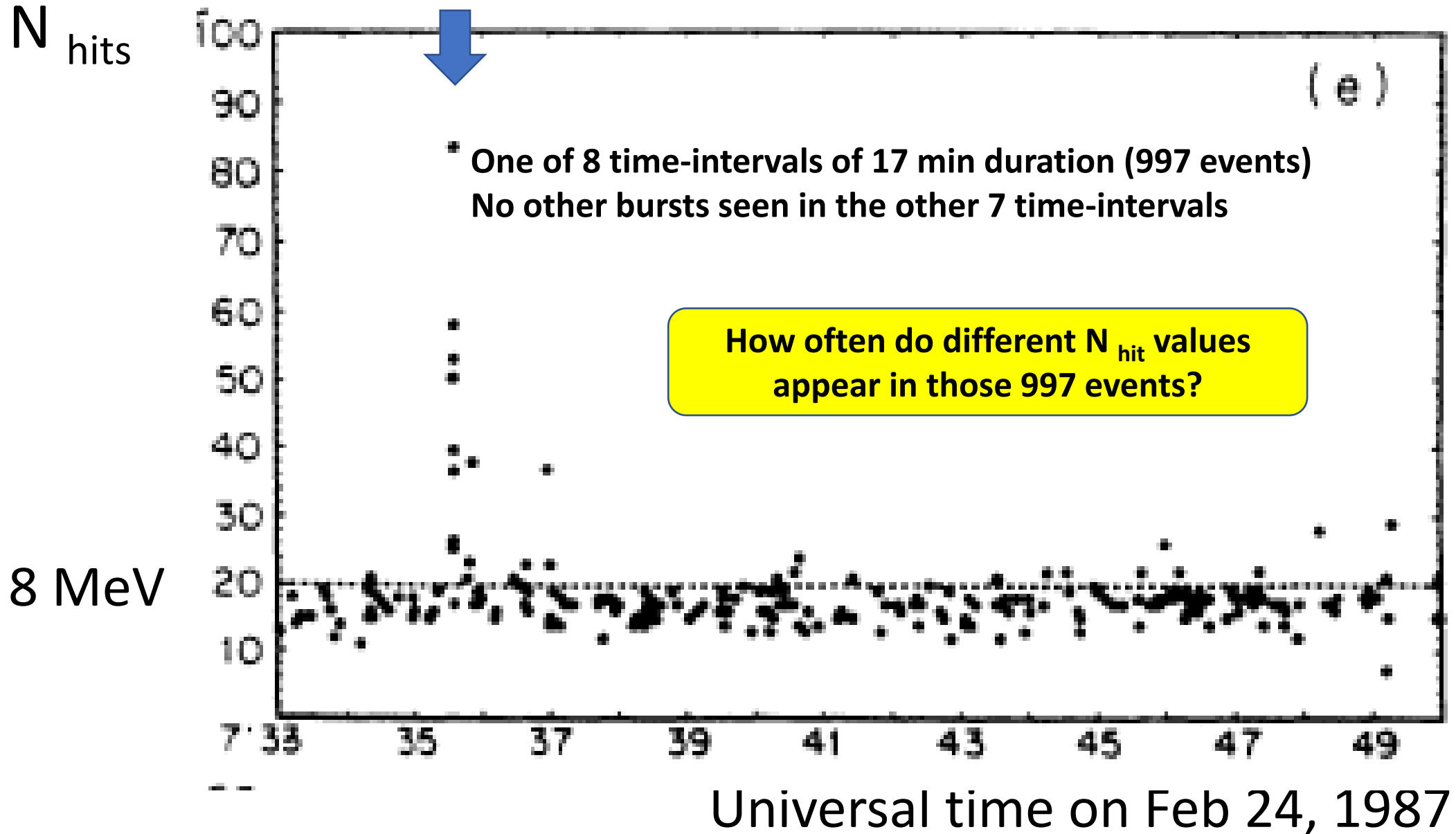


**Evidence for an 8 MeV line  
in the SN 1987A neutrino Spectrum  
and five reasons to expect one**

Robert Ehrlich  
BSM 2021 Conference

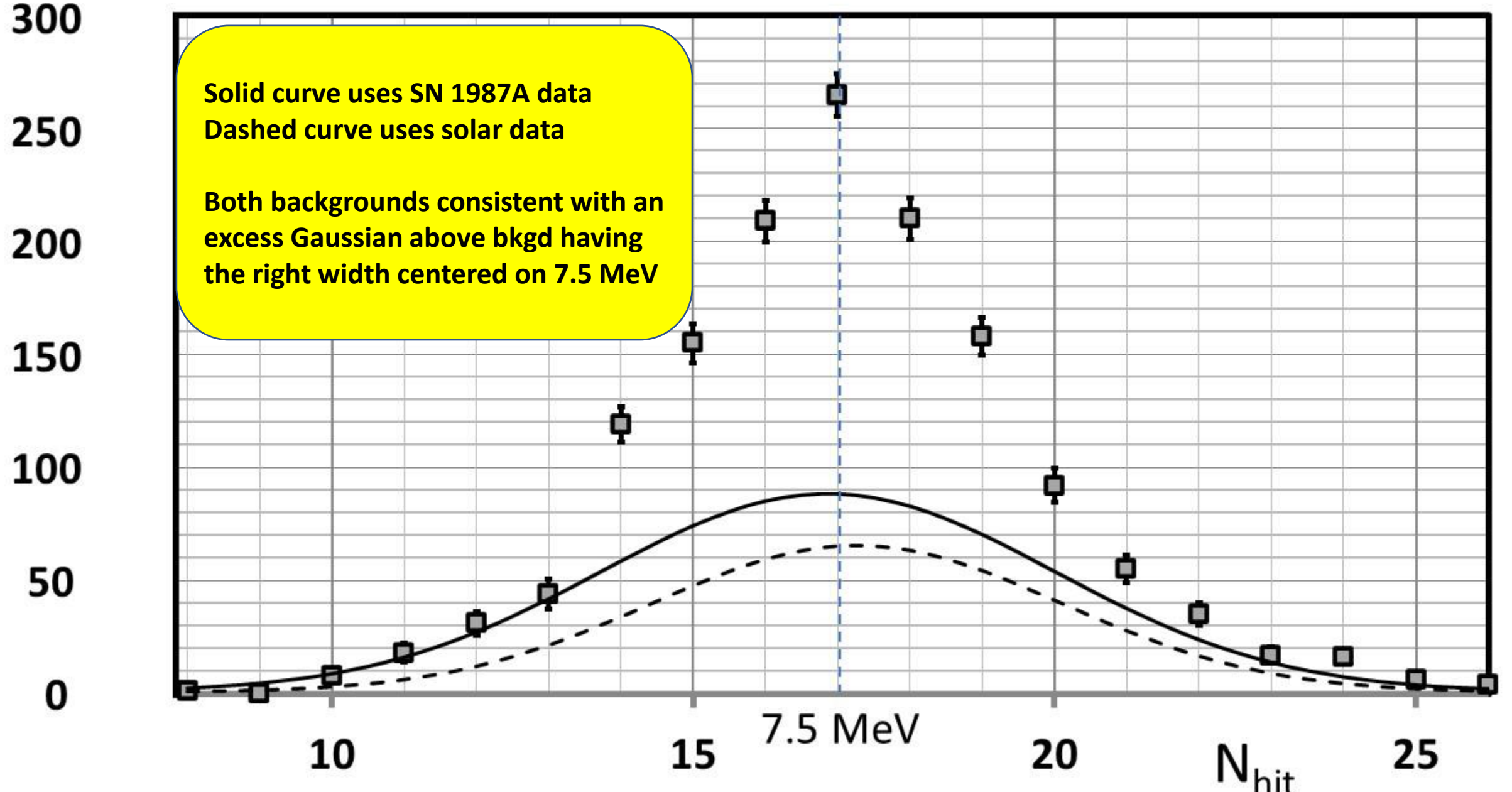
<https://arxiv.org/abs/2101.08128>

# Kamiokande-II data Feb 24, 1987A

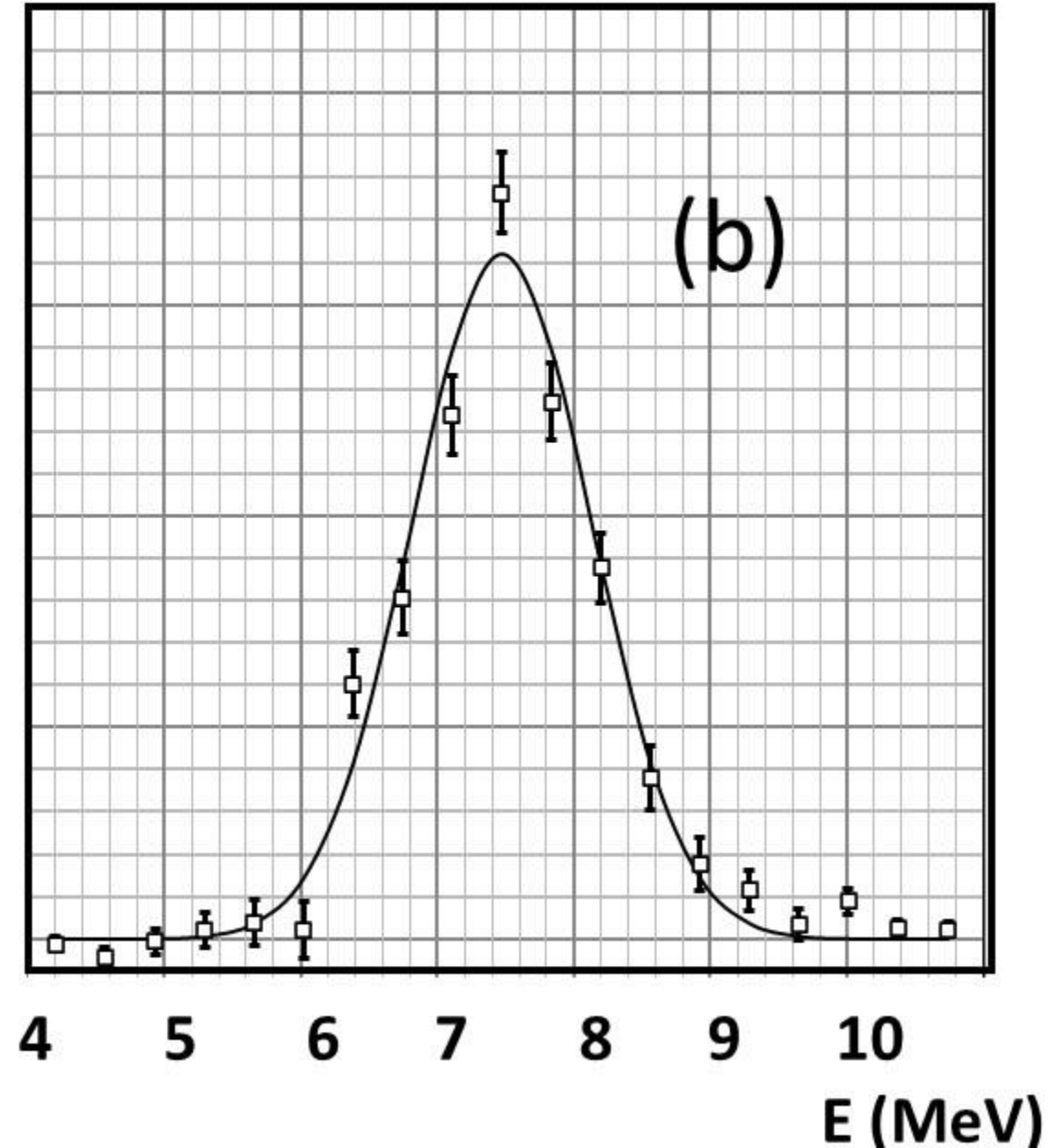
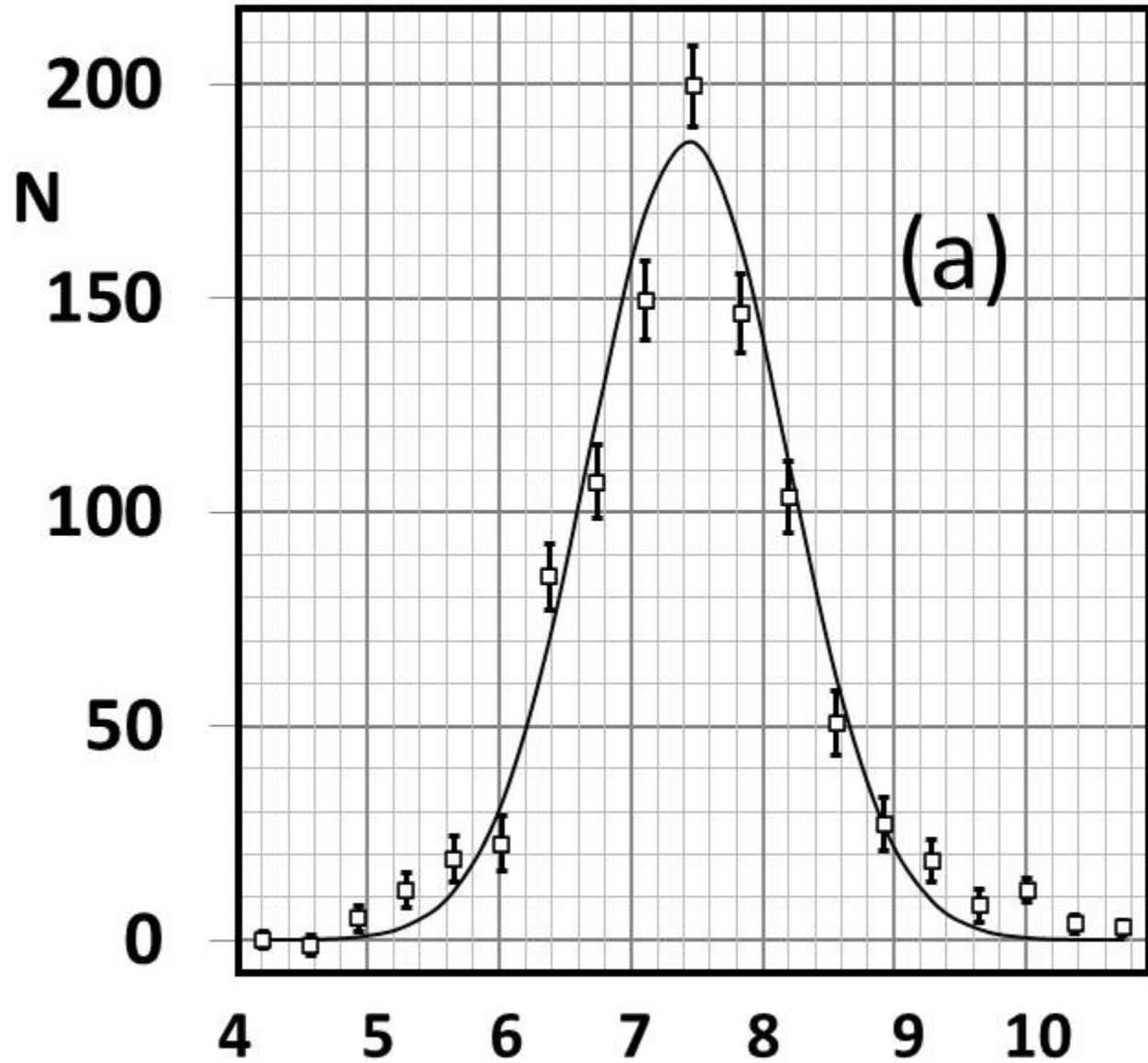


# events

Solid & dashed curves are two versions of the background



Background subtracted data consistent with a 7.5 MeV neutrino line.



# How to get an $\sim 8$ MeV neutrino line?

CDM annihilation of a DM particle of mass 8 MeV

$$XX \rightarrow Z' \rightarrow \nu\bar{\nu}$$

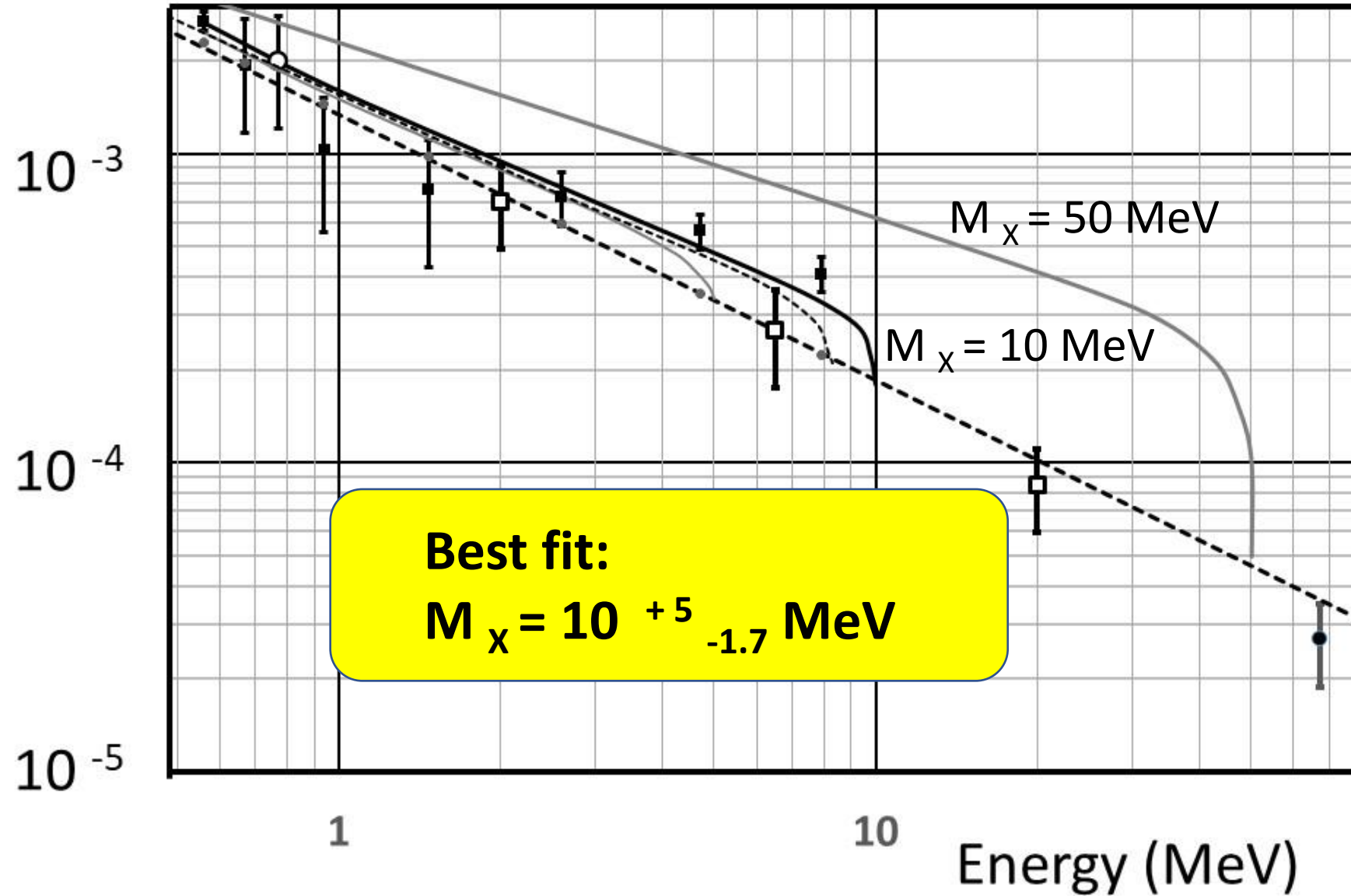
$Z'$  would be a 16 MeV mediator particle

-- the only other leptonic decay mode:  $Z' \rightarrow e^+e^-$

Evidence of 8 MeV  $X$  particle in Galactic Center gammas?

# Reason 1: Galactic center gamma ray spectrum

$E dF/dE$



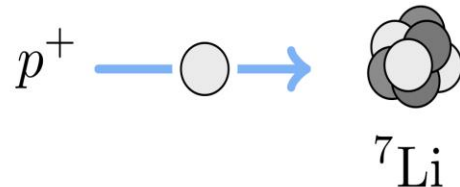
# Reason 2: The "Atomki experiment" suggests a new 17 MeV mass $Z'$ particle

Same anomaly seen in  $\text{He}^*$  decay

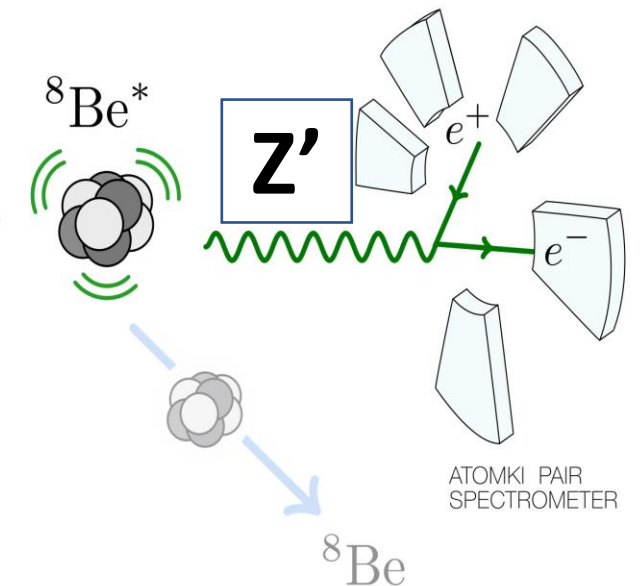
Existence of  $Z'$  Supported by B decay lepton and flavor anomaly and pion lepton decay anomaly

Would create 8.5 MeV neutrinos

$$Z' \rightarrow \nu \bar{\nu}$$



$$Z' \rightarrow e^+ e^-$$



# Reason 3: Light Dark matter in Supernovae

XX annihilation into leptons could provide the energy to prevent shock wave stalling.

Fayet et al. show  $m_\chi \sim 1-30$  MeV DM particles can play a significant role in the core-collapse.<sup>1</sup>

**Important** that if  $m_\chi < 10$  MeV the DM would cool on a time scale  $> 100$  times longer, i.e. hours not seconds.<sup>1</sup>

P. Fayet, D. Hooper, and G. Sigl, PRL 96, 211302 (2006).



# Reason 4: The ~5 hr early Mont Blanc SN 1987A burst

TABLE I – *Event number, time (UT), and preliminary visible energy (MeV) of the pulses in the burst detected on February 23rd, 1987.*

Event No.	Time (UT)	$E_\nu$ (MeV)
994	2 h 52 m 36 s.79	7
995	40.65	8
996	41.01	11
997	42.70	7
998	43.80	9

All 5 neutrinos have visible energy consistent with the average value 8 MeV

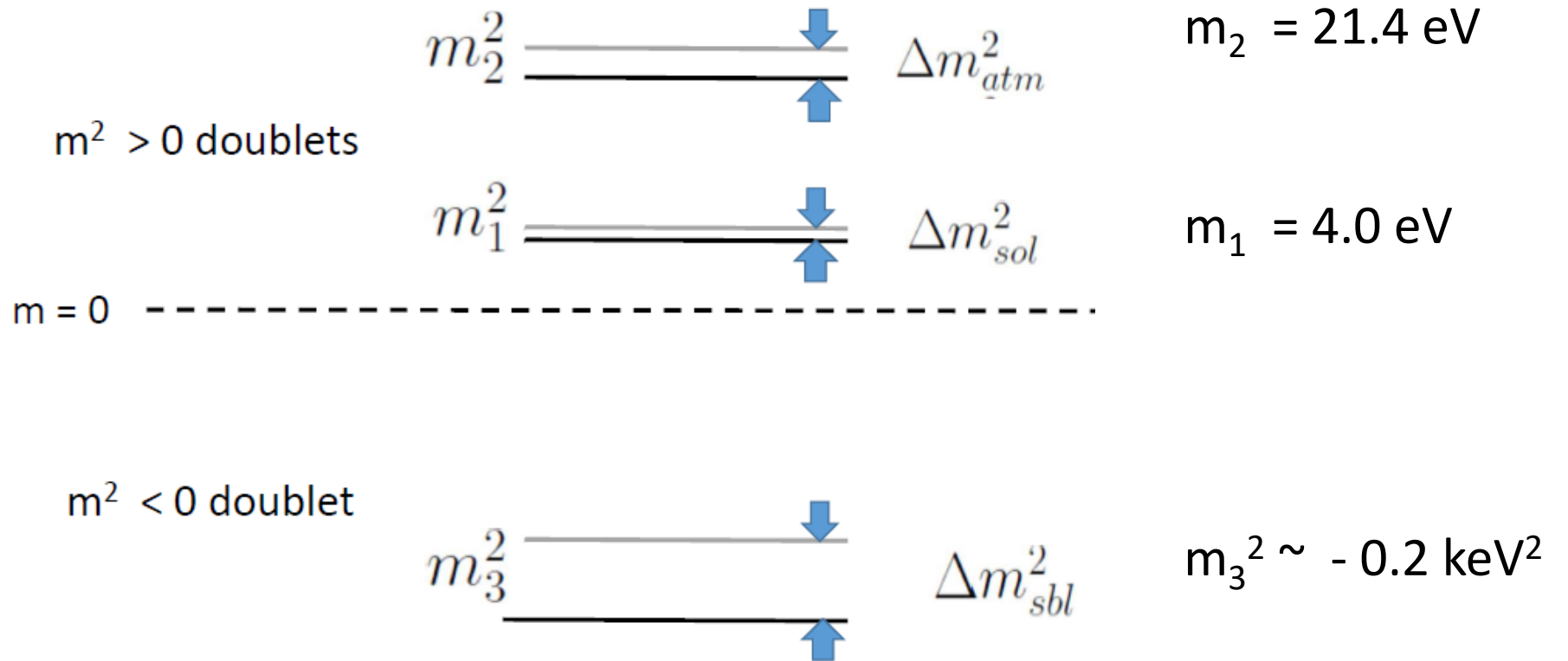
Neutrino energy would be 9.3 MeV (+/- 20%)

Aglietta explains why no such burst seen in Kamiokande detector

M. Aglietta et al., Europhys. Lett., **3** (12), 1321-1324 (1987).

# Reason 5: The "3 + 3" neutrino model (2013)<sup>1</sup>

3 active-sterile pairs having much larger masses than normally assumed



One pair are  
tachyons

If the Mont Blanc early burst really are tachyons, they **MUST** be monochromatic with energy  $E \sim 8 \text{ MeV}$ !

Ehrlich, R. "Tachyonic neutrinos and the neutrino masses," *Astropart. Phys.*, 41 (2013) 1-6,  
<http://arxiv.org/pdf/1204.0484.pdf>

# Possible contradictions to 8 MeV line

1. No time variation in strength of line seen over hours  
OK for DM which heats & cools very slowly
2. Impossible number of neutrinos above background ( $\sim 700$ )  
Barely possible given that DM models allow a substantial fraction of star being DM
3. Line not seen in diffuse supernova searches  
Those searches either had energy threshold  $> 8$  MeV or else looked for very short bursts

# Conclusions

There is evidence for an 8 MeV neutrino line from SN 1987A

May be able to see it in diffuse supernova searches using existing detectors

KATRIN will soon test the validity of  $3 + 3$  model