

ASPERA WP 2: Neutrino Mass

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Physics motivations

Experiments and projects

Project summaries & ressources

Physics case

- Absolute neutrino mass and neutrino mass hierarchy (**SDB, DBD**)
- Nature of neutrino : Dirac ($\nu \neq \bar{\nu}$) or Majorana ($\nu = \bar{\nu}$) (**DBD**)
- Leptonic number violation (**DBD**)
- Right-handed current interaction (**DBD**)
- CP violation in leptonic sector (**DBD**)
- Search of Supersymmetry and new particles (**DBD**)

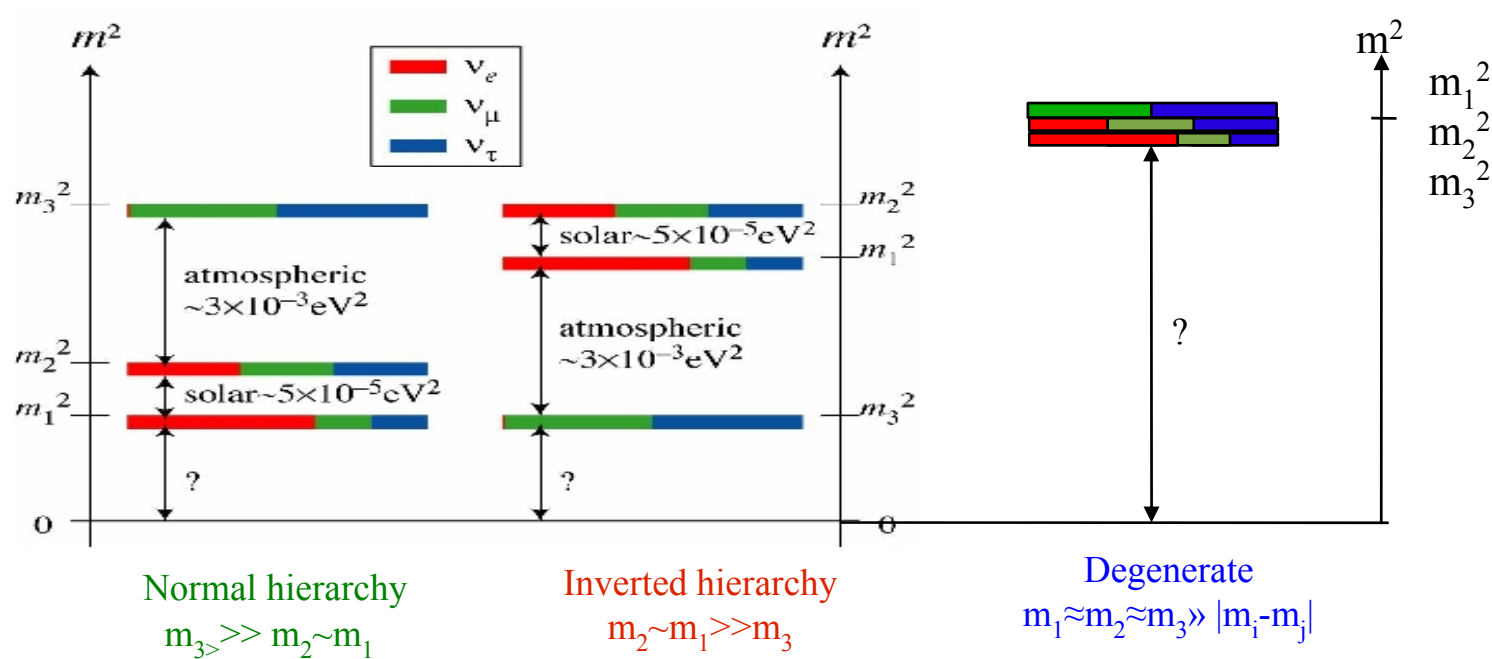
SDB: Single Beta Decay

DBD: Double Beta Decay

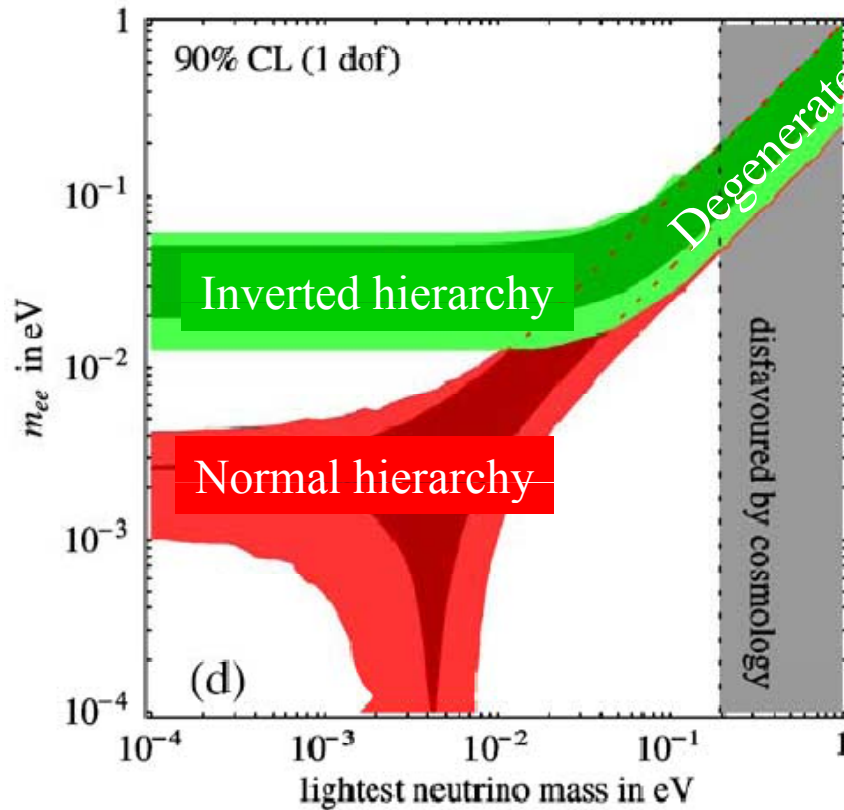
Absolute Neutrino mass

Process	Measured mass	Present limit	Futur
Beta decay	$m_\nu = \left[\sum U_{ei} ^2 m_i^2 \right]^{1/2}$	<2.3 eV	< 0.2 eV
Double beta decay	$ \langle m_\nu \rangle = \sum U_{ei}^2 m_i $	< 0.2 - 0.8 eV	< ~0.05 eV
Cosmology	$\sum m_i = m_1 + m_2 + m_3$	< ~1 eV	< ~0.1 eV

Mass Hierarchy



Mass Hierarchy



Degenerate: can be tested by SDB and DBD

Inverted hierarchy: tested by the next generation of $\beta\beta$ experiment

Normal hierarchy: seems inaccessible within 10 years

Nature of neutrino

Dirac neutrino particle \neq antiparticle 4 states $\nu_L, \nu_R, \bar{\nu}_L, \bar{\nu}_R$
Conservation of global leptonic number

Majorana neutrino $\nu = \bar{\nu}$ 2 states ν_L and ν_R
No conservation of global leptonic number

If ν_{Majorana} and CP violation in leptonic sector \rightarrow leptogenesis

WP2 roadmap

Currently Europe has leadership for both single and double beta decay experiments

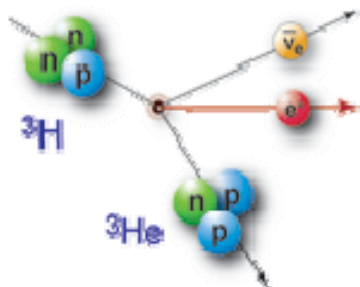
To improve the sensitivity on neutrino mass by at least a factor 10 within 10 years

To select the most sensitive techniques for the next generation of experiments

To support R&D to develop new techniques

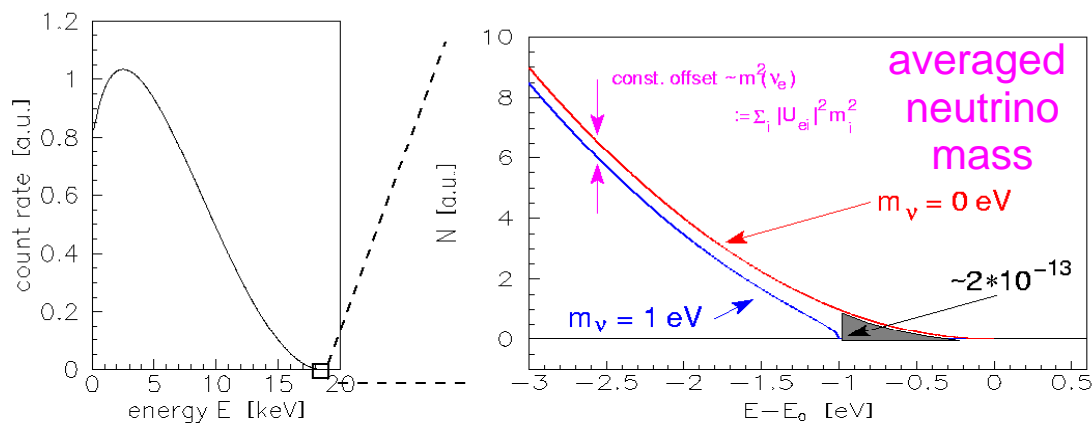
Single beta decay

$$(A,Z) \rightarrow (A,Z+1) + e^- + \bar{\nu}_e$$



$$dN/dE \sim [(E_0 - E_e)^2 - m_{\nu_i}^2]^{1/2}$$

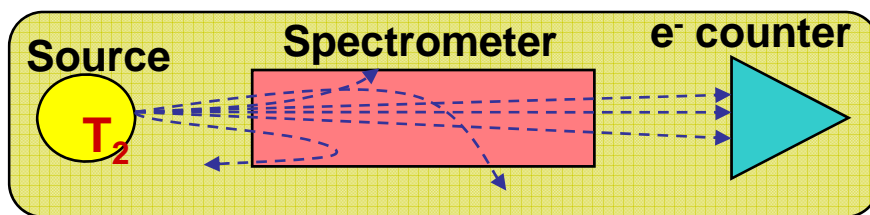
$$m_{\nu_e}^2 = \sum |U_{ei}|^2 m_i^2$$



Extraction of m_{ν} is not model dependant

Two possible approaches

KATRIN Spectrometer



Source (^3He) \neq detector

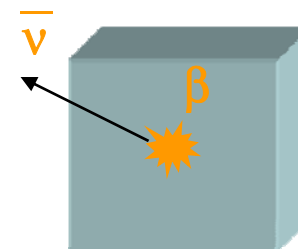
High activity source $\sim 2\text{Ci/s}$

Energy resolution: 0.92 eV

Integrated beta energy spectrum

Size limit

MARE Calorimeter



Bolometer

β -source (^{187}Re) = detector

Low activity source $\sim 1\text{ Bq/mg}$

Energy resolution: 5 eV

Measure differential spectrum DN/DE

Can be extended

Different systematics  complementary techniques

KATRIN

Neutrino mass

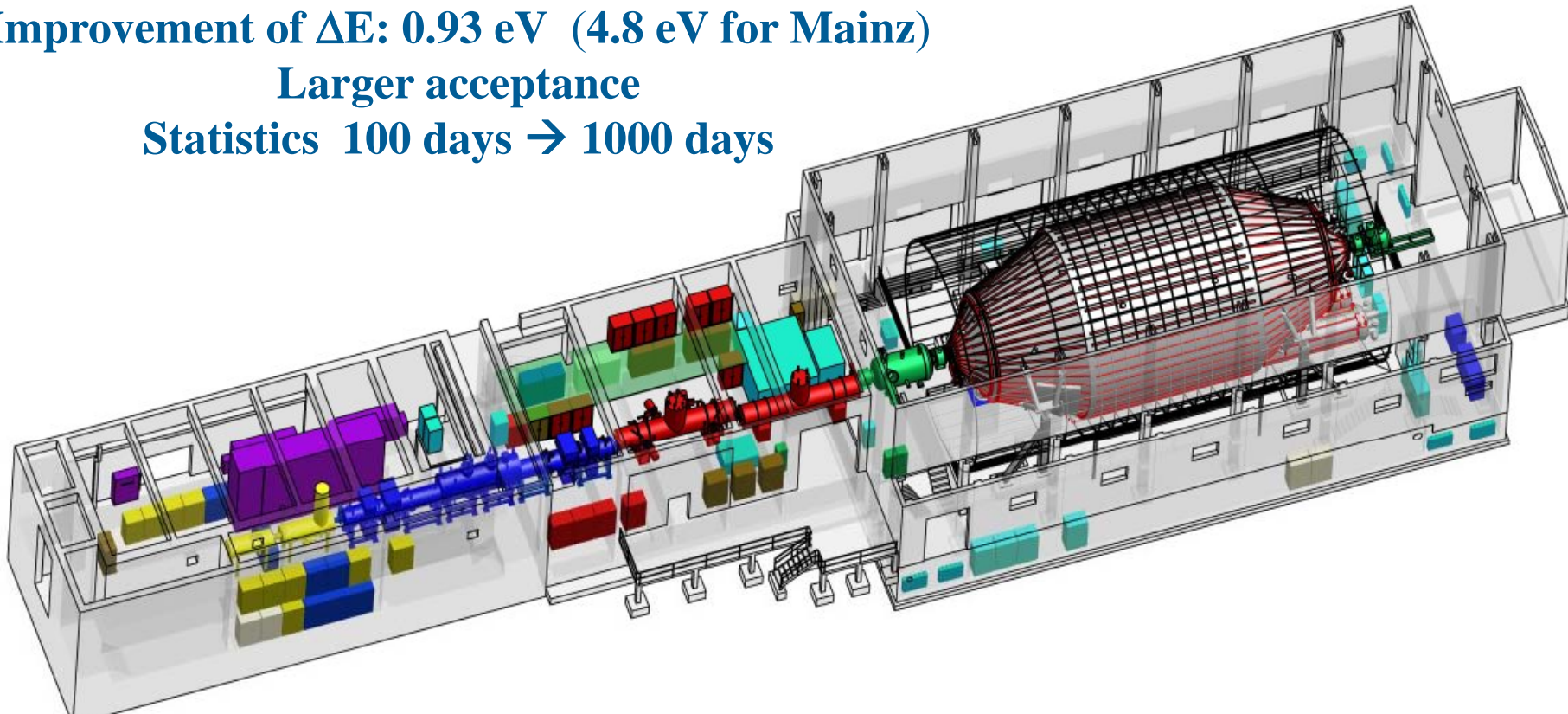
Czech Republic – Germany – Russian Federation – United Kingdom – United States

2007: 120 Collaborators from 16 institutions including 30 PhD and diploma students

Improvement of ΔE : 0.93 eV (4.8 eV for Mainz)

Larger acceptance

Statistics 100 days \rightarrow 1000 days



KATRIN

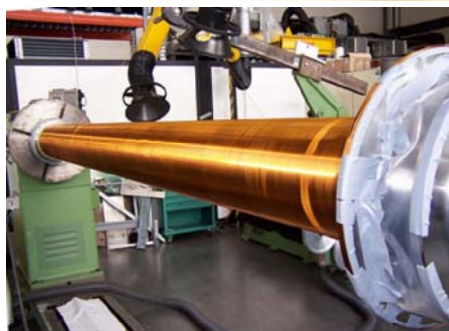
Neutrino mass

Tritium source



one of the most complex cryostats ever built:

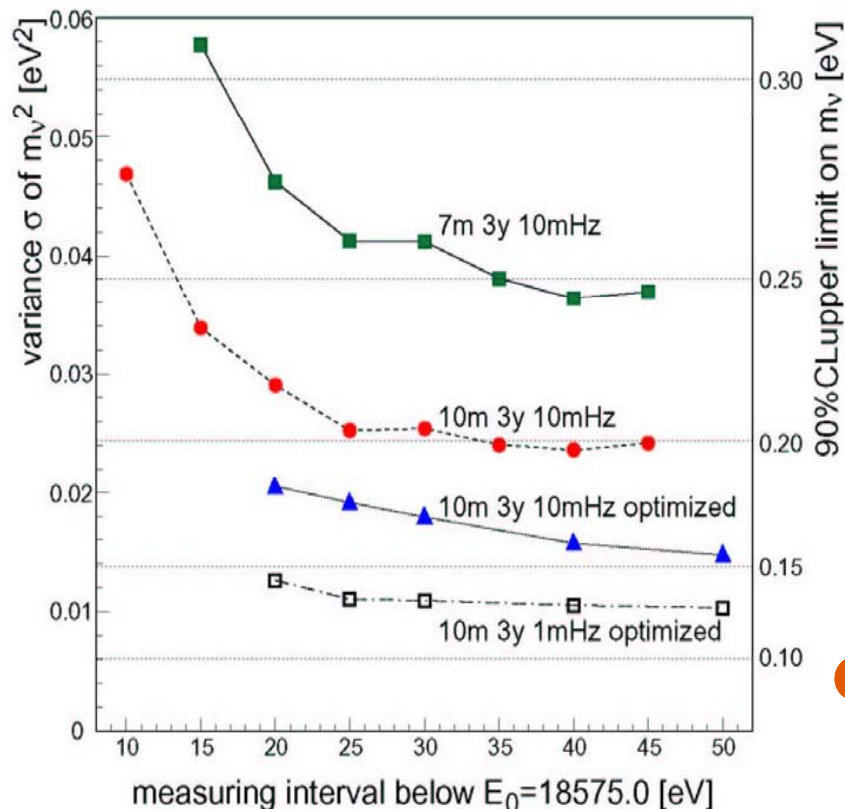
- 42 sub-assemblies
- 12 cryo circuits
- 6 working fluids
- > 500 sensors
- 7 magnet modules



Main spectrometer



KATRIN



Sensitivity (90% CL)
 $m(\nu) < 0.2 \text{ eV}$

Discovery potential
 $m(\nu) = 0.35 \text{ eV} (5\sigma)$

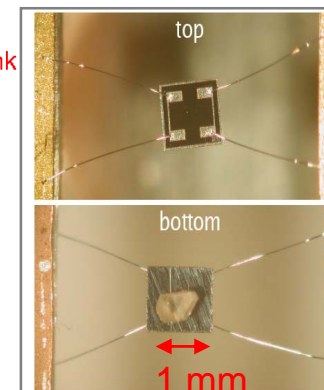
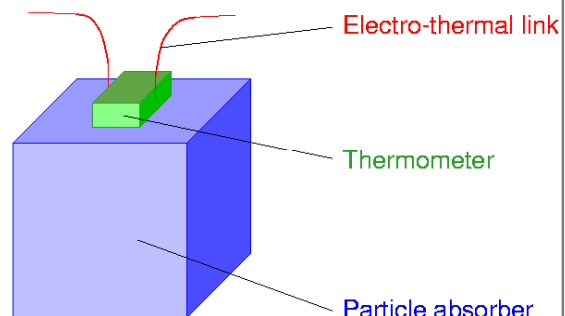
Commissioning and operation: 2010
Final results 2015-2016

MARE

Italy, USA, Germany (10 physicist in Europe)

MicroBolometers of ArReO4

$$^{187}\text{Re } Q_{\beta} = 2.47 \text{ keV}$$



MARE-I: 300 detectors

Energy resolution ~20 eV

$$\langle m_{\nu} \rangle < 2 - 4 \text{ eV (2010)}$$

MARE – II : 50000 detectors

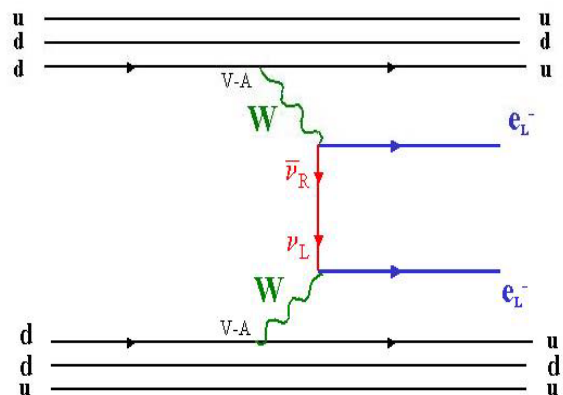
Energy resolution ~5 eV

Need new techniques

$$\langle m_{\nu} \rangle < 0.2 \text{ eV (2015)}$$

Neutrinoless double beta decay

Neutrino mass



Light neutrino exchange



$$\Delta L = 2$$

Lepton number violation

Majorana neutrino ($\nu=\bar{\nu}$)

Massive neutrino

Phase space factor

Nuclear matrix element

$$T_{1/2}^{-1} = F(Q_{\beta\beta}, Z) |M^{0\nu}|^2 \langle m_{\nu} \rangle^2$$

$$\langle m_{\nu} \rangle = m_1 |U_{e1}|^2 + m_2 |U_{e2}|^2 \cdot e^{i\alpha} + m_3 |U_{e3}|^2 \cdot e^{i\beta}$$

$|U_{ei}|$: mixing matrix elements

α et β : Majorana phases

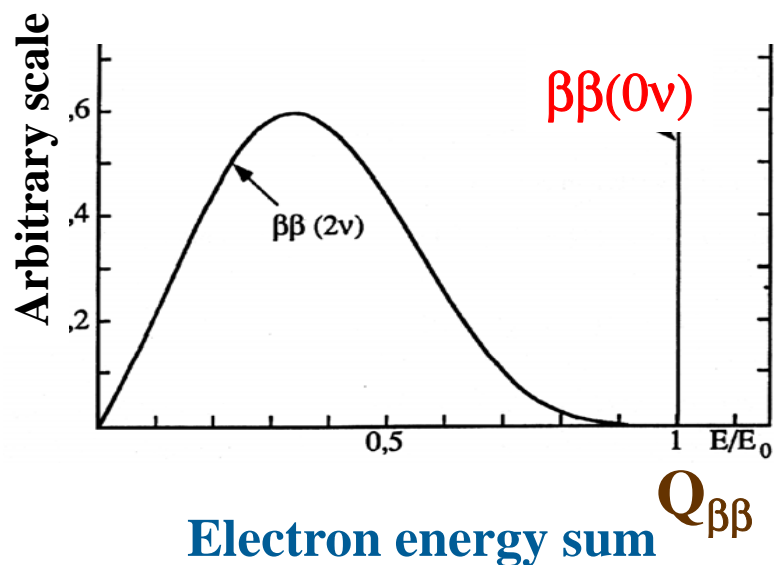
Other possible process :

V+A current : $\langle m_{\nu} \rangle, \langle \lambda \rangle, \langle \eta \rangle$

Majoron emission : $\langle g_M \rangle$

Supersymmetry : $\lambda'_{111}, \lambda'_{113}$

$\beta\beta(0\nu)$ observables



Angular distribution

Individual electron energy

Allow to distinguish
the mechanism

Background : natural radioactivity, radon, neutrons, muons, $\beta\beta(2\nu)$

Double beta decay

Why several nuclei and why several techniques are required ?

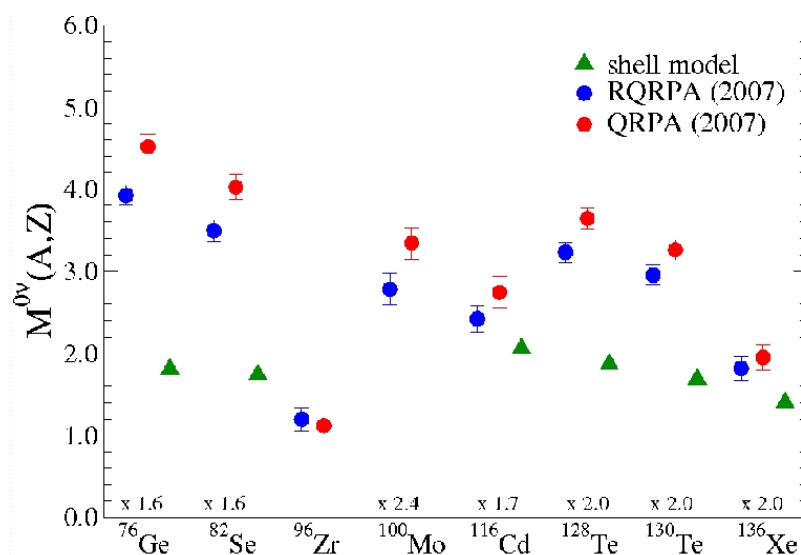
Neutrino mass

Isotope	$Q_{\beta\beta}$ (MeV)	Isotopic abundance (%)	$G_{0\nu}(\text{yr}^{-1}) \times 10^{25}$
^{48}Ca	4.271	0.187	2.44
^{76}Ge	2.040	7.8	0.24
^{82}Se	2.995	9.2	1.08
^{96}Zr	3.350	2.8	2.24
^{100}Mo	3.034	9.6	1.75
^{116}Cd	2.802	7.5	1.89
^{130}Te	2.528	33.8	1.70
^{136}Xe	2.479	8.9	1.81
^{150}Nd	3.367	5.6	8.00

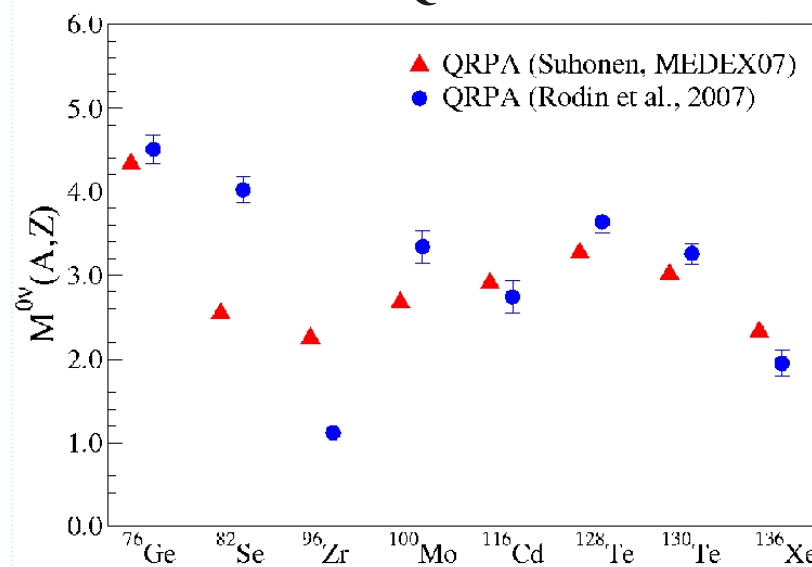
$$T_{1/2}^{-1} = F(Q_{\beta\beta}^5, Z) |M^{0\nu}|^2 \langle m_\nu \rangle^2$$

$\langle m_\nu \rangle$ is model dependant

Shell Model (Poves et al) - QRPA



Two different QRPA calculations



A lot of improvements have been done but still discrepancies
 Uncertainties for extraction of $\langle m_\nu \rangle$

With background:

$$T_{1/2}^{0\nu}(\text{y}) > \frac{\ln 2 \cdot \mathcal{N}}{k_{\text{C.L.}}} \cdot \frac{\epsilon}{A} \cdot \sqrt{\frac{M \cdot t}{N_{\text{Bckg}} \cdot \Delta E}}$$

M: masse (g)
 ϵ : efficiency
 $k_{\text{C.L.}}$: Confidence level
 \mathcal{N} : Avogadro number
 t: time (y)
 N_{Bckg} : Background events ($\text{keV}^{-1} \cdot \text{g}^{-1} \cdot \text{y}^{-1}$)
 ΔE : energy resolution (keV)

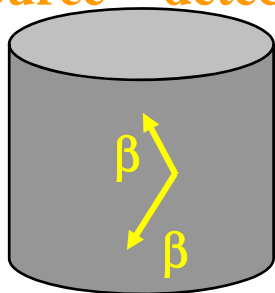
Today, no technique able to optimize all the parameters

Calorimeter

Semi-conductors

Bolometers

Source = detector

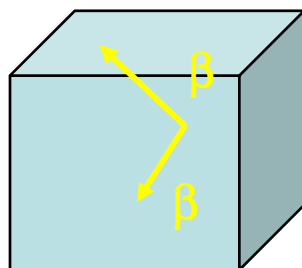


$\epsilon, \Delta E$

Calorimeter

(Loaded) Scintillator

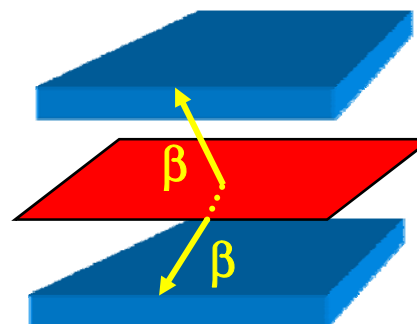
Source = detector



ϵ, M

Tracko-calorimeter

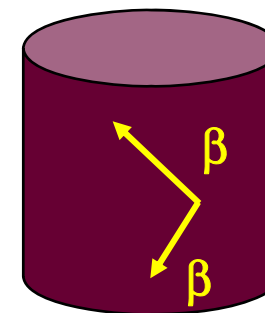
Source \neq detector



N_{Bckg} , isotope choice

Xe TPC

Source = detector



$\epsilon, M, (N_{\text{Bckg}})$

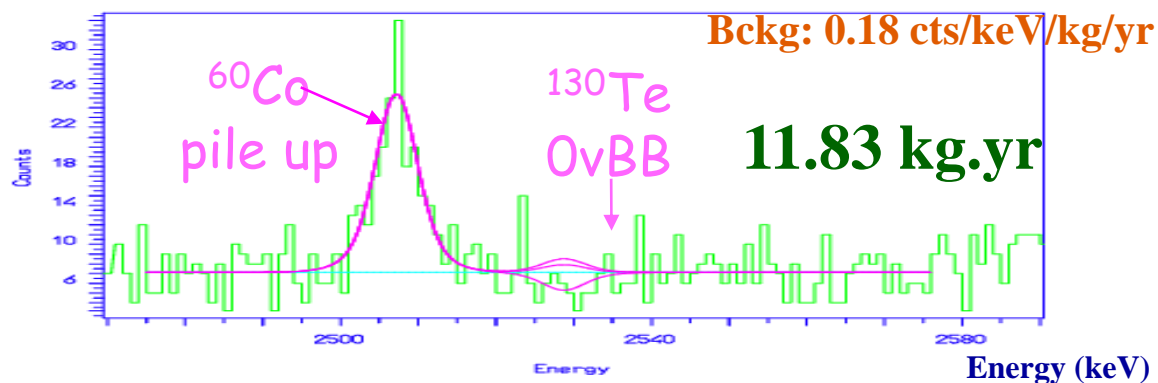
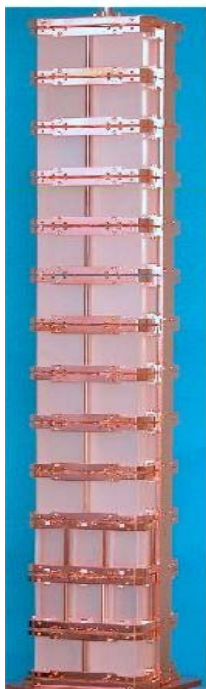
Neutrino mass

Name	Nucleus	Method	Location	European Members	Others
<i>Running experiments</i>					
CUORICINO	^{130}Te	bolometric	LNGS	IT, NL, ES	US
NEMO-3	^{82}Se	tracking	Frejus	FR, CZ, UK	US, RU, JP
<i>Construction funding</i>					
CUORE	^{130}Te	bolometric	LNGS	IT, NL, ES	US
GERDA	^{76}Ge	ionization	LNGS	DE, BE, IT, PO	RU
<i>Substantial R&D funding</i>					
COBRA	^{116}Cd , ^{130}Te	scintillation	LNGS	UK, DE	-
EXO	^{136}Xe	tracking	WIPP	CH	US, RU, CAN
MAJORANA	^{76}Ge	ionization	SNOLAB or DUSEL	-	US
SuperNEMO	^{82}Se	tracking	t.b.d.	FR, CZ, UK, SL	US, RU, JP
<i>R&D and/or conceptual design, and other decay modes</i>					
CANDLES	^{48}Ca	scintillation	Oto Lab	-	JP
CARVEL	^{48}Ca	scintillation	Solotvina		UKR, RU, US
DCBA	^{150}Nd	tracking	t.b.d.	-	JP
MOON	^{100}Mo	tracking	t.b.d.	-	JP
SNO++	^{150}Nd	scintillation	SNOLAB	-	CAN, US + ...
TGV	^{106}Cd , ^{48}Ca	electron capture, R&D	Frejus	FR, CZ	RU

Cuoricino

Italy, USA, Spain (12 scientists in Europe)

Bolometer of TeO_2 (~10 kg of ^{130}Te)
Running at gran Sasso since 2003



$T_{1/2} > 3. \cdot 10^{24}$ yr (90% CL)

$\langle m_\nu \rangle < 0.2 - 1$ eV (90% CL)

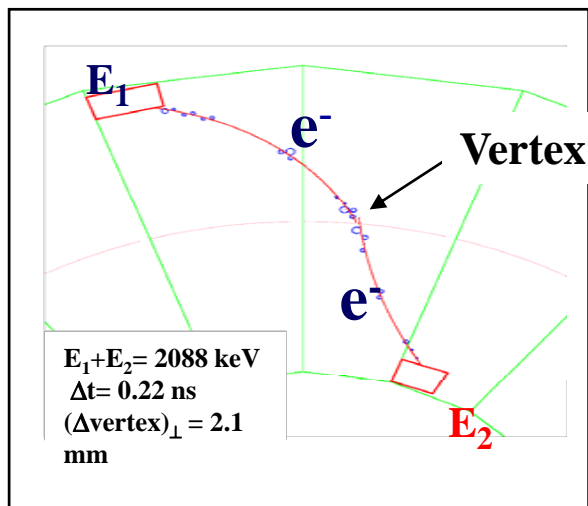
Expected final sensitivity ~2009: $T_{1/2} > 6. \cdot 10^{24}$ yr $\langle m_\nu \rangle < 0.1 - 0.7$ eV

NEMO 3

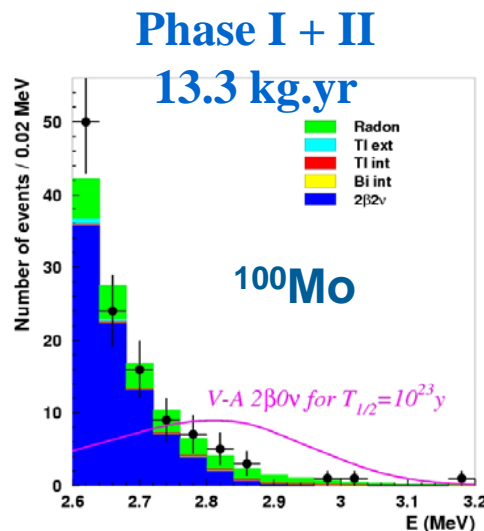
Neutrino mass

France, UK, Russia, Spain, USA, Czech Republic, Ukraine, Finland (30 scientists in Europe)

Tracko-calorimeter running at Modane since 2003
Multi-isotopes 7 kg of ^{100}Mo and 1 kg ^{82}Se



$\beta\beta$ events



$T_{1/2}(\beta\beta 0\nu) > 5.8 \cdot 10^{23} \text{ yr (90 \% C.L.)}$ $\langle m_{\nu} \rangle < 0.6 - 1.3 \text{ eV}$

Expected in 2009 $T_{1/2}(\beta\beta 0\nu) > 2 \cdot 10^{24} \text{ yr (90 \% CL)}$ $\langle m_{\nu} \rangle < 0.3 - 0.7 \text{ eV}$

Roadmap for double beta decay search

Current running experiment Cuoricino (bolometers ^{130}Te) and NEMO 3 tracko-calorimeters have ~ 10 kg of enriched isotopes sensitivities $\sim 0.2 - 0.7$ eV

Next generation with 100 – 200 kg enriched source, sensitivities $\sim 30 - 100$ meV
validation of background for 1 ton detector

In Europe, 3 complementary techniques: Semi-conductors, bolometers and tracko-calorimeters

Several nuclei can be studied: ^{76}Ge (check of HM claim), ^{130}Te , ^{82}Se (and ^{150}Nd ?)

R&D to improve technique or to develop new techniques

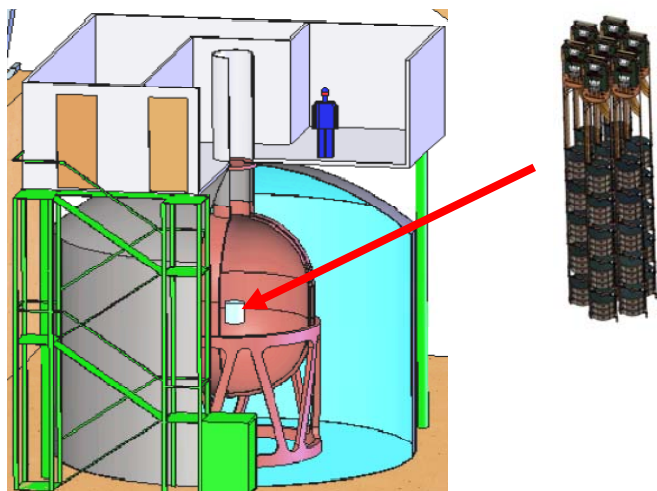
In case of signal: must be checked with other nucleus and techniques. Need of tracko-calorimeters to determine the process

GERDA

Neutrino mass

Germany, Italy, Belgium, Russia (27 scientists in Europe)

Ge diodes in liquid nitrogen
Background suppression by removing matter and active veto



PHASE I: 17.9 kg of enriched ^{76}Ge

In 1 year of data if $B=10^{-2}$ cts/keV/kg/yr

Start 2009 at Gran Sasso, results 2010

$$T_{1/2} > 3 \cdot 10^{25} \text{ yr} \quad \langle m_\nu \rangle < 250 \text{ meV}$$

PHASE II: 40 kg of enriched ^{76}Ge

if $B=10^{-3}$ cts/keV/kg/an

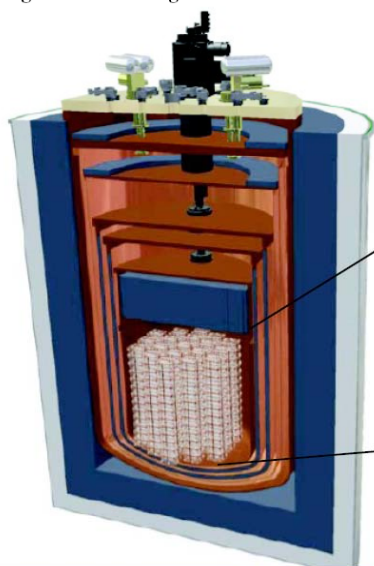
$$T_{1/2} > 2 \cdot 10^{26} \text{ yr in 3 years of data} \quad \langle m_\nu \rangle < 110 \text{ meV}$$

Phase I + II already funded

CUORE

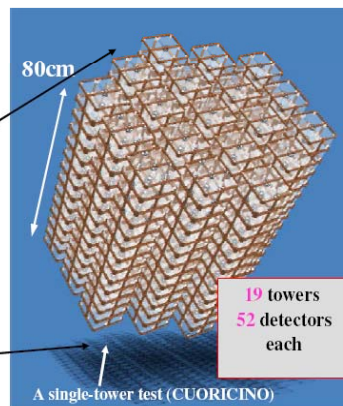
Italy, USA, Spain (12 scientists in Europe)

Single dilution refrigerator -10 mk



for Rare Events

- $\beta\beta\nu$, Cold Dark Matter, Axion searches
proposal hep/ph 0501010



Bolometer of TeO_2 (^{130}Te 203 kg)

Background rejection improvement

Five years of data

$$N_{\text{bckg}} = 0.01 \text{ cts.keV}^{-1}.\text{kg}^{-1}.\text{yr}^{-1}$$

$$T_{1/2} > 2.1 \cdot 10^{26} \text{ yr} \quad \langle m_\nu \rangle < 30 - 170 \text{ meV}$$

Detector is in construction in Gran Sasso

Commissioning: 2010

Operation: 2011

SuperNEMO

France, UK, Russia, Spain, USA, Japan, Czech Republic, Ukraine, Finland (30 scientists in Europe)

20 modules of a tracko-calorimeter

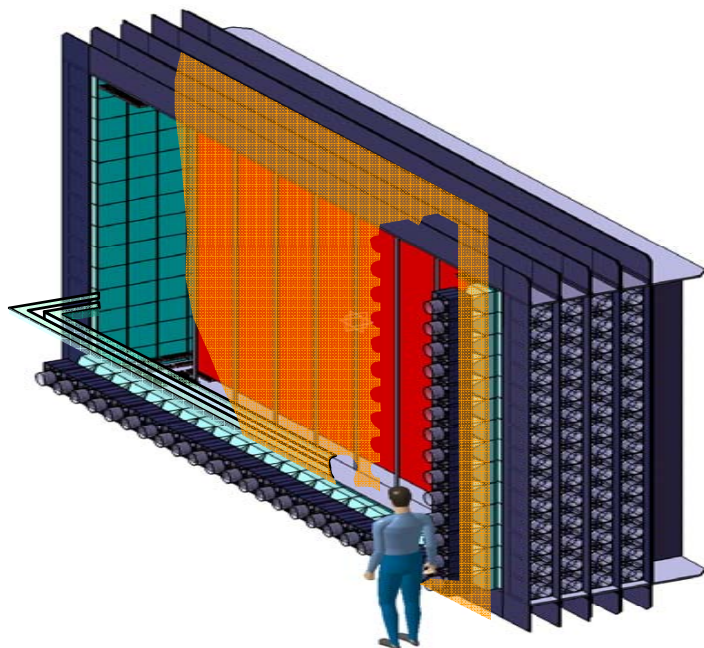
100 kg of ^{82}Se or ^{150}Nd

3 years R&D program: improvement of energy resolution

Increase of efficiency

Background reduction

.....



5 years of data

$T_{1/2} > 2 \cdot 10^{26}$ yr $\langle m_\nu \rangle < 50 - 90$ meV

R&D → 2009

TDR : 2009

First modules in Canfranc in 2011

HM

Cuoricino

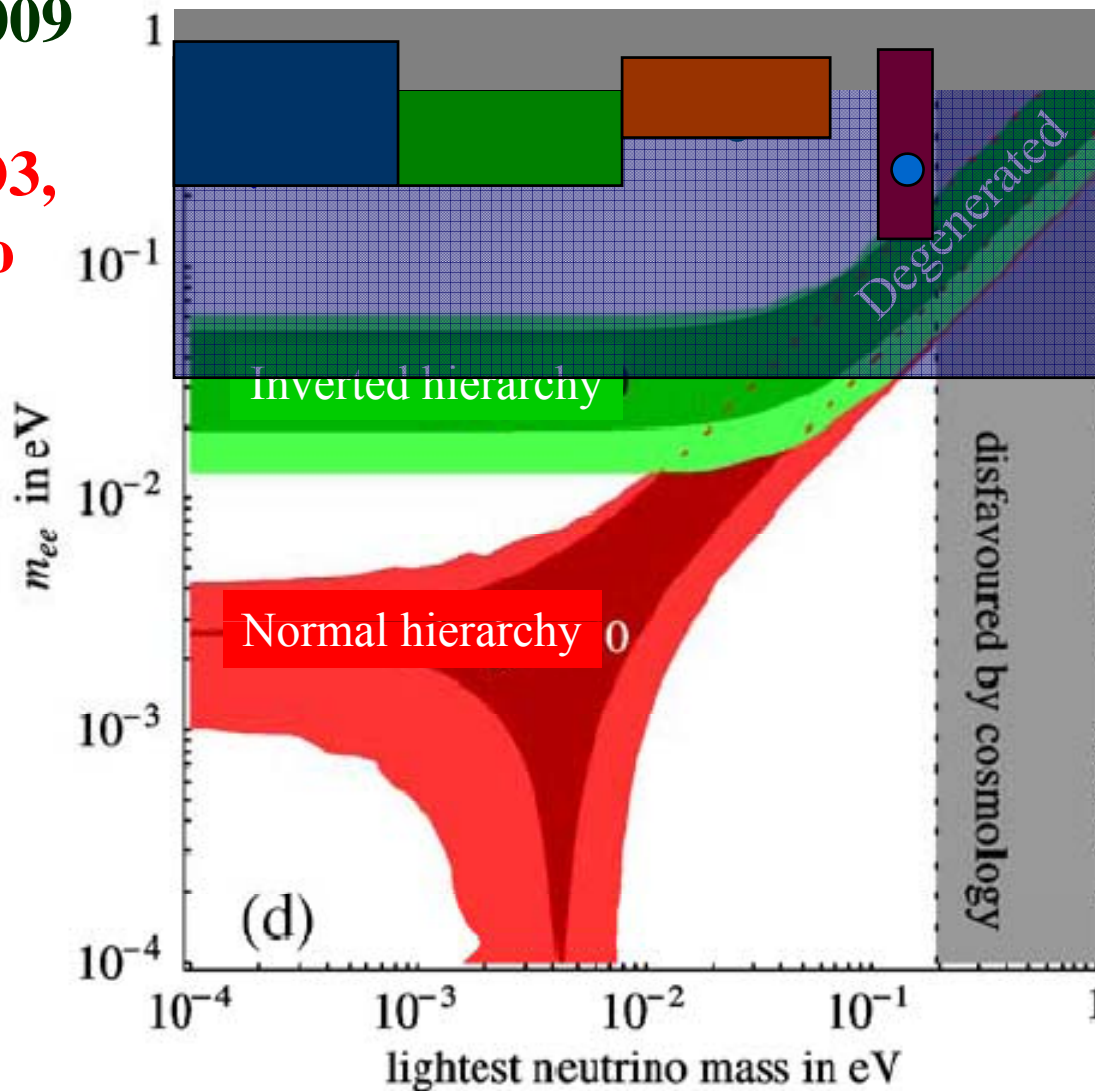
NEMO3

**Klapdor
claim**

Neutrino mass

Limits in 2009

**HM, NEMO3,
Cuoricino**



**Expected limits
2009 – 2015**

**CUORE, GERDA,
SuperNEMO,
Majorana, EXO,**

Strategy for double beta decay search

To try to cover the inverted hierarchy \rightarrow to go to 1 ton detectors

Which nuclei ?

Which techniques ?

R&D needed after 100 kg ? (it was the case to go from 1 kg to 10 kg
and from 10 kg to 100 kg)

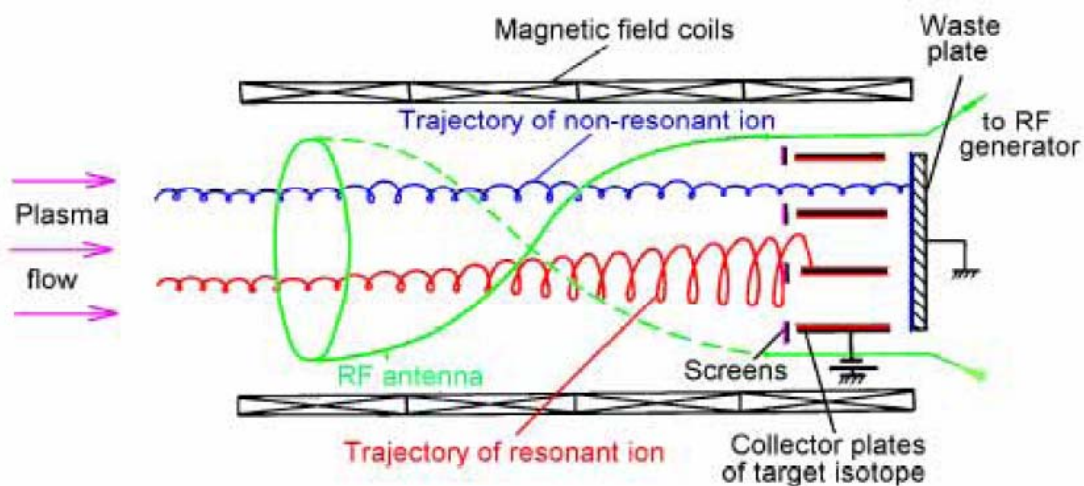
Difficult to produce today a calendar beyond 2015

Cost of the enriched product > 50 M€ for 1 ton

> 100 M€ per experiment

If Europe support at the level of 50 % 3 experiments $\rightarrow 150 - 200$ M€

Ion cyclotron resonance

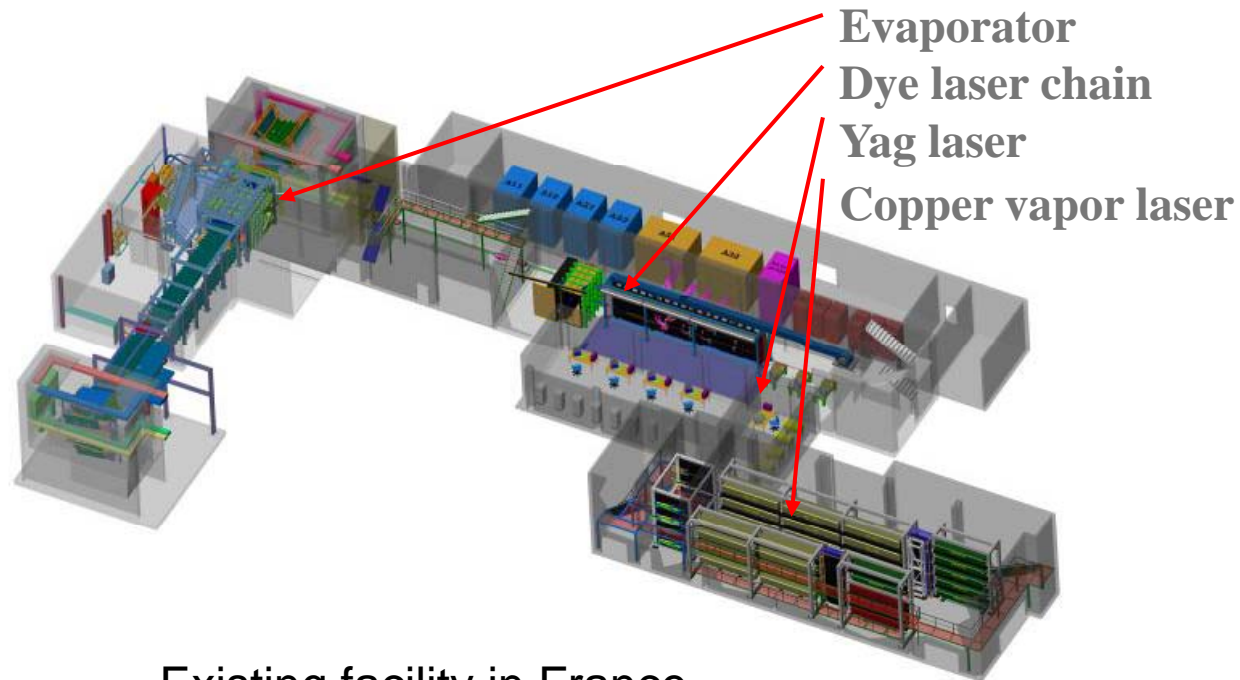


Allow production of ^{48}Ca , ^{76}Ge , ^{82}Se , ^{150}Nd , ^{100}Mo ,...

Prototype built by CEA has worked for small quantities (mg)

Extrapolation possible in principle to produce 10 – 100 kg

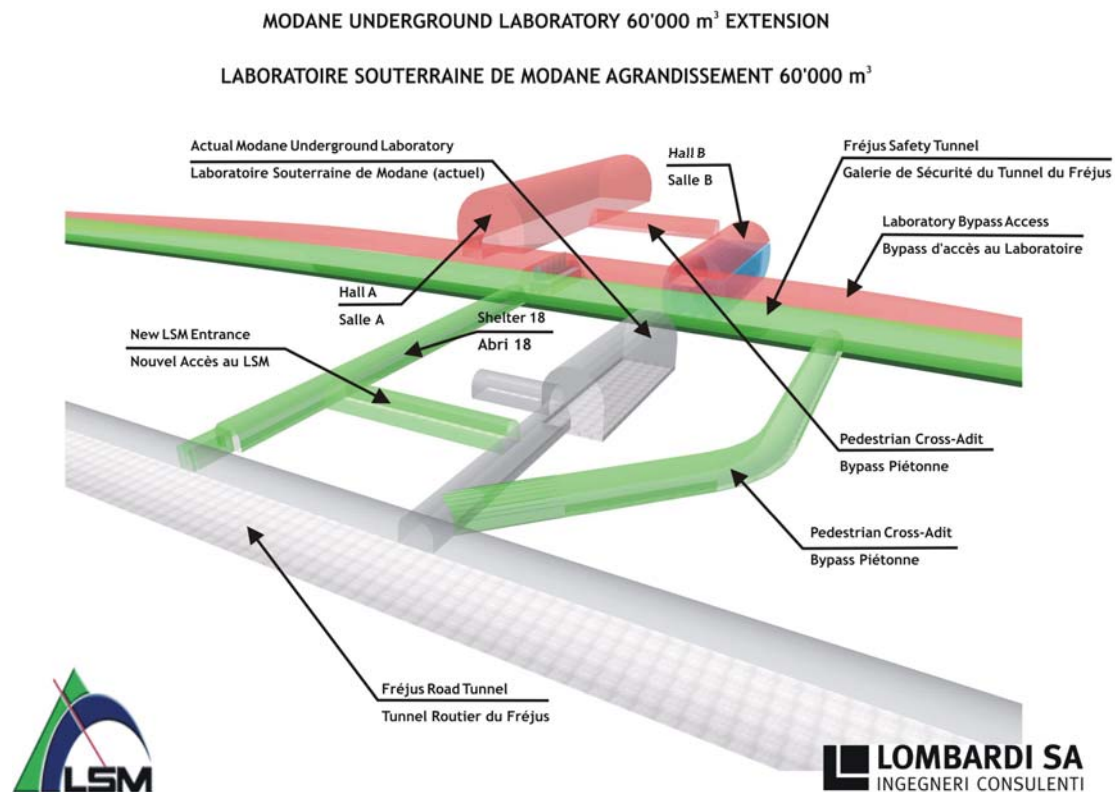
Laser isotopique separation



Existing facility in France

Could produced ^{150}Nd in large quantities

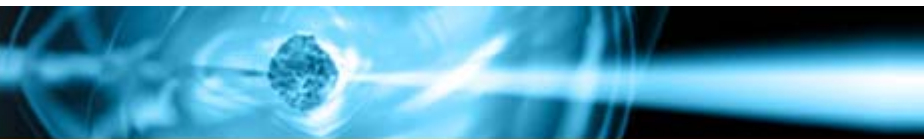
Modane underground laboratory extension



Excavation of safety gallery 2008 - 2011

	2008	2009	2010	2011	2012	2013	2014	2015	2016	2017	2018
Single beta decay											
KATRIN	Construction	Construction	Commissioning	Operation	Operation	Operation	Operation	Operation	Operation		
MARE II	Construction	Construction	Construction	Construction	Construction	Operation	Operation	Operation	Operation		
Double beta decay											
CUORE	Construction	Construction	Construction	Commissioning	Operation	Operation	Operation	Operation	Operation		
GERDA Phase I + II	Construction	Commissioning	Operation	Operation							
SuperNEMO Canfranc	Construction	Construction	Construction	Construction	Operation	Operation	Operation	Operation	Operation	Operation	Operation
SuperNEMO Modane					Construction	Construction	Commissioning	Operation	Operation	Operation	Operation
Infrastructures											
Modane extension LSM	Commissioning	Commissioning	Commissioning	Construction	Commissioning	Operation	Operation	Operation	Operation	Operation	Operation
ICR	Construction	Construction	Construction	Construction	Commissioning	Operation	Operation	Operation	Operation	Operation	Operation

Phases	
1. R&D	Light Green
2. Construction	Dark Green
3. Commissioning	Blue
4. Operation	Red

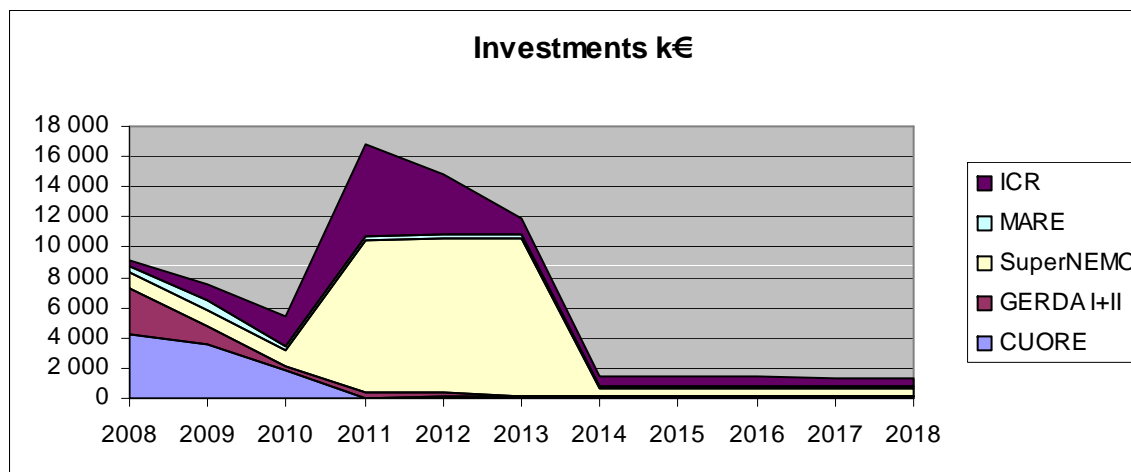


Neutrino mass

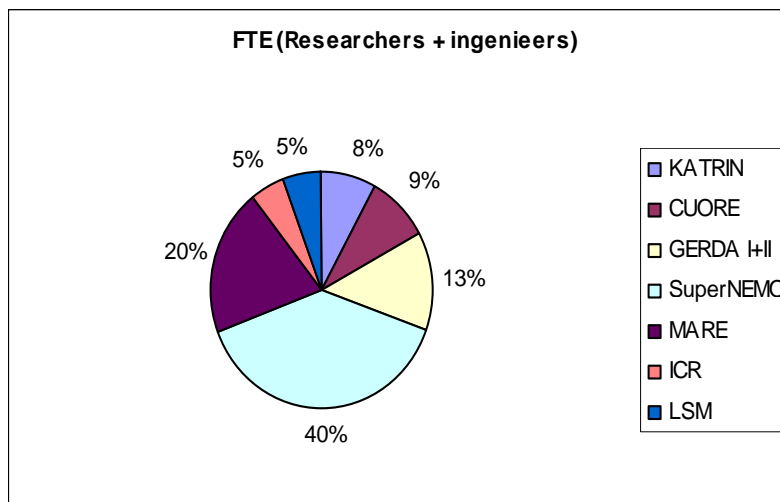
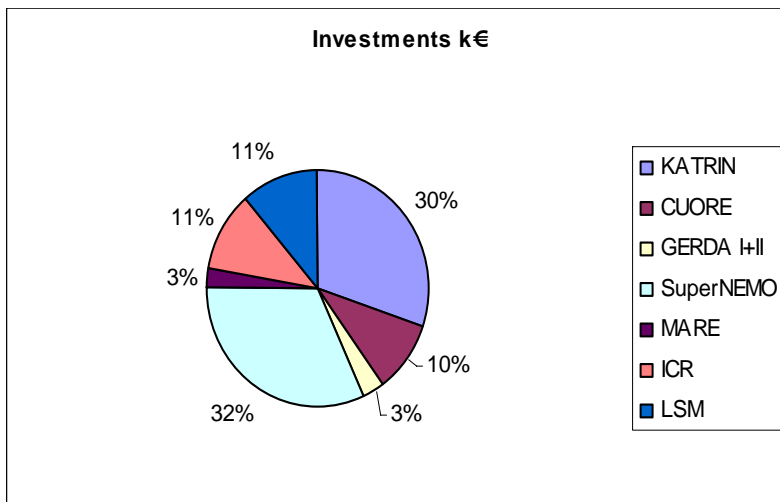
Experiment	k€	
main source		milestones
•KATRIN: Astrop.Phys.	40.000	- operation 2010
•MARE I: Astrop.Phys.	5.450	R&D<2011; construction=2011; operation 2016
•CUORE: Astrop. Phys	13.900	contruction + commissioning<2012; operation 2017
•GERDA I+II: Astrop. Phys	5.250	contruction + commissioning<2009; operation 2009
•SuperNEMO: Astrop. Phys	46.000	R&D <2009 construction <2013 Operation 2011

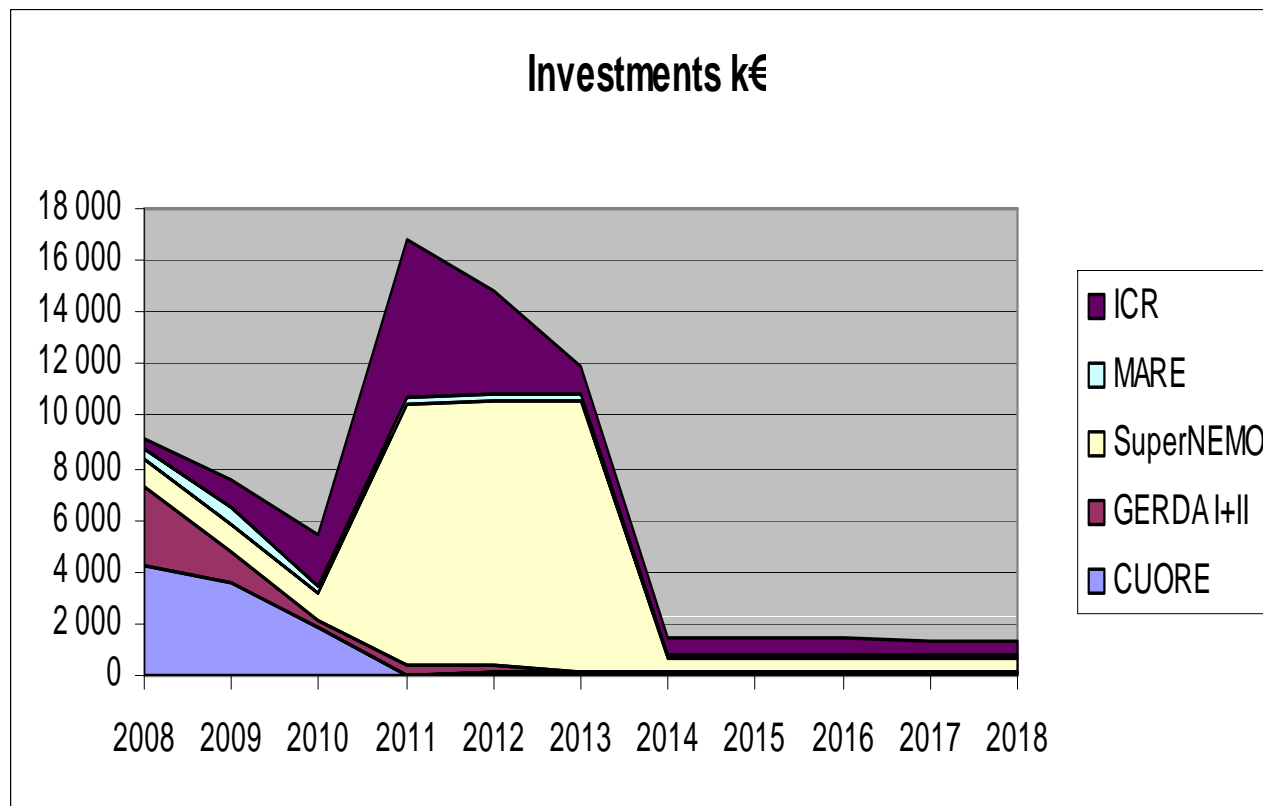
Investment

	<i>total</i>		<i>Europe</i>		<i>presently involved scientists</i>	
	Investmen	FTE	Investmen	FTE	total	ASPERA
KATRIN	40 000	5 000	35 000	5 000		
MARE	5 450	9 205	2 100	4 920	13	10
CUORE	13900	5240	9 500	3 680	16	12
GERDA I+II	5 250	9 840	5 250	7 980	35	27
SuperNEMO	45820	26775	33 000	21 300	36	30
ICR	17500	3200	14 500	3 200	4	4
	127 920	59 260	99 350	46 080	104	83

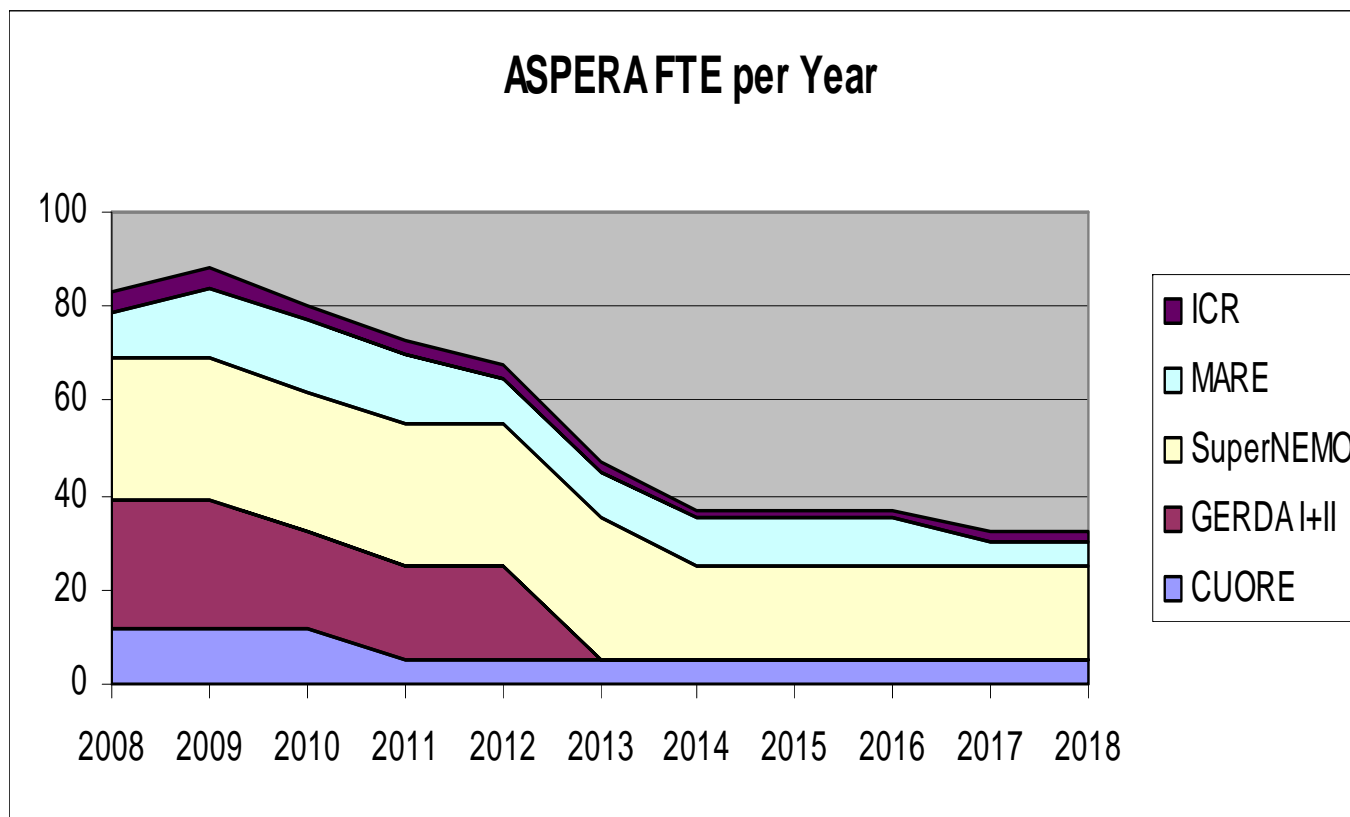


Neutrino mass





Need R&D investments beyond 2015 to improve and develop new techniques



Plans for 2008 - 2018

Plans up to 2015 are clear:

Single beta decay KATRIN (funded) and MARE (R&D)

Double beta decay: 2 experiments funded GERDA and CUORE (calorimeters)
1 R&D SuperNEMO (tracko-calo)

Results expected in 2013 - 2015

Need significant effort for R&D beyond this phase

Future plans beyond 2015 will depend on the results of 100 kg phase

