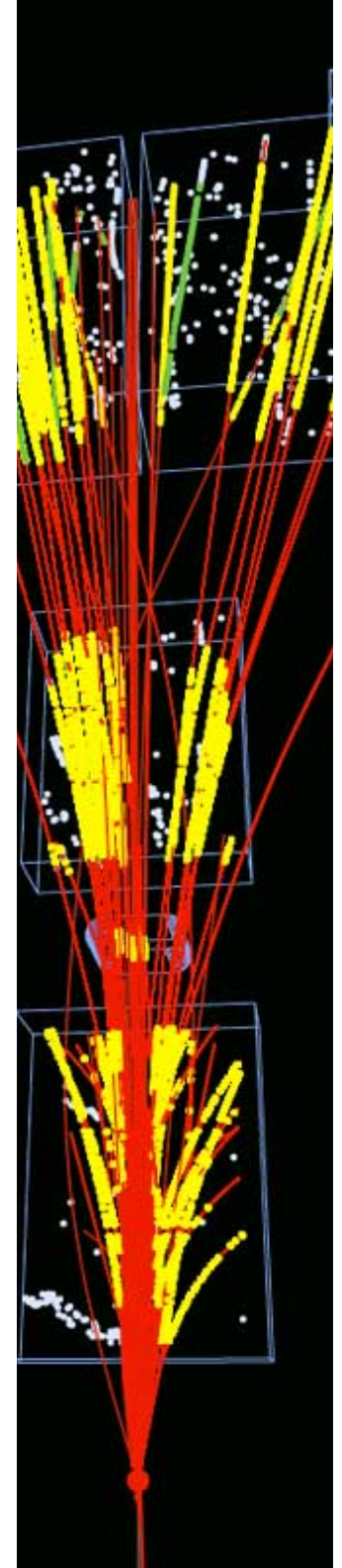


# Composition of spectator matter as a key to centrality determination

Aleksandr Svetlichnyi<sup>\*</sup>, Igor Pshenichnov, Roman Nepeyvoda

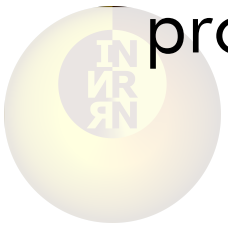
MIPT(NRU), INR RAS

<sup>\*</sup>) [aleksandr.svetlichnyy@phystech.edu](mailto:aleksandr.svetlichnyy@phystech.edu)

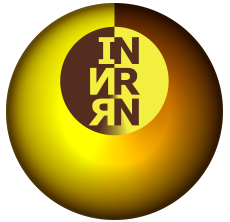


# Outline

- Modeling spectator formation: two ways
- Excitation of spectator matter: particle-hole model and alternatives
- Our model: Abrasion-Ablation Monte Carlo for Colliders (AAMCC)
- Physics of spectators: dependence on impact parameter
- Characterization of spectator matter: fragments and nucleons
- Centrality determination on the basis of various properties of spectator matter



# Modeling spectator formation

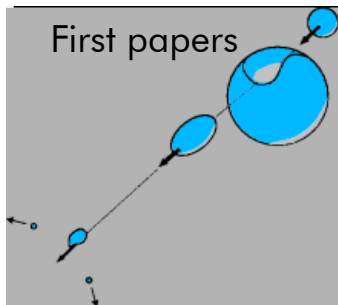


# Two ways to model spectator matter

## Participant-spectator picture

Abrasion-ablation models, cascade models (ABRABLA, DCM-SMM, LAQGSM-SMM, DPMJET-GEM etc.):

- Interacting (wounded) nucleons and spectator nucleons are distinguished. All the latter are assumed to be inside a nuclear residue (prefragment);
- A good prescription for calculating the excitation energy of the prefragment is necessary to model properly its further decay;
- A set of realistic prefragment decay models have to be involved.

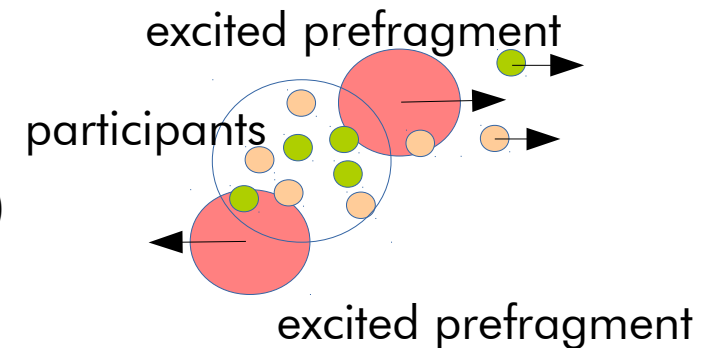


Abrasion-ablation model is still popular:

R. Thies et al. (R3B Collaboration)  
Phys. Rev. C **93** (2016) 054601

K. Mazurek et al., Phys. Rev. C **97** (2018)  
024604

and many other papers...



J. Gosset, H.H. Gutbrod,  
W.G. Meyer et al., PRC **16** (1977) 629

J. Hüfner, K. Schäfer, B. Schürmann,  
PRC **12** (1975) 1888

# Two ways to model spectator matter

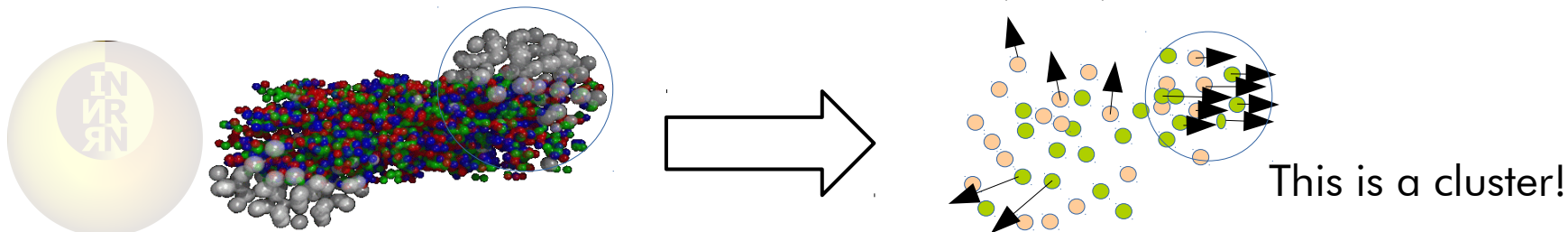
## Building fragments from individual nucleons:

Quantum molecular dynamics models (with their own assumptions): QMD, JQMD, NMD, UrQMD, PHQMD and others.

- No need to label explicitly participants and spectators and introduce prefragments.
- An algorithm to define a group of individual nucleons as a fragment (cluster) has to be developed (SACA, MST)<sup>1,2</sup>. More easy to build light fragments (aka coalescence).
- The time when QMD simulation is completed and fragments have to be defined is considered as a free parameter.
- A spontaneous nucleon emission from clusters (evolution of initial Fermi distribution to Boltzmann one) can not be avoided, but becomes less important at higher energies.

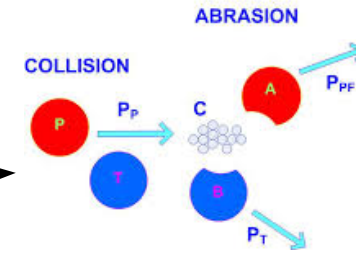
<sup>1</sup>) T. Ogawa, T. Sato, S. Hashimoto et al., PRC **98** (2018) 024611

<sup>2</sup>) J. Aichelin, E. Bratkovskaya, A. Le Fèvre et al. PRC **101** (2020) 044905

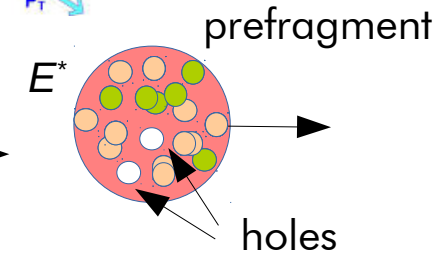


# Estimation of prefragment excitation energy: several methods

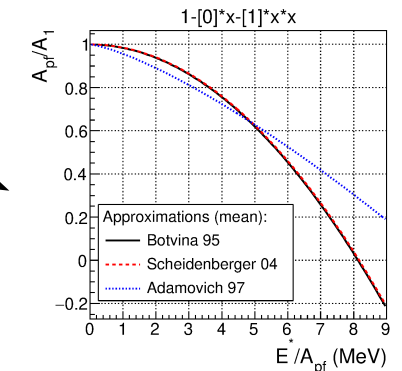
From prefragment geometry ("clean-cut"): excess of surface energy + empirical terms<sup>1,2)</sup>



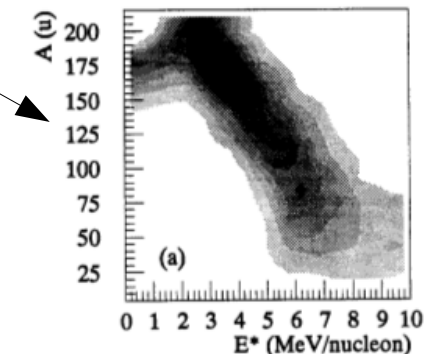
From particle-hole model: abraded nucleons create holes in nuclear cores of colliding nuclei<sup>3,4)</sup>



By inventing phenomenological correlations between prefragment excitation energy per nucleon and its mass<sup>5,6)</sup>



By extracting from measured events by finding the distribution which provides an optimum description of data. A recursive method has been used.<sup>7)</sup>



1) L.F. Oliveira, R. Donangelo, J.O. Rasmussen, PRC 19 (1979) 826

2) K. Mazurek et al., Phys. Rev. C 97 (2018) 024604

3) J.-J. Gaimard K.-H. Schmidt, NPA 531 (1991) 709

4) C. Scheidenberger, I.P. K. Sümmerer et al., PRC 70 (2004) 01492

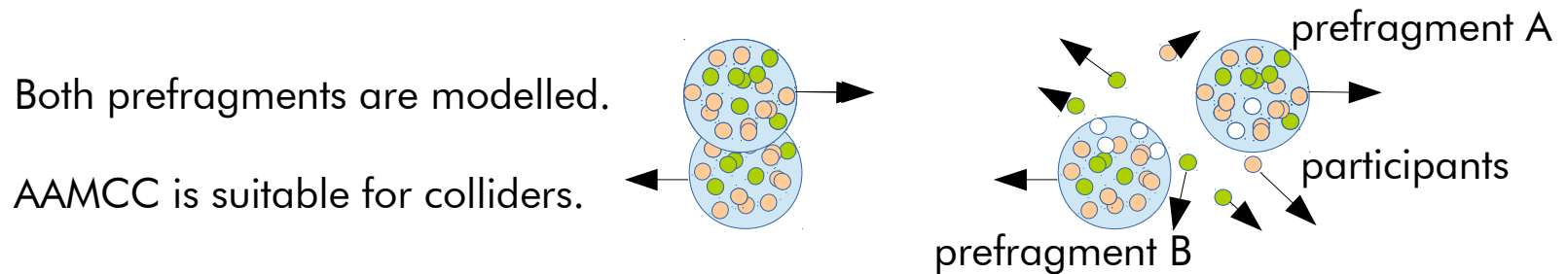
5) A.S. Botvina, I.N. Mishustin, M. Begemann-Blaich et al., NPA 584 (1995) 737

6) M.I. Adamovich, M.M. Aggarwal, Y.A. Alexandrov et al., Z. Phys. A 359 (1997) 277

7) P. Désesquelles et al., NPA 604 (1996) 183

# AAMCC

- Our model **Abrasion-Ablation Monte Carlo for Colliders (AAMCC)**<sup>1)</sup> written in C++ is based on the famous Glauber Monte Carlo v.3.0<sup>2)</sup> and models of decays of excited nuclei from Geant4 toolkit<sup>3)</sup> (G4Evaporation, G4SMM, G4FermiBreakUp).
- Glauber MC is de facto a standard tool adopted by all major experiments on relativistic AA collisions (ALICE, CMS, ATLAS, STAR, BRAHMS etc.).
- It is possible to take into account the difference between proton and neutron density through the use of GlauberMC v3.0.
- We tested and improved<sup>4)</sup> G4SMM ( $E^*/A_{pf} > 3$  MeV) and G4FermiBreakUp (the latter is for explosive decays of  $Z < 9, A < 19$  nuclei).



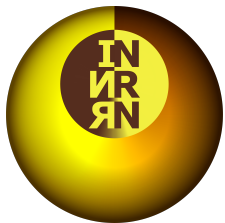
<sup>1)</sup> A. S., I.Pshenichnov. Bull. RAS: Phys. **84** (2020) 1103

<sup>2)</sup> C. Loizides, J.Kamin, D. d'Enterria, PRC **97** (2018) 054910

<sup>3)</sup> J.M. Quesada, V. Ivanchenko, A. Ivanchenko et al., Prog. Nucl. Sci. Tech. **2** (2011)

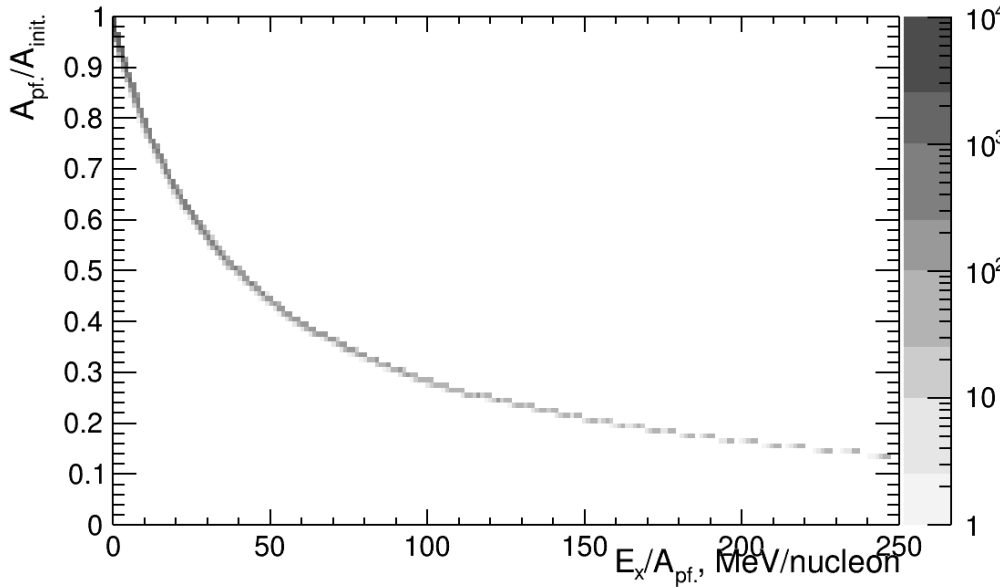
936

<sup>4)</sup> I.Pshenichnov., A.S. Botvina, I. Mishustin, W. Greiner, NIMB **268** (2010) 604



# AAMCC

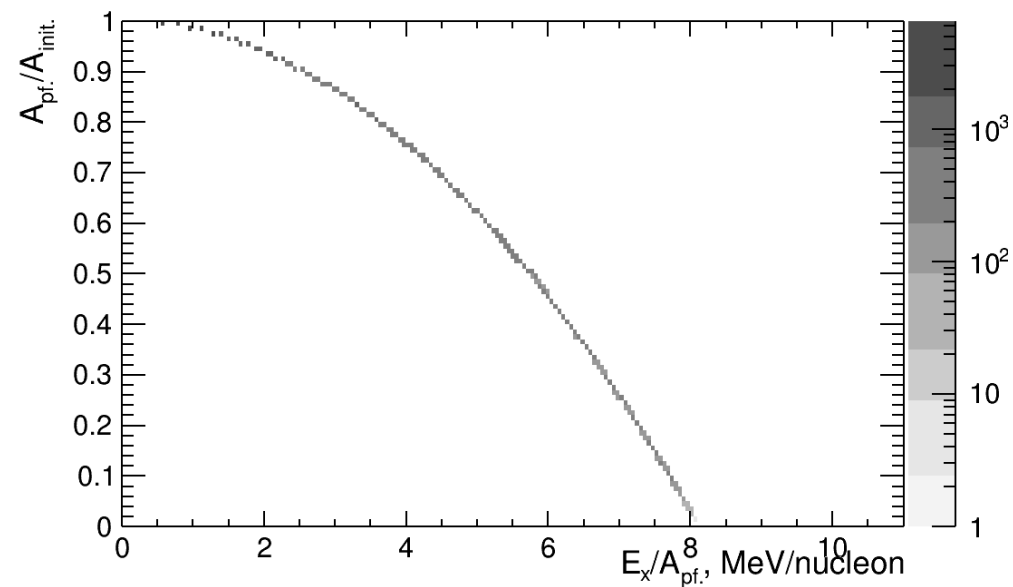
Two kinds of correlation between excit. energy and prefragment volume are implemented in AAMCC



$$\rho_e(E_x, a) = \frac{g_0^a}{a!(a-1)!} E_x^{a-1}$$

Level density in the particle-hole model:  
Ericson formula

C. Scheidenberger et al. – PRC 70, 014902 (2004)



$$1 - a/A = 1 - 0.015[E_x/(A - a)]^2$$

Empirical approximation by ALADIN  
collaboration

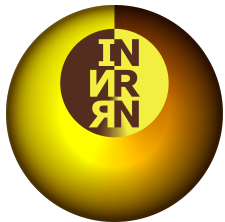
A. S. Botvina et al. – Nucl. Phys. A 584, 737 (1995)

$E_x$  – excitation energy

$a$  – number of removed nucleons

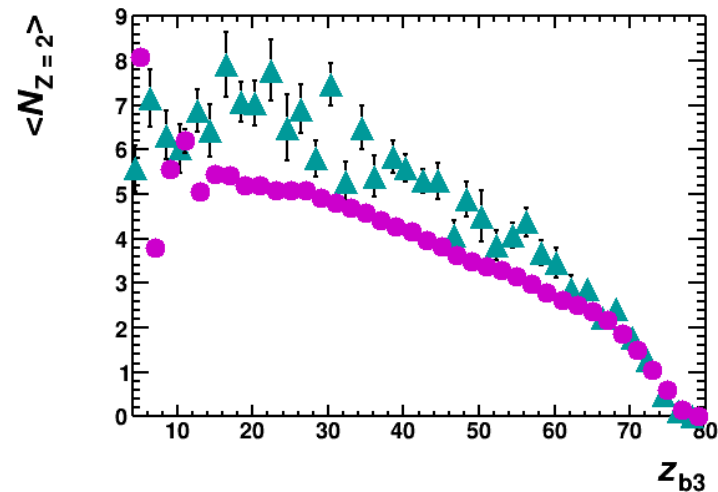
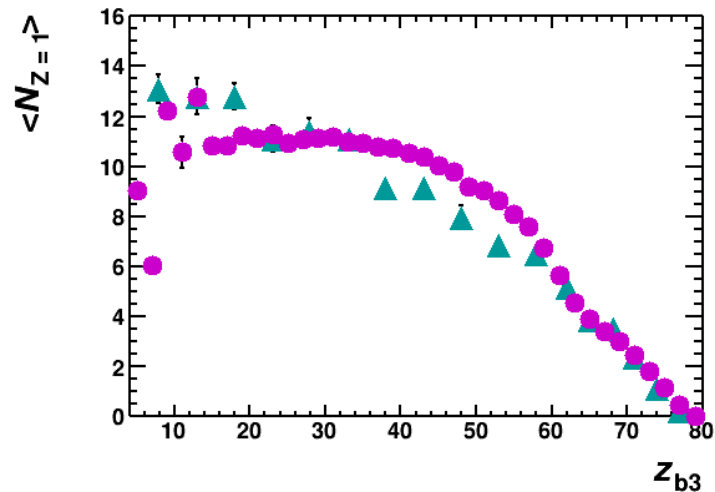
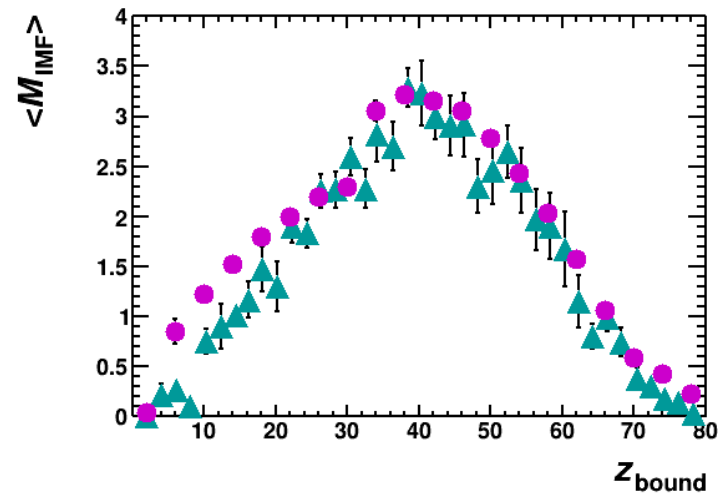
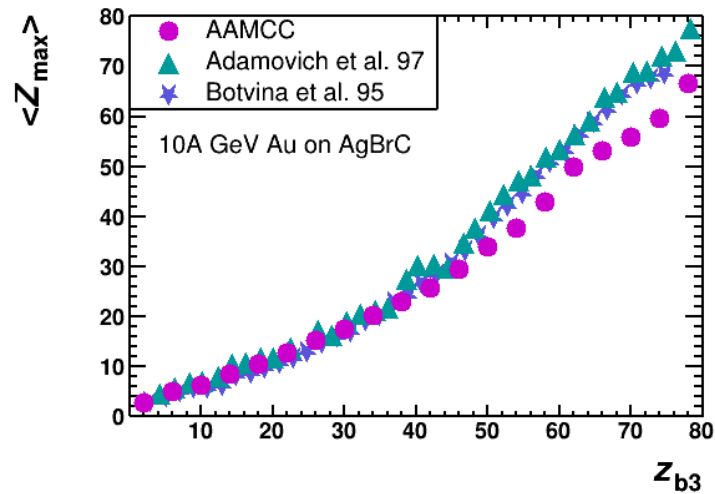
$A$  – mass number of the initial nucleus

$g_i$  – model parameters  $g_0 \approx 16 \text{ MeV}^{-1}$ ,  $g_1 \approx 0.7 \text{ MeV}^{-2}$





# AAMCC: comparison with experiment



EMU-01/12 collaboration – ZPA 359, 277 (1997)



Good description of these data in general  
with ALADIN approximation

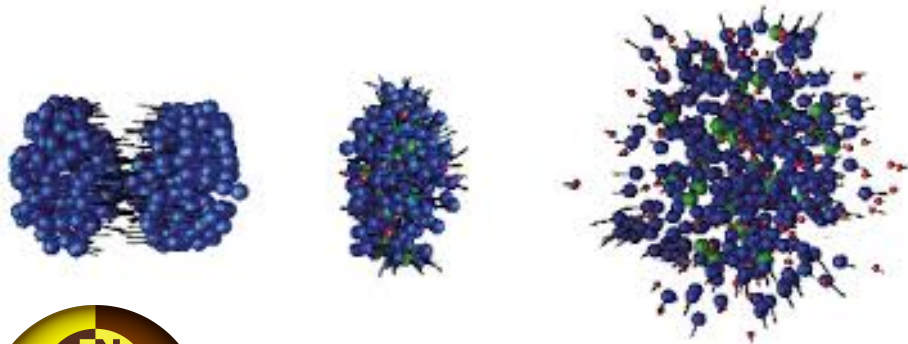
# Physics of spectators: dependence on impact parameter

- Central events: decay of a hot system
- Semi-central events: multifragment explosion
- Peripheral events: dominance of evaporation

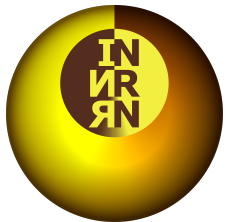


# Central collisions: decay of hot systems

- Small prefragment, very different from initial nucleus – particle-hole model is not applicable!
- Excitation energy is comparable to the total binding energy decaying system ( $\sim 8$  MeV/nucleon)
- Statistical decay of a hot system is calculated via Fermi Break Up model

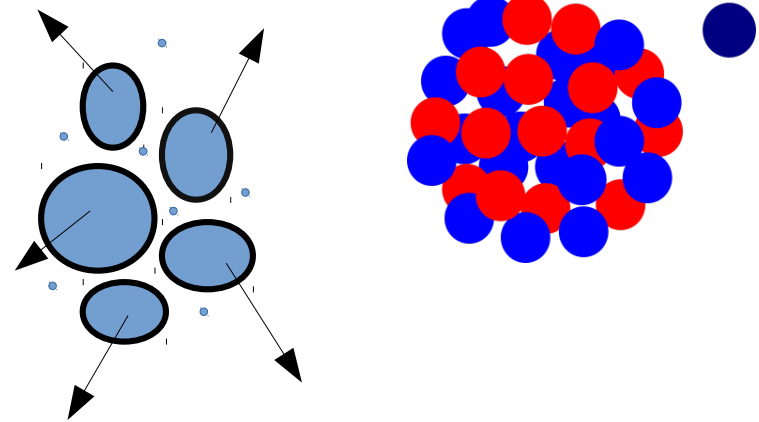
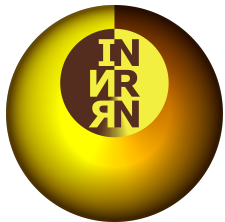


**Mostly nucleons, deuterons,  
tritons and alphas in the final state**



# Semi-peripheral collisions: two concurrent decay modes

- Prefragment volume is relatively large
- Number of participants is lower than in central collisions, but excitation energy still exceeds 3 MeV/nucleon.
- Nuclear multifragmentation comes into effect in this case. Fragments of intermediate mass ( $3 < Z < 30$ ) are multiply produced.
- Multifragmentation competes with evaporation at 3-4 MeV/nucleon.
- Decay is handled by two competing models: nuclear multifragmentation and evaporation.



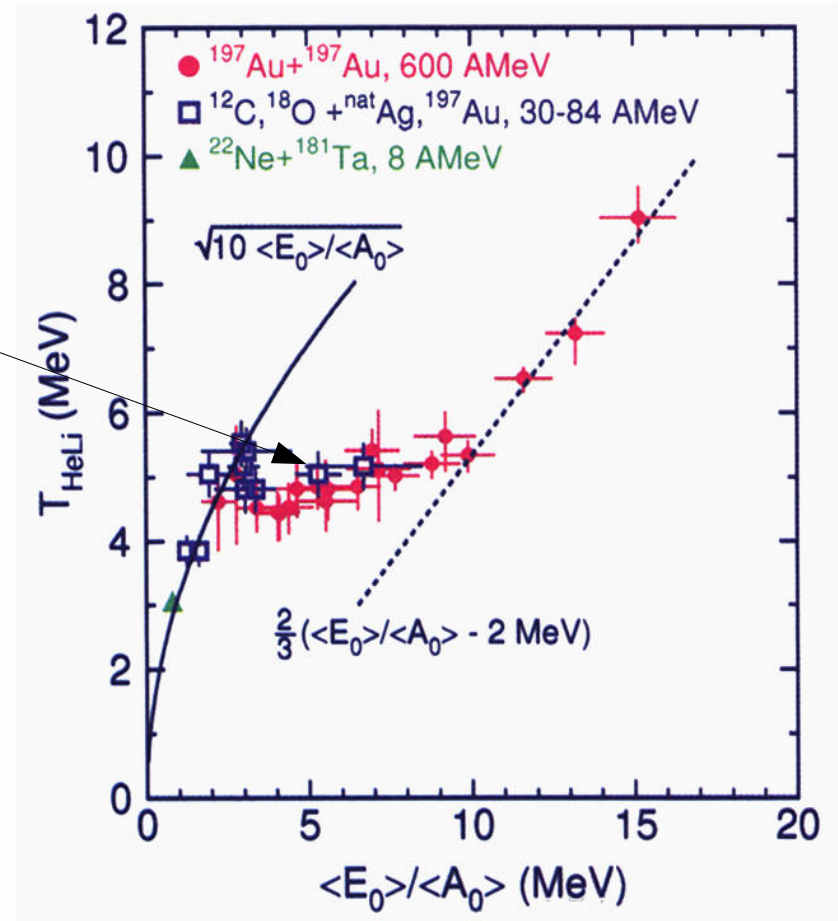
# Semi-peripheral collisions: multifragment explosion

Excitation energy 3-8 MeV/nucleon  
corresponds to nuclear matter  
liquid-gas phase transition.

Coexistence of liquid and  
gas phases of nuclear matter.

Studied in details by J. Bondorf, I. Mishustin,  
A. Botvina, W. Trautman and other groups.  
in 80-90s,

- J.P. Bondorf, R. Donangelo, I.N. Mishustin, et al.,  
Nucl. Phys. A443 (1985) 321; A444 (1985) 460;
- W. Trautman et al., Nucl. Phys. A538 (1992), 473;
- J.P. Bondorf, A.S. Botvina, A.S. Iljinov, I.N. Mishustin,  
K. Sneppen, Phys. Rep. 257 (1995) 133;
- A.S. Botvina et al., Nucl. Phys. A 584 (1995) 737;
- J.P. Bondorf et al., Phys. Rep. 257 (1995) 133;



J. Pochodzalla, T. Mohlenkamp, T.  
Rubehn, et al, Phys. Rev. Let. 75  
(1995) 1040

# Semi-peripheral collisions: modeling multifragment explosion

To describe multifragmentation a model named Statistical Model of Multifragmentation (SMM) was designed by Botvina et al. Its main assumptions are:

- A decaying system is in thermal equilibrium. It is characterized by:
  - neutron and proton numbers;  $N_0, Z_0$
  - excitation energy;  $E^*$
  - decay volume.  $V(M_f) = (1 + \kappa(M_f)) \cdot V_0$
- Microcanonical description for small systems ( $A < 20$ ), macrocanonical description for heavy.
- Decay volume is increasing with the multiplicity approximately following isobarometric line.

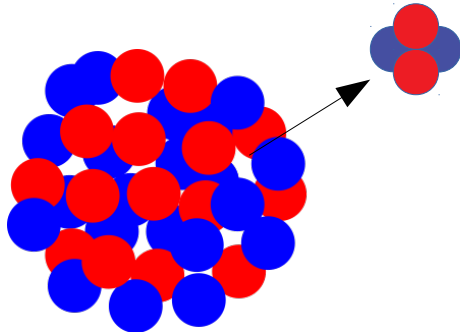


J.P. Bondorf, R. Donangelo, I.N. Mishustin, et al., Nucl. Phys. A443 (1985) 321; A444 (1985) 460;  
J.P. Bondorf, A.S. Botvina, A.S. Iljinov, I.N. Mishustin, K. Sneppen, Phys. Rep. 257 (1995) 133

# Semi-peripheral collisions: evaporation at high excitation

- Excitation energy is comparable to Coulomb barrier.
- Sequential evaporation of protons and neutrons as a result of high excitation
- Evaporation of alpha-particles as well as deuterons and tritons are available decay channel

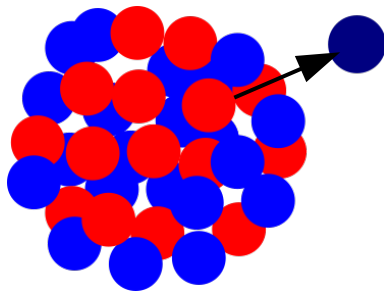
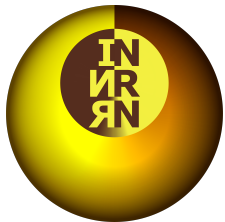
There are some alphas, deuterons, tritons in addition to multiple neutrons and protons. Evaporation process leaves single heavy residual nucleus.



# Peripheral collisions: dominance of evaporation

- Prefragment volume in peripheral collision are relatively big.
- Due to low numbers of participants, excitation energy of prefragment is also low ( $<3$  MeV/nucleon).
- Because of Coloumb barrier mostly neutrons are evaporated from heavy nuclei.

**Spectator matter composition is represented  
by a heavy fragment and many nucleons,  
mostly neutrons**





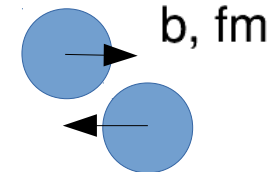
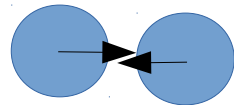
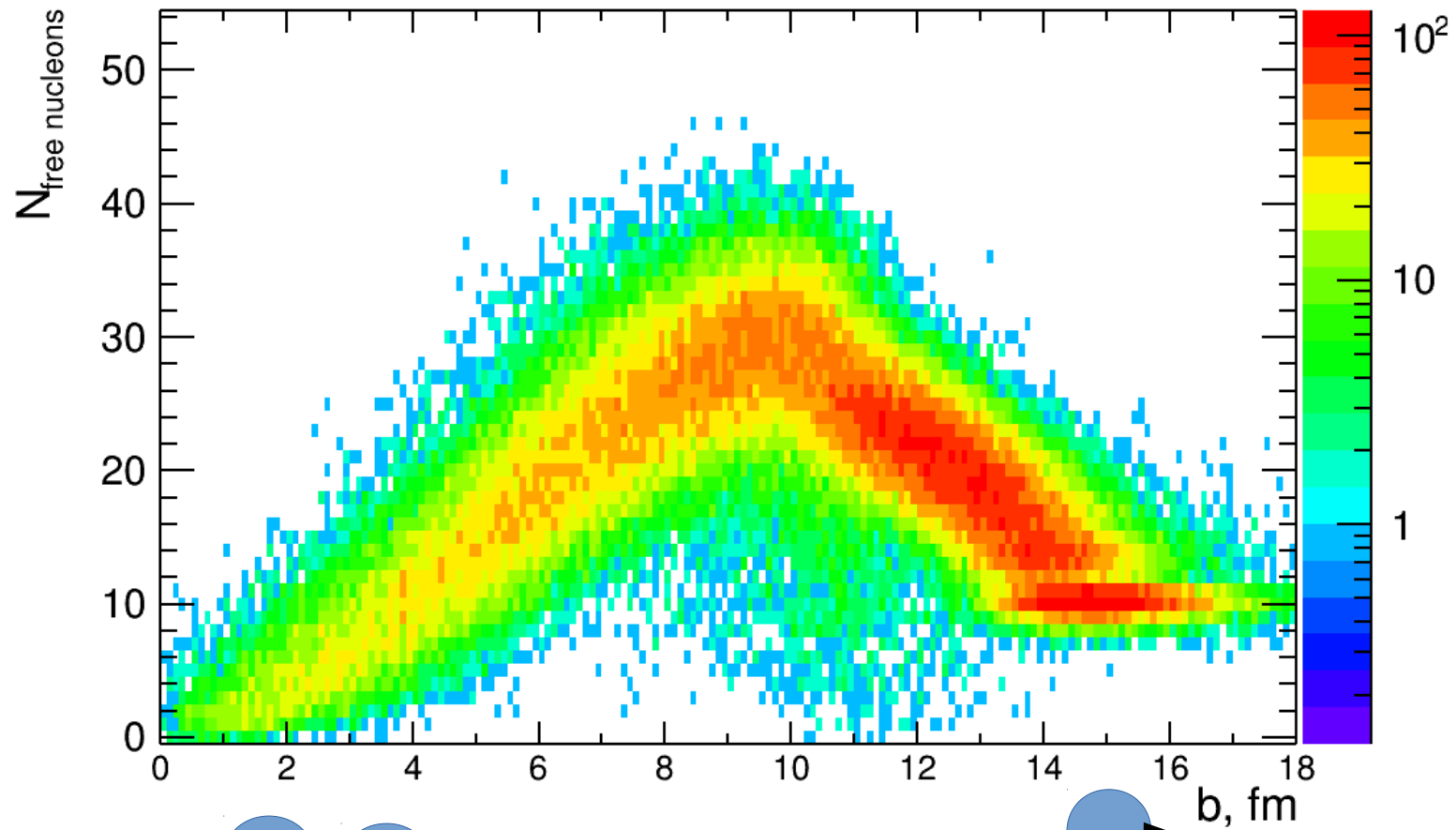
# Composition of spectator matter

- Unbound nucleons
  - Fraction of free nucleons in spectator matter
  - Absolute number of free nucleons
  - n/p-ratio for free nucleons
- Spectator fragments
  - Yields of light fragments
  - Number of charged fragments per spectator nucleon
  - Sum of squares of fragment charge per spectator nucleon

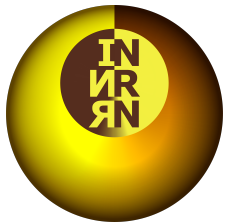


# Free nucleons vs. impact parameter

158A GeV, PbPb

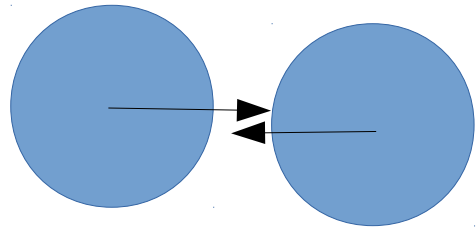


Not all nucleons are released.  
Its number is lower than 50.

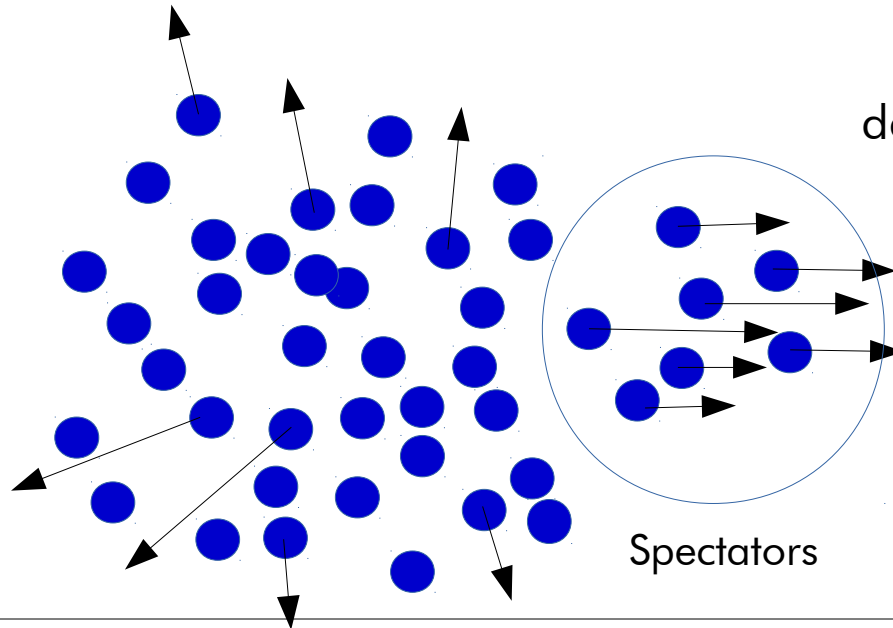


# Free nucleons: rise and fall

Central collision



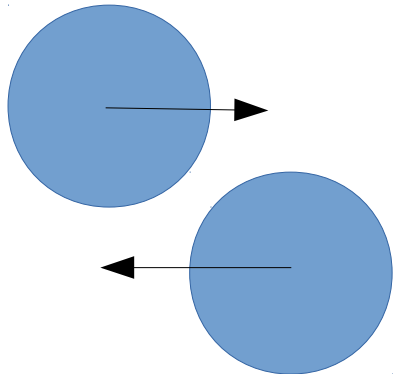
Participants



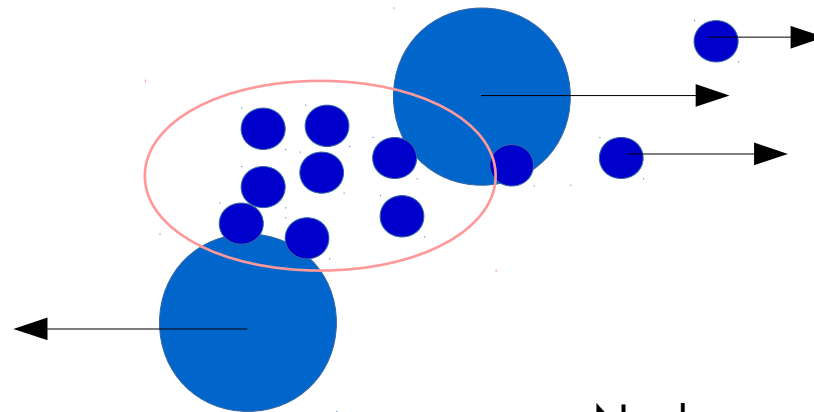
Free nucleons are formed in decays of **highly excited system**.

Spectators

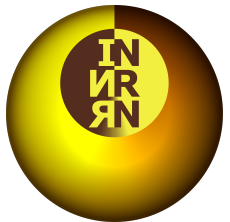
Peripheral collision



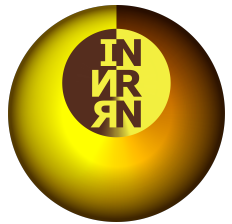
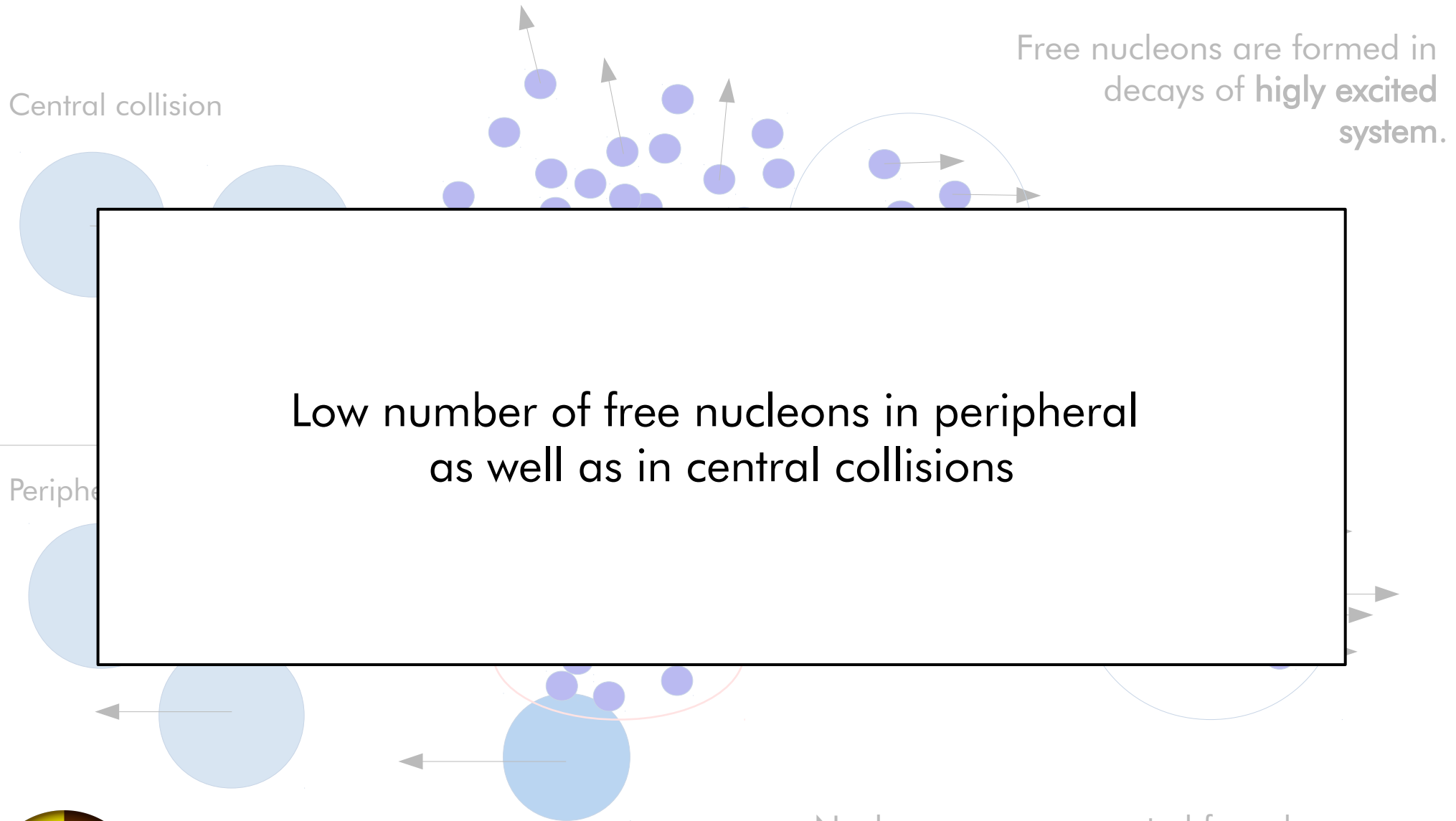
Spectator prefragment



Nucleons are evaporated from heavy residue with low excitation.

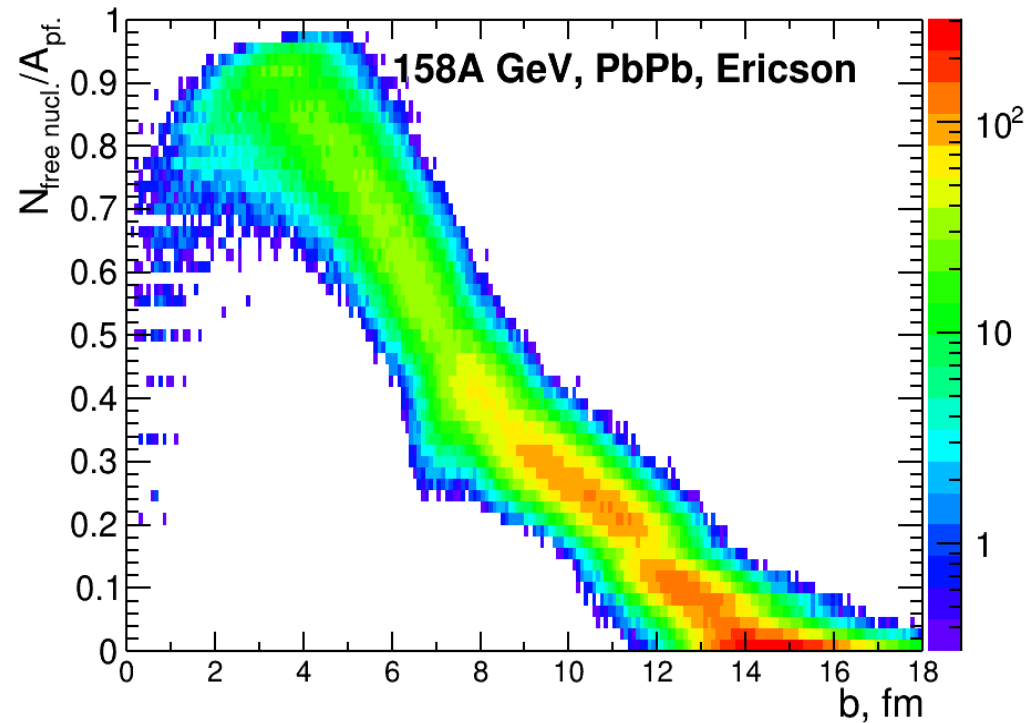
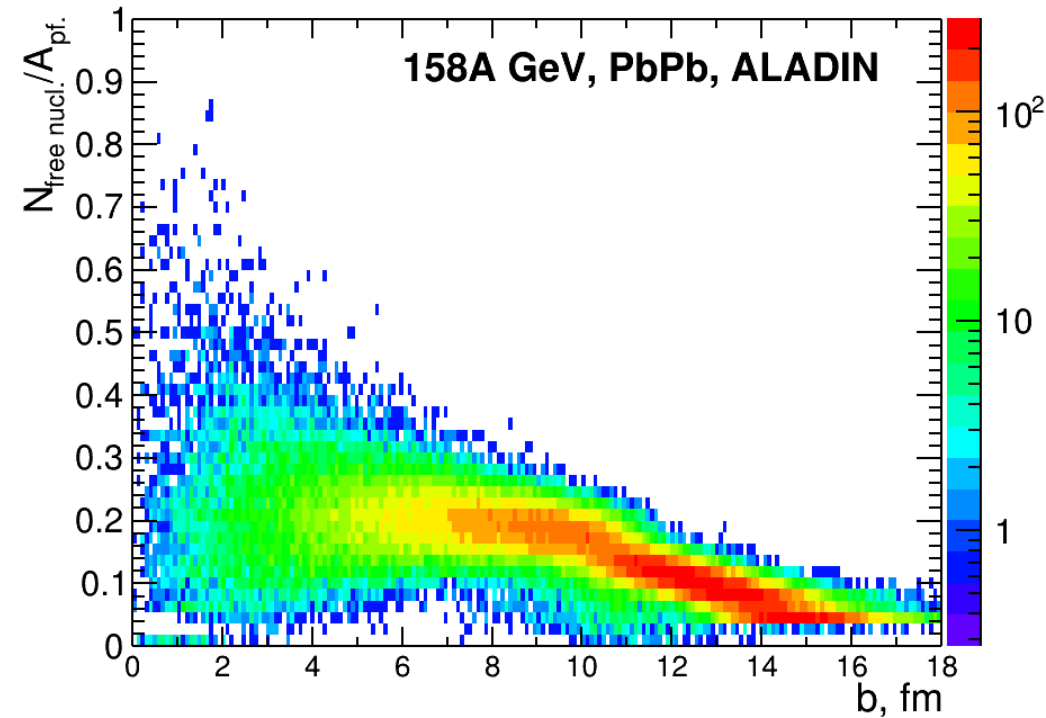


# Free nucleons: rise and fall

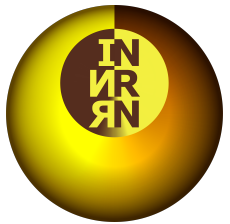


# Fraction of free nucleons vs. impact parameter

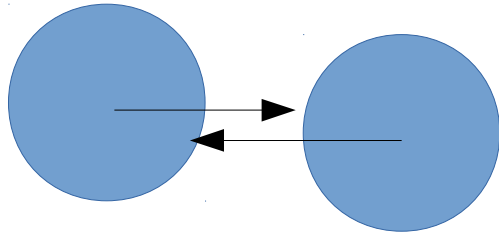
Fraction of free nucleons decreases with impact parameter



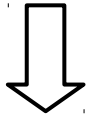
Even in the most central events alphas and deuterons are formed in addition to free nucleons.



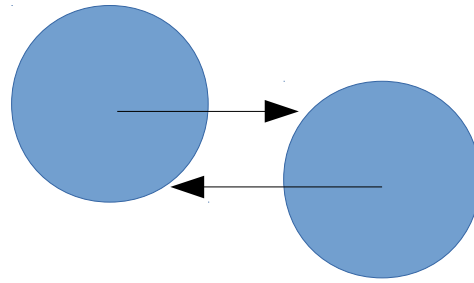
# Free nucleons fraction vs. impact parameter



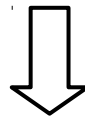
Light prefragment,  
very high excitation.



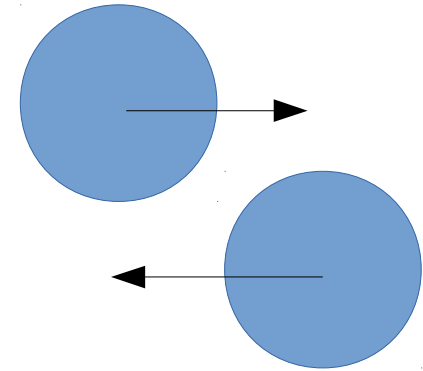
Multifragment  
explosion with  
a few intermediate  
and light mass  
fragments.  
Most of nucleons  
are free.



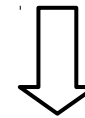
Medium-weight prefragment,  
higher excitation.



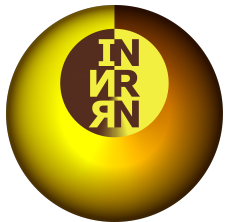
Evaporating a lot of  
nucleons, deuterons  
and  $\alpha$ -particles  
living massive residue.  
Approximately half of  
nucleons remain bound.



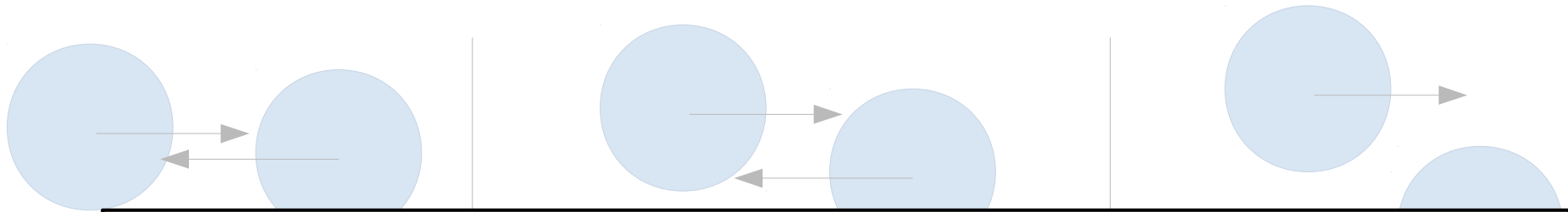
Heavy prefragment,  
low excitation.



Evaporating  
neutrons, leaving  
a heavy residue.  
Most of nucleons  
remain bound.



# Free nucleons fraction vs. impact parameter



Lig  
very

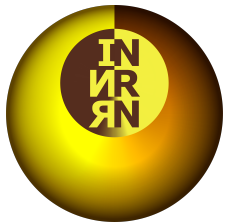
Monotonic decrease of free nucleon fraction with impact parameter.

M  
e  
a fe

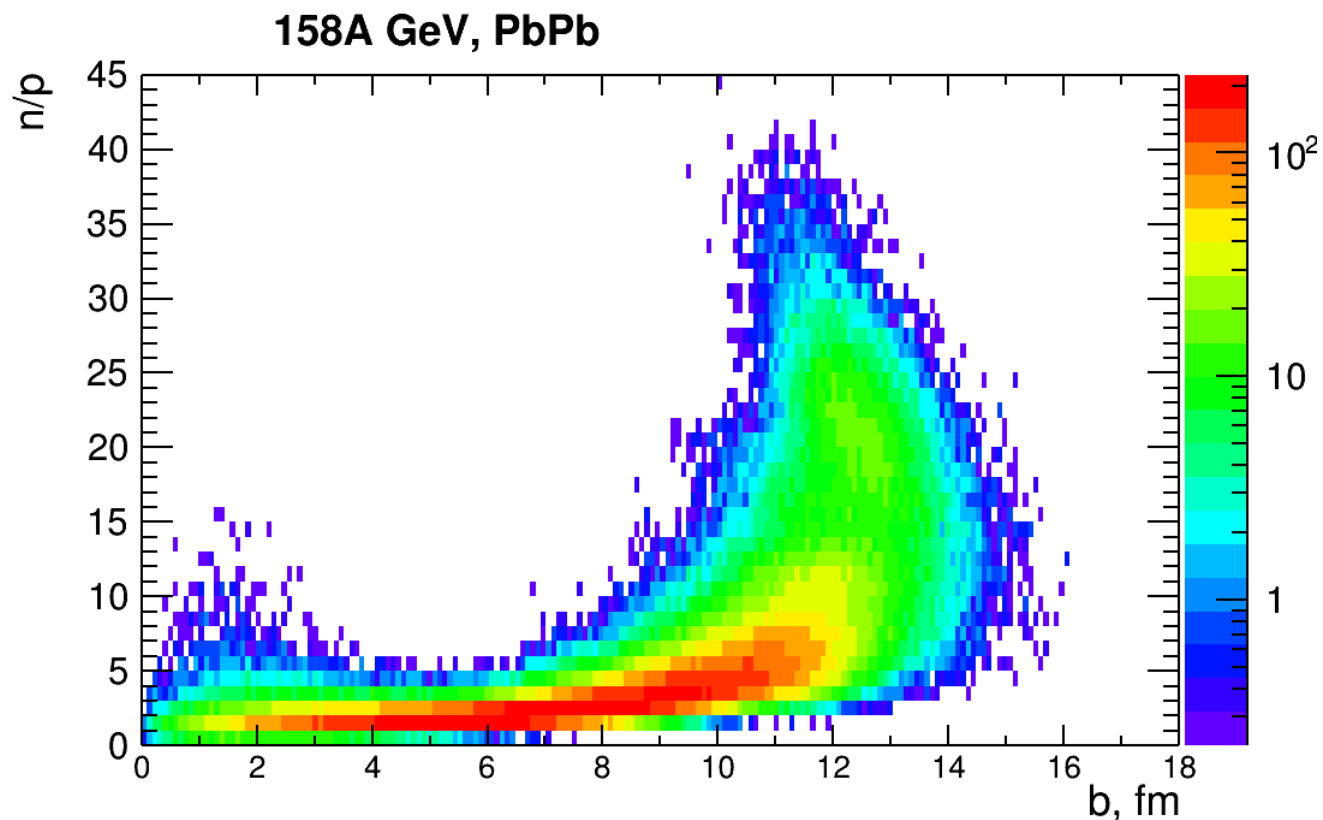
and light mass fragments.  
Most of nucleons are **free**.

living **massive residue**.  
Approximately **half** of nucleons remain **bound**.

Most of nucleons remain **bound**.



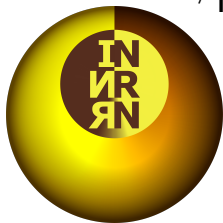
# n/p-ratio for free nucleons



b, fm	$2.1 \pm 0.9$	$3.7 \pm 0.9$	$5.4 \pm 0.7$	$6.9 \pm 0.7$	$8.4 \pm 0.7$
Experiment <sup>*)</sup>	$1.3 \pm 0.3$	$1.4 \pm 0.3$	$1.7 \pm 0.2$	$1.9 \pm 0.3$	$1.9 \pm 0.3$
AAMCC, Ericson	$1.9 \pm 0.8$	$1.8 \pm 0.6$	$1.9 \pm 0.5$	$2.3 \pm 0.7$	$3.3 \pm 0.8$

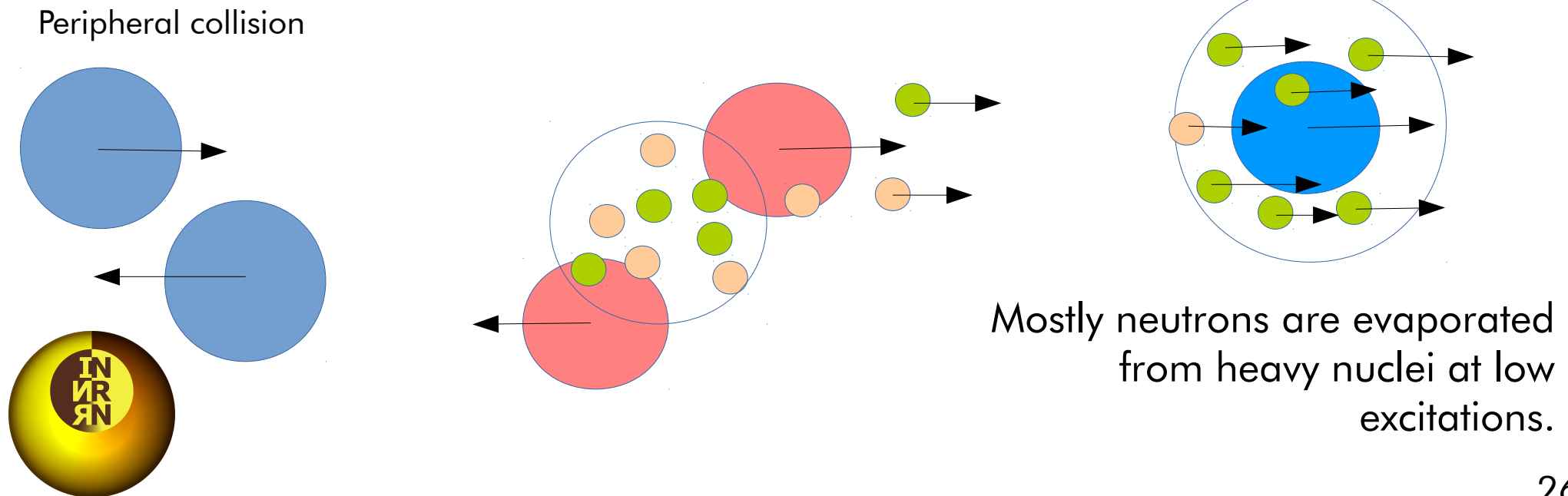
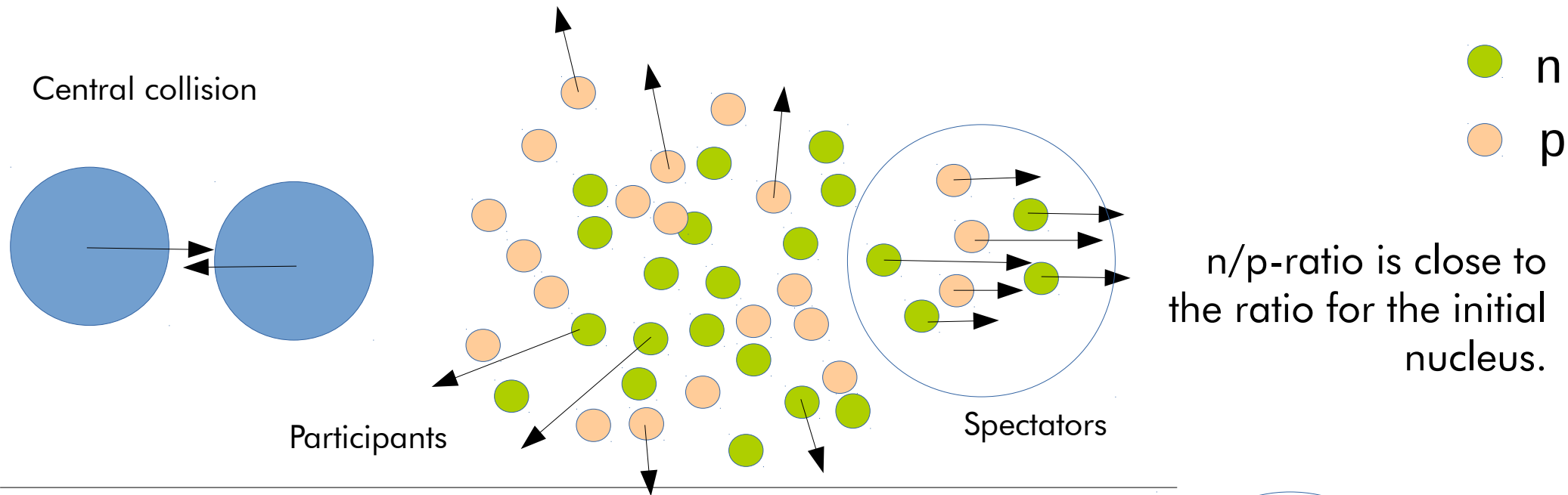
<sup>\*)</sup> H. Appelshauser, J. Bachler, S.J. Bailey et al., EPJ A 2, 383 (1998)

**Below 10 fm n/p-ratio is close to one of the primary nucleus.  
A dramatic increase above 10 fm is observed.**

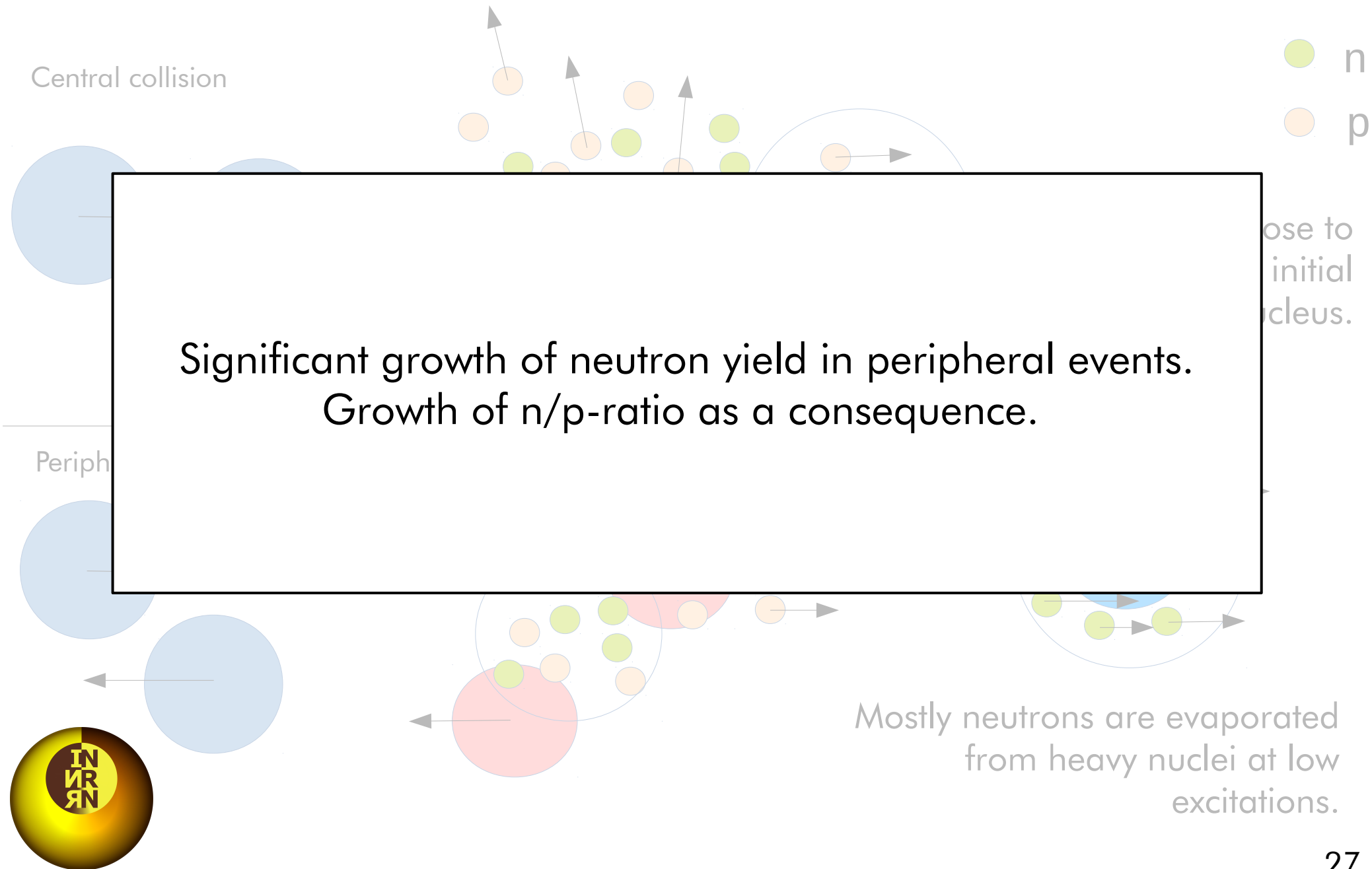




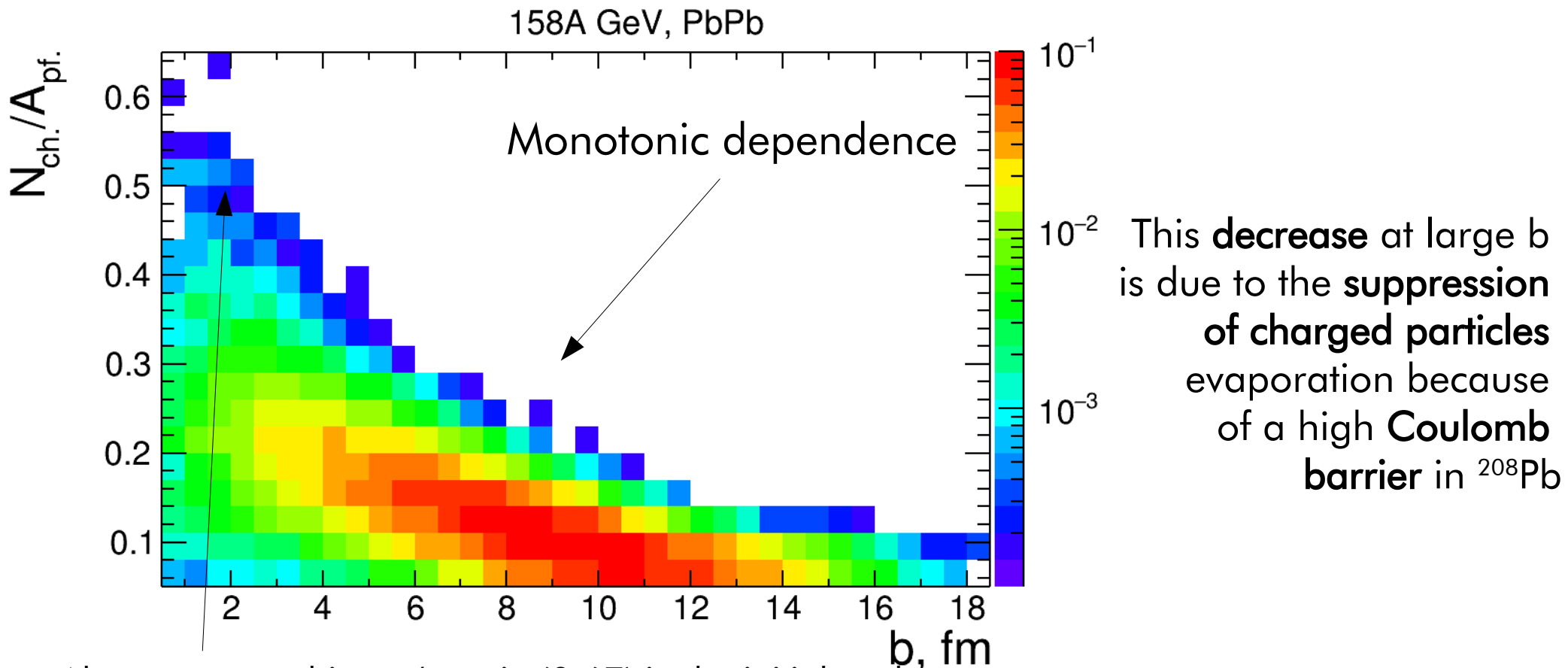
# n/p-ratio for free nucleons



# n/p-ratio for free nucleons



# Number of charged fragments per nucleon of spectator matter



This **decrease** at large  $b$  is due to the **suppression of charged particles** evaporation because of a high **Coulomb barrier** in  $^{208}\text{Pb}$

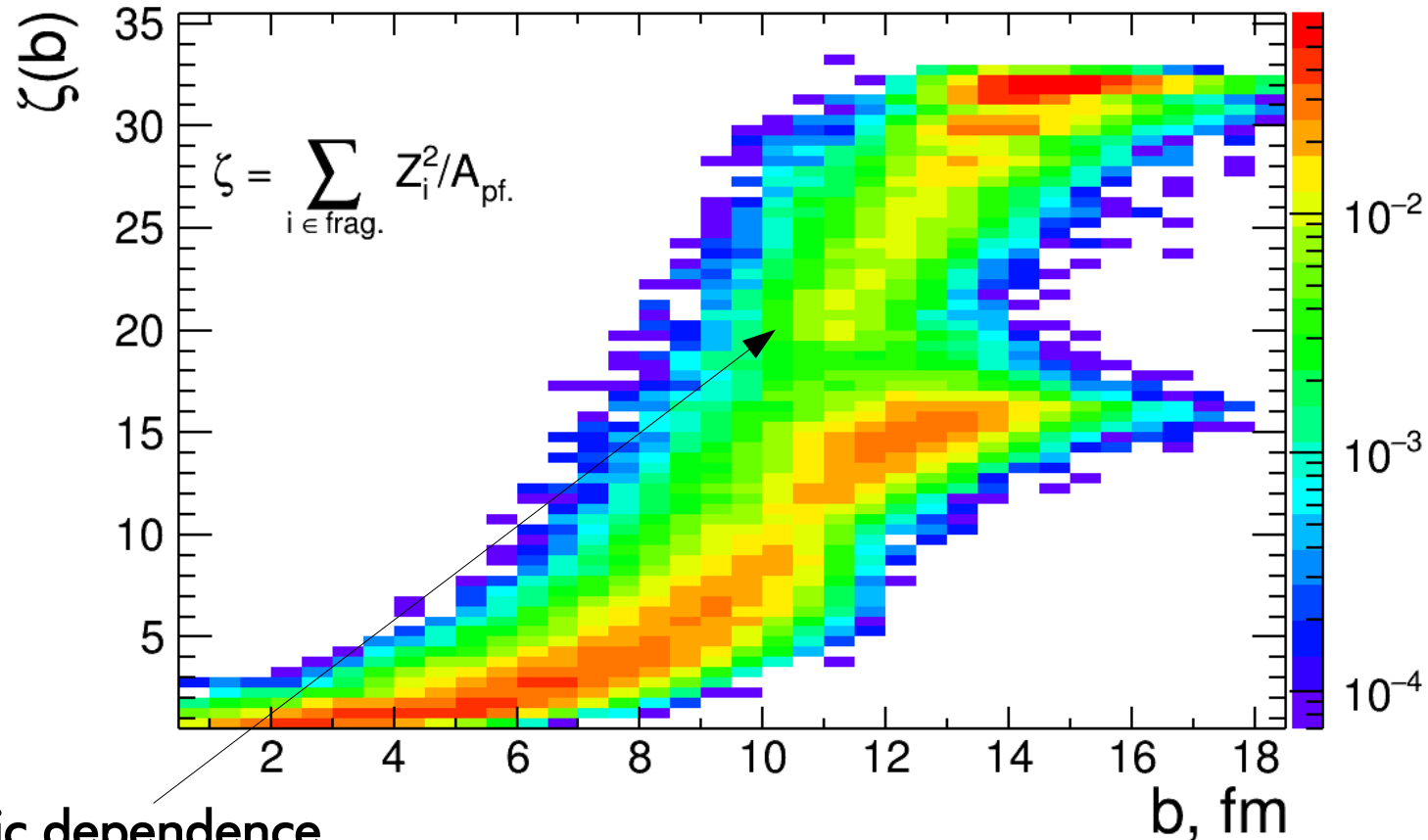
Almost approaching p/n-ratio (0.67) in the initial nucleus because almost all n and p are free in central collisions



If a charge-sensitive forward detector is available in addition to ZDC, this ratio can be also considered for centrality determination

# Sum of squares of fragment charges per spectator nucleon

158 GeV, PbPb



Monotonic dependence

Can be linked to the relation between signals of forward scintillating detector and forward calorimeter

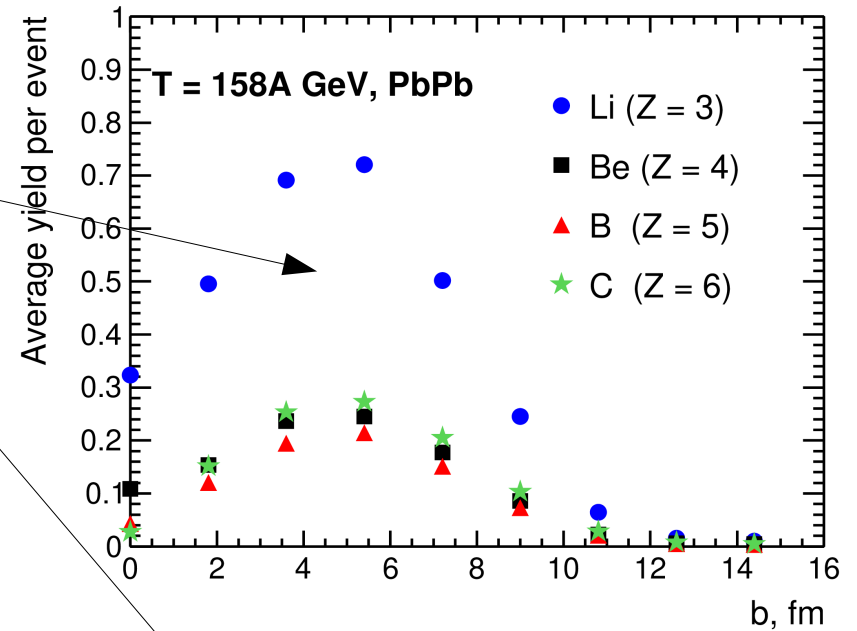
If a charge-sensitive forward detector is available in addition to ZDC, this ratio can be also considered for centrality determination



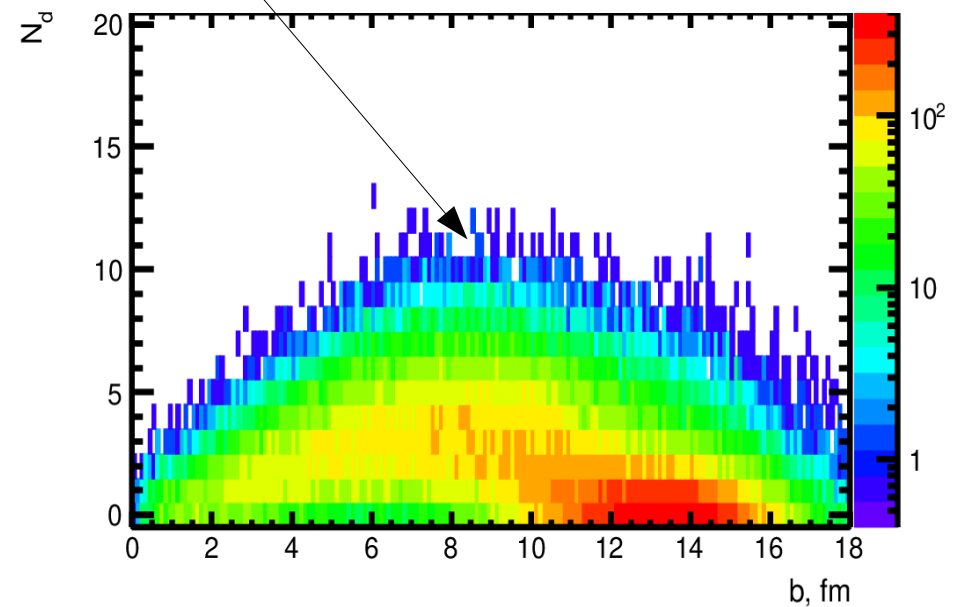
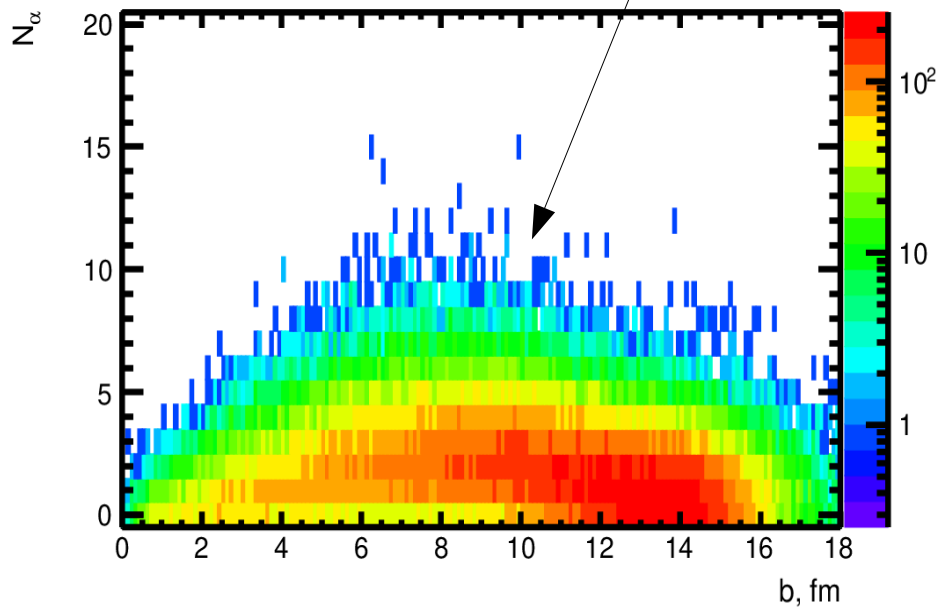
# Light fragment yield vs. impact parameter

Maximum in semi-central collisions

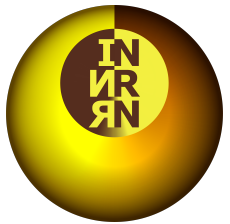
Quite high multiplicities of deuterons and alphas in semi-central collisions.



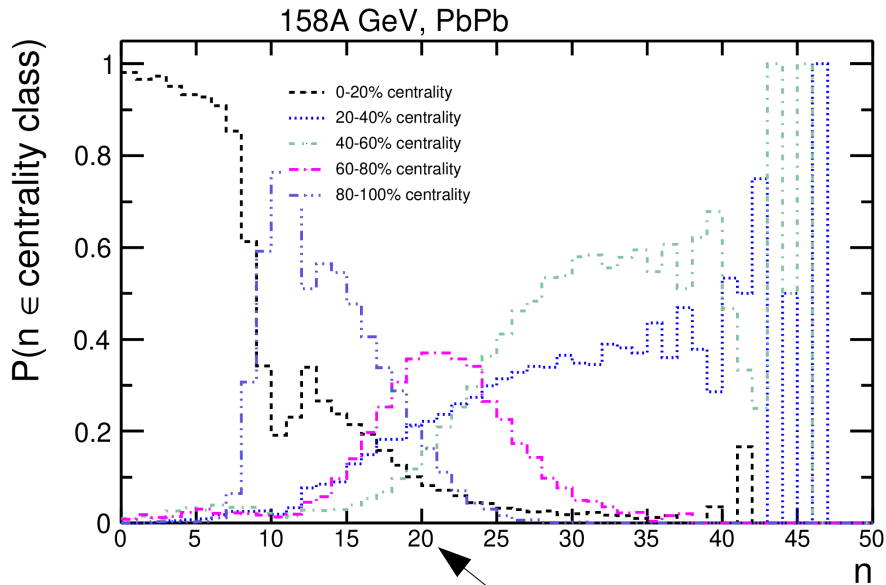
158A GeV, PbPb



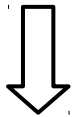
# How spectators can help in centrality determination



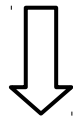
# Centrality 20-40%



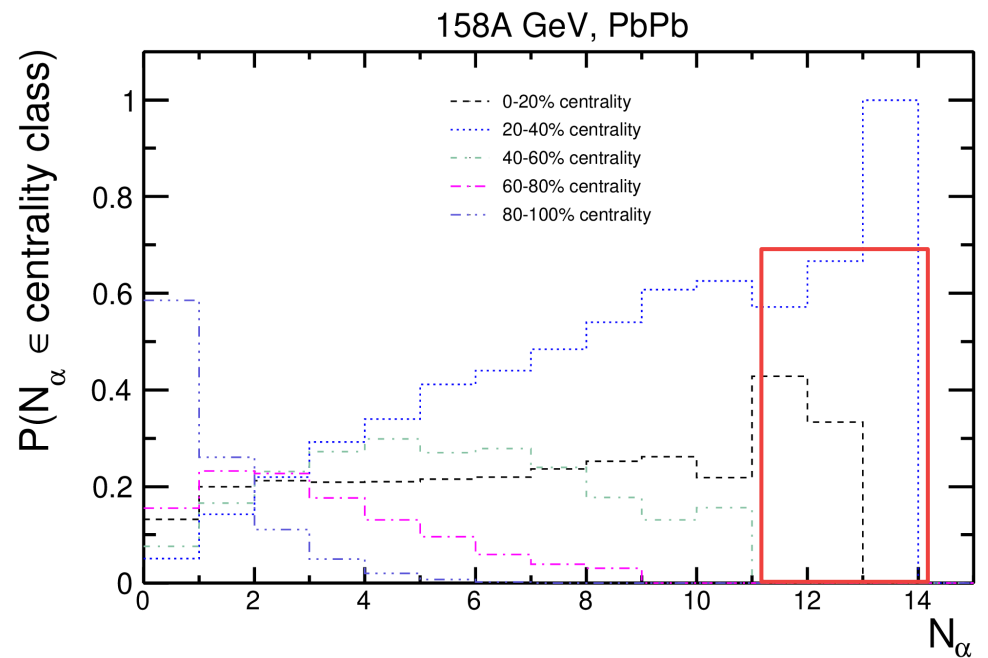
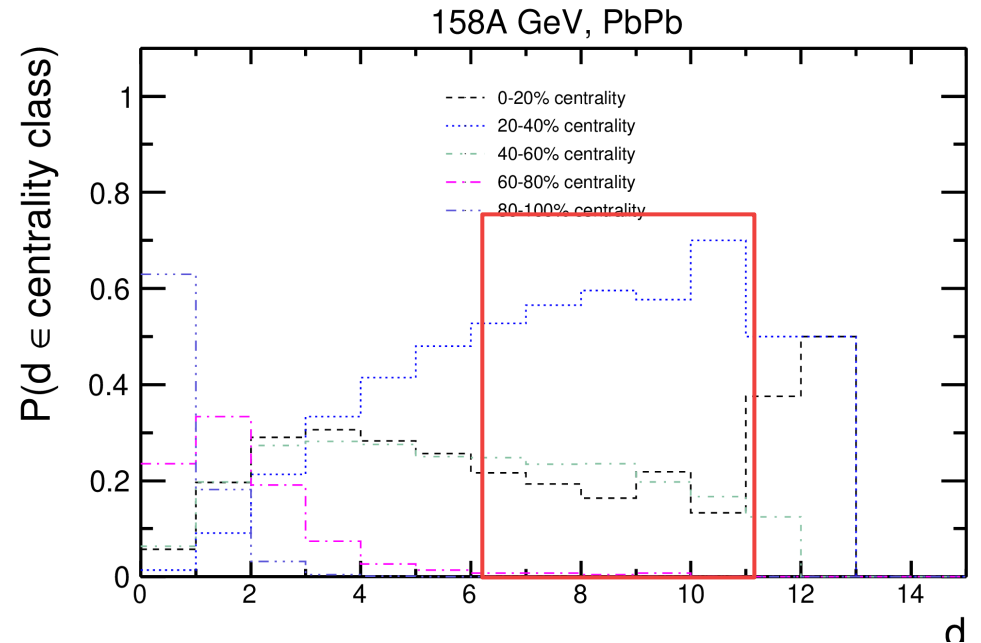
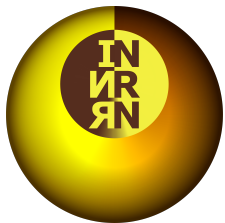
Let's consider event with  $\sim 20$  neutrons



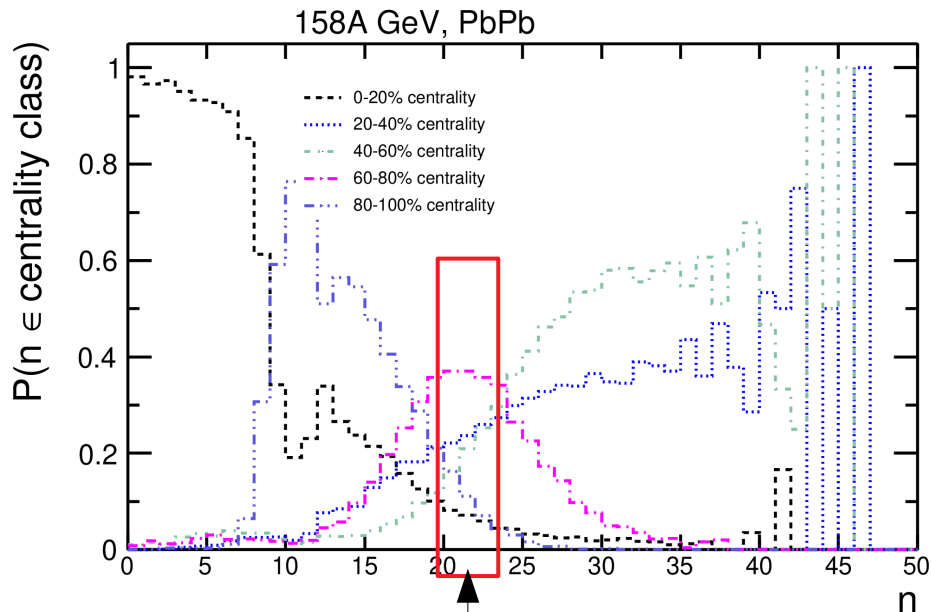
If number of deuterons is  $> 6$ ,  
and number of alphas is  $\sim 13$ .



Centrality 20-40%

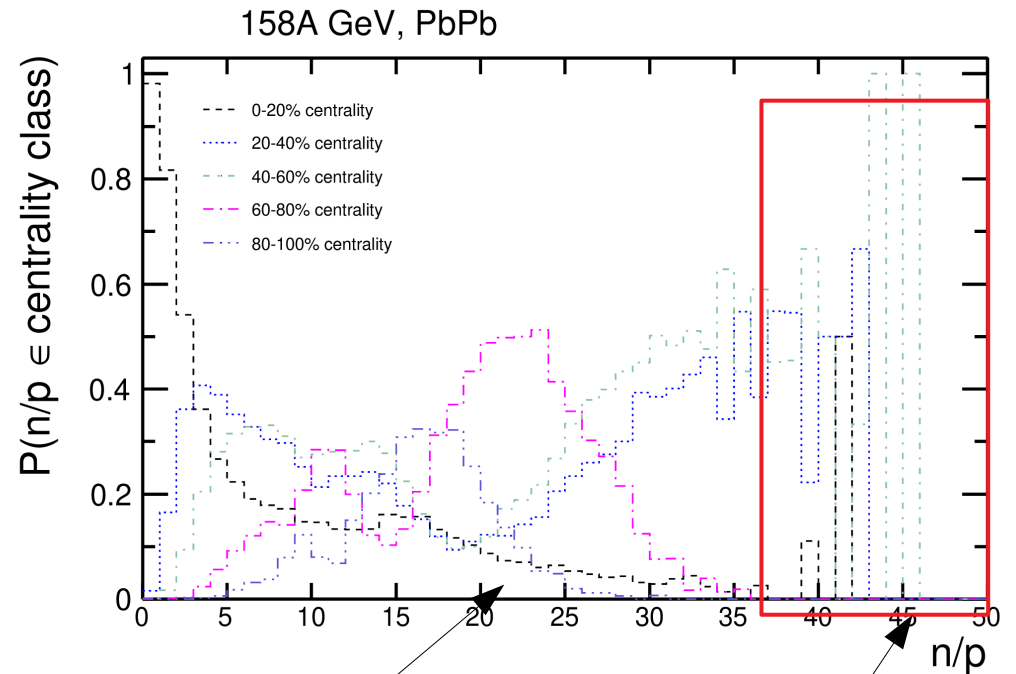


# Centrality 40-60% and 60-80%



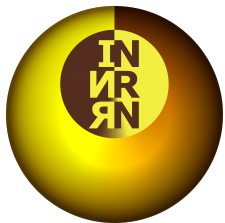
Let's consider event with  $\sim 23$  neutrons

With equal probability of  $\sim 0.3$  this event can be in each of three centrality classes: 0-20%, 40-60% or 60-80%



In the case of  $n/p > 35$  it is very probable that the event is in 40-60% centrality bin.

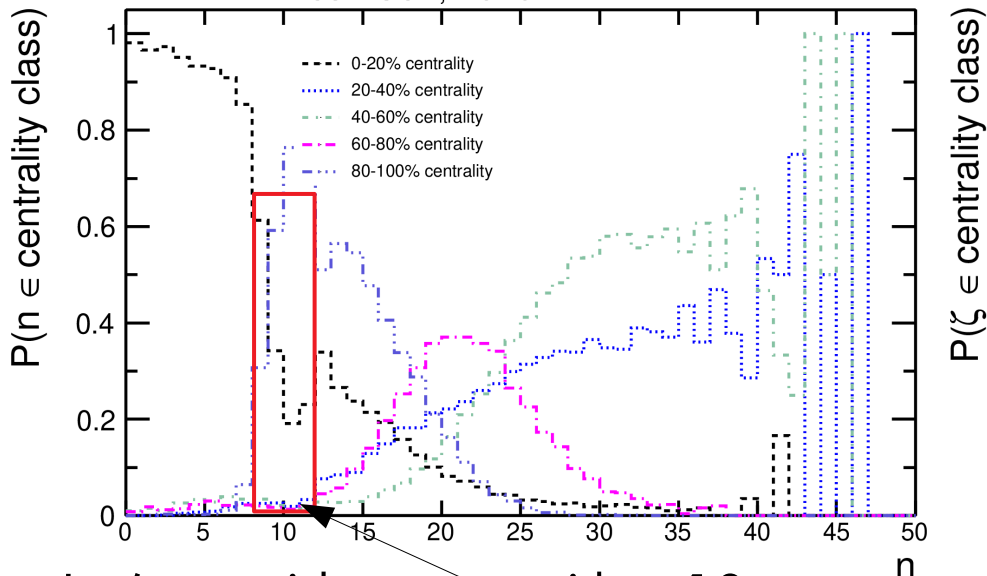
In the opposite case it is in 60-80% bin.





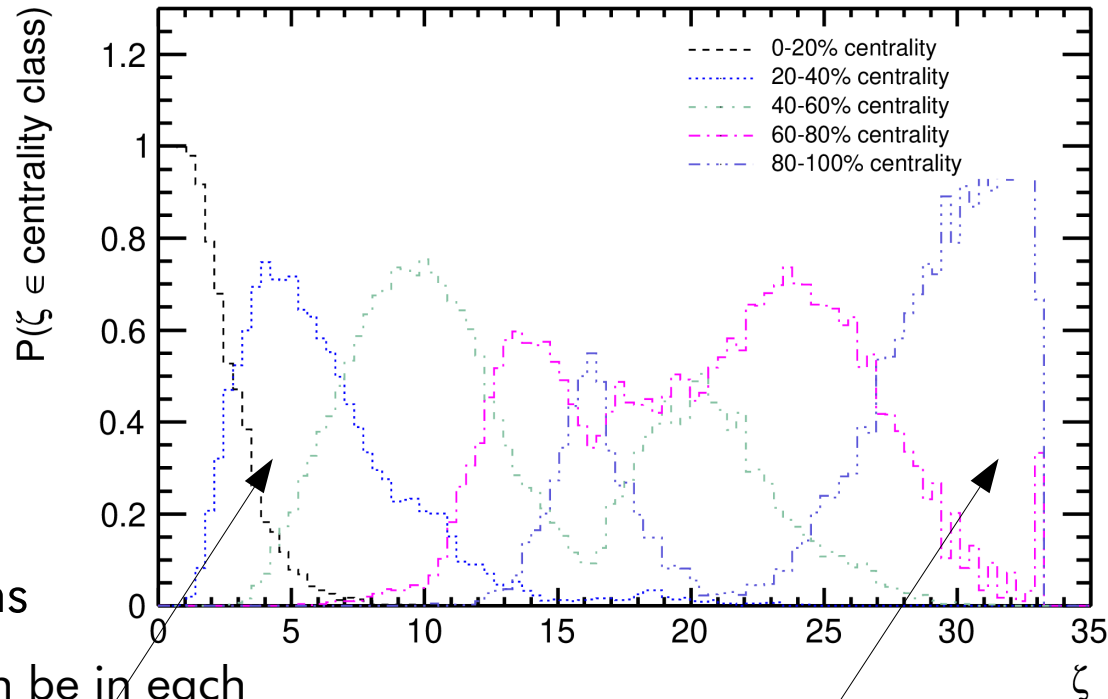
# Scintillator in front of calorimeter

158A GeV, PbPb



Let's consider event with  $\sim 10$  neutrons

158A GeV, PbPb



With equal probability of  $\sim 0.5$  this event can be in each of two centrality classes: 0-20% or 80-100%

$$\zeta = \sum_{i \in frag.} \frac{Z_i^2}{A_{pf.}}$$

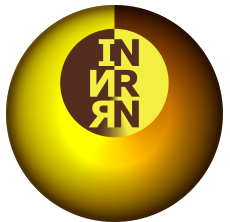
In the opposite case  $\zeta < 1$  it is in 0-20% bin.

In the case of  $\zeta > 30$  it is very probable that the event is in 80-100% centrality bin.



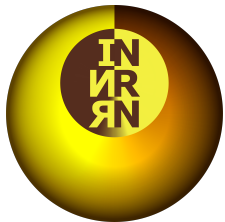
# Classification

Centrality	Main characteristic	Supporting 1	Supporting 2
0-20%	$n_{neutr.}$	$\zeta = \sum_{i \in frag.} \frac{Z_i^2}{A_{pf.}}$	-
20-40%	$n_{neutr.}$	$n_\alpha, n_d$	-
40-60%	$n_{neutr.}$	$n/p$	$n_\alpha, n_d$
60-80%	$n_{neutr.}$	$n/p$	$\zeta = \sum_{i \in frag.} \frac{Z_i^2}{A_{pf.}}$
80-100%	$n_{neutr.}$	$\zeta = \sum_{i \in frag.} \frac{Z_i^2}{A_{pf.}}$	-



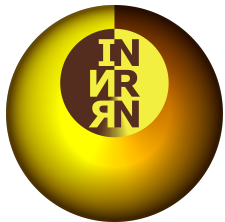
# Summary

- Composition of spectator matter depends on impact parameter
  - Central collisions: nucleons, deuterons, tritons and alphas
  - Semi-peripheral collisions: a few intermediate mass fragments and nucleons along with light fragments, competition of multifragmentation and evaporation
  - Peripheral collisions: neutrons and a heavy residue in evaporation events
- Characteristics useful for centrality determination are:
  - number of neutrons (a traditional one);
  - n/p-ratio of free nucleons;
  - multiplicities of alphas and deuterons
  - sum of squares of fragment charges per spectator nucleon
  - number of charged fragments per nucleon of spectator matter

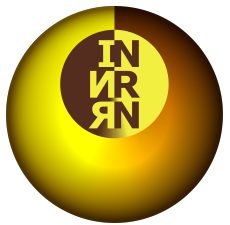
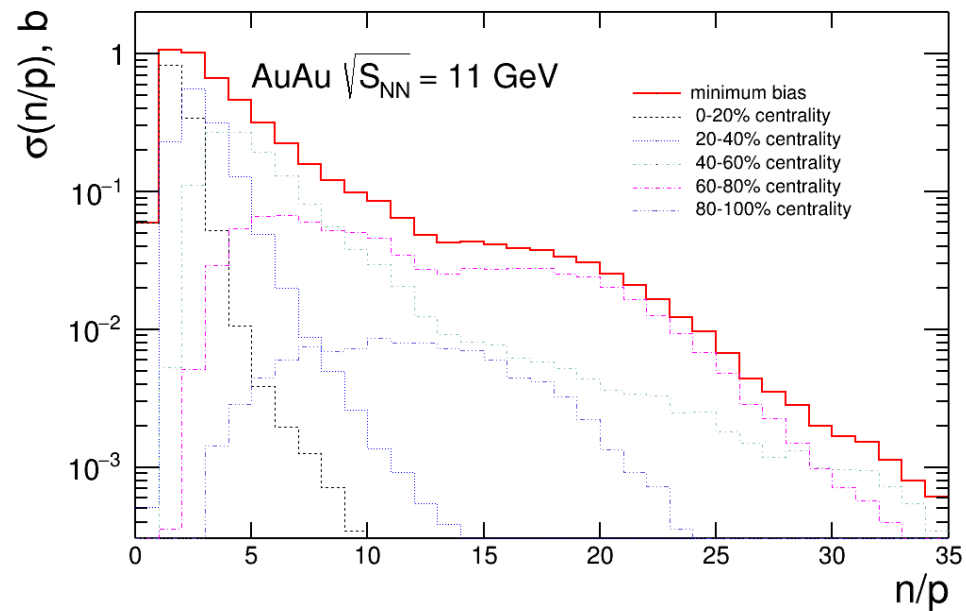
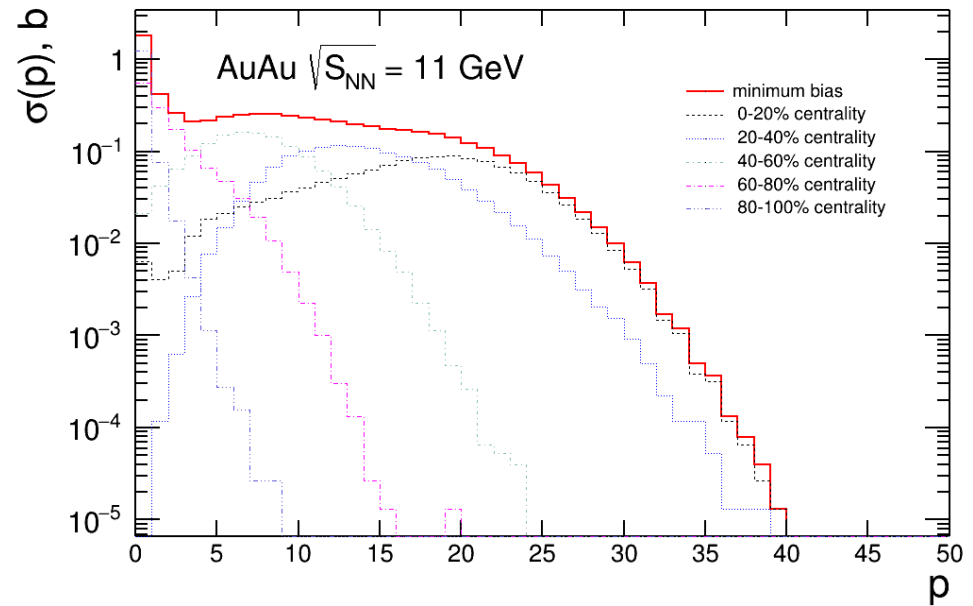
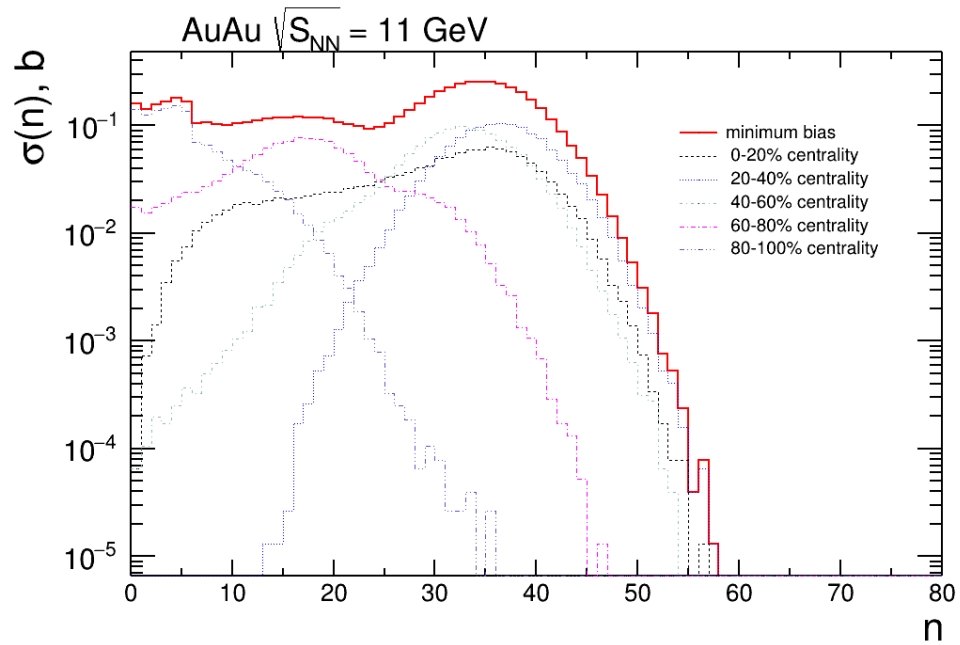


# Thank you for attention!

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# Distributions: free nucleons



# Distributions: alphas and deuterons

