

# Ultralight scalars in leptonic observables

**Pablo Escribano**

IFIC – CSIC / U. Valencia

In collaboration with

**Avelino Vicente**

10.1007/JHEP03(2021)240 [2008.01099]

**BSM - 2021**



VNIVERSITAT  
ID VALÈNCIA



GENERALITAT  
VALENCIANA



**CSIC**  
CONSEJO SUPERIOR DE INVESTIGACIONES CIENTÍFICAS

# 1. Introduction

Several modern experiments recently started taking data

New searches for lepton flavor violating processes

motivated by

- Experimental observation of neutrino flavor oscillations



More precise measurements of lepton flavor conserving observables

- Charged leptons anomalous magnetic moments.

Improved measurements expected at the Muon g-2 experiment

What type of new physics can be probed?

⇒ We will concentrate on ultralight scalars that couple to charged leptons

## 2. The effective Lagrangian

The general interaction between charged leptons and the real scalar is given by

Effective Lagrangian

$$\mathcal{L}_{\ell\ell\phi} = \phi \bar{\ell}_\beta \left( S_L^{\beta\alpha} P_L + S_R^{\beta\alpha} P_R \right) \ell_\alpha + \text{h.c.}$$

Possible flavor combinations:  $\beta\alpha = \{ee, \mu\mu, \tau\tau, e\mu, e\tau, \mu\tau\}$

$$\begin{array}{l} m_\phi \ll m_e \\ \text{in practice} \\ m_\phi \rightarrow 0 \end{array}$$

Full Effective Lagrangian

$$\mathcal{L} = \mathcal{L}_{\ell\ell\phi} + \mathcal{L}_{\ell\ell\gamma} + \mathcal{L}_{4\ell}$$

Dipole operators:  $\mathcal{L}_{\ell\ell\gamma} = \frac{em_\alpha}{2} \bar{\ell}_\beta \sigma^{\mu\nu} \left[ (K_2^L)^{\beta\alpha} P_L + (K_2^R)^{\beta\alpha} P_R \right] \ell_\alpha F_{\mu\nu} + \text{h.c.}$

4-fermion operator  $\mathcal{L}_{4\ell} = \sum_{\substack{I=S,V,T \\ X,Y=L,R}} (A_{XY}^I)^{\beta\alpha\delta\gamma} \bar{\ell}_\beta \Gamma_I P_X \ell_\alpha \bar{\ell}_\delta \Gamma_I P_Y \ell_\gamma + \text{h.c.}$

$$\Gamma_S = 1, \Gamma_V = \gamma_\mu \text{ and } \Gamma_T = \sigma_{\mu\nu}$$

[ Porod, Staub, Vicente, 1405.1434 ]

# 3. Bounds on the couplings: FCC

## Stellar cooling

The production and emission of the scalar inside stars or in supernovae may constitute a powerful cooling mechanism

Sloan Digital Sky Survey and SuperCOSMOS Sky Survey

$$\Rightarrow \text{Im } S^{ee} < 2.1 \times 10^{-13}$$

[ Calibbi, Redigolo, Ziegler, Zupan, 2006.04795 ]

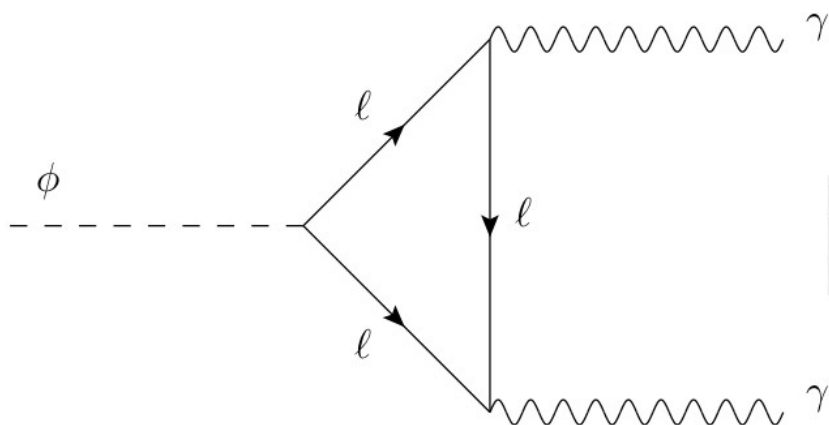
Supernova SN 1987A

$$\Rightarrow \text{Im } S^{\mu\mu} < 2.1 \times 10^{-10}$$

[ Croon, Elor, Leane, McDermott, 2006.13942 ]

$$\Rightarrow \text{Re } S^{\beta\beta} \lesssim [\text{Im } S^{\beta\beta}]_{\text{max}}$$

## Scalar interacting with a pair of photons



OSQAR experiment

[ Ballou *et al.*, 2006.13942 ]

$$\left| \sum_{\beta} \frac{\text{Re } S^{\beta\beta}}{m_{\beta}} \right|^2 < 3.8 \times 10^{-8} \text{ GeV}^{-2} \quad \left| \sum_{\beta} \frac{\text{Im } S^{\beta\beta}}{m_{\beta}} \right|^2 < 1.7 \times 10^{-8} \text{ GeV}^{-2}$$

$$\Rightarrow S^{ee} \lesssim 10^{-7} \text{ and } S^{\mu\mu} \lesssim 10^{-5}$$

Limits only valid if this diagram is the only contribution to the coupling to photons

### 3. Bounds on the couplings: FVC

$$\boxed{\ell_\alpha \rightarrow \ell_\beta \phi} \quad \Gamma(\ell_\alpha \rightarrow \ell_\beta \phi) = \frac{m_\alpha}{32\pi} |S^{\beta\alpha}|^2$$
$$\boxed{\alpha = \mu} \quad |S^{\beta\alpha}| = \left( |S_L^{\beta\alpha}|^2 + |S_R^{\beta\alpha}|^2 \right)^{1/2}$$

**TRIUMPH** [ Jodidio *et al.*, Phys.Rev.D 37, 237 ]

[ Hirsch, Vicente, Meyer, Porod, 0902.0525 ]

$$\longrightarrow \text{BR}(\mu \rightarrow e \phi) \lesssim 10^{-5}$$

$$\Rightarrow |S^{e\mu}| < 5.3 \times 10^{-11}$$

$$\boxed{\alpha = \tau}$$

**ARGUS** [ Albrecht *et al.*, Z. Phys. C 68, 25-28 ]

$$\frac{\text{BR}(\tau \rightarrow e \phi)}{\text{BR}(\tau \rightarrow e \nu \bar{\nu})} < 0.015$$

$$\frac{\text{BR}(\tau \rightarrow \mu \phi)}{\text{BR}(\tau \rightarrow \mu \nu \bar{\nu})} < 0.026$$

$$|S^{e\tau}| < 5.9 \times 10^{-7}$$

$$|S^{\mu\tau}| < 7.6 \times 10^{-7}$$

These limits are weaker than those for muon decays, but still very stringent. They are expected to be improved at Belle II.

# 4. Phenomenology: $\mu \rightarrow e \gamma$ vs $\mu \rightarrow 3 e$

$$\ell_\alpha \rightarrow \ell_\beta \gamma$$

$$\Gamma(\ell_\alpha \rightarrow \ell_\beta \gamma) = \frac{e^2 m_\alpha^5}{16 \pi} \left[ \left| (K_2^L)^{\beta\alpha} \right|^2 + \left| (K_2^R)^{\beta\alpha} \right|^2 \right]$$

$$\ell_\alpha^- \rightarrow \ell_\beta^- \ell_\beta^- \ell_\beta^+$$

$$\Gamma_\phi(\ell_\alpha^- \rightarrow \ell_\beta^- \ell_\beta^- \ell_\beta^+) =$$

$$\frac{m_\alpha}{512\pi^3} \left\{ \left( \left| S_L^{\beta\alpha} \right|^2 + \left| S_R^{\beta\alpha} \right|^2 \right) \left\{ \left| S^{\beta\beta} \right|^2 \left( 4 \log \frac{m_\alpha}{m_\beta} - \frac{49}{6} \right) - \frac{2}{6} \left[ (S^{\beta\beta*})^2 + (S^{\beta\beta})^2 \right] \right\} \right. \\ \left. - \frac{m_\alpha^2}{6} \left\{ S_L^{\beta\alpha} S^{\beta\beta} A_{LL}^{S*} + 2 S_L^{\beta\alpha} S^{\beta\beta*} A_{LR}^{S*} + 2 S_R^{\beta\alpha} S^{\beta\beta} A_{RL}^{S*} + S_R^{\beta\alpha} S^{\beta\beta*} A_{RR}^{S*} \right. \right. \\ \left. - 12 \left( S_L^{\beta\alpha} S^{\beta\beta} A_{LL}^{T*} + S_R^{\beta\alpha} S^{\beta\beta*} A_{RR}^{T*} \right) - 4 \left( S_R^{\beta\alpha} S^{\beta\beta} A_{RL}^{V*} + S_L^{\beta\alpha} S^{\beta\beta*} A_{LR}^{V*} \right) \right. \\ \left. \left. + 6e^2 \left[ S_R^{\beta\alpha} S^{\beta\beta} (K_2^L)^{\beta\alpha*} + S_L^{\beta\alpha} S^{\beta\beta*} (K_2^R)^{\beta\alpha*} \right] + \text{c.c.} \right\} \right\}$$

$$S^{\beta\beta} = S_L^{\beta\beta} + S_R^{\beta\beta*}$$

**NEW!**

# 4. Phenomenology: $\mu \rightarrow e \gamma$ vs $\mu \rightarrow 3 e$

## Simplified Effective Lagrangian

$$\mathcal{L}_{\text{LFV}}^{\text{simp}} = \frac{e m_\alpha (K_2^L)^{\beta\alpha}}{2} \bar{l}_\beta \sigma^{\mu\nu} P_L l_\alpha F_{\mu\nu} + S_L^{\beta\alpha} \phi \bar{l}_\beta P_L l_\alpha + \text{h.c.}$$

It only includes left-handed photonic dipole and scalar-mediated operators.

Here we assume the dipole contributions to be independent from the scalar induced ones.

$$l_\alpha^- \rightarrow l_\beta^- l_\beta^- l_\beta^+$$

$$\begin{aligned} \Gamma(l_\alpha^- \rightarrow l_\beta^- l_\beta^- l_\beta^+) = & \\ & \frac{m_\alpha}{512\pi^3} \left\{ |S_L^{\beta\alpha}|^2 \left\{ |S_L^{\beta\beta}|^2 \left( 4 \log \frac{m_\alpha}{m_\beta} - \frac{49}{6} \right) - \frac{2}{6} \left[ (S_L^{\beta\beta*})^2 + (S_L^{\beta\beta})^2 \right] \right\} \right. \\ & \left. + m_\alpha^4 e^4 |K_2^L|^2 \left( \frac{16}{3} \log \frac{m_{l_\alpha}}{m_{l_\beta}} - \frac{22}{3} \right) \right\} \end{aligned}$$

# 4. Phenomenology: $\mu \rightarrow e \gamma$ vs $\mu \rightarrow 3 e$

## Simplified Effective Lagrangian

$$\mathcal{L}_{\text{LFV}}^{\text{simp}} = \frac{e m_\alpha}{2} (K_2^L)^{\beta\alpha} \bar{l}_\beta \sigma^{\mu\nu} P_L l_\alpha F_{\mu\nu} + S_L^{\beta\alpha} \phi \bar{l}_\beta P_L l_\alpha + \text{h.c.}$$

We will make use of the parametrization [ Gouvea, Vogel, 1303.4097 ]

$$e (K_2^L)^{\beta\alpha} \equiv \frac{1}{(\kappa + 1)\Lambda^2}, \quad S_L^{\beta\alpha} \equiv m_\alpha \frac{\kappa}{(\kappa + 1)\Lambda}, \quad S_L^{\beta\alpha} = S_L^{\beta\beta}$$

- $\Lambda$  is dimensionful. Energy scale at which the coefficients are induced.
- $\kappa$  is dimensionless. Accounts for the relative intensity of both interactions.

$\kappa \ll 1$   $\longrightarrow$  Dipole operator dominates

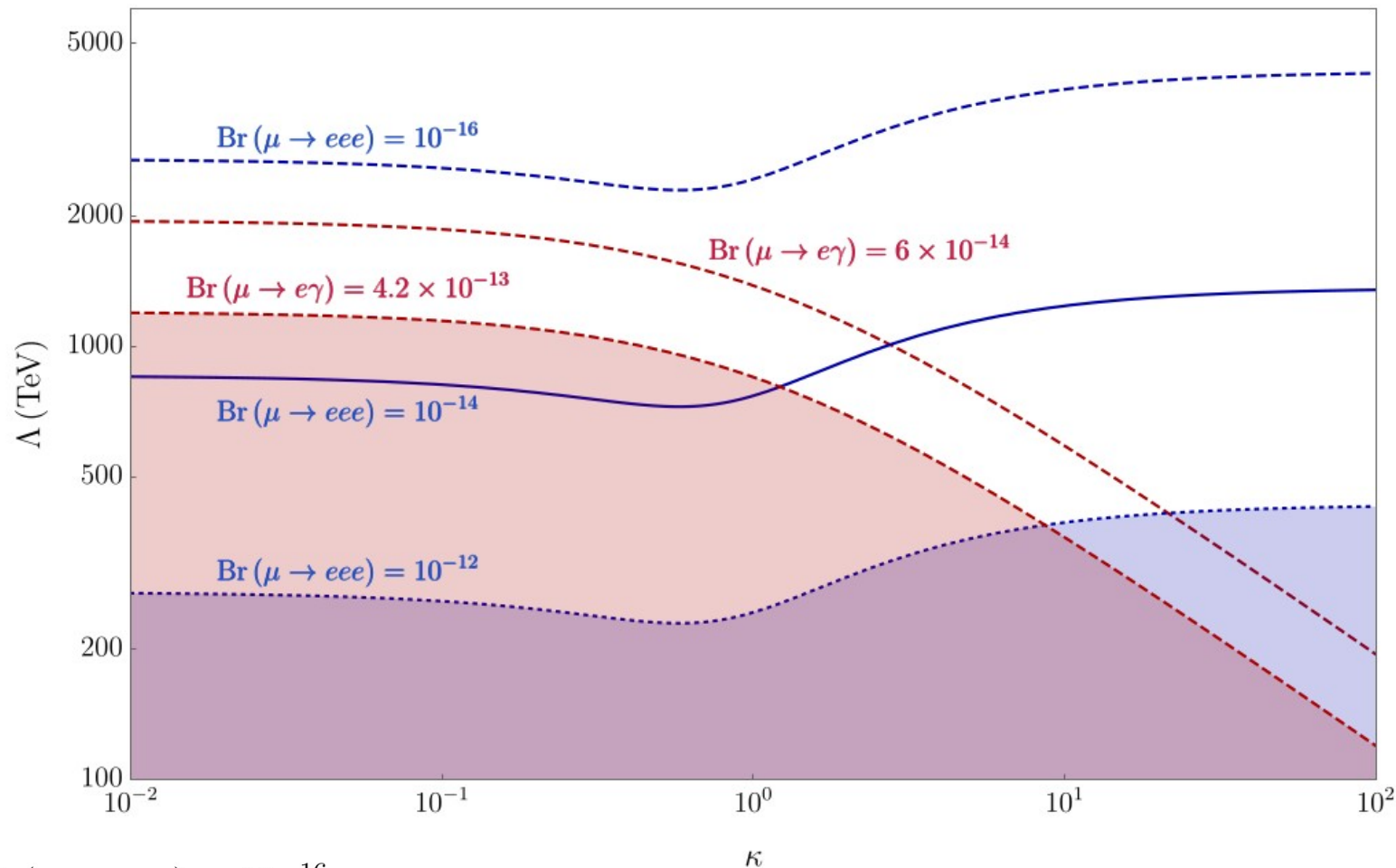
$\kappa \gg 1$   $\longrightarrow$  Scalar mediated contribution dominates

$$\Rightarrow \Gamma(l_\alpha \rightarrow l_\beta \gamma) = \frac{m_\alpha^5}{16\pi} \frac{1}{(\kappa + 1)^2 \Lambda^4}$$

$$\Rightarrow \Gamma(l_\alpha^- \rightarrow l_\beta^- l_\beta^- l_\beta^+) = \frac{m_\alpha^5}{512\pi^3} \left[ \frac{\kappa^2}{(\kappa + 1)^4 \Lambda^4} \left( 4 \log \frac{m_\alpha}{m_\beta} - \frac{53}{6} \right) + \frac{e^2}{(\kappa + 1)^4 \Lambda^4} \left( \frac{16}{3} \log \frac{m_{l_\alpha}}{m_{l_\beta}} - \frac{22}{3} \right) \right]$$



# 4. Phenomenology: $\mu \rightarrow e \gamma$ vs $\mu \rightarrow 3 e$



$\text{Br}(\mu \rightarrow eee) = 10^{-16}$   
 $\text{Br}(\mu \rightarrow e\gamma) = 6 \times 10^{-14}$



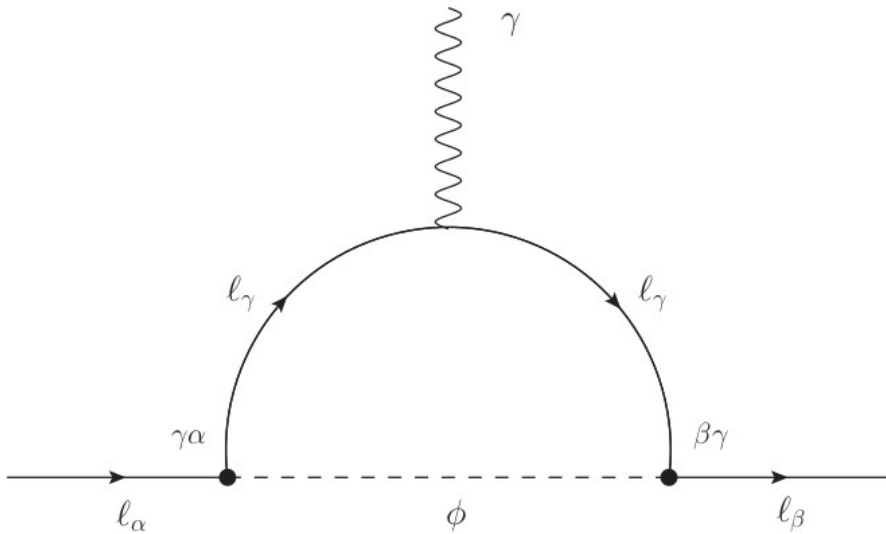
Future sensitivities for the MEG-II and Mu3e experiments

- If  $\kappa \gg 1$  and  $\text{Br}(\mu \rightarrow eee) > 10^{-16} \rightarrow \Lambda < 4000 \text{ TeV}$
- If  $\kappa \ll 1$  and  $\text{Br}(\mu \rightarrow e\gamma) > 10^{-14} \rightarrow \Lambda < 2000 \text{ TeV}$

The search for the scalar mediated contribution in Mu3e will be very constraining in all the parameter space

# 4. Phenomenology: $\mu \rightarrow e \gamma$ vs $\mu \rightarrow 3 e$

Generation of dipole operators by loops involving the ultralight scalar as shown here



- We assume that the scalar provides the dominant contribution to dipole operators.
- Only the electron diagonal coupling and the electron – muon couplings are allowed not to be zero and are considered to be real.

## Approximated expressions

$$(K_2^L)^{e\mu} = \frac{S^{ee}}{96\pi^2 m_\mu^3} \left\{ 3 m_\mu S_R^{e\mu} + m_e \left( -6 S_L^{e\mu} + 2 \pi^2 S_L^{e\mu} + 3 S_R^{e\mu} \right) + 3 m_e S_L^{e\mu} \log \left( -\frac{m_e^2}{m_\mu^2} \right) \left[ 1 + \log \left( -\frac{m_e^2}{m_\mu^2} \right) \right] \right\}$$

$$(K_2^R)^{e\mu} = \frac{S^{ee}}{96\pi^2 m_\mu^3} \left\{ 3 m_\mu S_L^{e\mu} + m_e \left( -6 S_R^{e\mu} + 2 \pi^2 S_R^{e\mu} + 3 S_L^{e\mu} \right) + 3 m_e S_R^{e\mu} \log \left( -\frac{m_e^2}{m_\mu^2} \right) \left[ 1 + \log \left( -\frac{m_e^2}{m_\mu^2} \right) \right] \right\}$$

**NEW!**

## 4. Phenomenology: $\mu \rightarrow e \gamma$ vs $\mu \rightarrow 3 e$

$$R_{\alpha\beta} = \frac{\text{BR}(\ell_\alpha \rightarrow \ell_\beta \ell_\beta \ell_\beta)}{\text{BR}(\ell_\alpha \rightarrow \ell_\beta \gamma)}$$

- Scenario 1:  $S_L^{e\mu} = 0$  or  $S_R^{e\mu} = 0$

$$R_{\mu e}^{(1)} \approx \frac{4\pi r}{3\alpha} \frac{12 \log r - 53}{|\log(-r)|^4 + r} \approx 3.2 \cdot 10^4$$

- Scenario 2:  $S_L^{e\mu} = S_R^{e\mu}$

$$R_{\mu e}^{(2)} \approx \frac{4\pi r}{3\alpha} \frac{12 \log r - 53}{|\log^2(-r) + \sqrt{r}|} \approx 1.9 \cdot 10^4 \quad r = \frac{m_\mu^2}{m_e^2}$$

- Scenario 3:  $S_L^{e\mu} = -S_R^{e\mu}$

$$R_{\mu e}^{(3)} \approx \frac{4\pi r}{3\alpha} \frac{12 \log r - 53}{|\log^2(-r) - \sqrt{r}|} \approx 1.1 \cdot 10^5$$

⇒  $R \gg 1$  in all the cases, since the decay in the numerator is induced at tree level, while the one in the denominator takes place at loop order.

⇒ The three scenarios lead to different predictions for the ratio and could allow us to determine the nature of the scalar if both processes are observed.

# 4. Phenomenology: AMM and EDM

## Anomalous magnetic moments (AMM)

$$\Delta a_e = a_e^{\text{exp}} - a_e^{\text{SM}} = (-87 \pm 36) \times 10^{-14}$$

$$\Delta a_\mu = a_\mu^{\text{exp}} - a_\mu^{\text{SM}} = (27.1 \pm 7.3) \times 10^{-10}$$

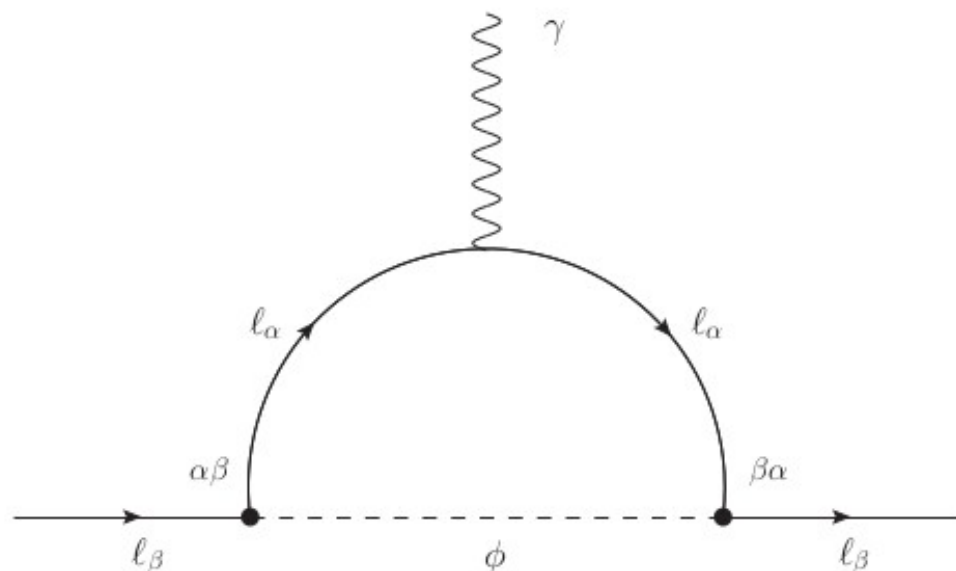
$$a_\beta = \frac{g_\beta - 2}{2}$$

## Electric dipole moments (EDM)

$$|d_e| < 1.1 \times 10^{-29} \text{ e cm}$$

$$|d_\mu| < 1.5 \times 10^{-19} \text{ e cm}$$

## Lepton conserving contributions

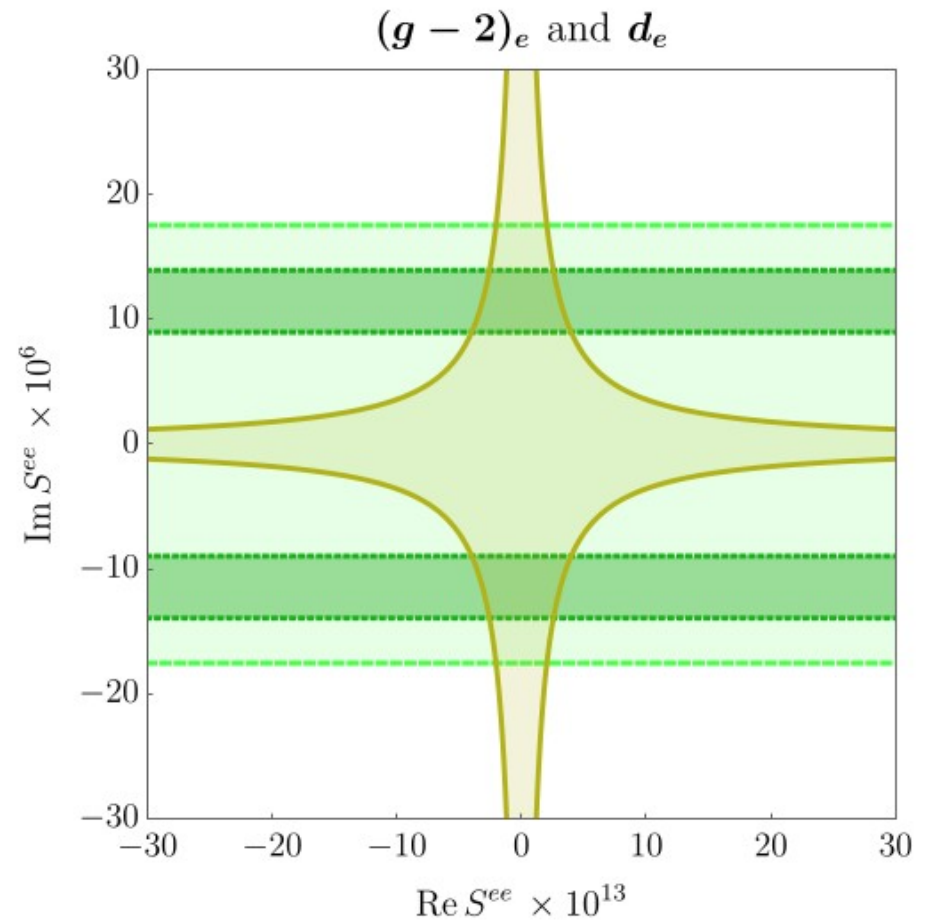
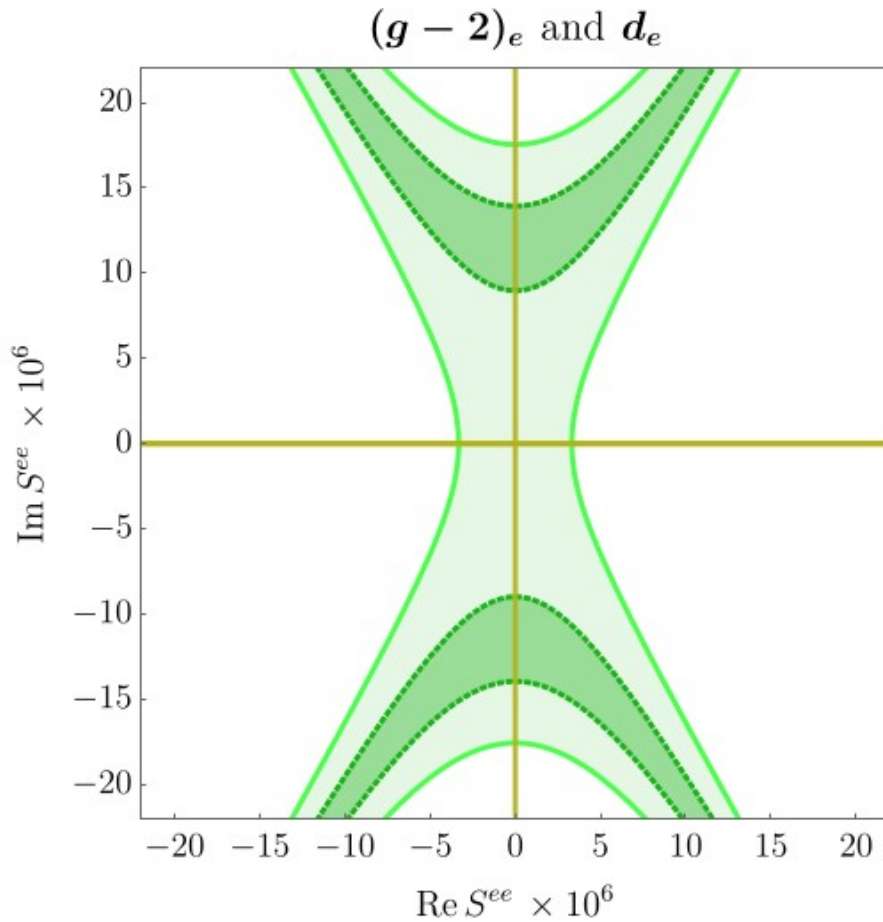


$$\Delta a_\alpha = \frac{1}{16\pi^2} [3 (\text{Re } S^{\alpha\alpha})^2 - (\text{Im } S^{\alpha\alpha})^2]$$

$$d_\alpha = -\frac{e}{8\pi^2 m_\alpha} (\text{Re } S^{\alpha\alpha}) (\text{Im } S^{\alpha\alpha})$$

Too high values for the off-diagonal couplings would be necessary to fully explain the anomalies, and these are excluded by LFV observations.

# 4. Phenomenology: AMM and EDM

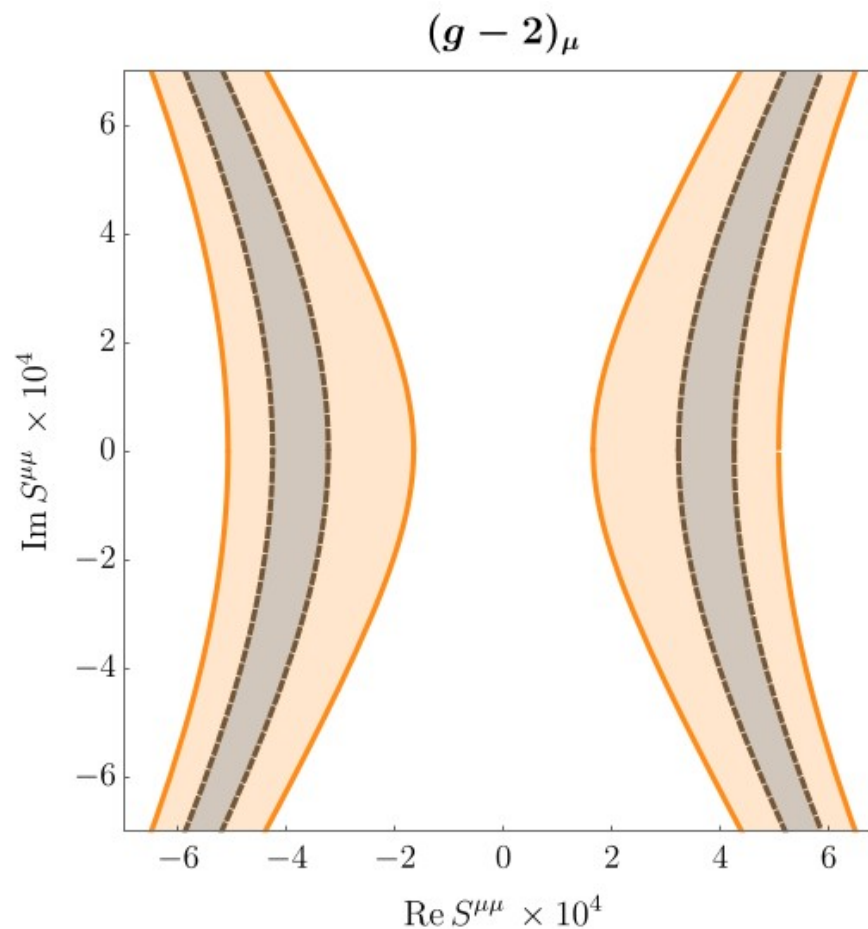
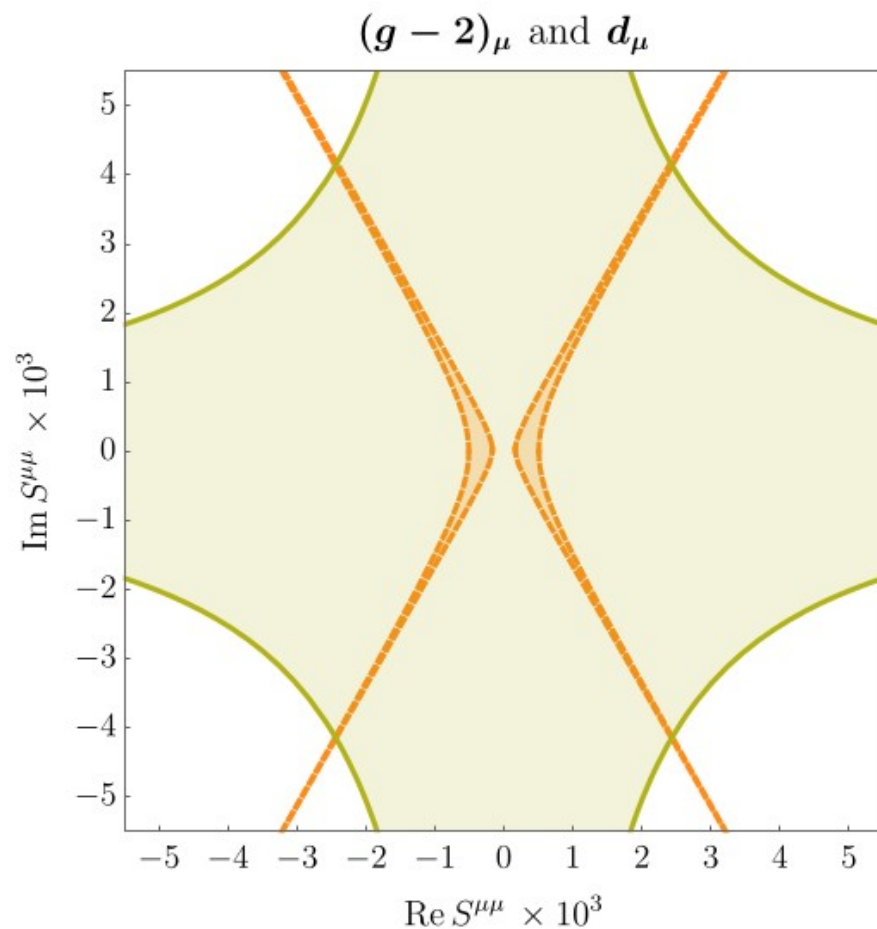


⇒ The electron EDM strongly constraints the coupling, making it essentially purely real or purely imaginary.

⇒ We can find regions that explain the g-2 anomaly compatible with the bound for the EDM

- Even with  $S^{ee} = 0$ , we stay in the  $3 \sigma$  region.
- We achieve agreement at the  $1 \sigma$  level if  $\text{Re}(S^{ee}) \leq 10^{-13}$  and  $\text{Im}(S^{ee}) \sim 10^{-5}$

# 4. Phenomenology: AMM and EDM



⇒ The muon EDM does not impose strong restrictions on the parameter space.

⇒ Larger couplings,  $S^{\mu\mu} \sim 10^{-4}$ , are necessary to explain the muon anomaly.

In both cases, the required values for the couplings are in conflict with the current bounds. A mechanism to suppress the processes from which the bounds are derived would be necessary for the scalar to be able to provide an explanation to the anomalies.

## 6. Summary and discussion

- ⇒ Ultralight scalars appear in a wide variety of SM extensions.
  - As very light states
  - As exactly massless Goldstone bosons
- ⇒ We have explored the impact of ultralight scalars adopting a model independent general approach, taking into account scalar and pseudoscalar interactions.
- ⇒ We have briefly reviewed the current bounds on the scalar-charged leptons couplings.
- ⇒ We have obtained analytical expressions for several observables and we explored some phenomenological aspects of this scenario.
- ⇒ We have shown that an explanation to the  $g-2$  anomalies can be possible if some mechanisms exist to suppress the processes from which the diagonal bounds are obtained.
- ⇒ Also, the observables discussed in the paper are complementary.
- ⇒ The scalar could also couple to quarks, opening many hadronic and semi-leptonic channels.