FREEZE-IN PRODUCED DARK MATTER IN THE ULTRA-RELATIVISTIC REGIME

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Beyond Standard Model - from Theory to Experiments March 30th, 2021

in collaboration with Jacopo Ghiglieri (arXiv 2012.09083, JCAP03(2021)075)

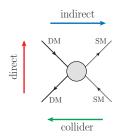




Department of Physics

MOTIVATION AND INTRODUCTION

FREEZE-IN PRODUCTION MECHANISM



- DM as a particle: many candidates see review G. Bertone 2016
- $\bullet\,$ non-interacting with photons, absolutely stable or long-lived $\sim \tau_{\rm Universe}$
- Any model has to comply with

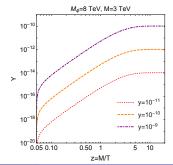
 $\Omega_{\rm DM} h^2(M_{\rm DM}, M_{\rm DM'}, \alpha_{\rm DM}, \alpha_{\rm SM}) = 0.1200 \pm 0.0012$

FREEZE-IN MECHANISM J. McDonald (2002)

- DM never reach thermal equilibrium
- DM from decay and/or annihilations of equilibrated species
- for a simple model $\mathcal{L}_{int} = -y\phi\bar{\chi}\chi$, $\phi \to \chi\chi$

$$rac{dn_{\chi}}{dt} + 3Hn_{\chi} = \langle \Gamma_{\phi
ightarrow \chi \chi}
angle n_{\phi}^{
m eq} , \ Y = n_{\chi}/s$$

• $\Omega_{\text{DM}} h^2 = \frac{M}{\text{GeV}} \frac{Y_{\text{fin}}}{3.645 \times 10^{-9}}$



MOTIVATION AND INTRODUCTION

FRAMING THE PARTICLE PHYSICS MODEL

• Simplified DM models:

 \Rightarrow capture the d.o.f. and parameters needed to study DM phenomenology

- χ Majorana fermion singlet, $\chi \equiv$ DM particle
- η is charged under QCD and U(1)_Y, $\eta \equiv$ mediator with $M_{\eta} = M + \Delta M$

$$\mathcal{L} = \mathcal{L}_{SM} + \frac{1}{2} \bar{\chi} \left(i \partial \!\!\!/ - M \right) \chi + (D_{\mu} \eta)^{\dagger} D^{\mu} \eta - M_{\eta}^{2} \eta^{\dagger} \eta - \lambda_{2} (\eta^{\dagger} \eta)^{2} - \lambda_{3} \eta^{\dagger} \eta \phi^{\dagger} \phi - y \eta^{\dagger} \bar{\chi} a_{R} q - y^{*} \bar{q} a_{L} \chi \eta$$

same model and freeze-in see M. Garny and J. Heisig (2018) and G. Bélanger et al (2018)

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WITHIN THE FREEZE-IN $y \leq \mathcal{O}(10^{-8})$ and for T > 150 GeV

- we shall address the high-temperature dynamics
- $\bullet\,$ multiple soft scatterings and $2\to 2$ process

 $T \gg M_{\eta}$ can be very important even for renormalizable interactions

PRODUCTION RATE AND RATE EQUATION

GENERAL APPROACH

- Given a field χ weakly coupled to a an equilibrated bath, with internal couplings g
 - [T. Asaka, M. Laine and M. Shaposhnikov (2006), M. Laine and A. Vuorinen (2017), D. Bödeker, M. Sangel and M. Wörmann (2016)]
- at leading order in y and all orders in g one can prove D. Bödeker, M. Sangel and M. Wörmann (2016)

$$\dot{f}_{\chi}(t,\boldsymbol{k}) = \Gamma(k)[n_{\mathrm{F}}(k^{0}) - f_{\chi}(t,\boldsymbol{k})], \quad \Gamma(k) = \frac{|y|^{2}}{2k^{0}} \int d^{4}X \, e^{iK \cdot X} \langle [J(X), J(0)] \rangle$$

• $f_{\chi}(t, \mathbf{k})$ is the single-particle phase-space distribution; J made of bath fields



$$\Gamma(k) = \frac{|y|^2}{k^0} \mathrm{Im} \Pi_R$$

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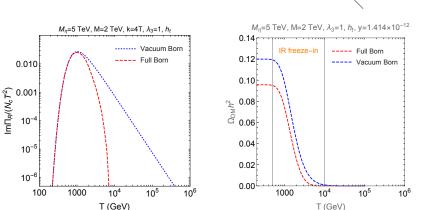
- when doing perturbative expansions ⇒ Boltzmann equation is recovered
- general framework to include: resummation and NLO computations, non-perturbative and thermal effects

IN-VACUUM VERSUS FINITE THERMAL MASSES

• Born rate $\eta \rightarrow \chi + q$ with and without thermal masses (recall m_q is purely thermal)

for thermal masses see also L. Darmé, A. Hryczuk, D. Karamitros and L. Roszkowski (2019)

$$\dot{n}_{\text{DM}} + 3Hn_{\text{DM}} = 2|y|^2 \int_k \frac{n_{\text{F}}(k^0)}{k^0} \text{Im}\Pi_R^{\text{Born}}$$



HIGH-TEMPERATURES $\pi T \gg M_{\eta}$

- all the particles are seen as massless,
- momenta of external particles $p \sim \pi T$,
- particles are ultra-relativistic, and gT is a soft scale

 \Rightarrow collinear kinematics \approx high T



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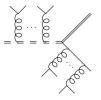
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n soft scatterings: LPM resummation

L. Landau and I. Pomeranchuk (1953) and A. B. Migdal (1956) see e.g. J. Ghiglieri and G. D. Moore (2014) for a review

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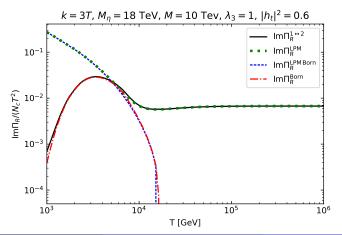
• at $T \gg M_\eta$ three effective processes contribute to the production of χ

 $\begin{array}{c|c} \eta \rightarrow \chi + q, \quad q \rightarrow \chi + \eta, \quad q + \eta \rightarrow \chi \end{array}$

LPM RESULTS

• prescription for any temperature (inspired from I. Ghisoiu and M. Laine (2014))

$$\mathrm{Im}\Pi_{R}^{1\leftrightarrow2} = \mathrm{Im}\Pi_{R}^{\mathrm{LPM}} - \mathrm{Im}\Pi_{R}^{\mathrm{LPM Born}} + \mathrm{Im}\Pi_{R}^{\mathrm{Born}}$$



$2 \rightarrow 2$ scatterings



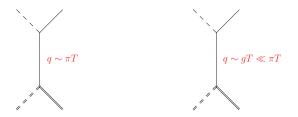
- Considered by M. Garny and J. Heisig (1809.10135) for T ≤ M (possibly some issues with IR of some processes)
- for s/t and u/t contributions from **both hard and soft momentum regions** D. Besak and D. Bodeker (2012), J. Ghiglieri and M. Laine (2016)

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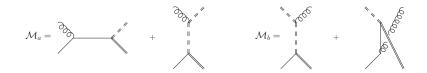


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$$\begin{split} \mathrm{Im} \Pi_{R}^{2\leftrightarrow 2} &= \frac{2}{(4\pi)^{3}k} \int_{k}^{\infty} \mathrm{d}q_{+} \int_{0}^{k} \mathrm{d}q_{-} \Big\{ \big[n_{\mathrm{F}}(q_{0}) + n_{\mathrm{B}}(q_{0}-k) \big] N_{c} \big(Y_{q}^{2} g_{1}^{2} + C_{F} g_{3}^{2} + |h_{q}|^{2} \big) \Phi_{s2} \Big\} \\ &+ \frac{2}{(4\pi)^{3}k} \int_{0}^{k} \mathrm{d}q_{+} \int_{-\infty}^{0} \mathrm{d}q_{-} \Big\{ \big[1 - n_{\mathrm{F}}(q_{0}) + n_{\mathrm{B}}(k-q_{0}) \big] N_{c} \big(Y_{q}^{2} g_{1}^{2} + C_{F} g_{3}^{2} + |h_{q}|^{2} \big) \Phi_{t2} \\ &- \Big[n_{\mathrm{B}}(k) + \frac{1}{2} \Big] N_{c} \big(Y_{q}^{2} g_{1}^{2} + C_{F} g_{3}^{2} + |h_{q}|^{2} \big) \frac{k\pi^{2} T^{2}}{q^{2}} \Big\} + N_{c} \frac{m_{q}^{2}}{16\pi} \Big[n_{\mathrm{B}}(k) + \frac{1}{2} \Big] \ln \left(1 + \frac{4k^{2}}{m_{q}^{2}} \right) \end{split}$$

Ultra-relativistic regime: LPM and 2 \rightarrow 2

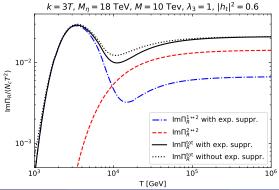
SUMMARY OF THE RATES

• Phenomenological switch off the high-temperature processes (lack of NLO rates)

$$\kappa(M_{\eta}) = rac{3}{\pi^2 T^3} \int_0^\infty dp \, p^2 \, n_{
m B}(E_{\eta}) [1 + n_{
m B}(E_{\eta})] \,,$$

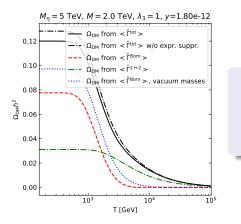
$$\Pi_R^{\rm tot} = {\rm Im} \Pi_R^{1\leftrightarrow 2} + {\rm Im} \Pi_R^{2\leftrightarrow 2}$$

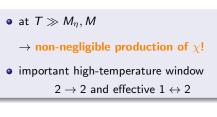
• $\operatorname{Im}\Pi_{R}^{1\leftrightarrow 2} = (\operatorname{Im}\Pi_{R}^{\operatorname{LPM}} - \operatorname{Im}\Pi_{R}^{\operatorname{LPM}\operatorname{Born}})\kappa(M_{\eta}) + \operatorname{Im}\Pi_{R}^{\operatorname{Born}}$



Ultra-relativistic regime: LPM and 2 \rightarrow 2

Benchmark point P0





- Born rate with vacuum masses \Rightarrow 20% reduction of $\Omega_{\rm DM} h^2$ with respect to $\Pi_R^{\rm tot}$
- $\bullet~$ 30% when including thermal masses but excluding 2 \rightarrow 2 and effective 1 \leftrightarrow 2
- estimation of theoretical error: LPM with and without $\kappa(M_\eta)$, here $\sim 10\%$ effect

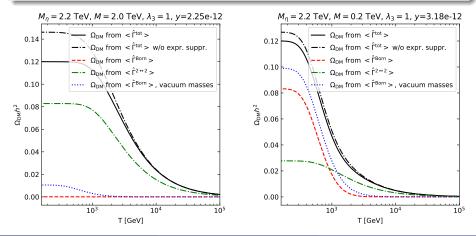
Ultra-relativistic regime: LPM and 2 \rightarrow 2

LARGE AND SMALL MASS SPLITTINGS

• the smaller $\Delta M/M$ the larger the effect of thermal masses, LPM and 2 \rightarrow 2

Left plot: $\Delta M/M = 0.1$; Right plot: $\Delta M/M = 10$

• also other parameters of the model are relevant (h_q, λ_3)



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SUMMARY

- we studied the impact of the ultra-relativistic regime on the production of a feebly interacting DM particle
- Before: in renormalizable models bulk DM population produced at $T \sim M$
- Our work: this is not always the case high-temperature 1 ↔ 2, 2 → 2 can give O(1) contribution

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- simplified dark matter model: χ Majorana fermion DM and η mediator charged under SU(3) \otimes U(1)_Y
- freeze-in with large impact from 1 \leftrightarrow 2, 2 \rightarrow 2

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• similar effects can affect other models if DM comes from particles in equilibrium

• Our main uncertainty comes from the lack of NLO rates (state-of-art M.Laine (2013))

THE BORN RATE WITH VANISHING THERMAL MASSES

• Let us look at the model at hand

$$\begin{pmatrix} \frac{\partial}{\partial t} - Hk_i \frac{\partial}{\partial k_i} \end{pmatrix} f_{\chi}(t, \mathbf{k}) = \Gamma(k) [n_{\rm F}(k^0) - f_{\chi}(t, \mathbf{k})],$$

$$\Gamma(k) = \frac{|y|^2}{k^0} {\rm Im} \Pi_R = \frac{|y|^2}{2k^0} {\rm Tr} \left\{ \not\!\!{k} a_R \left[\rho(\mathcal{K}) + \rho(-\mathcal{K}) \right] a_L \right\},$$

• Retarded correlator, Euclidean correlator and spectral function are connected

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$$\Pi^{E}(K) \equiv \operatorname{Tr}\left\{i \not k \left[\int_{X} e^{iK \cdot X} a_{R} \langle (\eta^{\dagger} q)(X)(\bar{q}\eta)(0) \rangle a_{L}\right]\right\}$$
$$= N_{c} \int_{p} T \sum_{n} \frac{-i \not P a_{L}}{p_{n}^{2} + E_{q}^{2}} \frac{1}{(p_{n} + k_{n})^{2} + E_{\eta}^{2}}$$

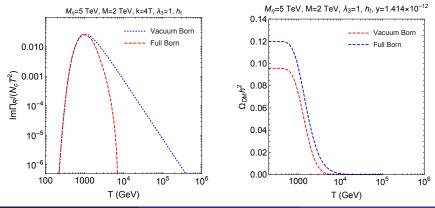
ullet with $E_q = |oldsymbol{p}| = p$ and $E_\eta = \sqrt{(oldsymbol{p}+oldsymbol{k})^2 + M_\eta^2}$

see M. Laine and A. Vuorinen (2017)

IN-VACUUM VERSUS FINITE THERMAL MASSES

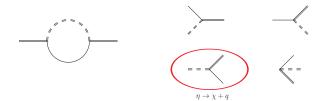
• Scalar mass:
$$\mathcal{M}_\eta^2=\mathcal{M}_\eta^2+m_\eta^2$$
, for $\eta o\chi+q$ and $E_p=\sqrt{p^2+m_q^2}$

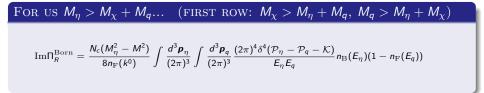
$$\mathrm{Im}\Pi^{\mathrm{Born}}_{\mathrm{R},\eta\to\chi q} = \frac{N_c}{16\pi k} \int_{\rho_{\mathrm{min}}}^{\rho_{\mathrm{max}}} dp [\mathcal{M}_{\eta}^2 - M^2 - m_q^2 - 2k^0 (E_p - p)] [n_{\mathrm{B}}(k^0 + E_p) + n_{\mathrm{F}}(E_p)]$$



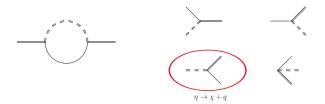
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BORN TERM AND BOLTZMANN EQUATION





BORN TERM AND BOLTZMANN EQUATION



For us $M_{\eta} > M_{\chi} + M_{q}$... (first row: $M_{\chi} > M_{\eta} + M_{q}, M_{q} > M_{\eta} + M_{\chi}$)

$$\mathrm{Im}\Pi_{R}^{\mathrm{Born}} = \frac{N_{c}(M_{\eta}^{2} - M^{2})}{8n_{\mathrm{F}}(k^{0})} \int \frac{d^{3}\boldsymbol{p}_{\eta}}{(2\pi)^{3}} \int \frac{d^{3}\boldsymbol{p}_{q}}{(2\pi)^{3}} \frac{(2\pi)^{4}\delta^{4}(\mathcal{P}_{\eta} - \mathcal{P}_{q} - \mathcal{K})}{E_{\eta}E_{q}} n_{\mathrm{B}}(E_{\eta})(1 - n_{\mathrm{F}}(E_{q}))$$

•
$$n_{\mathrm{DM}}=2\int_{\pmb{k}}f_{\chi}(t,\pmb{k}), \ \text{for} \ \eta o \chi+q$$

$$\begin{split} \dot{n}_{\rm DM} + 3H n_{\rm DM} &= 2|y|^2 \int_k \frac{n_{\rm F}(k^0)}{k^0} {\rm Im} \Pi_R \\ &= 2|y|^2 N_c (M_\eta^2 - M^2) \int_{\pmb{p}_\eta, \pmb{p}_q, \pmb{k}} \frac{(2\pi)^4 \delta^4 (\mathcal{P}_\eta - \mathcal{P}_q - \mathcal{K})}{8E_\eta E_q \, k^0} n_{\rm B}(E_\eta) \left[1 - n_{\rm F}(E_q)\right] \end{split}$$

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FIRST IMPROVEMENT: THERMAL MASSES

 at high temperatures, πT ≫ M_η, repeated interactions within the bath change the dispersion relations ⇒ asymptotic masses

see also L. Darmé, A. Hryczuk, D. Karamitros and L. Roszkowski (2019)

• for $T > T_c \simeq 150$ GeV the quarks only have

$$m_q^2 = \frac{T^2}{4} (g_3^2 C_F + Y_q^2 g_1^2 + |h_q|^2)$$

• for the colored scalar

$$m_{\eta}^{2} = \left(\frac{g_{3}^{2}C_{F} + Y_{q}^{2}g_{1}^{2}}{4} + \frac{\lambda_{3}}{6}\right)T^{2}$$

• no thermal mass correction for χ since $y \ll g$

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$$m_{\rm B}^2 = \left(\frac{n_{\rm S}}{6} + \frac{5n_{\rm G}}{9} + \frac{Y_q^2 N_c}{3}\right) g_1^2 T^2 , \quad m_g^2 = \left(\frac{N_c}{3} + \frac{n_{\rm G}}{3} + \frac{1}{6}\right) g_3^2 T^2$$

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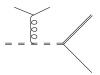
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$$m_{\eta}^{2}(M_{\eta}, p) = \frac{Y_{q}^{2}g_{1}^{2} + C_{F}g_{3}^{2}}{2\pi^{2}} \int_{0}^{\infty} dq \left\{ \frac{n_{\rm B}(E_{\eta}(q))}{E_{\eta}(q)} \left[q^{2} + \frac{M_{\eta}^{2}q}{2p} \ln\left(\frac{(p+q)^{2}}{(p-q)^{2}}\right) \right] + 2 q n_{\rm B}(q) \right\}$$
$$+ \frac{\lambda_{3}}{\pi^{2}} \int_{0}^{\infty} dq q n_{\rm B}(q)$$

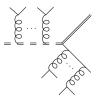
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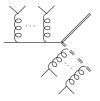


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$$m_g^2 = \left(\frac{N_c}{3} + \frac{n_G}{3} + \frac{1}{6}\right)g_3^2T^2$$

Long formation time $\sim \frac{1}{g^2 T}$ see e.g. J. Ghiglieri and G. D. Moore (2014)

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- all the particles are seen as massless,
- all the momenta of extarnal particles are $p \sim \pi T$,
- particles are ultra-relativistic, and gT is a soft scale
 ⇒ collinear kinematics ≈ high T



- connection between LPM resummation and $\text{Im}\Pi_R^{\text{LPM}} \Rightarrow \chi$ self-energy
- notation and computational setting from

$$\hat{H}\equiv-rac{M^2}{2k_0}+rac{m_q^2-
abla_{\perp}^2}{2E_q}+rac{\mathcal{M}_{\eta}^2-
abla_{\perp}^2}{2E_\eta}+i\Gamma(y)\,,\quad y\equiv|m{y}_{\perp}|$$

$$\Gamma(y) = \frac{T}{2\pi} g_1^2 Y_q^2 \left[\ln\left(\frac{m_{\rm B}y}{2}\right) + \gamma_E + K_0(m_{\rm B}y) \right] + \frac{T}{2\pi} g_3^2 C_F \left[\ln\left(\frac{m_g y}{2}\right) + \gamma_E + K_0(m_g y) \right]$$

NUMERICAL STRATEGY AND LPM BORN

• \hat{H} enters the inhomogeneous equations for the functions g(y) and f(y) $(\hat{H} + i0^+)g(y) = \delta^{(2)}(y)$, $(\hat{H} + i0^+)f(y) = -\nabla_{\perp}\delta^{(2)}(y)$

$$\begin{split} \mathrm{Im} \Pi_{\mathrm{R}}^{\mathrm{LPM}} &= -\frac{N_c}{8\pi} \int_{-\infty}^{+\infty} dE_q \int_{-\infty}^{+\infty} dE_\eta \, \delta(k_0 - E_q - E_\eta) [1 - n_{\mathrm{F}}(E_q) + n_{\mathrm{B}}(E_\eta)] \\ & \frac{k_0}{E_\eta} \lim_{\mathbf{y} \to 0} \left\{ \frac{M^2}{k_0^2} \mathrm{Im}[g(\mathbf{y})] + \frac{1}{E_q^2} \mathrm{Im}[\nabla_\perp \cdot f(\mathbf{y})] \right\} \end{split}$$

NUMERICAL STRATEGY AND LPM BORN

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Once $E_{\eta} = k^0 - E_q$ by the δ

- 2 $E_q < 0$: this corresponds to the effective $1 \rightarrow 2$ process $\eta \rightarrow q\chi$
 - \Rightarrow LPM-Born with thermal masses is the n = 0 limit (no scatterings)
- **(a)** $E_q > k^0$: this corresponds to the effective $1 \to 2$ process $q \to \eta \chi$.

DM ENERGY DENSITY AND SUPER-WIMP CONTRIBUTION

$$(\Omega_{
m DM}h^2)_{
m obs.} = (\Omega_{
m DM}h^2)_{
m freeze-in} + (\Omega_{
m DM}h^2)_{
m super-WIMP}$$

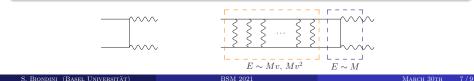
- η particles stays in chemical equilibrium till late times $T \sim M_\eta/25$
- there is a populations of $\eta,$ as governed by freeze-out, which decays into χ $_{\rm M.~Garny,~J.~Heisig~(2018)}$

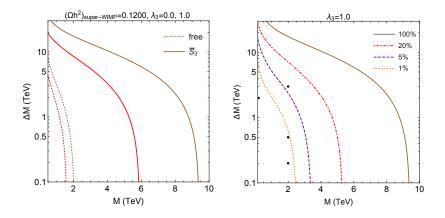
$$(\Omega_{ ext{DM}} h^2)_{ ext{super-WIMP}} = rac{M}{M_\eta} (\Omega h^2)_\eta \,.$$

. .

$$rac{dn_\eta}{dt} + 3Hn_\eta = -\langle \sigma_{
m eff} v
angle ig(n_\eta^2 - n_{\eta,
m eq}^2 ig) \,, \quad \langle \sigma_{
m eff} v
angle = rac{c_3 ar{S}_3 + c_4 C_{
m F} ar{S}_4}{N_c}$$

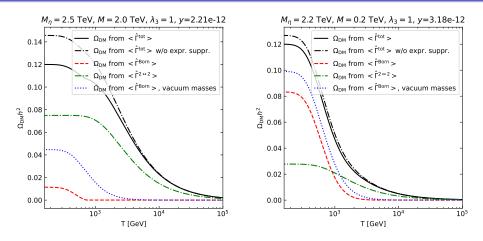
A. Mitridate et al (2017), S. B. and M. Laine (2018), S. B. and S. Vogl (2019), J. Hartz and K. Petraki (2018)





• P0 (M = 2.0 Tev, $M_{\eta} = 5.0$ TeV); P1 (M = 0.2 Tev, $M_{\eta} = 2.2$ TeV); P2 (M = 2.0 Tev, $M_{\eta} = 2.5$ TeV); P3 (M = 2.0 Tev, $M_{\eta} = 2.2$ TeV);

LARGE AND SMALL MASS SPLITTINGS II



- for $M_{\eta} = 2.2$ TeV the freeze-in production stops fairly close to $T_c \simeq 150$ GeV \Rightarrow follow DM production in the SM broken phase
- ullet CMS analysis provides us with $M_\eta > 1250$ GeV cms-pas-exo-16-036