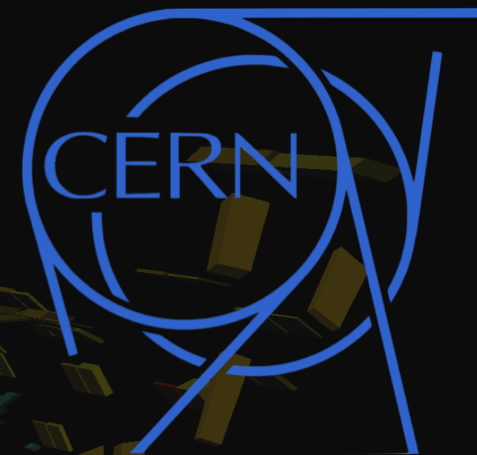




UC SANTA CRUZ



Searches for Supersymmetry

Dr. Giordon Stark 
Lepton Photon 2021 Manchester
Jan 11th, 2022
 giordonstark.com



Run: 300800

Event: 2418777995

2016-06-04 03:47:03 **Q1**

if you can read this, you're too close

 indico.cern.ch/e/949705/

"SUSY is just around the corner."
— Carlos Wagner



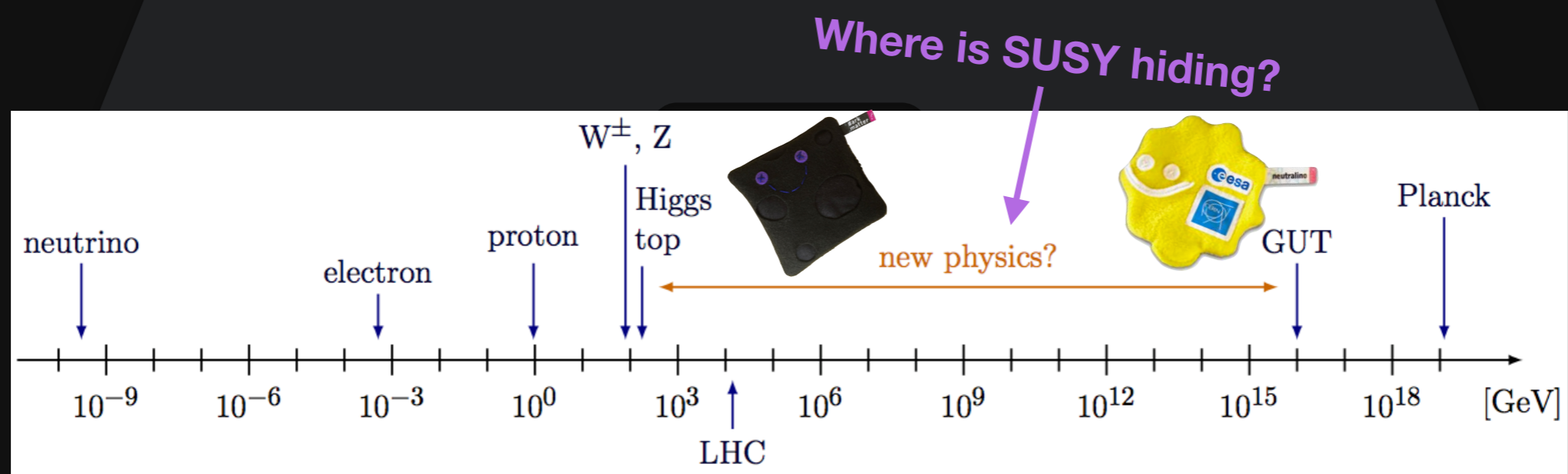
Overview of Today

- A highlight of searches for **new physics inspired by SUSY** across R-Parity Conserving (**RPC**) R-Parity Violating (**RPV**) scenarios
 - Strong,
 - 3rd Generation (**3G**),
 - Electroweak (**EWK**), and
 - Long-Lived Particles (**LL** or **LLP**)
- and how these results fit into a global picture

KEY

PRODUCTION
LUMI
FIT

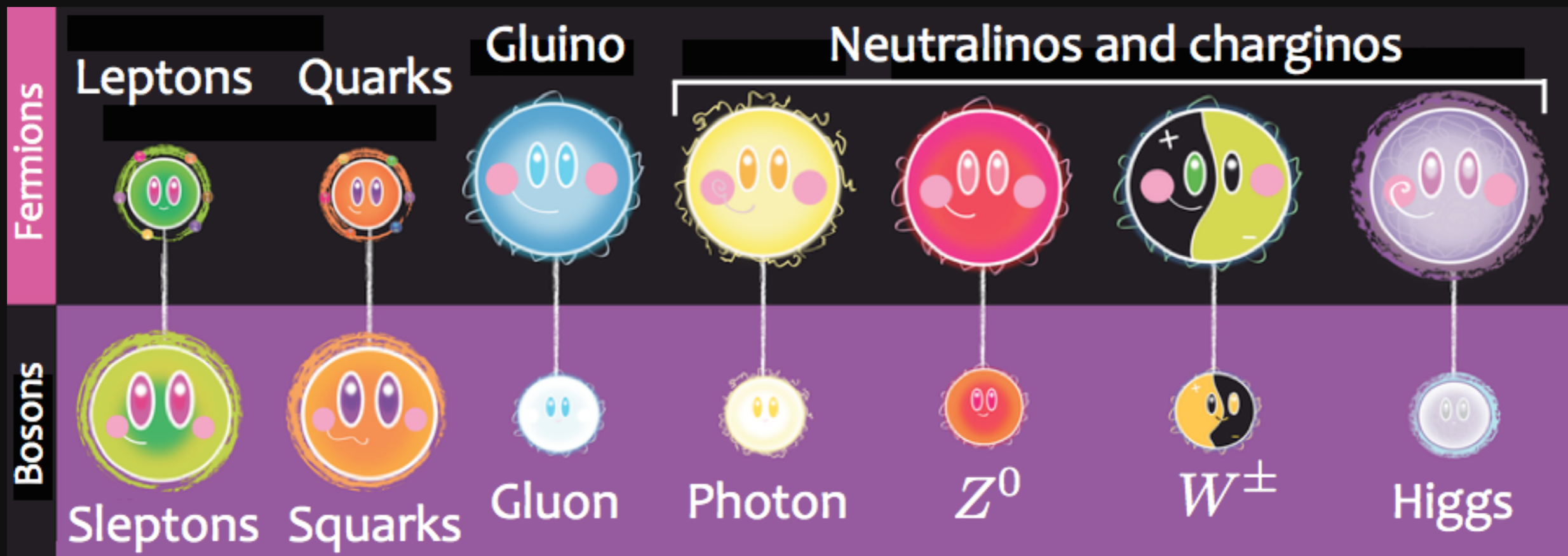
Strong (\tilde{g})
3G (\tilde{t}, \tilde{b})
EWK ($\tilde{\ell}, \tilde{\nu}, \tilde{\chi}_{1\pm}$)
RPVLL (λ', λ'')
36 fb ⁻¹
80 fb ⁻¹
139 fb ⁻¹
cut and count
multi-bin



! Symmetry between fermions and bosons

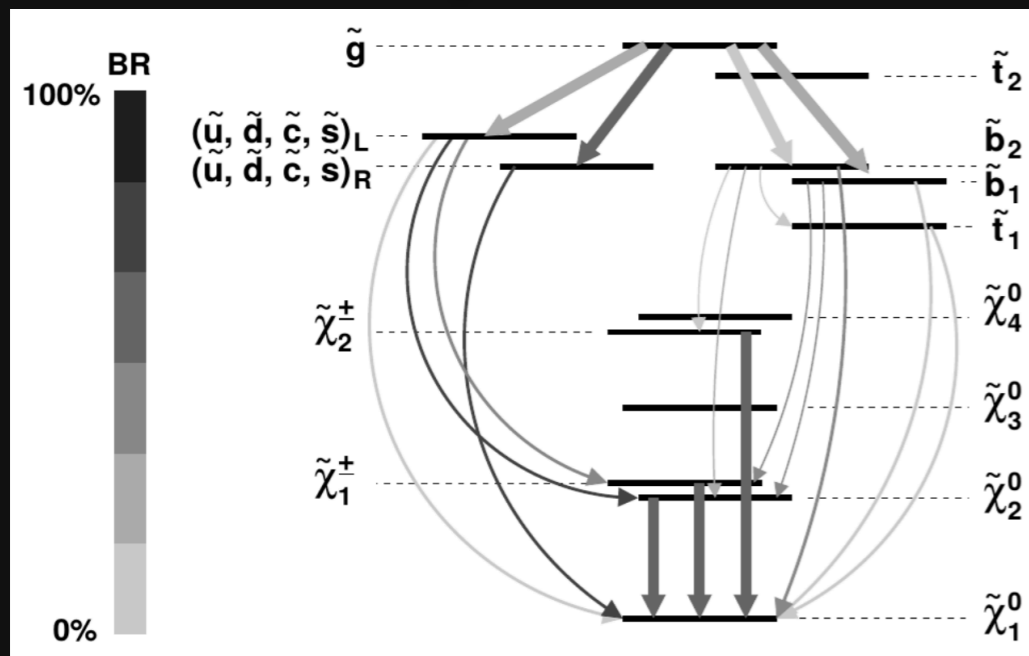
What is supersymmetry?

- **A set of theories** that predicts new boson (fermionic) partners for the fermions (bosons) of the Standard Model — each with spin differing by 1/2 unit
 - Stable lightest supersymmetric particle (LSP) in R-parity conserving SUSY theories are possible dark matter candidates, Higgs boson mass stabilized, and possible unification of inverse gauge couplings
- When undergoing electroweak symmetry breaking, the higgsinos and electroweak gauginos mix (binos, winos, higgsinos)
 - neutral higgsinos and neutral electroweak gauginos mix to form **neutralinos**
 - charged higgsinos and charged electroweak gauginos mix to form **charginos**

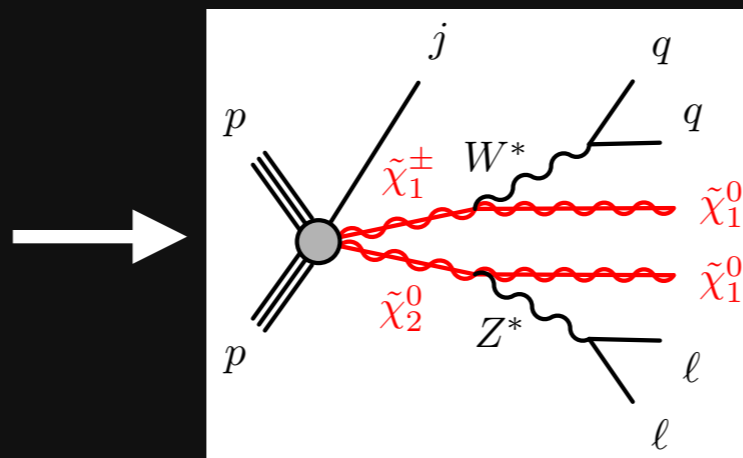
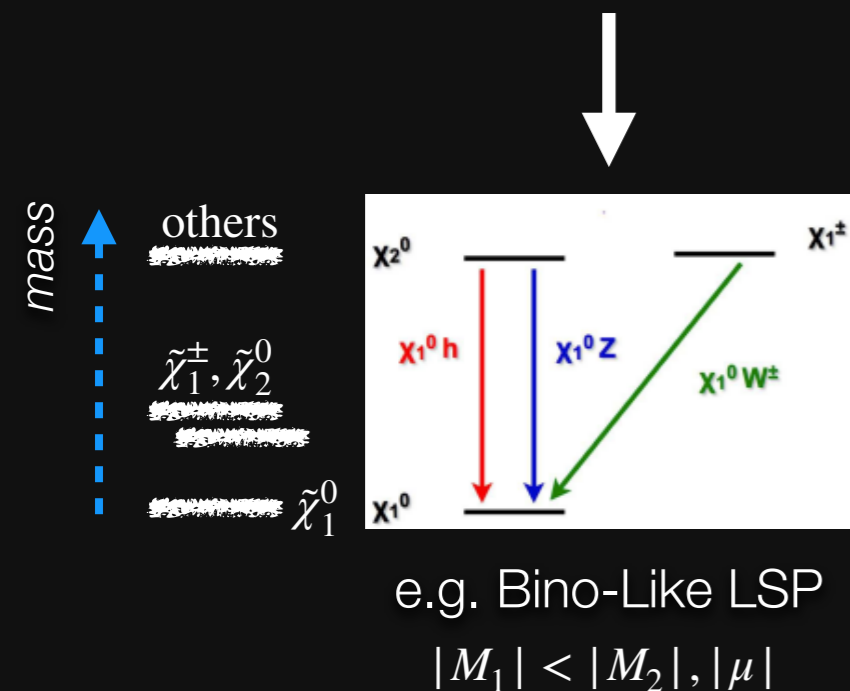


SUSY "Simple" Models

124 free parameters!



- Make some assumptions to reduce phase-space and focus on specific decay chains or specific simplified models (typically 3 or 4 parameters)
- **Pro:** (mostly) orthogonal searches for simplified models are easy to combine and re-interpret with alternative models
- **Con:** not physically realizable in the sense that nature won't give us clean, 100% decays



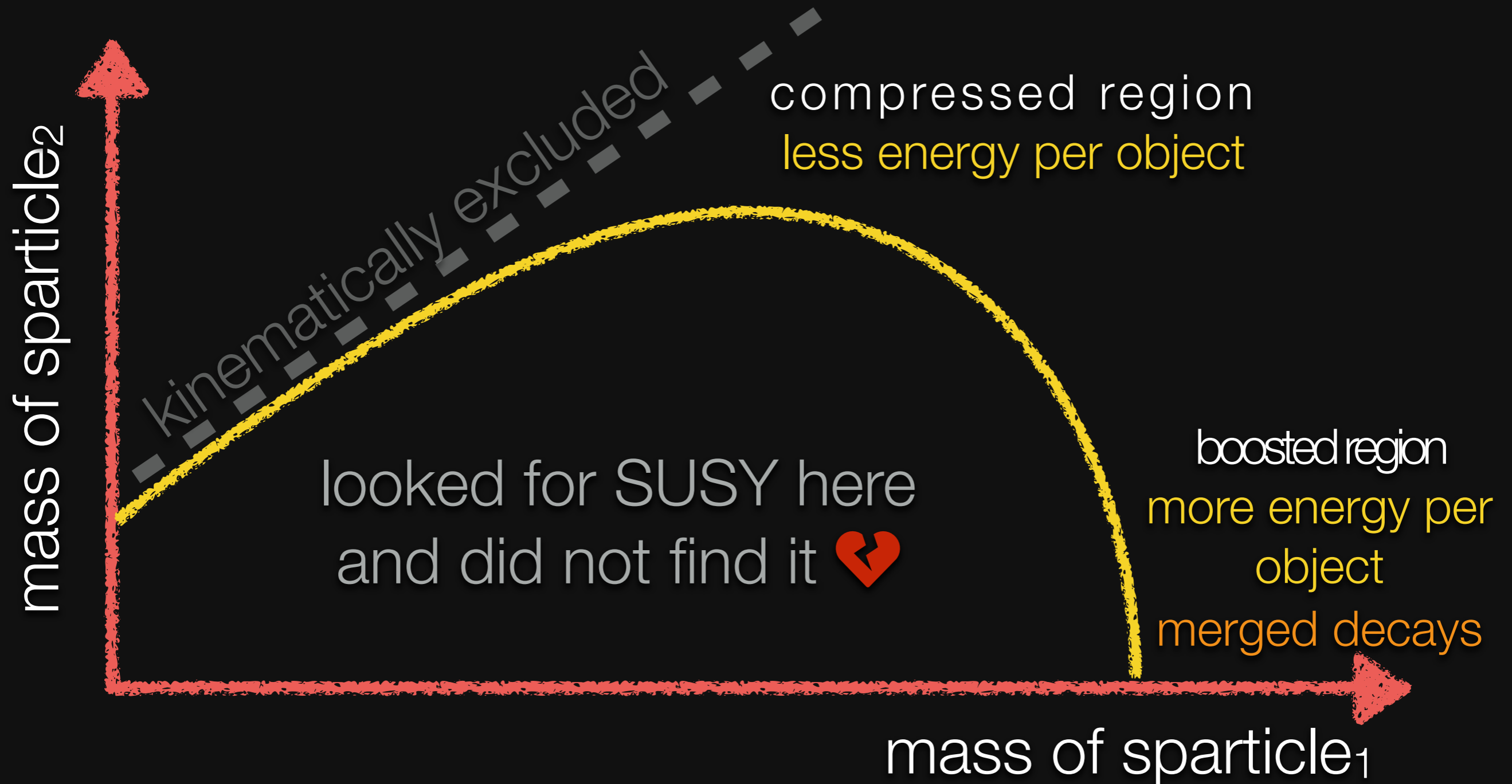
(this is just one example model, there's many others!)
<https://feynman.docs.cern.ch/>

Favorite Discriminating Variables

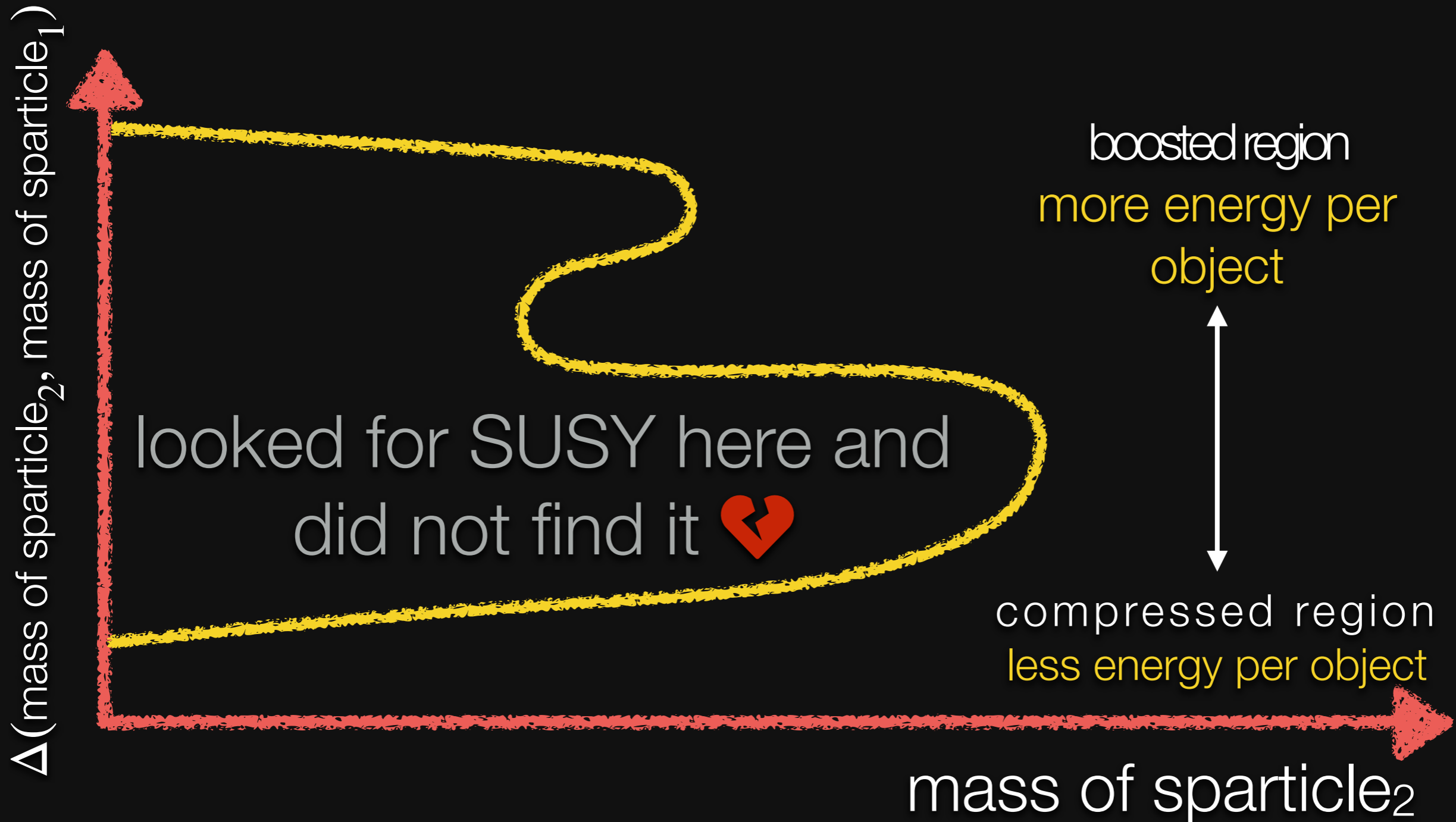
Event selection observables sensitive to features of supersymmetry models are classified as:

- ✦ **Missing momentum-type**: sensitive to the properties of the invisible states
 - ❓ **how many neutralinos in the event? (*RPC: LSP escapes detection*)**
- ✦ **Energy scale-type**: sensitive to the overall energy scale of the event
 - ❓ **what is the mass of the gluino? (*Strong: can reach high mass scales*)**
- ✦ **Energy structure-type**: sensitive to the structure of the visible energy
 - ❓ **how is the energy of the decay partitioned across the final state visible/invisible objects? (*e.g. decay angle between LSP and jets*)**

Interpreting Results



Interpreting Results (“dM”)



Searching for SUSY

2L0J, W/Z 3L, soft 2L

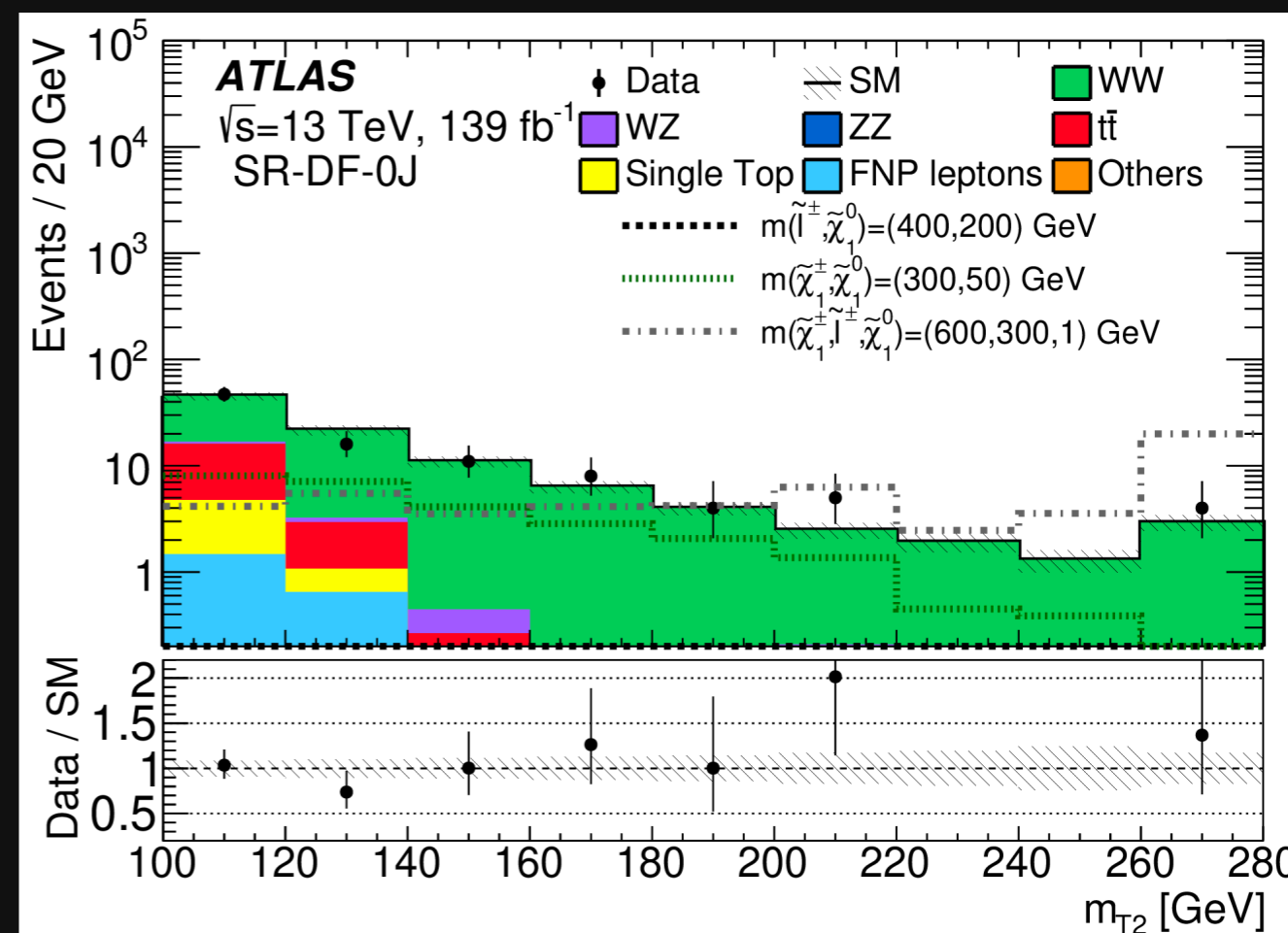
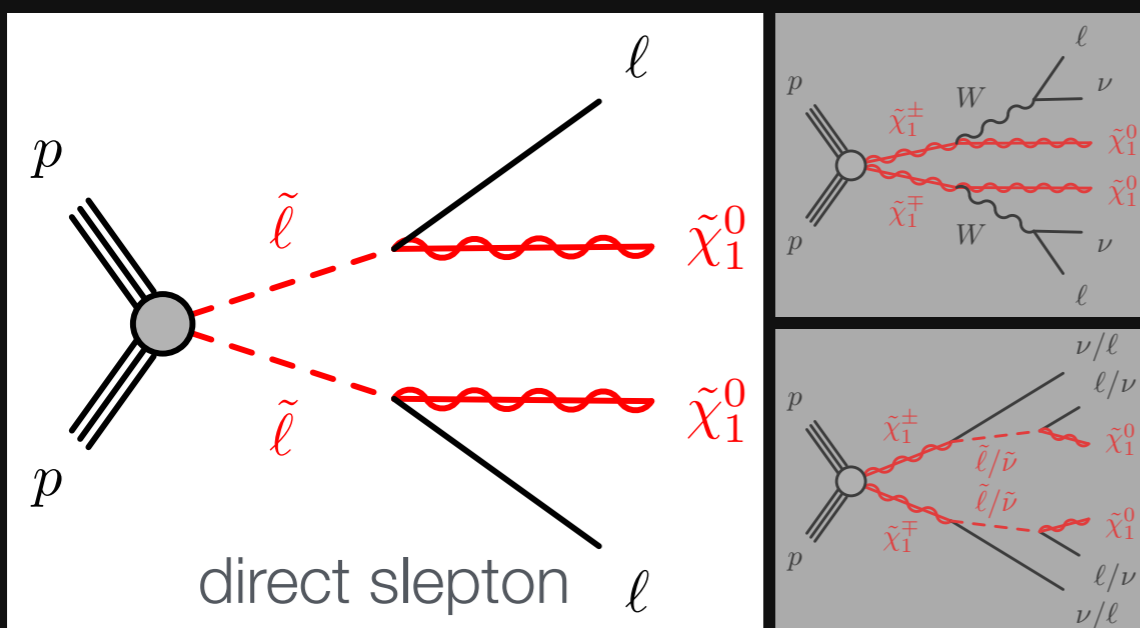
regions with different/same flavor electrons/muons

139 fb⁻¹

cut and count

multi-bin

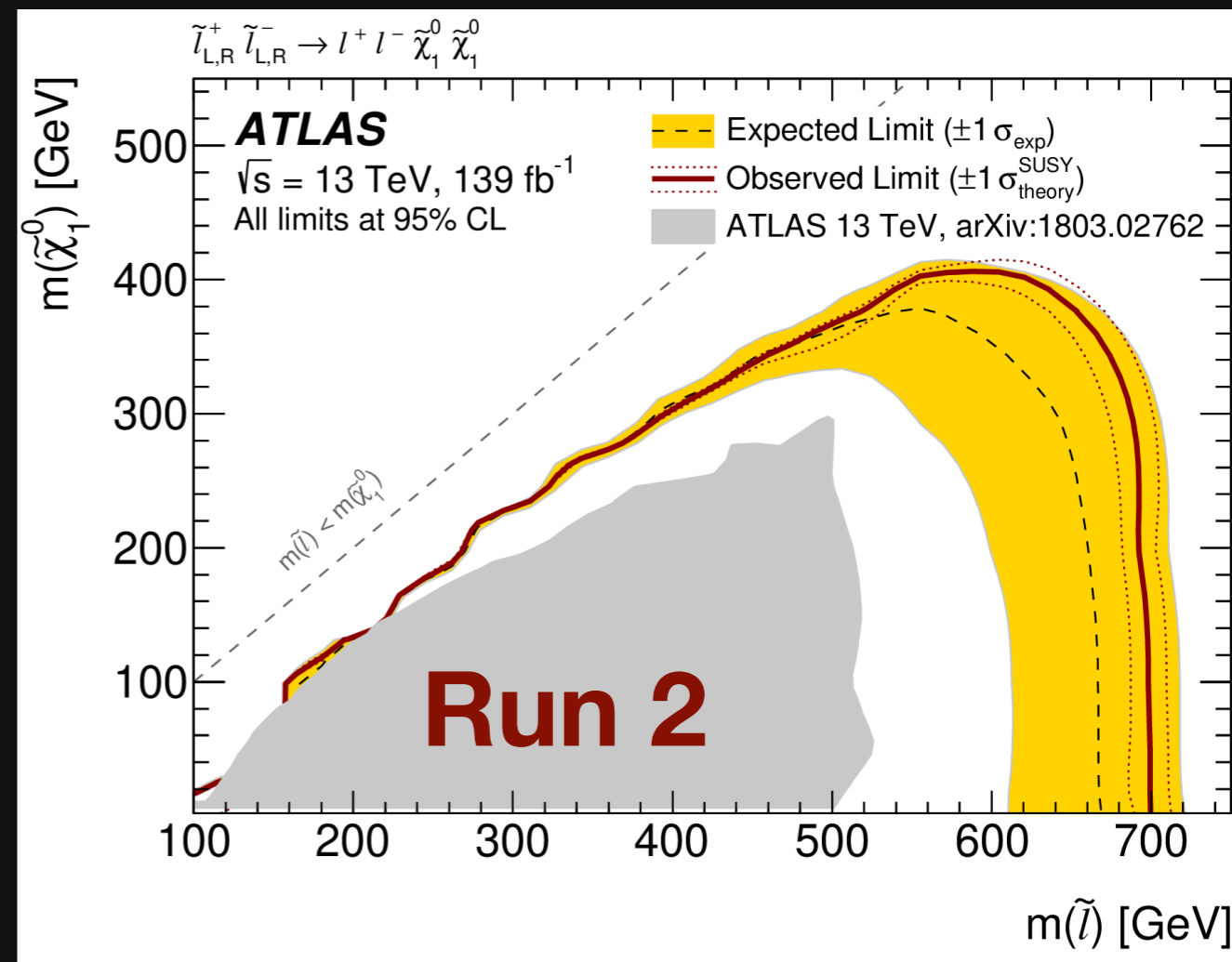
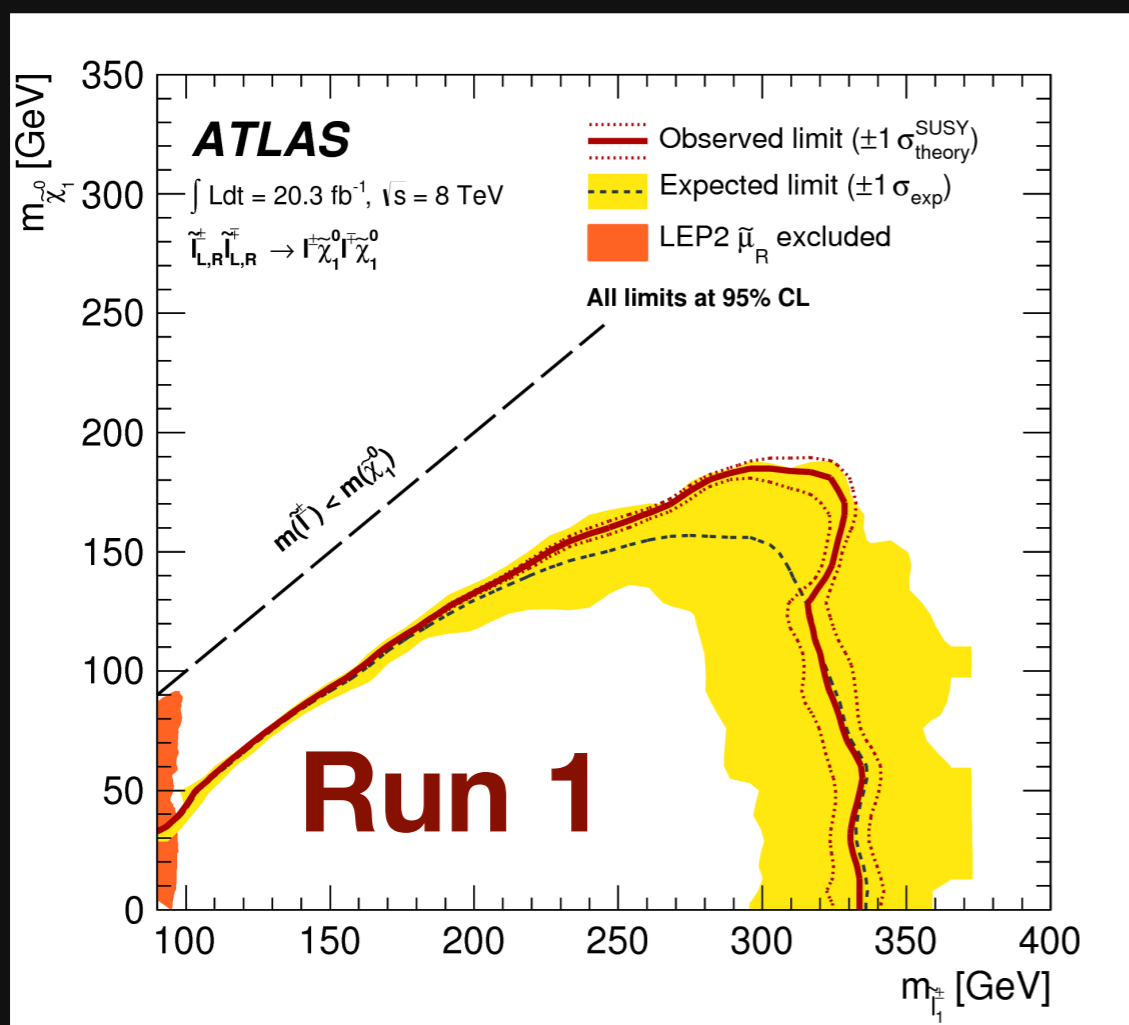
Sleptons: 2 lepton, 0 jet



different-flavor, 0-jet signal region

- MODEL PARAMETERS: $\tilde{\ell}, \tilde{\nu}, \tilde{\chi}_{1^0}, \tilde{\chi}_{1^\pm}$
- FINAL STATE: =2 leptons (OS), MET
- THREE SIGNAL SCENARIOS (9 exclusion regions):
 - $\tilde{\ell} \rightarrow \ell \tilde{\chi}_{1^0}, \ell \in [e, \mu]$ (direct slepton)
 - $\tilde{\chi}_{1^\pm} \rightarrow W \tilde{\chi}_{1^0}$
 - $\tilde{\chi}_{1^\pm} \rightarrow \tilde{\ell} \nu / \tilde{\nu} \ell, \tilde{\ell} / \tilde{\nu} \rightarrow \tilde{\chi}_{1^0} \ell / \nu$
- DOMINANT BACKGROUNDS: $t\bar{t}, Wt, WW, WZ, ZZ$
- CHALLENGE: theory uncertainties of diboson, top

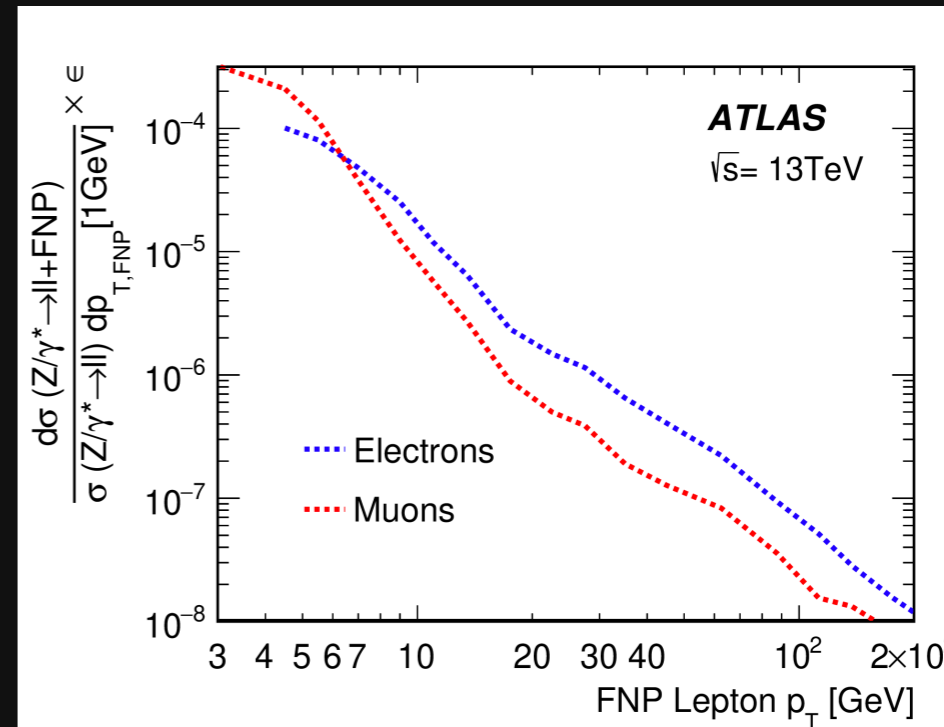
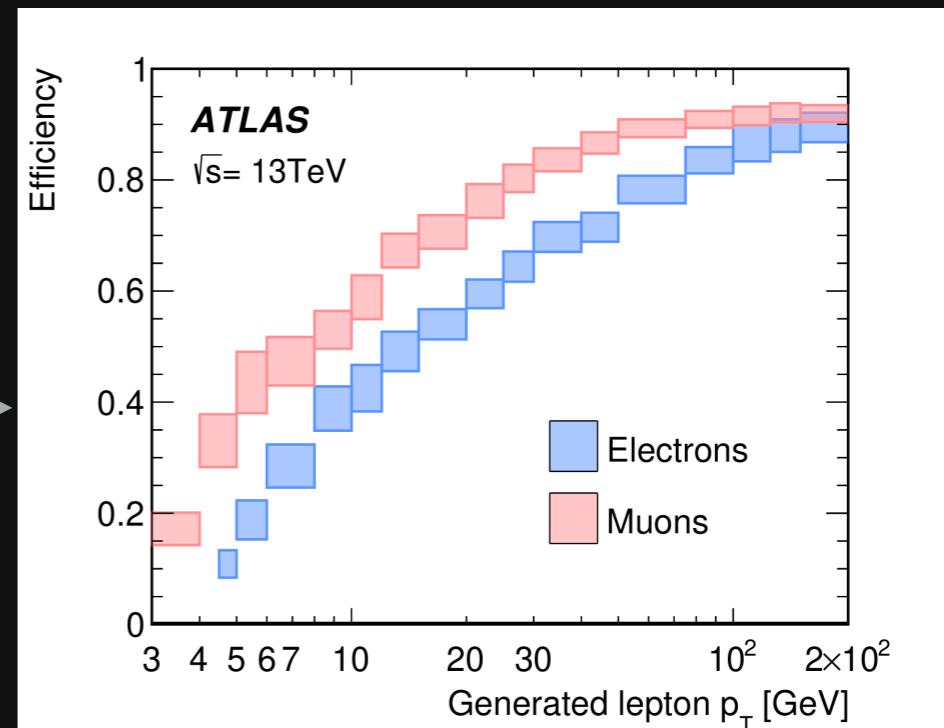
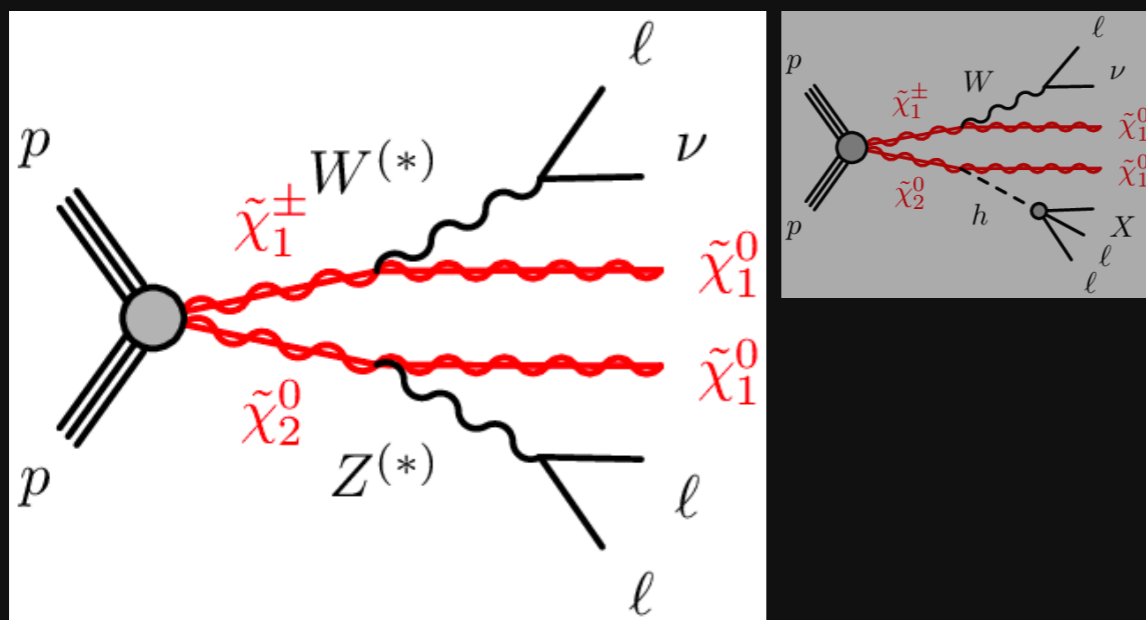
Sleptons: 2 lepton, 0 jet



Optimized for direct sleptons
(used in g-2 global fit)

Sensitivity increased from 350 to 700 GeV

W/Z 3-leptonic



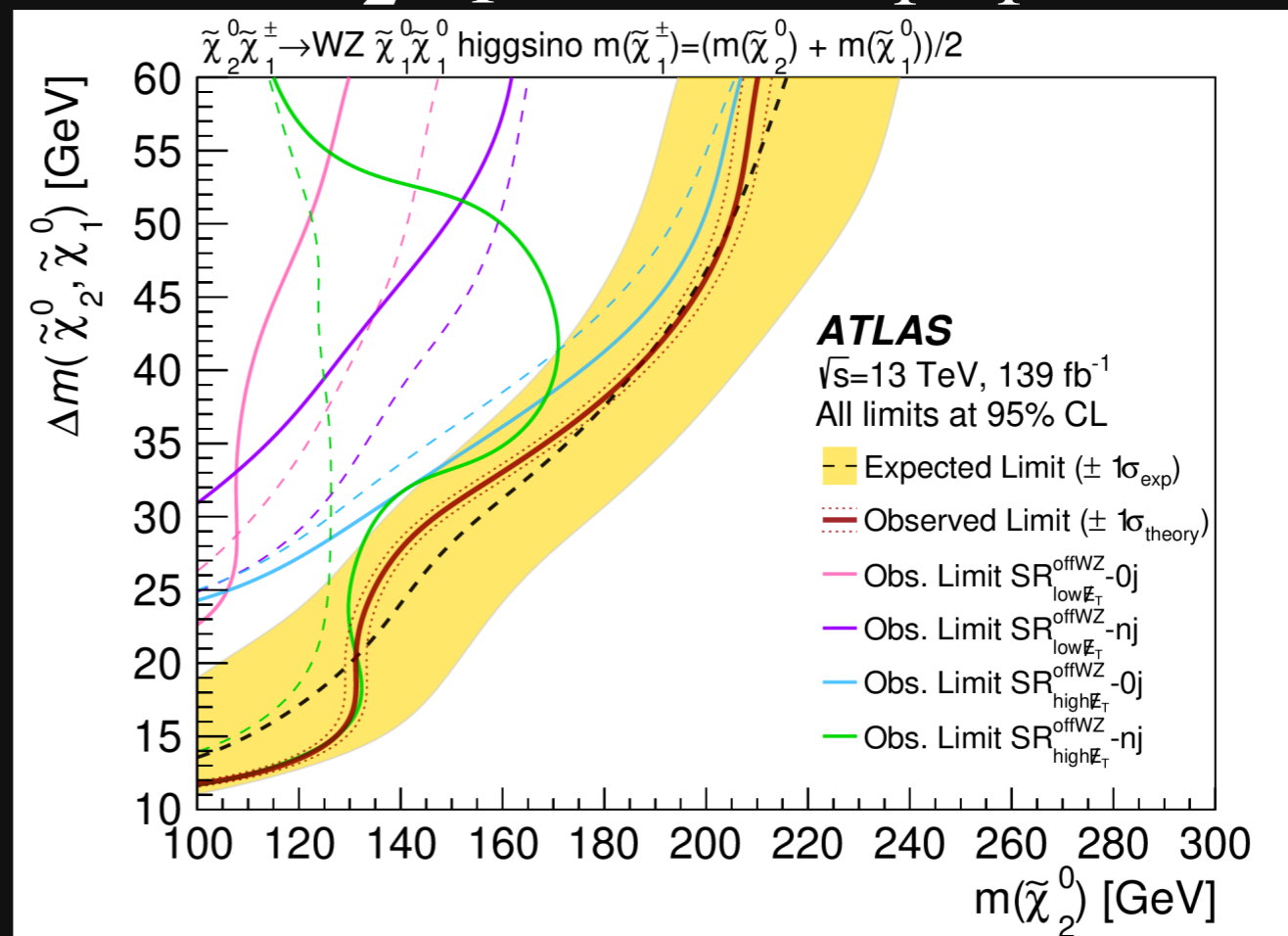
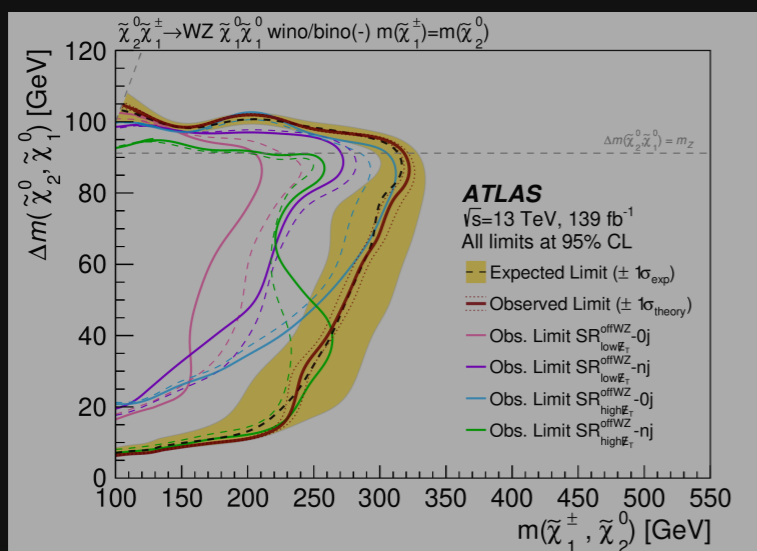
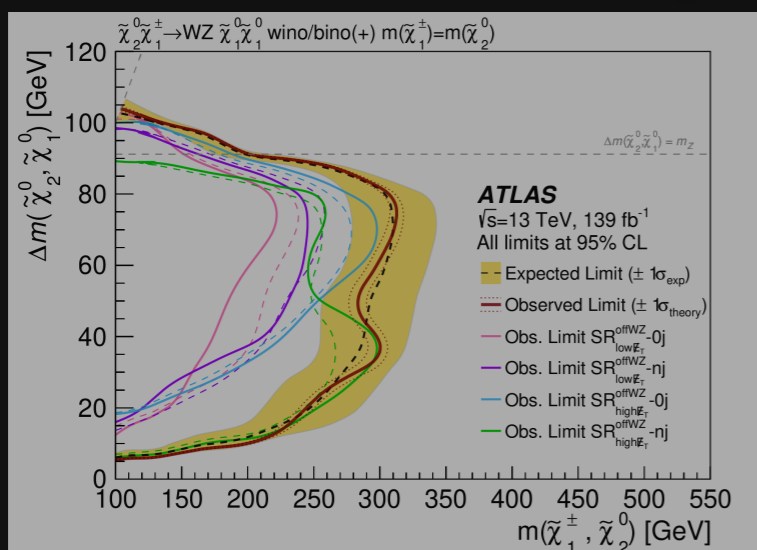
signal efficiency

Z+jets(+FNP) efficiency

- MODEL PARAMETERS: $\tilde{\chi}_1^\pm, \tilde{\chi}_2^0, \tilde{\chi}_1^0$
- FINAL STATE: 3 leptons (OS-SF) + MET
- THREE SIGNAL SCENARIOS (9 exclusion regions):
 - $|M_1| < |M_2| \ll |\mu|, m(\tilde{\chi}_1^\pm) = m(\tilde{\chi}_2^0)$
 - wino/bino(+) with $|M_1| \times |M_2| > 0$
 - $|M_1| < |M_2| \ll |\mu|, m(\tilde{\chi}_1^\pm) = m(\tilde{\chi}_1^0)$
 - wino/bino(-) with $|M_1| \times |M_2| < 0$
 - $(\tilde{\chi}_2^0, \tilde{\chi}_1^\pm, \tilde{\chi}_1^0), m(\tilde{\chi}_1^\pm) = \frac{1}{2} [m(\tilde{\chi}_2^0) + m(\tilde{\chi}_1^0)]$
 - all of the above split up into on-shell W/Z or off-shell W*/Z*
- DOMINANT BACKGROUNDS: SM W/Z, $t\bar{t}$, Z + jets with FNP lepton
- CHALLENGE: suppressing FNP leptons (dedicated BDT)

W/Z 3-leptonic

$$\tilde{\chi}_2^0 \tilde{\chi}_1^\pm \rightarrow WZ \tilde{\chi}_1^0 \tilde{\chi}_1^0$$



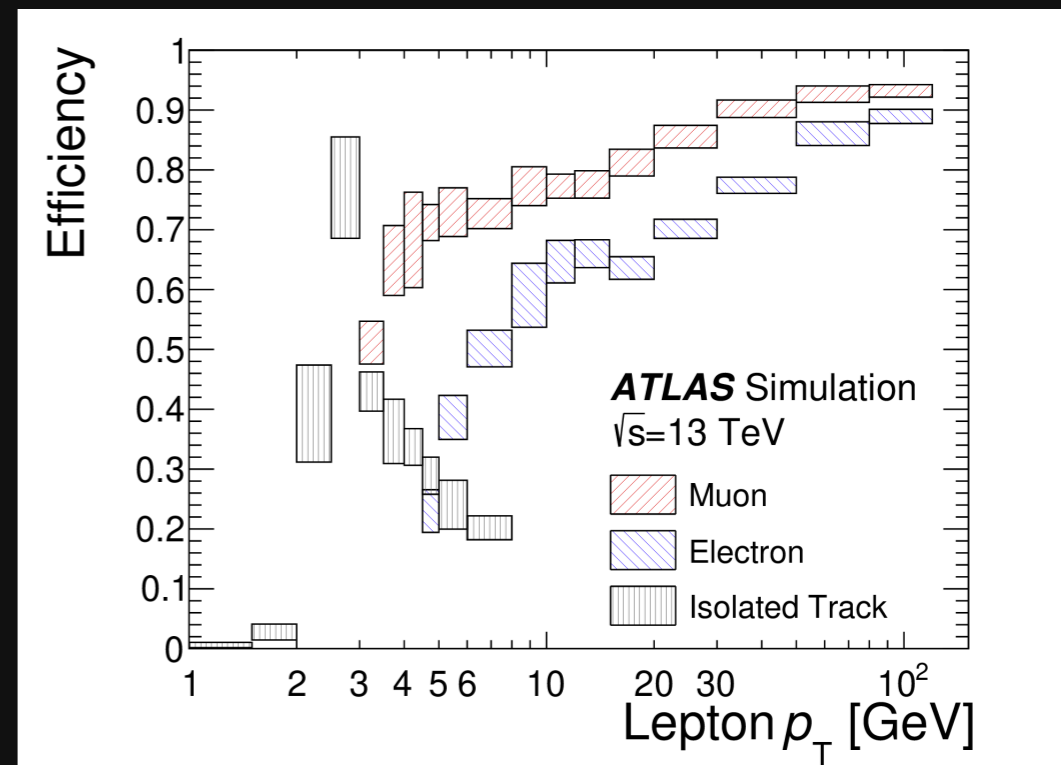
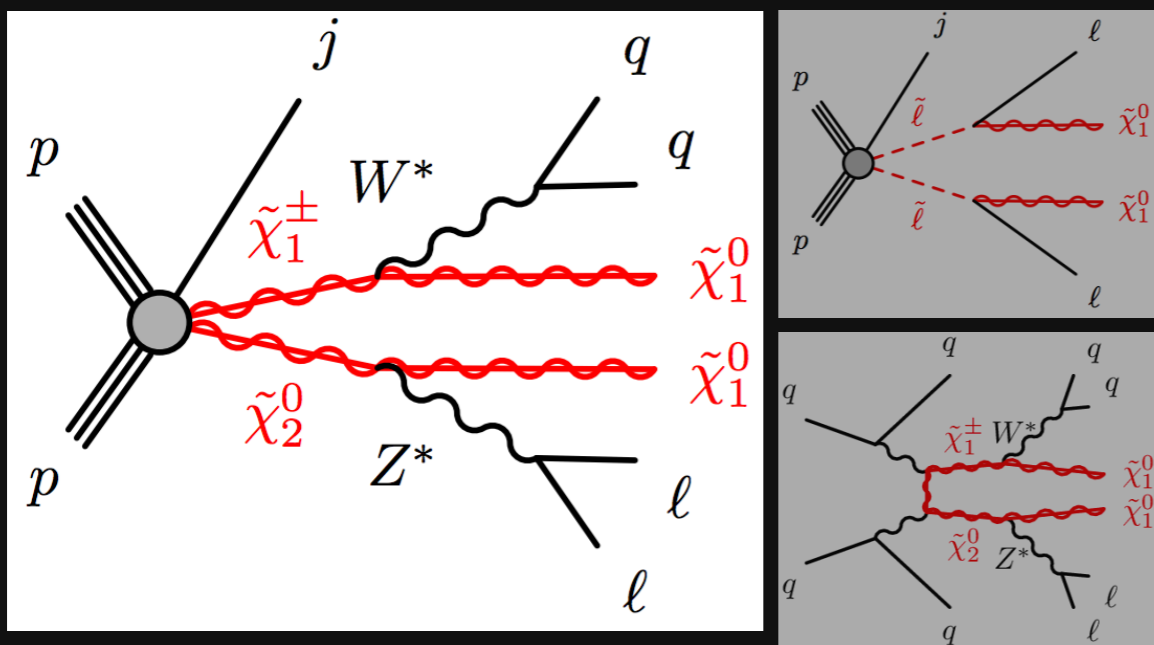
Wino/Bino

Higgsino

Sensitivity for $\Delta m > 12$ GeV

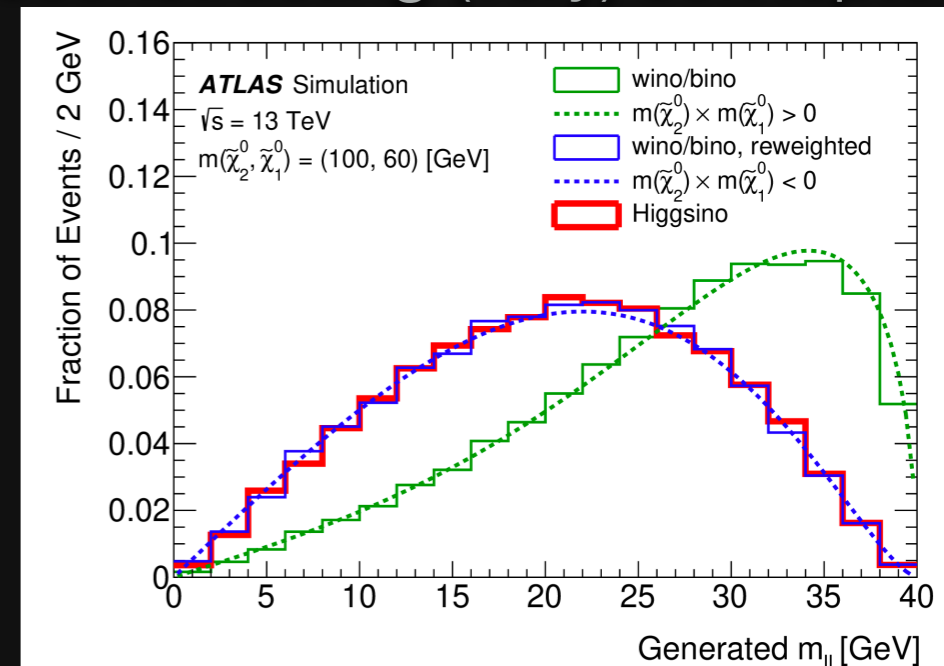
(see next few slides for statistical combination)

Soft 2-leptonic

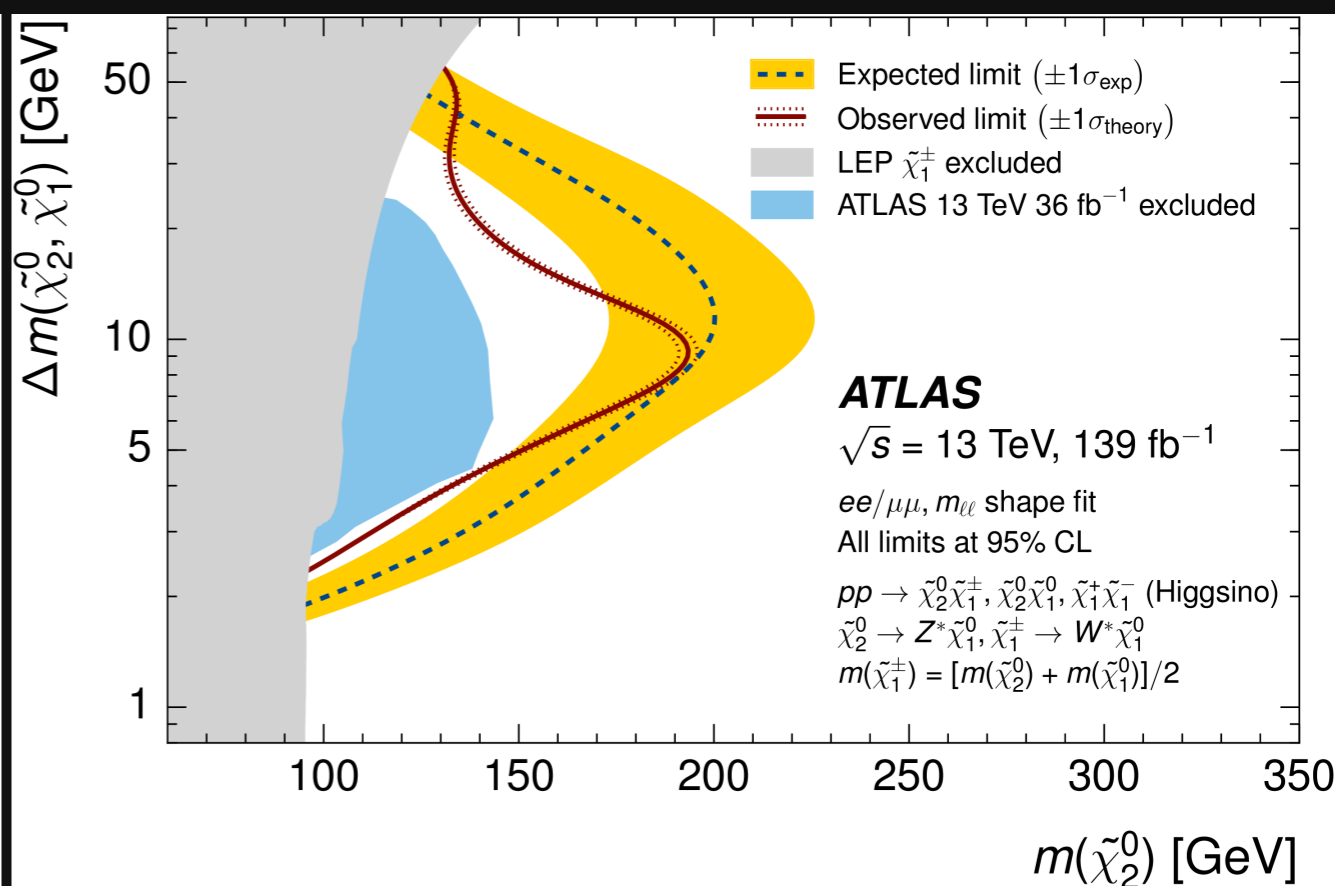
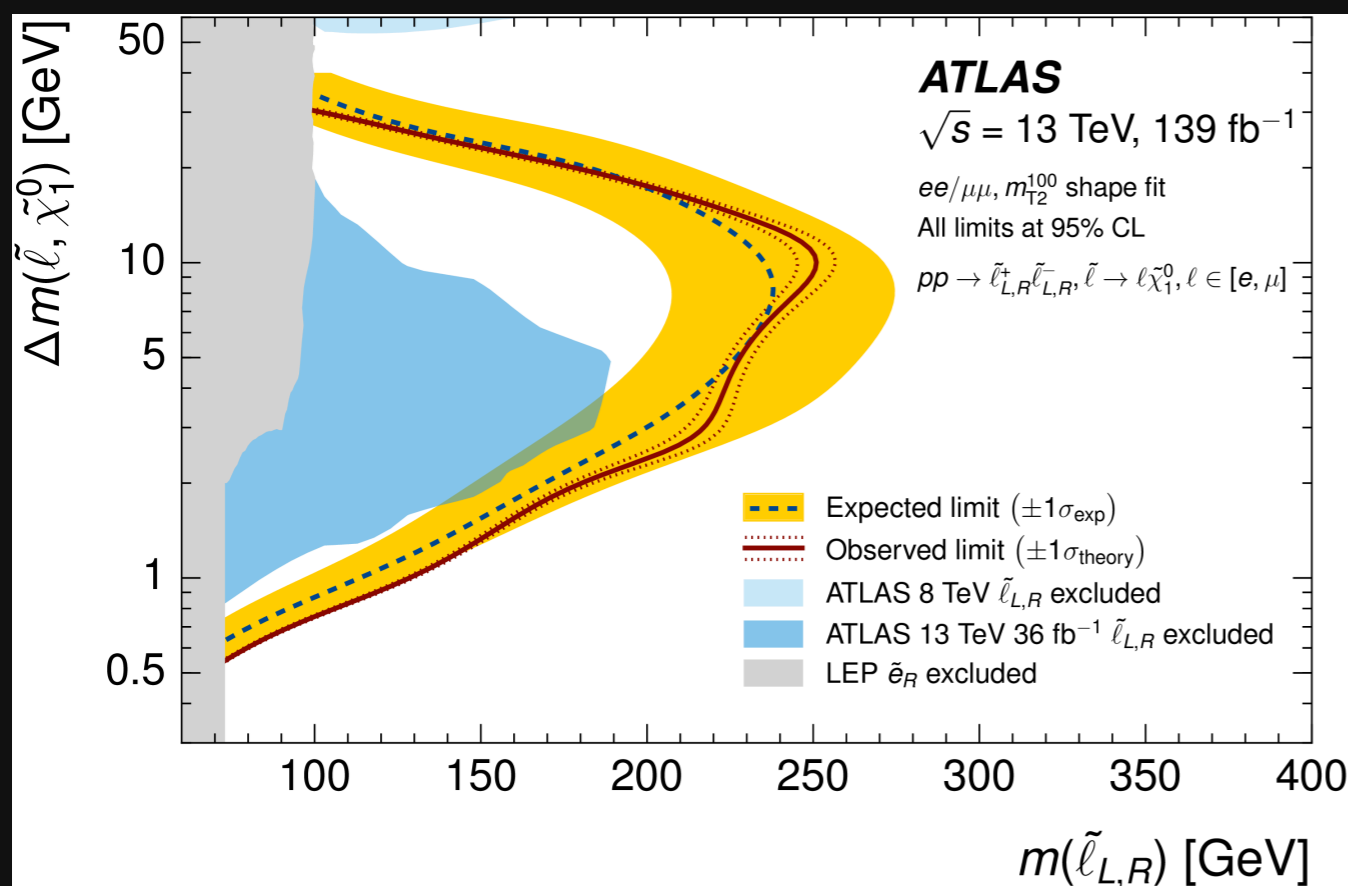


reconstructing (very) soft leptons

- MODEL PARAMETERS: $\tilde{\chi}_2^0, \tilde{\chi}_1^0, \tilde{\chi}_1^\pm, \tilde{\ell}$
- FINAL STATE: =2 leptons [or 1 lepton+1 track] (OS-SF), MET, ISR jet
- FOUR SIGNAL SCENARIOS (6 multi-bin regions):
 - $\tilde{\chi}_2^0 \rightarrow Z^* \tilde{\chi}_1^0, \tilde{\chi}_1^\pm \rightarrow W^* \tilde{\chi}_1^0$ (with ISR or VBF)
 - here similar to W/Z 3-leptonic [three scenarios]
 - $\tilde{\ell} \rightarrow \ell \tilde{\chi}_1^0, \ell \in [e, \mu]$ (sleptons, bino-like LSP)
- DOMINANT BACKGROUNDS: $t\bar{t}$, singletop, $Z(\rightarrow \tau\tau) + \text{jets}$, diboson, multi-jet fake/nonprompt leptons
- CHALLENGE: compressed-mass search, fake/non-prompt lepton background



Soft 2-leptonic



Mass-degenerate Sleptons
 Useful in a $g-2$ global fit

Higgsinos
 Useful in a statistical combination

Sensitive to 200 GeV Higgsinos at mass splitting of 10 GeV

Statistical Reinterpretations

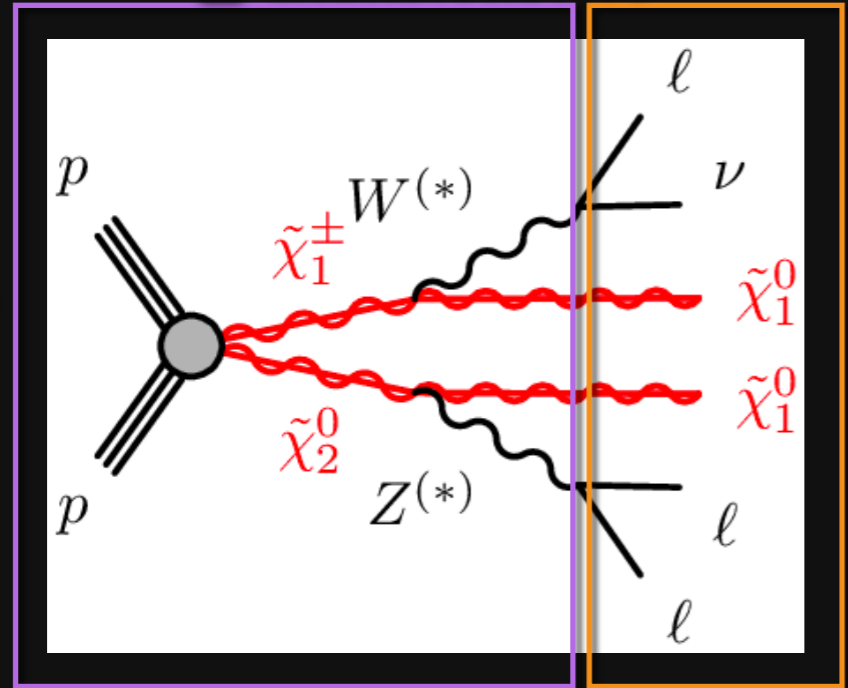
Combinations and global fits

E.G.: Stat. Combination (I)

- Multiple analyses with different **signatures** can still target the same **model**
 - See Zach Marshall's talk on differences between signature and model: <https://indico.cern.ch/event/1023573/contributions/4400586/>
- Goal: combine multiple searches to paint a tapestry of our sensitivity to Higgsino/Wino-Bino models (production) which decay to on-shell/off-shell Standard Model bosons (W/Z)

[ATLAS-CONF-2020-015](#)
[arXiv:2106.01676](#)

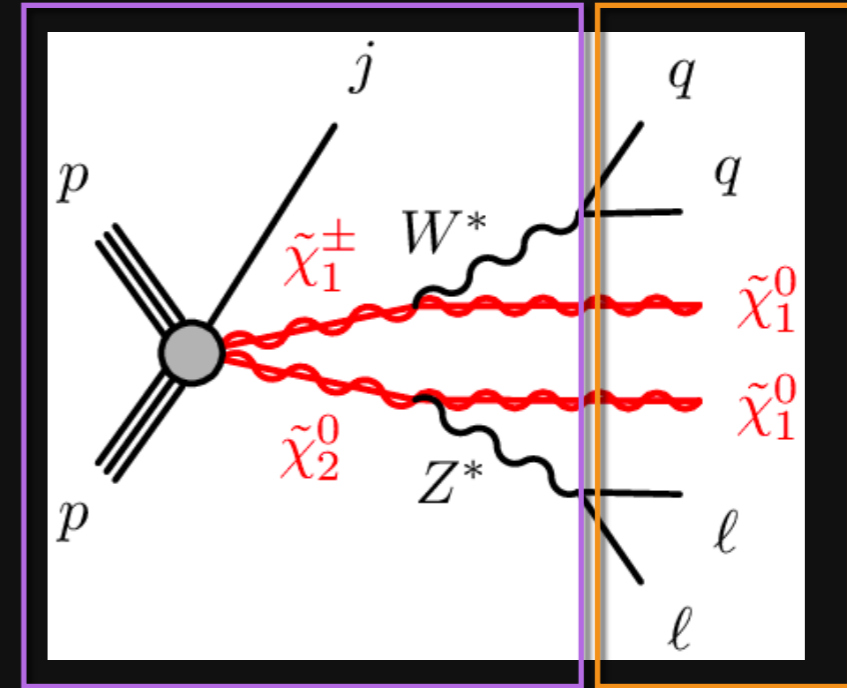
W/Z 3-leptonic



off-shell: $\Delta m(\tilde{\chi}_2^0, \tilde{\chi}_1^0) \in [5, 90]$ GeV
 on-shell: $\Delta m(\tilde{\chi}_2^0, \tilde{\chi}_1^0) \geq 90$ GeV
 signature: $3\ell + 0j + E_T^{\text{miss}}$

[arXiv:1911.12606](#)

soft 2L

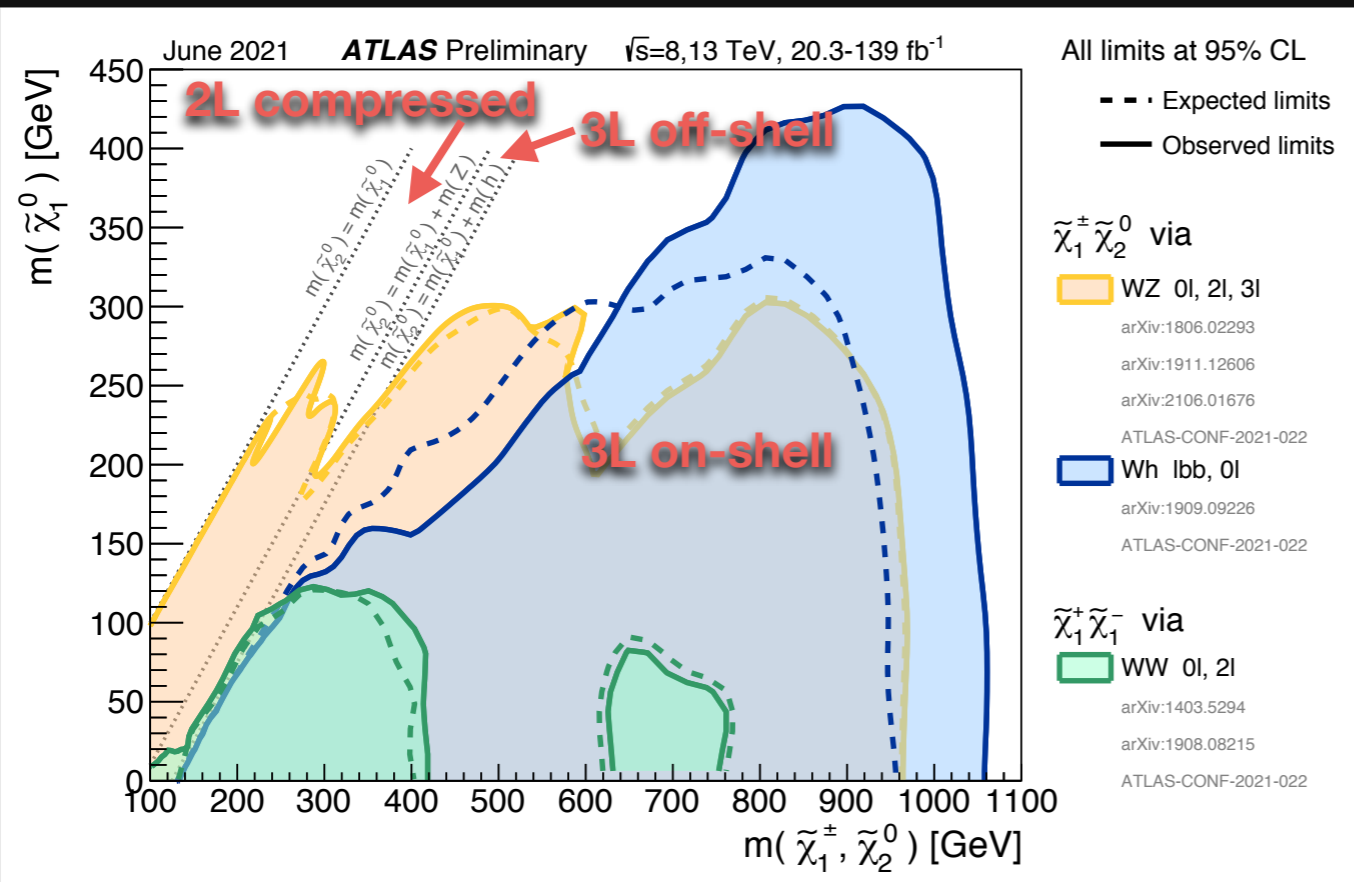


soft 2L: $\Delta m(\tilde{\chi}_2^0, \tilde{\chi}_1^0) \in [1, 60]$ GeV
 signature: $2\ell + 3j + E_T^{\text{miss}}$

E.G.: Stat. Combination (II)

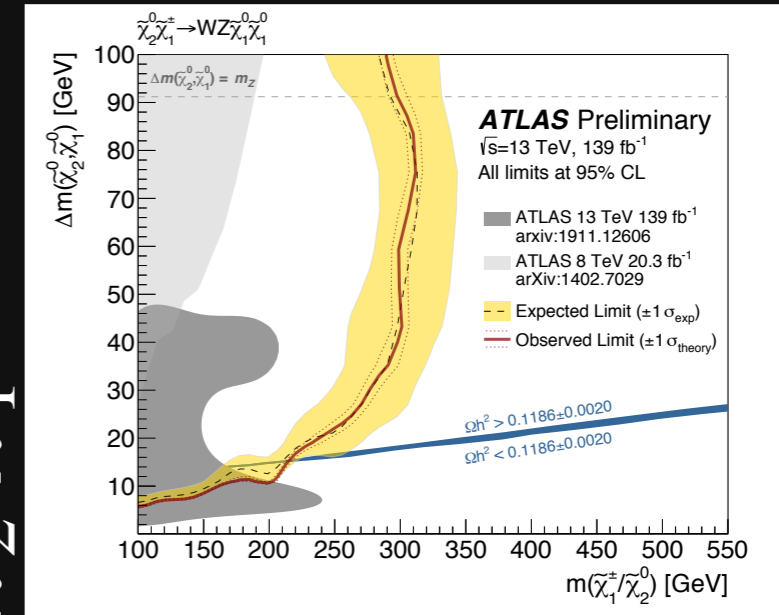
- Multiple analyses with different **signatures** can still target the same **model**
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- Goal: combine multiple searches to paint a tapestry of our sensitivity to Higgsino/Wino-Bino models (production) which decay to on-shell/off-shell Standard Model bosons (W/Z)

$m(\tilde{\chi}_1^0)$

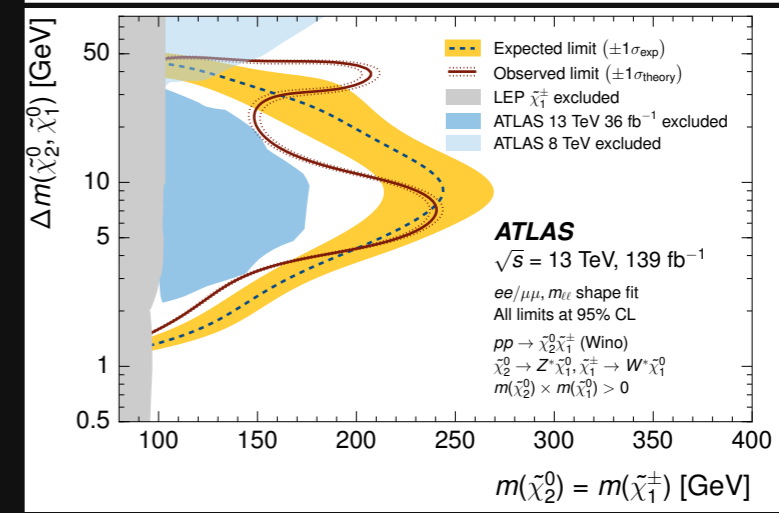


$m(\tilde{\chi}_1^\pm, \tilde{\chi}_2^0)$

$\Delta m(\tilde{\chi}_2^0, \tilde{\chi}_1^0)$



3L

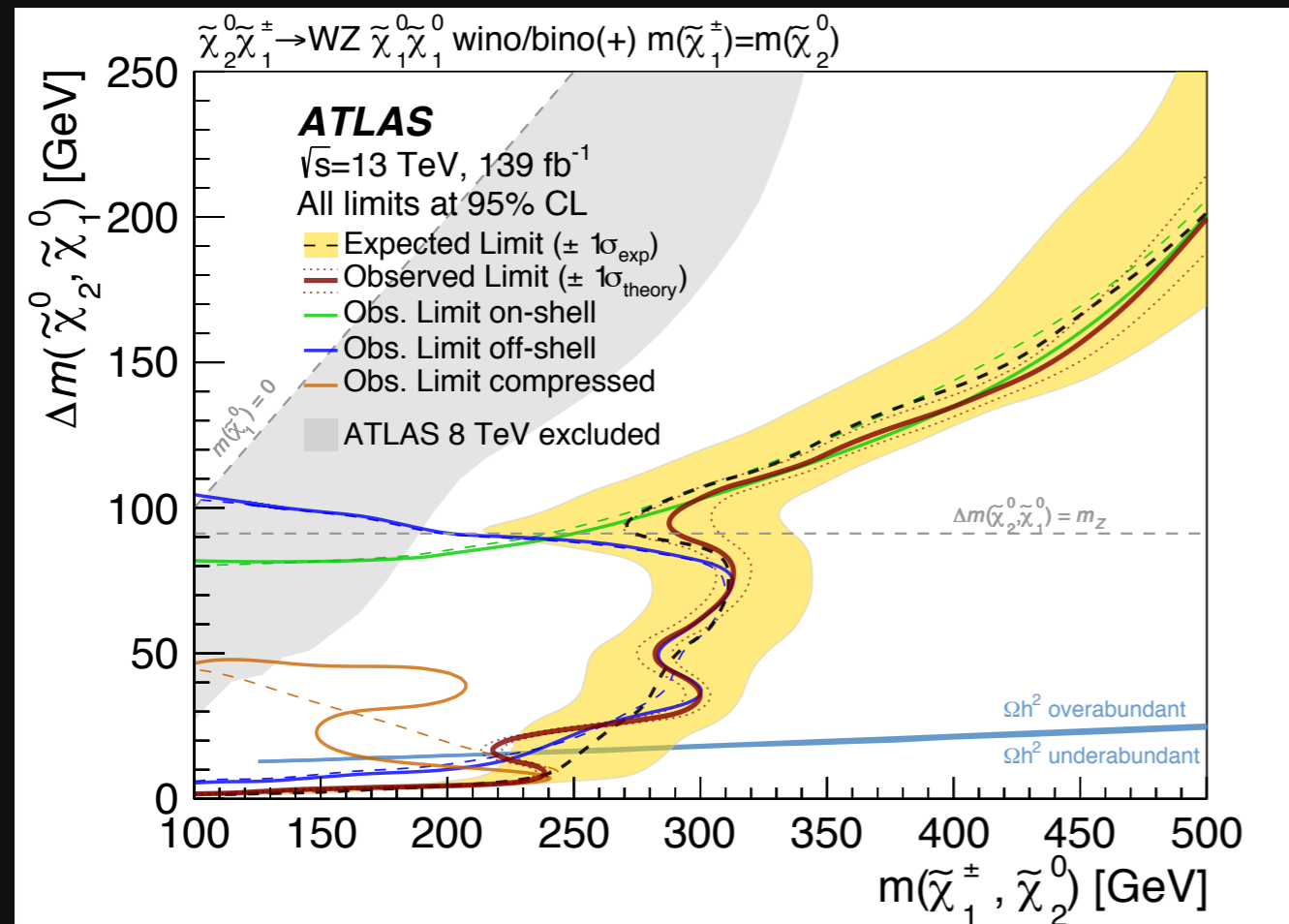


2L

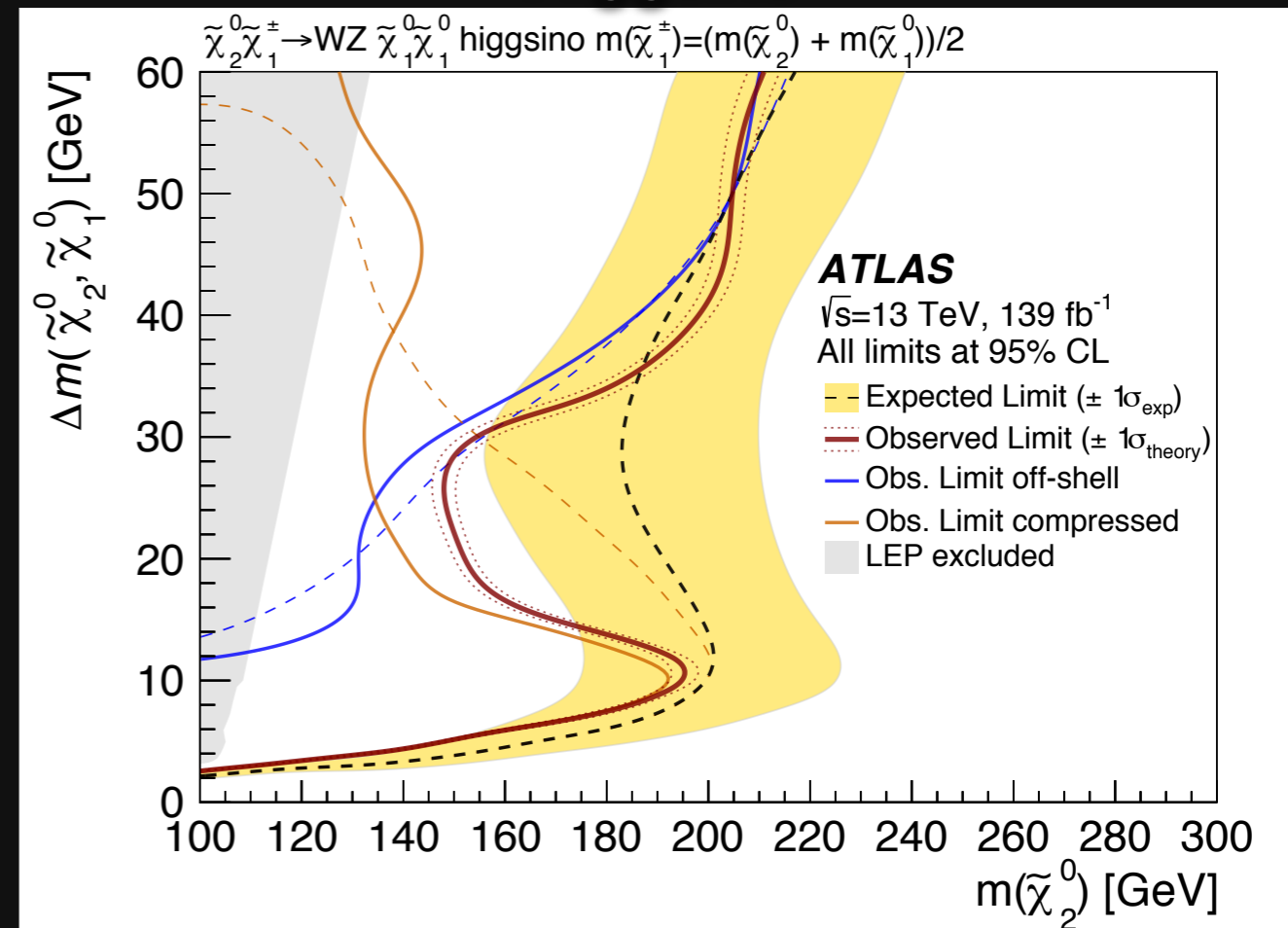
E.G.: Stat. Combination (III)

- Experimental uncertainties were correlated if both analyses used the same object definitions
- Modeling uncertainties were correlated if both analyses used the same background and estimation procedure
- Scale factors (e.g. data-driven normalization) are not correlated as they probe different phase spaces

Wino-Bino



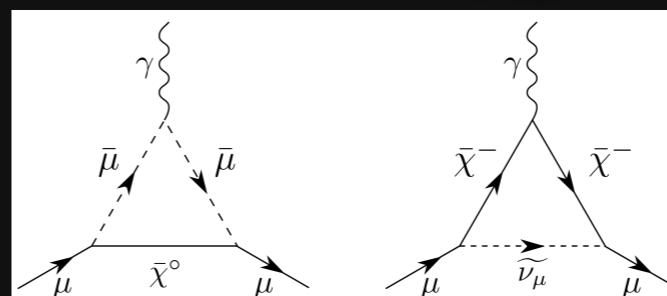
Higgsino



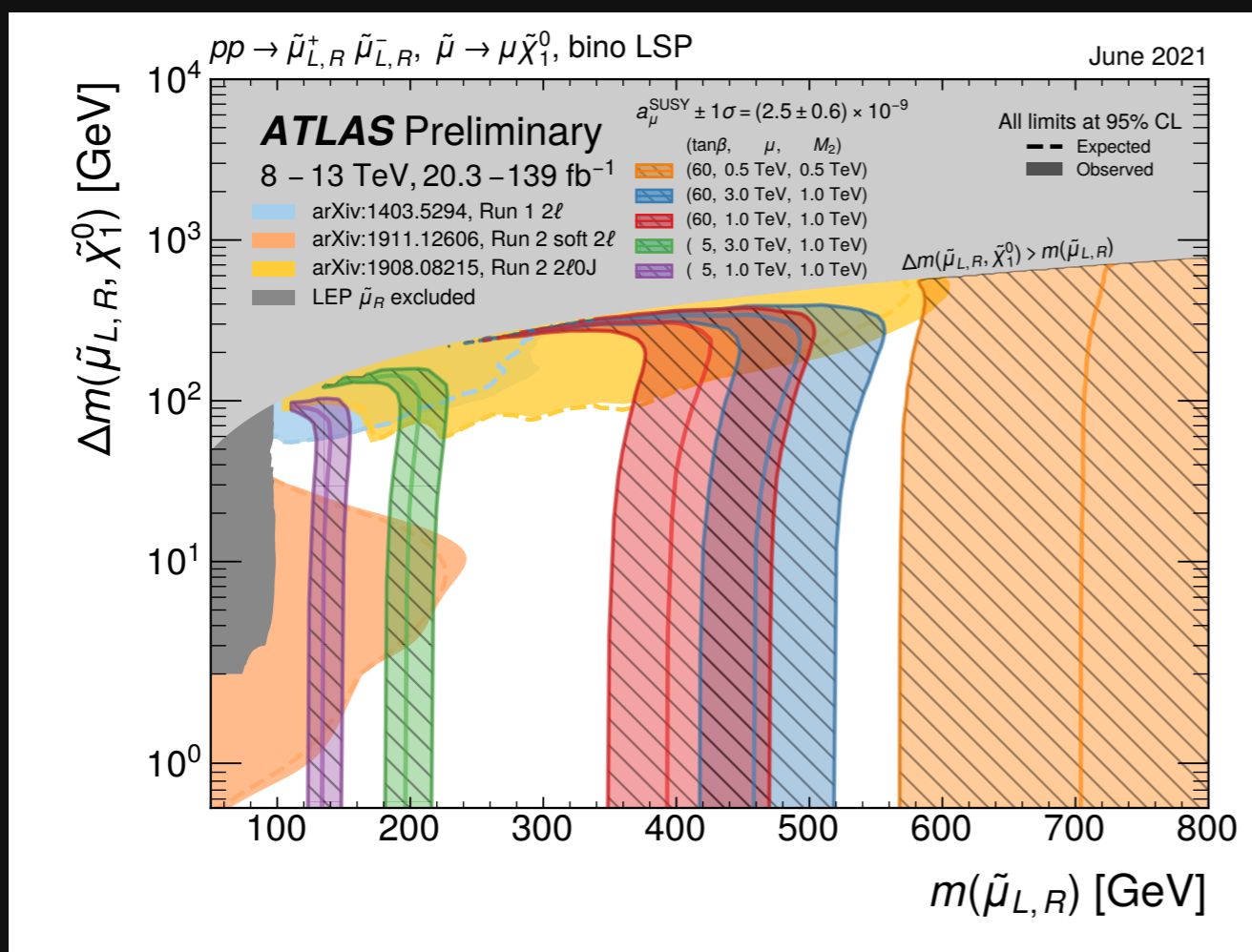
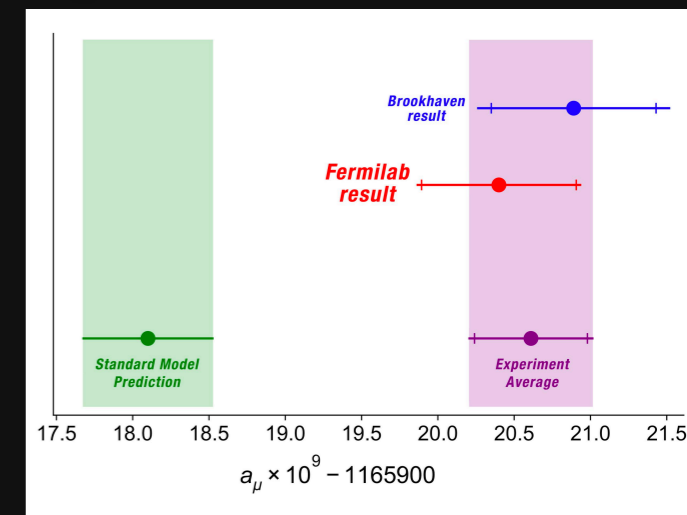
Example: global fit (g-2)

- Can measurements performed by other experiments be used to constrain or inform our search programme?

Yes!



$$\tilde{\mu}^\pm / \tilde{\chi}^0 \text{ or } \tilde{\chi}^\pm / \tilde{\nu}_\mu$$

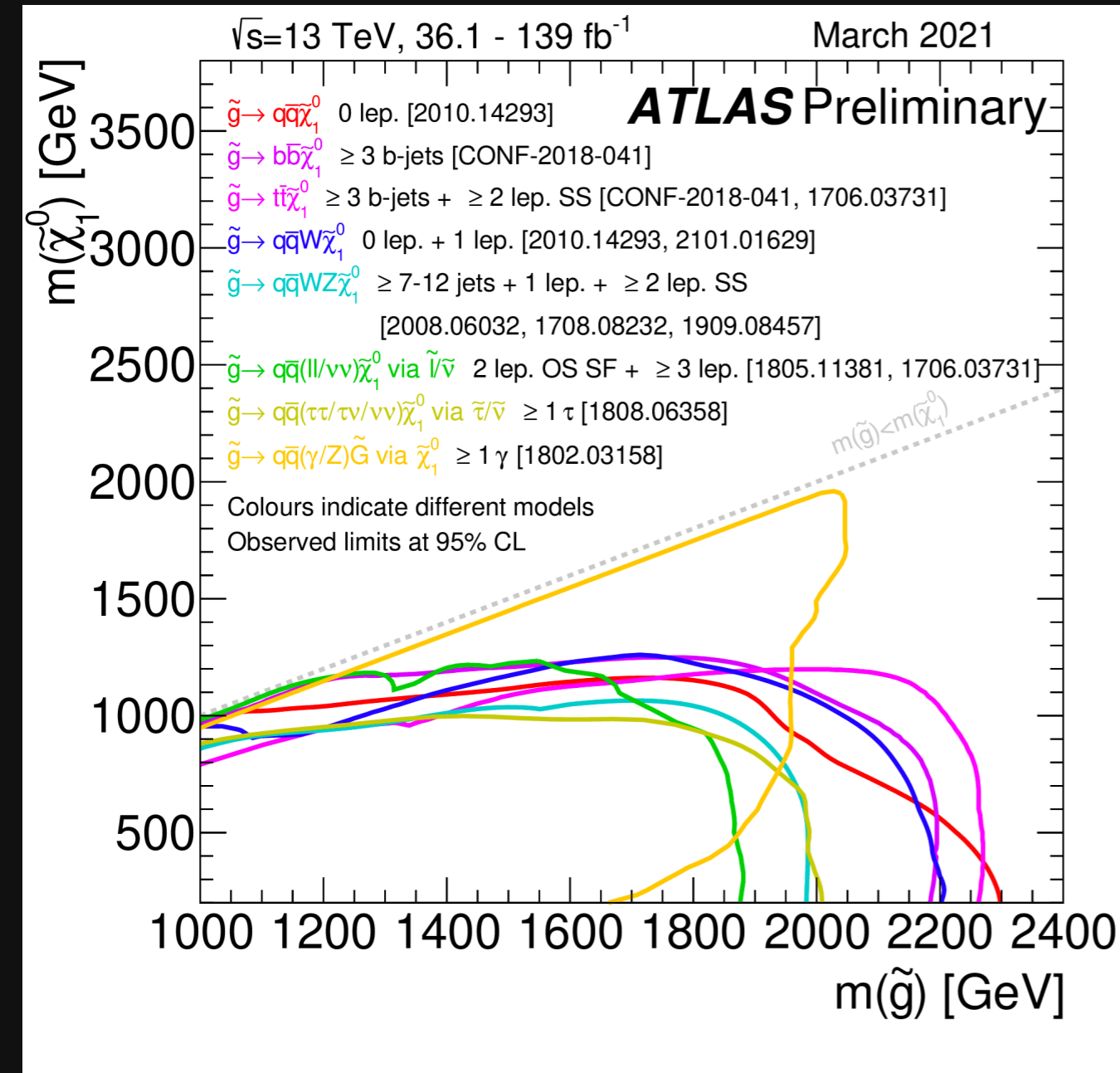


April 27th, 2021

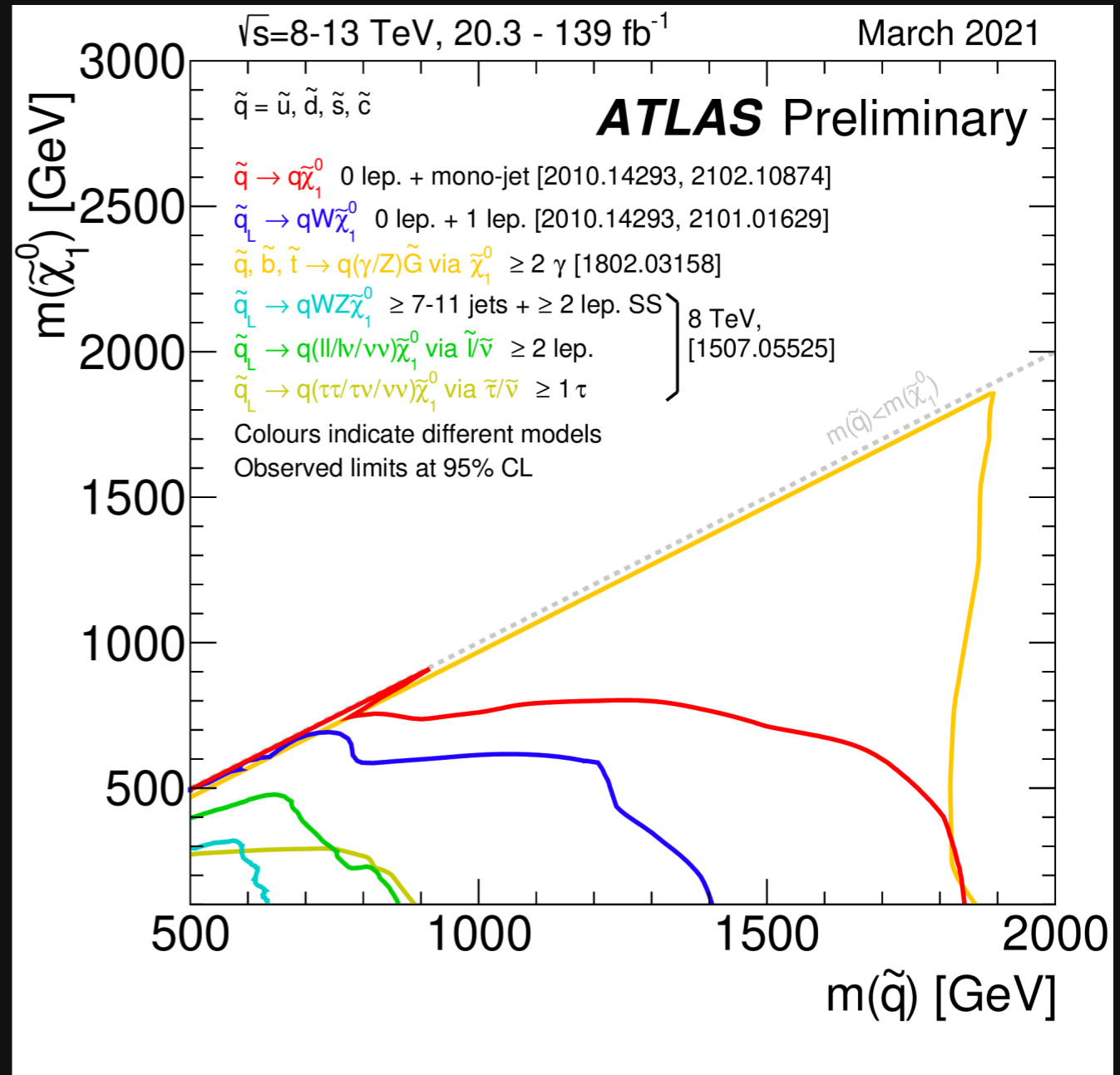
- Compute exclusion limits in the $(\tilde{\mu}, \tilde{\chi}_1^0)$ mass plane for different analyses that are sensitive to the smuon model with muon/bino-like $\tilde{\chi}_1^0$ signature
- Hatched bands are compatible with the observed g-2 anomaly measured by Fermilab and BNL experiments to $\pm 1\sigma$, for a variety of pMSSM parameters

Light smuon/light neutralino needed!

Gluino/Squark Summary

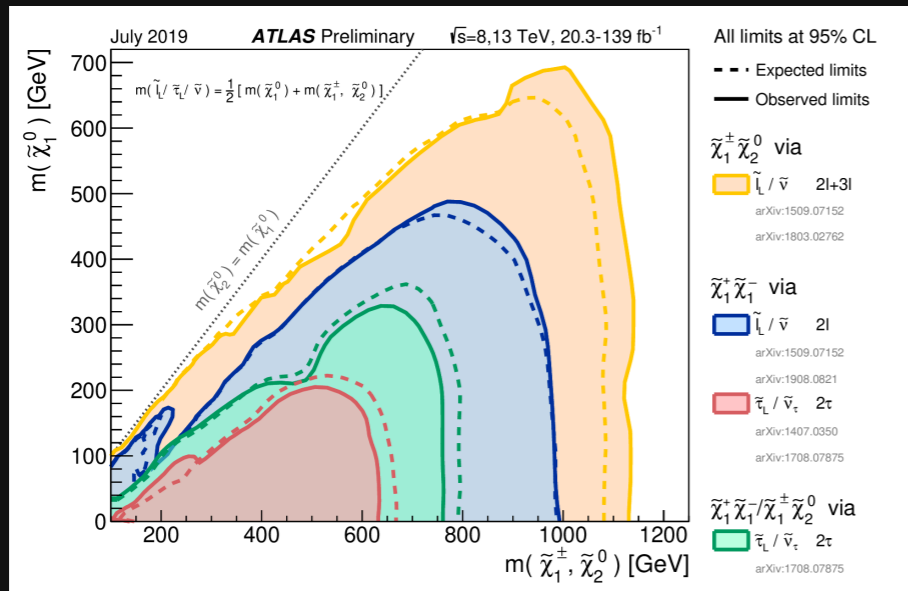


$\tilde{g} \rightarrow \text{LSP or } \tilde{G}$

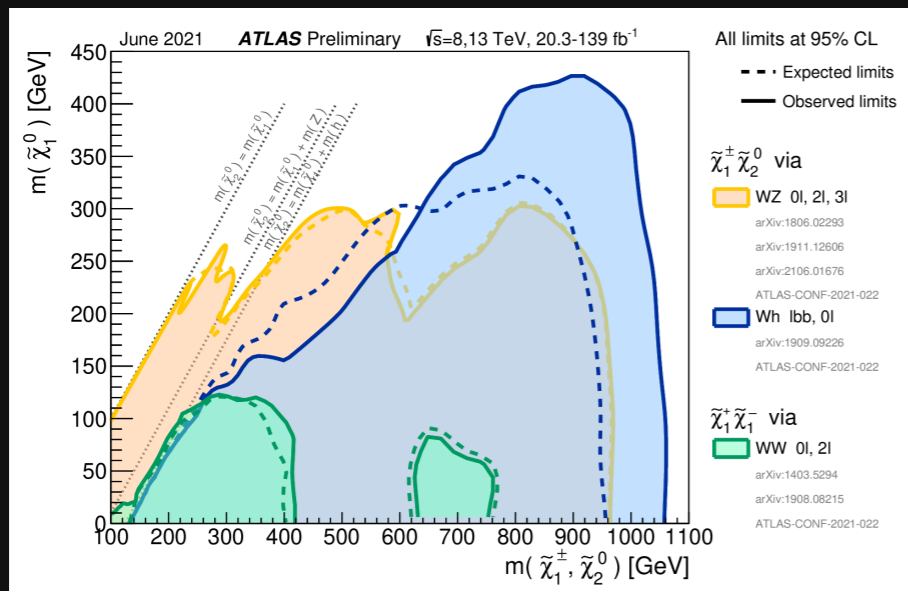


$\tilde{q} \rightarrow \text{LSP or } \tilde{G}$

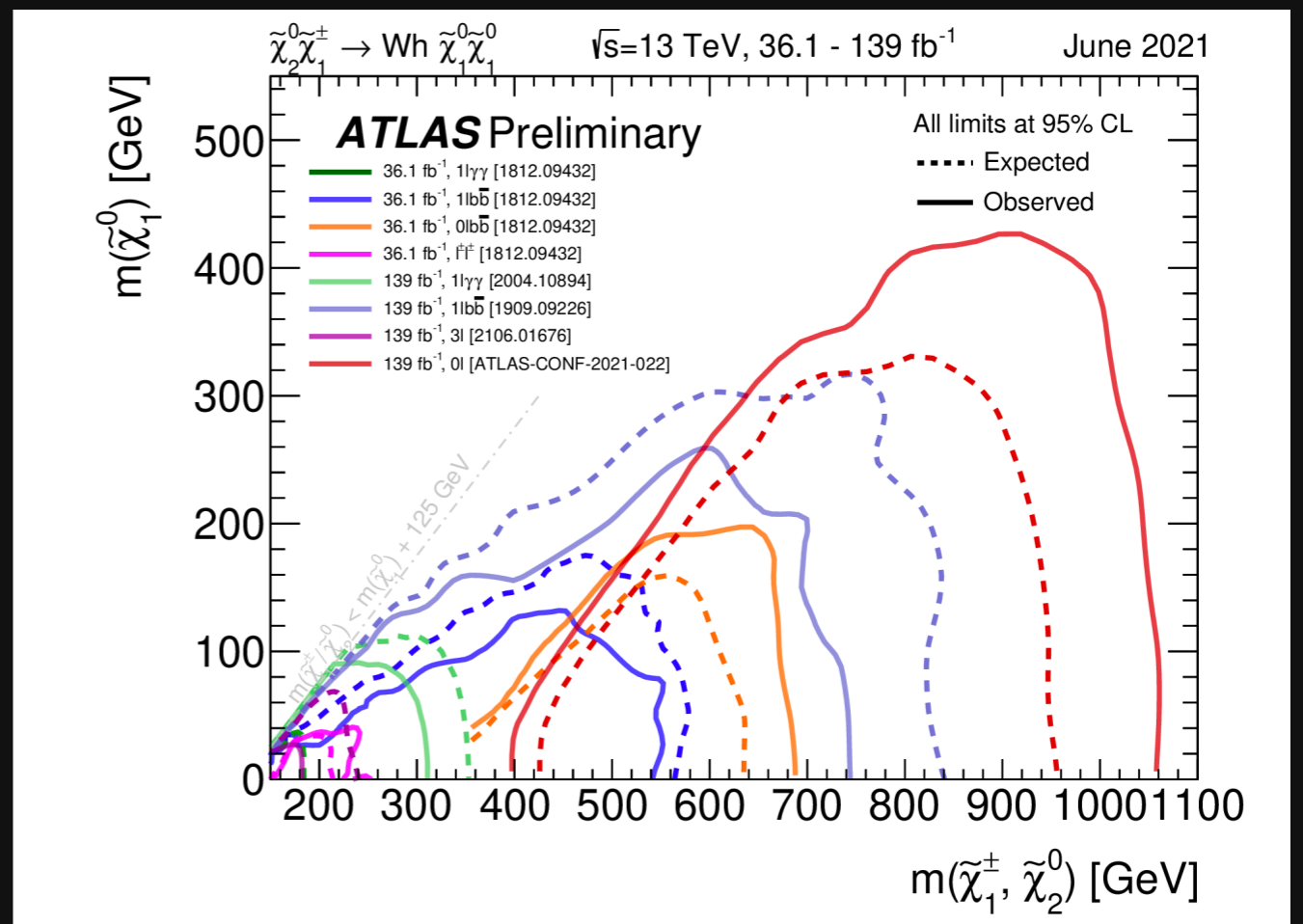
C1N2/C1C1 Summary



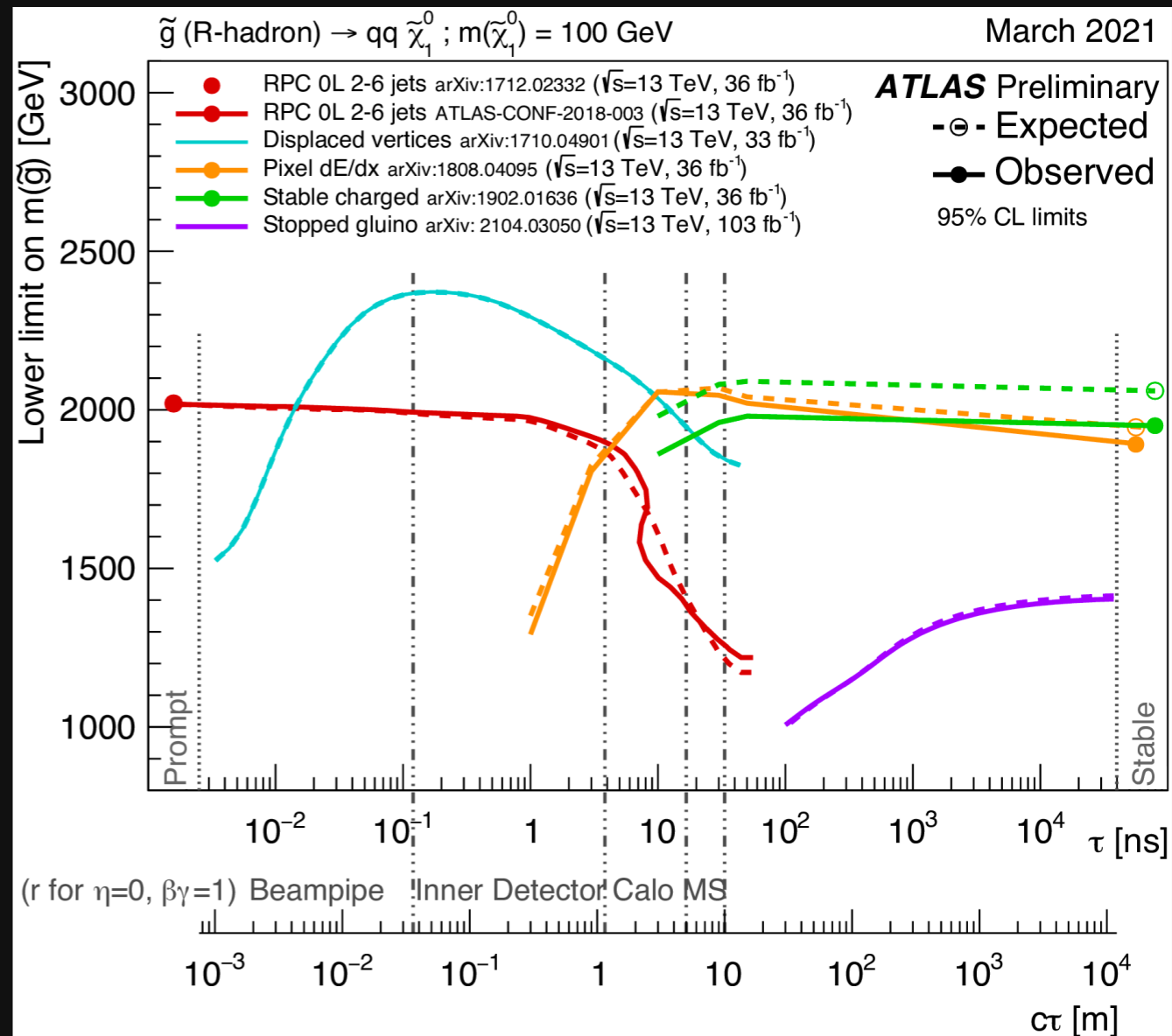
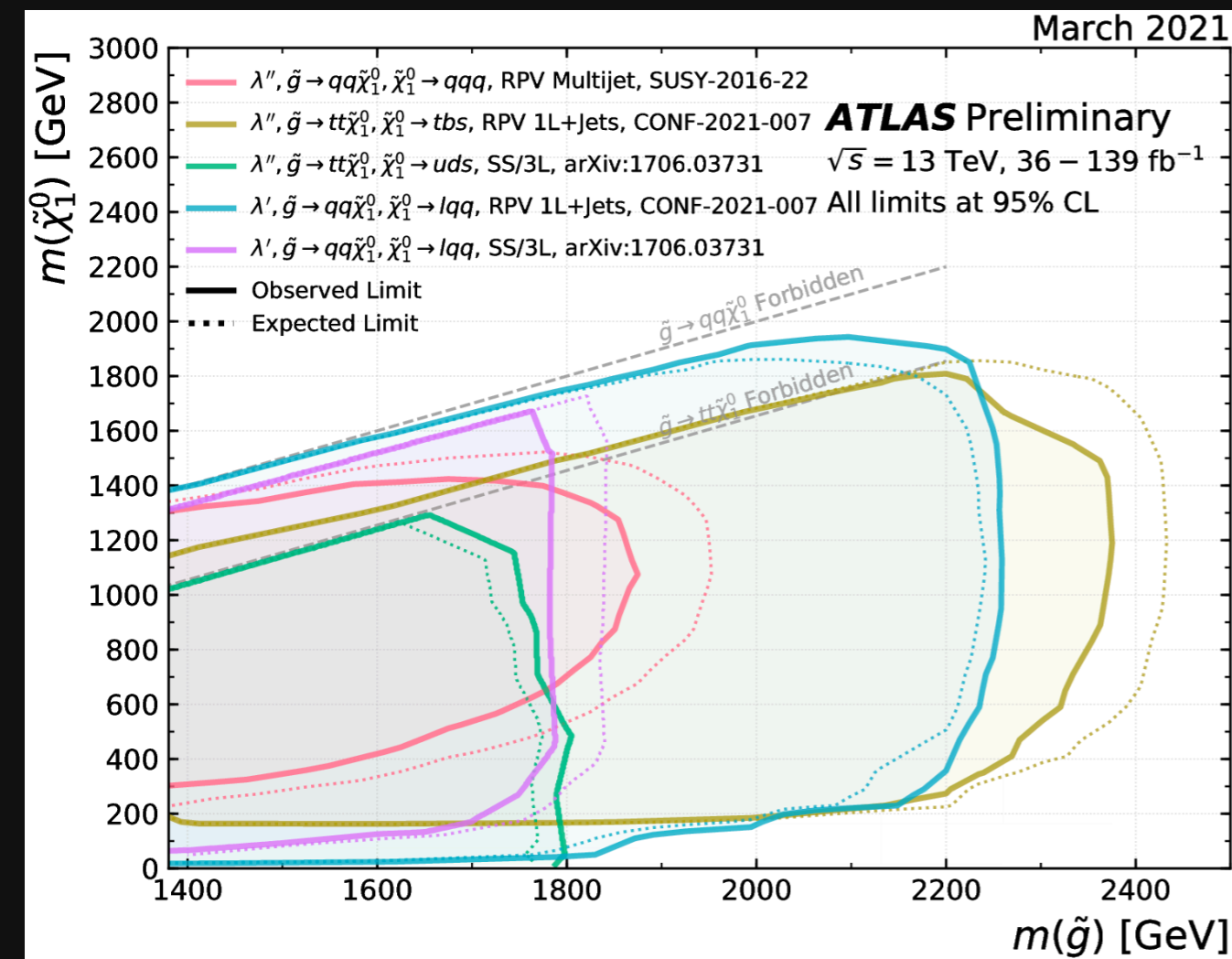
$$\tilde{\chi} \rightarrow \ell$$



$$\tilde{\chi} \rightarrow W/Z/h$$



Gluino (II) Summary



$\tilde{g} \rightarrow \text{LSP} \rightarrow \text{SM}$

$\tilde{g} \text{ R-hadron} \rightarrow \text{LSP}$

ATLAS SUSY Searches* - 95% CL Lower Limits

June 2021

ATLAS Preliminary

$\sqrt{s} = 13$ TeV

Model	Signature	$\int \mathcal{L} dt$ [fb ⁻¹]	Mass limit	Reference	
Inclusive Searches	$\tilde{q}\tilde{q}, \tilde{q} \rightarrow q\tilde{\chi}_1^0$	0 e, μ mono-jet	2-6 jets 1-3 jets E_T^{miss} 139 36.1	\tilde{q} [1x, 8x Degen.] 1.0 1.85 \tilde{q} [8x Degen.] 0.9	$m(\tilde{\chi}_1^0) < 400$ GeV $m(\tilde{q}) - m(\tilde{\chi}_1^0) = 5$ GeV 210.14293 2102.10874
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^0$	0 e, μ	2-6 jets E_T^{miss} 139	\tilde{g} 2.3 \tilde{g} Forbidden 1.15-1.95	$m(\tilde{\chi}_1^0) = 0$ GeV $m(\tilde{\chi}_1^0) = 1000$ GeV 210.14293 210.14293
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}W\tilde{\chi}_1^0$	1 e, μ	2-6 jets 139	\tilde{g} 2.2	$m(\tilde{\chi}_1^0) < 600$ GeV 2101.01629
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}(\ell\ell)\tilde{\chi}_1^0$	$ee, \mu\mu$	2 jets E_T^{miss} 36.1	\tilde{g} 1.2	$m(\tilde{g}) - m(\tilde{\chi}_1^0) = 50$ GeV 1805.11381
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}WZ\tilde{\chi}_1^0$	0 e, μ	7-11 jets E_T^{miss} 139	\tilde{g} 1.97	$m(\tilde{\chi}_1^0) < 600$ GeV 2008.06032
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^0$	SS e, μ	6 jets E_T^{miss} 139	\tilde{g} 1.15	$m(\tilde{g}) - m(\tilde{\chi}_1^0) = 200$ GeV 1909.08457
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow t\tilde{t}\tilde{\chi}_1^0$	0-1 e, μ SS e, μ	3 b 6 jets E_T^{miss} 79.8 139	\tilde{g} 2.25 \tilde{g} 1.25	$m(\tilde{\chi}_1^0) < 200$ GeV $m(\tilde{g}) - m(\tilde{\chi}_1^0) = 300$ GeV ATLAS-CONF-2018-041 1909.08457
3 rd gen. squarks direct production	$\tilde{b}_1\tilde{b}_1$	0 e, μ	2 b E_T^{miss} 139	\tilde{b}_1 1.255 \tilde{b}_1 0.68	$m(\tilde{\chi}_1^0) < 400$ GeV 10 GeV $< \Delta m(\tilde{b}_1, \tilde{\chi}_1^0) < 20$ GeV 2101.12527 2101.12527
	$\tilde{b}_1\tilde{b}_1, \tilde{b}_1 \rightarrow b\tilde{\chi}_2^0 \rightarrow bh\tilde{\chi}_1^0$	0 e, μ 2 τ	6 b 2 b E_T^{miss} 139 139	\tilde{b}_1 Forbidden 0.23-1.35 \tilde{b}_1 0.13-0.85	$\Delta m(\tilde{\chi}_2^0, \tilde{\chi}_1^0) = 130$ GeV, $m(\tilde{\chi}_1^0) = 100$ GeV $\Delta m(\tilde{\chi}_2^0, \tilde{\chi}_1^0) = 130$ GeV, $m(\tilde{\chi}_1^0) = 0$ GeV 1908.03122 ATLAS-CONF-2020-031
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow t\tilde{\chi}_1^0$	0-1 e, μ	≥ 1 jet E_T^{miss} 139	\tilde{t}_1 1.25	$m(\tilde{\chi}_1^0) = 1$ GeV 2004.14060, 2012.03799
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow Wb\tilde{\chi}_1^0$	1 e, μ	3 jets/1 b E_T^{miss} 139	\tilde{t}_1 Forbidden 0.65	$m(\tilde{\chi}_1^0) = 500$ GeV 2012.03799
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow \tilde{\tau}_1 b\nu, \tilde{\tau}_1 \rightarrow \tau\tilde{G}$	1-2 τ	2 jets/1 b E_T^{miss} 139	\tilde{t}_1 Forbidden 1.4	$m(\tilde{\tau}_1) = 800$ GeV ATLAS-CONF-2021-008
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow c\tilde{\chi}_1^0 / \tilde{c}\tilde{c}, \tilde{c} \rightarrow c\tilde{\chi}_1^0$	0 e, μ 0 e, μ	2 c mono-jet E_T^{miss} 36.1 139	\tilde{c} 0.85 \tilde{t}_1 0.55	$m(\tilde{\chi}_1^0) = 0$ GeV $m(\tilde{t}_1, \tilde{c}) - m(\tilde{\chi}_1^0) = 5$ GeV 1805.01649 2102.10874
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow t\tilde{\chi}_2^0, \tilde{\chi}_2^0 \rightarrow Z/h\tilde{\chi}_1^0$	1-2 e, μ	1-4 b E_T^{miss} 139	\tilde{t}_1 0.067-1.18	$m(\tilde{\chi}_2^0) = 500$ GeV 2006.05880
	$\tilde{t}_2\tilde{t}_2, \tilde{t}_2 \rightarrow \tilde{t}_1 + Z$	3 e, μ	1 b E_T^{miss} 139	\tilde{t}_2 Forbidden 0.86	$m(\tilde{\chi}_1^0) = 360$ GeV, $m(\tilde{t}_1) - m(\tilde{\chi}_1^0) = 40$ GeV 2006.05880
EW direct	$\tilde{\chi}_1^\pm \tilde{\chi}_2^0$ via WZ	Multiple ℓ /jets $ee, \mu\mu$	≥ 1 jet E_T^{miss} 139 139	$\tilde{\chi}_1^\pm / \tilde{\chi}_2^0$ 0.96 $\tilde{\chi}_1^\pm / \tilde{\chi}_2^0$ 0.205	$m(\tilde{\chi}_1^0) = 0$, wino-bino $m(\tilde{\chi}_1^\pm) - m(\tilde{\chi}_1^0) = 5$ GeV, wino-bino 2106.01676, ATLAS-CONF-2021-022 1911.12606
	$\tilde{\chi}_1^\pm \tilde{\chi}_1^\mp$ via WW	2 e, μ	E_T^{miss} 139	$\tilde{\chi}_1^\pm$ 0.42	$m(\tilde{\chi}_1^0) = 0$, wino-bino 1908.08215
	$\tilde{\chi}_1^\pm \tilde{\chi}_2^0$ via Wh	Multiple ℓ /jets	E_T^{miss} 139	$\tilde{\chi}_1^\pm / \tilde{\chi}_2^0$ Forbidden 1.06	$m(\tilde{\chi}_1^0) = 70$ GeV, wino-bino 2004.10894, ATLAS-CONF-2021-022
	$\tilde{\chi}_1^\pm \tilde{\chi}_1^\mp$ via $\tilde{\ell}_L/\tilde{\nu}$	2 e, μ	E_T^{miss} 139	$\tilde{\chi}_1^\pm$ 1.0	$m(\tilde{\ell}, \tilde{\nu}) = 0.5(m(\tilde{\chi}_1^\pm) + m(\tilde{\chi}_1^0))$ 1908.08215
	$\tilde{\tau}\tilde{\tau}, \tilde{\tau} \rightarrow \tau\tilde{\chi}_1^0$	2 τ	E_T^{miss} 139	$\tilde{\tau}$ [$\tilde{\tau}_L, \tilde{\tau}_{R,L}$] 0.16-0.3 0.12-0.39	$m(\tilde{\chi}_1^0) = 0$ 1911.06660
	$\tilde{\ell}_{L,R}\tilde{\ell}_{L,R}, \tilde{\ell} \rightarrow \ell\tilde{\chi}_1^0$	2 e, μ $ee, \mu\mu$	0 jets ≥ 1 jet E_T^{miss} 139 139	$\tilde{\ell}$ 0.7 $\tilde{\ell}$ 0.256	$m(\tilde{\chi}_1^0) = 0$ $m(\tilde{\ell}) - m(\tilde{\chi}_1^0) = 10$ GeV 1908.08215 1911.12606
	$\tilde{H}\tilde{H}, \tilde{H} \rightarrow h\tilde{G}/Z\tilde{G}$	0 e, μ 4 e, μ 0 e, μ	≥ 3 b 0 jets ≥ 2 large jets E_T^{miss} 36.1 139 139	\tilde{H} 0.13-0.23 0.29-0.88 \tilde{H} 0.55 \tilde{H} 0.45-0.93	$\text{BR}(\tilde{\chi}_1^0 \rightarrow h\tilde{G}) = 1$ $\text{BR}(\tilde{\chi}_1^0 \rightarrow Z\tilde{G}) = 1$ $\text{BR}(\tilde{\chi}_1^0 \rightarrow Z\tilde{G}) = 1$ 1806.04030 2103.11684 ATLAS-CONF-2021-022
	Long-lived particles	Direct $\tilde{\chi}_1^+ \tilde{\chi}_1^-$ prod., long-lived $\tilde{\chi}_1^\pm$	Disapp. trk	1 jet E_T^{miss} 139	$\tilde{\chi}_1^\pm$ 0.66 $\tilde{\chi}_1^\pm$ 0.21
Stable \tilde{g} R-hadron		Multiple	36.1	\tilde{g} 2.0	1902.01636, 1808.04095
Metastable \tilde{g} R-hadron, $\tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^0$		Multiple	36.1	\tilde{g} [$\tau(\tilde{g}) = 10$ ns, 0.2 ns] 2.05 2.4	$m(\tilde{\chi}_1^0) = 100$ GeV 1710.04901, 1808.04095
$\tilde{\ell}\tilde{\ell}, \tilde{\ell} \rightarrow \ell\tilde{G}$		Displ. lep	E_T^{miss} 139	$\tilde{\ell}, \tilde{\mu}$ 0.7 $\tilde{\tau}$ 0.34	$\tau(\tilde{\ell}) = 0.1$ ns $\tau(\tilde{\ell}) = 0.1$ ns 2011.07812 2011.07812
RPV	$\tilde{\chi}_1^\pm \tilde{\chi}_1^\mp / \tilde{\chi}_1^0, \tilde{\chi}_1^\pm \rightarrow Z\ell \rightarrow \ell\ell\ell$	3 e, μ	139	$\tilde{\chi}_1^\pm / \tilde{\chi}_1^0$ [BR(Z τ)=1, BR(Z e)=1] 0.625 1.05	Pure Wino 2011.10543
	$\tilde{\chi}_1^\pm \tilde{\chi}_1^\mp / \tilde{\chi}_2^0 \rightarrow WW/Z\ell\ell\ell\nu\nu$	4 e, μ	0 jets E_T^{miss} 139	$\tilde{\chi}_1^\pm / \tilde{\chi}_2^0$ [$\lambda_{133} \neq 0, \lambda_{12k} \neq 0$] 0.95 1.55	$m(\tilde{\chi}_1^0) = 200$ GeV 2103.11684
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow q\tilde{q}\tilde{\chi}_1^0$	4-5 large jets	36.1	\tilde{g} [$m(\tilde{\chi}_1^0) = 200$ GeV, 1100 GeV] 1.3 1.9	Large λ'_{112} 1804.03568
	$\tilde{t}\tilde{t}, \tilde{t} \rightarrow t\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow t\tilde{b}s$	Multiple	36.1	\tilde{t} [$\lambda'_{323} = 2e-4, 1e-2$] 0.55 1.05	$m(\tilde{\chi}_1^0) = 200$ GeV, bino-like ATLAS-CONF-2018-003
	$\tilde{t}\tilde{t}, \tilde{t} \rightarrow b\tilde{\chi}_1^\pm, \tilde{\chi}_1^\pm \rightarrow b\tilde{b}s$	$\geq 4b$	139	\tilde{t} Forbidden 0.95	$m(\tilde{\chi}_1^0) = 500$ GeV 2010.01015
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow b\tilde{s}$	2 jets + 2 b	36.7	\tilde{t}_1 [qq, bs] 0.42 0.61	1710.07171
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow q\tilde{\ell}$	2 e, μ 1 μ	2 b DV 36.1 136	\tilde{t}_1 0.4-1.45 \tilde{t}_1 [$1e-10 < \lambda'_{23k} < 1e-8, 3e-10 < \lambda'_{23k} < 3e-9$] 1.0 1.6	$\text{BR}(\tilde{t}_1 \rightarrow b\tilde{e}/b\tilde{\mu}) > 20\%$ $\text{BR}(\tilde{t}_1 \rightarrow q\tilde{\mu}) = 100\%, \cos\theta_t = 1$ 1710.05544 2003.11956
	$\tilde{\chi}_1^\pm / \tilde{\chi}_2^0 / \tilde{\chi}_1^0, \tilde{\chi}_{1,2}^0 \rightarrow t\tilde{b}s, \tilde{\chi}_1^\pm \rightarrow b\tilde{b}s$	1-2 e, μ	≥ 6 jets 139	$\tilde{\chi}_1^0$ 0.2-0.32	Pure higgsino ATLAS-CONF-2021-007

*Only a selection of the available mass limits on new states or phenomena is shown. Many of the limits are based on simplified models, c.f. refs. for the assumptions made.

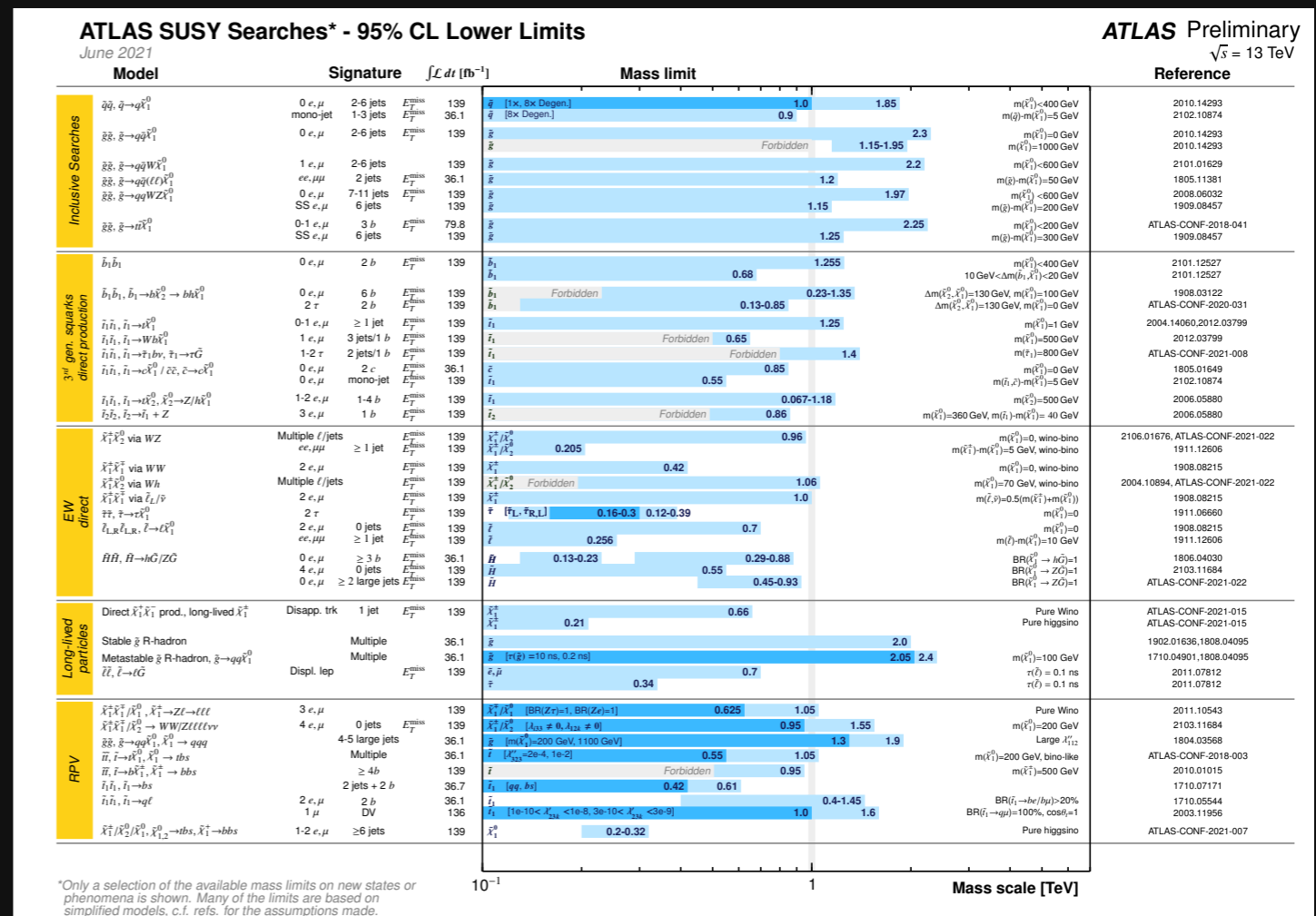
10⁻¹

1

Mass scale [TeV]

Conclusion

- Gearing up for the 3rd operation of the LHC and the ATLAS detector in March 2022
- ATLAS has a robust SUSY search programme
 - Take advantage of likelihood preservation and reinterpretation to expand coverage
- Identify uncovered areas and find new rocks to look under for SUSY



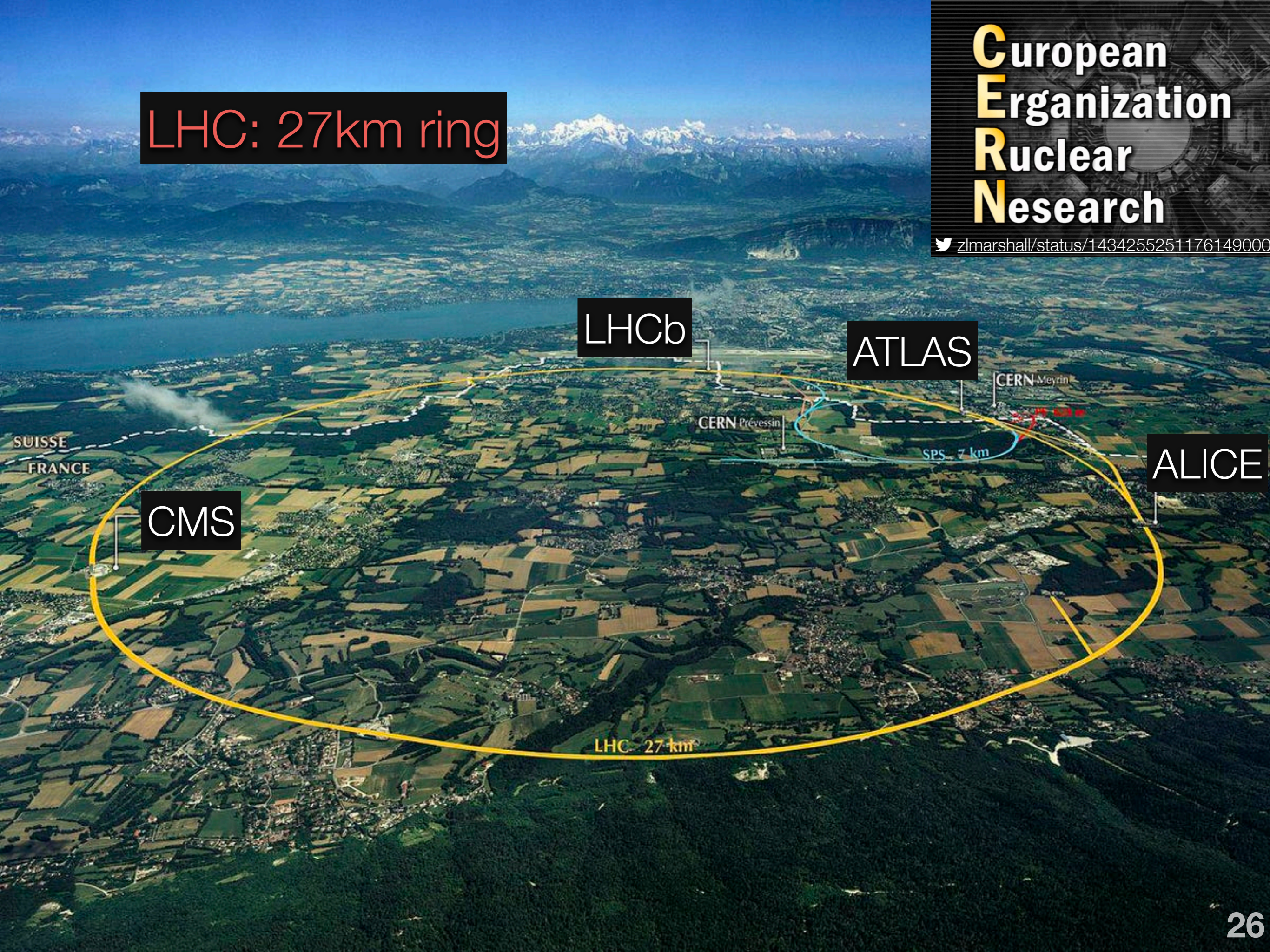
Backup

Rama lama lama ka dinga da dinga dong

LHC: 27km ring

European Organization Nuclear Research

[zlmarshall/status/1434255251176149000](https://twitter.com/zlmarshall/status/1434255251176149000)



LHCb

ATLAS

CERN Meyrin

CERN Prévessin

SPS 7 km

SUISSE
FRANCE

CMS

ALICE

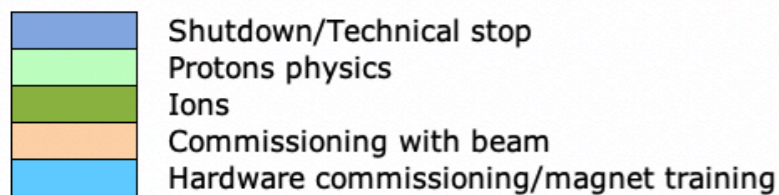
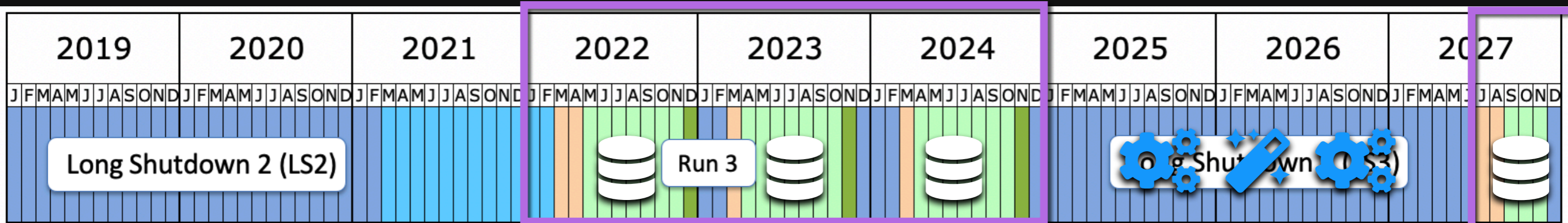
LHC 27 km

LHC Schedule

No data collection

Run 3

Run 4+

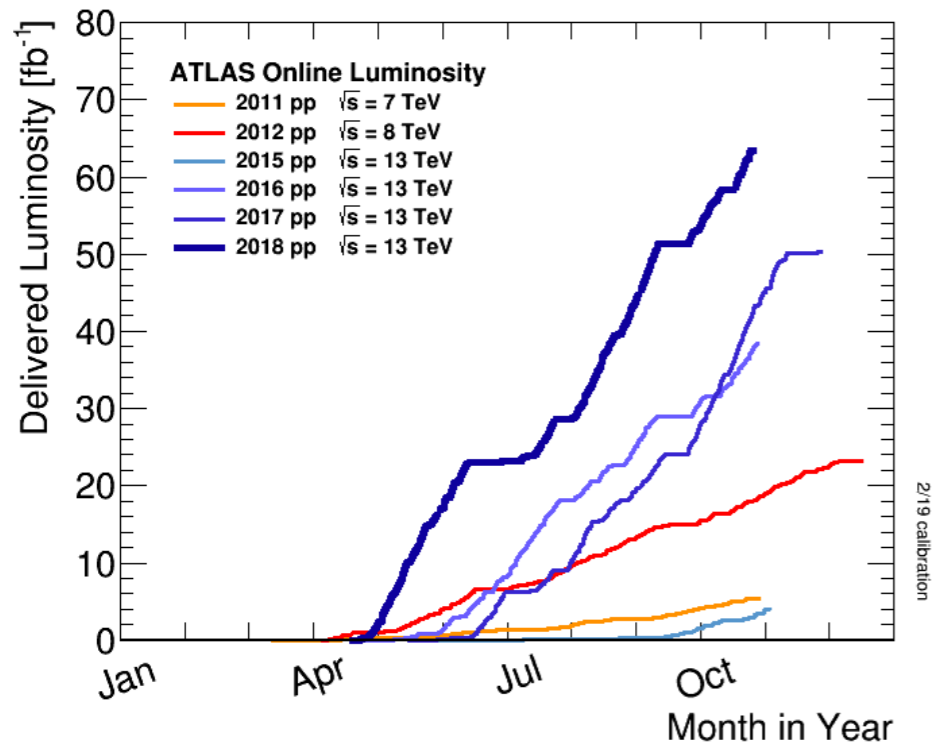


Here be dragons 

Experiments currently shutdown, **no new data**  until early 2022!

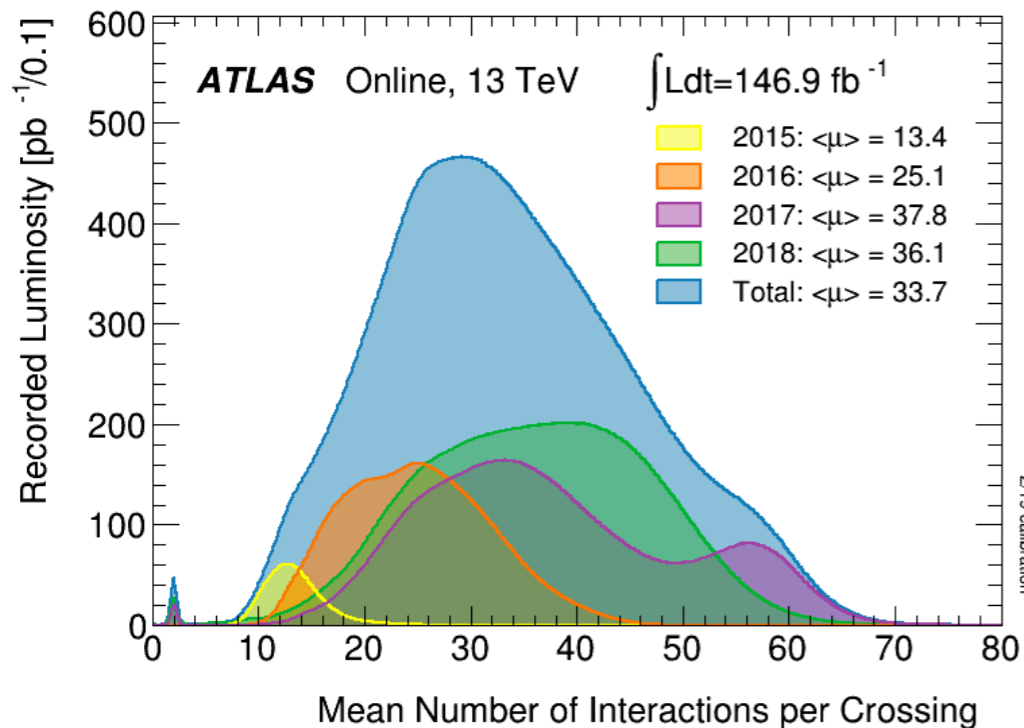
- ✦ Focus on doing more with what we have
 - ✦ clever techniques to find new physics (SUSY?) in existing data
 - ⚠ **global fits and large scale combinations to determine future directions**
- ✦ **Finalized calibrations** on physics objects (electrons, muons, jets, photons) and pushed **object definitions to lower energies**

The Large Hadron Collider



A proton-proton collider at 13 TeV center-of-mass energy

- ✦ For 2015-2018 operation:
 - ✦ Operating peak luminosity: $2.10 \times 10^{34} \text{ cm}^{-2} \text{ s}^{-1}$
 - ✦ Proton bunch spacing: **25 ns** — 40 million crossings per second
 - ✦ 2808 bunches colliding in ATLAS
 - ✦ **146.9 ifb of data delivered**
 - ✦ Up to 70 collisions per bunch crossing (**pileup!**) — billions of collisions per second!



❓ *Why that peak at low $\langle \mu \rangle$? ATLAS does many special runs. This one was for precision W-physics.*

Beyond the Standard Model

What is dark matter?

Where did all the antimatter go?

Why does the standard model look the way it does?

Why is the weak force so much stronger than gravity? (Hierarchy problem)

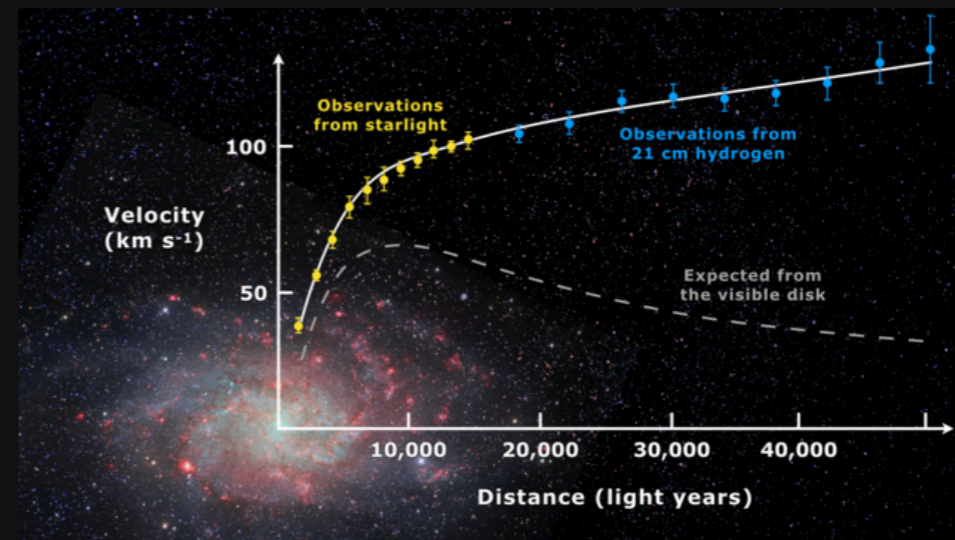
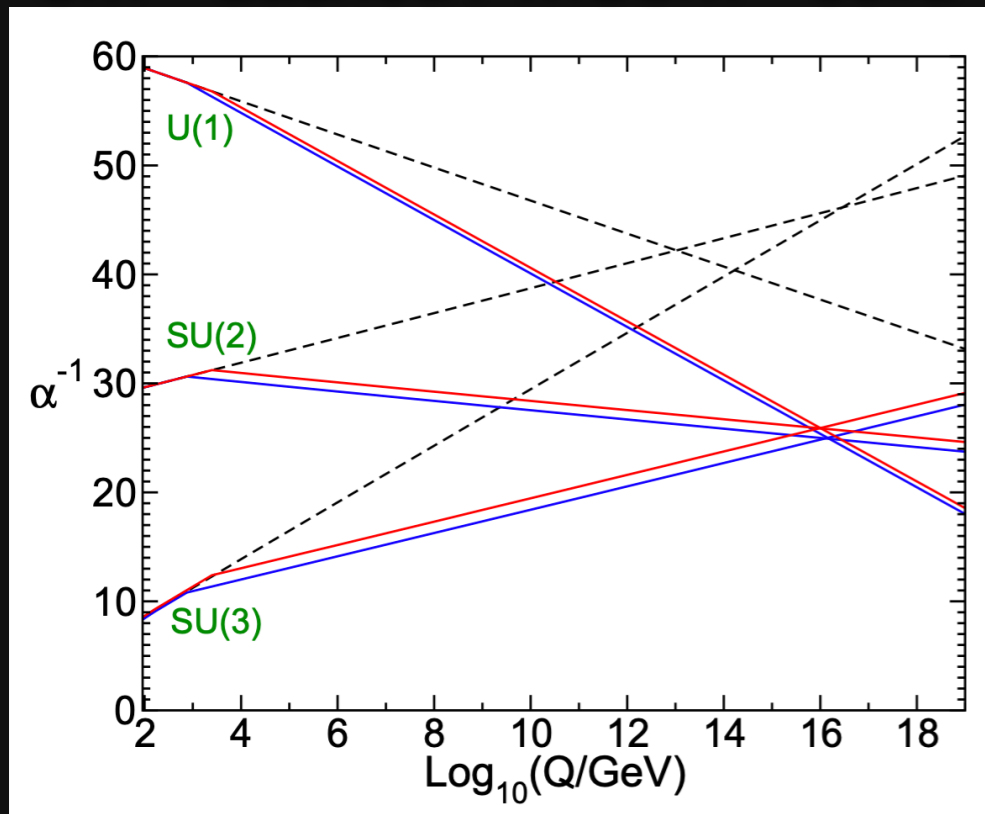
Supersymmetry (SUSY) is a framework with good theoretical motivations in which theorists can study BSM physics



Supersymmetry (SUSY) is a set of benchmark models to help experimentalists answer these questions!

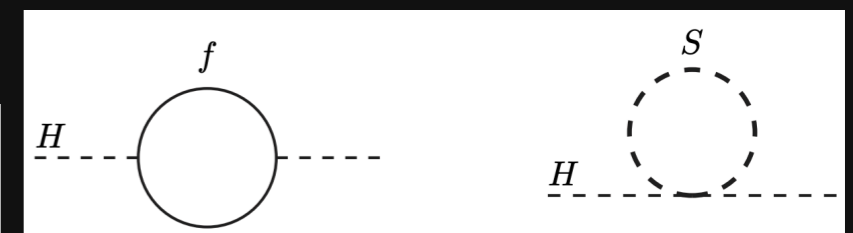
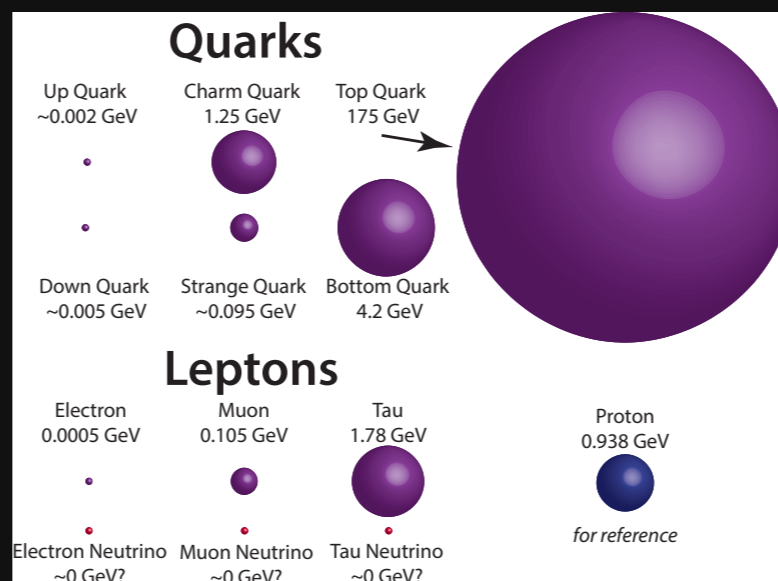
Beyond the Standard Model

SM - dashed / MSSM - solid



Cold Dark Matter candidates?

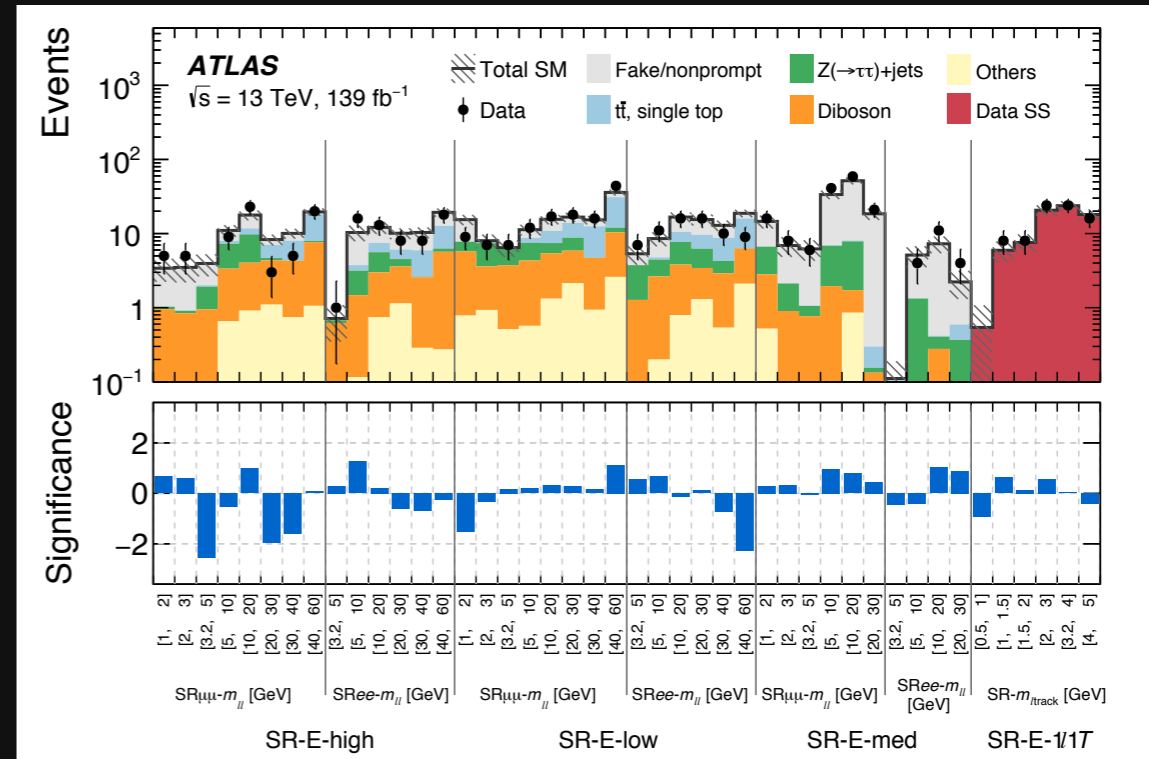
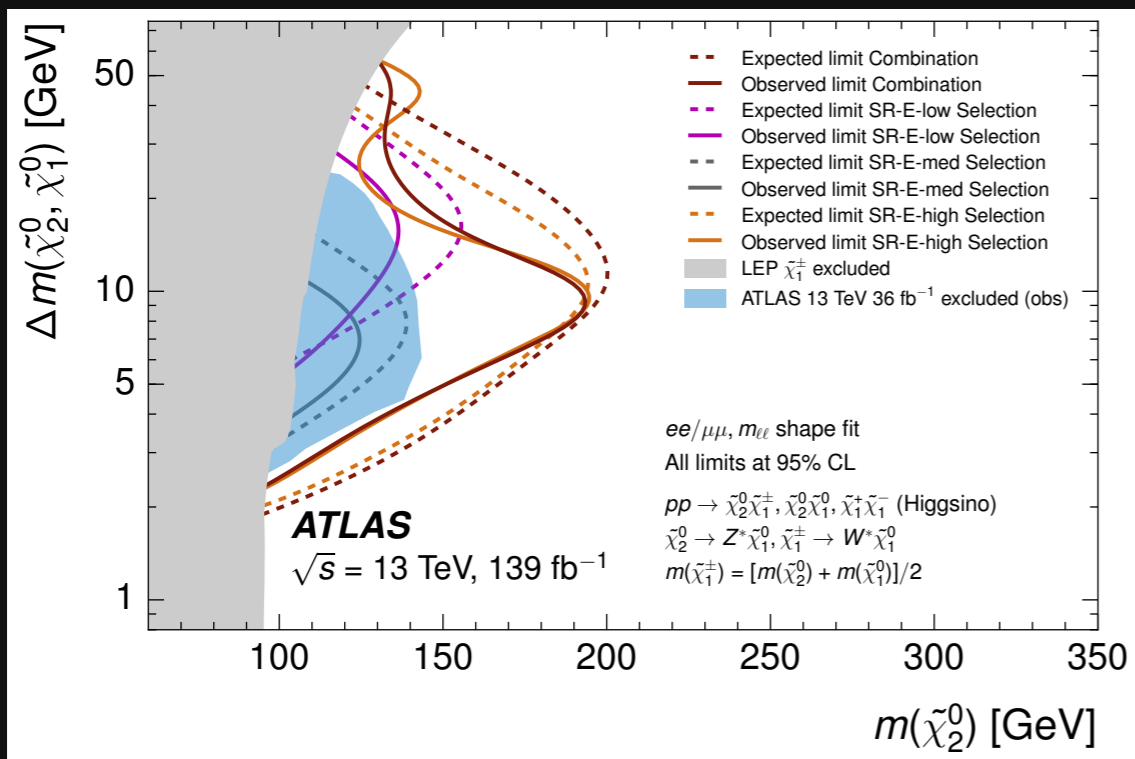
Unification of couplings at LHC energy scale?



Hierarchy Problem

Higgsino combination

- Why is the observed weaker than the expected?
- SR-E-high and SR-E-low both have worse observed than predicted in that region. $\Delta m \approx 20$ GeV corresponds with $m_{\ell\ell} \approx 10$ GeV and there are some correlated (small) excesses across the bins in that phase-space.



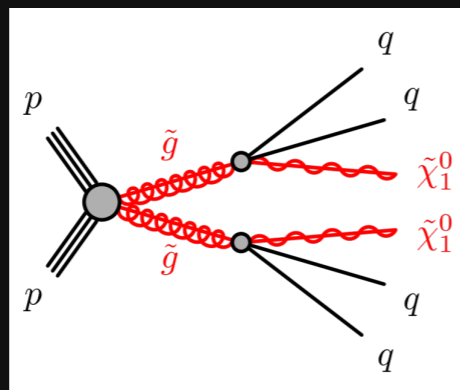
RPC-RPV Spectrum

$$\mathcal{W}_{\text{RPV}} = \frac{\lambda_{ijk}}{2} L_i L_j \bar{E}_k + \lambda'_{ijk} L_i Q_j \bar{D}_k + \frac{\lambda''_{ijk}}{2} \bar{U}_i \bar{D}_j \bar{D}_k + \kappa_i L_i H_u,$$

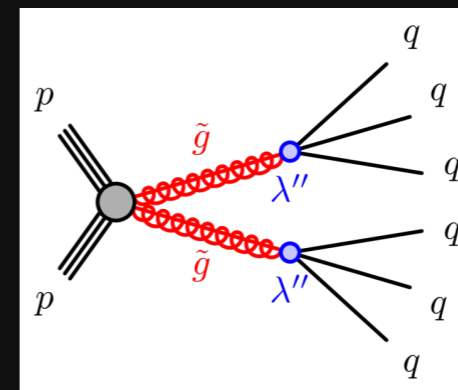
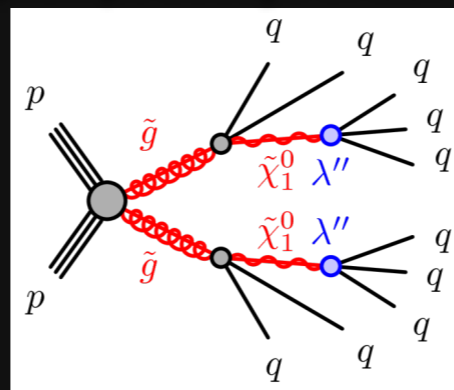
RPV terms in the SUSY superpotential

- ✦ All simplified models fall somewhere on the RPC-RPV spectrum... either
 - ✦ set all terms large or ...
 - ✦ set all terms to 0 (conserve "SUSYness", RPC) and require the lightest SUSY particle (LSP) to be stable → dark matter candidates?

long-lived LSPs
displaced decays?



prompt decays of LSPs



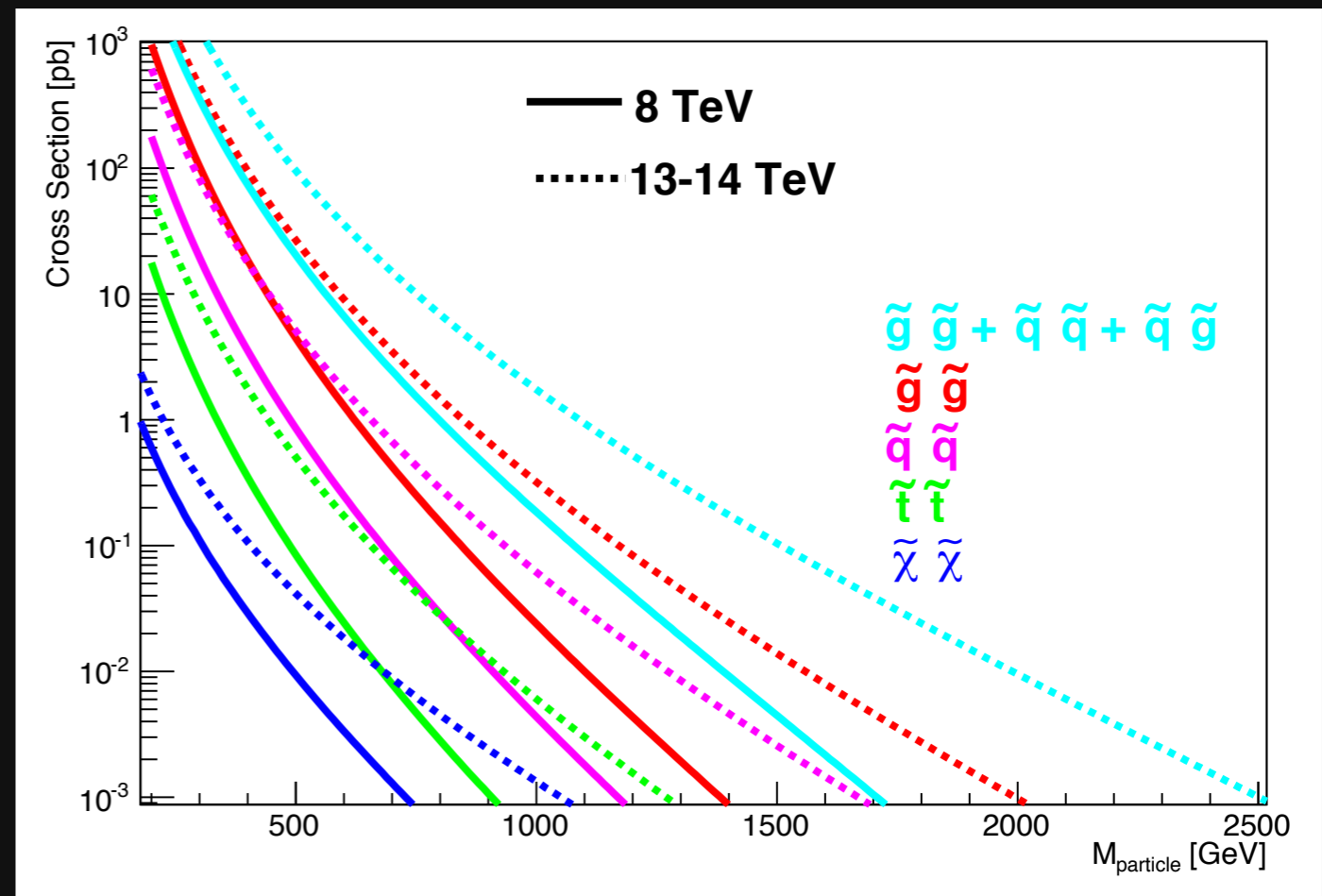
prompt decays
of gluinos

increasing λ'' →

! Naturalness motivates light gluinos and stops

Searching for SUSY

- ✦ **Gluinos**, because of their strong color coupling, have the highest theoretical cross-section of the sparticles found at the LHC
 - ✦ The upgrade of LHC from 8 TeV to 13 TeV also provides an order of magnitude increase in the theoretical cross-section
- ✦ Theoretical cross-sections are shown for:
 - ✦ **total strong production**
 - ✦ **gluino production**
 - ✦ **total squark production**
 - ✦ **heavy squark**
 - ✦ **electroweak**



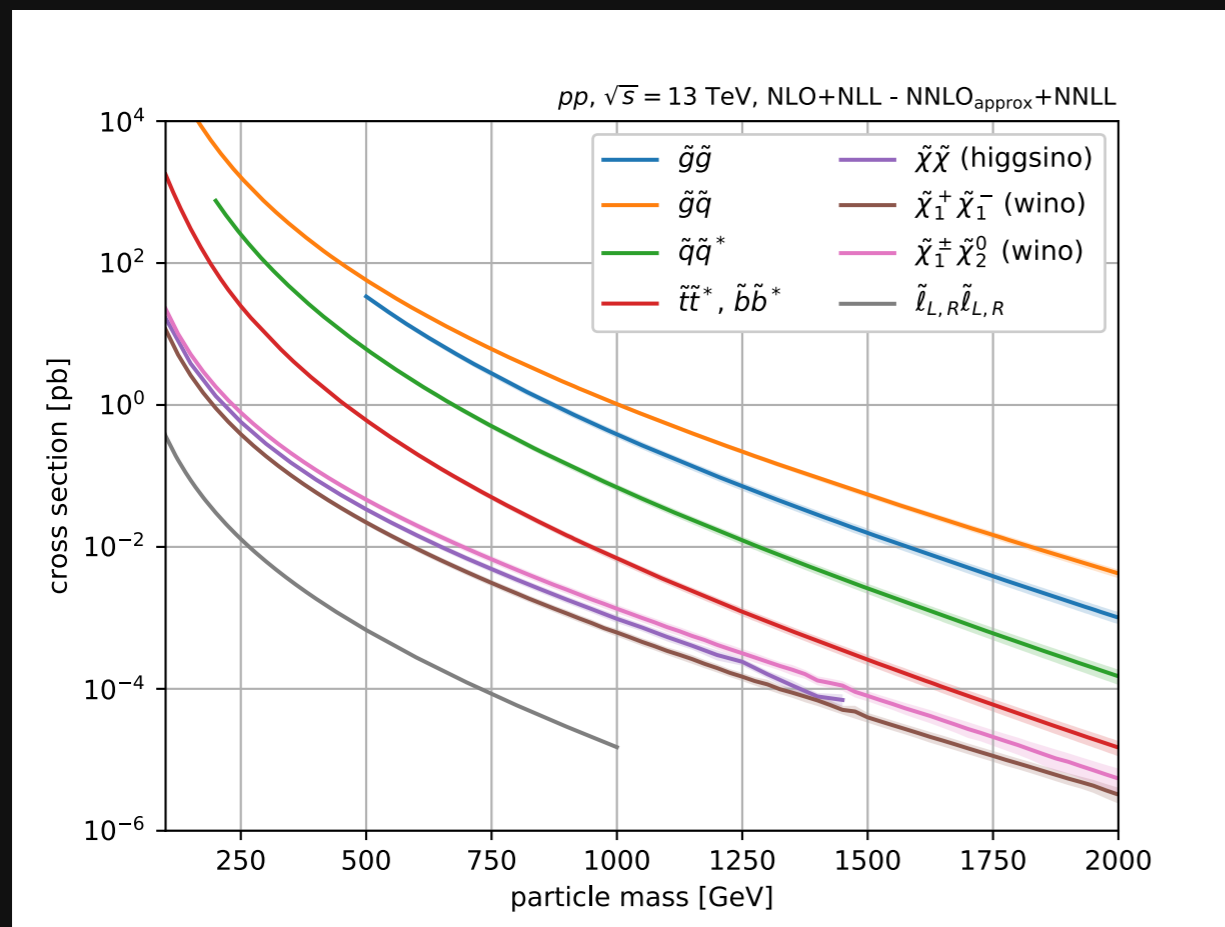
[\[1407.5066\]](#)

[\[1206.2892\]](#)

Search for strongly-produced sparticles!

(electroweak states may be first detected if high mass limits on strong production)

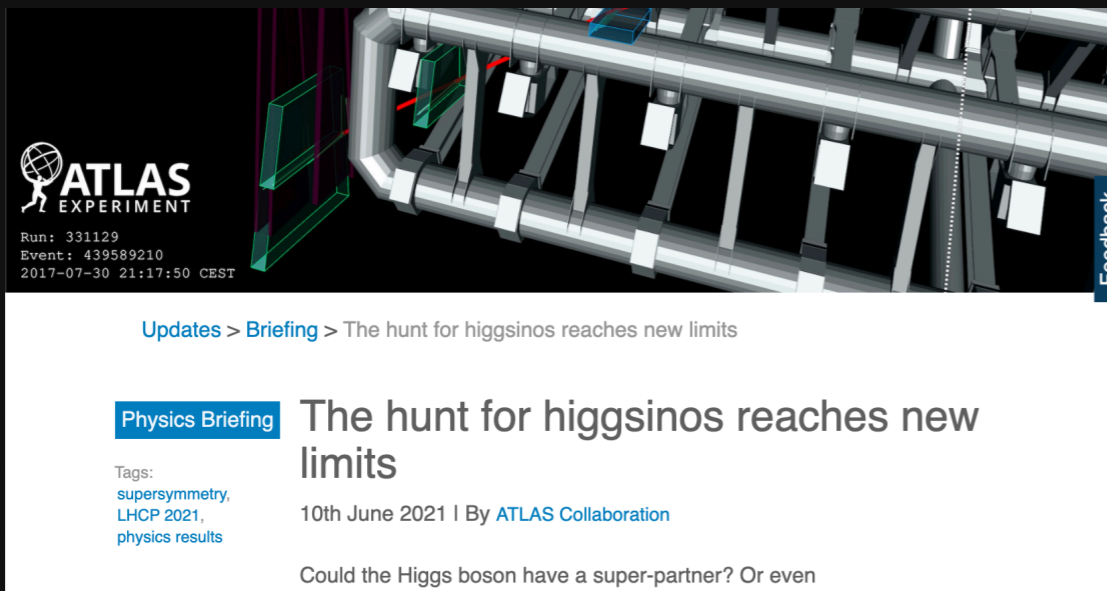
Cross Sections



- ✦ LHC 13 TeV well-motivated to search for SUSY (some searches are possible for the first time!)
- ✦ Typically: Strong/3G will have larger cross-sections than electroweak (EWK)

 Motivate: search for light sparticles

Technical: combine... how?



▪ Create the workspaces: 3L on-shell, 3L off-shell, and 2L

- `pyhf patchset apply --name CN_WZ_200_190 bkg_onshell.json onshell_patchset.json --output-file CN_WZ_200_100_onshell.json`
- `pyhf patchset apply --name CN_WZ_200_190_offshell bkg_offshell.json offshell_patchset.json --output-file CN_WZ_200_190_offshell.json`
- `pyhf patchset apply --name 200p0_190p0 EWKinos_bkgonly.json EWKinos_patchset.json --output-file CN_WZ_200_190_2L.json`

▪ Rename modifiers (“systematics”) in a workspace to (de)correlate nuisance parameters

- `pyhf rename -m mu_Ztt_lowmet mu_CMP_Ztt_lowmet -m EG_RESOLUTION_ALL COMB_EG_RESOLUTION_ALL -m ...`

▪ Combine the workspaces

- `pyhf combine --join outer CN_WZ_200_190_onshell.json CN_WZ_200_190_offshell.json --output-file combined_CN_WZ_200_190_3L.json`
- `pyhf combine --join outer CN_WZ_200_190_3L.json CN_WZ_200_190_2L.json --output-file combined_CN_WZ_200_190L.json`

▪ Perform inference

- `pyhf cls combined_CN_WZ_200_100.json`

g-2 references

- Some papers on it
 - <https://arxiv.org/abs/2104.03302> (pheno)
 - <https://arxiv.org/abs/2104.03294> (MSSM)
- RPV with LLE/LQD
 - <https://arxiv.org/abs/1511.08874>
 - <https://arxiv.org/abs/2104.03294>

light SUSY particles → **g-2 contributions:**

$\tilde{\mu}_R$	$\tilde{\mu}_L/\tilde{\nu}_\mu$	\tilde{B}	\tilde{W}	\tilde{h}
	✓		✓	✓
	✓	✓		✓
✓		✓		✓
✓	✓	✓		

$$a_\mu^{\text{WHL}} = \frac{\alpha_2}{4\pi} \frac{m_\mu^2}{M_2\mu} \tan\beta \cdot f_C\left(\frac{M_2^2}{m_{\tilde{\nu}_\mu}^2}, \frac{\mu^2}{m_{\tilde{\nu}_\mu}^2}\right) - \frac{\alpha_2}{8\pi} \frac{m_\mu^2}{M_2\mu} \tan\beta \cdot f_N\left(\frac{M_2^2}{m_{\tilde{\mu}_L}^2}, \frac{\mu^2}{m_{\tilde{\mu}_L}^2}\right)$$

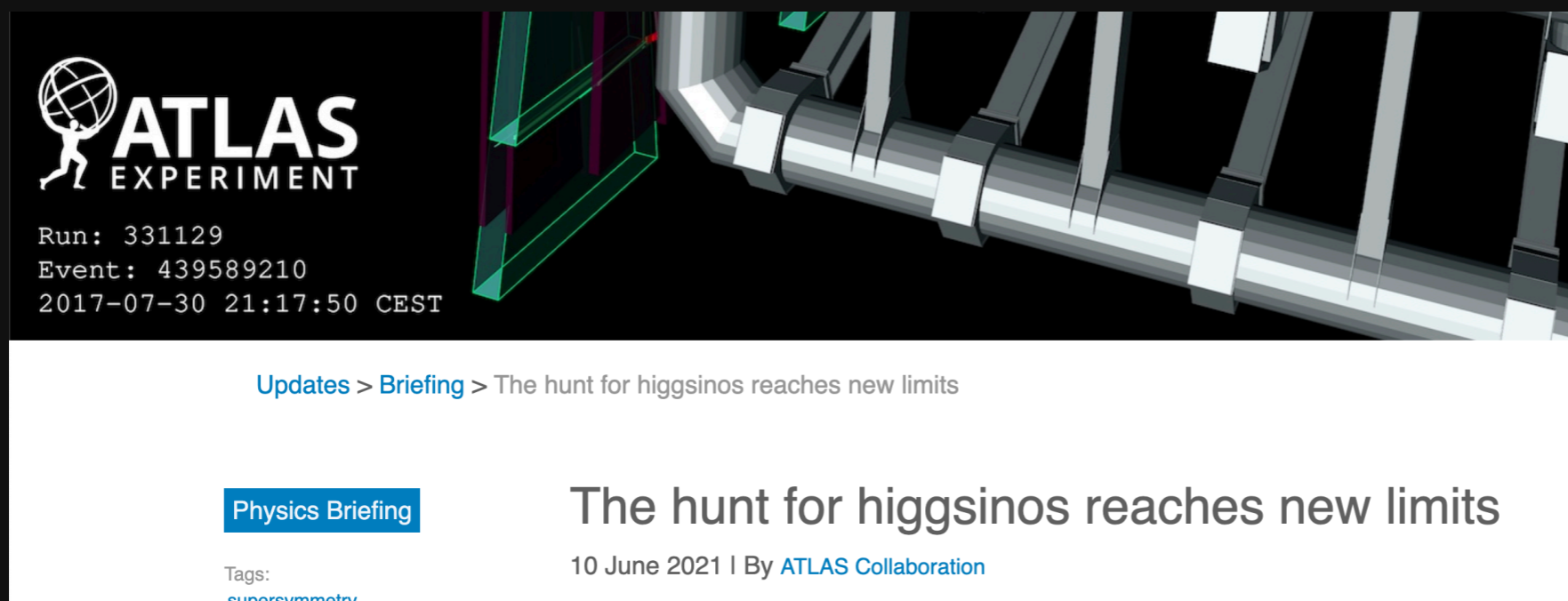
$$a_\mu^{\text{BHL}} = \frac{\alpha_Y}{8\pi} \frac{m_\mu^2}{M_1\mu} \tan\beta \cdot f_N\left(\frac{M_1^2}{m_{\tilde{\mu}_L}^2}, \frac{\mu^2}{m_{\tilde{\mu}_L}^2}\right)$$

$$a_\mu^{\text{BHR}} = -\frac{\alpha_Y}{4\pi} \frac{m_\mu^2}{M_1\mu} \tan\beta \cdot f_N\left(\frac{M_1^2}{m_{\tilde{\mu}_R}^2}, \frac{\mu^2}{m_{\tilde{\mu}_R}^2}\right)$$

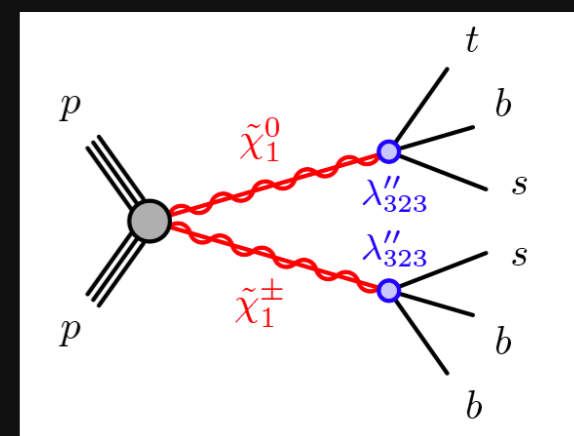
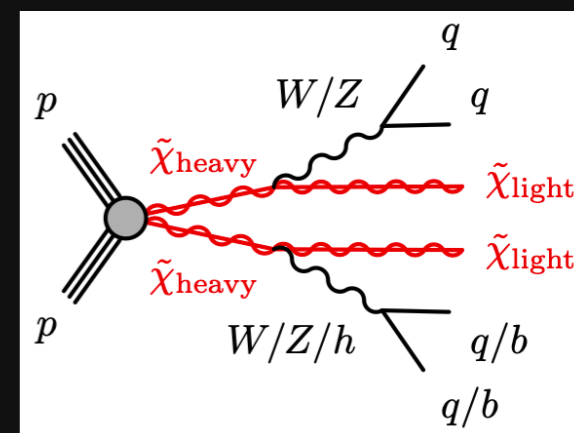
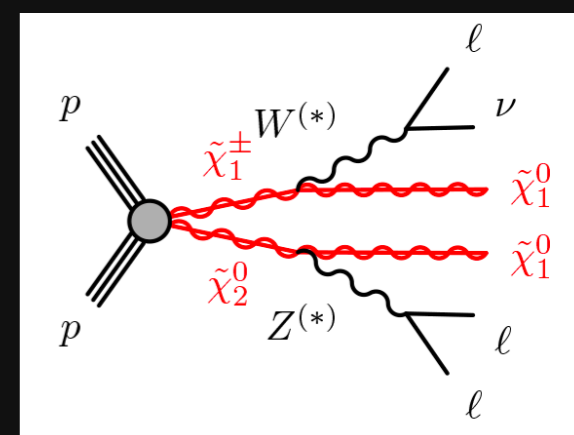
$$a_\mu^{\text{BLR}} = \frac{\alpha_Y}{4\pi} \frac{m_\mu^2 M_1\mu}{m_{\tilde{\mu}_L}^2 m_{\tilde{\mu}_R}^2} \tan\beta \cdot f_N\left(\frac{m_{\tilde{\mu}_L}^2}{M_1^2}, \frac{m_{\tilde{\mu}_R}^2}{M_1^2}\right),$$

credit: B. Hooberman

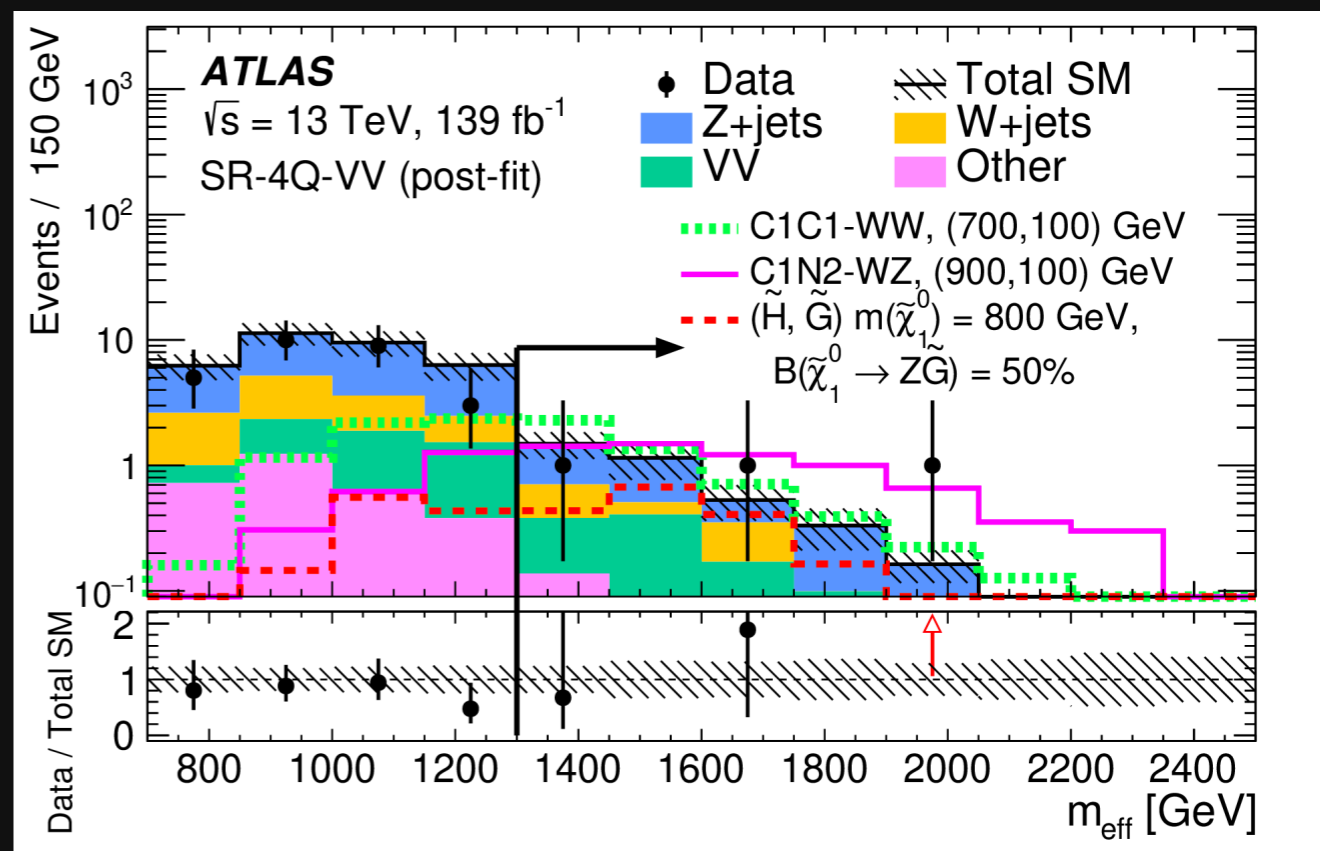
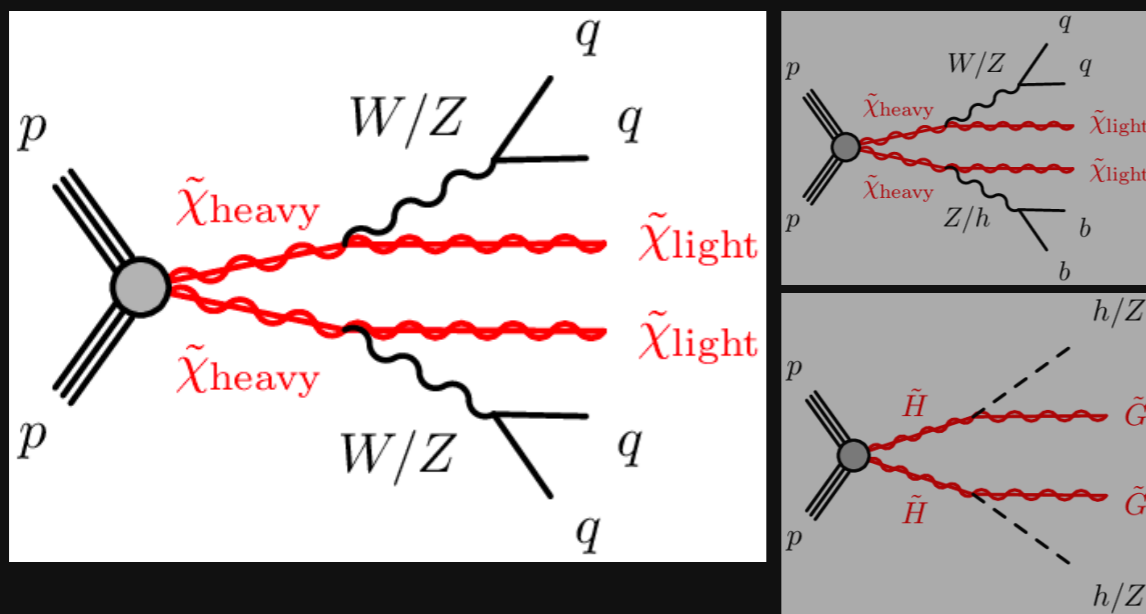
Brief sidebar



- If we consider SUSY RPC models with a Higgsino LSP, heavier Higgsinos are theoretically predicted to decay only into the LSP and soft SM particles
- Two different avenues for searching, each with unique challenges:
 - very **heavy** (hadronic) Higgsinos, decay into lighter SUSY particles and **boosted** SM bosons
 - very **light** (leptonic) Higgsinos, decay into **soft** leptons (poor reconstruction efficiency!)

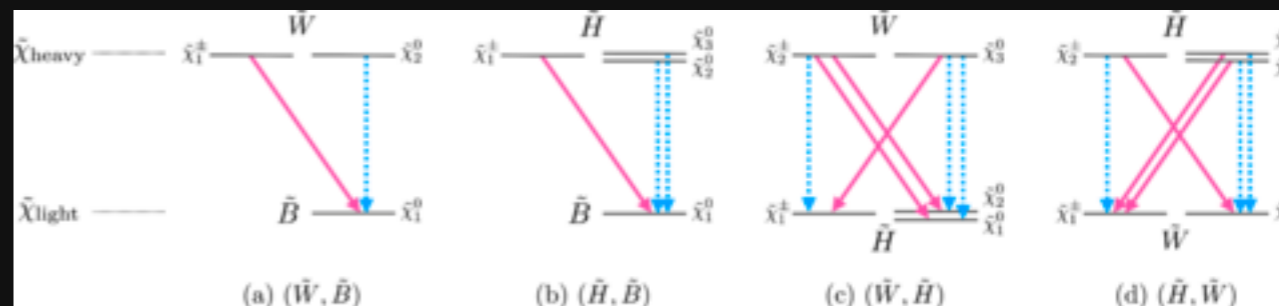


W/Z Hadronic



- MODEL PARAMETERS: $\tilde{W}, \tilde{H}, \tilde{B}$
- FINAL STATE: boosted jets, MET, no leptons
- THREE SIGNAL SCENARIOS (10 exclusion regions):
 - $(\tilde{\chi}_h, \tilde{\chi}_l) \in (\tilde{W}, \tilde{H}, \tilde{B})$ (RPC neutralino LSP)
 - $\tilde{H} \rightarrow h/Z + \tilde{G}$ (GGM, Gravitino LSP)
 - $\tilde{H} \rightarrow h/Z + \tilde{a}$ (Naturalness axino LSP)
- DOMINANT BACKGROUNDS: $Z(\rightarrow \nu\nu) + \text{jets}$, $W(\rightarrow \ell\nu) + \text{jets}$, VV
- CHALLENGE: “fake boson jet” estimation, semi-data-driven approach, accidental substructure

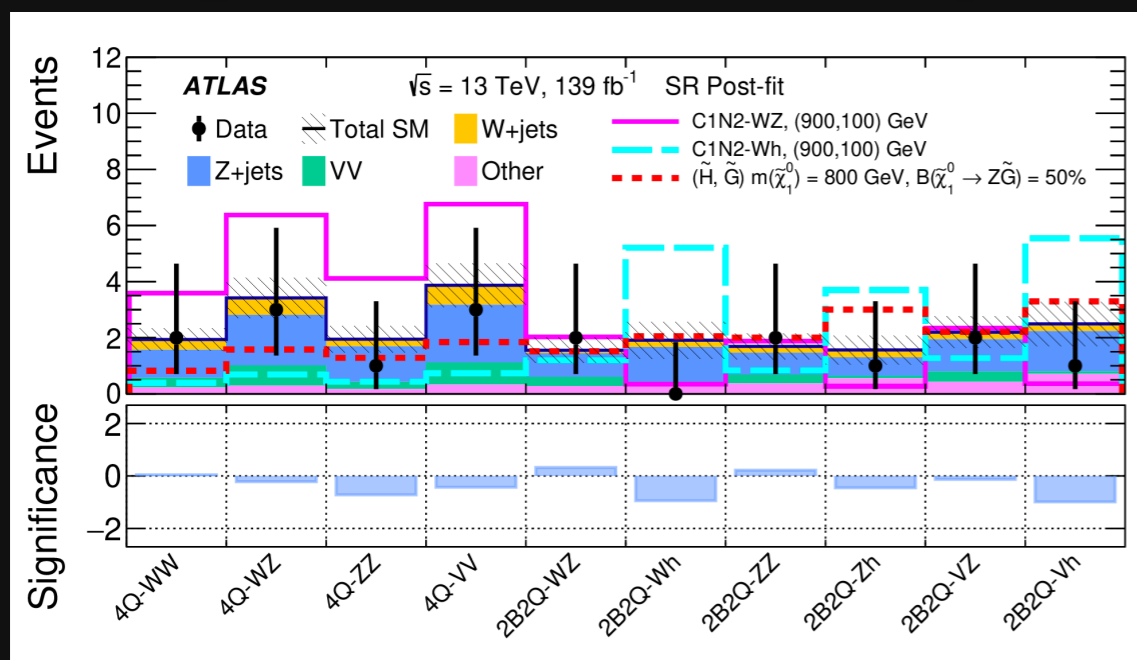
2 large-R jets with at most 1 b -jet to target the $qqqq$ final state (not $bbqq$)



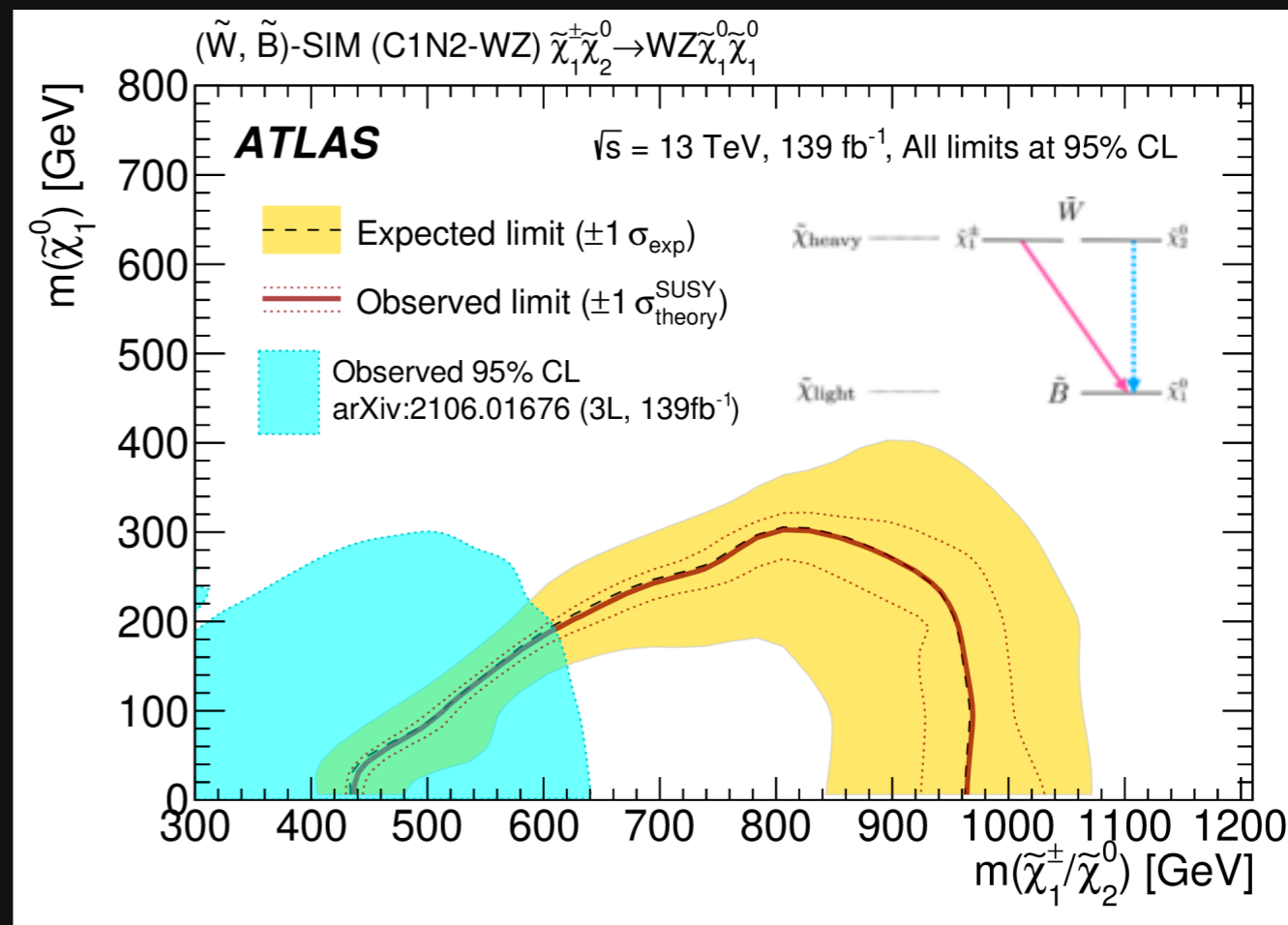
W/Z Hadronic

$$\tilde{\chi}_2^0 \tilde{\chi}_1^\pm \rightarrow WZ \tilde{\chi}_1^0 \tilde{\chi}_1^0$$

No significant excess found in any SR



Significance



$$(\tilde{\chi}_{\text{heavy}}, \tilde{\chi}_{\text{light}}) = (\tilde{W}, \tilde{B})$$

Improve sensitivity from $\Delta m \sim [300, 600] \rightarrow [450, 1000]$ GeV

(also did DM global fit, and MSSM reinterpretations)