

A minimal non-abelian dark sector (work in progress)

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Introduction

The existence of dark matter was confirmed by many independent observations through its gravitational effect.

Properties of DM

- Neutral
- Non-relativistic
- Stable
- Weakly interacting with SM particles

$$\Omega_{\text{DM}} h^2 = 0.120 \pm 0.001$$

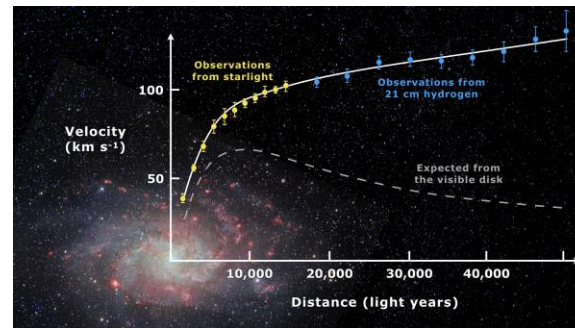
N. Aghanim et al. (Planck), 2018, arXiv: 1807.06209

E. Aprile et al. (XENON1T), 2018, arXiv: 1705.06655

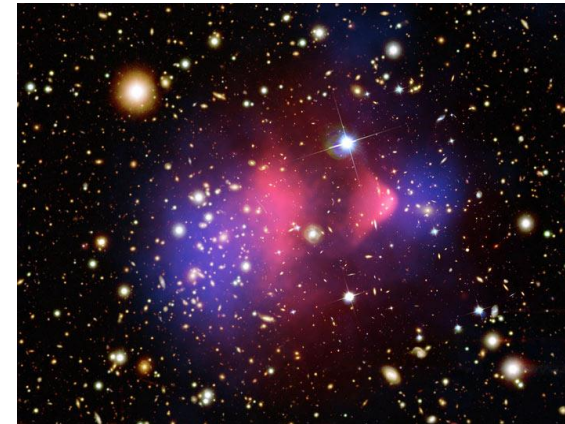
"Galaxy rotation curve." Wikipedia, 28 June 2021.

A minimal non-abelian dark matter

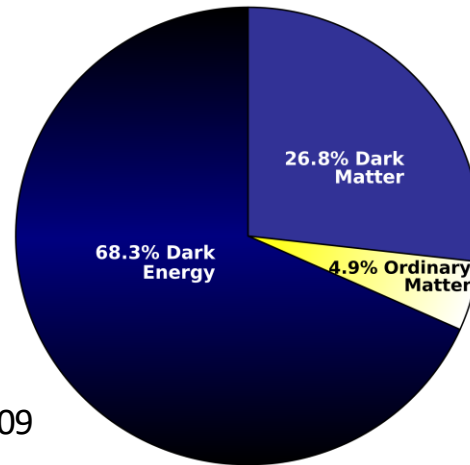
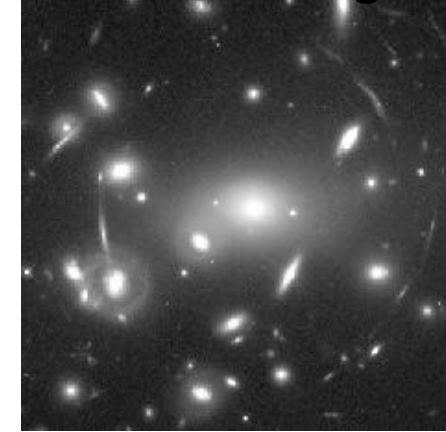
Galaxies rotation



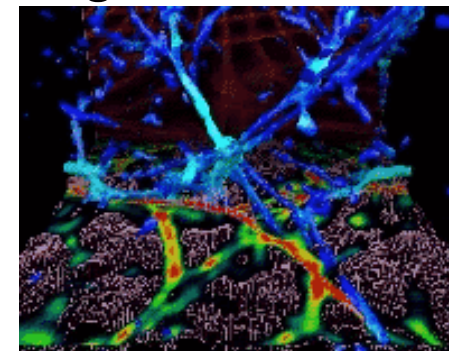
Bullet cluster



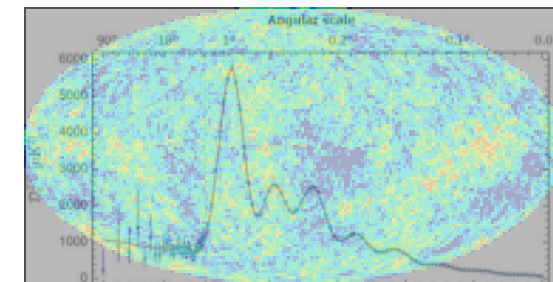
Grav. lensing



Large scale structure



CMB: WMAP & PLANCK



Introduction

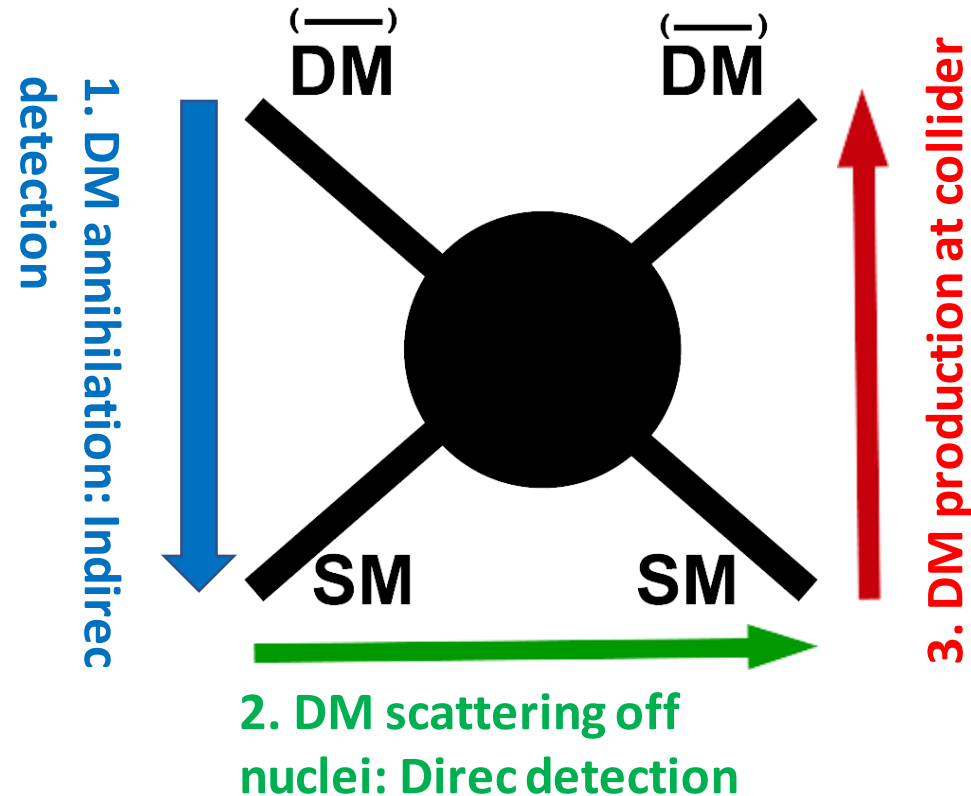
How can we observe DM?

Signatures from neutralino annihilation in halo, core of the Earth and Sun

- photons,
- Anti-protons
- positrons,
- Neutrinos

Neutrino telescopes:

- Amanda
- Icecube
- Antares



Signature from energy deposition from nuclei recoil: LUX, XENON, WARP,

LHC signatures

- mono-jet
- mono-photon
- mono-Z
- mono Higgs

- VBF+MET
- soft leptons+MET
-

New particles and DM candidates

Goal: to create a **vector model** of dark matter with **fermion propagator** based on a new SU(2) group. This is the minimal setup.

New particles

- 1 Three gauge boson from a new $SU(2)_D$ ($V_{D\pm}^0, Z'$).
- 2 A scalar H_D : giving mass to new gauge boson.
- 3 D-fermions ($t_D, b_D, c_D \dots, \tau_D, \nu_{\tau D}, \mu_D \dots$).
- 4 A scalar Φ : giving mass to SM particles and D-fermions.

DM candidates (electrically neutral and carried D-charge)

- 1 scalars: $h1_{D+}^0, h2_{D+}^0$.
- 2 fermions(D-neutrinos): $\nu_{\tau D}, \nu_{\mu D}, \nu_{eD}$.
- 3 vectors: $V_{D\pm}^0$.

***We have three kinds of DM candidates!**

New Lagrangian

$$\mathcal{L}_{\text{gauge}} = -\frac{1}{4} V^{\mu\nu} V_{\mu\nu}$$

$$\mathcal{L}_{\text{scalar}} = |D_\mu H|^2 + |D_\mu \Phi|^2 + |D_\mu H_D|^2 - V(H, \Phi, H_D)$$

$$\mathcal{L}_{\text{D-fermion}} = \bar{F}_L i \not{D} F_L + \bar{f}_R i \not{D} f_R + \text{Yukawa terms}$$

$$D = \partial_\mu - (igW_\mu^a T^a + ig' Y B_\mu) - (ig_D V_\mu^a T_D^a)$$

The mass of vector bosons

$$M_W^2 = \frac{1}{4} g^2 (v^2 + v_\phi^2), \quad M_Z^2 = \frac{1}{4} (g^2 + g'^2) (v^2 + v_\phi^2),$$

$$M_Z'^2 = \frac{g_D^2}{4} v_D^2, \quad M_V^2 = \frac{1}{4} g_D^2 (4v_\phi^2 + v_D^2)$$

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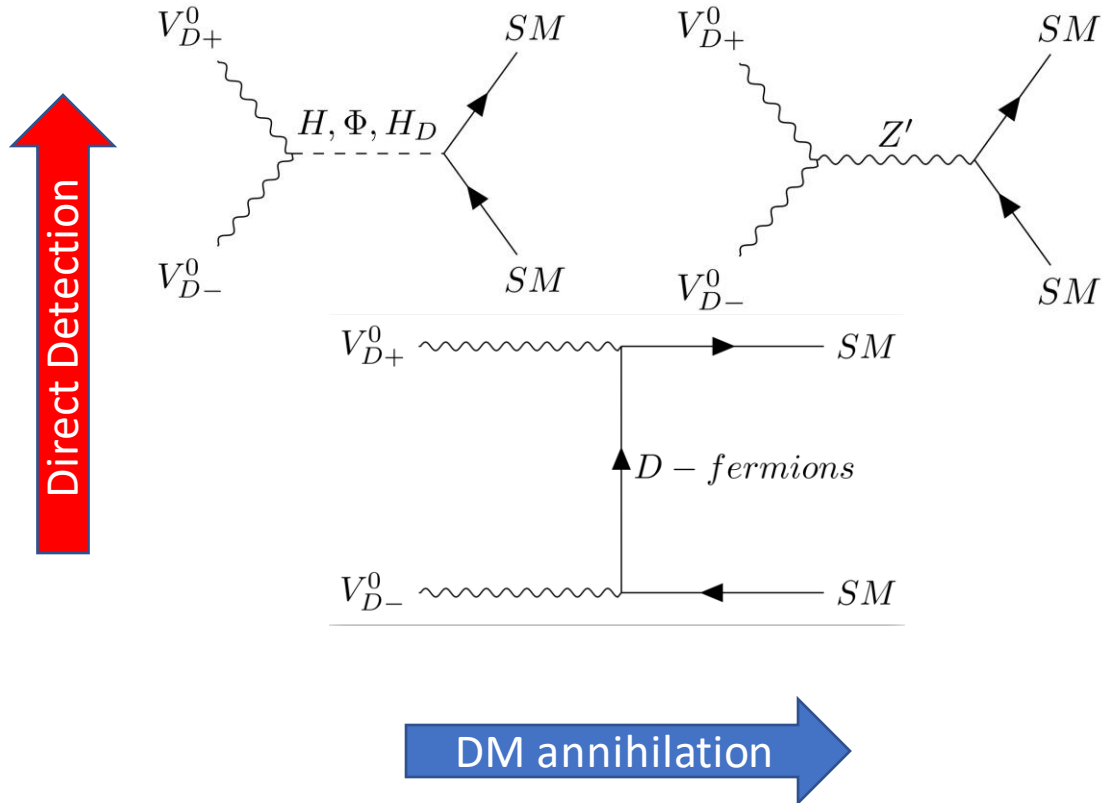
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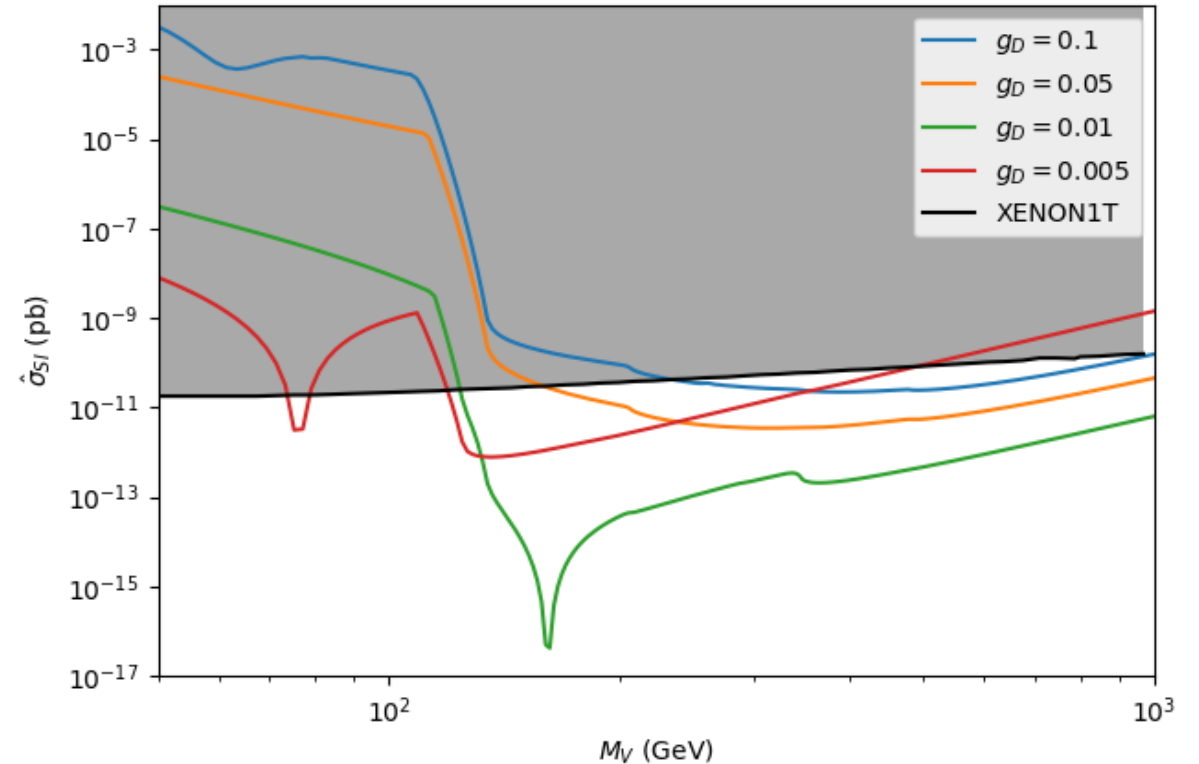
Results

Example: Vector dark matter **Packages:** LanHEP, CalcHEP, micrOMEGAs.

Feynman diagrams



Spin-independent cross section off proton



Future plan

- Dark matter (collider & non-collider) phenomenology
- Z' phenomenology ($Z' \rightarrow$ leptons)
- We can explain neutrino mass (Interaction of D-neutrinos and SM-neutrinos)
- Matter-antimatter asymmetry (new source of CP violation from D-fermion sector)

Thank you

Backup

Input parameters

We have 19 input parameter (independent) and 16 dependent ones

These are used in model implementation

And scanning in **micrOMEGA**

Scalar masses



Scalar couplings



Scalar couplings



Scalar couplings



```
gd = 0.005;  
DeltaM = 1;  
MassPhi = 2.357873E+02 ;  
MassHD = 3.460817E+02 ;  
Massp0 = 1.344139E+02 ;  
Masspp = 2.077473E+02 ;  
Masspm = 4.811826E+02 ;  
lbHPhiHD = 6.049728E+00 ;  
lbPhiHD2 = -3.384986E-01 ;  
lbPhi2 = -1.979089E+00 ;  
lbHPhi1 = -3.114088E+00 ;  
MasstD = 5.917625E+02 ;  
MassbD = 4.206233E+03 ;  
MasslD = 7.014723E+03 ;  
MassnD = 1.651420E+03 ;  
SinA12 = 1.918818E-01 ;  
SinA13 = -8.982609E-03 ;  
SinA23 = -9.894726E-01 ;
```