

# Neutral Current Neutrino Interactions at FASER $\nu$

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# Neutral Current Cross-Section at FASER $\nu$

- FASER $\nu$  will give us a unique opportunity to measure  $\nu$  cross-section in the  $\sim [100\text{GeV} - \text{few TeV}]$  range. Charged Current (CC) cross-sections were studied earlier.<sup>2</sup>
- Here we present an analysis strategy to identify and reconstruct Neutral Current (NC) interactions and hence constrain neutral current  $\nu$  cross-sections.

$\nu$  NC studies face two main obstacles at FASER $\nu$  :

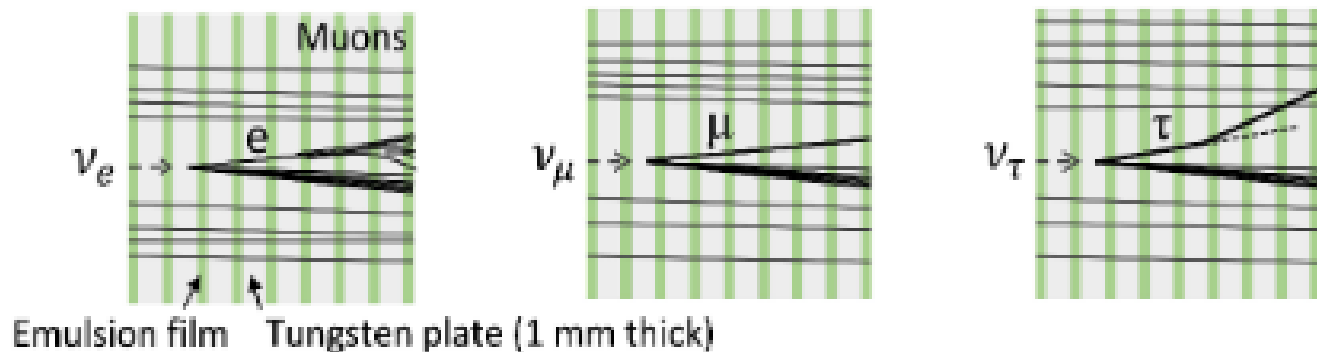
- The missing energy in the final state (carried away by the  $\nu$ ) makes event energy reconstruction very difficult. This is a problem shared by all  $\nu$  NC studies.
- The main background for NC events at FASER $\nu$  are
  - CC events (*one person's treasure is another's background*). This is a less severe problem.
  - Neutral Hadrons (NH), mainly induced by  $\mu$ 's.

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<sup>2</sup>arXiv:1908.02310

# Backgrounds

CC:



**NH:** Apart from the  $\nu$ 's we are interested in only,  $\mu$ 's can travel all the way through rock to the FASER $\nu$  detector. The  $\mu$ 's interact with the rock in front of the detector and within the detector producing NHs. These  $\mu$  induced NHs are our most dominant background. NH interactions look very similar to our signal events.

NH= $n, \bar{n}, \Lambda, \bar{\Lambda}, K_{L,S}, \pi$

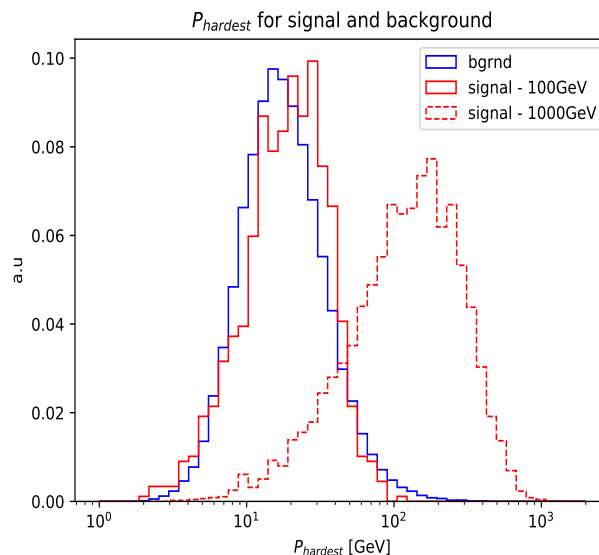
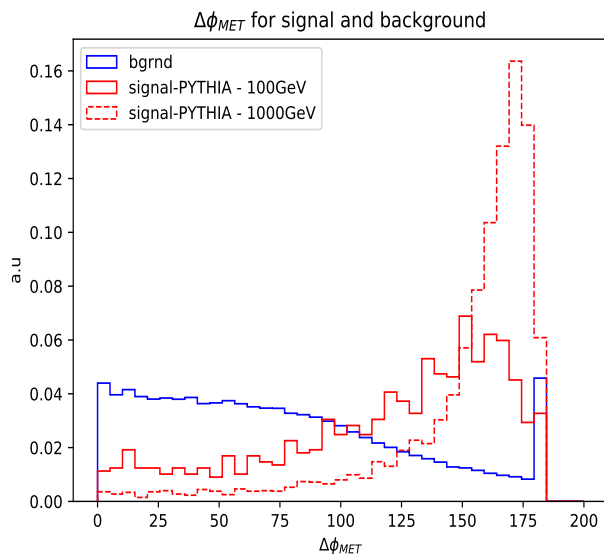
# Event Generation and NN training

- *Event Generation:* We use Pythia to simulate  $\nu$ -W and NH-W collision. Other generators were compared with Pythia and were in agreement.
- *Event Selection:* We select events with  $\geq 5$  charged tracks, each charged track has energy  $\geq 1$  GeV, and  $\theta < \pi/4$ .
- *Detector Simulation:*
  - Track momentum and energy smearing.
  - Identifies each visible track as electron, photon or a normal track.
  - Determines if the track interacts within the detector.
- *NN training:* We use 2 NN's:
  - Classifier N/W: Distinguishes signal(NC) and background(NH) events.
  - Regression N/W: Estimates the incoming particle energy. Only on identified signal events.

# Observables - Signal vs Background

We use a total of 10 observables to characterize an event.

$\Delta\phi_{MET}$  = The azimuthal angle between the reconstructed missing transverse momentum and the nearest track.



# Neural Network Results

The trained n/w's predict on a separate data set of signal and background events. First the classifier network classifies events into signal and background. Only the events classified as signal are passed into the regression n/w for energy estimation.

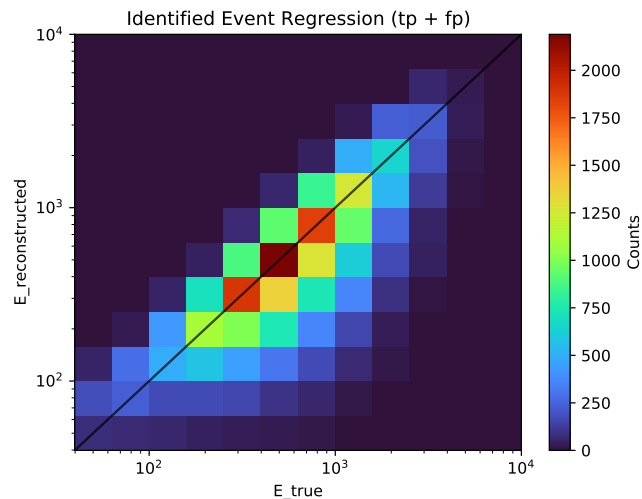
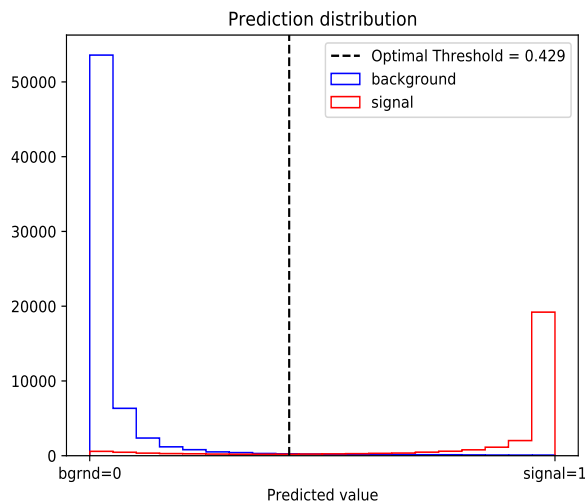
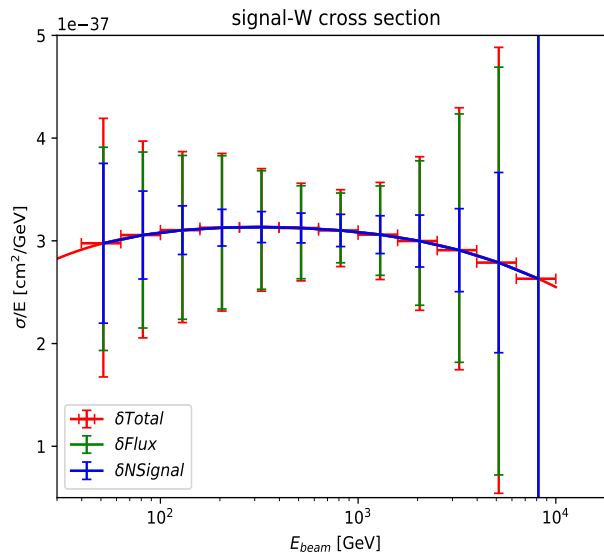
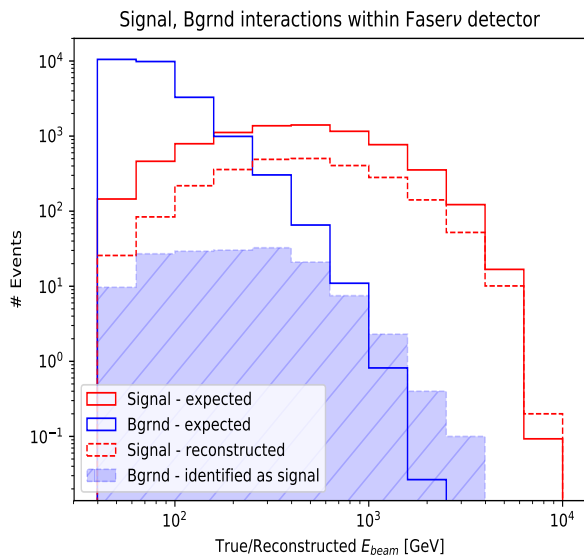


Figure: Results of the (**left**) classifier and (**right**) regression n/w.

# Cross-Section Results

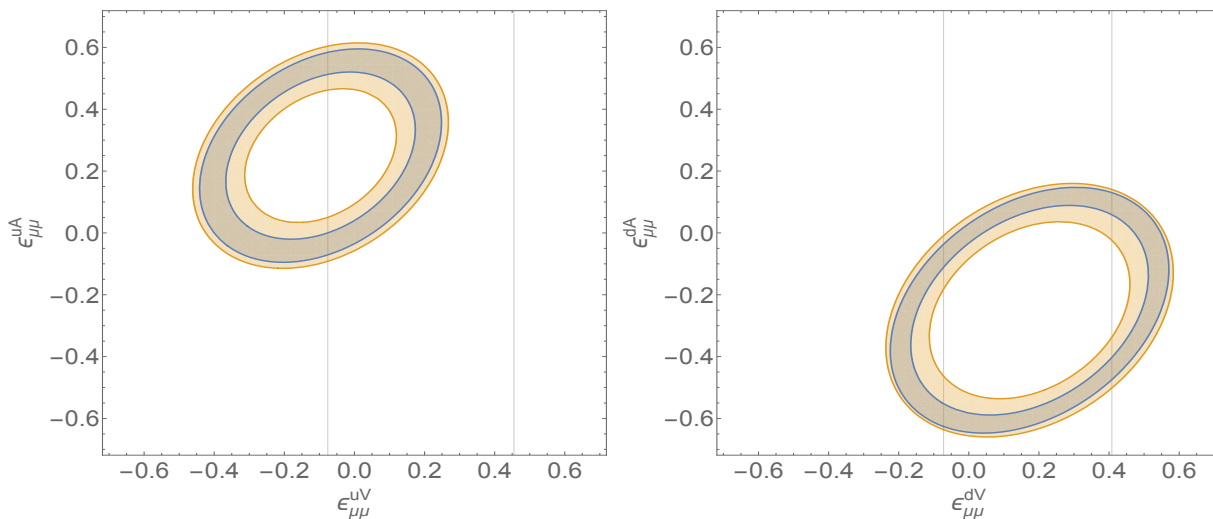
O/p of the NN's gives us the number of reconstructed events in each energy bin. This gives us size of statistical uncertainty on  $\nu$  NC interaction cross-section. The other source of uncertainty is the one on incoming flux.



Other uncertainties: NH flux, simulation.

# Future Work - Constraining NSI (an example)

$$\mathcal{L} = -2\sqrt{2}G_F \sum_{f,\alpha,\beta,X} \epsilon_{\alpha,\beta}^{f,X} (\bar{\nu}_\alpha \gamma^\mu P_L \nu_\beta) (\bar{f} \gamma_\mu P_X f)$$



**Figure:** Comparison of bounds on NSI couplings from CHARM(400GeV)(orange) and FASER $\nu$  (grey) for **(left:)** up quark and **(right:)** down quark in the Vector-Axial vector coupling plane. Vertical lines are bounds from oscillations and COHERENT that constrain only vector NSI.

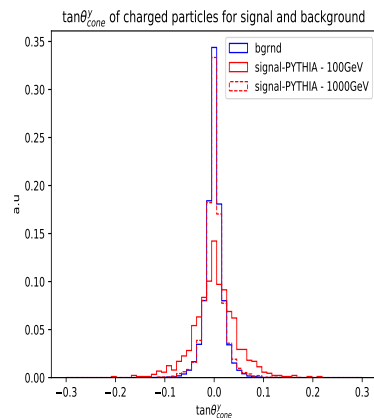
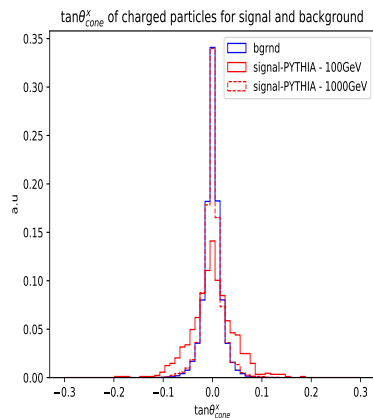
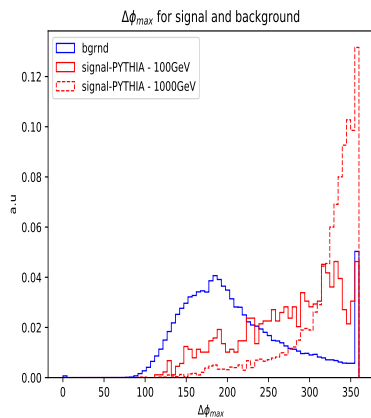
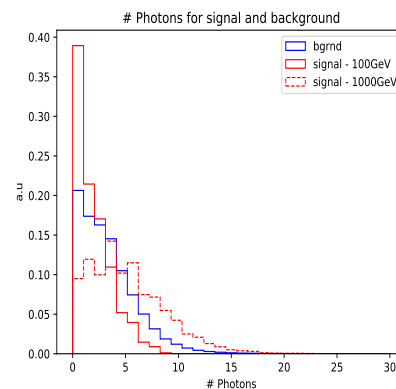
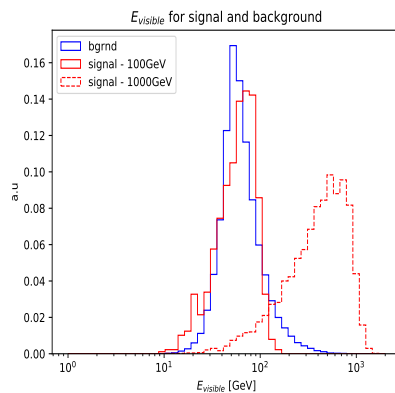
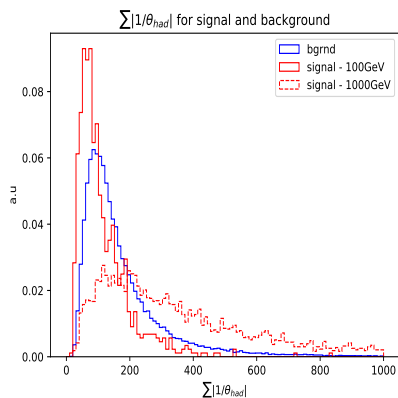


- FASER $\nu$  can study CC and NC events at unprobed energies.
- We show here a strategy to overcome the usual difficulties with NC studies using machine learning.
- Both event identification and energy reconstruction were done to constrain  $\nu$  NC cross-section.
- This sensitivity to NC interactions can be used to do various physics, eg: constraining NSI, light mediators etc.

# Backup Slides-Observables

- $n_{ch} \sim \log E_{had}$
- $n_{\gamma} \sim n_{\pi^0} \sim \log E_{had}$
- $\sum E_{ch} + \sum E_{\gamma} \sim E_{had}$
- $p_{hard} \sim E_{had}$
- $\sum |1/\theta_{had}| \sim E_{had}$
- $\tan \theta_{cone}^S = (\sum p_{T,i})/(\sum p_i) \sim H_T/E_{had}$
- $\tan \theta_{cone}^V = (\sum \vec{p}_{T,i})/(\sum p_i) \sim \vec{p}_T/E_{had}$
- Largest Azimuthal Gap: The largest difference in azimuthal angle between two neighbouring tracks,  $\Delta\phi_{max}$ .
- Track-MET-Angle: The azimuthal angle between the reconstructed missing transverse momentum,  $\vec{p}_T^{\#}$  and the nearest track,  $\Delta\phi_{MET}$ .

# Backup Slides-Other Observables



# Backup Slides - Comparing various NHs

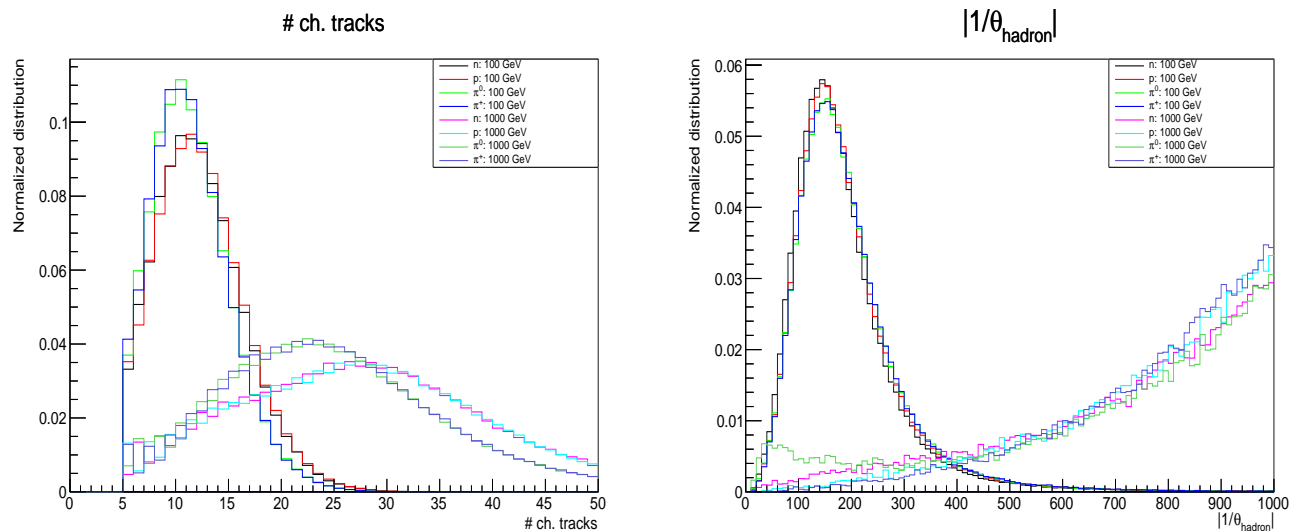
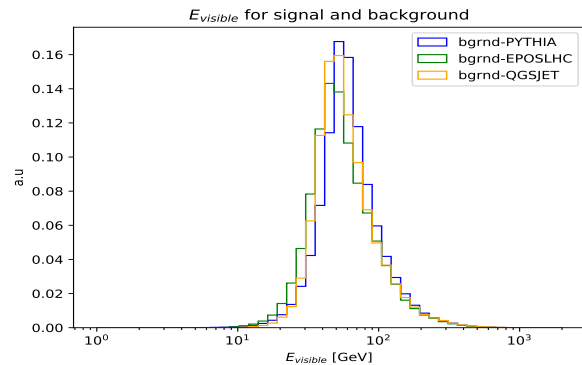
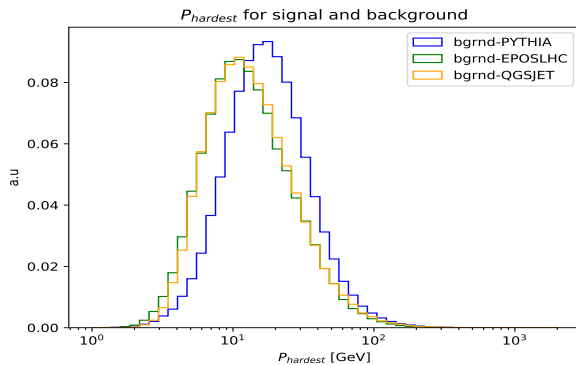
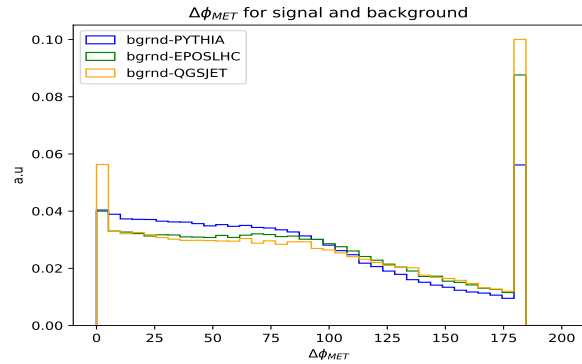
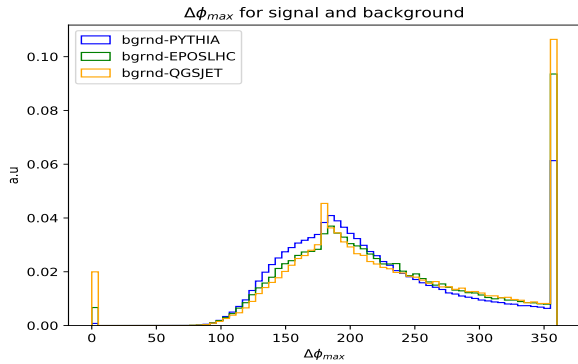
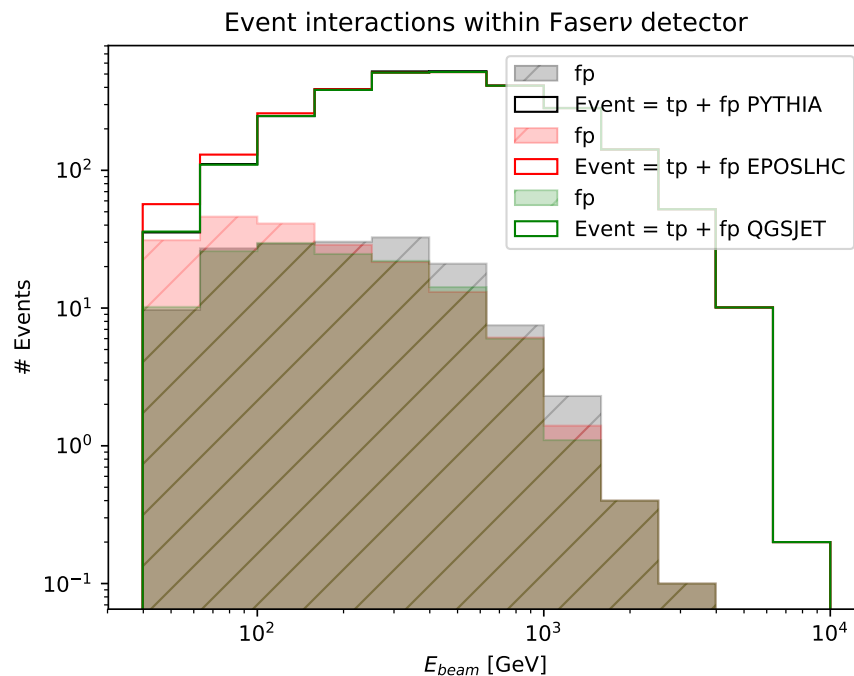


Figure: Comparison of (**left**) charged track multiplicity and (**right**)  $\sum |1/\theta_{\text{had}}|$  for  $n, p, \pi^0, \pi^+$  at 100,1000 GeV.

# Backup Slides - Background Observables: Pythia vs EPOSLHC vs QGSJET



# Backup Slides - Prediction with Different Backgrounds: Pythia vs EPOSLHC vs QGSJET



# Backup Slides - Neural Network Results

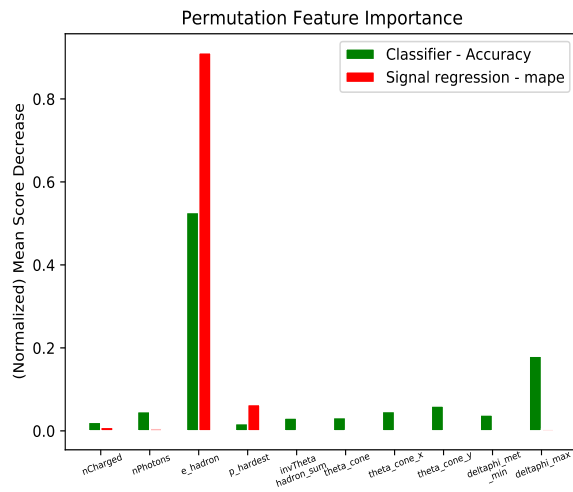
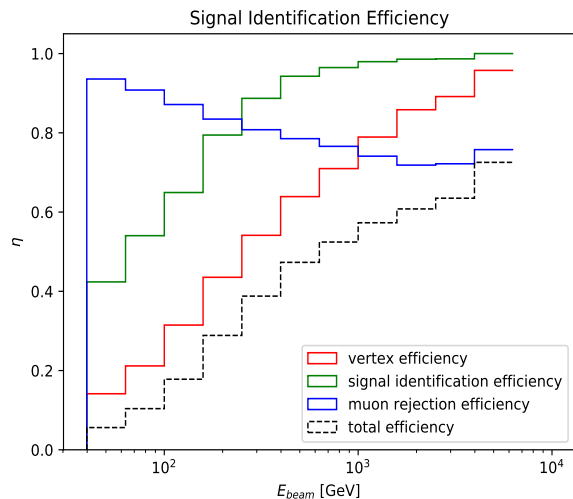


Figure: (**Left**) Signal identification efficiency and (**right**) feature importance.