



Neutrino oscillations and CPT violation

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**Beyond Standard Model:
From Theory to Experiment
(BSM-2021)**

BSM
March 29-
April 2, 2021

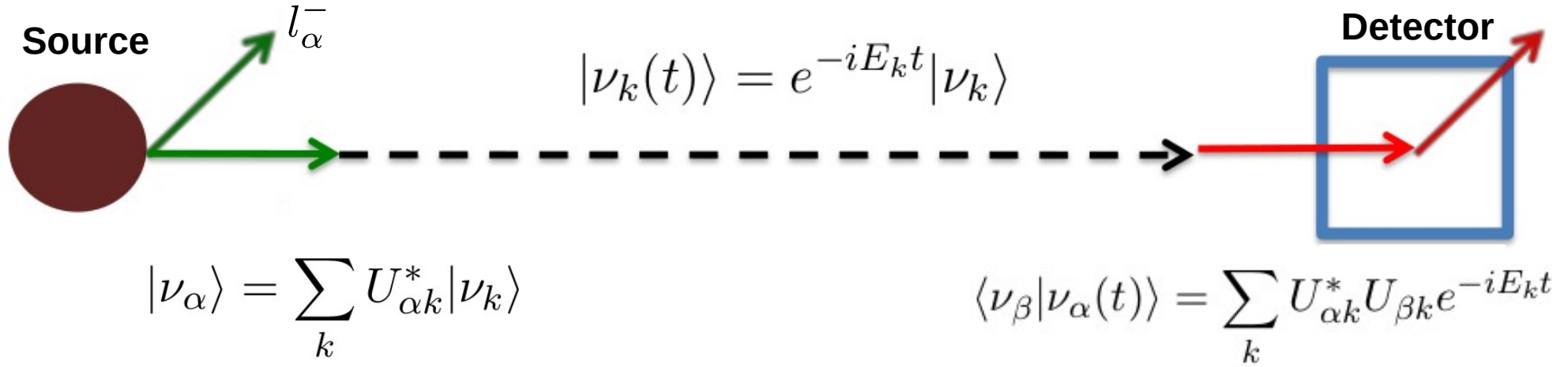
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SEZIONE DI TORINO

March 31st 2021

Neutrino oscillations



Three-neutrino oscillations

Neutrino mixing matrix

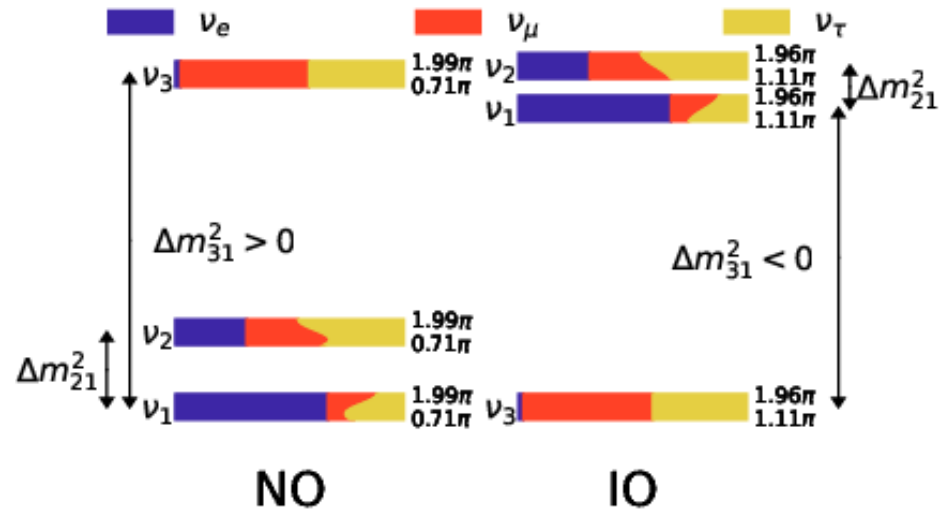
$$U = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

Three mixing angles $\theta_{12}, \theta_{13}, \theta_{23}$

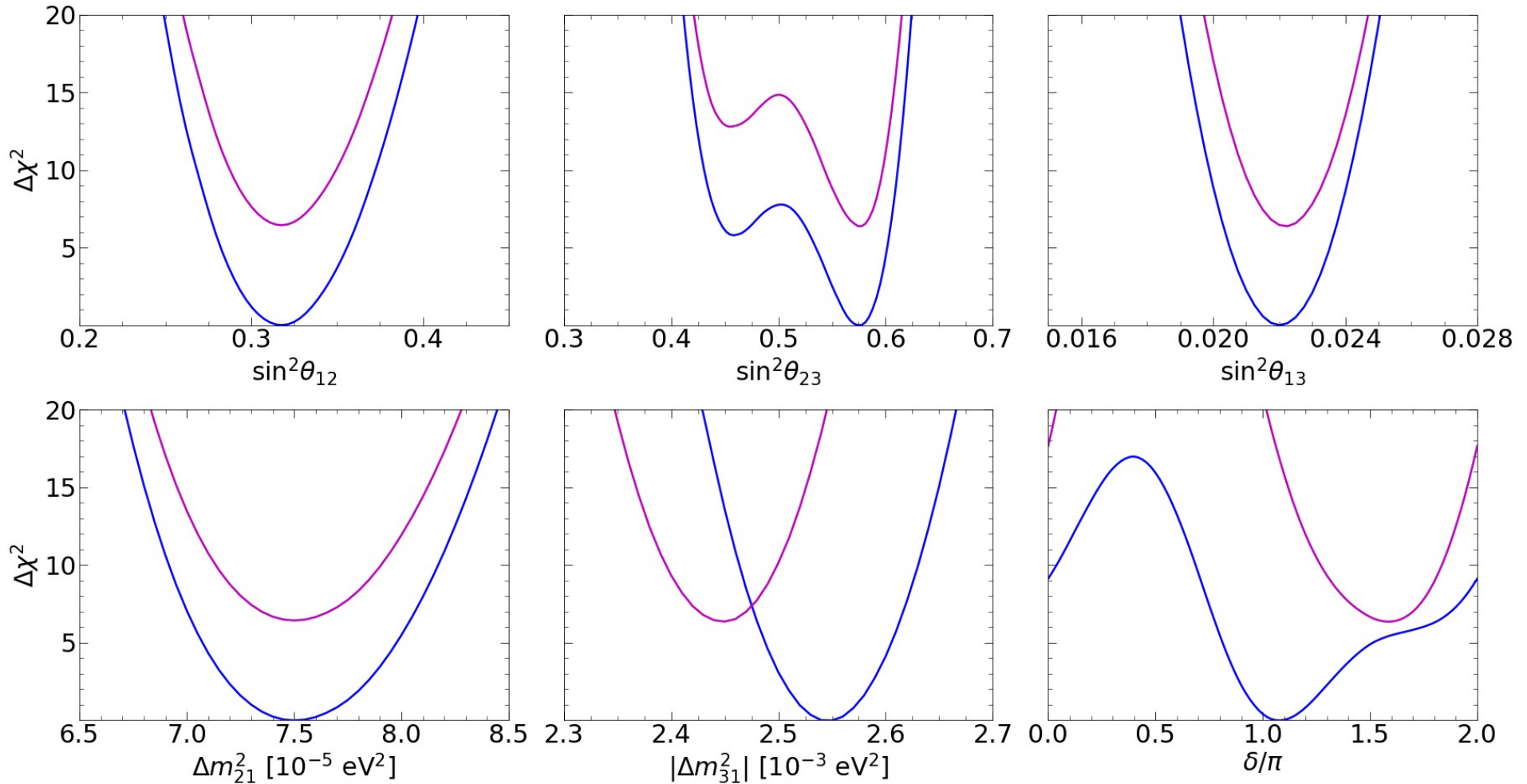
1 Dirac + 2 Majorana CP-phases

Three masses m_1, m_2, m_3 for which two orderings are possible

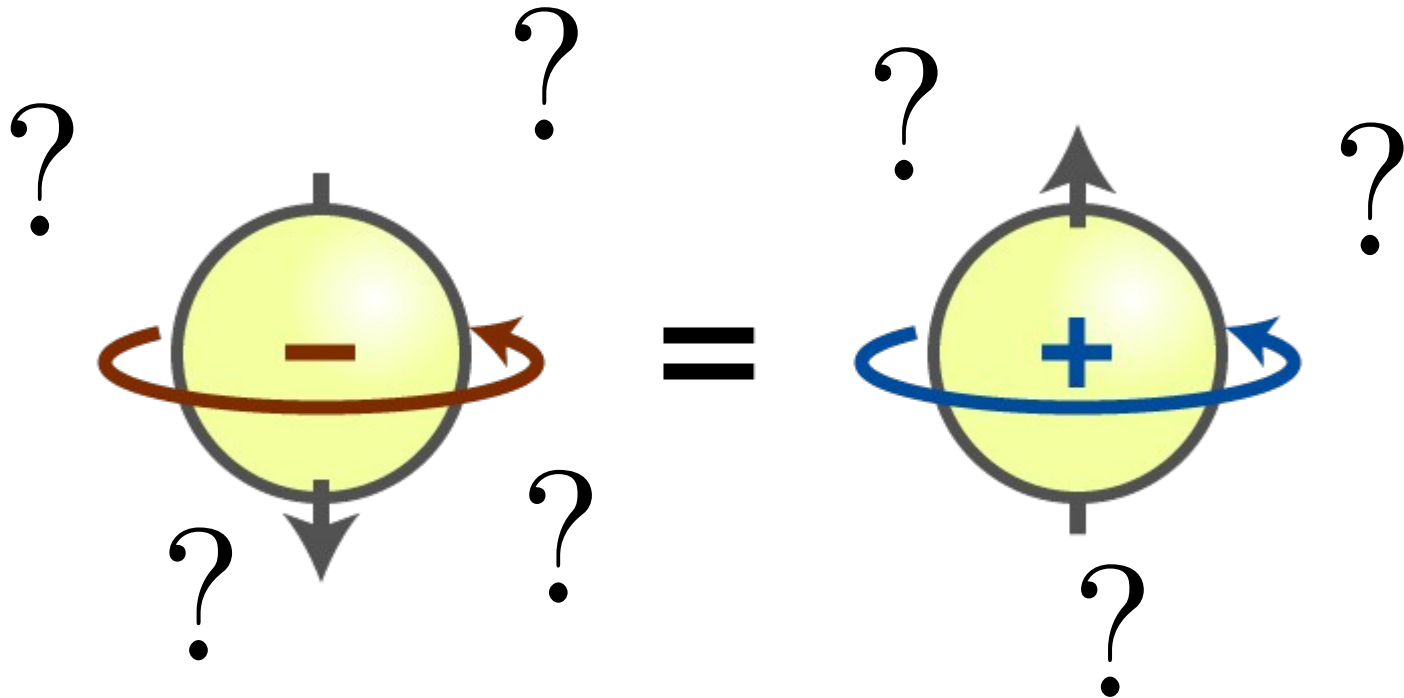
Oscillations are only sensitive to mass splittings



Three-neutrino oscillations



CPT violation



CPT violation

Assume that neutrinos oscillate with parameters

$$\Delta m_{21}^2, \Delta m_{31}^2, \theta_{12}, \theta_{13}, \theta_{23}, \delta$$

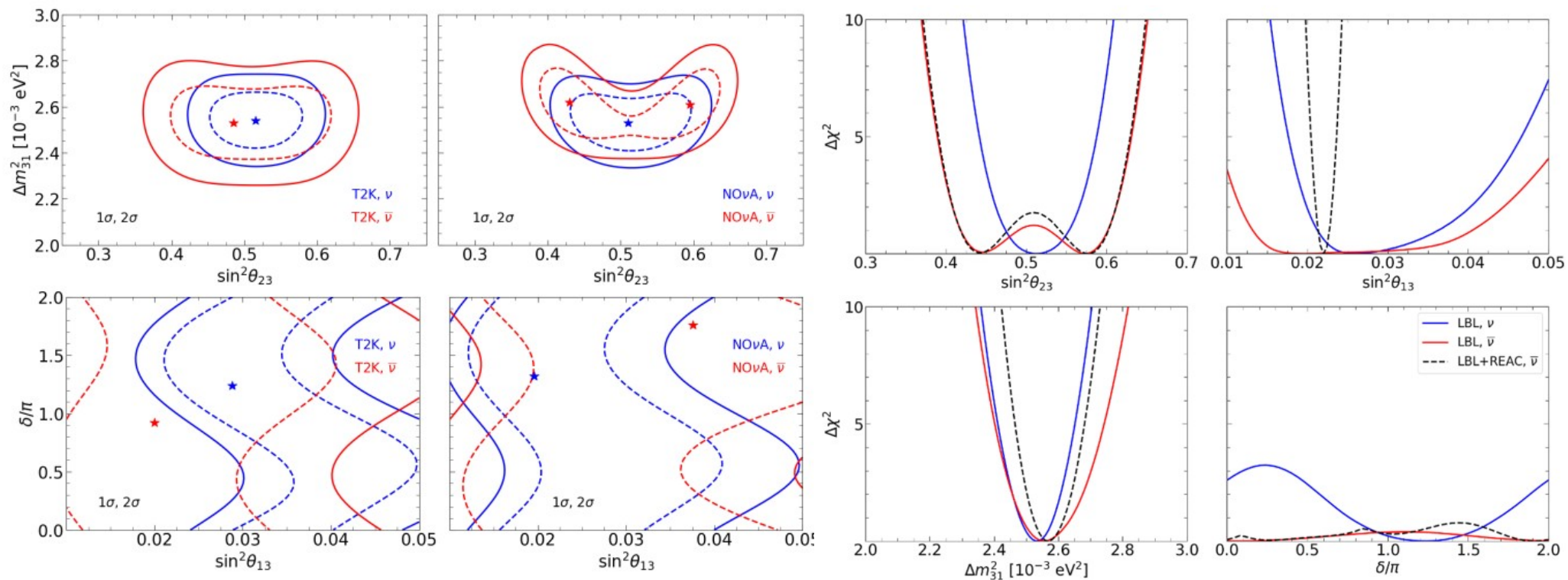
while the antineutrinos oscillate with a new set of parameters

$$\Delta \bar{m}_{21}^2, \Delta \bar{m}_{31}^2, \bar{\theta}_{12}, \bar{\theta}_{13}, \bar{\theta}_{23}, \bar{\delta}$$

Different parameters for neutrinos and antineutrino would indicate a violation of CPT symmetry

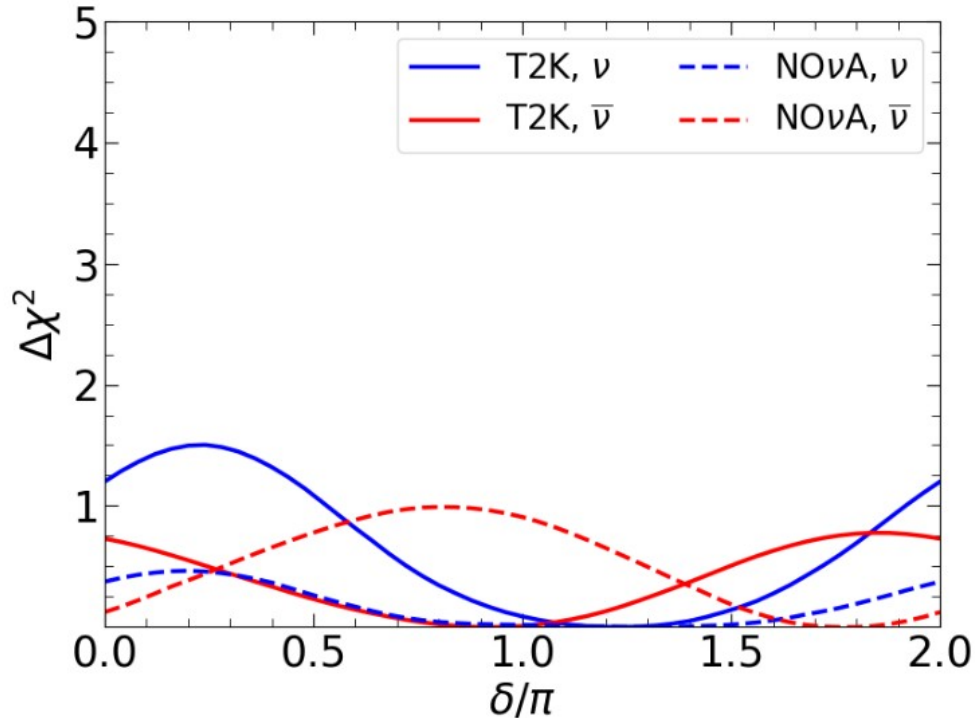
Current experiments

Current experiments



Barenboim, Ternes, Tórtola, 2005.05975, JHEP 2020

Current experiments

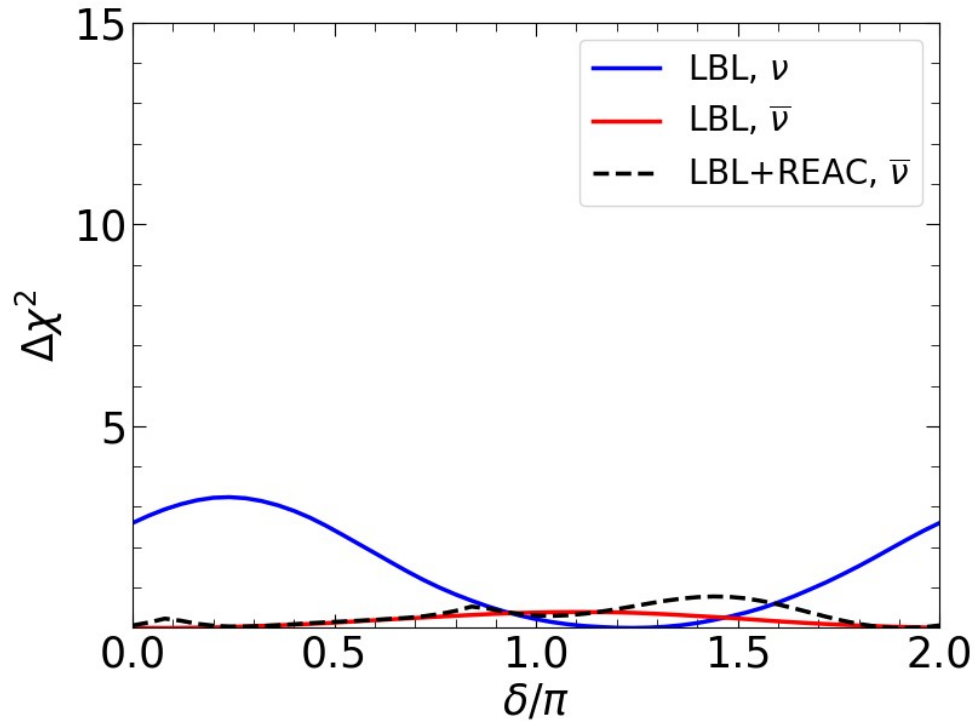


If CPT is not conserved a measurement of CP violation is currently not possible

The different event spectra could be explained with different reactor mixing angles

Barenboim, Ternes, Tórtola, 2005.05975, JHEP 2020

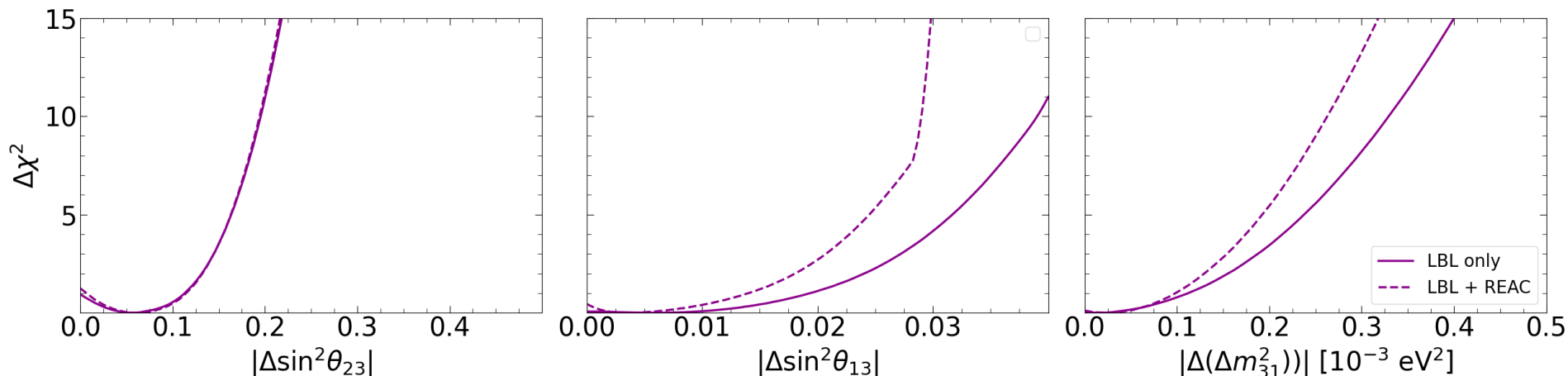
Current experiments



Combining accelerator
and reactor data does
not improve the
situation

Barenboim, Ternes, Tórtola, 2005.05975, JHEP 2020

Current experiments

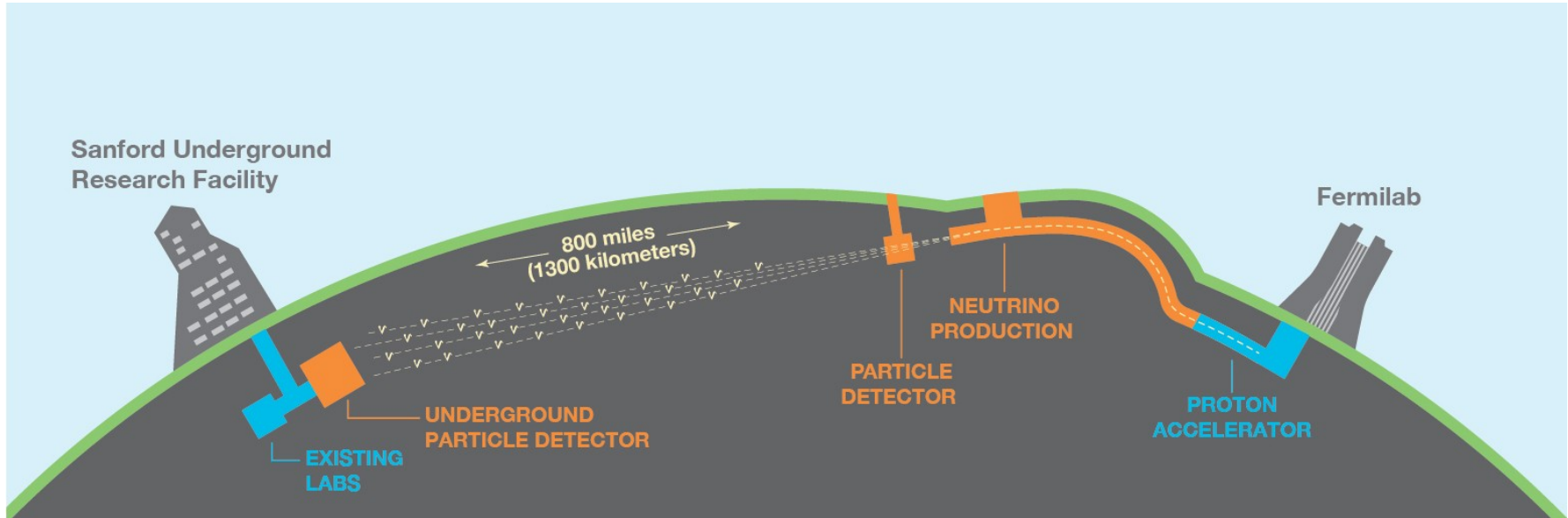


The same data allows us to bound CPT-violating neutrino oscillation parameters

$$\begin{aligned} |\Delta m_{21}^2 - \Delta \bar{m}_{21}^2| &< 4.7 \times 10^{-5} \text{ eV}^2, \\ |\Delta m_{31}^2 - \Delta \bar{m}_{31}^2| &< 2.5 \times 10^{-4} \text{ eV}^2, \\ |\sin^2 \theta_{12} - \sin^2 \bar{\theta}_{12}| &< 0.14, \\ |\sin^2 \theta_{13} - \sin^2 \bar{\theta}_{13}| &< 0.029, \\ |\sin^2 \theta_{23} - \sin^2 \bar{\theta}_{23}| &< 0.19. \end{aligned}$$

Barenboim, Ternes, Tórtola, 2005.05975, JHEP 2020

Future perspectives at DUNE

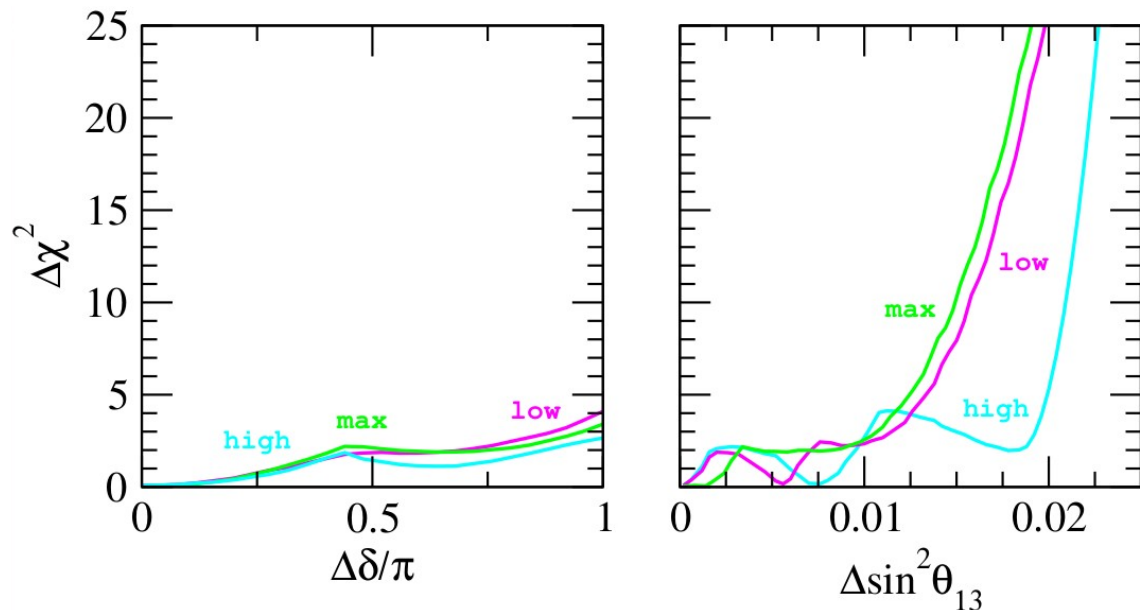


Future perspectives at DUNE

How will this improve with DUNE?

Poor sensitivity to CP phases and reactor angles

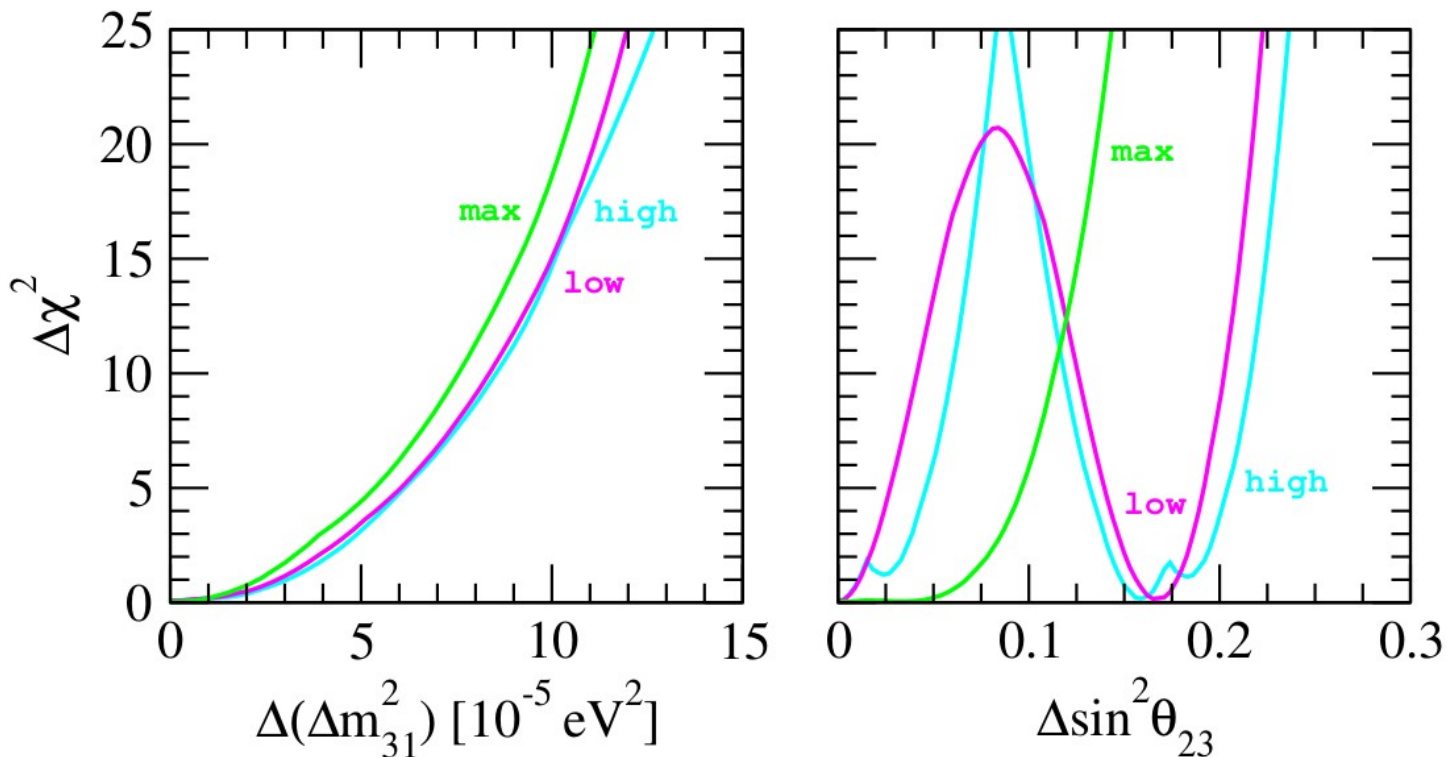
Barenboim, Ternes, Tórtola, 1712.01714, PLB 2018



Valencia - 2017 Global Fit,
1708.01186, PLB 2018

parameter	value
Δm_{21}^2	$7.56 \times 10^{-5} \text{eV}^2$
Δm_{31}^2	$2.55 \times 10^{-3} \text{eV}^2$
$\sin^2 \theta_{12}$	0.321
$\sin^2 \theta_{23}$	0.43, 0.50, 0.60
$\sin^2 \theta_{13}$	0.02155
δ	1.50π

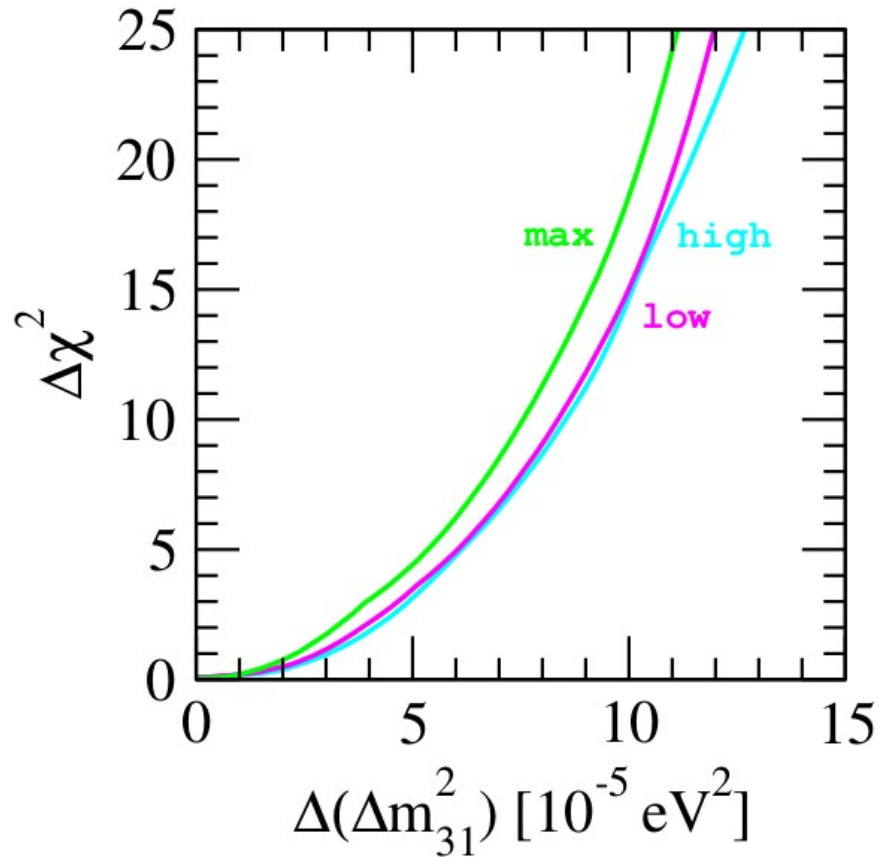
Future perspectives at DUNE



Good improvement
for the atmospheric
mass splitting and
mixing angle

Barenboim, Ternes, Tórtola, 1712.01714, PLB 2018

Future perspectives at DUNE

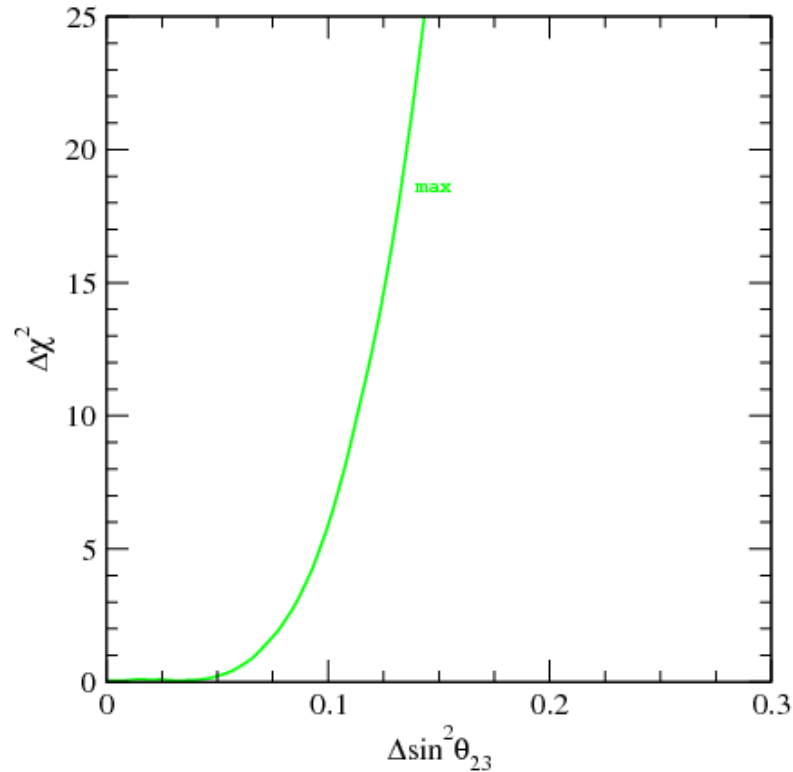


We could obtain

$$\Delta(\Delta m_{31}^2) < 8.1 \times 10^{-5} \text{ eV}^2$$

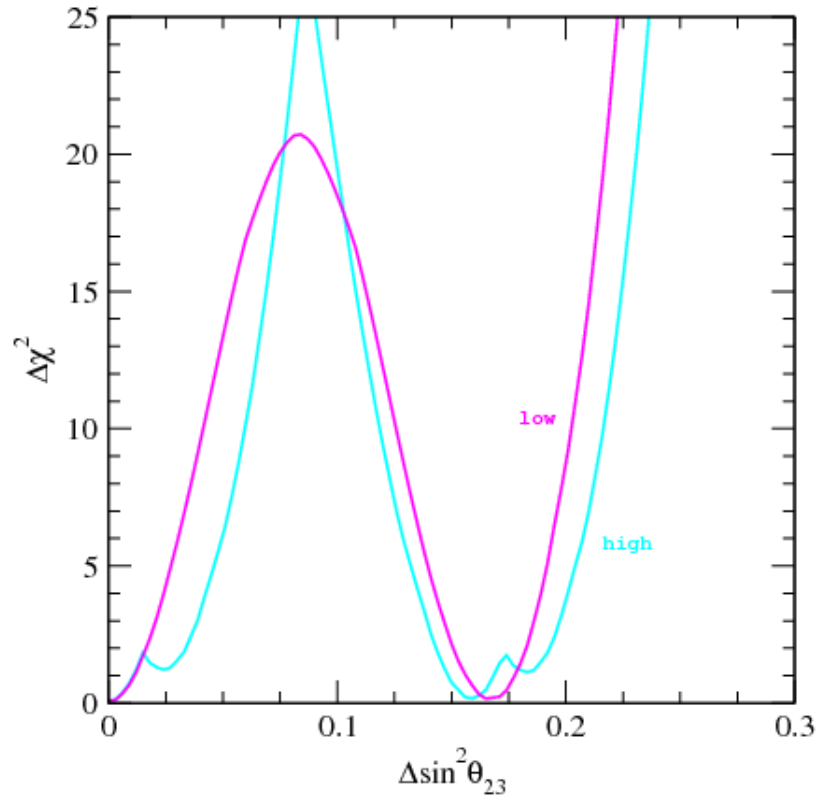
at 3σ C.L.

Future perspectives at DUNE



For the atmospheric angle we obtain increasing sensitivity for maximal mixing

Future perspectives at DUNE



For the atmospheric angle we obtain increasing sensitivity for maximal mixing

For the other values instead it increases and then decreases again

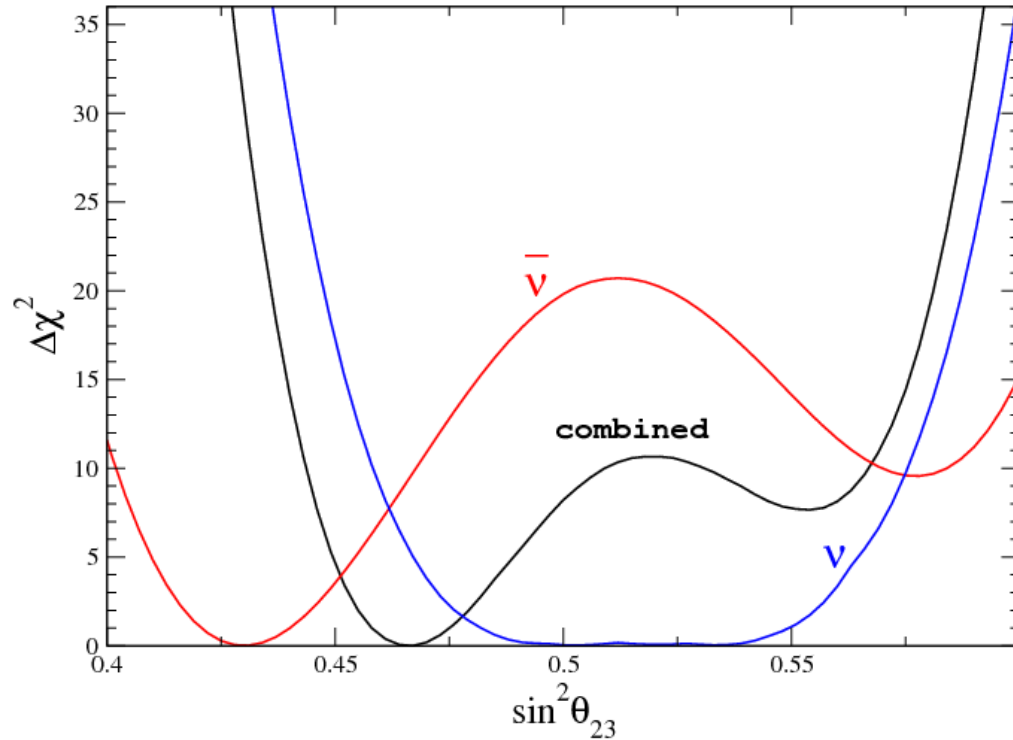
This is due to degenerate solutions

Obtaining impostor solutions



Obtaining impostor solutions

Assume $\sin^2 \theta_{23} = 0.5$, $\sin^2 \bar{\theta}_{23} = 0.43$



The combined best fit value is now

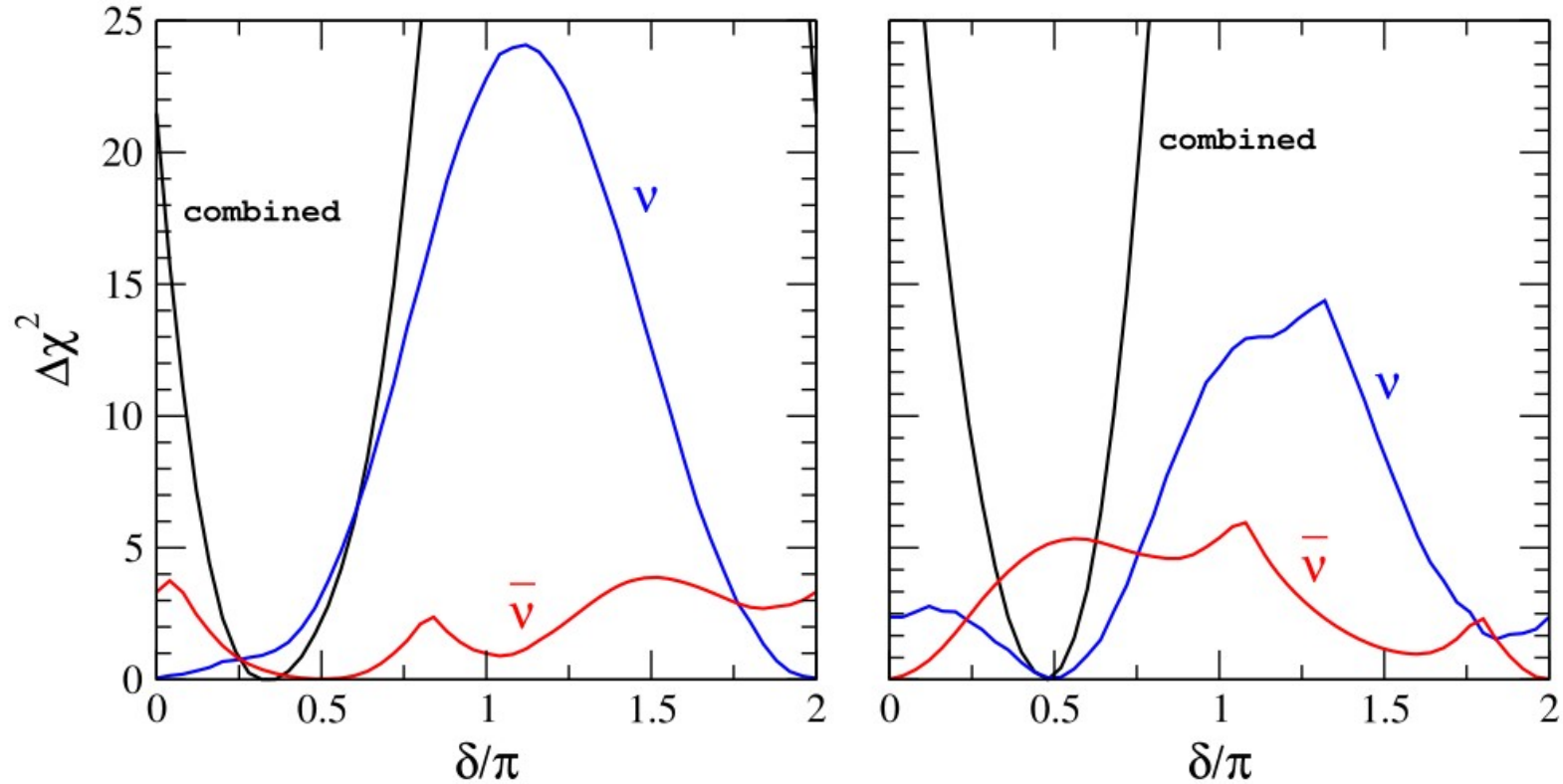
$$\sin^2 \theta_{23}^{\text{comb}} = 0.467$$

The real true values are disfavored at close to 3σ and more than 5σ confidence levels

Barenboim, Ternes, Tórtola, 1712.01714, PLB 2018

Obtaining impostor solutions

This can also happen for other parameters

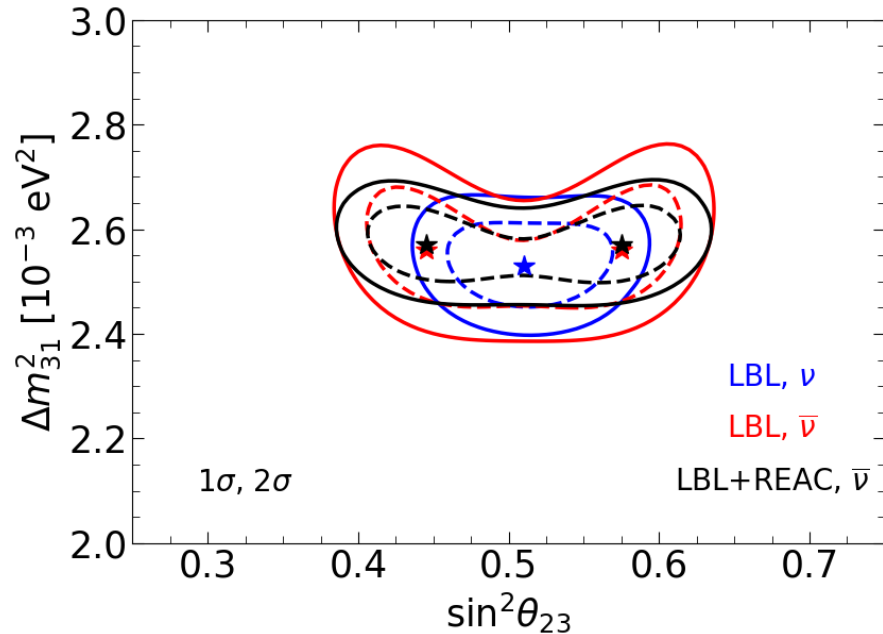


Barenboim, Ternes, Tórtola, 1712.01714, PLB 2018

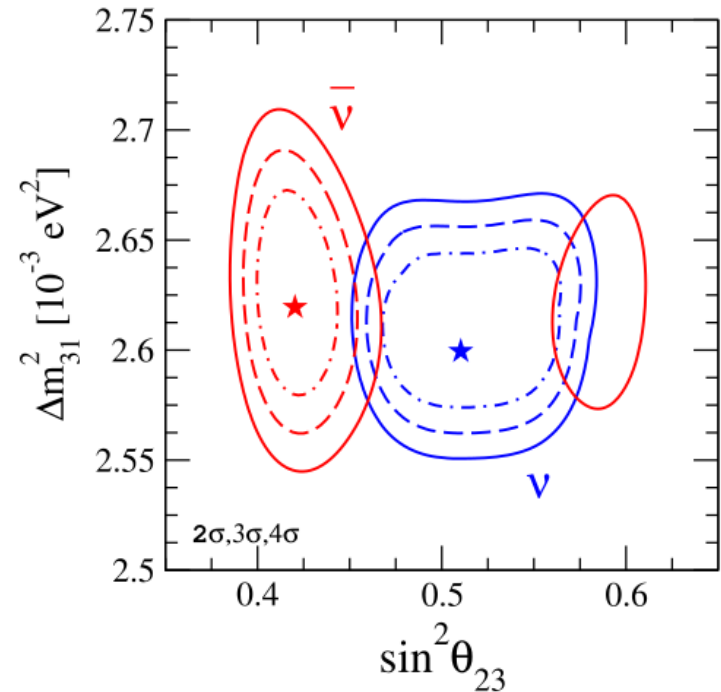
Observing CPT violation

Observing CPT violation

What if the values obtained in T2K and NOvA turn out to be true?

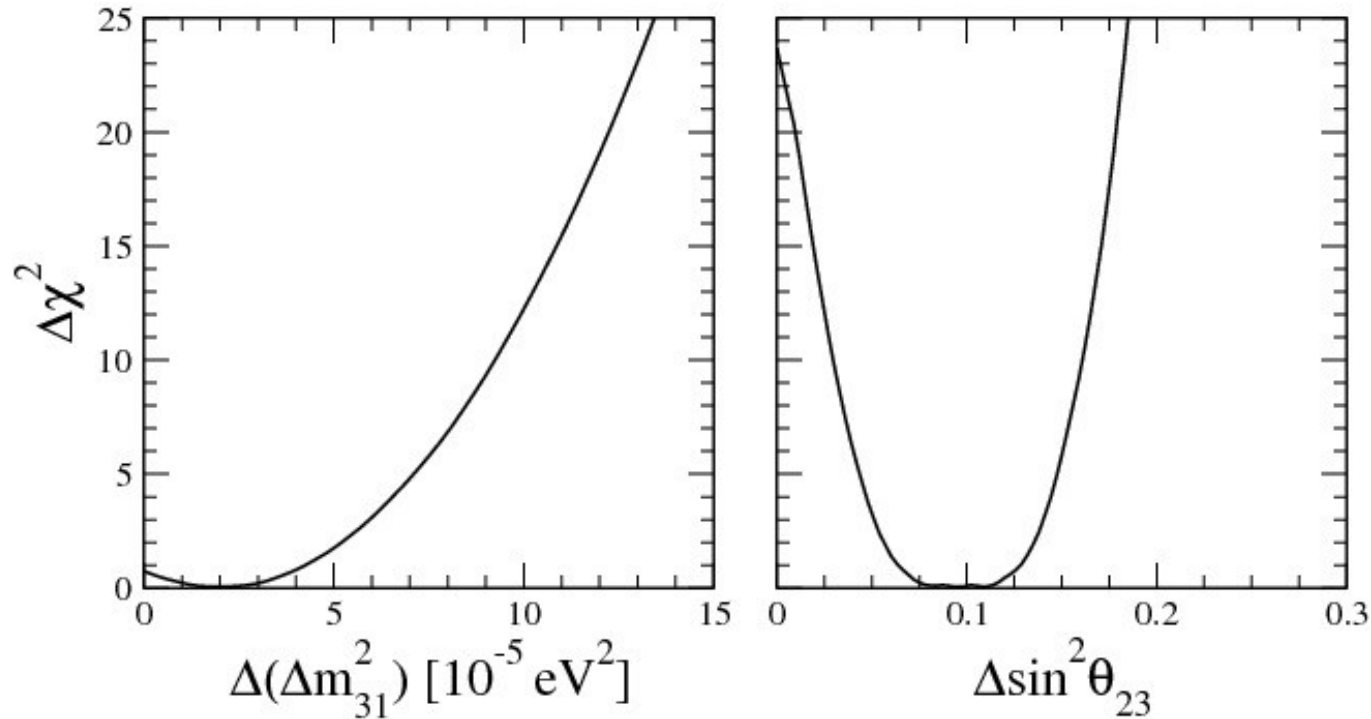


Barenboim, Ternes, Tórtola,
2005.05975, JHEP 2020



Barenboim, Ternes, Tórtola
1712.01714, PLB 2018

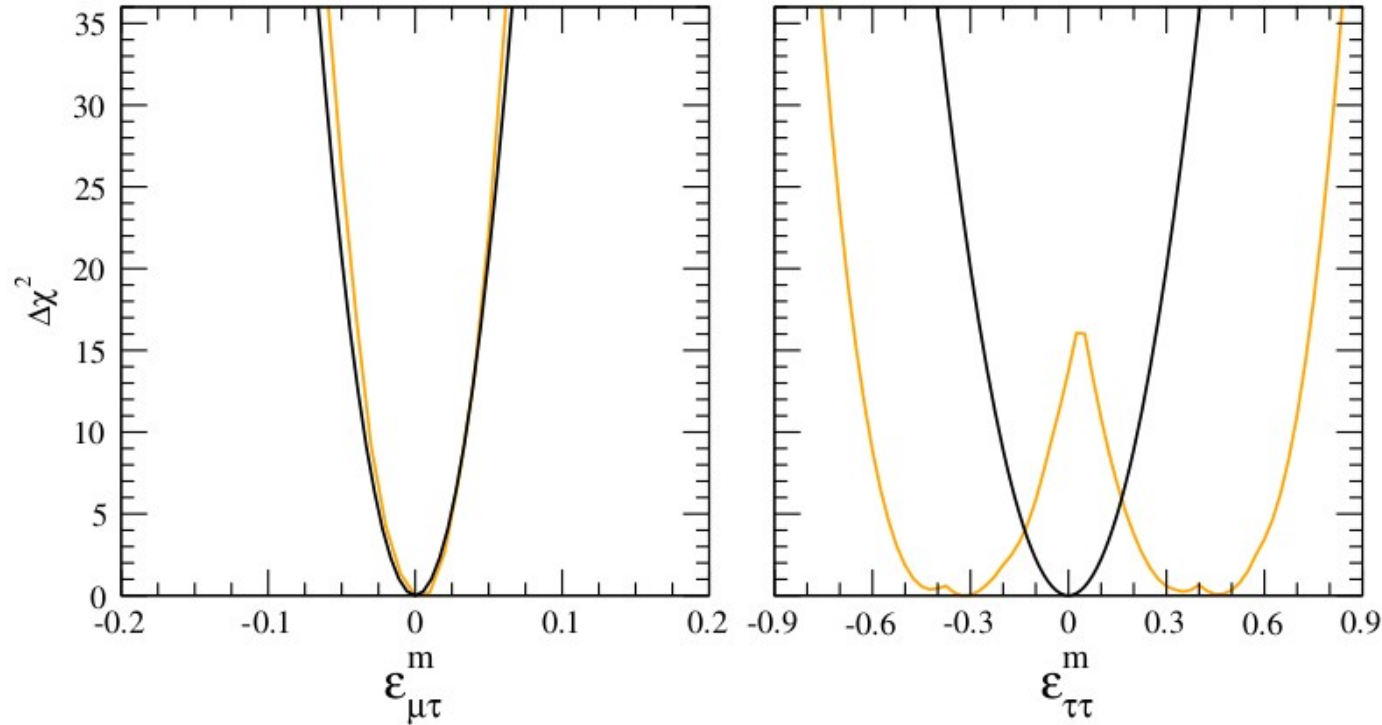
Observing CPT violation



If the different best fit values obtained for neutrino and antineutrino oscillations turned out to be true, DUNE could observe CPT violation at very high significance

Barenboim, Ternes, Tórtola, 1712.01714, PLB 2018

Observing CPT violation



Need to be careful to disentangle CPT violation and neutrino non standard interactions, since they can mimic CPT effects

Barenboim, Ternes, Tórtola, 1804.05842, EPJC 2019

Conclusions

Determination of neutrino oscillation parameters worsens when CPT is not conserved

DUNE will improve the current bounds on CPT violation in the neutrino sector

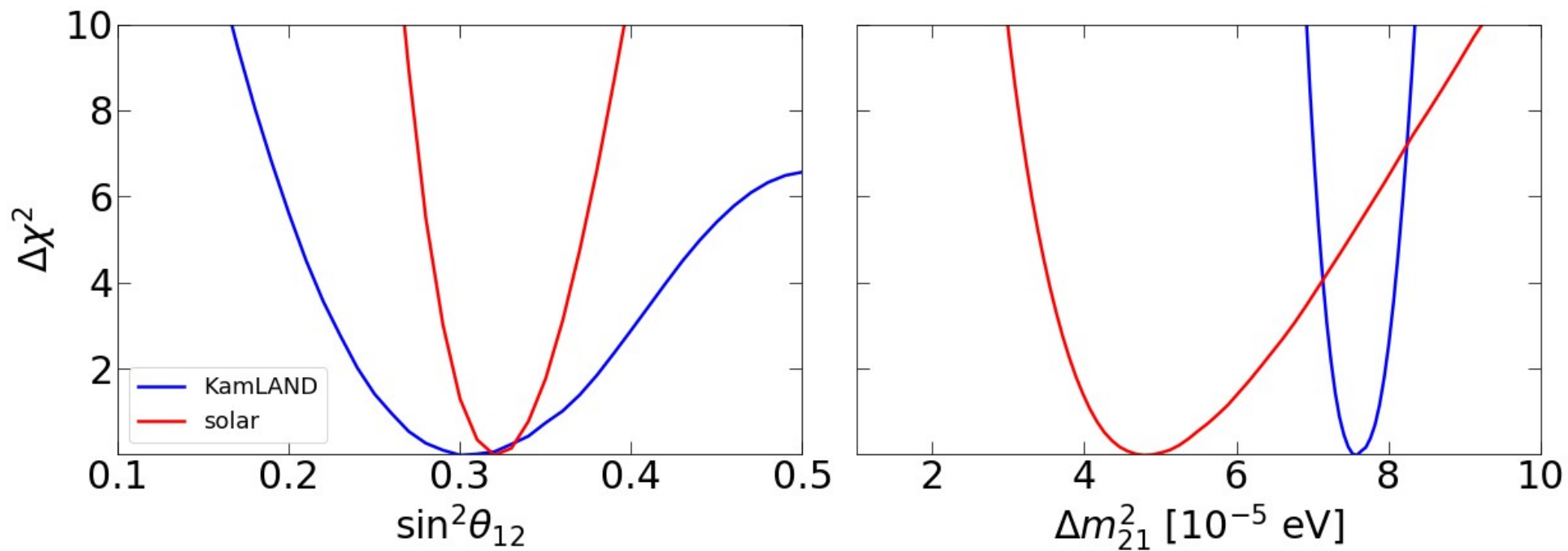
Impostor solutions can arise in the determination of oscillation parameters when combining channels

DUNE might distinguish CPT violation from NSI

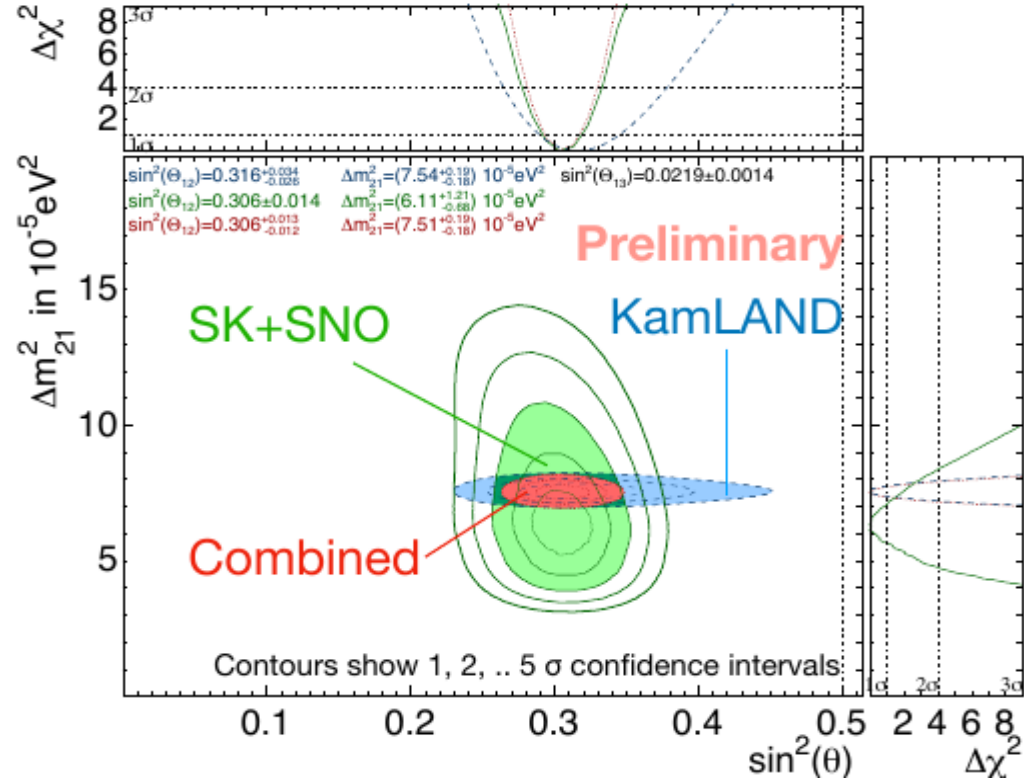
Thanks!



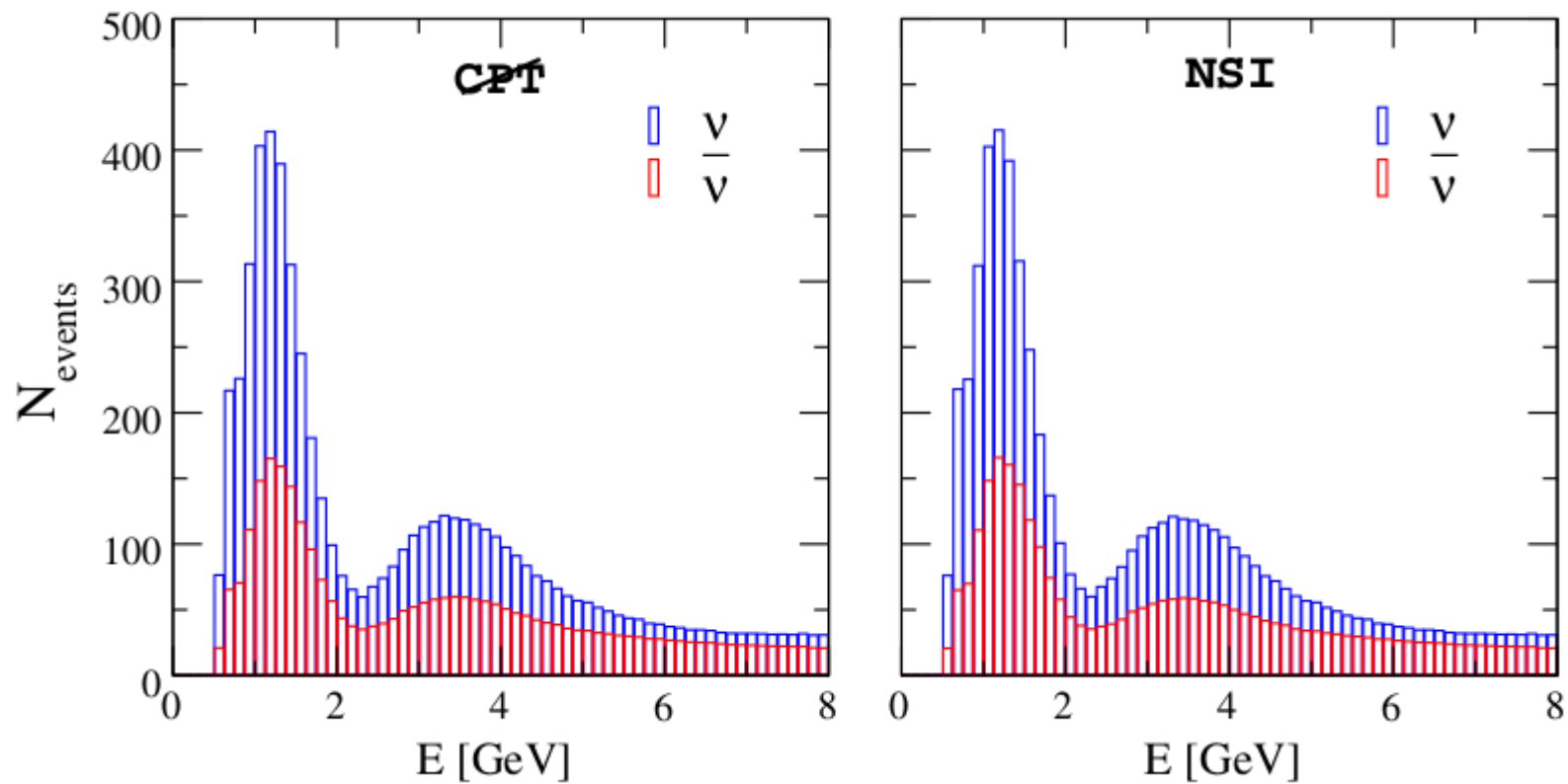
Solar sector



Solar sector



CPT-NSI



CPT-NSI

$$H = \frac{1}{2E} \left[U \begin{pmatrix} 0 & 0 \\ 0 & \Delta m^2 \end{pmatrix} U^\dagger + A \begin{pmatrix} \epsilon_{\mu\mu}^m & \epsilon_{\mu\tau}^{m*} \\ \epsilon_{\mu\tau}^m & \epsilon_{\tau\tau}^m \end{pmatrix} \right]$$

$$\Delta m_{\nu}^2 \cos 2\theta_{\nu} = \Delta m^2 \cos 2\theta + \epsilon_{\tau\tau}^m A,$$

$$\Delta m_{\nu}^2 \sin 2\theta_{\nu} = \Delta m^2 \sin 2\theta + 2\epsilon_{\mu\tau}^m A,$$

$$\Delta m_{\bar{\nu}}^2 \cos 2\theta_{\bar{\nu}} = \Delta m^2 \cos 2\theta - \epsilon_{\tau\tau}^m A,$$

$$\Delta m_{\bar{\nu}}^2 \sin 2\theta_{\bar{\nu}} = \Delta m^2 \sin 2\theta - 2\epsilon_{\mu\tau}^m A,$$