





BlackHawk v1.2: A tool for computing Black Hole evaporation Tutorial session

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1 Introduction

This is a (short) tutorial designed for the attendants of the TOOLS 2020 conference. Following this tutorial, you will be able to launch BlackHawk for two test runs (instantaneous and time-dependent). BlackHawk can be downloaded at

https://blackhawk.hepforge.org

Please make sure that you always use the latest version of the code as bugs are continuously solved by updates. The information given in this tutorial is very concise, more information can be found in the BlackHawk manual (A. Arbey and J. Auffinger, Eur. Phys. J. C79 (2019) 693, arXiv:1905.04268 [gr-qc]).

2 Installation

To install BlackHawk, you need to download the package on the webpage and to un-tar it (using the terminal or your favorite archive software). You obtain a folder called blackhawk_v1.2. You now need to move into this folder within your terminal to install it. For this, just execute the command

make

in the root folder. This may take several minutes. You just created the library libblackhawk.a which contains all the code routines.

3 First run: instantaneous spectra

3.1 Launching the run

First, we are going to use the instantaneous version of BlackHawk. You need to compile the program called BlackHawk_inst by executing the command

make BlackHawk_inst

This should be shorter than the first compilation. Let's now have a look at the parameters.txt file. It contains all the run parameters, some of which are of no interest for us for now. We will now modify some parameters, we *never* modify things out of this parameter file:

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- destination_folder = test1 this will be our fist test
- BHnumber = 1 we will simulate a single BH mass
- Mmin = 1e+10 the BH mass will be 10^{12} g (corresponding to a temperature $T \sim 1$ TeV)
- anumber = 1 we will simulate a single BH spin
- amin = 0 the BH spin will be $a^* = 0$ (Schwarzschild BH)
- spectrum_choice = 0 Dirac distribution for masses
- spectrum_choice_a = 0 Dirac distribution for spins
- Enumber = 1000 1000 energies for optimal precision in the output
- Emin = 5 the lower primary energy will be 5 GeV
- Emax = 1e+5 the higher primary energy will be 10⁵ GeV
- grav = 1 we will compute the graviton spectrum
- primary_only = 0 we will compute the hadronized spectra as well
- hadronization_choice = 2 we will keep "cosmologically stable" particles

Now that the parameters are set, we can run BlackHawk_inst by executing

./BlackHawk_inst.x parameters.txt

the parameters file being given as an argument. This is what you should see:



The program has estimated the amount of RAM and static memory used by the run you are launching. To go on just type y in the asked area. The launch should then go to its end without issue.

3.2 Exploitation of the results

We now want to have a look at the results. In the root folder, there is a subfolder called /results. Go into that folder and you will find your destination_folder /test1 there. This folder contains 4 files:

- BH_spectrum.txt contains the BH mass and spin distribution
- instantaneous_primary_spectra.txt contains the primary SM particles spectra
- instantaneous_secondary_spectra.txt contains the secondary hadronized spectra
- parameters.txt contains a copy of your parameters file, in case you erased or modified it in the root folder

We can now plot the results, using Pyhton for example. For this purpose, proposition of plotting scripts are included in BlackHawk. You can of course modify them. Go into the subfolder /scripts/visualization_scripts and open plot_inst.py. You must now enter your personal path to the root folder of BlackHawk in the script, and the name of your run, as shown below:



You may have to modify (or comment with #) the lines 115, 116, 215 and 216 as the titles are formatted in a way that may cause bugs. By running the cells 1 (package importations in Python) and 2, you must obtain the two figures below:



showing the primary photon and electron/positron spectra, and the secondary photon, electron/positron and proton/anti-proton spectra. The wiggles on the secondary photon spectrum come from the energy discretization. The figures are saved in the /result/test1 subfolder.

You may play a little bit with this tutorial, for example by modifying the BH mass (don't forget to adapt the energy ranges Emin and Emax in that case) or spin, or display other particle spectra with the visualization scripts.

4 Second run: time-dependent spectra

4.1 Launching the run

Then, we are going to launch the time-dependent version of BlackHawk. You need to compile the program called BlackHawk_tot by executing the command

make BlackHawk_tot

in the root folder. We will have the same parameters compared to the first run at the exception of destination_folder = test2 (this will be our second test). There is no conflict between the instantaneous and time-dependent files but this is clearer this way. Now that the parameters are set, we can run BlackHawk_tot by executing

./BlackHawk_tot.x parameters.txt

the parameters file being given as an argument. The run should be somewhat longer, with more information displayed on the screen: BlackHawk is computing the same information as before, but for all the BH evolution until evaporation, in a number of timesteps.

4.2 Exploitation of the results

We now want to have a look at the results. Go into the /result folder and you will find your destination_folder /test2 there. This folder contains 5 types of files:

- BH_spectrum.txt contains the BH mass and spin distribution
- *_primary_spectrum.txt files contain the primary SM particles spectra
- *_secondary_spectrum.txt files contain the secondary hadronized spectra
- life_evolutions.txt contains the timestep of the BH mass and spin integration
- parameters.txt contains a copy of your parameters file, in case you erased or modified it in the root folder

You now need to open the visualization script plot_tot.py. You must still enter your personal path to the root folder of BlackHawk in the script, and the name of your run, as shown below:



By running cells 1, 2, 3, 5 and 6, you should obtain the instantaneous primary photon spectrum at the first timestep. This should be equal to the result of the instantaneous computation of the first part of this tutorial!



By running cell 7, you should obtain the rate of emission of photons at a fixed energy as a function of time:



You may play a little bit with this tutorial, for example by modifying the BH mass (don't forget to adapt the energy ranges Emin and Emax in that case) or spin, or display other particle spectra with the visualization scripts.

4.3 BH evolution

One secondary output of BlackHawk_tot is the global time evolution of the BHs. It is contained in the file life_evolutions.txt. You may want to have a look at the evolution of both mass and spin for different BH masses and spins in this file.

5 Tables: internal information

Finally, BlackHawk contains a number of tabulated quantities that an have an interest in other areas of physics. For example, the hadronization tables are of interest for the particle DM search community. You may have a look at those tables...

6 Conclusion

Thank you and have a good time using BlackHawk! You may contact me if you have any problem at my email address.

