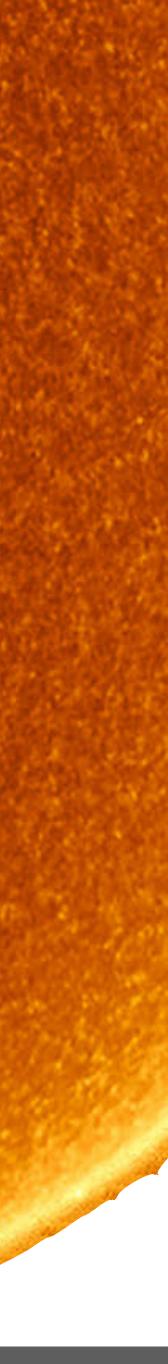
# Solar WIMP Search with the IceCube Neutrino Observatory

Jeffrey Lazar Beyond Standard Model: From Theory to Experiment Zewail City of Science and Technology







# Outline

- Background
- $\chi aro\nu$  software
- Current analysis





# DM Capture and Indirect Detection



XXX



# Two Ways to Measure $\sigma_{N\gamma}$

## Annihilation rate $-C_A N^2$

= 0

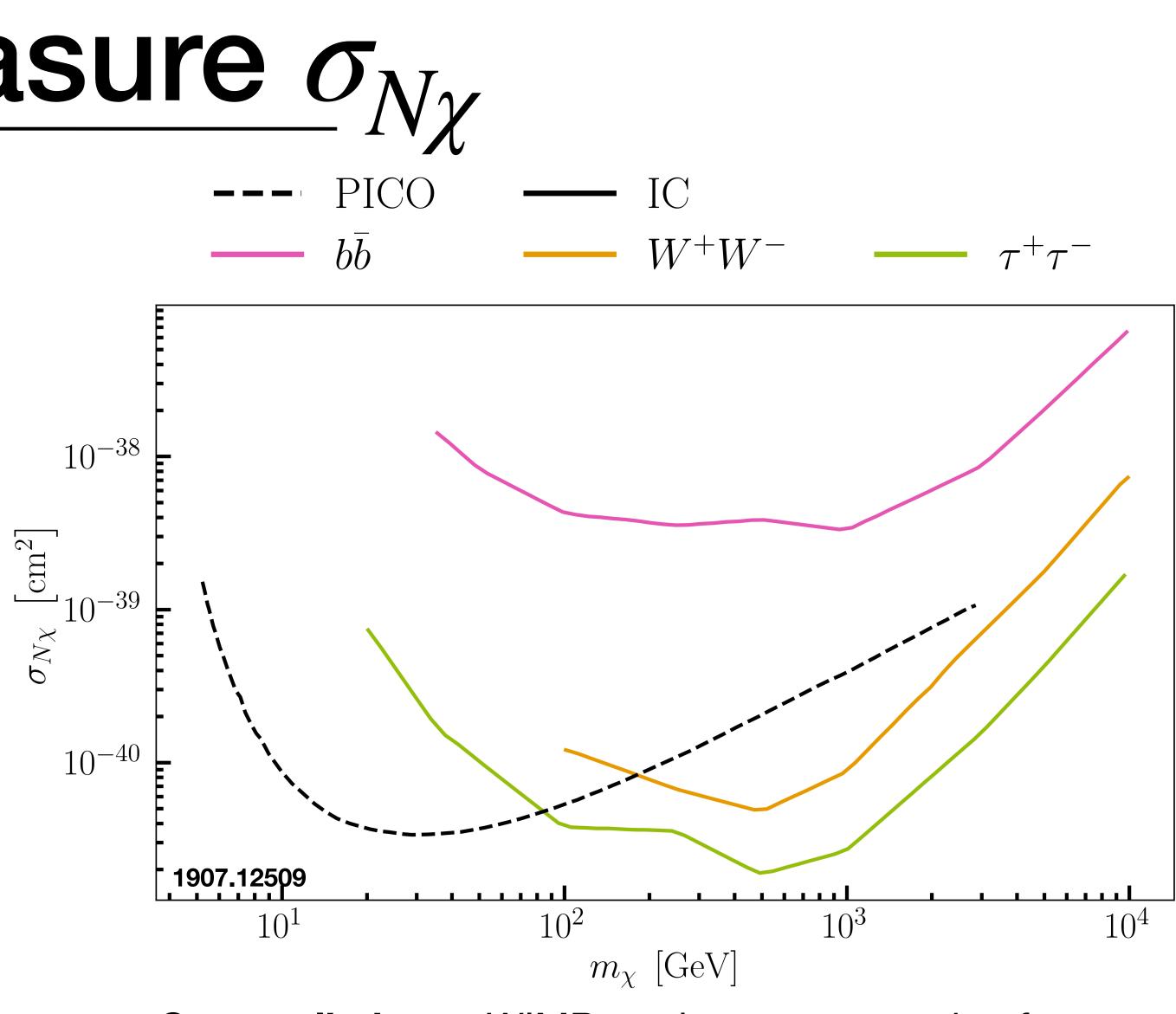
Capture rate: Proportional to  $\sigma_{\gamma N}$ 

Annihilation rate: Sets rate of WIMP conversion to neutrinos

**Evaporation rate: Negligible for WIMP** masses above a few GeV

 $\implies \frac{1}{2} = C_C \propto \sigma_{\chi N}$ 



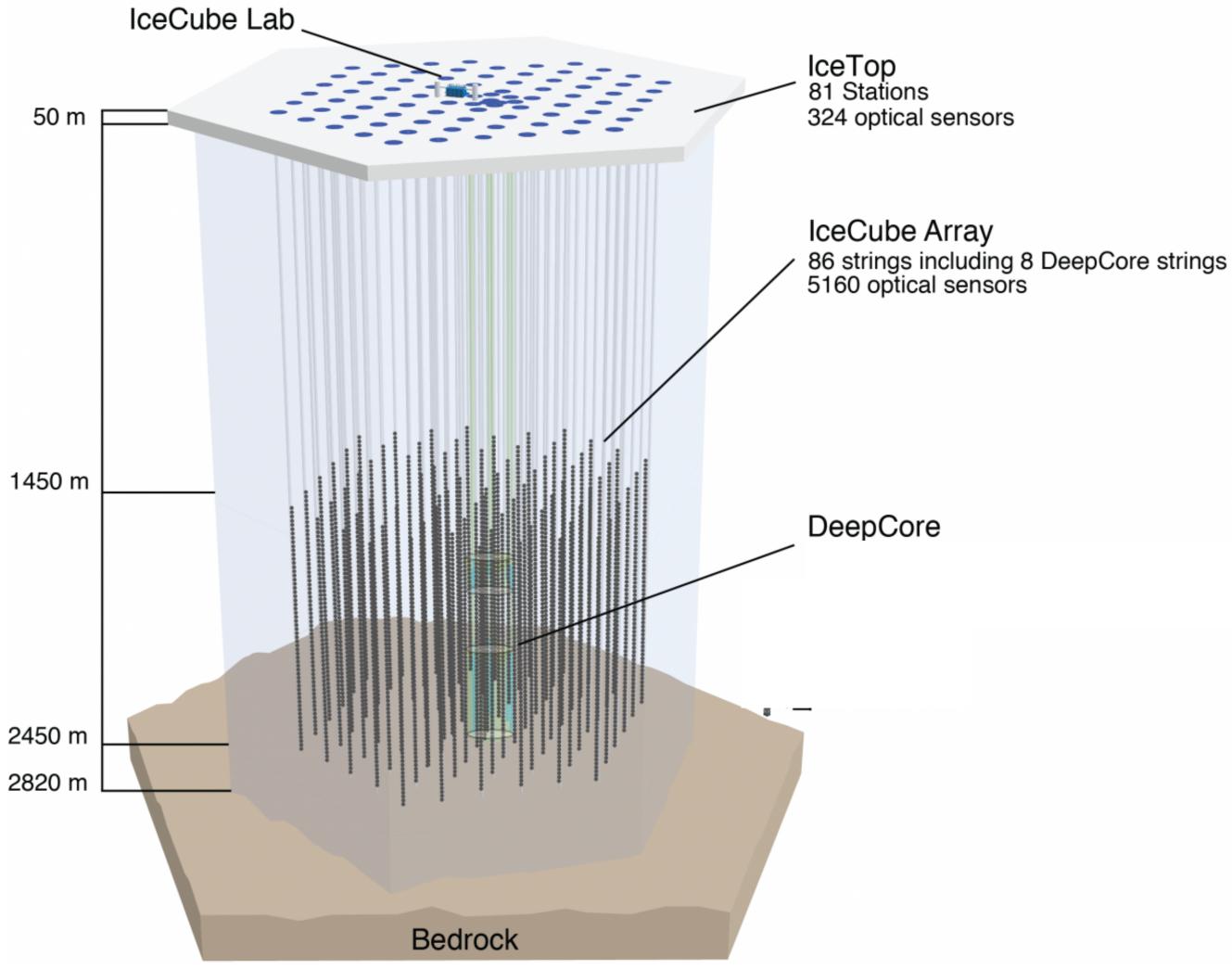


**Current limits** on WIMP-nucleon cross section from IceCube and PICO





# The IceCube Neutrino Observatory





- 1 km<sup>3</sup> instrumented ice at the geographic south pole
- Digital optical module (DOM) detects light created by charged byproducts of neutrino interaction



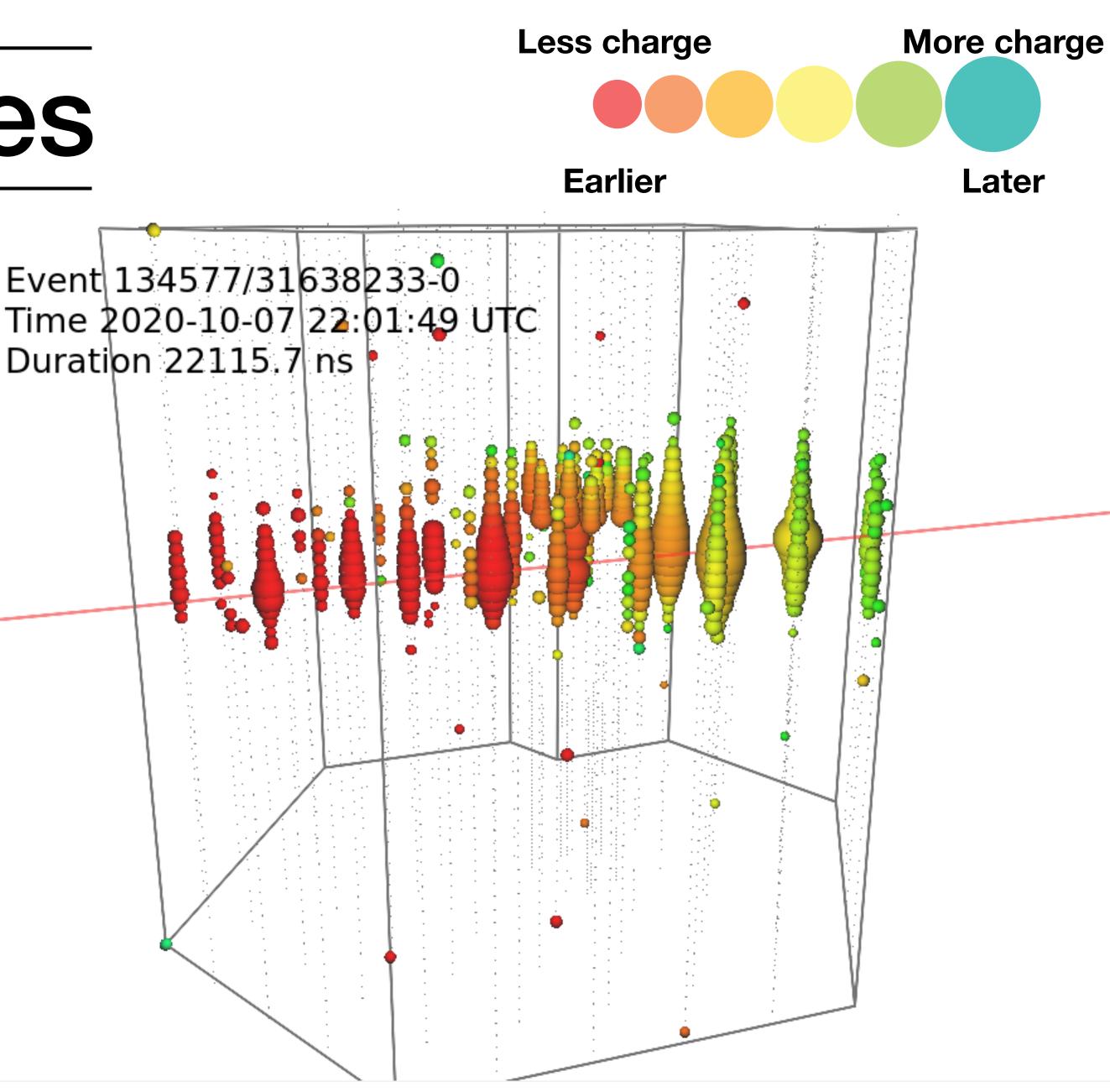


# **Event Morphologies**

Track

 $u_{\mu} CC$ 





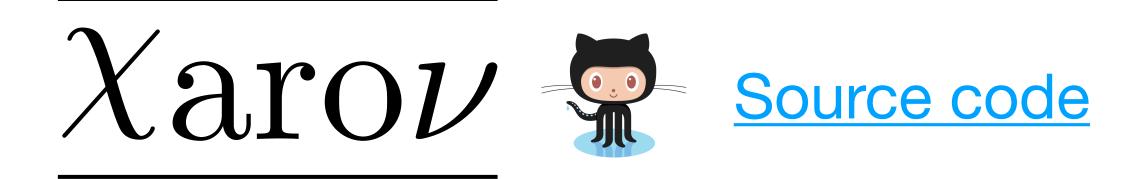


# Outline

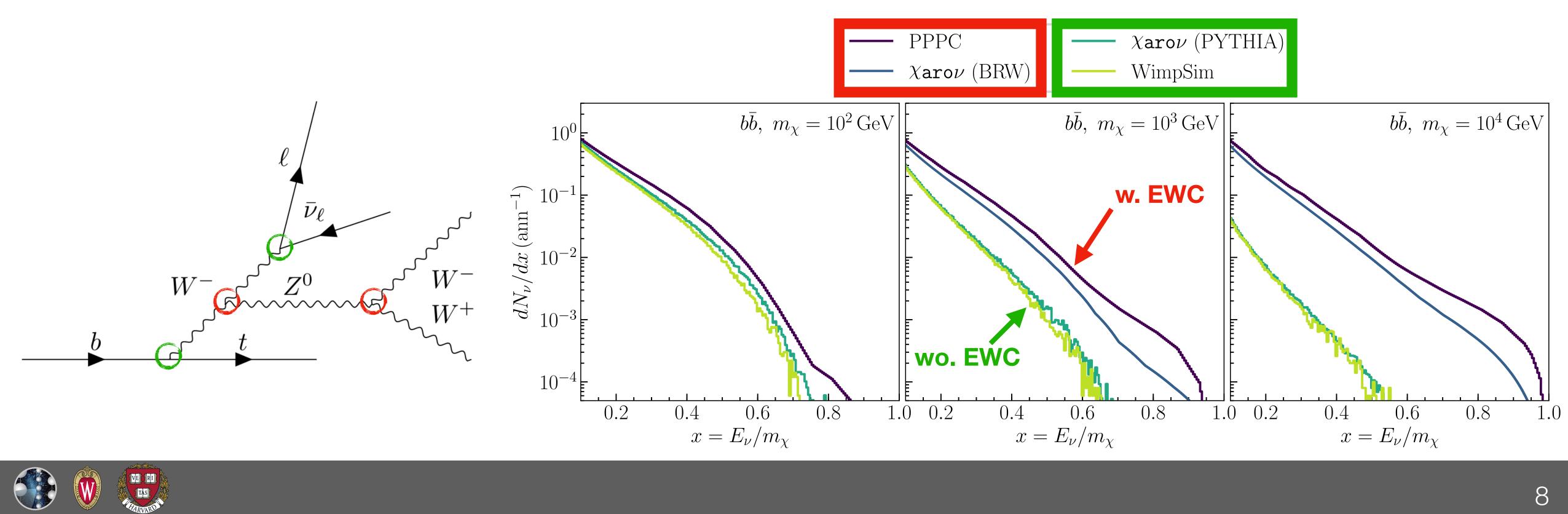
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- Software package for calculating neutrino yields from DM annihilation/decay. arXiv:2007:15010
- Couples PYTHIA8 to an updated calculation of EW correction (BRW calculation). arXiv:2007:15001







# Flux Generator Comparison

	Generation	EW corrections	Secluded DM	Long-lived particle stopped decay	Locations	Flux production	Propagation
WimpSim	PYTHIA 6.4				Earth, Sun	Read files or run Fortran scripts	Read files or run Fortran scripts with oscillation parameters
PPPC	PYTHIA 8.1 (+ GEANT4)				Galactic Halo, Sun	Read table in Mathematica	Read table in Mathematica
Charon wo/ BRW	PYTHIA 8.2				Galactic Halo, Sun, Earth or custom environment	Read table or run C++ script	Flexible propagation wit nuSQuIDs by allowing
Charon w/ BRW	DGLAP + PYTHIA 8.2				Galactic Halo, Sun, Earth	Read table	options of input fluxes, oscillation parameters, xsec







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### Improved All-Energy Analysis IceCube + DeepCore to cover WIMP mass range from 10 GeV to 10 TeV

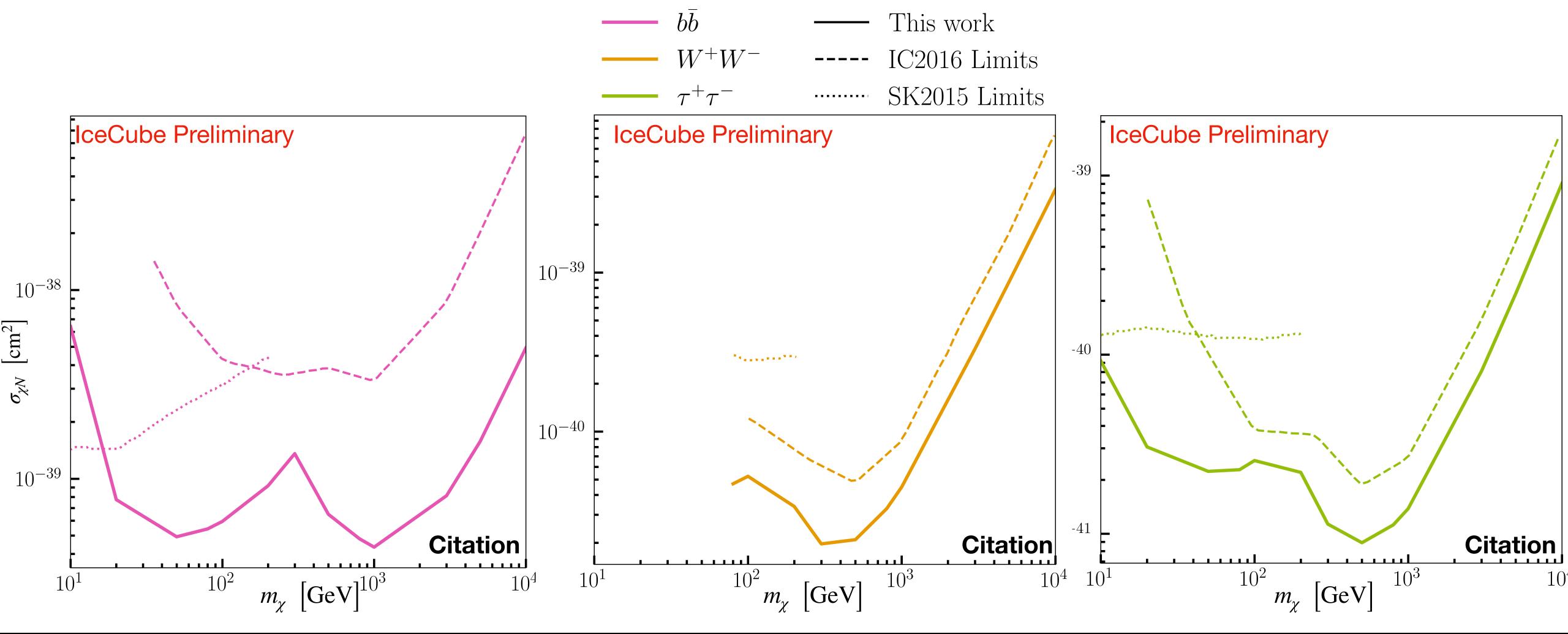
- Directional reconstruction challenging at lower-energies
- Cascade backgrounds 10x lower—> Include all flavors in analysis
- Developing new event selection to target SUSY WIMP region of interest

	mulation of 200 eV electron				ი თე <del>თ</del> იფიფიდითაფიადიი გ <i>იწიწი</i> ციციციციკიკი ი	° ° ° ° ° ° ° ° ° ° ° ° ° ° ° ° ° ° °
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	Simulation of 200 GeV muon

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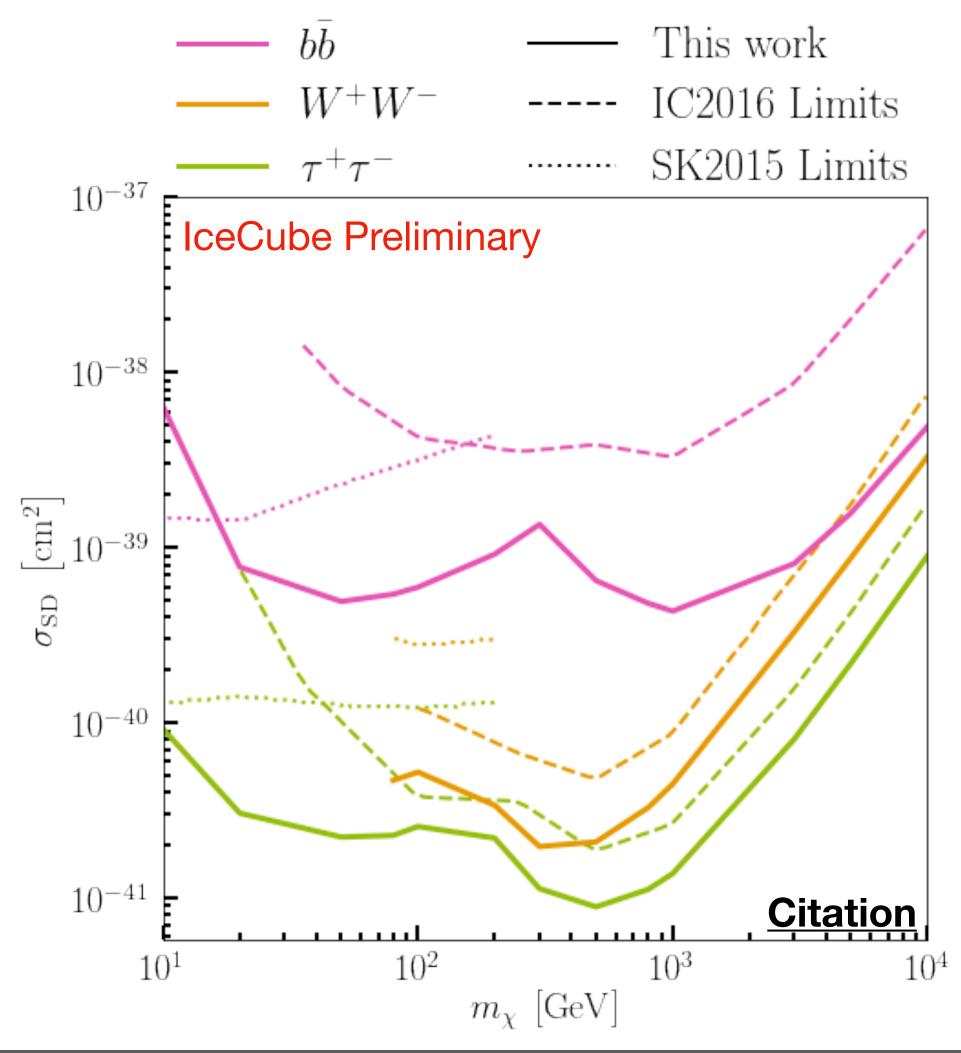
# Ten Years IC+DeepCore Sensitivity







# Status and Outlook



- World-leading sensitivities for almost entire range
- Currently working to further improve ~100 GeV range
- Stay tuned: results coming soon



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# Thank you to the Organizers and thank you for listening

## Question ??





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# Back Ups







# Rough Comparison to DD

