

# The Kinetic Mass of Heavy Quarks at Three Loops

RADCOR-LoopFest 2021

Matteo Fael | May 18, 2021

INSTITUTE FOR THEORETICAL PARTICLE PHYSICS - KIT KARLSRUHE

based on [Fael, Schönwald, Steinhauser, PRL 125 \(2020\) 052003, JHEP 10 \(2020\) 087](#)

- $jV_{cb}^{\text{inc}}j = (42:19 \quad 0:78) \quad 10^{-3}$

HFLAV 2019

- Global fits in the **kinetic scheme**

Bigi et al, PRD 56 (1997) 4017;

Gambino, Schwanda, Phys.Rev.D 89 (2014) 1, 014022

Alberti, Gambino, Healey, Nandi, Phys.Rev.Lett. 114 (2015) 6, 061802.

- Also 1S scheme

Hoang, Ligeti, Manohar, Phys.Rev.Lett. 82 (1999) 277

Bauer, Ligeti, Manohar, Trott, Phys.Rev.D 70 (2004) 094017

$$s_l = C_0 + C_2 \frac{2}{m_b^2} + C_G \frac{2}{m_b^2} + C_D \frac{3}{m_b^3} + C_{LS} \frac{3}{m_b^3} + \dots$$

- Wilson coefficients  $C_i$  are calculable in pQCD
- Non-perturbative HQE parameters  $\langle \bar{\psi} \psi \rangle$ ;  $G$ ;  $D$ ;  $LS$   $\langle \bar{b} \gamma_j O_i \psi \rangle$
- Global fits determine  $|V_{cb}|$ ;  $m_b$ ;  $m_c$  and the HQE parameters.

# Why three-loop correction to $m_b^{\text{kin}}$ ?

- Precision predictions need short mass schemes for better convergence of  $\alpha_s$  series.
- Precise conversion between  $\overline{m}_b$  and  $m_b^{\text{kin}}$ .
- Improve the SM prediction by including  $\alpha_s^3$  corrections in pQCD.

# The heavy-quark expansion

	tree	s	$\frac{2}{s}$	$\frac{3}{s}$	
1	3	3	3	NEW	<p>Jezebek, Kuhn, NPB 314 (1989) 1; Gambino et al., NPB 719 (2005) 77;                      Melnikov, PLB 666 (2008) 336; Pak, Czarnecki, PRD 78 (2008) 114015.                      MF, Schönwald, Steinhauser, hep-ph:2011.13654 ;</p>
$1=m_b^2$	3	3	!		<p>Alberti, Gambino, Nandi, JHEP 1401 (2014) 147;                      Mannel, Pivovarov, Rosenthal, PRD 92 (2015) 054025.                      Becher, Boos, Lunghi, JHEP 0712 (2007) 062.</p>
$1=m_b^3$	3	3			<p>Mannel, Pivovarov, PRD 100 (2019) 9.</p>
$1=m_b^{4;5}$	3				<p>Dassinger, Mannel, Turczyk, JHEP 0703 (2007) 087, JHEP 1011 (2010) 109                      MF, Mannel, Vos, JHEP 02 (2019) 177, JHEP 12 (2019) 067.</p>
$\bar{m}_b \quad m_b^{\text{kin}}$		3	3	NEW	<p>Bigi et al, PRD 56 (1997) 4017; Czarnecki, Melnikov, Uraltsev, PRL 80 (1998) 3189.                      MF, Steinhauser, Schönwald, PRL 125 (2020) 5; PRD 103 (2021) 014005</p>

# The heavy-quark expansion

K. Schönwald's talk on Thursday  
see also M. Steinhauser and A. Czarnecki talks

	tree	s	$\frac{2}{s}$	$\frac{3}{s}$	
1	3	3	3	NEW	<p>Jezabek, Kuhn, NPB 314 (1989) 1; Gambino et al., NPB 719 (2005) 77; Melnikov, PLB 666 (2008) 336; Pak, Czarnecki, PRD 78 (2008) 114015. MF, Schönwald, Steinhauser, hep-ph:2011.13654 ;</p> <p>Alberti, Gambino, Nandi, JHEP 1401 (2014) 147; Mannel, Pivovarov, Rosenthal, PRD 92 (2015) 054025. Becher, Boos, Lunghi, JHEP 0712 (2007) 062.</p> <p>Mannel, Pivovarov, PRD 100 (2019) 9.</p> <p>Dassinger, Mannel, Turczyk, JHEP 0703 (2007) 087, JHEP 1011 (2010) 109 MF, Mannel, Vos, JHEP 02 (2019) 177, JHEP 12 (2019) 067.</p>
$1=m_b^2$	3	3	!		<p>Bigi et al, PRD 56 (1997) 4017; Czarnecki, Melnikov, Uraltsev, PRL 80 (1998) 3189. MF, Steinhauser, Schönwald, PRL 125 (2020) 5; PRD 103 (2021) 014005</p>
$1=m_b^3$	3	3			
$1=m_b^{4;5}$	3				
$\bar{m}_b \quad m_b^{\text{kin}}$		3	3	NEW	

A short distance mass for  $B \rightarrow X_c \gamma$

$$s_l = \frac{G_F^2 |V_{cb}|^2 (m_b^{\text{OS}})^5}{192 \pi^3} f(0.25) \left[ 1 + 1.78 \frac{s}{m_b^2} + 13.1 \frac{s^2}{m_b^4} + 163.3 \frac{s^3}{m_b^6} + \mathcal{O}\left(\frac{s^4}{m_b^8}\right) \right]$$

See: Bigi, Shifman, Uraltsev, Vainshtein PRD 50 (1994) 2234; Beneke, Braun, NPB 426 (1994) 301;  
Ball, Beneke, Braun, PRD 52 (1995) 3929; Melnikov, van Ritbergen, PLB 482 (2000) 99.

■ Mass scheme change:  $m_b^{\text{OS}} \rightarrow m_b \left( 1 + c \frac{s}{m_b^2} \right)$

$$s_l / (m_b)^n \left[ 1 + (nc + a_1) \frac{s}{m_b^2} + \frac{n(n+1)}{2} c^2 \frac{s^2}{m_b^4} + nc a_1 + a_2 \frac{s^2}{m_b^4} + \dots \right]$$

■ Can we resum the power enhanced  $(\frac{s}{m_b^2})^k$  terms (with  $n = 5$ )?

# Meson-quark mass relation

$$m_b = M_B - \frac{E_B}{2m_b} + \dots$$

- $E_B$ : the B-meson binding energy.
- $\dots$ : the kinetic energy induced by the residual motion of the heavy quark.

The relevant parameter in  $\Gamma_{sl}$  is  $m_b^5$ , not  $M_B^5$ :

$$\Gamma_{sl} \propto \frac{G_F^2 |V_{cb}|^5}{192 \cdot 3} (M_B - E_B)^5$$



# The kinetic mass

$$m_b^{\text{kin}}(\mu) = m_b^{\text{OS}} \left[ \bar{m}_b(\mu) \right]_{\text{pert}} \frac{[\bar{m}_b^2(\mu)]_{\text{pert}}}{2m_b^{\text{kin}}(\mu)} \dots$$

Bigi, Shifman, Uraltsev, Vainshtein, PRD 56 (1997) 4017.  
 see also: Czarniecki, Melnikov, Uraltsev, PRL 80 (1998) 3189;  
 Gambino, JHEP 09 (2011) 055;

- In pQCD, we can peel off the IR renormalon from the on-shell mass identifying:

$$m_b(\mu) \neq m_b^{\text{kin}}(\mu) \quad \bar{M}_B \neq m_b^{\text{OS}}$$

$$\bar{m}_b(\mu) \neq [\bar{m}_b(\mu)]_{\text{pert}} \quad [\bar{m}_b^2(\mu)] \neq [\bar{m}_b^2(\mu)]_{\text{pert}}$$

# The Small Velocity Sum Rules

- How to give an operative definition of  $\bar{\epsilon}$  and  $\bar{\epsilon}^2$ ?
- Moments of the excitation energy:

$$I_n(q^2) = \sum_{d!} \frac{d^n}{d!} \frac{d}{dq^2}$$

with  $d! = E_{X_c} M_D$  and  $q = p_0 + p$

# The Small Velocity Sum Rules

Take the limit where the  $X_c$ 's velocity is small:  $|\mathbf{v}| = |\mathbf{q}| = m_c$  1:

$$I_0(\mathbf{q}^2) = |\mathbf{q}| \frac{G_F^2 V_{cb}^2}{8} (m_b - m_c)^2 + \mathcal{O}(|\mathbf{v}|^2); \frac{\text{QCD}}{m_b}$$

$$I_1(\mathbf{q}^2) = I_0 \frac{v^2}{2} + \mathcal{O}(|\mathbf{v}|^3); \frac{\text{QCD}^2}{m_b^2}$$

$$I_2(\mathbf{q}^2) = I_0 \frac{v^2}{3} + \mathcal{O}(|\mathbf{v}|^3); \frac{\text{QCD}^3}{m_b^3}$$

# The Small Velocity Sum Rules

$$[\overline{(\quad)}]_{\text{pert}} = \lim_{v \rightarrow 0} \lim_{m_b \rightarrow 1} \frac{2}{v^2} \frac{\int_0^R d!! W(!; v)}{\int_0^R d! W(!; v)}$$

$$[{}^2(\quad)]_{\text{pert}} = \lim_{v \rightarrow 0} \lim_{m_b \rightarrow 1} \frac{3}{v^2} \frac{\int_0^R d!! {}^2 W(!; v)}{\int_0^R d! W(!; v)}$$

- Consider only soft radiation QCD  $m_b$

# The Small Velocity Sum Rules

$$[\bar{(\quad)}]_{\text{pert}} = \lim_{v \rightarrow 0} \lim_{m_b \rightarrow 0} \frac{2}{v^2} \frac{\int_0^R d! \int_0^R d! W(!; v)}{\int_0^R d! W(!; v)}$$

$$[{}^2(\quad)]_{\text{pert}} = \lim_{v \rightarrow 0} \lim_{m_b \rightarrow 0} \frac{3}{v^2} \frac{\int_0^R d! \int_0^R d! {}^2 W(!; v)}{\int_0^R d! W(!; v)}$$

# SV Limit \$ Threshold Limit

- Excite the heavy quark,  
but just a bit :::

$$y = s \frac{m_b^2}{2m_b!} m_b^2$$

- SV Limit corresponds to 1 Particle Threshold limit !
- Factorization:

$$W(!; \nu) ' H U(!; \nu)$$

# Ingredients for $m_b^{\text{kin}}$ at $O(\frac{3}{s})$

- $W(!; \forall)$  up to  $O(\frac{3}{s})$
- Imaginary part of forward scattering amplitudes:

$$W(!; \forall) = W_{\text{el}}(\forall) (!) + \frac{\forall^2}{!} W_{\text{real}}(!) (!) + O(\forall^4; \frac{!}{m_b})$$

- Threshold expansion via method of regions:  $y = s - m_b^2$ .

[Beneke, Smirnov, NPB 522 \(1998\) 321; Smirnov Springer Tracts Mod. Phys. 250 \(2010\)](#)

# Expansion by Regions

- For one heavy particle threshold, there are two regions:

see also: [Smirnov Springer Tracts Mod. Phys. 250 \(2010\)](#)

hard (h):  $k_i \gg m_b$

ultra-soft (us):  $k_i \ll m_b$

(u) (h)

(uu) (uh) (hh)

(uuu) (uuh) (uhh) (hhh)



# Expansion by regions

- For one heavy particle threshold, there are two regions:

see also: [Smirnov Springer Tracts Mod. Phys. 250 \(2010\)](#)

$$\begin{aligned} \text{hard (h): } & k_i \gg m_b \\ \text{ultra-soft (us): } & k_i \ll m_b \end{aligned}$$

(u) (h)

(uu) (uh) (hh)

(uuu) (uuh) (uhh) (hhh)

- All hard regions don't contribute: no imaginary part.
- After renormalization and decoupling  $\frac{(n_l + n_h)!}{s^{n_l + n_h}} \rightarrow \frac{(n_l)!}{s^{n_l}}$ , only all ultra-soft part remains

# Expansion by regions

- For one heavy particle threshold, there are two regions:

see also: [Smirnov Springer Tracts Mod. Phys. 250 \(2010\)](#)

$$\begin{aligned} \text{hard (h): } & k_i \gg m_b \\ \text{ultra-soft (us): } & k_i \ll m_b \end{aligned}$$

(u) (h)

(uu) (uh) (hh)

(uuu) (uuh) (uhh) (hhh)

- All hard regions don't contribute: no imaginary part.
- After renormalization and decoupling  $\frac{(n_l + n_h)!}{s^{n_l + n_h}} \rightarrow \frac{(n_l)!}{s^{n_l}}$ , only all ultra-soft part remains

# For the statistics . . .

- 4, 66 and 1586 diagrams at one, two and three loops
- Partial fractioning and mapping between families implemented in FORM thanks to code LIMIT.

[Herren, PhD thesis, KIT, 2020](#)

- FIRE and LiteRed reduction of integral families:

[Smirnov, Chuharev, hep-ph/1901.07808](#); [Lee, hep-ph/1212.2685](#).

- $f^2; 2g$  in the  $f(uu); (uh)g$  regions at two loops
- $f^4; 4; 3g$  in the  $f(uuu); (uuh); (uhh)g$  regions at three loops
- New master integrals:
  - 3 in the  $(uu)$  region
  - 20 in the  $(uuu)$  region
- Heavy quark form factors up to  $O(\frac{2}{s})$  (static limit)

[Lee, Smirnov, Smirnov, Steinhauser, JHEP 1805 \(2018\) 187](#); [Blümlein, Marquard, Rana, PRD 99 \(2019\) 016013](#)

$$I_1^{2|}$$

$$I_2^{2|}$$

$$I_3^{3|}$$

- The new master integrals contains linear-massive propagators:

$$I_2^{2|} = \int d^d k_1 d^d k_2 \frac{1}{k_1^2 (k_1 - k_2)^2 (2k_1 - p - y)(2k_2 - p)}$$

## ■ Mellin-Barnes method

MBpackage

Czakon, *Comput. Phys. Commun.* 175 (2006) 559;

Smirnov<sup>2</sup>, *EPJC* 62 (2009) 445.

## ■ PSLQ

## ■ Analytic summation of residues

HarmincSums

[www3.risc.jku.at/research/combinat/software/HarmonicSums/](http://www3.risc.jku.at/research/combinat/software/HarmonicSums/)

## ■ Differential equations in auxiliary variable

Kotikov, *PLB* 254 (1991), 158

Gehrmann, Remiddi, *NPB* 580 (2000) 485

Henn, *PRL* 110 (2013), 251601.

# The kinetic mass

MF, Schönwald, Steinhauser, PRL 125 (2020) 052003

- The mass relation is written in terms of  $\binom{n_l}{s}$ .
- $n_l =$  number of massless quarks,  $l = \log(2 = s)$ .

Charm mass effects:  $0 < m_c, m_c \ll m_b$

- Charm mass effects to the  $\overline{\text{MS}}$ -on-shell mass relation known up to  $\mathcal{O}(\frac{3}{s})$ .  
[Davydychev, Grozin, PRD 59 \(1999\) 054024](#); [Broadhurst, Gray, Schlicher, Z. Phys. C 52 \(1991\) 111](#);  
[Bekavac, Grozin, Seidel, Steinhauser, JHEP 10 \(2007\) 006](#); [MF, Schönwald, Steinhauser, JHEP 10 \(2020\) 087](#).
- No charm mass effects known for the kinetic-on-shell mass relation.
- We assume  $|j| \ll m_c^2; m_b^2$ , i.e. no cut through a charm loop

# The kinetic- $\overline{\text{MS}}$ mass relation for bottom

- Input values:

$${}^{(5)}_s(M_Z) = 0:1179 \quad \overline{m}_c(3 \text{ GeV}) = 0:993 \text{ GeV} \quad \overline{m}_b(\overline{m}_b) = 4:163 \text{ GeV}$$

- The charm quark wants to be treated as heavy .

scheme	${}^{(n_r)}_s$	$m_c$ in $\overline{\text{MS}}$ -OS	$m_c$ in kin-OS	
(A)	3	3	-	$m_b^{\text{kin}}(1 \text{ GeV}) = 4163 + 248 + 80 + 30 = 4520 \text{ MeV}$
(B)	4	3	3	$4163 + 259 + 78 + 26 = 4526 \text{ MeV}$
(C)	4	3	7	$4163 + 259 + 84 + 41 = 4547 \text{ MeV}$
(D)	3	7	7	$4163 + 248 + 81 + 30 = 4521 \text{ MeV}$

MF, Schönwald, Steinhauser, Phys.Rev.D 103 (2021) 014005



# Scheme conversion uncertainty

- Half of  $\frac{3}{s}$  correction

$$m_b^{\text{kin}} \cdot 15 \text{ MeV}$$

- Large- $\frac{4}{s}$  term at

$$m_b^{\text{kin}} \cdot 8 \text{ MeV}$$

Compared with:

- scheme conversion uncertainty at two-loops:  $m_b^{\text{kin}} = 30 \text{ MeV}$  [Gambino, JHEP 09 \(2011\) 055](#)
- $m_b^{\text{kin}}(1 \text{ GeV}) = 4554 \cdot 18 \text{ MeV}$  from B !  $X_c \cdot \text{global } t$  [HFLAV 2019](#)

# Conclusions

- We computed the  $O(\frac{3}{s})$  relation between the kinetic and the on-shell mass.
- We studied finite charm mass effects in  $m_b^{\text{kin}}$ : the charm wants to be heavy!
- Scheme conversion uncertainty is reduced by a factor of 2.
- Mass formula is necessary to improve  $B \rightarrow X_c \ell \bar{\nu}$  prediction in the SM.
- Our results are crucial for future extractions of  $|V_{cb}|$  and  $m_b$  from Belle-II data.

Spare

## Let's include radiative corrections ...

$$I_n(\mathfrak{q}^2) = \int_{|\mathfrak{q}|}^{q_0^{\max}} dq_0 \int^n \frac{d \text{ tree}}{dq_0 d\mathfrak{q}^2} + \int_{q_0^{\max}}^{q_0^{\max}} dq_0 \int^n \frac{d \text{ s}}{dq_0 d\mathfrak{q}^2} + \int_{|\mathfrak{q}|}^{q_0^{\max}} dq_0 \int^n \frac{d \text{ s}}{dq_0 d\mathfrak{q}^2}$$

- We introduce a **Wilsonian cutoff** , with  $\Lambda_{\text{QCD}}$   $M_B$ , to separate gluons with
  - $\Lambda_{\text{QCD}} < M_B$  that belong to the non-perturbative regime,
  - $\Lambda_{\text{QCD}} > M_B$  that can be described in pQCD.

## Let's include radiative corrections ...

$$I_1(q^2) = \int_{|q|}^{q_0^{\max}} dq_0 \left[ \frac{d_{\text{tree}}}{dq_0 dq^2} + \int_{q_0^{\max}}^{q_0^{\max}} dq_0 \left[ \frac{d_s}{dq_0 dq^2} + \int_{|q|}^{q_0^{\max}} dq_0 \frac{d_s}{dq_0 dq^2} \right] \right]$$

use this to define in the SV limit  $\bar{\Lambda}(\mu)$

- We introduce a **Wilsonian cutoff**, with  $\Lambda_{\text{QCD}} < \mu < M_B$ , to separate gluons with
  - $\mu < \Lambda_{\text{QCD}}$  that belong to the non-perturbative regime,
  - $\mu > \Lambda_{\text{QCD}}$  that can be described in pQCD.

# $m_b^{\text{kin}}(1 \text{ GeV})$ from $\overline{m}_b(\overline{m}_b)$

scheme	$(n_f)_s$	$m_c$ in $\overline{\text{MS}}\text{-OS}$	$m_c$ in kin-OS	in MeV
(A)	3	3	—	$4163 + 248 + (81 + 7_{m_c} + 12_{\text{dec}} 20_{n_c})$ $+ (30 + 14_{m_c} + 16_{\text{dec}} 30_{n_c} 1_{n_c} \text{ dec} + 0:4_{m_c} \text{ dec})$ $= 4163 + 248 + 80 + 30 = 4520$
(B)	4	3	3	$4163 + 259 + (88 + 7_{m_c} + 5_{m_c}^{\text{kin}} 22_{n_c})$ $+ (34 + 16_{m_c} + 10_{m_c}^{\text{kin}} 34_{n_c})$ $= 4163 + 259 + 78 + 26 = 4526$

$m_b^{\text{kin}}(1 \text{ GeV})$  from  $\bar{m}_b(\bar{m}_b)$

in MeV

	$(n_f)$ s	$m_c$ in $\overline{\text{MS}}\text{-OS}$	$m_c$ in kin-OS	
(C)	4	3	7	$4163 + 259 + (99 + 7_{m_c} 22_{n_c})$ $+ (59 + 16_{m_c} 34_{n_c})$ $= 4163 + 259 + 84 + 41 = 4547$

(D)	3	7	7	$4163 + 248 + 81 + 30 = 4521$
-----	---	---	---	-------------------------------

# Charm quark mass

- $m_c^{\text{kin}}(0.5 \text{ GeV})$ :

$$m_c^{\text{kin}}(0.5 \text{ GeV}) = 993 + 191 + 100 + 52 \text{ MeV} = 1336 \text{ MeV};$$

$$m_c^{\text{kin}}(0.5 \text{ GeV}) = 1099 + 163 + 76 + 34 \text{ MeV} = 1372 \text{ MeV};$$

$$m_c^{\text{kin}}(0.5 \text{ GeV}) = 1279 + 84 + 30 + 11 \text{ MeV} = 1404 \text{ MeV};$$

- $m_c^{\text{kin}}(1 \text{ GeV})$ :

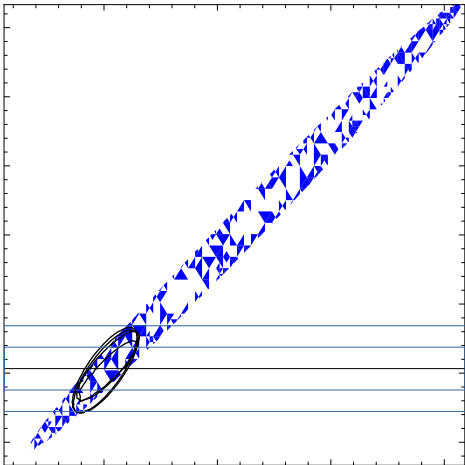
$$m_c^{\text{kin}}(1 \text{ GeV}) = 993 + 83 + 35 + 14 \text{ MeV} = 1125 \text{ MeV};$$

$$m_c^{\text{kin}}(1 \text{ GeV}) = 1099 + 37 + 2 + 3 \text{ MeV} = 1135 \text{ MeV};$$

$$m_c^{\text{kin}}(1 \text{ GeV}) = 1279 + 73 + 61 + 17 \text{ MeV} = 1128 \text{ MeV};$$

where from top to bottom  $s = 3 \text{ GeV}; 2 \text{ GeV}$  and  $\bar{m}_c$  for  $\bar{m}_c(s)$  and  $\bar{m}_c^{(3)}(s)$ .





Original plot from  
[Gambino, Schwanda, PRD 89 \(2014\) 014022](#)  
Horizontal error bands superimposed by MF