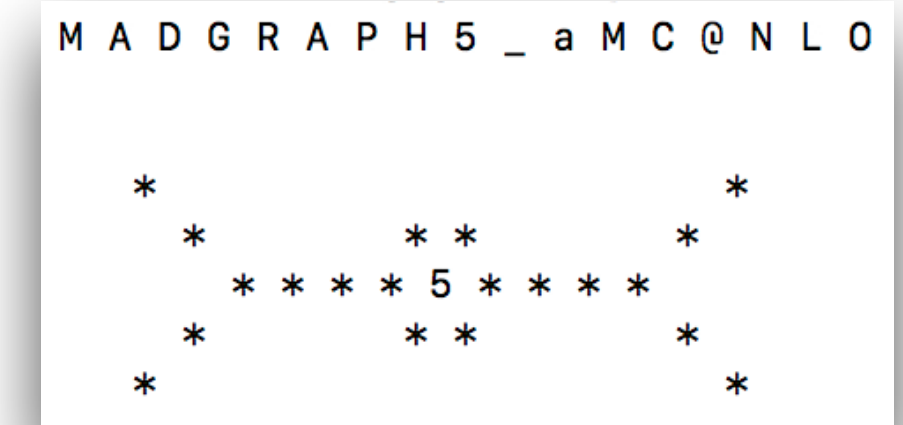
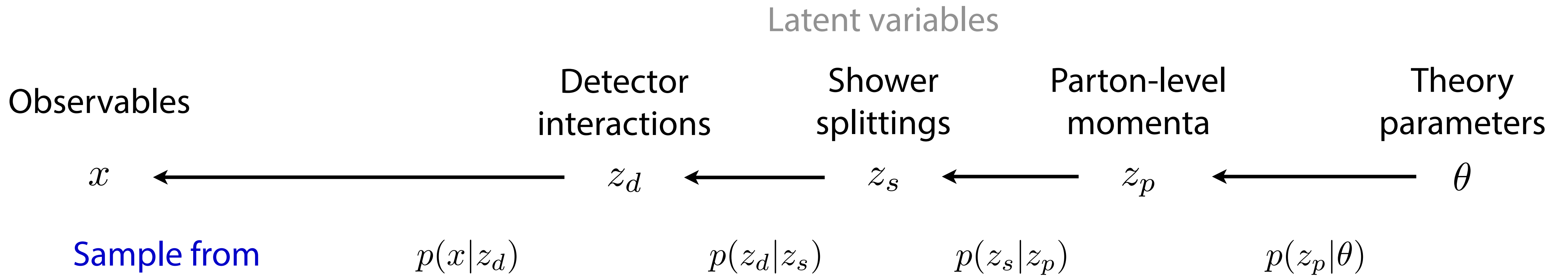


[M. Cacciari, G. Salam, G. Soyez 0802.1189]



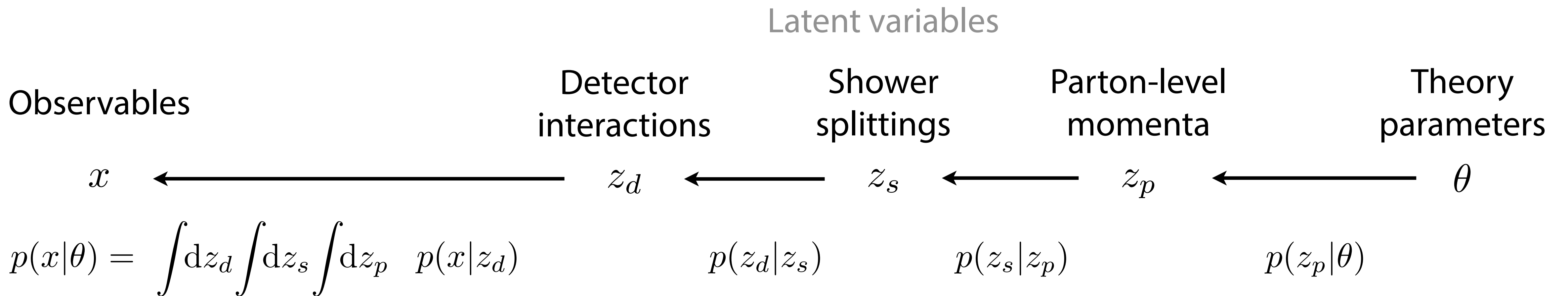


# Modeling particle physics processes



←————— Prediction (simulation)

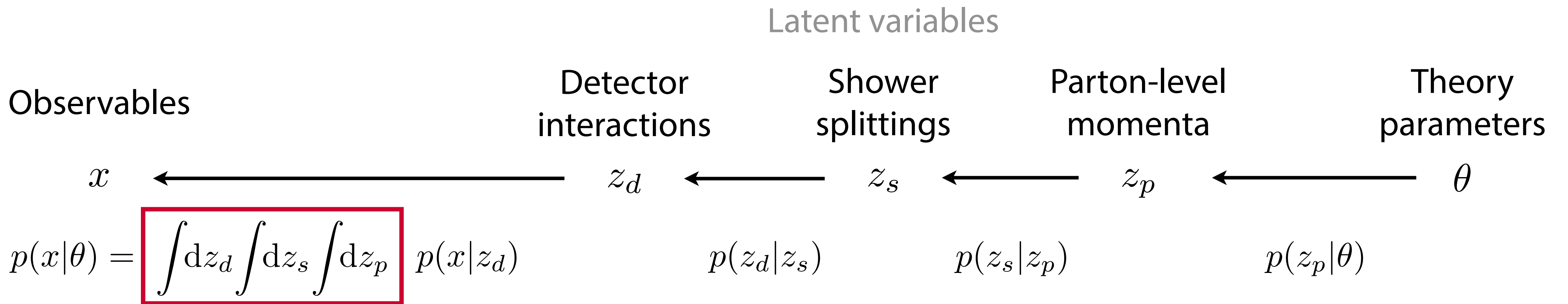
# Modeling particle physics processes



Inference



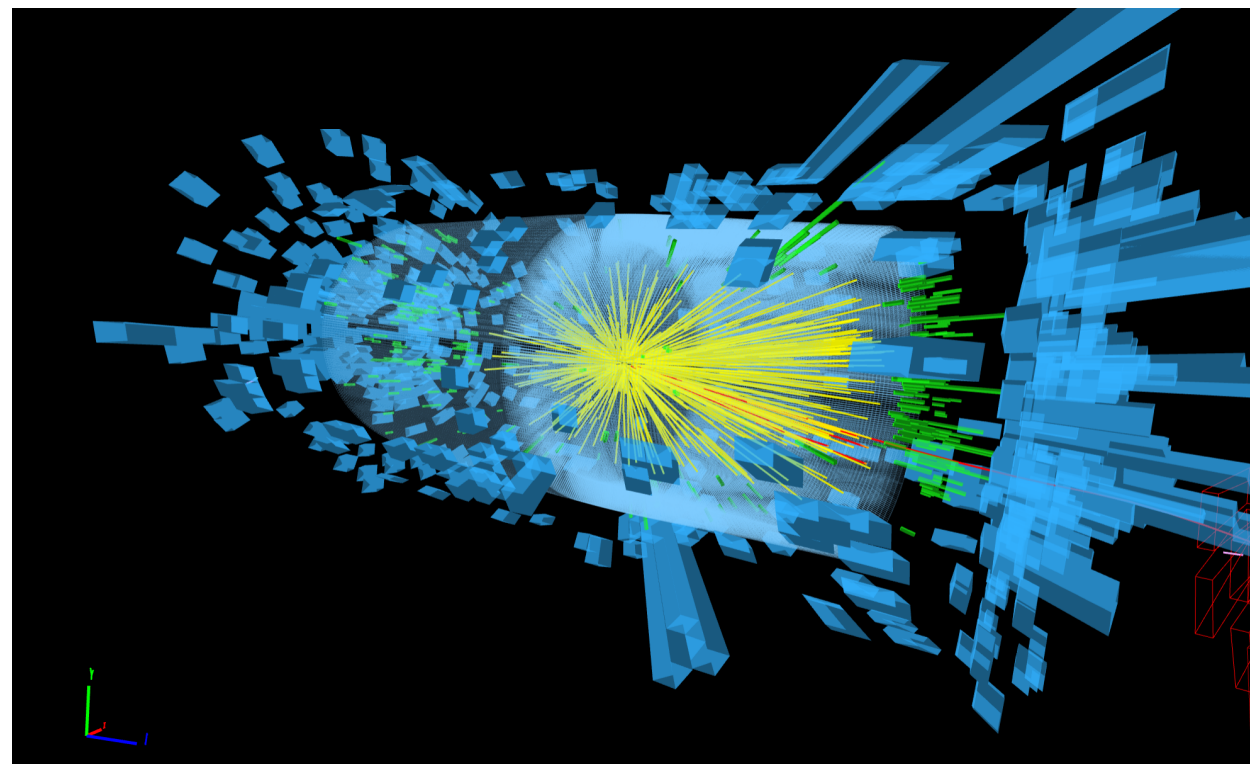
# Modeling particle physics processes



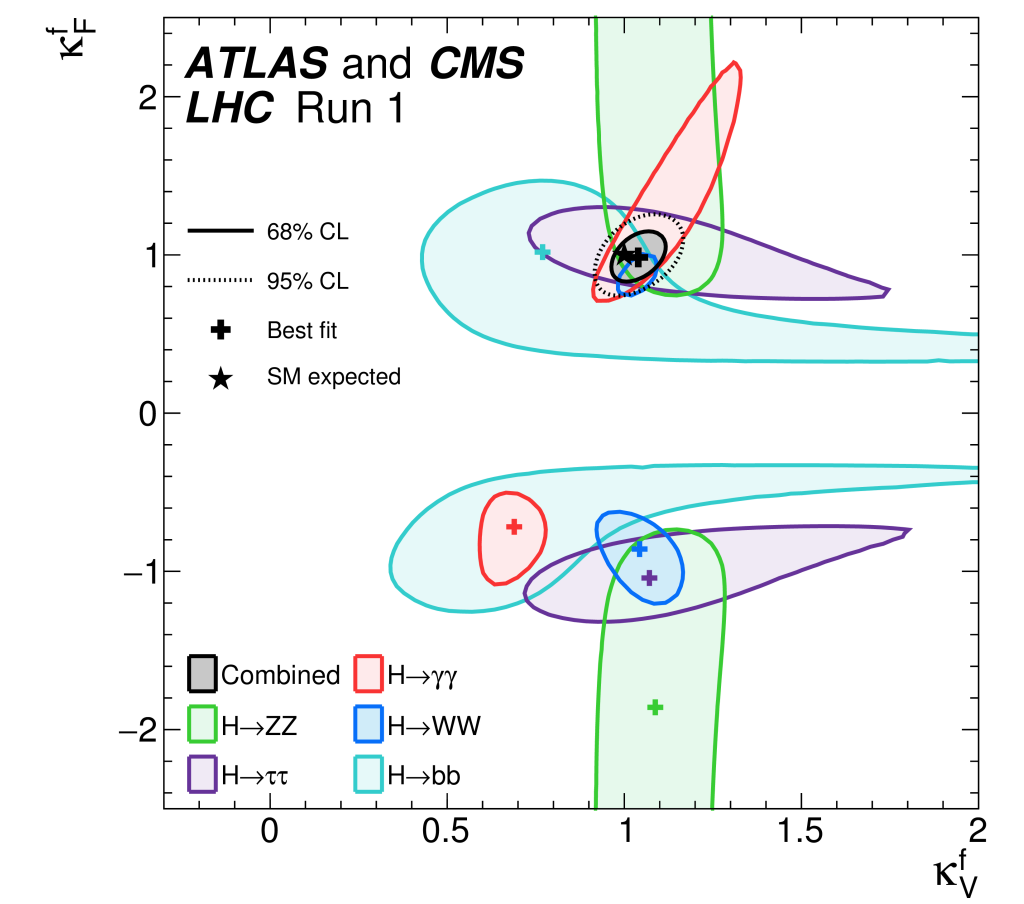
It's infeasible to calculate the integral over this enormous space!

(More subtle: We cannot sample from  $p(z|x, \theta)$  efficiently.)

Inference

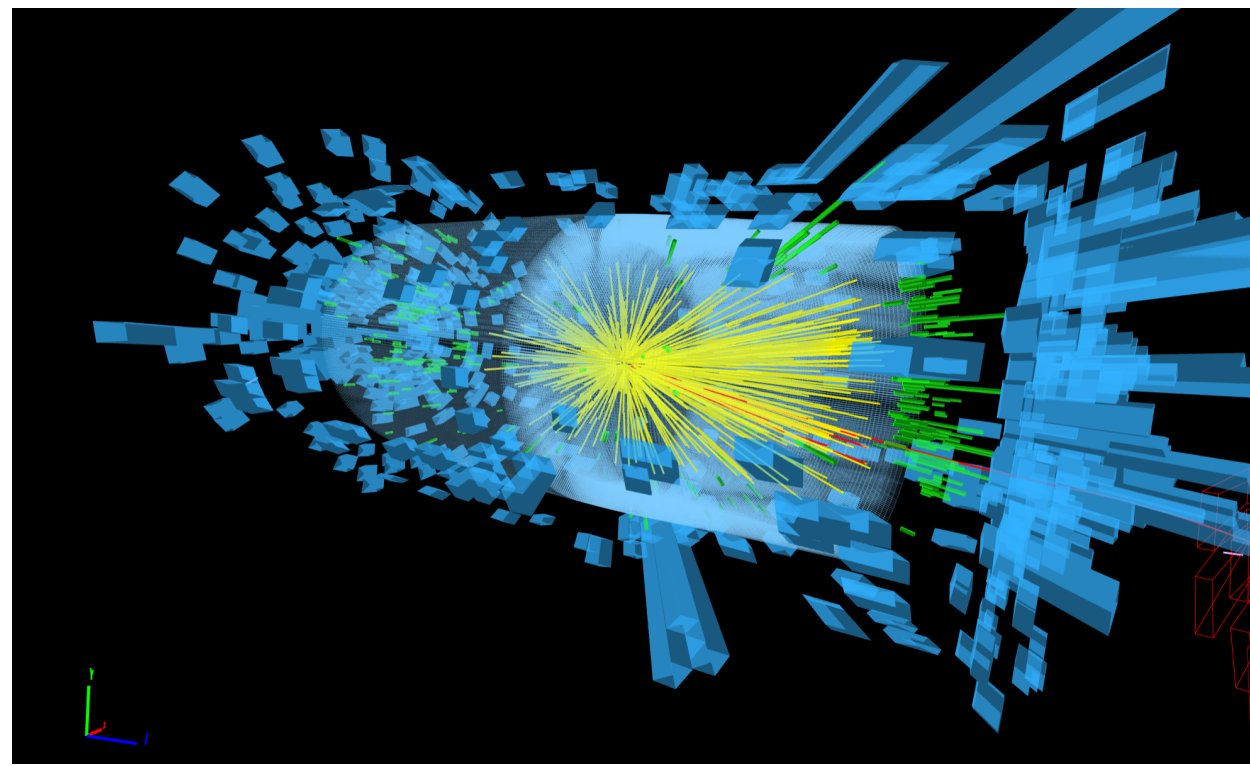


High-dimensional  
event data  $x$



Constraints on  
parameters  $\theta$





High-dimensional event data  $x$

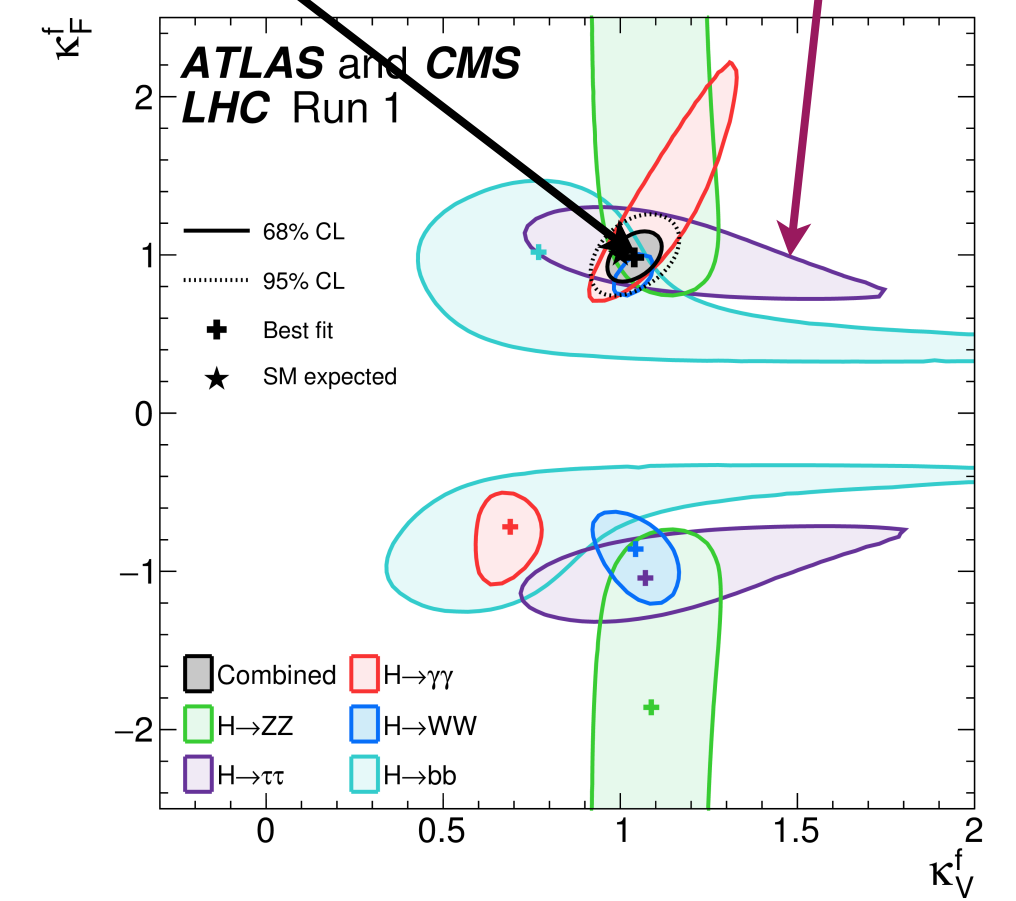


Likelihood function  
 $p(x|\theta)$

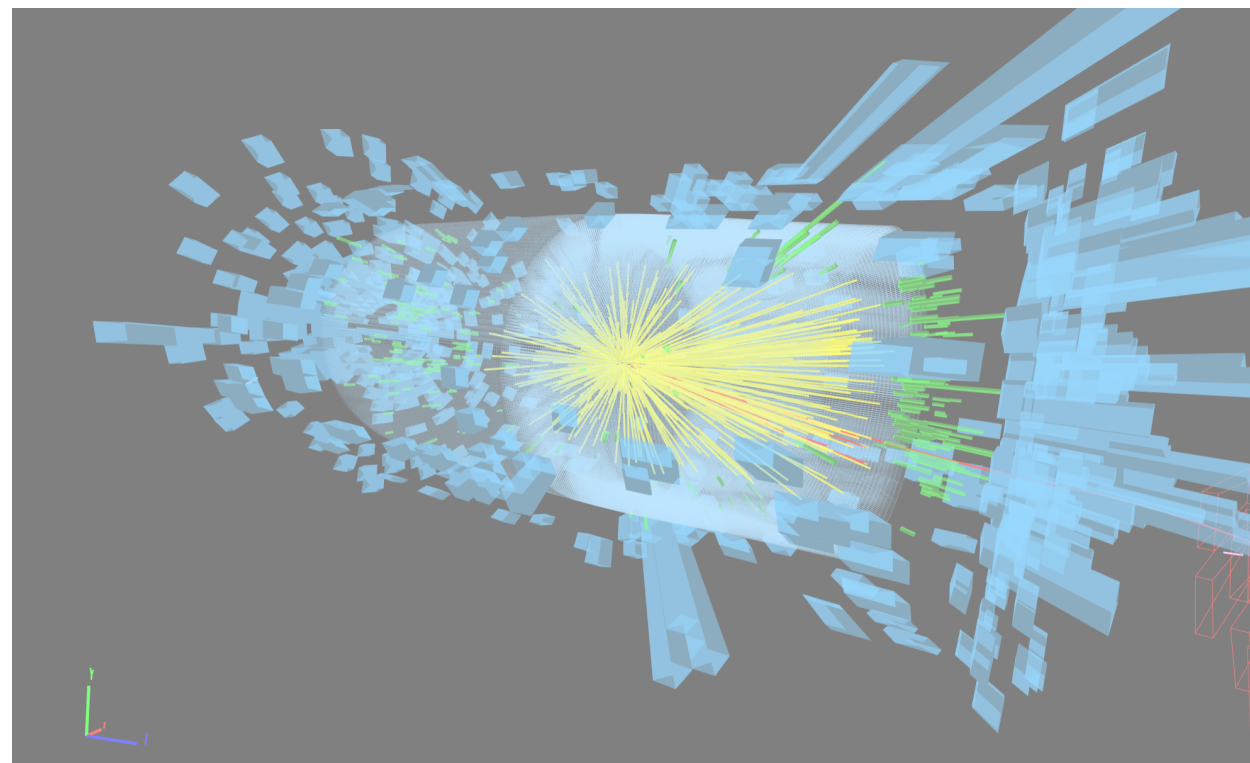


Maximum-likelihood estimator

Confidence limits based on likelihood ratio tests



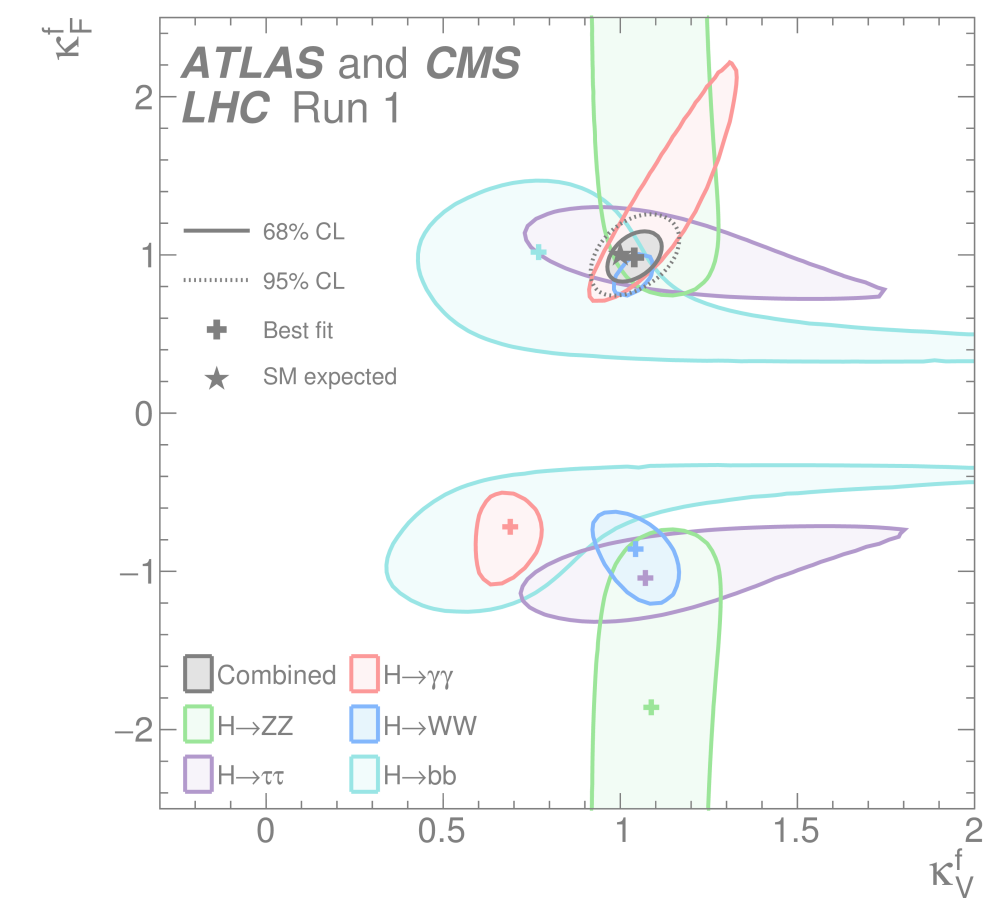
Constraints on parameters  $\theta$



High-dimensional event data  $x$



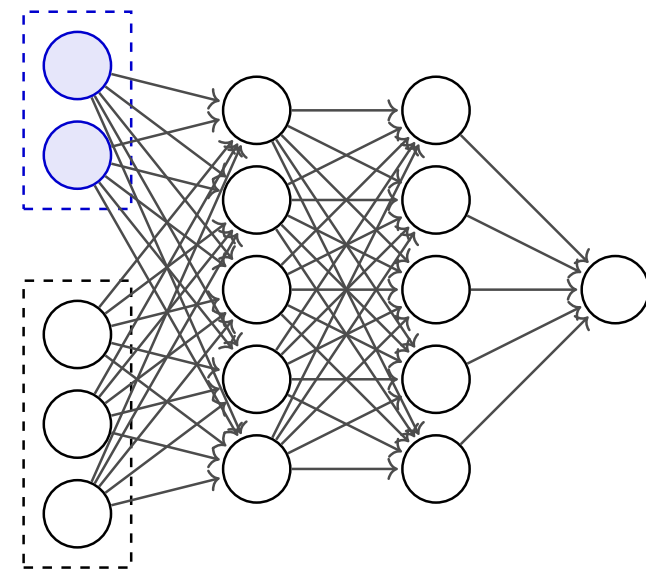
Likelihood function  
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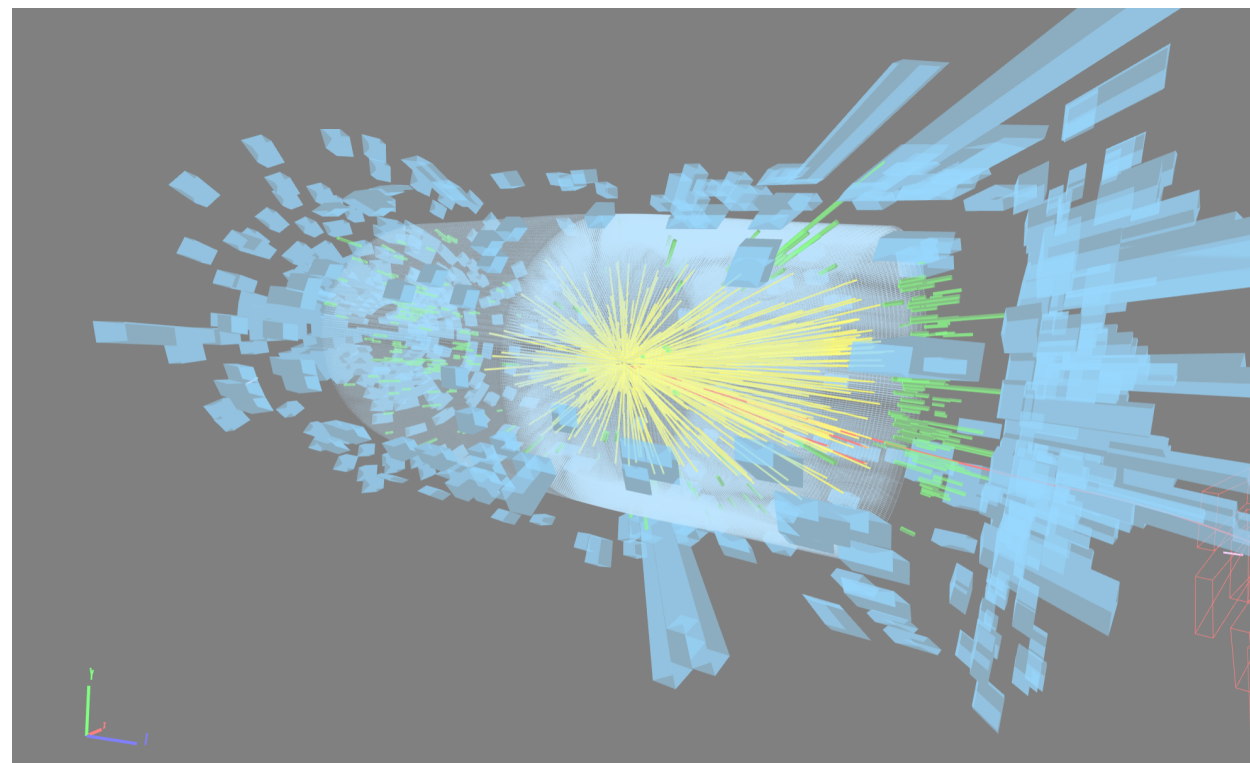
Constraints on parameters  $\theta$

Surprisingly, when we want to use high-dimensional data and have to deal with the detector response, we do not have a good way to calculate the likelihood.

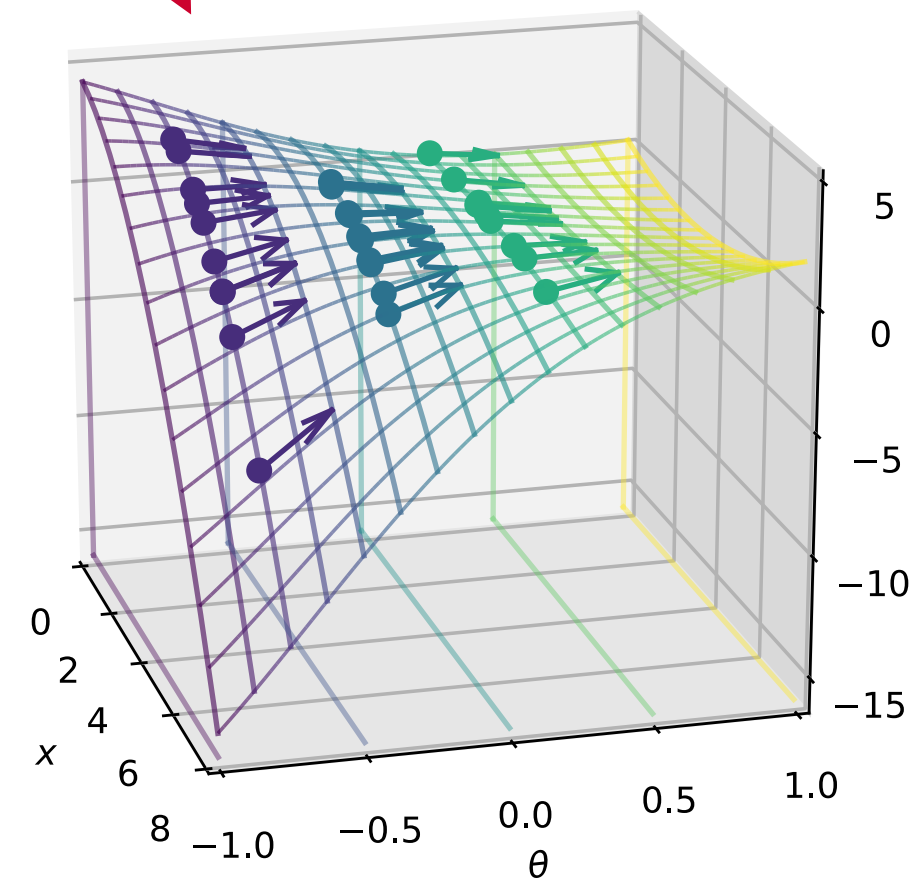




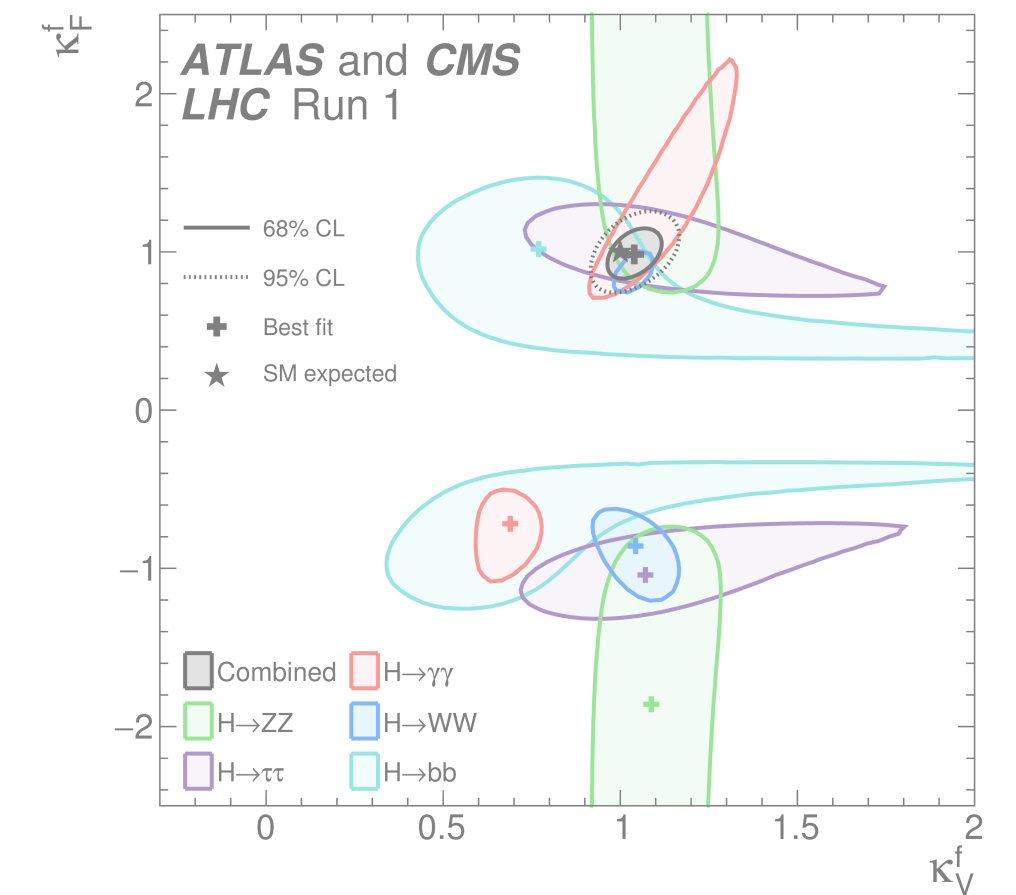
Machine learning



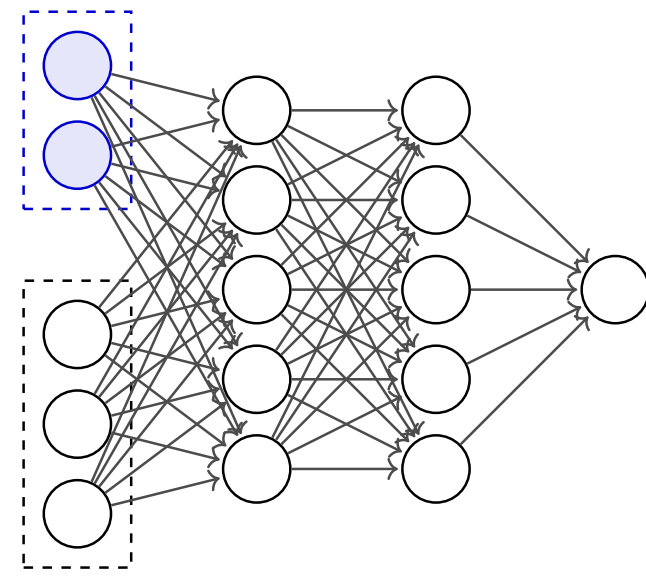
High-dimensional event data  $x$



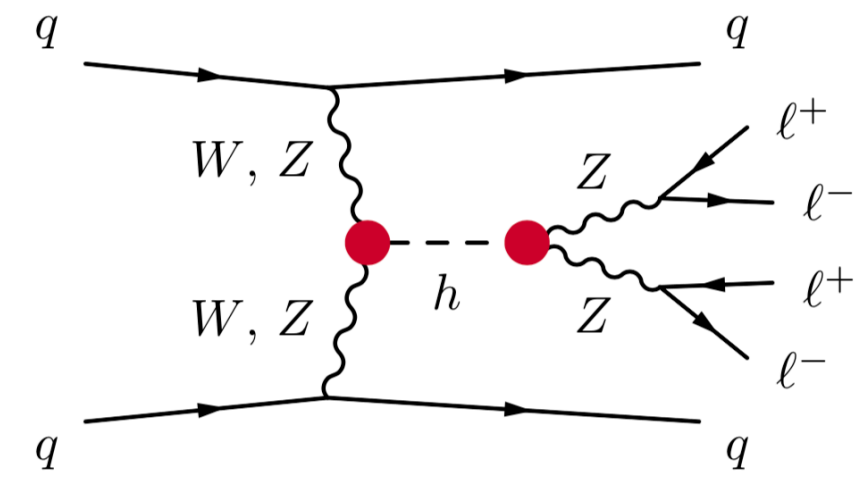
Estimator of the likelihood  $p(x|\theta)$



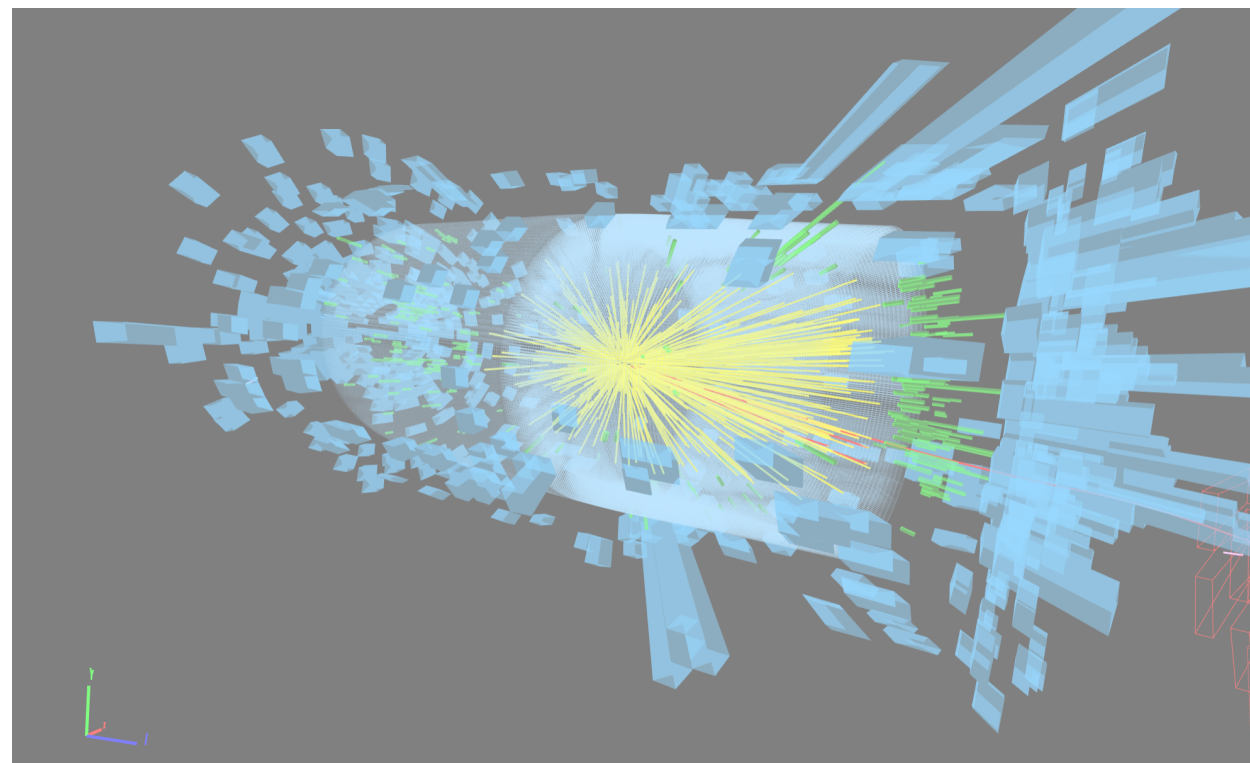
Constraints on parameters  $\theta$



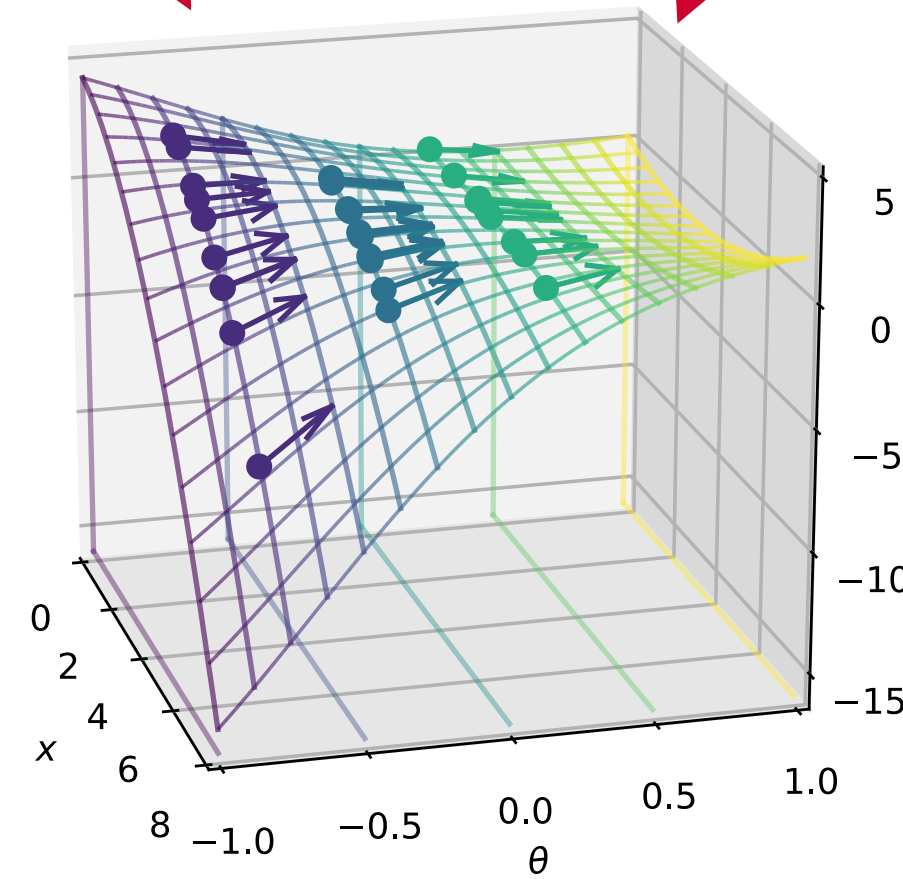
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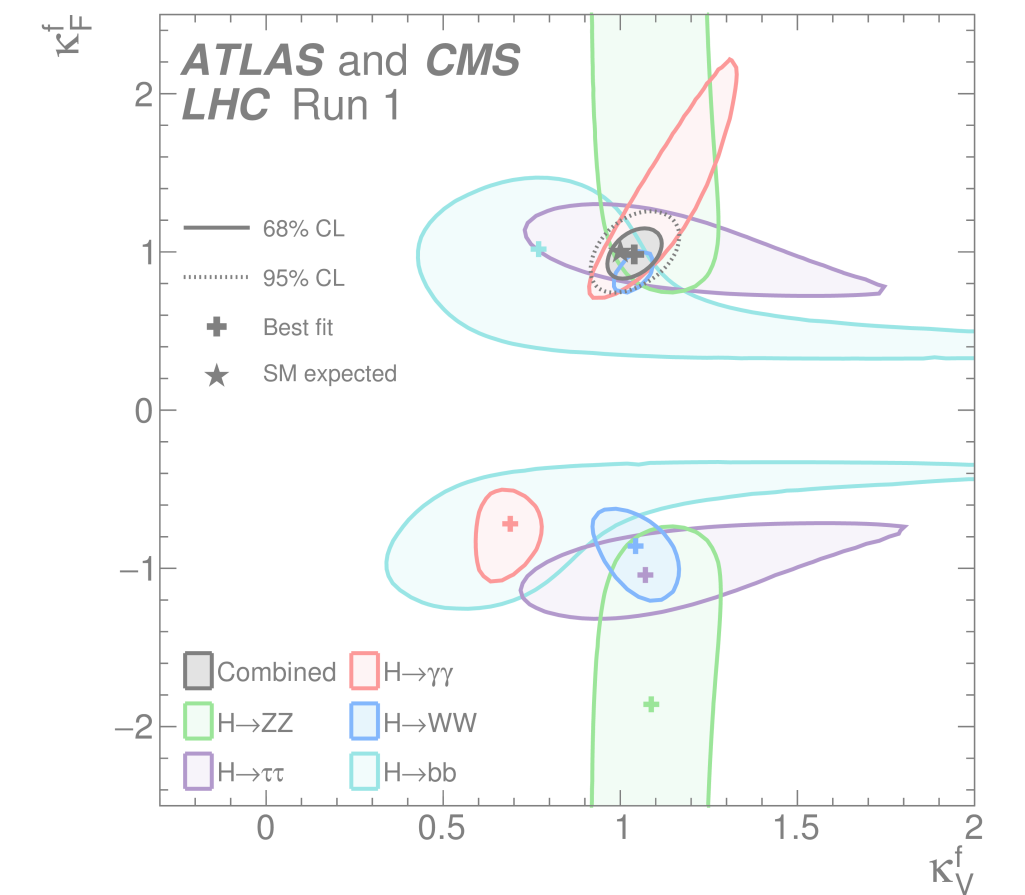
Physics insight:  
matrix element information



High-dimensional event data  $x$



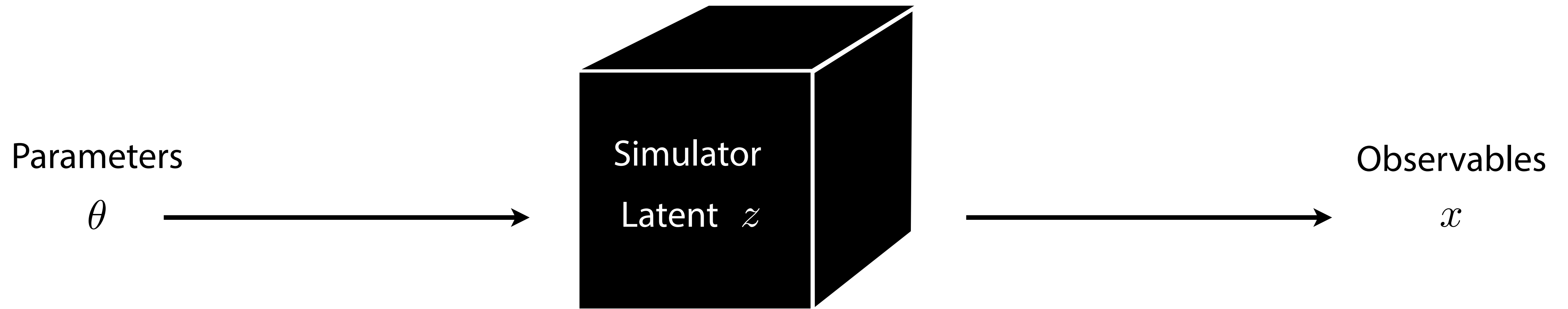
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Constraints on parameters  $\theta$

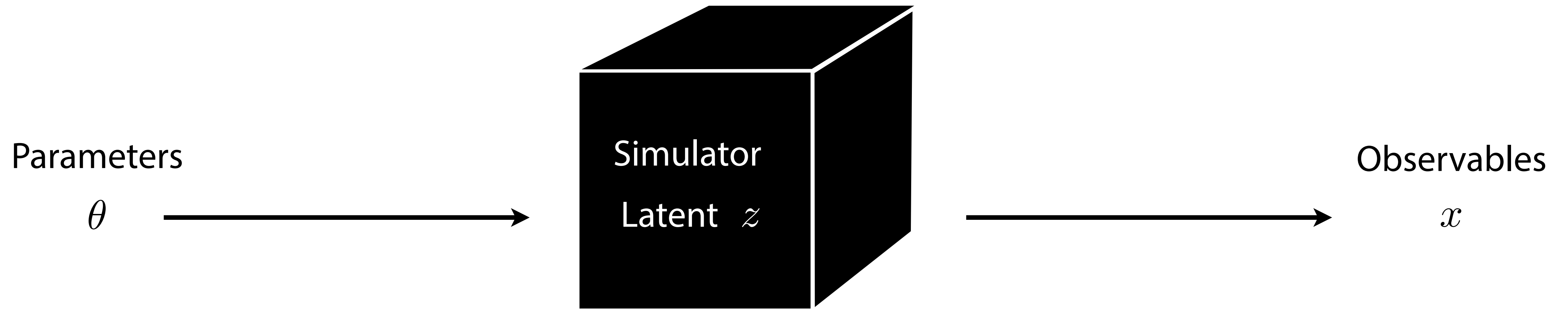


# Simulation-based ("likelihood-free") inference



- Prediction:
- Well-understood mechanistic model
  - Simulator can generate samples  $x \sim p(x|\theta)$

# Simulation-based ("likelihood-free") inference



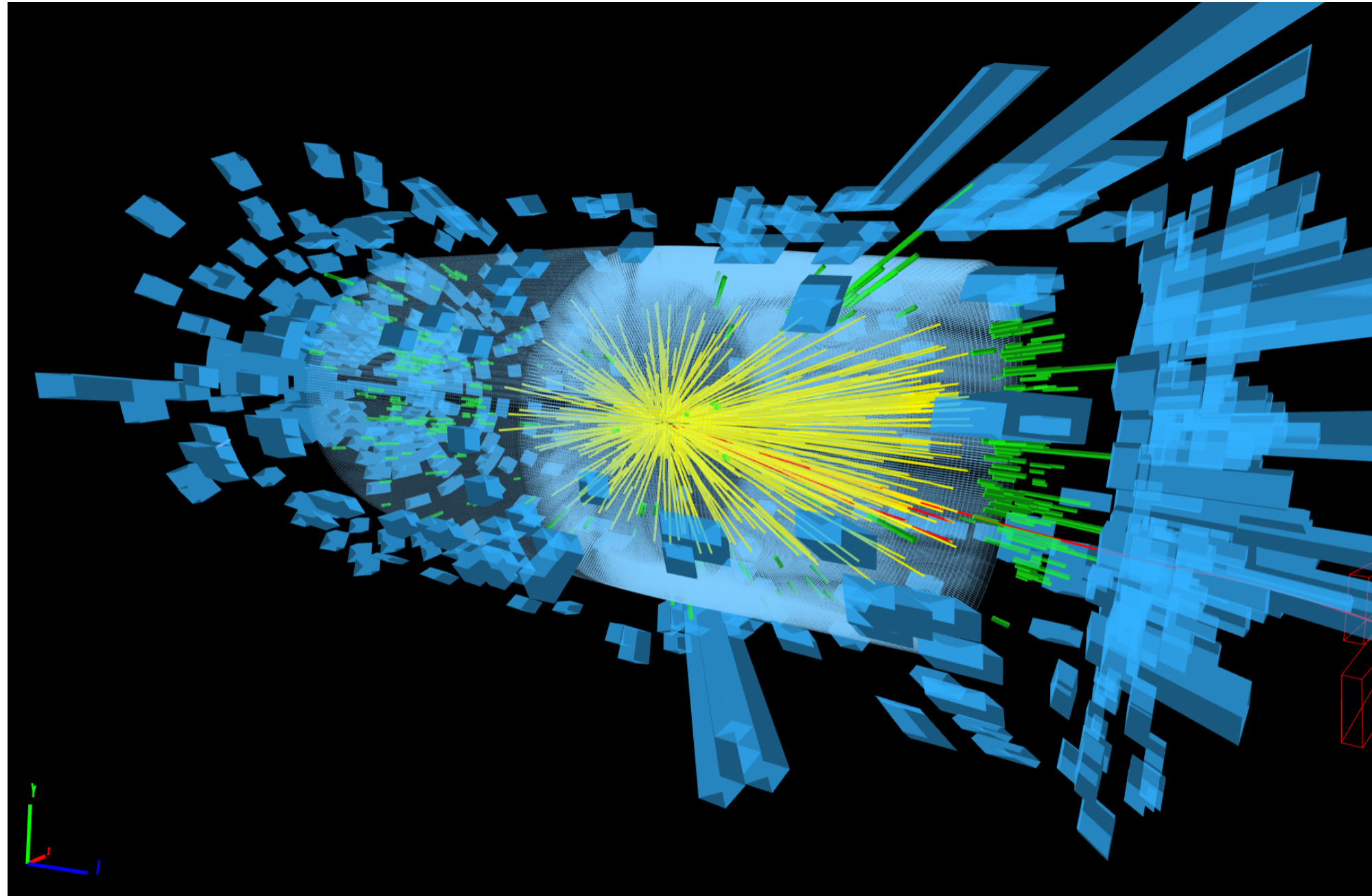
- Prediction:
- Well-understood mechanistic model
  - Simulator can generate samples  $x \sim p(x|\theta)$

- Inference:
- Likelihood  $p(x|\theta) = \int dz p(x, z|\theta)$  is intractable
  - Inference is challenging



Traditionally, inference is made possible with summary statistics.

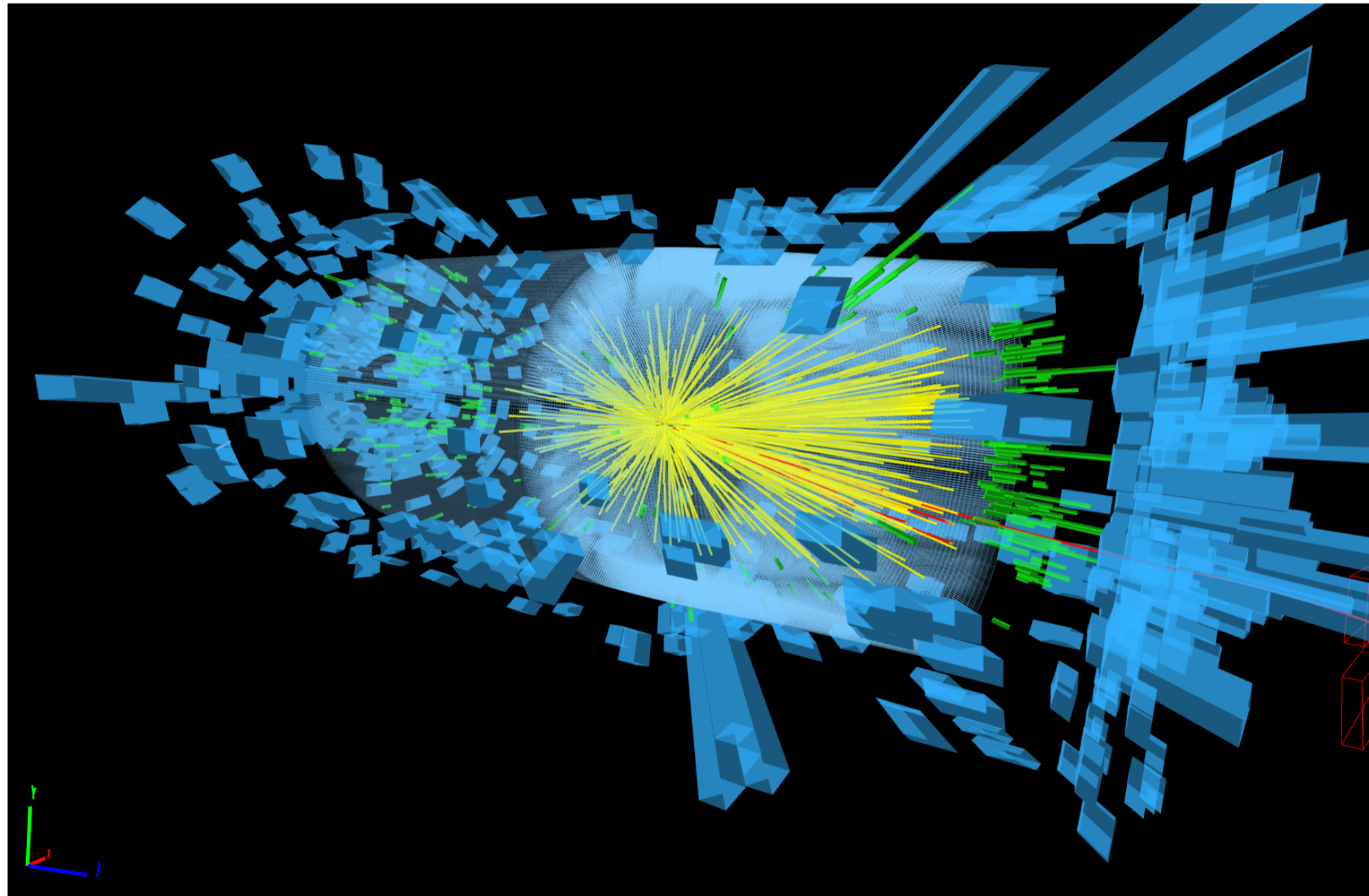
# Solve it with summary statistics



High-dimensional event data  $x$

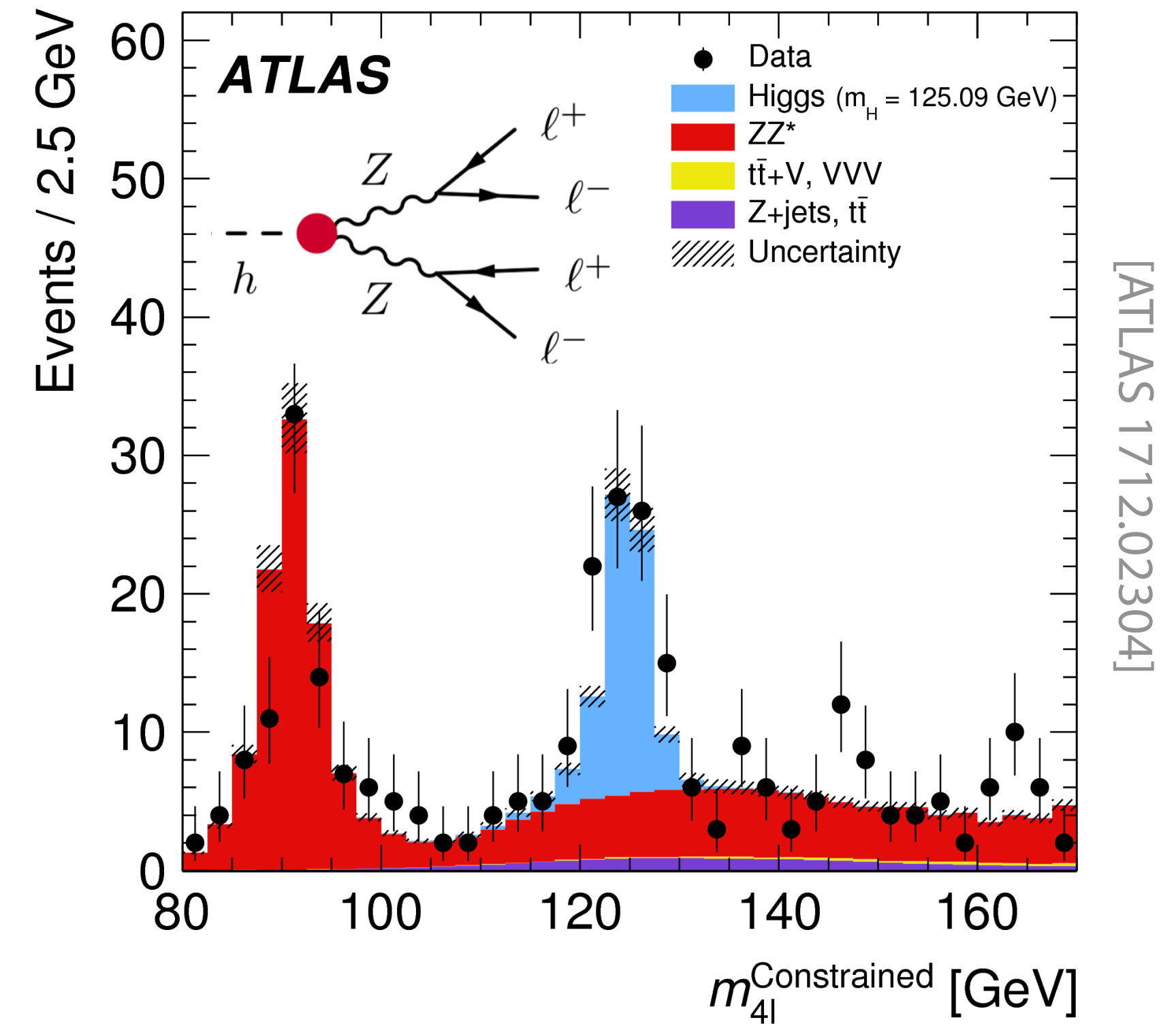
$p(x|\theta)$  cannot be calculated

# Solve it with summary statistics



High-dimensional event data  $x$

$p(x|\theta)$  cannot be calculated



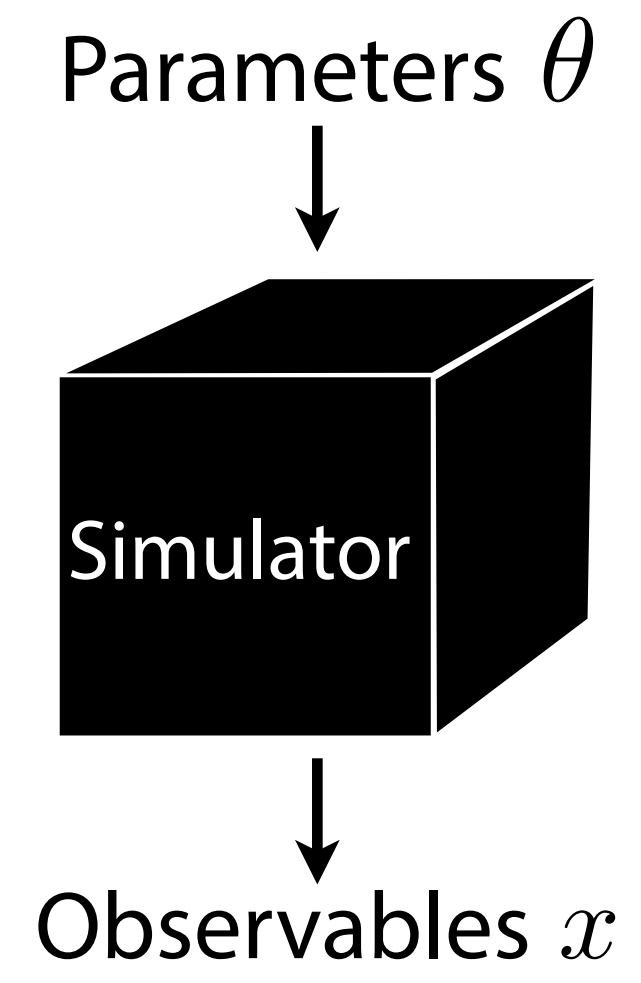
One or two summary statistics  $x'$

$p(x'|\theta)$  can be estimated  
with histograms, KDE, ...



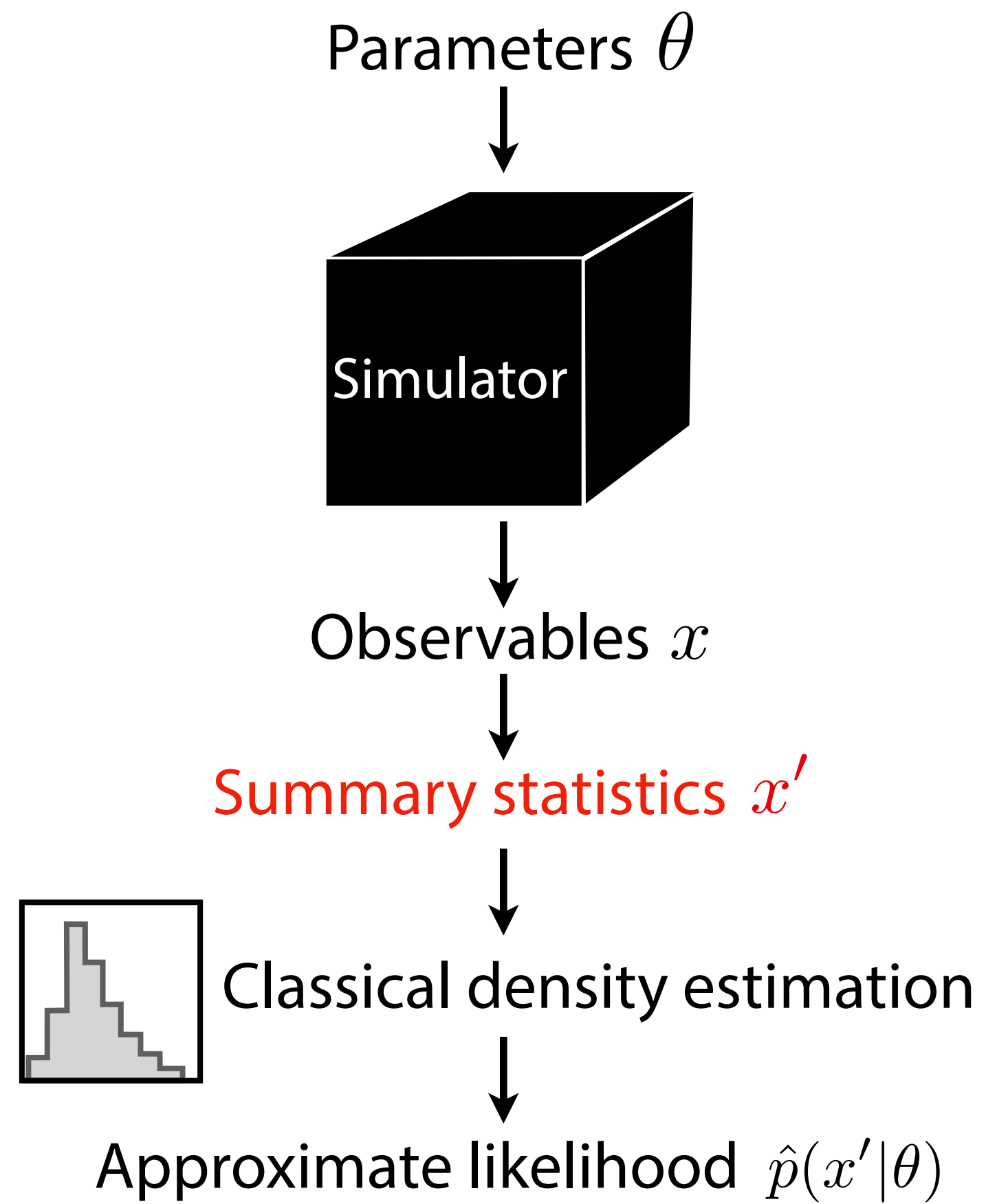
# Inference by estimating the likelihood

[e.g. P. Diggle, R. Gratton 1984]



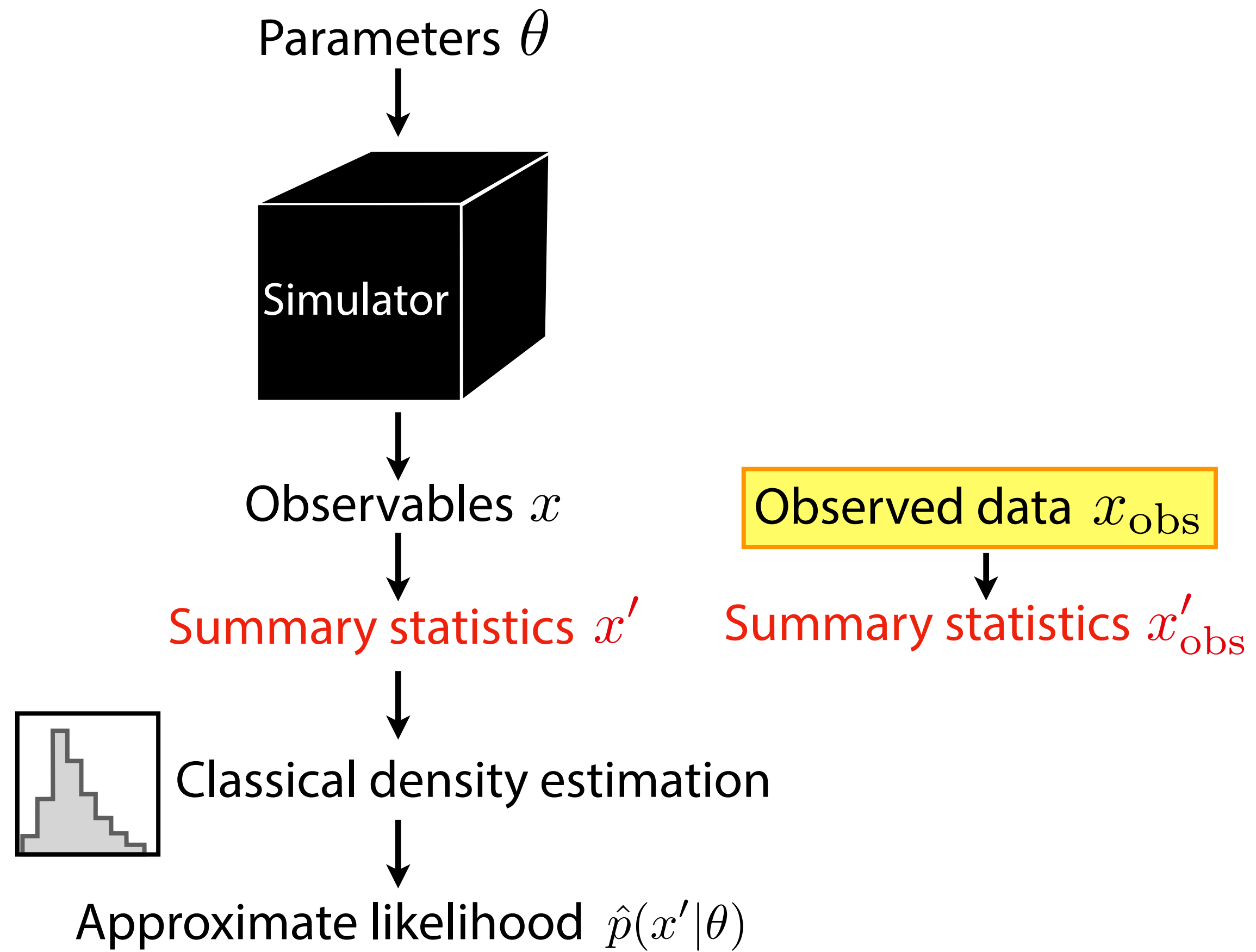
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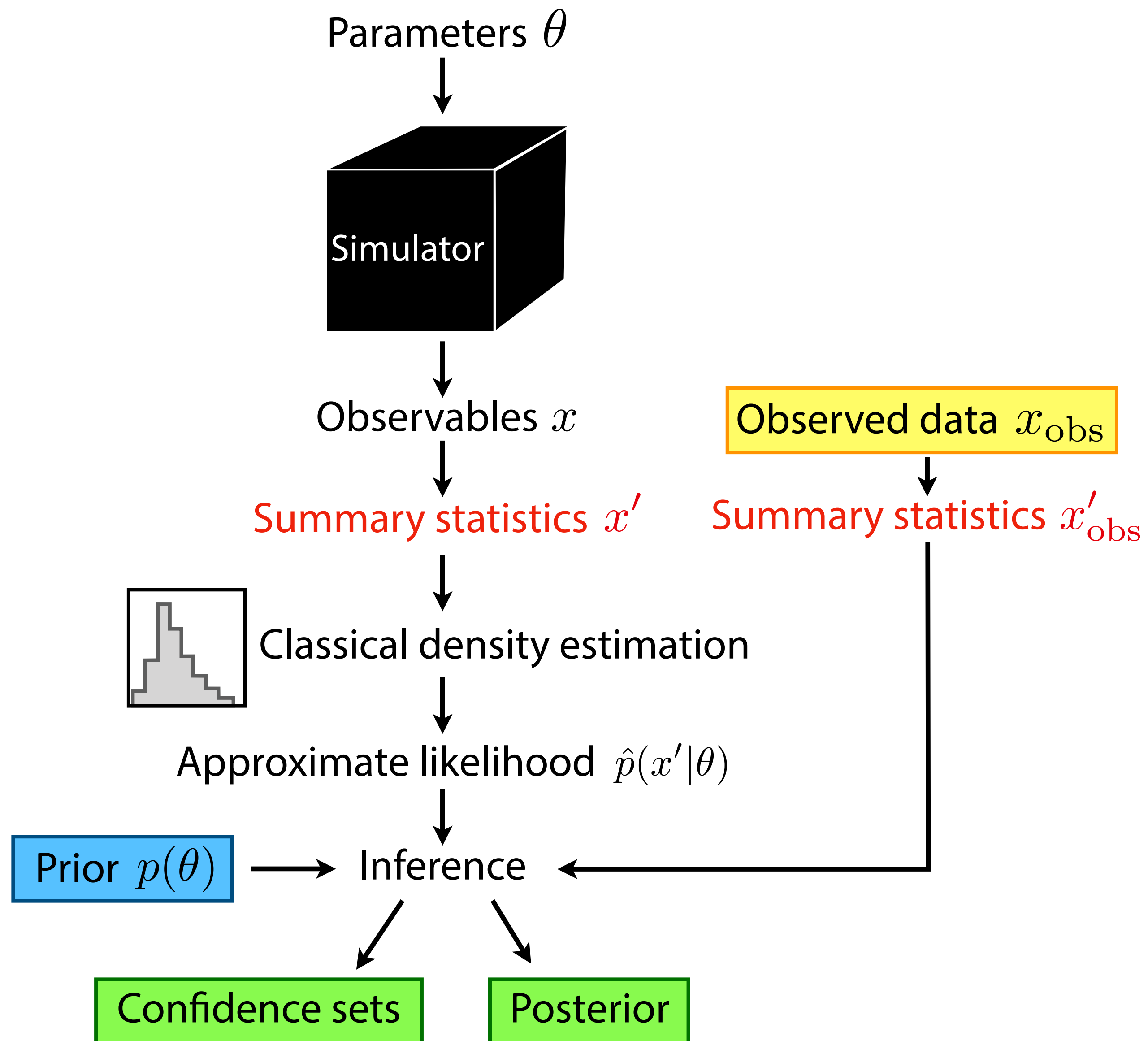
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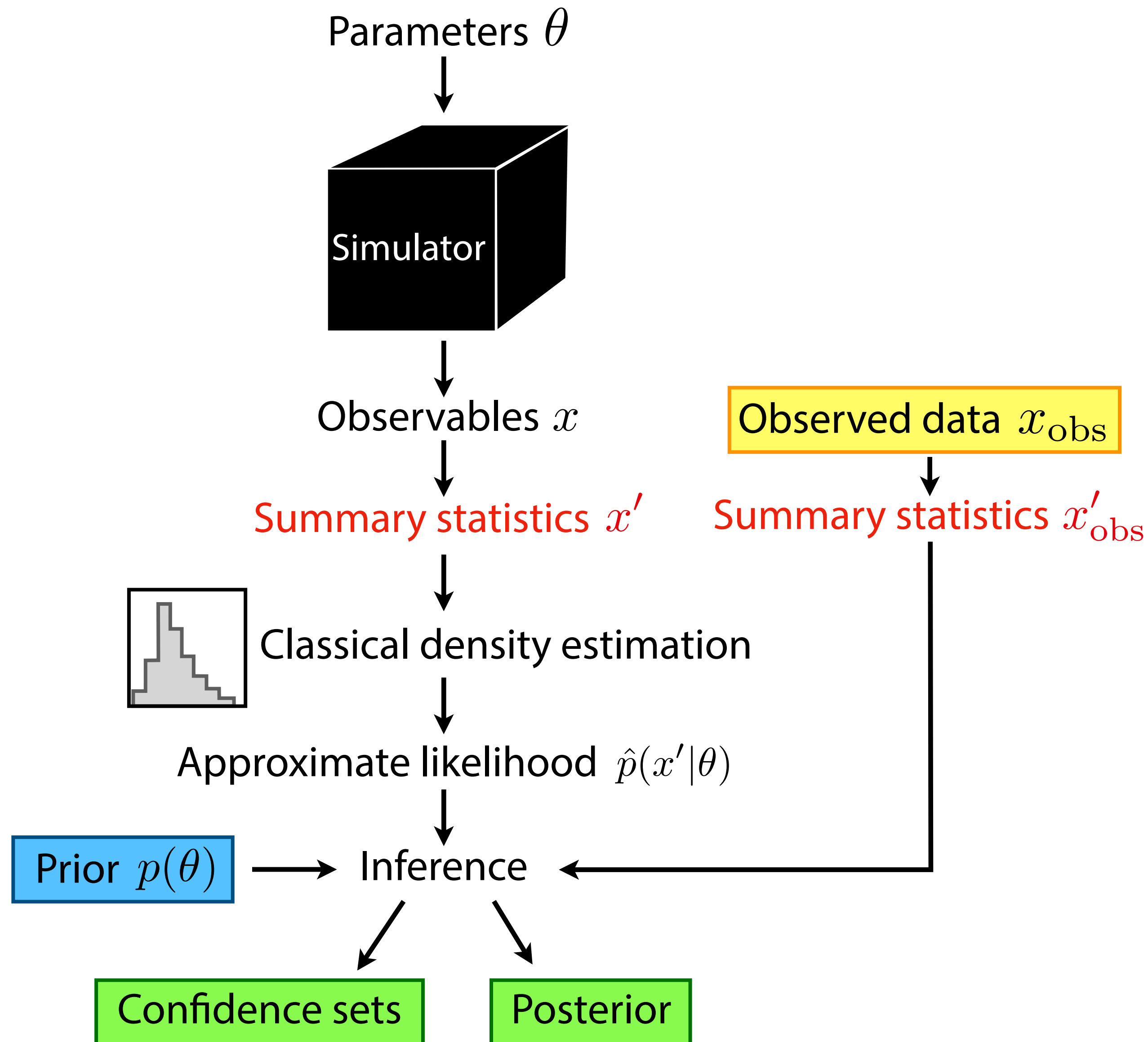
# Inference by estimating the likelihood

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# Inference by estimating the likelihood

[e.g. P. Diggle, R. Gratton 1984]



- Compression to summary statistics loses information & reduces quality of inference
- Curse of dimensionality: does not scale to more than a few summary statistics
- Related alternative: Approximate Bayesian Computation (ABC) [D. Rubin 1984]

# Summary statistics for LHC measurements?

- In many LHC problems there is no single good summary statistic: compressing to any  $x'$  loses information!

[JB, K. Cranmer, F. Kling, T. Plehn 1612.05261;  
JB, F. Kling, T. Plehn, T. Tait 1712.02350]

- Ideally: analyze all trustworthy high-level features (reconstructed four-momenta...), or some form of low-level features, including correlations

(“fully differential cross section”)

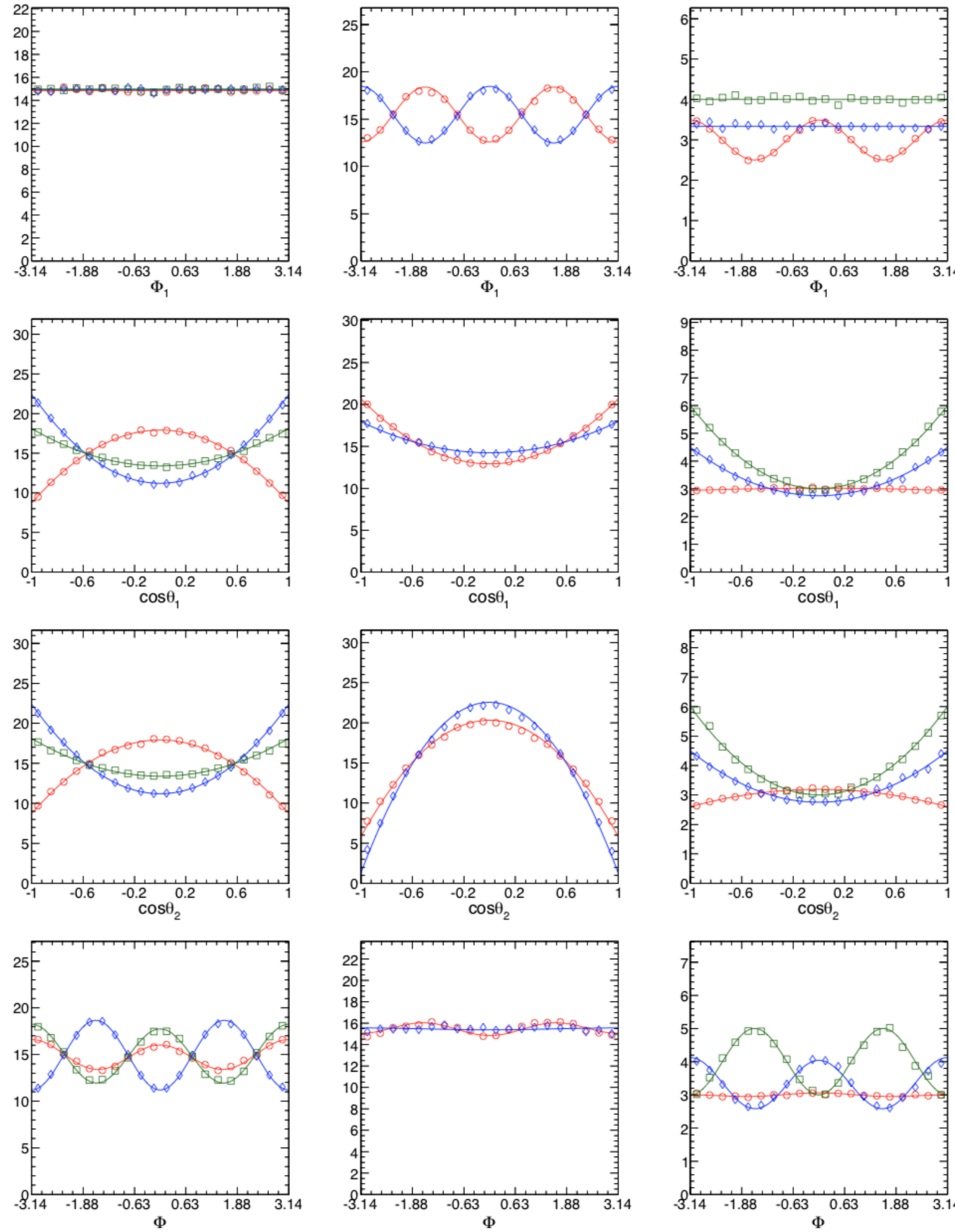
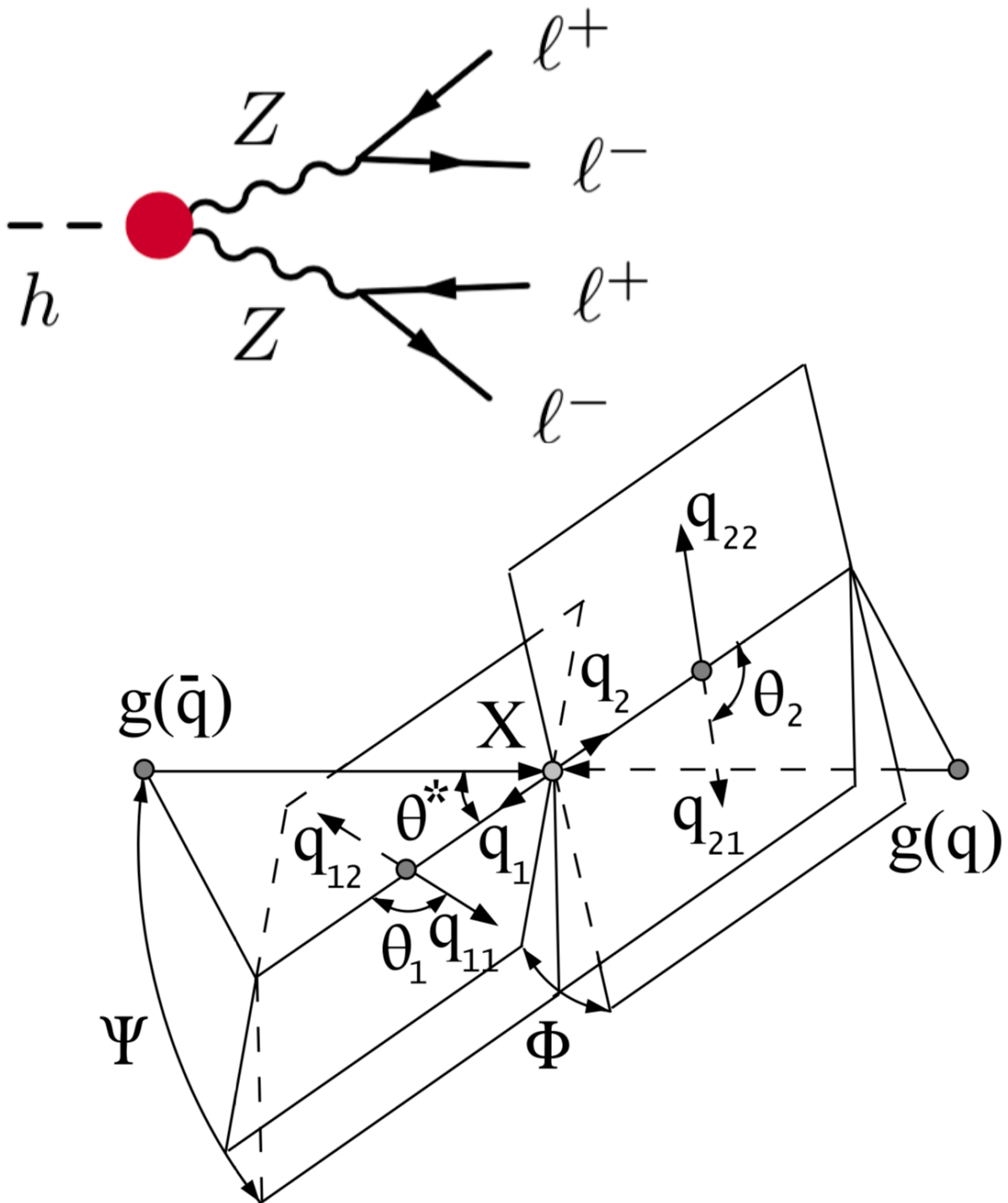


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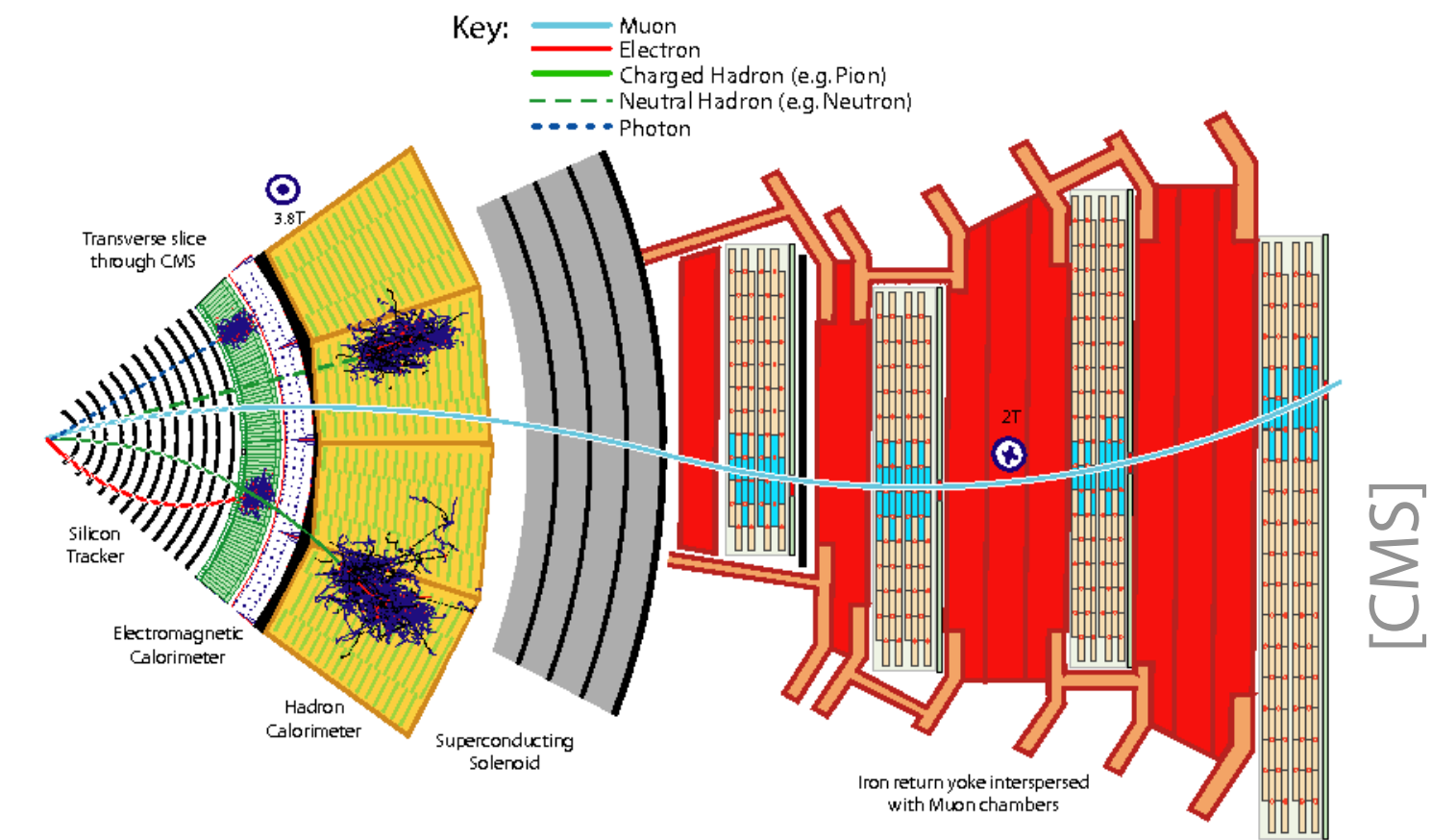


[Bolognesi et al. 1208.4018]

# Solve it by approximating the integral

- Problem: high-dimensional integral over **shower / detector trajectories**

$$p(x|\theta) = \int dz_d \int dz_s \int dz_p p(x|z_d) p(z_d|z_s) p(z_s|z_p) p(z_p|\theta)$$



# Solve it by approximating the integral

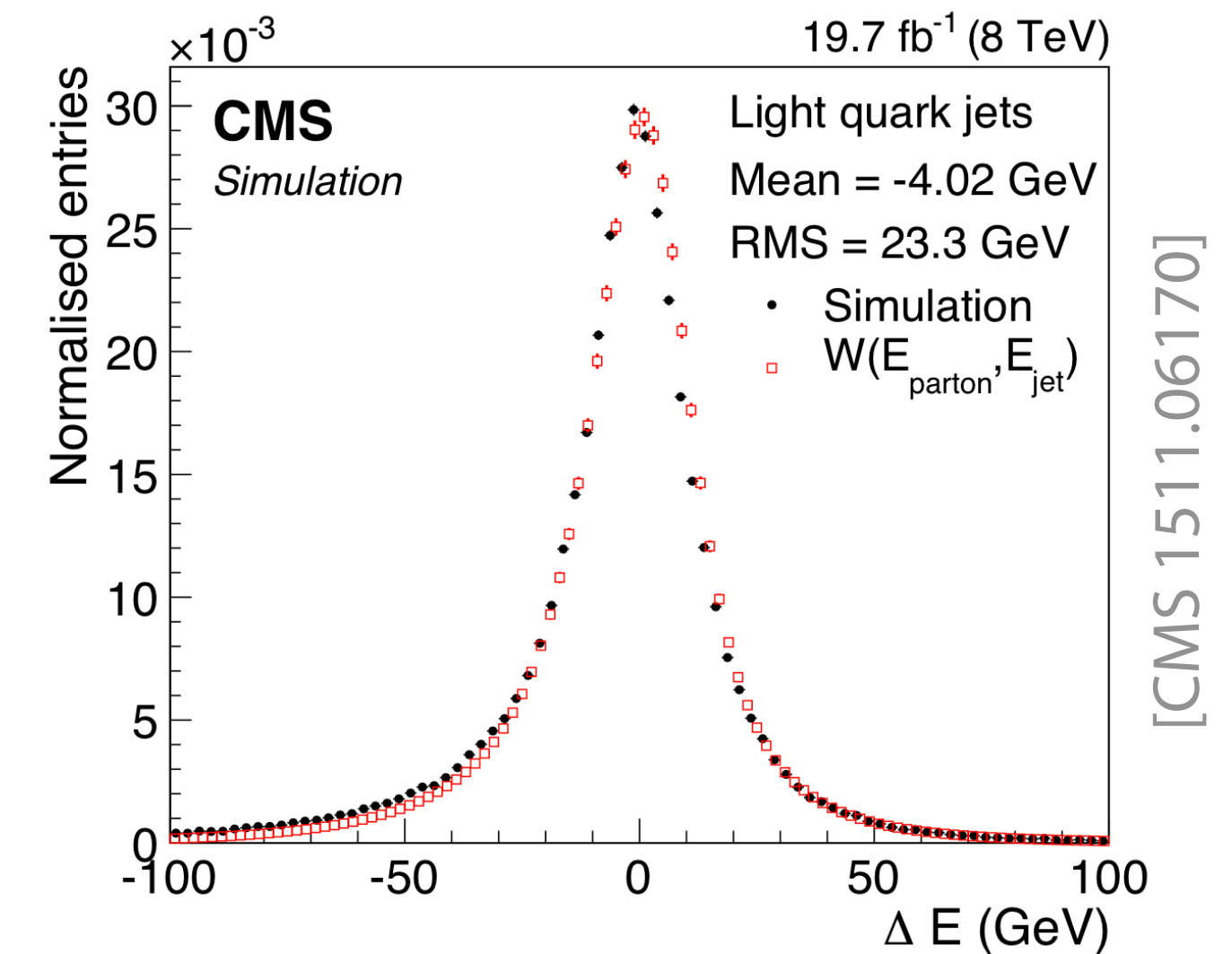
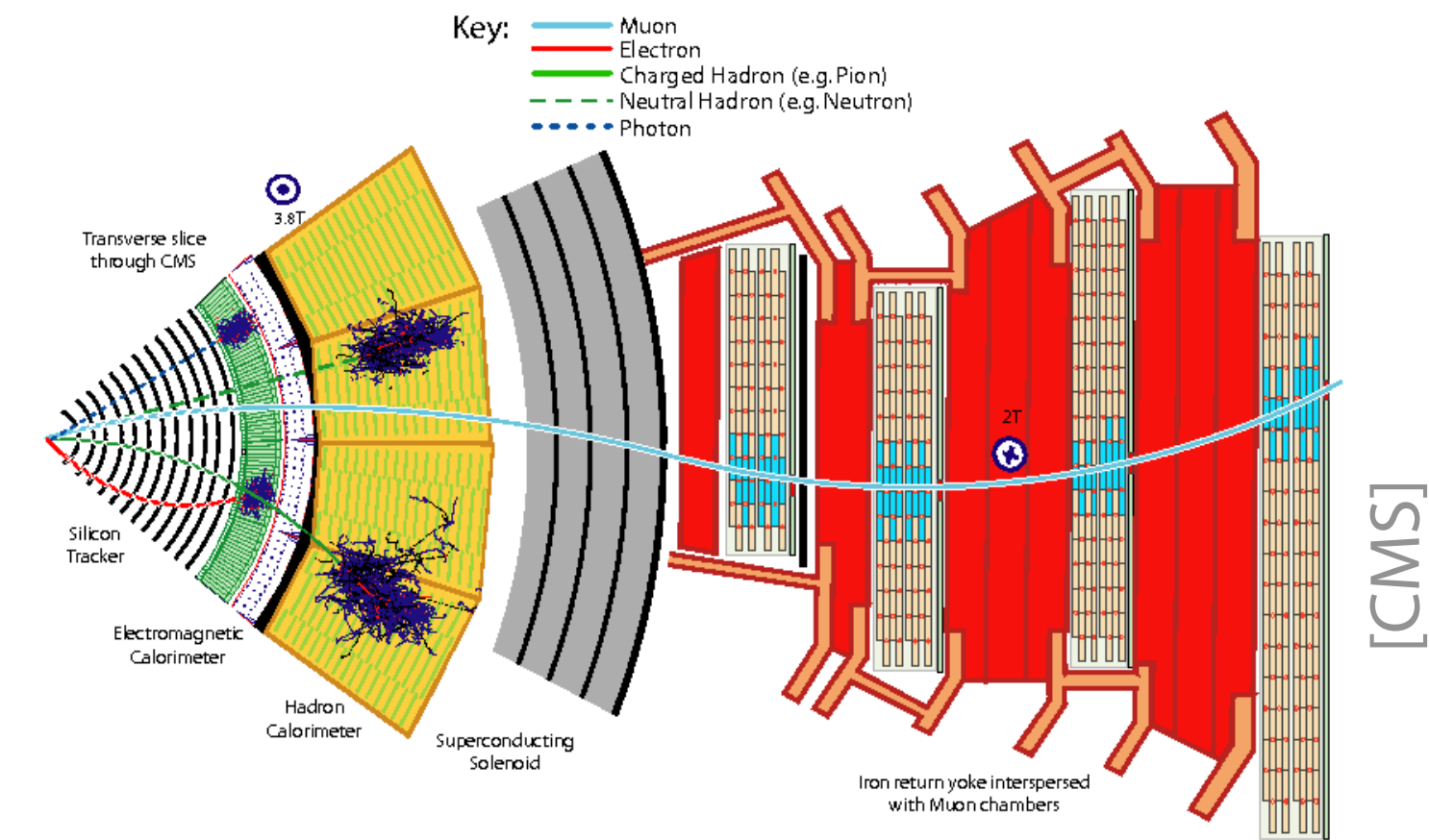
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- Matrix Element Method (and similarly Optimal Observables): [K. Kondo 1988]

- approximate **shower + detector effects** into **transfer function**  $\hat{p}(x|z_p)$
- explicitly calculate remaining integral

$$\hat{p}(x|\theta) = \int dz_p \hat{p}(x|z_p) p(z_p|\theta)$$





# Solve it by approximating the integral

- Problem: high-dimensional integral over **shower / detector trajectories**

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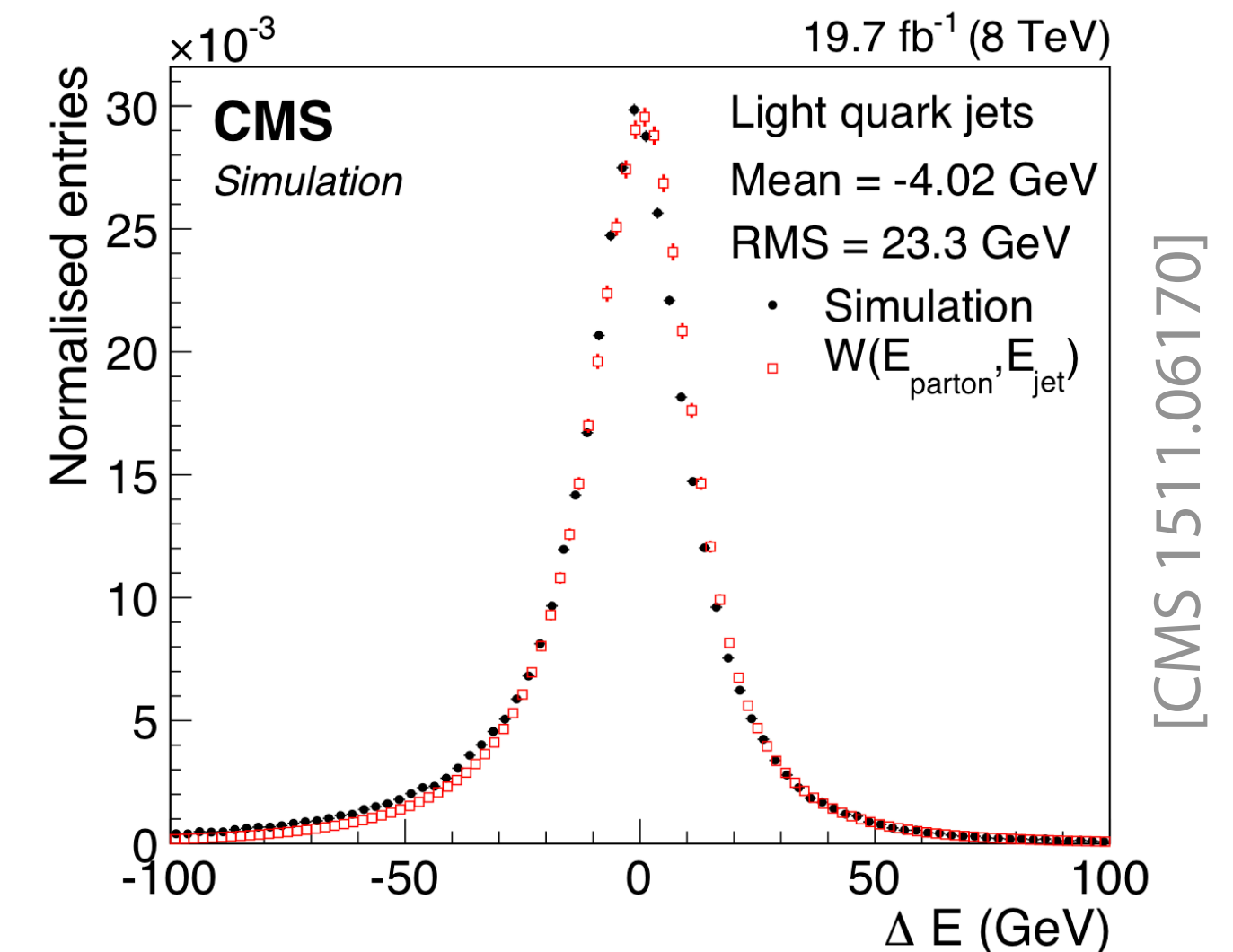
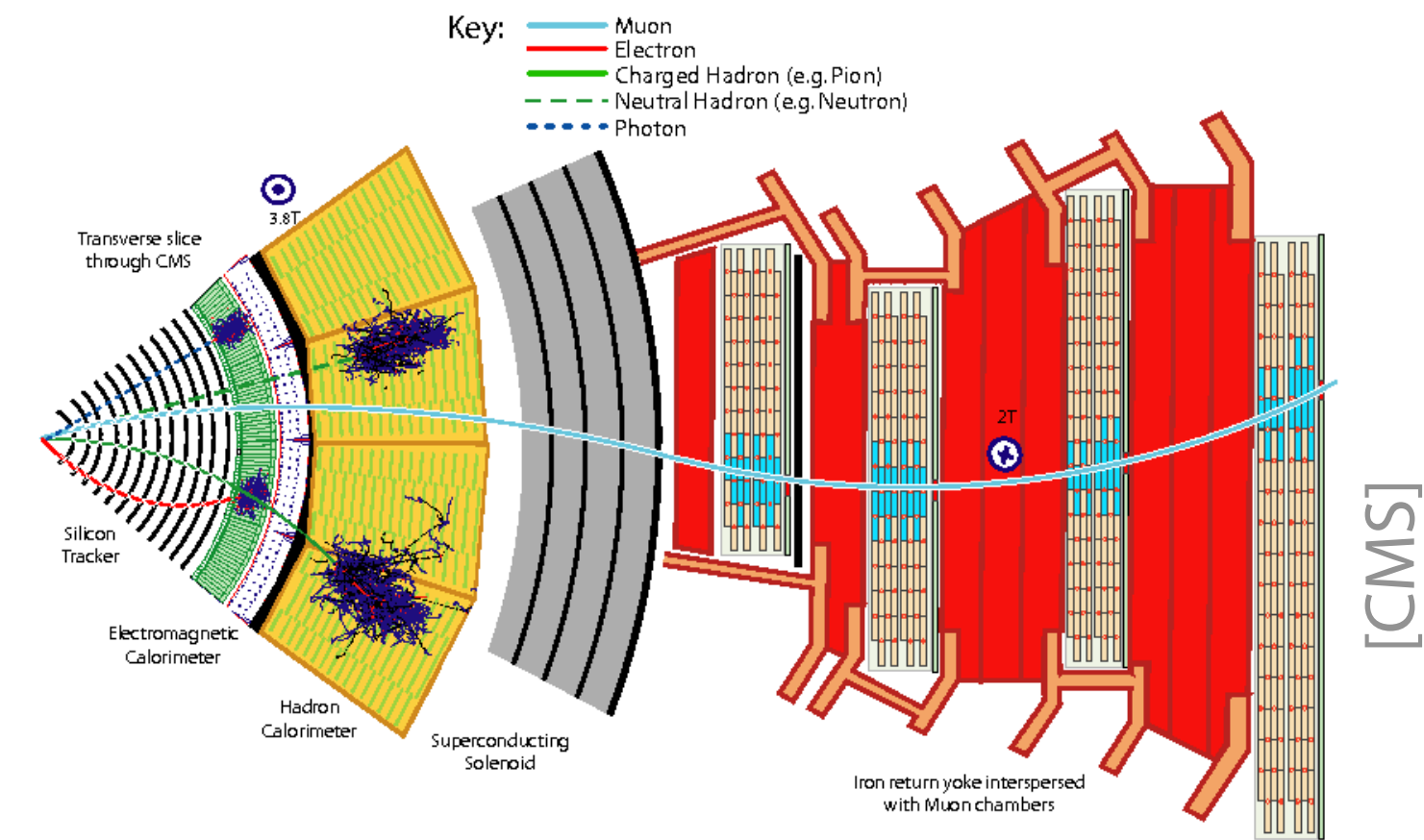
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⇒ Uses matrix-element information, no summary statistics necessary, but:

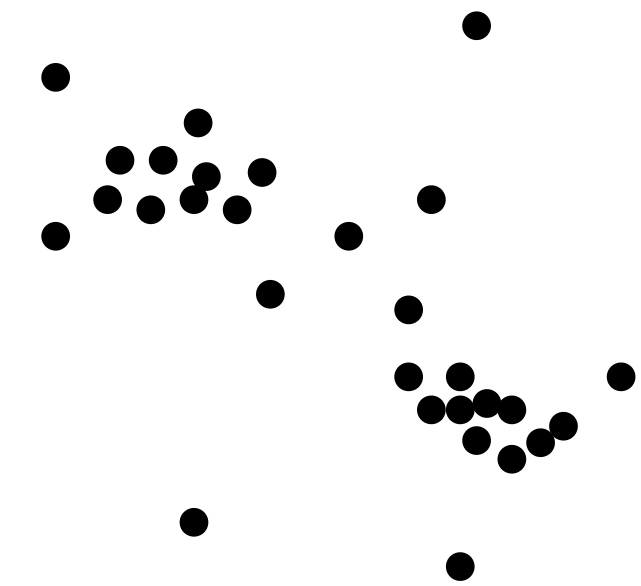
- ad-hoc transfer functions (what about extra radiation?)
- evaluation still requires calculating an expensive integral



But we can use machine learning  
to make inference possible.

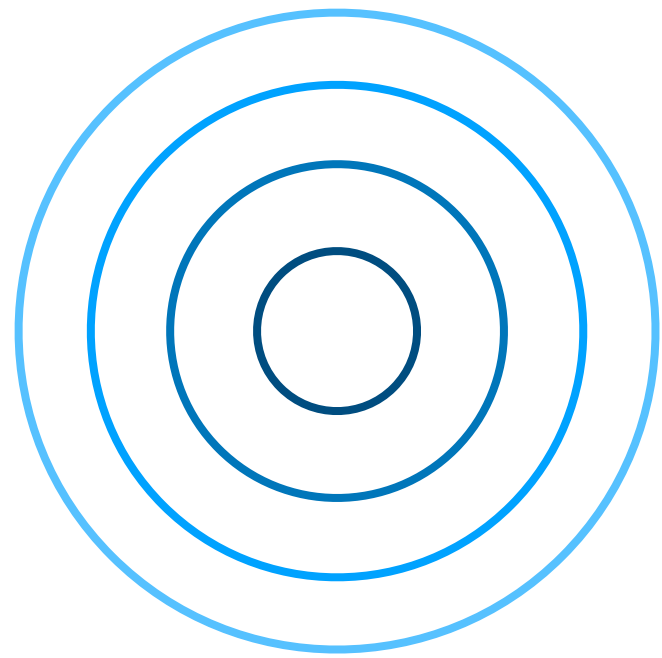
[K. Cranmer, JB, G. Louppe 1911.01429]

# High-dimensional density estimation with normalizing flows



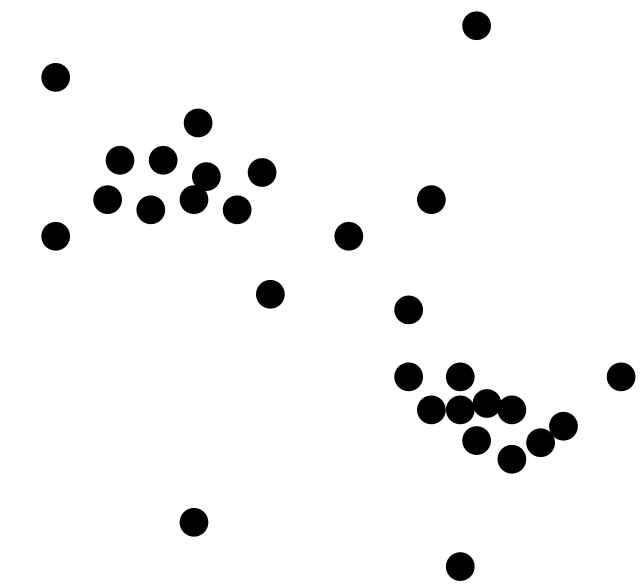


# High-dimensional density estimation with normalizing flows

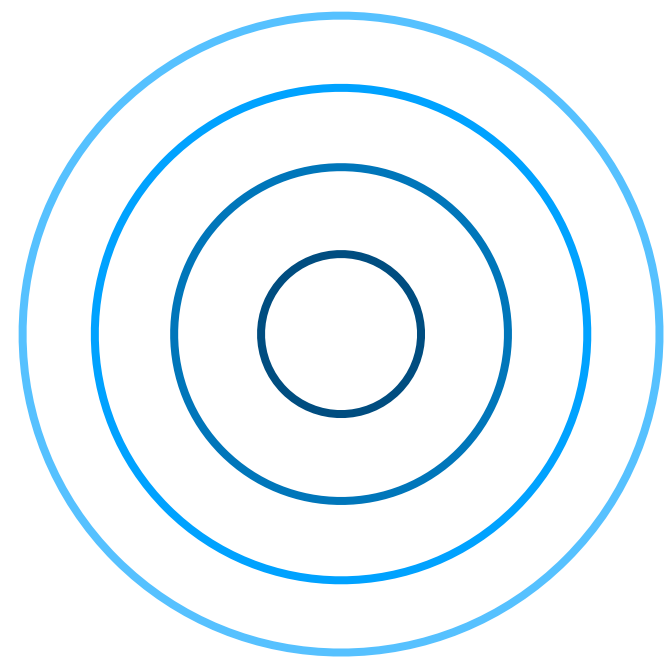


Simple base density

$$u \sim \pi(u)$$

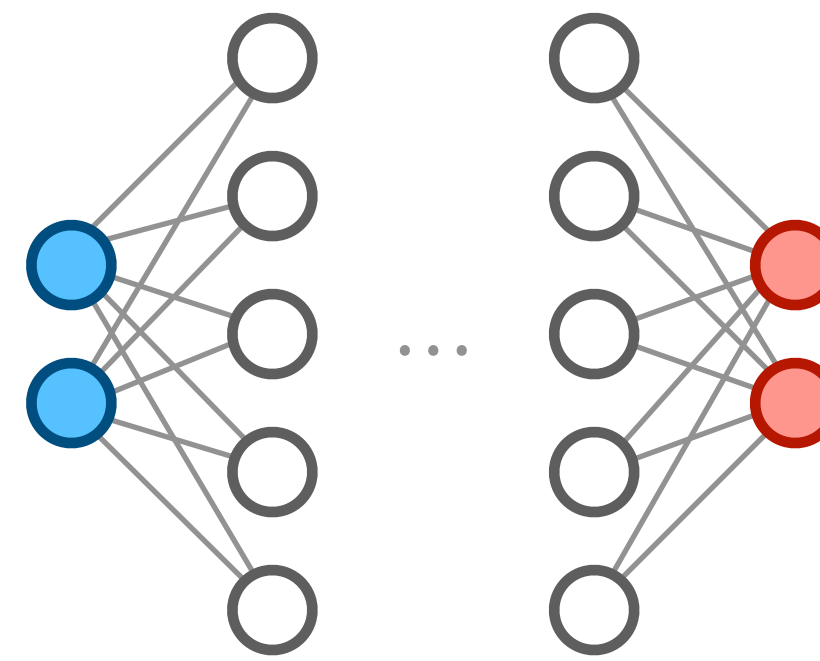


# High-dimensional density estimation with normalizing flows



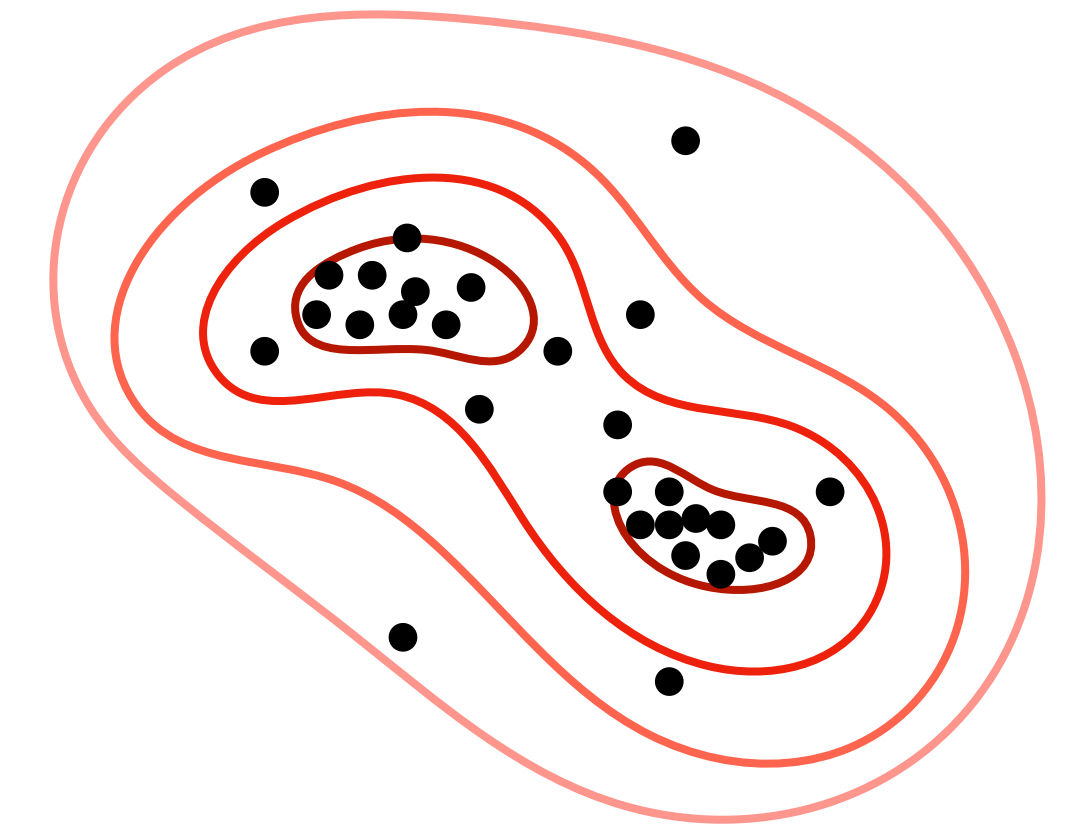
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NN: transformation  $x = f(u)$

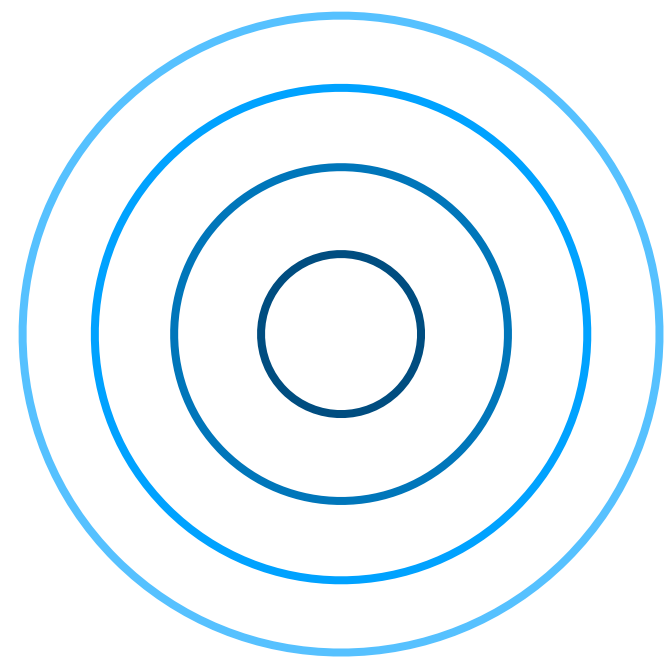
- one-to-one and invertible
- differentiable
- $f^{-1}$  and  $\det \nabla f$  are tractable



Target density is given by

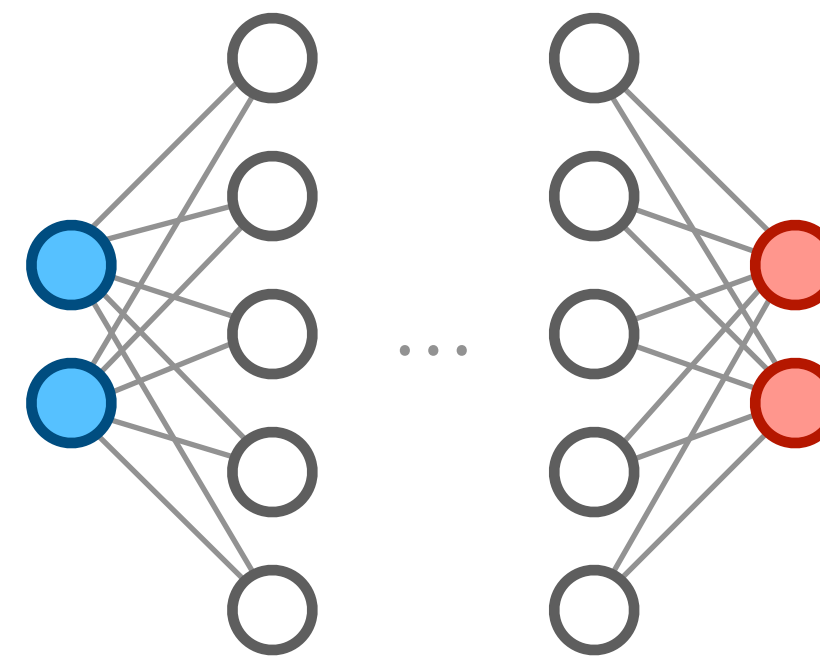
$$\hat{p}(x) = \pi(f^{-1}(x)) |\det \nabla f|^{-1}$$

# High-dimensional density estimation with normalizing flows



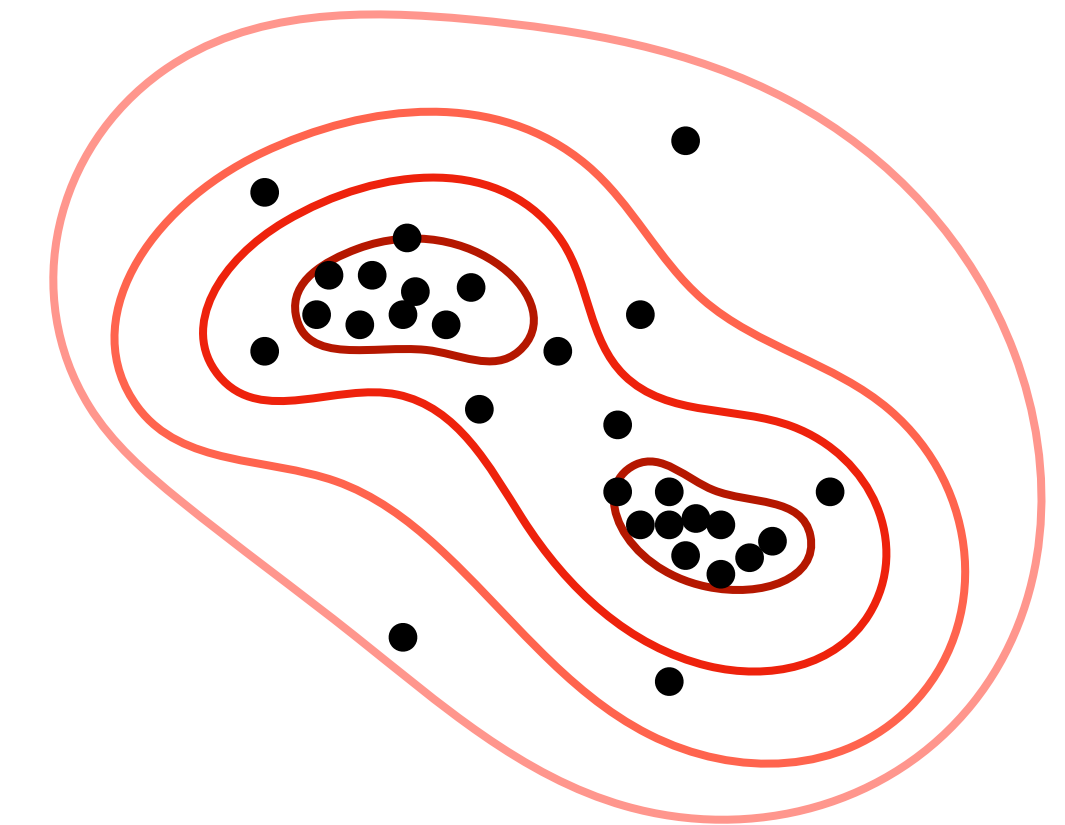
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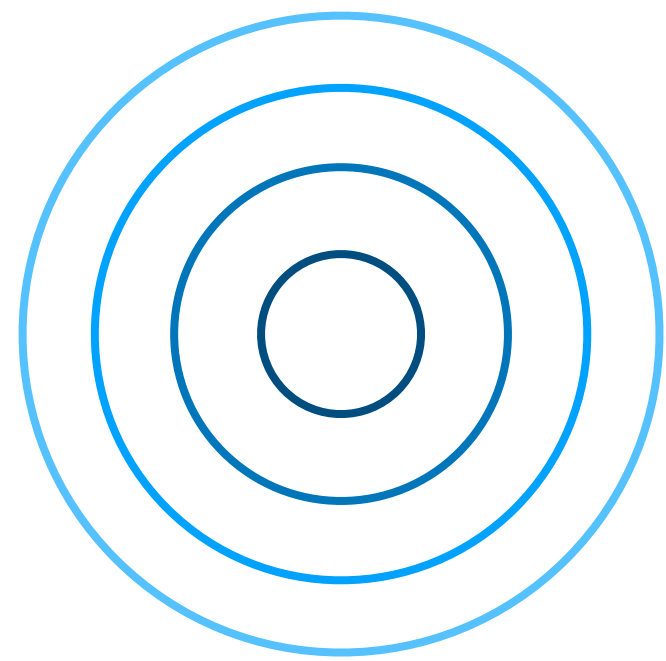
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Train transformation by  
maximizing  $\log \hat{p}(x)$

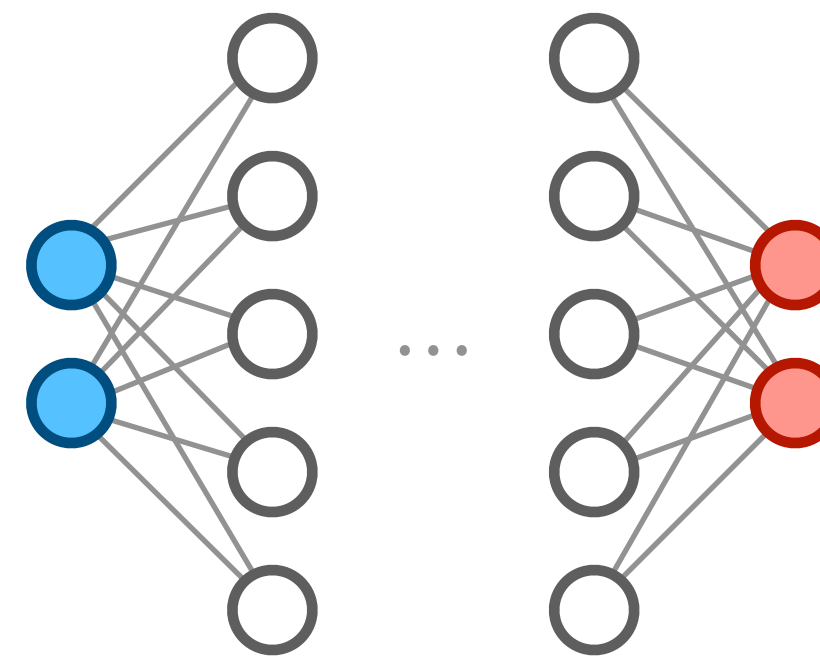


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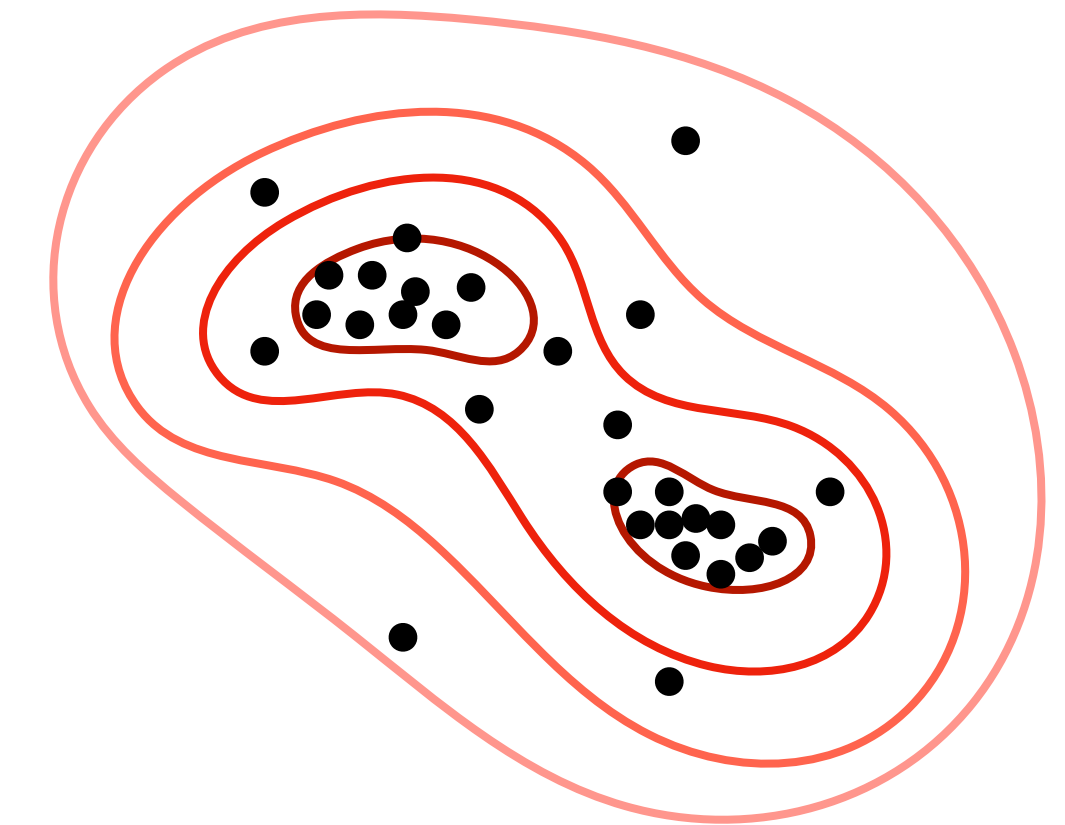
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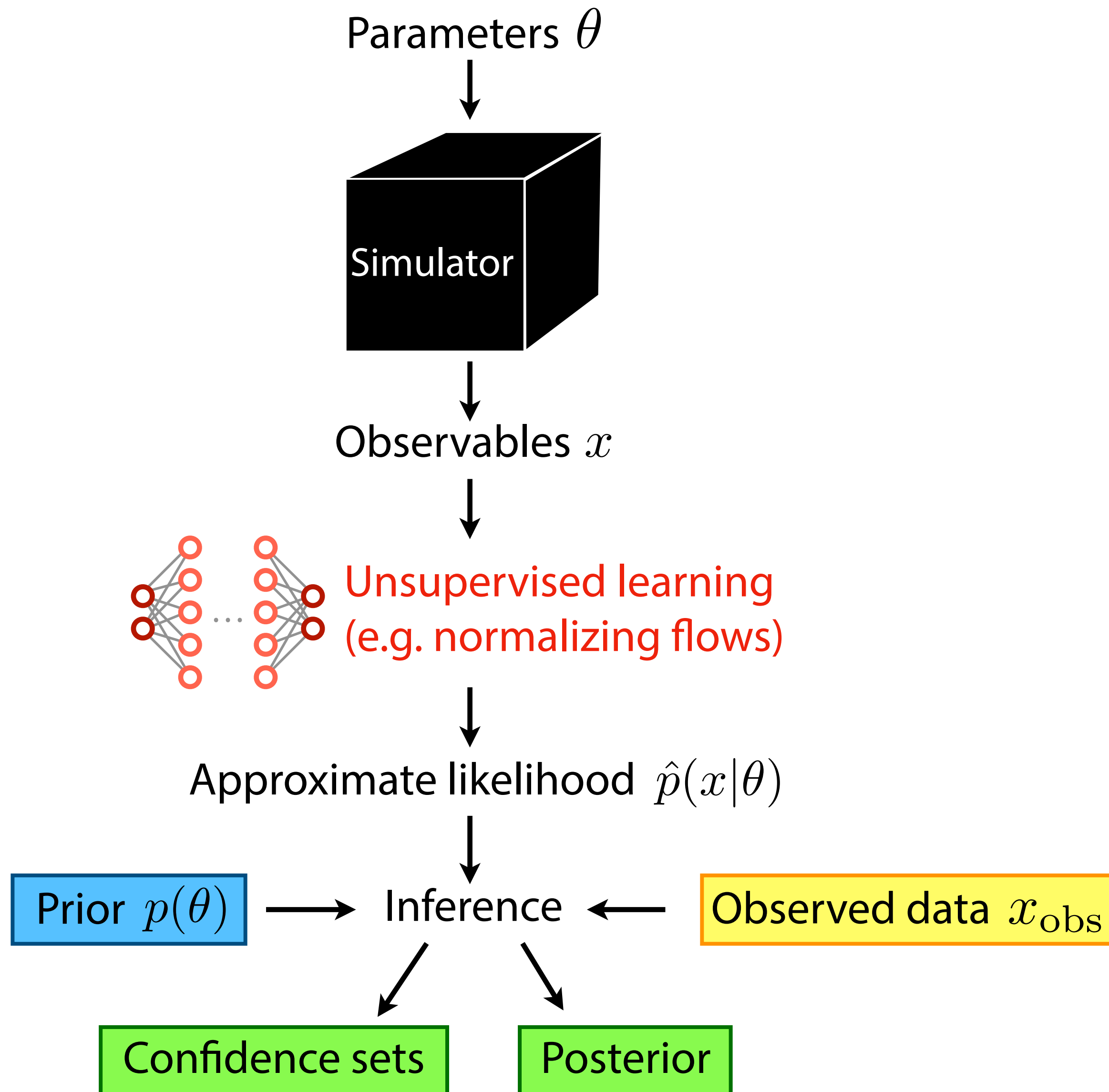
$$\hat{p}(x) = \pi(f^{-1}(x)) |\det \nabla f|^{-1}$$

Train transformation by  
maximizing  $\log \hat{p}(x)$

Transformation can depend on  $\theta$   
to model conditional density  $\log \hat{p}(x|\theta)$

# Inference with neural likelihood estimation

[K.C., Gilles Louppe <https://doi.org/10.5281/zenodo.198541> ;  
G. Papamakarios, D. Sterrart, I. Murray 1805.07226;  
J.-M. Lueckmann, G. Bassetto, T. Karaletsos, J. Macke 1805.09294]



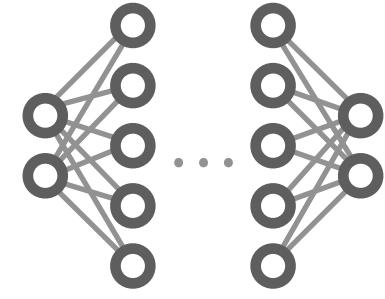
- Conditional neural density estimator (e.g. normalizing flow) as tractable surrogate for simulator likelihood
- Scales well to high-dimensional data (no compression to summary stats necessary)
- Amortized: After upfront simulation + training phase, inference is efficient for new data or prior
- Related alternative: learn posterior  $\hat{p}(\theta|x_{obs})$

[G. Papamakarios et al 1605.06376;  
J.-M. Lueckmann et al 1711.01861]

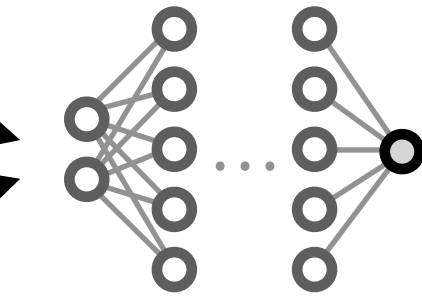
# The likelihood ratio trick

- Generative Adversarial Networks (GANs):

Generator  
(intractable likelihood  $p_{\text{gen}}(x)$ )



Examples  
(intractable likelihood  $p_{\text{true}}(x)$ )



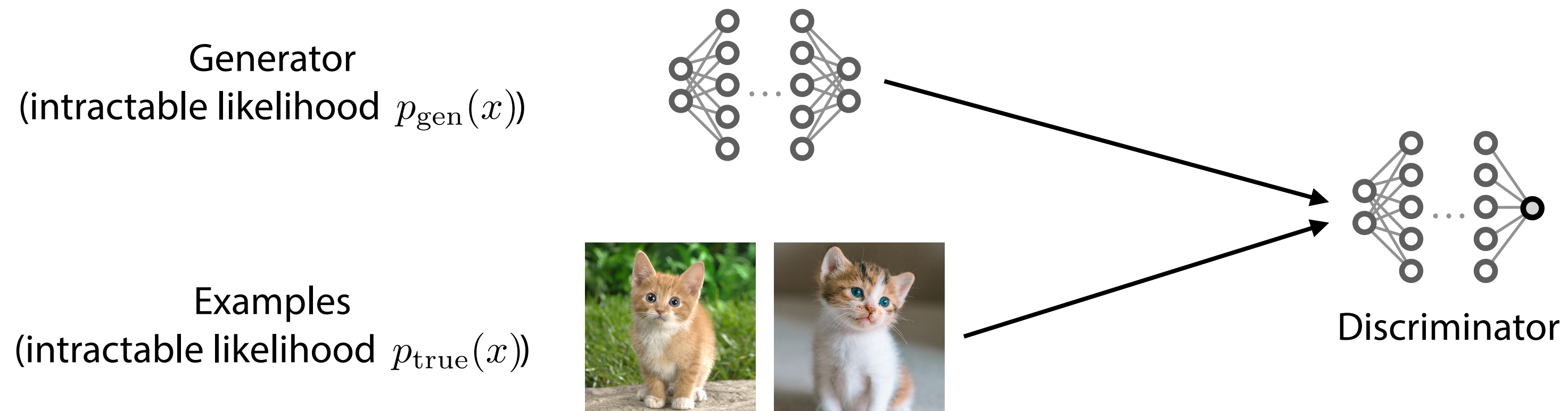
Discriminator

[I. Goodfellow et al. 1406.2661]



# The likelihood ratio trick

- Generative Adversarial Networks (GANs):



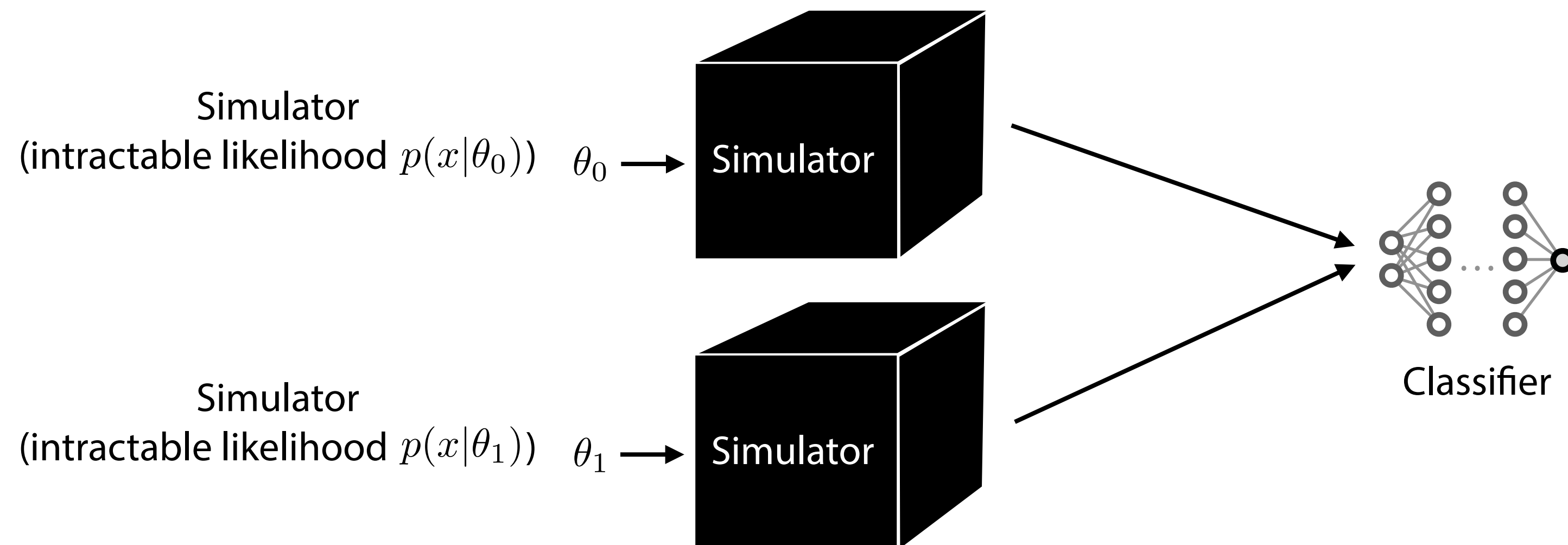
[I. Goodfellow et al. 1406.2661]

Discriminator learns decision function

$$s(x) \rightarrow \frac{p_{\text{true}}(x)}{p_{\text{gen}}(x) + p_{\text{true}}(x)}$$

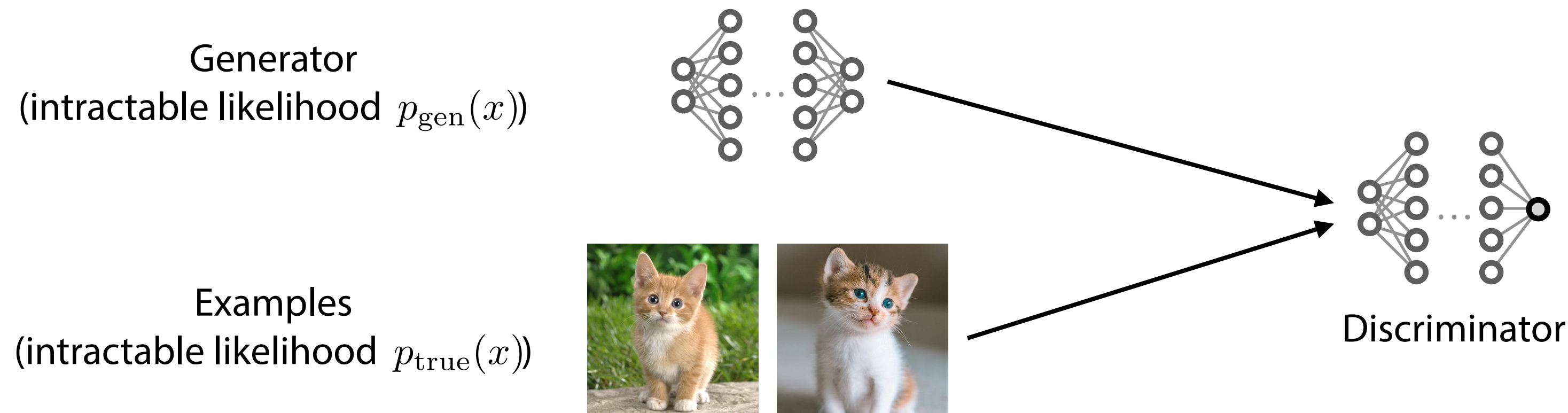
- Similarly, we can train a classifier between two sets of simulated samples

[K. Cranmer, J. Pavez, G. Louppe 1506.02169]



# The likelihood ratio trick

- Generative Adversarial Networks (GANs):



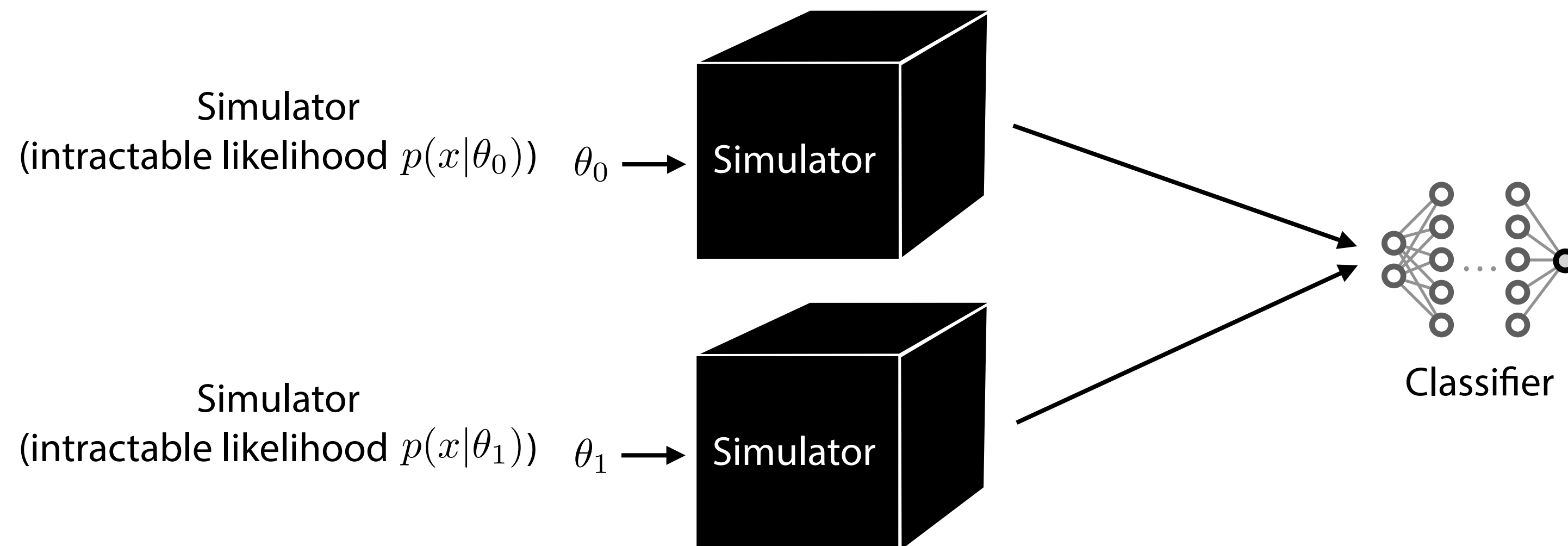
[I. Goodfellow et al. 1406.2661]

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[K. Cranmer, J. Pavez, G. Louppe 1506.02169]



Classifier learns decision function

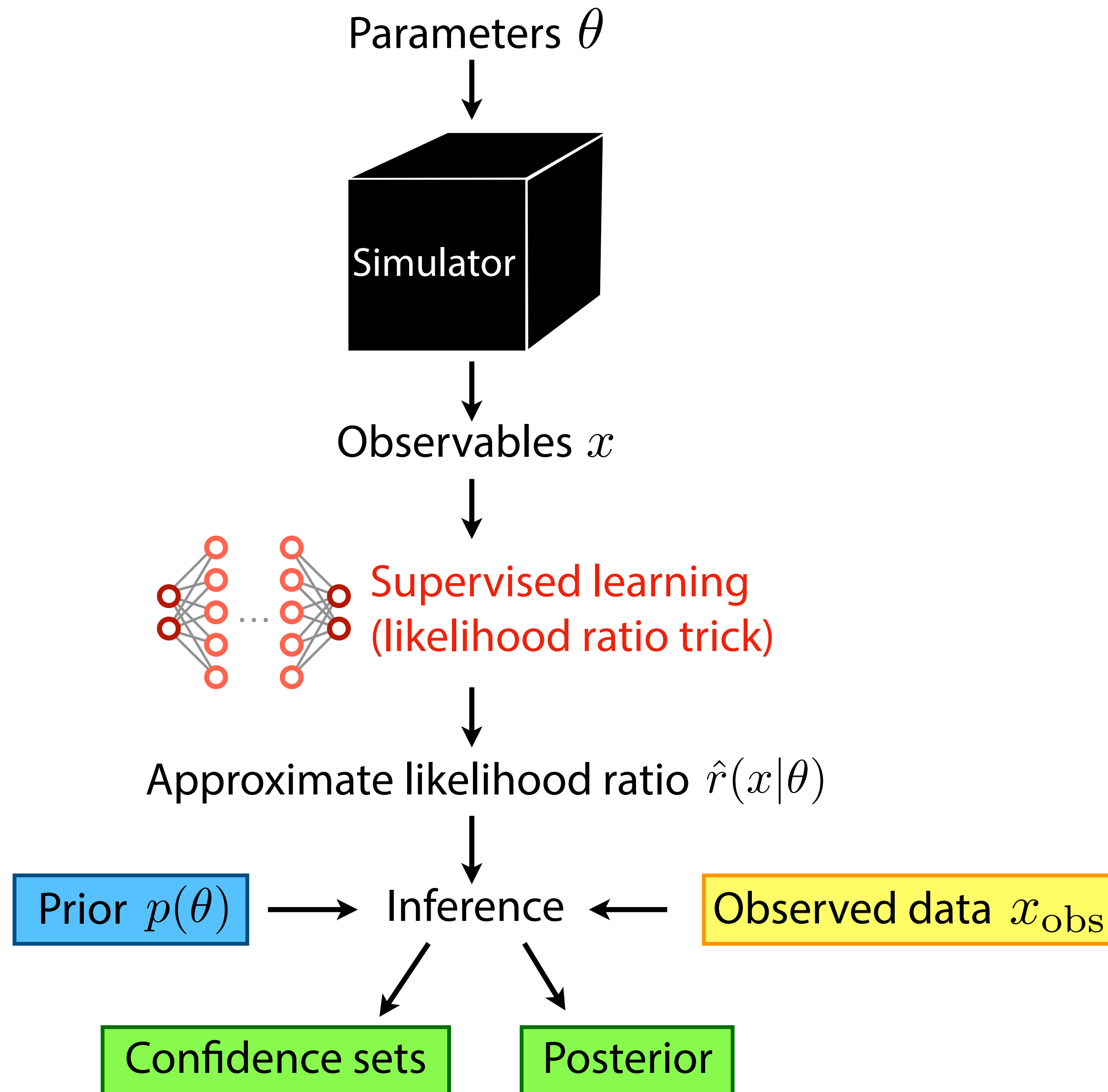
$$s(x) \rightarrow \frac{p(x|\theta_1)}{p(x|\theta_0) + p(x|\theta_1)}$$

$\Rightarrow$  Estimator for likelihood ratio

$$\hat{r}(x) = \frac{1 - s(x)}{s(x)} \rightarrow \frac{p(x|\theta_0)}{p(x|\theta_1)}$$

# Inference by likelihood ratio trick

[K. Cranmer J. Pavez, G. Louppe 1506.02169]



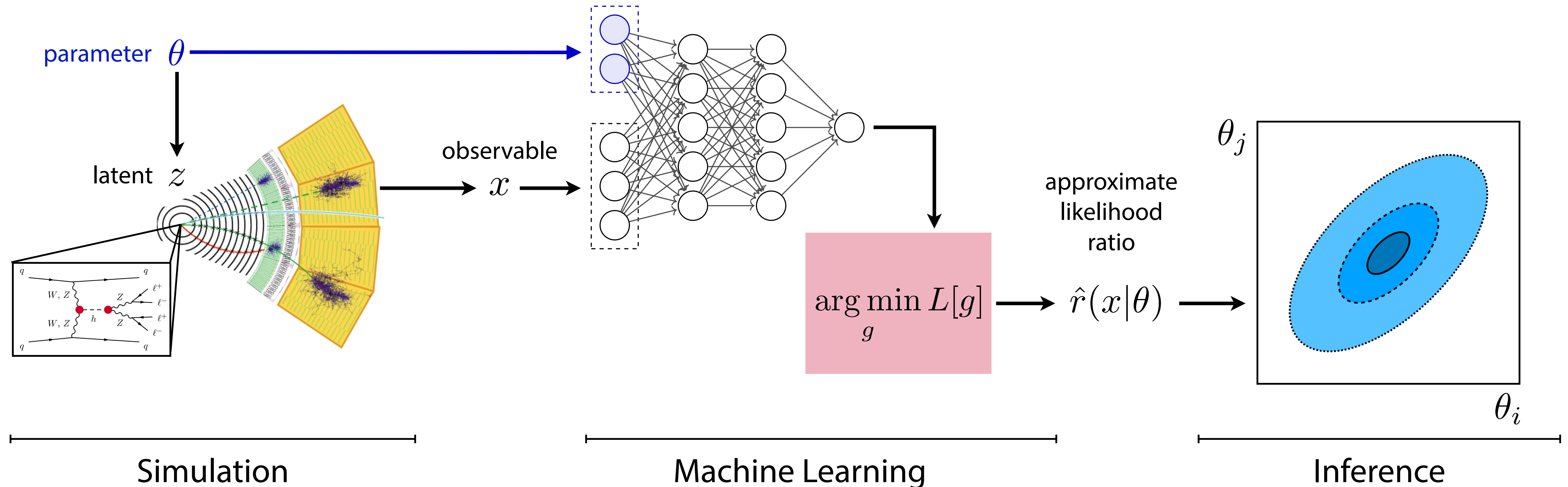
- For inference, likelihood and likelihood ratio are interchangeable
- Advantage: Learning the likelihood ratio can be a simpler task than learning the likelihood
- Disadvantage: Cannot sample from likelihood ratio



“Mining gold”:  
Physics insights can make these inference methods  
more efficient.

[JB, K. Cranmer, G. Louppe, J. Pavez 1805.00013, 1805.00020, 1805.12244]

# Learning with Simulated Data

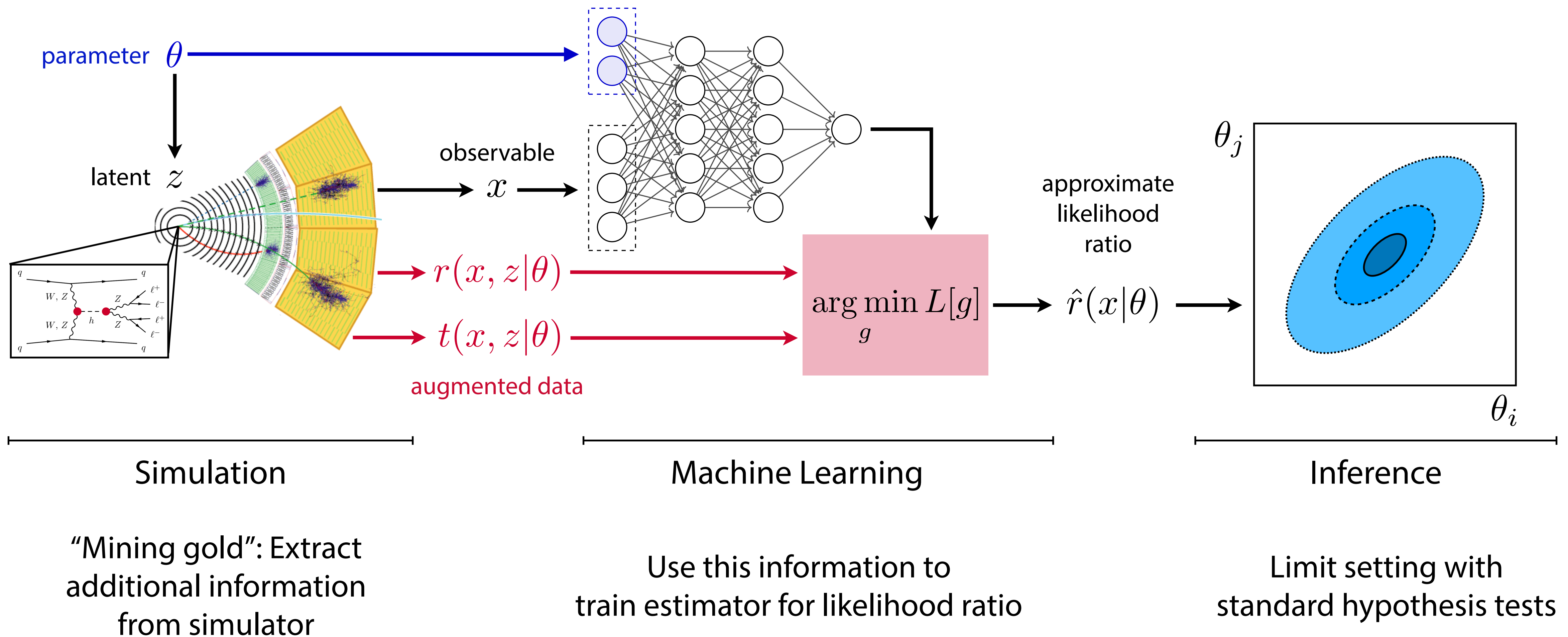


“Mining gold”: Extract additional information from simulator

Use this information to train estimator for likelihood ratio

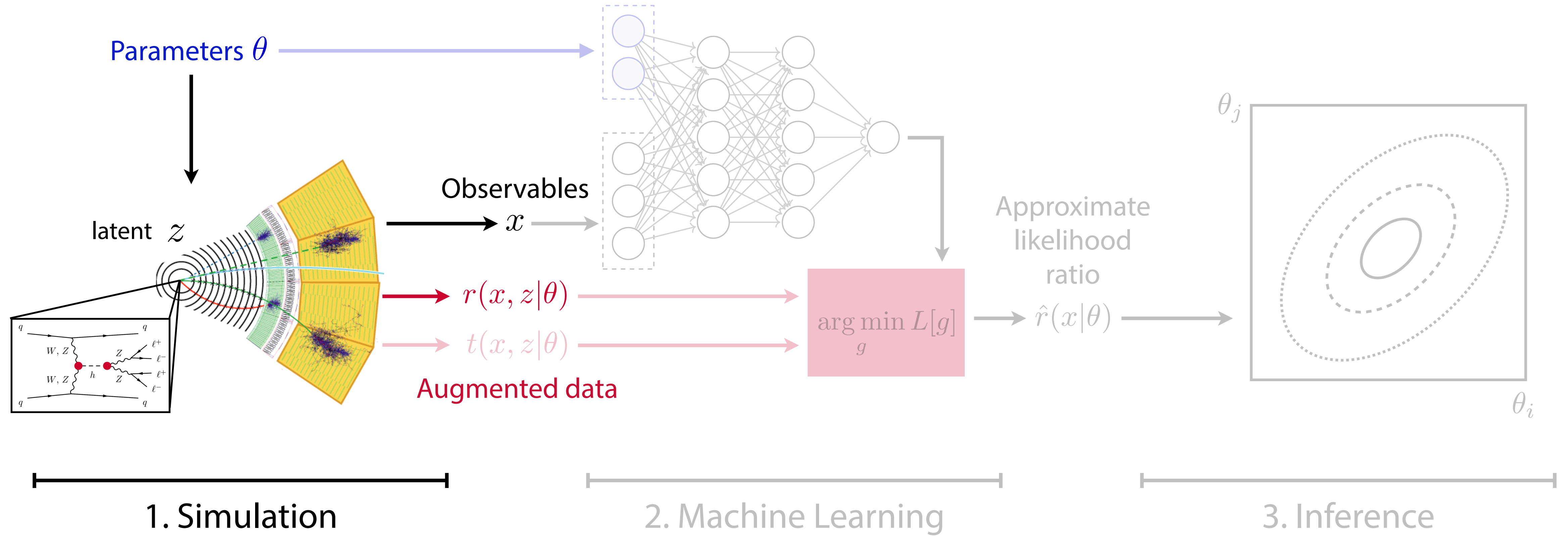
Limit setting with standard hypothesis tests

# Learning with Augmented Data

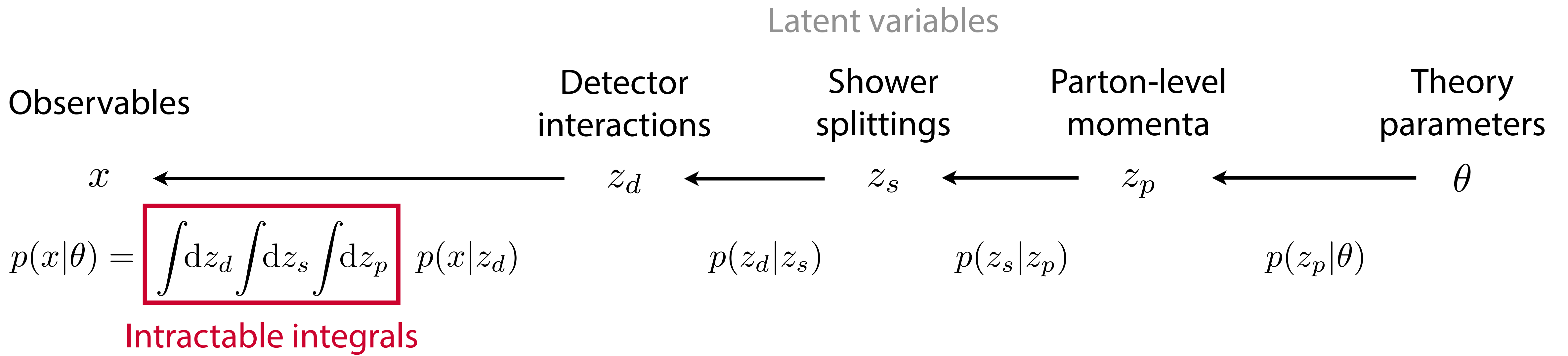




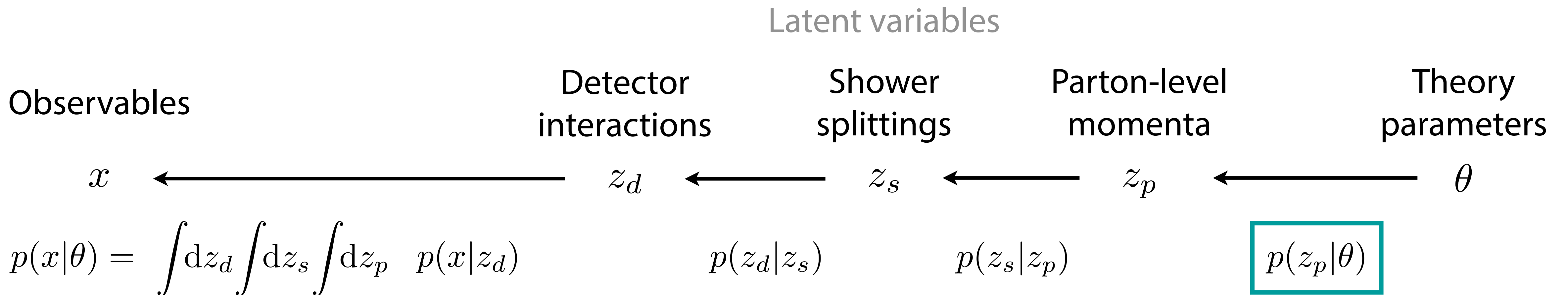
# Learning with Augmented Data



# Mining gold from the simulator



# Mining gold from the simulator



Parton-level likelihood is given by matrix element and can be evaluated!

⇒ For each simulated event, we can calculate the **joint likelihood ratio** which depends on the specific evolution of the simulation:

$$r(x, z|\theta_0, \theta_1) \equiv \frac{p(x, z_d, z_s, z_p|\theta_0)}{p(x, z_d, z_s, z_p|\theta_1)} = \frac{p(x|z_d)}{p(x|z_d)} \frac{p(z_d|z_s)}{p(z_d|z_s)} \frac{p(z_s|z_p)}{p(z_s|z_p)} \frac{p(z_p|\theta_0)}{p(z_p|\theta_1)} \sim \frac{|\mathcal{M}(z_p|\theta_0)|^2}{|\mathcal{M}(z_p|\theta_1)|^2}$$



# The value of gold

We can calculate the **joint likelihood ratio**

$$r(x, z|\theta_0, \theta_1) \equiv \frac{p(x, z_d, z_s, z_p|\theta_0)}{p(x, z_d, z_s, z_p|\theta_1)}$$



We want the **likelihood ratio function**

$$r(x|\theta_0, \theta_1) \equiv \frac{p(x|\theta_0)}{p(x|\theta_1)}$$

(“How much more likely is this simulated event, including all intermediate states, for  $\theta_0$  compared to  $\theta_1$ ?”)

(“How much more likely is the observation  $x$  for  $\theta_0$  compared to  $\theta_1$ ?”)

# The value of gold

We can calculate the **joint likelihood ratio**

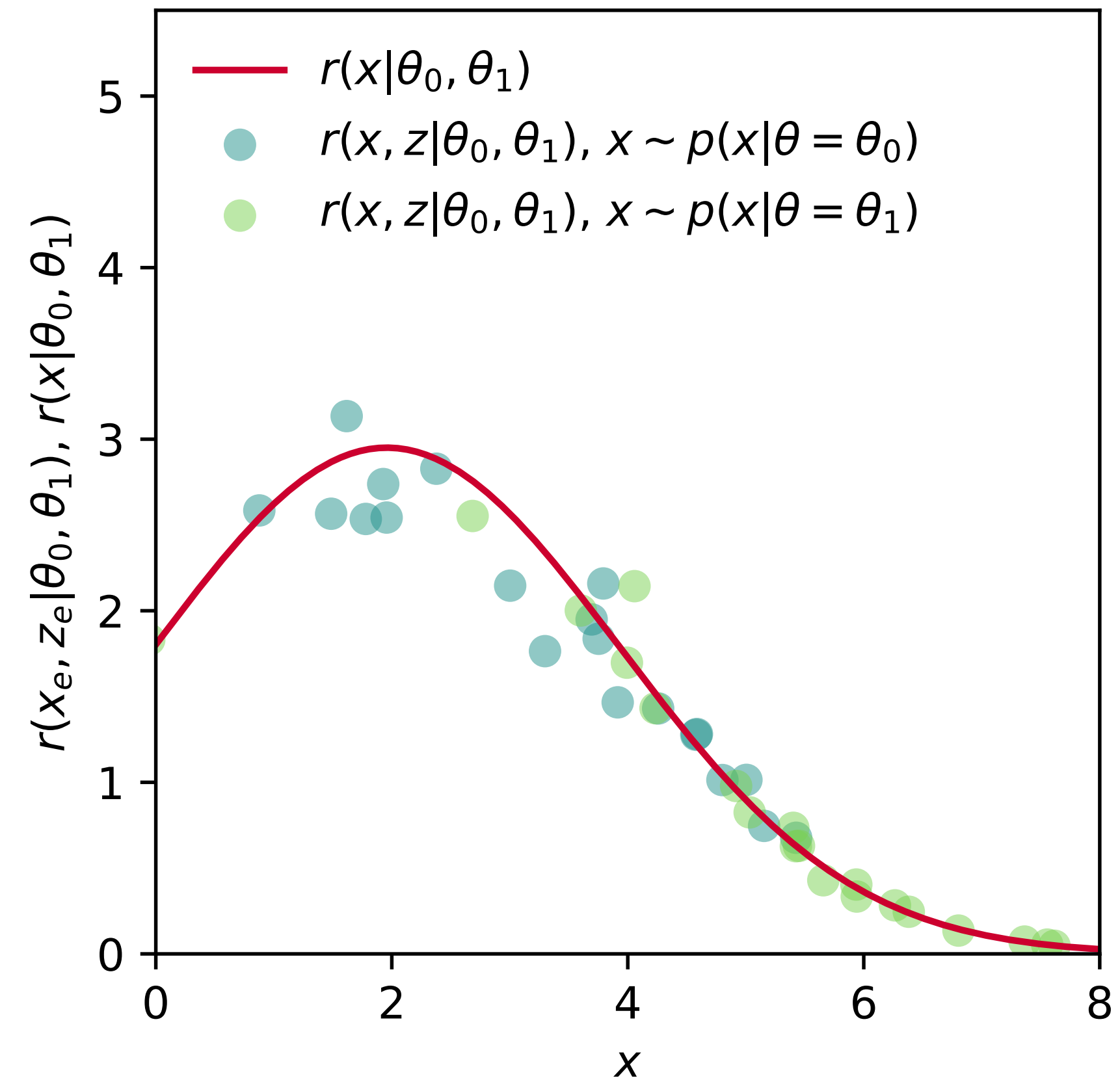
$$r(x, z|\theta_0, \theta_1) \equiv \frac{p(x, z_d, z_s, z_p|\theta_0)}{p(x, z_d, z_s, z_p|\theta_1)}$$



$r(x, z|\theta_0, \theta_1)$  are scattered around  $r(x|\theta_0, \theta_1)$

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We want the **likelihood ratio function**

$$r(x|\theta_0, \theta_1) \equiv \frac{p(x|\theta_0)}{p(x|\theta_1)}$$

With  $r(x, z|\theta_0, \theta_1)$ , we define a functional like

$$L_r[\hat{r}(x|\theta_0, \theta_1)] = \int dx \int dz p(x, z|\theta_1) \left[ (\hat{r}(x|\theta_0, \theta_1) - r(x, z|\theta_0, \theta_1))^2 \right]$$

It is minimized by

$$\mathbb{E}_{z \sim p(z|x, \theta_1)} [r(x, z|\theta_0, \theta_1)] = \arg \min_{\hat{r}(x|\theta_0, \theta_1)} L_r[\hat{r}(x|\theta_0, \theta_1)]!$$

(And we can sample from  $p(x, z|\theta)$  by running the simulator.)



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**.... and then magic ...**

$$\begin{aligned} \mathbb{E}_{z \sim p(z|x, \theta_1)} [r(x, z|\theta_0, \theta_1)] &= \int dz p(z|x, \theta_1) \frac{p(x, z|\theta_0)}{p(x, z|\theta_1)} \\ &= \int dz \frac{p(x, z|\theta_1)}{p(x|\theta_1)} \frac{p(x, z|\theta_0)}{p(x, z|\theta_1)} \\ &= r(x|\theta_0, \theta_1) ! \end{aligned}$$

# Machine learning = applied calculus of variations

So to get a good estimator of the likelihood ratio, we need to minimize a functional numerically:

The diagram illustrates the process of finding a good estimator of the likelihood ratio. It features three main components at the top: "Variational family", "Extremization", and "Functional with integral".

- An arrow labeled "Variational family" points from the top left to the  $\hat{r}(x|\theta_0, \theta_1)$  term in the equation below.
- An arrow labeled "Extremization" points from the top center to the  $\arg \min$  operator in the equation below.
- An arrow labeled "Functional with integral" points from the top right to the  $L_r[\hat{r}(x|\theta_0, \theta_1)]$  term in the equation below.

$$r(x|\theta_0, \theta_1) = \arg \min_{\hat{r}(x|\theta_0, \theta_1)} L_r[\hat{r}(x|\theta_0, \theta_1)]$$

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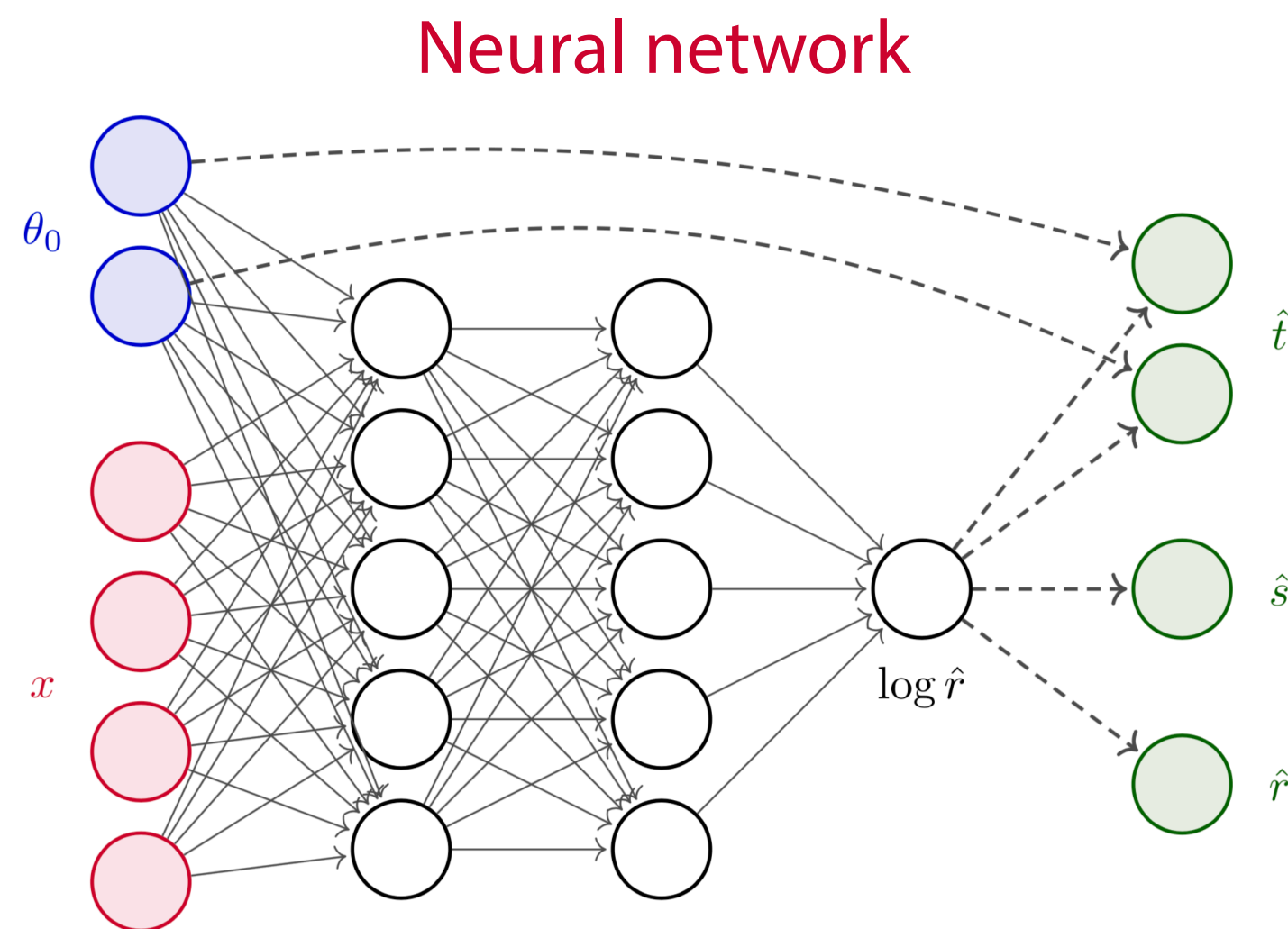
$$r(x|\theta_0, \theta_1) = \underset{\hat{r}(x|\theta_0, \theta_1)}{\operatorname{arg\,min}} L_r[\hat{r}(x|\theta_0, \theta_1)]$$

Variational family      Extremization      Functional with integral

This is where machine learning comes in!

Loss function with finite sum over samples

Stochastic gradient descent



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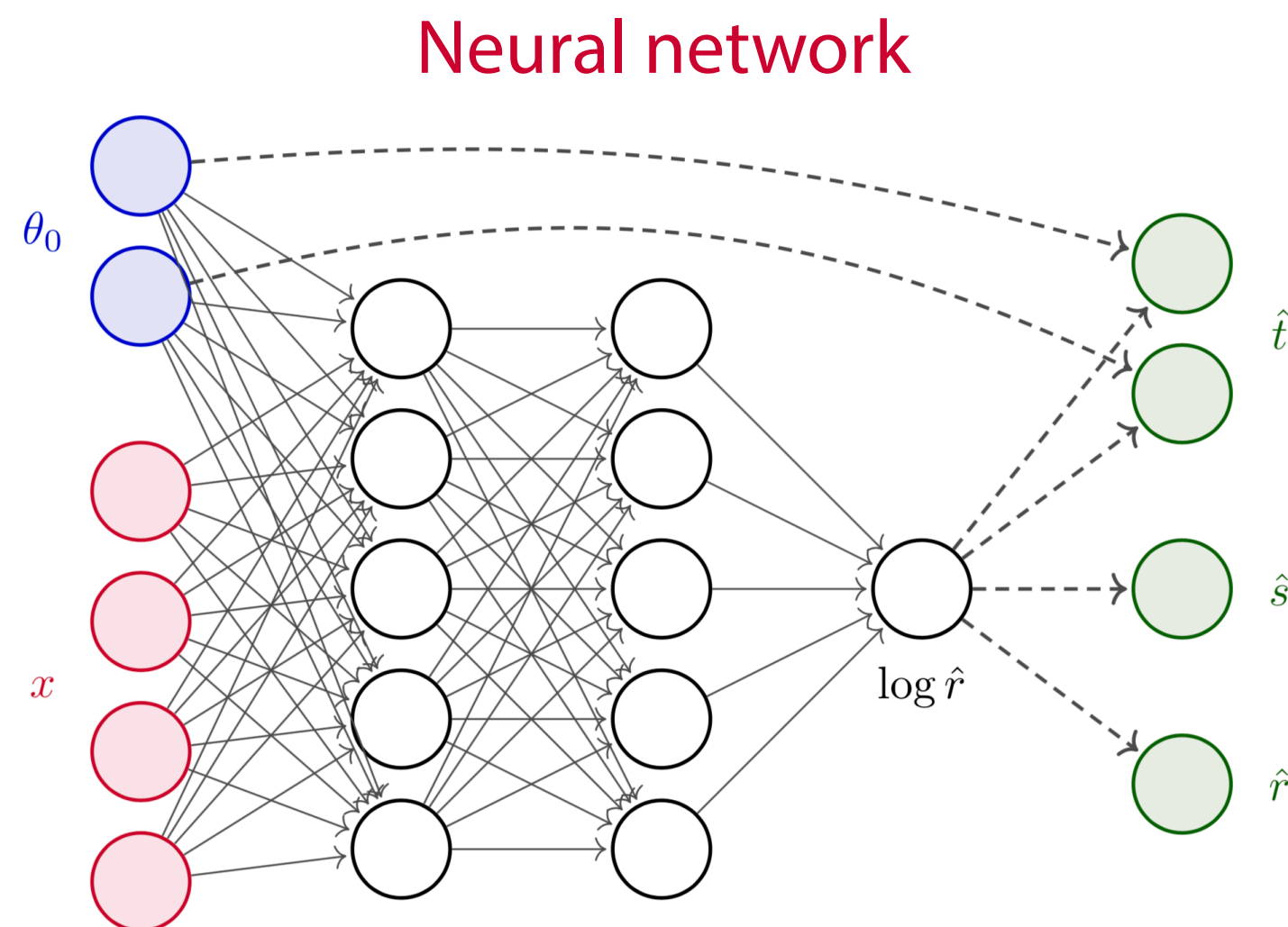
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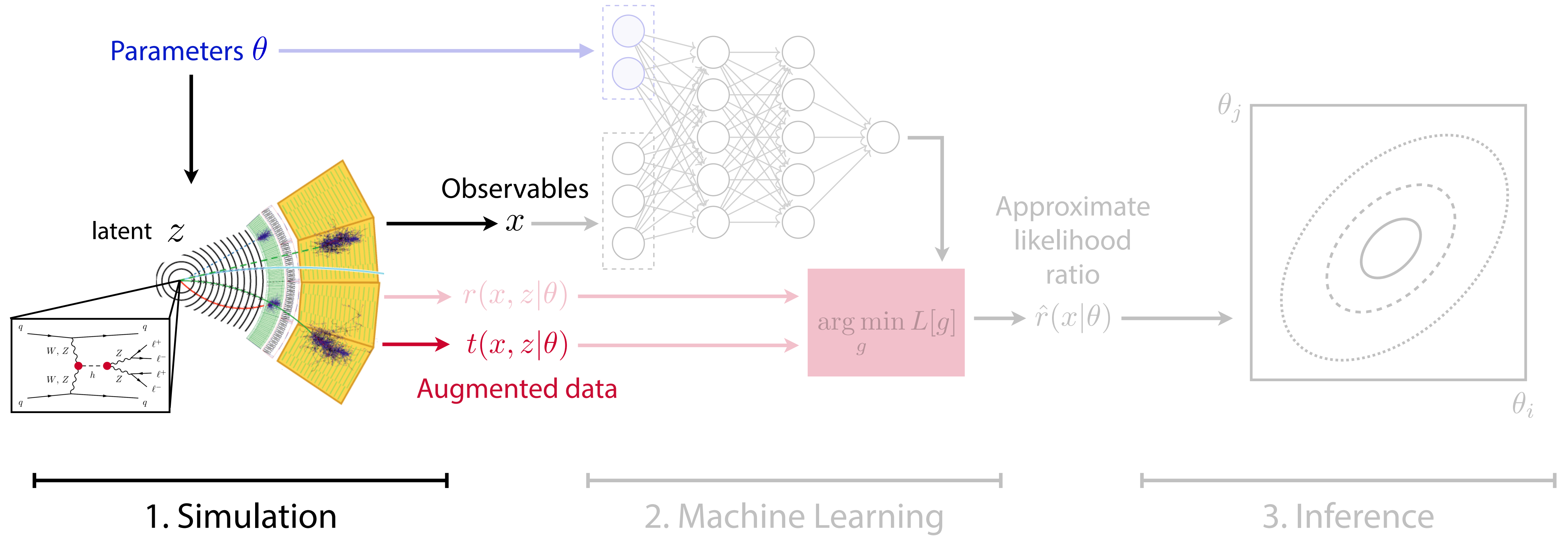
Stochastic gradient descent



A sufficiently expressive neural network efficiently trained in this way with enough data will learn the likelihood ratio function  $r(x|\theta_0, \theta_1)$ !

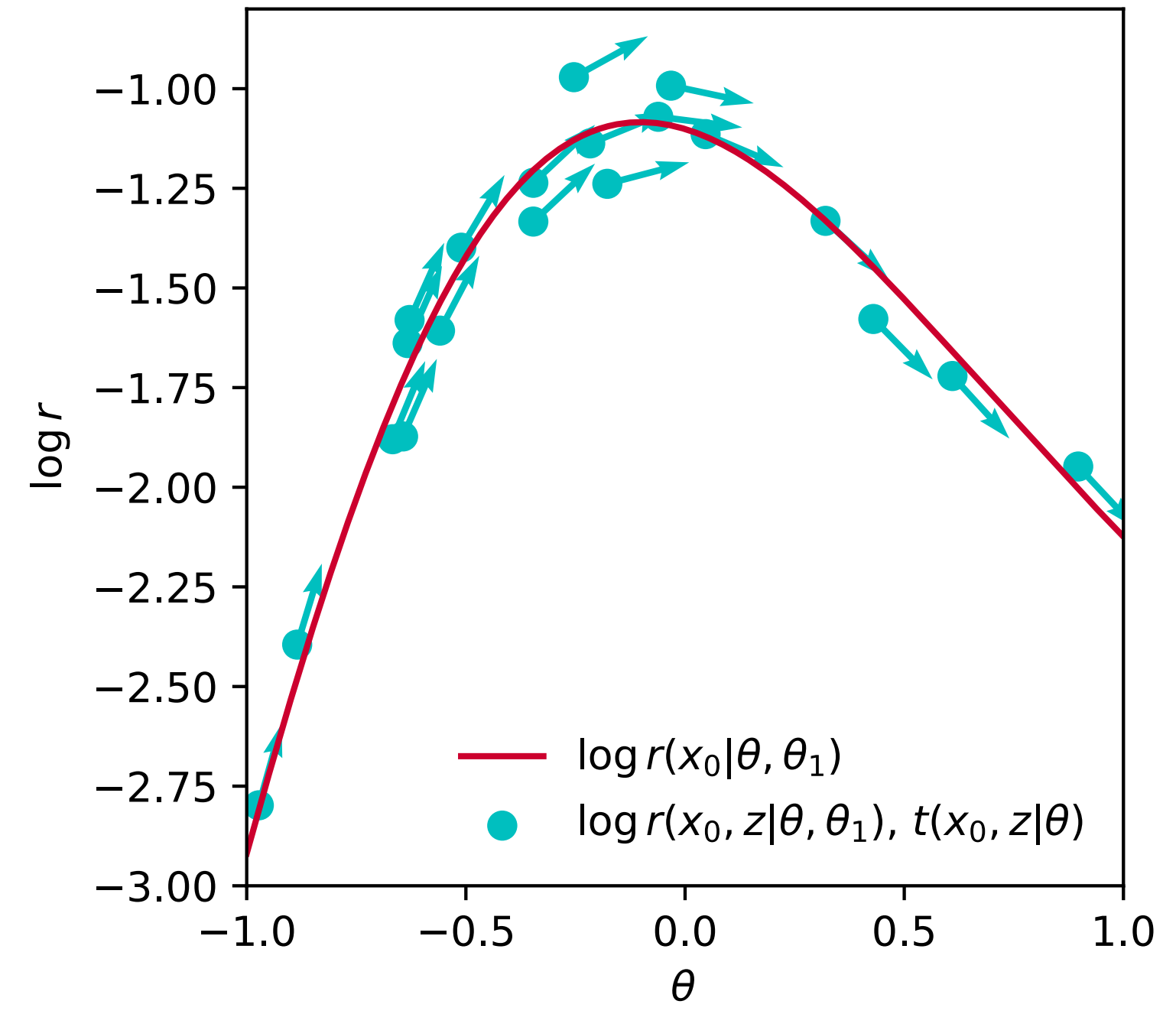
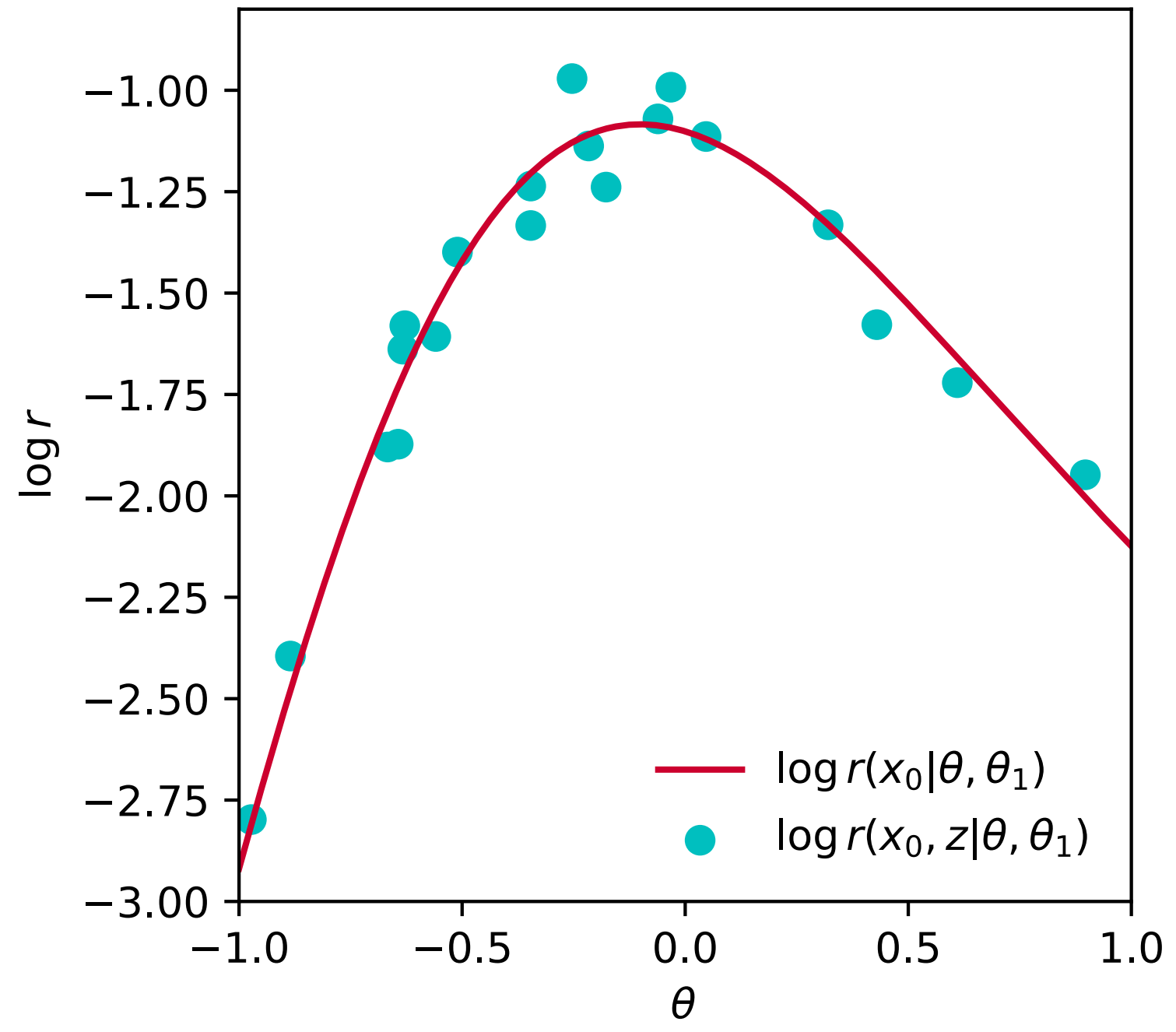


# Learning with Augmented Data



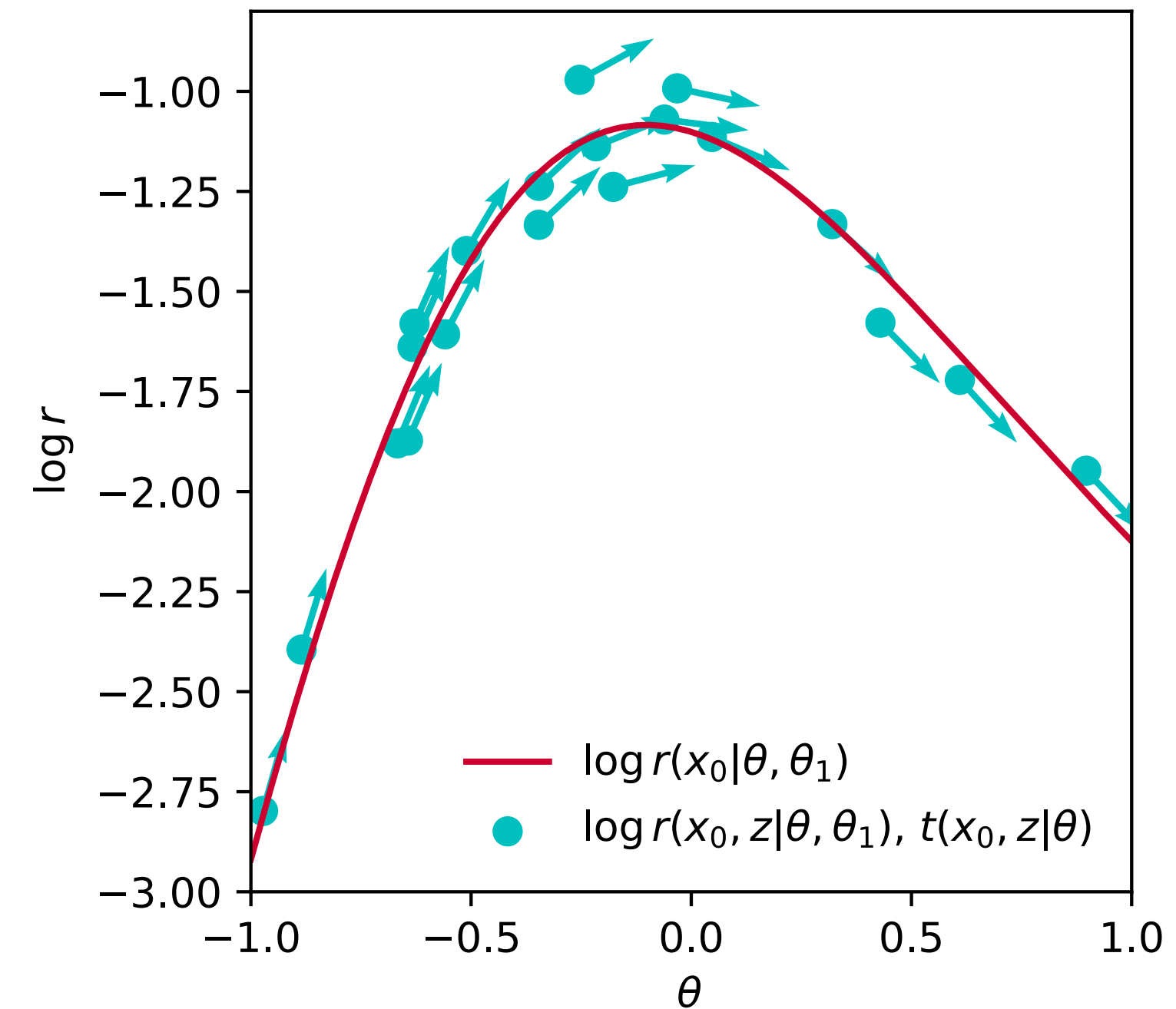
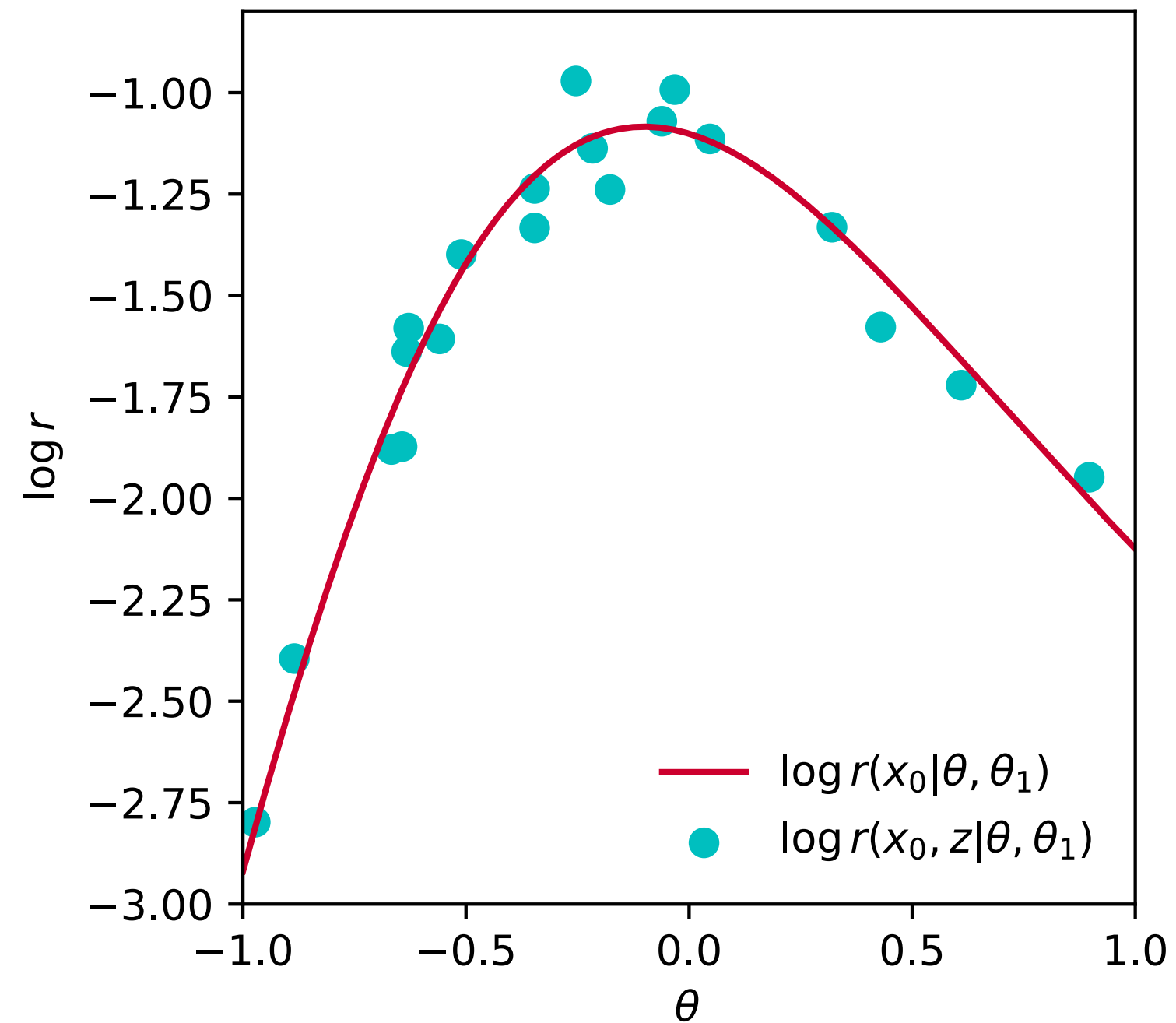
# One more piece: the score

- Knowing derivative often helps fitting:



# One more piece: the score

- Knowing derivative often helps fitting:



- In our case, the relevant quantity is the **score**  $t(x|\theta_0) \equiv \nabla_{\theta} \log p(x|\theta) \Big|_{\theta_0}$ .

- The score itself is intractable. But...

# Learning the score

Similar to the joint likelihood ratio, from the simulator we can extract the **joint score**

$$t(x, z|\theta_0) \equiv \nabla_{\theta} \log p(x, z_d, z_s, z_p|\theta) \Big|_{\theta_0}$$



We want the **score**

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$$t(x|\theta_0) \equiv \nabla_{\theta} \log p(x|\theta) \Big|_{\theta_0}$$

Given  $t(x, z|\theta_0)$ ,  
we define the functional

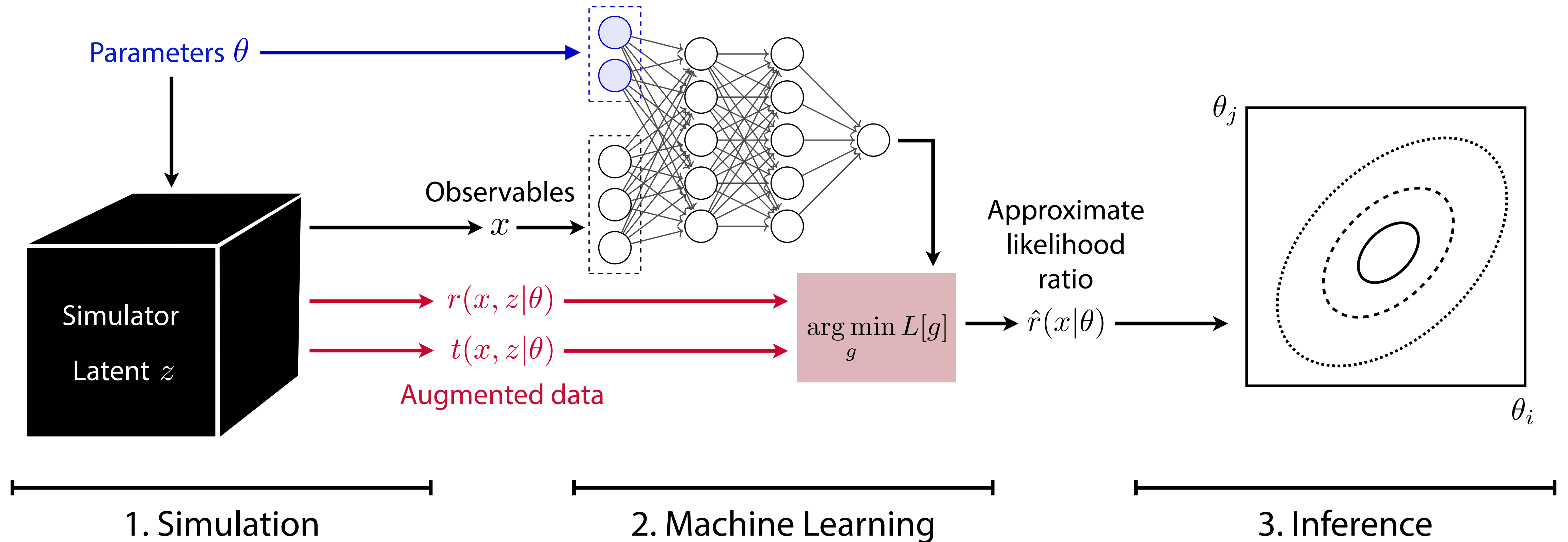
$$L_t[\hat{t}(x|\theta_0)] = \int dx \int dz p(x, z|\theta_0) \left[ (\hat{t}(x|\theta_0) - t(x, z|\theta_0))^2 \right].$$

One can show it is minimized by

$$t(x|\theta_0) = \arg \min_{\hat{t}(x|\theta_0)} L_t[\hat{t}(x|\theta_0)].$$

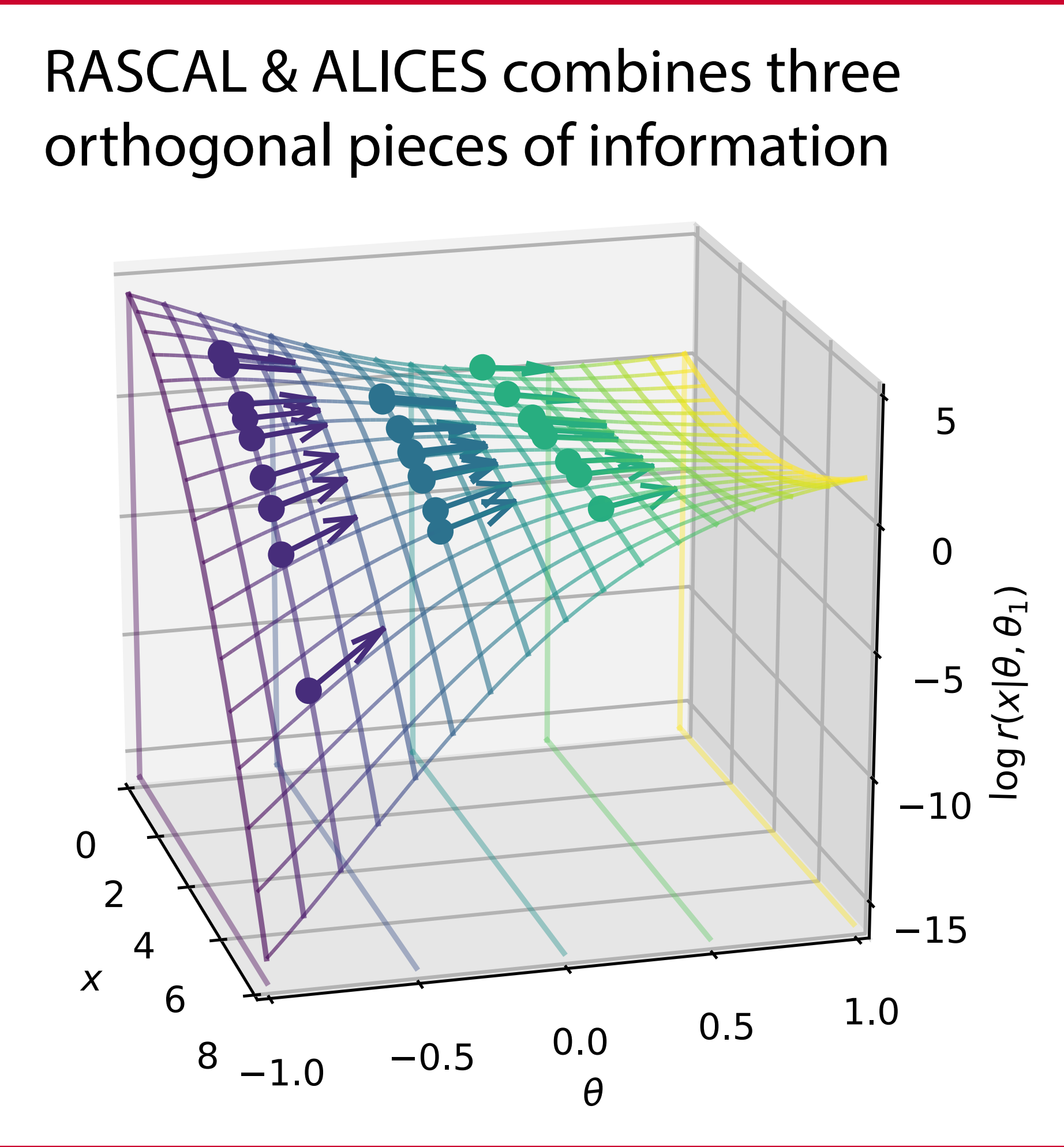
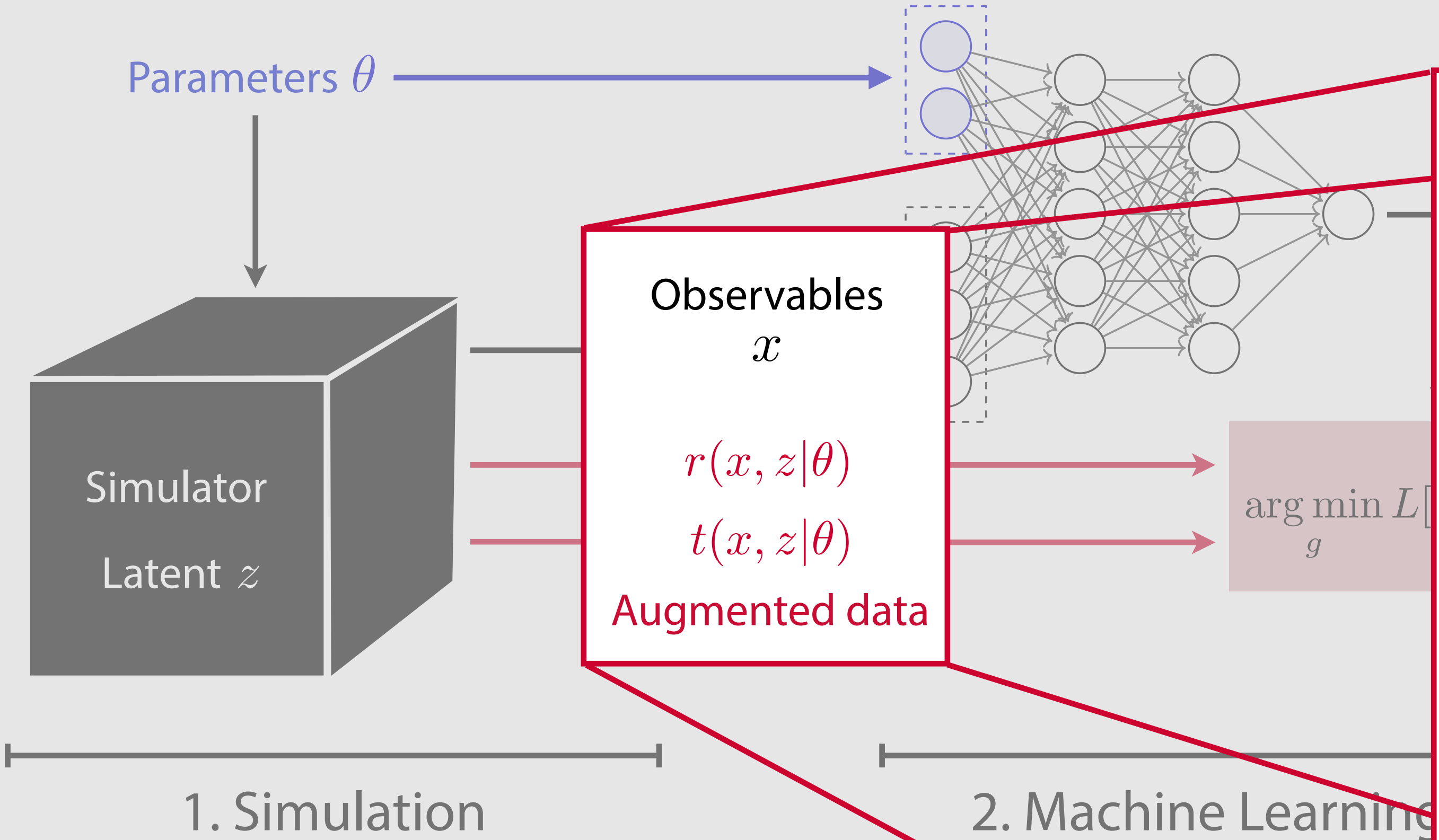
Again, we implement this minimization through machine learning.

# Putting the pieces together: RASCAL & ALICES



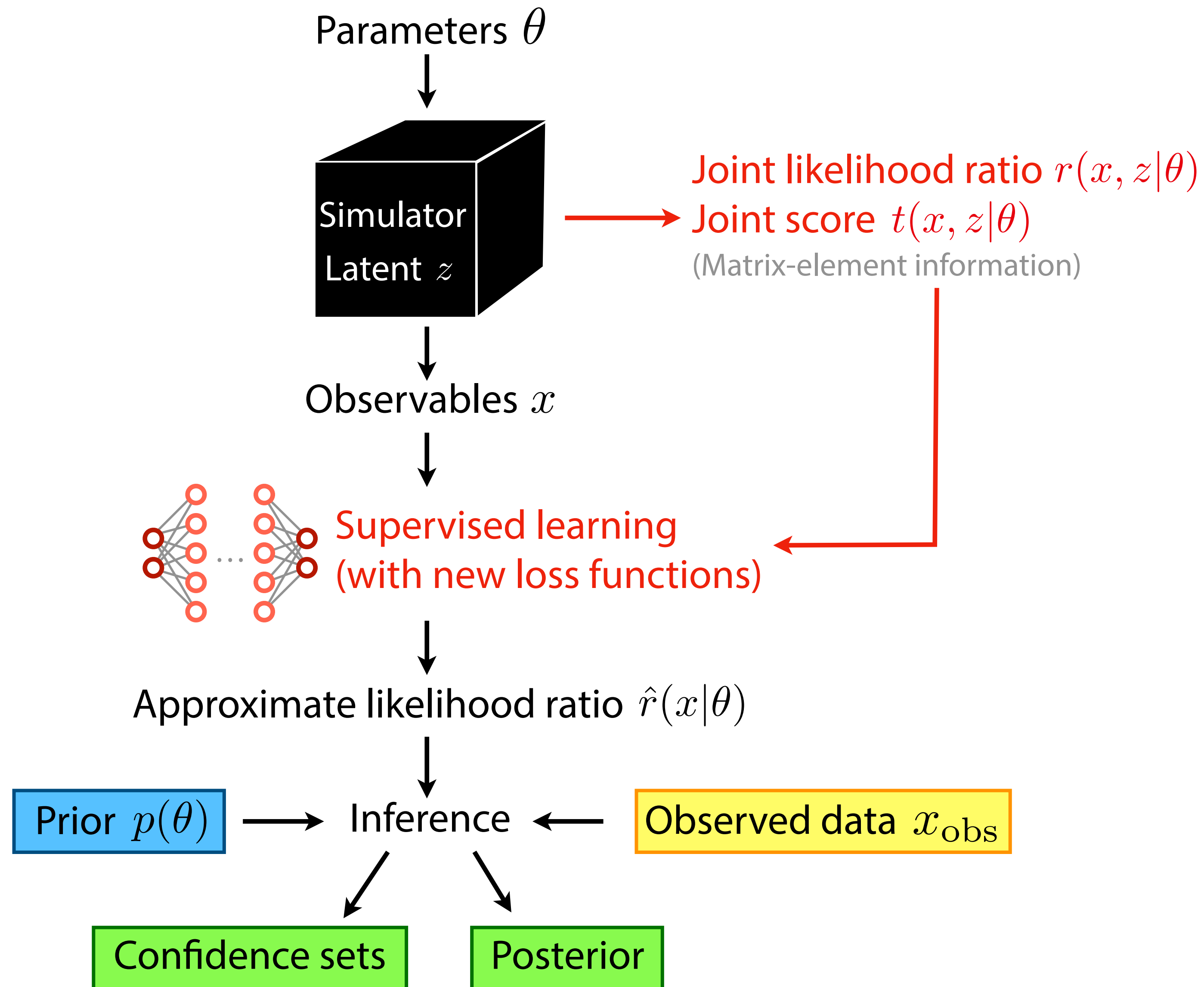
"The machine learning version of the Matrix Element Method"

# Putting the pieces together: RASCAL & ALICES



"The machine learning version of the Ma

# Mining gold: Summary



“The machine learning version of the Matrix Element Method”

- Scales to high-dimensional data (no summary statistics necessary)
- Uses matrix-element information to improve sample efficiency
- Neural networks learn effect of shower + detector (no transfer functions & works with ME+PS matching)
- Amortized (evaluation in  $\mu\text{s}$ )



**MadMiner automates all of these methods.**

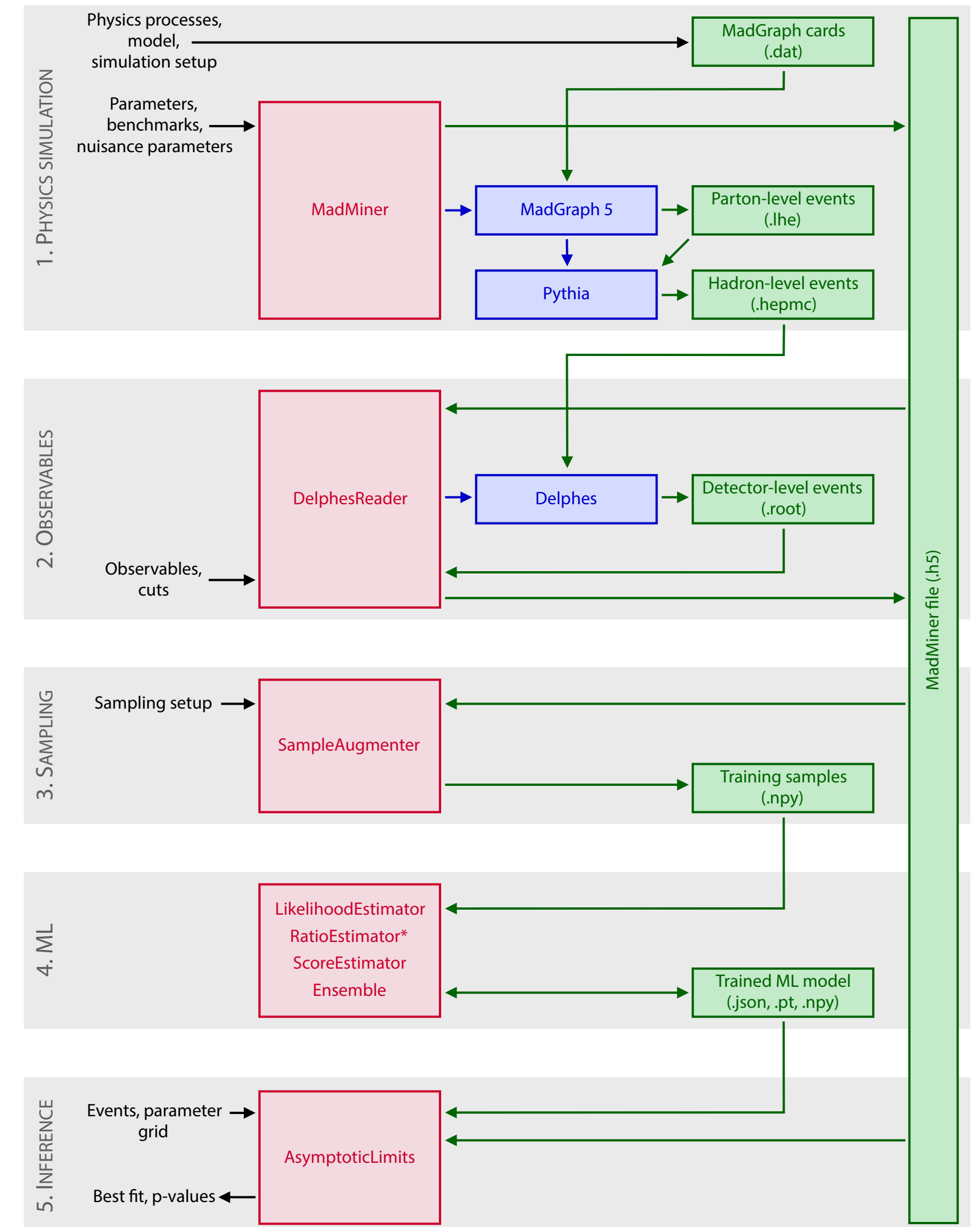
[JB, F. Kling, I. Espejo, K. Cranmer 1907.10621]

# Automation

New Python package **MadMiner** makes it straightforward to apply the new techniques to LHC problems

- Out of the box: Pheno-level analyses
  - MadGraph, Pythia, Delphes
  - Systematic uncertainties from PDF / scale variation
- Scalable to state-of-the-art experimental tools
  - Mostly requires bookkeeping of fully differential cross sections
- Modular interface
  - Extensive documentation
  - Embedded into Python / ML ecosystem

[JB, F. Kling, I. Espejo, K. Cranmer 1907.10621]



### MadMiner: Machine learning-based inference for particle physics

Johann Brehmer,<sup>1,\*</sup> Felix Kling,<sup>2,3,†</sup> Irina Espejo,<sup>1,‡</sup> and Kyle Cranmer<sup>1,§</sup>

<sup>1</sup>Center for Data Science and Center for Cosmology and Particle Physics,  
New York University, New York, NY 10003, USA

<sup>2</sup>Department of Physics and Astronomy, University of California, Irvine, CA 92697, USA

<sup>3</sup>SLAC National Accelerator Laboratory, 2575 Sand Hill Road, Menlo Park, CA 94025, USA

Precision measurements at the LHC often require analyzing high-dimensional event data for subtle kinematic signatures, which is challenging for established analysis methods. Recently, a powerful family of multivariate inference techniques that leverage both matrix element information and machine learning has been developed. This approach neither requires the reduction of high-dimensional data to summary statistics nor any simplifications to the underlying physics or detector response. In this paper we introduce MadMiner, a Python module

### MadMiner: Machine learning-based inference for particle physics

By Johann Brehmer, Felix Kling, Irina Espejo, and Kyle Cranmer

pypi package 0.6.3 build passing docs failing chat on gitter code style black License MIT DOI 10.5281/zenodo.1489147

arXiv 1907.10621

#### Introduction

Particle physics processes are usually modeled with complex Monte-Carlo simulations of the hard process, parton shower, and detector interactions. These simulators typically do not admit a tractable likelihood function: given a (potentially high-dimensional) set of observables, it is usually not possible to calculate the probability of these observables for some model parameters. Particle physicists usually tackle this problem of "likelihood-free inference" by hand-picking a few "good" observables or summary statistics and filling histograms of them. But this conventional

Repository and tutorials:  
[github.com/johannbrehmer/madminer](https://github.com/johannbrehmer/madminer)

Paper with detailed explanations:  
[1907.10621](https://arxiv.org/abs/1907.10621)

Installation:  
`pip install madminer`

API documentation:  
[madminer.readthedocs.io](https://madminer.readthedocs.io)

We can also machine learn optimal observables.

[JB, K. Cranmer, G. Louppe, J. Pavez 1805.00013, 1805.00020, 1805.12244]



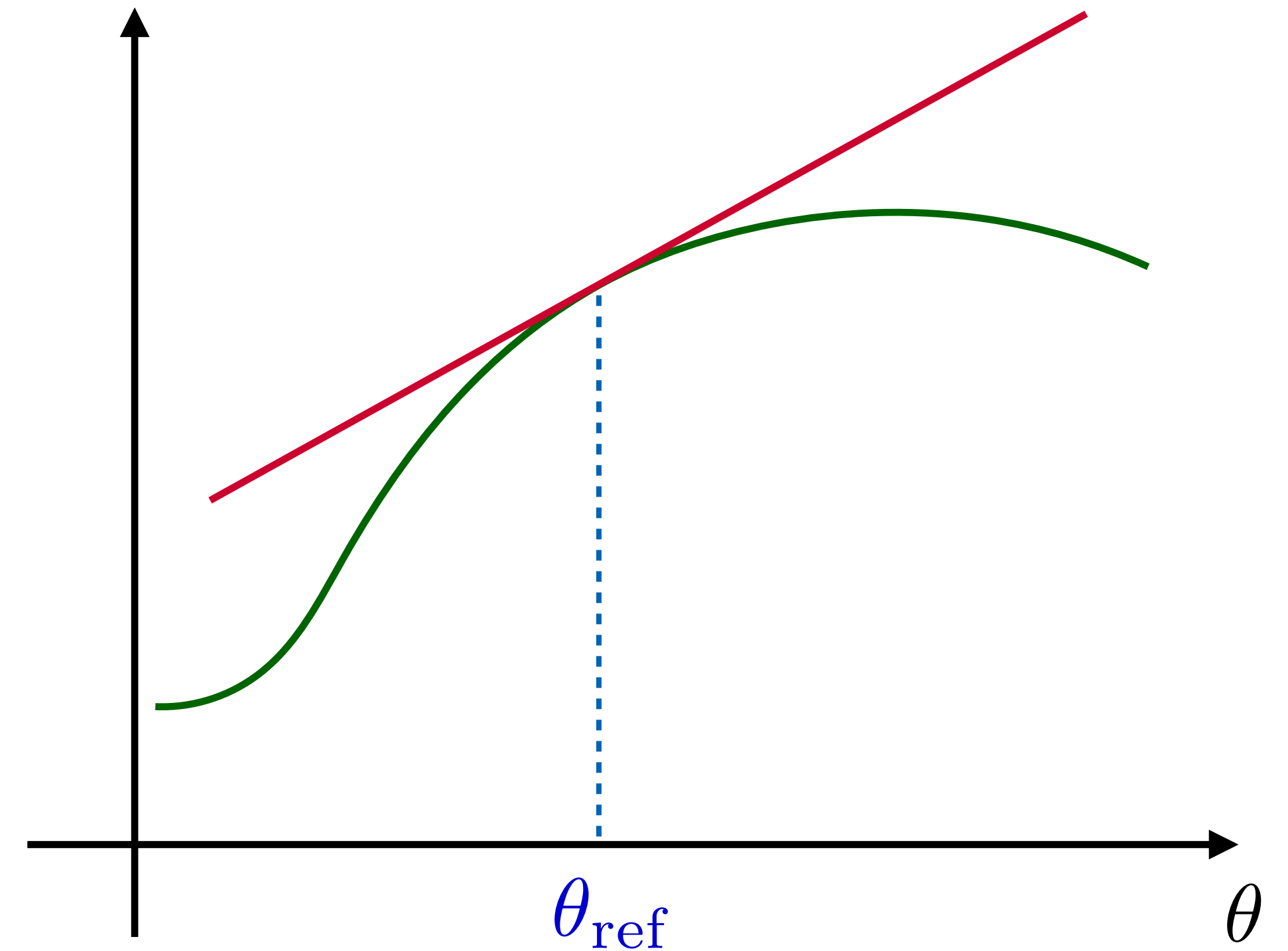


# The local model

[see also J. Alsing, B. Wandelt 1712.00012; J. Alsing, B. Wandelt, S. Freney 1801.01497; P. de Castro, T. Dorigo 1806.04743; J. Alsing, B. Wandelt 1903.01473]

Taylor expansion of  $\log p(x|\theta)$  around  $\theta_{\text{ref}}$ :

$$\begin{aligned}\log p(x|\theta) &= \log p(x|\theta_{\text{ref}}) \\ &+ \underbrace{\nabla_{\theta} \log p(x|\theta) \Big|_{\theta_{\text{ref}}}}_{\equiv t(x|\theta_{\text{ref}})} \cdot (\theta - \theta_{\text{ref}}) \\ &+ \mathcal{O}((\theta - \theta_{\text{ref}})^2)\end{aligned}$$



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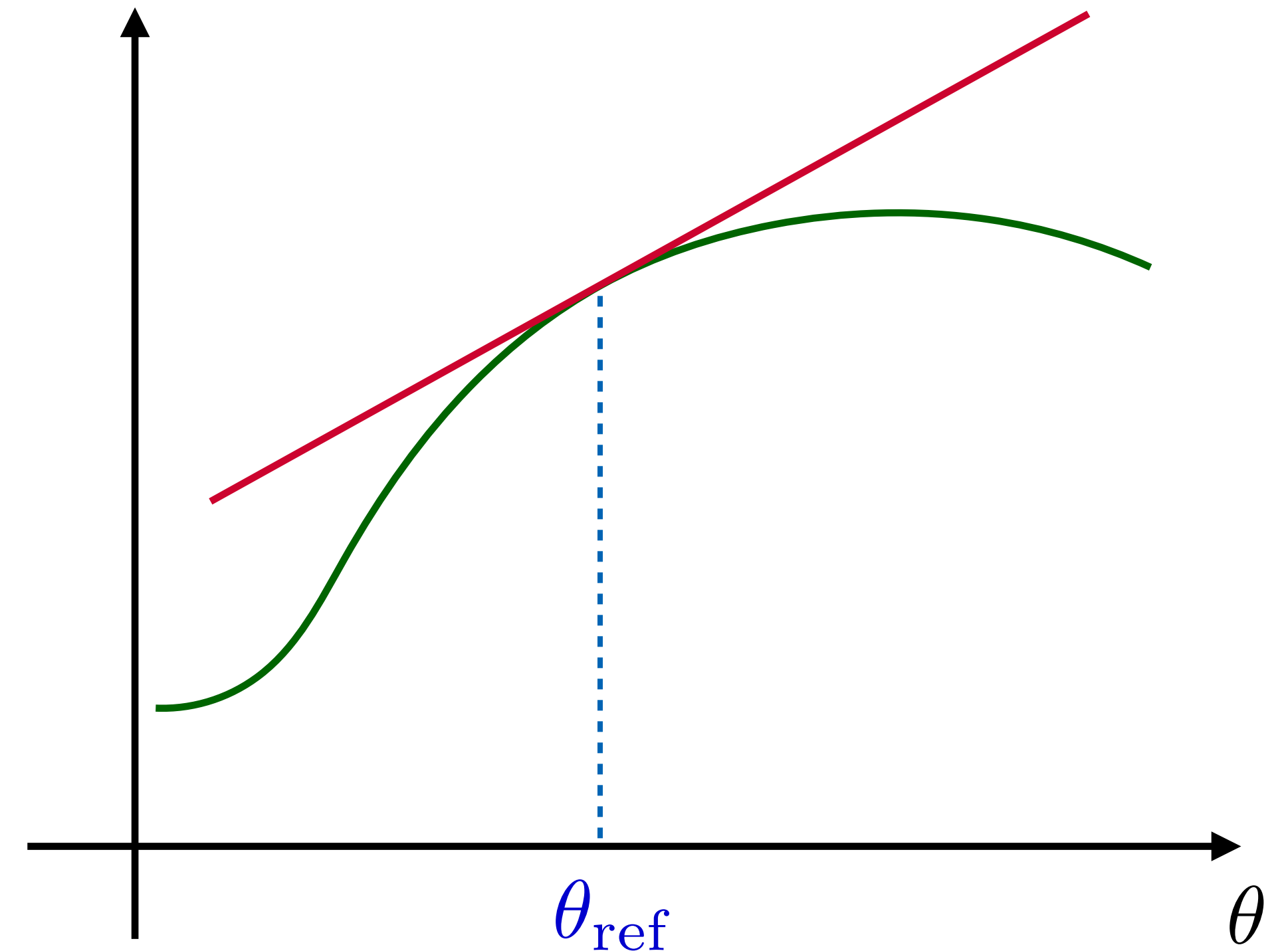
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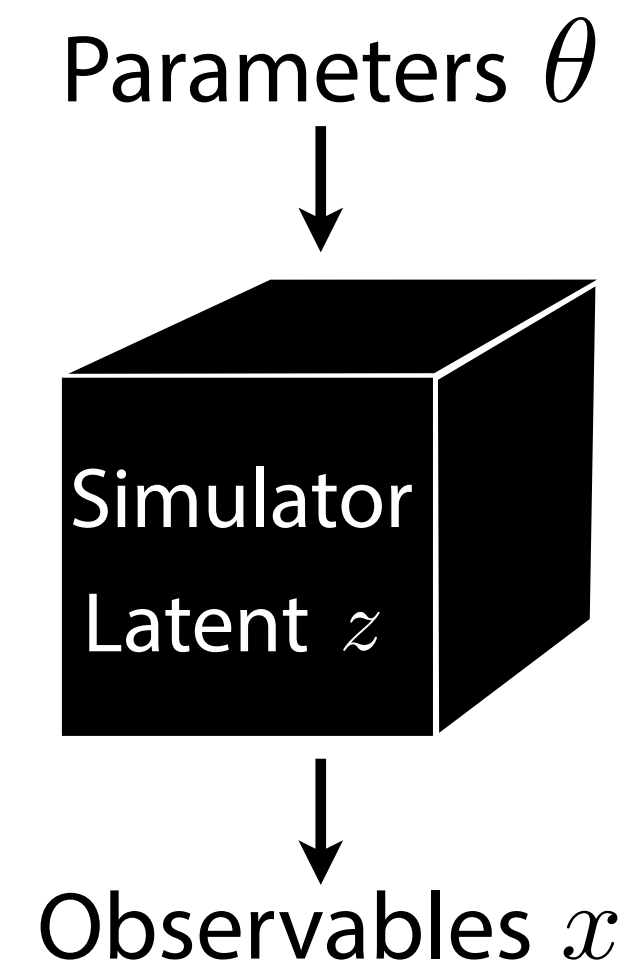
In the neighborhood of  $\theta_{\text{ref}}$  (e.g. close to the SM):

- the **score vector**  $t(x|\theta_{\text{ref}})$  components are sufficient statistics
- knowing  $t(x|\theta_{\text{ref}})$  is just as powerful as knowing the full function  $\log p(x|\theta)$
- $t(x|\theta_{\text{ref}})$  are the most powerful observables

The score itself is intractable. But we can use the same trick as for the likelihood ratio!

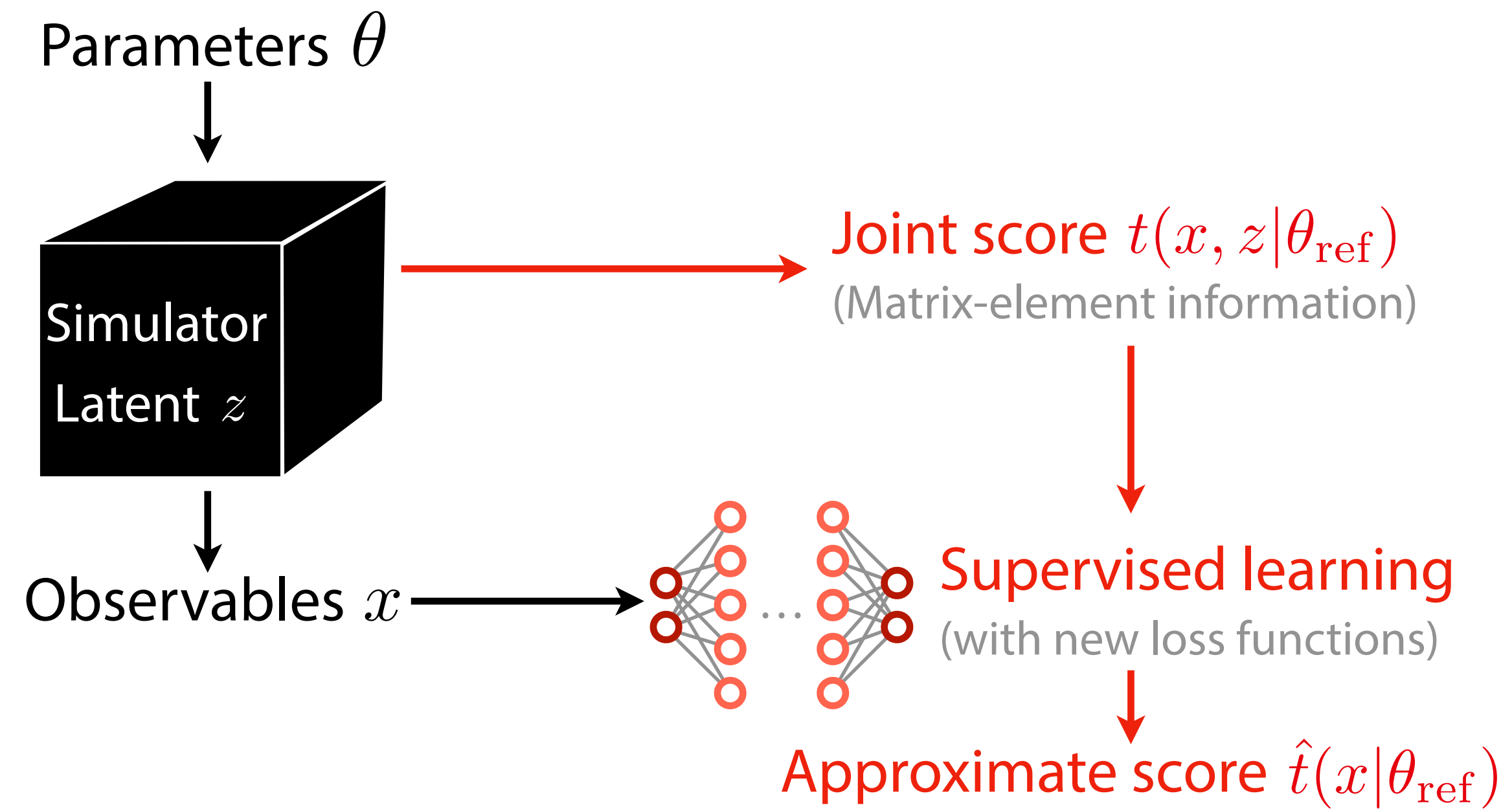


# Neural optimal observables (SALLY)

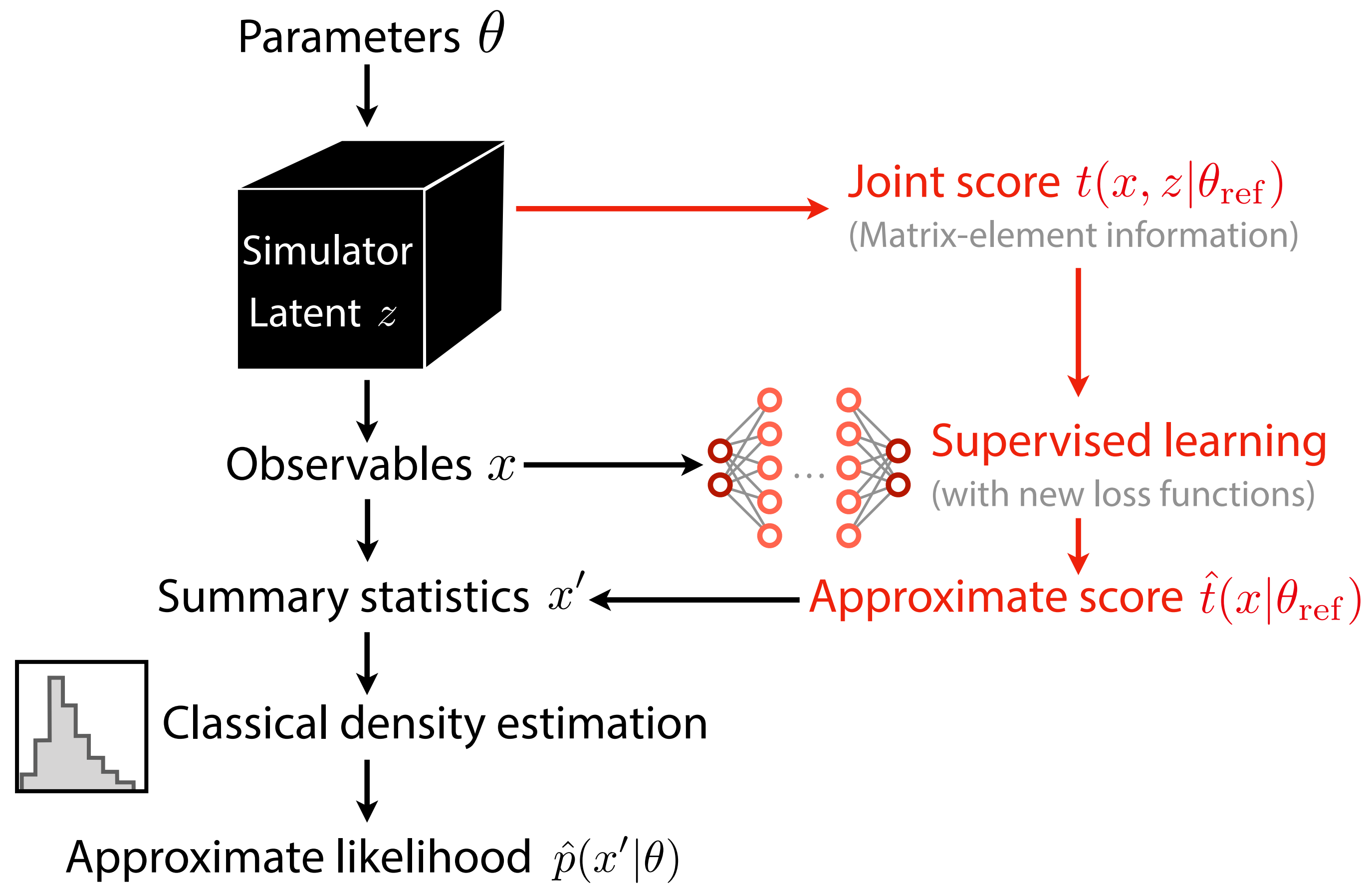




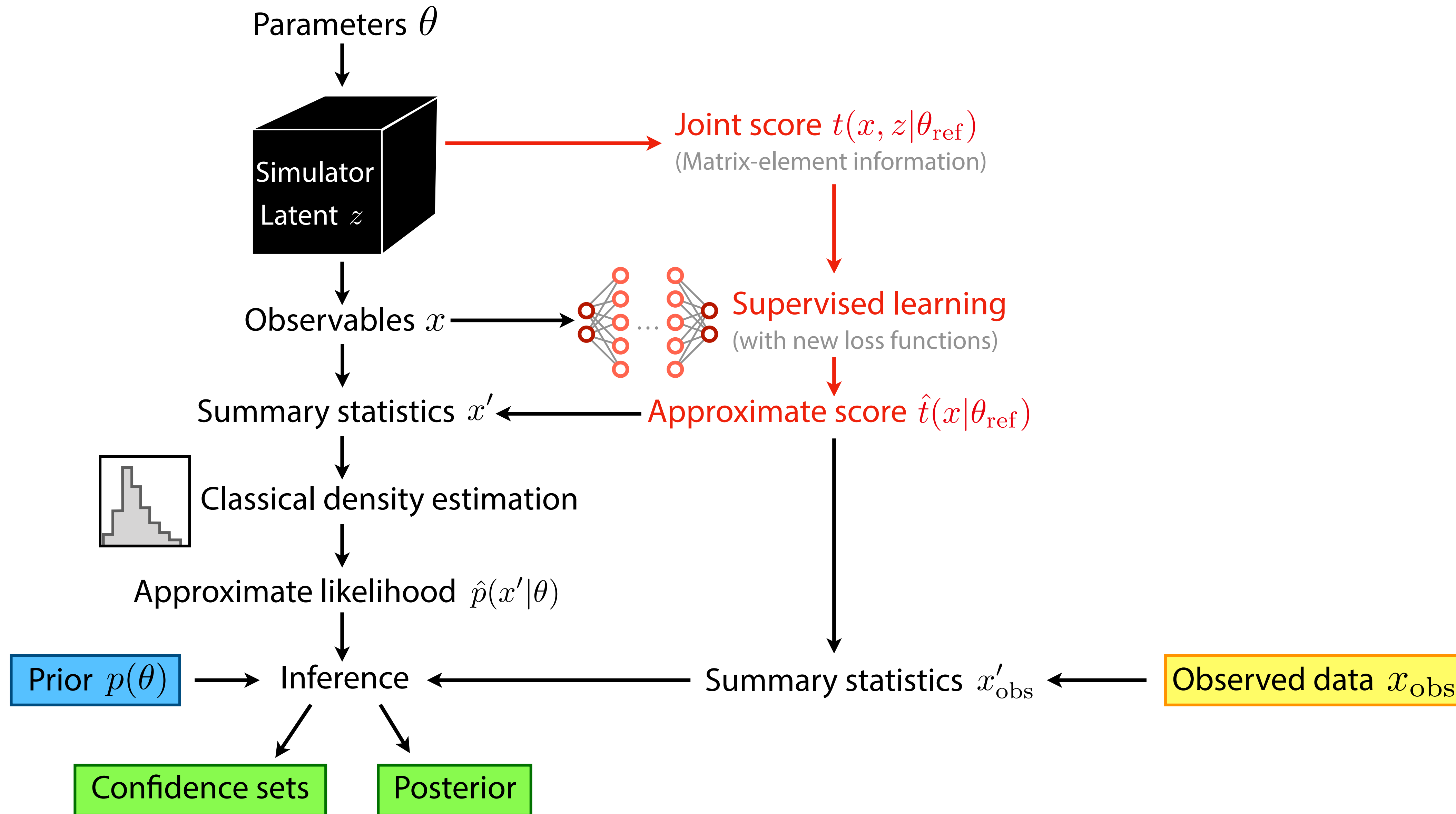
# Neural optimal observables (SALLY)



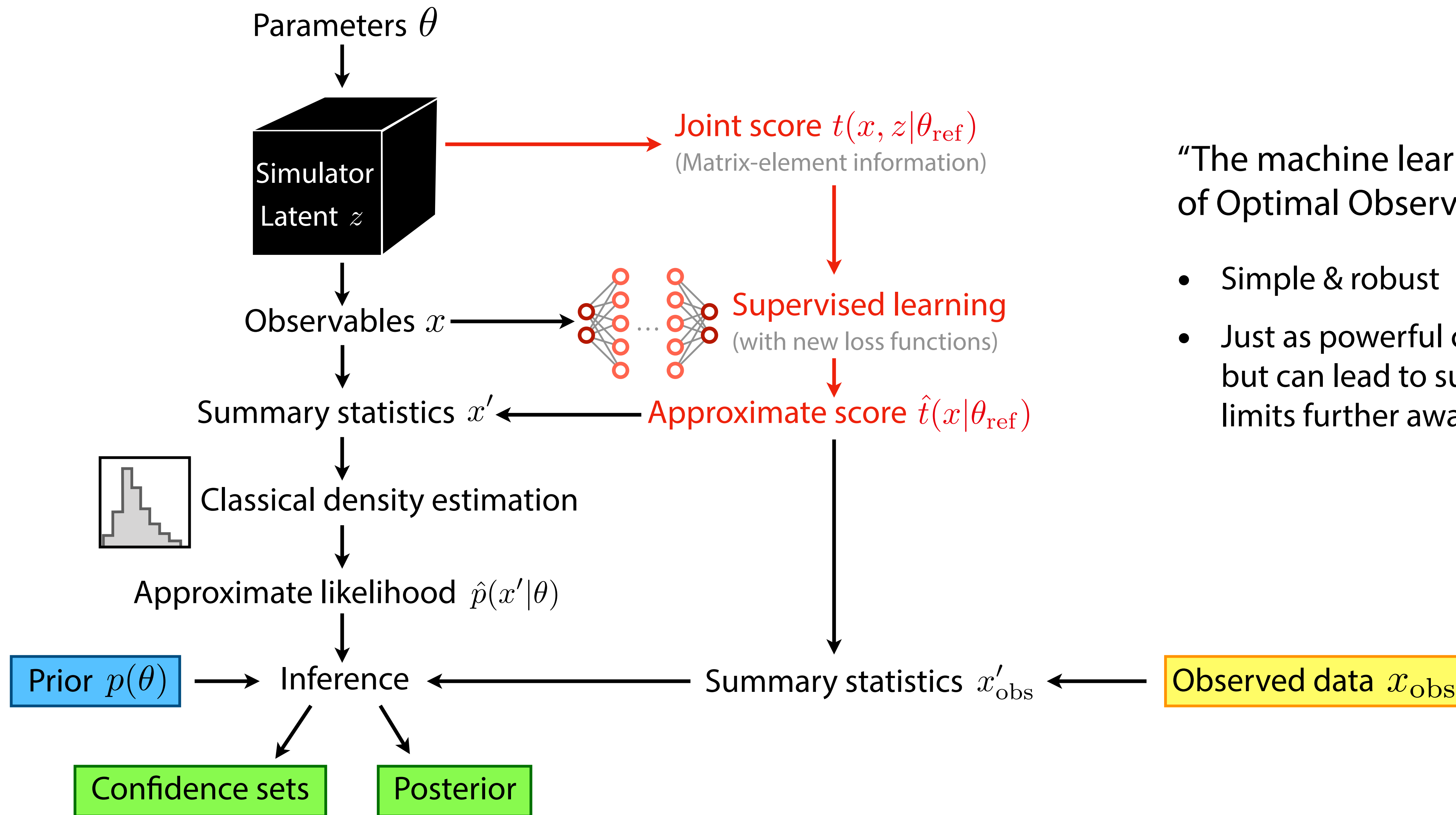
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# Neural optimal observables (SALLY)



“The machine learning version of Optimal Observables”

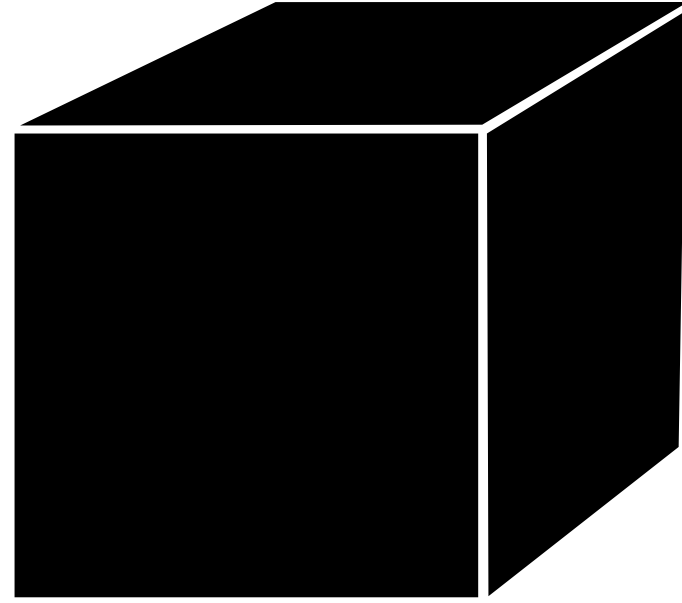
- Simple & robust
- Just as powerful close to  $\theta_{\text{ref}}$ , but can lead to suboptimal limits further away



# An incomplete wrap-up of simulation-based inference methods

Method	Approximations	Upfront cost	Eval
Summary statistics:			
Likelihood for summary stats (standard histograms)	Reduction to summary stats	Fast	Fast
Approximate Bayesian Computation	Reduction to summary stats	Depends	Depends
Matrix elements:			
Matrix Element Method	Transfer fns	Fast	Slow
Optimal Observables	Transfer fns, optimal only locally	Fast	Slow
Neural networks:			
Neural likelihood	NN	Needs many samples	Fast
Neural posterior	NN	Needs many samples	Fast
Neural likelihood ratio	NN	Needs many samples	Fast
Neural networks + matrix elements:			
Neural likelihood (ratio) + gold mining (RASCAL etc)	NN	Needs less samples	Fast
Neural optimal observables (SALLY)	NN, optimal only locally	Needs less samples	Fast

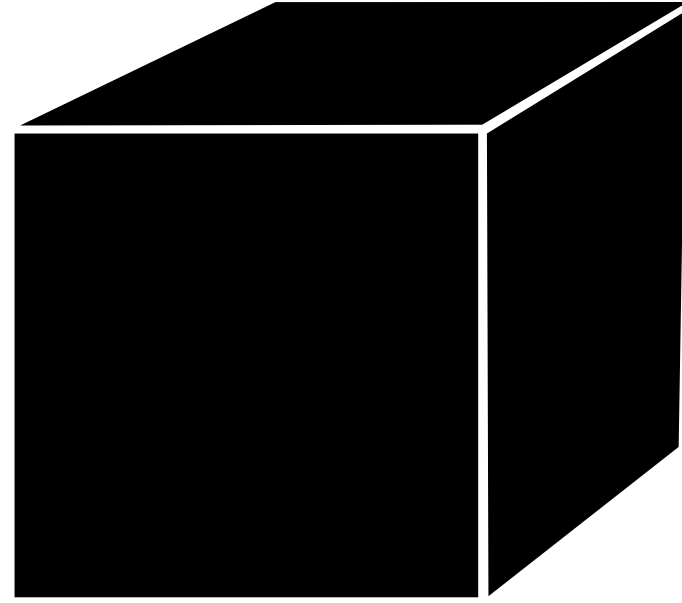
# Systematics



## Don't fully trust the simulator?

- Nuisance parameters to model systematic uncertainties
- Methods learn dependence both on parameters of interest and nuisance parameters. Then we can construct profile likelihood and “nuisance-hardened” score  
[J. Alsing, B. Wandelt 1903.01473;  
see also P. de Castro, T. Dorigo 1806.04743]
- Alternatively: Robustness to nuisance with adversarial training  
[G. Louppe, M. Kagan, K. Cranmer 1611.01046]

# Systematics



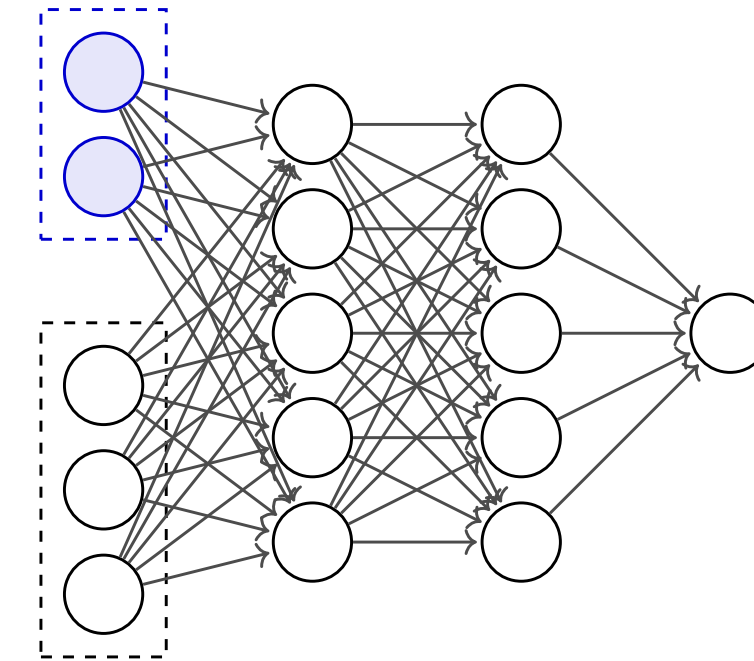
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[G. Louppe, M. Kagan, K. Cranmer 1611.01046]



## Don't blindly trust the neural network?

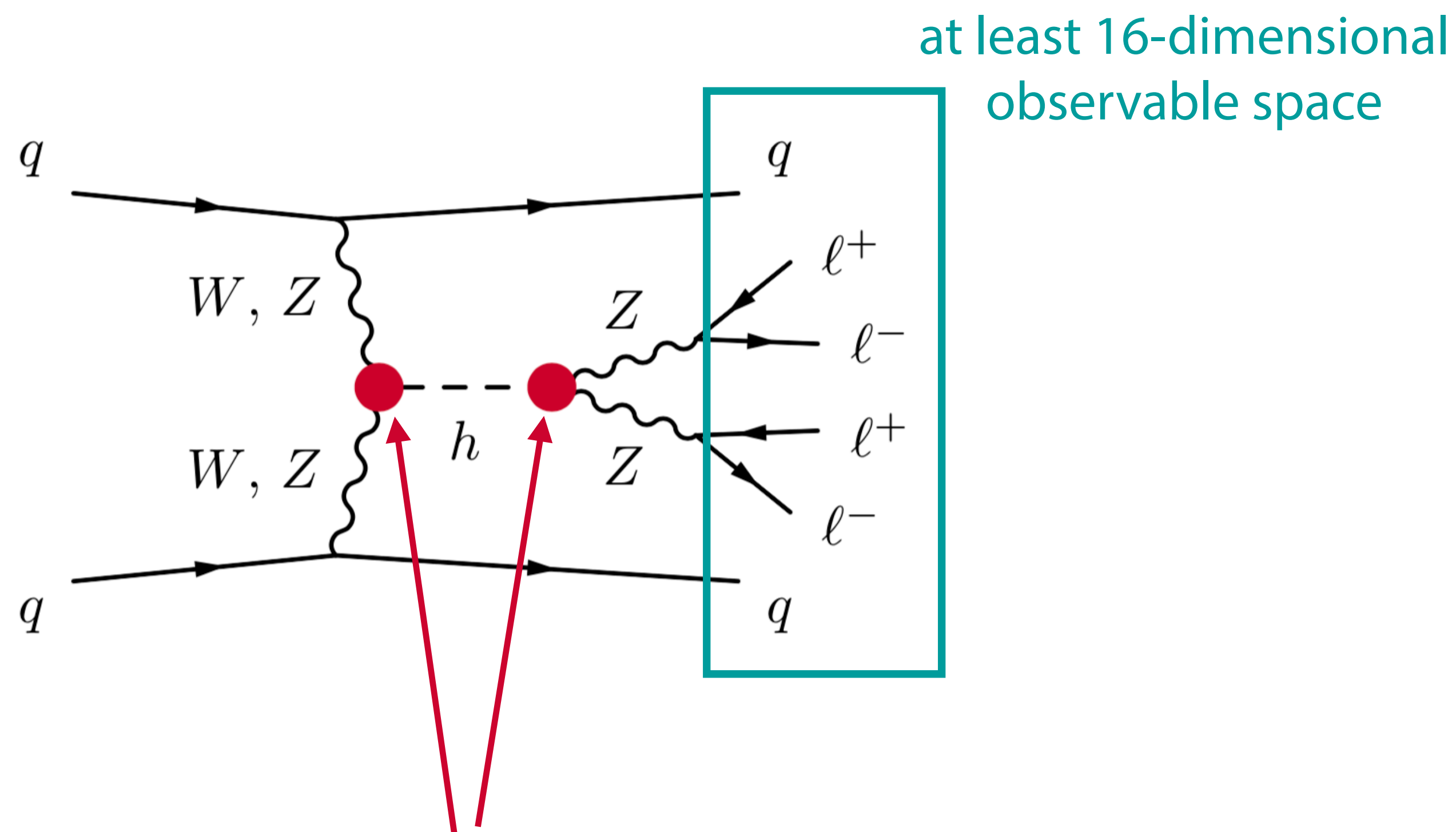
- Sanity checks: expectation values, “critic” tests
- Calibration / Neyman construction with toys  
(badly trained network can lead to suboptimal limits, but not to wrong limits)

These techniques let us constrain effective theories more effectively.



# Proof of concept: Higgs production in weak boson fusion

[JB, K. Cranmer, G. Louppe, J. Pavez  
1805.00013, 1805.00020]



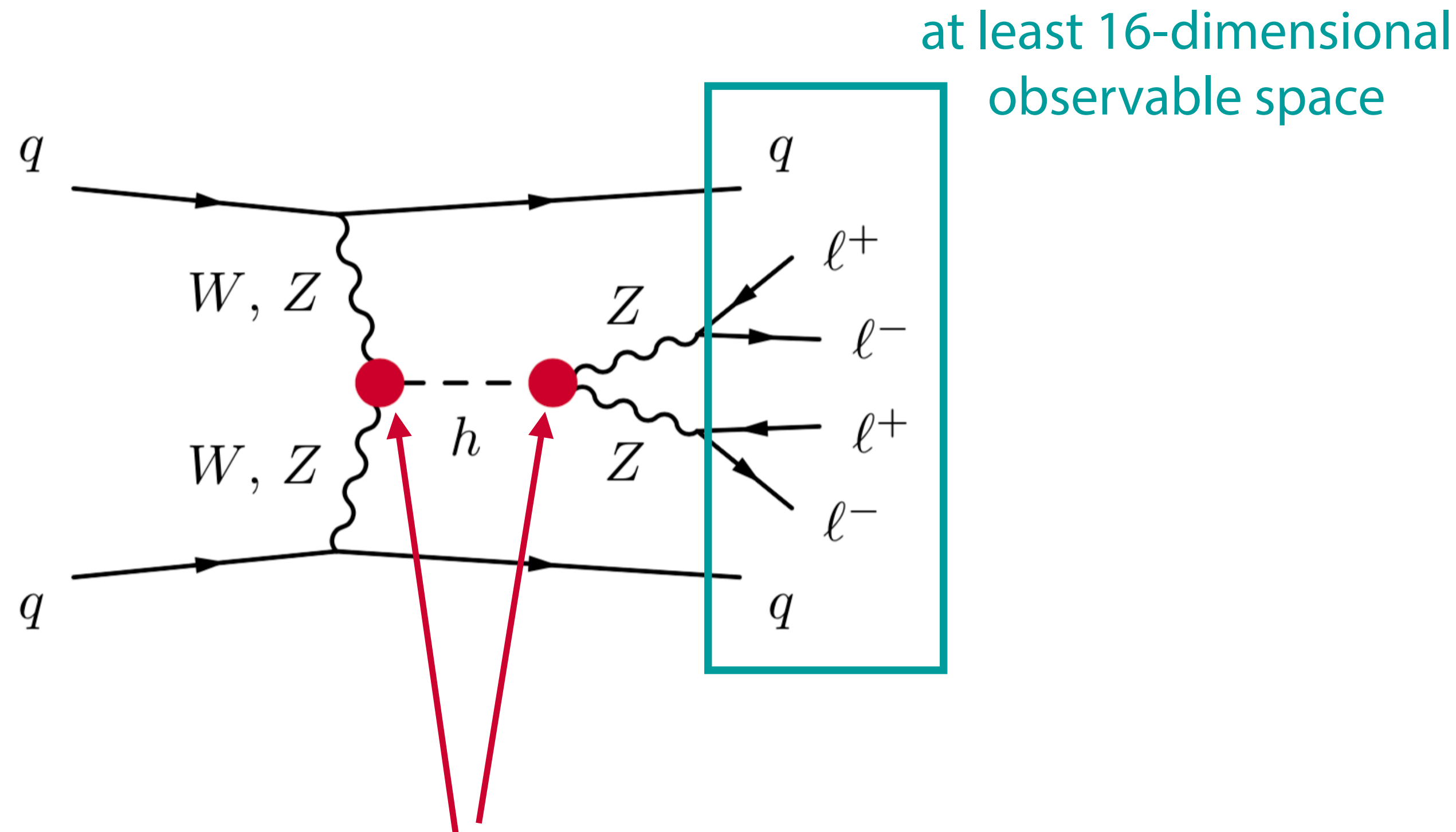
Exciting new physics might hide here!

We parameterize it with two EFT coefficients:

$$\mathcal{L} = \mathcal{L}_{\text{SM}} + \boxed{\frac{f_W}{\Lambda^2}} \underbrace{\frac{ig}{2} (D^\mu \phi)^\dagger \sigma^a D^\nu \phi W_{\mu\nu}^a}_{\mathcal{O}_W} - \boxed{\frac{f_{WW}}{\Lambda^2}} \underbrace{\frac{g^2}{4} (\phi^\dagger \phi) W_{\mu\nu}^a W^{\mu\nu a}}_{\mathcal{O}_{WW}}$$

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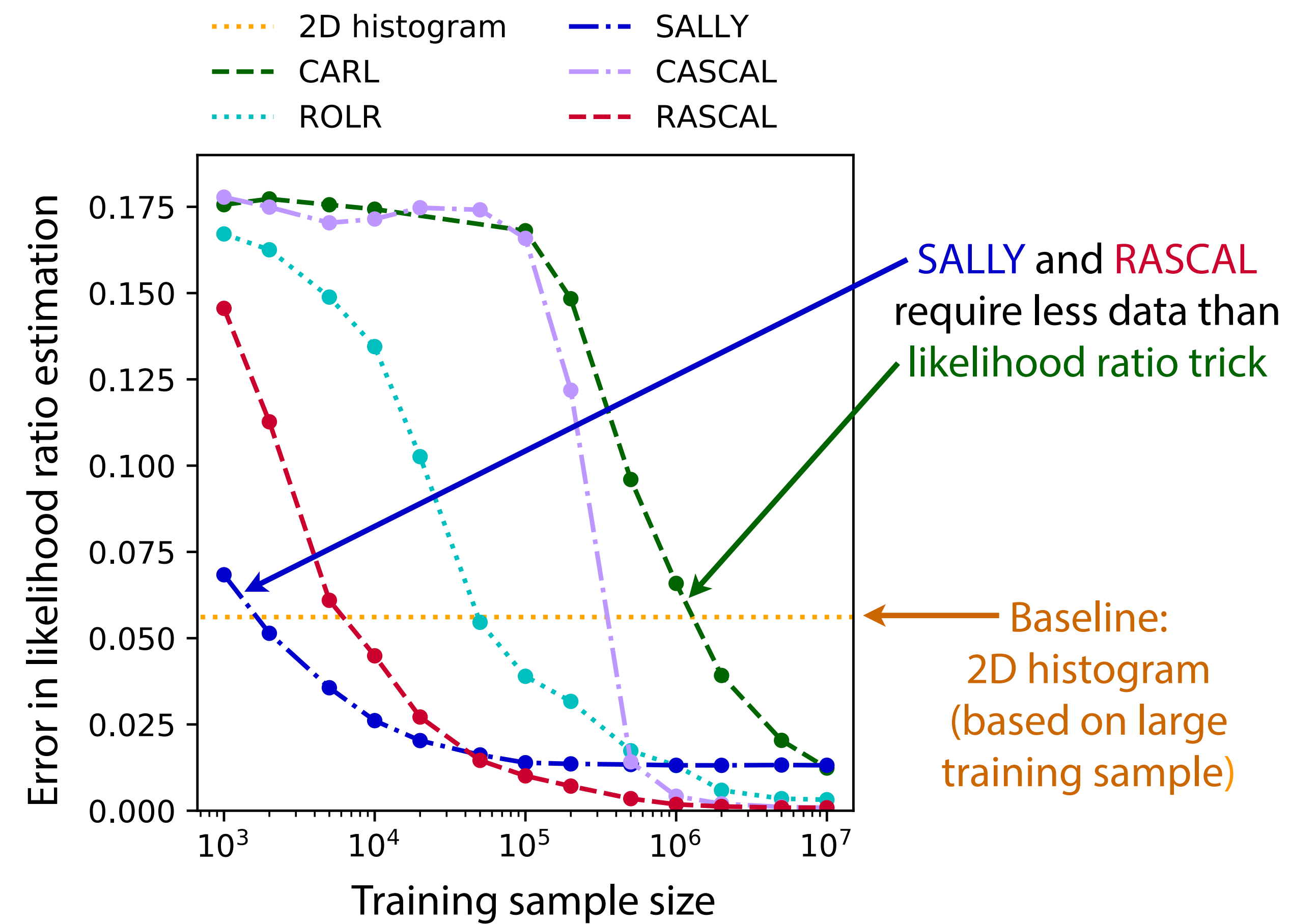
Goal: constrain the **two EFT parameters**

- new inference methods
- baseline: 2d histogram analysis of **jet momenta & angular correlations**

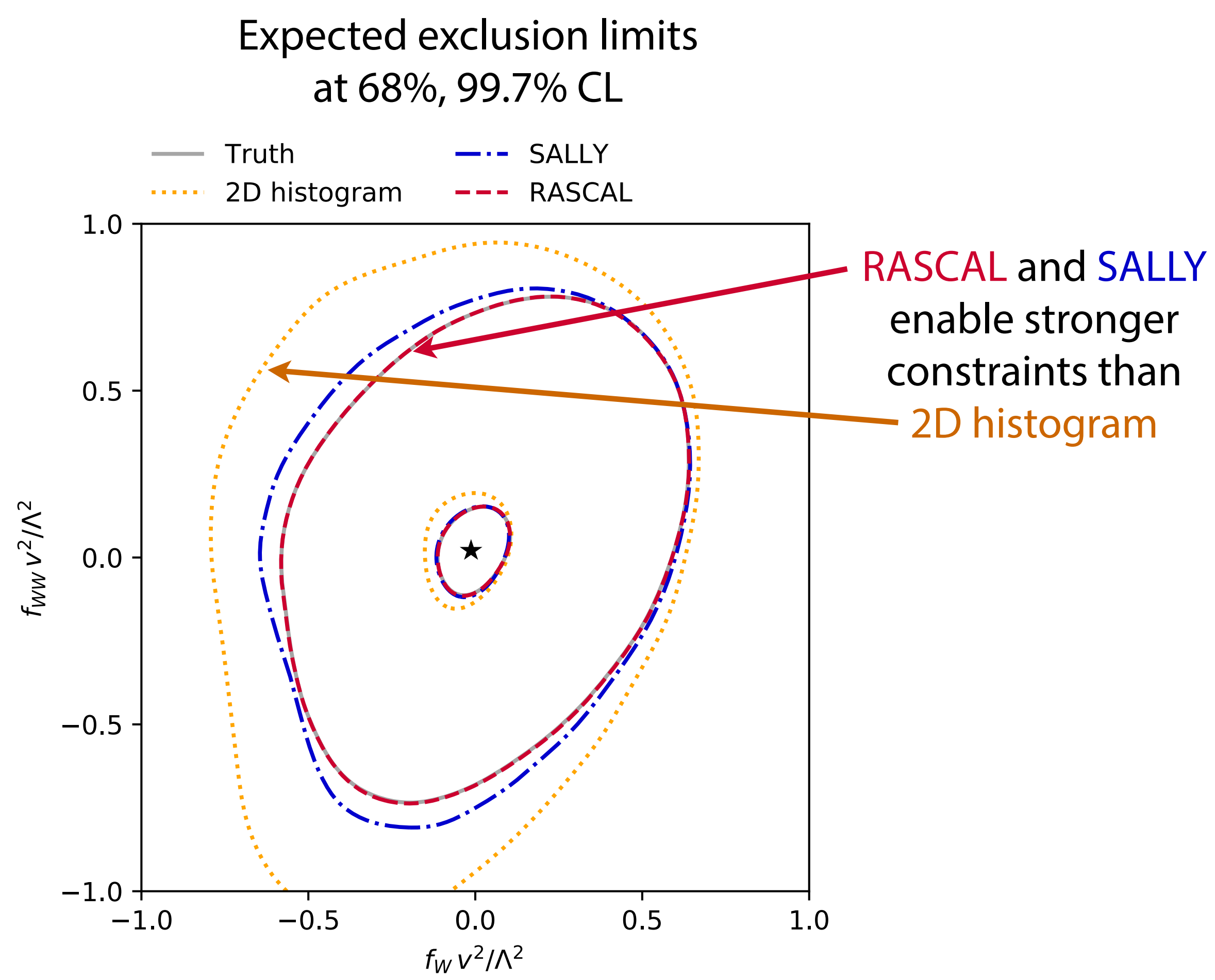
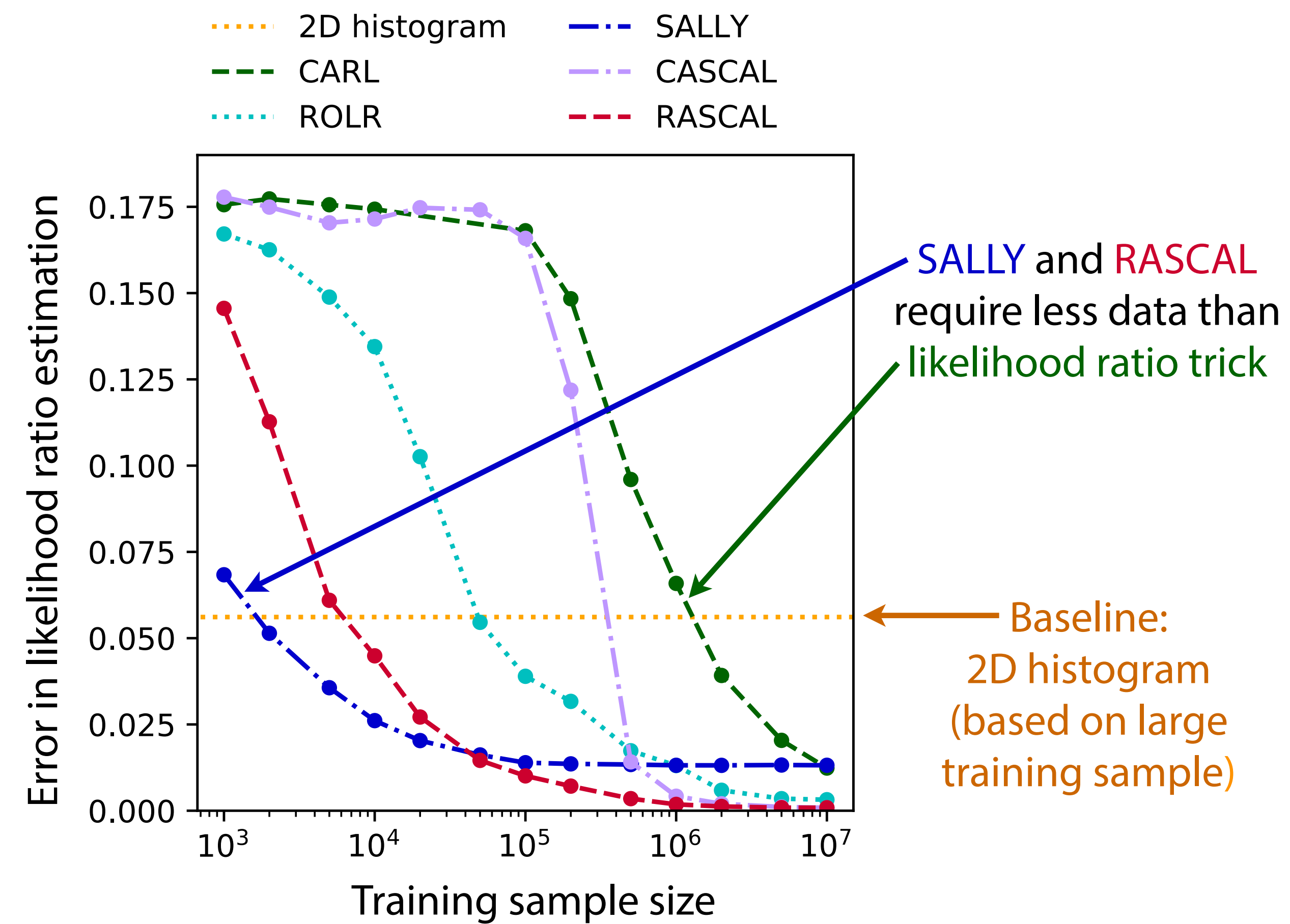
Two scenarios:

- Simplified setup in which we can compare to true likelihood
- "Realistic" simulation with approximate detector effects

# Proof of concept: Stronger constraints with less training data

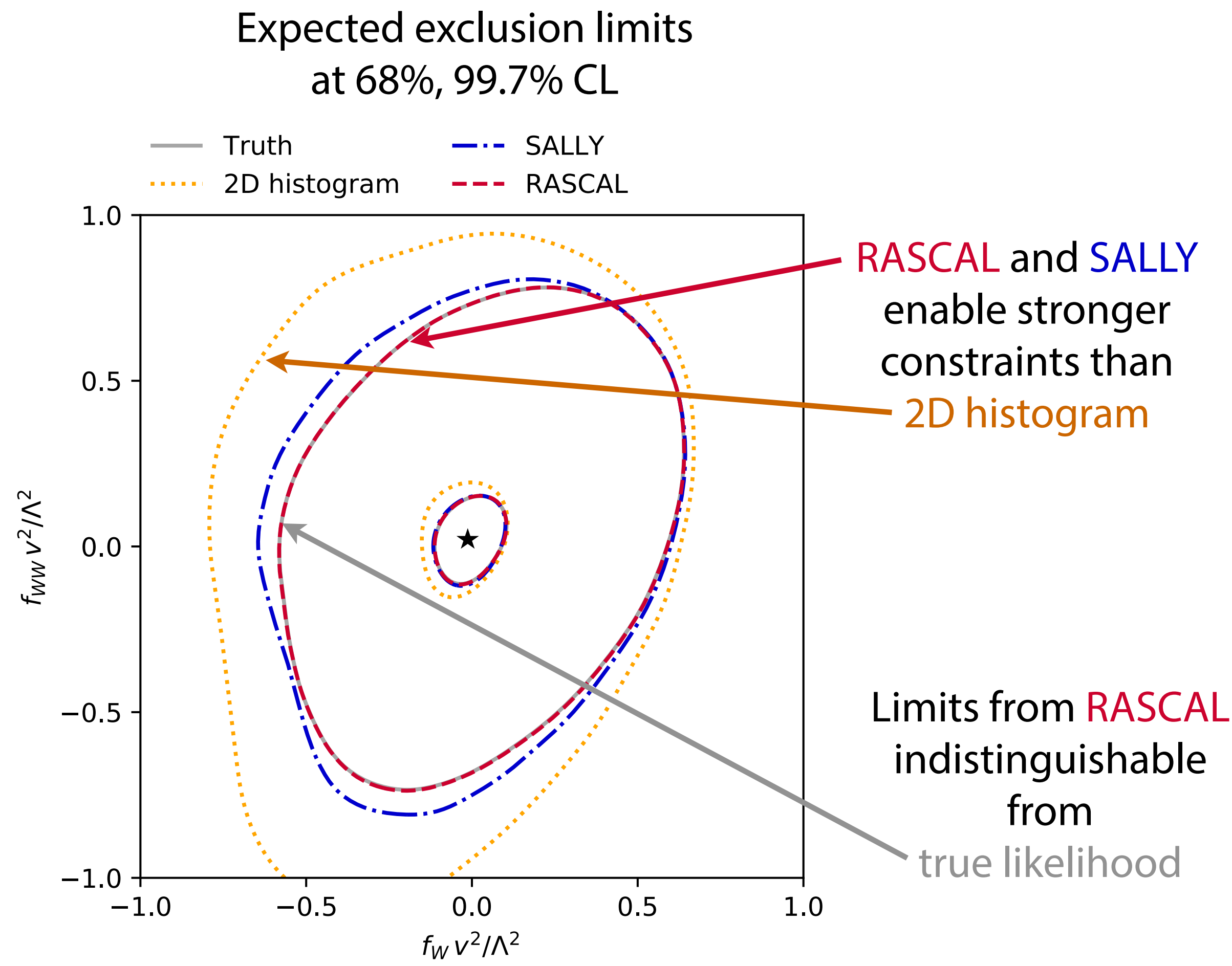
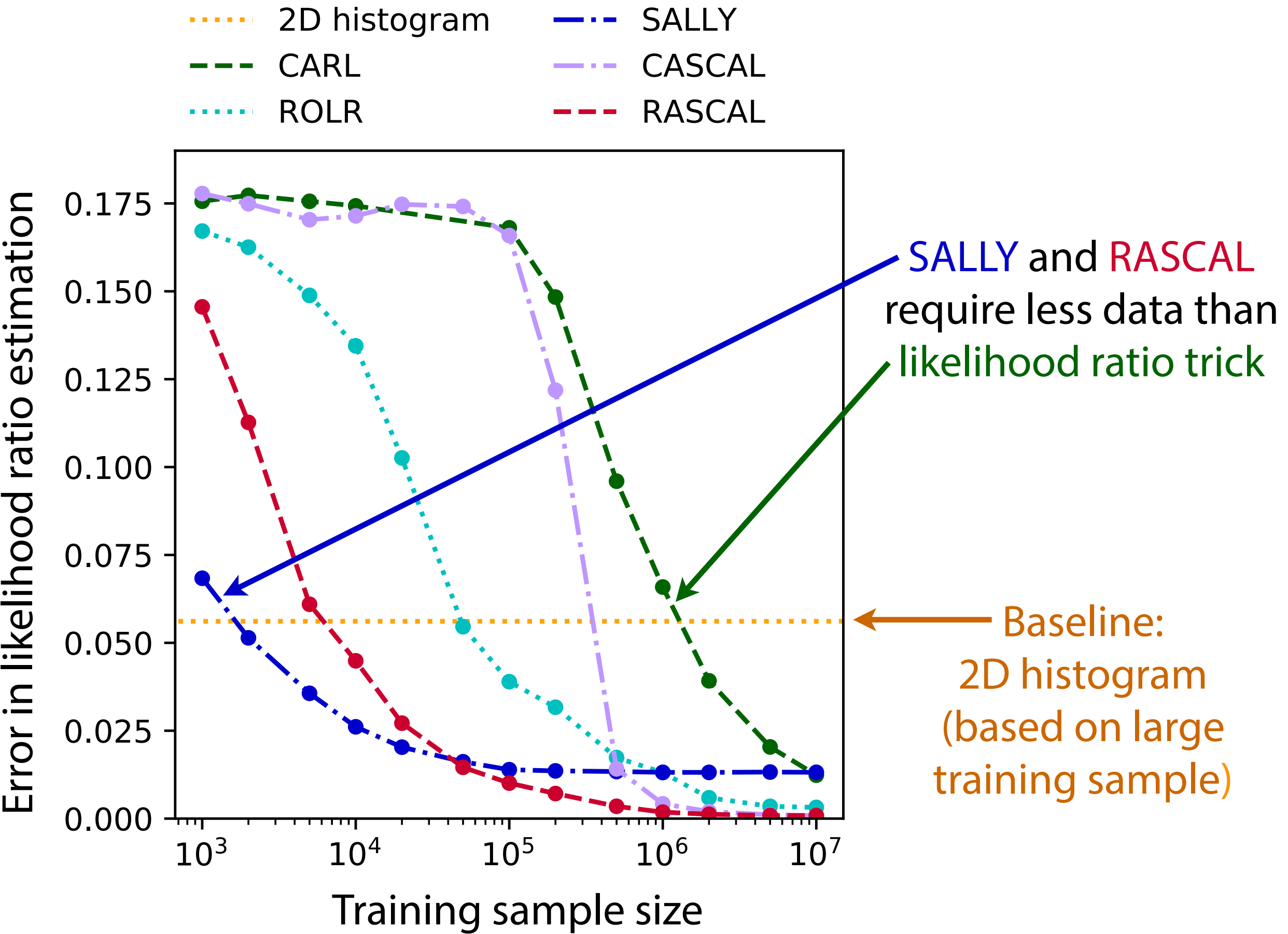


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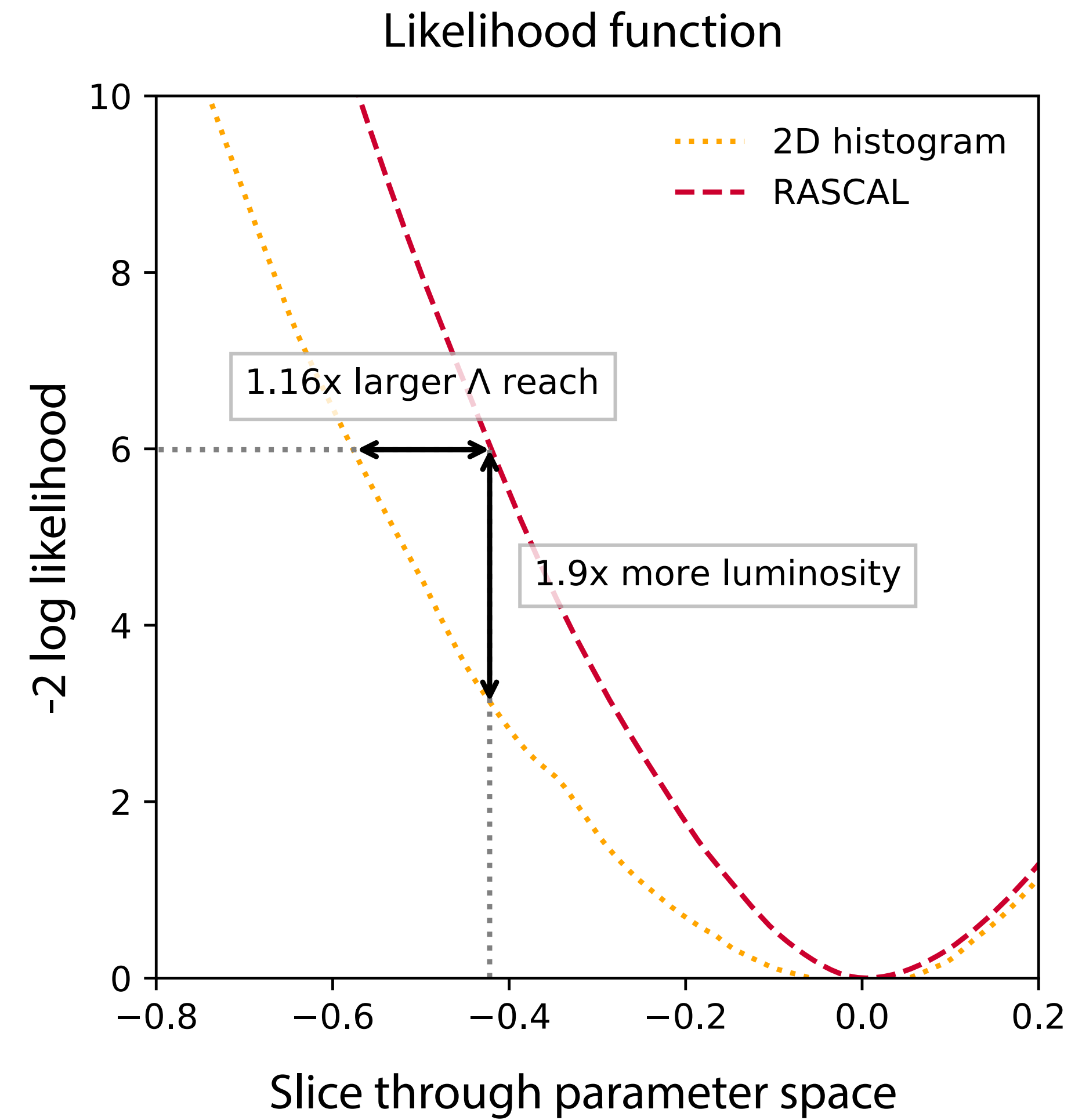
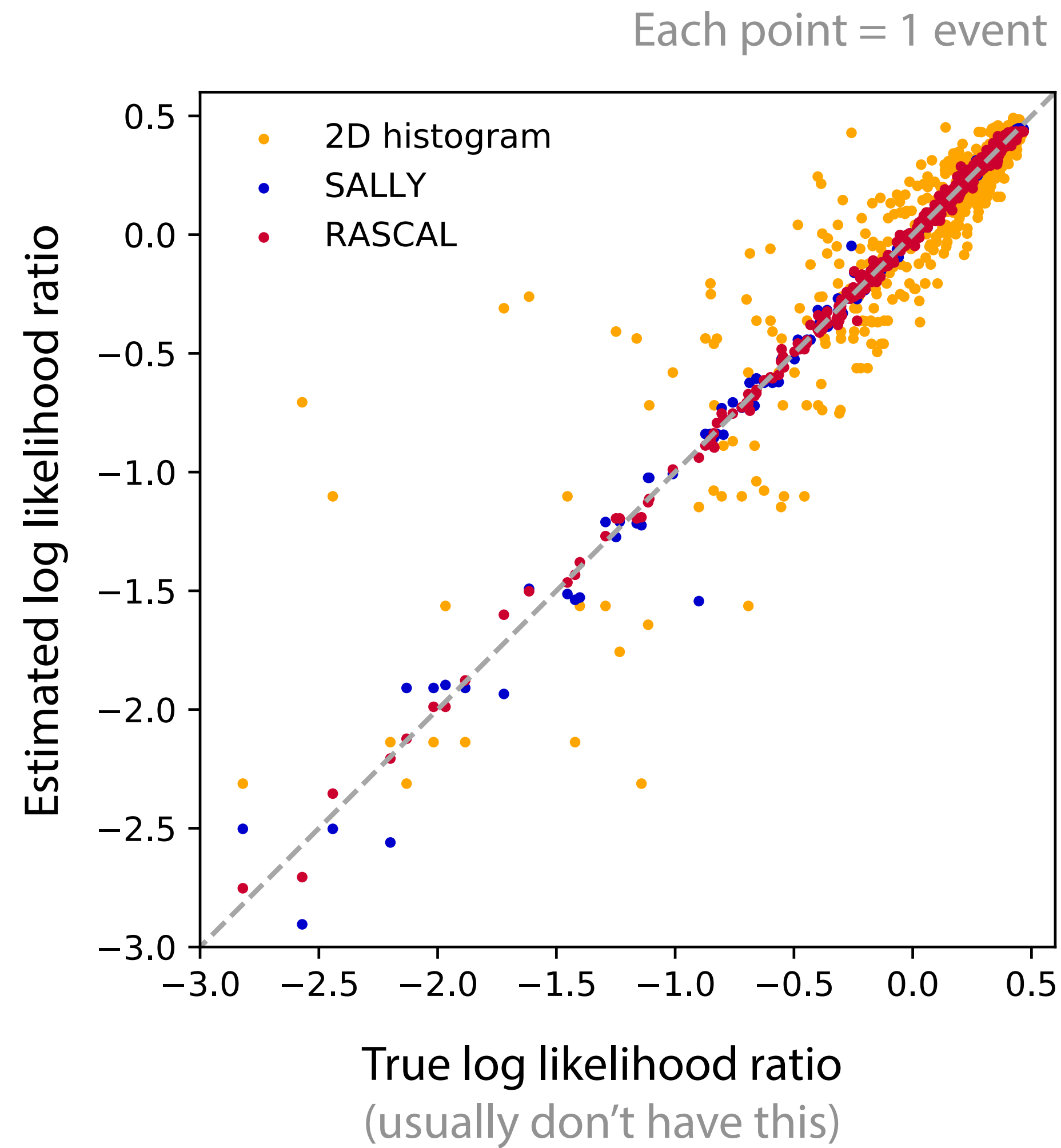




# Proof of concept: Stronger constraints with less training data



# Proof of concept: 16% greater reach ( $\sim 90\%$ more data)



# Constraining operators in ttH effectively

[JB, F. Kling, I. Espejo, K. Cranmer 1907.10621]

- Pheno-level analysis of

$$pp \rightarrow t\bar{t}h \rightarrow (bl^+) (\bar{b}l^-) (\gamma\gamma) E_T^{\text{miss}}$$

with MadGraph + Pythia + Delphes

- Inference on three EFT operators:

$$\mathcal{O}_u = -\frac{1}{v^2}(H^\dagger H)(H^\dagger \bar{Q}_L)u_R, \quad \mathcal{O}_G = \frac{g_s^2}{m_W^2}(H^\dagger H)G_{\mu\nu}^a G_a^{\mu\nu},$$

$$\mathcal{O}_{uG} = -\frac{4g_s}{m_W^2}y_u(H^\dagger \bar{Q}_L)\gamma^{\mu\nu}T_a u_R G_{\mu\nu}^a$$

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[JB, F. Kling, I. Espejo, K. Cranmer 1907.10621]

- Pheno-level analysis of

$$pp \rightarrow t\bar{t}h \rightarrow (bl^+) (\bar{b}l^-) (\gamma\gamma) E_T^{\text{miss}}$$

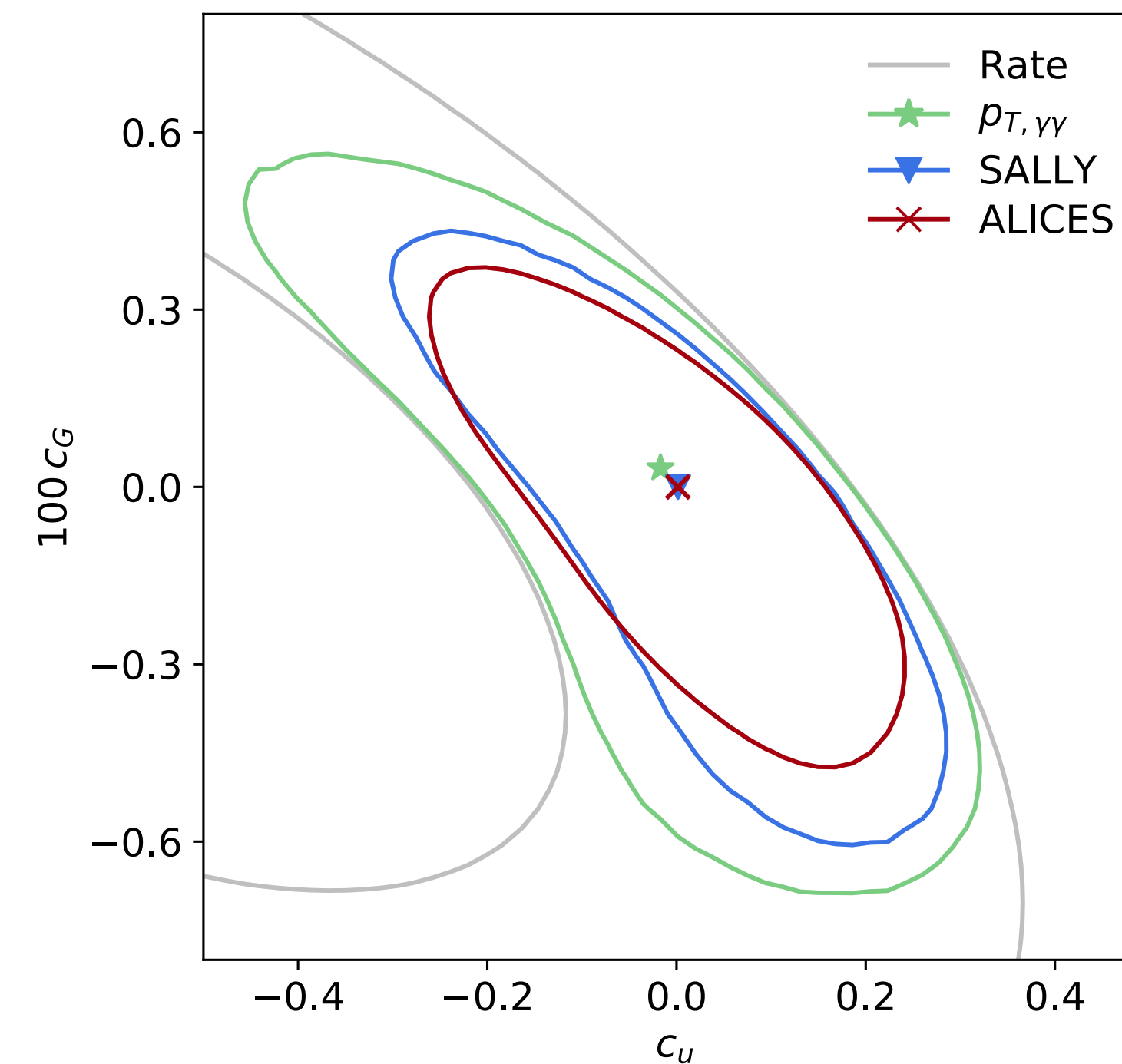
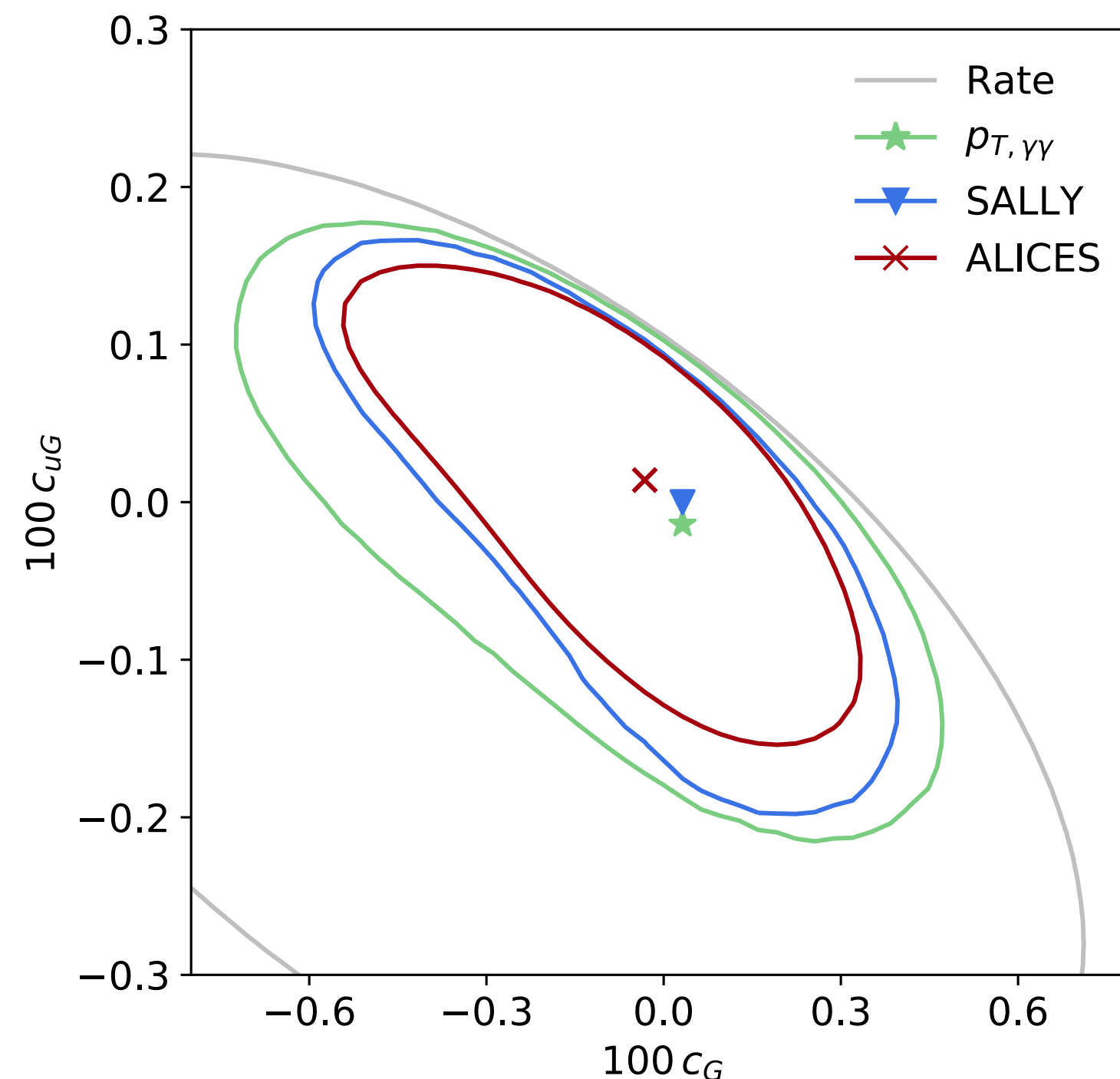
with MadGraph + Pythia + Delphes

- Inference on three EFT operators:

$$\mathcal{O}_u = -\frac{1}{v^2}(H^\dagger H)(H^\dagger \bar{Q}_L)u_R, \quad \mathcal{O}_G = \frac{g_s^2}{m_W^2}(H^\dagger H)G_{\mu\nu}^a G_a^{\mu\nu},$$

$$\mathcal{O}_{uG} = -\frac{4g_s}{m_W^2}y_u(H^\dagger \bar{Q}_L)\gamma^{\mu\nu}T_a u_R G_{\mu\nu}^a$$

- New **inference techniques** improve expected HL-LHC limits compared to **histogram baseline**:



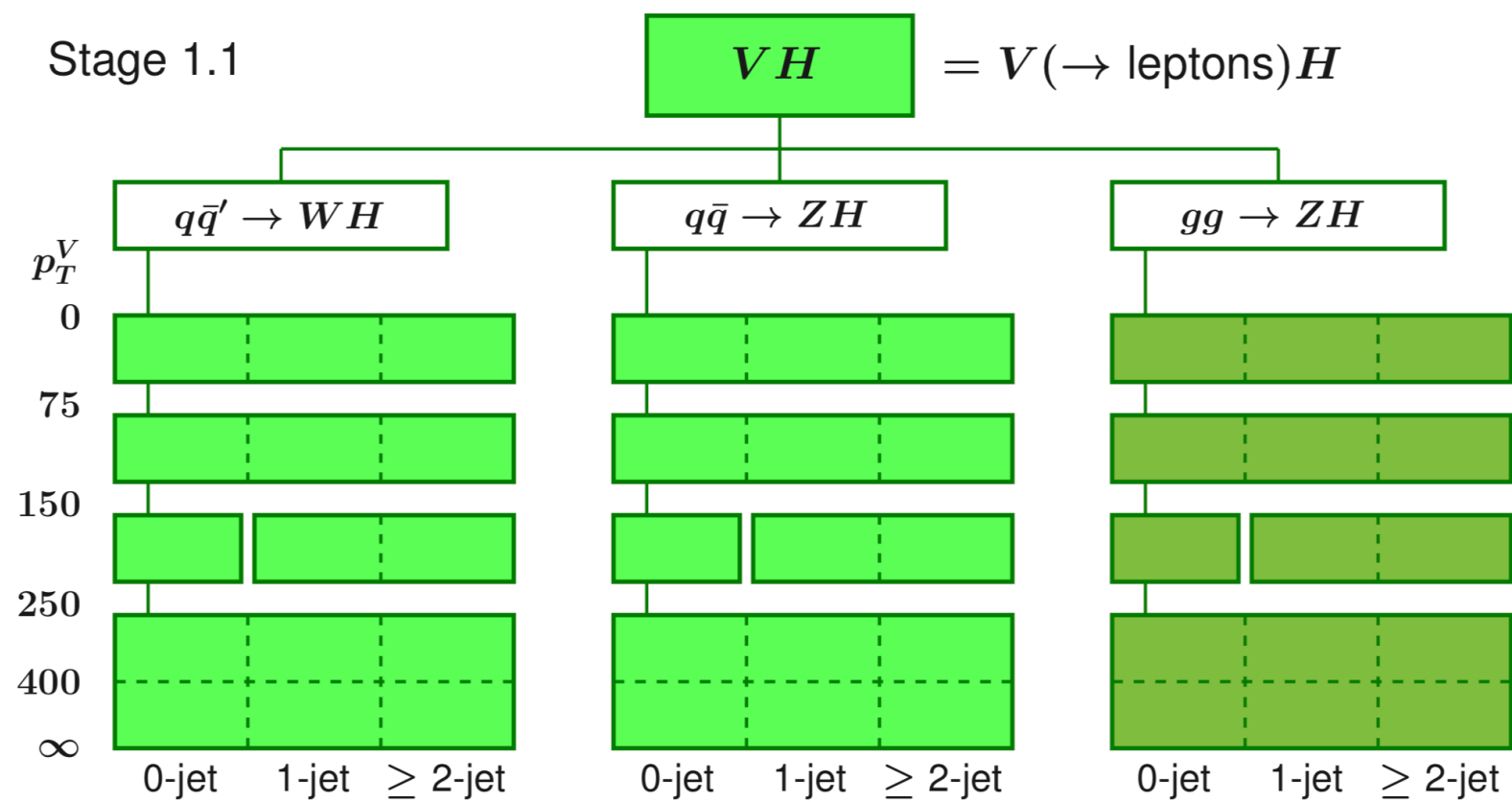


# Benchmarking STXS in WH

[JB, S. Dawson, S. Homiller, F. Kling, T. Plehn 1908.06980]

- Simplified Template Cross-Sections (STXS) define observable bins that are supposed to capture as much information on NP as possible

[N. Berger et al. 1906.02754; HXSWG YR4]



- Let's check! How much information on

$$\tilde{\mathcal{O}}_{HD} = \mathcal{O}_{H\Box} - \frac{\mathcal{O}_{HD}}{4} = (\phi^\dagger \phi) \Box (\phi^\dagger \phi) - \frac{1}{4} (\phi^\dagger D^\mu \phi)^* (\phi^\dagger D_\mu \phi)$$

$$\mathcal{O}_{HW} = \phi^\dagger \phi W_{\mu\nu}^a W^{\mu\nu a}$$

$$\mathcal{O}_{Hq}^{(3)} = (\phi^\dagger i \overleftrightarrow{D}_\mu^a \phi) (\bar{Q}_L \sigma^a \gamma^\mu Q_L),$$

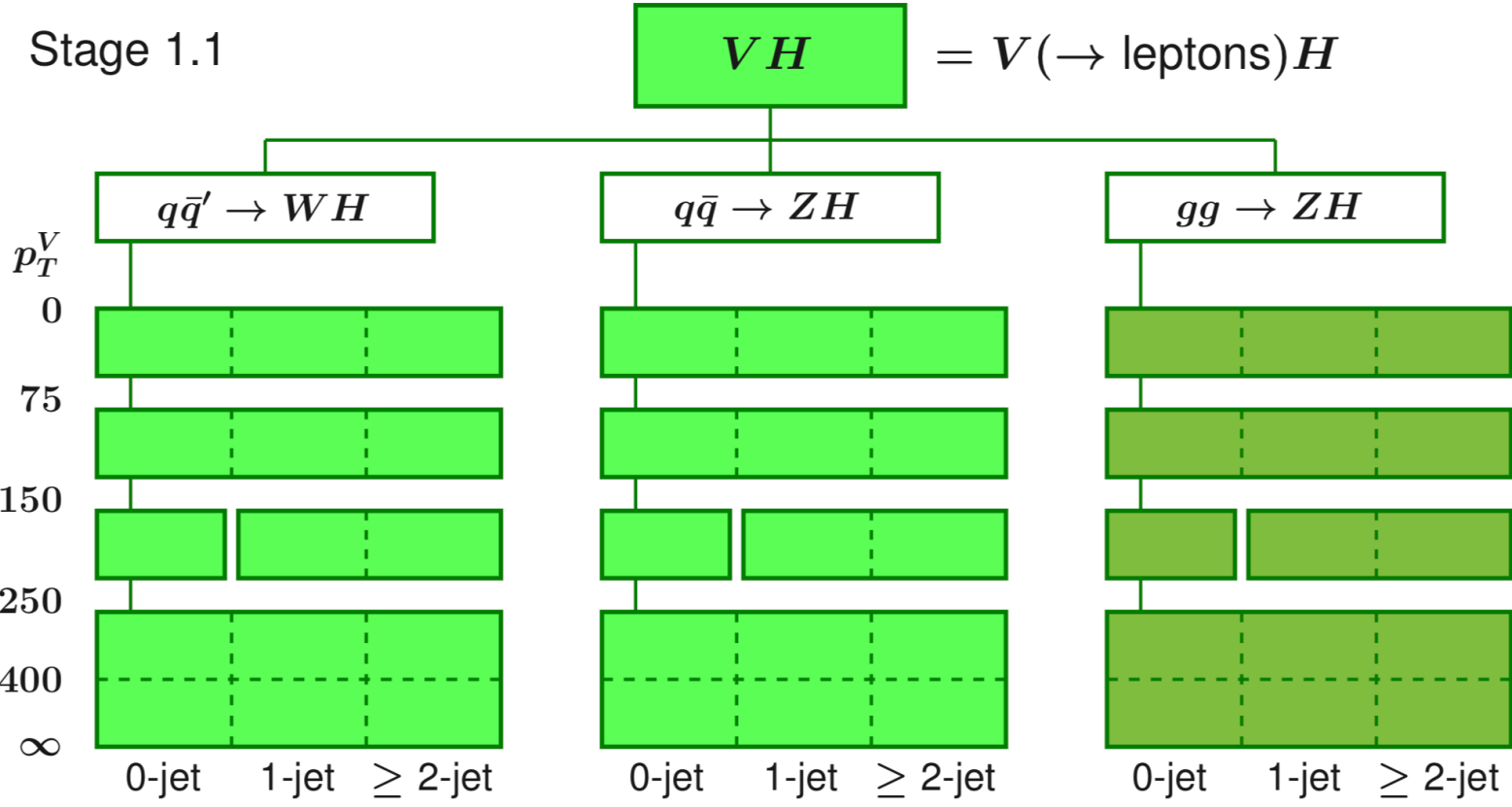
can we extract from  $pp \rightarrow WH \rightarrow \ell\nu b\bar{b}$  ?

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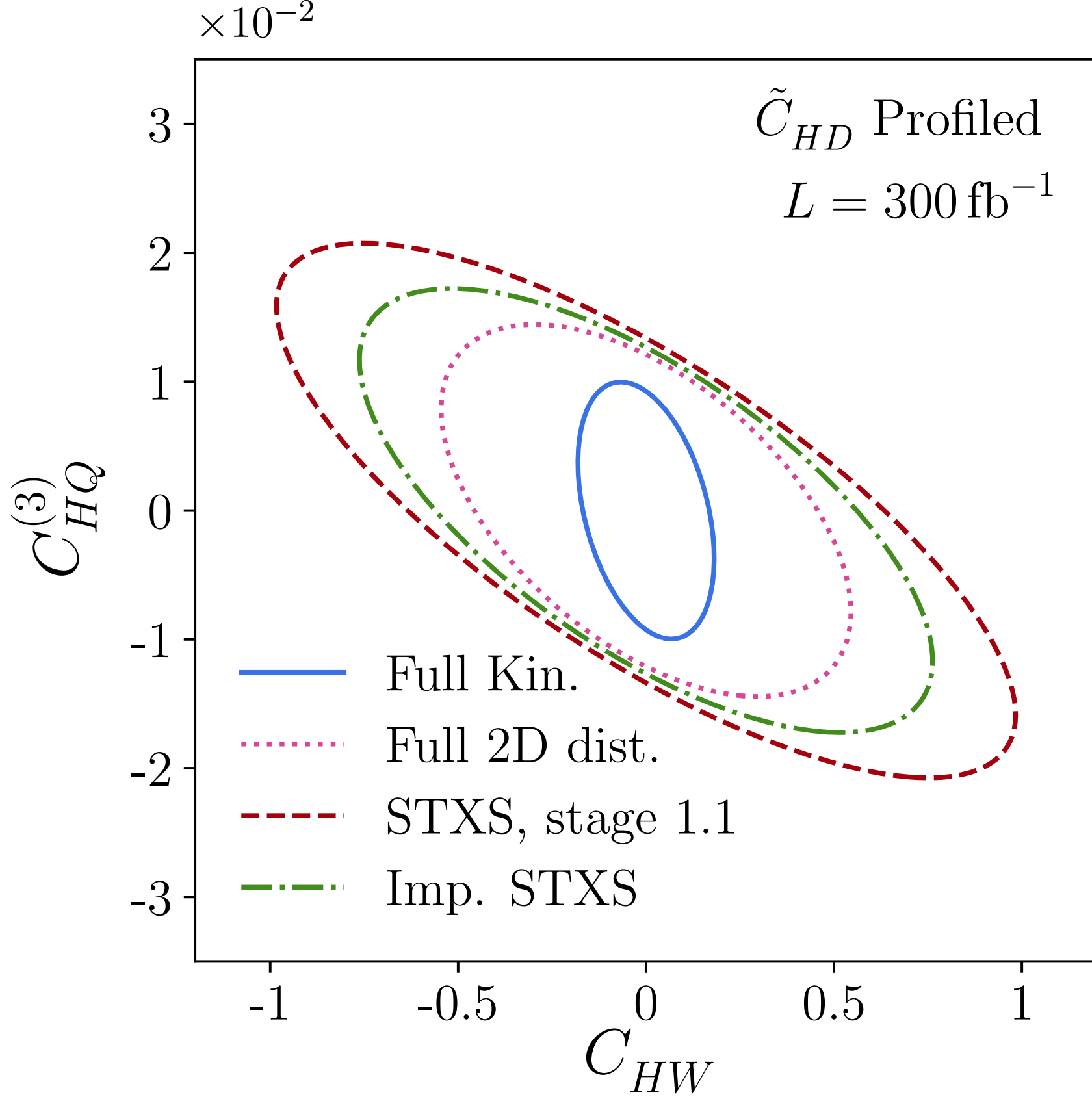
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- Results: **STXS** are indeed sensitive to operators, **adding a few more bins** improve them, but **a multivariate analysis** is still stronger

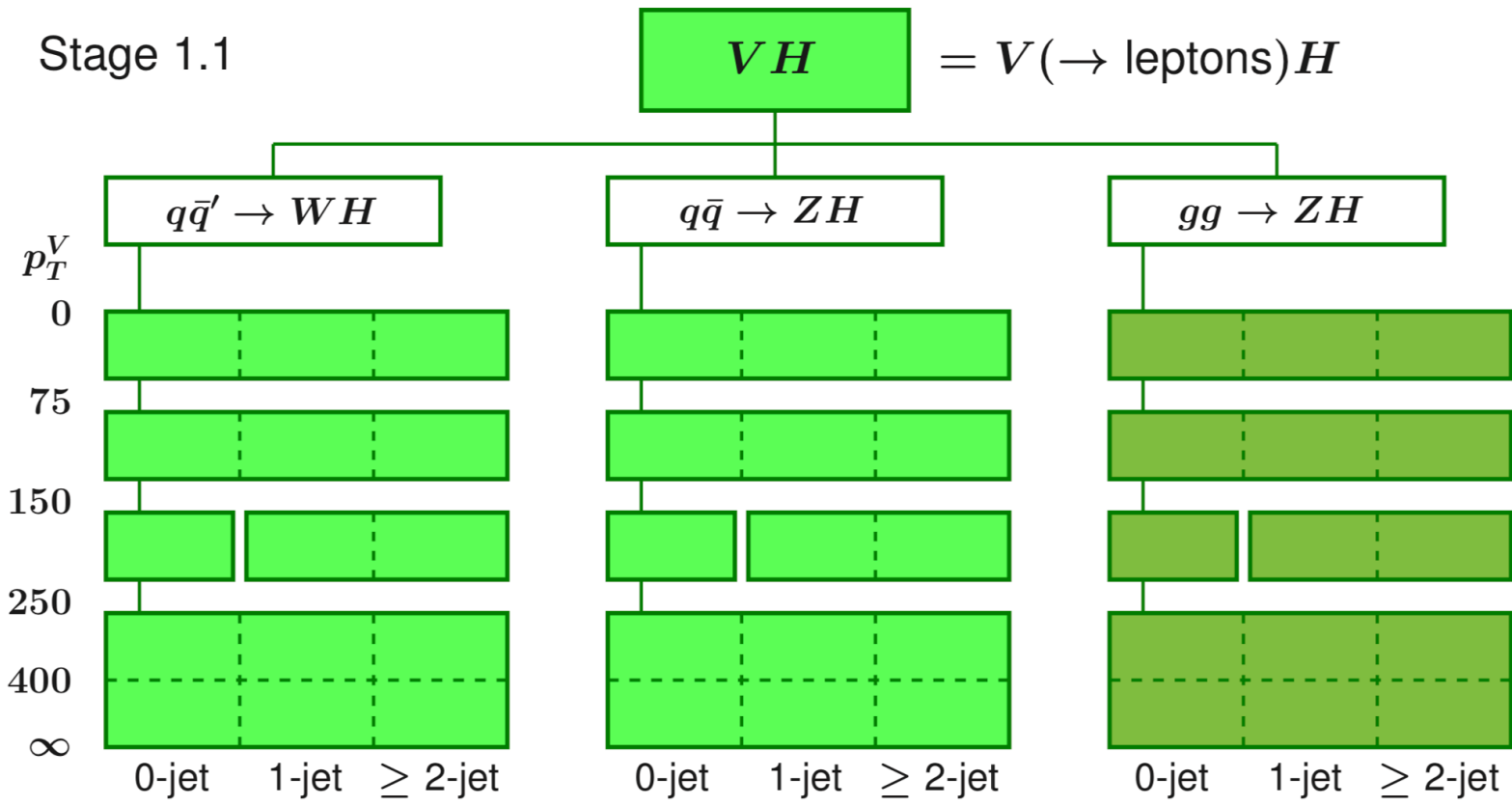


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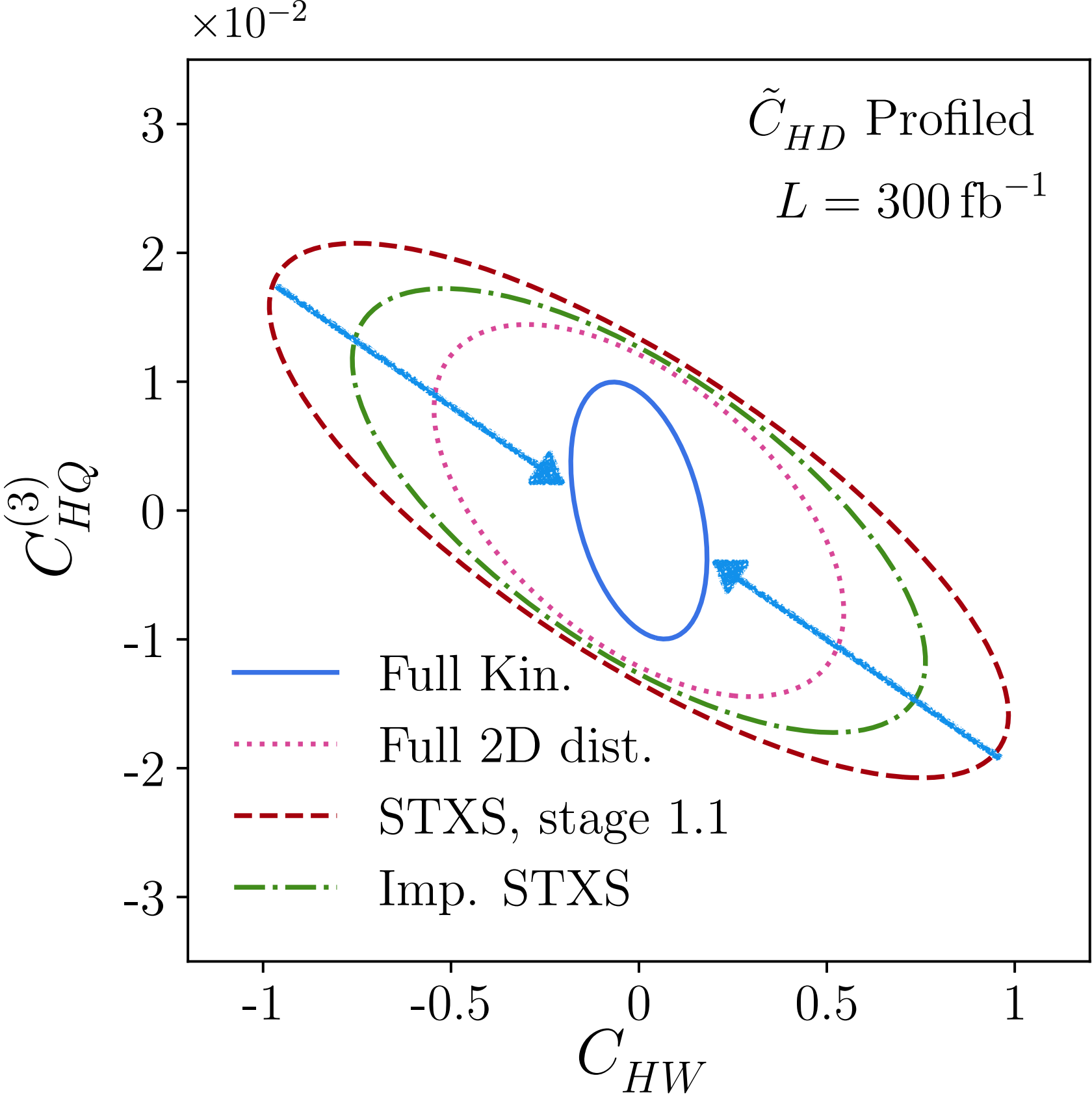
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can we extract from  $pp \rightarrow WH \rightarrow \ell\nu b\bar{b}$  ?



# Diboson production

- In inclusive observables, the interference between SM and new physics amplitudes vanishes

⇒ Reduced sensitivity to new physics

- “Diboson interference resurrection”:  
an **angular variable**  $\varphi$  can be constructed to be sensitive to this interference

[G. Panico, F. Riva, A. Wulzer 1708.07823;

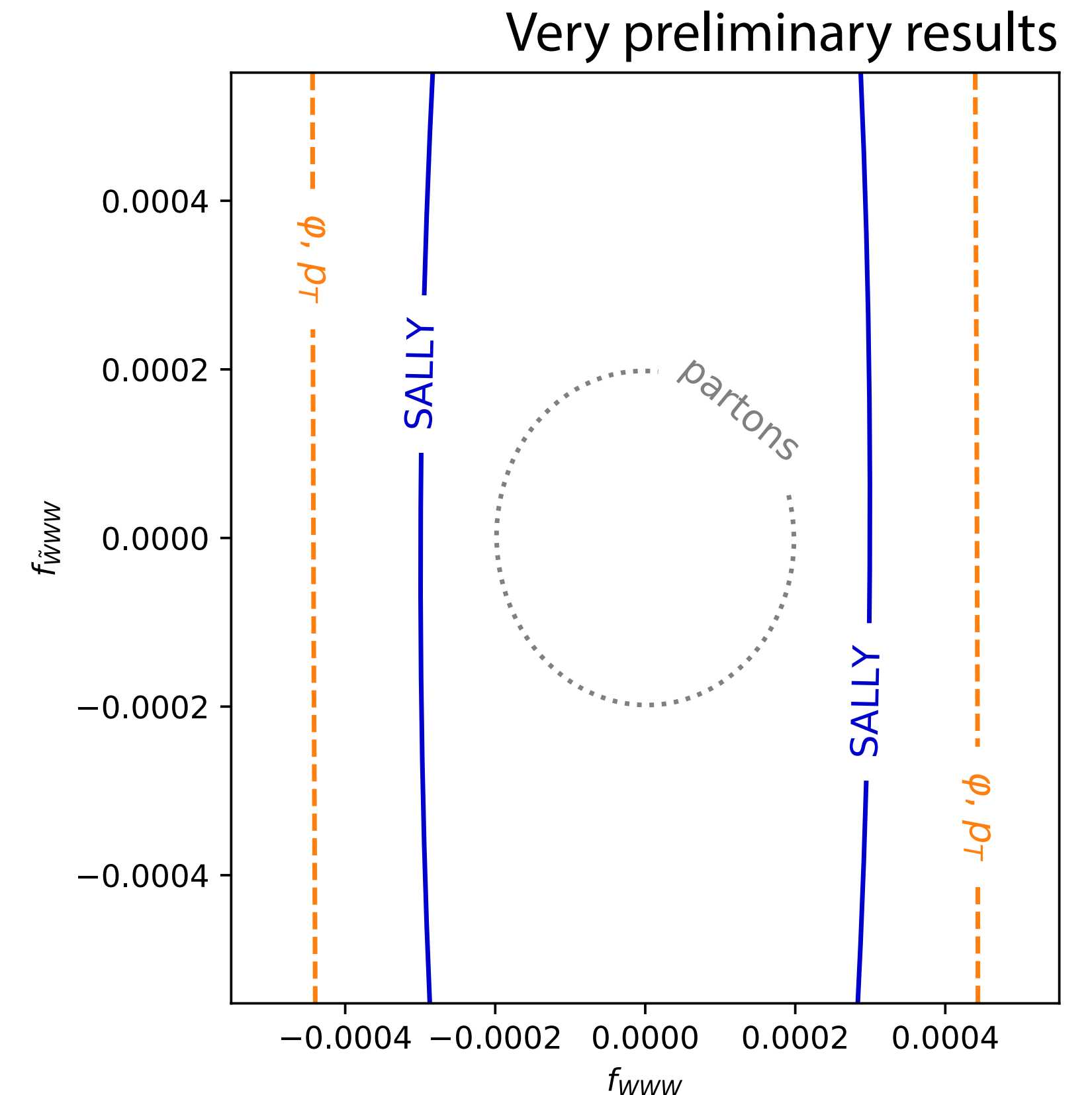
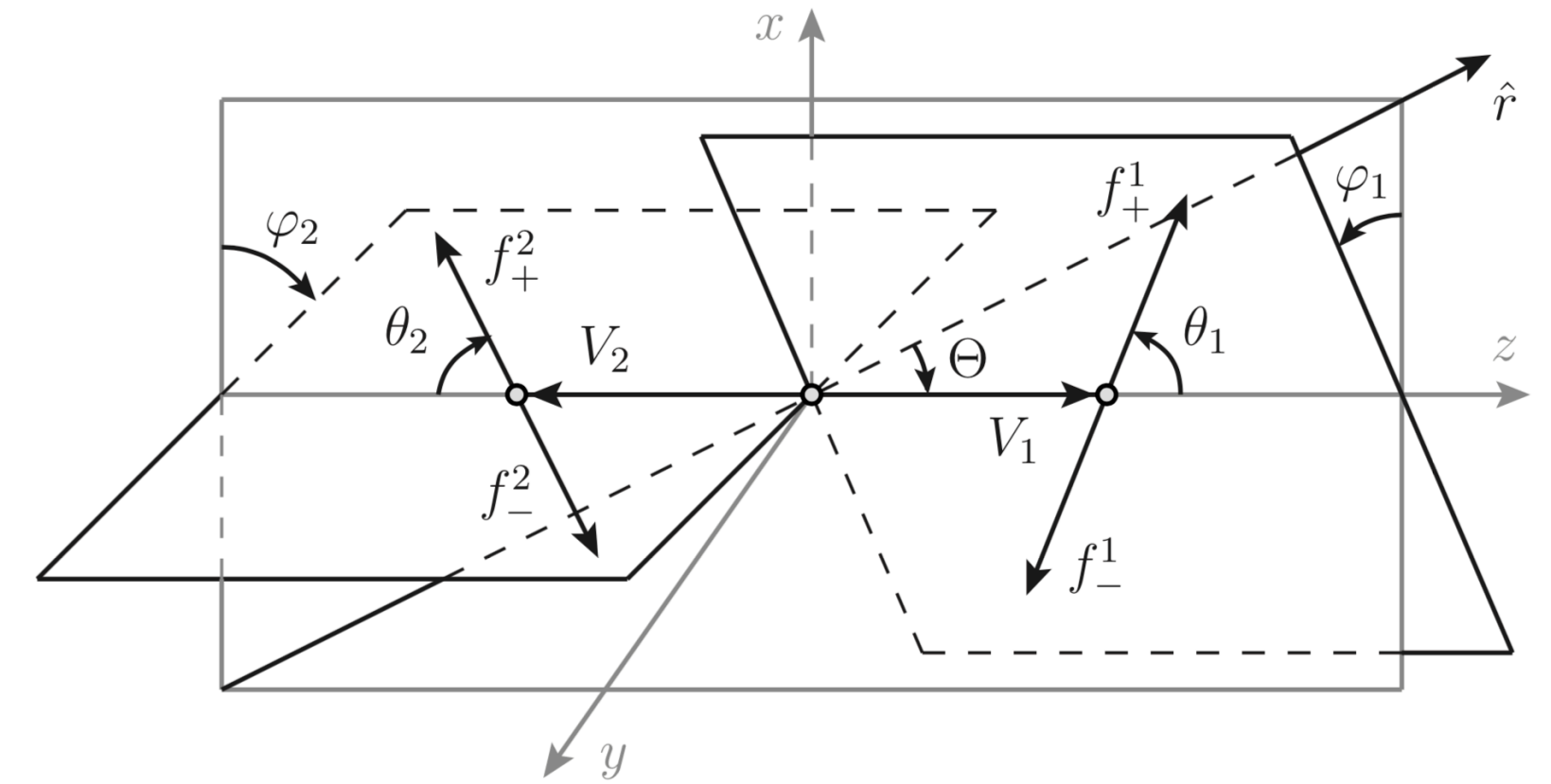
A. Azatov, D. Barducci, E. Venturini 1901.04821]

- We test the ML approach in EFT measurements in  $W\gamma \rightarrow \ell\nu\gamma$

[JB, K. Cranmer, M. Farina, F. Kling, D. Pappadopulo, J. Ruderman in progress]

**New:**  $WZ \rightarrow \ell\ell \ell\nu$  by Chen, Glioti, Panico, Wulzer [arXiv:2007.10356](https://arxiv.org/abs/2007.10356)

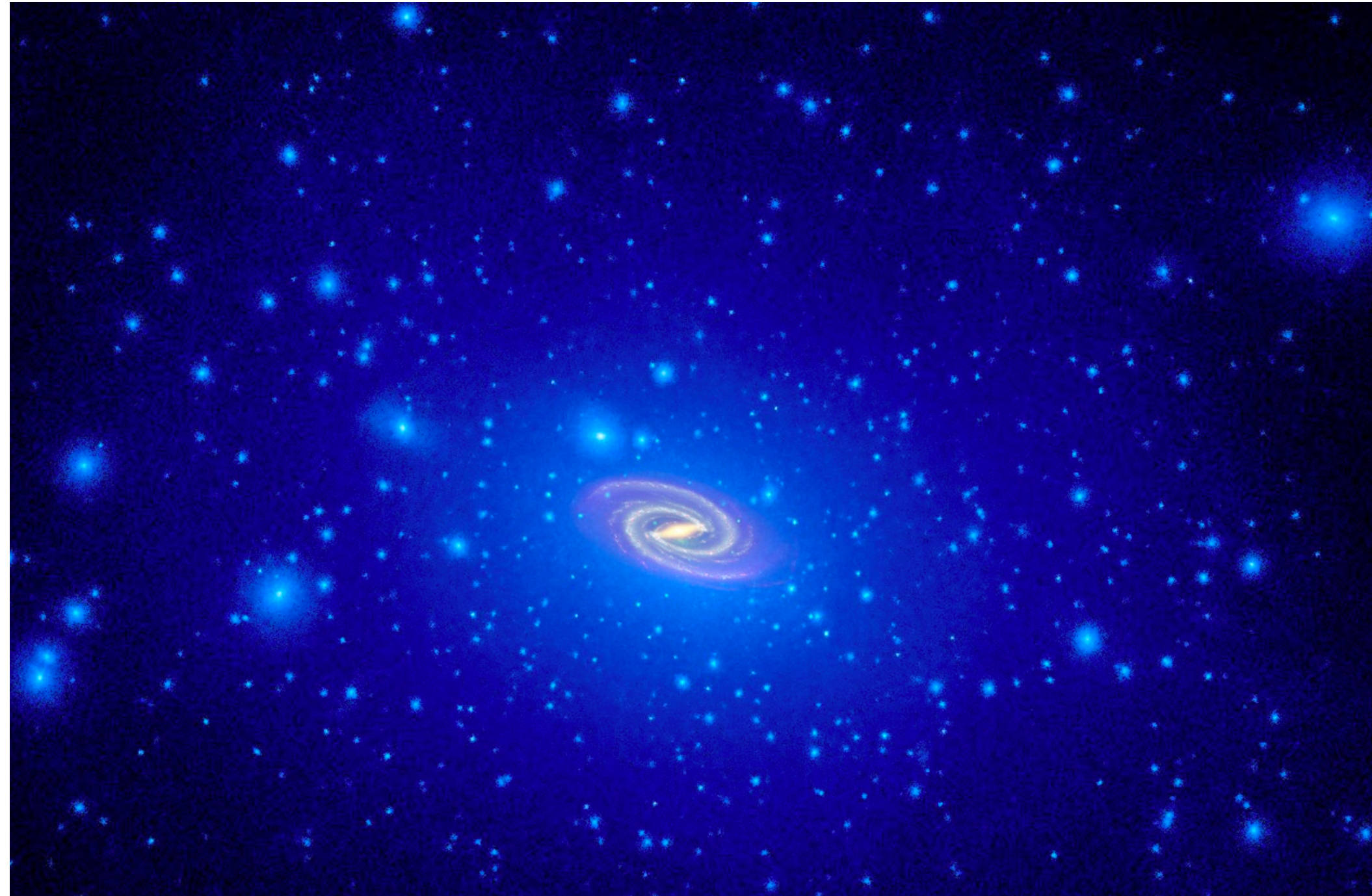
- Preliminary results: we can extract more information when we **analyze events** with **SALLY** than with **histograms of** **and standard observables**





These methods work beyond the LHC.

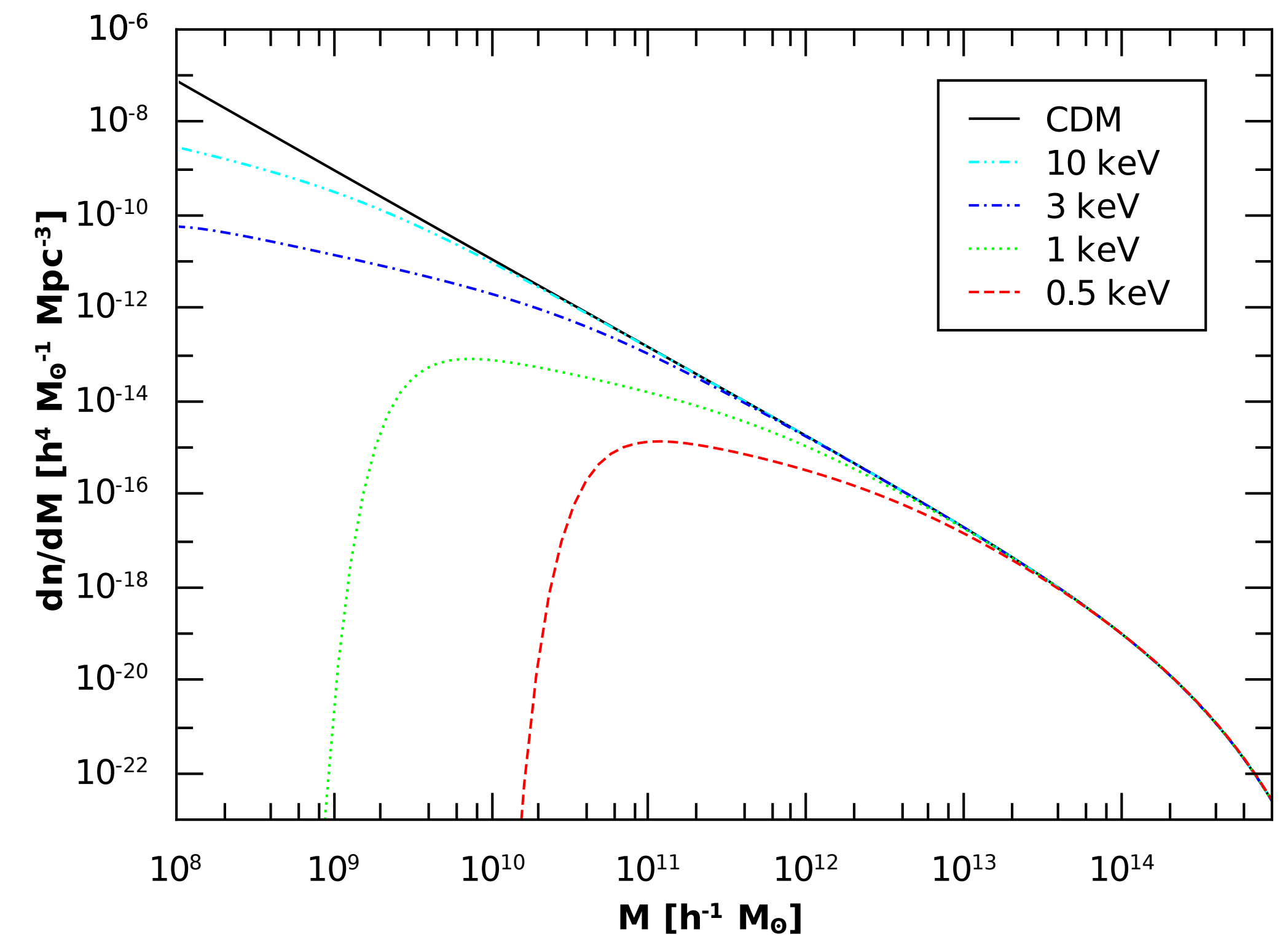
# DM small-scale structure as a probe of DM particle properties



[T. Brown, J. Tumlinson]



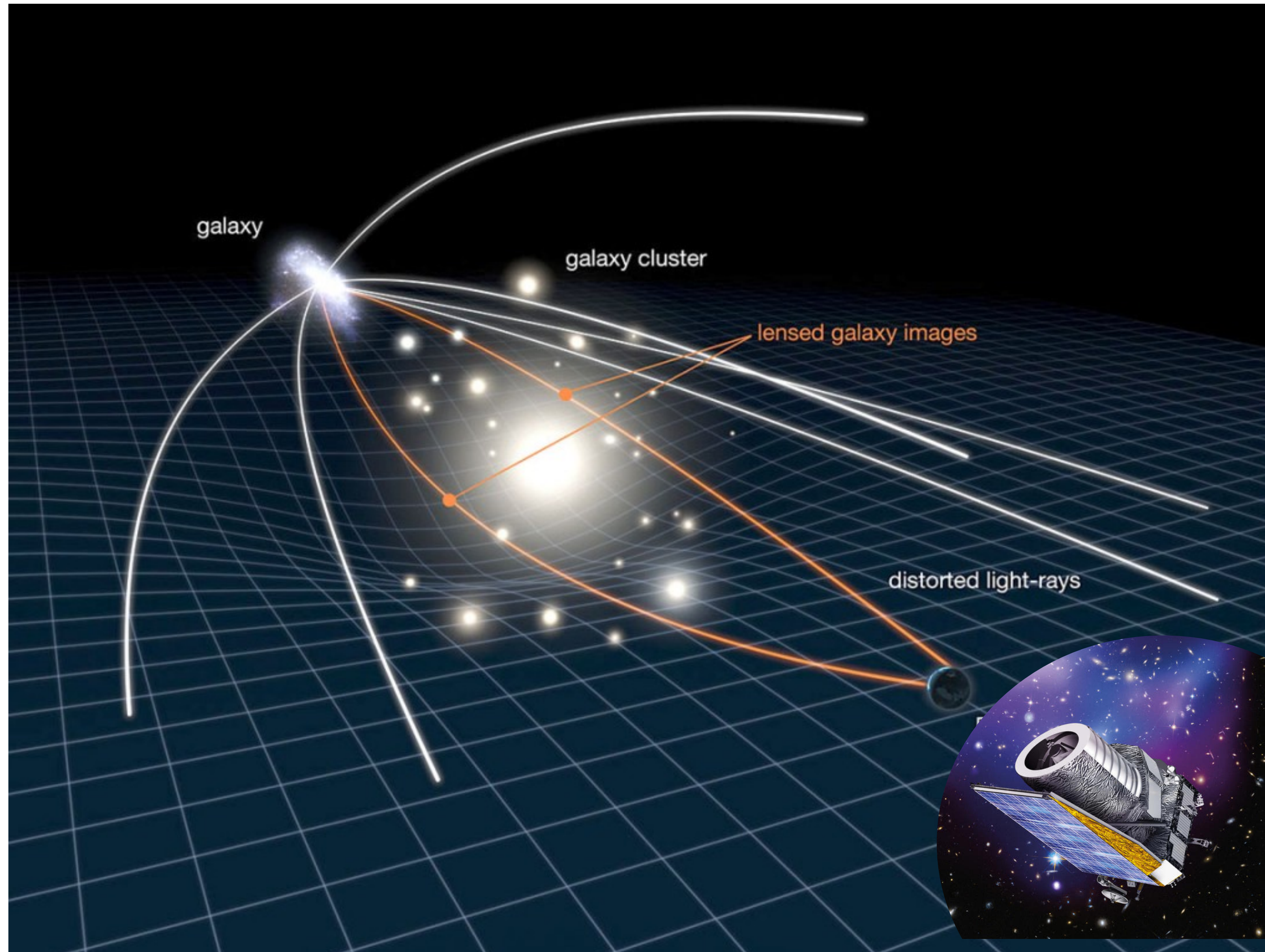
### Abundance of DM subhalos vs mass:



[R. Dunstan et al 1109.6291]



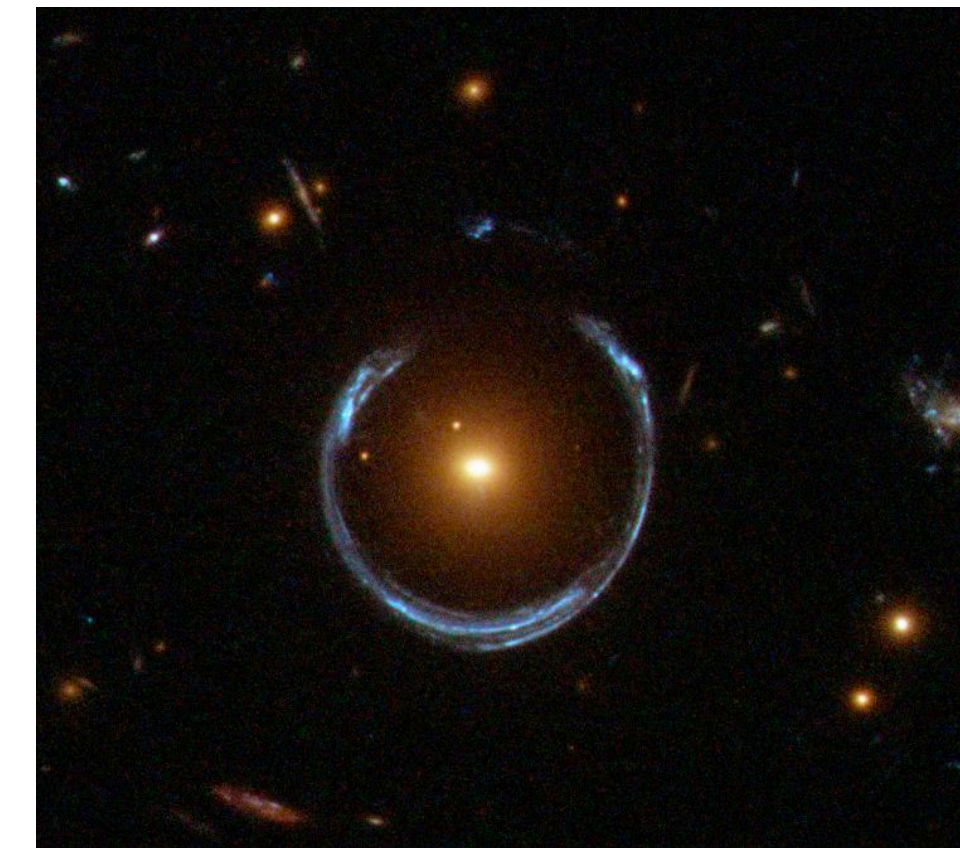
# Strong gravitational lensing



Multiple images  
of quasars

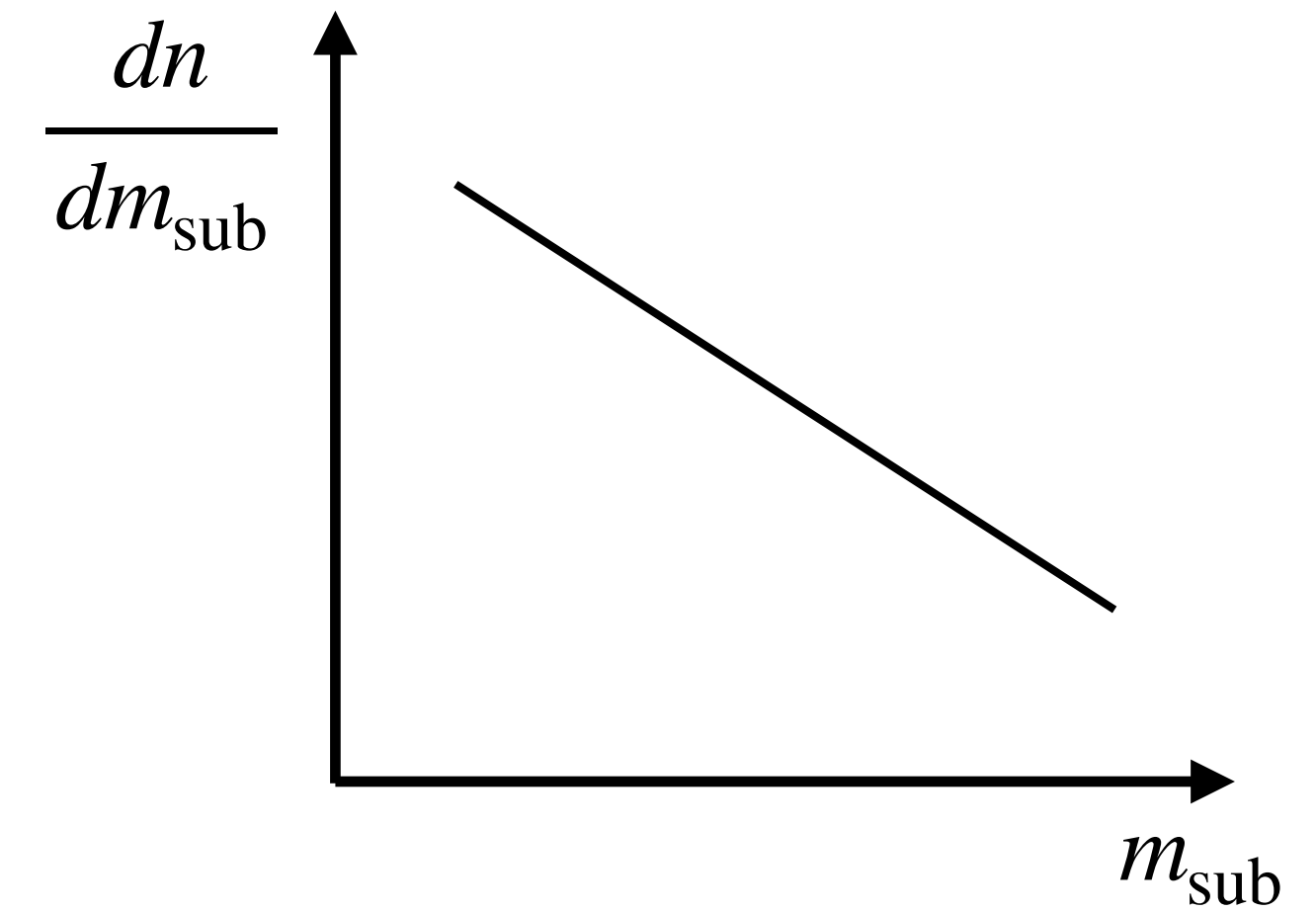
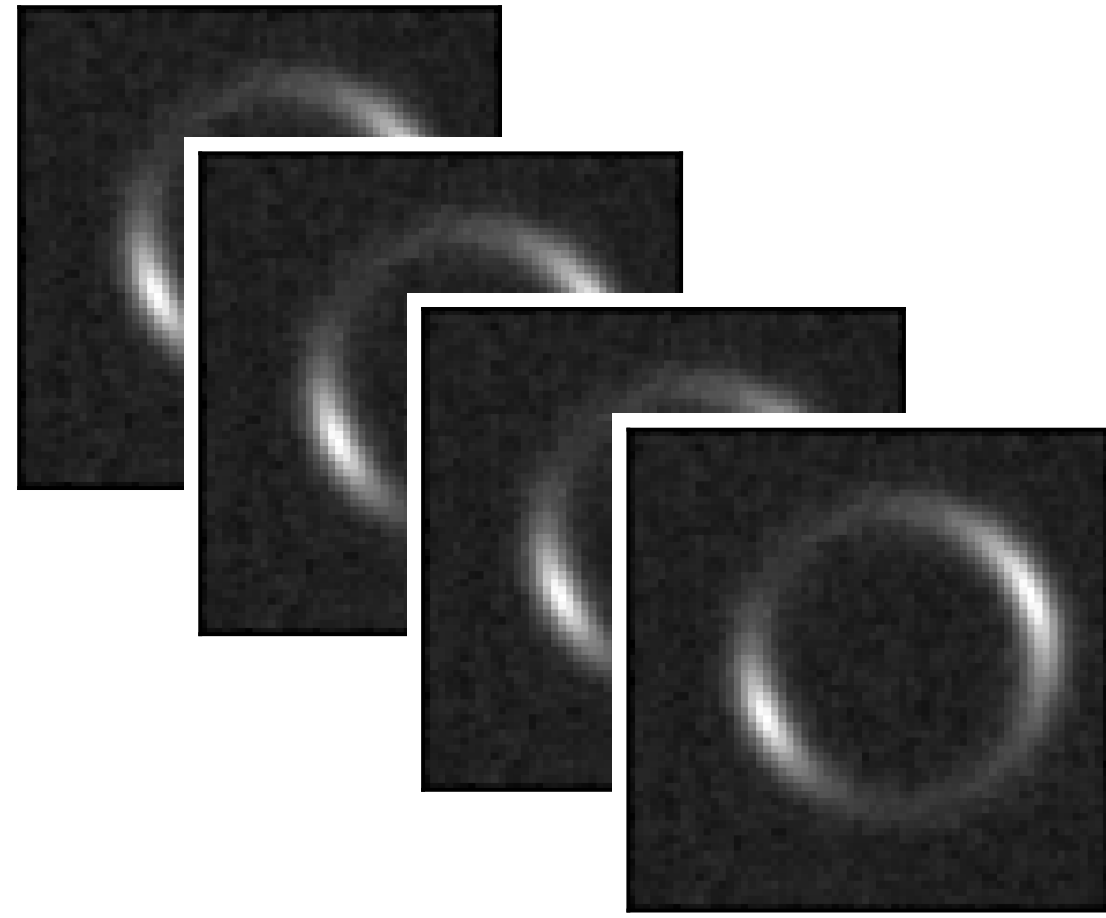


Extended arcs  
from galaxies





# Scalable inference for small subhalos

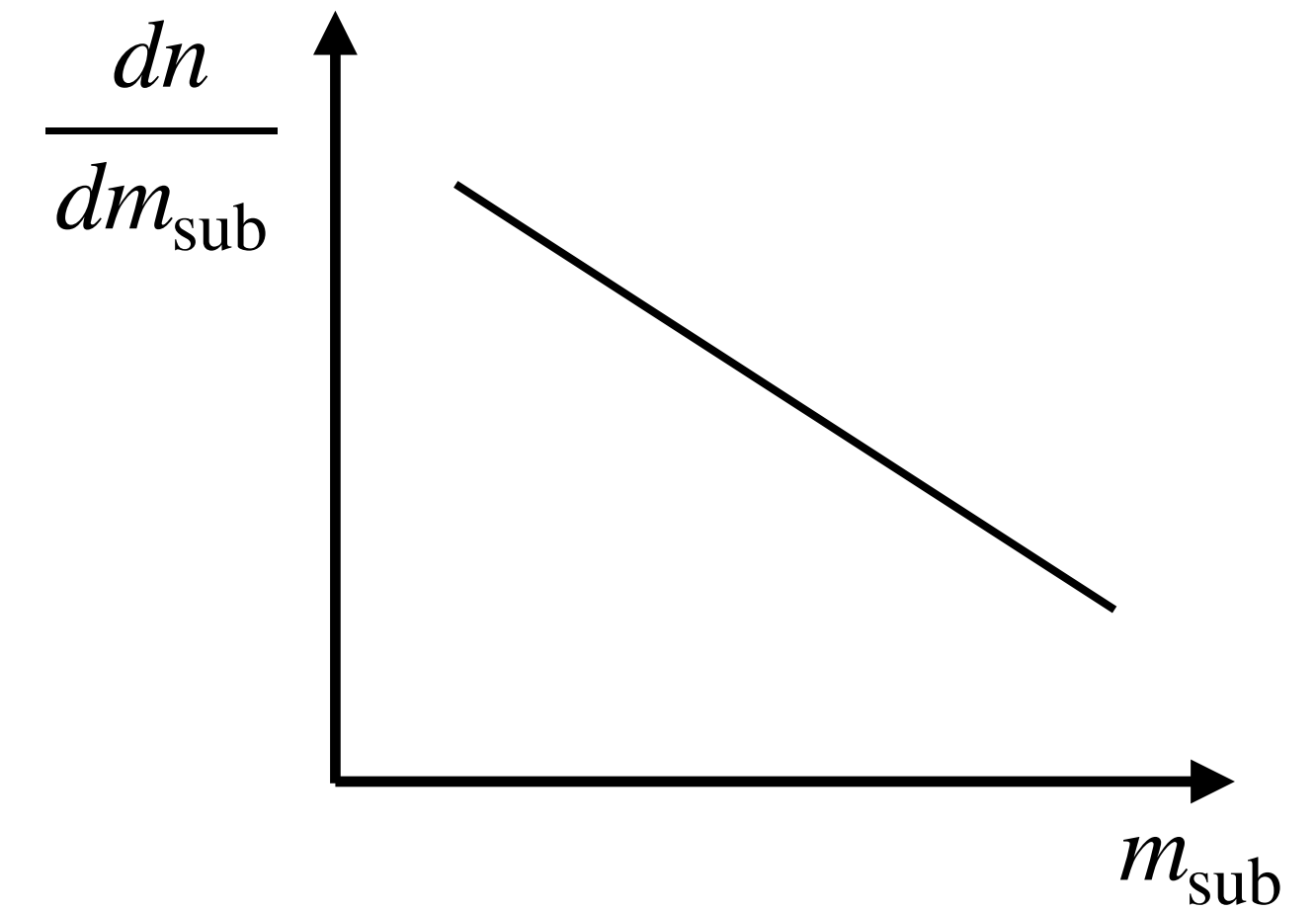
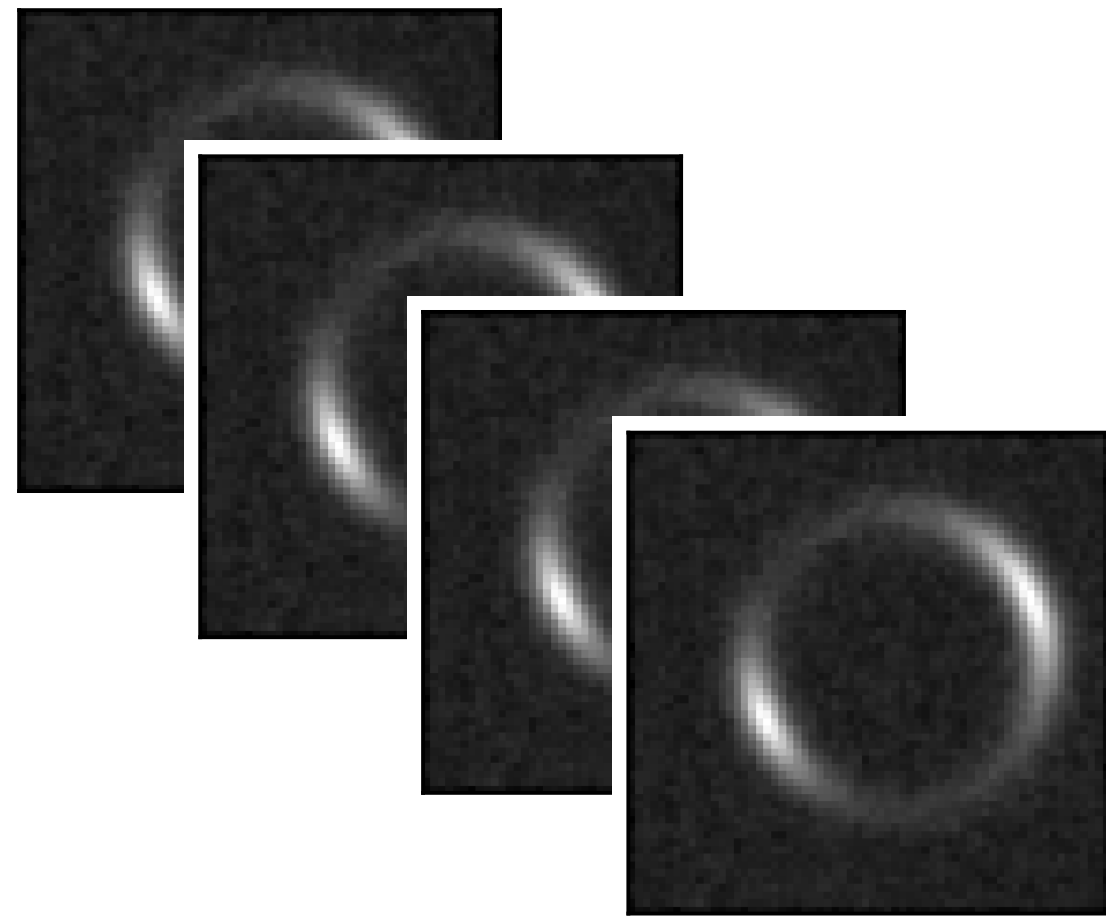


Future surveys (LSST, Euclid) are expected to deliver large samples of galaxy-galaxy strong lenses [Collett et al 1507.02657]

Goal: infer subhalo mass distribution through collective effects of many light subhalos



# Scalable inference for small subhalos



Future surveys (LSST, Euclid) are expected to deliver large samples of galaxy-galaxy strong lenses [Collett et al 1507.02657]

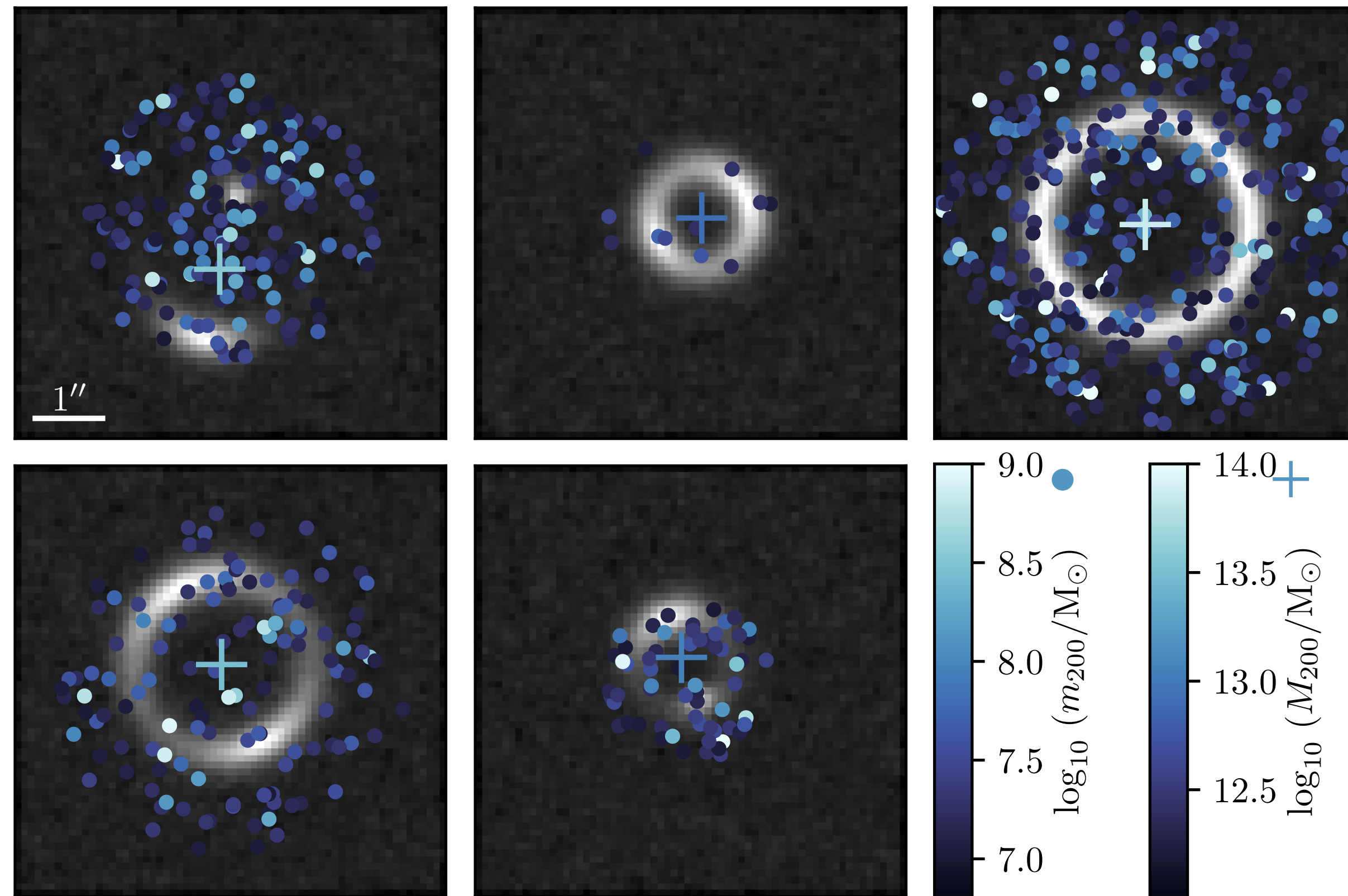
Goal: infer subhalo mass distribution through collective effects of many light subhalos

⇒ Need inference technique that

- scales to many lenses (fast evaluation)
- captures subtle effects in high-dimensional image data
- can deal with a large number of subhalos (latent variables)

# Proof-of-principle simulator

[following T. Collett 1507.02657]

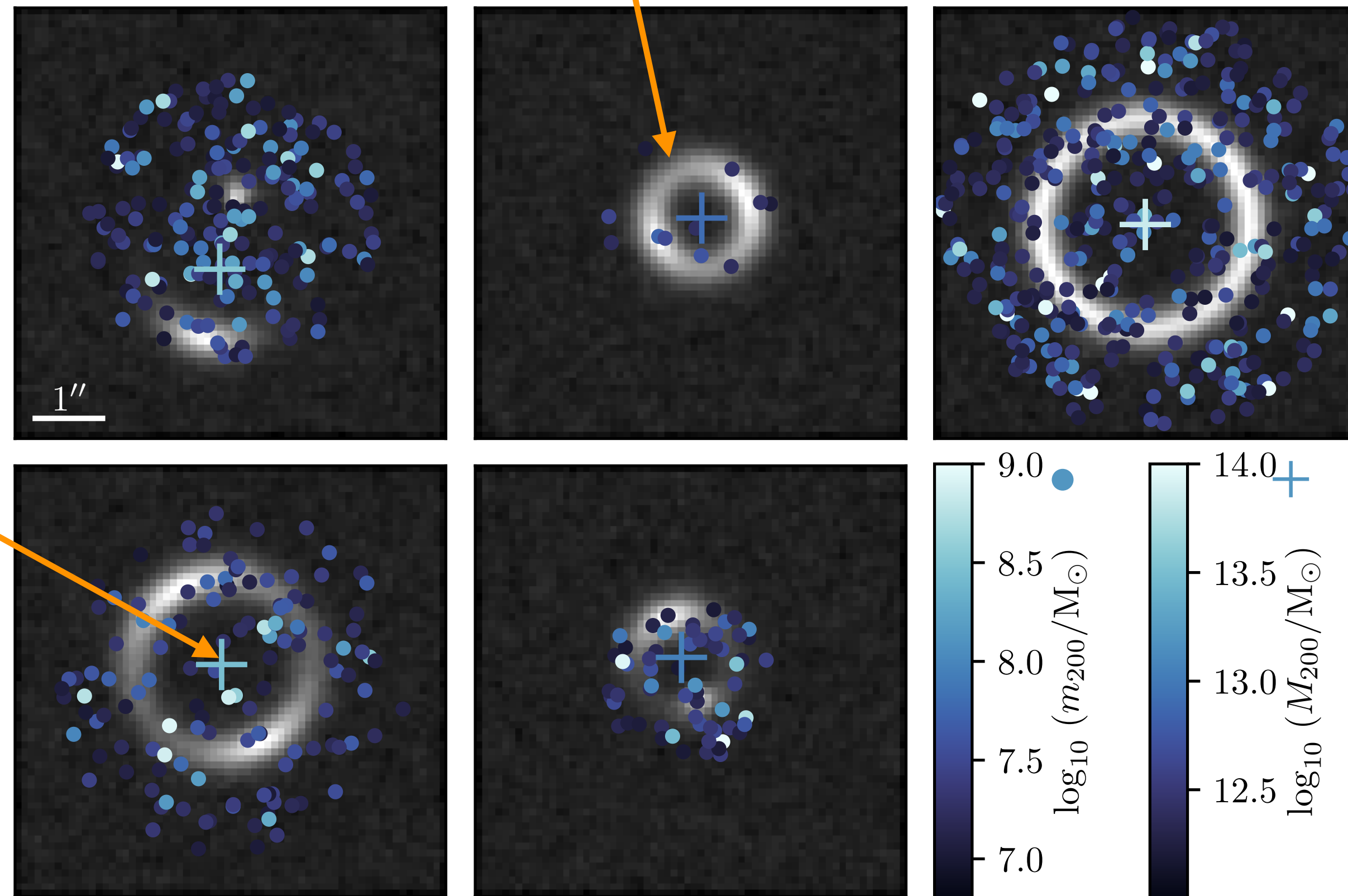


# Proof-of-principle simulator

[following T. Collett 1507.02657]

Spherical lensing host galaxies  
at redshift  $\sim 0.5 \dots 1$

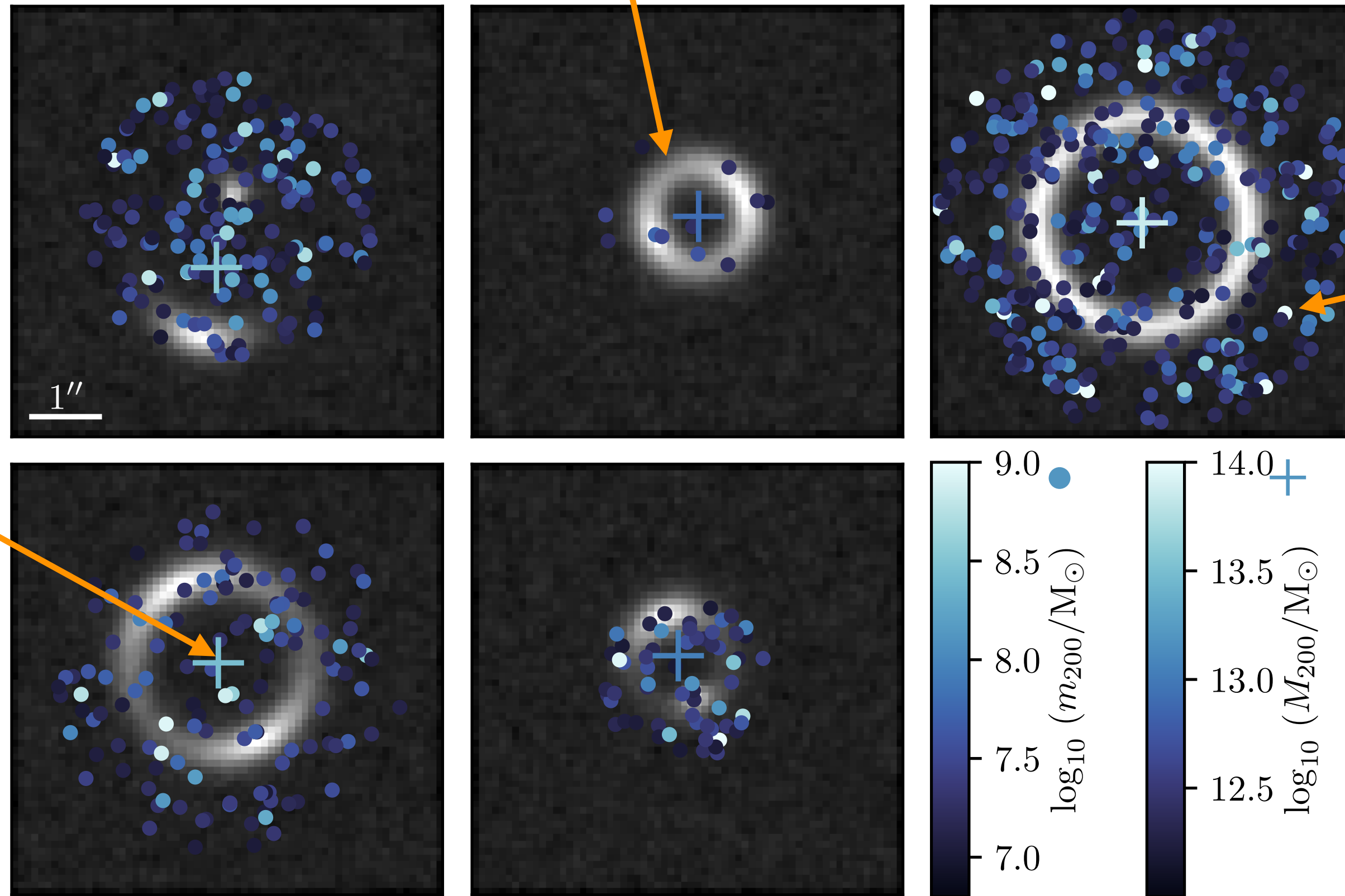
Extended galaxy sources  
at redshift 1.5



# Proof-of-principle simulator

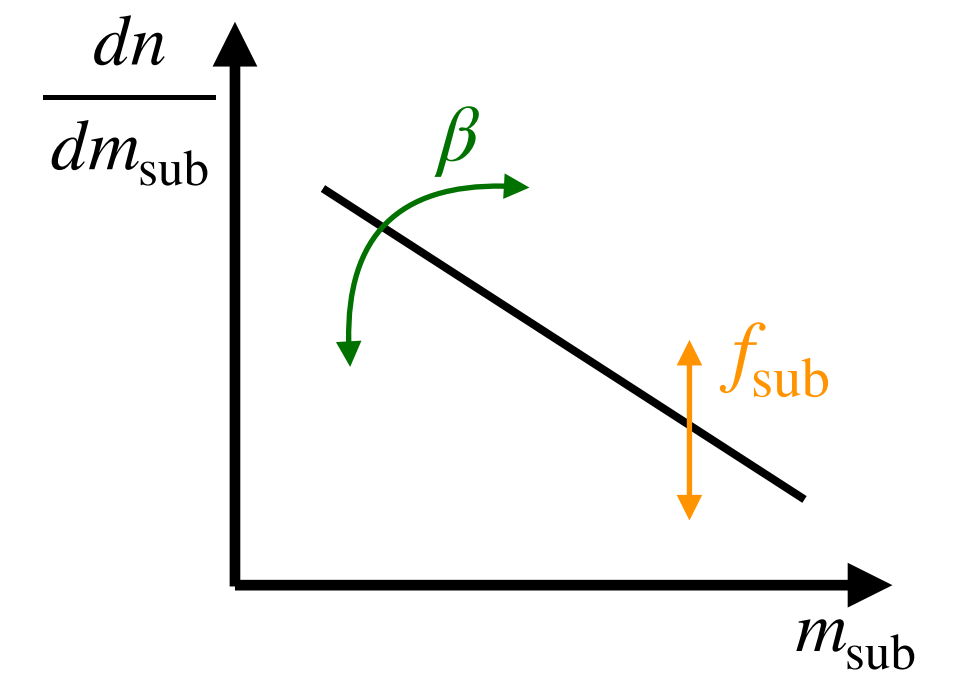
[following T. Collett 1507.02657]

Spherical lensing host galaxies  
at redshift  $\sim 0.5 \dots 1$



Extended galaxy sources  
at redshift 1.5

Subhalos follow mass  
distribution with two  
parameters



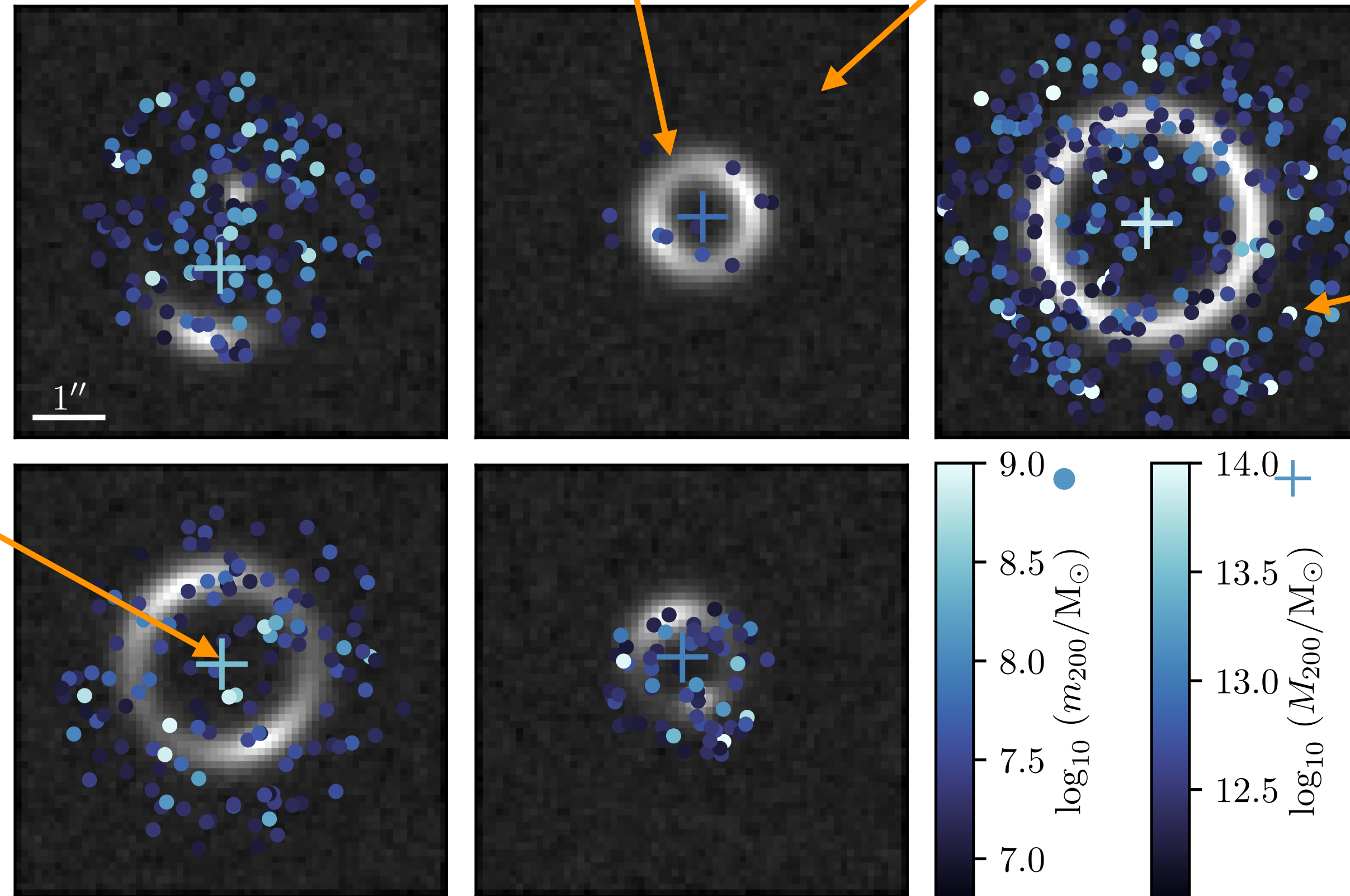


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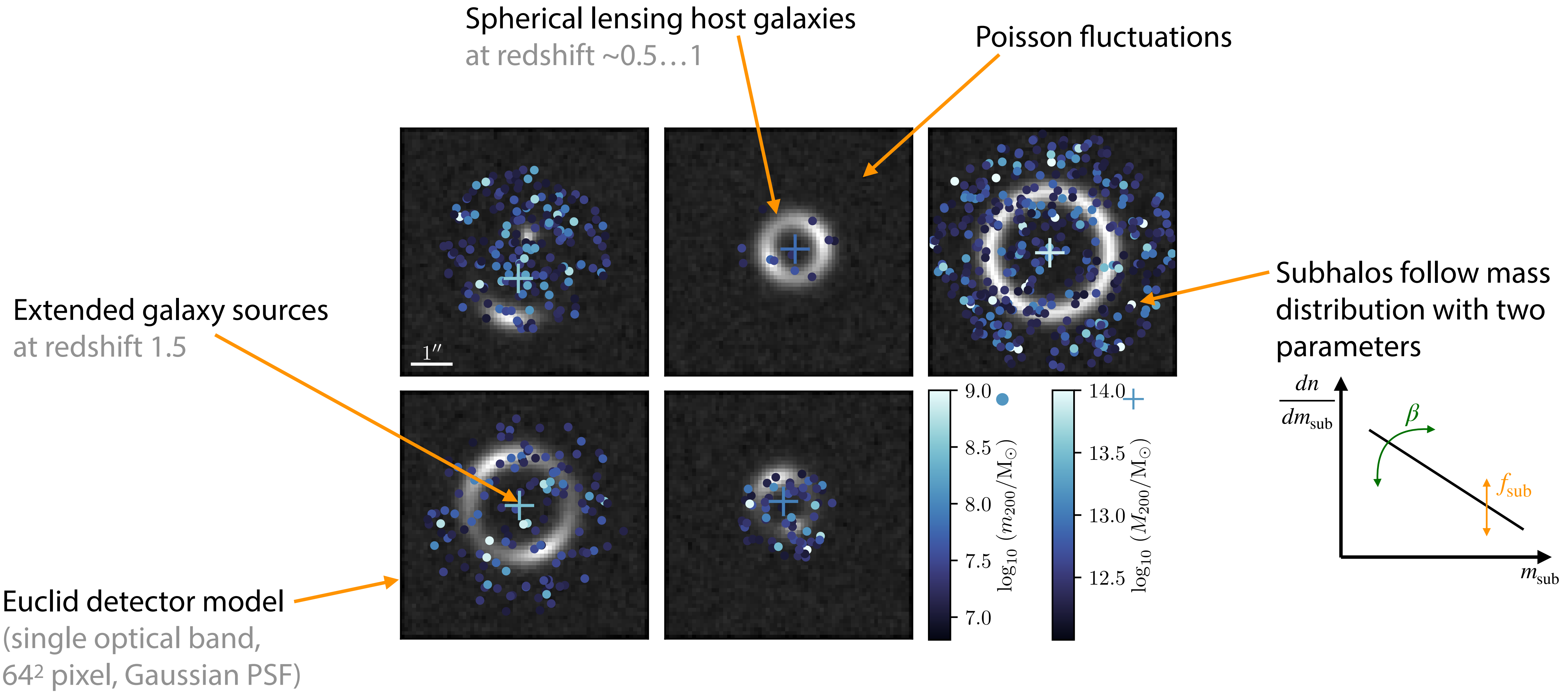
Spherical lensing host galaxies  
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Poisson fluctuations



# Proof-of-principle simulator

[following T. Collett 1507.02657]



# Bayesian inference

Watch how knowledge of the subhalos mass distribution improves as data comes in. See posterior for two parameters concentrate around true value used to generate mock data



Sid Mishra-Sharma



Johann Brehmer

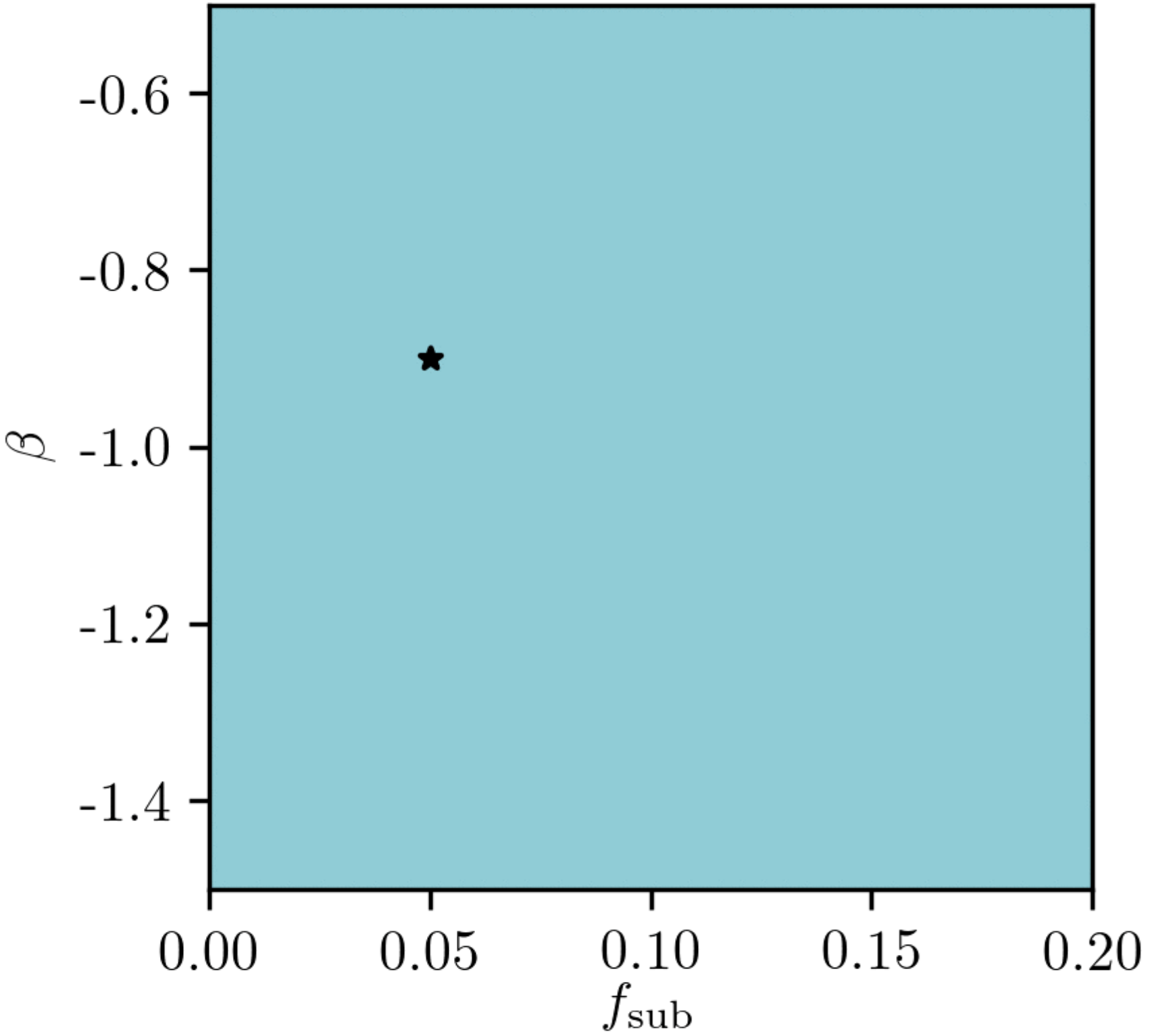
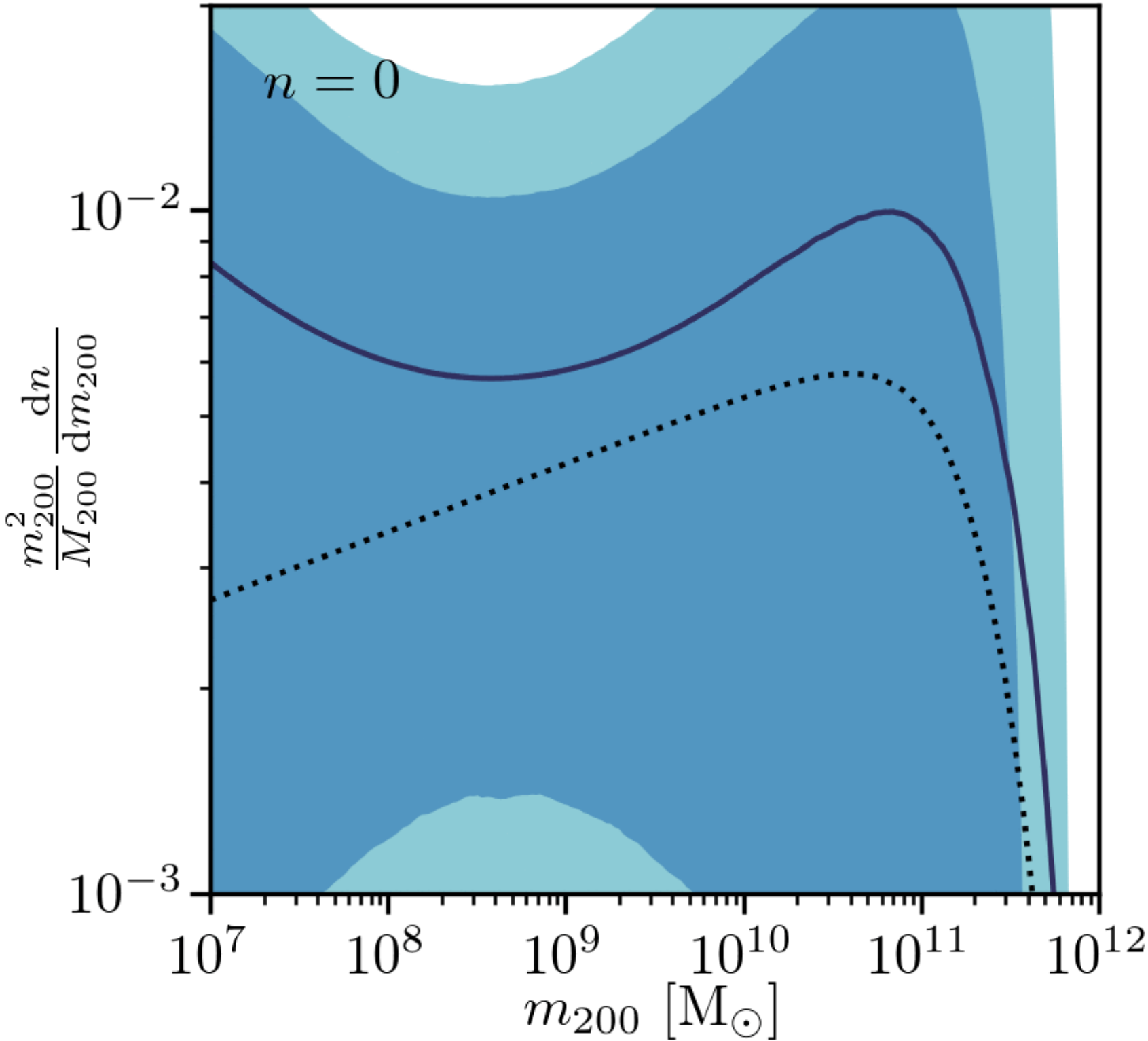


Gilles Louppe



Joeri Hermans

+ K.C. published in ApJ, 886 (2019) arXiv:1909.02005 ]



- $f_{\text{sub}}$ , defined as the fraction of the total dark matter halo mass contained in bound substructure in a given mass range
- The halo virial mass  $M_{200}$  describes the total mass contained with the virial radius  $r_{200}$ , defined as the radius within which the mean density is 200 times the critical density of the universe



# Bayesian inference

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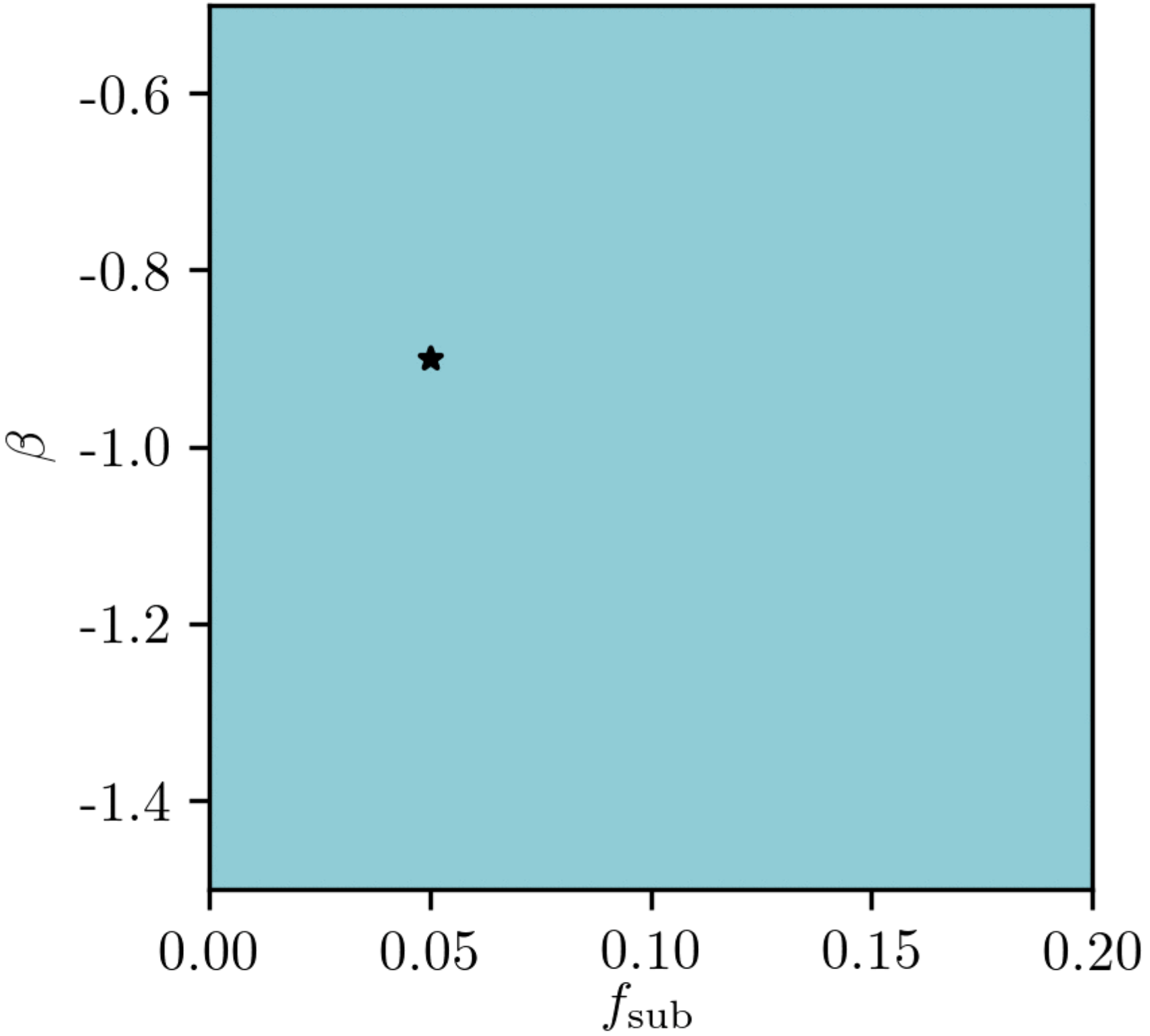
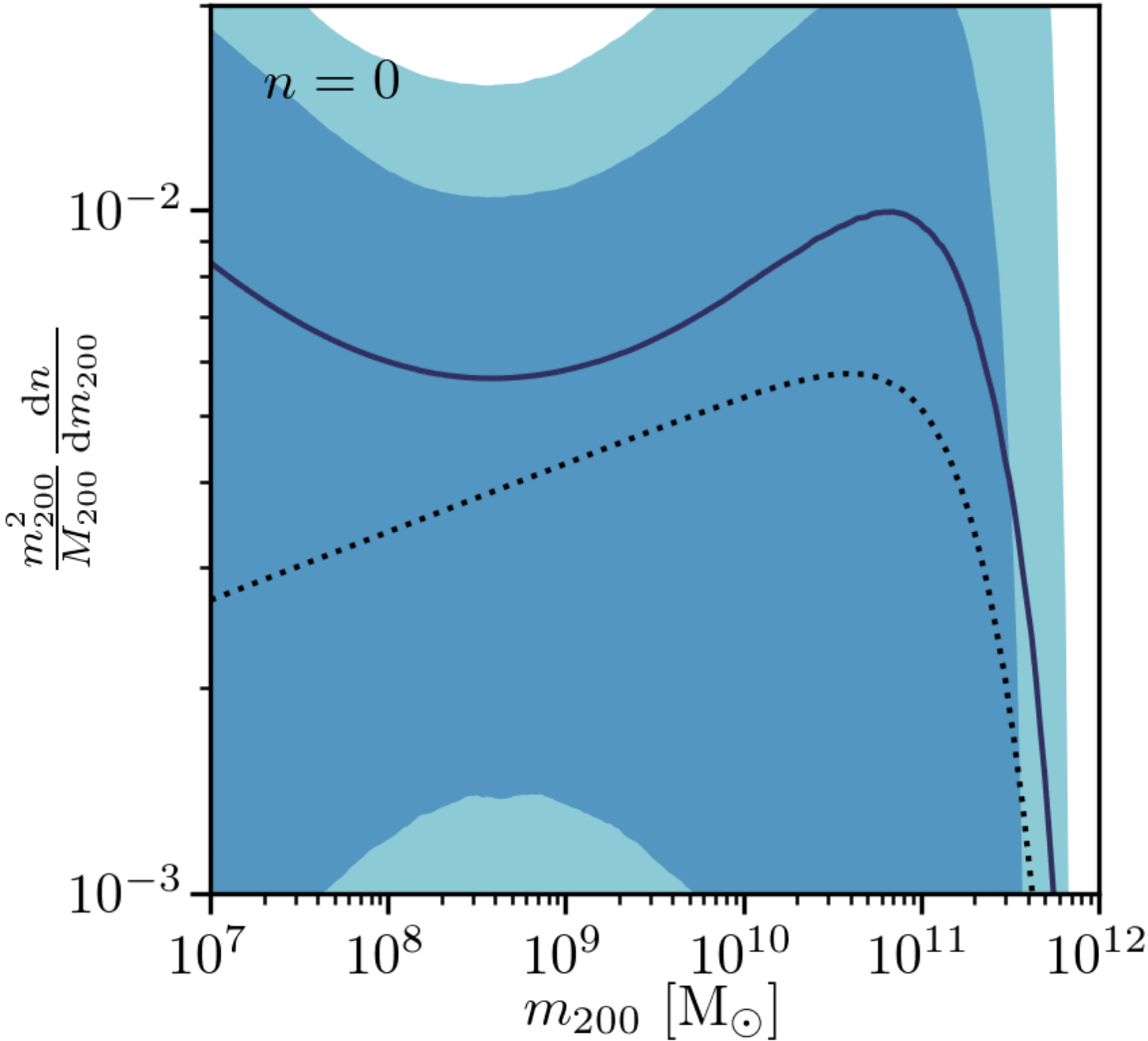


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# Conclusions

Information geometry is a useful concept for phenomenology and could be very useful for Snowmass studies

Simulation-based inference techniques powered by machine learning have the potential to greatly improve our sensitivity for a wide range of searches and measurements in particle physics

- In particular, it allows us to model the likelihood for a lower-level, high-dimensional representation of the data including complicated detector effects

# References

## **Opinionated review**

K. Cranmer, JB, G. Louppe:

“The frontier of simulation-based inference”

[1911.01429]

## **Do It Yourself (for LHC physics)**

JB, F. Kling, I. Espejo, K. Cranmer:

“MadMiner: Machine learning—based inference for particle physics”

[CSBS, 1907.10621, <https://github.com/diana-hep/madminer>]

## **Strong lensing**

JB, S. Mishra-Sharma, J. Hermans, G. Louppe, K. Cranmer

“Mining for Dark Matter Substructure: Inferring subhalo population properties from strong lenses with machine learning”

[ApJ, 1909.02005]

## **LHC HXSWG YR4 STXS**

JB, S. Dawson, S. Homiller, F. Kling, T. Plehn:

“Benchmarking simplified template cross sections in WH production”

[JHEP, 1908.06980]

## **Meh, the original trilogy was better**

JB, K. Cranmer, G. Louppe, J. Pavez:

“Constraining Effective Field Theories with machine learning”

[PRL, 1805.00013]

JB, K. Cranmer, G. Louppe, J. Pavez:

“A guide to constraining Effective Field Theories with machine learning”

[PRD, 1805.00020]

JB, G. Louppe, J. Pavez, K. Cranmer:

“Mining gold from implicit models to improve likelihood-free inference”

[PNAS, 1805.12244]

## **Follow-up with incremental improvements**

M. Stoye, JB, K. Cranmer, G. Louppe, J. Pavez:

“Likelihood-free inference with an improved cross-entropy estimator”

[NeurIPS workshop, 1808.00973]

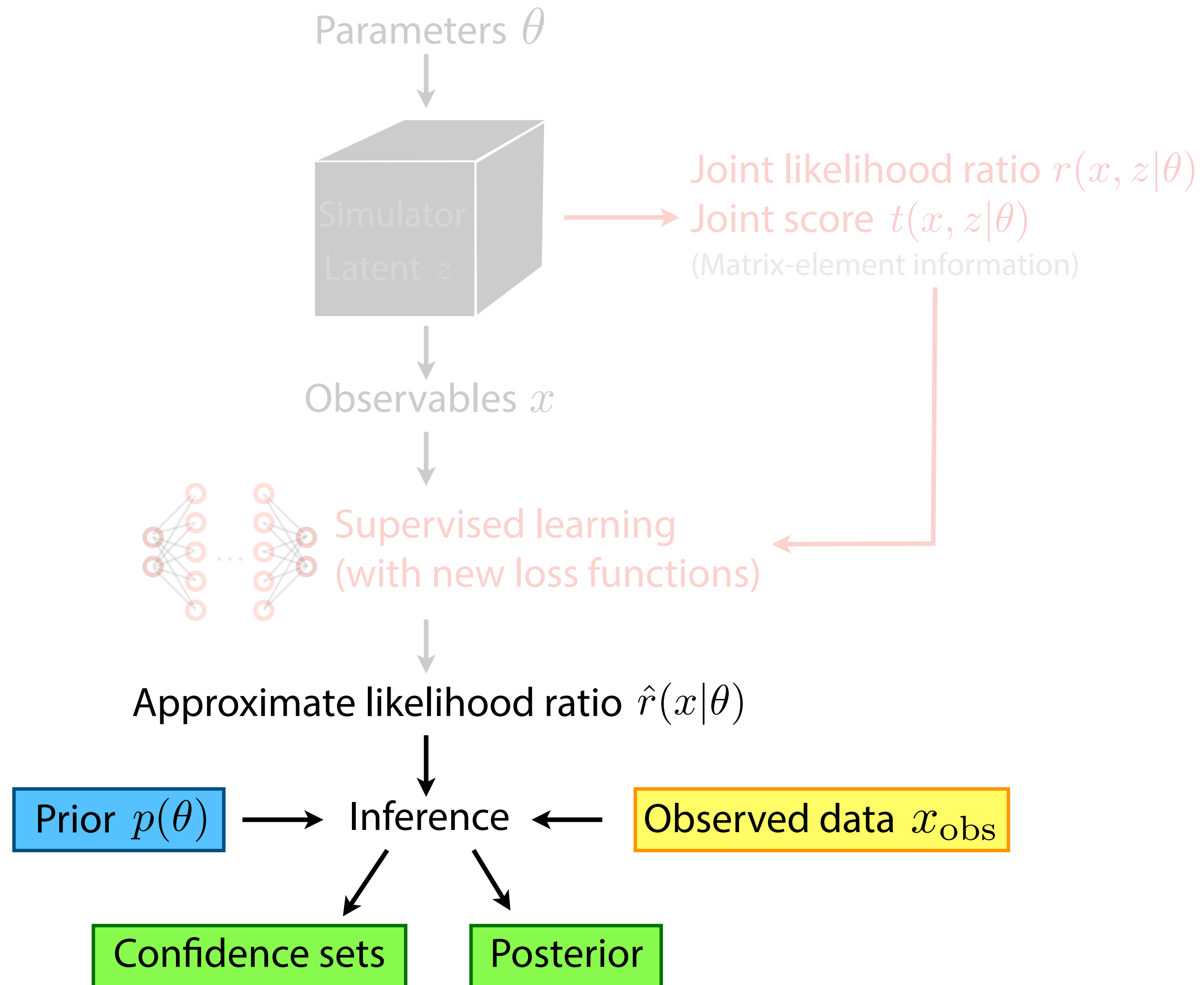
Backup

# Inference, Limits, & Fisher Information Geometry

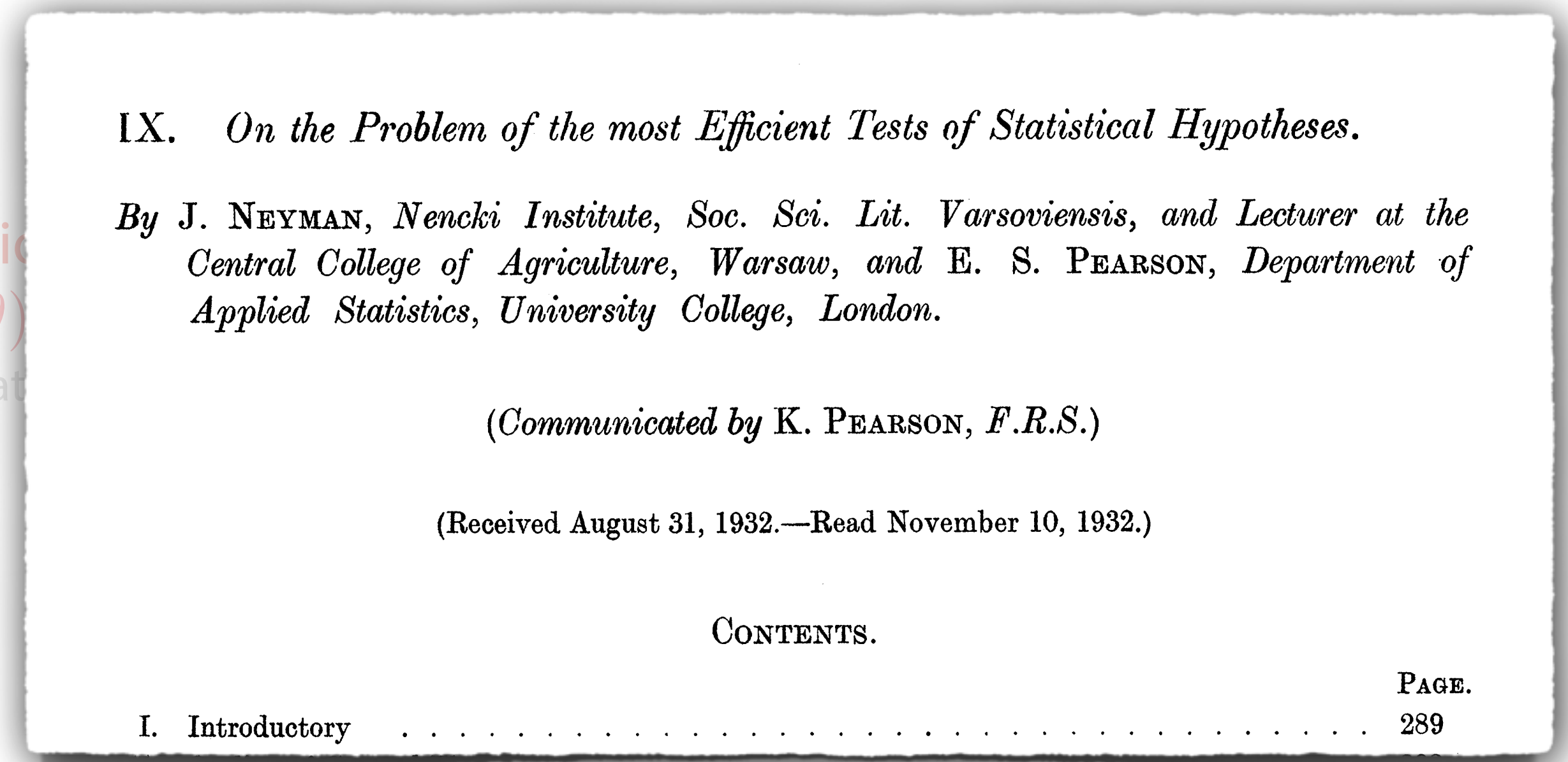
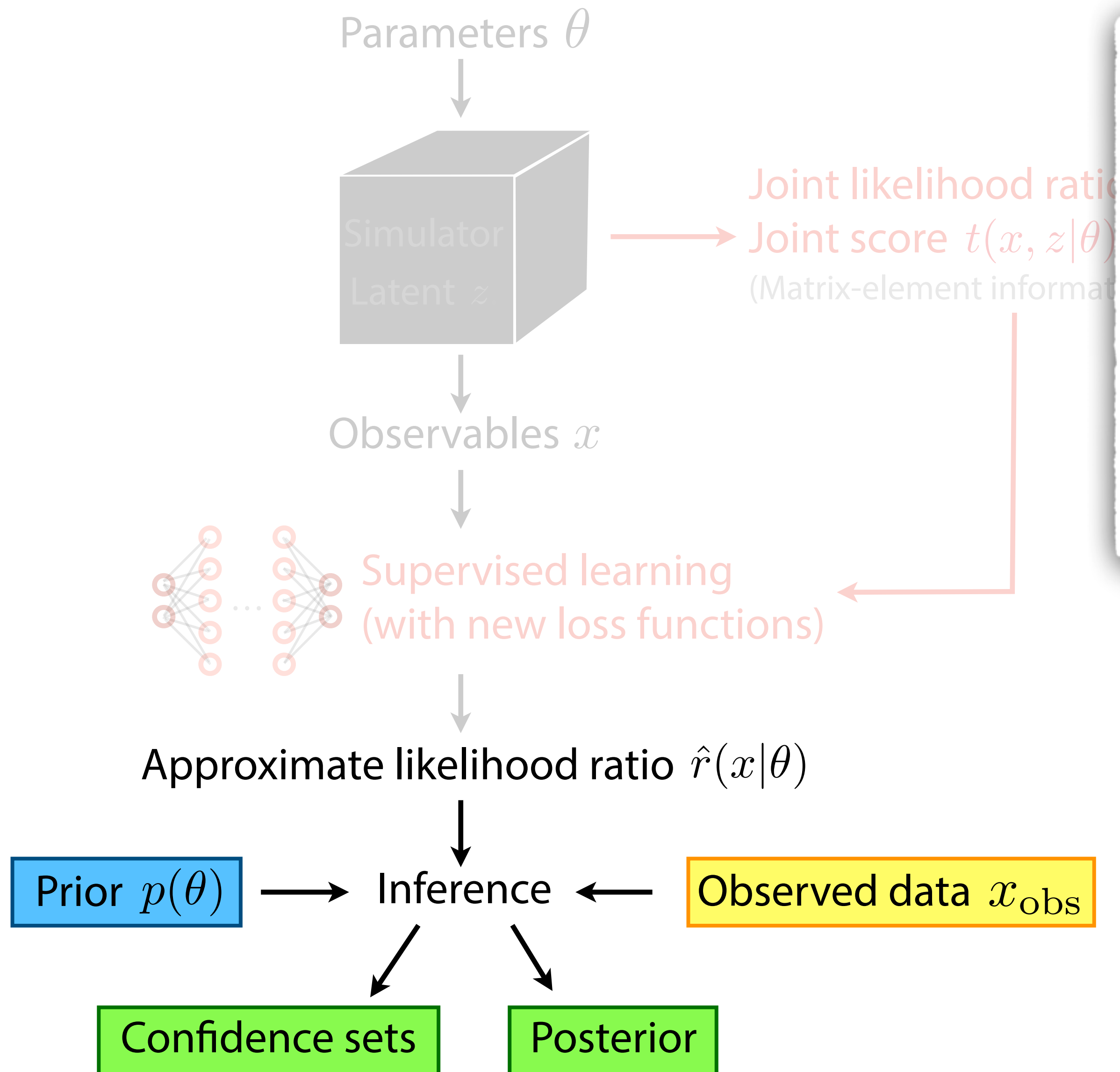
[JB, K. Cranmer, G. Louppe, J. Pavez 1805.00013, 1805.00020, 1805.12244]



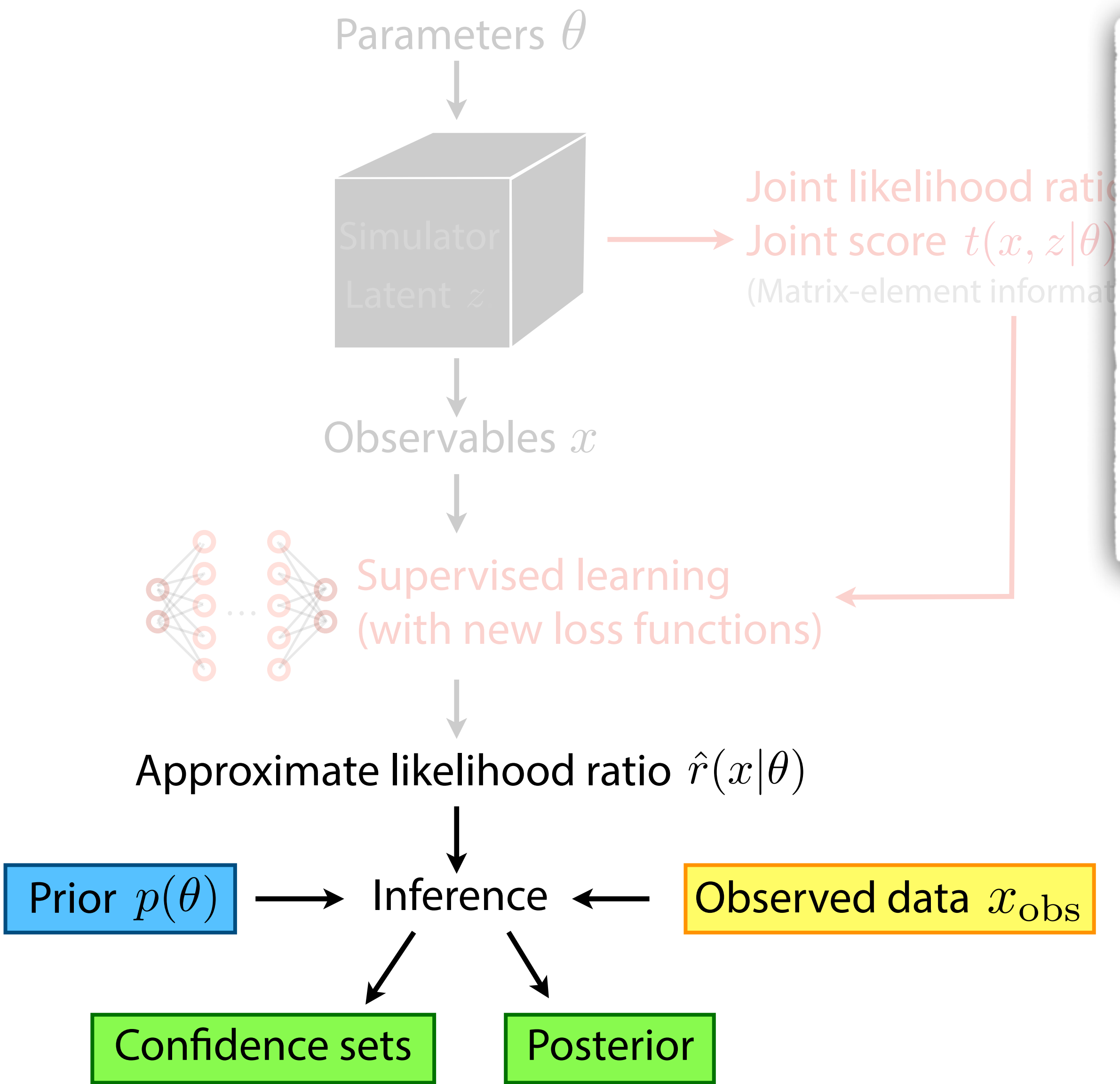
# Step 3: Inference



# Step 3: Inference



# Step 3: Inference



*IX. On the Problem of the most Efficient Tests of Statistical Hypotheses.*

By J. NEYMAN, *Nencki Institute, Soc. Sci. Lit. Varsoviensis, and Lecturer at the Central College of Agriculture, Warsaw,* and E. S. PEARSON, *Department of Applied Statistics, University College, London.*

(Communicated by K. PEARSON, F.R.S.)

(Received August 31, 1932.—Read November 10, 1932.)

Eur. Phys. J. C (2011) 71: 1554  
DOI 10.1140/epjc/s10052-011-1554-0

THE EUROPEAN PHYSICAL JOURNAL C

Special Article - Tools for Experiment and Theory

**Asymptotic formulae for likelihood-based tests of new physics**

Glen Cowan<sup>1</sup>, Kyle Cranmer<sup>2</sup>, Eilam Gross<sup>3</sup>, Ofer Vitells<sup>3,a</sup>

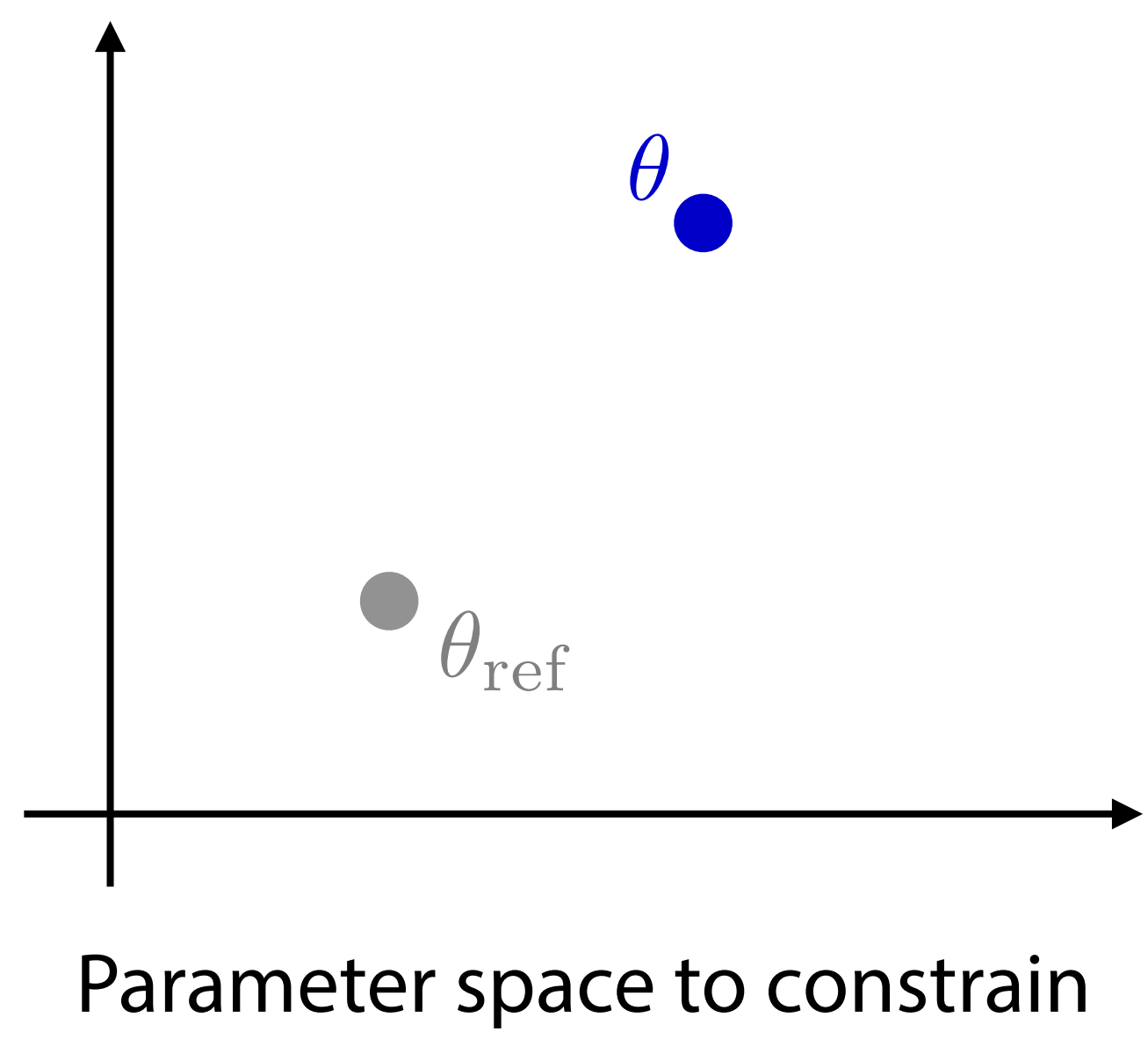
<sup>1</sup>Physics Department, Royal Holloway, University of London, Egham TW20 0EX, UK  
<sup>2</sup>Physics Department, New York University, New York, NY 10003, USA  
<sup>3</sup>Weizmann Institute of Science, Rehovot 76100, Israel

Received: 15 October 2010 / Revised: 6 January 2011 / Published online: 9 February 2011  
 © The Author(s) 2011. This article is published with open access at Springerlink.com

**Abstract** We describe likelihood-based statistical tests for use in high energy physics for the discovery of new phenomena and for construction of confidence intervals on model data sets by a single representative one, referred to here as the “Asimov” data set.<sup>1</sup> In the past, this method has been used and justified intuitively (e.g., [4, 5]). Here we provide

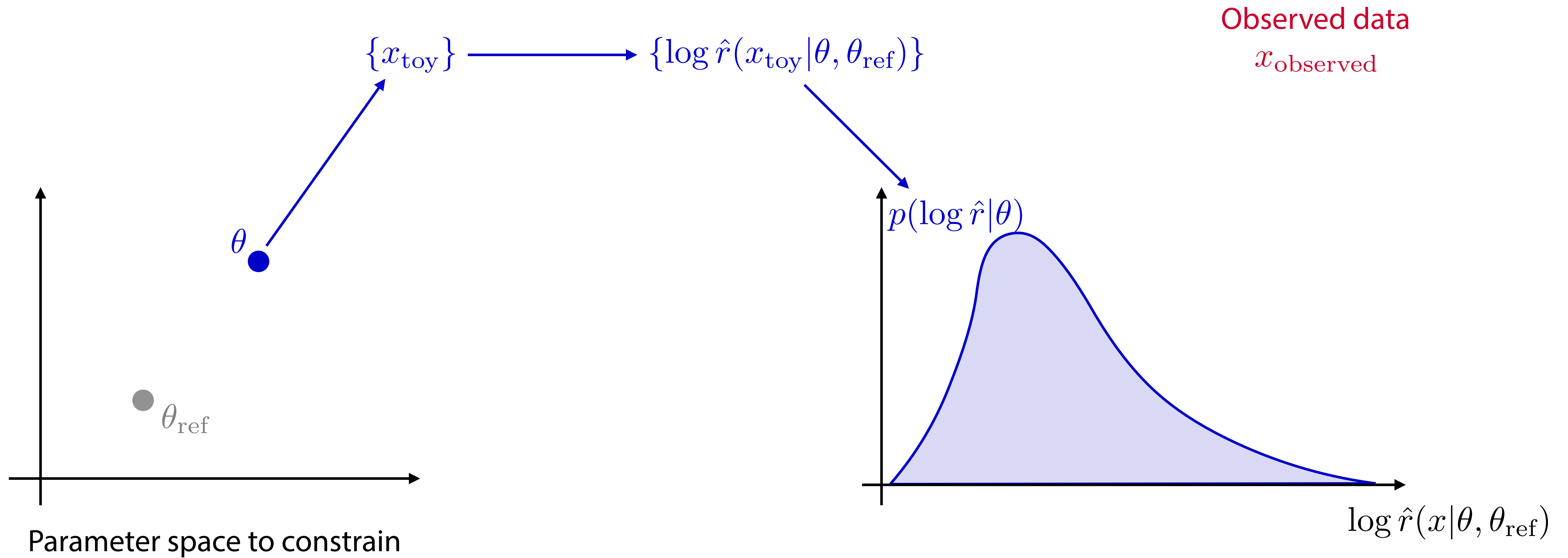
# Limit setting (frequentist, standard ATLAS / CMS practice)

Observed data  
 $x_{\text{observed}}$

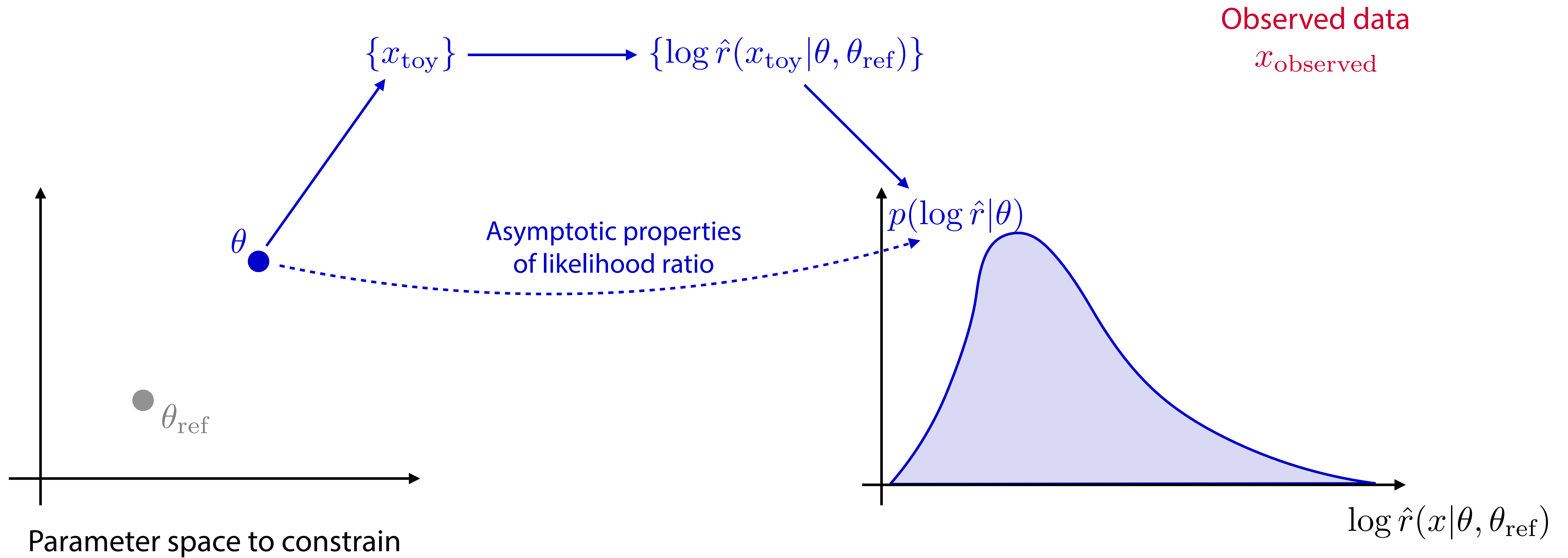




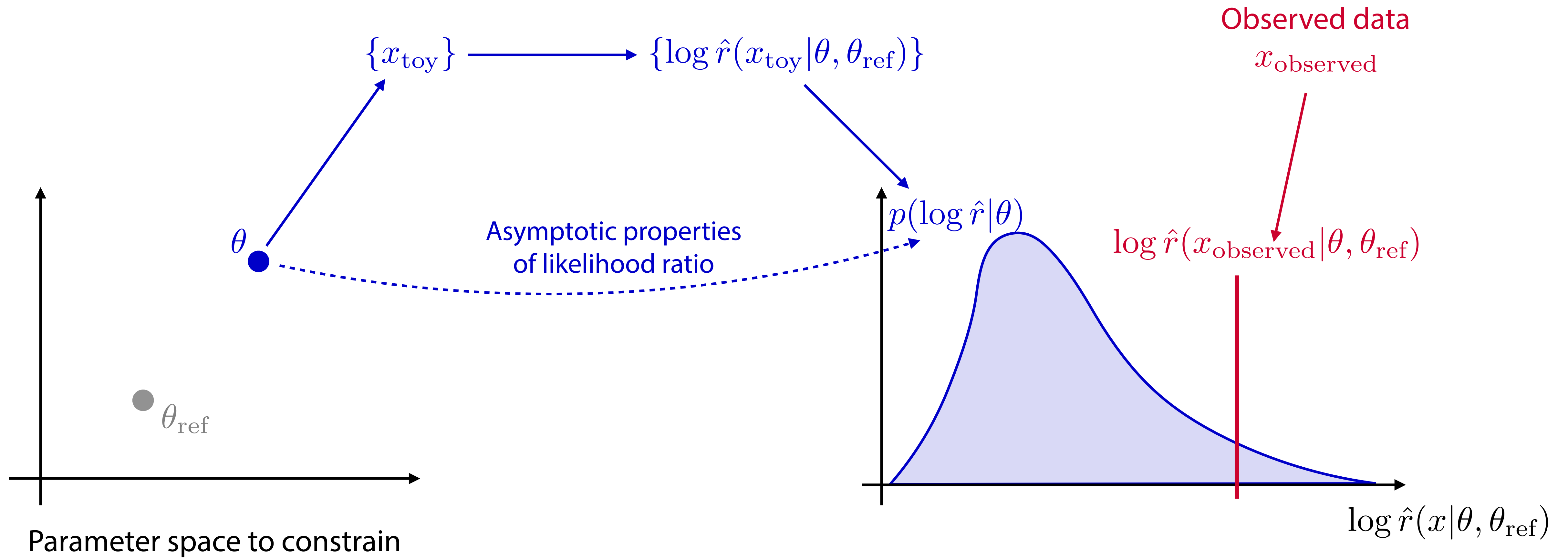
# Limit setting (frequentist, standard ATLAS / CMS practice)



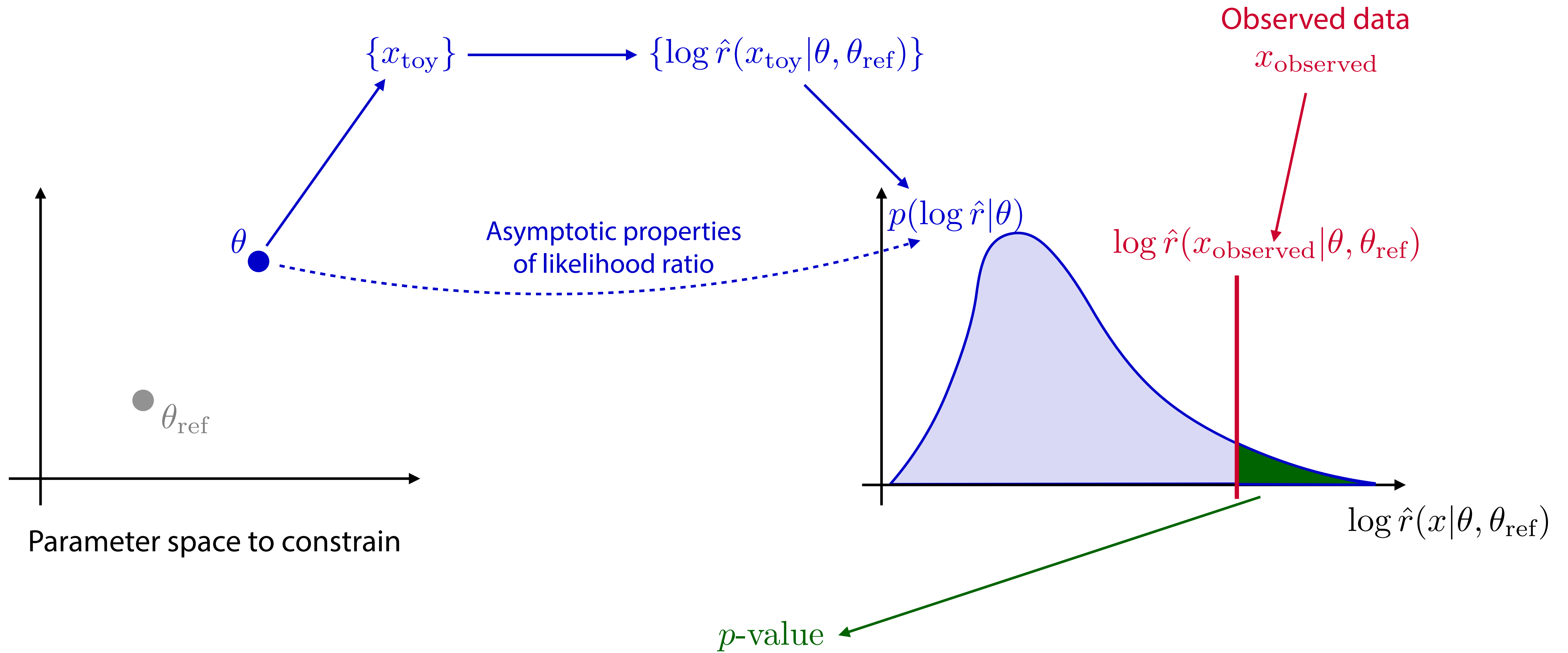
# Limit setting (frequentist, standard ATLAS / CMS practice)



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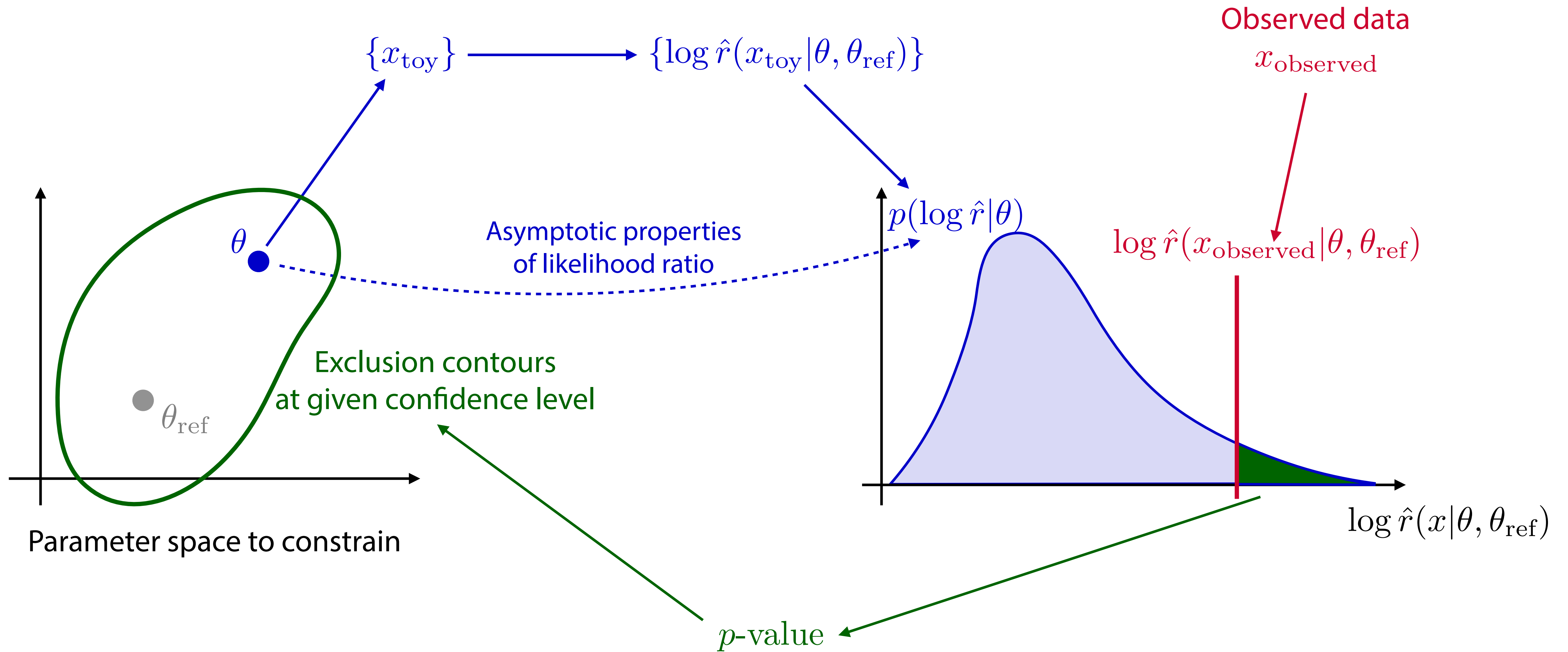


# Limit setting (frequentist, standard ATLAS / CMS practice)

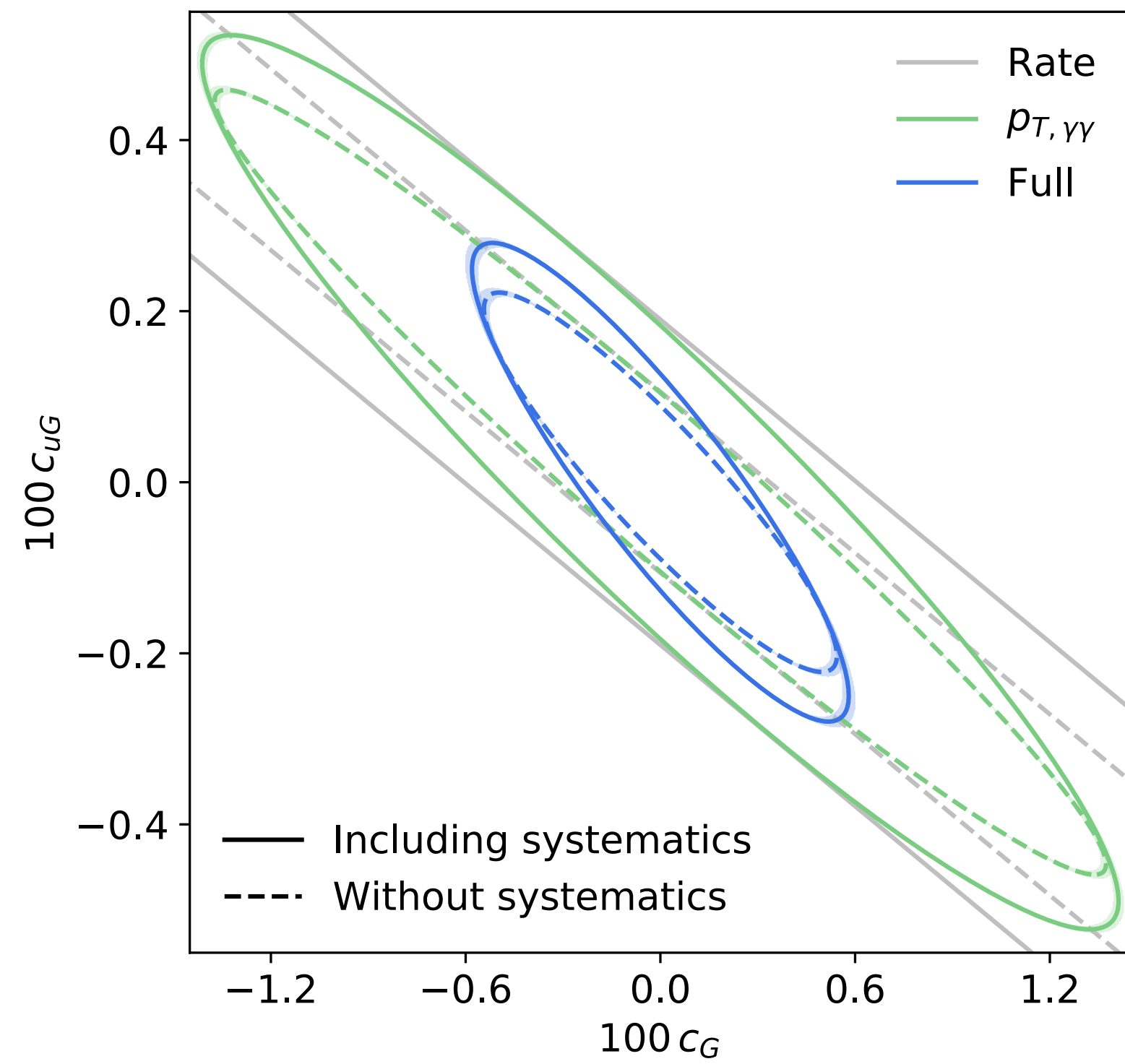




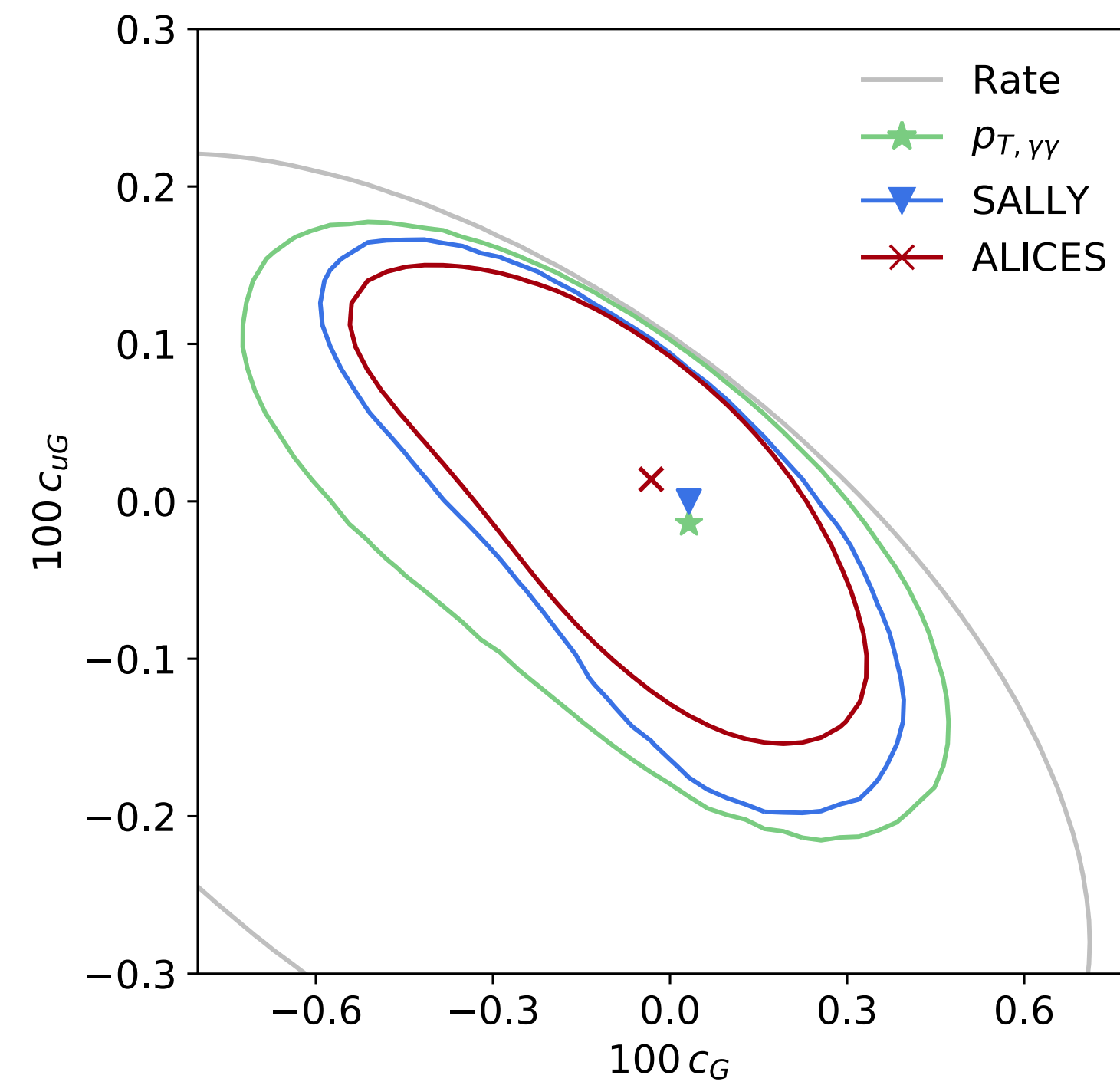
# Limit setting (frequentist, standard ATLAS / CMS practice)



# Types of Inference Results & Sensitivity Summaries



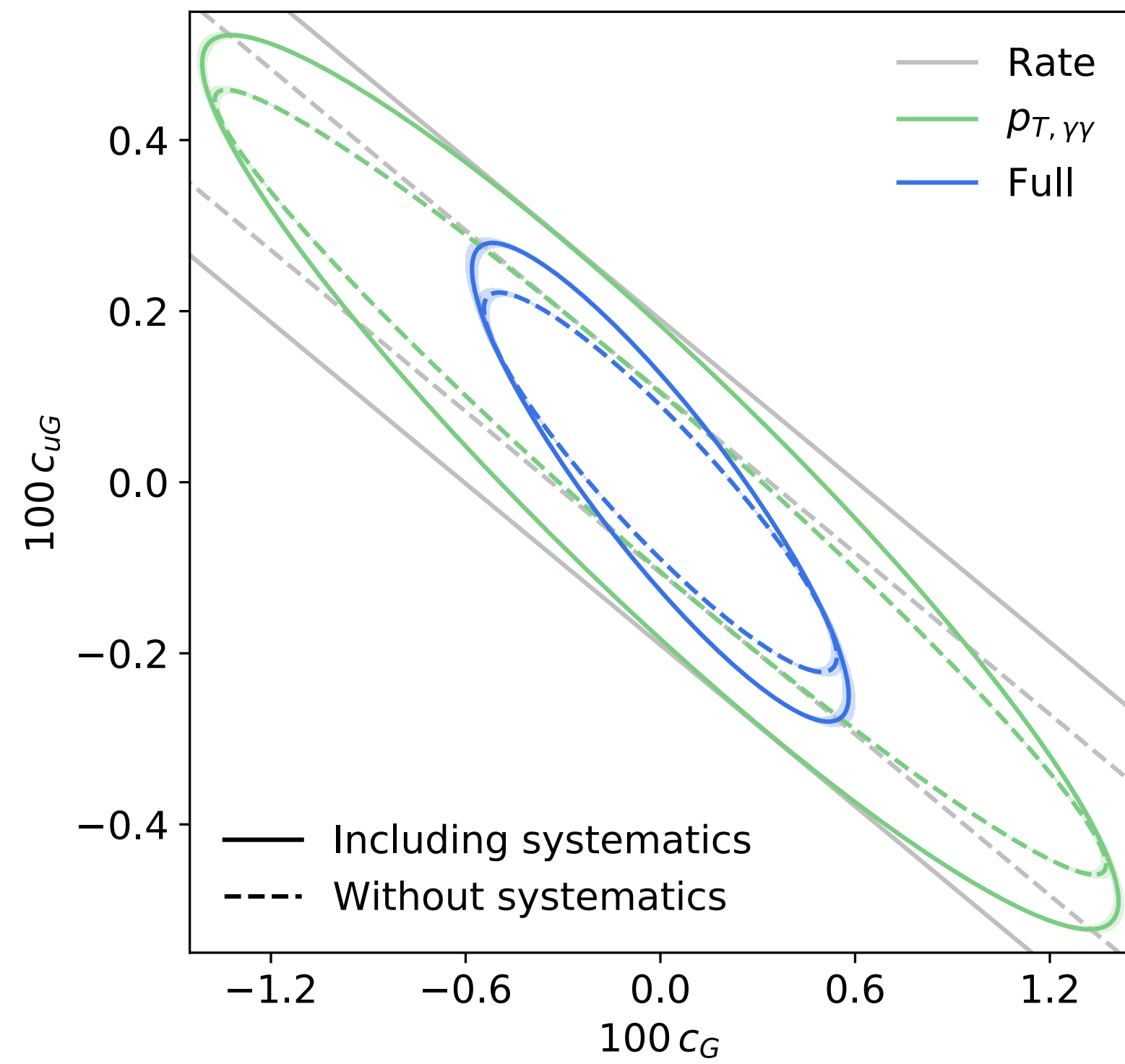
Elliptical Contours  
(with / without systematics)



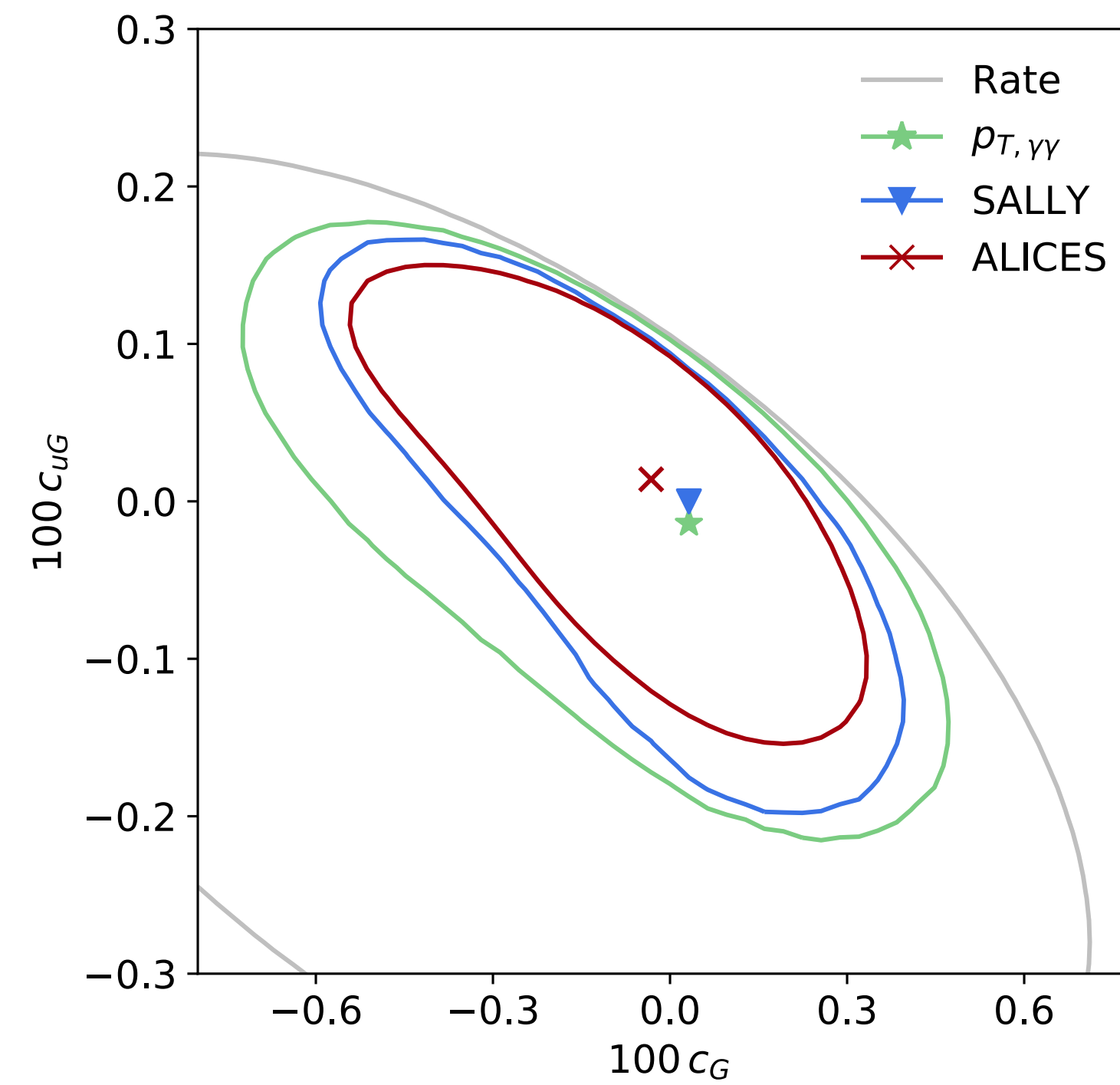
Non-Elliptical Contours

Fisher Information  
&  
Information Geometry

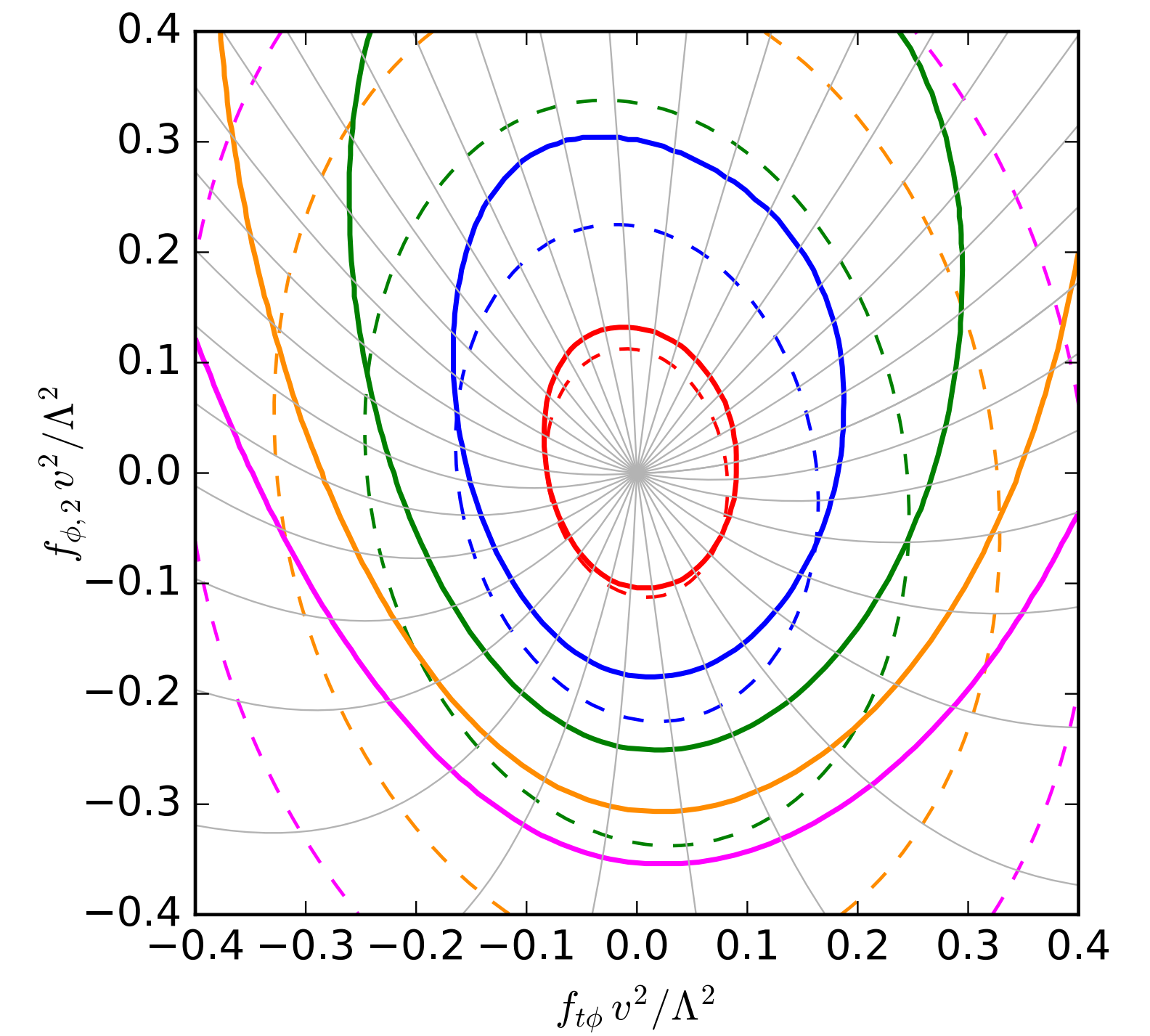
# Types of Inference Results & Sensitivity Summaries



Elliptical Contours  
(with / without systematics)



Non-Elliptical Contours



Fisher Information  
&  
Information Geometry

**MadMiner automates all of these methods.**

[JB, F. Kling, I. Espejo, K. Cranmer 1907.10621]

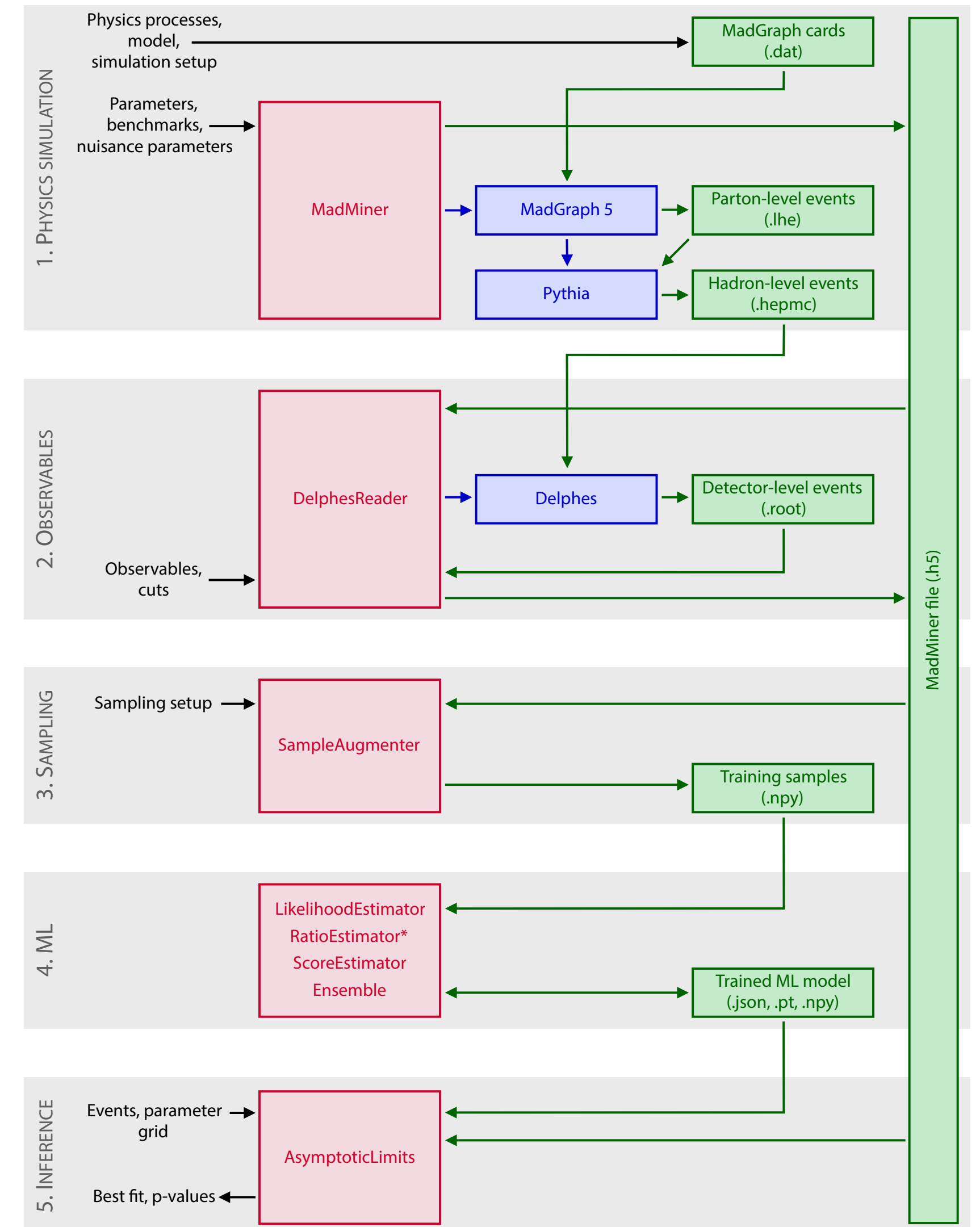


# Automation

New Python package **MadMiner** makes it straightforward to apply the new techniques to LHC problems

- Out of the box: Pheno-level analyses
  - MadGraph, Pythia, Delphes
  - Systematic uncertainties from PDF / scale variation
- Scalable to state-of-the-art experimental tools
  - Mostly requires bookkeeping of fully differential cross sections
- Modular interface
  - Extensive documentation
  - Embedded into Python / ML ecosystem

[JB, F. Kling, I. Espejo, K. Cranmer 1907.10621]



### MadMiner: Machine learning-based inference for particle physics

Johann Brehmer,<sup>1,\*</sup> Felix Kling,<sup>2,3,†</sup> Irina Espejo,<sup>1,‡</sup> and Kyle Cranmer<sup>1,§</sup>

<sup>1</sup>Center for Data Science and Center for Cosmology and Particle Physics,  
New York University, New York, NY 10003, USA

<sup>2</sup>Department of Physics and Astronomy, University of California, Irvine, CA 92697, USA

<sup>3</sup>SLAC National Accelerator Laboratory, 2575 Sand Hill Road, Menlo Park, CA 94025, USA

Precision measurements at the LHC often require analyzing high-dimensional event data for subtle kinematic signatures, which is challenging for established analysis methods. Recently, a powerful family of multivariate inference techniques that leverage both matrix element information and machine learning has been developed. This approach neither requires the reduction of high-dimensional data to summary statistics nor any simplifications to the underlying physics or detector response. In this paper we introduce MadMiner, a Python module

### MadMiner: Machine learning-based inference for particle physics

By Johann Brehmer, Felix Kling, Irina Espejo, and Kyle Cranmer

pypi package 0.6.3
build passing
docs failing
chat on gitter
code style black
License MIT
DOI 10.5281/zenodo.1489147

arXiv 1907.10621

#### Introduction

Particle physics processes are usually modeled with complex Monte-Carlo simulations of the hard process, parton shower, and detector interactions. These simulators typically do not admit a tractable likelihood function: given a (potentially high-dimensional) set of observables, it is usually not possible to calculate the probability of these observables for some model parameters. Particle physicists usually tackle this problem of "likelihood-free inference" by hand-picking a few "good" observables or summary statistics and filling histograms of them. But this conventional

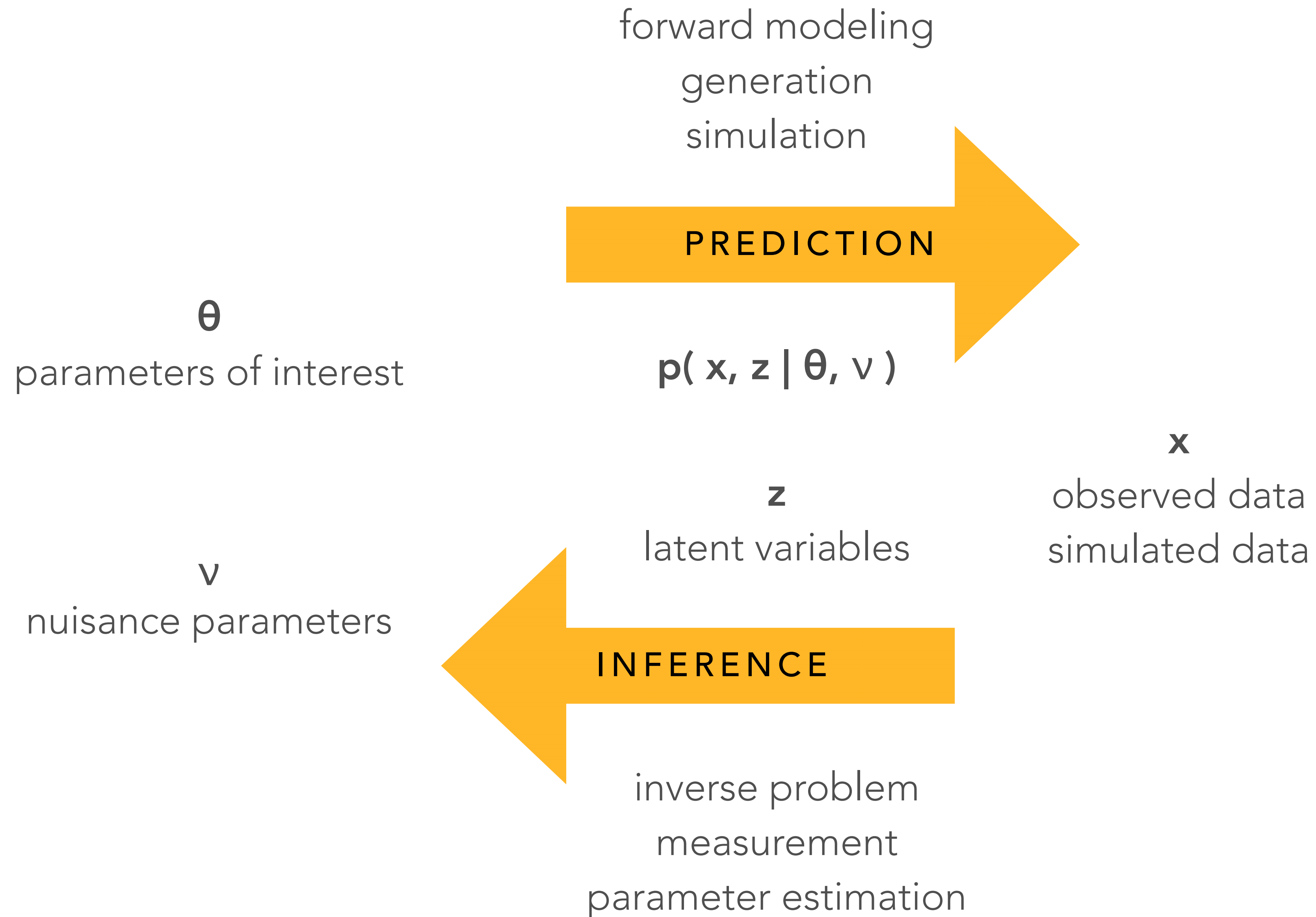
Repository and tutorials:  
[github.com/johannbrehmer/madminer](https://github.com/johannbrehmer/madminer)

Paper with detailed explanations:  
[1907.10621](https://arxiv.org/abs/1907.10621)

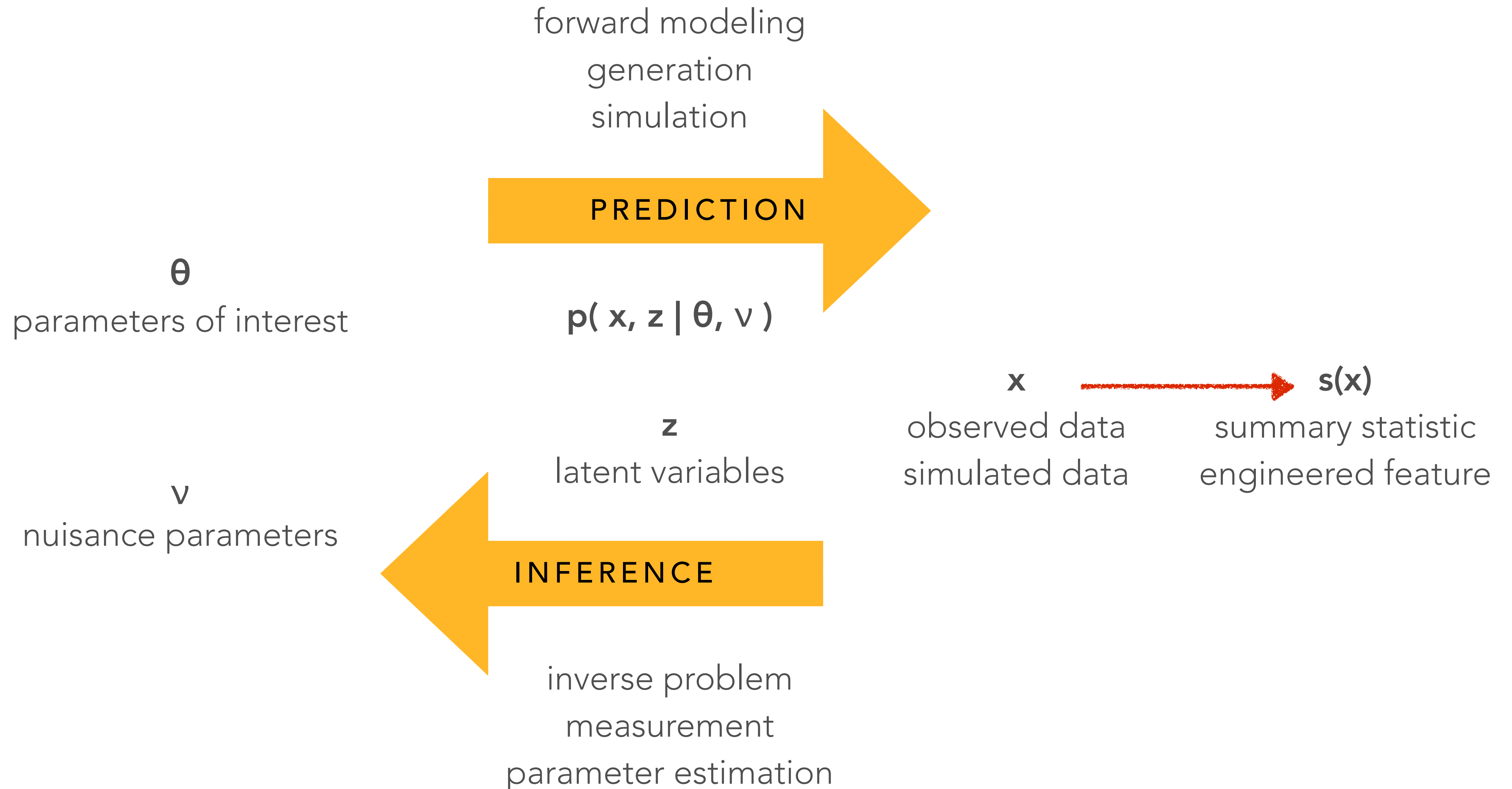
Installation:  
`pip install madminer`

API documentation:  
[madminer.readthedocs.io](https://madminer.readthedocs.io)

# Forward modeling and inverse problems



# Forward modeling and inverse problems



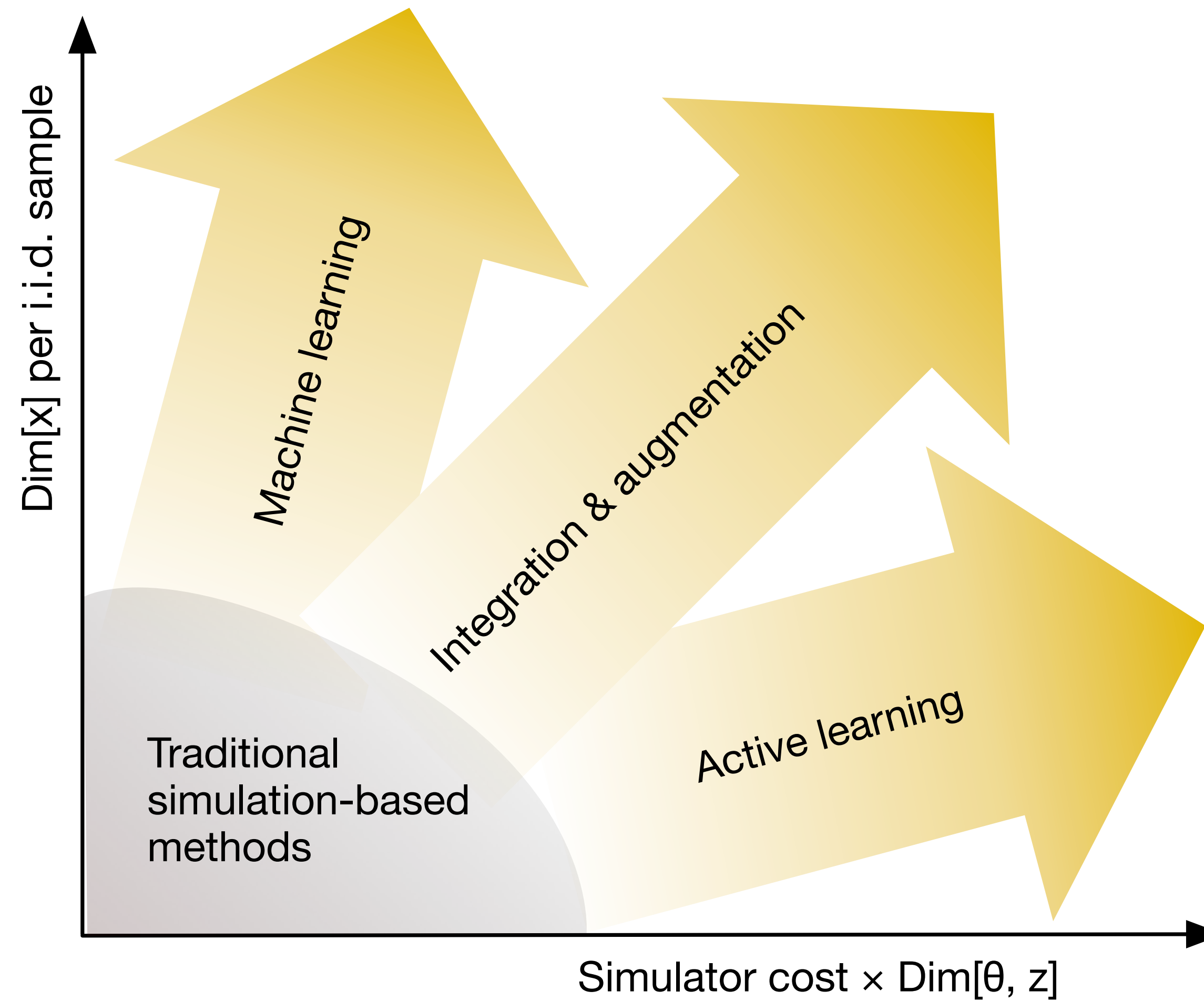


# The frontier of simulation-based inference

Kyle Cranmer<sup>a,b,1</sup>, Johann Brehmer<sup>a,b</sup>, and Gilles Louppe<sup>c</sup>

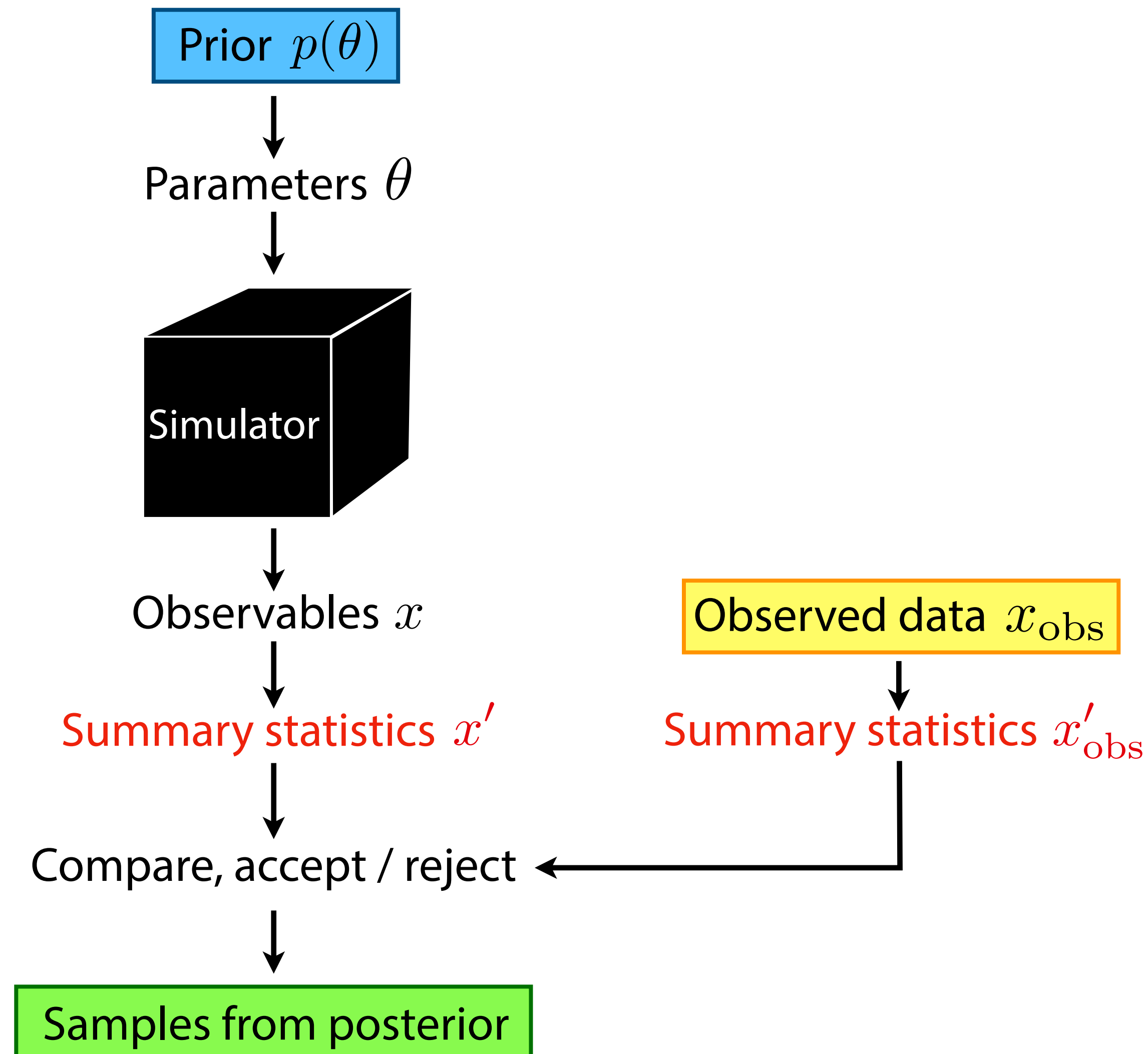
<sup>a</sup>Center for Cosmology and Particle Physics, New York University, USA; <sup>b</sup>Center for Data Science, New York University, USA; <sup>c</sup>Montefiore Institute, University of Liège, Belgium

April 3, 2020



# Approximate Bayesian Computation (ABC)

[D. Rubin 1984]

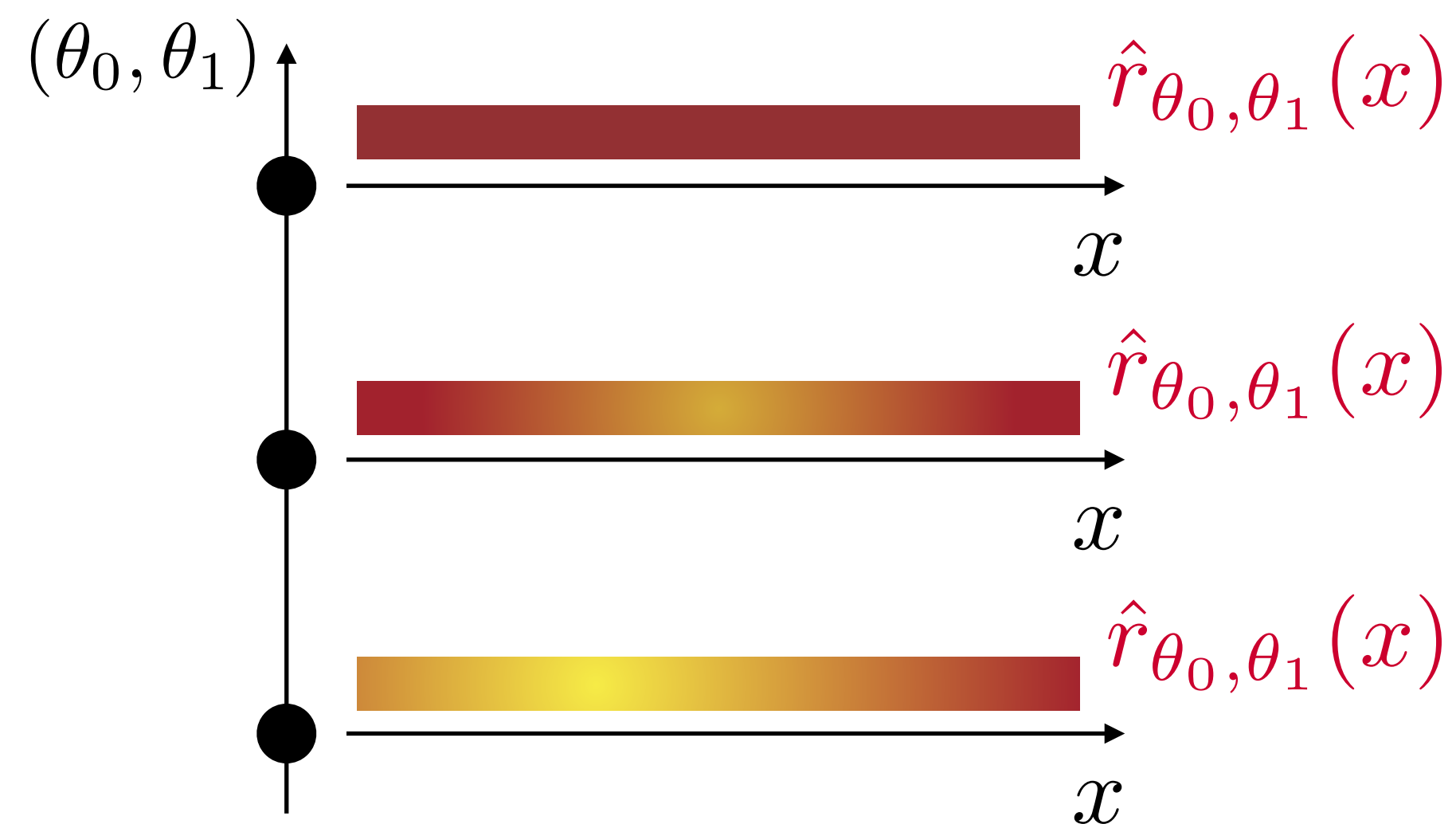


- Compression to summary statistics and acceptance threshold reduce quality of inference
- Rejection algorithm can be very sample inefficient

# Two types of likelihood ratio estimators

## A) Point by point:

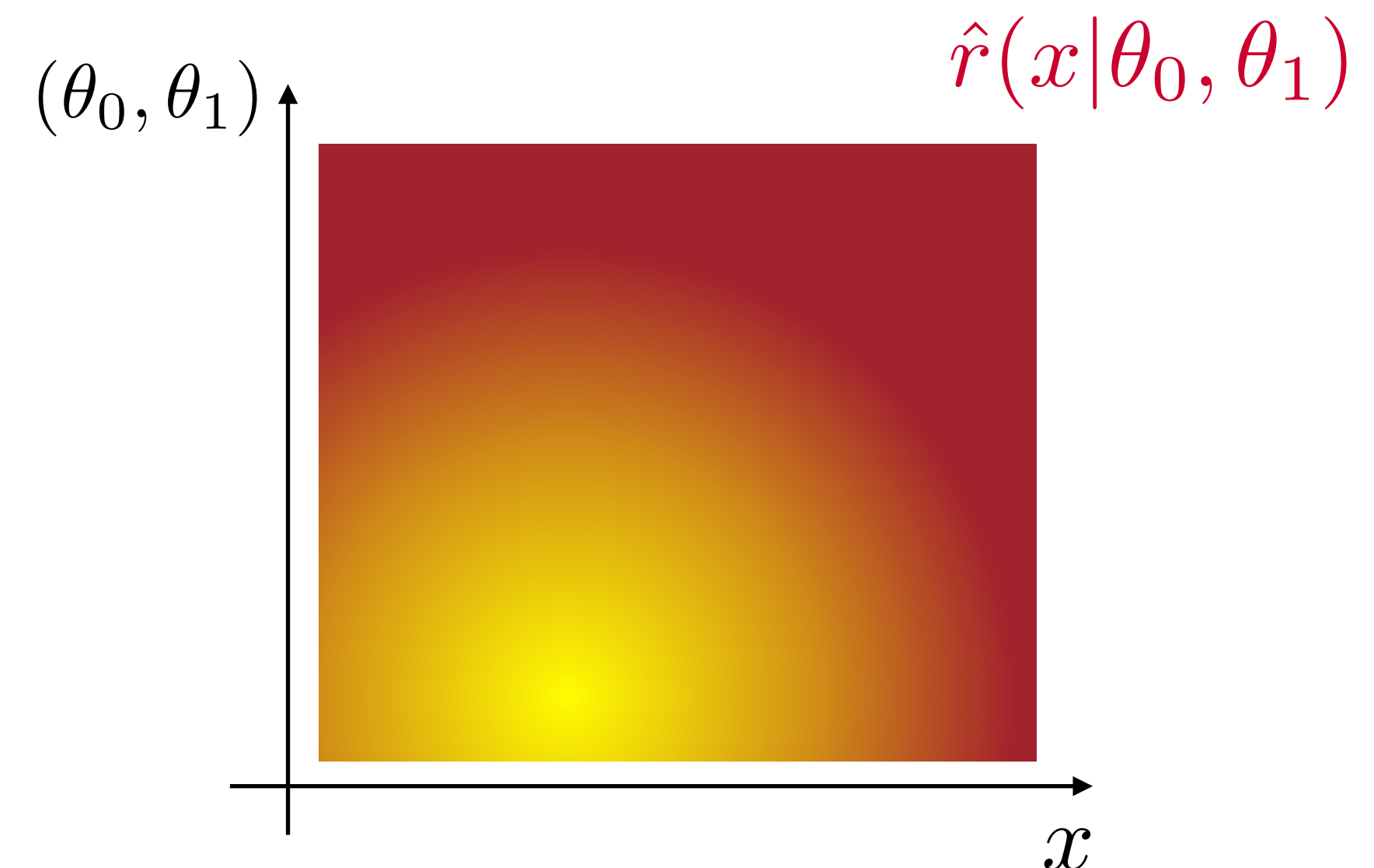
- first, define grid of parameter points  $\{(\theta_0, \theta_1)\}$
- for each combination  $(\theta_0, \theta_1)$ , create separate estimator  $\hat{r}_{\theta_0, \theta_1}(x)$
- final results can be interpolated between grid points



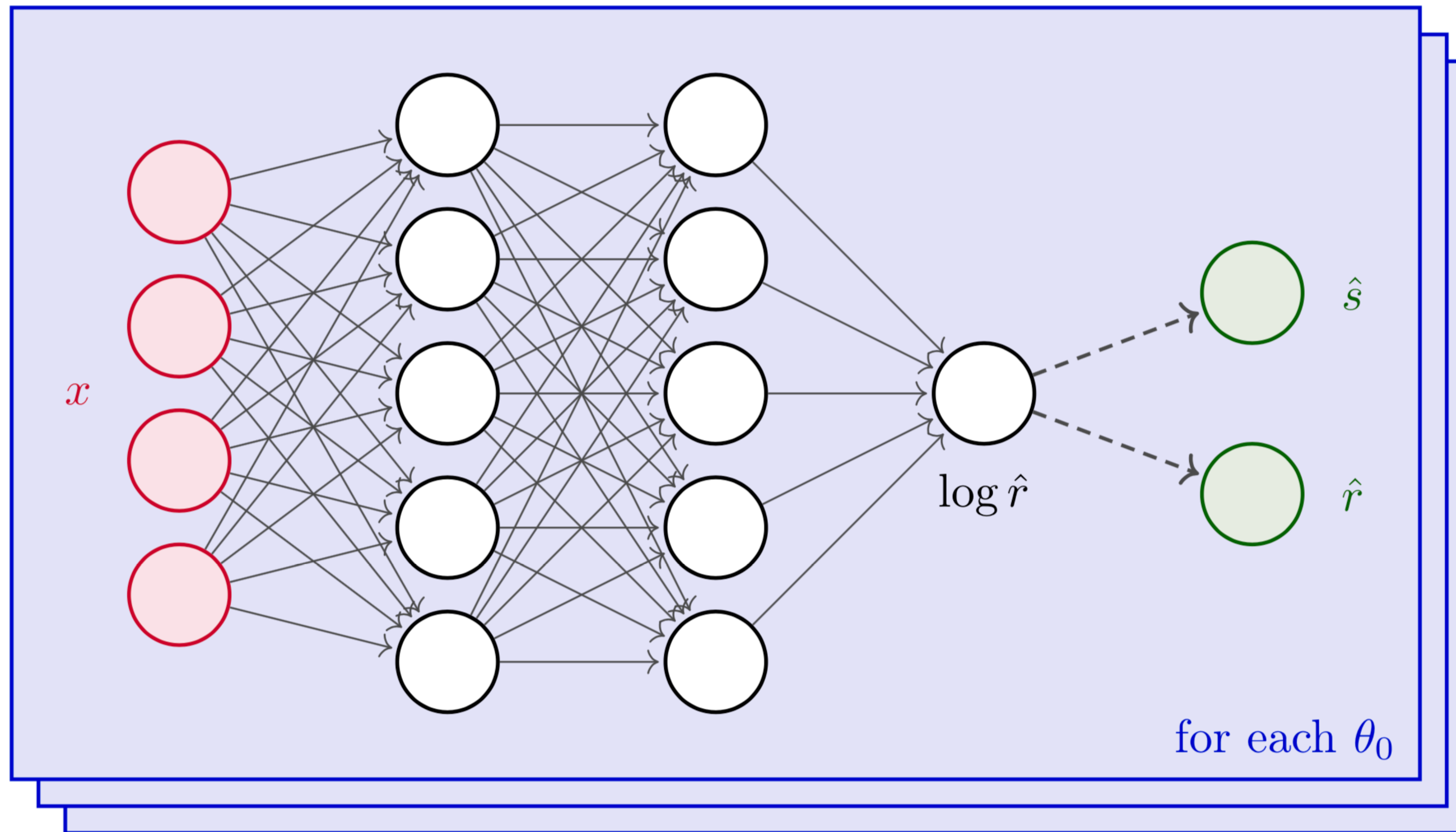
## B) Parameterized:

[K. Cranmer, J. Pavez, G. Louppe 1506.02169;  
P. Baldi et al. 1601.07913]

- create one estimator  $\hat{r}(x|\theta_0, \theta_1)$  that is a function of  $\theta_0$  and  $\theta_1$
- no further interpolation necessary
- “borrows information” from close points

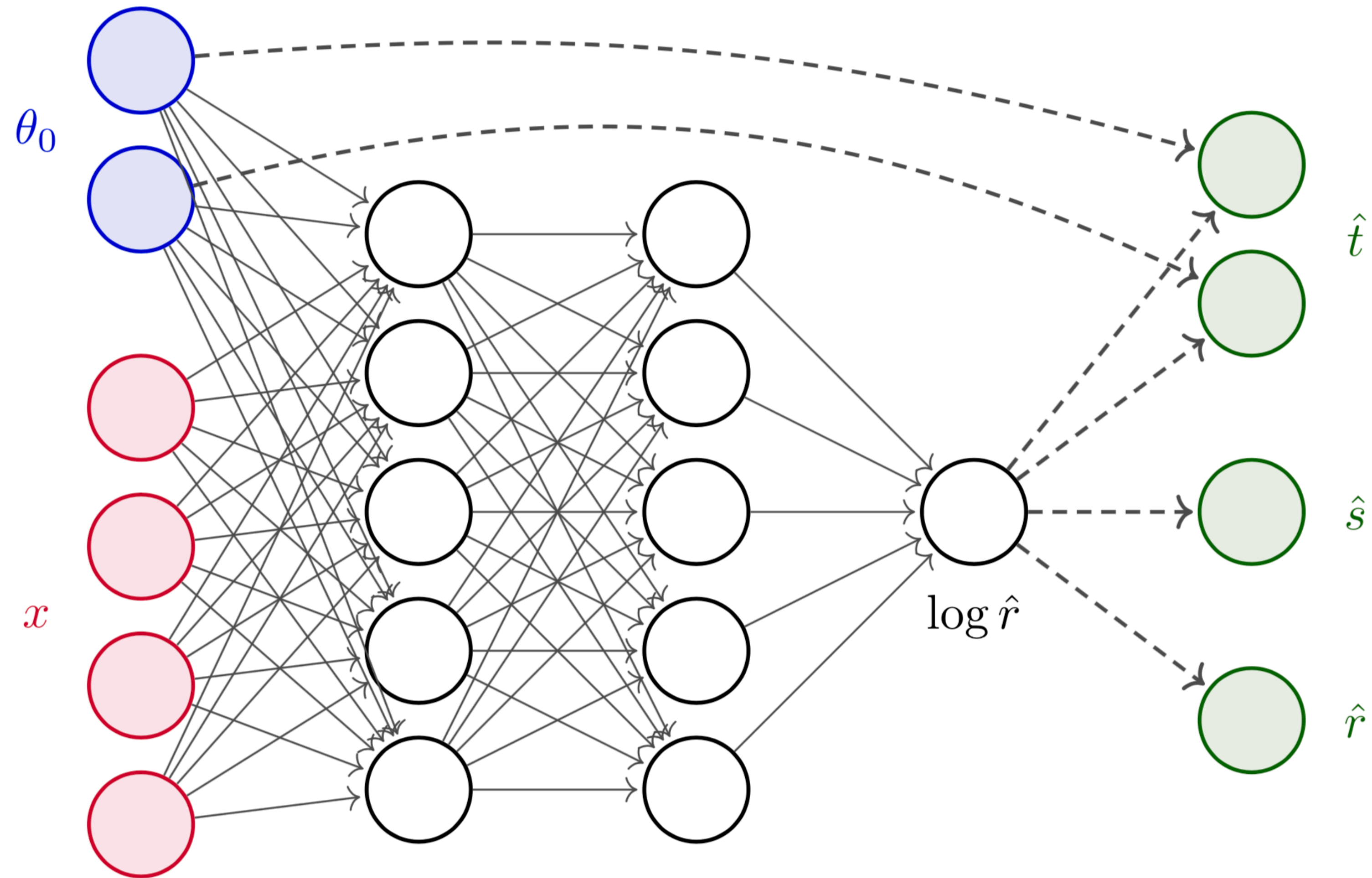


# Point by point

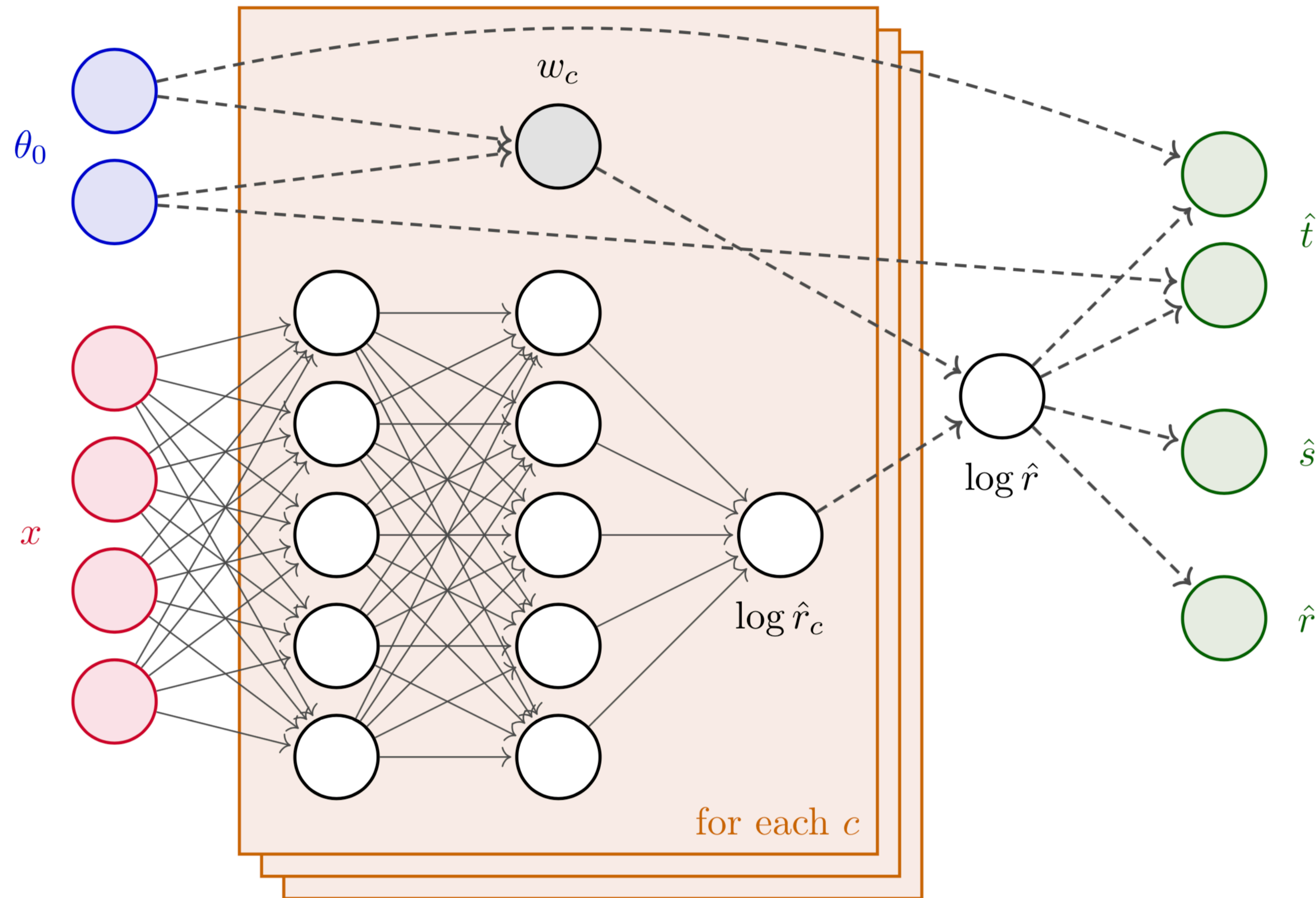




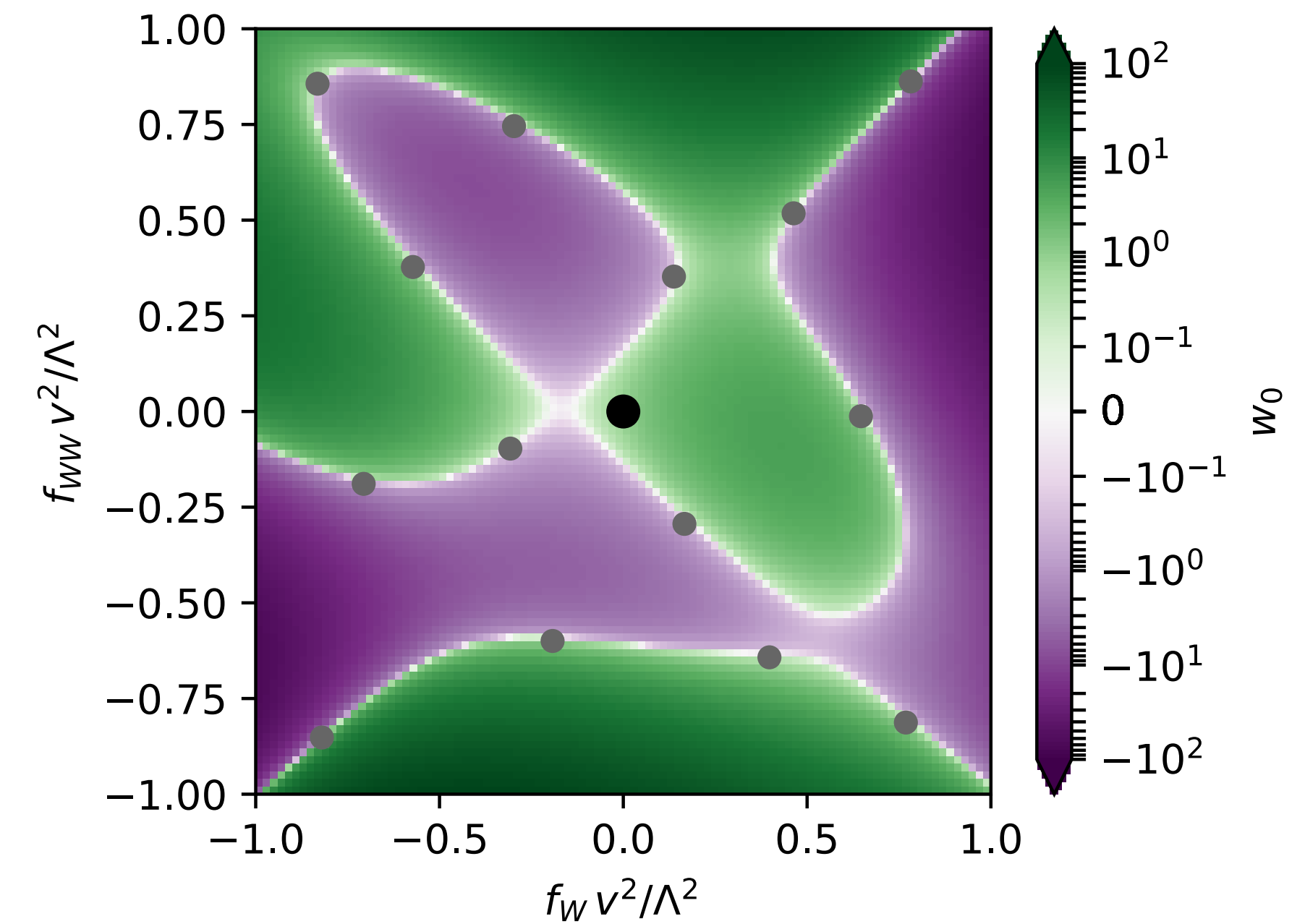
# (Physics-agnostic) parameterized estimators



# Morphing-aware parameterized estimators



$$\hat{r}(x|\theta_0, \theta_1) = \sum_c w_c(\theta_0) \hat{r}_c(x)$$

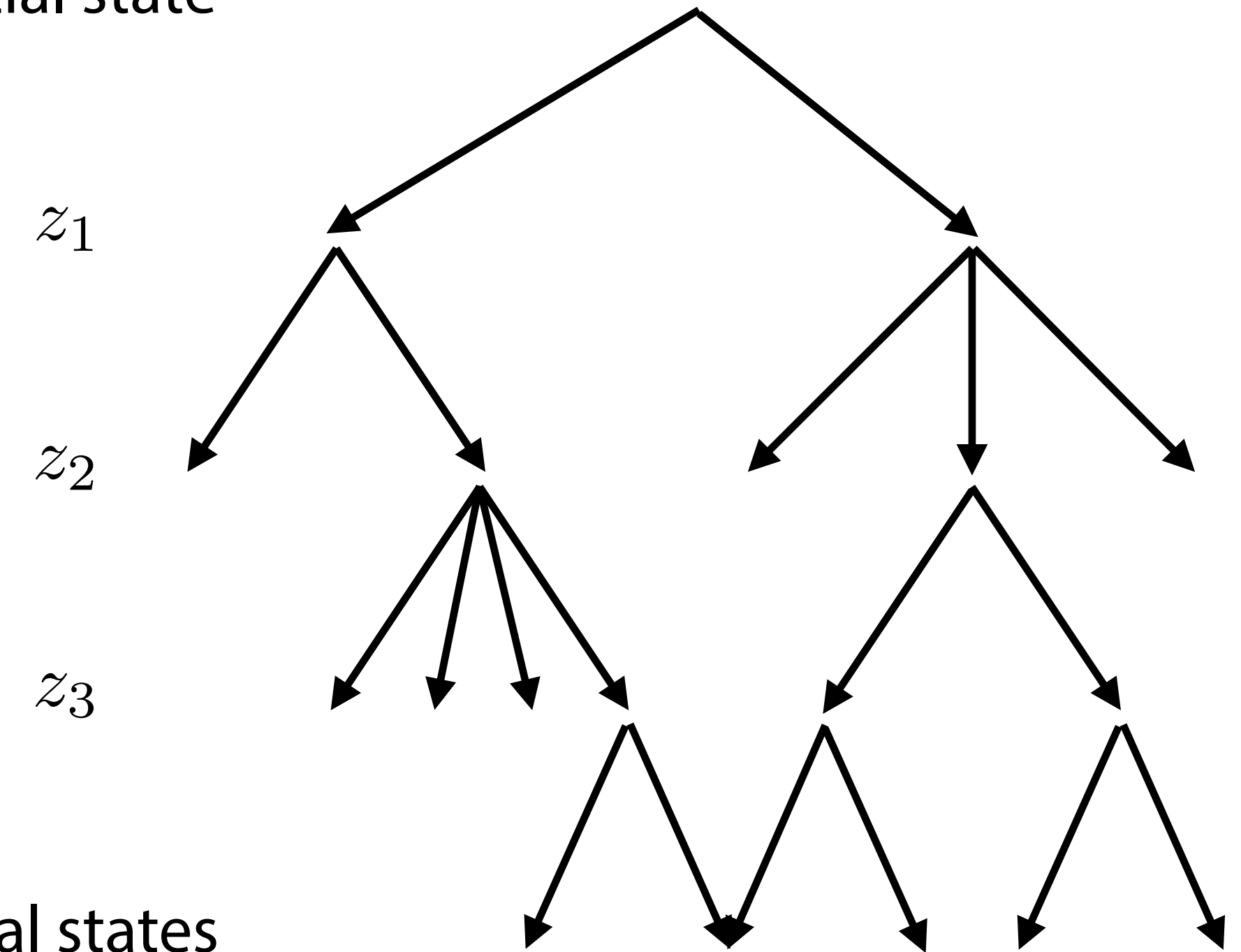


# Mining gold from any simulation

- Computer simulation typically evolve along a tree-like structure of successive random branchings
- The probabilities of each branching  $p_i(z_i|z_{i-1}, \theta)$  are often clearly defined in the code:

```
if random() > 0.1 + 2.5 * model_parameter:  
    do_one_thing()  
else:  
    do_another_thing()
```

Initial state



Final states



# Mining gold from any simulation

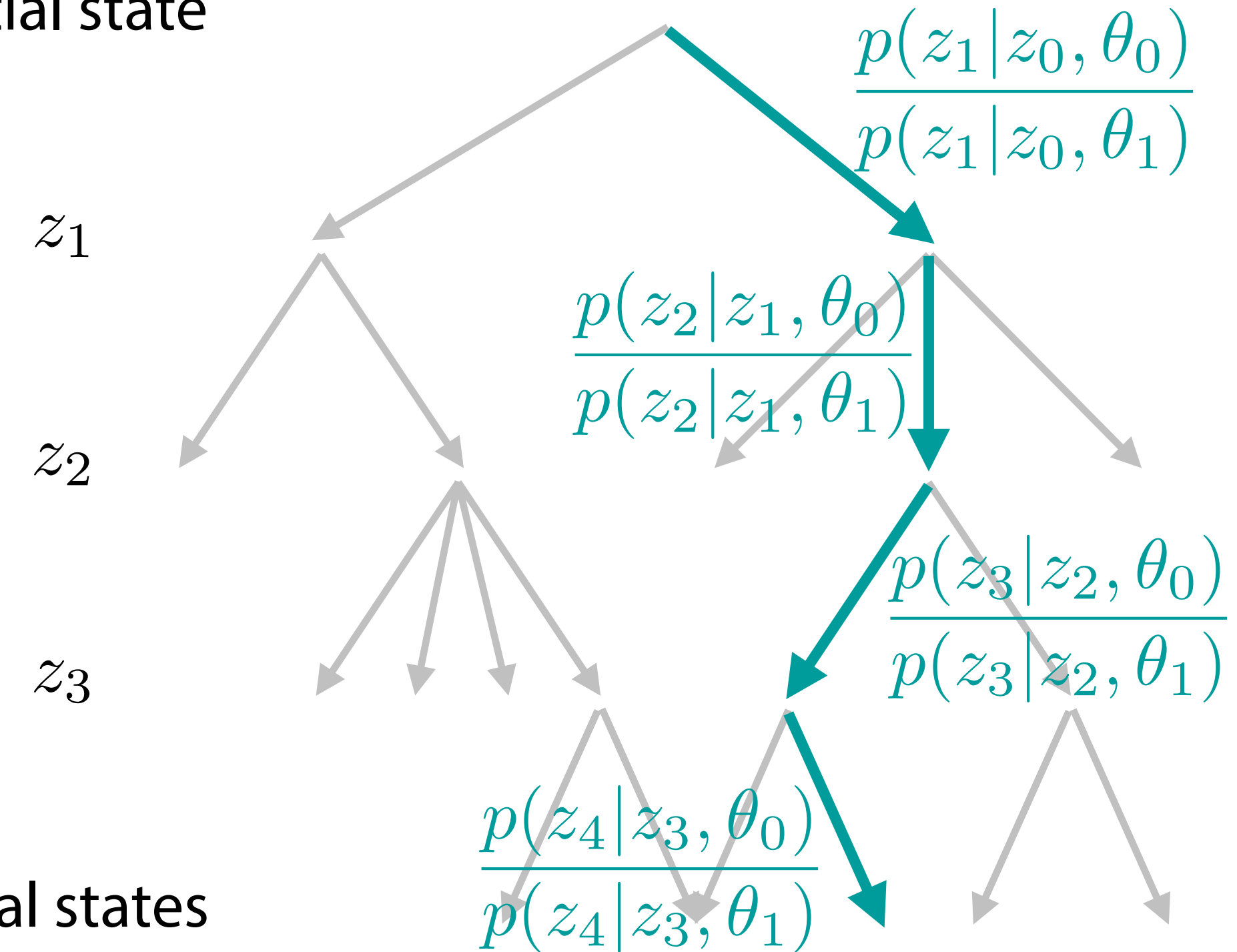
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```
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else:  
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```

- For each run of the simulator, we can calculate the probability of the chosen path for different values of the parameters, and the “joint likelihood ratio”:

$$r(x, z|\theta_0, \theta_1) = \frac{p(x, z|\theta_0)}{p(x, z|\theta_1)} = \prod_i \frac{p(z_i|z_{i-1}, \theta_0)}{p(z_i|z_{i-1}, \theta_1)}$$

Initial state





# Mining gold: A family of new inference techniques

Method	Simulate	Extract		NN estimates	Asympt. exact	Generative
		$r(x, z)$	$t(x, z)$			
ROLR	$\theta_0 \sim \pi(\theta), \theta_1$	✓		$\hat{r}(x \theta_0, \theta_1)$	✓	
CASCAL	$\theta_0 \sim \pi(\theta), \theta_1$		✓	$\hat{r}(x \theta_0, \theta_1)$	✓	
ALICE	$\theta_0 \sim \pi(\theta), \theta_1$		✓	$\hat{r}(x \theta_0, \theta_1)$	✓	
RASCAL	$\theta_0 \sim \pi(\theta), \theta_1$	✓	✓	$\hat{r}(x \theta_0, \theta_1)$	✓	
ALICES	$\theta_0 \sim \pi(\theta), \theta_1$	✓	✓	$\hat{r}(x \theta_0, \theta_1)$	✓	
SCANDAL	$\theta \sim \pi(\theta)$		✓	$\hat{p}(x \theta)$	✓	✓
SALLY	$\theta_{\text{ref}}$		✓	$\hat{t}(x \theta_{\text{ref}})$	in local approx.	
SALLINO	$\theta_{\text{ref}}$		✓	$\hat{t}(x \theta_{\text{ref}})$	in local approx.	

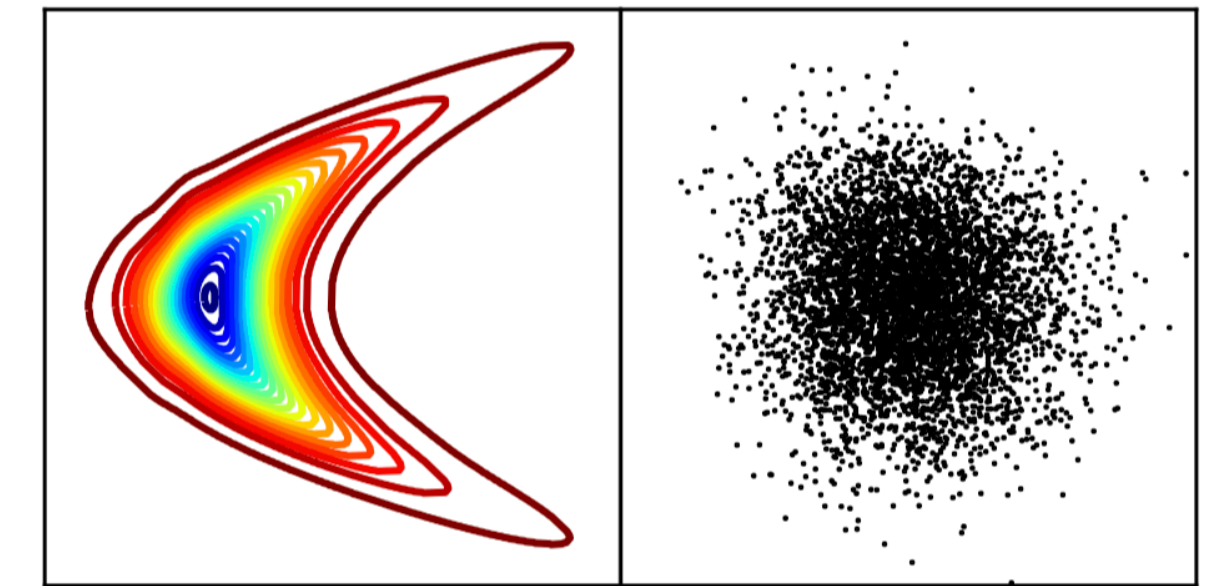
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Performance gains with cross-entropy-based loss  
 [M. Stoye, JB, K. Cranmer, G. Louppe, J. Pavez 1808.00973]

# Mining gold: A family of new inference techniques

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ALICES	$\theta_0 \sim \pi(\theta), \theta_1$	✓	✓	$\hat{r}(x \theta_0, \theta_1)$	✓	
SCANDAL	$\theta \sim \pi(\theta)$		✓	$\hat{p}(x \theta)$	✓	✓
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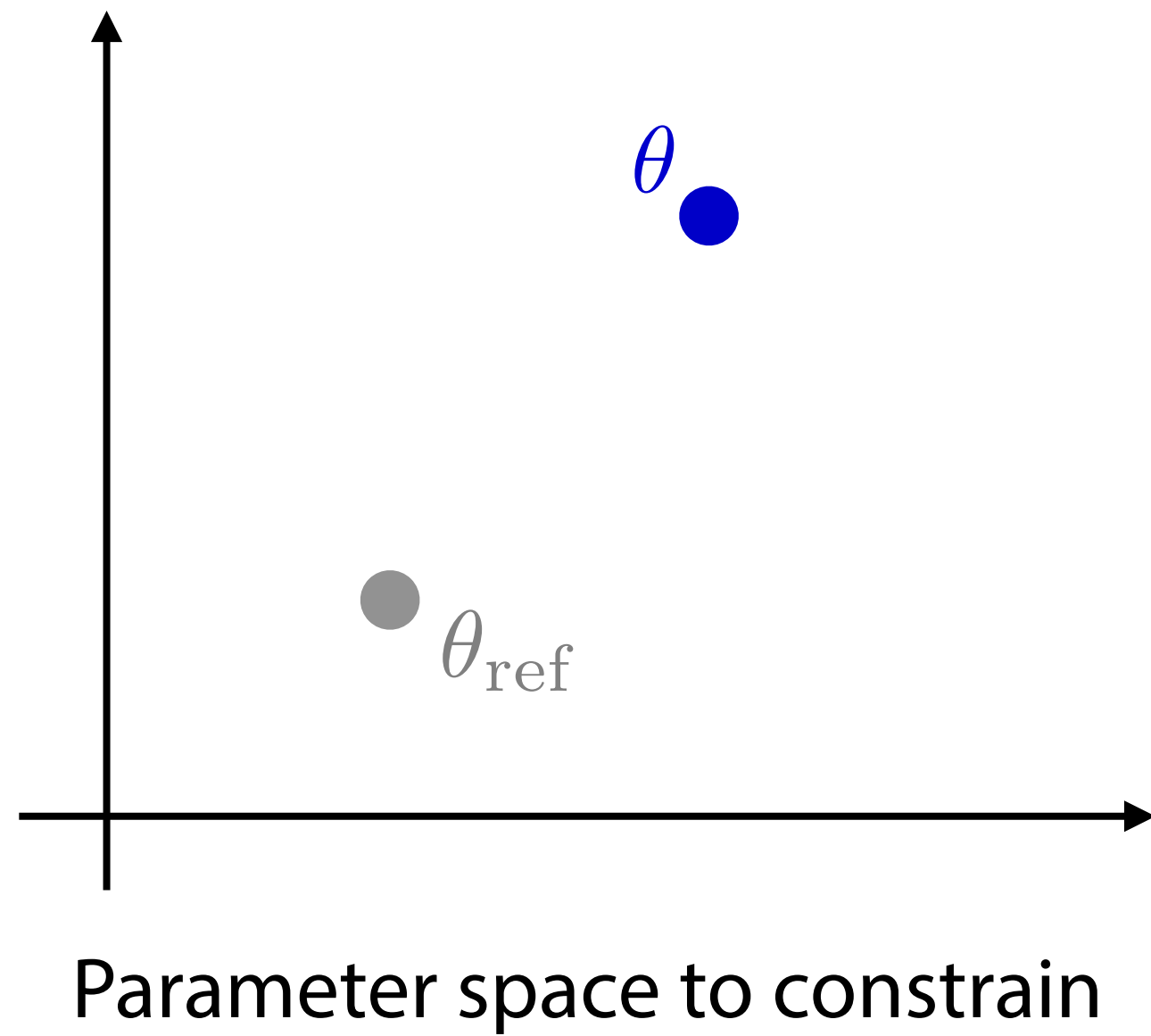
Combination with state-of-the-art conditional neural density estimators, e.g. normalizing flows

[everything by G. Papamakarios:  
G. Papamakarios, T. Pavlakou, I. Murray 1705.07057;  
G. Papamakarios, D. Sterratt, I. Murray 1805.07226; ...]

# Limit setting (frequentist, standard ATLAS / CMS practice)

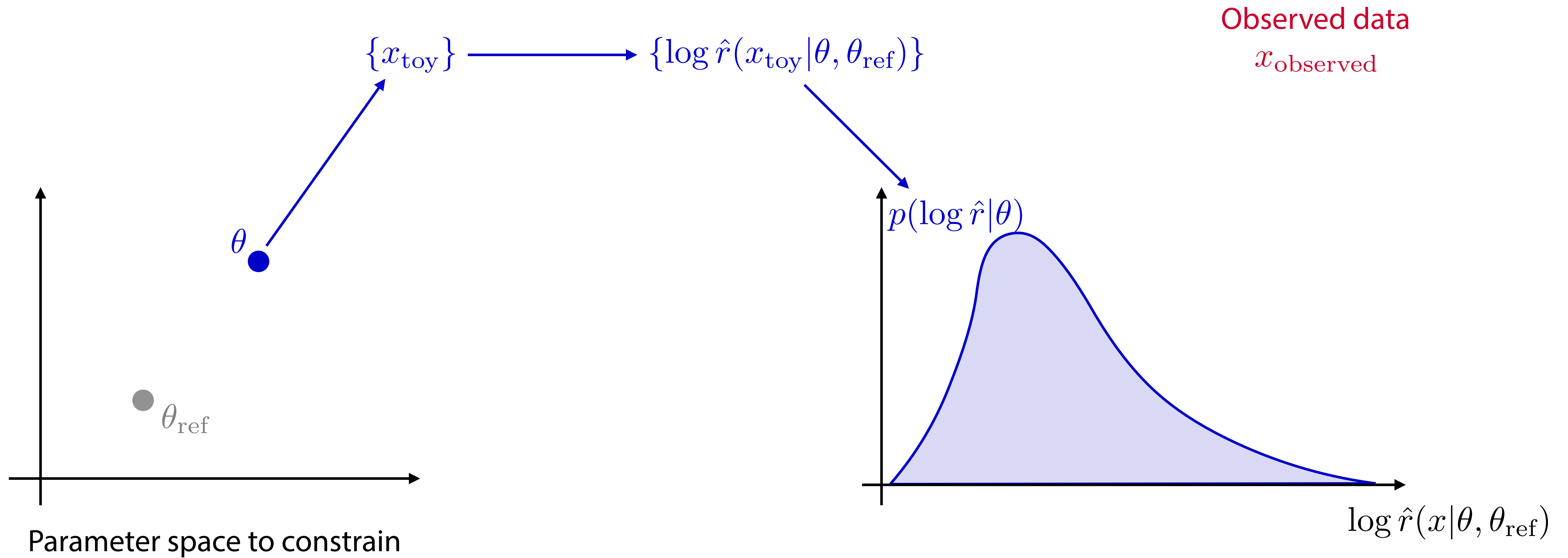
Observed data

$x_{\text{observed}}$

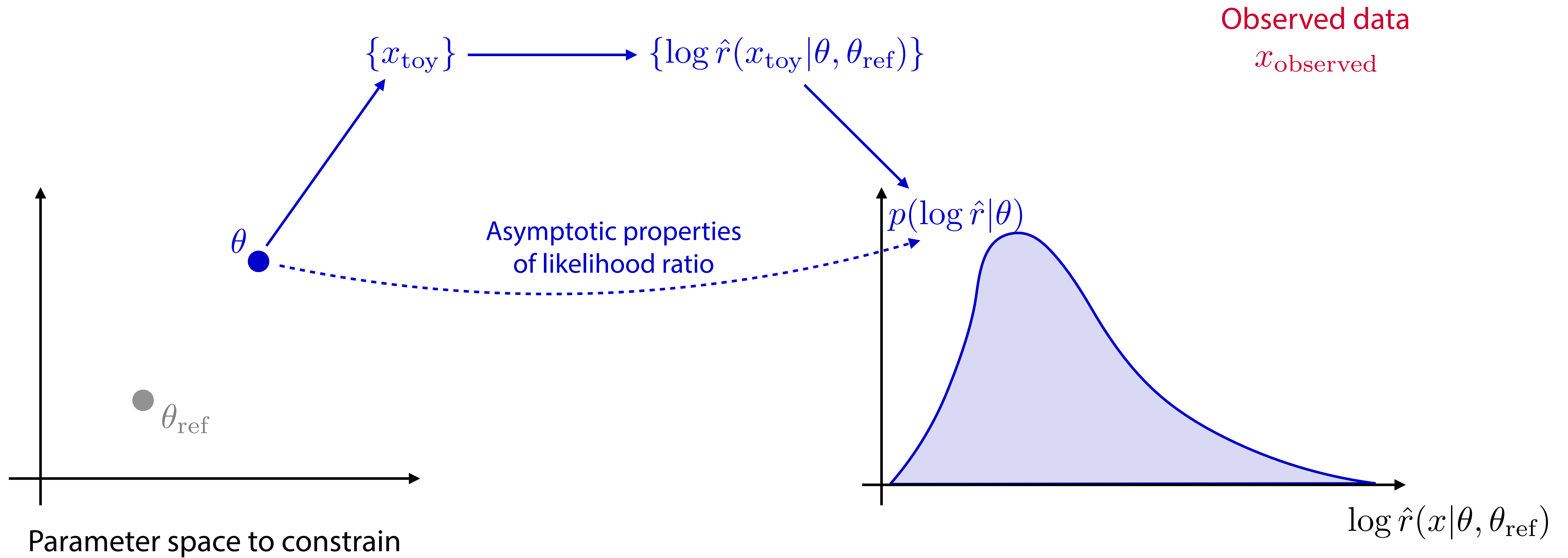




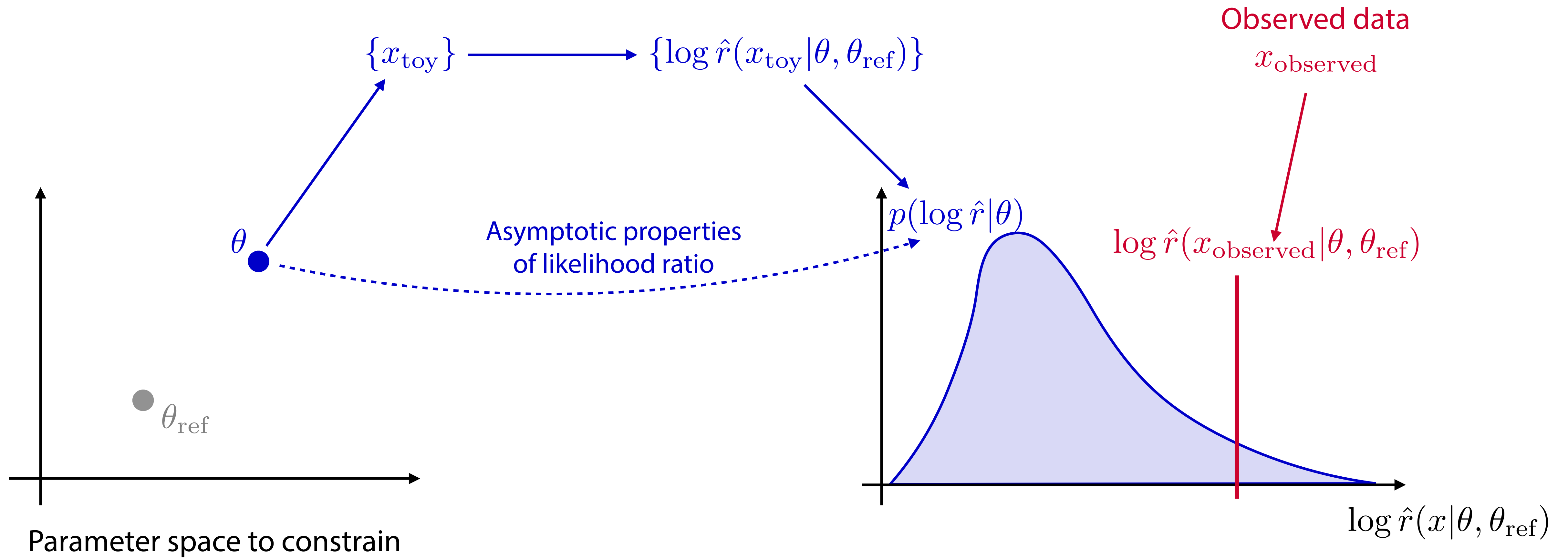
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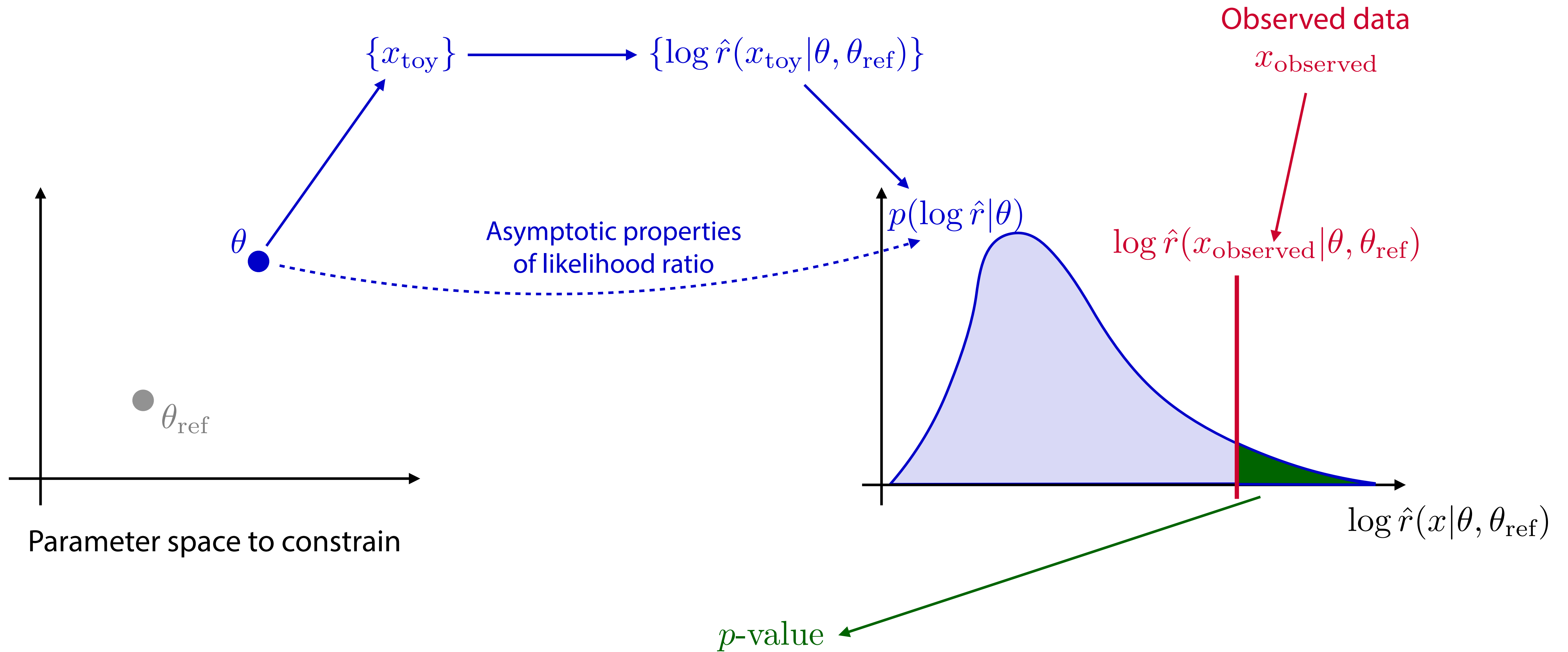
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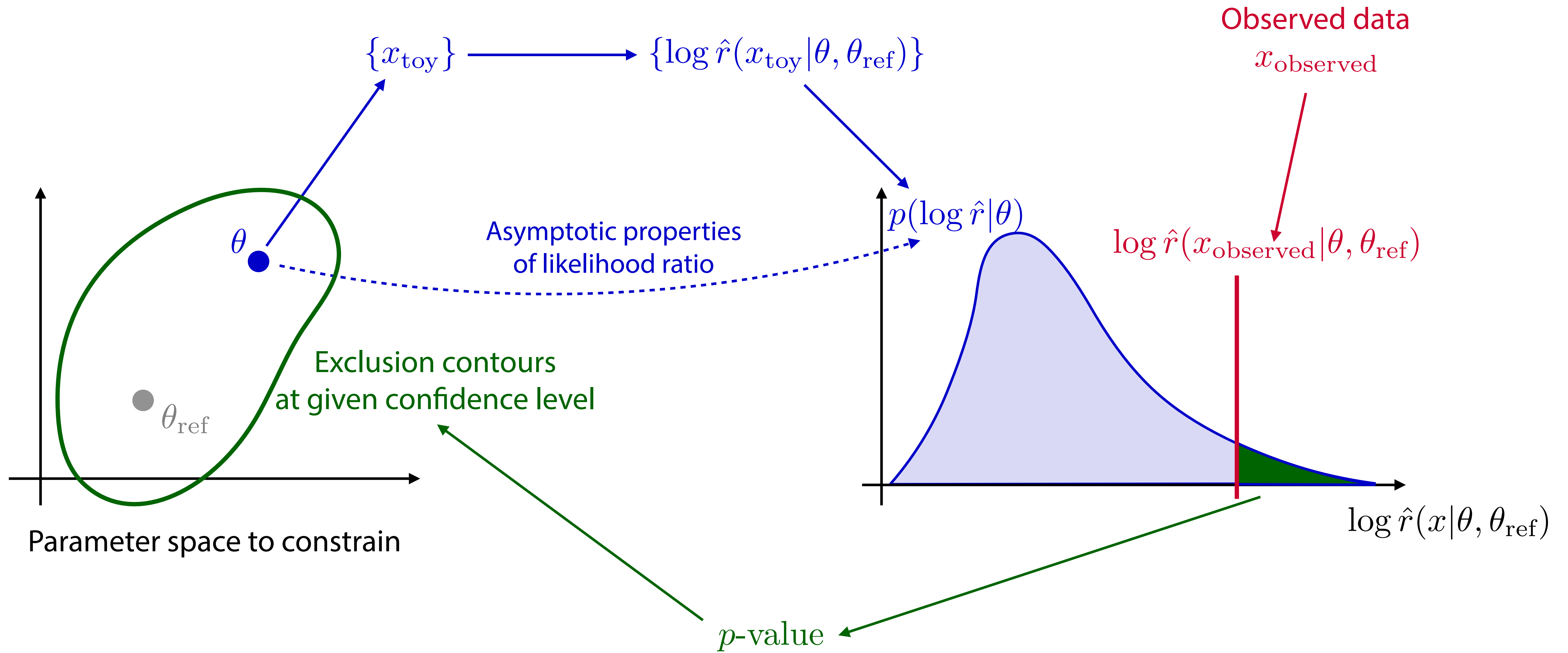


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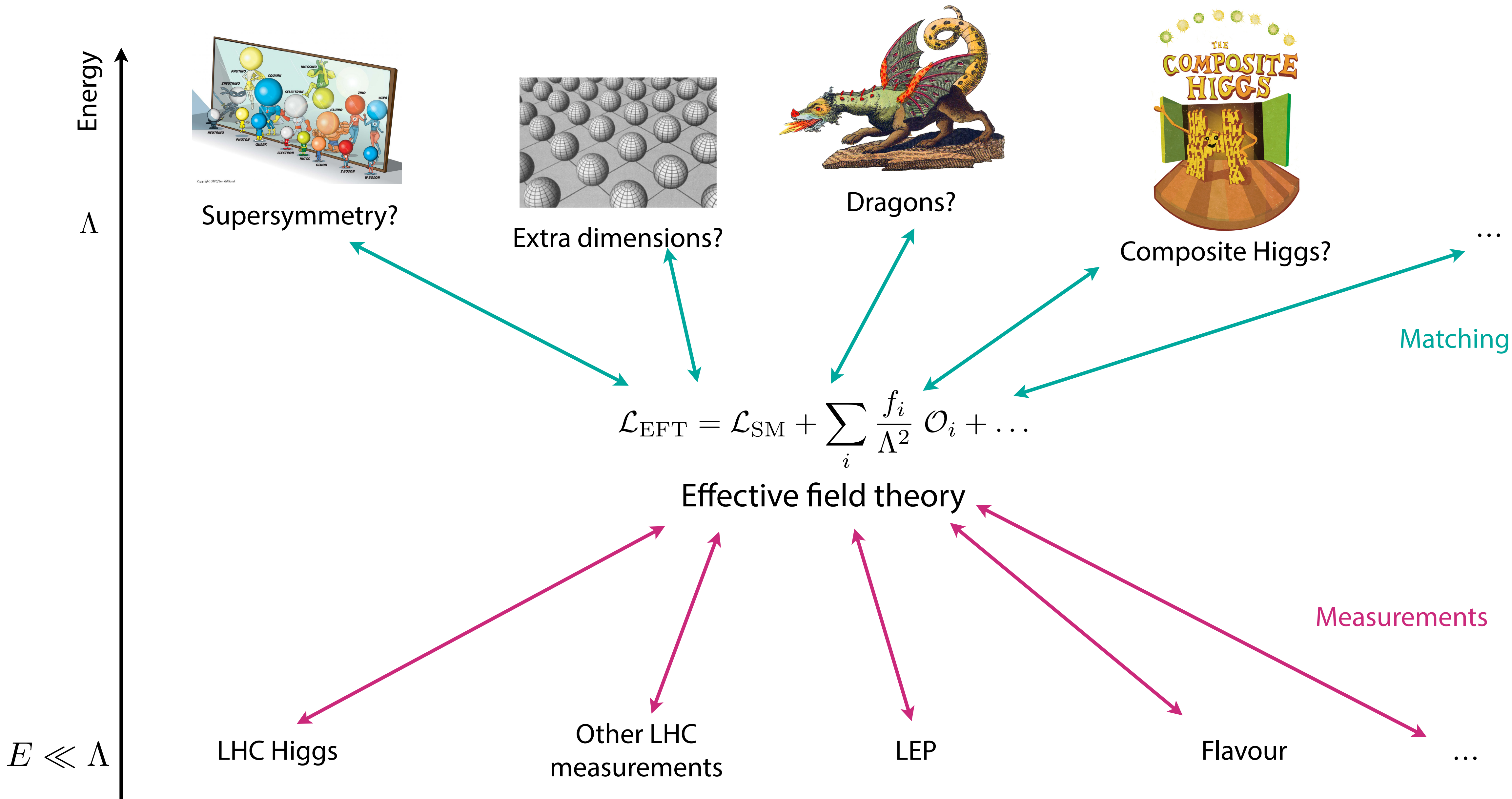


# Limit setting (frequentist, standard ATLAS / CMS practice)



# Effective field theory

[STFC / Ben Gililand, Sean Carroll, Friedrich Justin Bertuch 1806, symmetry]



## Operators

- all possible interactions between SM particles mediated by new physics
- fixed by SM particles + SM symmetries + expansion in  $1/\Lambda$ , independent of high-energy physics
- affect rates + kinematics

---

$\mathcal{O}_{\phi,1} = (D_\mu \phi)^\dagger \phi \phi^\dagger D^\mu \phi$	$\mathcal{O}_{GG} = (\phi^\dagger \phi) G_{\mu\nu}^a G^{\mu\nu a}$
$\mathcal{O}_{\phi,2} = \frac{1}{2} \partial_\mu (\phi^\dagger \phi) \partial^\mu (\phi^\dagger \phi)$	$\mathcal{O}_{BB} = -\frac{g'^2}{4} (\phi^\dagger \phi) B_{\mu\nu} B^{\mu\nu}$
$\mathcal{O}_{\phi,3} = \frac{1}{3} (\phi^\dagger \phi)^3$	$\mathcal{O}_{WW} = -\frac{g^2}{4} (\phi^\dagger \phi) W_{\mu\nu}^a W^{\mu\nu a}$
$\mathcal{O}_{\phi,4} = (\phi^\dagger \phi) (D_\mu \phi)^\dagger D^\mu \phi$	$\mathcal{O}_{BW} = -\frac{g g'}{4} (\phi^\dagger \sigma^a \phi) B_{\mu\nu} W^{\mu\nu a}$
	$\mathcal{O}_B = \frac{i g'}{2} (D^\mu \phi)^\dagger D^\nu \phi B_{\mu\nu}$
	$\mathcal{O}_W = \frac{i g}{2} (D^\mu \phi)^\dagger \sigma^a D^\nu \phi W_{\mu\nu}^a$

---

$$\mathcal{L}_{\text{EFT}} = \mathcal{L}_{\text{SM}} + \sum_i \frac{f_i}{\Lambda^2} \mathcal{O}_i + \dots$$

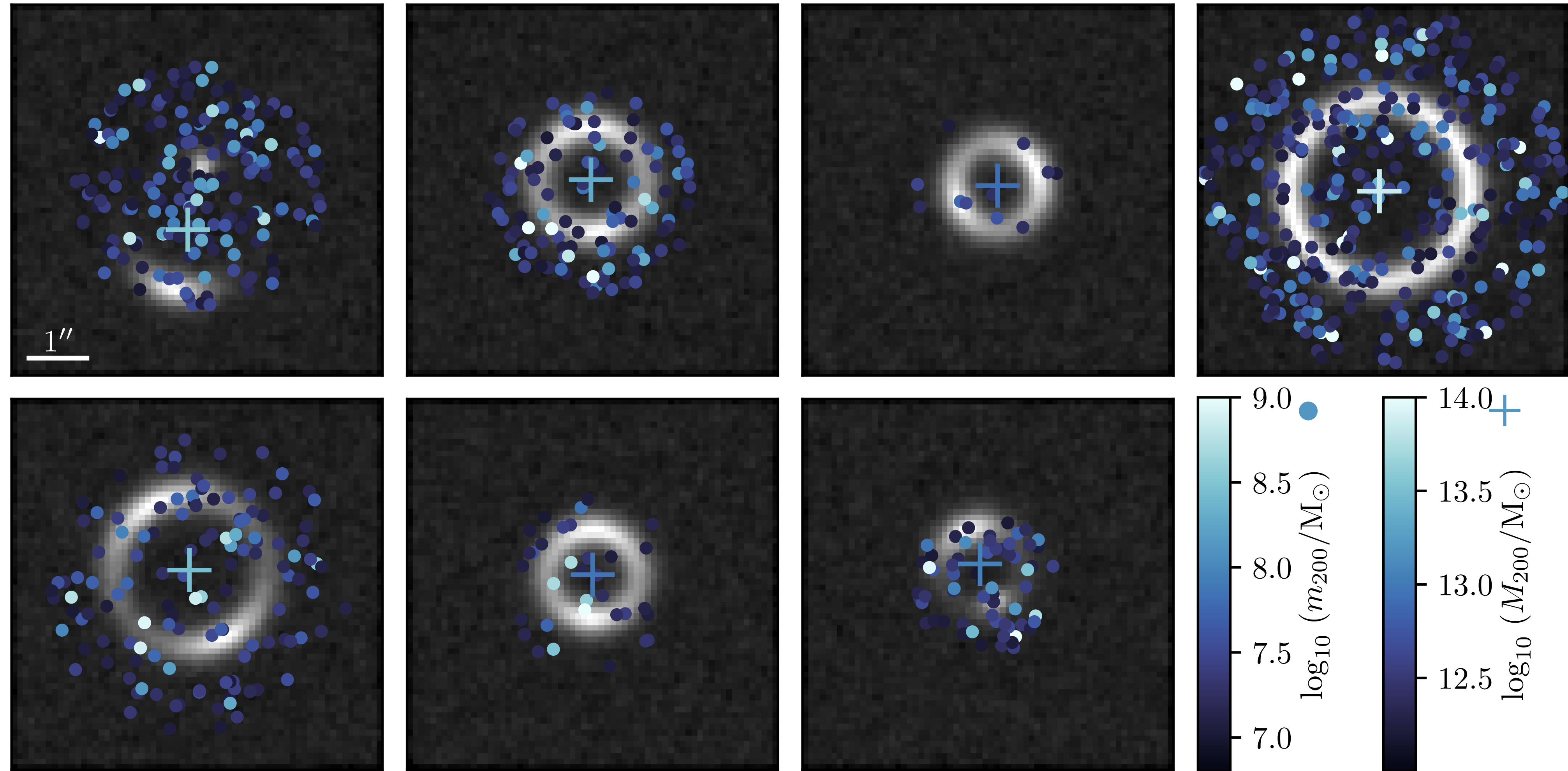
## Wilson coefficients

- precise measurement of these parameters is one of the most important goals of the LHC
- can be translated to high-energy physics parameters

## Higher-order terms

- suppressed by additional factors of  $E^2 / \Lambda^2$

# Strong lensing: simulated images





# Strong lensing: expected posterior

