

Probing heavy neutral leptons at the electron positron collider

Arindam Das
Osaka University

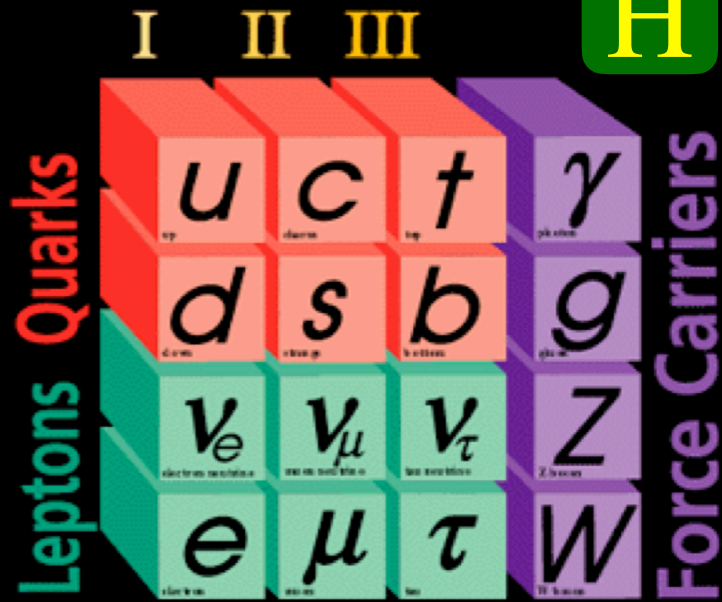


Searches for Long – Lived Particles October 28, 2020

The Standard Model of Particle Interactions

Three Generations of Matter

H



Over the decades experiments have found each and every missing pieces

Verified the facts that they belong to this family

Finally at the Large Hadron collider Higgs has been observed

→ Its properties must be verified

Strongly established with interesting shortcomings

Few of the very interesting anomalies :

Tiny neutrino mass and flavor mixings
Relic abundance of dark matter ...

Neutrino oscillation experiment : SNO, Super - K, etc .

Unkown

- Nature : Majorana/ Dirac
- Ordering : Normal/Inverted
- Nature of the mixing between the mass and the flavor eigenstates

2 SM can not explain them

Models of Neutrino to explain the origin of neutrino mass

There is a wide variety of neutrino mass models

The predicted models extend the SM minimally

At the tree level SM can be extended by Singlet fermions

→ Right handed neutrinos seesaw mechanism
inverse seesaw mechanism

Minkowski, Ramond, Slansky, Yanagida, Gell – Mann, Glashow, Mohapatra, Senjanovic

Linear, Hybrid

Alternative ideas extending the Standard Model

→ SU(2) triplet scalar : type – II seesaw

Schechter, Valle, Lazarides, Shafi, Wetterich, Mohapatra, Senjanovic

SU(2) triplet fermion : type – III seesaw

Foot, Lew, He, Joshi, Ma

One – loop and even at 2/3 – loop models also exist

→ For example : Ma – model, Zee – Model, Zee – Babu model, BNT, KNT, etc .

Babu, Leung, Hirsch, King, Nasri, Volkas Dev, Pilaftsis

→ Gauge extended : U(1), Left – Right

Pati, Salam; Mohapatra, Pati; Senjanovic, Mohapatra Buchmuller, Greub; Fileviez Perez, Han, Li; Deppisch, Desai, Valle; Kang, Ko, Li; Heeck, Teresi; Gluza, Chakraborty Keung, Senjanovic; Ferrari et . al . ; Nemevsek, Nesti, Senjanovic, Zhang; Chen, Dev, Mohapatra; Dev, Mohapatra, Zhang; Dev, Goswami, Mitra AD, Dev, Mohapatra; AD, Okada, Raut

Particle content

Dobrescu, Fox; Cox, Han, Yanagida; AD, Okada, Raut; AD, Dev, Okada;
Chiang, Cottin, AD, Mandal; AD, Takahashi, Oda, Okada

	$SU(3)_c$	$SU(2)_L$	$U(1)_Y$		$U(1)_X$
q_L^i	3	2	+1/6	x_q	$= \frac{1}{6}x_H + \frac{1}{3}x_\Phi$
u_R^i	3	1	+2/3	x_u	$= \frac{2}{3}x_H + \frac{1}{3}x_\Phi$
d_R^i	3	1	-1/3	x_d	$= -\frac{1}{3}x_H + \frac{1}{3}x_\Phi$
ℓ_L^i	1	2	-1/2	x_ℓ	$= -\frac{1}{2}x_H - x_\Phi$
e_R^i	1	1	-1	x_e	$= -x_H - x_\Phi$
H	1	2	+1/2	x'_H	$= \frac{1}{2}x_H$
N_R^i	1	1	0	x_ν	$= -x_\Phi$
Φ	1	1	0	x'_Φ	$= 2x_\Phi$

$m_{Z'} = 2 g_X v_\Phi$
 x_H, x_Φ will appear
the coupling with Z'

3 generations of
SM singlet right handed
neutrinos (anomaly free)

Charges **before**
the anomaly cancellations

Charges **after**
Imposing the
anomaly
cancellations

$U(1)_X$ breaking

$$\mathcal{L}_Y \supset - \sum_{i,j=1}^3 Y_D^{ij} \bar{\ell}_L^i H N_R^j - \frac{1}{2} \sum_{i=k}^3 Y_N^k \bar{N}_R^k c N_R^k + \text{h.c.},$$

$$m_D^{ij} = \frac{Y_D^{ij}}{\sqrt{2}} v_h$$

$$m_{N^i} = \frac{Y_N^i}{\sqrt{2}} v_\Phi$$

$$m_\nu = \begin{pmatrix} 0 & M_D \\ M_D^T & M_N \end{pmatrix} \quad m_\nu \simeq -M_D M_N^{-1} M_D^T$$

Seesaw mechanism

Direct interaction of the Right Handed Neutrinos through light – heavy mixing

Antusch, Drewes, AD, Dev, Fischer, Deppisch

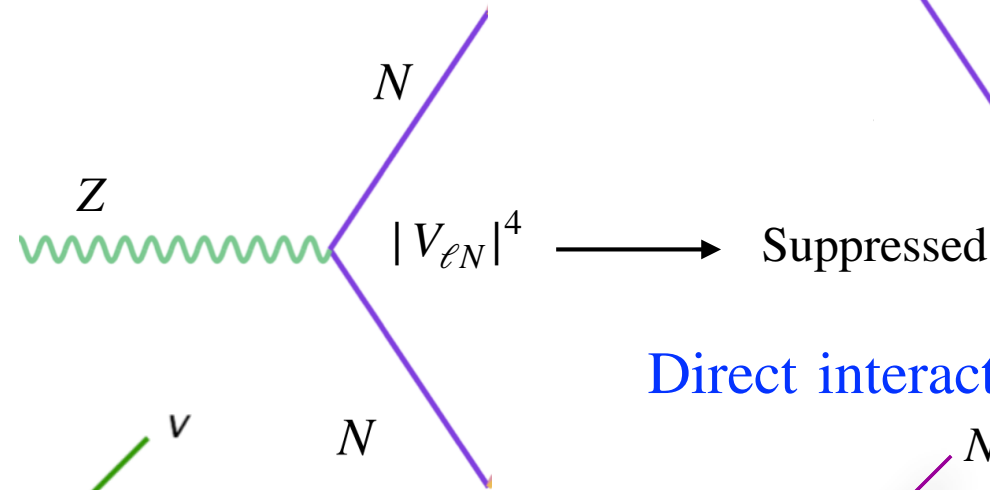
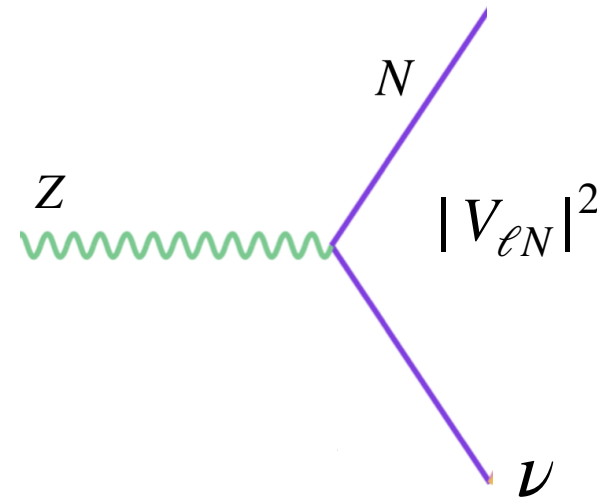
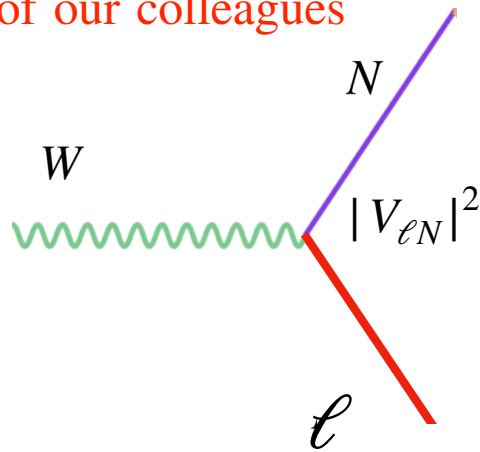
Flavor eigenstate can be expressed in terms of the mass eigenstate

and many of our colleagues

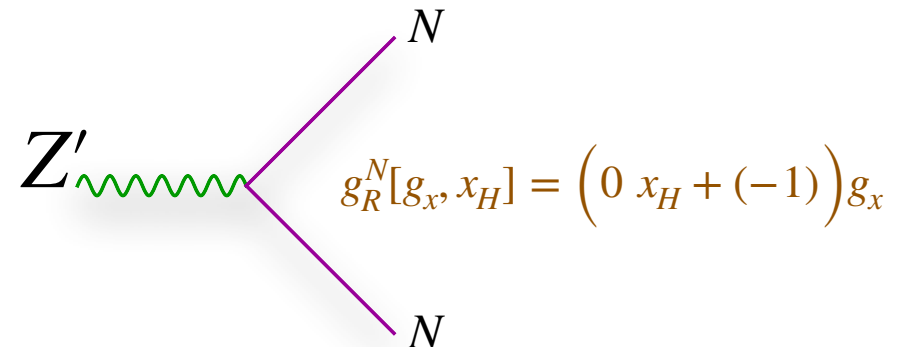
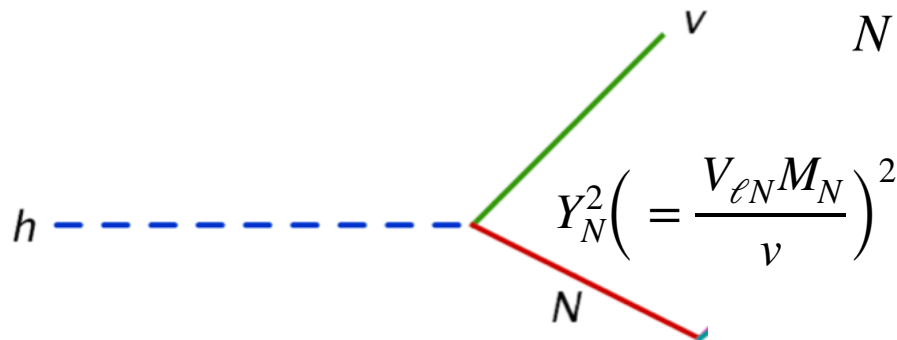
$$\nu_\ell \simeq U_{\ell m} \nu_m + V_{\ell n} N_n$$

PMNS matrix

$$M_D M_N^{-1}$$



Direct interaction of the RHNs



Properties of the model and phenomenology

New particles

Z' boson

Heavy Majorana Neutrino

$U(1)_X$ Higgs boson

Phenomenology

Z' boson production and decay

Z' boson mediated processes

Heavy neutrino production

$U(1)_X$ Higgs phenomenology : Vacuum Stability

Dark Matter

collider

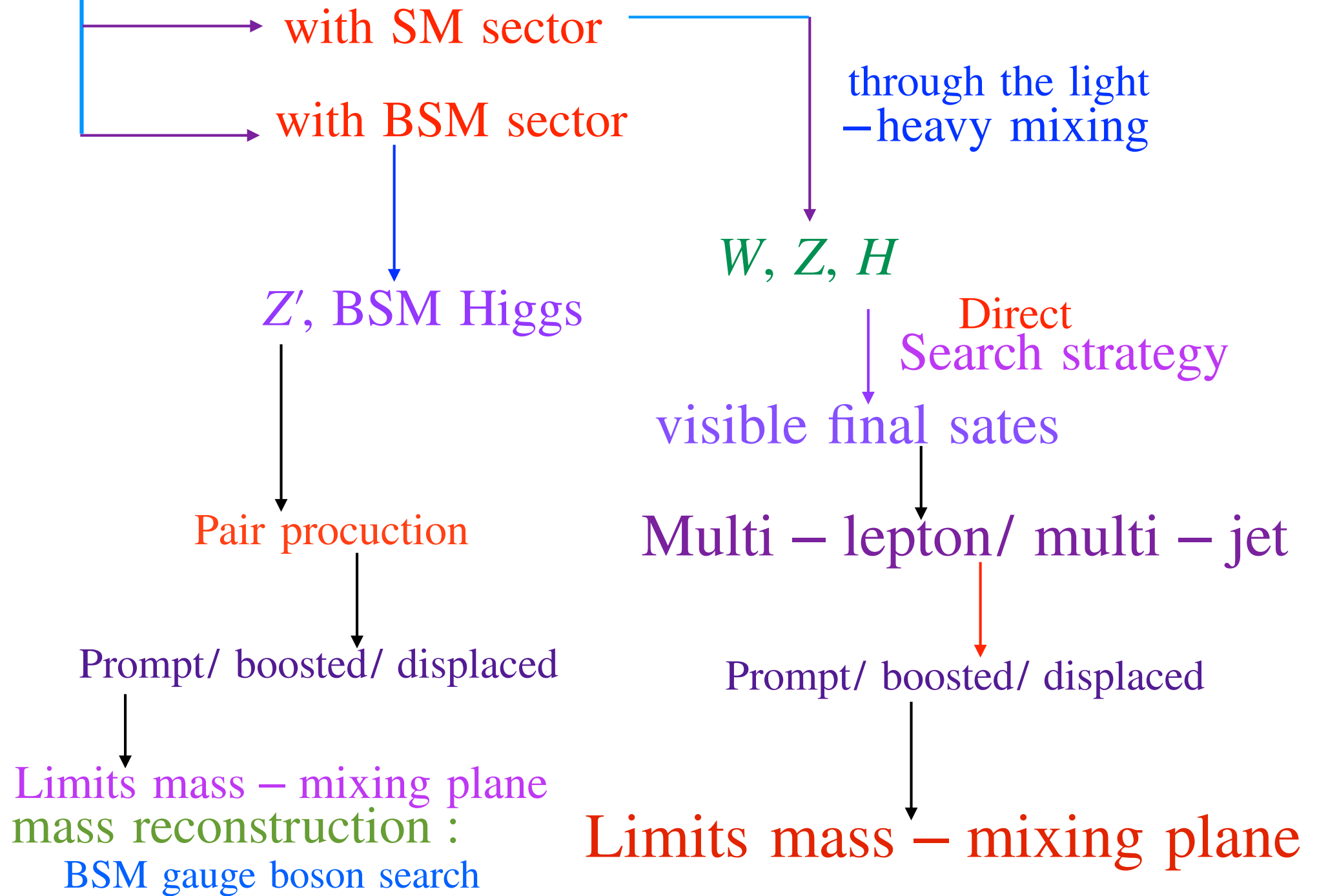
Leptogenesis and many more

Dev, Pilaftsis; Iso, Okada, Orikasa

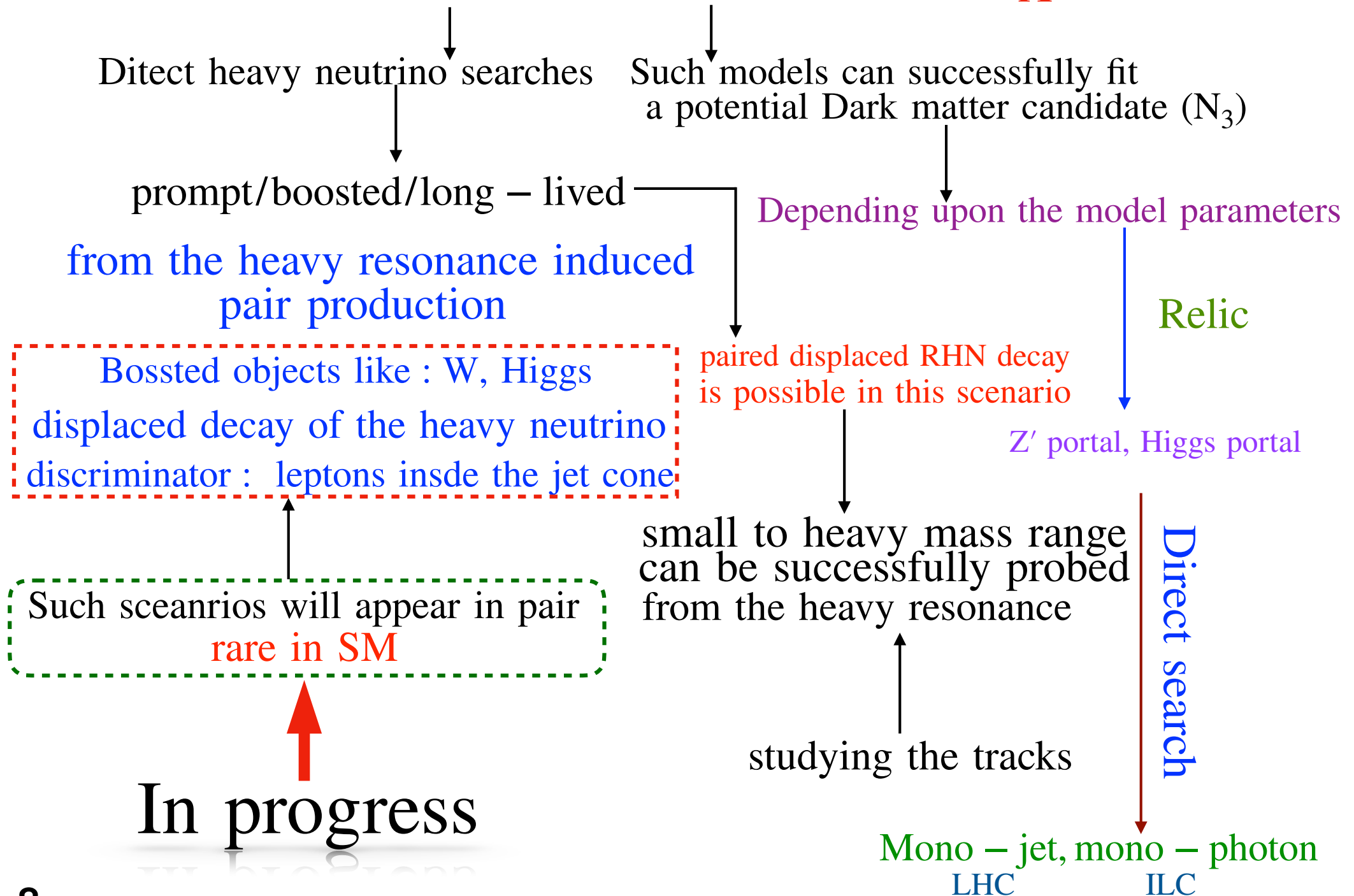
Orikasa, Okada, Yamada; Dev, Mohapatra, Zhang

Z' boson and heavy neutrino phenomenology

Heavy neutrino interactions



Other interesting aspects in the $U(1)_X$ scenario



Impression/s at the ILC

$$\sigma(e^+ e^- \rightarrow \bar{\nu}_\alpha N_i) = \sigma_{\text{ILC}} |\mathcal{R}_{\alpha i}(\delta, \rho, y)|^2$$

1207.3734

Using the general parameters and Casas – Ibarra conjecture

$$\begin{pmatrix} \cos(x + iy) & \sin(x + iy) \\ -\sin(x + iy) & \cos(x + iy) \end{pmatrix}$$

$$|\mathcal{R}_{\alpha i}(\delta, \rho, y)|^2$$

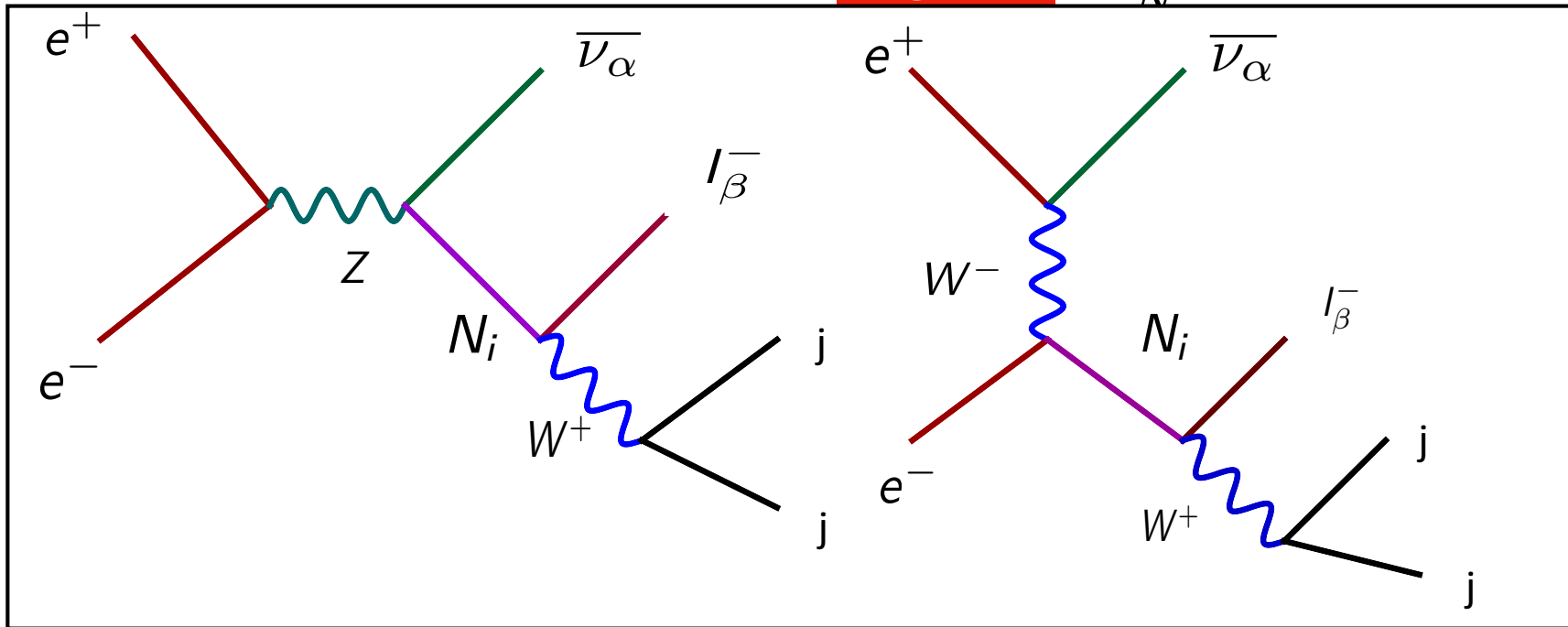
branching ratios

$$N_i \rightarrow \ell_\alpha^- W^+ / \nu_\alpha Z / \nu_\alpha h$$

Leading mode

$$\mathcal{N}^\dagger \mathcal{R} \simeq U_{\text{MNS}}^\dagger \mathcal{R} \text{ because } |\epsilon_{\alpha\beta}| \ll 1$$

Signals $M_N = 150 \text{ GeV}$

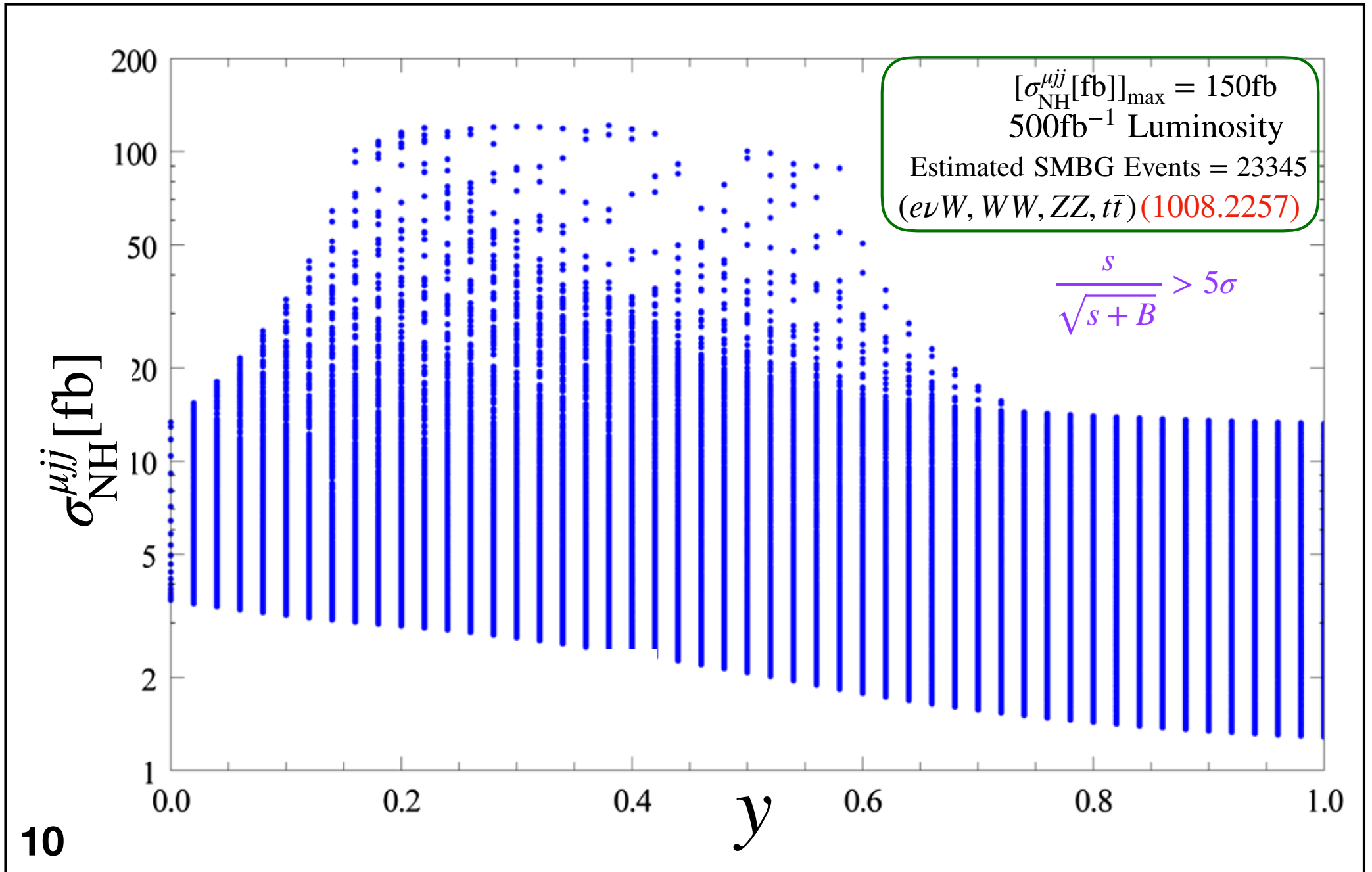


We scan over the phases and then general parameter to find the cross section as a function of the general parameter 'y', $-\pi < \delta, \rho < \pi, 0 < y < 1$

$e^+ e^- \rightarrow \nu N$, followed by the decay $N \rightarrow \ell W$ ($\ell = \mu$) $W \rightarrow q\bar{q}'$

$M_N = 150$ GeV $\sqrt{s} = 500$ GeV

1207.3734



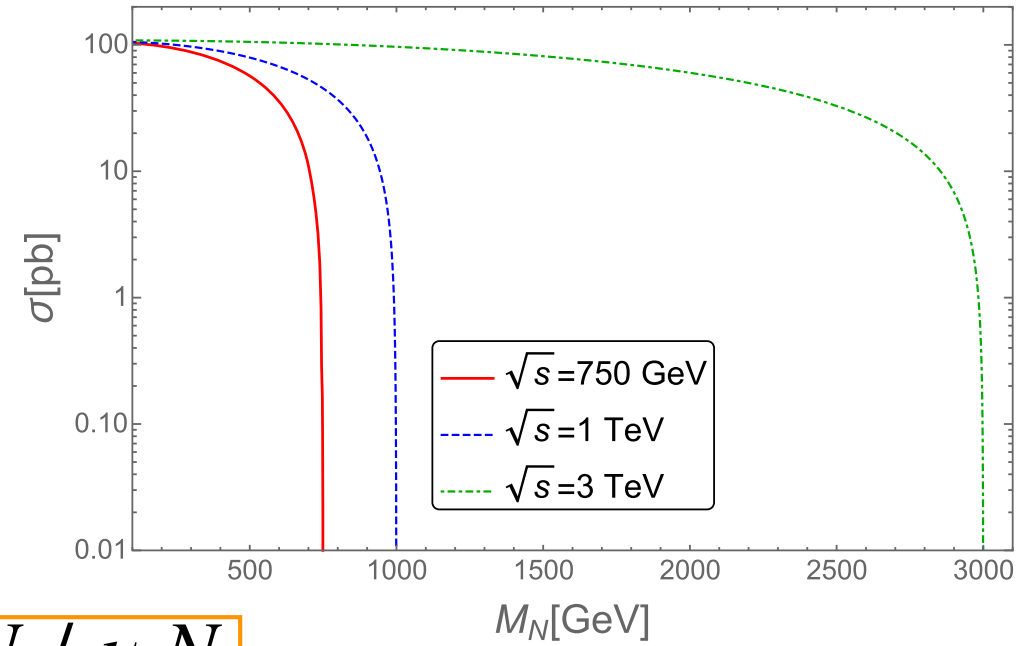
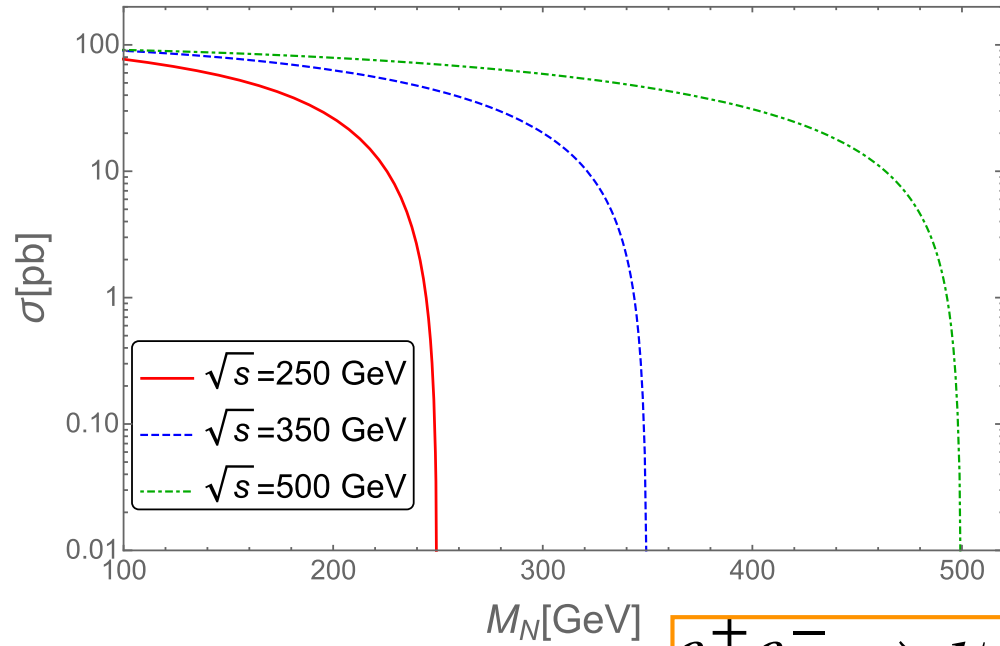
Production of the heavy neutrinos at the Linear Collider using fat jet

Normalized by mixing

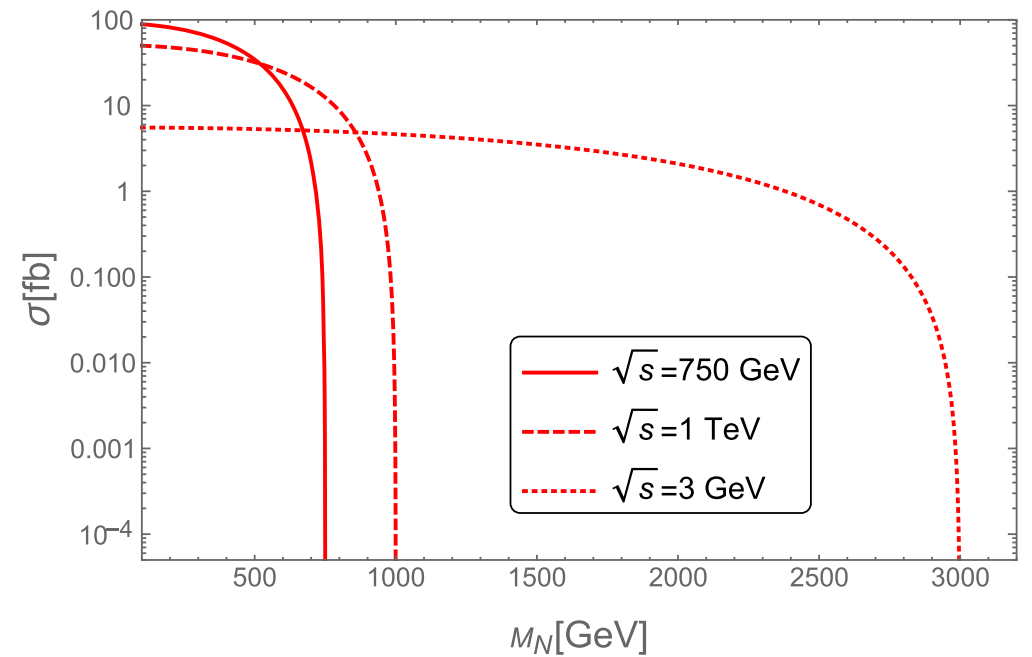
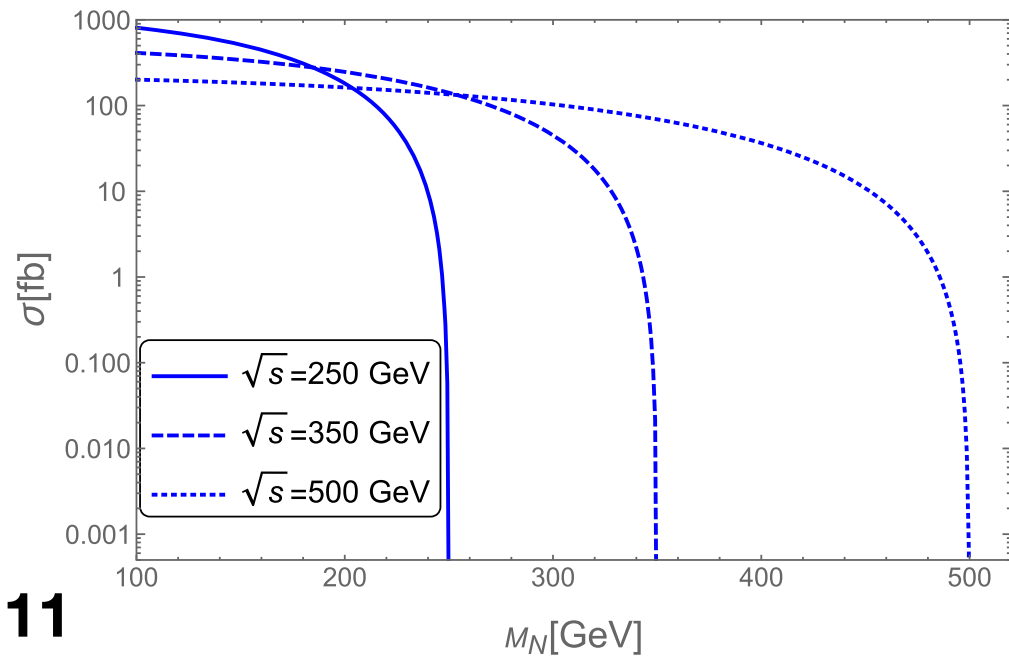
1811.04291

$$e^+e^- \rightarrow \nu_1 N_1$$

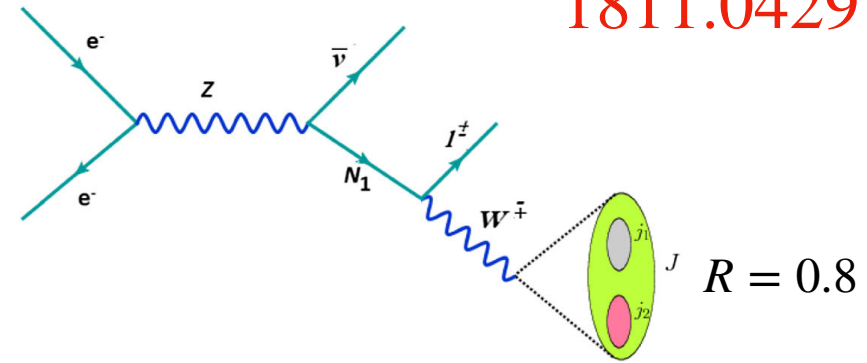
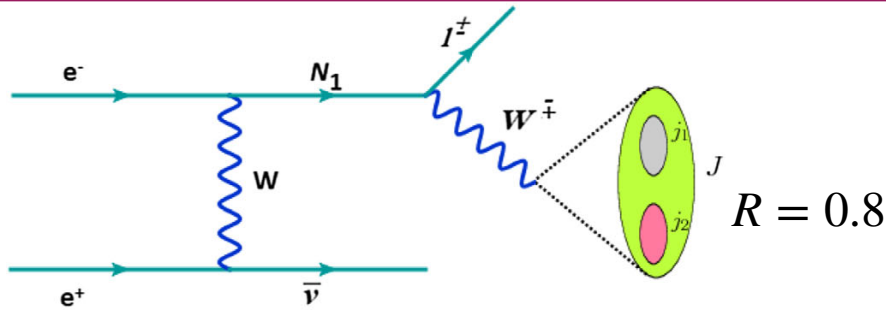
AD, Jana, Nandi, Mandal Chakraborty, Mitra, Shil
Banerjee, Dev, Ibarra, Mandal, Mitra Antusch, Fischer



$$e^+e^- \rightarrow \nu_2 N_2 / \nu_3 N_3$$



$e + J + p_T^{\text{miss}}$ final states at the linear colliders.



- Transverse momentum for fat-jet $p_T^J > 150$ GeV for M_N mass range 400 GeV-600 GeV and $p_T^J > 250$ GeV for M_N mass range 700 GeV-900 GeV.
- Transverse momentum for leading lepton $p_T^{e^\pm} > 100$ GeV for M_N mass range 400 GeV-600 GeV and $p_T^{e^\pm} > 200$ GeV for M_N mass range 700 GeV-900 GeV.
- Polar angle of lepton and fat-jet $|\cos \theta_e| < 0.85$, $|\cos \theta_J| < 0.85$.
- Fat-jet mass $M_J > 70$ GeV.

1 TeV e^-e^+ collider

- Transverse momentum for fat-jet $p_T^J > 250$ GeV for the M_N mass range 700 GeV-900 GeV and $p_T^J > 400$ GeV for M_N mass range 1 – 2.9 TeV.
- Transverse momentum for leading lepton $p_T^{e^\pm} > 200$ GeV for M_N mass range 700 – 900 GeV and $p_T^{e^\pm} > 250$ GeV for M_N mass range 1 – 2.9 TeV.
- Polar angle of lepton and fat-jet $|\cos \theta_e| < 0.85$, $|\cos \theta_J| < 0.85$.

3 TeV e^-e^+ collider

- Fat-jet mass $M_J > 70$ GeV.

Cut flow for the signal and background events for the final state $e^\pm + J + p_T^{\text{miss}}$ for $M_N = 500$ GeV at the $\sqrt{s} = 1$ TeV linear collider. The signal events are normalized by the square of the mixing.

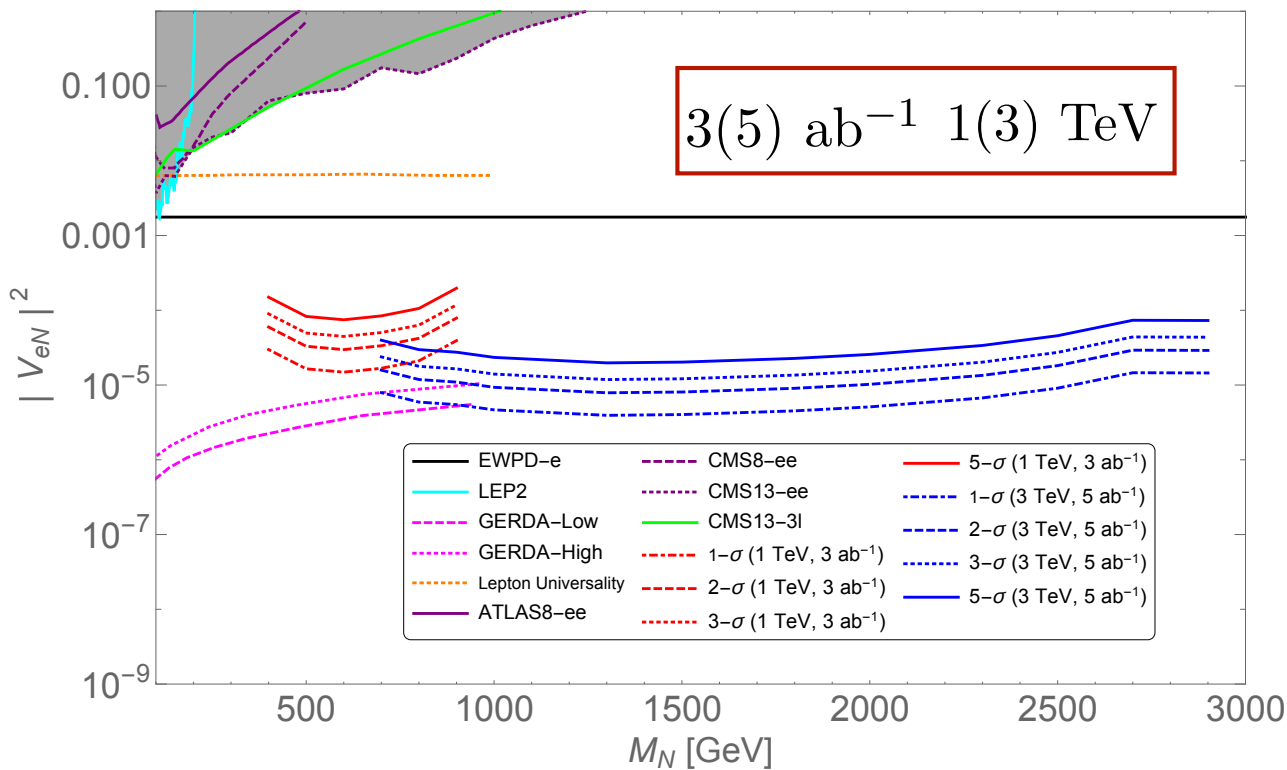
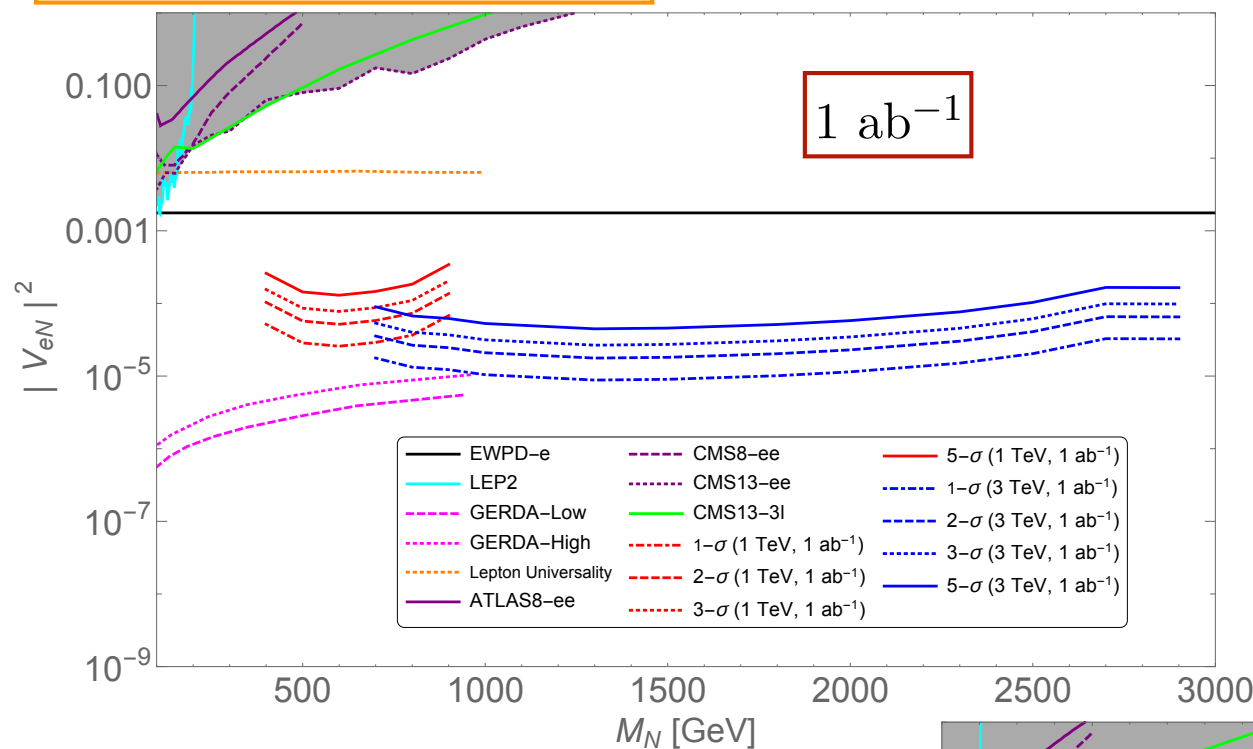
Cuts	Signal	Background				Total
		$\nu_e e W$	WW	ZZ	$t\bar{t}$	
Basic Cuts	12,996,200	201,586	72,244	7,200	4,300	285,330
$ \cos \theta_J \leq 0.85$	12,789,800	148,802	44,910	3,800	4,100	201,600
$ \cos \theta_e \leq 0.85$	12,671,800	79,008	40,574	2,800	3,900	126,280
$p_T^J > 150$ GeV	12,308,300	70,669	40,490	2,300	3,200	116,660
$M_J > 70$ GeV	10,923,100	62,303	37,043	2,100	2,300	103,700
$p_T^\ell > 100$ GeV	10,714,500	57,076	33,488	1,400	1,530	93,400

Cut flow for the signal and background events for the final state $e^\pm + J + p_T^{\text{miss}}$ for $M_N = 800$ GeV at the $\sqrt{s} = 1$ TeV linear collider. The signal events are normalized by the square of the mixing.

Cuts	Signal	Background				Total
		$\nu_e e W$	WW	ZZ	$t\bar{t}$	
Basic Cuts	8,684,990	201,586	72,244	7,200	4,300	285,330
$ \cos \theta_J \leq 0.85$	8,649,570	148,802	44,910	3,800	4,100	201,600
$ \cos \theta_e \leq 0.85$	8,618,420	79,008	40,574	2,800	3,900	126,280
$p_T^J > 250$ GeV	7,681,440	59,001	40,329	2,303	2,720	104,354
$M_J > 70$ GeV	7,176,280	53,990	36,997	2,187	2,282	95,437
$p_T^\ell > 200$ GeV	7,080,200	38,729	26,208	942	613	66,493

Mass-mixing limit plots

$\sqrt{s} = 1 \text{ TeV}$ (red band) and 3 TeV (blue band)



Alternative scenario under $U(1)_X$

AD, Okada, Raut

AD, Okada, Okada, Raut

	$SU(3)_c$	$SU(2)_L$	$U(1)_Y$	$U(1)_X$
q_{L_i}	3	2	1/6	$(1/6)x_H + (1/3)$
u_{R_i}	3	1	2/3	$(2/3)x_H + (1/3)$
d_{R_i}	3	1	-1/3	$-(1/3)x_H + (1/3)$
ℓ_{L_i}	1	2	-1/2	$(-1/2)x_H - 1$
e_{R_i}	1	1	-1	$-x_H - 1$
H	1	2	-1/2	$(-1/2)x_H$
$N_{R_{1,2}}$	1	1	0	-4
N_{R_3}	1	1	0	+5
H_E	1	2	-1/2	$(-1/2)x_H + 3$
Φ_A	1	1	0	+8
Φ_B	1	1	0	-10
Φ_C	1	1	0	-3

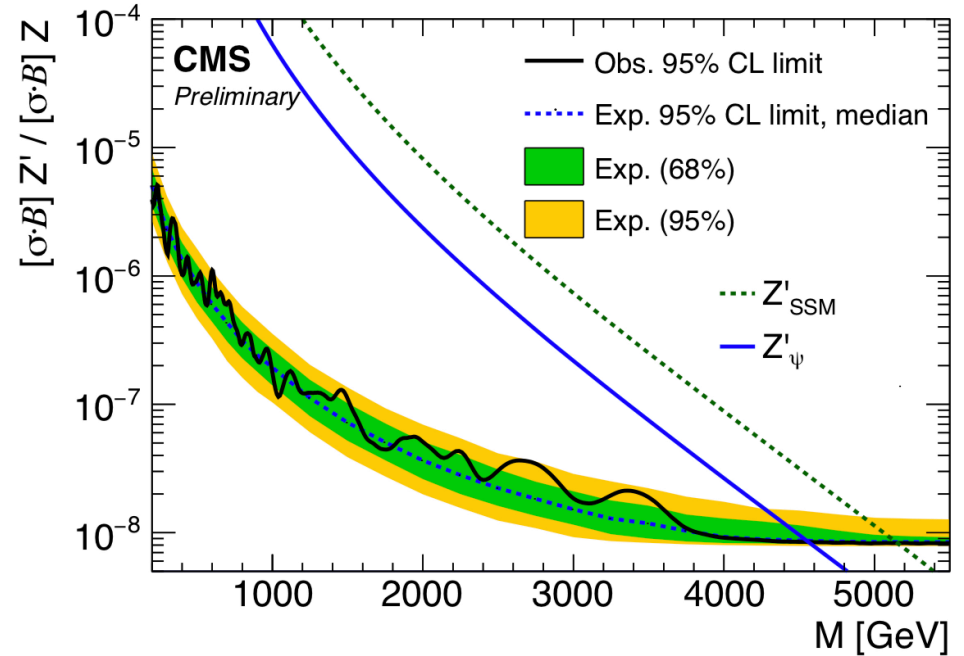
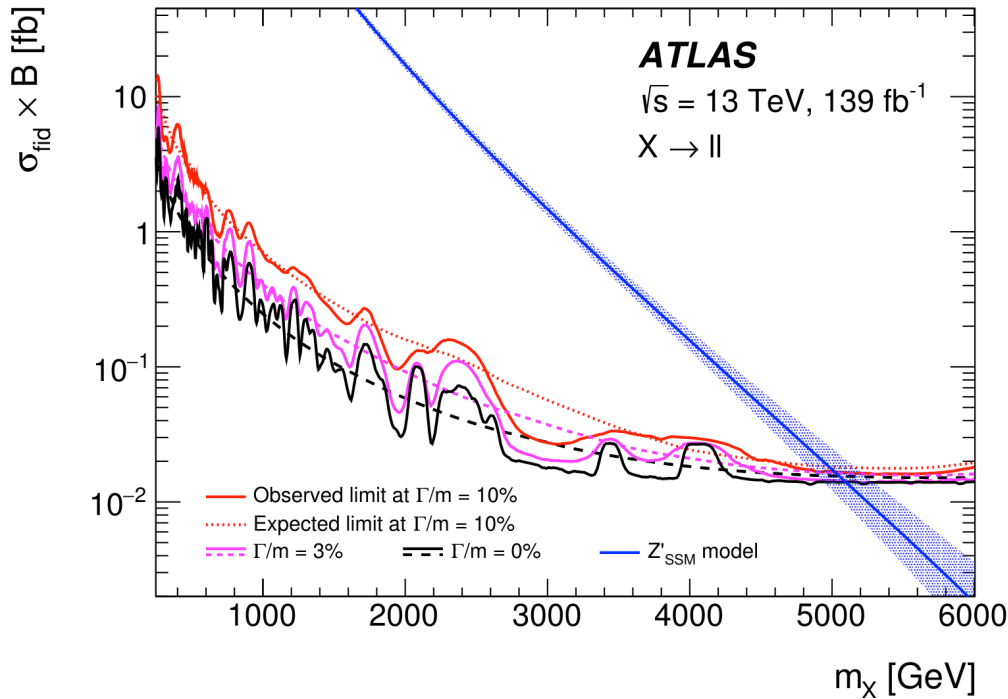
Possible alternative B - L, with $x_H = 0$

Detailed scalar sector study
In Progress

$$\mathcal{L}_Y \supset - \sum_{i=1}^3 \sum_{j=1}^2 Y_D^{ij} \overline{\ell}_L^i H_E N_R^j - \frac{1}{2} \sum_{k=1}^2 Y_N^k \Phi_A \overline{N_R^{kc}} N_R^k - \frac{1}{2} Y_N^3 \Phi_B \overline{N_R^{3c}} N_R^3 + \text{h.c.}$$

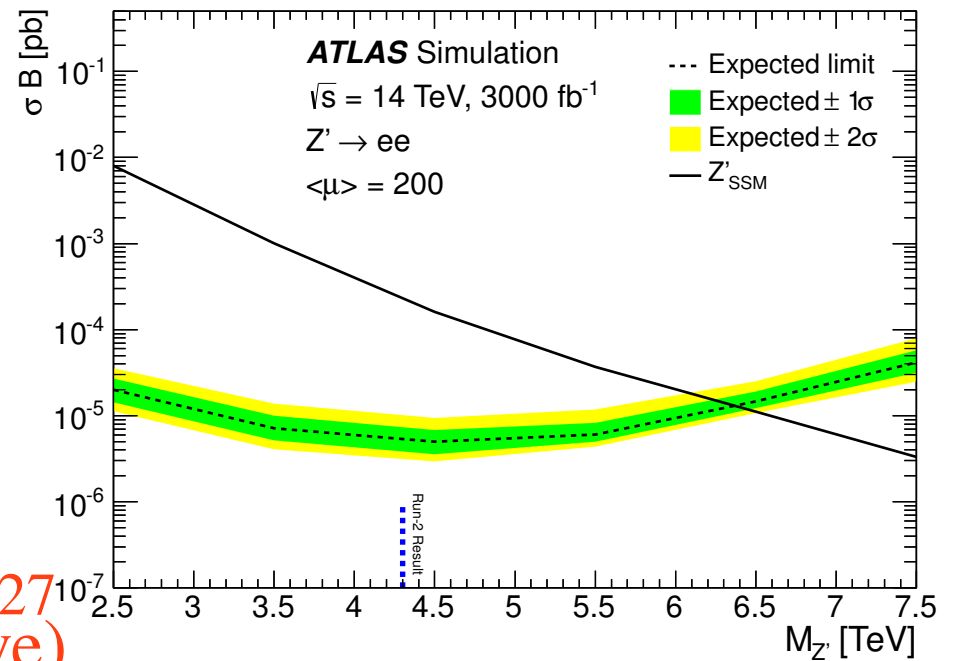
Bounds on the $U(1)_X$ gauge coupling

CMS PAS EXO – 19 – 019
 $ee(139 \text{ fb}^{-1}) + \mu\mu(140 \text{ fb}^{-1})$

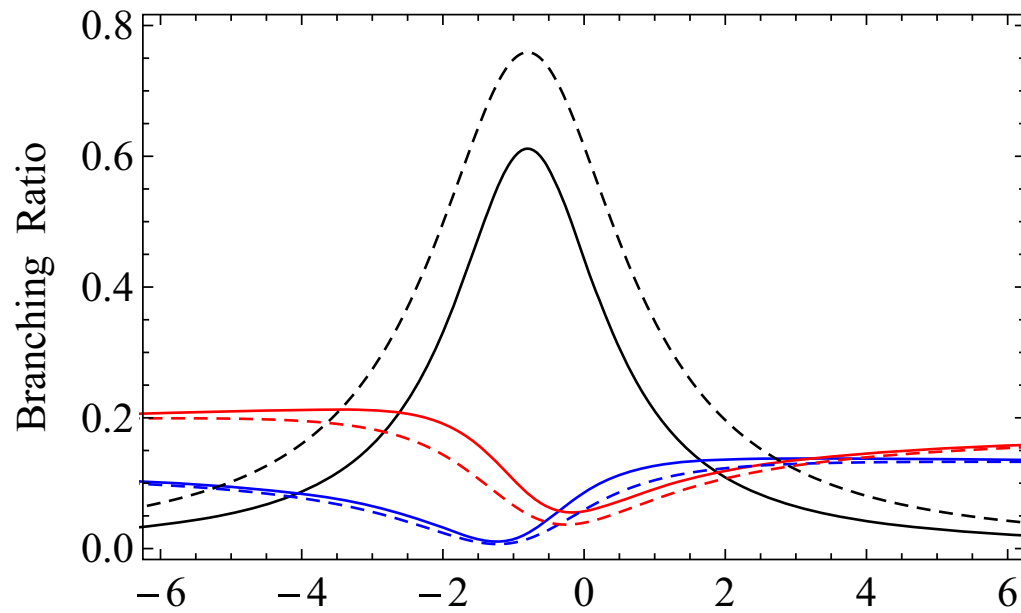


2ℓ , ATLAS : 1903.06248 (139 fb^{-1})

$$g' = \sqrt{\frac{\sigma_{\text{ATLAS/CMS}}^{\text{Observer}}}{\frac{\sigma_{\text{Model}}}{g_{\text{Model}}^2}}}$$



ATLAS – TDR – 027
 ee (Prospective)



$$m_{Z'} = 3 \text{ TeV.}$$

Solid

$$m_{N1} = m_{Z'}/4$$

$$m_{N2} > m_{Z'}/2.$$

Dashed

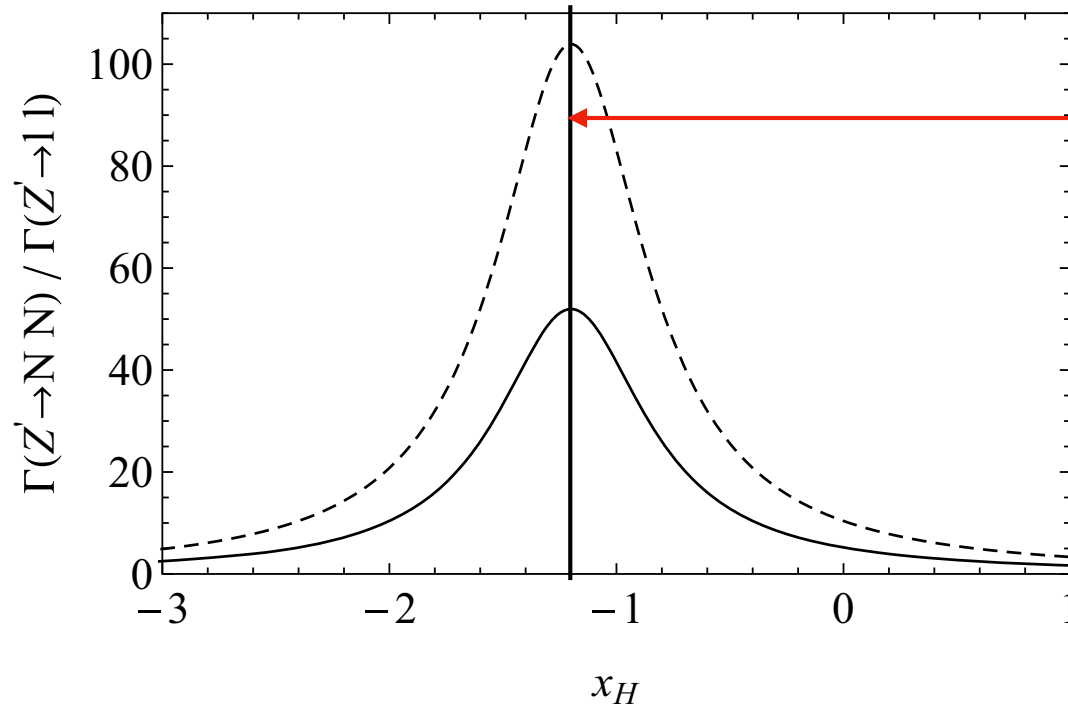
$$m_{N1,2} = m_{Z'}/4.$$

Top → bottom : Solid (Red, Black, Blue) x_H

Up and down quarks

Heavy neutrinos

Charged leptons



$$x_H = -1.2$$

Solid

$$m_{N1} = m_{Z'}/4$$

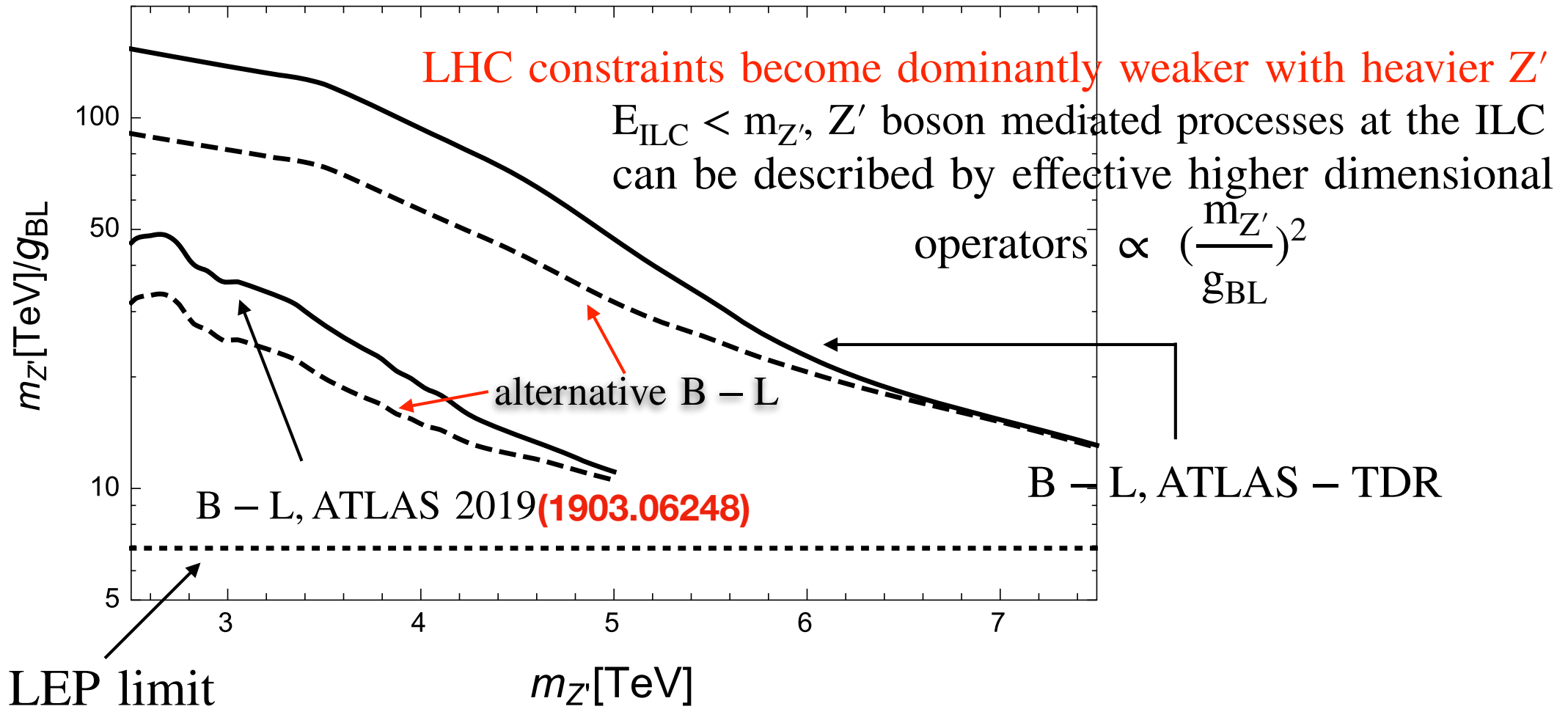
$$m_{N2} > m_{Z'}/2.$$

Dashed

$$m_{N1,2} = m_{Z'}/4.$$

$$\frac{\Gamma(Z' \rightarrow NN)}{\Gamma(Z' \rightarrow \bar{\ell}\ell)} = \frac{64}{8 + 12x_H + 5x_H^2} \left(1 - \frac{4m_N^2}{m_{Z'}^2}\right)^{3/2}$$

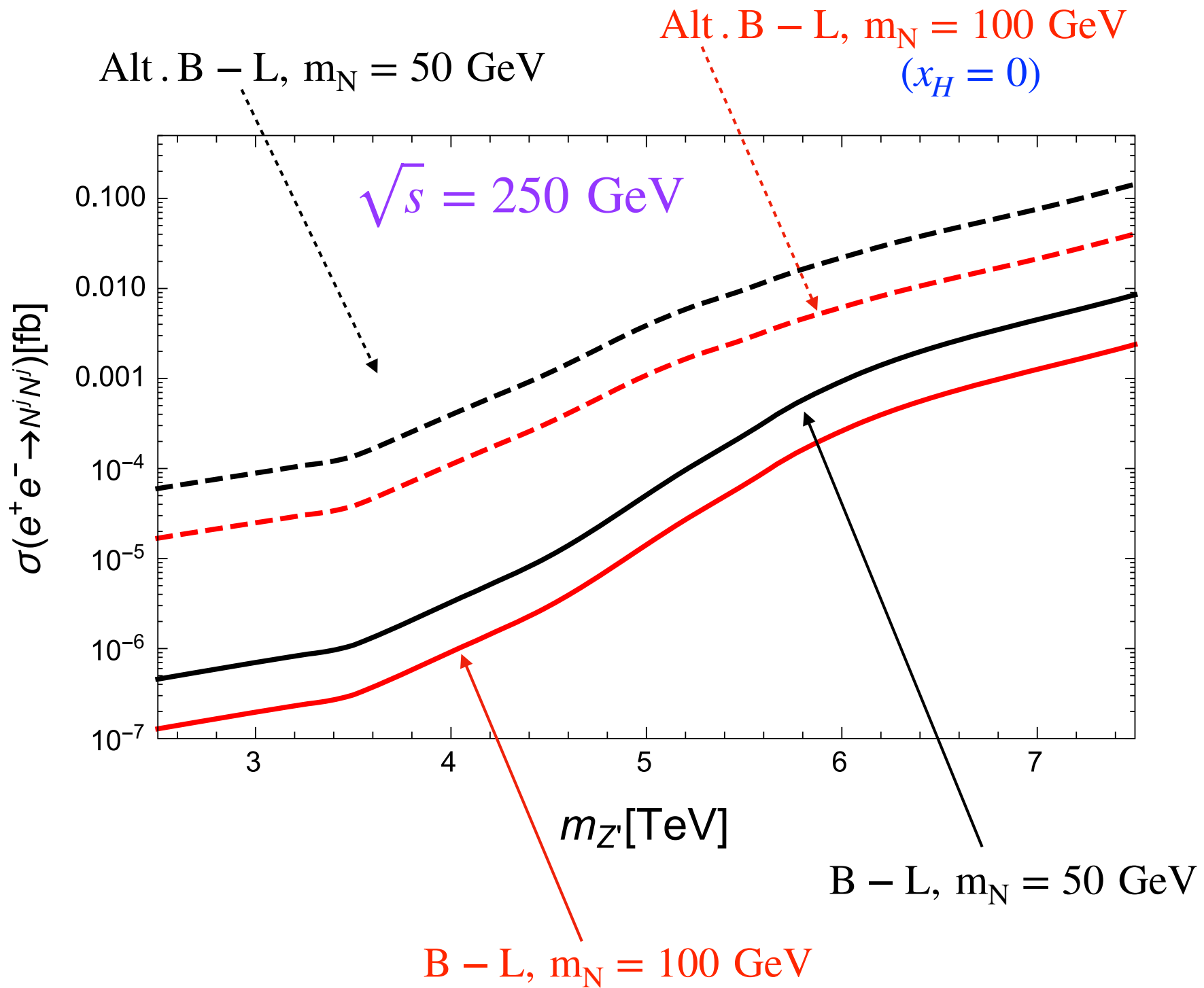
Production of the heavy neutrino at the ILC

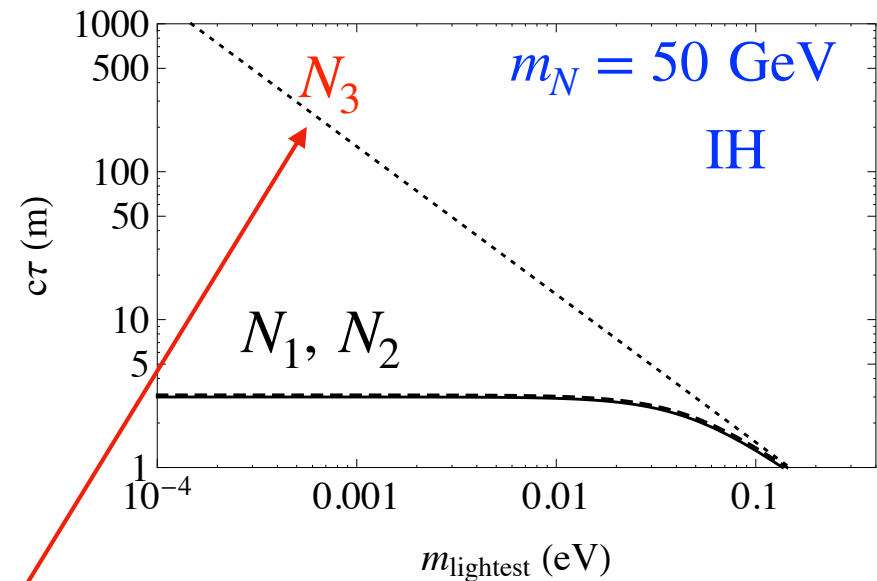
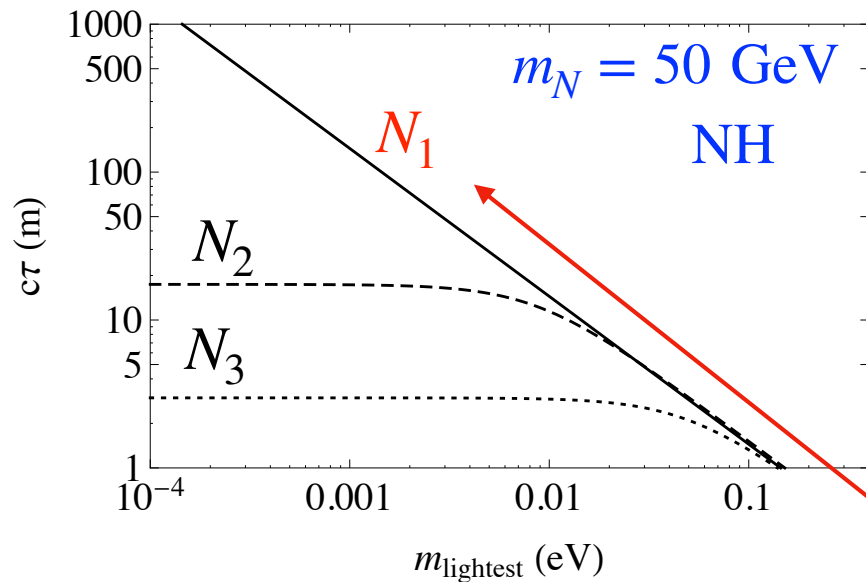


Dashed lines represent the Atl. B - L case

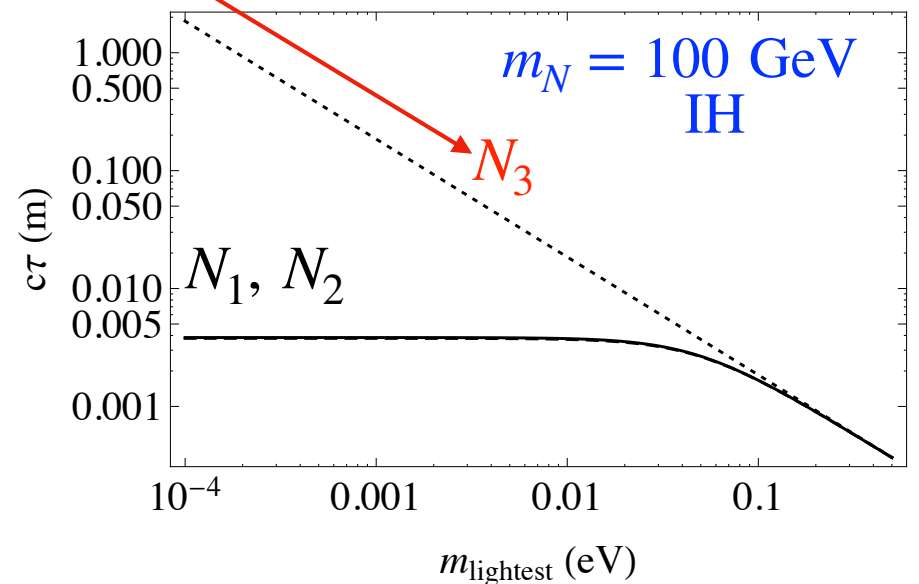
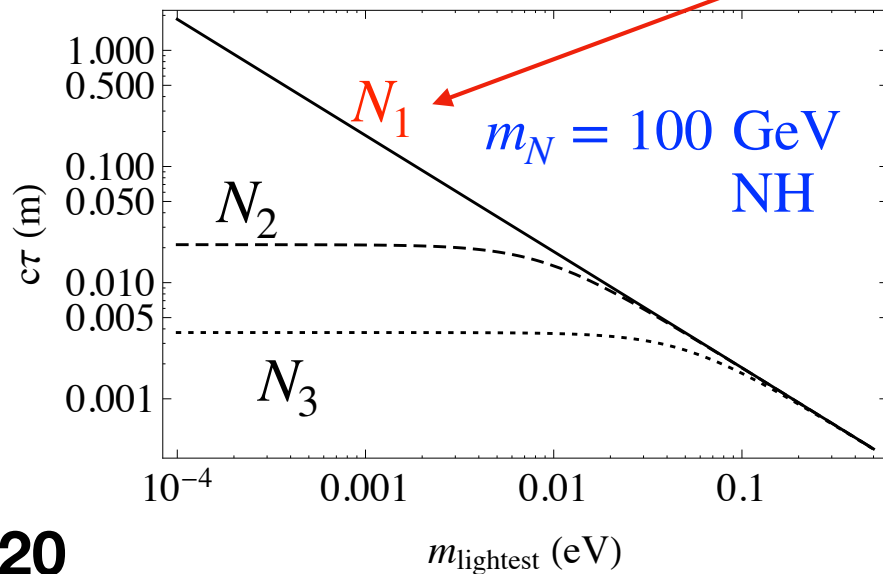
As a result ILC is a powerful machine to probe Z' beyond HL - LHC

$$\sigma(e^+e^- \rightarrow Z'^* \rightarrow N^i N^i) \simeq \frac{(Q_{Ni})^2}{24\pi} s \left(\frac{g_{\text{BL}}}{m_{Z'}}\right)^4 \left(1 - \frac{4m_{Ni}^2}{m_{Z'}^2}\right)^{\frac{3}{2}}$$





Longest lived RHN life time is inversely proportional to m_{lightest}
 $m_{\text{lightest}} \rightarrow 0$ leads to the long lived species as a potential DM candidate



Conclusions

SM can not explain the origin of the tiny neutrino mass .
We consider some benchmark models which can explain the origin of tiny neutrino mass .
These models are equipped with the heavy neutrinos under the simple extension of the SM .

These heavy neutrinos can mix with the light neutrinos .
Generalizing the mixings and reproducing the neutrino oscillation data we have studied the production of the RHNs : **prompt and boosted** .
We finally probed the light heavy mixings successfully **beyond the EWPD at the e^-e^+ colliders** .

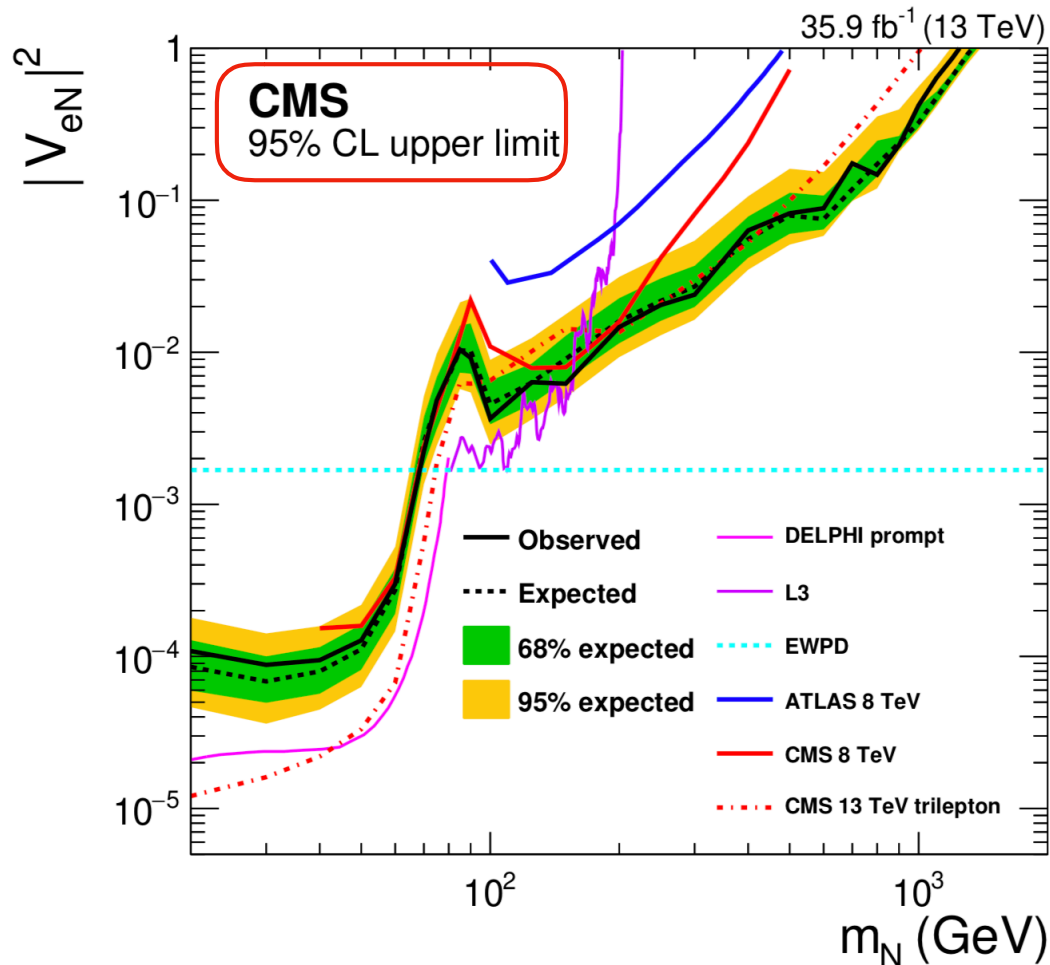
In these models there is a neutral BSM gauge boson Z' which directly interacts with the heavy neutrinos .
This could be a good source to study the long – lived RHNs .

Thank You



Back up slides

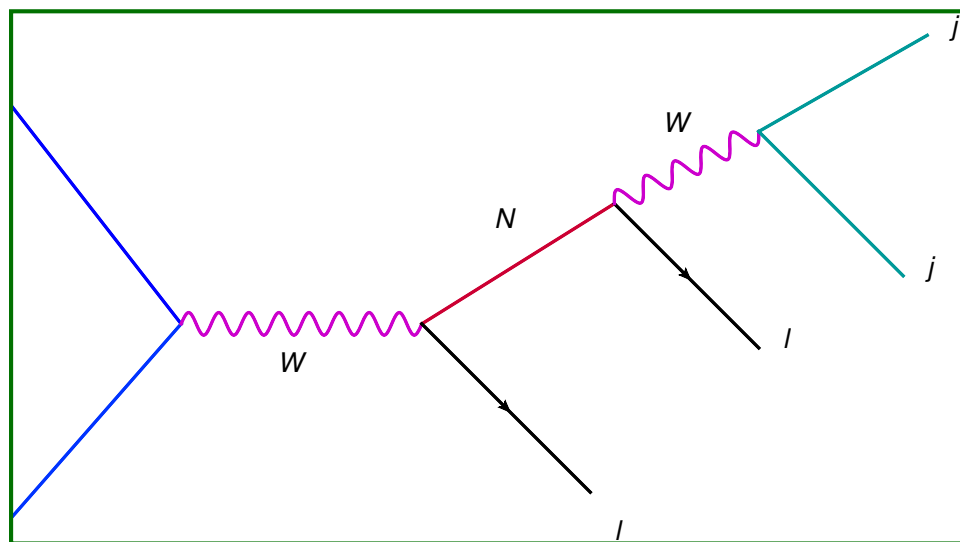
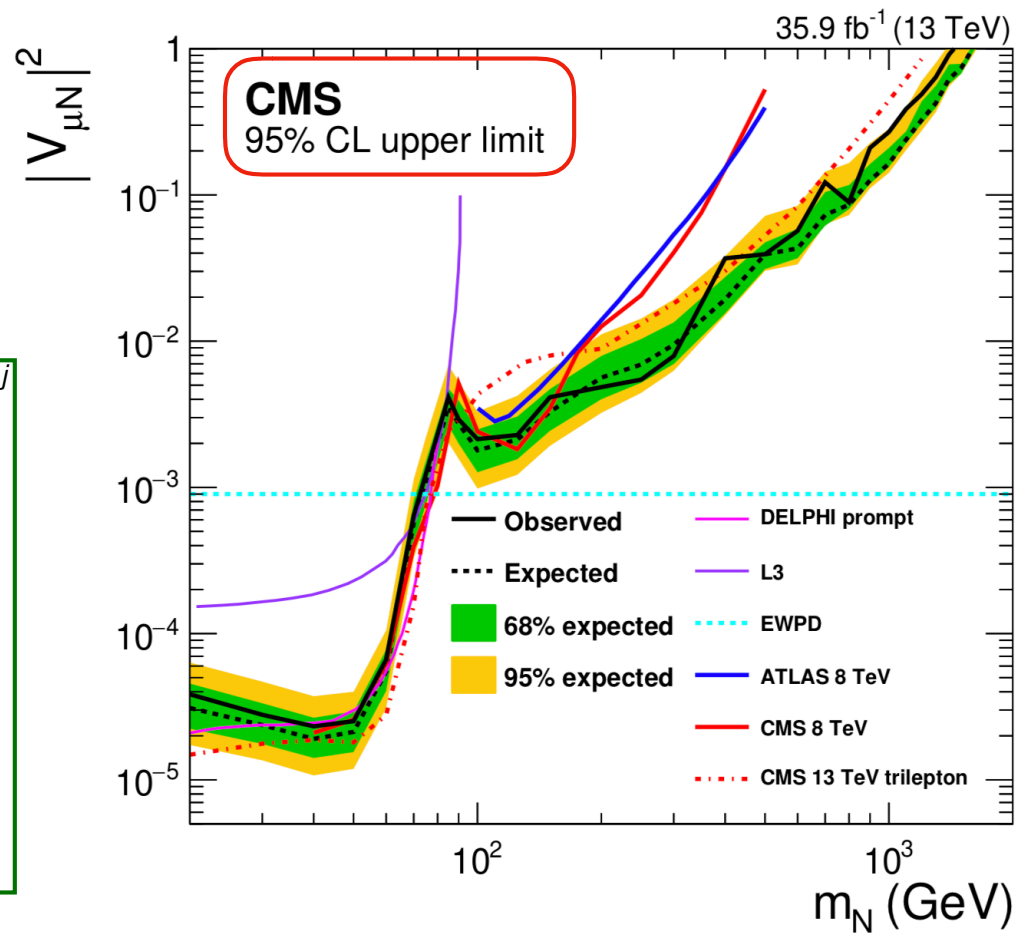
Current experimental limits



$\ell^\pm \ell^\pm + \text{jets}$

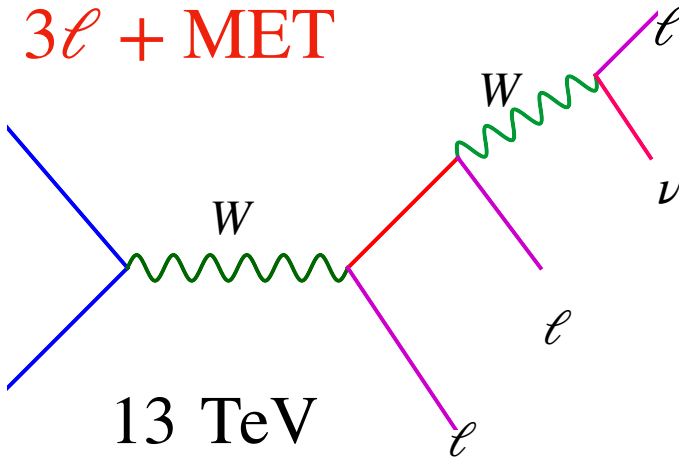
CMS Prompt

1806.10905 13 TeV, 35.9 fb⁻¹



Experimental limits

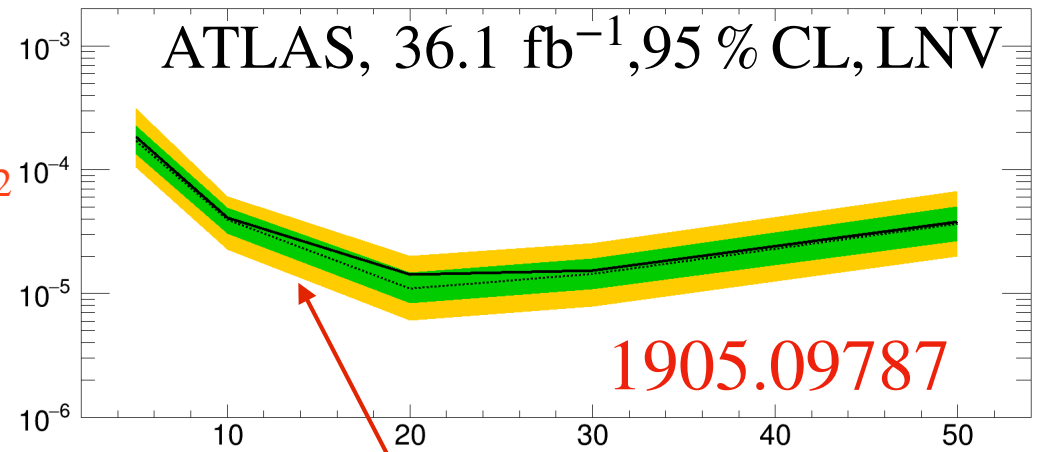
$3\ell + \text{MET}$



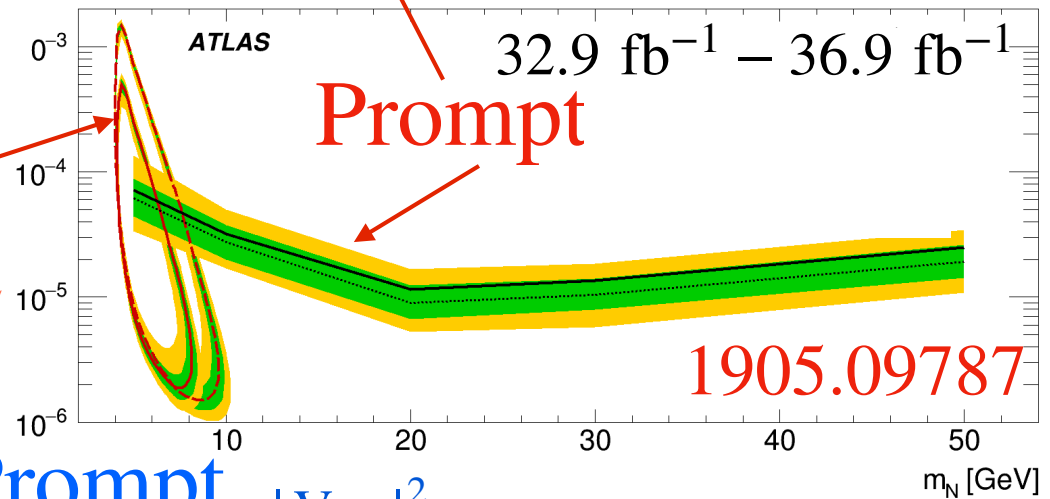
Displaced

— LNV
 - - - LNC

$|V_{eN}|^2$



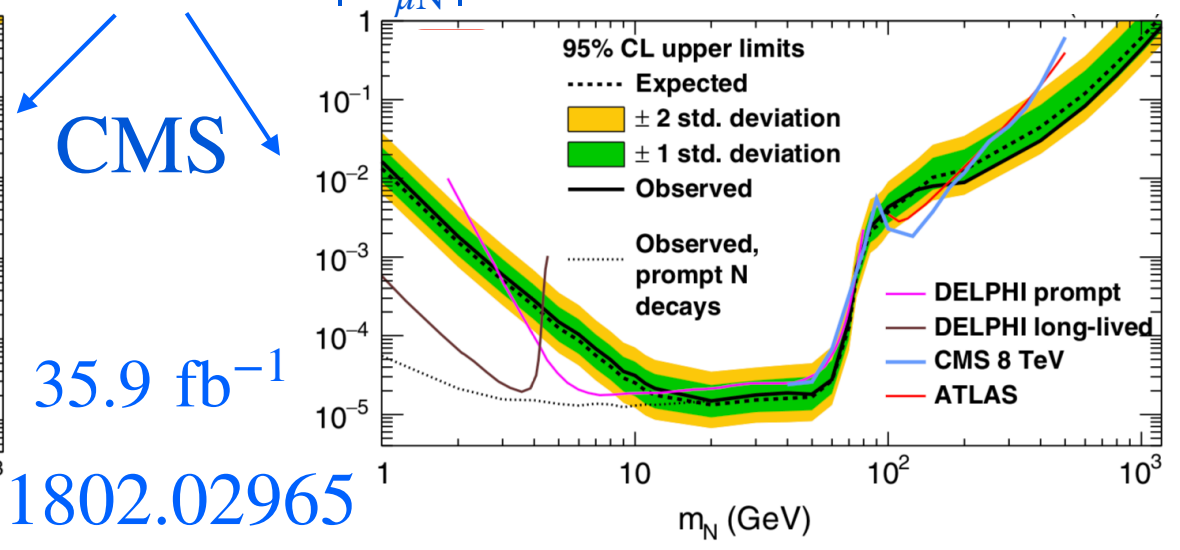
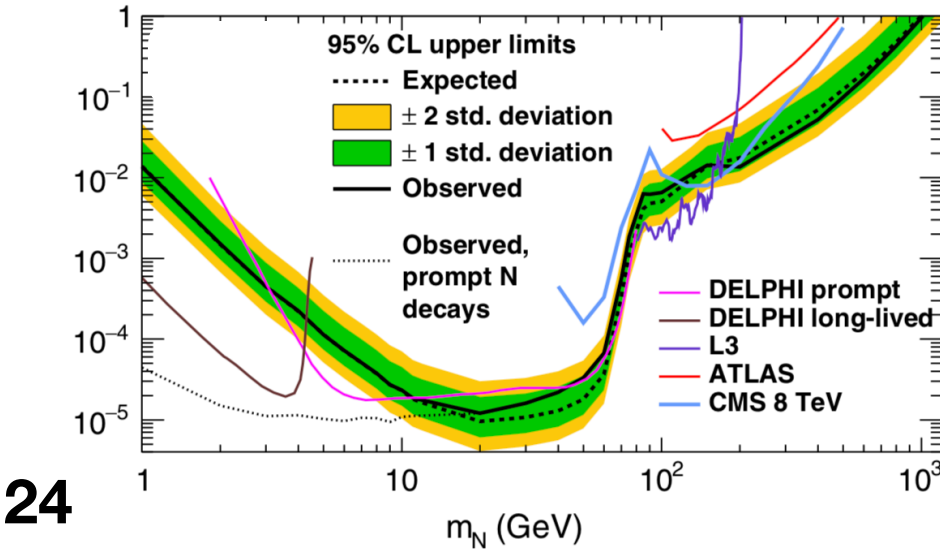
$|V_{\mu N}|^2$



$|V_{eN}|^2$

Prompt

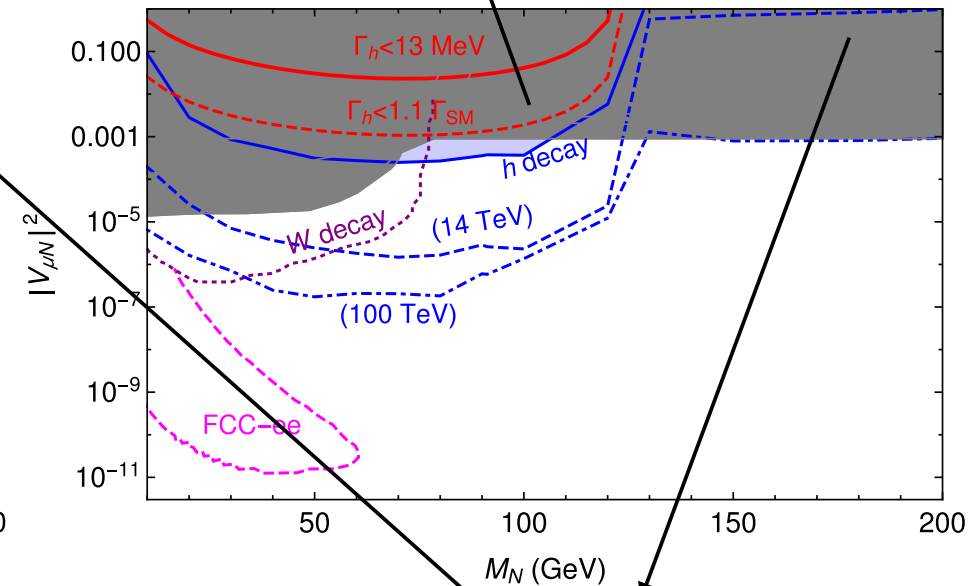
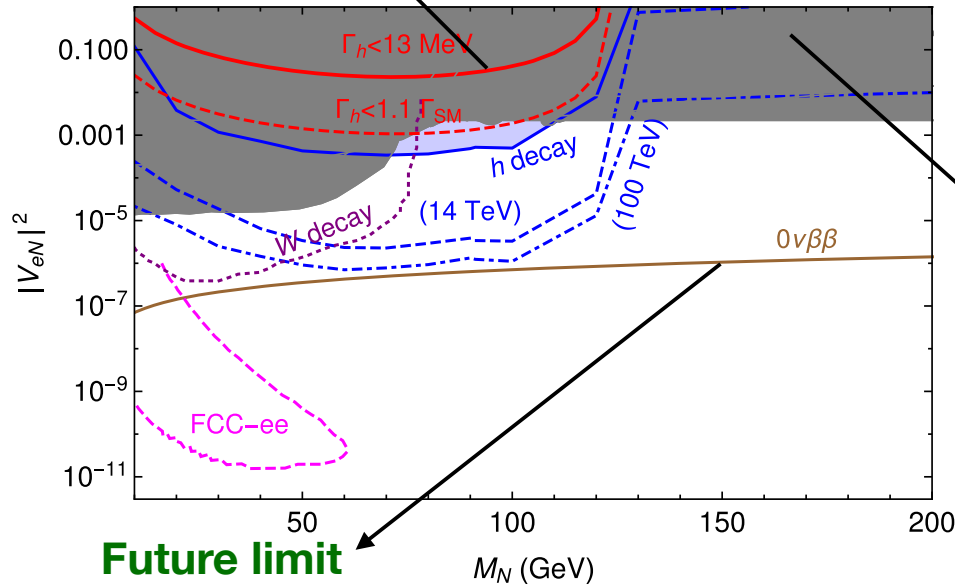
$|V_{\mu N}|^2$



$2\ell + p_T^{\text{miss}}$: bounds from the Higgs decay ($h \rightarrow N\nu, N \rightarrow 2\ell\nu$)

CMS, JHEP 09 (2016) 051: 7&8 TeV combined
 $H \rightarrow W W^*$, upper limit on Yukawa as well as mixing

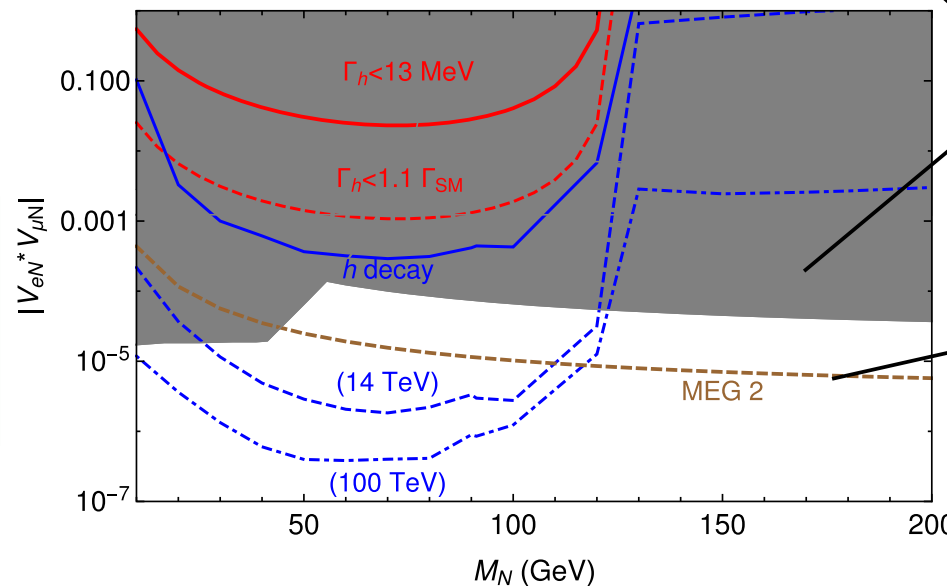
Future sensitivity can go down to
 10% precise result at pp collider:
 arXiv:1606.09408



Future limit considering Majorana heavy neutrinos only

FCC-ee : Limits from Z decay
 W-decay @LHC

Future limits



Excluded by LEP LHC,EWPD, LFV limits from CMS is also included in the lower panel

$\mu \rightarrow e\gamma$
 ~ future branching ratio $O(10^{-15})$

Type – III seesaw

SM + SU(2)_L triplet fermion

Franceschini, Hambye, Strumia Biggio, Bonnet Goswami, Poulou
 Jana, Okada, Raut; Biggio, Fernandez Martinez, Hernandez Garcia, Lopez Pavon
 Goswami, Vishnudath, Khan AD, Mandal, Modak; AD, Mandal;
 Bandyopadhyay, Jangid, Mitra

$$\mathcal{L} = \mathcal{L}_{\text{SM}} + \text{Tr}(\bar{\Psi} i \gamma^\mu D_\mu \Psi) - \frac{1}{2} M \text{Tr}(\bar{\Psi} \Psi^c + \bar{\Psi}^c \Psi) - \sqrt{2}(\bar{\ell}_L Y_D^\dagger \Psi H + H^\dagger \bar{\Psi} Y_D \ell_L)$$

$$\Psi = \begin{pmatrix} \Sigma^0/\sqrt{2} & \Sigma^+ \\ \Sigma^- & -\Sigma^0/\sqrt{2} \end{pmatrix} \text{ and } \Psi^c = \begin{pmatrix} \Sigma^{0c}/\sqrt{2} & \Sigma^{-c} \\ \Sigma^{+c} & -\Sigma^{0c}/\sqrt{2} \end{pmatrix}$$

$$-\mathcal{L}_{\text{mass}} = (\bar{e}_L \ \bar{\Sigma}_L) \begin{pmatrix} m_\ell & Y_D^\dagger v \\ 0 & M \end{pmatrix} \begin{pmatrix} e_R \\ \Sigma_R \end{pmatrix} + \frac{1}{2} (\bar{\nu}_L^c \ \bar{\Sigma}_R^0) \begin{pmatrix} 0 & Y_D^T \frac{v}{\sqrt{2}} \\ Y_D \frac{v}{\sqrt{2}} & M \end{pmatrix} \begin{pmatrix} \nu_L \\ \Sigma_R^{0c} \end{pmatrix} + h.c.$$

$$m_\nu \simeq -\frac{v^2}{2} Y_D^T M^{-1} Y_D = M_D M^{-1} M_D^T$$

$$\Gamma(\Sigma^\pm \rightarrow \nu W) = \frac{g^2 |V_{\ell\Sigma}|^2}{32\pi} \left(\frac{M^3}{M_W^2}\right) \left(1 - \frac{M_W^2}{M^2}\right)^2 \left(1 + 2\frac{M_W^2}{M^2}\right)$$

$$\Gamma(\Sigma^\pm \rightarrow \ell Z) = \frac{g^2 |V_{\ell\Sigma}|^2}{64\pi \cos^2 \theta_W} \left(\frac{M^3}{M_Z^2}\right) \left(1 - \frac{M_Z^2}{M^2}\right)^2 \left(1 + 2\frac{M_Z^2}{M^2}\right)$$

$$\Gamma(\Sigma^\pm \rightarrow \ell h) = \frac{g^2 |V_{\ell\Sigma}|^2}{64\pi} \left(\frac{M^3}{M_h^2}\right) \left(1 - \frac{M_h^2}{M^2}\right)^2,$$

$$\Gamma(\Sigma^\pm \rightarrow \Sigma^0 \pi^\pm) = \frac{2G_F^2 V_{ud}^2 \Delta M^3 f_\pi^2}{\pi} \sqrt{1 - \frac{m_\pi^2}{\Delta M^2}}$$

$$\Gamma(\Sigma^\pm \rightarrow \Sigma^0 e \nu_e) = \frac{2G_F^2 \Delta M^5}{15\pi}$$

$$\Gamma(\Sigma^\pm \rightarrow \Sigma^0 \mu \nu_\mu) = 0.12 \Gamma(\Sigma^\pm \rightarrow \Sigma^0 e \nu_e)$$

$$\Gamma(\Sigma^0 \rightarrow \ell^+ W) = \Gamma(\Sigma^0 \rightarrow \ell^- W) = \frac{g^2 |V_{\ell\Sigma}|^2}{64\pi} \left(\frac{M^3}{M_W^2}\right) \left(1 - \frac{M_W^2}{M^2}\right)^2 \left(1 + 2\frac{M_W^2}{M^2}\right)$$

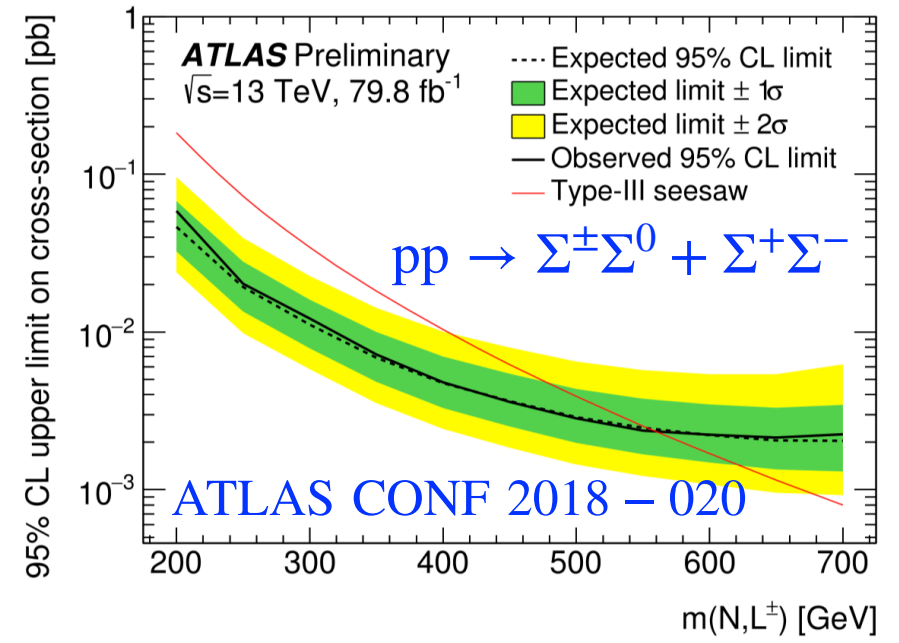
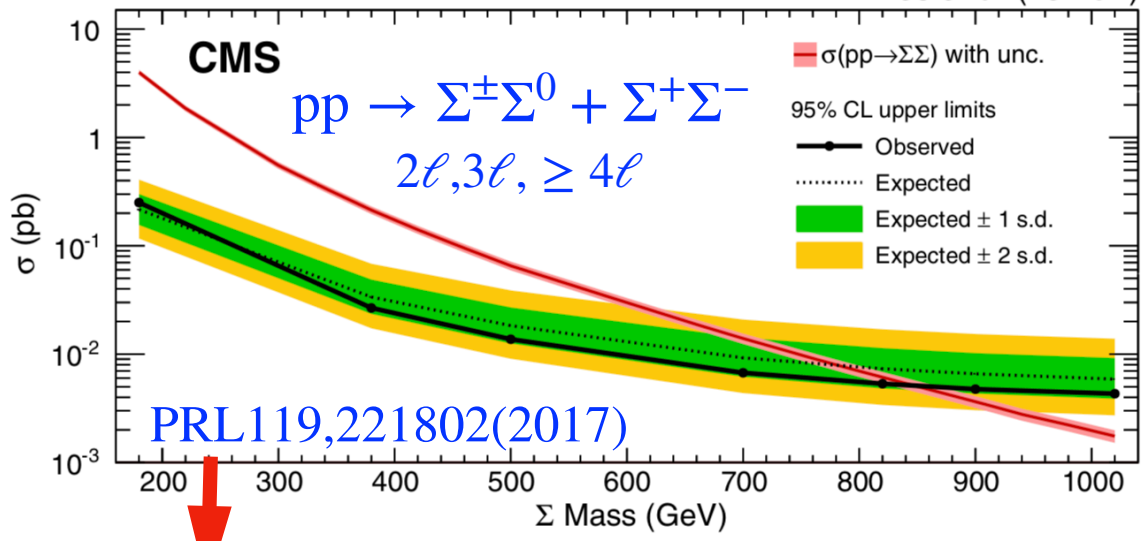
$$\Gamma(\Sigma^0 \rightarrow \nu Z) = \Gamma(\Sigma^0 \rightarrow \bar{\nu} Z) = \frac{g^2 |V_{\ell\Sigma}|^2}{128\pi \cos^2 \theta_W} \left(\frac{M^3}{M_Z^2}\right) \left(1 - \frac{M_Z^2}{M^2}\right)^2 \left(1 + 2\frac{M_Z^2}{M^2}\right)$$

$$\Gamma(\Sigma^0 \rightarrow \nu h) = \Gamma(\Sigma^0 \rightarrow \bar{\nu} h) = \frac{g^2 |V_{\ell\Sigma}|^2}{128\pi} \left(\frac{M^3}{M_h^2}\right) \left(1 - \frac{M_h^2}{M^2}\right)^2,$$

Experimental limits

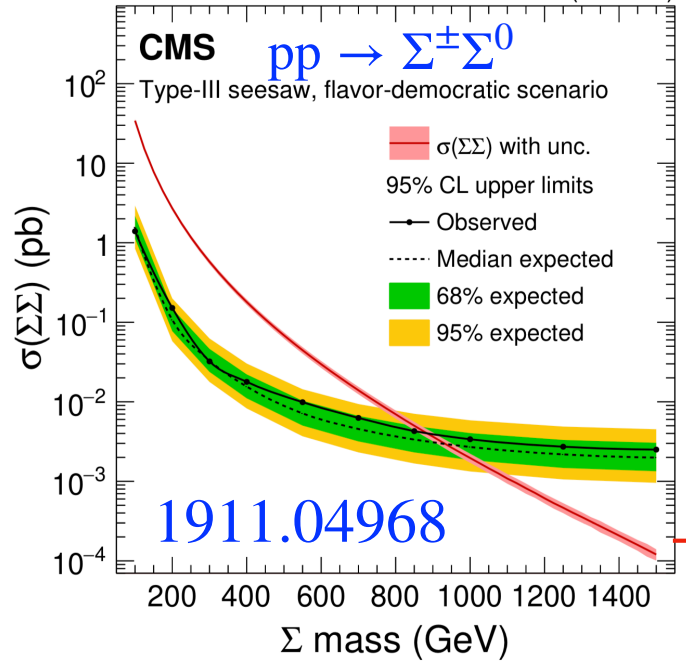
$$BR = B_\ell \propto \frac{|V_\ell|^2}{|V_e|^2 + |V_\mu|^2 + |V_\tau|^2} \quad B_e = B_\mu = B_\tau$$

Flavor – democratic scenario

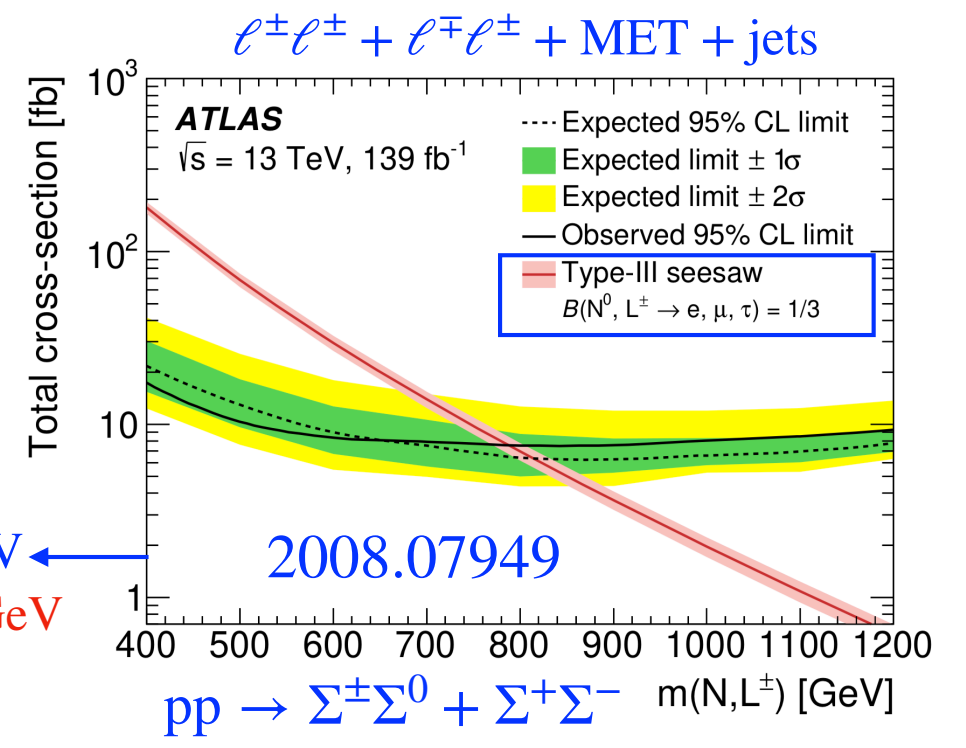


τ – phoic, $B_\tau = 0, M_\Sigma = 900 \text{ GeV}, 90\% \text{ CL}$
 (e, μ) – phoic, $B_{e+\mu} = 0, M_\Sigma = 390 \text{ GeV}, 90\% \text{ CL}$

$l^\pm l^\mp l^\pm l^\mp + \text{MET}$ 137 fb⁻¹ (13 TeV)

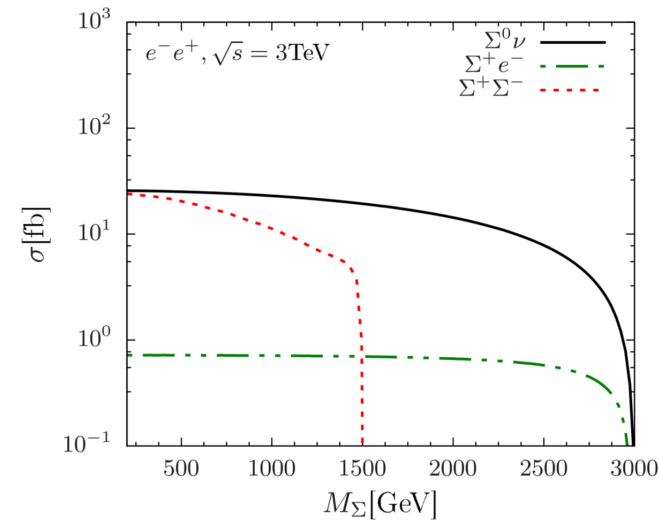
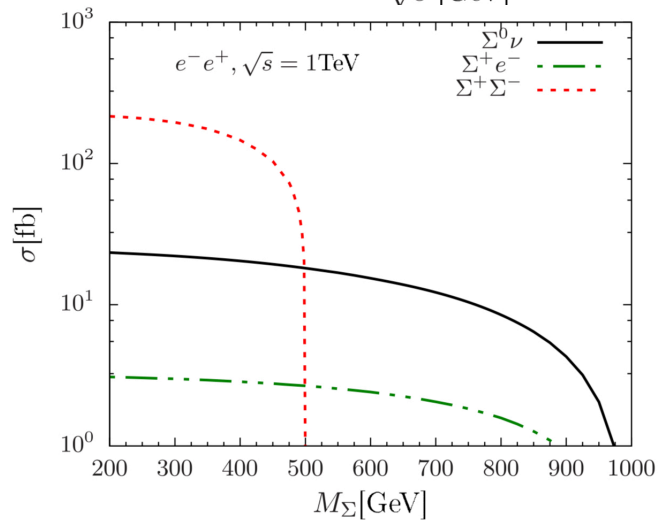
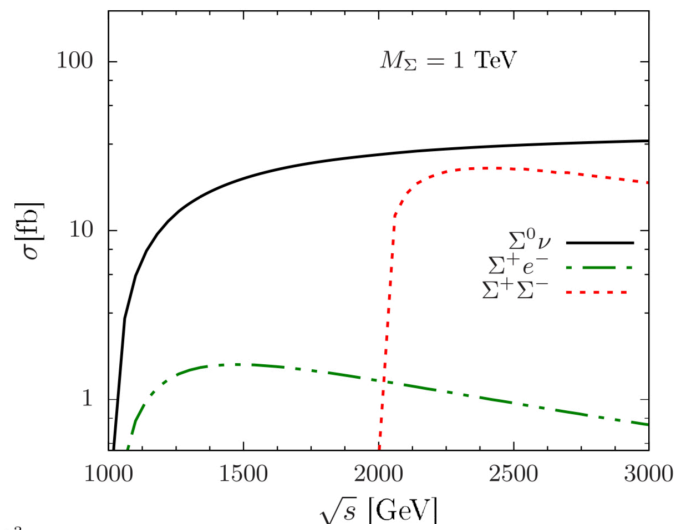
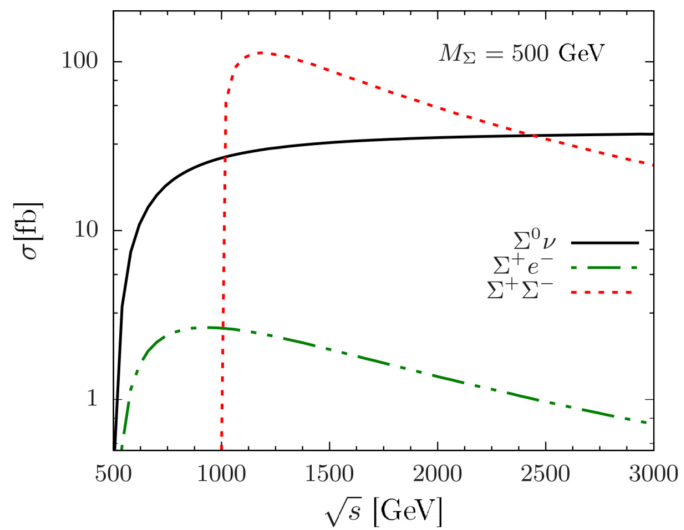
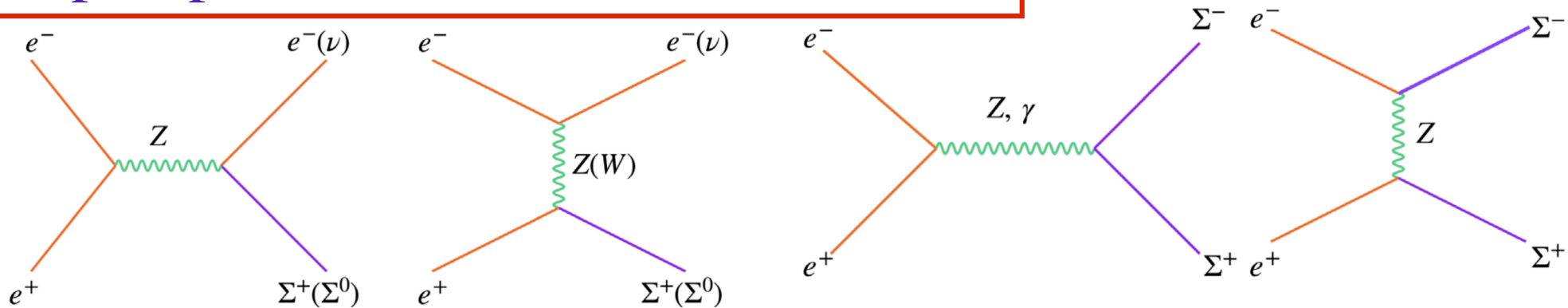


$M_\Sigma \leq 800 \text{ GeV}$
 $M_\Sigma \leq 900 \text{ GeV}$



Triplet production at the e^-e^+ collider

AD, Mandal, Modak;



Mass-mixing limit plots

AD, Mandal, Modak;

