# First Look at Pixel Endcap Cosmic Data

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### Outline

- Occupancy and noise studies.
- Clustering studies. effect of hot pixels.
- Tracking studies.
  - Tuning with data.
  - New geometry.
- Module efficiency.

### Introduction



# Occupancy (I)



Fraction of pixels readout per module.

Very low occupancy, but large variations among modules, even in events not synchronized by the cosmic trigger (noise) 4

### Hot Pixel Studies (I)



Fraction of events that a same pixel fires. If R>0.0001, pixel is masked as "hot" (1544 hot pixels)

# Hot Pixel Studies (II)



Selected pixels arise from the trigger BC, whereas hot pixels are uniformly distributed in Bcids.

### Hot Pixel Studies (III)





Distribution of number of hot pixels per module, per layer.

# Hot Pixel Topology



Very often, hot pixels are produced within "regions" involving several nearby pixels. But not all pixels in the hot-regions fire at the same time.

### **Problematic Modules**



Module 34-L2 shows a different behavior with respect to all other modules: Larger number of hot pixels, almost uniformly distributed among the module.

#### Hot Pixel Distributions



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#### Module Occupancy After Hot-Pixel Cleaning



Low signal for module 34-L2, at the level of noise.

# Clustering

Implemented cluster-finder algorithm, to be used after hot-pixel cleaning:

1<sup>st</sup> pass:

Select seed pixels with |Bcid-5|<=1 Attach all neighbors pixels if DeltaRow<=1, DeltaCol<=1.

2<sup>nd</sup> pass:

Merge pre-clusters sharing pixels.

In an event-by-event basis pixels are classified as "hot" or "good".

# Cluster Occupancy per Event



# Cluster ToT Distribution (I)



Pixel ToT from signal tracks should peak at ToT=30.

### Cluster ToT Distribution (II)



# Tracking Studies (I)

Implemented simple track-finder algorithm:

1<sup>st</sup> pass: select events with at least 1 cluster in each layer. Form all possible paths, and select the track with the smaller chi2/ndof.

2<sup>nd</sup> pass: exclude 1 hit at the time and evaluate the change in chi2/ndof of the fit.<sup>3</sup> Reject cluster if that improves the chi2 of the track.

Iterate until no more clusters can be removed or the number of attached clusters is 3.

Fitting is done with a 3D line parametrization.



-150

-20050

500

550

600

650

700 Z (mm)

# Tracking Studies (II)



Cluster errors "tuned" to 3.3mm, so that chi2 mean value is 1. Very low fraction of 4-hit tracks (5%, expected ~15-20%) Extremely large cluster error needed.

# Tracking Studies (Resolution)



Form track with 1 cluster in each layer, remove cluster in middle layer, re-fit track, and compute resolution as the difference between the cluster position at layer 1 and the track prediction. Very large tails -> geometry problem (see Su Dong's talk)

### Tracking with Fixed Geometry

Results using new geometry (phi flip between front/back modules in a same layer)

Resolutions are now close to 50um.





Event #16781070

Look for tracks with at least one doublet, and 1 cluster in a second layer. Form track, extrapolate to the 3<sup>rd</sup> layer, and compute an unbiased module cluster efficiency. Take into account module acceptance.



First look at data from cosmics run:

Very low noise (module occupancy  $\sim 10^{-7}$ )

Hot-pixels seems to be clustered in regions. Need further investigation.

Tracks found after geometry fix, and hot-pixel removal. ToT cluster distribution and number of overlapping clusters in tracks consistent with expectation.

Very first look at module efficiencies. Work in progress.